



Determination
of the LBNF
Neutrino Flux

Mary Bishai
(on behalf of
LBNF/DUNE)
Brookhaven
National Lab

The LBNF
Beamline

Flux Modeling
and
Uncertainties

Future had.
prod.
measurements

Near
Detector(s)
Flux
Measurements

Muon
Monitors

Summary

Determination of the LBNF Neutrino Flux

NuInt 2018, 15-19 October 2018, Gran Sasso Science Institute, Italy

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- 2** Flux Modeling and Uncertainties
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The LBNF Beamline

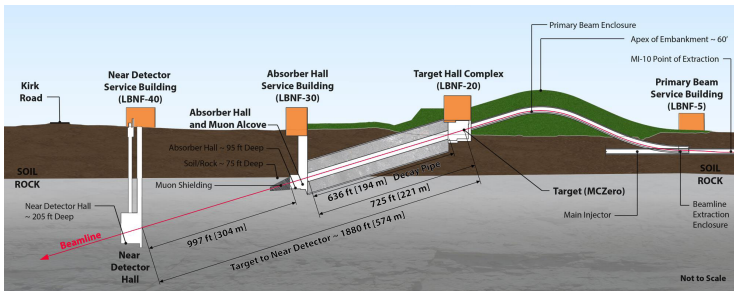
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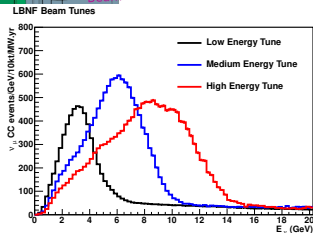
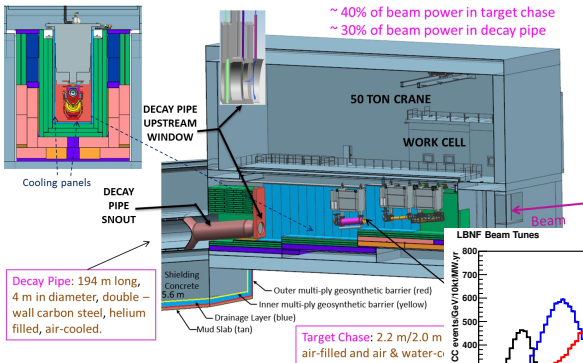
Summary



- Primary proton beam 60-120 GeV
- Initial 1.2 MW beam power, upgradable to 2.4 MW
- Embankment allows target complex to be at grade
- Wide-band configurable beam (on-axis) optimized for CP Violation sensitivity
- Decay pipe: 194m x 4m diameter, He filled

ND default: 574 m from target, FD: 1297 km

Initial conceptual design was of a *tunable wide-band* NuMI-style focusing:



LBNF has switched to CPV optimized focusing design with 3 horns

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Optimization of flux for Physics: CP Violation

Laura Fields

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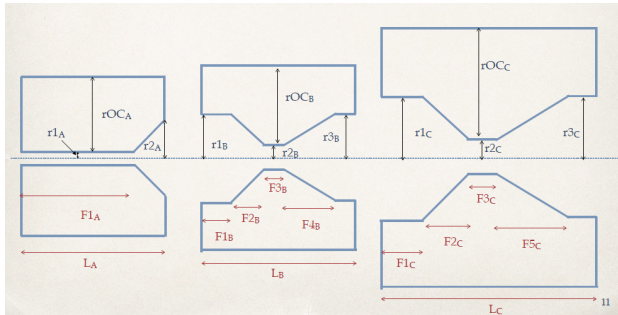
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- In Sep 2017 LBNF adopted a focusing design with 3 horns optimized using a *genetic algorithm* with the physics parameter to be measured (CPV sensitivity) used to gauge fitness.
- Target geometry is optimized at the same time, as well as proton beam energy with realistic Main Injector power profile (1.03 MW at 60 GeV to 1.2 MW at 120 GeV).



