

Fermi Large Area Telescope:

Long Term Public Analysis

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on behalf of the Fermi LAT collaboration



The Fermi mission

Two instruments :

The Large Area Telescope (LAT) : High Energy Gamma-ray Astrophysics
The Gamma-ray Burst Monitor (GBM) : focuses on GRBs and transients

Launched 11 June 2008 – LAT activated 25 June

Nominal Science observations started on 4 August 2008

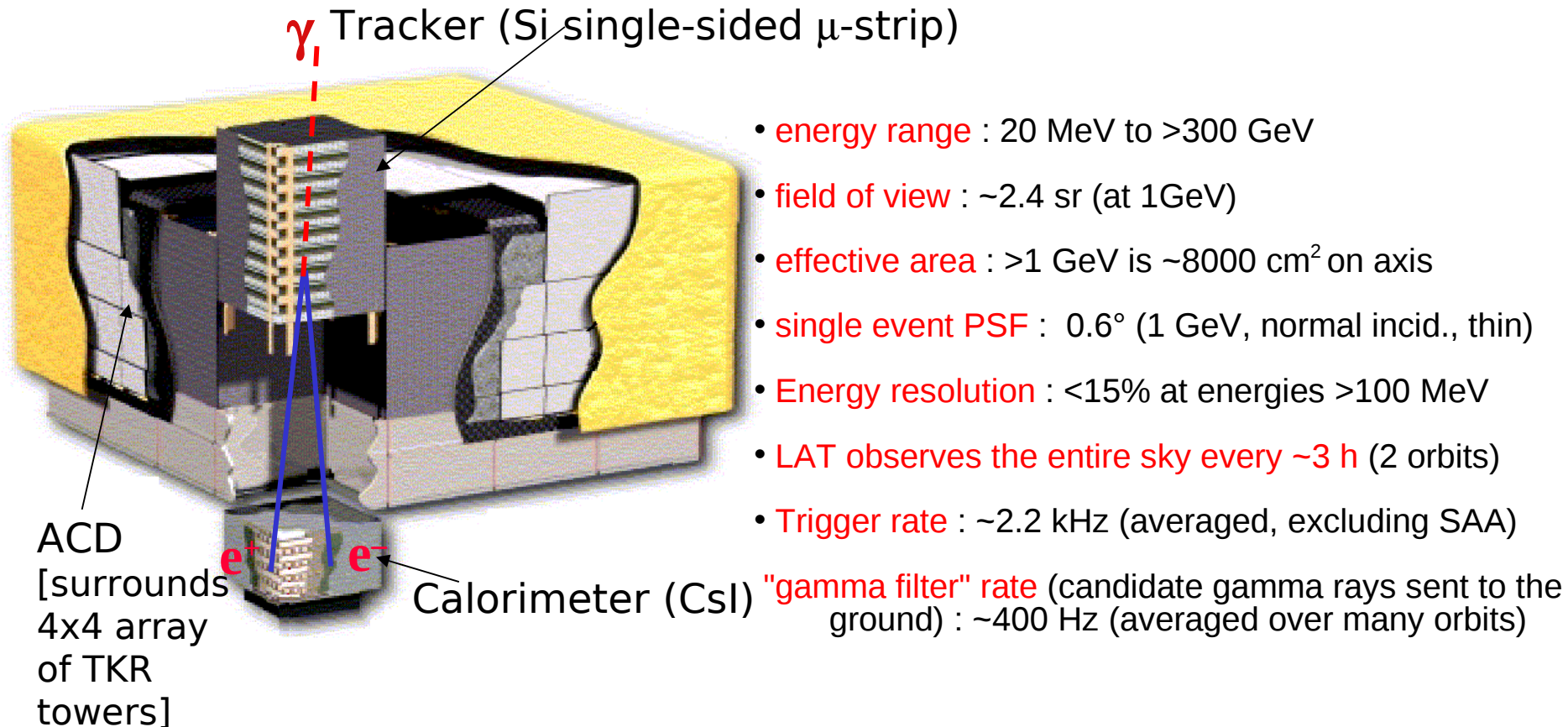
THIS TALK FOCUSES ON THE LAT PIPELINES



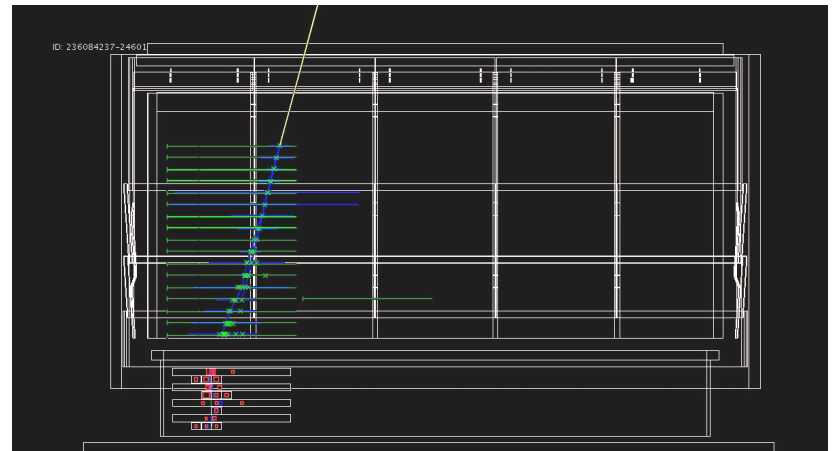
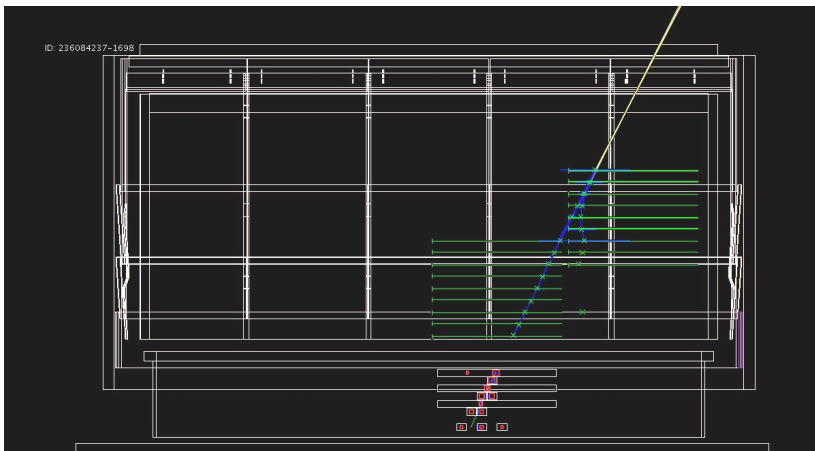
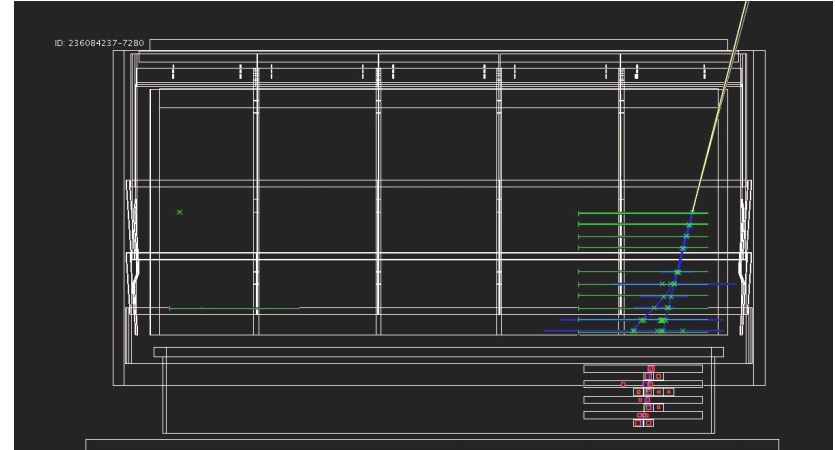
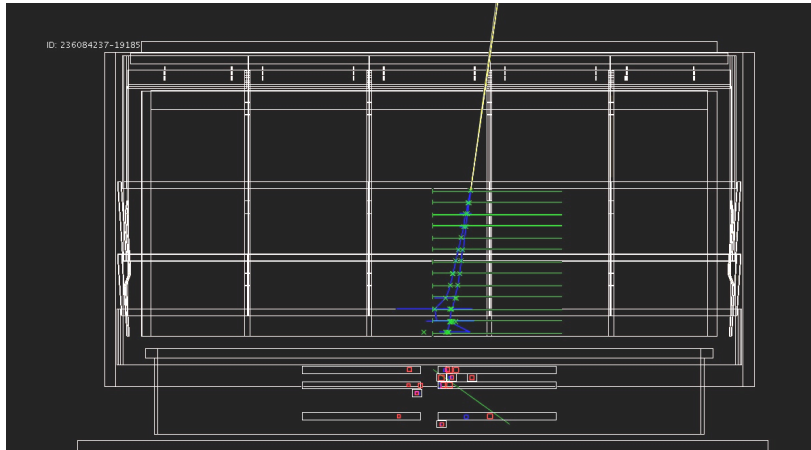
The Large Area Telescope (LAT)

- **A HEP instrument for Astrophysics!**

- Built and operated by a mixed HEP/Astro collaboration
- Managed at SLAC, Stanford University.



In Orbit: Single Events in the LAT



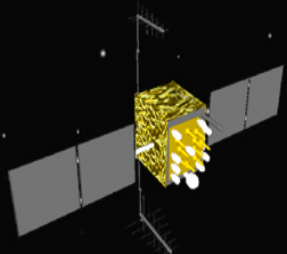
