

# Probing Baryogenesis with Displaced Vertices

- from the LHC to CLIC

Yanou Cui

University of California-Riverside

- Phys.Rev.D, 87,11603, YC and Raman Sundrum
- JHEP 1312 (2013) 067, YC
- JHEP 1502 (2015) 049, YC and Brian Shuve
- Ongoing cooperation with ATLAS LLP working group
- Phys.Rev. D94 (2016) YC and Takemichi Okui, Arash Yunesi
- Work in progress, YC and Aniket Joglekar, Brian Shuve, Yuhsin Tsai

*BSM Searches@CLIC meeting,  
Feb 21, 2018*

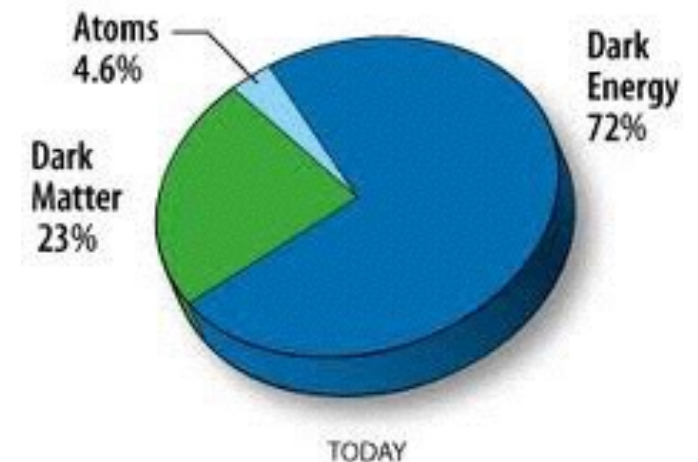
# Cosmological Motivation for LLP Searches

## - Baryogenesis from Metastable Weak-scale New Particle

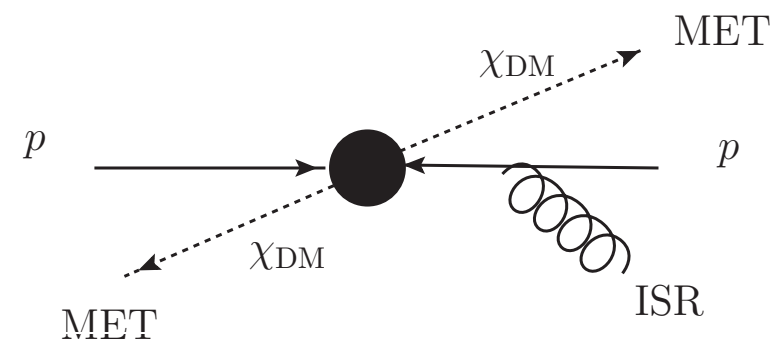


Could collider experiments shed light on prominent puzzles in modern cosmology?

$$\Omega_{\text{DM}} \approx 23\%, \Omega_{\text{B}} \approx 5\%, \Omega_{\text{B}} \sim \Omega_{\text{DM}}$$



- Familiar/well-studied case: WIMP dark matter ( $\Omega_{\text{DM}}$ )
  - Stable, weak-scale mass, can be produced within  $E_{\text{LHC}} = 14 \text{ TeV}$
  - Pair produced ( $Z_2$ ),
  - Invisible, MET + X



# Cosmological Motivation for LLP Searches

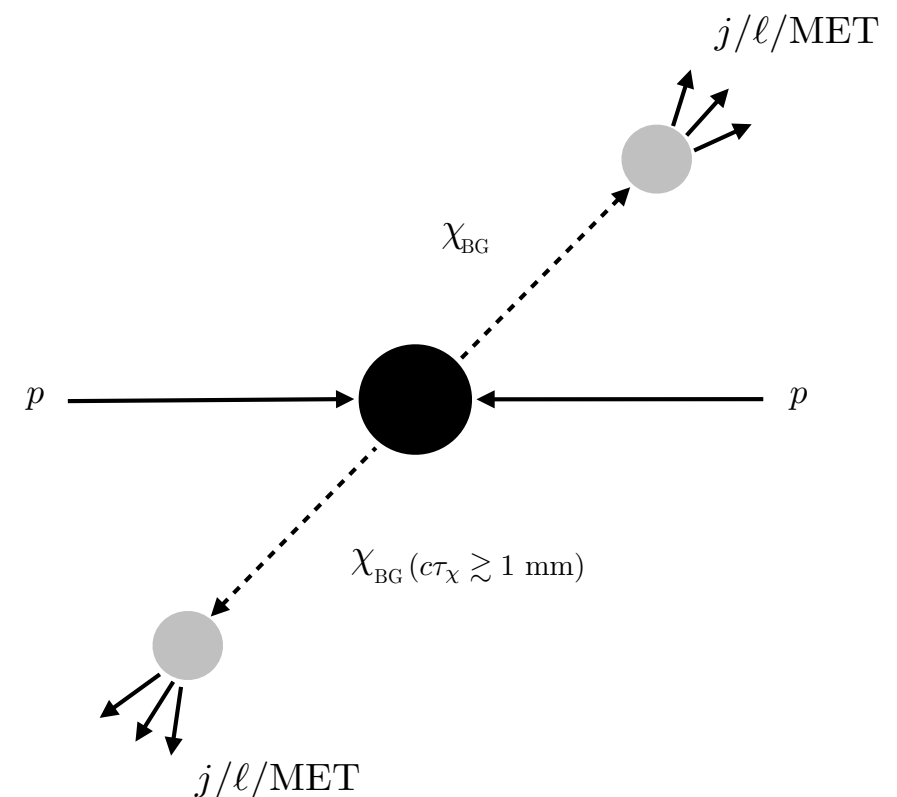
## - Baryogenesis from Metastable Weak-scale New Particle

- **New opportunity: baryogenesis**  
(address  $\Omega_B$ , possibly  $+\Omega_B \sim \Omega_{DM}$ )

- New metastable particle (baryon parent),  
weak scale mass

- Pair produced (approx.  $Z_2$ )

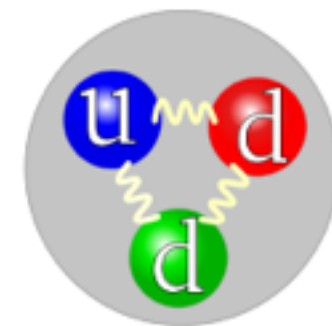
- **Displaced decay** to  $j/\ell/\text{MET}$   
by cosmological condition!



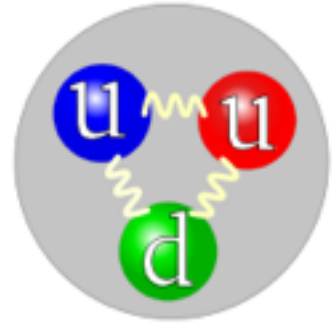
# Baryon $\Omega_B \approx 5\%$

## — The Unknown Aspects of the Known

- **Baryon**: proton, neutron  $\Rightarrow$  atoms, stars, ourselves!
- Where does  $\Omega_B$  come from?  
= Where do we ourselves come from?



NEUTRON  
Quark structure



PROTON  
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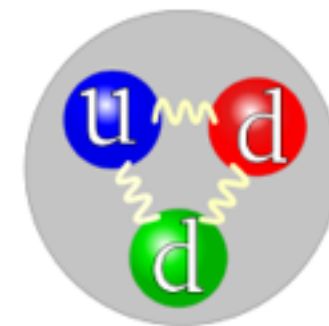
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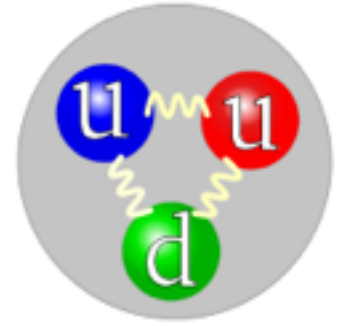
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**We do not know!**



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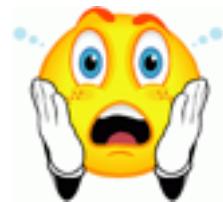
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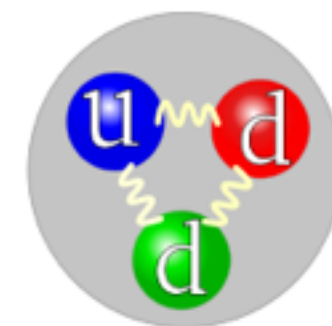
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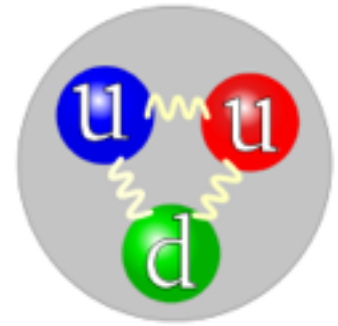
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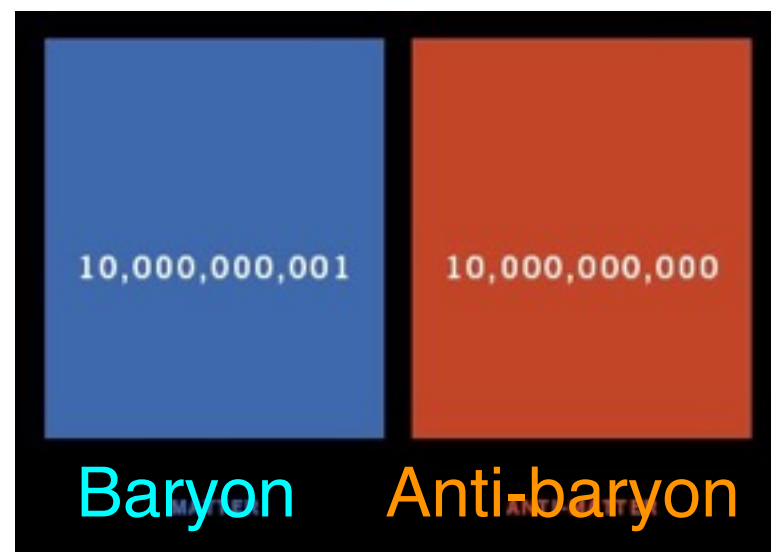
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**Initial  $B - \bar{B}$  asymmetry**