Optimized horn design with 297kA current :



Parameter	Value	Parameter	Value
Horn A Length (mm)	2218	Horn A F1 (% of length)	53
Horn A R1 (mm)	43	Horn A OC Radius (mm)	369
Horn A R2 (mm)	33		
Horn B Length (mm)	3932	Horn C Length (mm)	2184
Horn B R1 (mm)	159	Horn C R1 (mm)	284
Horn B R2 (mm)	81	Horn C R2 (mm)	131
Horn B R3 (mm)	225	Horn C R3 (mm)	362
Horn B F1 (% of length)	31	Horn C F1 (% of length)	20
Horn B F2 (% of length)	22	Horn C F2 (% of length)	9
Horn B F3 (% of length)	2	Horn C F3 (% of length)	7
Horn B F4 (% of length)	16	Horn C F4 (% of length)	35
Horn B OC Radius (mm)	634	Horn C OC Radius (mm)	634
Horn B Position (mm)	2956	Horn C Position (mm)	17806

Optimized target is 4λ (2m C) with $\sigma_{\text{beam}} = 2.7\text{mm}$, $E_p \sim 110\text{ GeV}$



Optimization of flux for Physics: CP Violation

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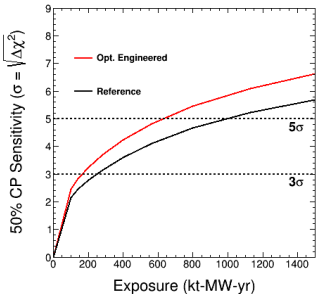
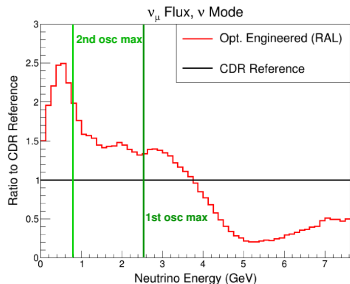
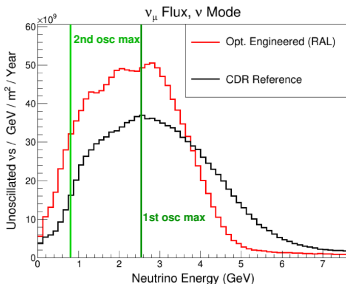
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Computationally advanced optimization techniques = significant gain in flux and CPV sensitivity from many small changes

Gain in sensitivity \equiv 70% increase in FD mass for goal of $\geq 3\sigma$ CPV sensitivity over 75% of δ_{CP} values.

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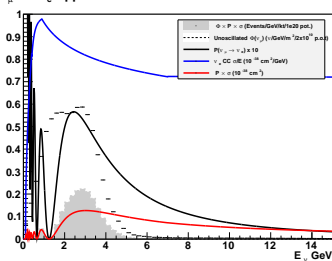
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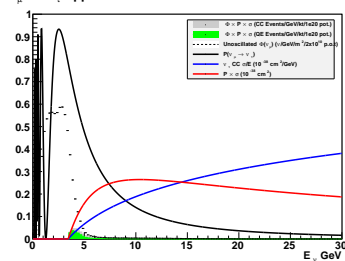
NuMI-like reference design could be tuned to higher energy to observe $\nu_\mu \rightarrow \nu_\tau$ with high statistics.

2015 two horn optimized design $E_p = 66$ GeV:

$\nu_\mu \rightarrow \nu_e$ Appearance at 1300 km



$\nu_\mu \rightarrow \nu_\tau$ Appearance at 1300 km



$\nu_\mu \rightarrow \nu_e$ 290 events

$\nu_\mu \rightarrow \nu_\tau$ 60 events

in 40 ktons, 1 year at 1.2 MW

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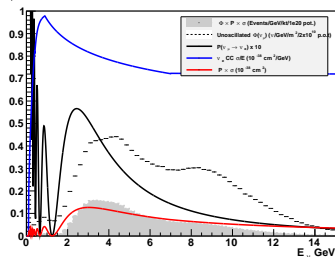
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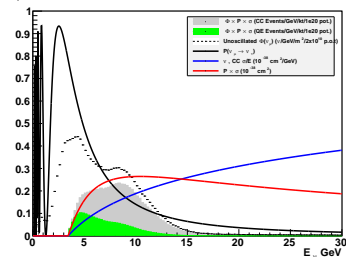
NuMI-like reference design could be tuned to higher energy to observe $\nu_\mu \rightarrow \nu_\tau$ with high statistics.

LBNF target -2m from horn 1, NuMI focusing 230 kA, horns 17m apart

$\nu_\mu \rightarrow \nu_e$ Appearance at 1300 km



$\nu_\mu \rightarrow \nu_\tau$ Appearance at 1300 km



$\nu_\mu \rightarrow \nu_e$ 330 events $\nu_\mu \rightarrow \nu_\tau$ 700 events
in 40 ktons, 1 year at 1.2 MW

Increase ν_τ appearance 10x !! Potential for large ν_τ data set.