The green crosses show the detected positions of the charged particles, the blue lines show the reconstructed track trajectories, and the yellow lines shows the candidate gamma-ray estimated direction. The red crosses show the detected energy depositions in the calorimeter.

0.25 CPU sec/event to reconstruct: downlink 500 Hz
Each event independent of others

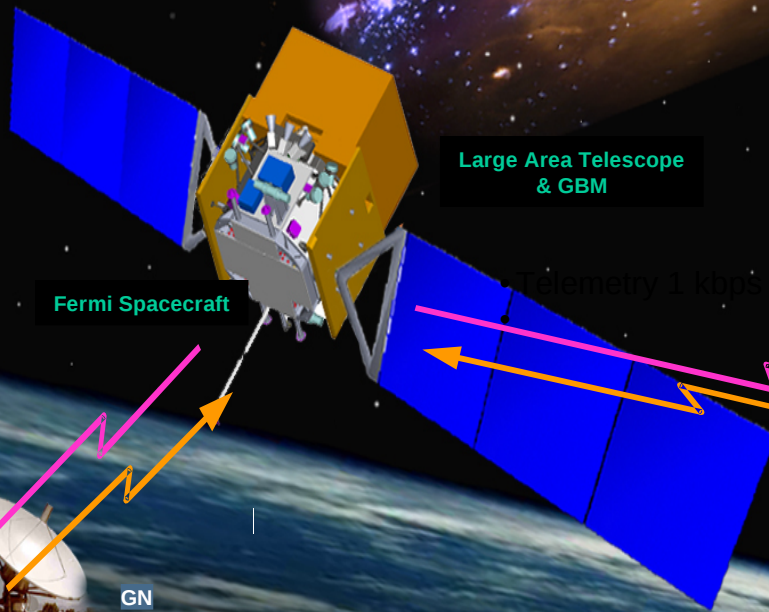
Science Accomplishments in a Nutshell

- **Mapping and measuring the entire GeV sky at a unprecedented angular and energy resolution and statistical accuracy**
 - Detected >40 pulsars, including all the EGRET ones and several radio quiet ones
- Detected over 1000 sources
 - First year catalogue in preparation
 - Bright source list of ~100 sources published 6 months ago
- Mapped the diffuse Galactic emission and measured its spectrum
- Measured the cosmic-ray electron spectrum
- Discovered flares from several AGN reported in ATels
- Detected the binary LSI+61 303, the LMC, and >10 GRBs already
- Detected the moon and the quiet sun (and earth)
- Measured the light-curve and spectrum of the Vela pulsar
- Detected two Galactic plane transients

