

The Short Baseline Near Detector

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Outline

- Short Baseline Physics
 - Detecting neutrinos with the SBND LArTPCs
 - Cross Section measurements in SBND
 - Current Status
 - Conclusions

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Sterile Neutrino Hints

 Recalculation of reactor neutrino fluxes and analysis of sources in gallium experiments (still) shows a disappearance effect.

<u>MiniBooNE</u>

Baseline 540 m E=[0 - 2] GeV L/E \approx 1 m/MeV







excess appearance of v_e and v_e : So called low energy excess. evidence for $v_{\mu} \rightarrow v_e$ oscillation at $\Delta m^2 \simeq 1 eV^2$?



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Global Fits

Dentler, Hernández-Cabezudo, Kopp, Machado, Maltoni, Martinez-Soler, Schwetz, arXiv:1803.10661.



Tension with experiments that observe no signal, especially recent measurements by **IceCube and MINOS+** leads to significant constraints on possible sterile neutrino parameters.





Full SBN running







The Short-Baseline Near Detector (SBND), will be located closest to the source of neutrinos.

It will characterize the beam before oscillations occur and address one of the dominant systematic uncertainties.

Planned start of operation 2020. 19/10/18 A. M. Szelc @ NuINT 2018, GSSI



SBND at a glance

112-ton (active volume) in two Liquid Argon Time Projection Chambers.

Four Anode Plane Assemblies and 2 Cathode Plane Assemblies.

4x4x5m Active Volume.

Cold readout electronics.

Membrane Cryostat.

Cosmic Ray Tagger









Booster Neutrino Beamline

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- A well known and understood system.
- Running stably for +10 years



LArTPC Operation





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Detecting neutrinos in a LArTPC

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- The golden channel in the SBN is ν_e appearance $(\nu_{\mu} \rightarrow \nu_e)$

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- Photon backgrounds make v_e appearance a particularly challenging game.
- The LArTPC and its bubble chamber-like data gives us strong background rejection tools.
- But it is also an excellent tool for precision crosssection measurements.





SBND cross-section physics

Charged Current	
$ u_{\mu}$ Inclusive	5,389,168
$ ightarrow 0\pi$	3,814,198
$\longrightarrow 0 ho$	27,269
$\longrightarrow 1 ho$	1,261,730
$\longrightarrow 2p$	1,075,803
$\longrightarrow \geq 3p$	1,449,394
$ ightarrow 1\pi^+ + X$	942,555
$ ightarrow 1\pi^- + X$	38,012
$ ightarrow 1\pi^0 + X$	406,555
$ ightarrow 2\pi + X$	145,336
$\rightarrow \geq 3\pi + X$	42,510
$ ightarrow K^+K^- + X$	521
$ ightarrow K^0 ar{K}^0 + X$	582
$ ightarrow \Sigma_c^{++} + X$	294
$ ightarrow \Sigma_c^+ + X$	98
$ ightarrow \Lambda_c^+ + X$	672
ν_e Inclusive	pprox 12,000
Neutral Current	
Inclusive	2,170,990
$ ightarrow$ 0 π	1,595,488
$ ightarrow 1\pi^{\pm} + X$	231,741
$ ightarrow \geq 2\pi^{\pm} + X$	343,760
$ ightarrow e(^-)$	374

SBND will see a huge event rate.

Enables precision measurements of neutrino cross-sections and nuclear effects.

Crucial for energy reconstruction in oscillation measurements.

A multitude of exclusive channels





SBND Cross-Sections

The University of Manchester





A. M. Szelc @ NuINT 2018, GSSI

SBND First planned Charged Current Constraint Charged Current



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SBND event rates for rare event searches

- Two proton events will no longer be "rare" (as seen in Raquel's talk).
- v. large sample of electron neutrinos.
- significant number of hyperons produced.
- Electron scattering measurements also possible.