Increase high energy ν_e appearance - good for NSI/Sterile searches

Flux components at near and far

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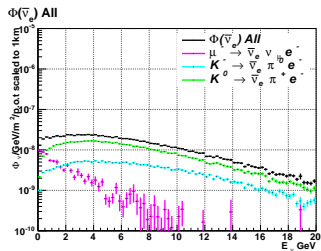
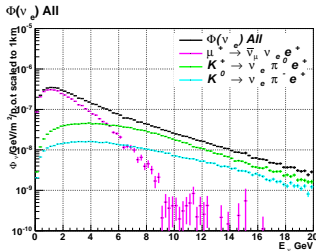
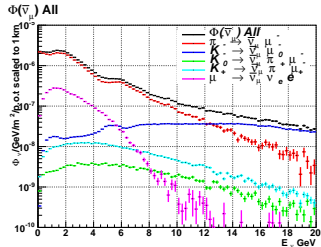
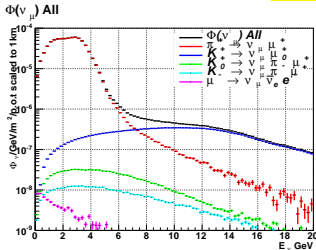
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FD 1300km



Baseline scaled to 1km from middle of decay channel



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The main sources of flux modeling uncertainties are:

- **Hadron production uncertainties:** driven by uncertainties in the hadron interaction models used to estimate hadron distributions exiting the target (prior to focusing) as well as secondary and tertiary interactions of hadrons with beamline material. **Fully evaluated for LBNF/DUNE using the ppx package developed for MINER ν A.**
- **Focusing uncertainties:** Dominated by horn material, geometry and magnetic field modeling as well as target geometry and density. Alignment of the neutrino beamline elements can also have large impact on ν flux. Includes proton counting uncertainties. **These uncertainties are assessed by simulating individual effects in Geant 4 and combining.**
- **Other beamline uncertainties:** Primarily uncertainties on the distribution of passive material in the beamline: for e.g. impact of Nitrogen in the target chase, decay pipe window thickness...etc. **Experience with NuMI indicates these are subdominant**

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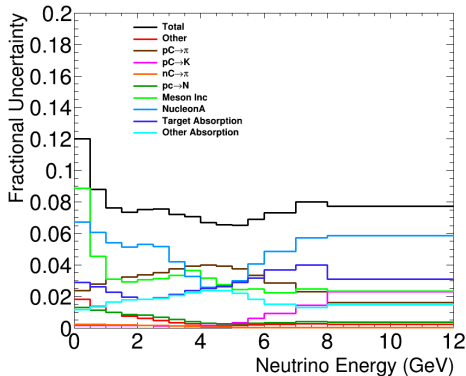
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Hadron Prod. Uncertainties

NA49/MIPP/older datasets used to constrain $pC \rightarrow \pi^\pm, K^\pm, n(p)X$
 Pion production by neutrons from data (assuming isospin symmetry)
 Nucleon incident interactions not covered by data

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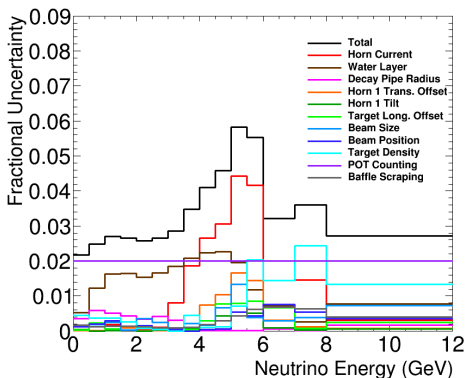
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Focusing Uncertainties

Detailed focusing uncertainties based on the NuMI experience in MINER ν A. **Detailed estimates for both 2015 NuMI-like design and CPV optimized design with simplified 2 horns.**

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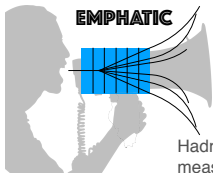
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A new hadron production experiment for improved neutrino flux predictions

Hadron production measurements to reduce flux uncertainties from secondary and tertiary interactions in the neutrino production target

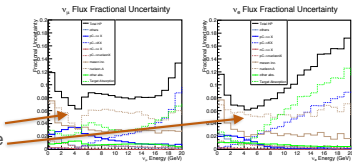
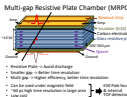
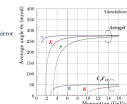
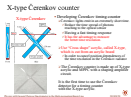
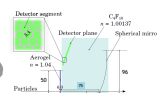
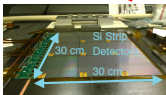
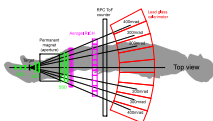
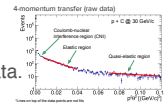


Table-top-sized experiment



2 week-long test run this past January already collected useful data.