Fermi MISSION ELEMENTS



GPS



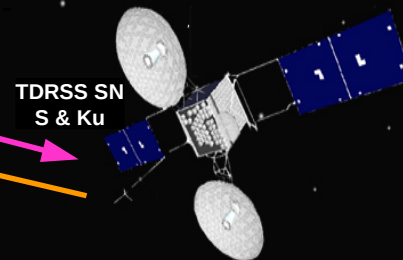
Fermi Spacecraft

Large Area Telescope & GBM

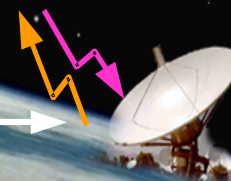
DELTA 7920H



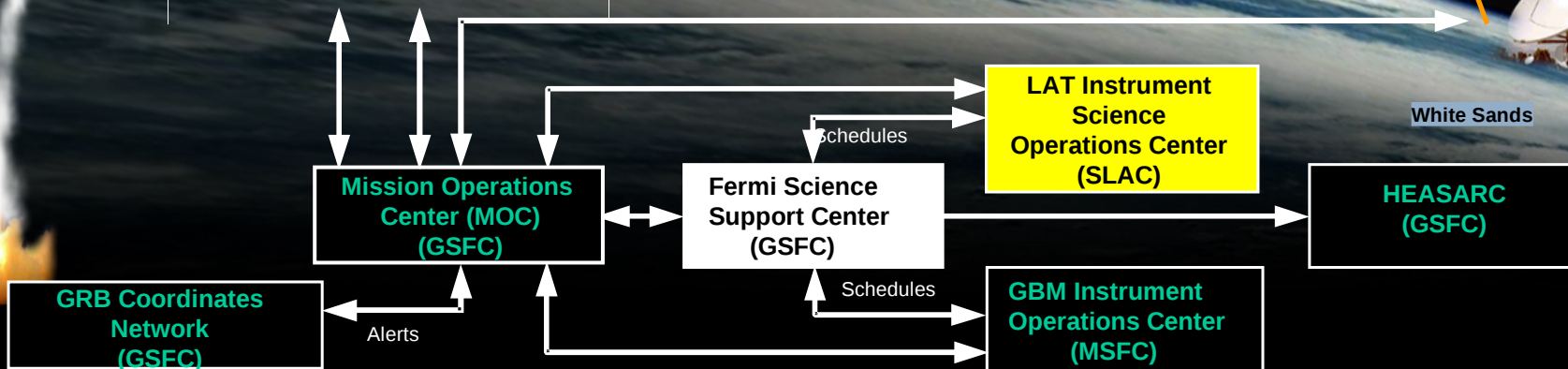
GN



TDRSS SN S & Ku



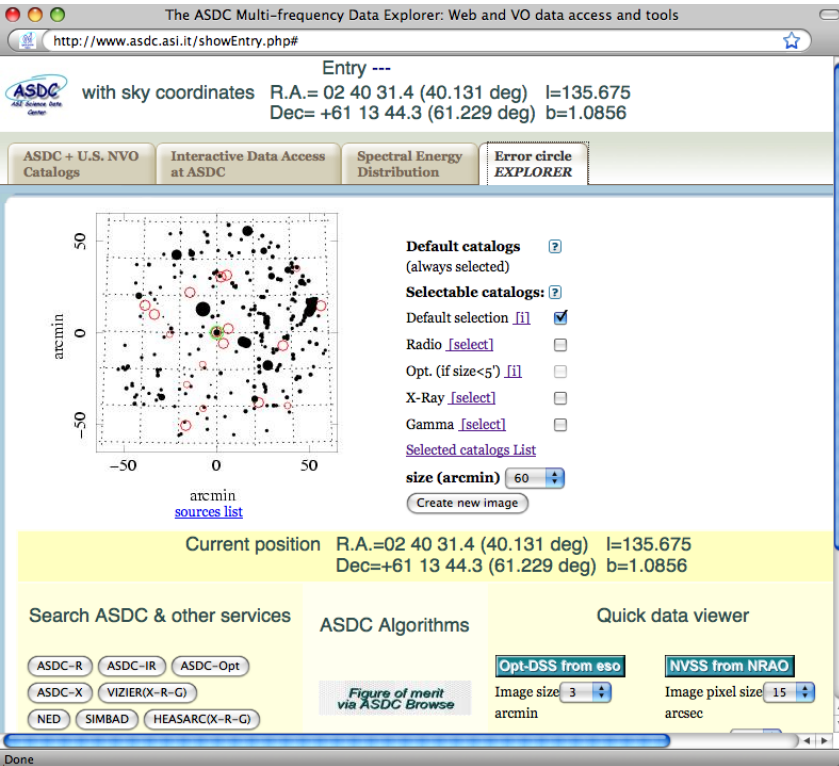
White Sands



Data Processing Flow

- Downlink from Goddard Space Flight Center (FASTCopy) ~8/day
 - 15 GB total daily
 - Half-pipe
 - Automatic response to downlink
 - Decode & repackage incoming data
 - Trigger Level 1 Processing
 - Level 1 Processing
 - Full event reconstruction: factor ~x50 expansion on raw data! 750 GB/day
 - Monitoring plots for Instrument Science Operations Center (> several 100s)
 - Transfer science summary files to Goddard Science Support Center - 200 MB/day
 - Trigger ASP
 - ASP (Automated Science Processing)
 - GRB and Flare detection
 - Spectral analysis
 - RSP (Routine Science Processing)
 - Automated Science Group processing
 - Final science analyses
- + annual bulk reprocessing
+ simulations

