

# Impact of cross section modeling on NOvA oscillation analyses

on behalf of the NOvA collaboration



Jeremy Wolcott  
Tufts University

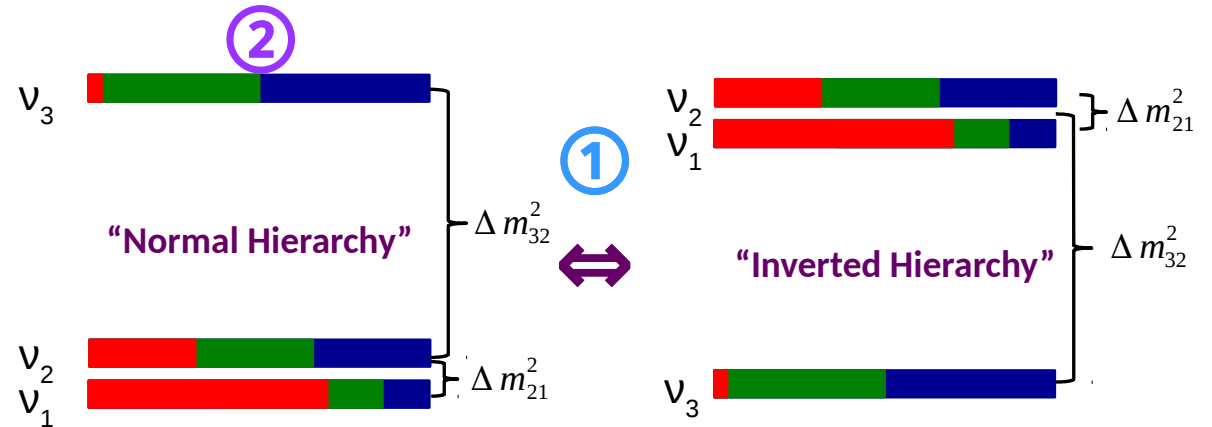
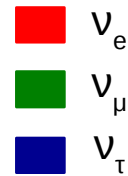
October 17, 2018  
NuInt 2018 (L'Aquila, Italy)



# NOvA: $\nu$ oscillation physics

① How are the mass eigenstates ordered?

② Is there a symmetry governing mixing between  $\nu_\mu$  and  $\nu_\tau$ ?

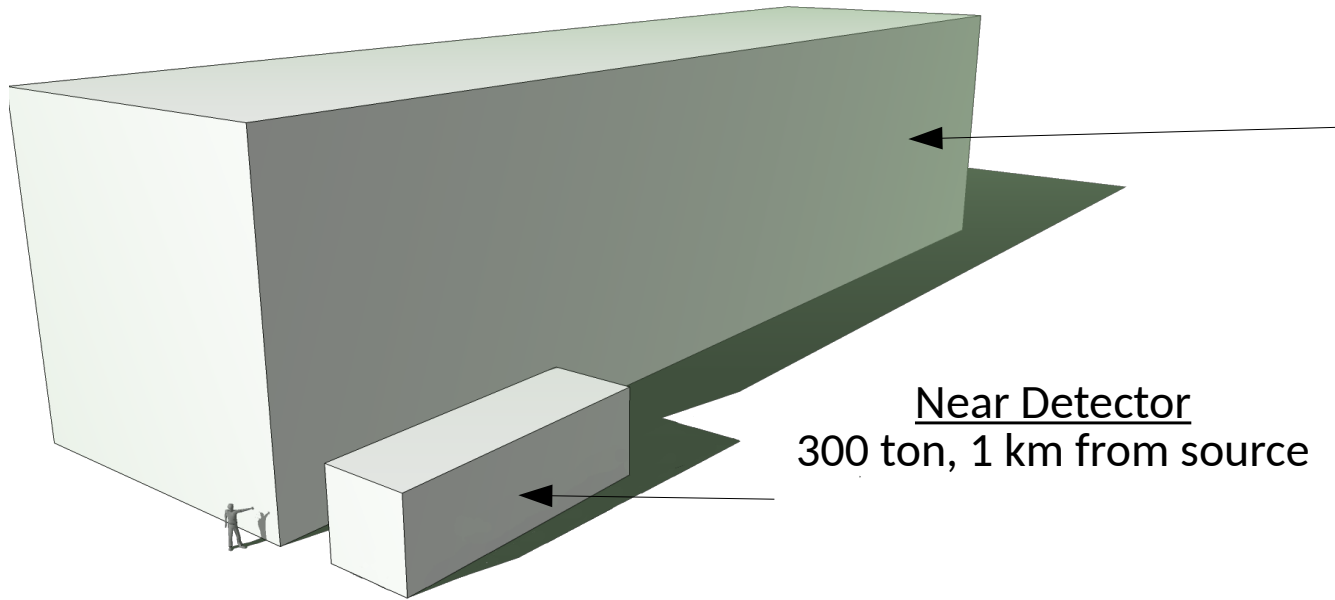


$$\mathbf{U} = \begin{pmatrix} 1 & 0 & \textcircled{2} 0 \\ 0 & \cos\theta_{23} & \sin\theta_{23} \\ 0 & -\sin\theta_{23} & \cos\theta_{23} \end{pmatrix} \begin{pmatrix} \textcircled{3} \cos\theta_{13} & 0 & \sin\theta_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -\sin\theta_{13}e^{i\delta} & 0 & \cos\theta_{13} \end{pmatrix} \begin{pmatrix} \cos\theta_{12} & \sin\theta_{12} & 0 \\ -\sin\theta_{12} & \cos\theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

③ Is there CP violation in leptons?

Measuring key parameters in oscillation physics

# NOvA: design considerations

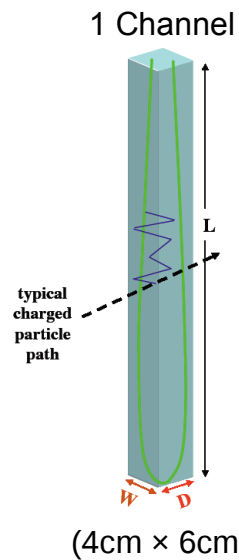


Far Detector  
14 kton, 810 km from source

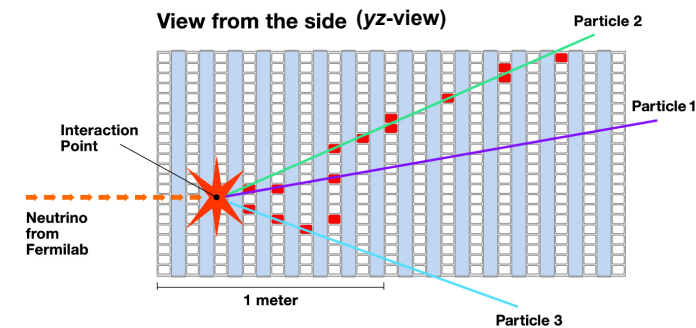
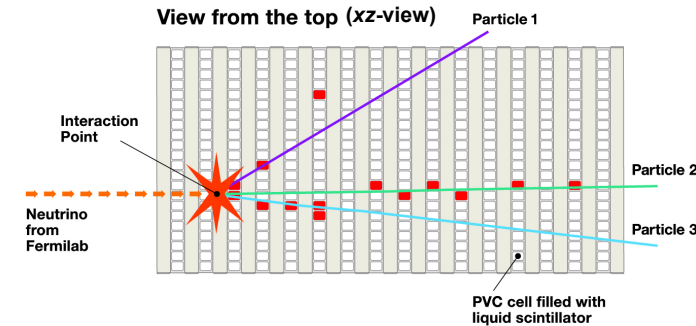
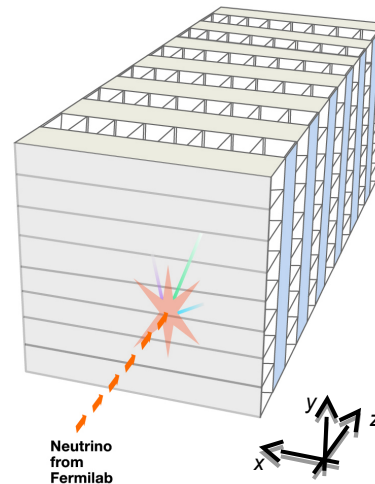
Near Detector  
300 ton, 1 km from source

**Functionally identical detectors**

**Sampling calorimeter detectors**

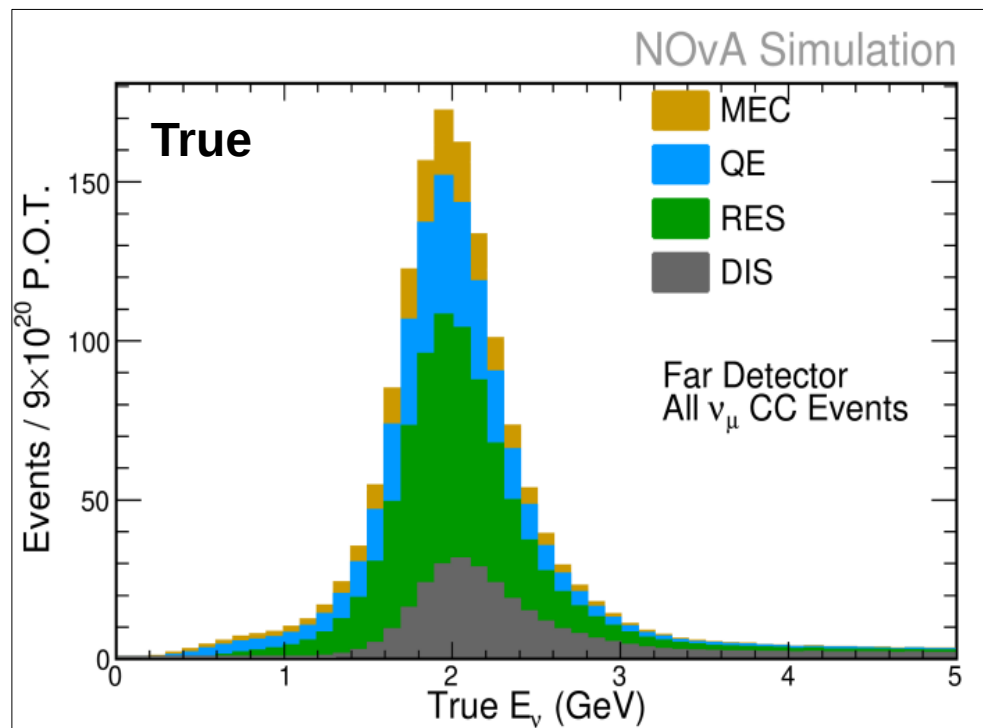


stacked in planes



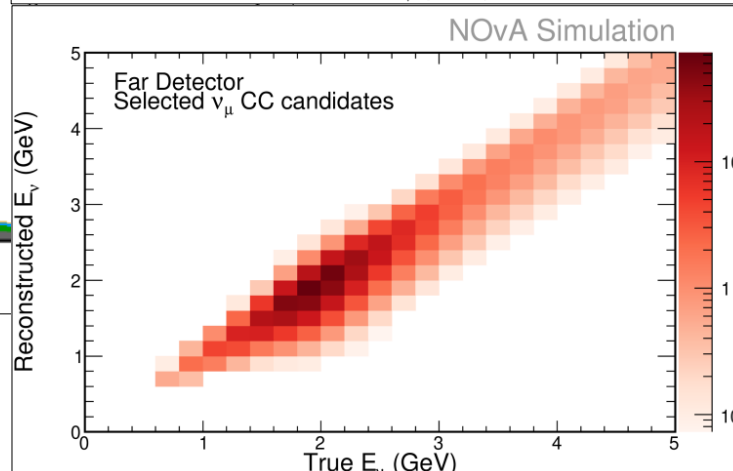
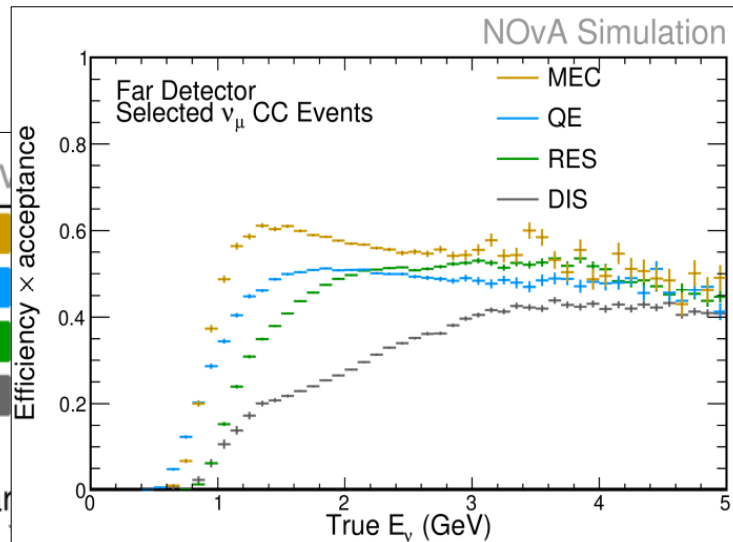
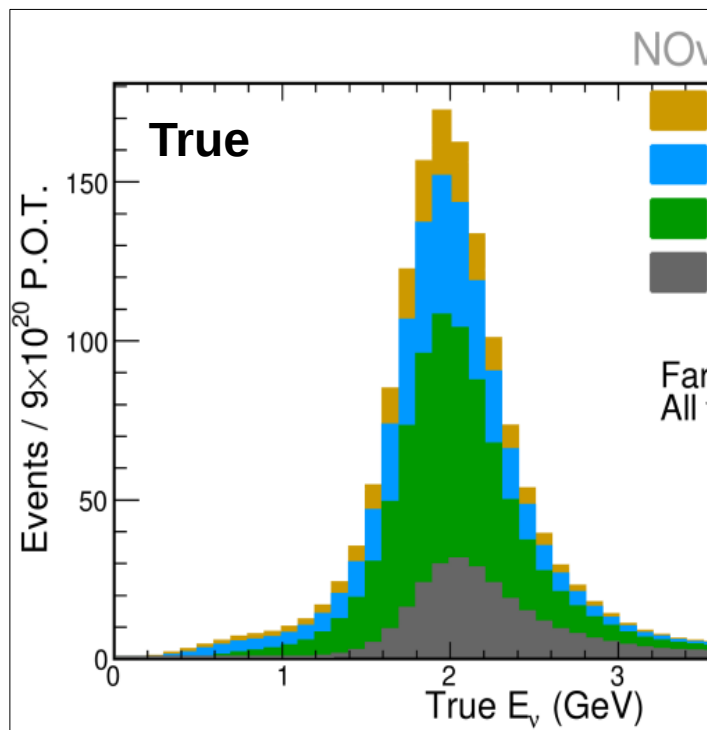
# How cross sections enter the story: energy reconstruction

- $P(\nu_\alpha \rightarrow \nu_\beta)$  depends on  $E_{\text{true}}$ , but detectors measure  $E_{\text{reco}}$
- Detectors/reconstruction have different sensitivities to different processes, which have different  $E_{\text{true}} \leftrightarrow E_{\text{reco}}$



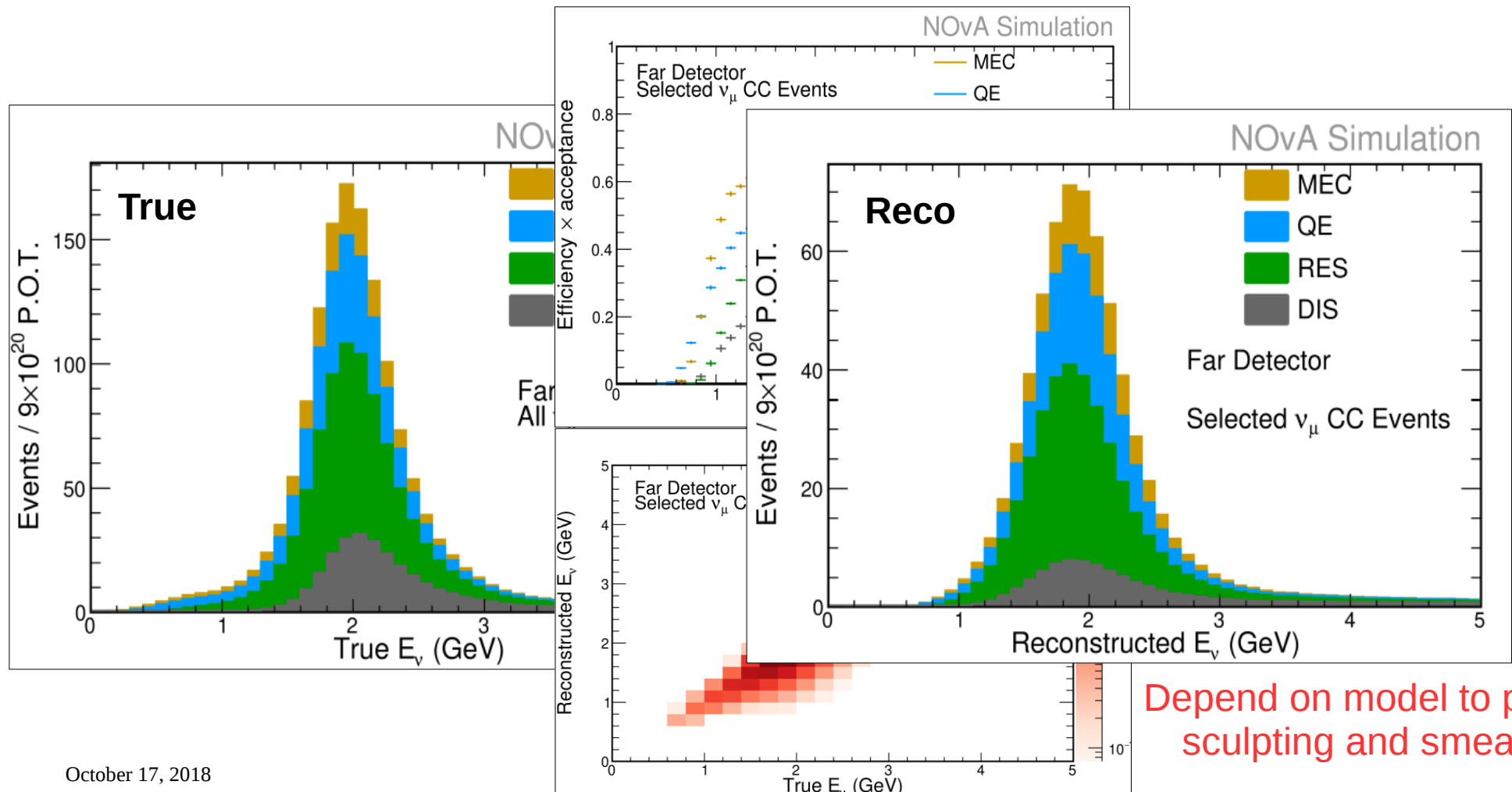
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# Model building

We adjust GENIE 2.12 (base model) based on **theory work and other experiments' data:** thanks to the hard work of many in this room!

(fuller discussion in J. Wolcott, [FNAL Neutrino Seminar, Apr. 23 2018](#); paper forthcoming)

## Free-nucleon model

Use **dipole axial FF** with

$$M_A^{QE} = 1.04 \pm 0.05 \text{ GeV}$$

based on error-weighted mean we calculated from bubble chamber data

[collected in PRD **93**, 113015]

(GENIE default: 0.99 GeV)

Z-expansion in our future...



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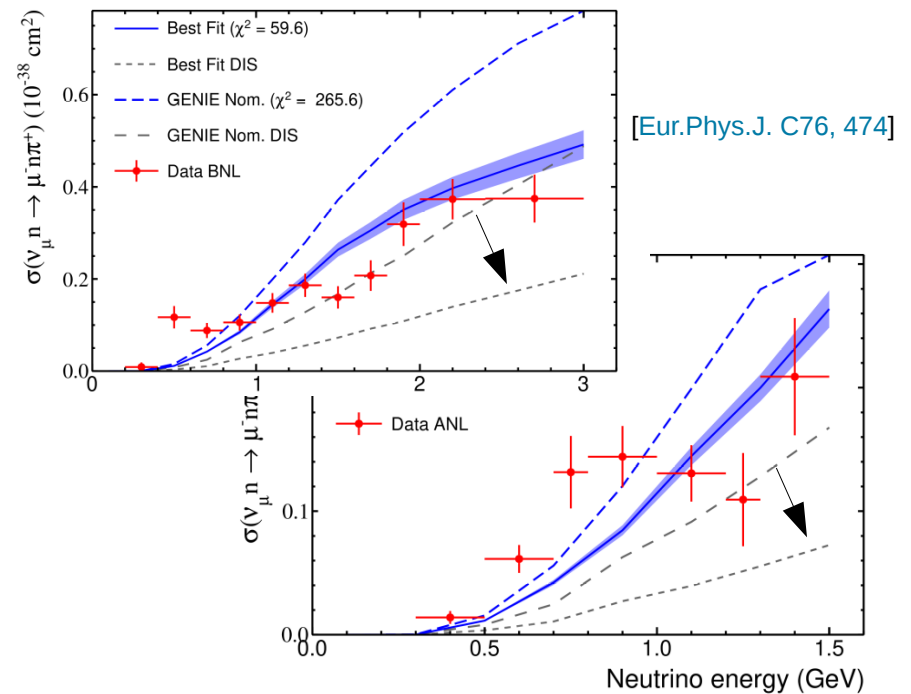
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Z-expansion in our future...



**Nonresonant  $1\pi^+$  production** from neutrons needs to be **reduced by ~50%** based on updated fits to free-nucleon data

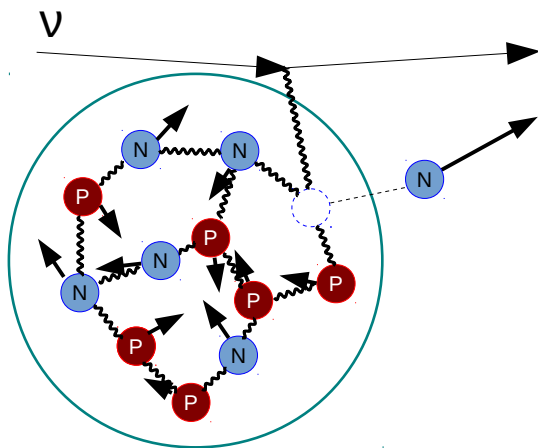


# Model building



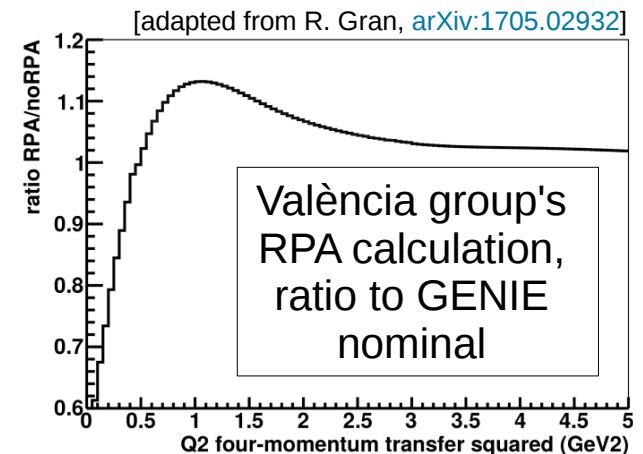
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**Effective nuclear “screening”** from collective excitations: treated with **RPA**.

We use Valencia group calculation for QE; **also use  $Q^2$  shape for RES** based on suppression noticed in external and NOvA data



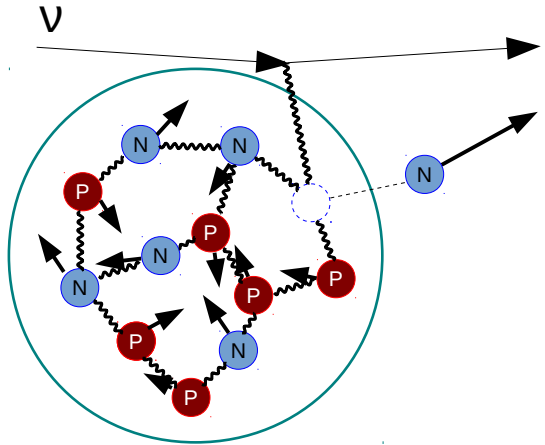
## Nuclear model

# Model building



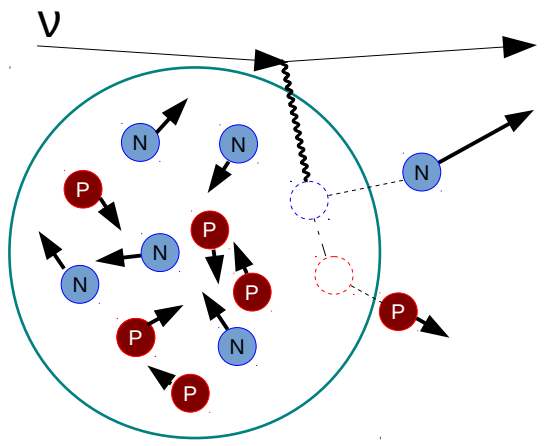
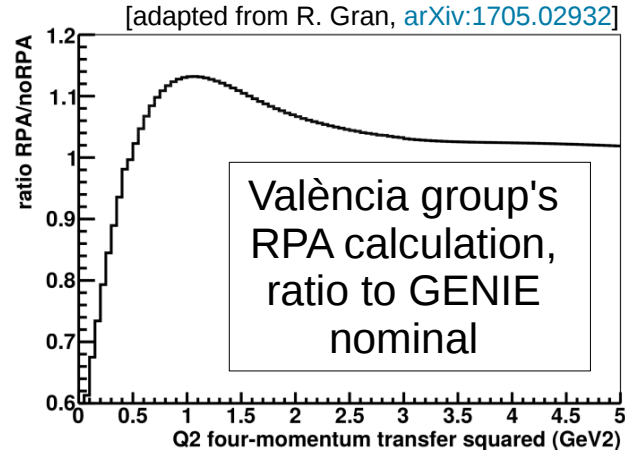
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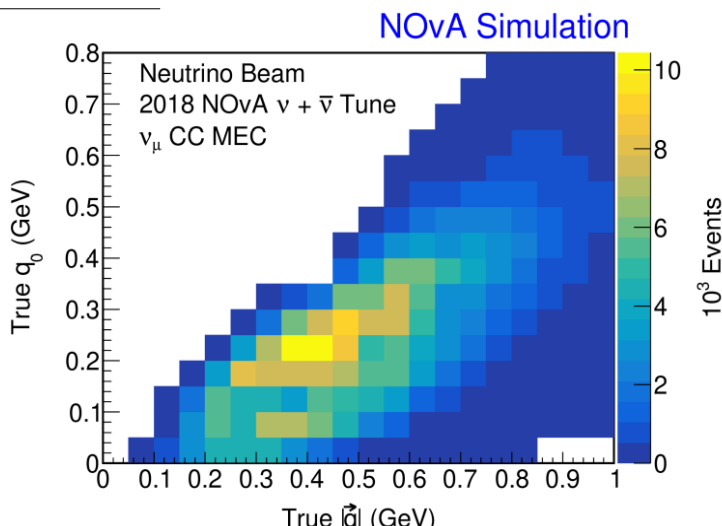
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**Multinucleon knockout (2p2h)**

We enable GENIE "Empirical MEC", retune it based on our data. Uncertainties from fits under different assumptions.