See posters by J. Paley, M. Pavin & T. Vladislavljevic, and T. Sugimoto!



A Spectrometer for Hadron Flux Measurements in LBNF?

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Proposal by Laura Fields at Fermilab:

The **LBNF Spectrometer** is a concept for a thick-target hadron production measurement after the focusing horns. It would involve a **replica of the LBNF target and horns** in an external beamline at Fermilab. In addition to hadron production in the target, the spectrometer would also **measure hadron production and absorption in the horns and the effects of the magnetic fields** in the horns.



Detector technology is always challenging

Need to get more people interested and involved to succeed

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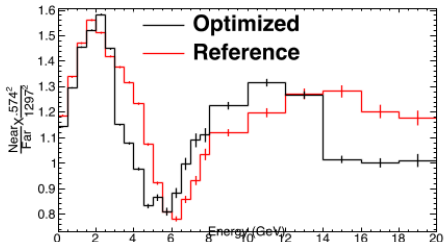
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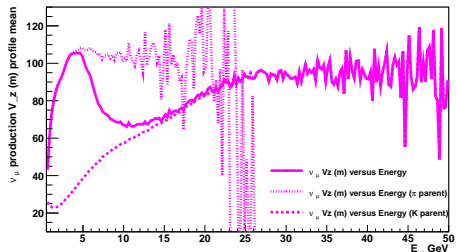
Simple ratio of near spectrum/far spectrum:

Neutrino parent decay location in decay pipe:

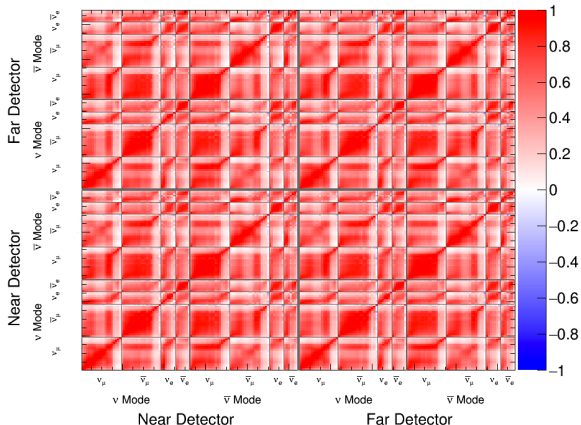
π/K decay kinematics and decay channel geometry are primary reason for strange shape of N/F ratio



ν_{μ} events at FD (1300km)



To correctly relate near to far fluxes - need to use a correlation matrix:



Flux correlation matrix comes from simulation and is highly correlated

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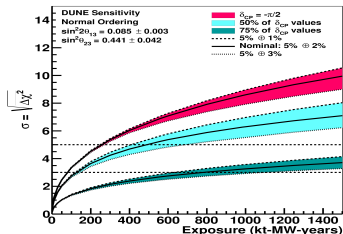
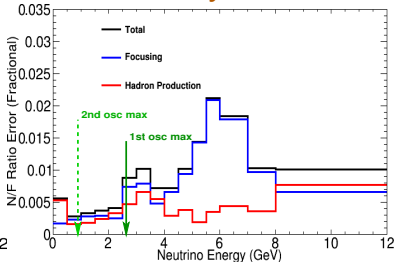
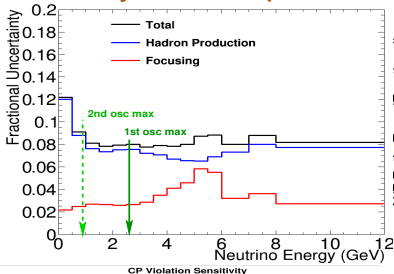
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Summary

Uncertainty on FD flux prediction Residual uncertainty on flux at FD



How well do we actually trust the simulation to correctly estimate the uncertainties on near → far extrapolation?