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 $\rightarrow el$

19/10/18

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231,741

343,760 374



Beyond SM Searches

- There is a rising interest in potential detection of unconventional neutrinosector and dark-sector physics signals in large-volume neutrino experiments
- The proximity to the beam target, large detector mass and relative detection isotropy makes the LAr TPC SBN detectors well suited for beyond the standard model searches.

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- Sub-GeV dark matter (with proton beam dump)
- Hidden-sector particles
- Exotic signatures



Direct detection



Removing Cosmic Backgrounds with Machine Learning

- Liquid argon detectors provide image-like quality event displays.
- Can use tools for "image-recognition" to remove cosmic ray background events.
- Crucial tool in identifying neutrino interactions.
- Using Multiplane URes-Net





SBND Status: Building

© Steinkamp Photography 05.23.17



• Building is ready and first detector components are being installed.







SBND Status:

CRT Dnstream

- The Universit of Mancheste
- Cosmic Ray Tagger will ensure cosmic and dirst background suppression.
- First modules already at Fermilab and taking data.





SBND Status: APA, CPA & Fieldcage

Two Anode Plane Assemblies have been finalized and shipped to Fermilab.

- The CPA has arrived in Fermilab.
- **US APA** UK APA

CPA

 Field cage has been assembled and is ready to ship.

FieldCage

MANCHESTER 1824 SBND Status: Assembly & Construction



- Assembly Site in the DAB building is ready.
- First order is testing APA alignment with mock-frames.
- The whole fieldcage will be assembled at DAB and transported to the detector building once the cryostat is ready.



MANCHESTER 1824 Scintillation Light Detection in SBND

- Important R&D aspect
- Scintillation light applications:
 - trigger, t₀
 - background rejection
 - calorimetry, particle ID

• Mounted on anode planes:

- PhotoMultiplier Tubes
- Waveguide bars
- ARAPUCA light traps
- Mounted on cathode planes:
 - WLS covered reflector foils





Light Collectors



Primary System: 8" PMTs read out at 2ns sampling.

Sandblasted and coated with TPB to enable VUV light detection.

Secondary System(s): - Dip-coated light guide bars

- ARAPUCA light traps



A. A Machado and E. Segreto, JINST 11 2016





Boosting Light Collection

- Adding wavelength-shifting surface at the cathode recovers a large fraction of light that would normally be lost.
- The SBND LDS enables new applications of argon scintillation light calorimetry, timing, drift position reconstruction.
- Enhancement expected especially at low energies.







