

The Initial Correlations of the Glasma Energy-Momentum Tensor

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with advisors:

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Rencontres QGP

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Étretat

[arXiv:1807.????](#)



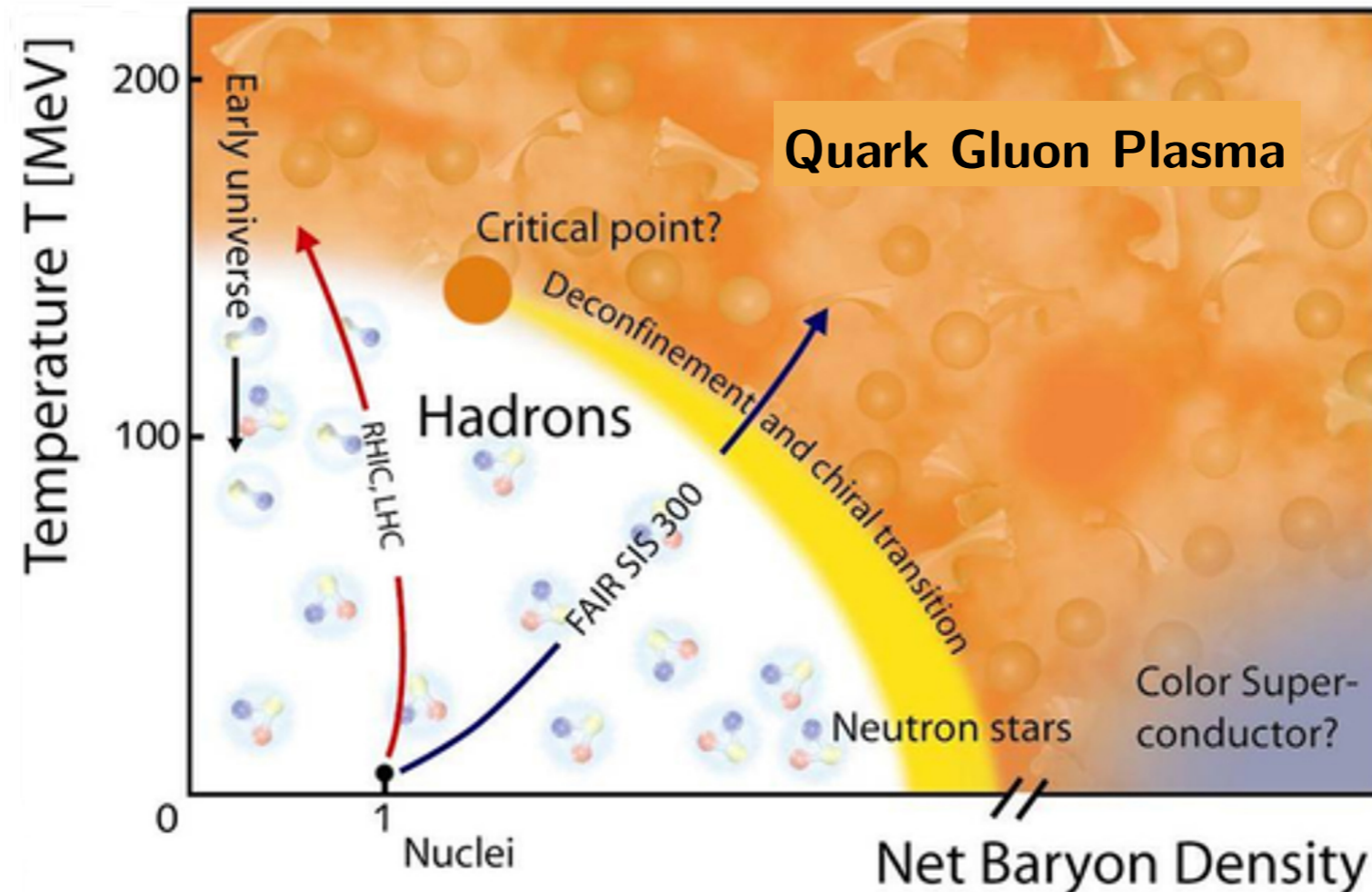
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Motivation and technical framework

