

LATTES baseline performance

Ruben Conceição

on behalf of the LATTES team

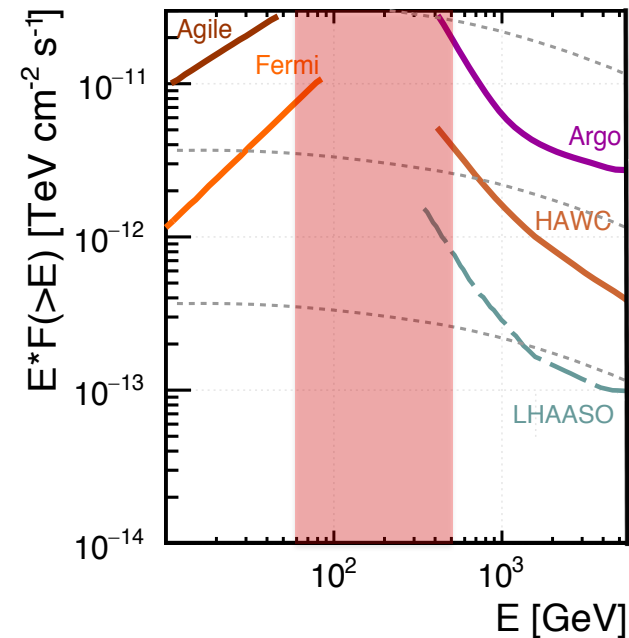
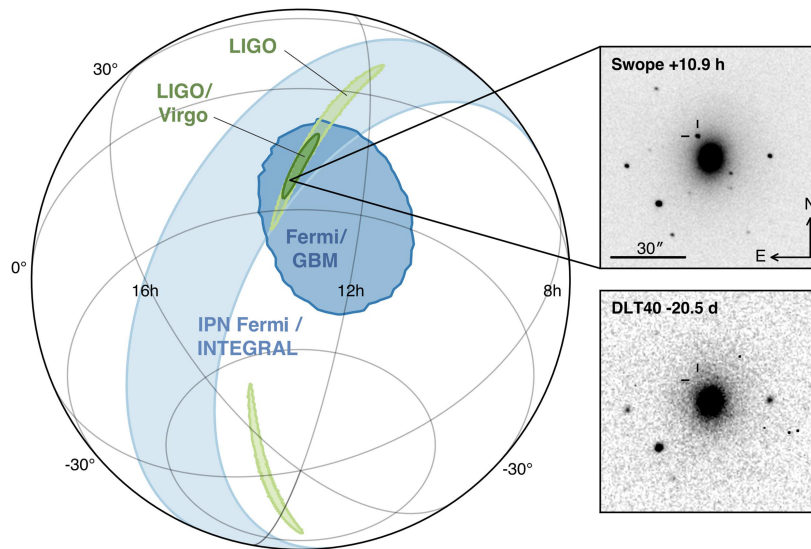


The era of multi-messenger observations



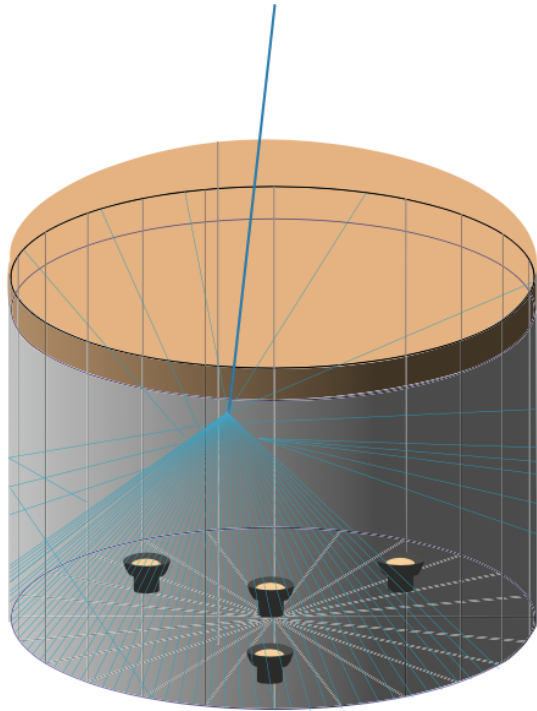
Joint publication of LIGO, VIRGO, INTEGRAL, Fermi, IceCube, Pierre Auger ...

Current wide FoV gamma-ray observatories

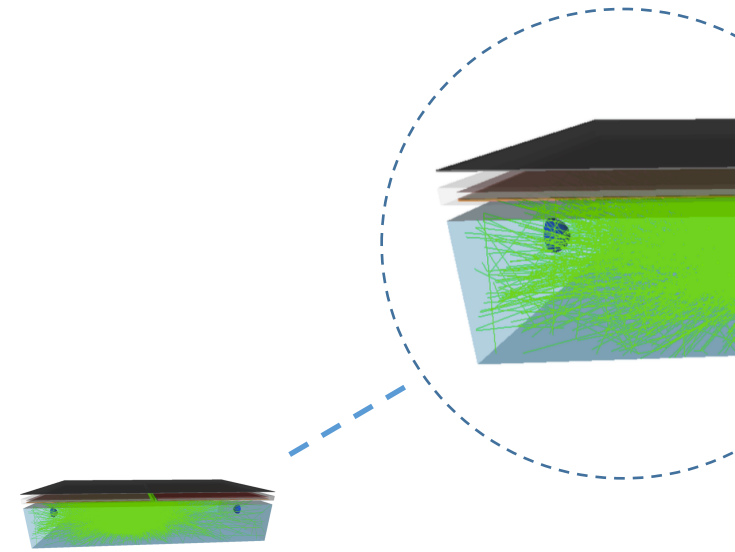


- ❖ Simultaneous observation of a Gravitational Wave + electromagnetic counterparts
- ❖ Study of transient phenomena in all energy windows is one of the main ingredients

Station: HAWC vs LATTES

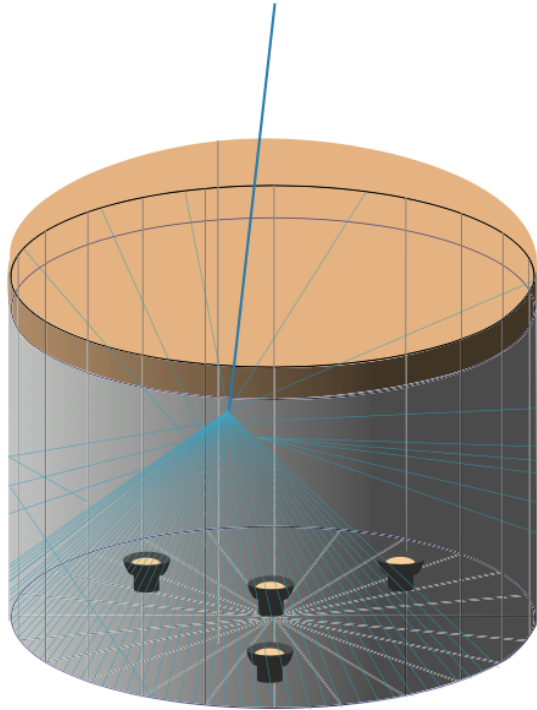


HAWC
(present detector)

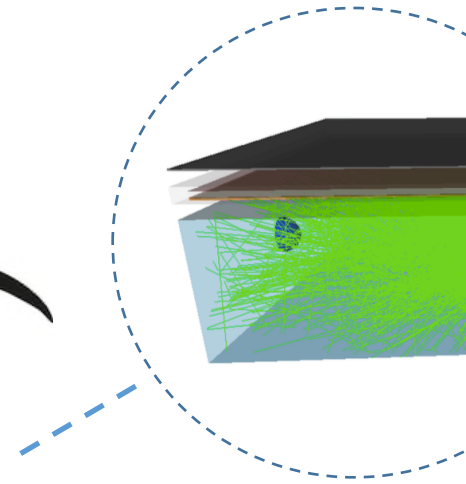
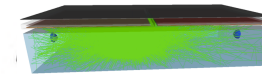


LATTES
(next generation)

Station: HAWC vs LATTES

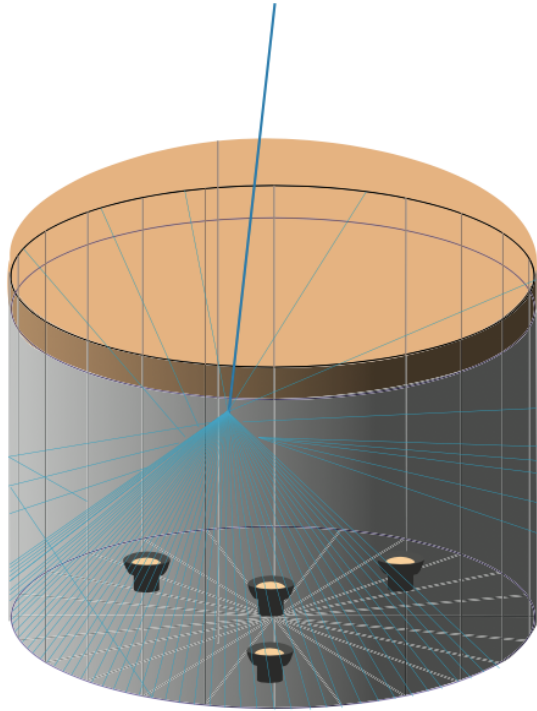


HAWC
(present detector)

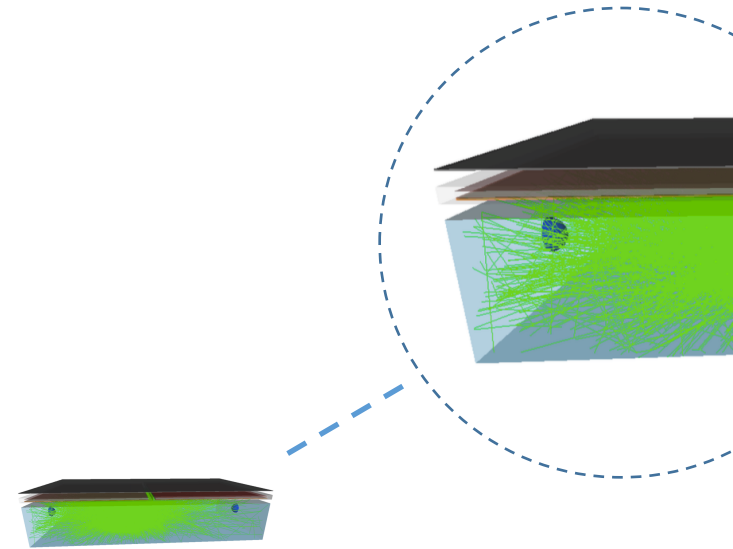


LATTES
(next generation)

Station: HAWC vs LATTES



HAWC
(present detector)



LATTES
(next generation)

Simulation Framework

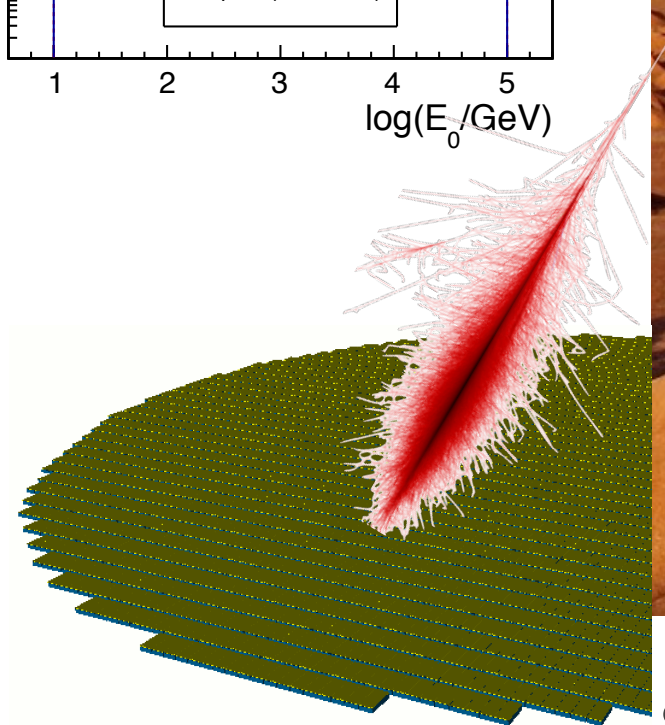
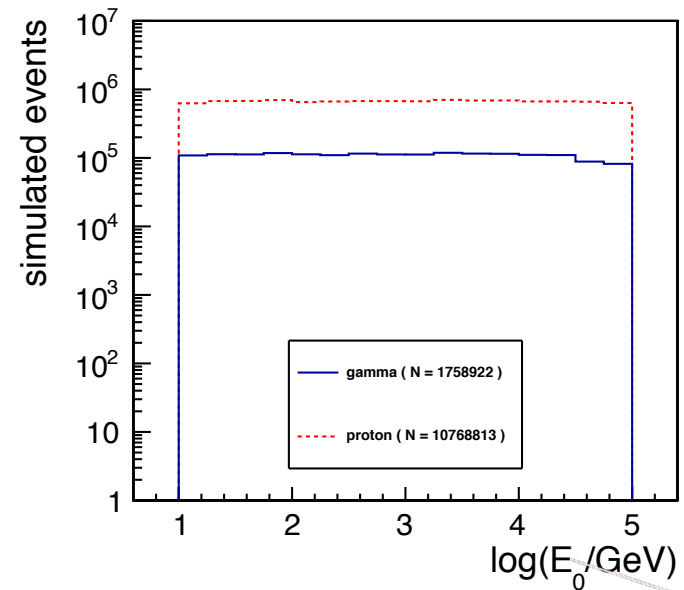
✧ End-to-end realistic simulation