**Nuclear model**



# Evaluating cross section uncertainties

Depend heavily on GENIE's reweight system...

## Primary process uncertainties

**QE:**  $M_A$ , Vector FF, Pauli supp...

**RES:**  $M_A$ ,  $M_V$ ,  $\Delta$  decay isotropy...

**DIS:** Bodek-Yang parameters, transition region (“non-resonant background” scale), ...

**COH:** Rein-Sehgal  $M_A$ ,  $R_0$ , ...

## Final-state model (hA) uncertainties

**Nucleon, pion** elastic, inelastic, chg ex., abs. reaction probabilities

**Hadron mean free paths**

(~50 reweight knobs in all)

...and build custom knobs for our growing library of GENIE 'adjustments':

MEC model for **2p2h**  
( $q^\mu$  shape,  $E_\nu$  shape, nn/np composition)

**RPA-QE** (based on València treatment; histograms from R. Gran)

**RES- $Q^2$**  (conservative “on” vs “off”)

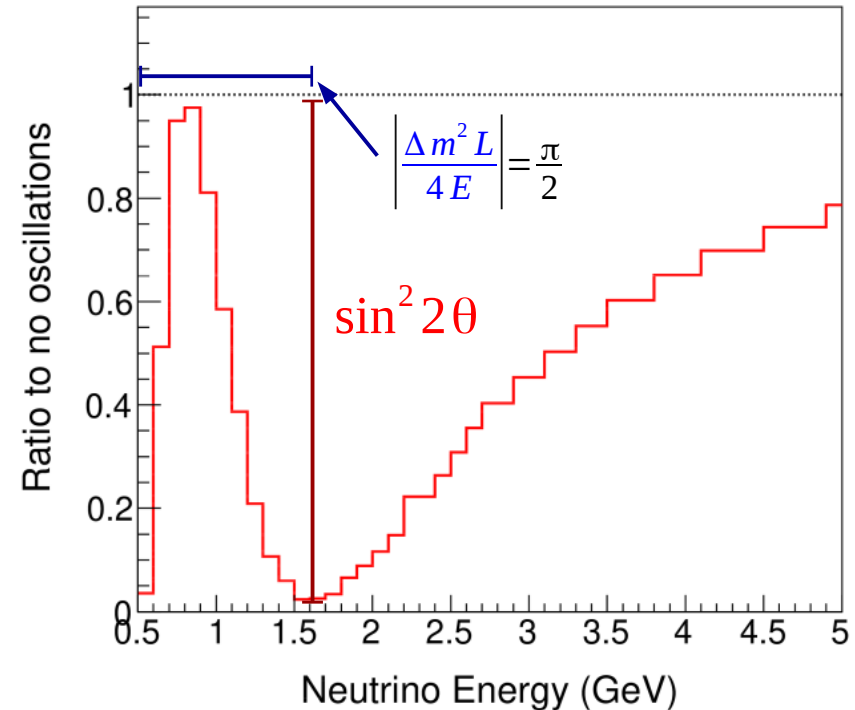
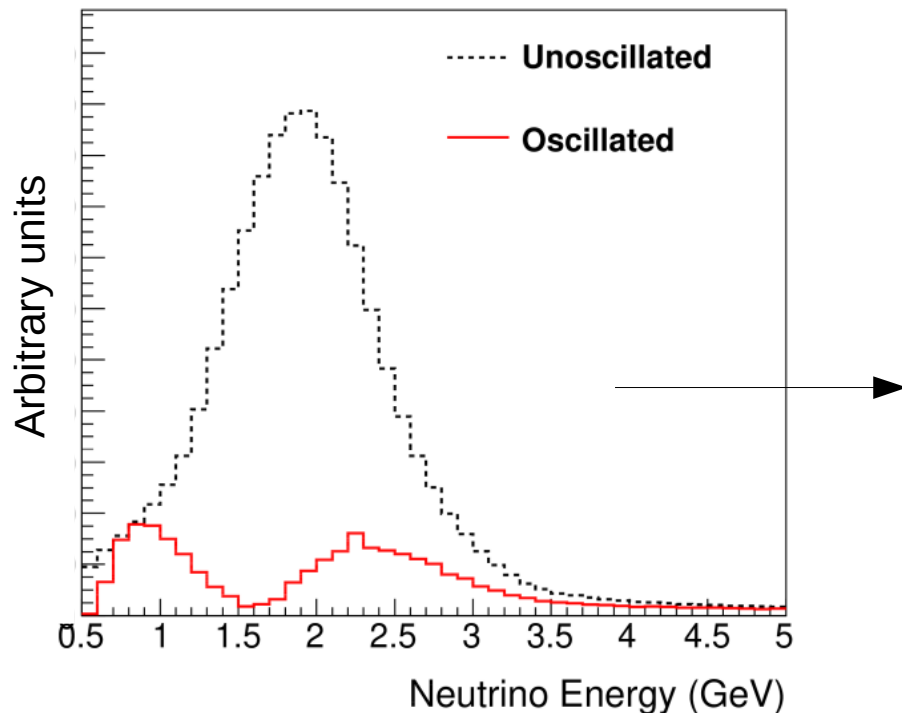
In practice:  
 $v_{\mu}$  disappearance  
[a worked example]

# $\nu_\mu$ disappearance

$$P_{\nu_\alpha \rightarrow \nu_\beta} \approx \sin^2 2\theta \sin^2 \left( \Delta m^2 \frac{L}{4E} \right)$$

How far away from the source you build your detector

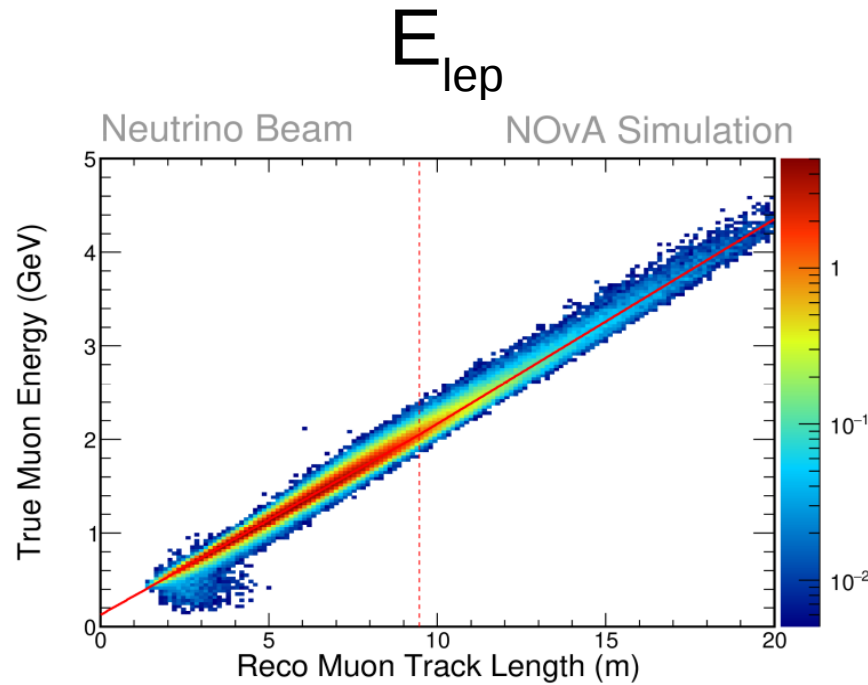
Energy spectrum of your neutrino beam



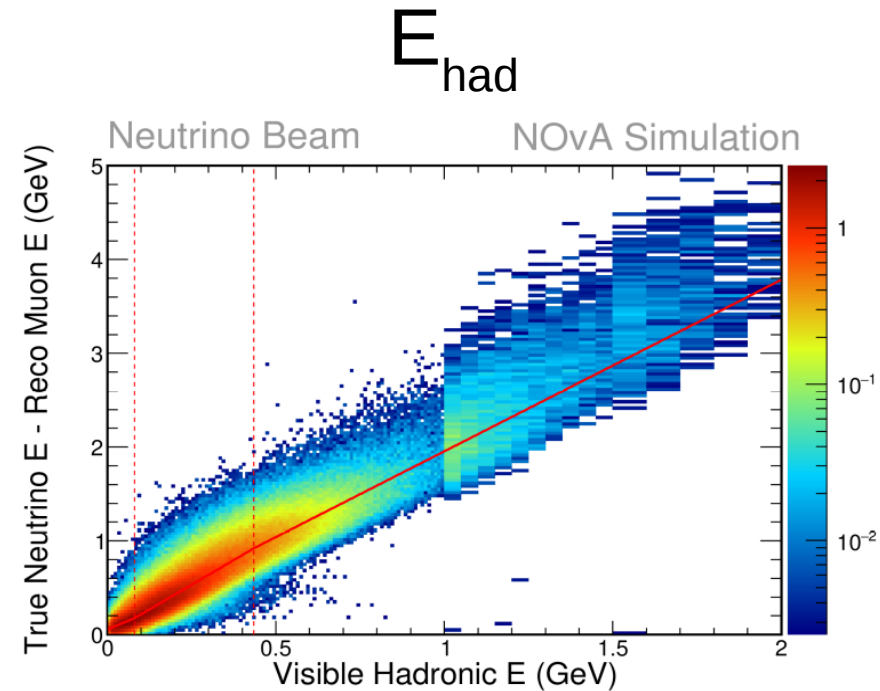
Goal: measure the location and strength of the “oscillation dip” relative to no-oscillations prediction

# $\nu_\mu$ disappearance: energy reconstruction

$$E_\nu =$$



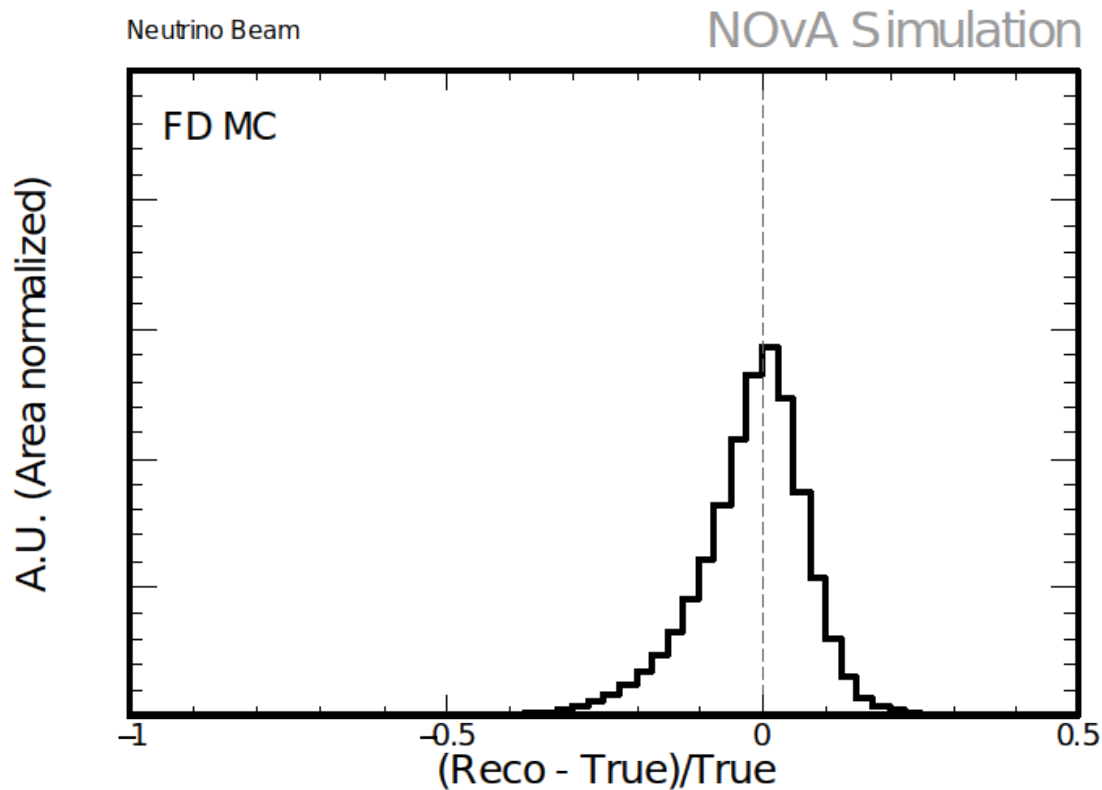
+



Calibrate muon track length to true  $E_\mu$ ,  
 then remaining visible energy to  
 (true  $E_\nu - \text{reco } E_\mu$ ).

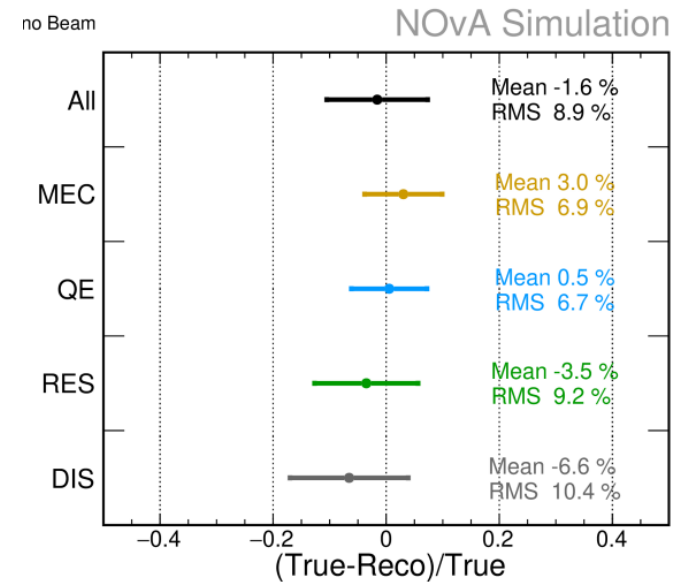
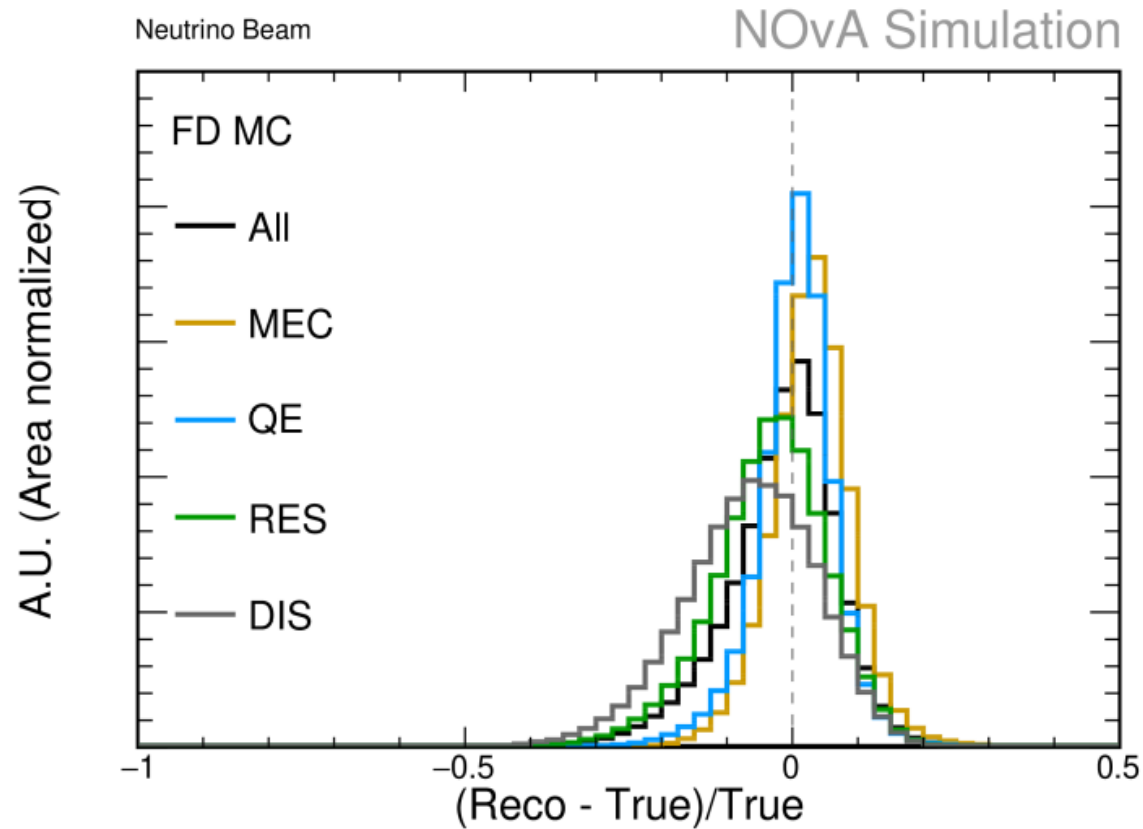
**Calorimetric (not kinematic) energy reconstruction**

# $\nu_\mu$ disappearance: energy reconstruction



**Nominal resolution  
on  $E_\nu \sim 9\%$ .**

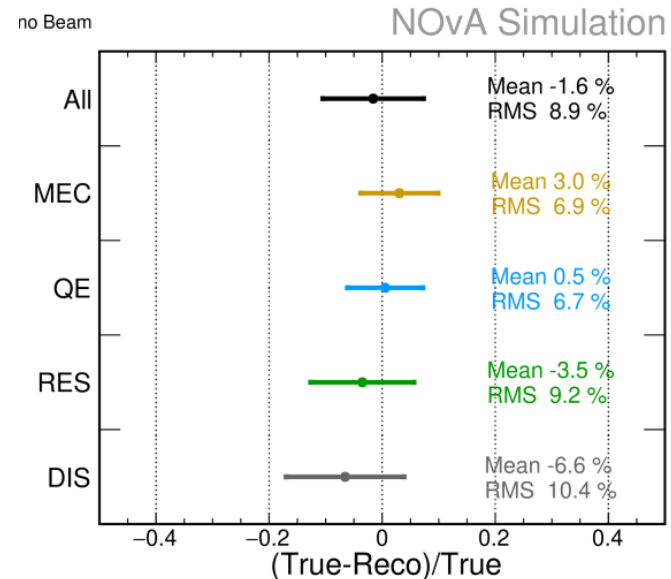
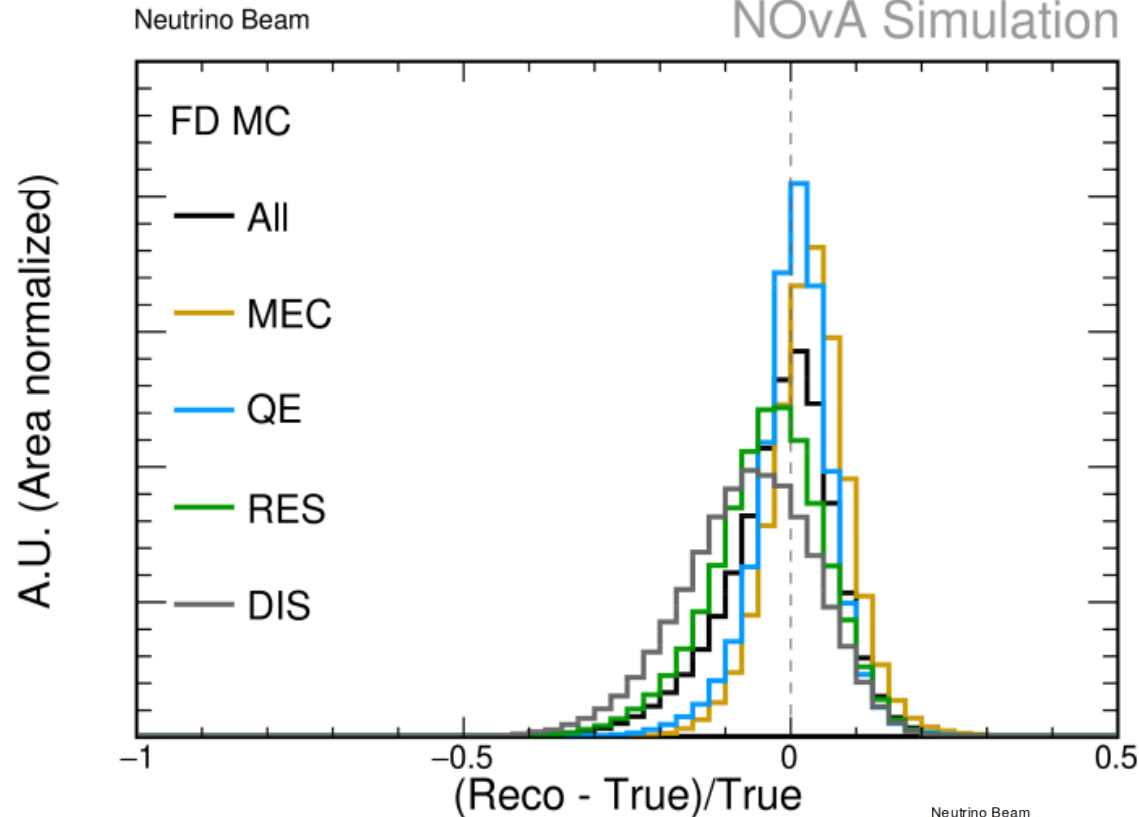
# $\nu_\mu$ disappearance: energy reconstruction



**Nominal resolution  
on  $E_\nu$  ~ 9%;  
different by reaction**

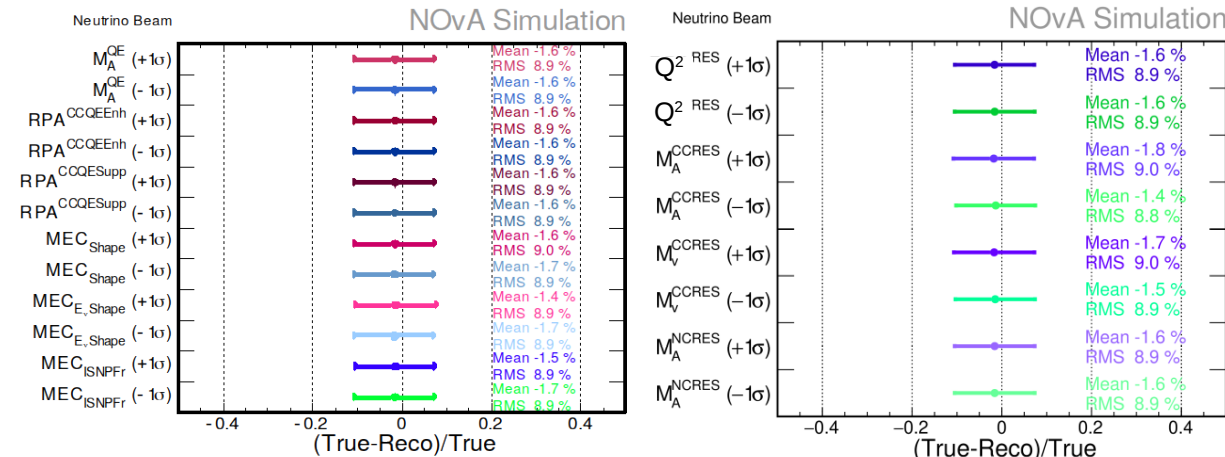


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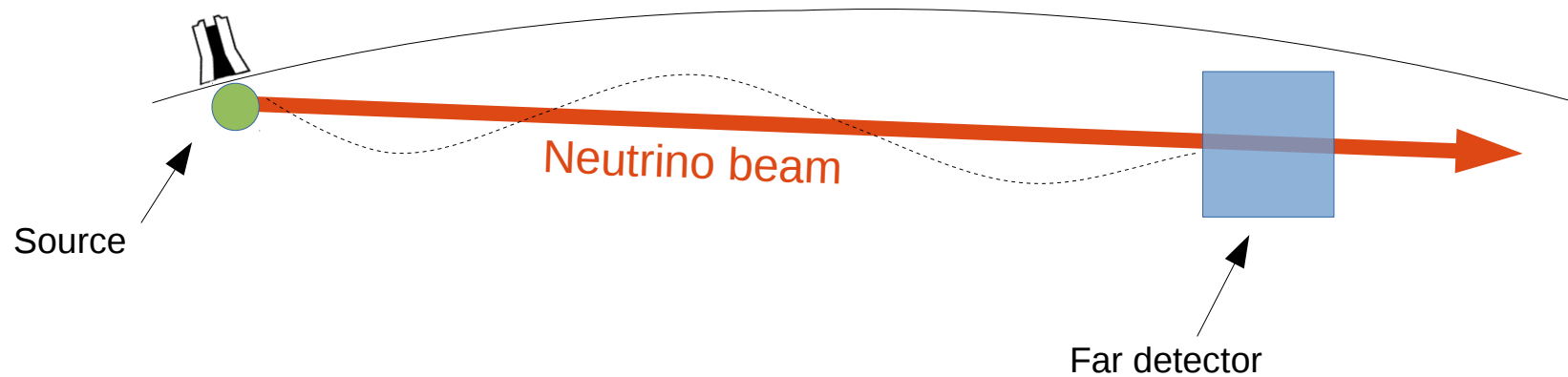


Nominal resolution  
on  $E_\nu \sim 9\%$ .