Automated Science Processing



The ASDC Multi-frequency Data Explorer: Web and VO data access and tools
http://www.asdc.asi.it/showEntry.php#

Entry ---
with sky coordinates R.A.= 02 40 31.4 (40.131 deg) l=135.675
Dec.= +61 13 44.3 (61.229 deg) b=1.0856

ASDC + U.S. NVO Catalogs Interactive Data Access at ASDC Spectral Energy Distribution Error circle EXPLORER

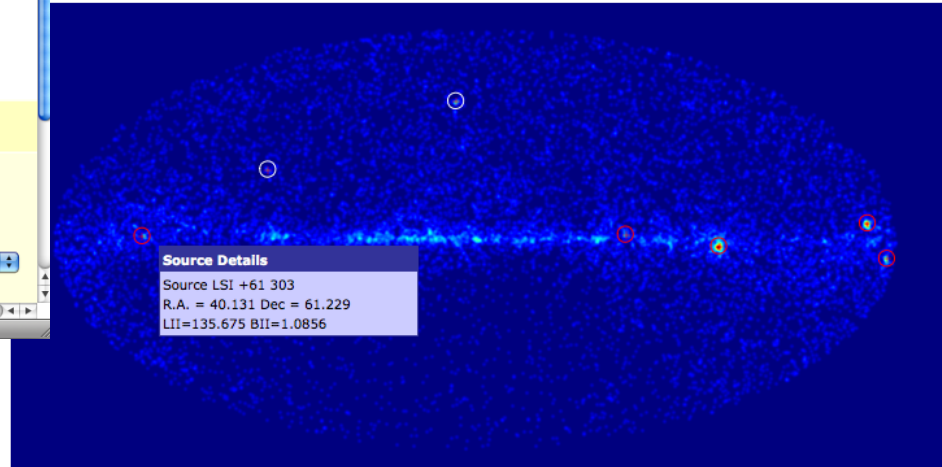
Default catalogs (always selected)
Selectable catalogs:
Default selection [f1](#)
Radio [\[select\]](#)
Opt. (if size<5') [\[f1\]](#)
X-Ray [\[select\]](#)
Gamma [\[select\]](#)
[Selected catalogs List](#)
size (arcmin)

Current position R.A.=02 40 31.4 (40.131 deg) l=135.675
Dec.=+61 13 44.3 (61.229 deg) b=1.0856

Search ASDC & other services
ASDC Algorithms Quick data viewer
ASDC-R ASDC-IR ASDC-Opt
ASDC-X VIZIER(X-R-G)
NED SIMBAD HEASARC(X-R-G)
Opt-DSS from eso NVSS from NRAO
Figure of merit via ASDC Browse
Image size arcmin
Image pixel size arcsec

Source Details
Source LSI +61 303
R.A. = 40.131 Dec = 61.229
l=l=135.675 Bll=1.0856

- Run every 6 hours and daily
- Track 23 public sources: get from FSSC
- Search for Flares
- Refine onboard GRB analysis
- Blind search for GRBs
- Connects to ASI Science Data Center for MWL correlation



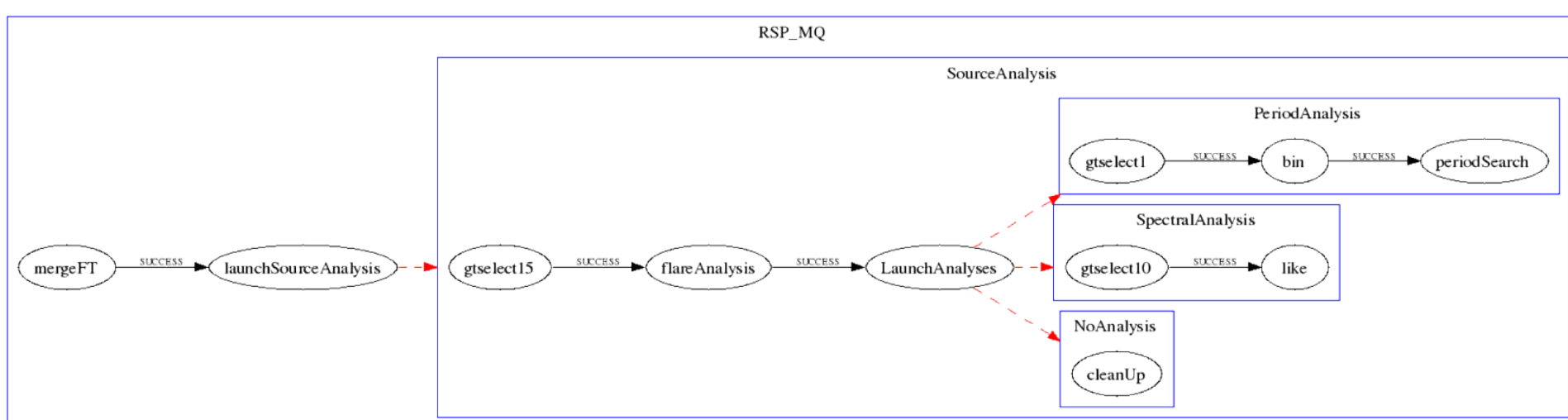
Start Time: 30/Oct/2008 18:00:00 - Stop Time: 31/Oct/2008 00:00:00