✧ Extensive Air Showers: **CORSIKA**

- ✧ v7.6400 with Fluka2011.2c
- ✧ More than 50 000 gamma/proton shower simulated randomly between 10 GeV – 300 TeV
- ✧ Gammas have a fixed zenith angle of 10 degrees
- ✧ Observation level at 5200 m of altitude

✧ Detector simulation: **Geant4**

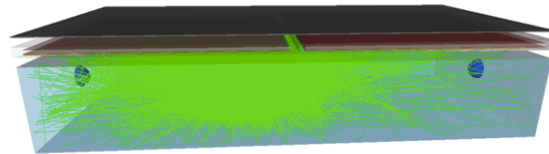
- ✧ v10.1.3
- ✧ Core array 20 000 m²
- ✧ Each shower is resampled 100 times over a big area containing all the array



Simulation Framework

✧ Reconstruction

- ✧ **First order analyses** with little optimization only to demonstrate principle



✧ Performance and sensitivity

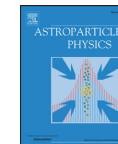
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Astroparticle Physics

journal homepage: www.elsevier.com/locate/astropartphys



Design and expected performance of a novel hybrid detector for very-high-energy gamma-ray astrophysics



P. Assis^{a,b}, U. Barres de Almeida^c, A. Blanco^d, R. Conceição^{a,b,*}, B. D'Ettorre Piazzoli^e,
A. De Angelis^{f,g,b,a}, M. Doro^{h,f}, P. Fonte^d, L. Lopes^d, G. Matthiaeⁱ, M. Pimenta^{b,a}, R. Shellard^c,
B. Tomé^{a,b}

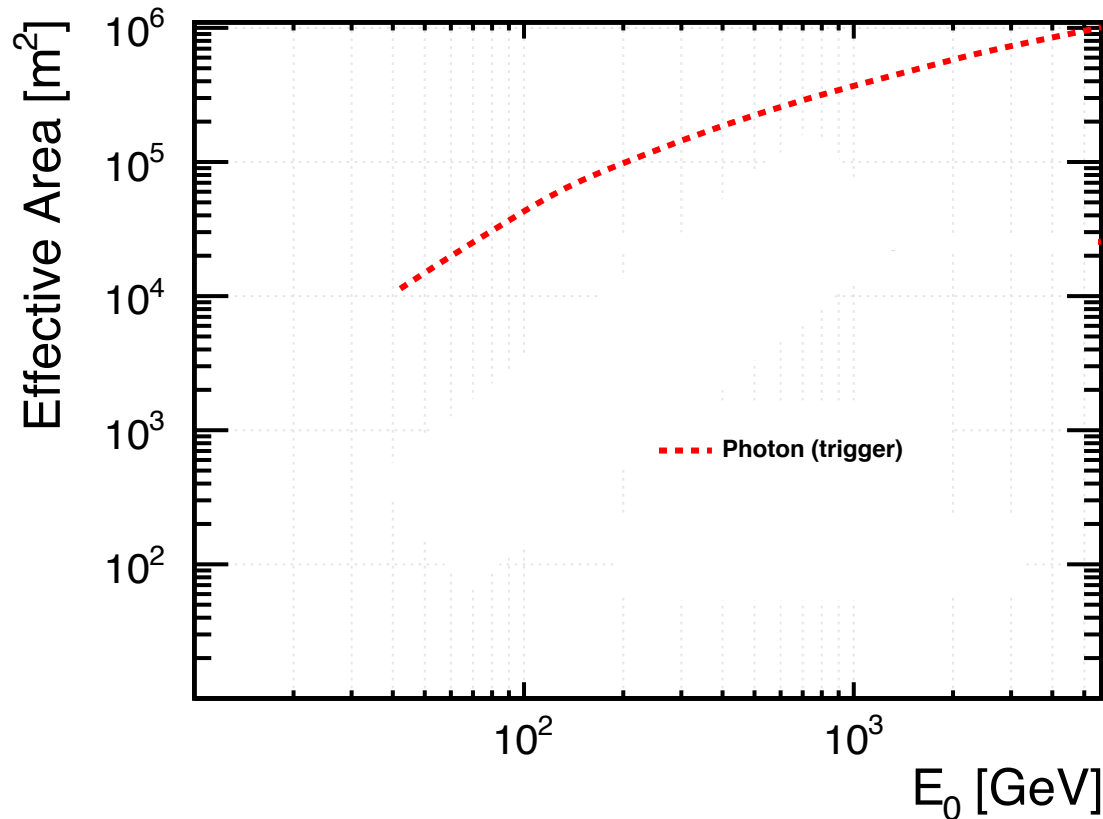
LATTES expected performance

- ✧ Trigger and effective area
- ✧ Core reconstruction
- ✧ Energy reconstruction
- ✧ Geometry reconstruction
- ✧ Gamma/hadron discrimination
- ✧ Sensitivity to steady sources

LATTES expected performance

- ✧ **Trigger and effective area**
- ✧ Core reconstruction
- ✧ Energy reconstruction
- ✧ Geometry reconstruction
- ✧ Gamma/hadron discrimination
- ✧ Sensitivity to steady sources

Trigger efficiency



- ✧ Use **WCD stations to trigger** at low energies
 - ✧ Trigger condition
 - ✧ Station: require more than 5 p.e. in each PMT
 - ✧ Event: require 3 triggered stations

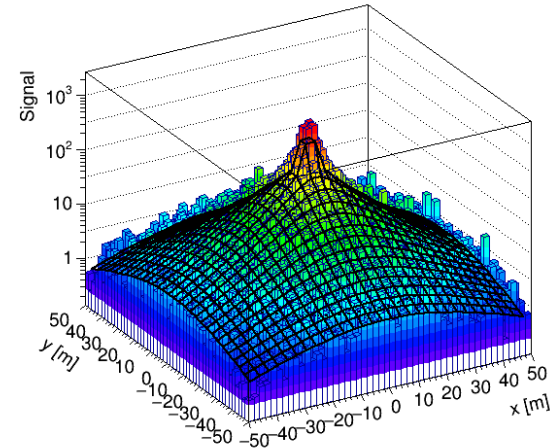
LATTES expected performance

- ✧ Trigger and effective area
- ✧ **Core reconstruction**
- ✧ Energy reconstruction
- ✧ Geometry reconstruction
- ✧ Gamma/hadron discrimination
- ✧ Sensitivity to steady sources

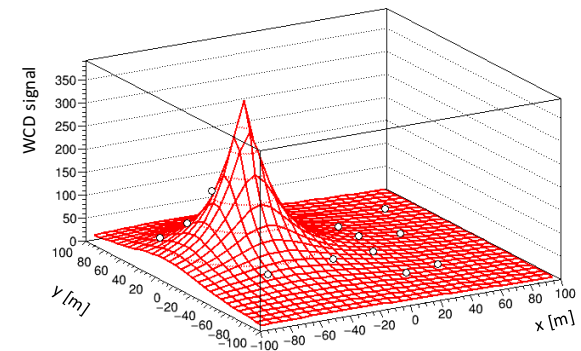
Shower core reconstruction

- ✧ Use the WCD signal
- ✧ Barycenter
 - ✧ Initial guess
 - ✧ Works but the core is always reconstructed inside the array
- ✧ Fit the WCD LDF
 - ✧ Fit photon average LDF to fix the shape
 - ✧ Function inspired in HAWC
 - ✧ Nearly no evolution with energy
 - ✧ Use this form to find the maximum, i.e. the shower core

Average LDF



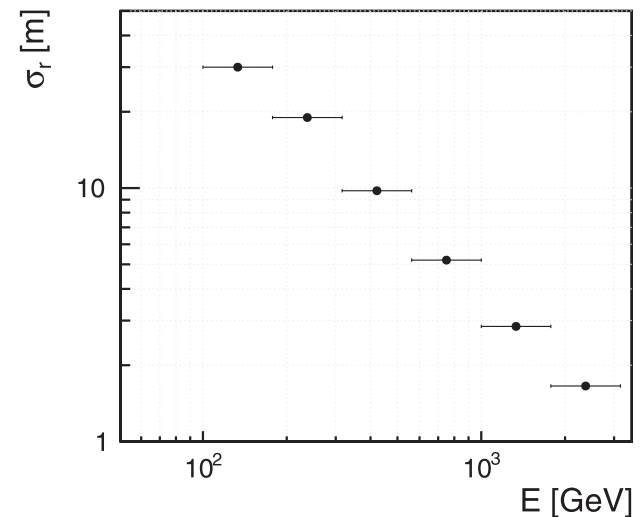
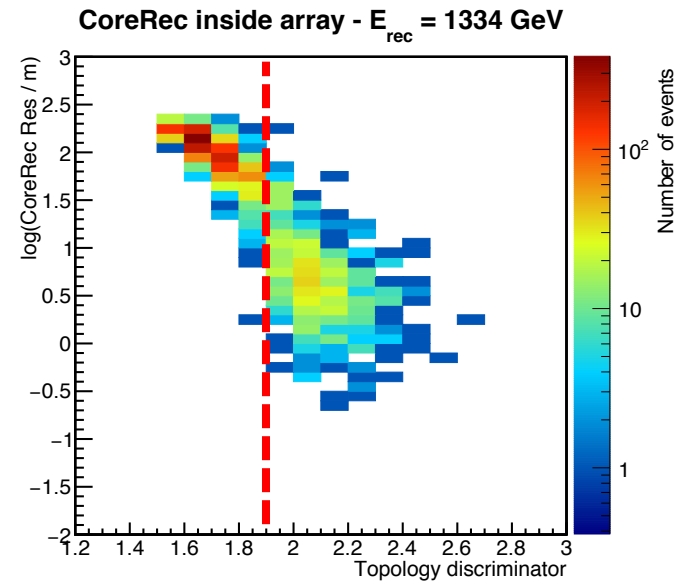
Single event



$$S_i = S(A, \vec{x}, \vec{x}_i) = A \left(\frac{1}{2\pi\sigma^2} e^{-|\vec{x}_i - \vec{x}|^2 / 2\sigma^2} + \frac{N}{(0.5 + |\vec{x}_i - \vec{x}| / R_m)^3} \right)$$

Shower core reconstruction

- ✧ Test whether the shower is inside/outside the array
 - ✧ Explore LDF topology
 - ✧ Is maximum observed inside of array?
 - ✧ Currently exploring the quality of the fit
 - ✧ Fixed cut for all energies
- ✧ Resolution better than 10 meters for showers above 300 GeV



LATTES expected performance

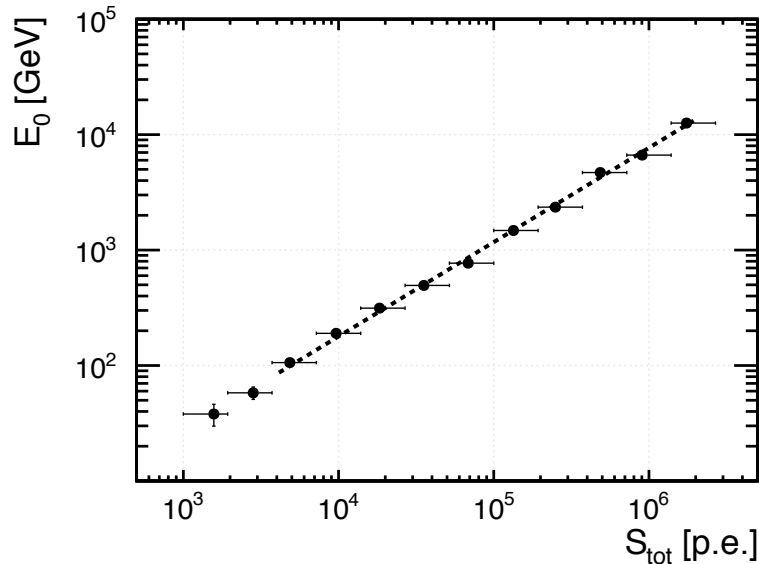
- ✧ Trigger and effective area
- ✧ Core reconstruction
- ✧ **Energy reconstruction**
- ✧ Geometry reconstruction
- ✧ Gamma/hadron discrimination
- ✧ Sensitivity to steady sources