Despite sculpting effect,  
calorimeter-style detectors ensure  
**cross section systematics  
don't significantly  
degrade energy resolution**



# Near detectors

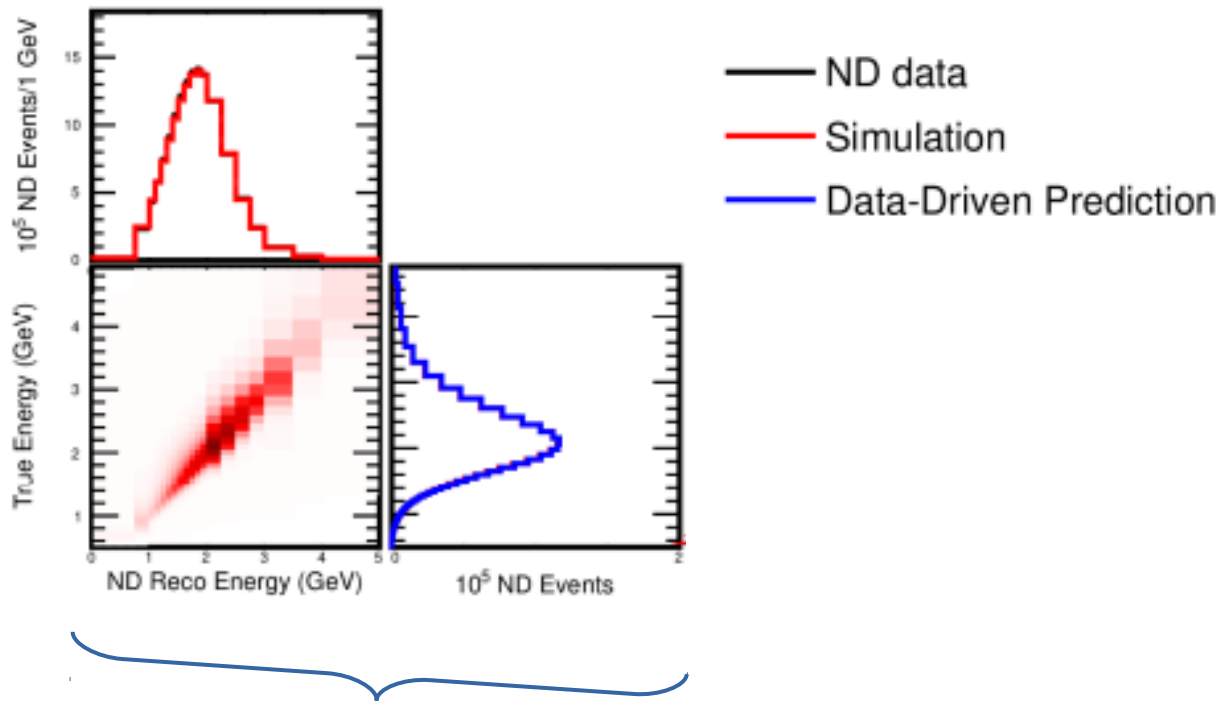


Want to measure **oscillation probability**.  
Many other variables...

$$N(E_{\nu}^{rec}) = \Phi(E_{\nu}^{true}) \times P_{osc}(E_{\nu}^{true}) \times \sigma(E_{\nu}^{true}, A) \times R(E_{\nu}^{true}) \times \epsilon(\dots)$$

# $\nu_\mu$ disappearance: “extrapolation”

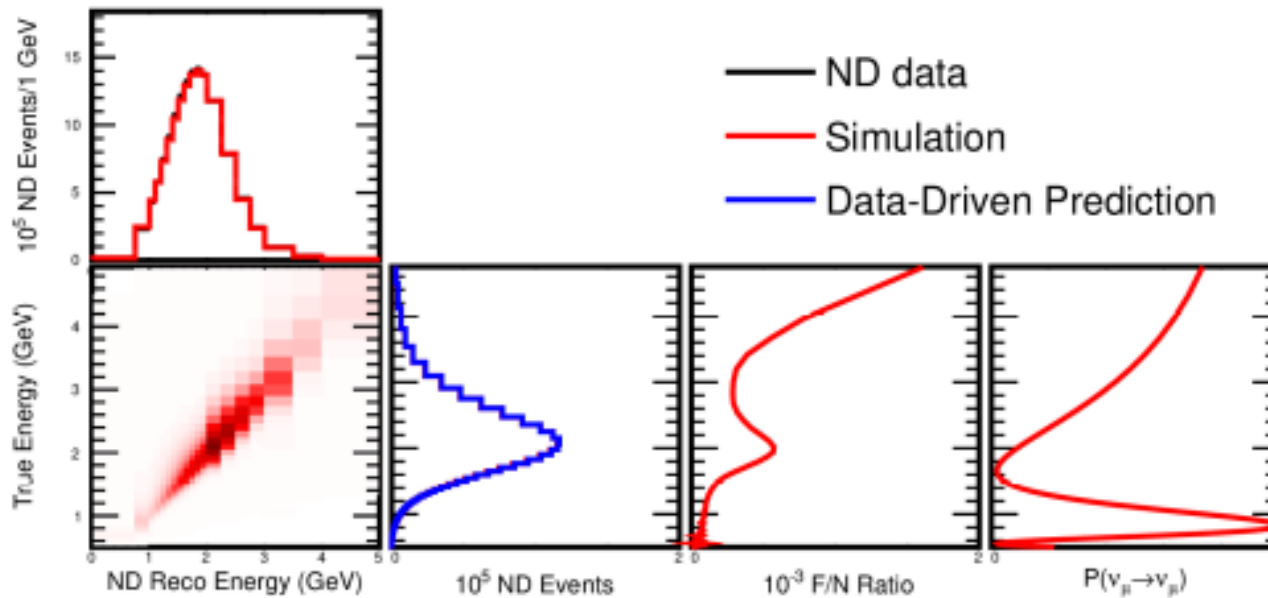
To produce a data-driven prediction at FD, based on ND:



True energy distribution is corrected so that reconstructed data & MC agree at the ND...

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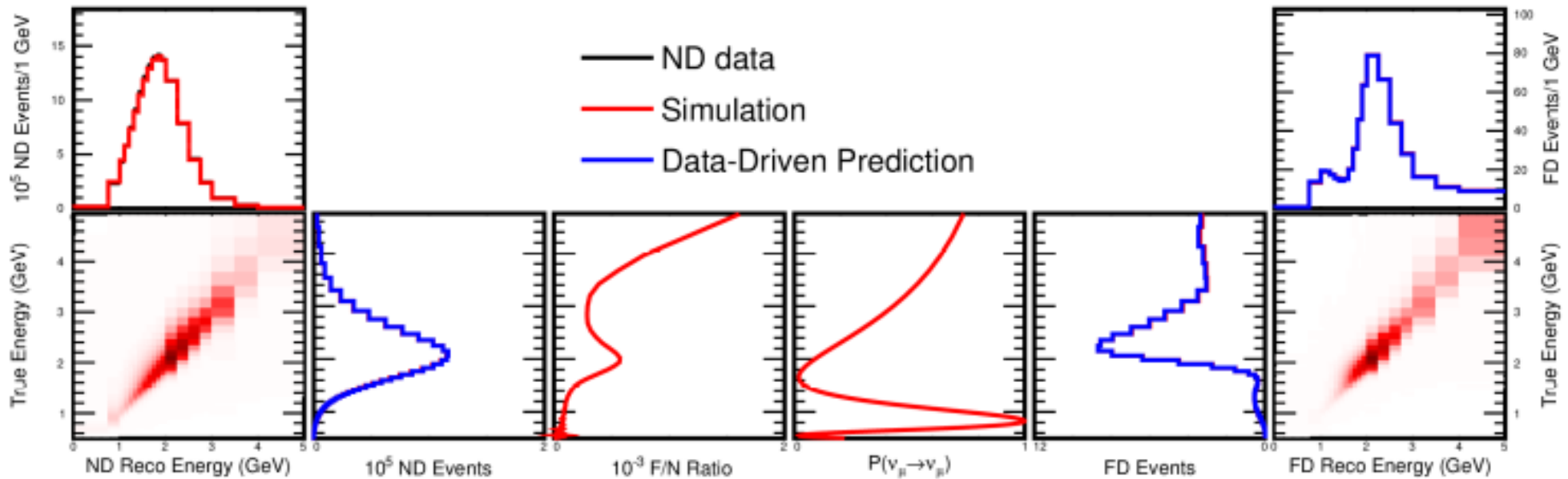


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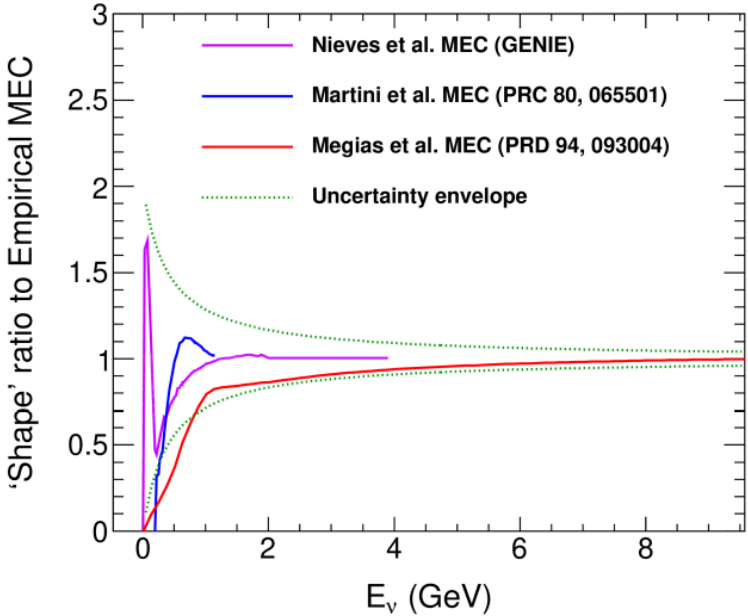
... and “extrapolated” reconstructed energy distribution computed to compare to data

# Illustrating XS systematics: MEC

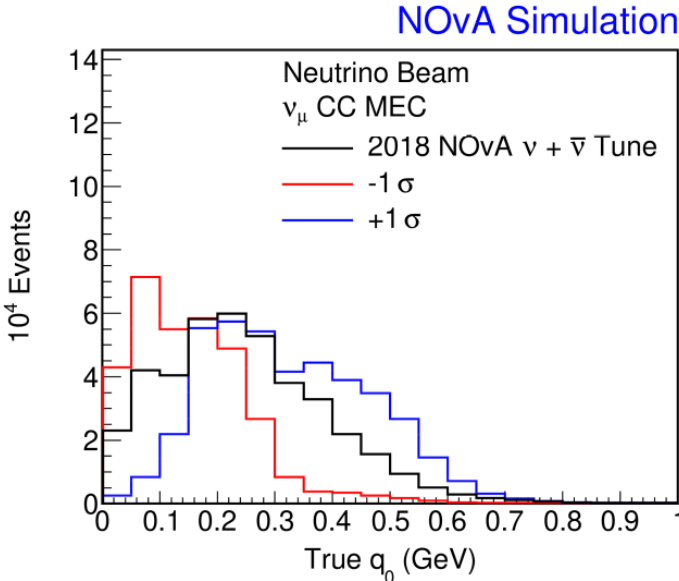
Examine this procedure through the lens of reaction of interest:

2p2h  
via  
Meson Exchange Currents  
(GENIE 'Empirical MEC' w/ ND tuning)

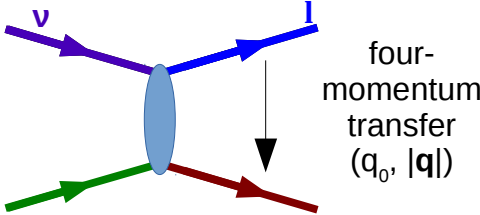
Illustrate behavior through two different knobs:



Neutrino energy dependence  
(brackets theoretical models)



Four-momentum transfer dependence  
(bounds our fits)

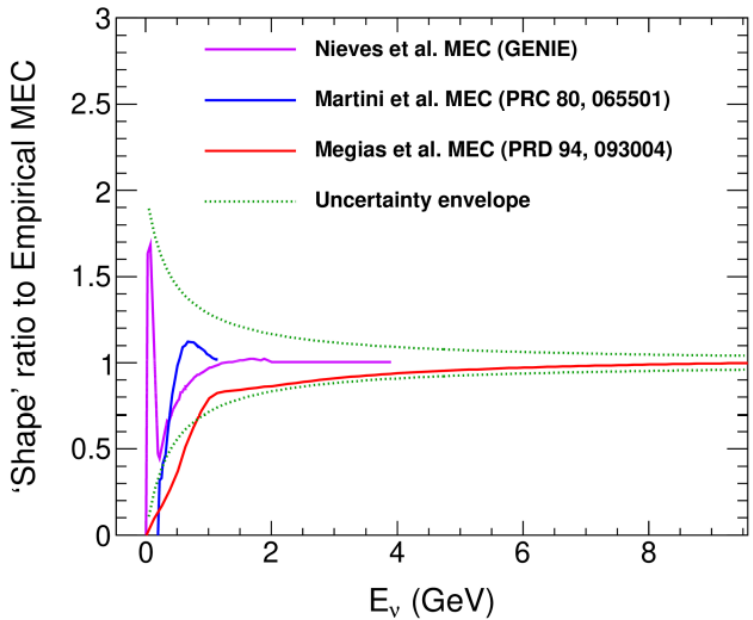


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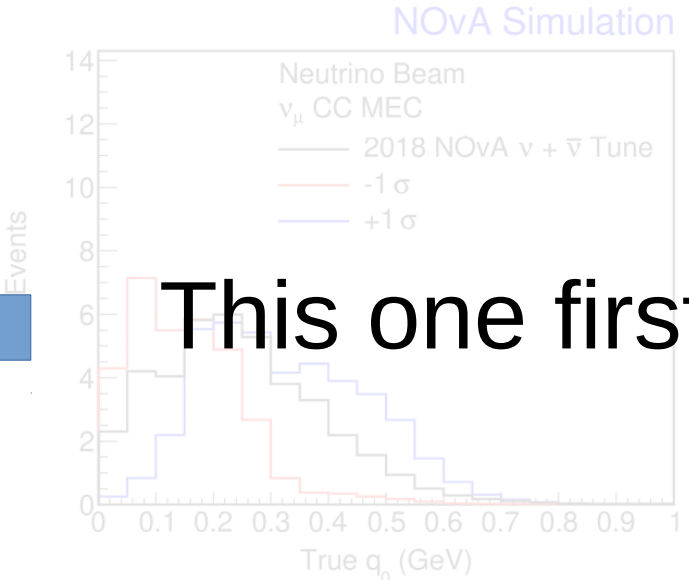
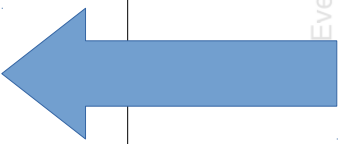
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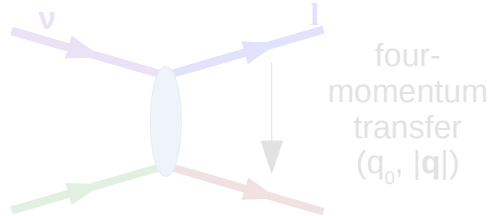


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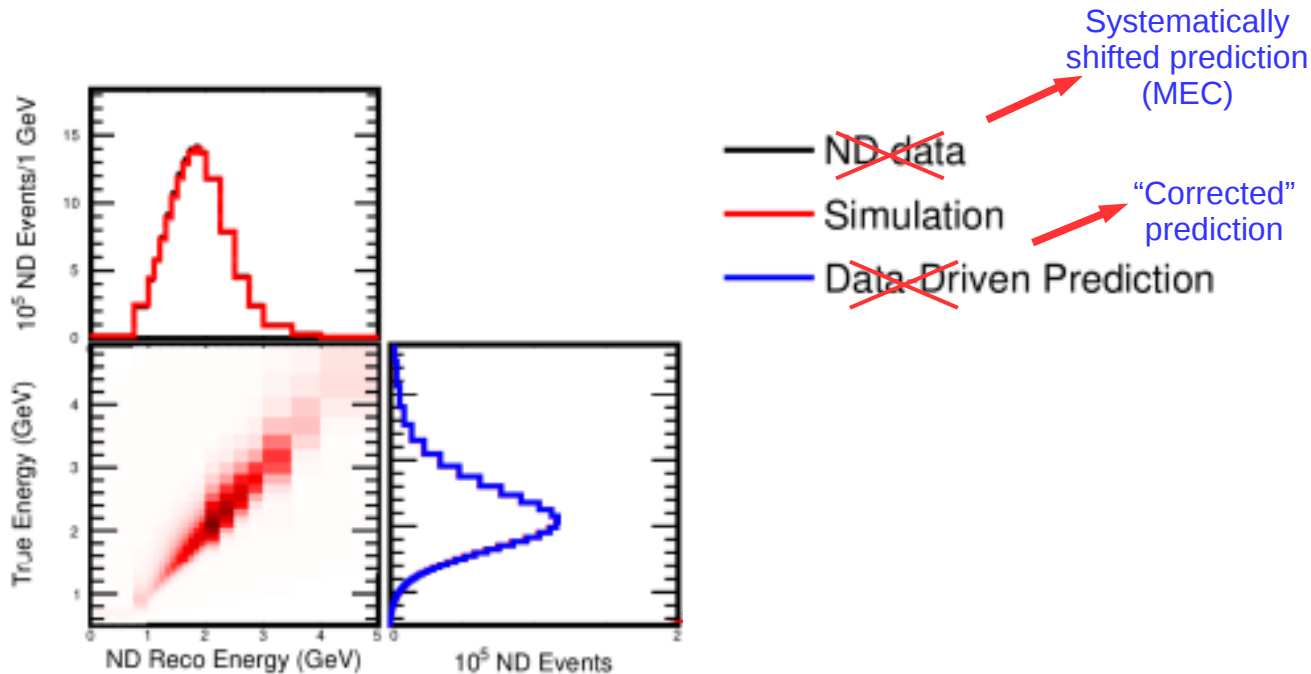
This one first

Four-momentum transfer dependence  
(bounds our fits)



# Testing extrapolation

To examine the effect of extrapolation:

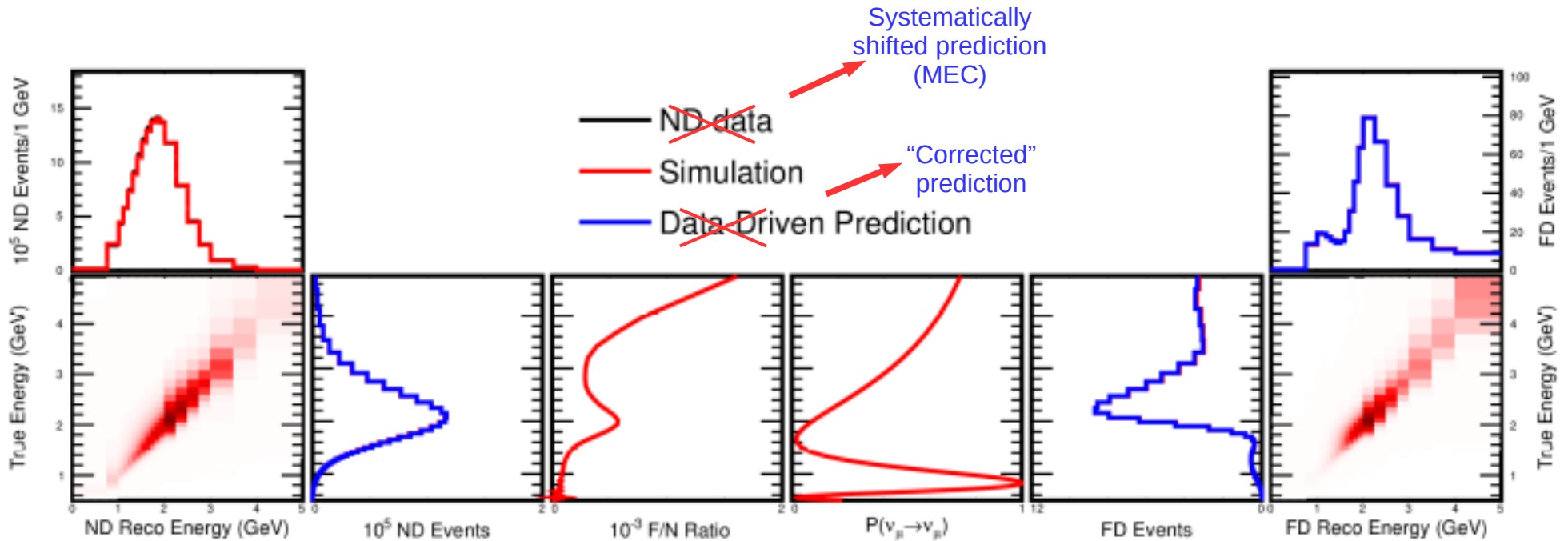


- ① Replace “ND Data” with “ND prediction under systematic shift”.  
(Asks: “if data exhibits this effect, and we use baseline simulation, how well does extrapolation compensate?”)