Detections

Source Name	ASDC Tools	RA	DEC	l	b	Flux	Flare Test Type	Flare Type Value	Tentative Associations (1 degree)
ASPJ124300-621500	ASDC Tools	190.75	-62.25	301.9498	0.6057	1.560E-6 +/- 3.970E-7	Chi2	14	J1245-6238
ASPJ150146+112224	ASDC Tools	225.4409	11.3734	12.072	55.6009	1.310E-6 +/- 2.870E-7	Chi2	0	J150339+104836

RSP – ongoing Growth Industry

- Leverage Pipeline to support Science Groups' needs
 - Automated spectral fitting and light curve creation
 - Pulsar implemented already
 - Binaries, GRB, AGN in progress
- Should be flexible to span times between ASP and catalogue
 - Run weekly in most cases
 - Also handle reprocessed data



Principal Computing Resources

- **SLAC compute farm (Fermi allocation)**
 - 1200 CPUs in batch farm running LSF (+ peak loads of >2000)
 - 600TB disk = 100 TB NFS + 500 TB xroot – 32 TB Raid 10 Sun thumpers (to be replaced by thors)
 - 350 TB tapes in silo (HPSS)
 - Making transition to higher density tapes
 - **Storage model strategy is to keep latest versions of data on disk; all else on tape.**
- **Lyon compute farm (CCIN2P3, France)**
 - 600 CPUs in batch farm running BQS
 - few TB disk allocated for all Fermi uses (transfer generated files to SLAC)
 - seamlessly used by pipeline from SLAC
 - **Currently used only for Monte Carlo**
- **Ongoing effort to also distribute simulation tasks on the grid (EGEE) : Existing VO with 10 active nodes which ran Fermi simulations successfully**

Data Storage and Catalog

Show: MC beam test obsolete

Folders

- ASP
- Data
 - Flight
 - LEOScience
 - Level1
 - LCI
 - LPA
 - ACDPEDSANALYZER
 - ACDPLOTS
 - CAL
 - CALGAINSANALYZER
 - CALHIST
 - CALHISTALARM
 - CALPEDSANALYZER
 - CALTREND
 - COMPAREDFM
 - DIGI
 - DIGIHIST
 - DIGIHISTALARM
 - DIGITREND
 - DIGITRENDALARM
 - FASTMONERROR
 - FASTMONHIST

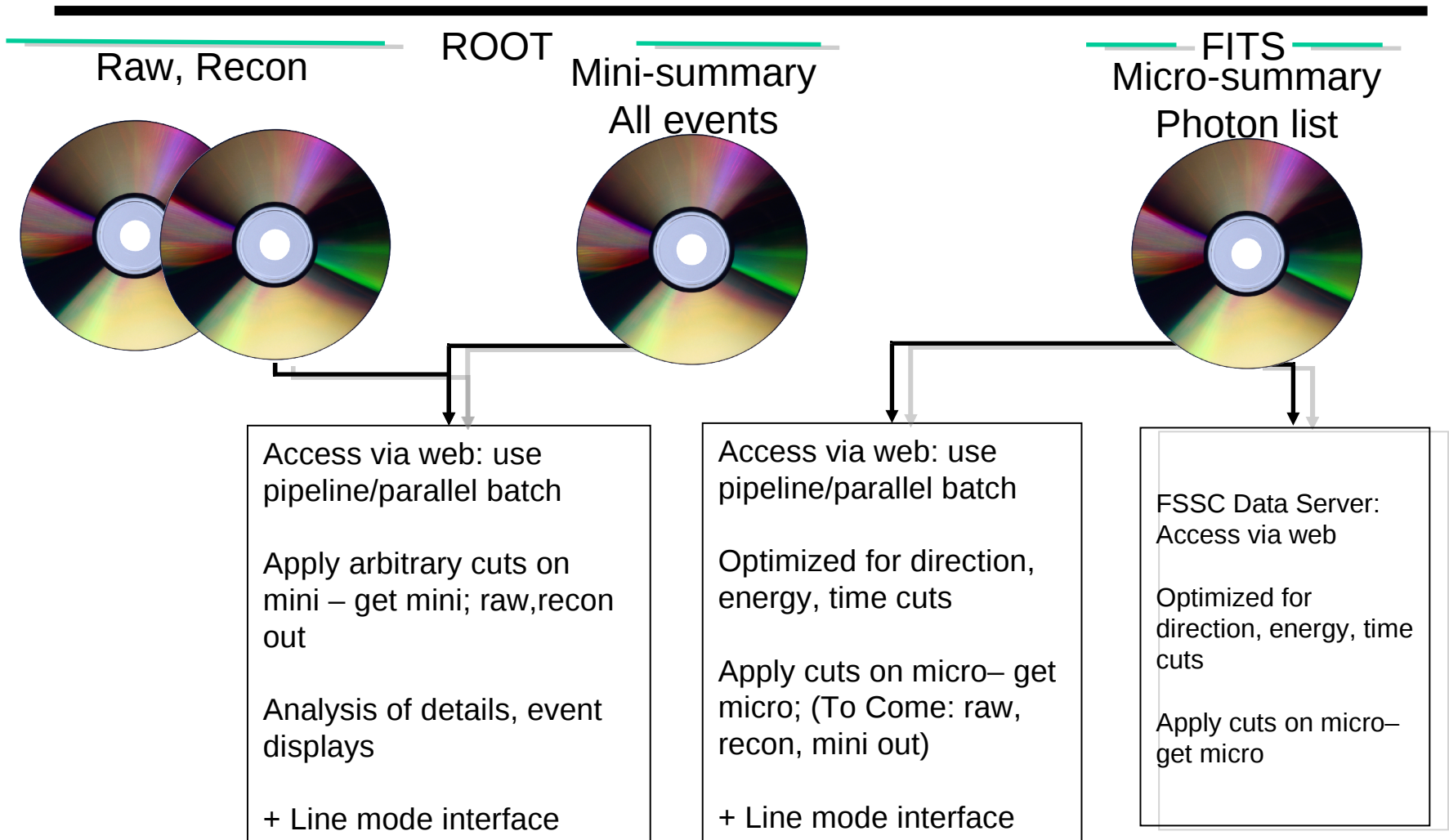
Folder /Data/Flight/Level1/LPA

Output from Level 1 processing of on-orbit data [Edit description](#)