Near Detector Flux Measurement Strategies

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Technique	Flavor	Absolute normalization	Relative flux $\Phi(E_\nu)$	Near Detector requirements
NC Scattering $\nu_\mu e^- \rightarrow \nu_\mu e^-$	ν_μ	2.5%	$\sim 5\%$	e^- ID θ_e Resolution e^-/e^+ Separation
Inverse muon decay $\nu_\mu e^- \rightarrow \mu^- \nu_e$	ν_μ	3%		μ^- ID θ_μ Resolution 2-Track ($\mu+X$) Resolution μ energy scale
CC QE $\nu_\mu n \rightarrow \mu^- p$ $Q^2 \rightarrow 0$	ν_μ	3 – 5%	5 – 10%	D target p Angular resolution p energy resolution Back-Subtraction
CC QE $\bar{\nu}_\mu p \rightarrow \mu^+ n$ $Q^2 \rightarrow 0$	$\bar{\nu}_\mu$	5%	10%	H target Back-Subtraction
Low-ν_0	ν_μ		2.0%	μ^- vs μ^+ E_μ -Scale Low- E_{Had} Resolution
Low-ν_0	$\bar{\nu}_\mu$		2.0%	μ^- vs μ^+ E_μ -Scale Low- E_{Had} Resolution
Low-ν_0	$\nu_e/\bar{\nu}_e$	1-3%	2.0%	e^-/e^+ Separation (K_L^0)
CC	ν_e/ν_μ	<1%	$\sim 2\%$	e^- ID & μ^- ID p_e/p_μ Resolution
CC	$\bar{\nu}_e/\bar{\nu}_\mu$	<1%	$\sim 2\%$	e^+ ID & μ^+ ID p_e/p_μ Resolution
Low-ν_0/CohPi	$\bar{\nu}_\mu/\nu_\mu$	$\sim 2\%$	$\sim 2\%$	μ^+ ID & μ^- ID p_μ Resolution E_{Had} Resolution

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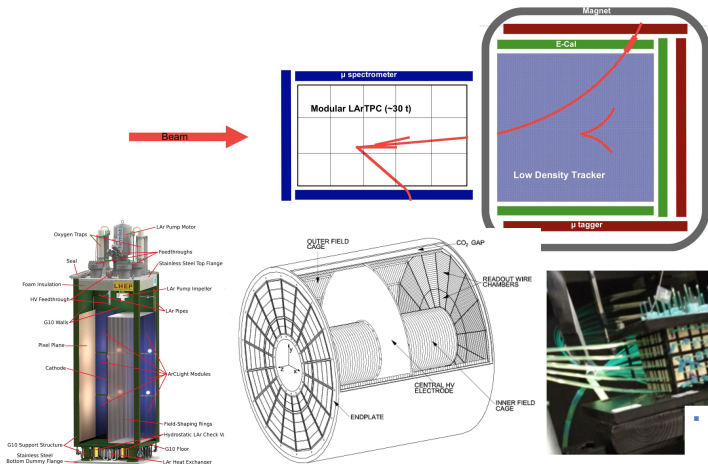
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Summary

An unmagnetized LArTPC with followed by a low density tracker embedded in a large $\sim 0.5T$ magnet with EM sampling calorimeters.



Segmented LArTPC

ALICE-like GARTPC

3D scint

5x5 cm³ prototype beam test @CERN, Oct 2017



DUNE ND Capabilities

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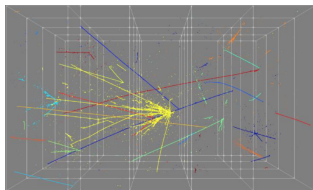
Summary

LArTPC: Allows for high statistics measurements on Ar and ND/FD detector systematics constraints. **With 25 ton FV \Rightarrow 37M CC ν_μ interactions/year (1MW beam)**

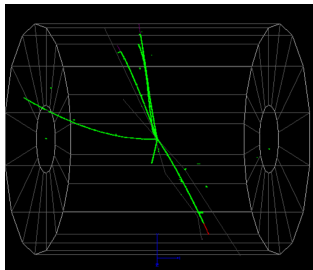
GArTPC: allows for precision measurement of ν -Ar vertex activity using Ar-CH₄. Low threshold improves low- ν on Ar. Could allow for different gas mixtures (H???).