# Testing extrapolation

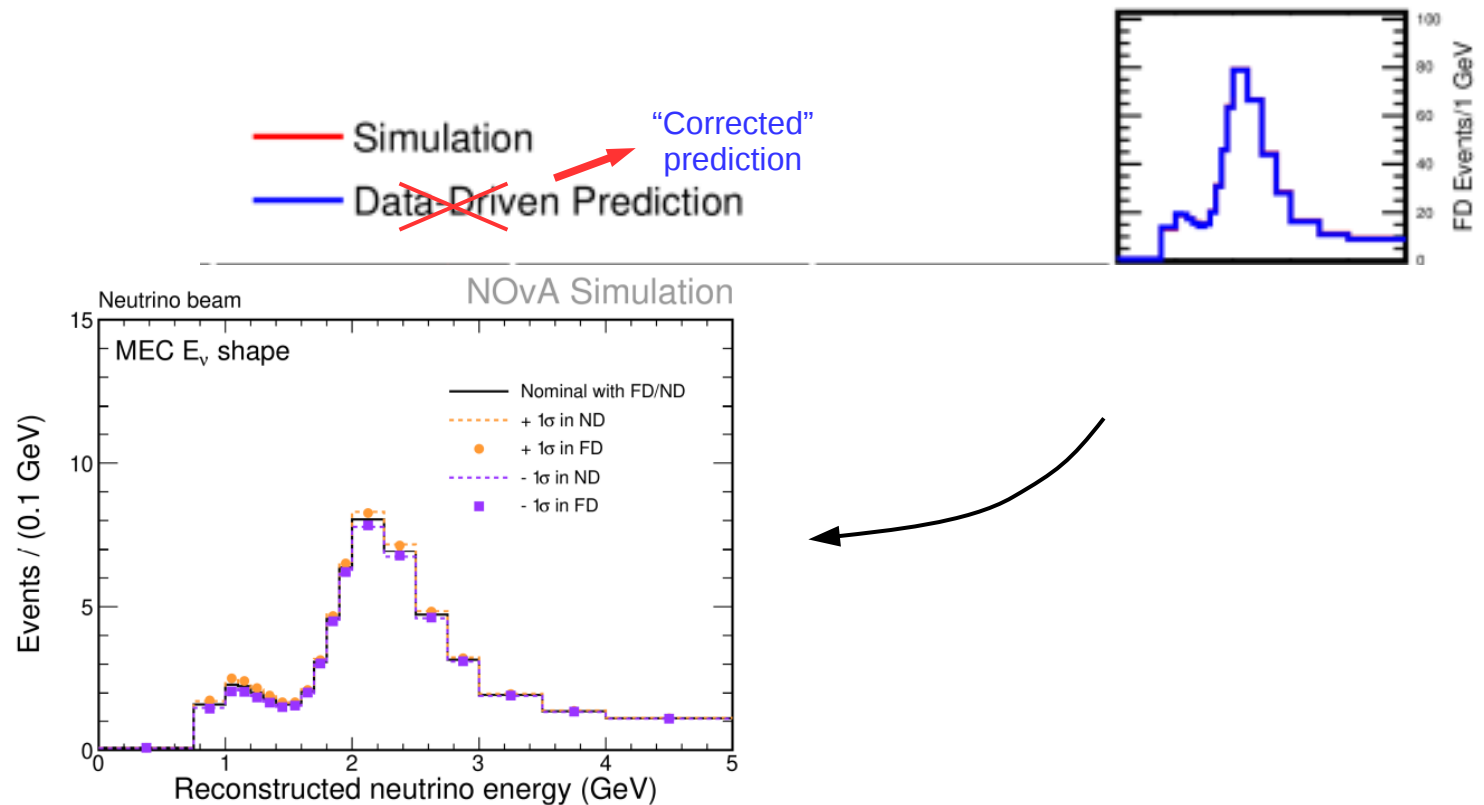
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② Transport “corrected” prediction through extrapolation process

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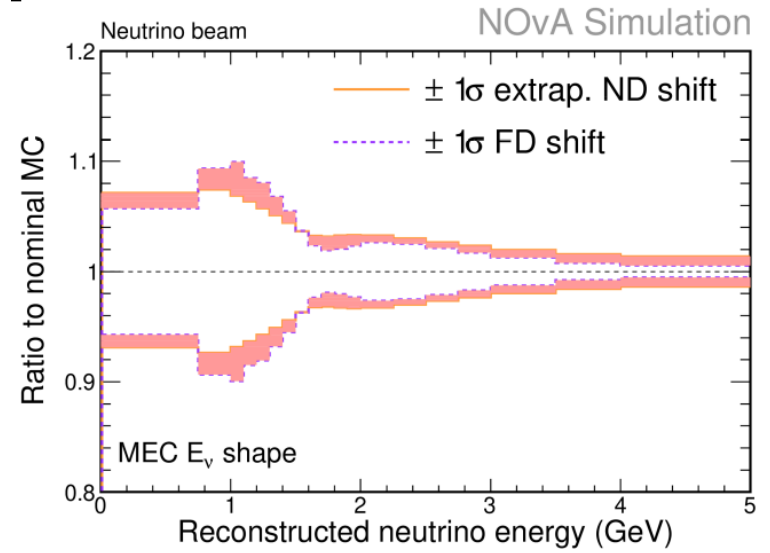
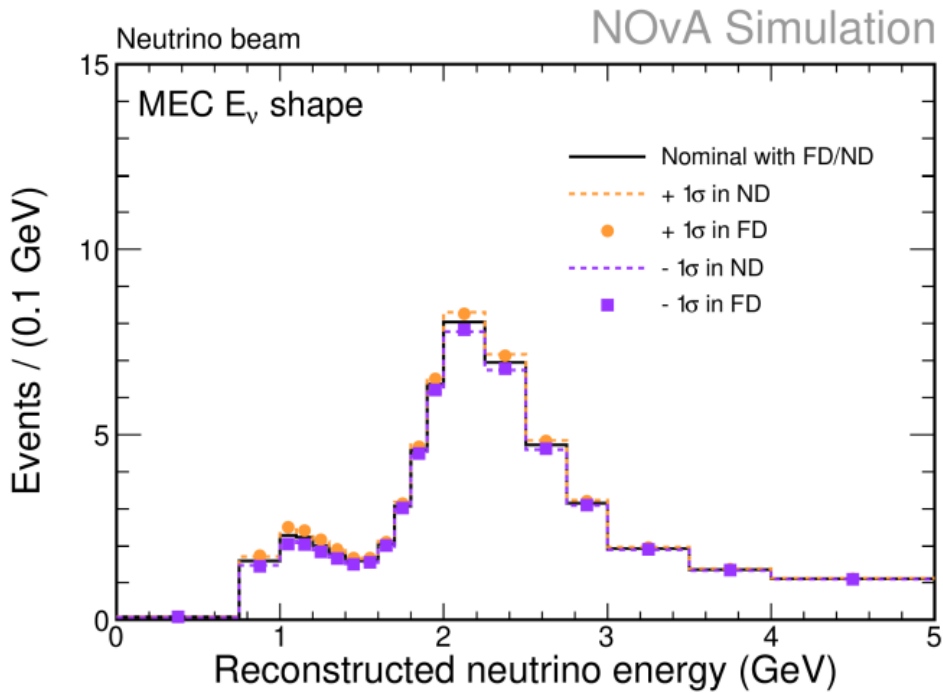


- ③ Compare “extrapolated” FD prediction to prediction obtained by varying FD directly.

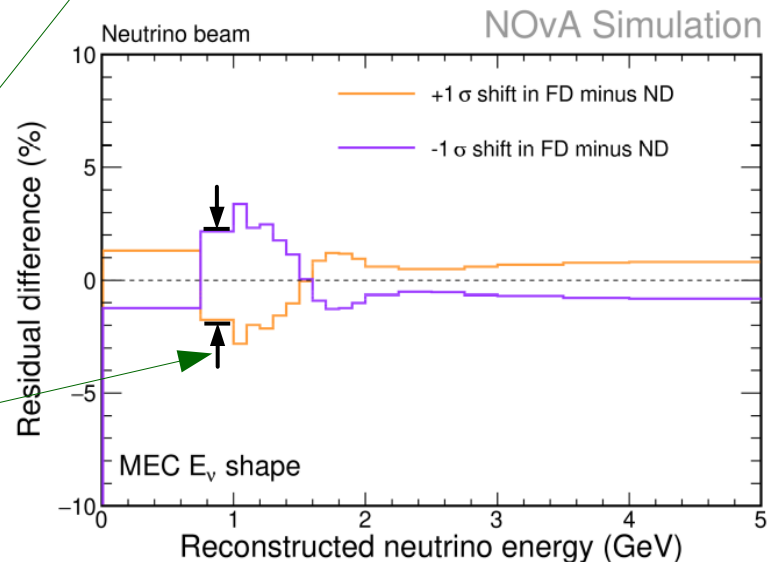
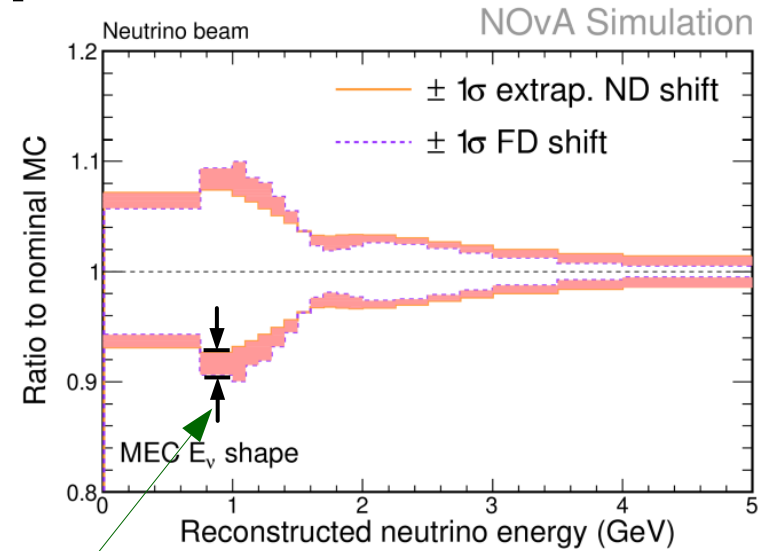
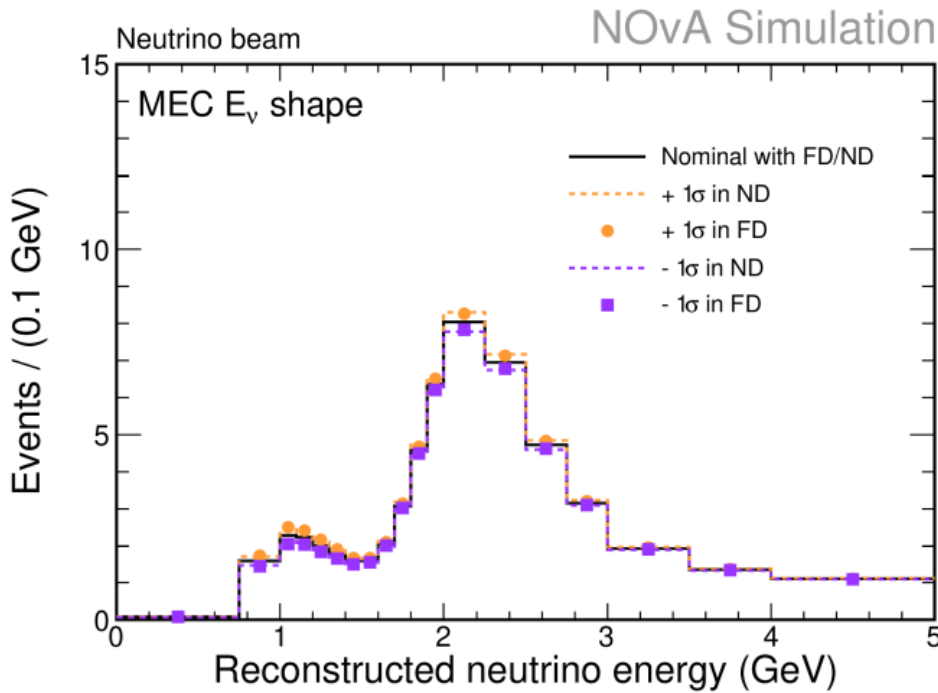


If they match, extrapolation perfectly 'cancels' the effect.

# Testing extrapolation

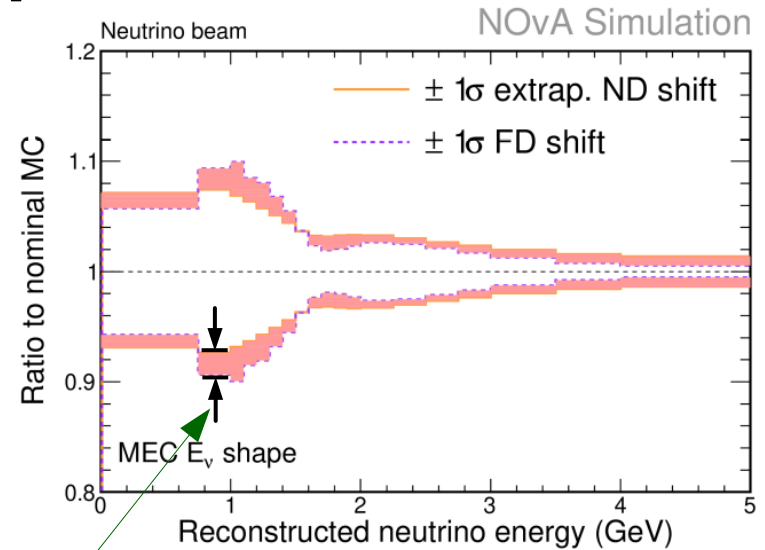
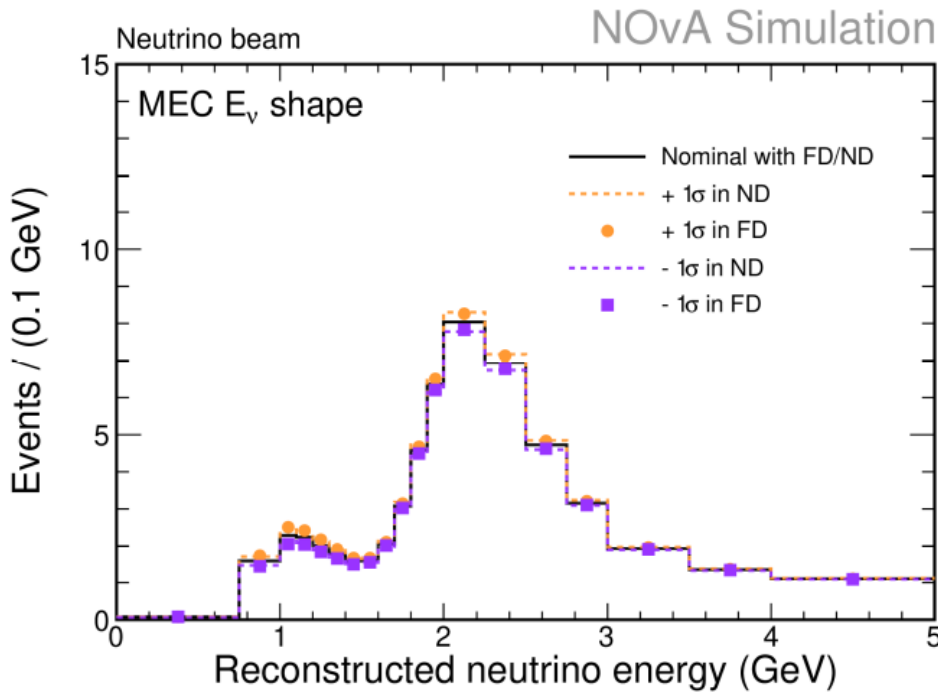


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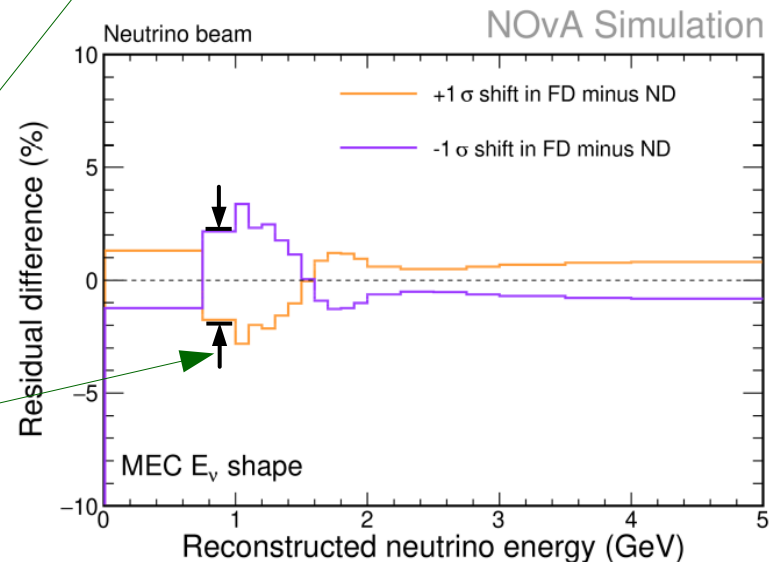


**Only a few percent *residual* effect of this MEC syst after extrapolation: the rest was canceled by the procedure.**

# Testing extrapolation

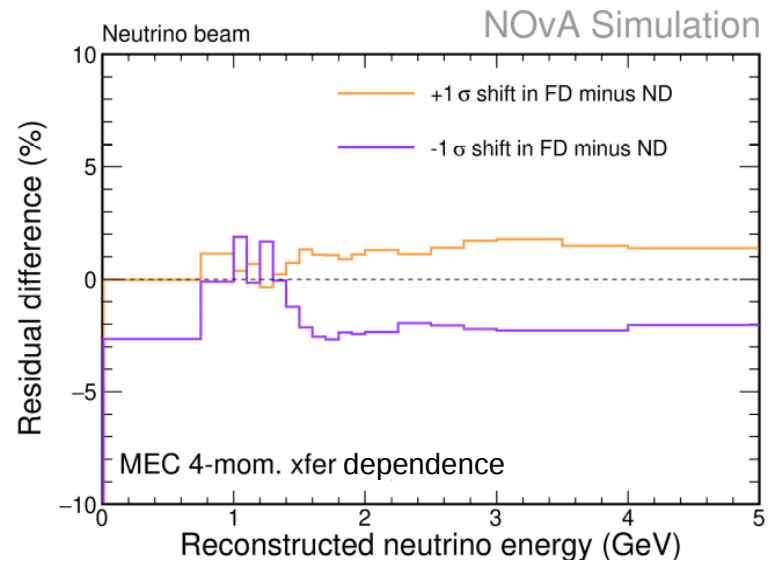
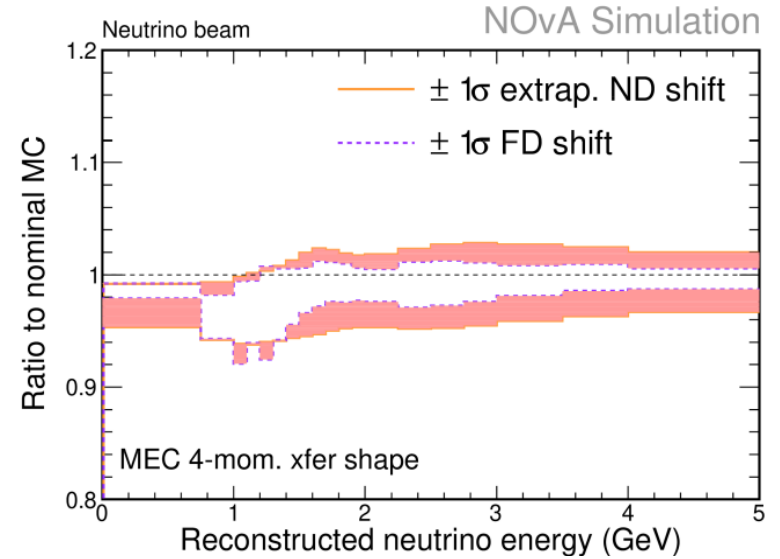
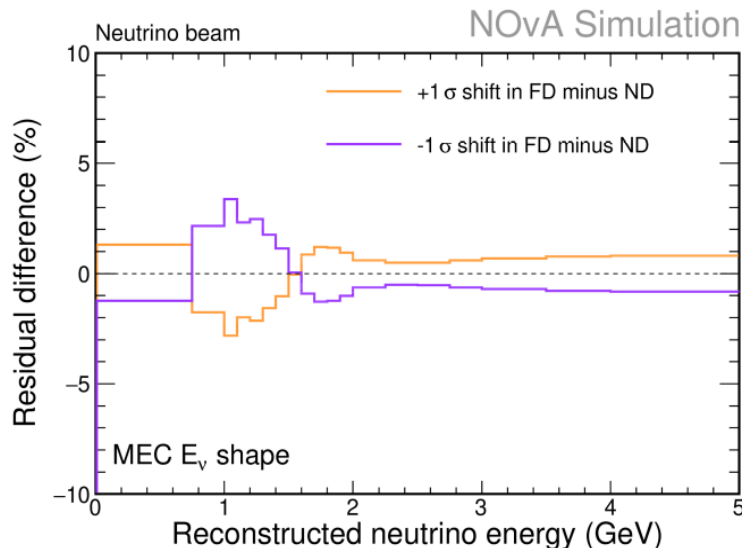
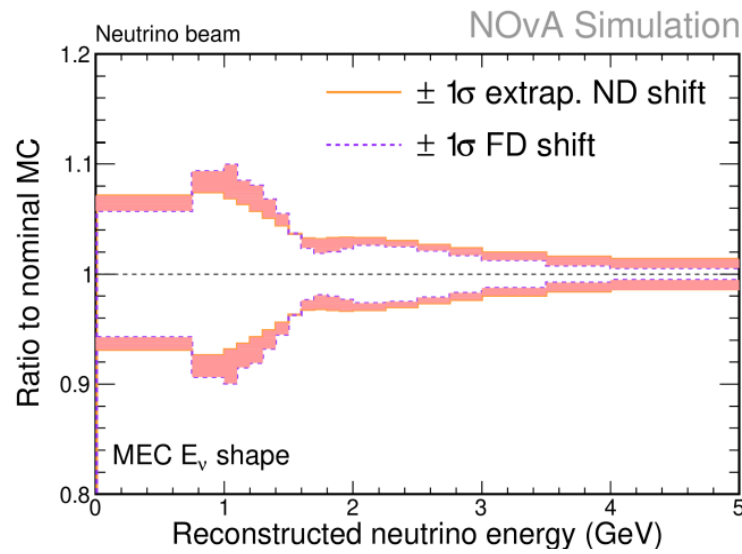


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Though extrapolation procedure can't remove *all* effect of cross section uncertainties like MEC, extrapolation significantly reduces sensitivity to XS systs

# Testing extrapolation

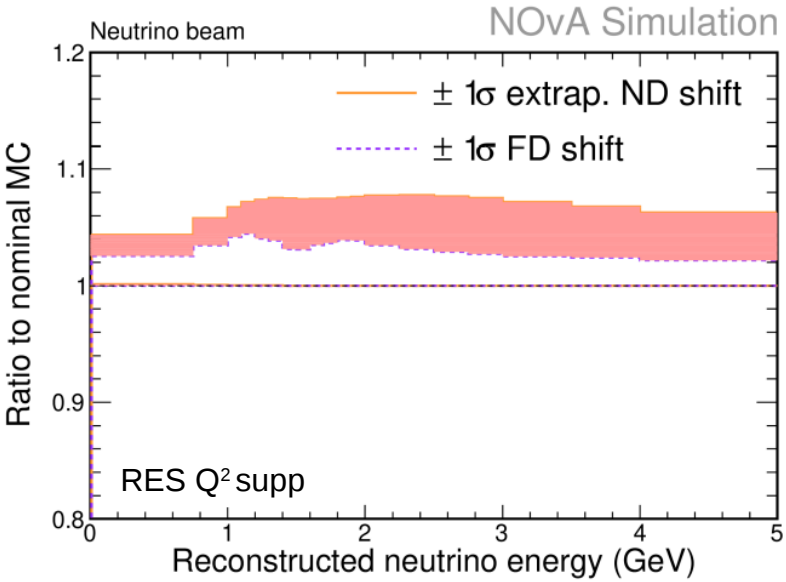
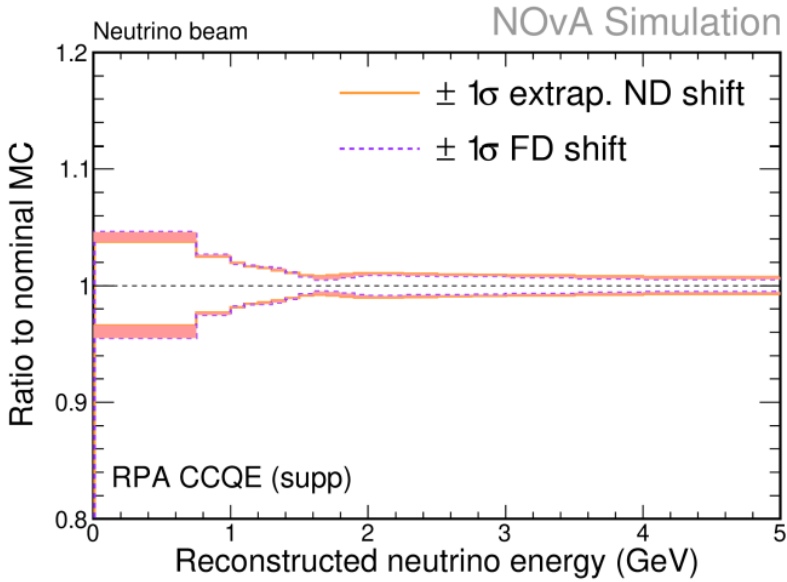


↑ Neutrino energy dependence

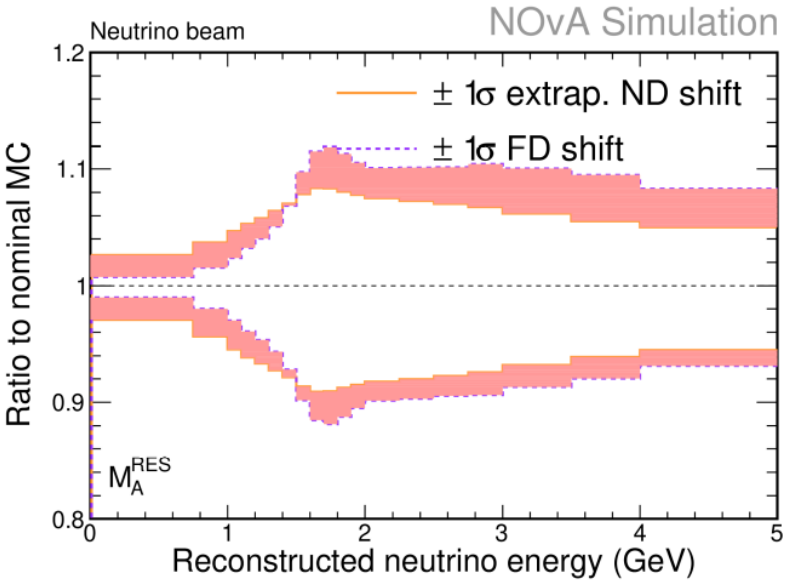
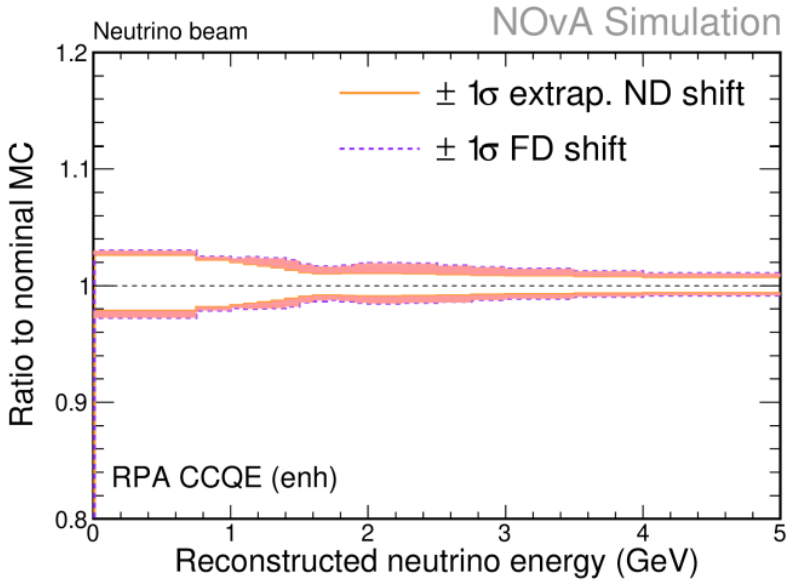
↓ Four-momentum transfer dependence

Far/Near extrapolation works best with *neutrino energy* systems,  
 but we derive benefit from it for the other shape dependence as well