Energy reconstruction

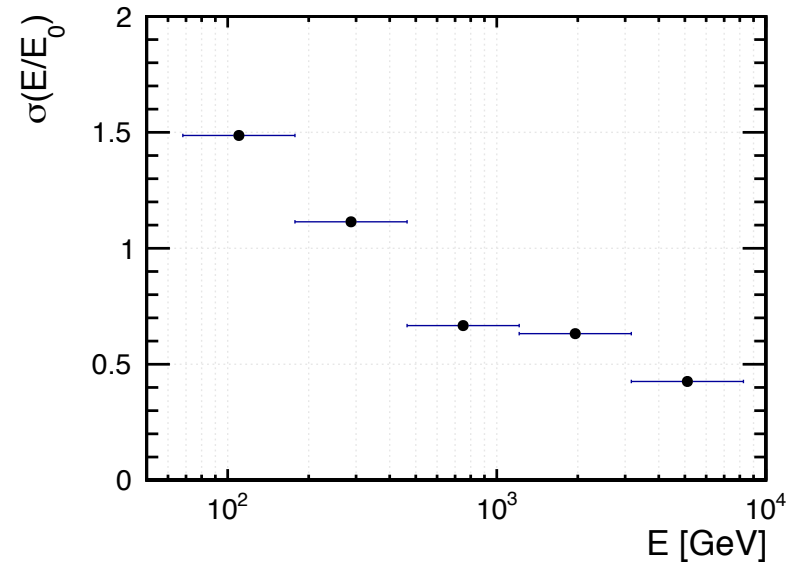
$E_0 \rightarrow$ Simulated energy

$E \rightarrow$ Reconstructed energy

Energy Calibration



Energy Resolution



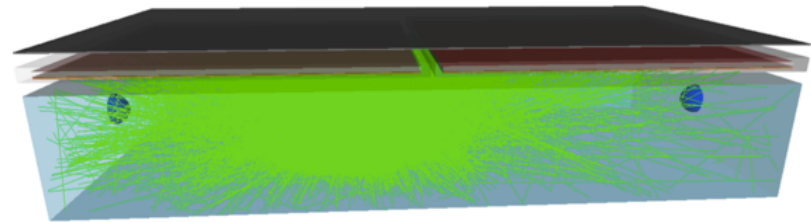
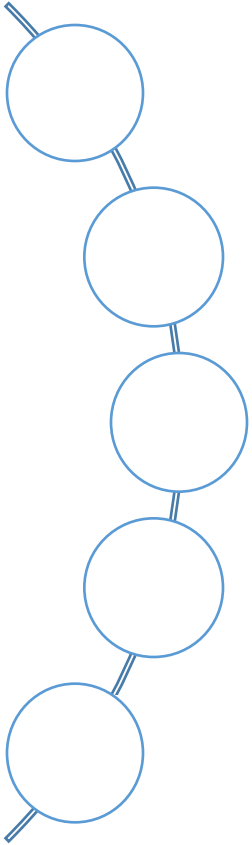
- ✧ Use as **energy estimator** the **total signal** recorded by **WCDs**
 - ✧ Use only shower cores reconstructed inside array
- ✧ Energy resolution at low energy dominated by shower fluctuations

LATTES expected performance

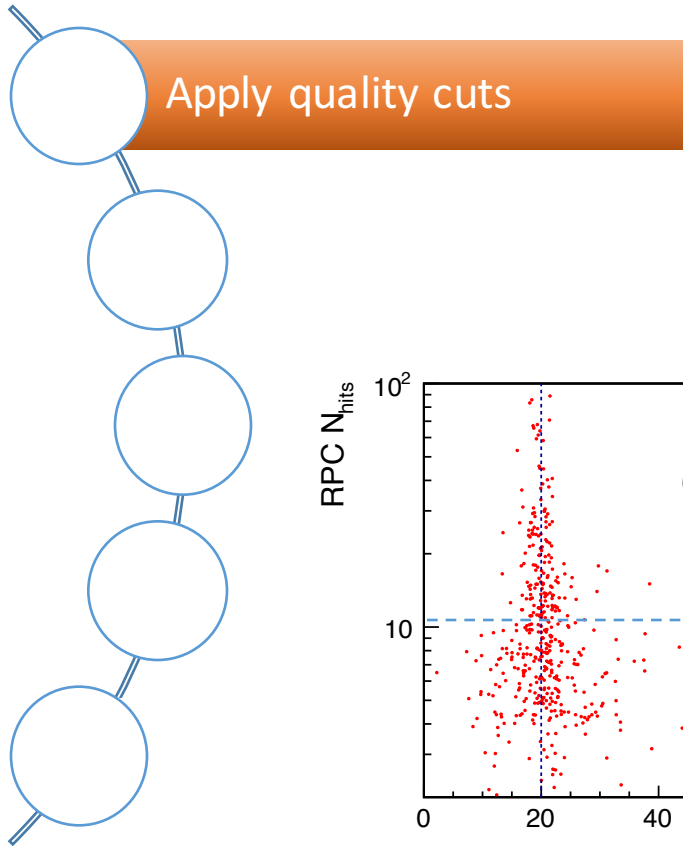
- ✧ Trigger and effective area
- ✧ Core reconstruction
- ✧ Energy reconstruction
- ✧ **Geometry reconstruction**
- ✧ Gamma/hadron discrimination
- ✧ Sensitivity to steady sources

Reconstruction of shower geometry

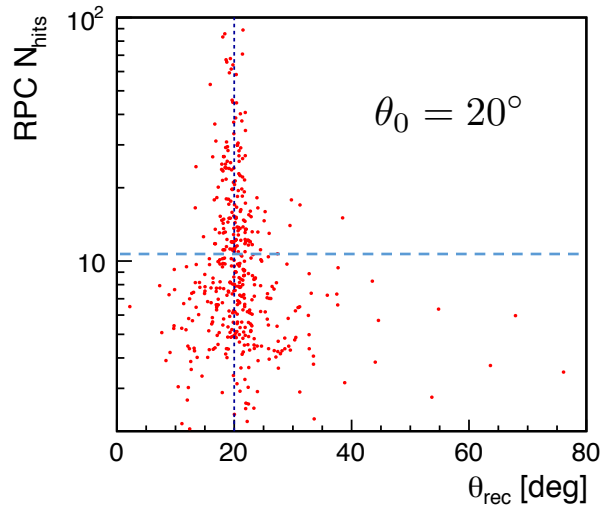
- ✧ Use **RPC hit time** information
 - ✧ Take advantage of high spatial and time resolution
 - ✧ Used time resolution of 1 ns



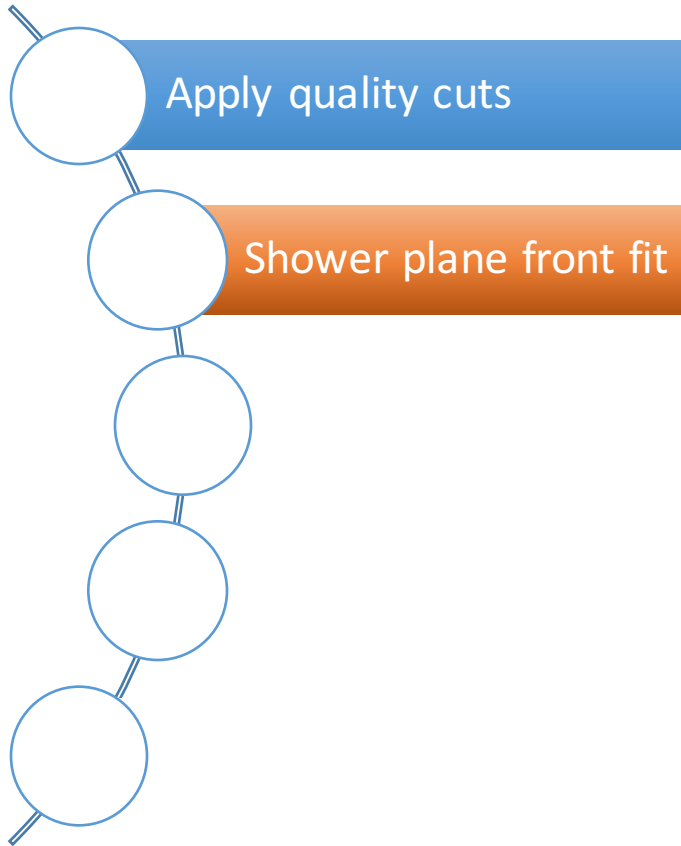
Reconstruction of shower geometry



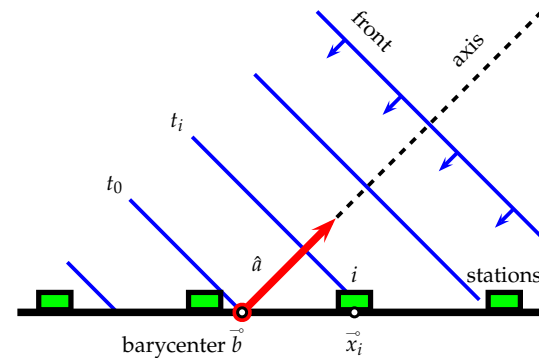
- ✧ Use **RPC hit time** information
 - ✧ Apply previous shower rec quality cuts
 - ✧ Apply cuts on the number of registered hits on the RPCs
 - ✧ Consider only RPCs in triggered WCD stations



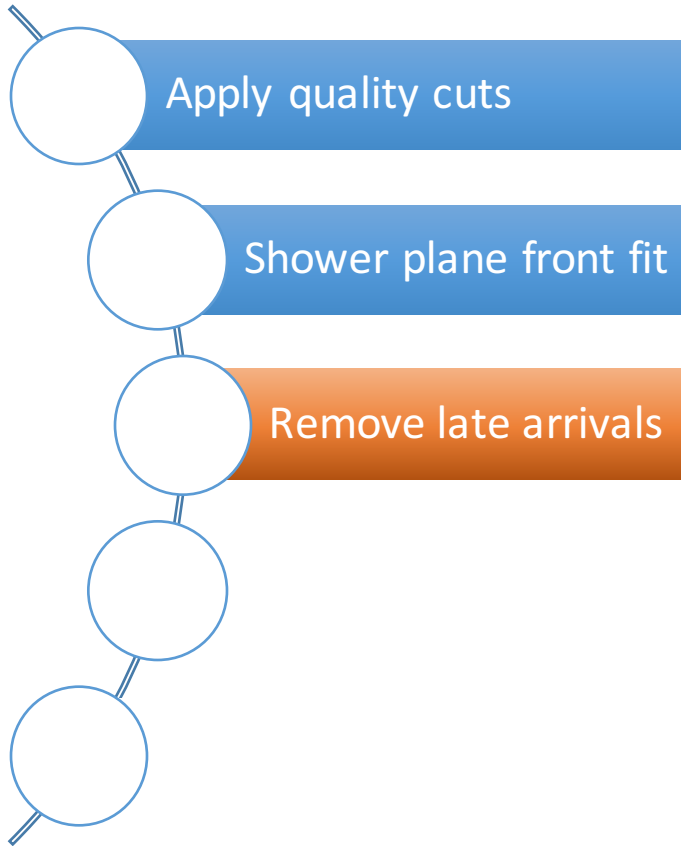
Reconstruction of shower geometry



- ✧ Use **RPC hit time** information
- ✧ Perform shower reconstruction
- ✧ Use shower front plane approximation
- ✧ Analytical procedure