Name	Type	Files	Events	Size	Created (UTC)	Links
RECON	Group	1931	3,803,647,787	52.7 TB	25-Jun-2008 16:43:00	Files
CAL	Group	1954	3,790,237,100	13.3 TB	25-Jun-2008 16:35:11	Files
SVAC	Group	1953	3,848,209,867	9.0 TB	25-Jun-2008 16:29:03	Files
DIGI	Group	1954	3,855,037,479	8.0 TB	25-Jun-2008 15:22:31	Files
FASTMONTUPLE	Group	1954	0	3.7 TB	25-Jun-2008 15:34:54	Files
MERIT	Group	1954	3,852,358,312	2.9 TB	25-Jun-2008 16:25:29	Files
GCR	Group	1954	3,852,208,291	92.6 GB	25-Jun-2008 16:32:57	Files
FASTMONTREND	Group	1954	0	58.9 GB	25-Jun-2008 15:57:51	Files
LS1	Group	1954	69,293,586	50.7 GB	25-Jun-2008 16:29:02	Files
DIGITREND	Group	1954	0	45.8 GB	25-Jun-2008 15:25:58	Files
MAGIC7HP	Group	1732	0	37.2 GB	08-Jul-2008 18:20:31	Files
CALHIST	Group	1954	0	26.4 GB	25-Jun-2008 15:32:55	Files
TKRANALYSIS	Group	1953	0	23.8 GB	25-Jun-2008 16:40:49	Files
RECONTREND	Group	1954	0	23.7 GB	25-Jun-2008 16:39:44	Files
RECONHIST	Group	1953	0	16.4 GB	25-Jun-2008 16:42:37	Files
LS3	Group	1954	0	14.7 GB	25-Jun-2008 16:29:02	Files
MAGIC7	Group	250	0	7.4 GB	25-Jun-2008 15:13:53	Files

- Data reside in xroot (some NFS, and some, temporarily, in AFS)
- Catalogue gives file location and user-supplied metadata

Accessing the Data



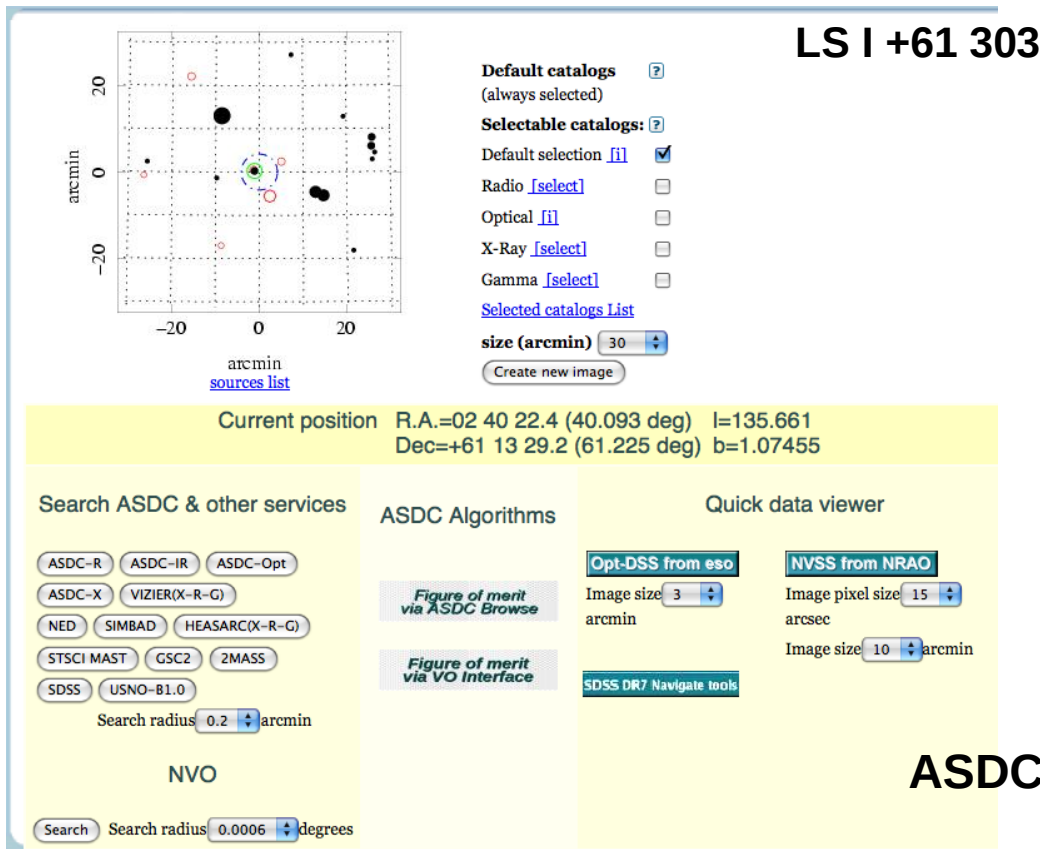
Pointing/Livetime history vital, distributed as well