$$\eta_B = (n_B - n_{\bar{B}})/n_\gamma \sim 10^{-10}$$



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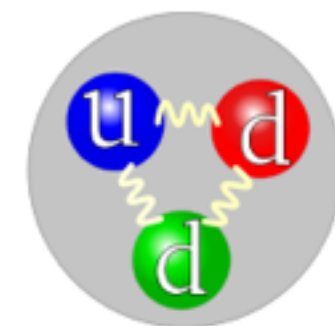
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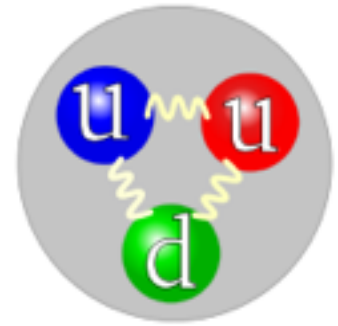
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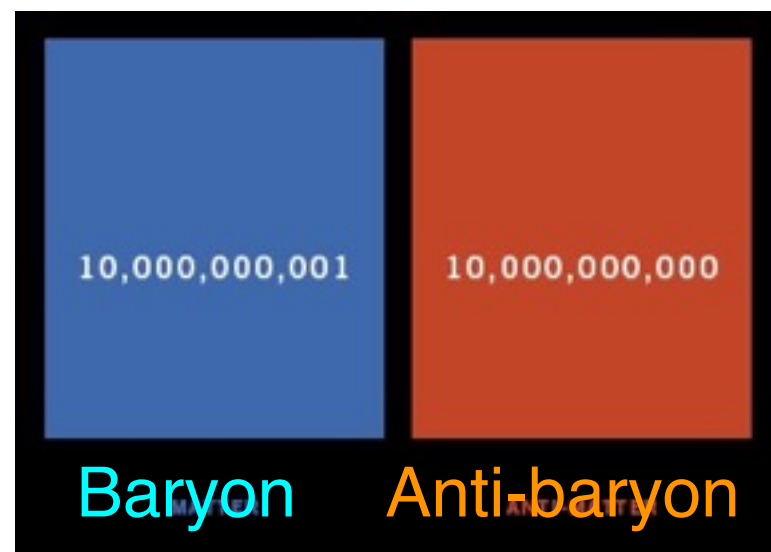
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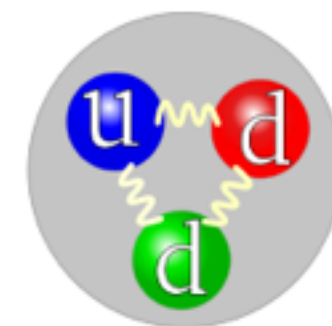
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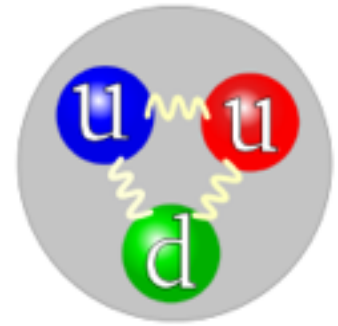
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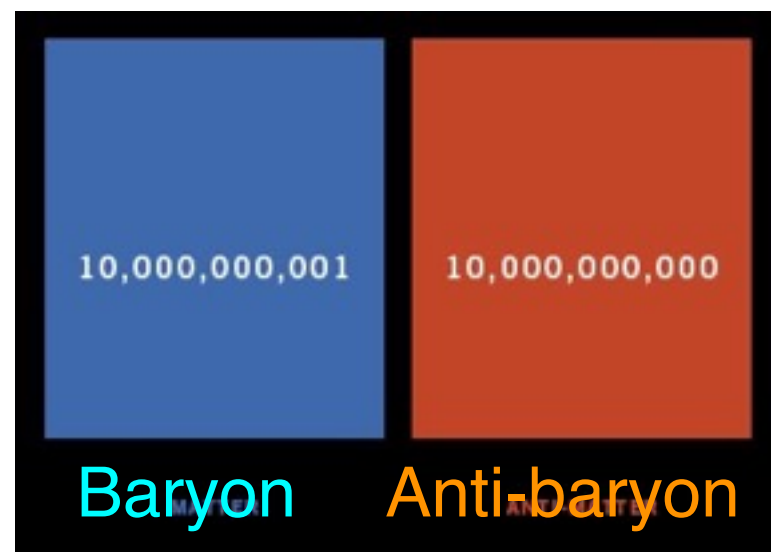
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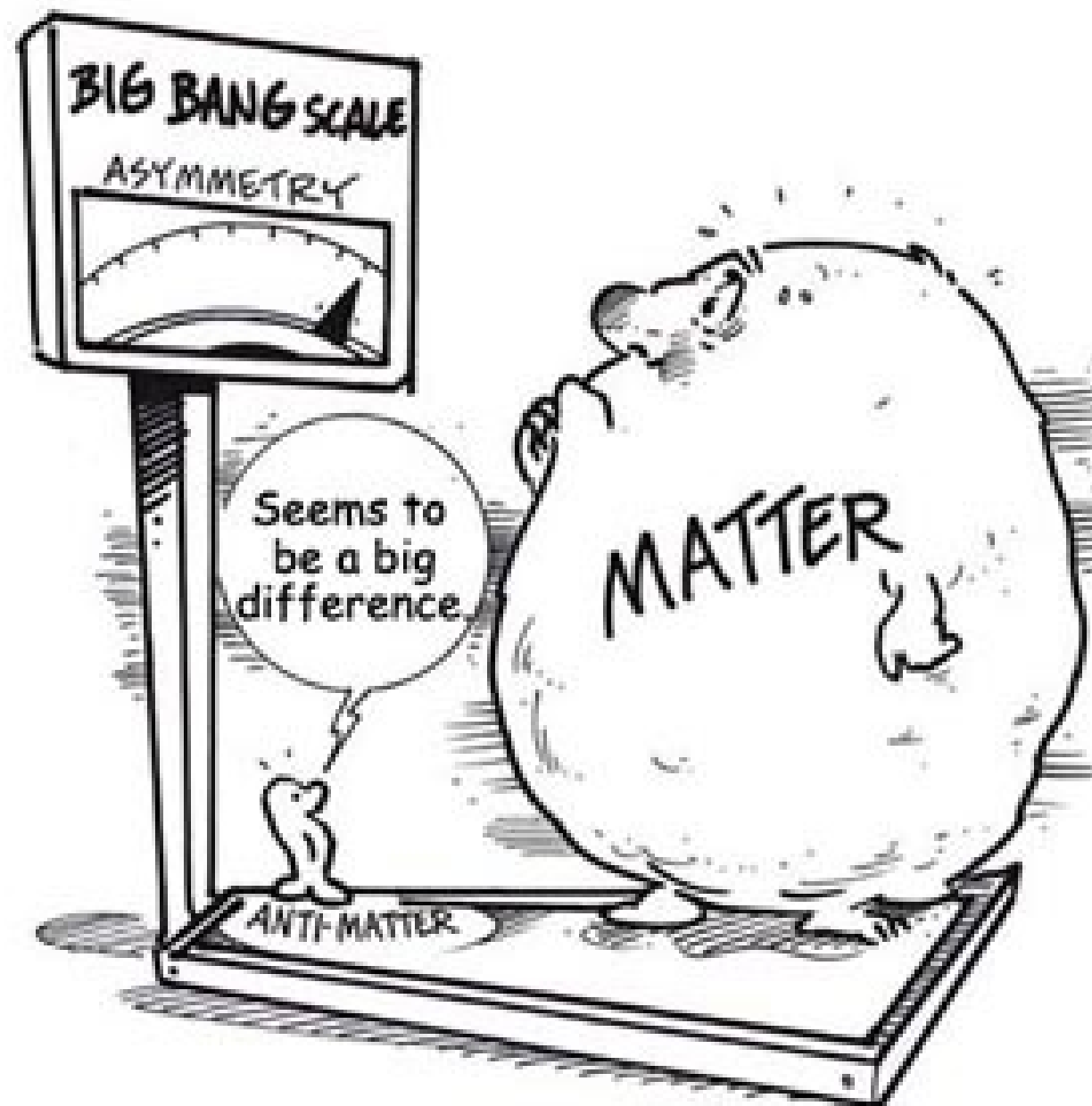


# Baryogenesis

## - the Origin of the Baryon Asymmetry

The Universe starts with  $B = 0$ ,  $\rightarrow$  ?  $B \neq 0$

$B - \bar{B}$  asymmetry



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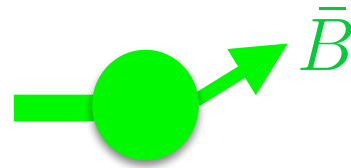
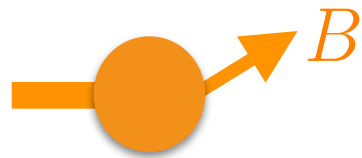
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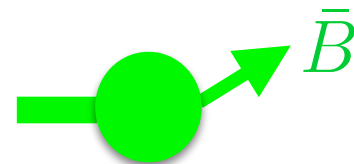
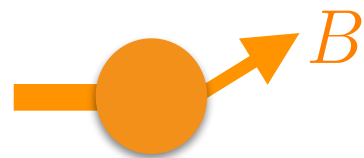
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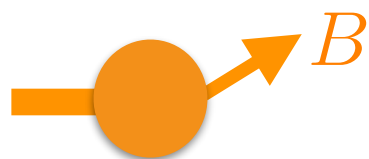
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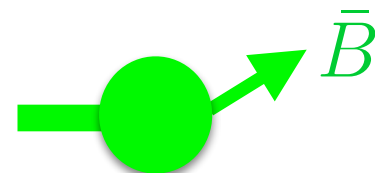
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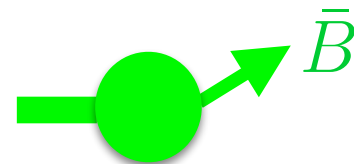
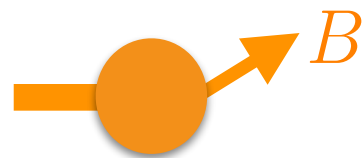
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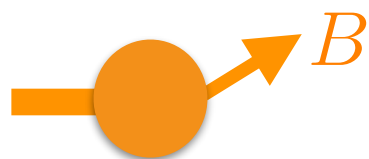
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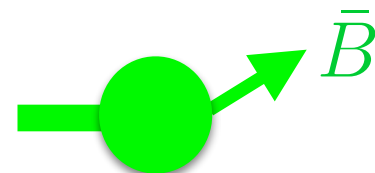
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- Require departure from equilibrium!