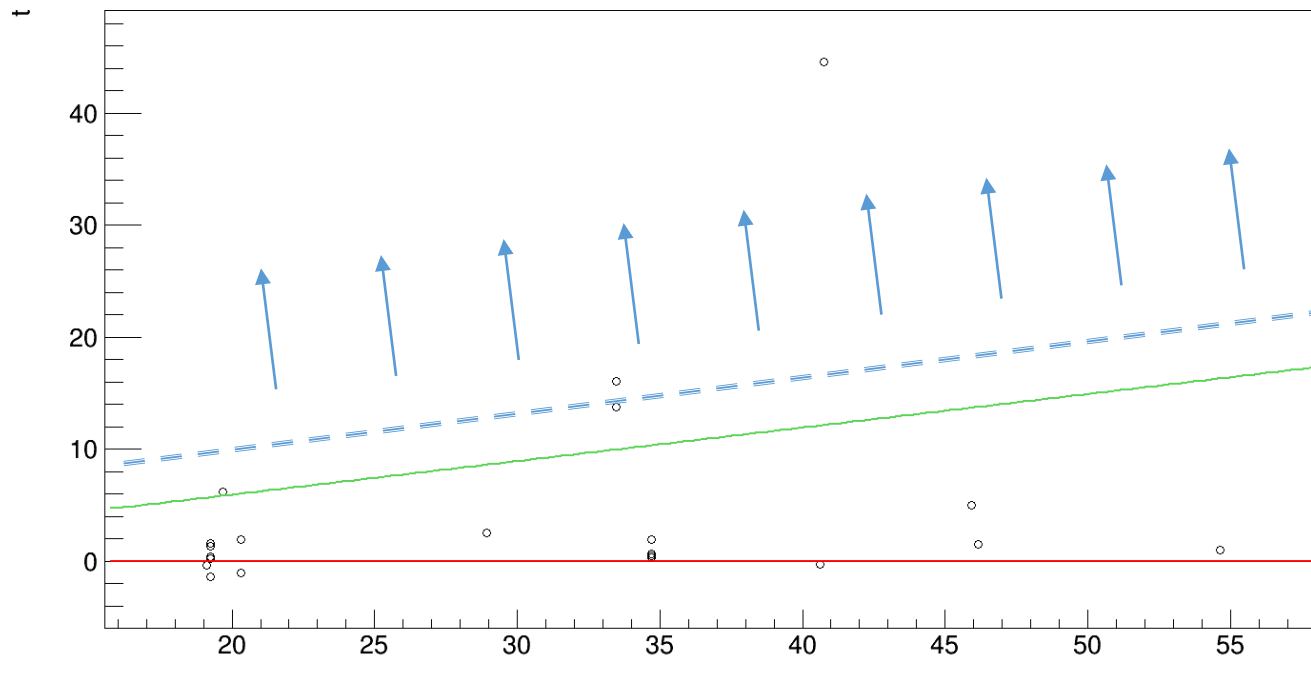
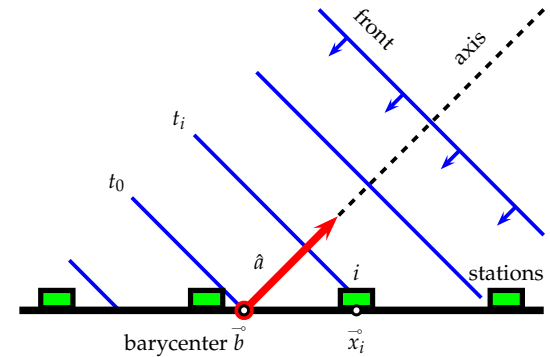
Reconstruction of shower geometry



- ✧ Use **RPC hit time** information
 - ✧ Identify late arrivals with respect to Rec Shower Front
 - ✧ Mainly low energy electrons that lost correlation with shower front

Removal of late arrivals

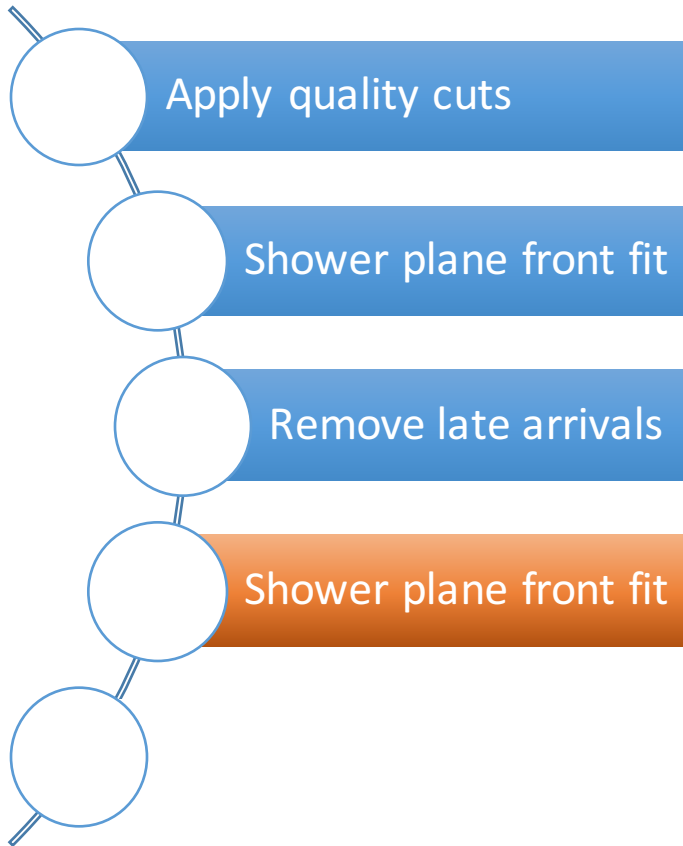
- ✧ Example of a vertical gamma shower
- ✧ Plot depicts arrival time (ns) distance to simulated shower core (m)



$$\theta_{rec} = 5.1^\circ$$

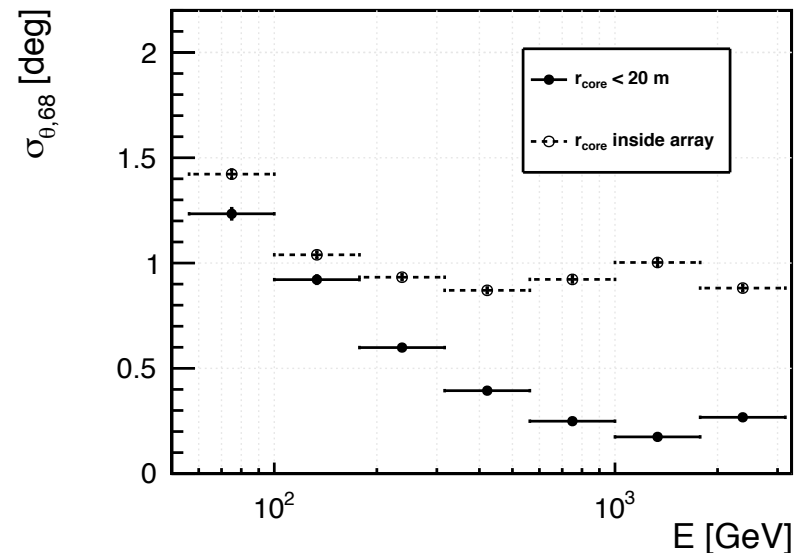
$$\theta_0 = 0^\circ$$

Reconstruction of shower geometry

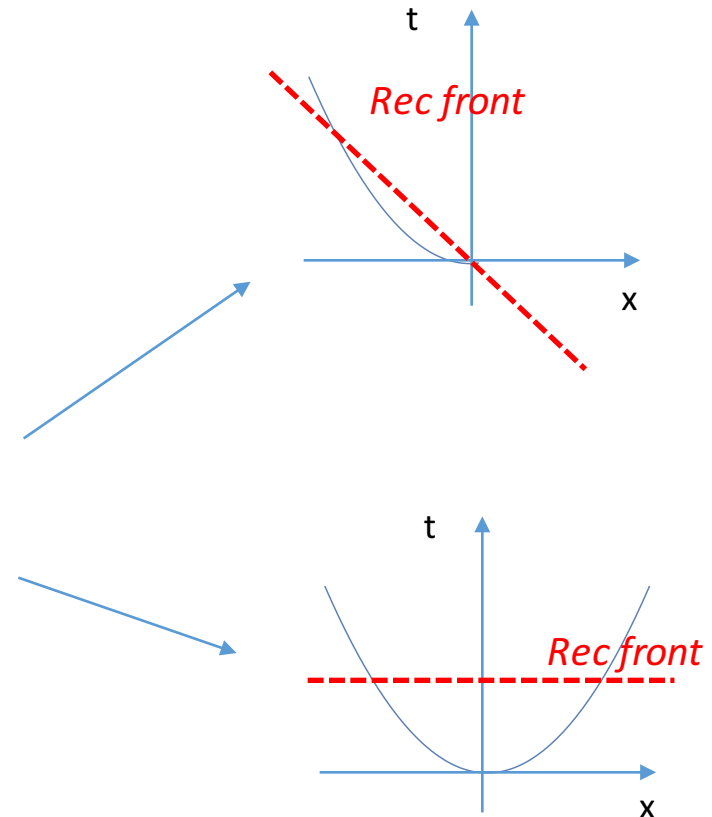
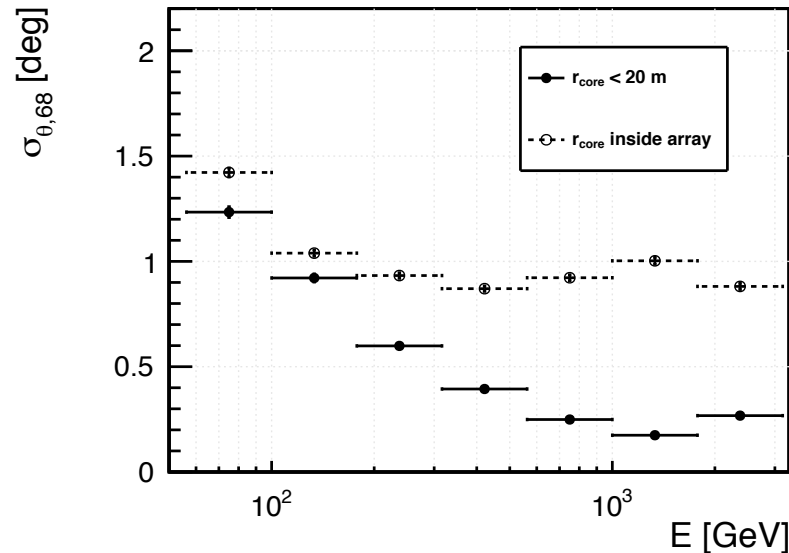


- ✧ Use **RPC hit time** information
 - ✧ Repeat fit without arrivals
 - ✧ Initial guess for next step

γ – showers; $\theta = 10^\circ$



Impact of shower curvature

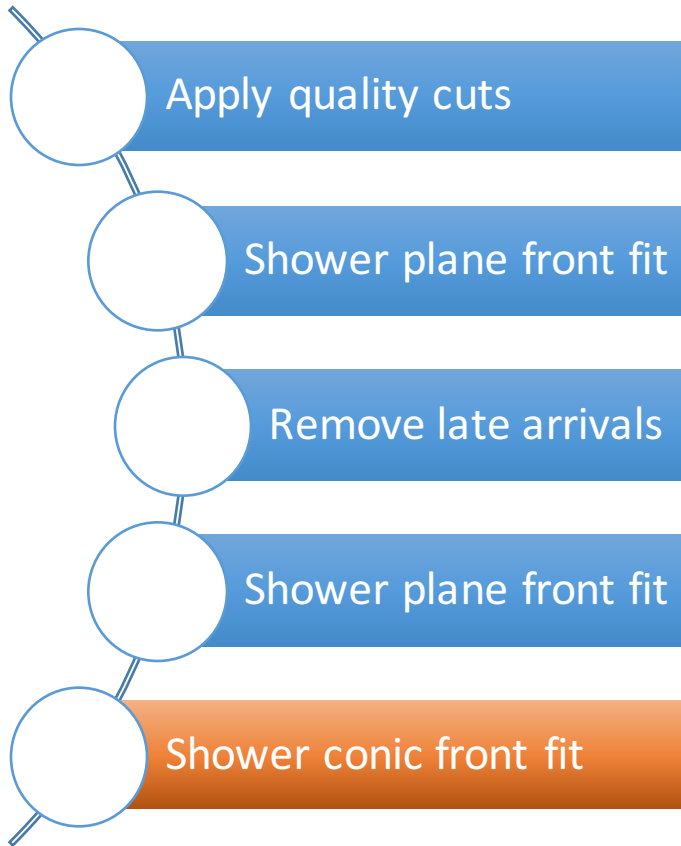


Center of the array Border of the array

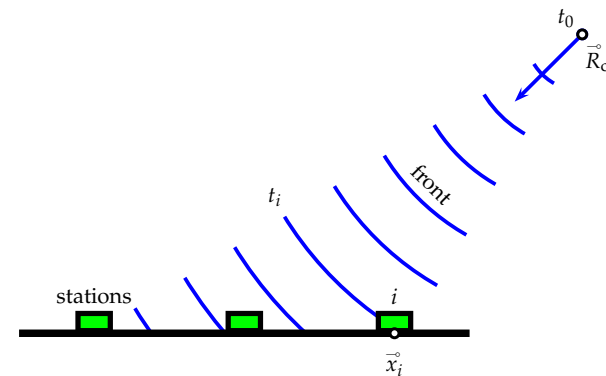
Solution: implement a conic fit instead of fitting a plane

$$\chi^2 = \sum (c \cdot (T_n - T_0) - X_n \cdot l - Y_n \cdot m - R_n \cdot \alpha)^2$$