3DST: Increases interaction statistics in magnetized volume \sim **6 t fiducial mass \Rightarrow 9M ν_μ events/yr.** Beam flux measurements on a lighter target. Contain EM showers from $\nu - e$ scattering when combined with ECAL (\sim 1000 events/yr).

Challenge: how to integrate with GArTPC.



ND LArTPC simulation



ND GArTPC simulation

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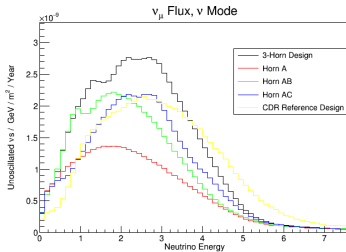
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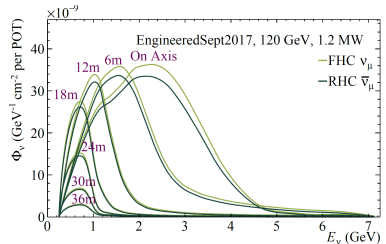
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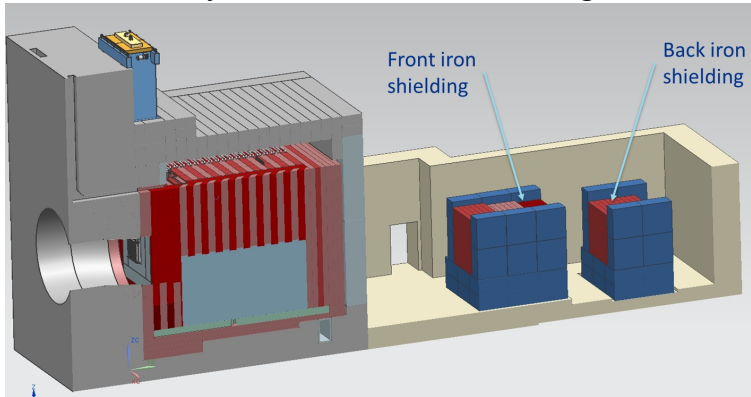
ν_μ flux different horn configs



$\nu_\mu/\bar{\nu}_\mu$ off-axis fluxes at ND

More flux and focusing constraints could be obtained by combining information from **ND off-axis measurements (DUNE-PRISM) and varying on-axis beam tunes (with varying horn current for e.g.)**. Studies are under way.

Layout of Muon Alcove and Shielding



High intensity makes it difficult to measure μ spectrum accurately. With a 2.4 MW beam, the absorber thickness is too large to sample the lower energy muons. But these systems play an essential role in monitoring *flux stability*

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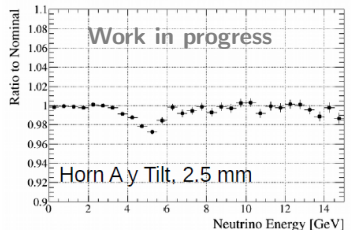
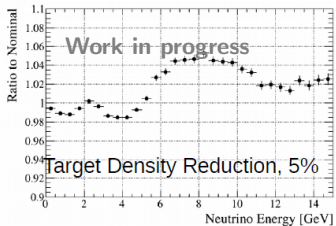
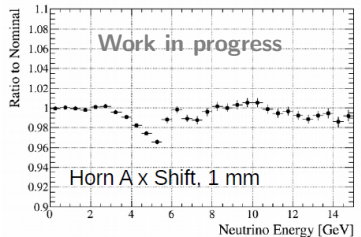
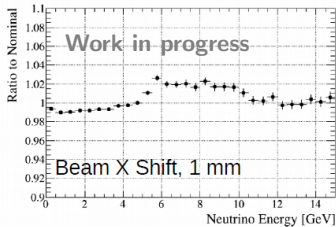
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ν Spectrum Changes