The SBND Collaboration

Updated September 2018

Including both scientific and technical personnel

Spokespeople Argonne National Lab: Z. Djurcic, R. Dharmapalan, G. Drake, M. Goodman, S. Magill University of Bern: Y. Chen, A. Ereditato, R. Hänni, I. Kreslo, D. Lorca, M. Lüthi, F. Piastra, J. Sinclair, M. Weber Brookhaven National Lab: M. Bass, M. Bishai, H. Chen, J. Farrell, J. Fried, S. Gao, J. Joshi, D. Lissauer, X. Qian, V. Radeka, E. Raguzin, C. Thorn, A. Timilsina, E. Worcester, M. Worcester, B. Yu, J. Zhang University of Campinas - UNICAMP: C. Escobar, E. Kemp, M. Guzzo, P. Holanda, M. Nunes, L. Santos, E. Segreto CERN: S. Bertolucci, J. Bremer, U. Kose, D. Mladenov, M. Nessi, F. Noto University of Chicago: A. Mastbaum, K. Miller, R. Northrop, G. Putnam, D.W. Schmitz Colorado State University: R. LaZur, M. Mooney, I. Terrazas Columbia University: L. Camilleri, C. Chi, D. Cianci, J. Crespo, V. Genty, G. Karagiorgi, M. Ross-Lonergan, M.H. Shaevitz, B. Sippach, K. Sutton Federal University of ABC - UFABC: A. Machado, C. Moura, L. Paulucci, L. Quintino Federal University of Alfenas – UFAL: M. dos Santos, G. Valdiviesso Federal University of Rio de Janeiro: C. Bonifazi Federal University of Sao Carlos – UFSCAR: F. Marinho Fermilab: W. Badgett, L. Bagby, B. Baller, R. Castillo Fernandez, F. Cavanna, S. Dixon, J. Estrada, M. Geynisman, H. Greenlee, C. James, W. Ketchum, M.J. Kim, D. Montanari, B. Norris, O. Palamara*, Z. Pavlovic, R. Rameika, B. Rebel, A. Schukraft, S. Shetty, M. Stancari, A. Stefanik, T. Strauss, D. Torretta, M. Toups, P. Wilson, G.P. Zeller, J. Zennamo Harvard University: C. Adams, R. Guenette Illinois Institute of Technology: I. Lepetic, B. Littlejohn Indiana University: S. Mufson Kansas State University: G. Horton-Smith Lancaster University: A. Blake, D. Brailsford, I. Mercer, J. Nowak, P.N. Ratoff University of Liverpool: C. Andreopoulos, S. Dennis, J. Henzerling, R. Jones, K. Mavrokoridis, N. McCauley, D. Payne, A. Roberts, M. Roda, P. Sutcliffe, J. Tena-Vidal, C. Touramanis Los Alamos National Lab: J. Boissevain, G. Garvey, E. Huang, W.C. Louis, K. Rielage, T. Thornton, R.G. Van de Water University of Manchester: V. Basque, A. Bitadze, J. Evans, J. Freestone, A. Furmanski, D. Garcia Gamez, O. Goodwin, P. Guzowski, C. Hill, K. Mistry, J. Pater, S. Söldner-Rembold, A.M. Szelc University of Michigan: C. Barnes, R. Fitzpatrick, J. Mousseau, B. Roe, J. Spitz MIT: J.M. Conrad, J. Moon New Mexico State University: R. Cooper Pacific Northwest National Lab: E. Church University of Pennsylvania: N. Barros, J. Klein, D. Rivera, R. Van Berg University of Puerto Rico: K. Matias, H. Mendez, S. Santana University of Sheffield: D. Barker, T. Brooks, T. Gamble, V.A. Kudryavtsev, M. Malek, N. McConkey, J. Mercer, F. Mouton, N. Spooner, M. Wright University of Sussex: C. Griffith, I. de Icaza Astiz Syracuse University: A. Bhat, P. Hamilton, G. Pulliam, O. Rodrigues, M. Soderberg University of Tennessee, Knoxville: S. Gollapinni, A. Mogan, W. Tang, G. Yarbrough University of Texas, Arlington: J. Asaadi, A. Chatterjee, A. Falcone, Z. Williams, J. Yu Tufts University: T. Wongjirad University College London: M. Cascella, A. Holin, R. Nichol, D. Waters Virginia Tech: C. Mariani

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The SBND Collaboration



196 Total Collaborators

166 Scientific Collaborators (faculty/scientists, postdocs, students)

36 Institutions

23 US Institutions

5 DOE national laboratories 18 US universities University of Puerto Rico

13 International Institutions

CERN

6 UK universities

- 1 Swiss university
- 5 Brazilian universities





Summary

- SBND construction is well on its way and the detector will come online in early 2020.
- As the SBN near detector it plays a key role in enabling the full scale SBN oscillation analysis.
- But more importantly it will provide a wealth of cross-section data on argon.
- Stay tuned!



Thank You for your Attention

Physics reach

Constraints on the flux and cross-sections from the near detector lead to a powerful combined exclusion region. LSND parameter space excluded at 5σ .

In addition, SBN can also perform v_{μ} disappearance searches. Would confirm an oscillation interpretation of any observed v_{e} appearance signal.





Fit from S. Gariazzo et al., arXiv:1703.00860

Electron- γ separation in LAr





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