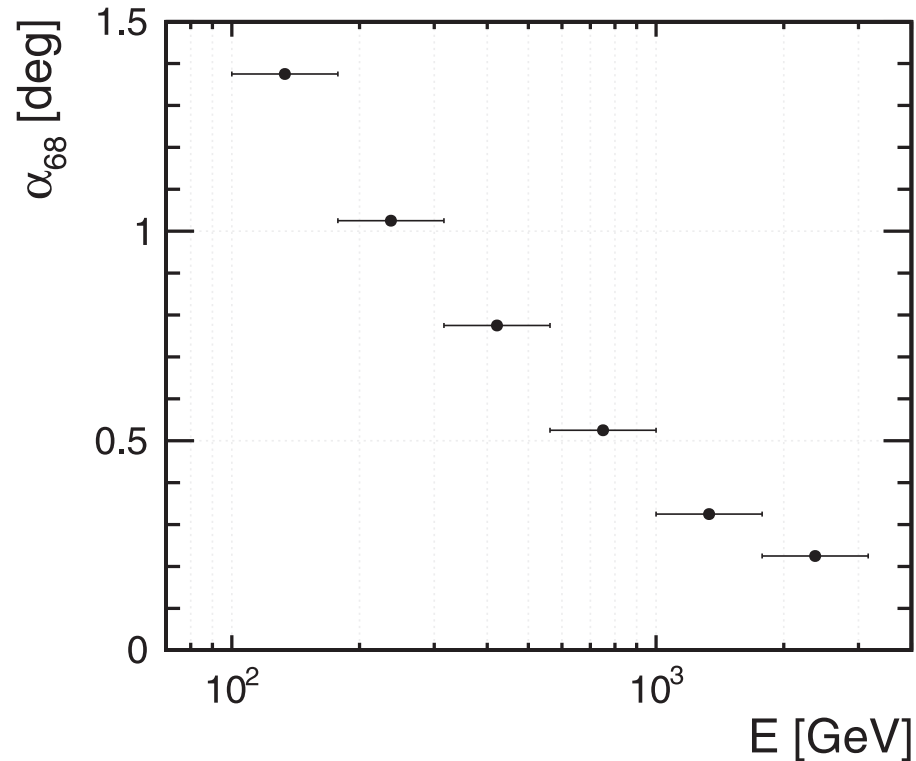
Reconstruction of shower geometry



- ✧ Use **RPC hit time** information
 - ✧ Fit the shower geometry using a shower conic front model
 - ✧ Depends on core position



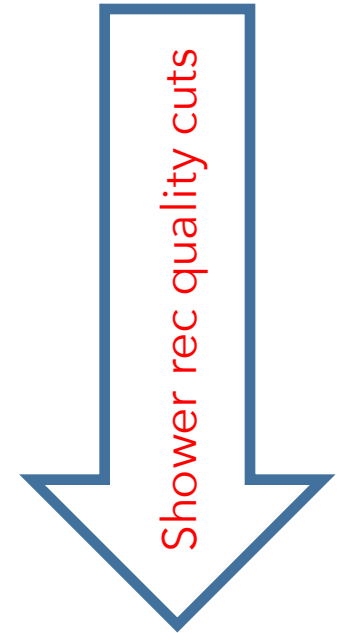
Shower geometry reconstruction



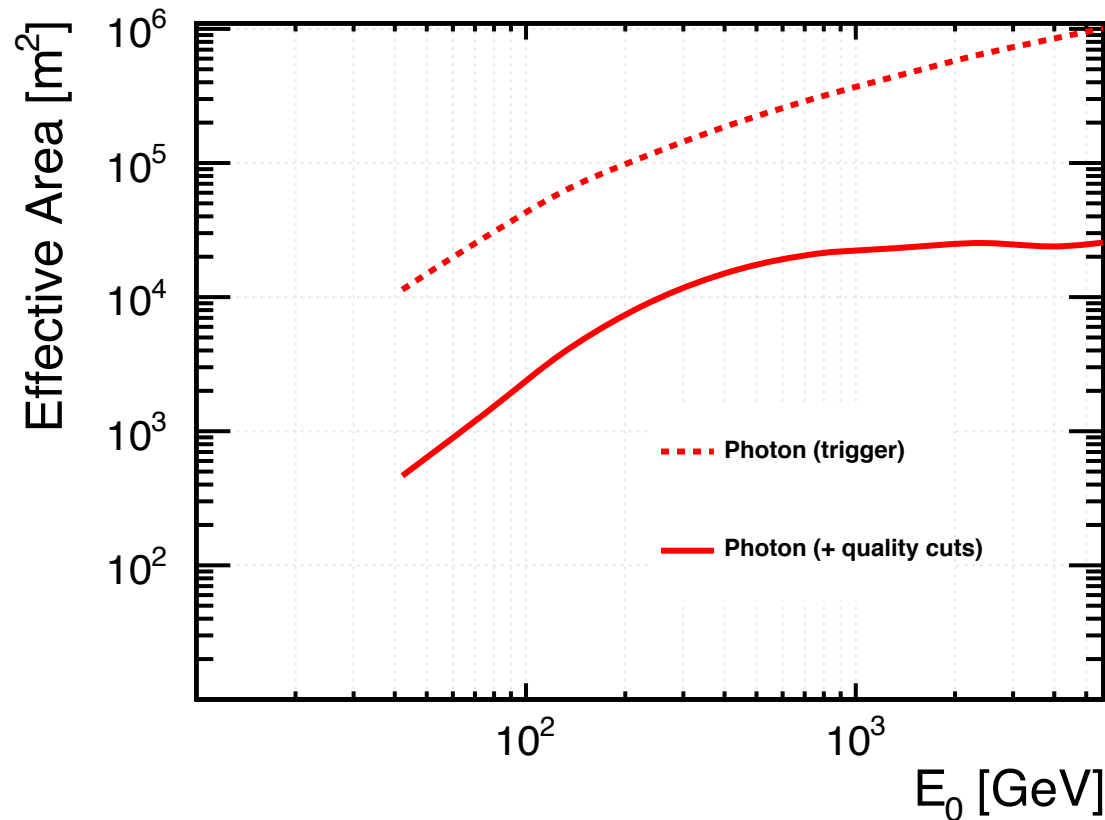
A good angular resolution can be achieved for all events reconstructed inside the array

LATTES expected performance

- ✧ Trigger and effective area
- ✧ Core reconstruction
- ✧ Energy reconstruction
- ✧ Geometry reconstruction
- ✧ Gamma/hadron discrimination
- ✧ Sensitivity to steady sources



Effective Area



Even applying all quality cuts LATTES
gets an effective area of $\sim 1000 \text{ m}^2$ for $E = 100 \text{ GeV}$

LATTES expected performance

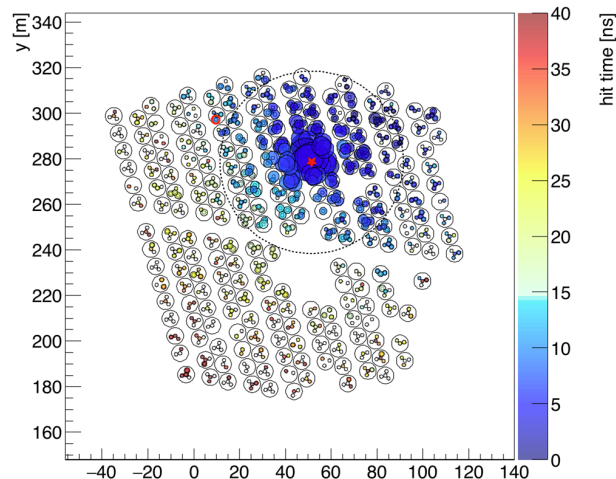
- ✧ Trigger and effective area
- ✧ Core reconstruction
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- ✧ Geometry reconstruction
- ✧ **Gamma/hadron discrimination**
- ✧ Sensitivity to steady sources

Looking for high- p_T sub-showers

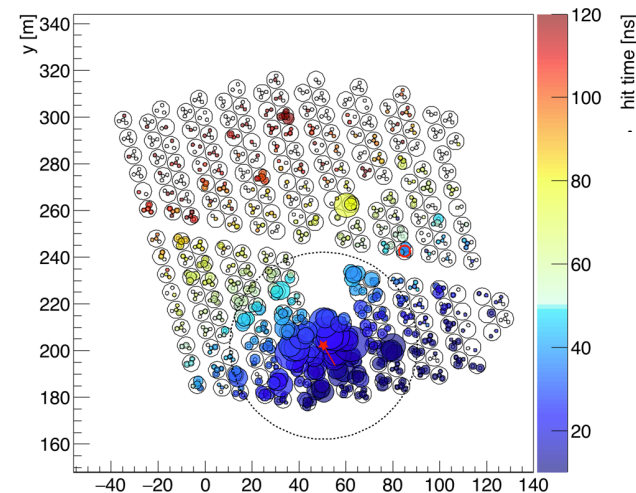
✧ HAWC g/h discrimination

- ✧ Look for high signal far away from the shower core (> 40 m)
- ✧ Take advantage of height of the tank to distinguish muons from electrons

gamma shower

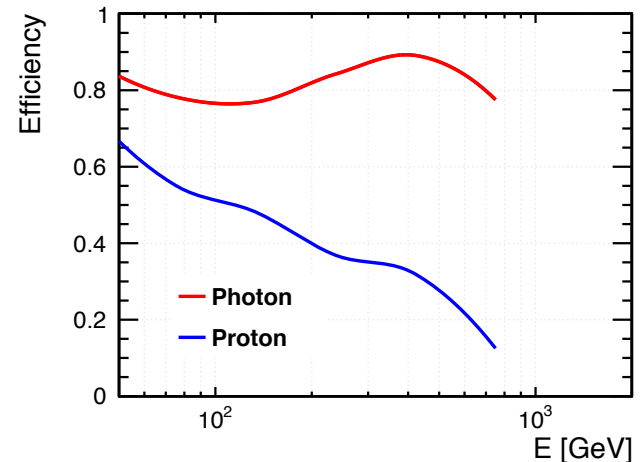
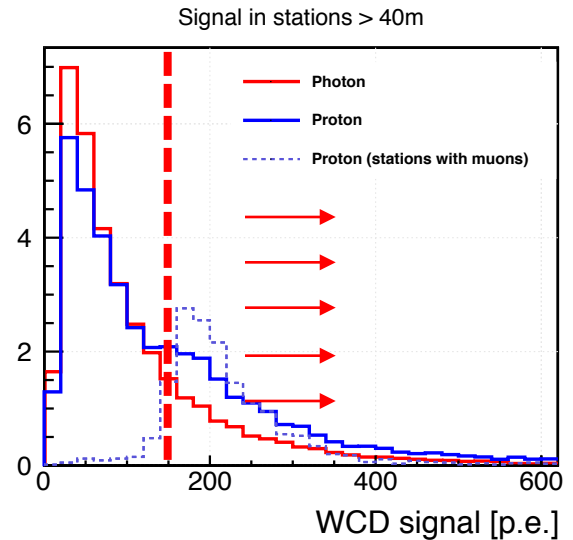


hadron shower

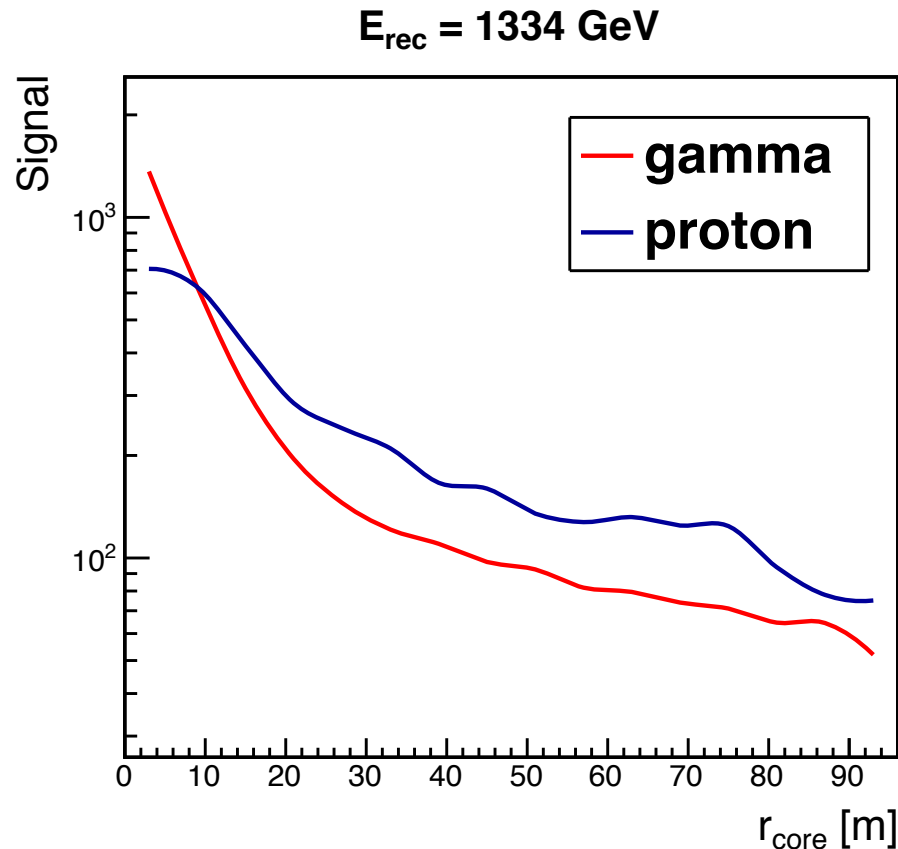


Looking for high p_t sub-showers

- ✧ LATTES g/h discrimination
 - ✧ Use only stations with a distance above 40 m
 - ✧ S40: sum all WCD stations signal
 - ✧ S40_high: sum all WCD stations that have a signal above the muon energy threshold
 - ✧ Compute S40_high / S40
 - ✧ Not optimized...

