# Charged Current Neutral Pion Production at MicroBooNE

Joel Mousseau University of Michigan 12th International Workshop On Neutrino-Nucleus Interactions In the Few GeV Region L'Aquila, Italy October 2018.

## **Sources of Photon Production**

- At BNB energies, most originate from Decay of  $\Delta$  to  $\pi^0$ .
- Sources we classify as background:
  - Radiative delta decay  $(\Delta \rightarrow \gamma + N)$
  - π<sup>0</sup> from re-scattering (charge exchange)
  - Higher mass resonances.
- Further sources:
  - Anomalous (constrained by radiative delta decay).
  - Nuclear de-excitation (sub 10 MeV).



## **Photon Detection with LArTPCs**



- Electron-gamma separation one of the hallmarks of LArTPCs.
- Measure energy deposition along a particle's trajectory.
- Gammas pair-produce into two electrons, dE / dx. profile follows 2 electrons.
- Challenge is reconstructing showers from gammas

## CC $\pi^0$ Production with Neutrinos



#### ANL

#### MINERvA

MiniBooNE

- Many energy ranges and target media explored.
- Nuclear effects expected to scale as A<sup>2/3</sup> cross-section as N.
- Argon largest nucleus measured to date!

## CC $\pi^0$ Production with Neutrinos



- Focus on MiniBooNE: same beam line as MicroBooNE.
- MiniBooNE signal:  $v_{\mu}$  + CH  $\rightarrow \mu^{-}$  +  $\pi^{0}$  + *plus no other mesons.*
- MicroBooNE:  $v_{\mu}$  + CH  $\rightarrow \mu^{-}$  +  $\pi^{0}$  + X

## **Event Selection**



## **Event Selection**



## **Charge Scale Calibration**



- Use a combination of cosmic and neutrino induced muons to calibrate dQ/dx and dE/dx.
- Gross correction of position dependent detector response.
- Performed on collection plane only.



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See MicroBooNE

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## **Event Selection**





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## **Shower Reconstruction**



- Threshold for distinguishing track / shower hits about 50 MeV.
- Consequence of high purity, and track-nature of low energy showers.
- Results in a lower efficiency of reconstructing low-energy showers, but high purity.

## **Event Selection**

Trigger	<ul> <li>Given low efficiency of second shower, we split the analysis into a <i>single</i> shower and <i>two</i> shower selection.</li> </ul>
Tracking Reco.	Cosmic Rejection Shower Recc
<ul> <li>At least 1 shower is higher efficiency.</li> <li>At least 2 shower used as cross check.</li> </ul>	Ar MicroBooNE Simulation Preliminary $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} + Ar \rightarrow \mu + (1 \pi^{0} \rightarrow \gamma \gamma) + X$ $v_{\mu} +$

## **Single Shower Selection**



- Measure conversion length of single photon to ensure it is a photon.
- Exclude first bin (large number of non-photons).
- Measured conversion length agrees with Simulation.
- Single shower Efficiency (Purity) 17% (53%).

## **Single Shower Selection**



- Final set of selection cuts removes cosmic background
- If two showers in the event, the higher energy shower is selected.
- Overlap exists between 1 and 2 shower samples.
- Each cut ~85% efficient.

## **Two Shower Selection**



- Using two showers with similar cuts correct for missing energy from clustering and hit-thresholding using MC.
- Plot diphoton invariant mass, confirm  $\pi^0$  mass in data and MC.
- Two shower efficiency (purity): 6% (64%).

## **Single Shower Backgrounds**



- Remaining backgrounds are cosmic, and other resonant events contributing to the single π<sup>0</sup> production.
- Estimate these backgrounds directly from GENIE (45% of events).

## **Systematic Uncertainties**

Type	% Error	Affected Measurement
Flux	16%	Flux division, Background Estimation
<b>Cross-Section</b>	17%	Background Estimation Efficiency Correction
<b>Detector Modeling</b>	<b>21%</b>	Background Estimation Efficiency Correction
TOTAL	31%	

## **Systematic Uncertainties**

Туре	% Error	Affected Measurement
<b>Detector Modeling</b>	<b>21%</b>	Background Estimation Efficiency Correction

• Major sources of detector modeling uncertainty:

- Induced charge on neighboring wires.
- Diffusion of charge as it propagates along the drift direction.
- Modeling of scintillation light.
- Improvements will come from:
  - Newer MicroBooNE simulation and calibration.
  - Improved shower reco. to detect lower energy showers lead to higher efficiency.

#### **Results (At Least 1 Shower)** MicroBooNE Preliminary 1.62e20 POT 3.5 GENIE Default + Emp. MEC 25 3.0 **GENIE** Alternative Flux (Arbitrary Scale) (σ)<sub>Φ</sub> (10<sup>-38</sup> cm<sup>2</sup>/Ar) 1.5 1.0 $\sigma(E_{\nu}) (10^{-38} \text{ cm}^2/\text{Ar})$ 20 Genie **Alternative** 15 **Uses BS** and 10 hA 2014 5 0.5 0.0 0 One Shower 500 1000 1500 2000 2500 3000 0 Selection Neutrino energy (MeV) $\mathrm{cm}^2$ $\sigma^{ u_{\mu}\mathbf{C}\mathbf{C}\pi^{0}}$ $= (1.94 \pm 0.16 \text{ [stat.]} \pm 0.60 \text{ [syst.]}) \times 10^{-38}$

#### First ever measurement on Ar.

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 $\operatorname{Ar}$ 

## Results



- Compare our result on Argon, to ANL and MiniBooNE results on C and D.
- Starting to probe differences between how different models predict A scaling.
- Currently, lack sensitivity to differences in FSI modeling.

## **Future Measurements**

- Large investment in improving MicroBooNE's detector simulation and reconstruction.
- New image recognition, machine learning techniques promise better shower reconstruction, ability to detect lower energy showers
- Much more data in further MicroBooNE runs (x8 more).
- Enables a differential measurement of  $\pi^0$  in  $\pi^0$  and  $\mu^-$  variables.





## Conclusions

- MicroBooNE has performed a world's first measurement of neutrino induced π<sup>0</sup> production on argon.
- Measured with both single and two photon sub-samples.
  - Both measurements consistent.
  - Select single photon as primary result due to larger efficiency.
- Currently see reasonable agreement with GENIE's A scaling.
- Sensitivity limited by detector model and reconstruction.
  - Improving detector model, including signal modeling.
  - Improving reconstruction, including the ability to detect low energy showers.

# Thank you for Your Attention!



### All MicroBooNE Public Notes Available Here: http://microboone.fnal.gov/public-notes/

## **Backup Slides**

## **Energy Correction and Mass Peak**



- Corrections applied:
  - Add energy from hits below threshold.
  - Add energy from clusters mis-id as track or cosmic.
- Both corrections derived from MC.
- Do *not* correct for uncontained clusters.

## **One and Two Shower Comparison**