Thermal equilibrium + CPT symmetry

$\Rightarrow \langle B \rangle_{\text{eq}} = 0$



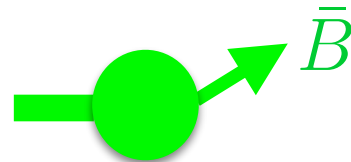
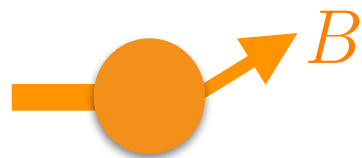


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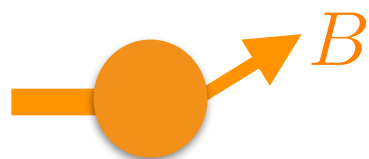
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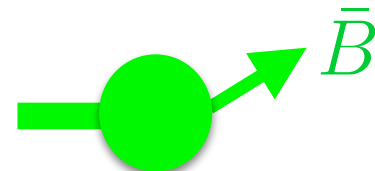
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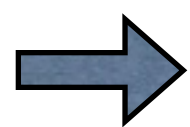


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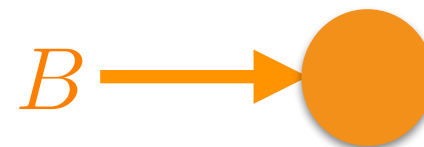
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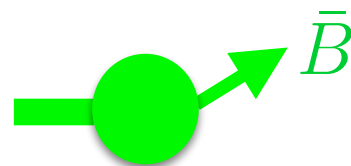
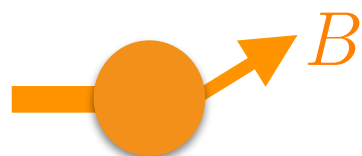


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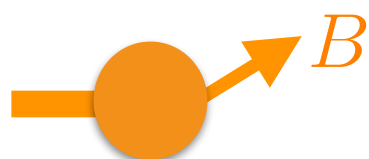
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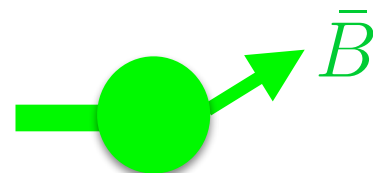
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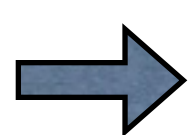


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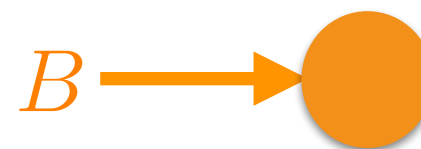
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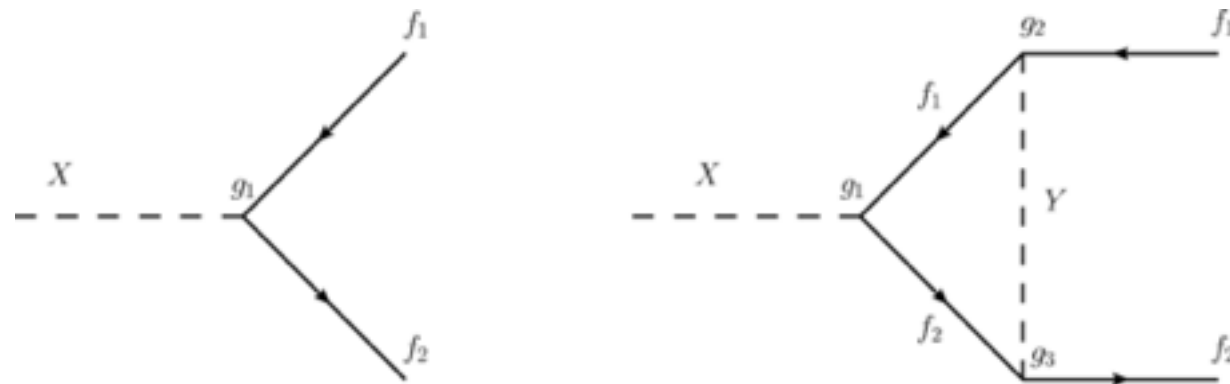
- ❖ Most existing BG mechanisms: high M or/and T, direct experimental test impossible (contrast: WIMP DM for  $\Omega_{\text{DM}}$ )



# Baryogenesis from Out-of-Equilibrium Decay

## A general class of baryogenesis models

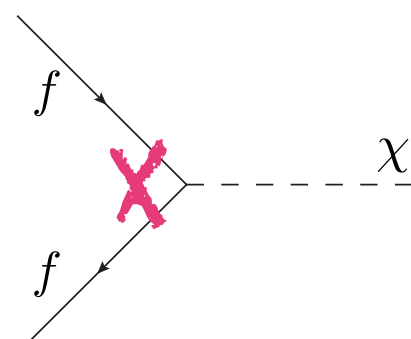
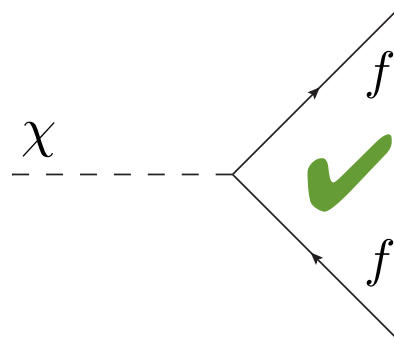
- Assume a massive neutral particle  $\chi$
- Baryon asymmetry can be produced in its decay (B-, CP-violation)



$$\Gamma(\chi \rightarrow f) \neq \Gamma(\chi \rightarrow \bar{f})$$

$$n_f - n_{\bar{f}} \neq 0$$

- Typically, the inverse processes efficiently erase the asymmetry
- But, if  $\chi$  is **long-lived**, and **decays only after  $T_f < M_\chi$** :  $\Gamma_\chi < H(T = M_\chi)$

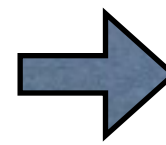


Inverse decay: Boltzmann suppressed  $(H: \text{Hubble expansion rate})$

$$e^{-M_\chi/T_{\text{decay}}}$$



**Out-of-equilibrium decay**



**Sakharov conditions ✓**

# Baryogenesis from Out-of-Equilibrium Decay

**An intriguing observation** (YC, Sundrum 2012; YC, Shuve, 2014)

- If  $\chi$  has **weak scale** mass:

$$\Gamma_\chi < H(T = M_\chi) \quad \longleftrightarrow \quad c\tau_\chi \gtrsim \text{mm}$$

- A **generic connection** between **cosmological slow rates at  $T \sim 100 \text{ GeV}$**  and **displaced vertices at colliders!**