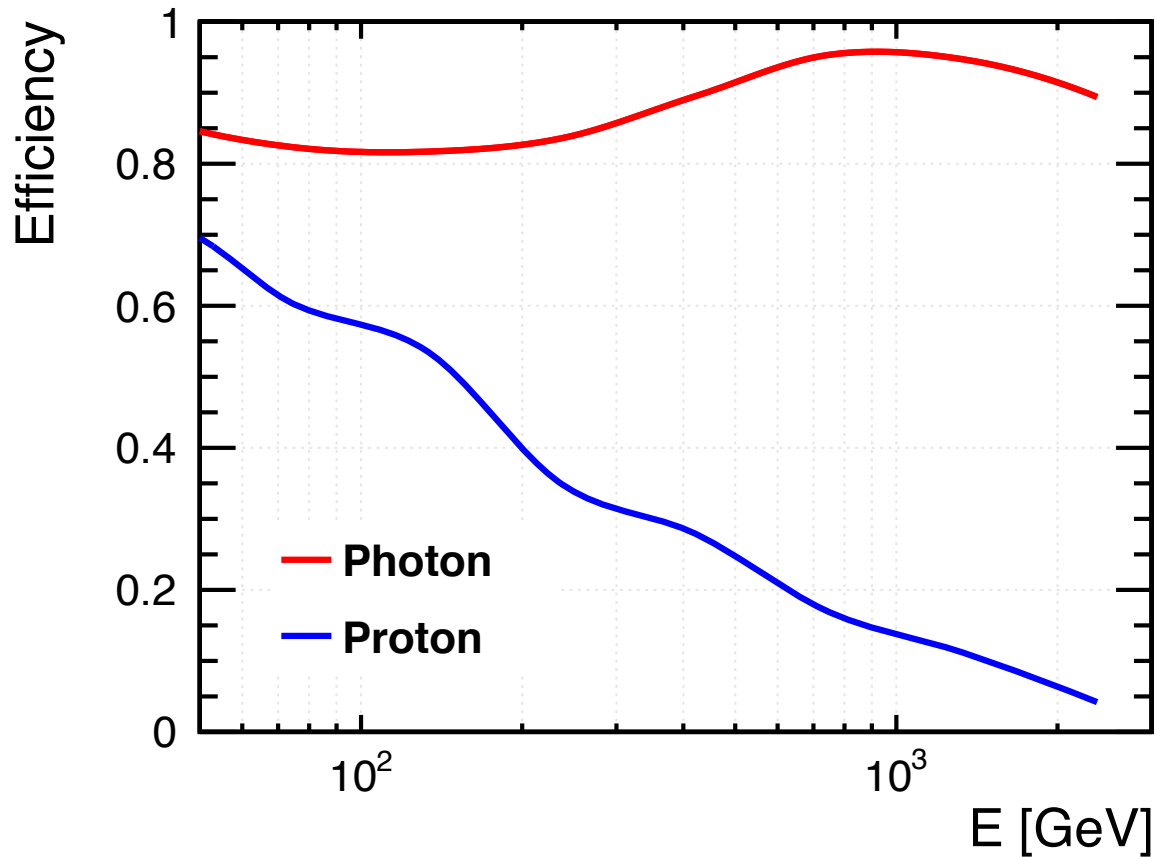


High-energy discrimination strategy



- ✧ Lateral distribution function (LDF)
 - ✧ LDF of gamma showers is more steep than the LDF of hadron showers

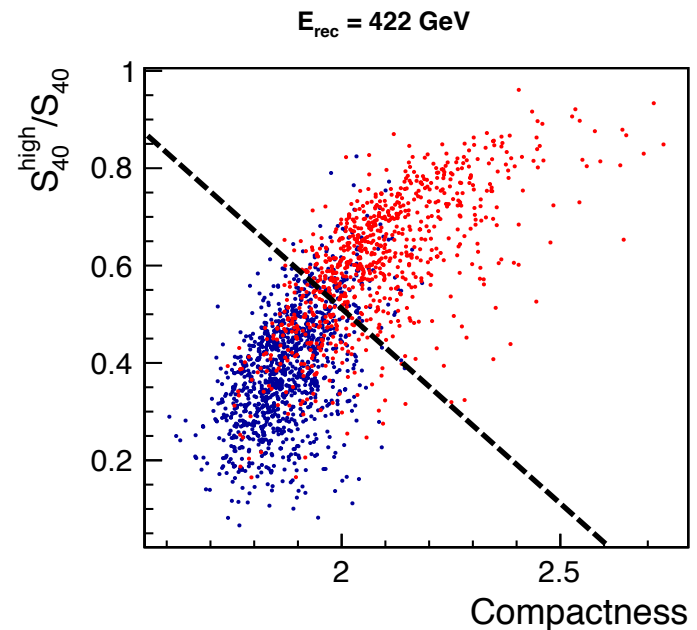
High-energy discrimination strategy



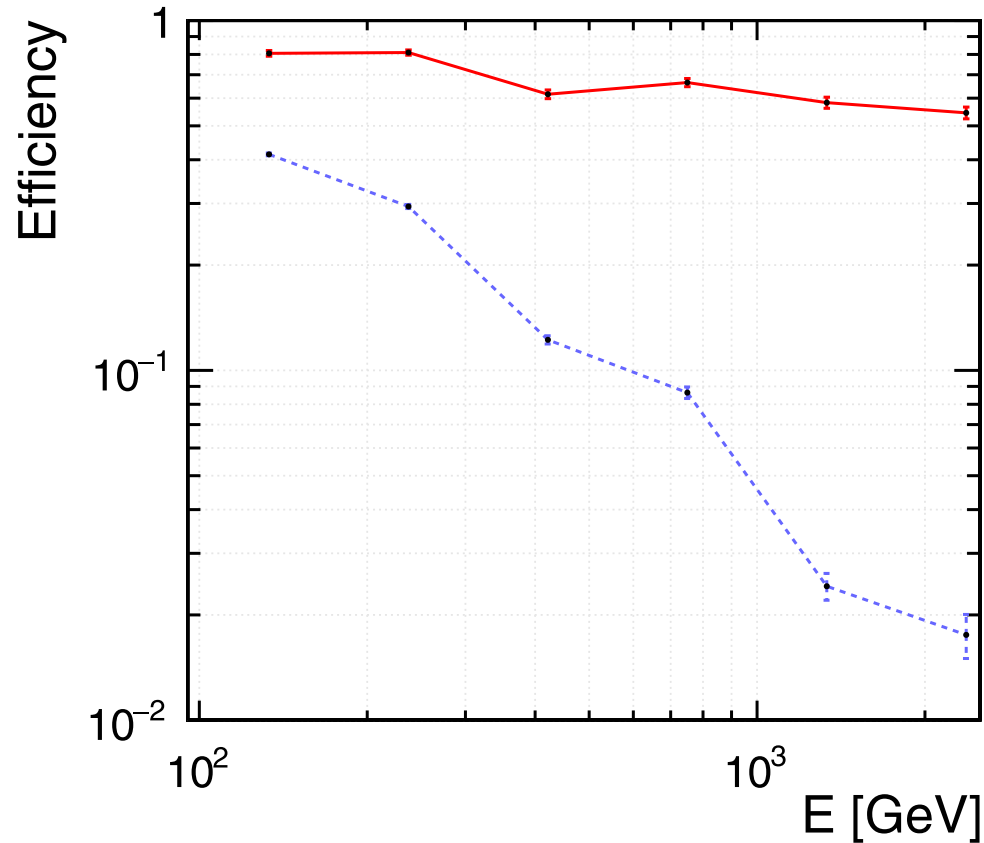
Shower **compactness** discrimination variable allows for a good background rejection which increases with energy

Combine information

- ✧ Fisher discriminant analysis to combine the two variables
 - ✧ $S_{40}^{\text{high}}/S_{40}$
 - ✧ Compactness
 - ✧ $S/\sqrt{B} = 6$ (at 2 TeV)
- ✧ LATTES MVA toolkit created
 - ✧ ROOT::TMVA
 - ✧ TinyXML
 - ✧ Python / C++
- ✧ Can easily be extended to:
 - ✧ add more discrimination variable
 - ✧ use higher-order methods BDT, ANN...



LATTES g/h discrimination

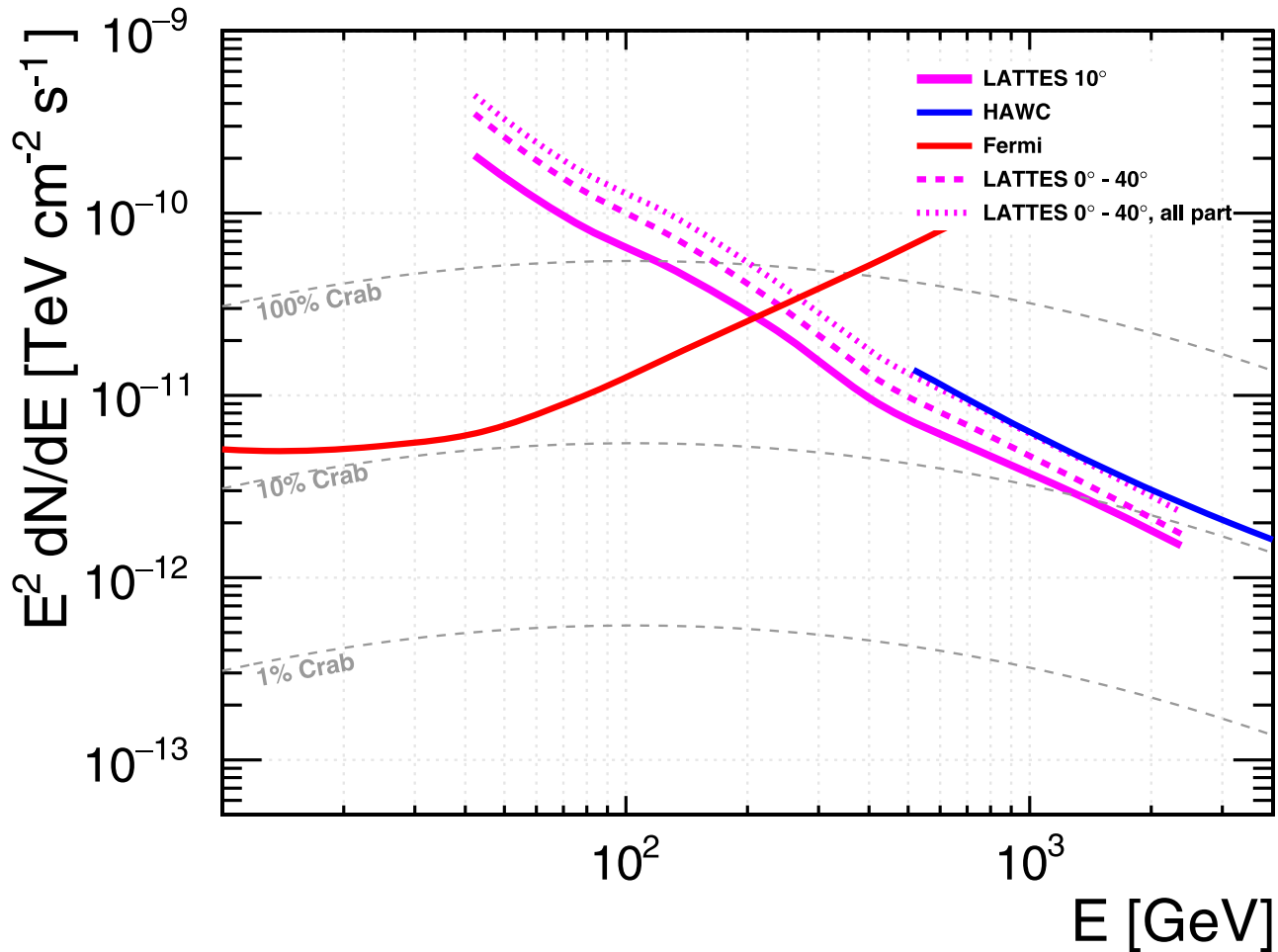


Although not optimized gamma/discrimination results are already very encouraging

LATTES expected performance

- ✧ Trigger and effective area
- ✧ Core reconstruction
- ✧ Energy reconstruction
- ✧ Geometry reconstruction
- ✧ Gamma/hadron discrimination
- ✧ **Sensitivity to steady sources**