NASA Astro Culture on Facilities

- telescopes are “facilities” :
 - driven by user groups with scheduled observations
 - not the property of the instrument teams
 - community organises multi-wavelength campaigns

TeV telescopes are the exception so far: more like HEP experiments

LS I +61 303



The screenshot shows the Fermi LAT web interface for the source LS I +61 303. It features a central plot of sources in Galactic coordinates (l, b) with axes ranging from -20 to 20 arcmin. To the right of the plot are controls for selecting catalogs (Default, Radio, Optical, X-Ray, Gamma) and a size filter set to 30 arcmin. Below the plot, the current position is given as R.A. = 02 40 22.4 (40.093 deg) and Dec = +61 13 29.2 (61.225 deg). The interface also includes a search bar for ASDC and other services, with buttons for various catalogs like ASDC-R, VIZIER, NED, STSCI, SDSS, etc. There are also sections for 'ASDC Algorithms' and 'Quick data viewer' with options for image size and pixel size.

The collaboration is expected to deliver ~ the same dataset that it would use in private, and an up-to-date version of the Science Analysis Software Toolkit (based on on planned features)

ASDC

The FSSC : portal for the user community

•Manages High-Level Science data only



GODDARD
SPACE FLIGHT CENTER

+ NASA Home page
+ GSFC Home page
+ Fermi Home page

SEARCH Fermi:

+ GO

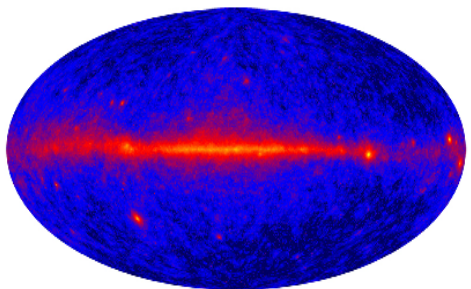


Fermi

Science Support Center

HOME
RESOURCES
PROPOSALS
DATA
HEASARC
HELP
SITE MAP

The Fermi Science Support Center (FSSC) runs the guest investigator program, creates and maintains the mission time line, provides analysis tools for the scientific community and archive and serves the Fermi data. This web site is the portal to Fermi for all guest investigators.



This all-skyview from Fermi reveals bright emission in the plane of the Milky Way (center), bright pulsars and super-massive black holes.
Credit: NASA/DOE/International LAT Team

Look into the "Resources" section for finding schedules, publications, useful links etc. The "Proposals" section is where you will be able to find the relevant information and tools to prepare and submit proposals for guest investigator projects. At "Data" you will be able to access the Fermi databases and find the software to analyse them. Address all questions and requests to the helpdesk in "Help".

- ### Quicklist
- 2009 Fermi Symposium
 - 11th COSPAR Capacity-Building Workshop
 - GLAST Fellowship Program
 - Fermi Guest Investigator Program
 - Fermi Sky Blog

News


December 3, 2009
Data Analysis Workshop Series
 The FSSC will be hosting a series of hands-on data analysis workshops for anyone interested in getting started with Fermi data analysis. A discussion of the Fermi GI program will follow the session. The current workshop schedule is:

- Stanford University (Dec 8)
- Fermilab (Dec 17)
- AAS Meeting in DC (Jan 7)
- Boston University (Jan 11)

Refer to the appropriate workshop web site for more details.

November 2, 2009
NASA's Fermi Telescope Detects Gamma-Ray From "Star Factories" in Other Galaxies

Nearby galaxies undergo a furious pace of star formation and so emit lots of gamma rays, say astronomers using NASA's Fermi Gamma-ray Space Telescope. Two so-called "starburst" galaxies, plus a satellite of our own Milky Way galaxy, represent a new category of gamma-ray-emitting objects detected both by Fermi and ground-based observatories.

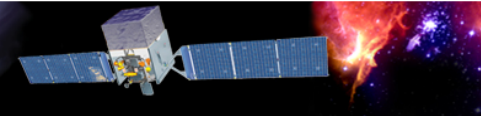


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Data Policy

Data Access

Data Analysis

Newsletter

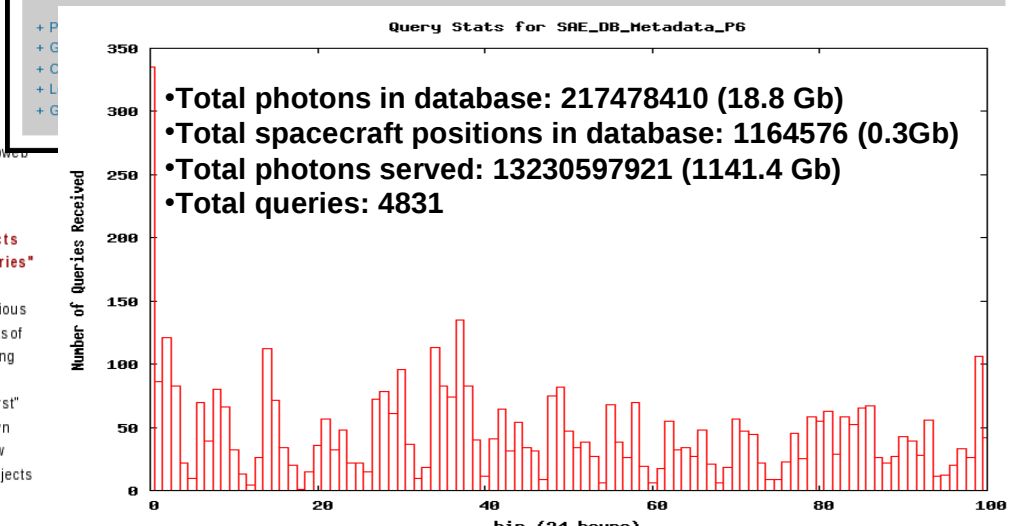
FAQ

Fermi Data

This is the portal to the Fermi data and the software to analyze them. Before the data or software are released, they are described here.