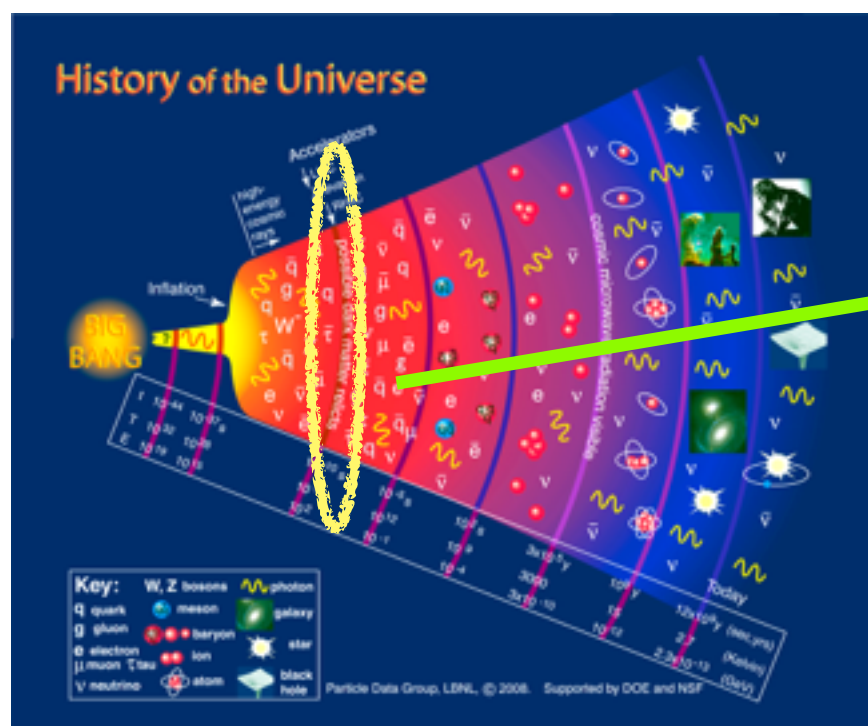
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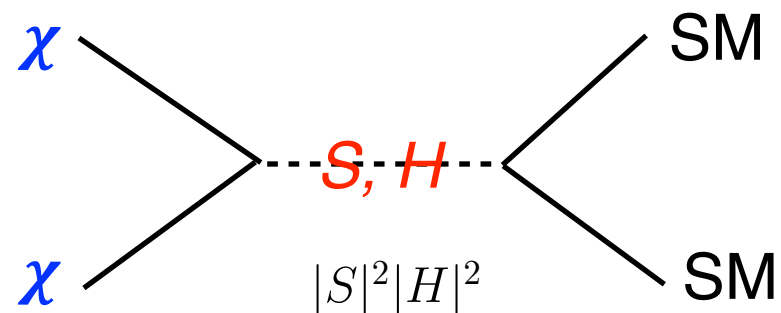
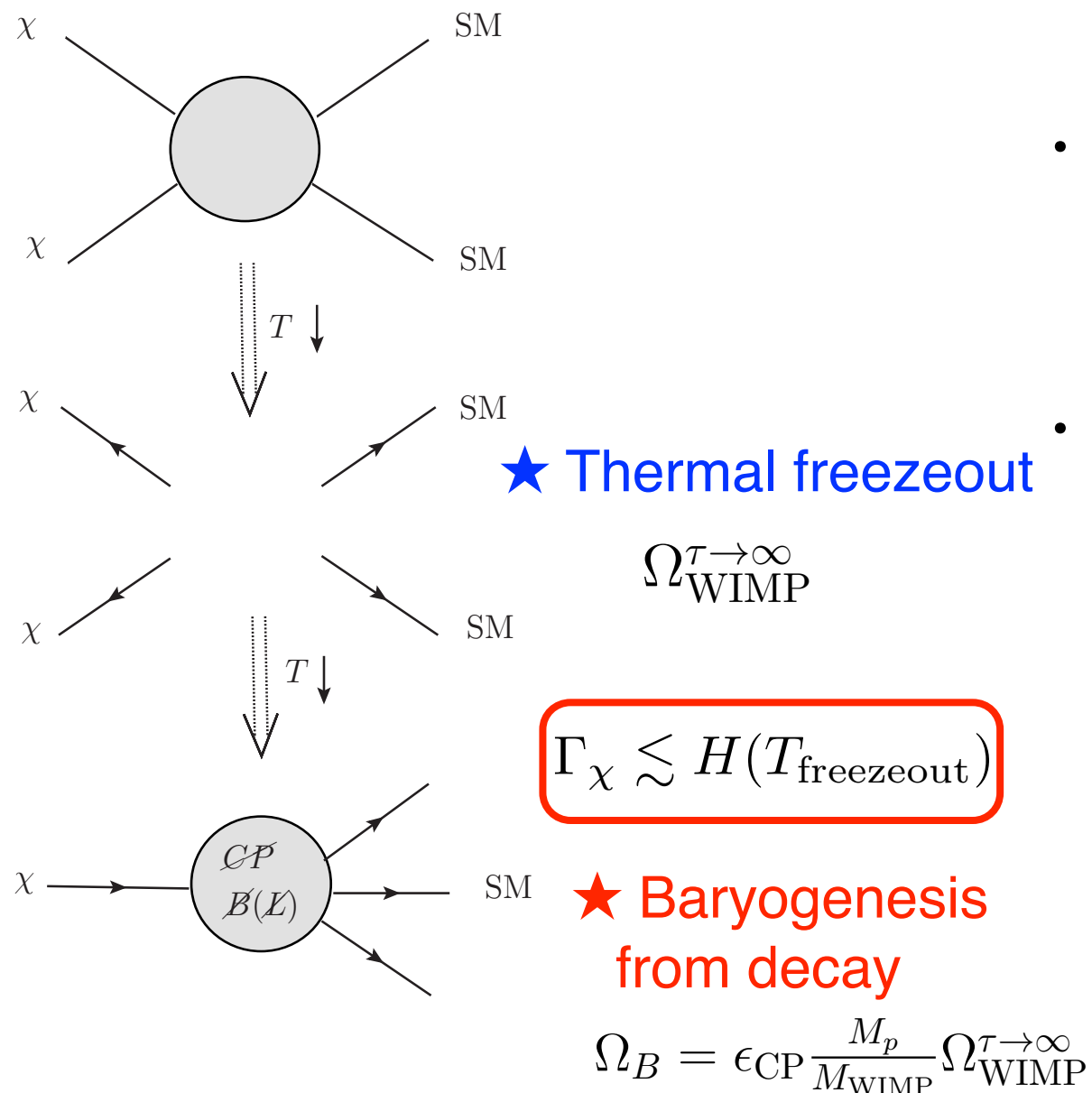
**Our universe around EW phase transition was just slightly bigger than LHC tracking resolution!**

# Baryogenesis from Meta-stable WIMP Decay

## - concrete, motivated model examples

(YC and Sundrum 2012; YC 2014; YC and Okui, Yunesi 2016)

- WIMP miracle prediction for  $\Omega_B$   
+ **new** path addressing  $\Omega_B \sim \Omega_{DM}$
- General mechanism, easy to embed in RPV SUSY (*natural or split*)
- Thermal annihilation of WIMP through e.g. Higgs or a singlet scalar  $S$



**Pair production of long-lived baryon parent WIMP at the LHC!**

# Simplified Model Approach for LHC Pheno

(YC and Shuve, 2014)

- Classify production modes (analogy to DM search @LHC!)
- Classify decay modes (unlike DM search)

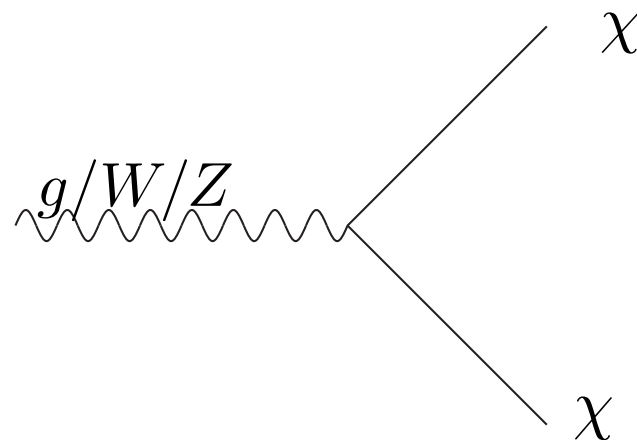
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Charged under SM gauge interactions:

wino/gluino-like (states in interference loop)



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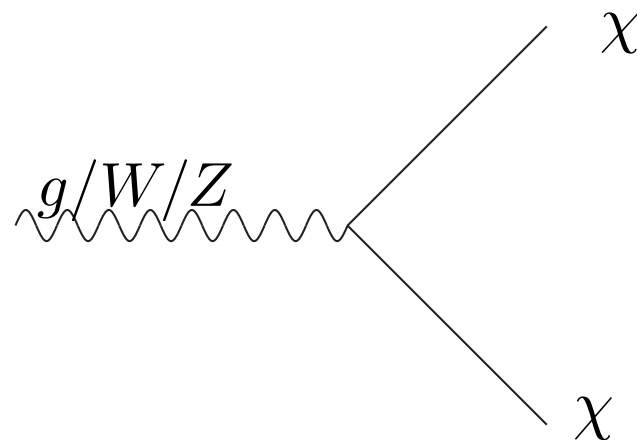
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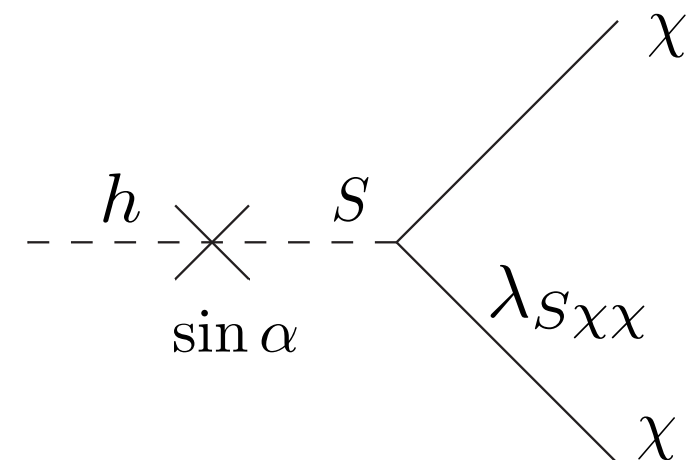
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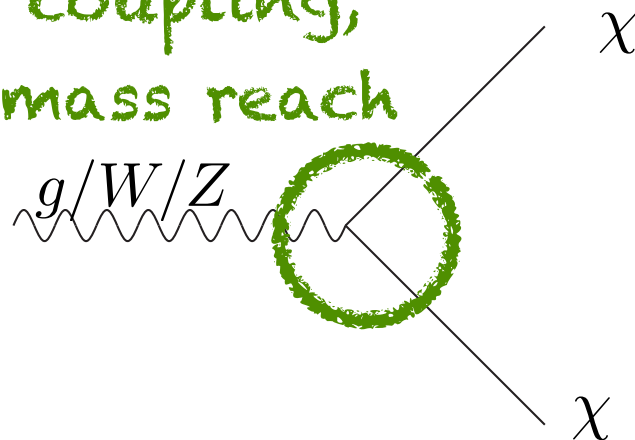
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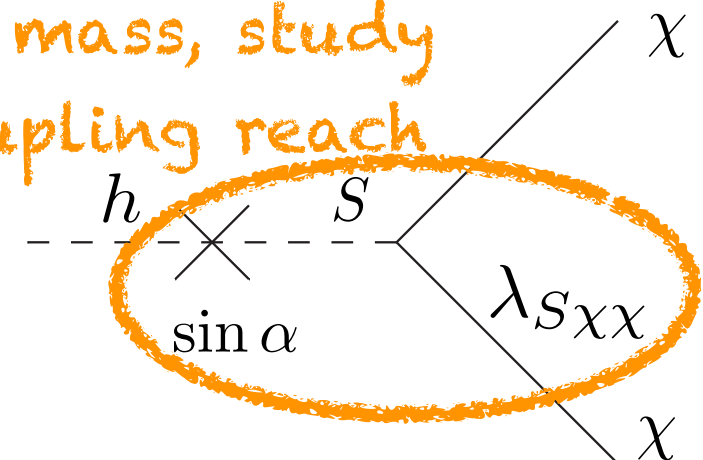
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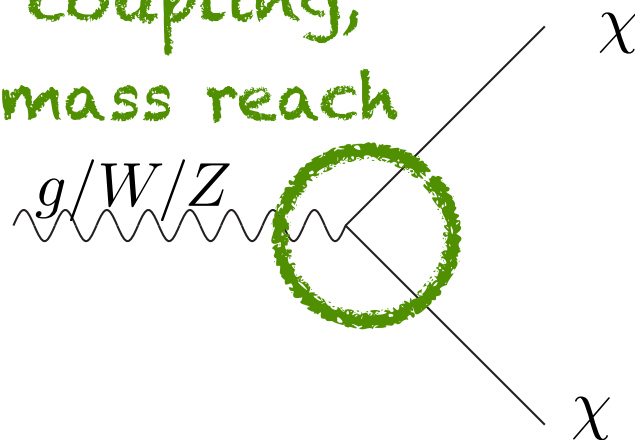
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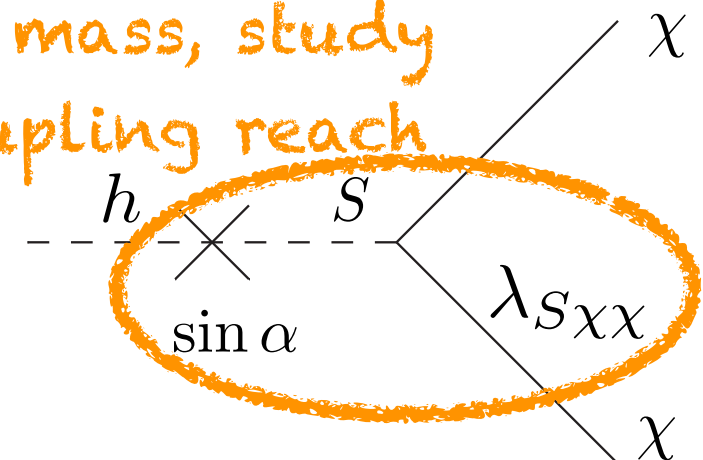
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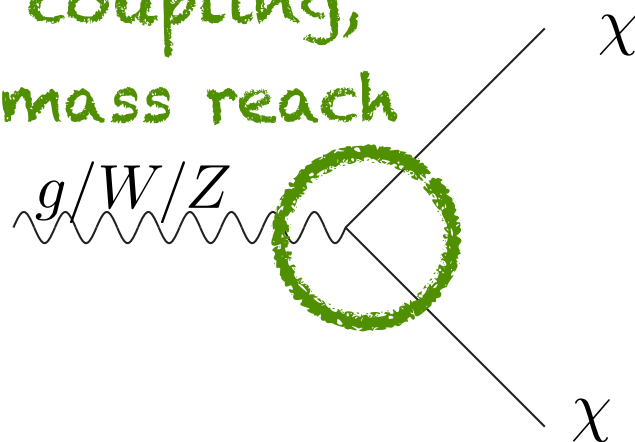
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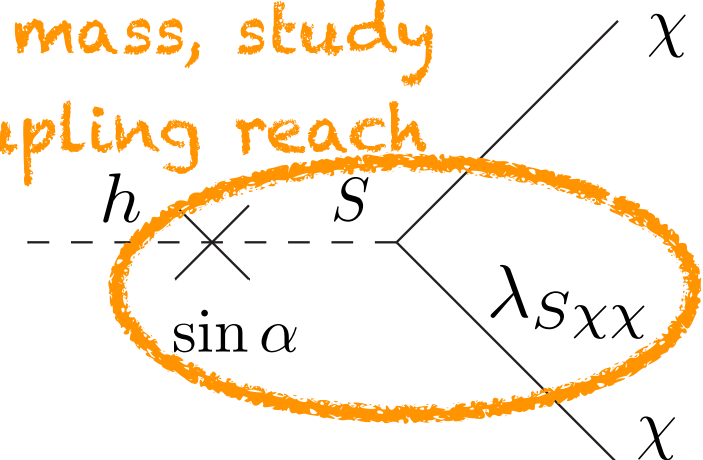
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## Lepton number violating:

$$\chi \rightarrow L_i Q_j \bar{d}_k$$

$$\chi \rightarrow L_i L_j \bar{E}_k$$

# Recast Existing LHC Searches

- Focus on displaced decay in tracking volume

Near lower bound  $c\tau_\chi \gtrsim \text{mm}$  , better sensitivity to wide lifetime range, easier to model with theorists' tools!

(decay in other parts of detector important too!)

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- Two concrete examples (light-flavour only):

**Baryon number violating:**

$$\chi \rightarrow 3q$$

displaced jets (all-hadronic)

CMS, arXiv:1411.6530

**Lepton number violating:**

$$\chi \rightarrow \ell + 2q$$

displaced muon + tracks

ATLAS-CONF-2013-092

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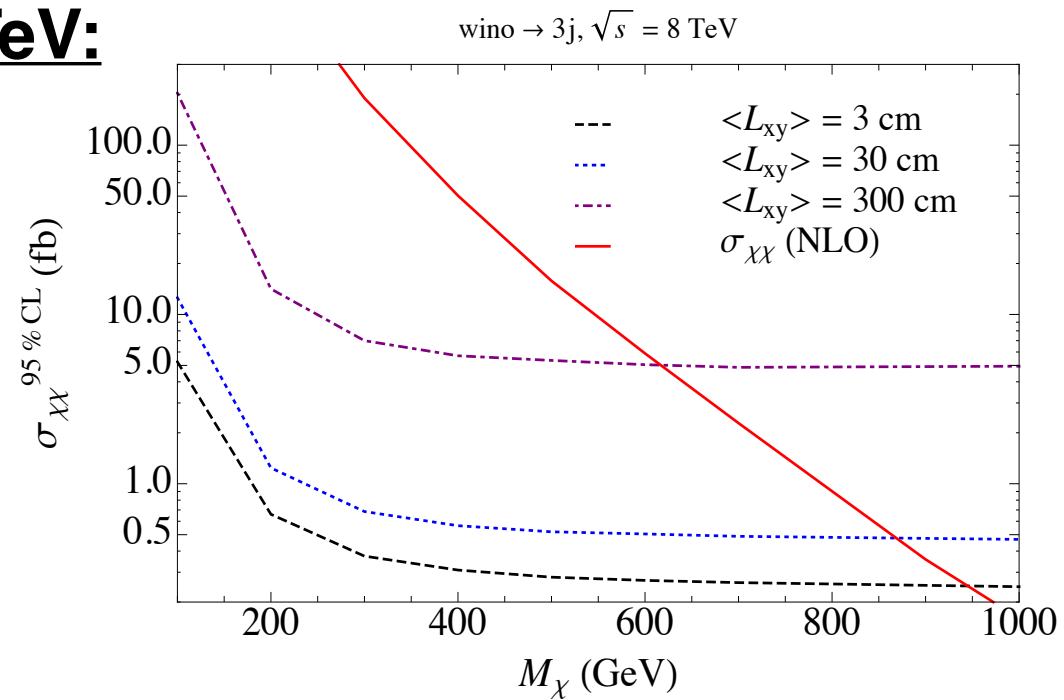
- What is the coverage for our simplified models based on benchmarks chosen by the collaborations?
- What advice can we provide for general experimental improvement?

# Fully hadronic displaced vertices

CMS displaced dijet, arXiv:1411.6530

*wino*

**8 TeV:**

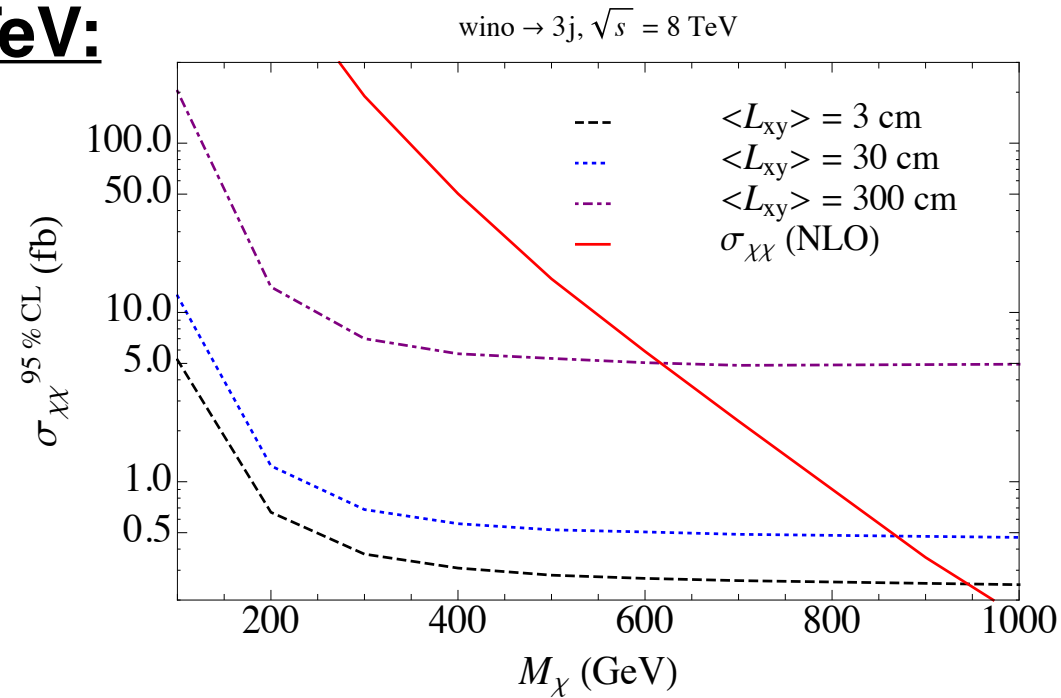


# Fully hadronic displaced vertices

CMS displaced dijet, arXiv:1411.6530

*wino*

**8 TeV:**



*singlet-like (Higgs portal)*

We studied a challenging case:

$M_\chi = 150$  GeV, moderately off-shell!