Sensitivity to steady sources

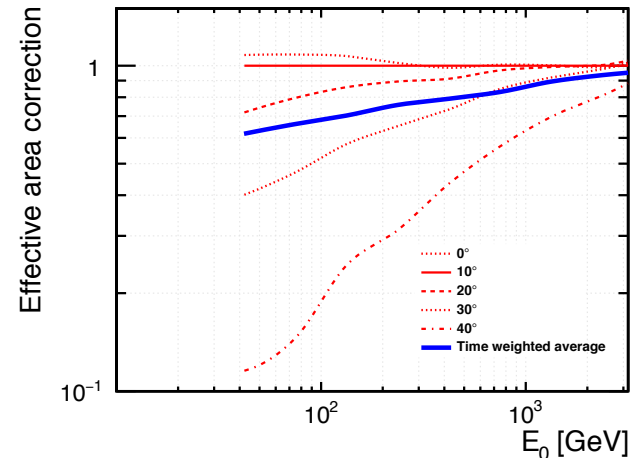
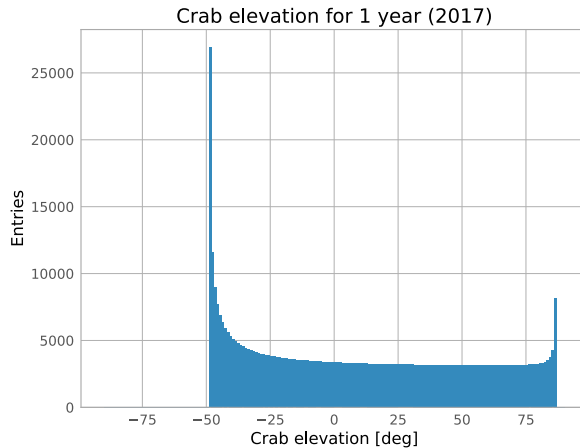
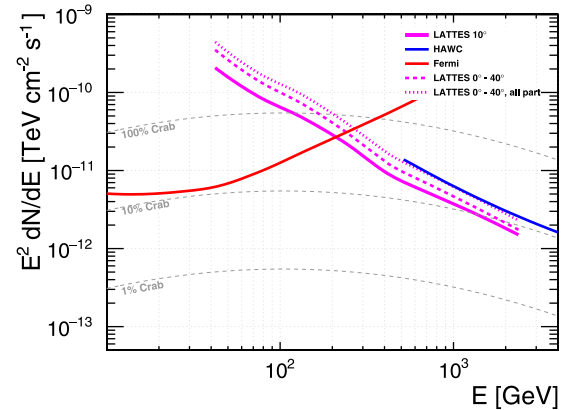


✧ *Full line*: full MC calculation for a source at 10 degrees in zenith

Sensitivity to steady sources

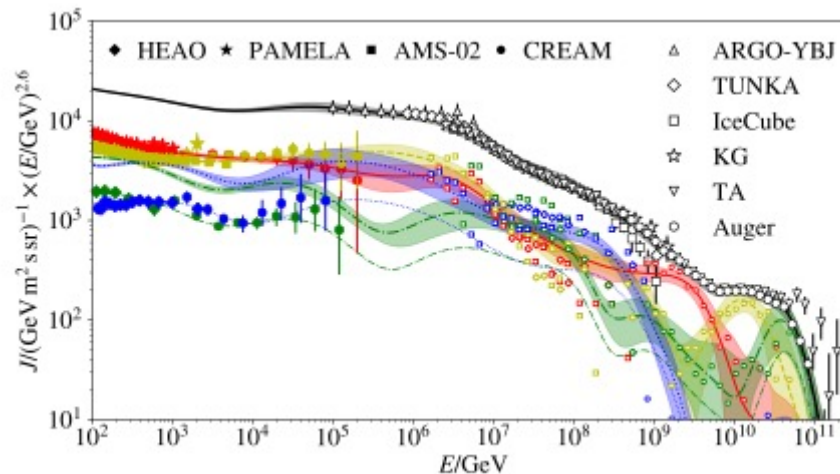
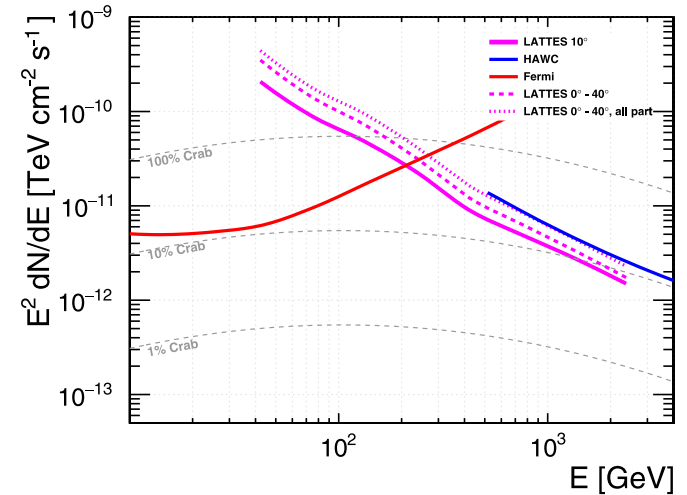
✧ *Dashed line*: Crab transit as seen by HAWC

✧ Degradation of effective area with zenith angle estimated from electromagnetic energy at ground



Sensitivity to steady sources

- ✧ *Dotted line: CR all-spectrum*
 - ✧ Additional elements (He, N, Fe...)
 - ✧ Assume that LATTES cannot distinguish gammas from irons



Dembinski et al (2017)

Summary

- ✧ LATTES shower reconstruction performance has been evaluated yielding very good results
 - ✧ Shower trigger (effective area)
 - ✧ Shower core reconstruction
 - ✧ Shower energy reconstruction
 - ✧ Shower geometry reconstruction
 - ✧ Gamma/hadron discrimination
- ✧ LATTES capabilities are far from being fully explored
 - ✧ Possible improvements already identified
 - ✧ Sparse array to veto far away high-energy showers (main background source) - see Bernardo's Talk
 - ✧ Use RPC patterns to discriminate g/h
 - ✧ Better assess LATTES ability to reconstruct
 - ✧ Inclined showers
 - ✧ Heavier primaries induced showers
 - ✧ ...

Acknowledgements



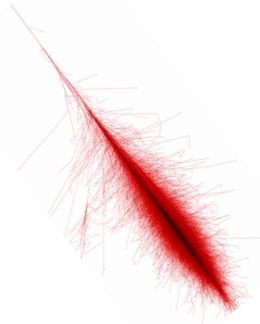
**REPÚBLICA
PORTUGUESA**



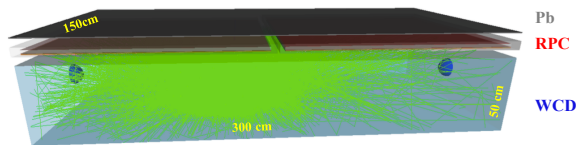
**TÉCNICO
LISBOA**

Backup slides

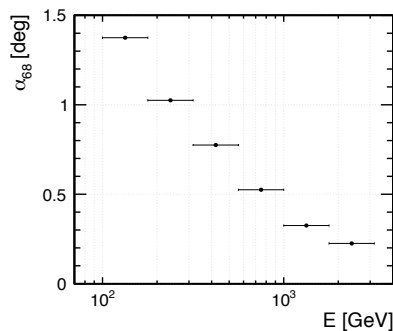
Towards LATTES sensitivity...



Shower simulation
(CORSIKA)



Detector simulation
(Geant4)

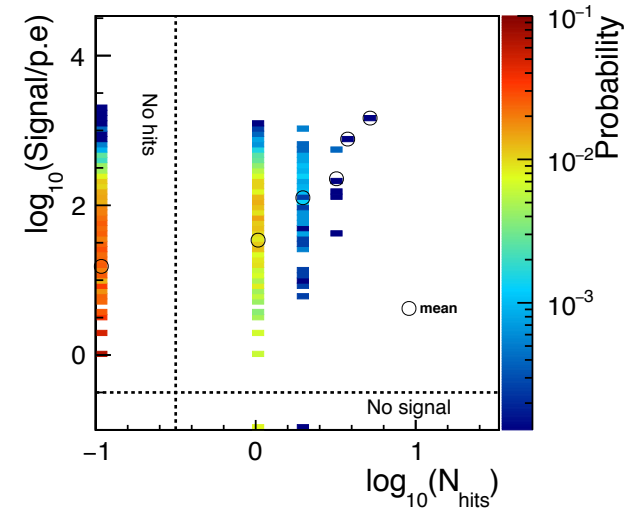


Shower reconstruction
(LATTESrec)

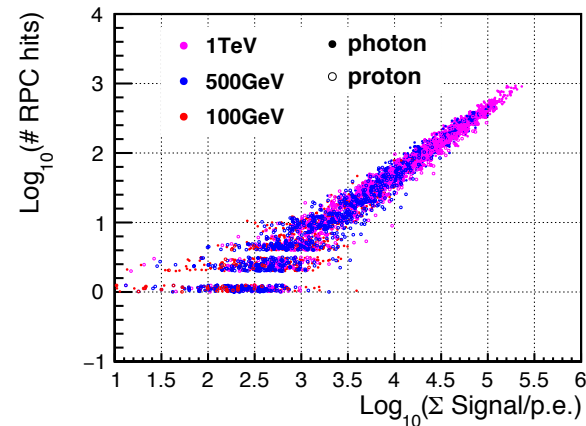
LATTES: a hybrid detector

- ❖ **Thin lead plate**
 - ❖ To convert the secondary photons
 - ❖ Improve geometric reconstruction
- ❖ **Resistive Plates Chamber**
 - ❖ Sensitive to charged particles
 - ❖ Good time and spatial resolution
 - ❖ Improve geometric reconstruction
 - ❖ Explore shower particle patterns at ground
- ❖ **Water Cherenkov Detector**
 - ❖ Sensitive to secondary photons and charged particles
 - ❖ Measure energy flow at ground
 - ❖ Improve trigger capability
 - ❖ Improve gamma/hadron discrimination

WCD vs RPC (station level)

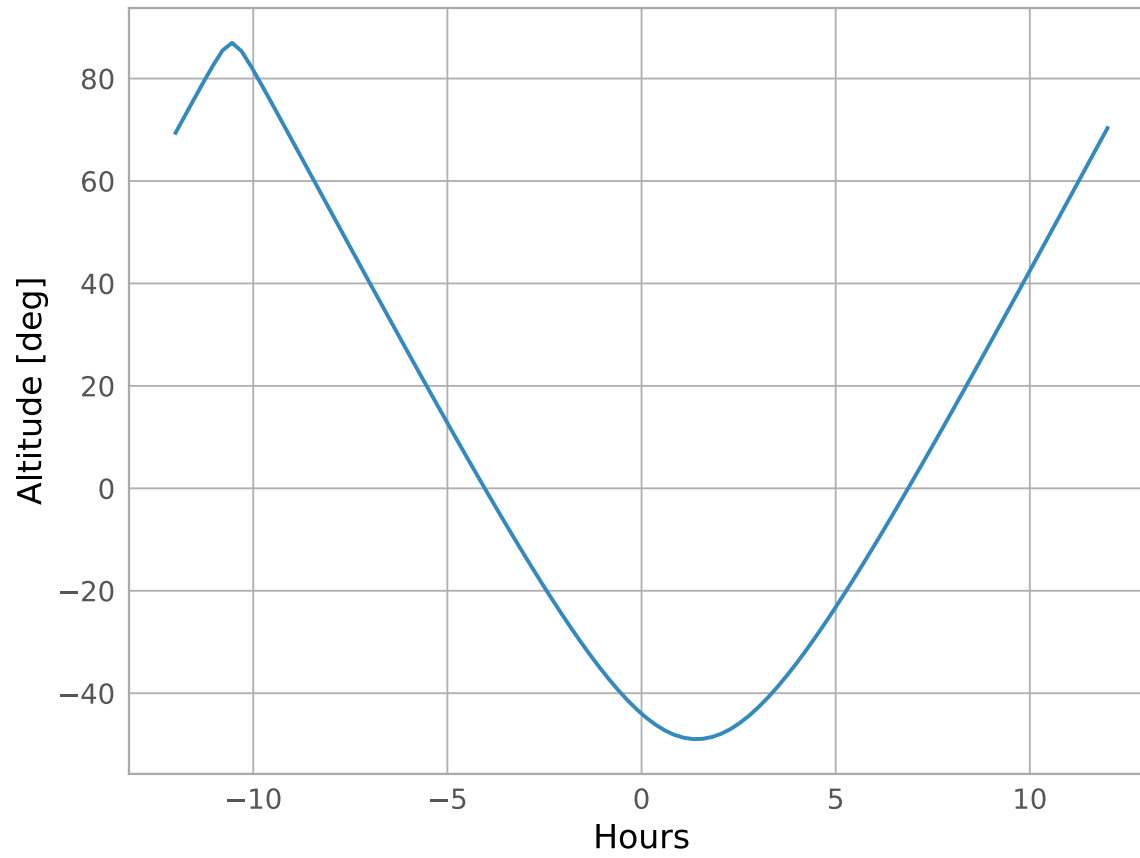


Complementarity



Inter-calibration

Crab



Accidentals contamination

Considering a time window D , the mean number of stations that randomly trigger within D is :

$$n_s = N_s \times R \times D$$

with N_s the # of stations in the array and R the single station trigger rate.