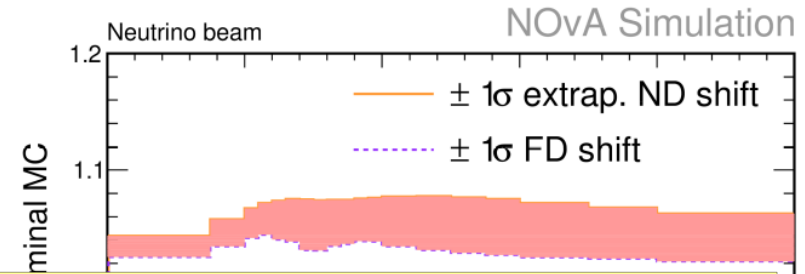
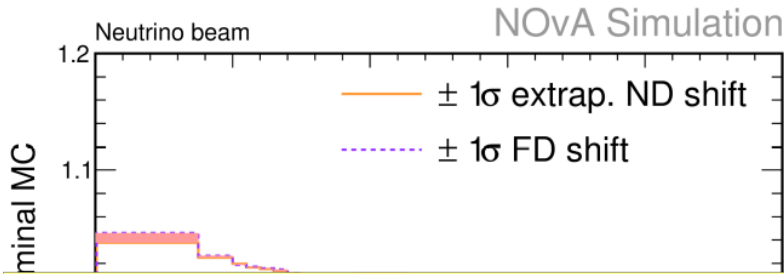
# Other important XS uncertainties



The story is similar for other important cross section uncertainties

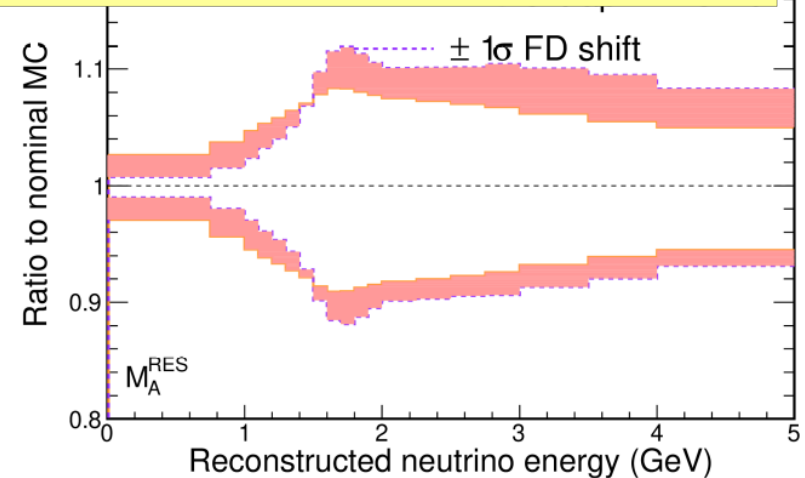
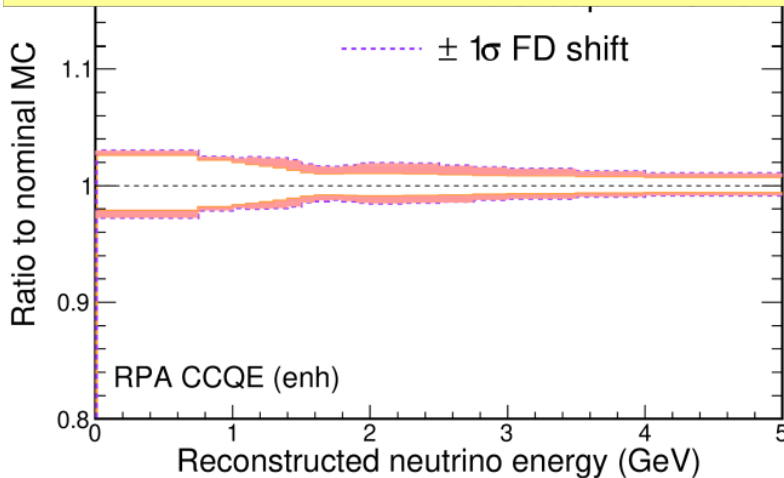


# Other important XS uncertainties



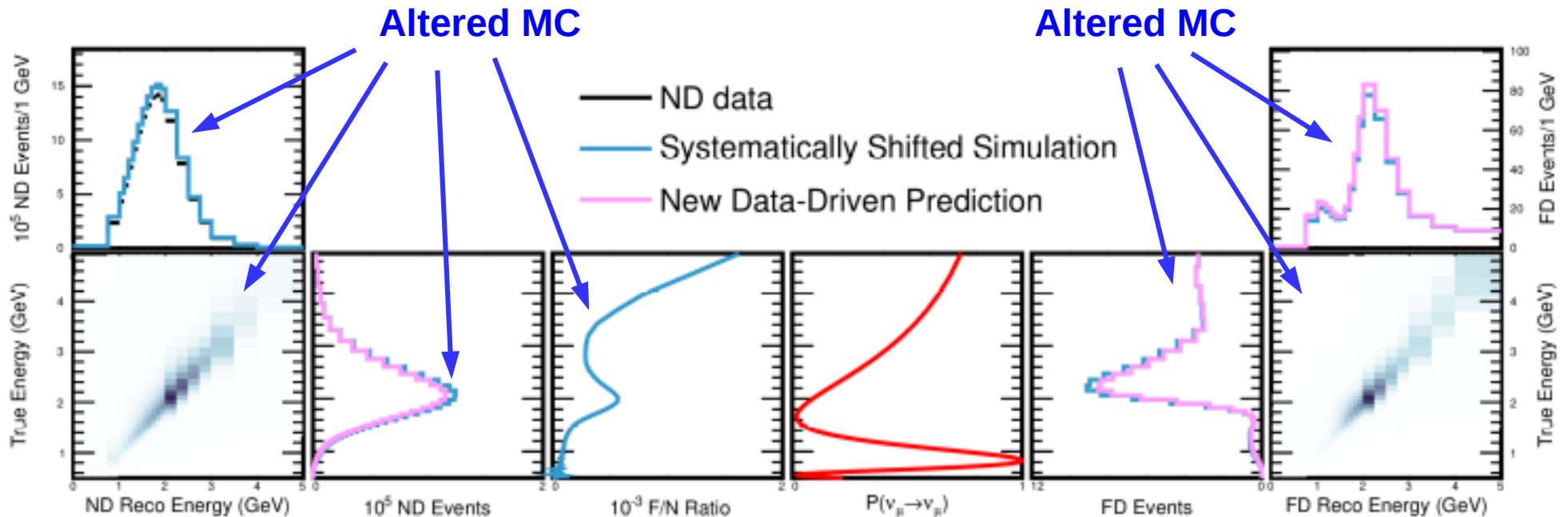
This illustrates how extrapolation responds to “unknown unknowns” in the *data*.

We do the “inverse” to handle “known unknowns” using our MC...





# “Extrapolation” and uncertainties



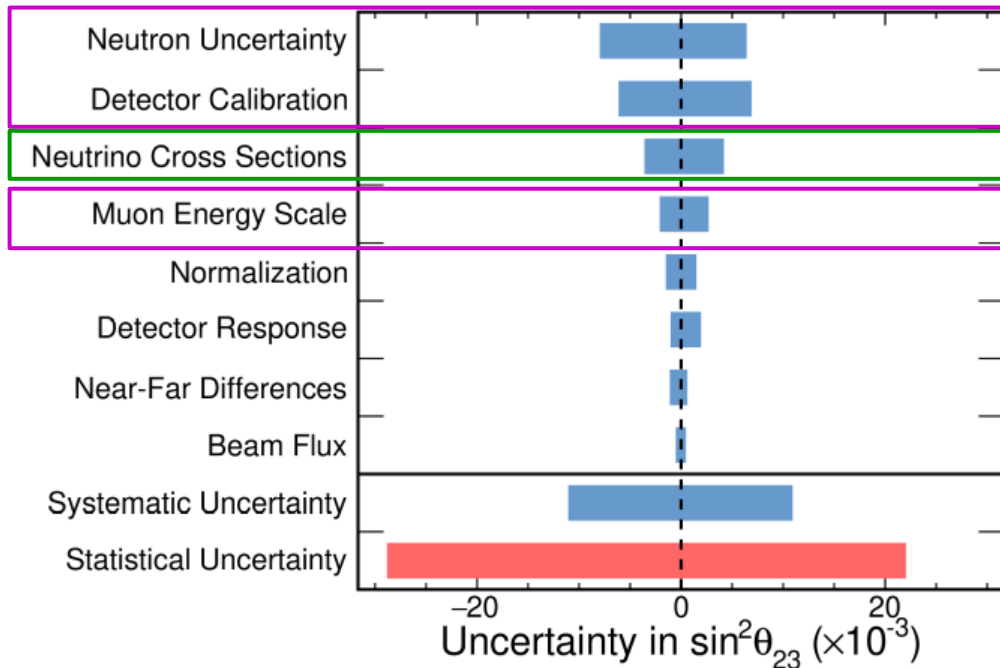
We simulate the effect of our cross section systematics' *residual* effect after extrapolation

**by re-doing the entire analysis for each systematic**  
(each of which can affect multiple both signal & bknd)

and use the difference to extrapolated nominal MC  
as nuisance parameters in our oscillation fits

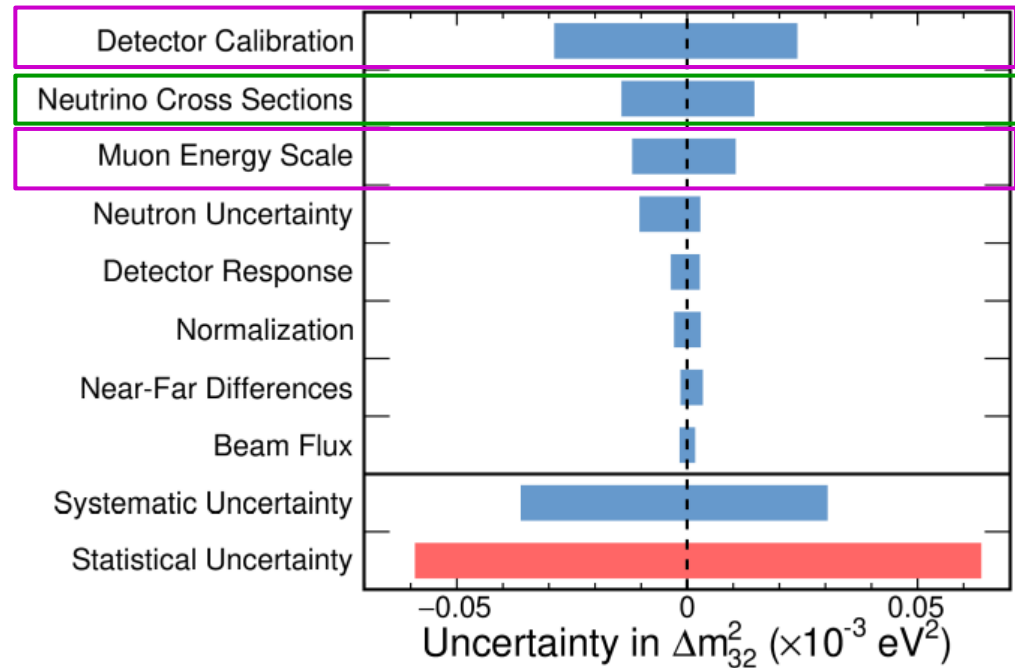
# Effect on analysis

NOvA Preliminary



(Uncertainty on joint  $\nu + \bar{\nu}$ ,  $\nu_\mu + \nu_e$  fit)

NOvA Preliminary



Cross section systematics are not dominant systematic uncertainties due to detector design & power of extrapolation.

But... dedicated **test beam program** (data taking in 2019) will likely drive **detector response** uncertainty down in the future, so soon enough **cross sections** will likely be atop the list...

Now:  
 $\nu_e$  appearance

# $\nu_e$ appearance

$$P(\overset{(-)}{\nu}_\mu \rightarrow \overset{(-)}{\nu}_e) \approx \sin^2 2\theta_{13} \sin^2 \theta_{23} \frac{\sin^2(A-1)\Delta}{(A-1)^2}$$

$$- \overset{(+)}{2\alpha} \sin \theta_{13} \sin \delta_{CP} \sin 2\theta_{12} \sin 2\theta_{23} \frac{\sin A\Delta}{A} \frac{\sin(A-1)\Delta}{A-1} \sin \Delta$$

$$+ 2\alpha \sin \theta_{13} \cos \delta_{CP} \sin 2\theta_{12} \sin 2\theta_{23} \frac{\sin A\Delta}{A} \frac{\sin(A-1)\Delta}{A-1} \cos \Delta$$

Where:  $\alpha = \frac{\Delta m_{21}^2}{\Delta m_{31}^2}$   $\Delta = \Delta m_{31}^2 \frac{L}{4E}$   $A = \overset{(-)}{+} G_f N_e \frac{L}{\sqrt{2}\Delta}$

Besides the dependence on the mixing parameters, we learn about the mass ordering (via  $\alpha$ ) and  $\delta_{CP}$ ...

# $\nu_e$ appearance

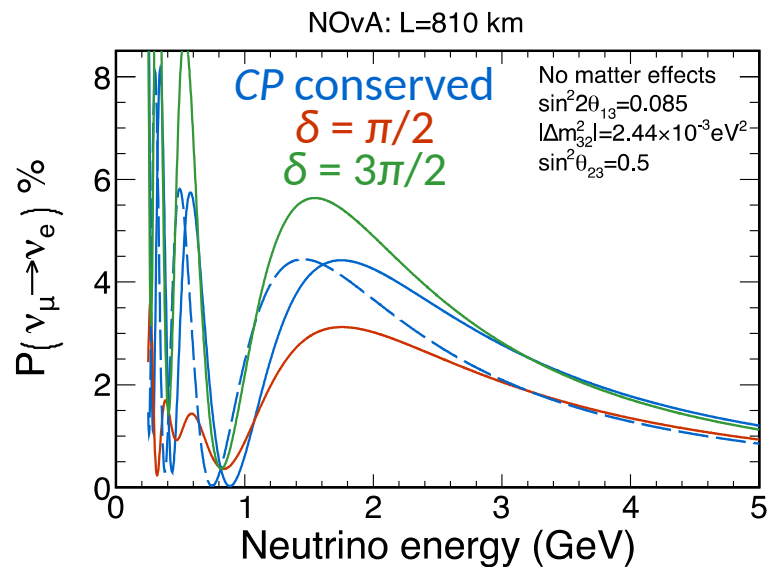
$$P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) \approx \sin^2 2\theta_{13} \sin^2 \theta_{23} \frac{\sin^2(A-1)\Delta}{(A-1)^2}$$

$$- 2\alpha \sin \theta_{13} \sin \delta_{CP} \sin 2\theta_{12} \sin 2\theta_{23} \frac{\sin A\Delta \sin(A-1)\Delta}{A(A-1)} \sin \Delta$$

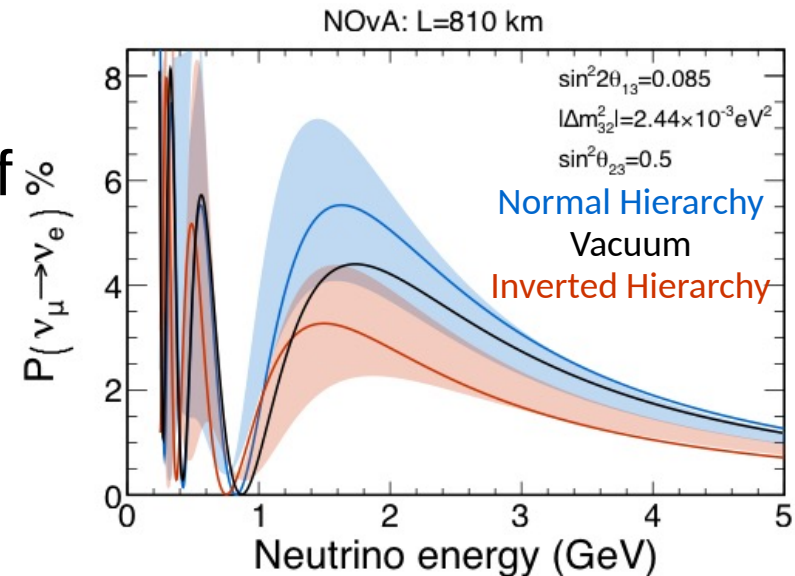
$$+ 2\alpha \sin \theta_{13} \cos \delta_{CP} \sin 2\theta_{12} \sin 2\theta_{23} \frac{\sin A\Delta \sin(A-1)\Delta}{A(A-1)} \cos \Delta$$

Where:  $\alpha = \frac{\Delta m_{21}^2}{\Delta m_{31}^2}$     $\Delta = \Delta m_{31}^2 \frac{L}{4E}$     $A = \begin{matrix} (-) \\ + \end{matrix} G_f N_e \frac{L}{\sqrt{2}\Delta}$

Besides the dependence on the mixing parameters, we learn about the mass ordering (via  $\alpha$ ) and  $\delta_{CP}$ ...

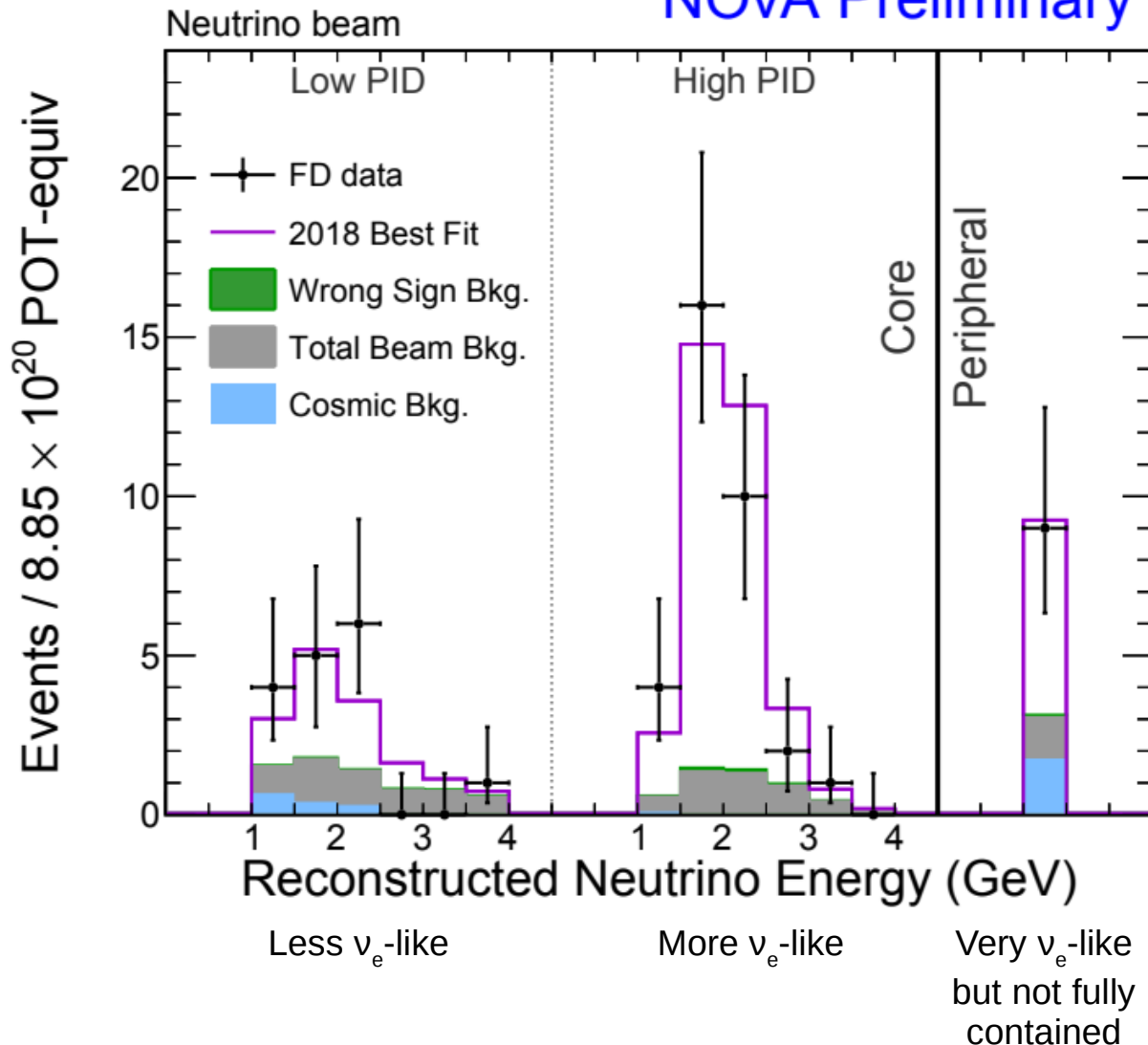


...but smallness of  $\theta_{13}$  makes it a very challenging measurement



# $\nu_e$ appearance

NOvA Preliminary



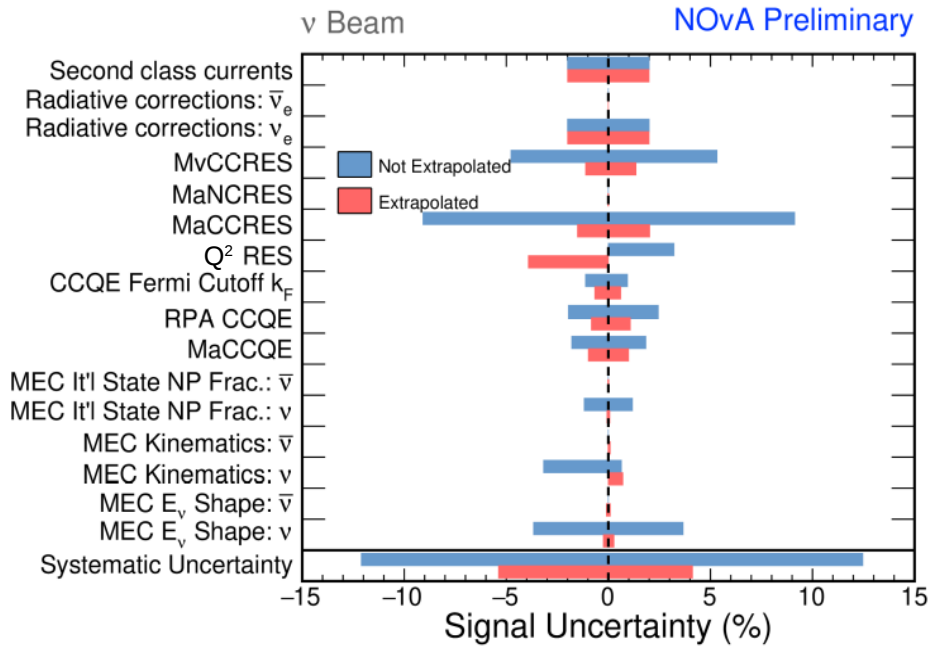
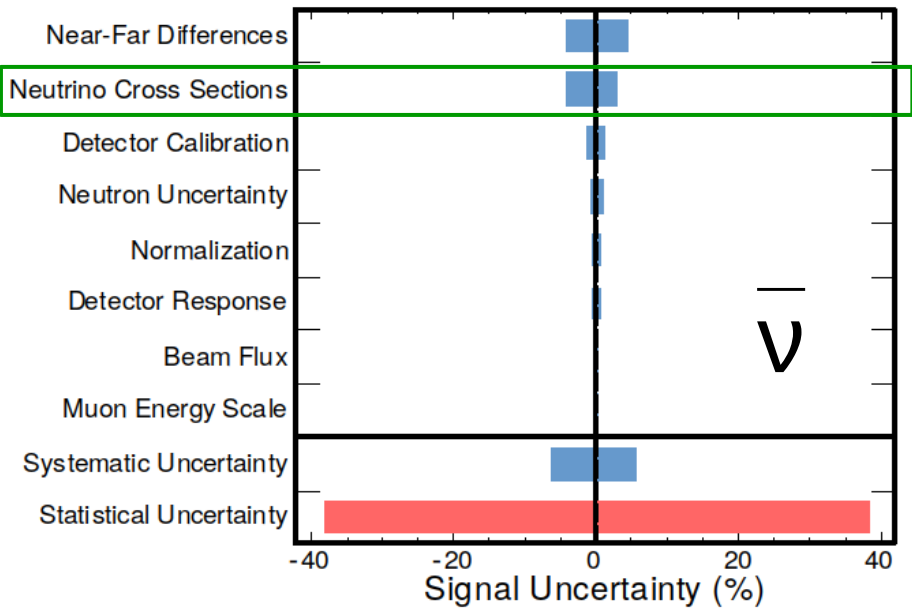
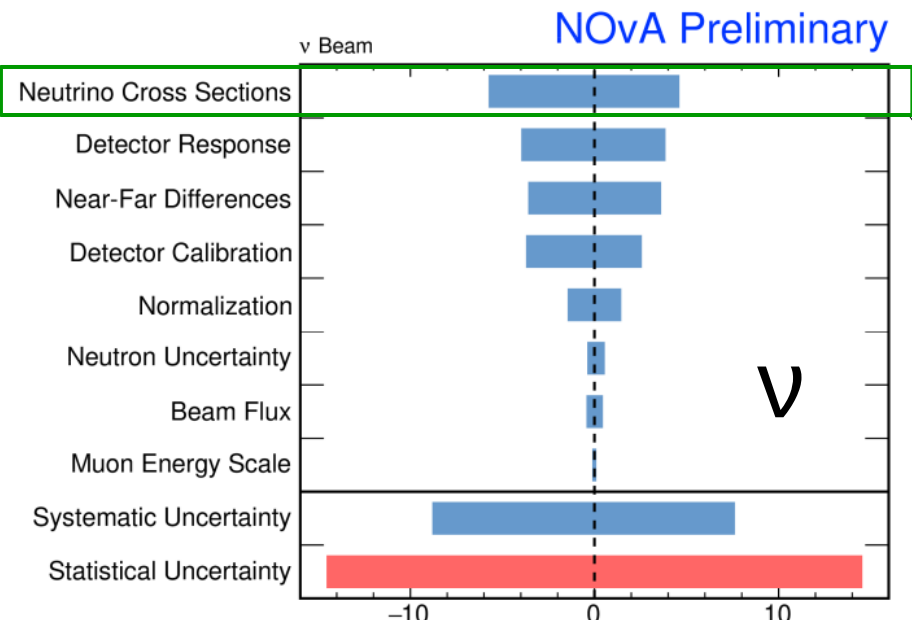
## Added challenges:

- **No signal at ND**
  - Need to assume relationship between  $\nu_\mu$  and  $\nu_e$  XSs
  - Different acceptance,  $\nu_\mu$  ND vs.  $\nu_e$  FD
- **Nontrivial beam backgrounds which oscillate differently**
  - **Beam  $\nu_e$**  oscillate very little over this L/E
  - **$\nu_\mu$**  almost entirely disappear
  - **NC** doesn't change due to oscillations (assume no steriles)



(Need to disentangle  
[“decompose”] before applying  
Far/Near makes any sense.)