- [Data Access](#) - the Fermi science data
- [Data Analysis](#) - the software to analyze Fermi data
- [Data Policy](#) - a summary of the policies governing the release of Fermi science data

As a resource for cycle-2 proposers, the Fermi LAT consortium, through an agreement with NASA, has provided a [LAT Bright Source List](#)



NASA Archives

NASA's HEASARC: Observatories

ACTIVE MISSIONS	PAST MISSIONS	UPCOMING MISSIONS	COMPARISON OF MISSION CAPABILITIES
-----------------	---------------	-------------------	------------------------------------

ACTIVE MISSIONS	
AGILE	
Chandra	
Fermi (formerly GLAST)	
INTEGRAL	
Ross XTE	
Suzaku	
Swift	
XMM-Newton	
PAST MISSIONS	
ANS	ARIEL V
ASCA	BBXRT
BeppoSAX	CGRO
COPERNICUS	COS-B
DXS	EINSTEIN
EUVE	EXOSAT
GINGA	GRANAT
HAKUCHO	HEAO-1
HEAO-3	HETE-2
OSO-7	OSO-8
ROSAT	SAS-2
SAS-3	TENMA
UHURU	VELA 5B

High Energy Astrophysics Observatories

These Web pages describe all the high energy astronomy observatories, or "missions," ever launched. The links to the left are dedicated to the most notable past and present X-ray and gamma-ray astronomy missions. The pages include

- a mission overview,
- technical information on the instrumentation,
- a bibliography and
- a gallery of images.


They also contain a description of the available data if they are present in the HEASARC archive. Information on other satellites carrying either an X-ray or a gamma-ray telescope is available on the [All Missions](#) pages.

- | | |
|--|--|
| <p>Upcoming Missions</p> <ul style="list-style-type: none"> • ASTROSAT • MAXI | <p>Other Resources</p> <ul style="list-style-type: none"> • All Missions (info & images) • All Missions by Time • All Missions by Energy • Comparison of Mission Capabilities • Future Missions • Images, Spectra, and Light Curves |
|--|--|

<http://heasarc.gsfc.nasa.gov/>

Latest News

- [GLAST](#)

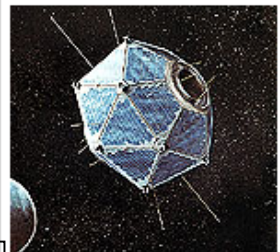


[launched June 11, 2008](#)

- [GLAST renamed to Fermi Gamma-ray Space Telescope](#) (Aug 26, 2008)

[More News](#)

Spacecraft Image of the Week



Vela 5B
DoD/AEC mission
May 1969 - June 1979


High Level Analysis

- **Output of a telescope (for “event” data) is:**
 - Location on sky, time, energy, quality
 - Very simple output tuple!
 - Implemented in Root and FITS
 - Where instrument and celestial analysis overlap
 - Public data makes sense in astrophysics
 - NEED data from multiple missions to understand celestial sources
- **NASA mandates that all its space missions use **FITS** as a data format**
 - In use for 25+ years
 - Format fully documented and files self documenting
 - File headers are an integral part of the format
 - Interface library supplied for popular languages
 - And that the data be made public
 - LAT negotiated one year hiatus on doing this. All existing & ongoing data went public on August 25.
 - Funds a Science Support Center to interface to the public
 - Instrument teams not asked to do this
 - 10+ FTEs for Fermi (LAT+GBM)
 - Charter is to support the Fermi data “forever”

More NASA Mandates

- **NASA-funded missions required to create a Project Data Management Plan prior to launch to cover the life of the mission**
 - **Specify data format details**
 - **Down to tuple column names**
 - **Under configuration control**
 - **Responsibilities across the Mission ground elements**
- **For high energy astrophysics, NASA further mandates their OGIP standards**
 - **FTOOLS are a notable example**
 - **Utilities do almost any manipulation you can think of**
 - **Adhere to the IRAF parameter file interface**
- **Agreement with NASA that a specified set of high level analysis tools would be jointly developed by the instrument teams and the FSSC to be used by both the team and the wider scientific community.**

Is FITS All Rosy?

- Nothing is.... 
- No structures/objects in the files (think ROOT....)
 - Images or tables (BIN ok, and good compression also)
- The analysis model that goes with it is atomic operations with files used as the messengers
 - Linked together with scripts (called pipelines)
 - Designed for interactive use
 - Not made for (parallel) batch processing
 - Very easy to step on the parameters files
- OTOH, FITS headers are nice
 - Intrinsic to the files : all tools respect the header
 - Carries processing history etc etc
 - Killer App? : very mature astrophysical libraries available
 - crucial WCS conversion/display capabilities

Remaining Possible Issues

- **Long-term support of code and systems**
- **Dealing with up to 250 TB data per year**
 - Large file servers mean big exposure to a failure
 - Data spread across many servers
 - Tape backup?
 - Multiple versions – annual reprocessings?
- **Scaling of databases**
 - Thinning/truncating trending data
 - “hot” expansion of db disk space
- **Maximising uptime**
 - Goal is to not miss flares on 12 hour time scales
 - For 10 years!

Summary

- **LAT is an amalgam of HEP (Root) and Astro (FITS)**
- **NASA worries about public data and long term analysis capability**
 - **We are riding their coat-tails**
 - **We have no plan for the Root data and assume at end of mission that it more or less dies**
 - **A price is paid in terms of functionality**
- **Our data have gone public in August 2008**
 - **2009 Fermi Symposium showed that the community is already active in looking at the data**
 - **For the collaboration, strong but healthy tension between scientific return in a competitive world, and bringing further improvements to the data (reconstruction, calibration, etc...)**