For LATTES $N_s = 3600$ and R was estimated from MC simulations to be of the order of 500 Hz; taking $D \sim 200$ ns yields :

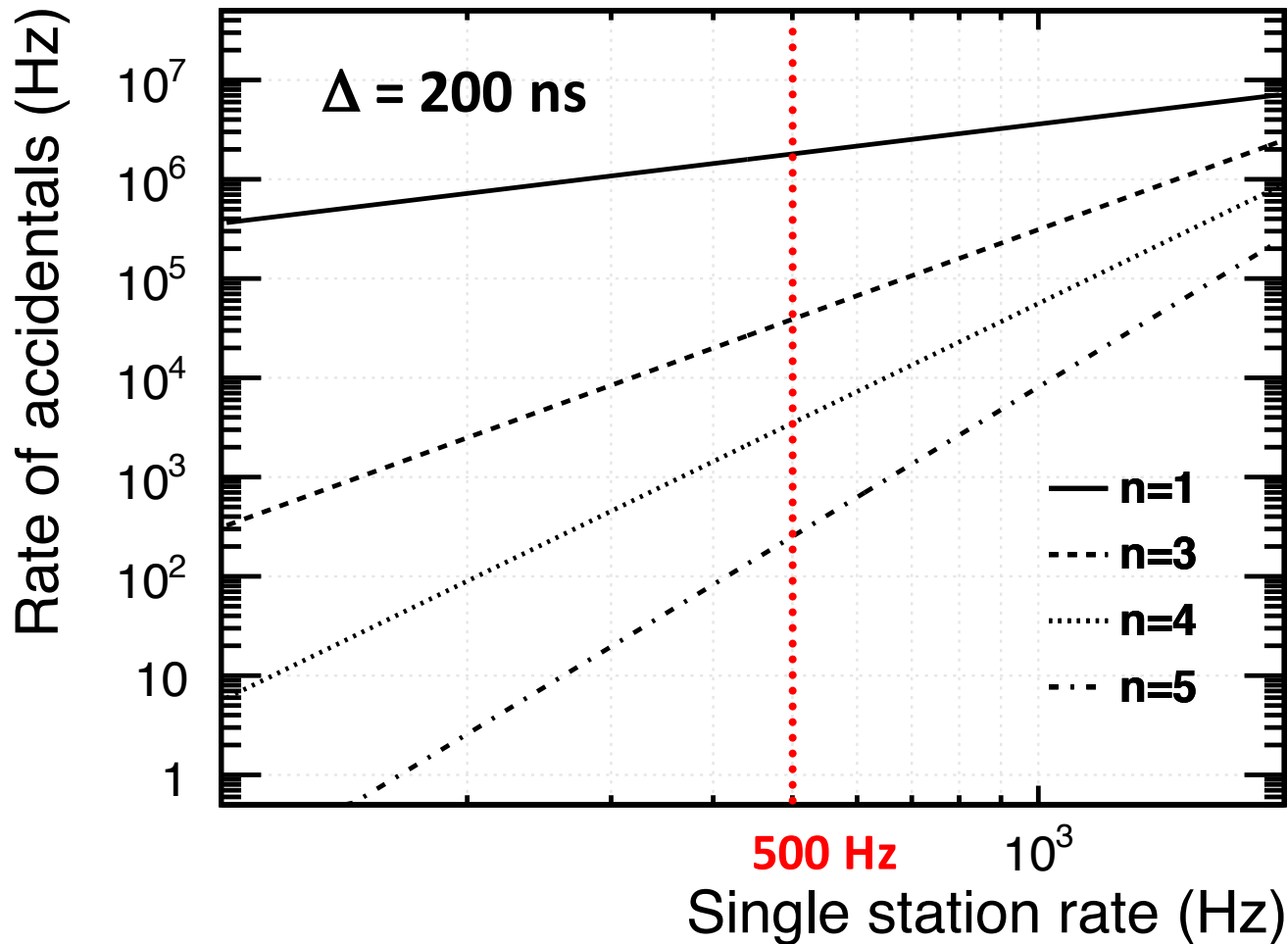
$$n_s \sim 0.4$$

to be compared with the minimum of stations required in a shower trigger, $n_s=3$.

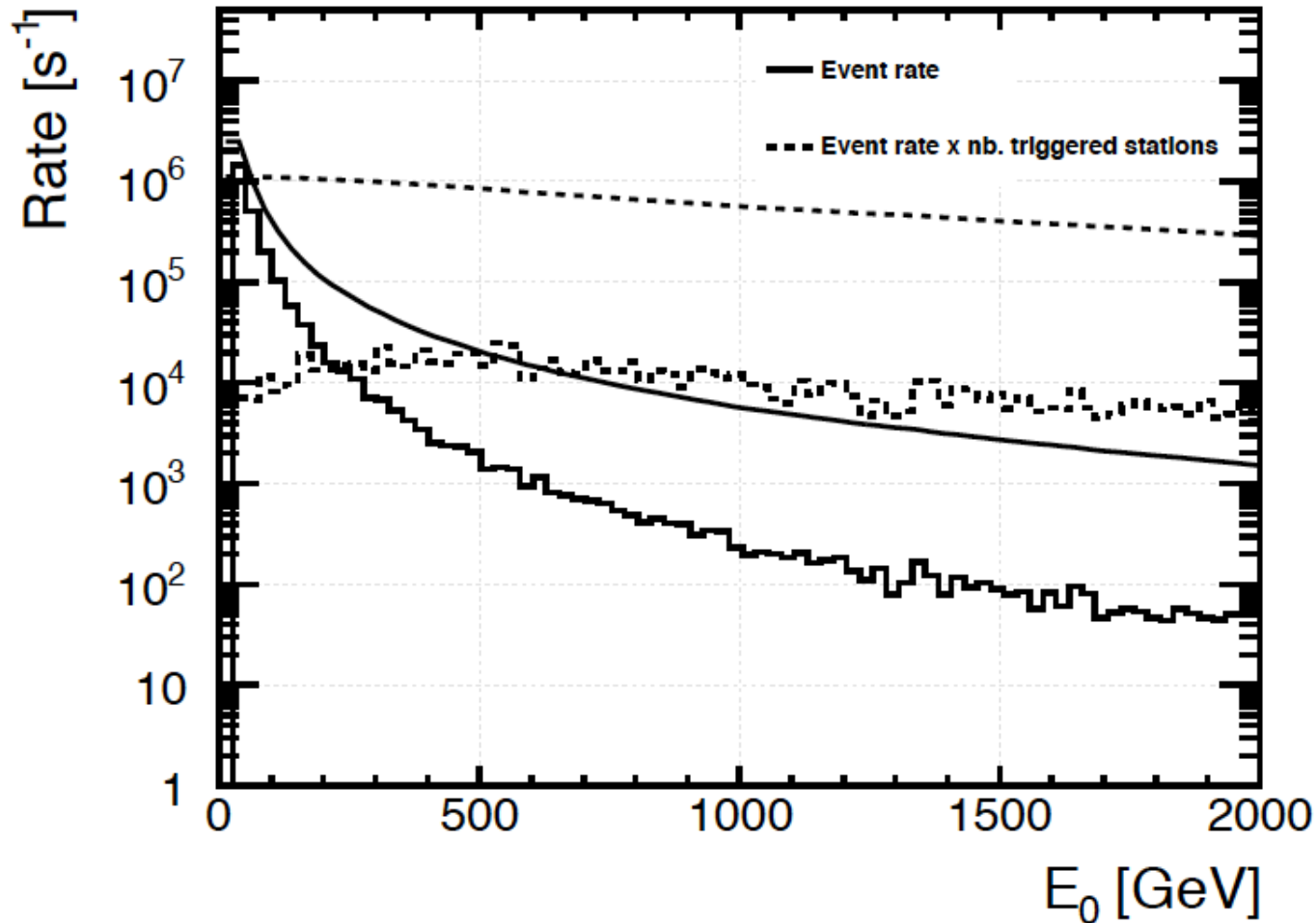
In any case a detailed MC simulation of the impact of the accidentals should be performed !

Random triggers

Rate of n-fold random coincidences in LATTES as a function of the single station trigger rate

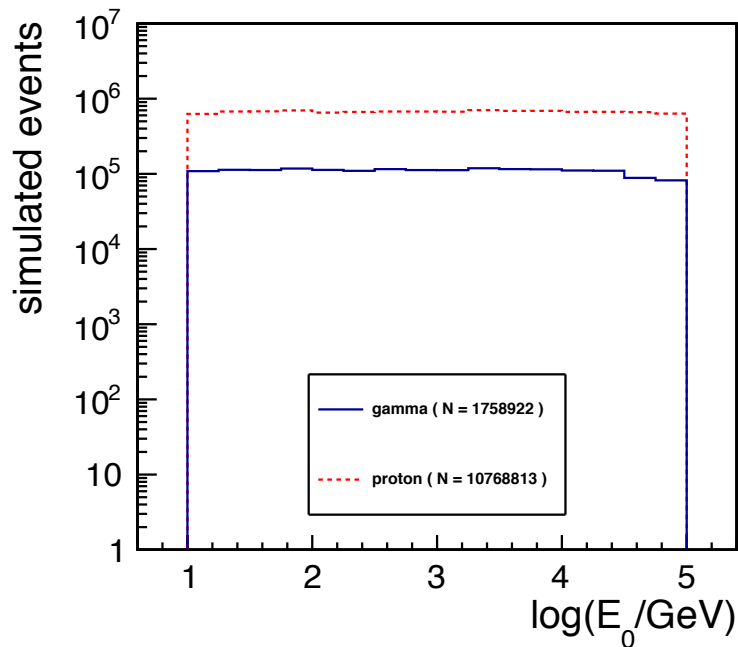


Cosmic rays and station trigger rate

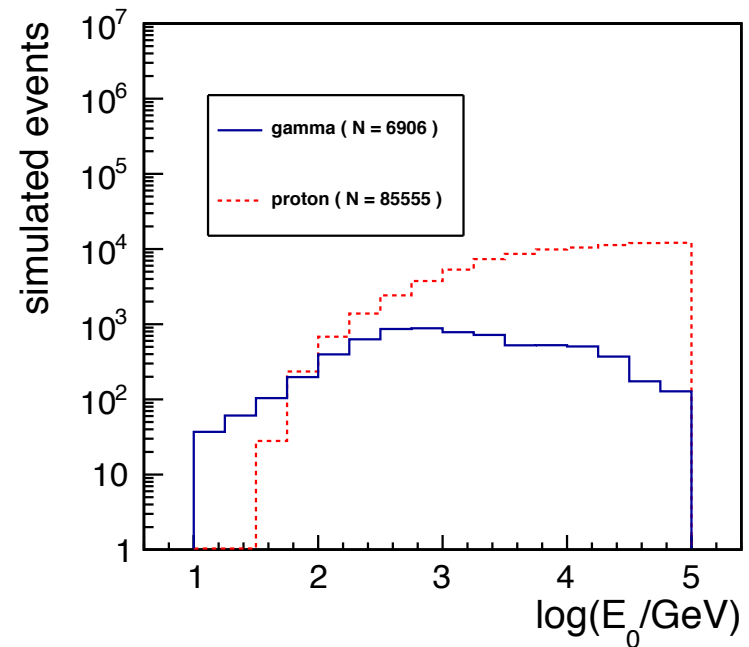


Reconstruction efficiency

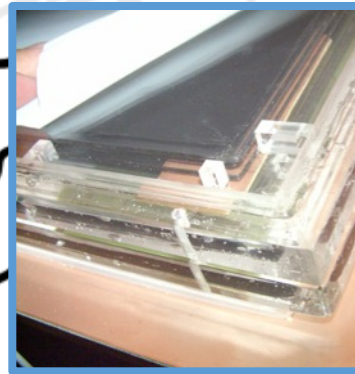
Before quality cuts



After quality cuts



Ongoing developments and tests on RPCs, electronics and read-out systems

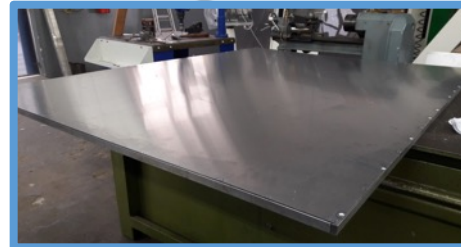


ents

DAQ Engineering prototype

RPC based muon hodoscope for precise studies of the Auger WCD

Construction and Assembling



RPCs in the field @ Auger



RPC hodoscope



R. Conceição