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The LBNF Beamline

Flux Modeling and Uncertainties

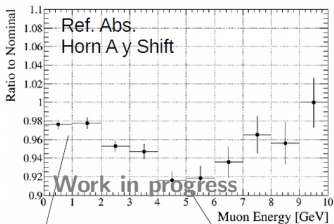
Future had. prod. measurements

Near Detector(s) Flux Measurements

Muon Monitors

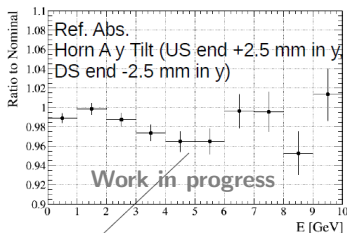
Summary

μ Spectrum Changes



Reduction in total flux

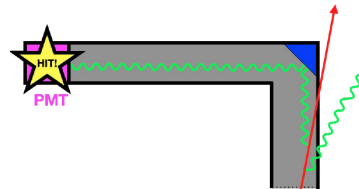
Shape changes at max near 5 GeV



Changes are v. small - need novel detector concepts

- **Array of ionization detectors:** Measures muon beam center and intensity. Spill by spill monitoring of beam stability. Both diamond and silicon under study
- **Threshold gas Cherenkov detector (R&D):** Uses signal intensity at different gas pressure and angles to extract rough muon spectrum.
- **Stopped muon counters (R&D):** separate stations with steel shielding in between could measure muon flux at several energies. Better measurement of beam flux spectrum and composition.

Gas Cherenkov counter concept:



Prototype in NuMI beamline:



Currently only ionization detectors included in the beam design.

Determination
of the LBNF
Neutrino Flux

Mary Bishai
(on behalf of
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Brookhaven
National Lab

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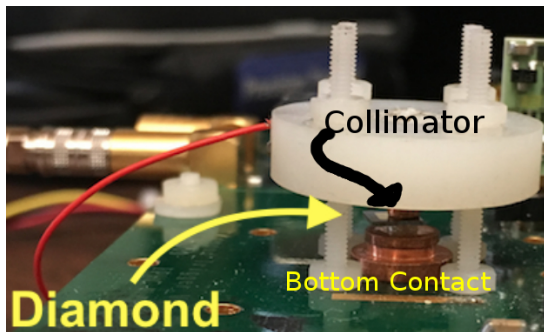
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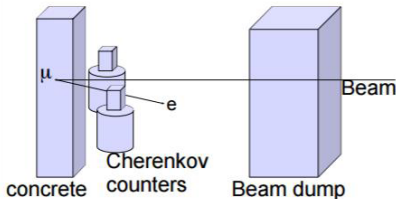
Use polycrystal chemical vapor deposition (pCVD) diamond - detects ionizing radiation when a large voltage potential (1V per μm of thickness) is applied across two sides of the diamond. Diamond is radiation hard.

pCVD detector prototype installed in NuMI during 2018 shutdown.

From K. Hiraide, *Muon monitor using the decay electrons*, NBI2003 Workshop



Strategy

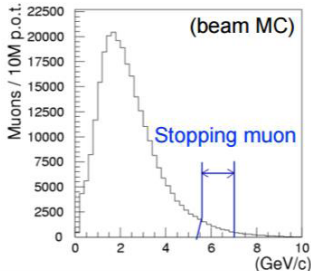


- Counting the decay electrons from muons stopping at the wall of μ -pit
- Measuring spatial and time distributions of events

- Energy loss of muons in the beam dump
- Range of electrons in the concrete



We can measure muons of
5.2~7.0 GeV/c
 by counting the decay electrons



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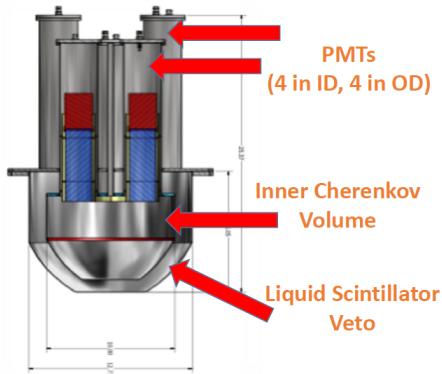
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Prototypes being commissioned with cosmics.



Summary and Conclusions

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Summary

- The next generation of long-baseline experiments requires determination of the flux at the **1-2% level**.
- LBNF is a **configurable wide-band beam** \Rightarrow requires beam flux/spectrum measurements over a large range of energies.
- Very high intensity beams (MW class) are needed \Rightarrow challenging near detector designs. Difficult to keep the same technology near and far. **For LBNF/DUNE the near detector concept is a combination of many different technologies**
- Focusing uncertainties dominate the residual uncertainties in the near to far extrapolation at LBNF/DUNE \Rightarrow determination of the hadron production from the target *is necessary but not sufficient* for a-priori calculations of the neutrino flux. **Do we need a spectrometer following the horns?**
- **Measurements of the muon flux after the absorber is difficult but necessary to monitor the tertiary beam stability. R&D is ongoing on new technologies.**