# $\nu_e$ appearance



**Cross section systematics still under control for now, but can't help  $\nu_e/\nu_\mu$  diff with ND... (Even more stat limited for  $\bar{\nu}$ .)**

We expect to continue to benefit from ongoing work by this audience as well to keep them that way...

# Cross section uncertainties: future needs

- **Refinement of 1p1h initial state prediction**
  - Unified nuclear model treatment in GENIE (move past RFG+RPA?)
  - Study binding/separation/removal energy situation (A. Bodek, [arXiv:1801.07975](https://arxiv.org/abs/1801.07975))...
- **More/better models for multinucleon knockout in GENIE**
  - Need theory models that describe inclusive data at  $E_\nu > 1$  GeV. SuSA-MEC in GENIE soon?
  - Need to study effect of anisotropic ejection of nucleon pairs
- **Nuclear models for *inelastic* processes**
  - Examine alternatives to GENIE hA FSI
  - What's causing low- $Q^2$  suppression? (“On-off” syst one of our largest)
- **$\nu_e/\nu_\mu$  for inelastic processes**
- **More antineutrino data**
- ...



# Summary

- NOvA relies on **strong internal constraints on cross section uncertainties** for its oscillation program
  - Calorimeter design minimizes *a priori* impact
  - Functionally identical detectors enable major cancellation of residual errors in oscillation analyses
- Comprehensive program underway to ensure all **relevant cross section issues are considered**
  - Necessary ingredients in base model
  - Appropriate uncertainties
- We look forward to **continuing the conversation:**
  - Continued development of models & systematic treatments
  - New measurements of cross sections
  - Neutrino oscillation physics results!

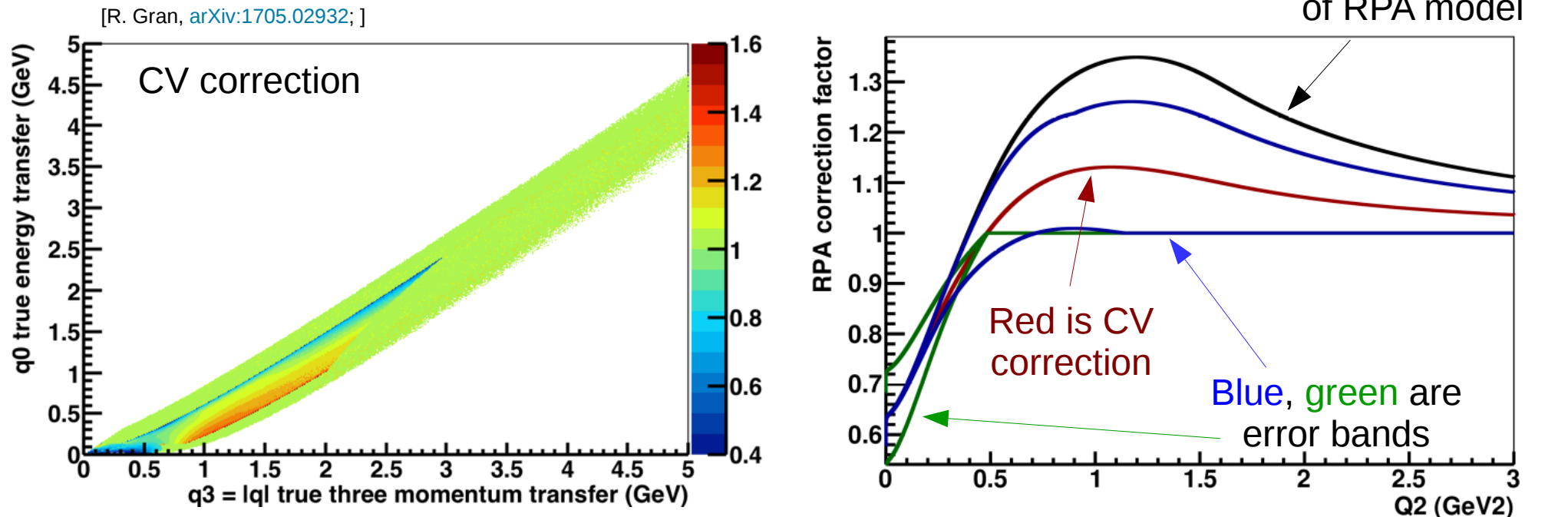


[NOvA 2018 at University of Austin]

Thank you for your attention!

# Overflow

# Modeling the nucleus: collective effects (RPA)

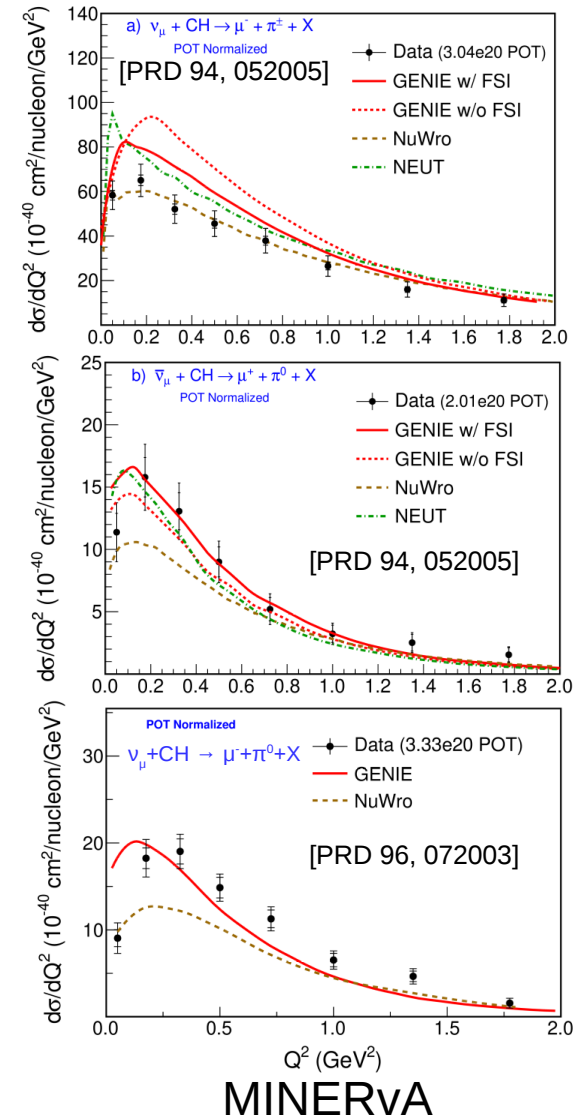
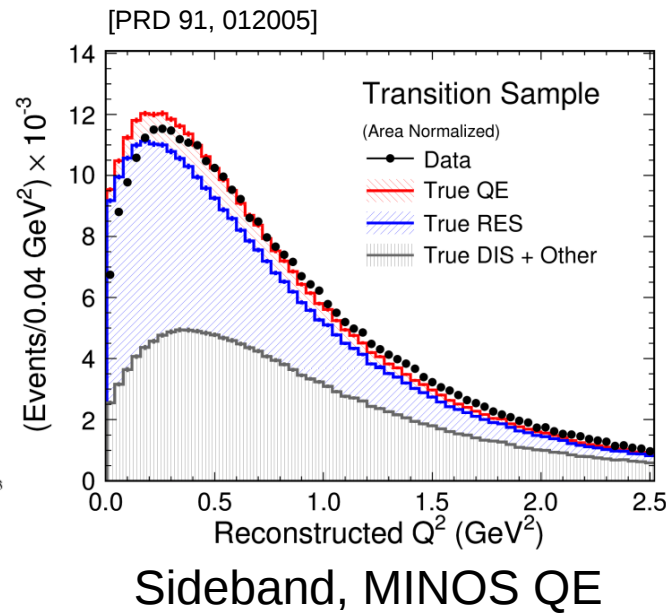
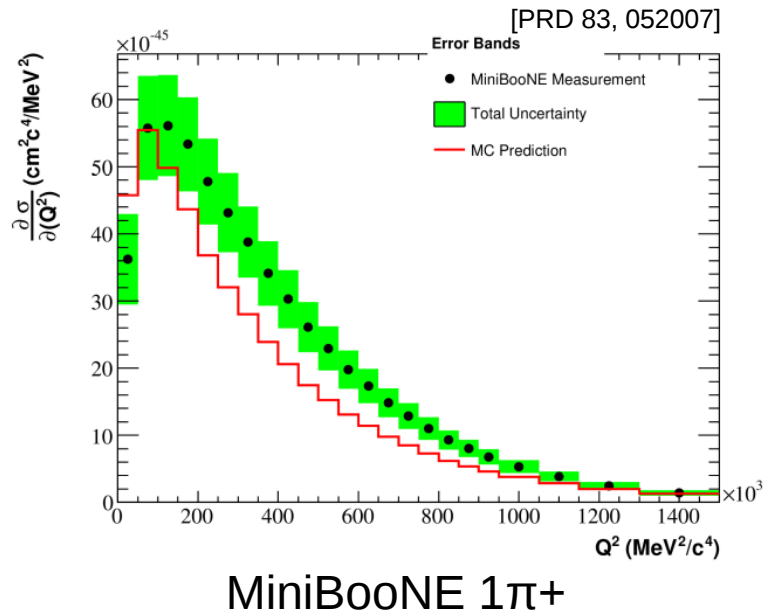


**Rik Gran's work (originally for MINERvA) to extend the València RPA CCQE effect (PRC 70, 055503) to a correction for GENIE's central value and his work to extend the uncertainties in the model to higher energies (PLB 638, 325, PRD 88, 113007) naturally work reasonably well for NOvA**

—————▶ we apply using Rik's code

# Modeling the nucleus: collective effects (RPA)

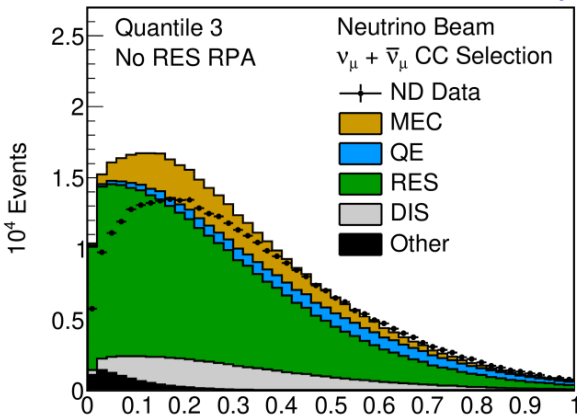
- Should  $\Delta$  production also be affected?
- Energy transfers are large ( $>100$  MeV), so perhaps not “traditional” RPA...
- But suppression at low  $Q^2$  appears in many data sets: MiniBooNE, MINOS, MINERvA



# Modeling the nucleus: collective effects (RPA)

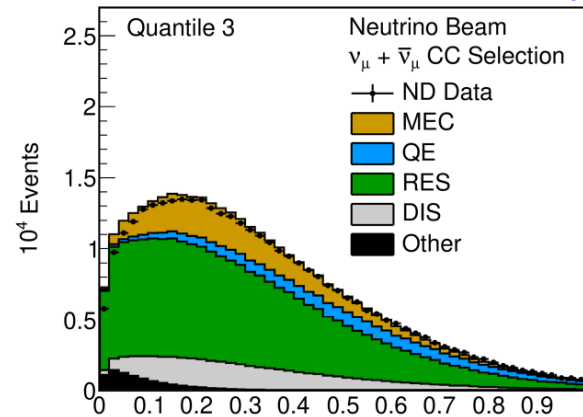
- Should  $\Delta$  production also be affected?
- Not surprising that *some* initial-state nuclear effects might influence resonances (relative to RFG)...

NOvA Preliminary

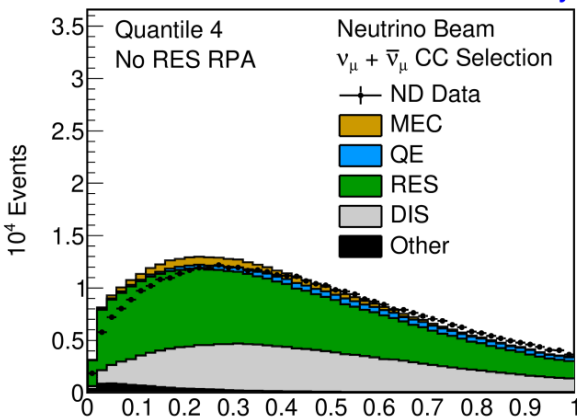


“Quantiles” are divided by hadronic energy fraction:  $\text{reco } E_{\text{had}} / \text{reco } E_{\nu}$ .

NOvA Preliminary

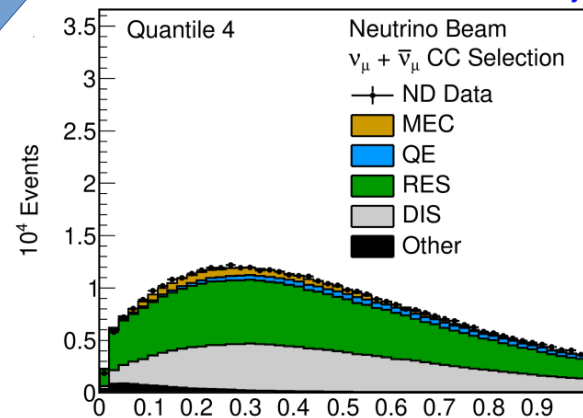


NOvA Preliminary



Quantiles 3 & 4 are RES-rich regions of  $\nu_{\mu}$  candidate sample.

NOvA Preliminary

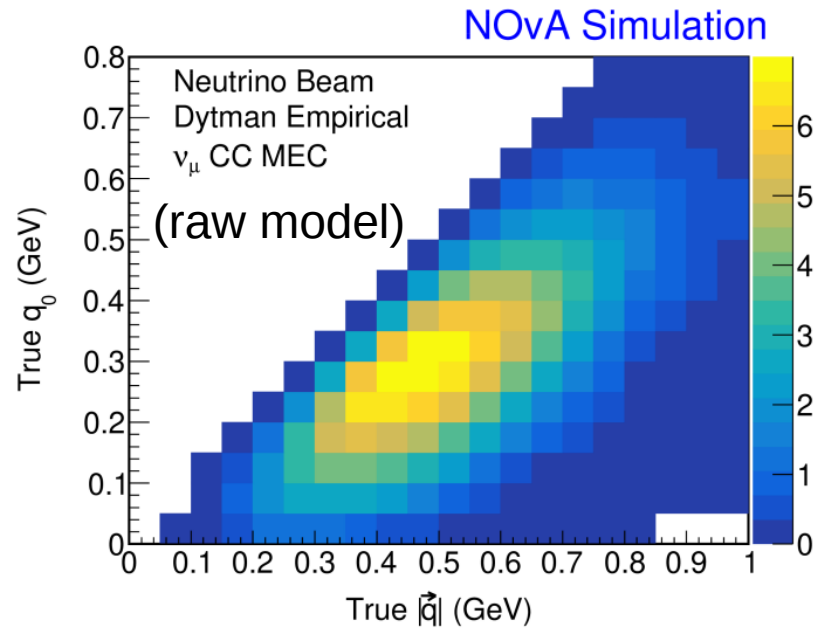


We apply the  $Q^2$ -based RPA weight from QE to resonant production as a stand-in for whatever nuclear effect it may be (w/ unmodified version as uncertainty variation)

# Modeling the nucleus: tuning 2p2h-MEC

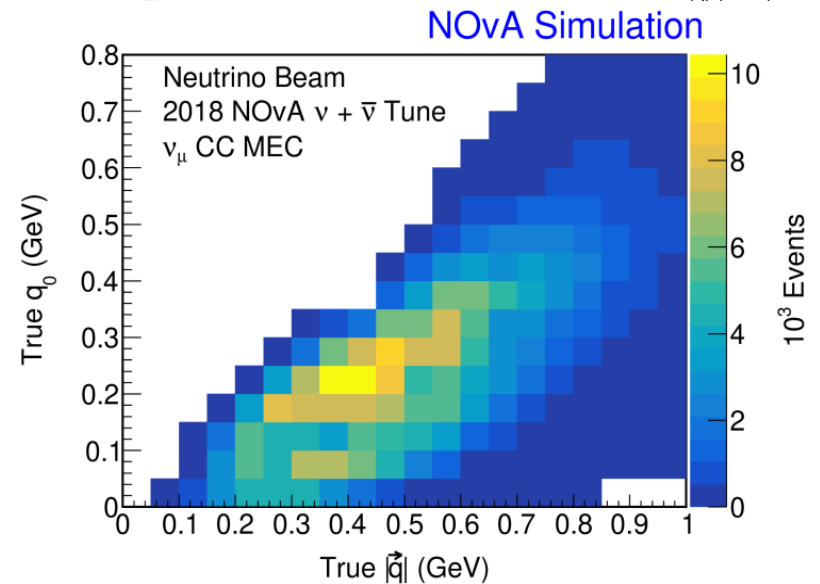
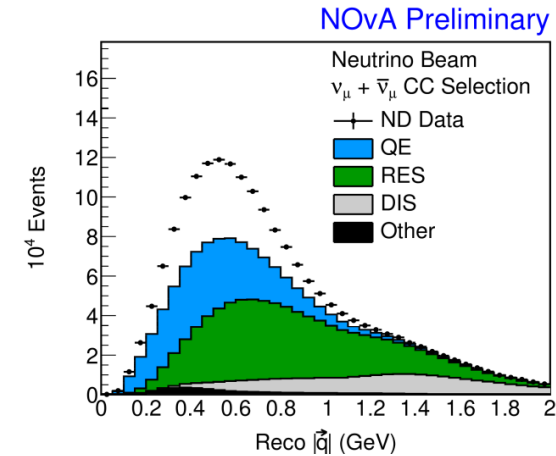
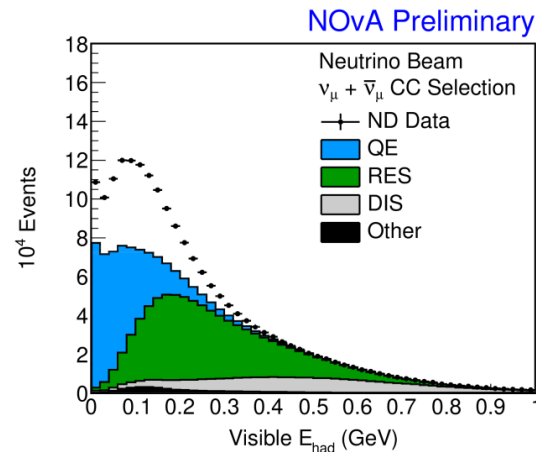
Our tuning is done in a two-dimensional space of the four-momentum transfer variables:

**energy transfer  $q_0$   
and  
momentum transfer  $|q|$**



fit a weight factor for each cell in this plot

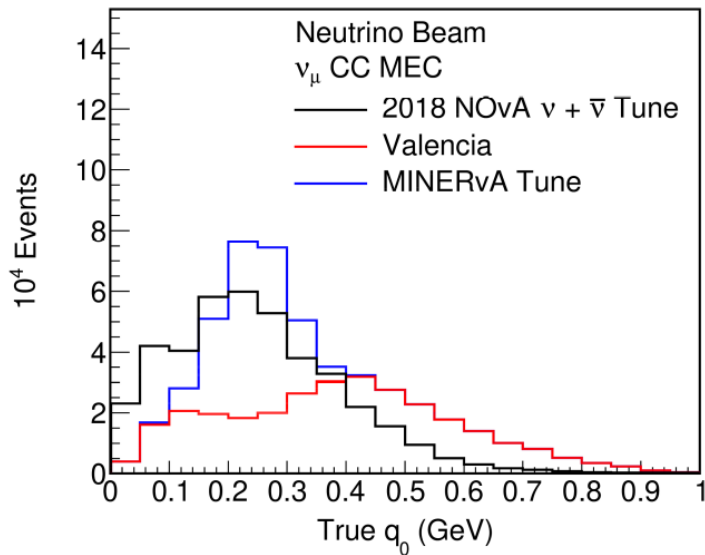
Fit in 2D space of nearest observables:  
Visible  $E_{\text{had}}$  ( $\sim q_0$ ) and reco  $|q|$



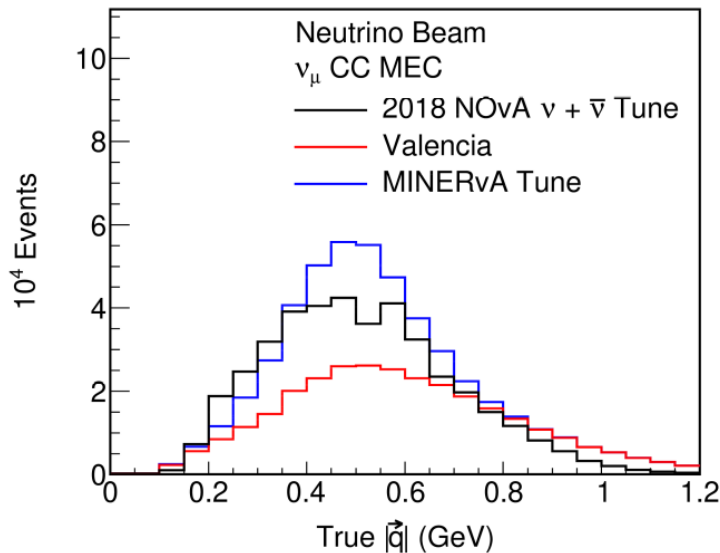
Resulting MEC shape

# Modeling the nucleus: tuning 2p2h-MEC

NOvA Simulation



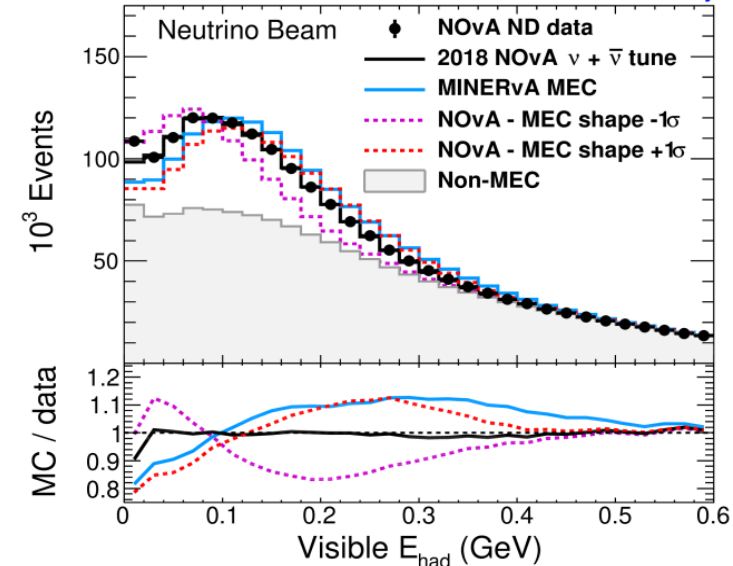
NOvA Simulation



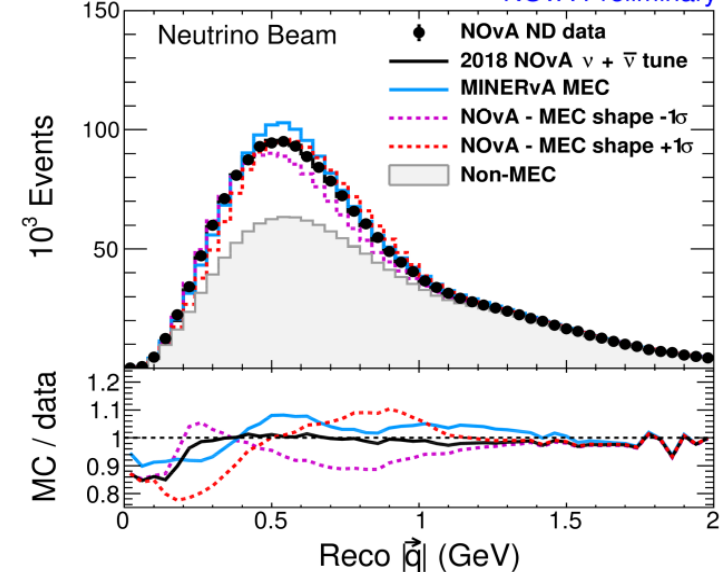
MINERvA carried out a tuning procedure similar in spirit to ours (though with fewer degrees of freedom) using their data (PRD 116, 071802) which they kindly shared with us (private communication).