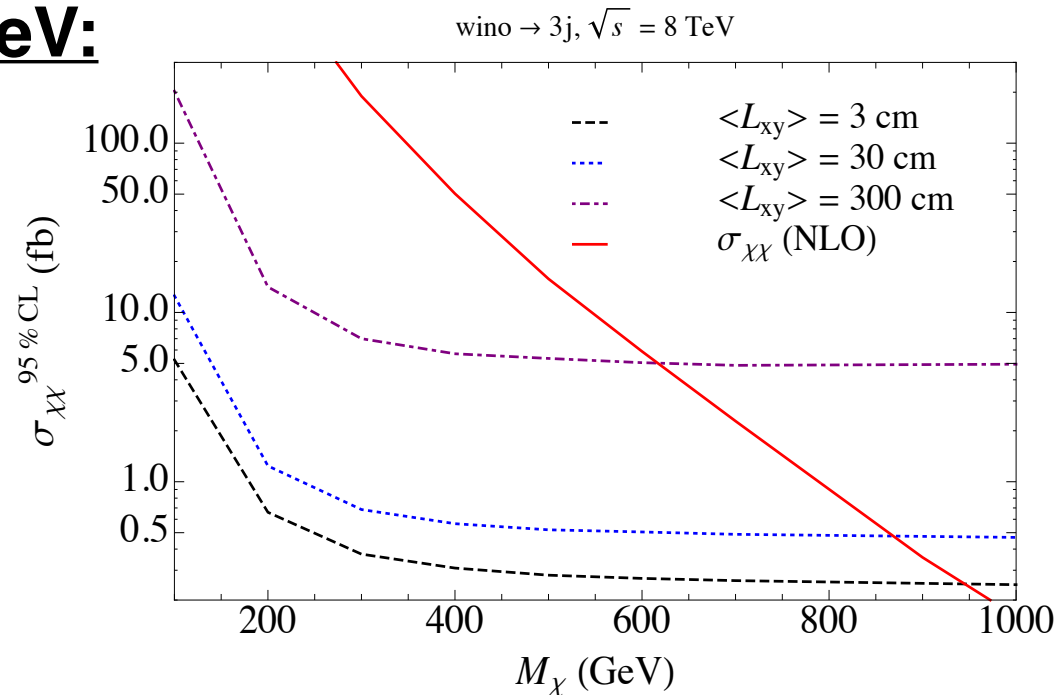
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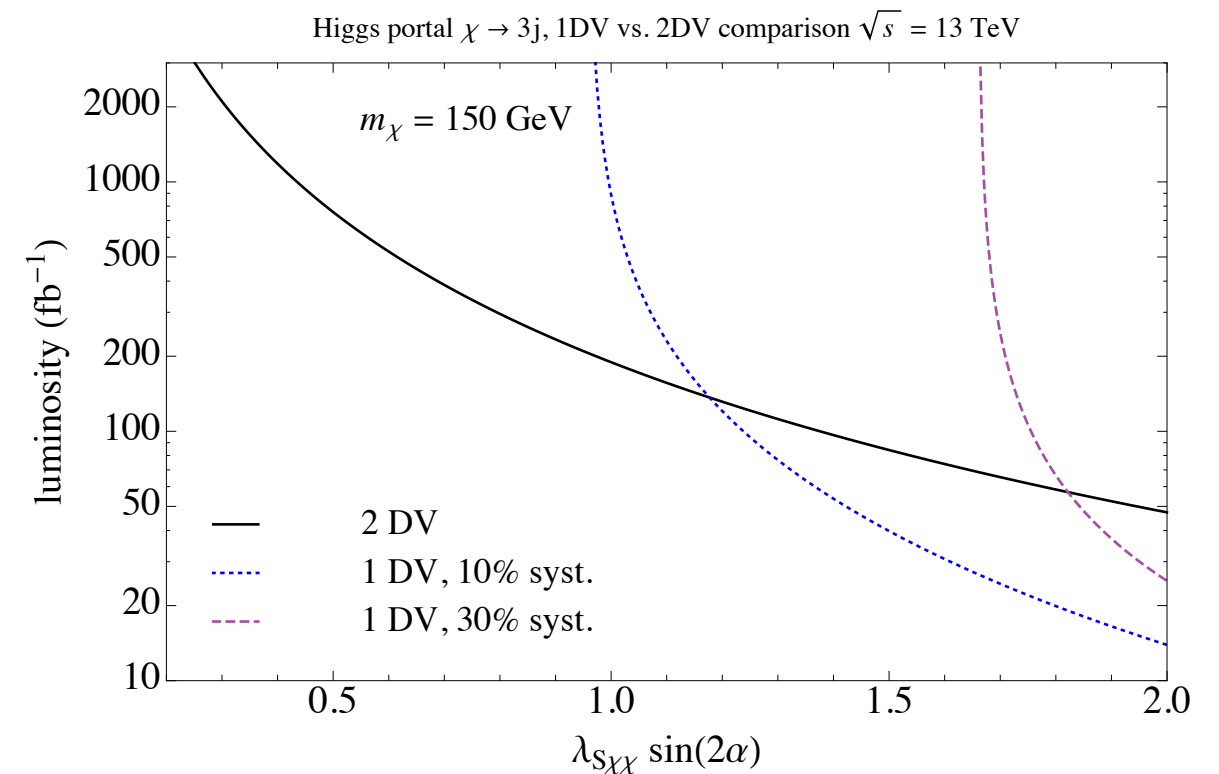
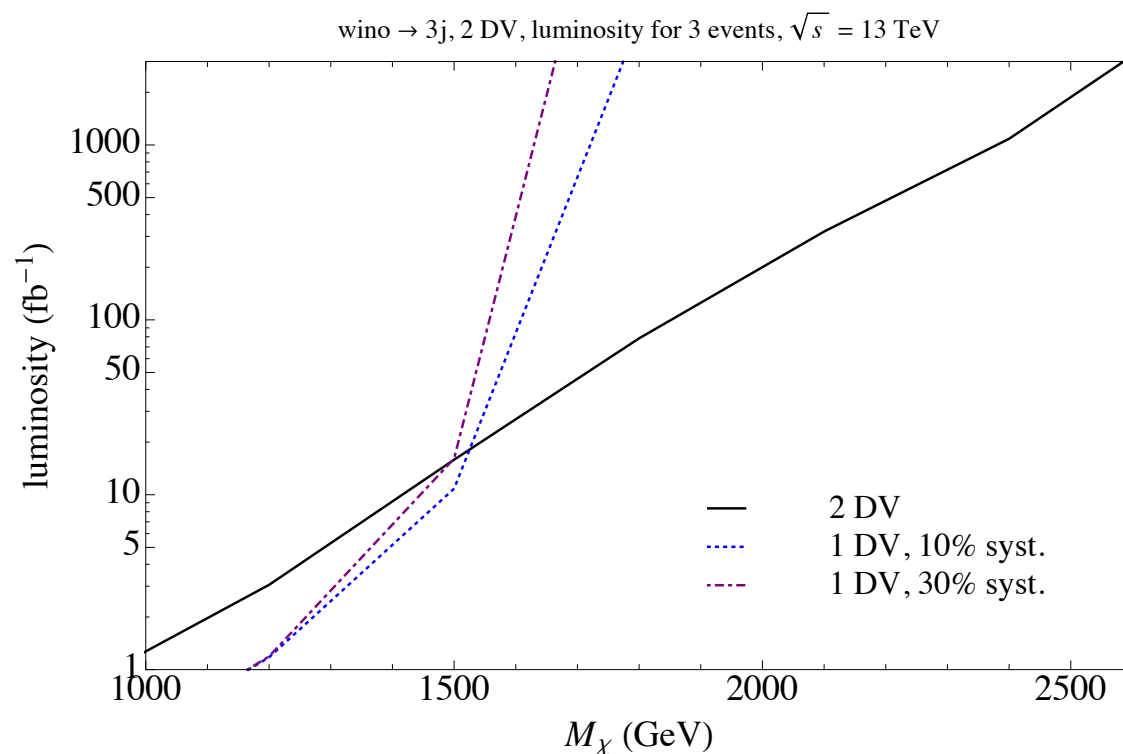
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$L_{xy} = 3 \text{ cm}$