It is not dissimilar to the  $1\sigma$  error band we arrive at (details on error construction next slide)

NOvA Preliminary



NOvA Preliminary





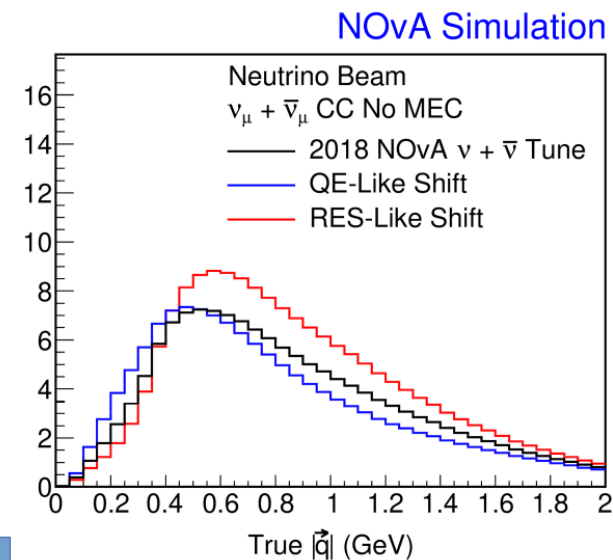
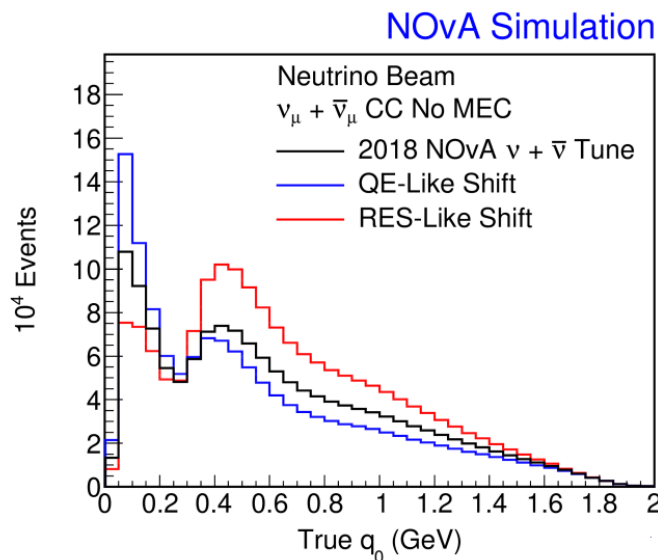
# Modeling the nucleus: 2p2h-MEC uncertainties

## Non-MEC base

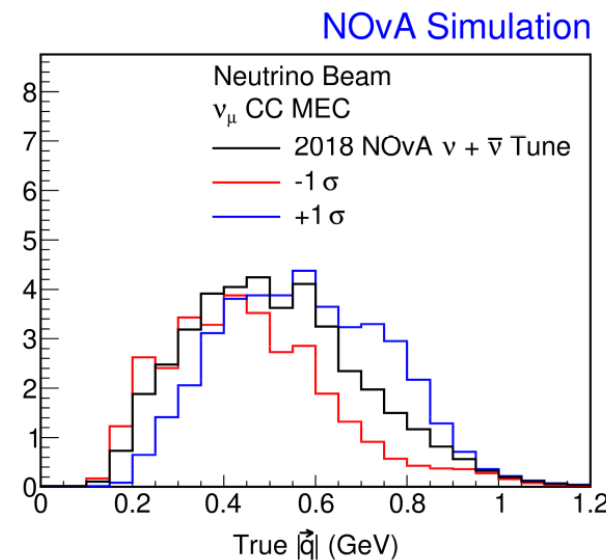
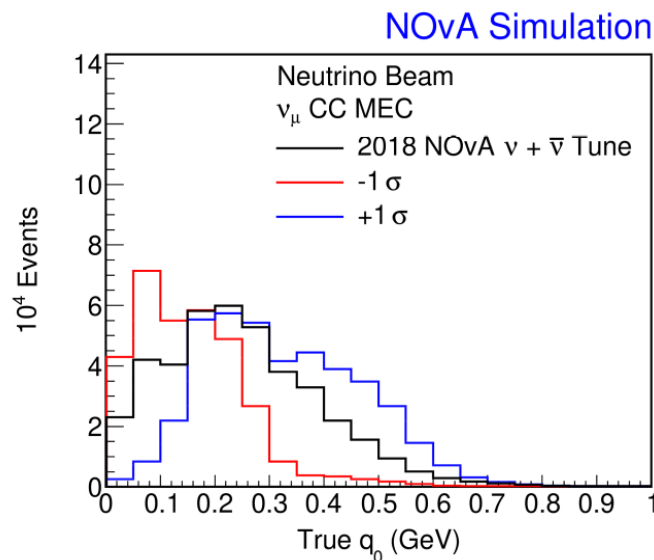
Two alternate fits:

Choose combinations of uncertainties to push initial MC more towards QE or RES

Knob	"QE-like" shift	"RES-like" shift
QE MA	+1 $\sigma$ (+5%)	-1 $\sigma$ (-5%)
QE RPA low-Q <sup>2</sup>	+1 $\sigma$	-1 $\sigma$
QE RPA high-Q <sup>2</sup>	+1 $\sigma$	-1 $\sigma$
QE Pauli Supp.	-1 $\sigma$	+1 $\sigma$
RES MA	-1 $\sigma$	+1 $\sigma$
RES MV	-1 $\sigma$	+1 $\sigma$
RES RPA	on (CV)	off

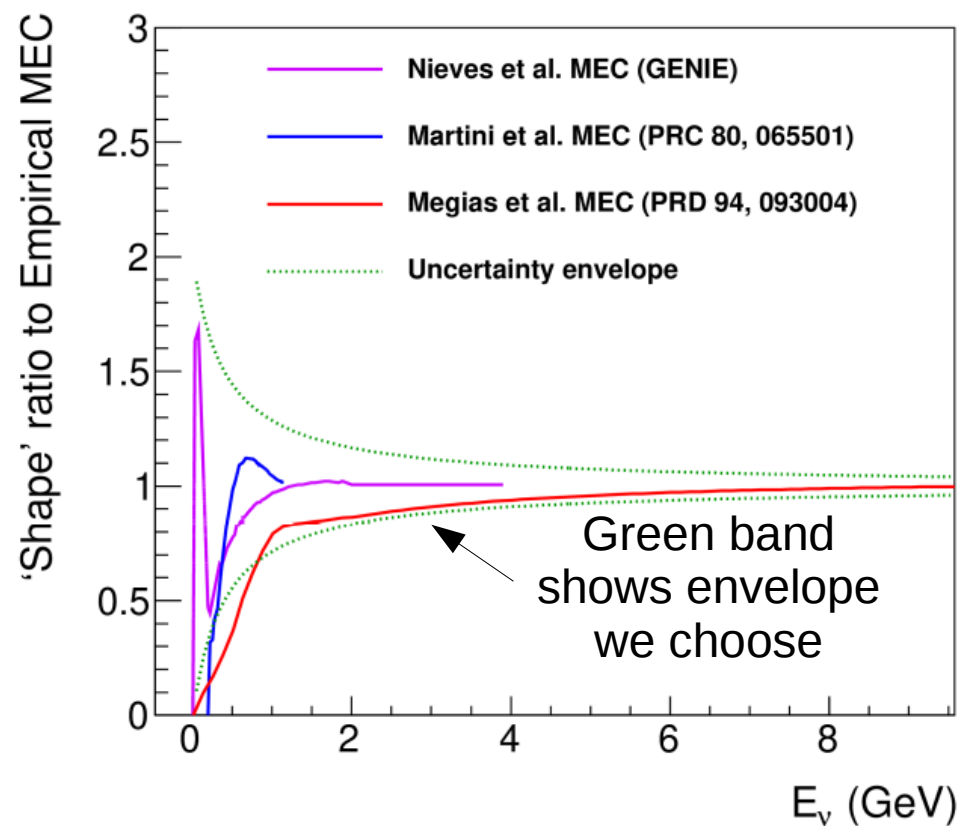
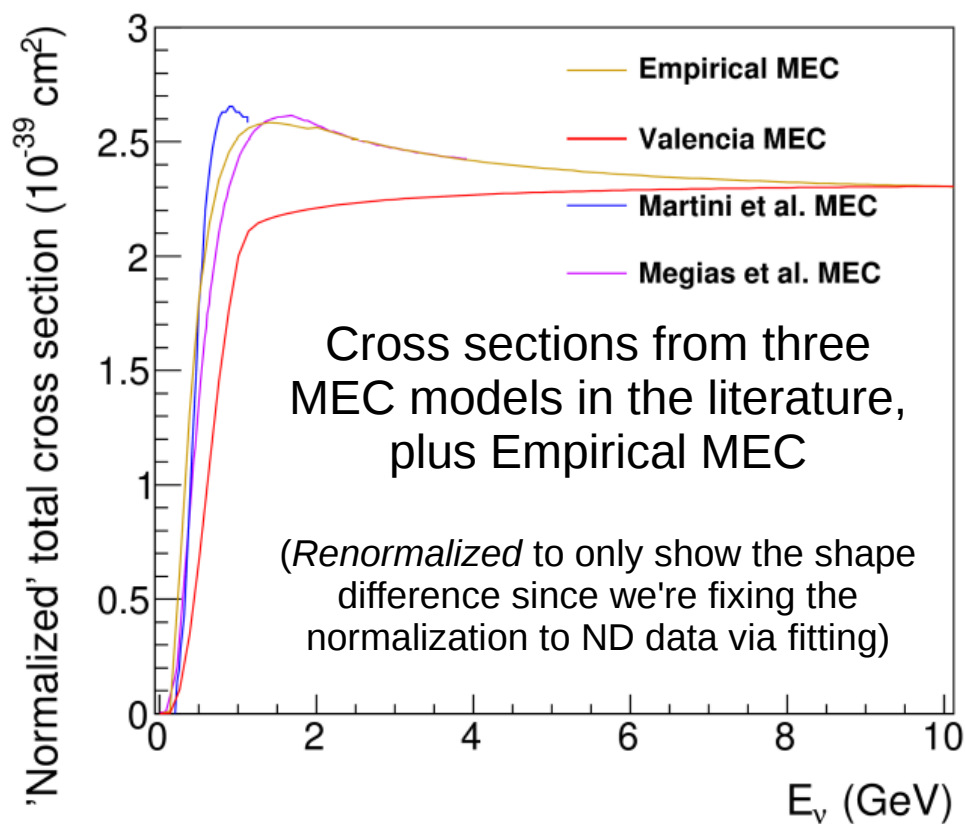


## Fitted MEC



# Modeling the nucleus: 2p2h-MEC uncertainties

## Cross section $E_\nu$ shape



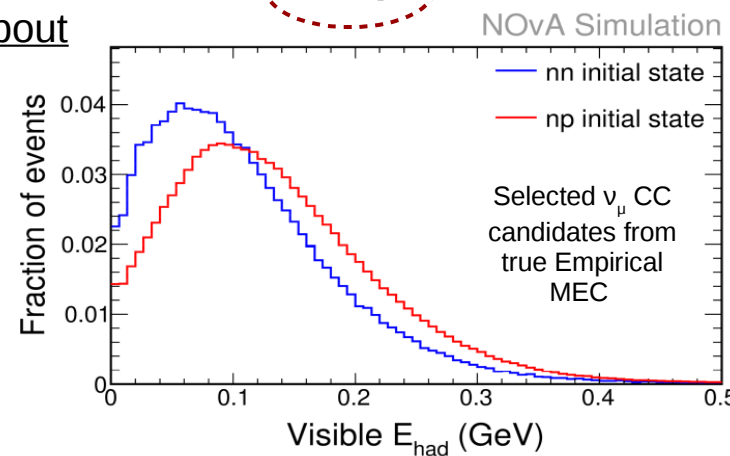
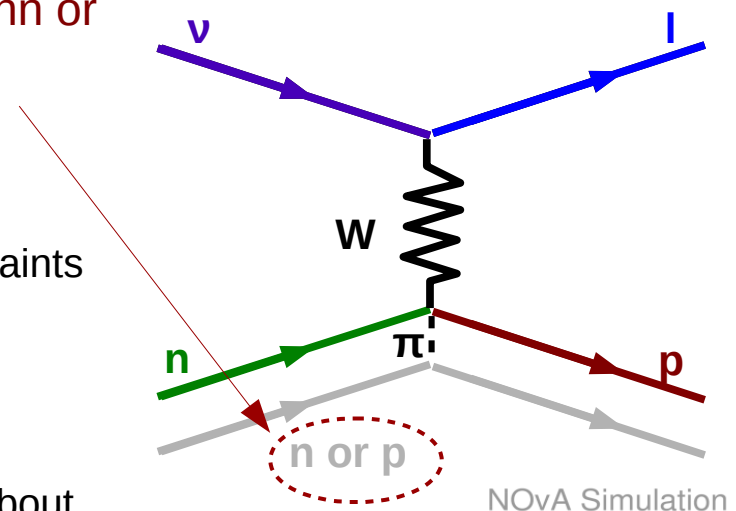
**Choose an envelope that more or less encloses the shapes of the predictions for our “ $\pm 1\sigma$ ” uncertainty**

# Modeling the nucleus: 2p2h-MEC uncertainties

## nn-np initial state composition

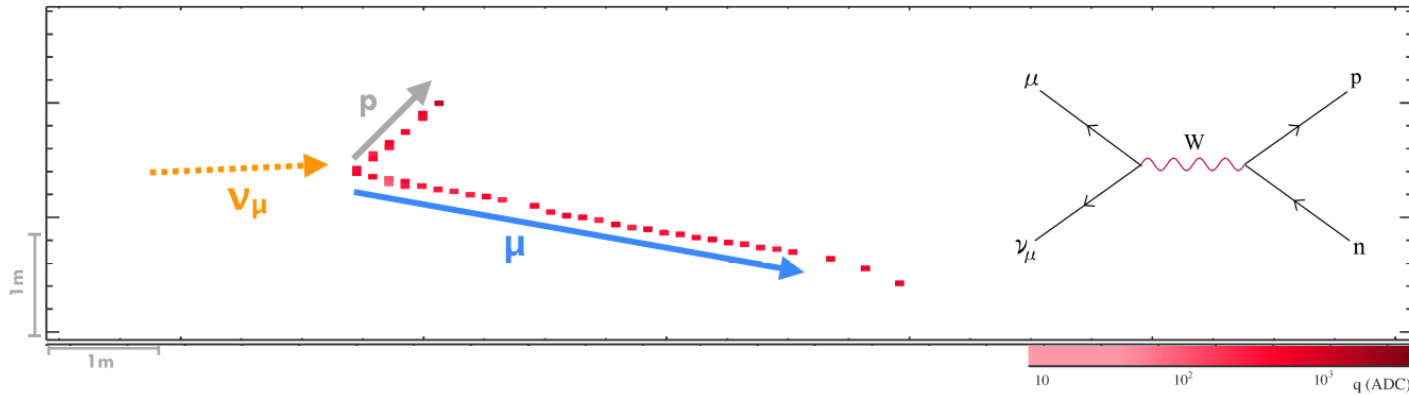
- Diagrams for  $\nu$  CC 2p2h allow two nucleon “pairs” in initial state: nn or np ( $\bar{\nu}$  has np or pp)
- Challenging to measure the real composition in data
  - LAr will help eventually?
  - MINERvA has made valiant efforts in the meantime, but not strong constraints on the *value* of the ratio (yet?)
- **Stuck with theory for now**
  - València prediction (via GENIE):  $\sim 70\%$  np/(nn+np).
  - SuSA prediction (PRC 94, 054610), detailed study: “The [np/nn] ratio is about 5-6 [i.e., np/(nn+np)  $\sim$  80-90%] for a wide range of neutrino energies.”
  - Empirical MEC default is 80%

We choose  $0.7 \leq \frac{np}{(np+nn)} \leq 0.9$  at  $1\sigma$ .



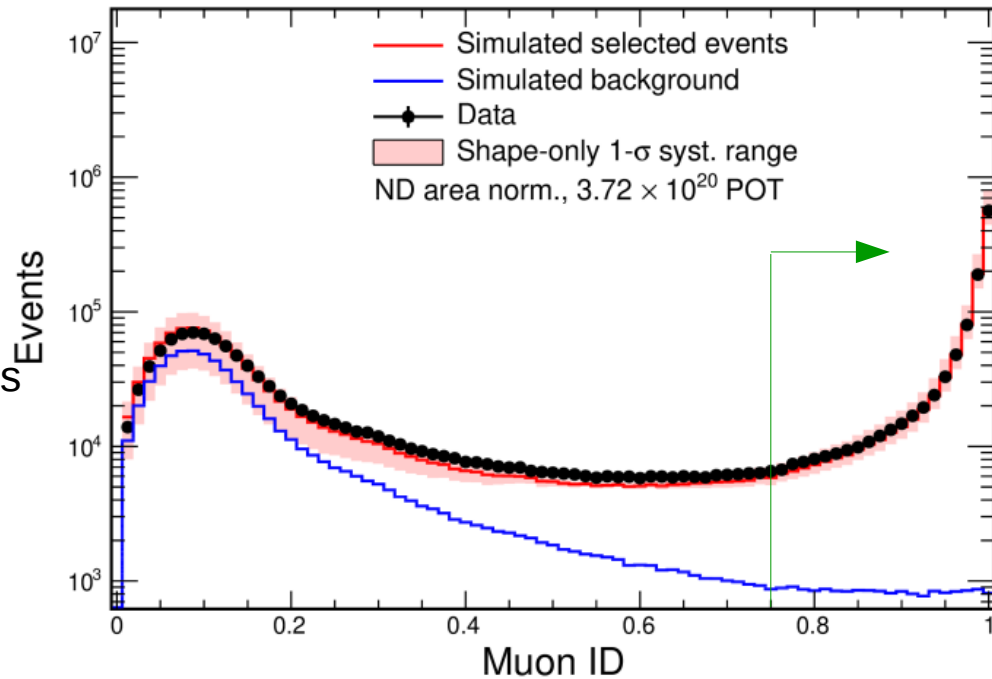
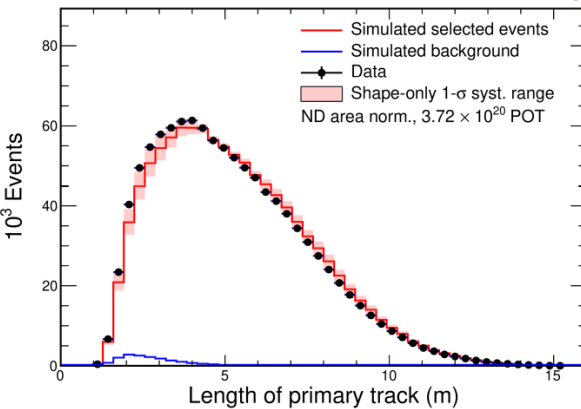
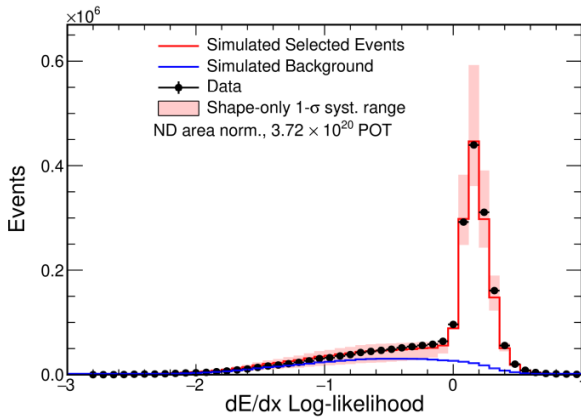
(It doesn't matter much; GEANT says we get  $\sim$ similar response)

# $\nu_\mu$ disappearance: selection



**kNN-based  $\nu_\mu$  CC classifier uses 4 inputs:**

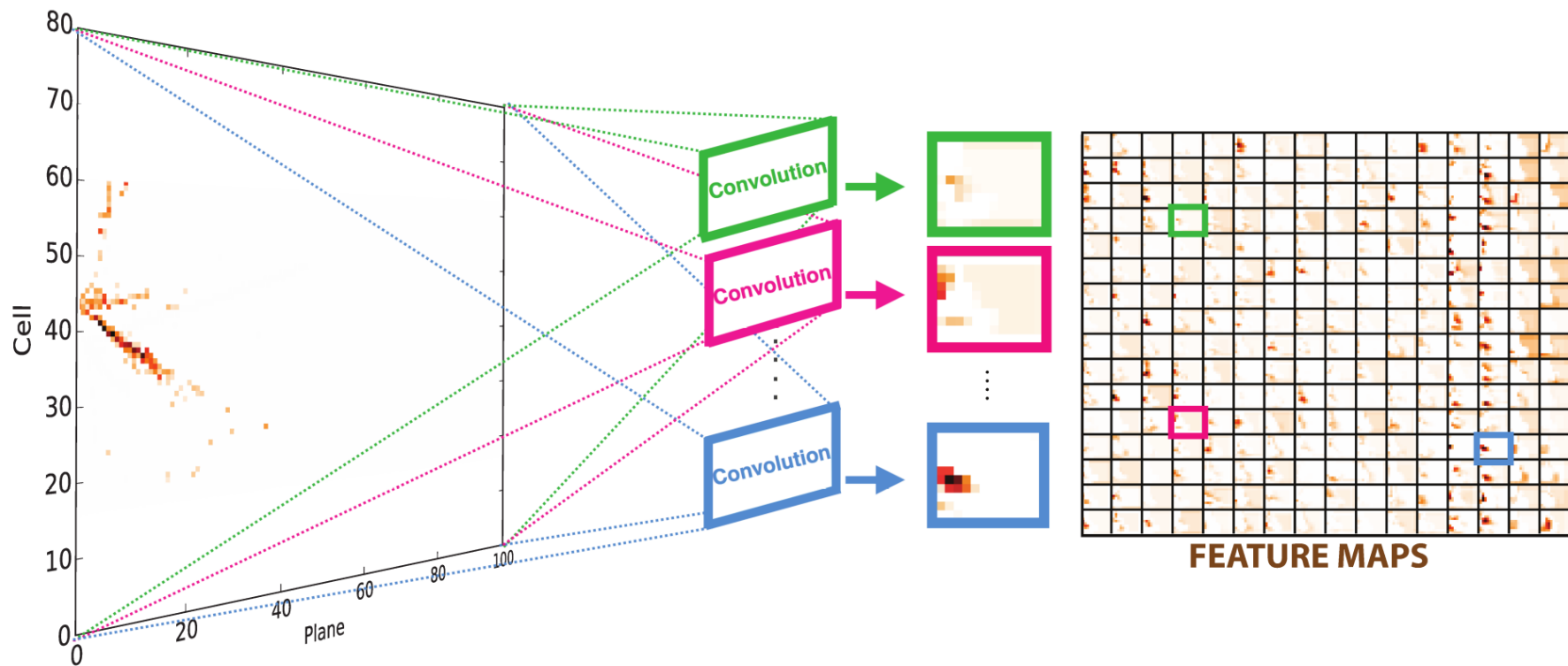
- Track length
- $dE/dx$
- Multiple scattering
- Fraction of track planes consistent w/ single particle  $dE/dx$



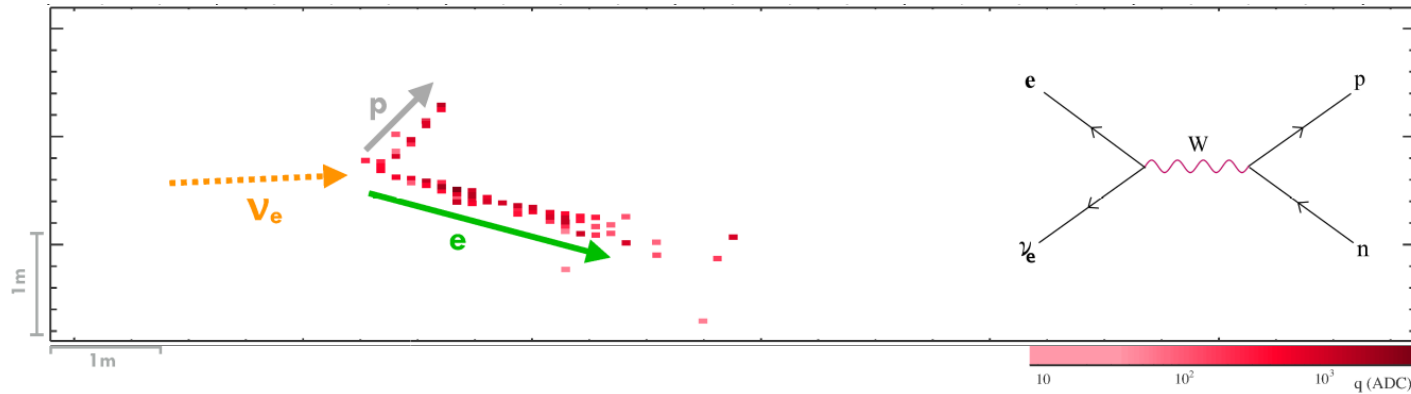
# $\nu_e$ appearance: selection

Event selection via a “Convolutional Neural Network”:

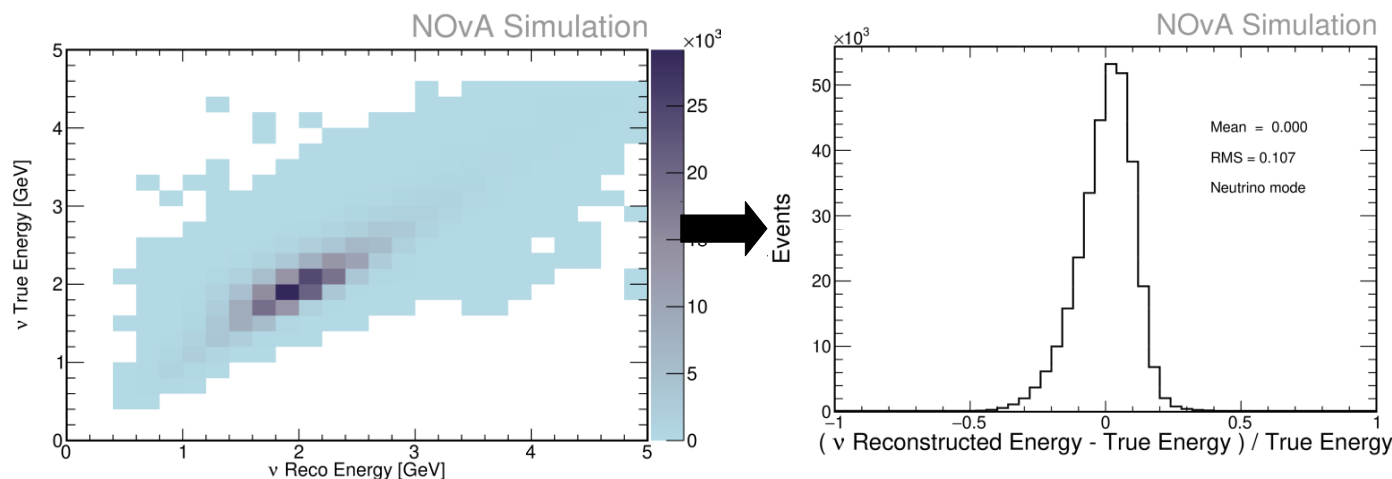
energy deposition patterns treated as images, algorithm extracts representative abstract features by applying learned filters



# $\nu_e$ appearance: selection & reconstruction

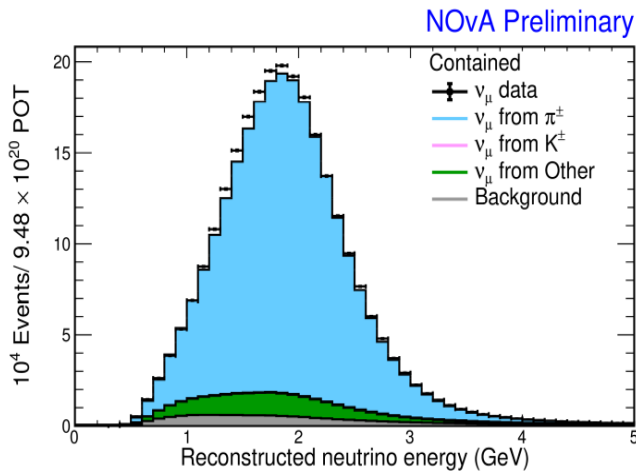
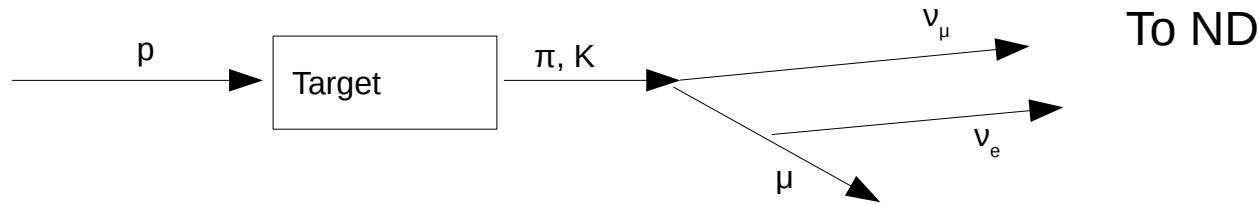


**Convolutional neural network** selects events via transformations applied to energy deposits treated as images

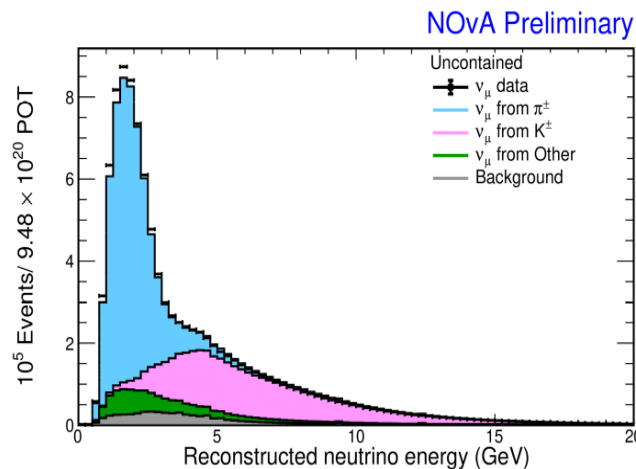
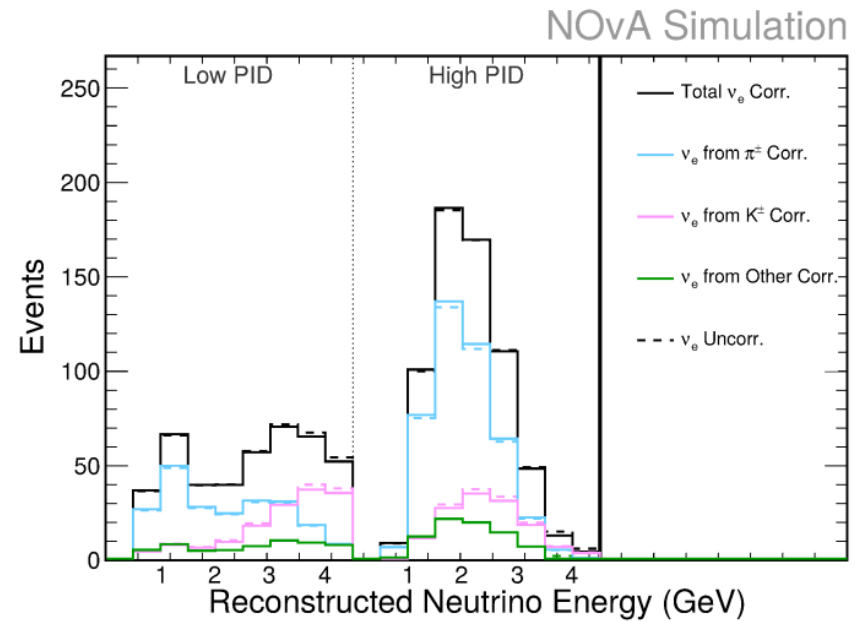


**Energy estimator** is quadratic function of  $E_e$  and  $E_{had}$ .  
~11% resolution

# $\nu_e$ appearance: constraining beam $\nu_e$ bknd



Assign discrepancies in ND  $\nu_\mu$  contained and uncontained samples to flux; **derive corrections according to parent mesons** (which also result in beam  $\nu_e$ )

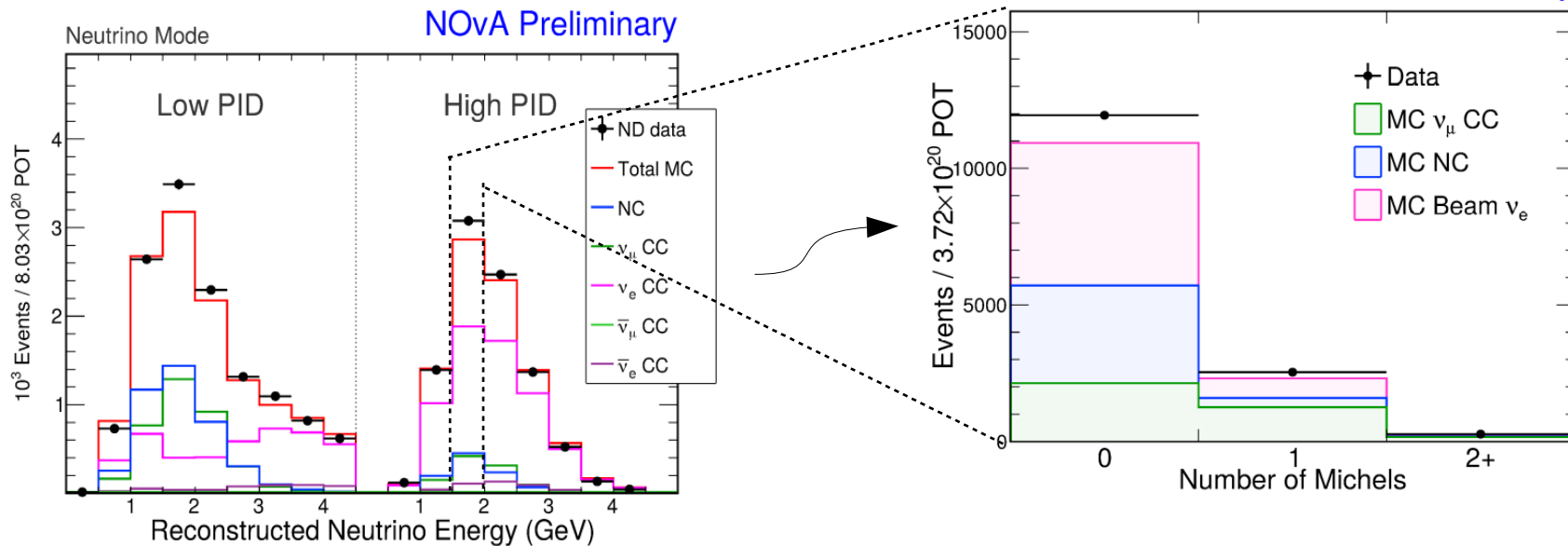


Pion-ancestor neutrinos are corrected in bins of parent ( $p_z$ ,  $p_T$ ). Average  $\sim +2\%$

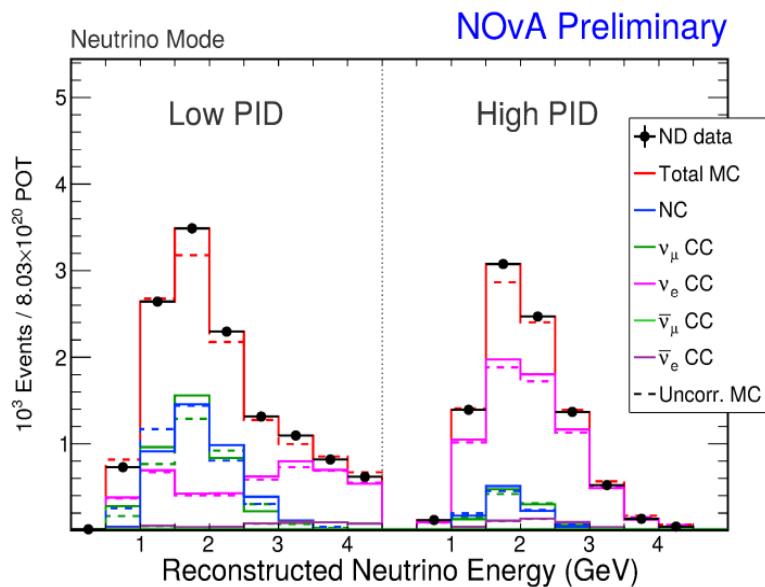
Kaon-ancestor neutrinos get a single weight:  $-6.3\%$

# $\nu_e$ appearance: constraining $\nu_\mu$ CC/NC ratio

NOvA Preliminary



Examine distribution of Michel electrons in each bin of ND  $\nu_e$  selected sample after beam  $\nu_e$  constraint (prev slide)

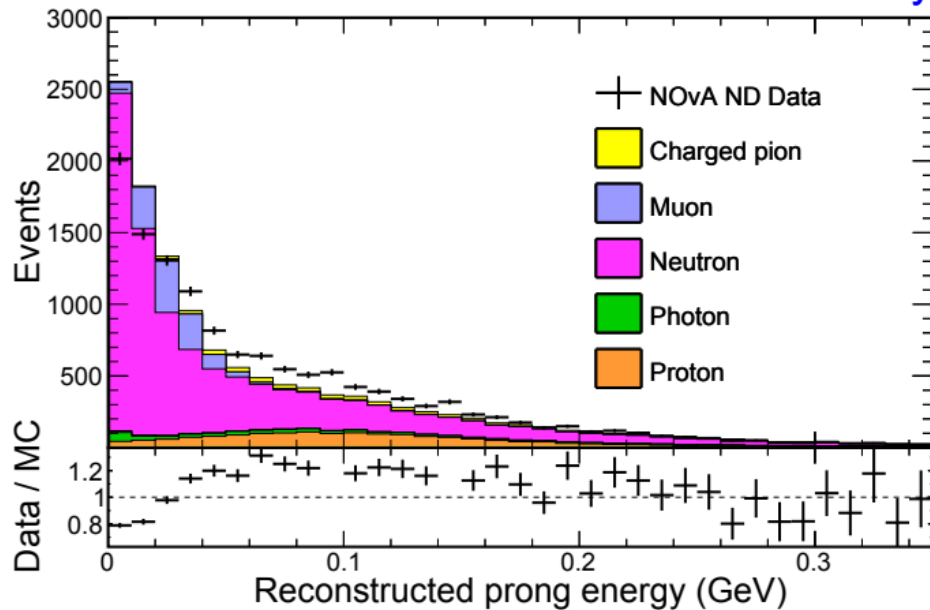


Fit these 18 distributions to determine  $\nu_\mu$  CC / NC corrections in each bin

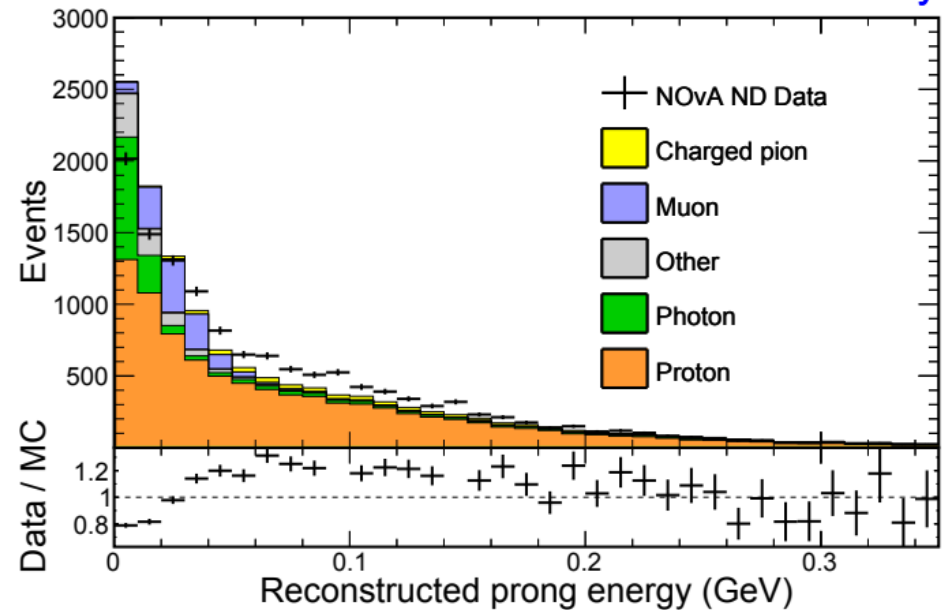


# Neutron response

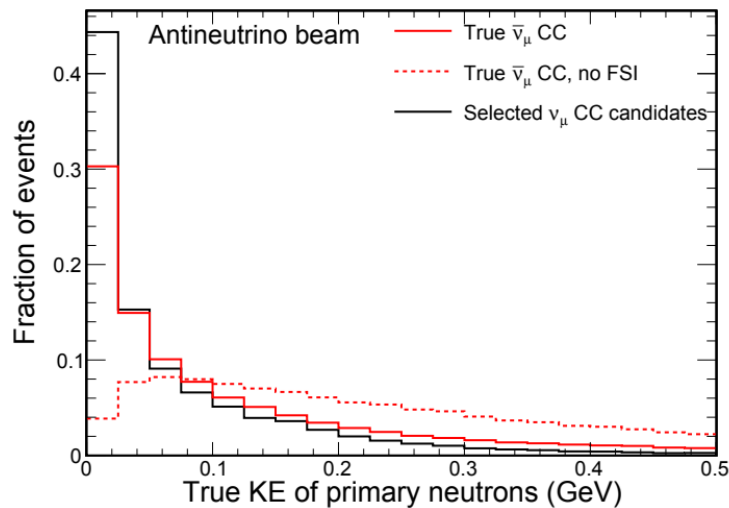
NOvA Preliminary



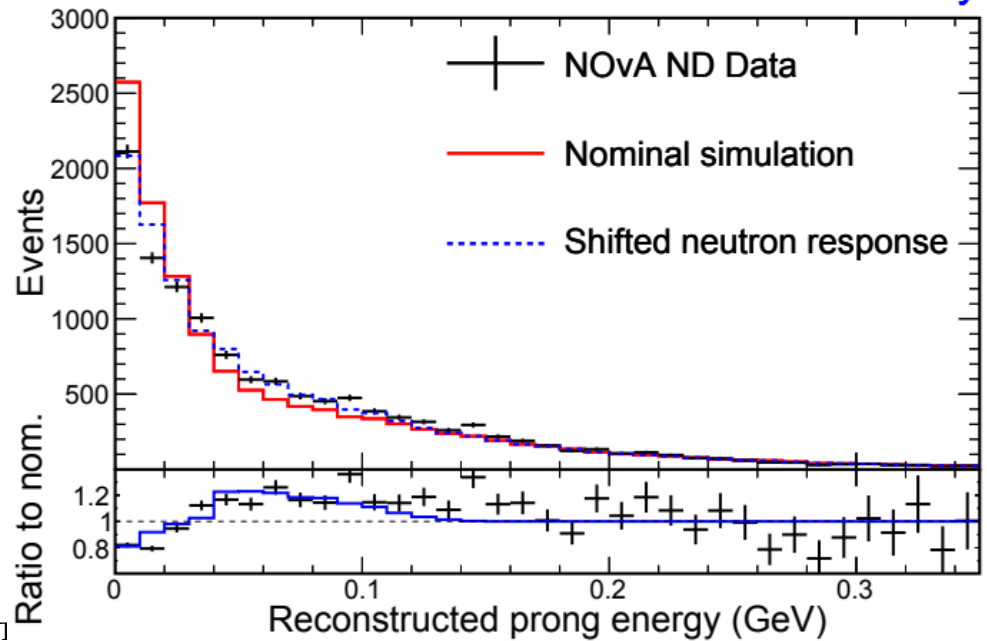
NOvA Preliminary



NOvA Simulation

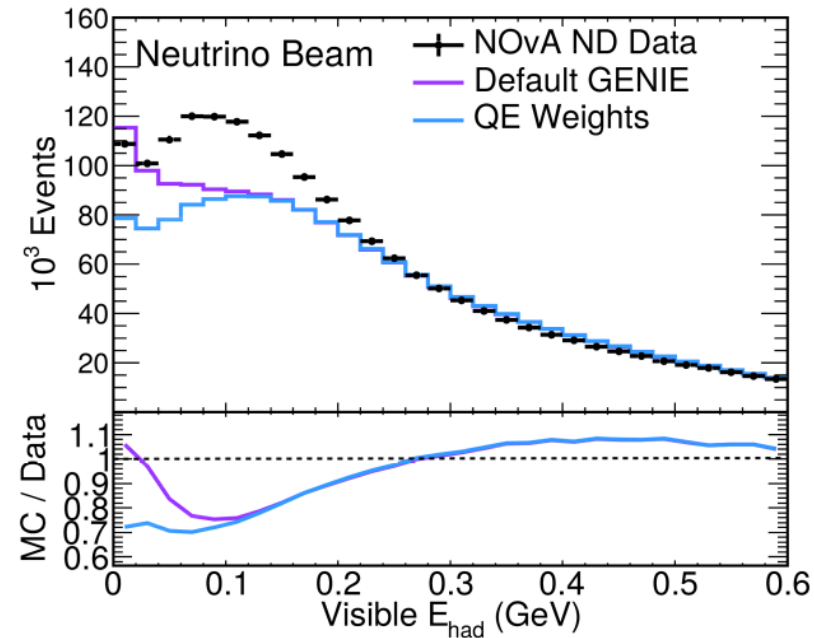


NOvA Preliminary



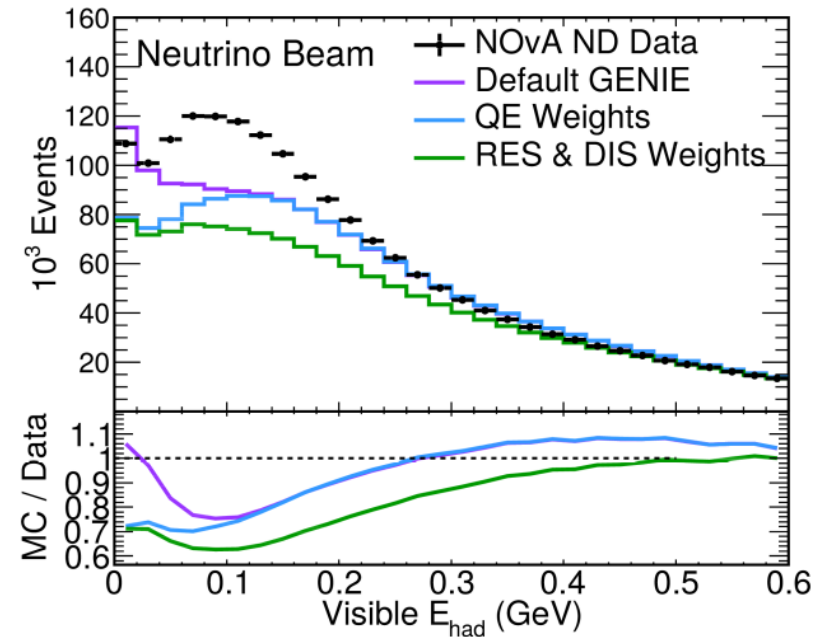
# Cross section uncertainties: future needs

- More certainty about **1p1h initial state**
  - **RPA treatments differ in sophistication** - how much detail do we need?
  - **Uncertainties** (from València) **still large**, not completely canceled by extrapolation
  - **Binding/separation/removal energy** situation (A. Bodek, [arXiv:1801.07975](https://arxiv.org/abs/1801.07975))...



# Cross section uncertainties: future needs

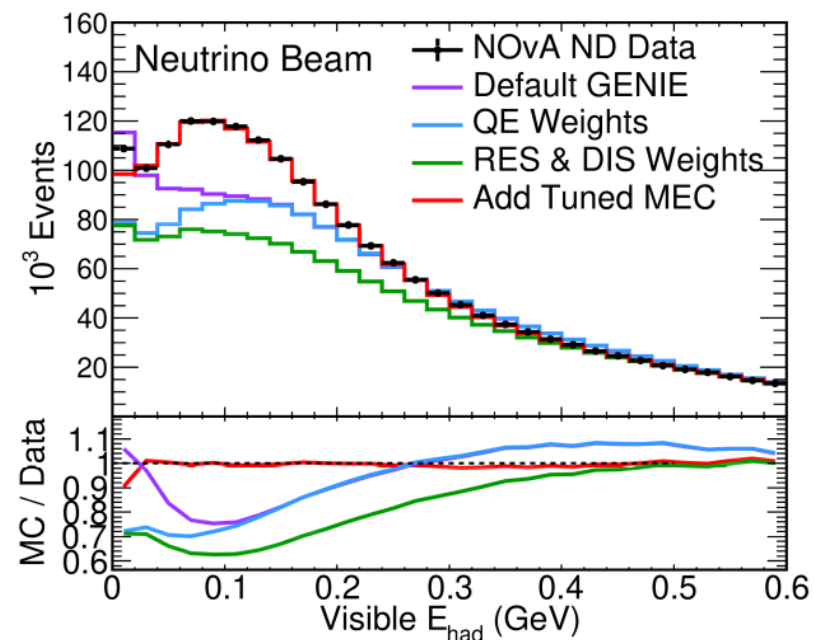
- **Nuclear models for *inelastic* processes** as well as QE
  - **Apparent low- $Q^2$  suppression relative to RFG suggests nuclear effects important**
    - Suggested by data (MINERvA+MINOS ND+MiniBooNE+NOvA ND)... no theory basis (yet?)
    - “On-off” treatment for syst one of our largest
  - Inelastic continuum at low  $E_\nu$ 
    - What does “shallow” inelastic scattering on carbon at  $E_\nu = 2$  GeV look like?
      - How does it interfere with RES? → GENIE uncertainties large
      - Free nucleon data helps only so much
    - Does **diffractive scattering from H** matter? How close are models?
- **$\nu_e/\nu_\mu$  differences for inelastic processes**
  - Current uncertainties are *ad hoc*



[NOvA has cross section measurements in progress which will help address some of these questions:  
see L. Aliaga's and M. Judah's talks, next]

# Cross section uncertainties: future needs

- More/better **models for multinucleon knockout** in GENIE
  - València model agrees poorly with MINERvA, NOvA ND data; **no alternatives in current versions** (but good things on the horizon?)
  - GENIE **assumes nucleon pair ejected isotropically**. Need (“semi-inclusive”) models for what should *really* happen
  - Empirical tuning procedure doesn't prescribe **correlations between  $\nu$  and  $\bar{\nu}$**  – so left uncorrelated...



# Cross section uncertainties: future needs

- **Much more  $\bar{\nu}$  scattering data**
  - Every issue mentioned above applies also for antineutrinos, only there are **fewer data constraints**
  - Abundance of fast neutrons an interesting challenge for calorimetry: **final-state particle measurements** especially helpful