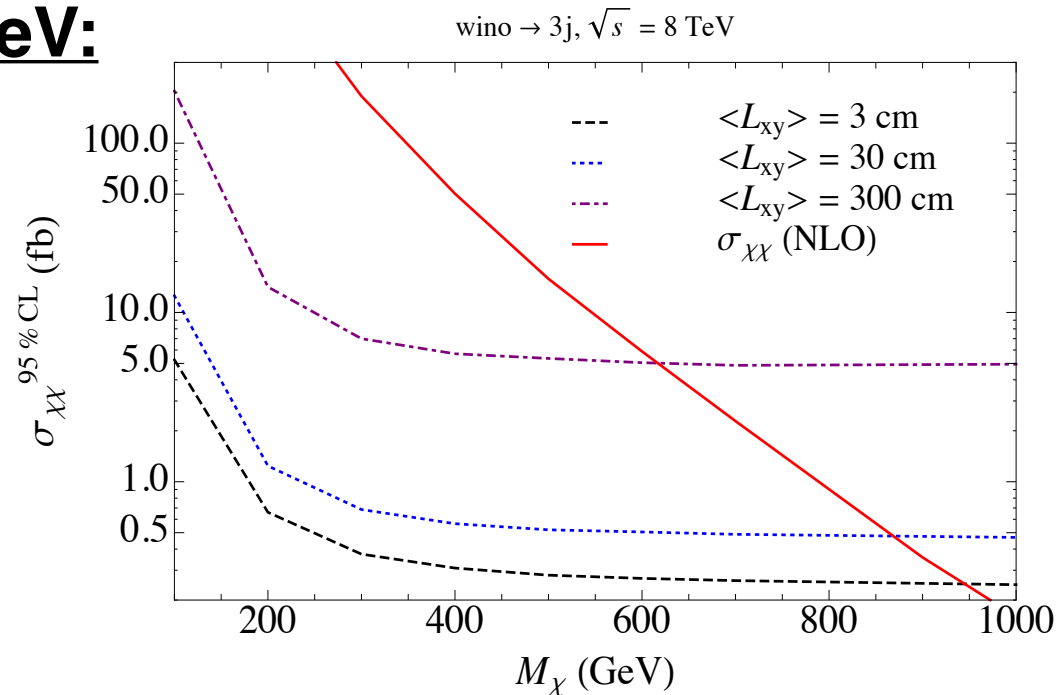


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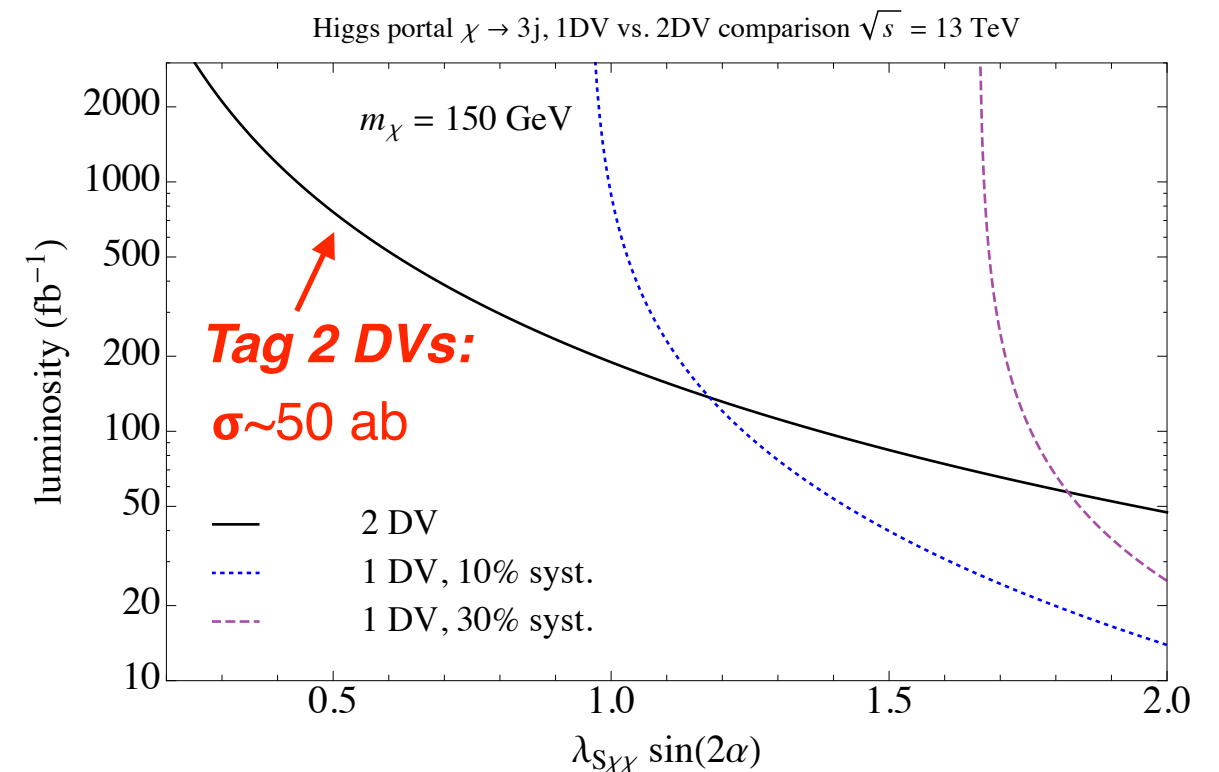
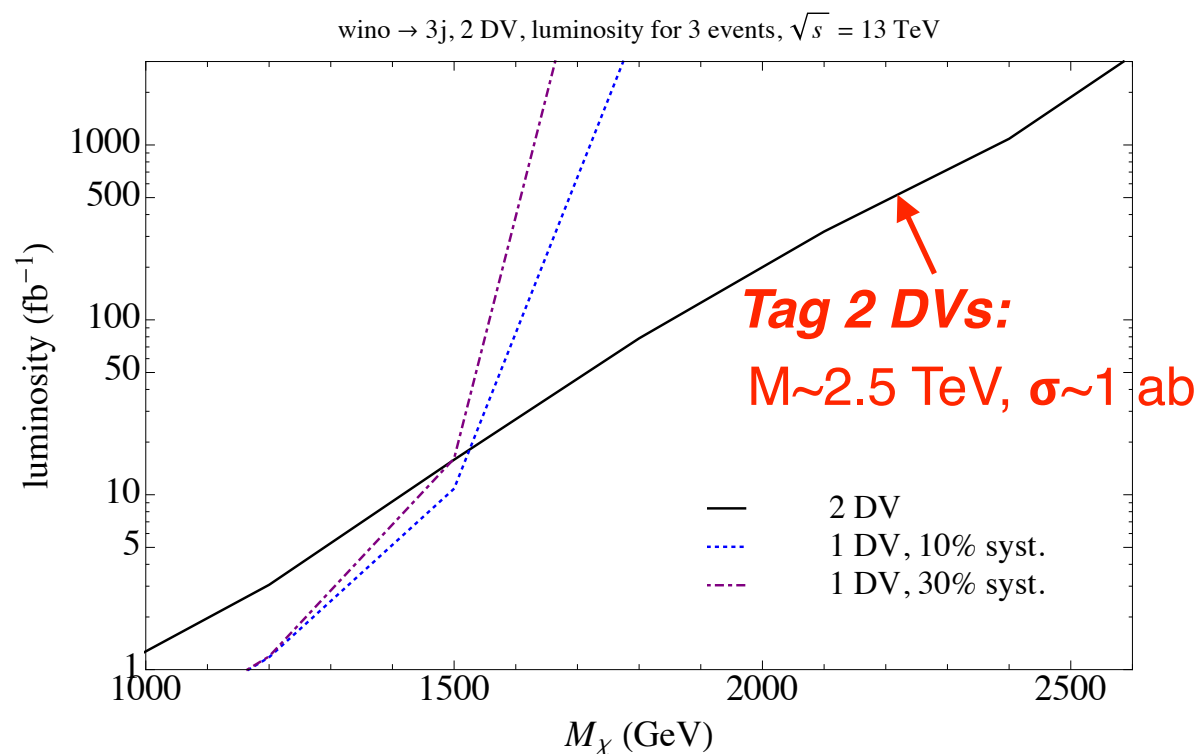
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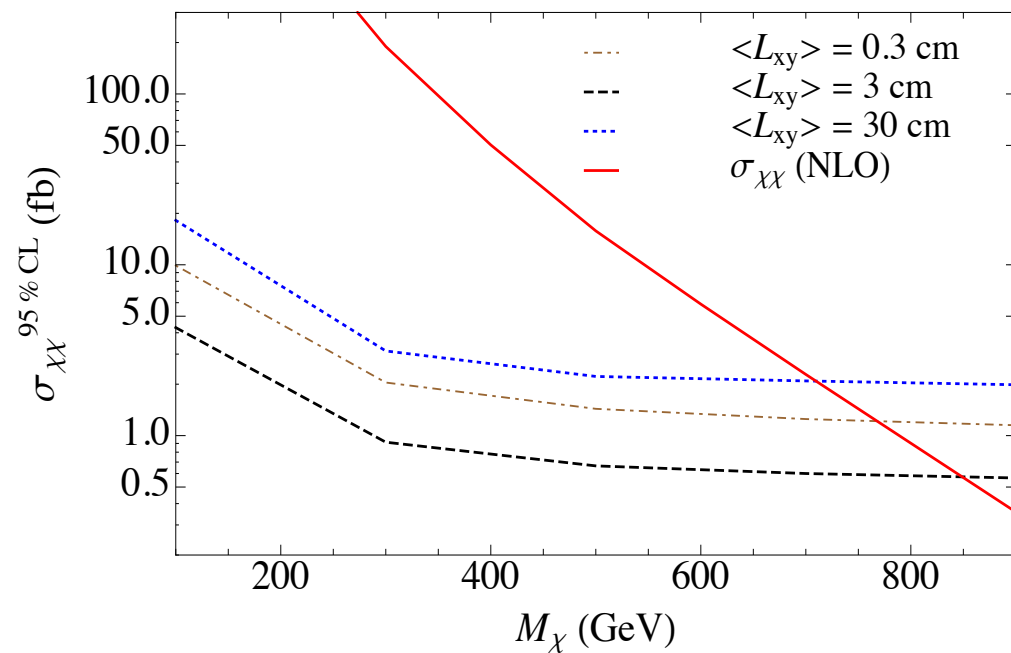


# Displaced muon + Tracks

ATLAS-CONF-2013-092

*wino*

wino  $\rightarrow \mu + \text{tracks}$ ,  $\sqrt{s} = 8 \text{ TeV}$



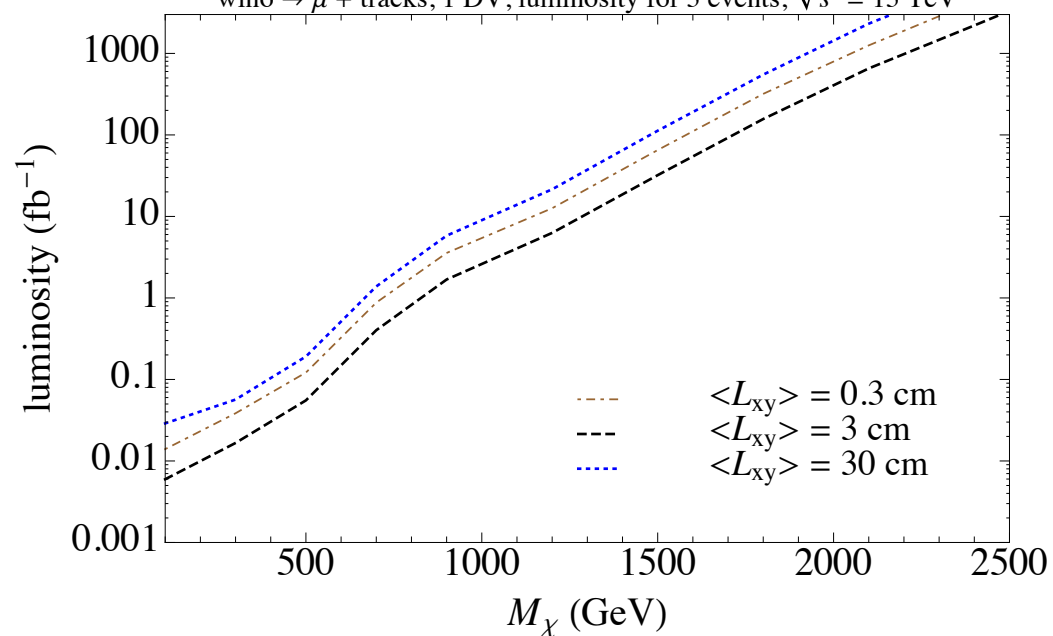
*8 TeV*

*13 TeV:*

**Tag 1 DV**  
 **$M \sim 2.5 \text{ TeV}$**

*(lower bkg  
than all-  
hadronic)*

wino  $\rightarrow \mu + \text{tracks}$ , 1 DV, luminosity for 3 events,  $\sqrt{s} = 13 \text{ TeV}$

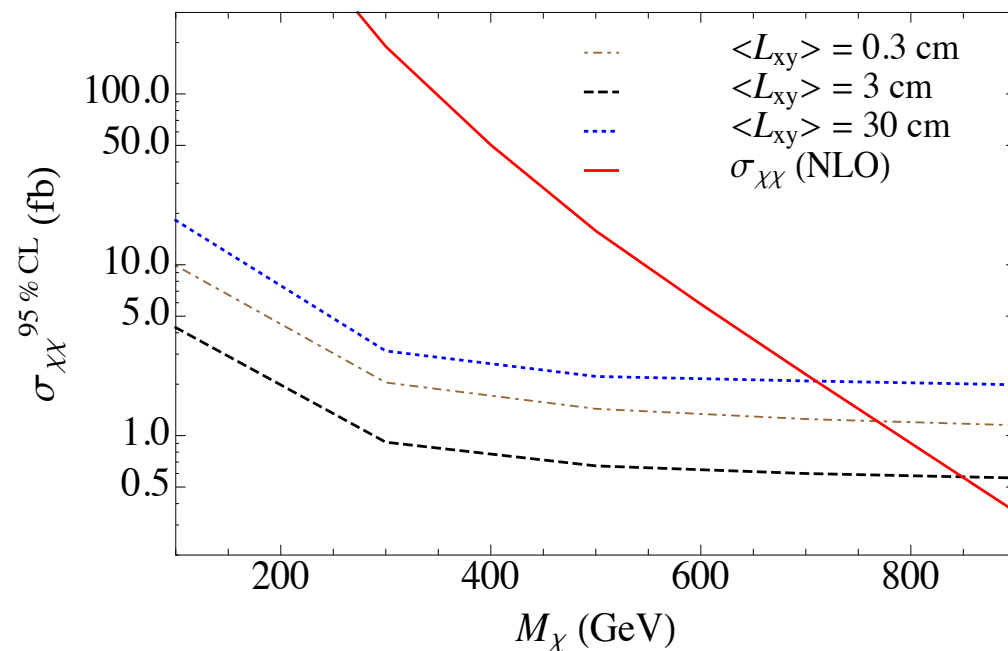


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ATLAS-CONF-2013-092

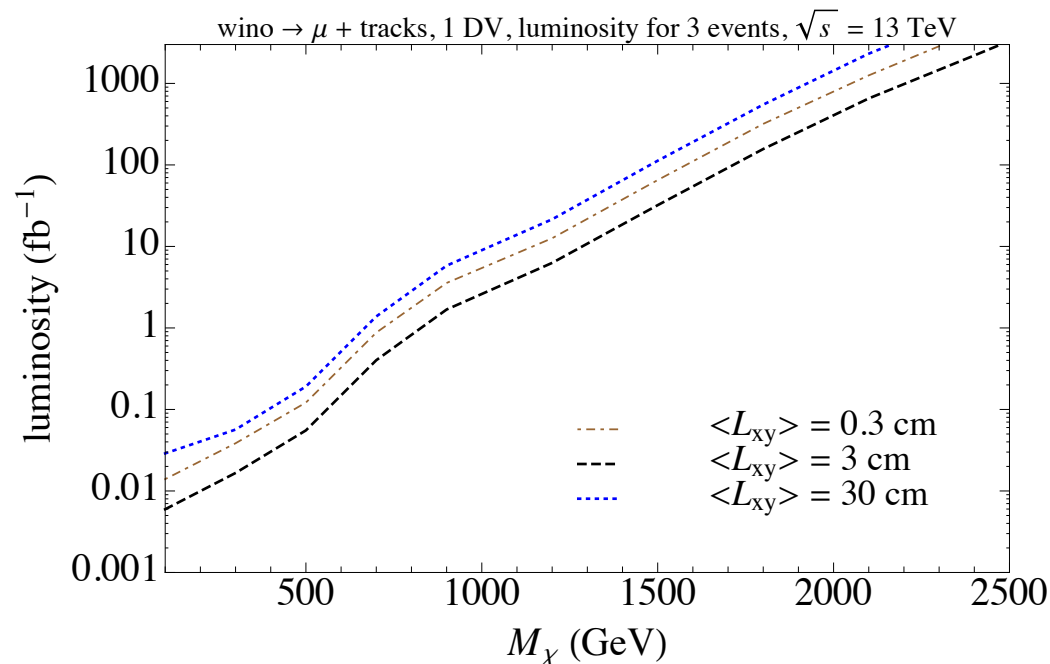
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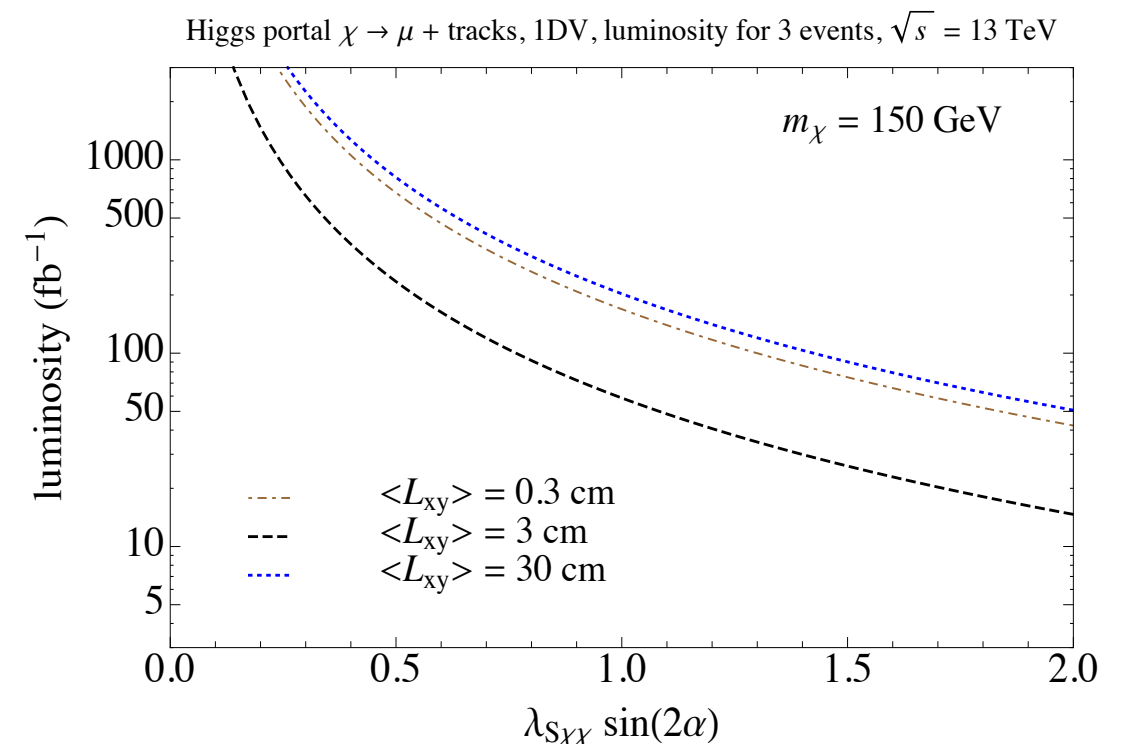


*singlet (Higgs portal)*

(singlet-like,  $M_\chi = 150 \text{ GeV}$ )

No bound @ 8 TeV 20  $\text{fb}^{-1}$

- 13 TeV:  $\sigma_S \sim 50 \text{ ab}$  for  $L_{xy} \sim 1 \text{ cm}$   
(Tag 1 DV)



# Opportunities for CLIC?

— *Some Preliminary Thoughts*

LHC has its limitations for DV searches,  
there are opportunities for CLIC!

- **A general challenge at the LHC:** trigger for low mass all-hadronic DV events (overwhelmed by large QCD background at a hadron collider!)  
e.g.  $L1 H_T$  trigger at 13 TeV:  $H_T \gtrsim 500 \text{ GeV} \rightarrow$  a big loss of signals from Higgs-portal production

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— **This is particularly relevant for WIMP BG!**

Out of equilibrium decay of low mass WIMP ( $\lesssim 100 \text{ GeV}$ ) typically occurs well after EW phase transition when sphaleron process ( $\Delta L \leftrightarrow \Delta B$ ) shuts off  $\rightarrow$

Has to be B-violating decay, i.e. hadronic final states!

# Opportunities for CLIC?

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- **Another general challenge at the LHC:** **ultra long-lived particles** (simply MET? motivate MATHUSLA...)
  - **Relevance for WIMP BG:** wider range of lifetimes (from weak scale to BBN), light WIMPs typically decay later

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- Optimizing the opportunity at the LHC to overcome/relieve these challenges: worthwhile ongoing effort (*e.g. work in progress: YC with Joglekar, Shuve, Tsai*)
- **CLIC has a clear edge vs. LHC!**  
Lepton collider, much cleaner for hadronic states, should be much easier to trigger low-mass events



# Opportunities for CLIC?

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## Examples of CLIC's potential advantages over LHC:

- Singlet-like WIMP: limited sensitivity even after HL-LHC
  - Hadronically-decaying light WIMPs produced from light  $Z'$
  - Hadronically-decaying light WIMPs produced from  $H$ -portal (caveat: the heavier Higgs in the S-H system may be easier to probe with the LHC)
- Wino-like WIMP: HL-LHC has excellent coverage for 1.5 TeV (or higher) mass wino with  $L_{xy} \sim 1$  cm, hard to beat; but not for much longer  $L_{xy}$ !
  - Advantages with CLIC: MET+mono- $\gamma$ , suppressed bkg from jet punch-through for decay in  $\mu$ -spectrometer

*Dedicated consideration/design with CLIC certainly helps!*

# Summary/Outlook

- **Baryogenesis from metastable weak scale particle decay:**
  - **A robust cosmological motivation** for DV searches
  - Exciting opportunity to reproduce and study the early universe BG @ collider experiments today! (cf. WIMP DM search)
- **WIMP baryogenesis:** a motivated example, new mechanism addressing  $\Omega_B (+) \Omega_B \sim \Omega_{DM}$ , natural embedding in SUSY
- Good reaches at HL-LHC for certain models/parameter regions, but still great potential for improvement with CLIC!