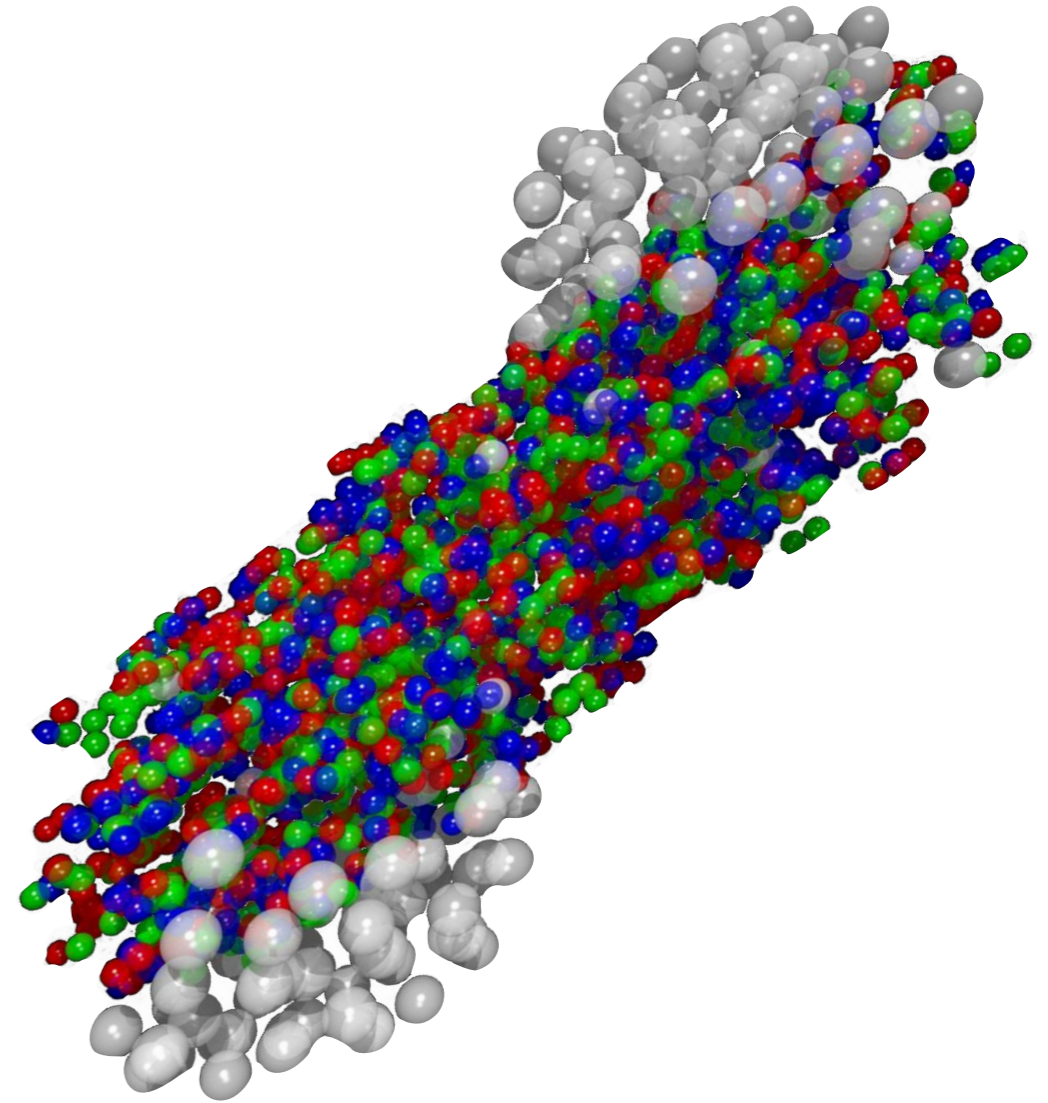
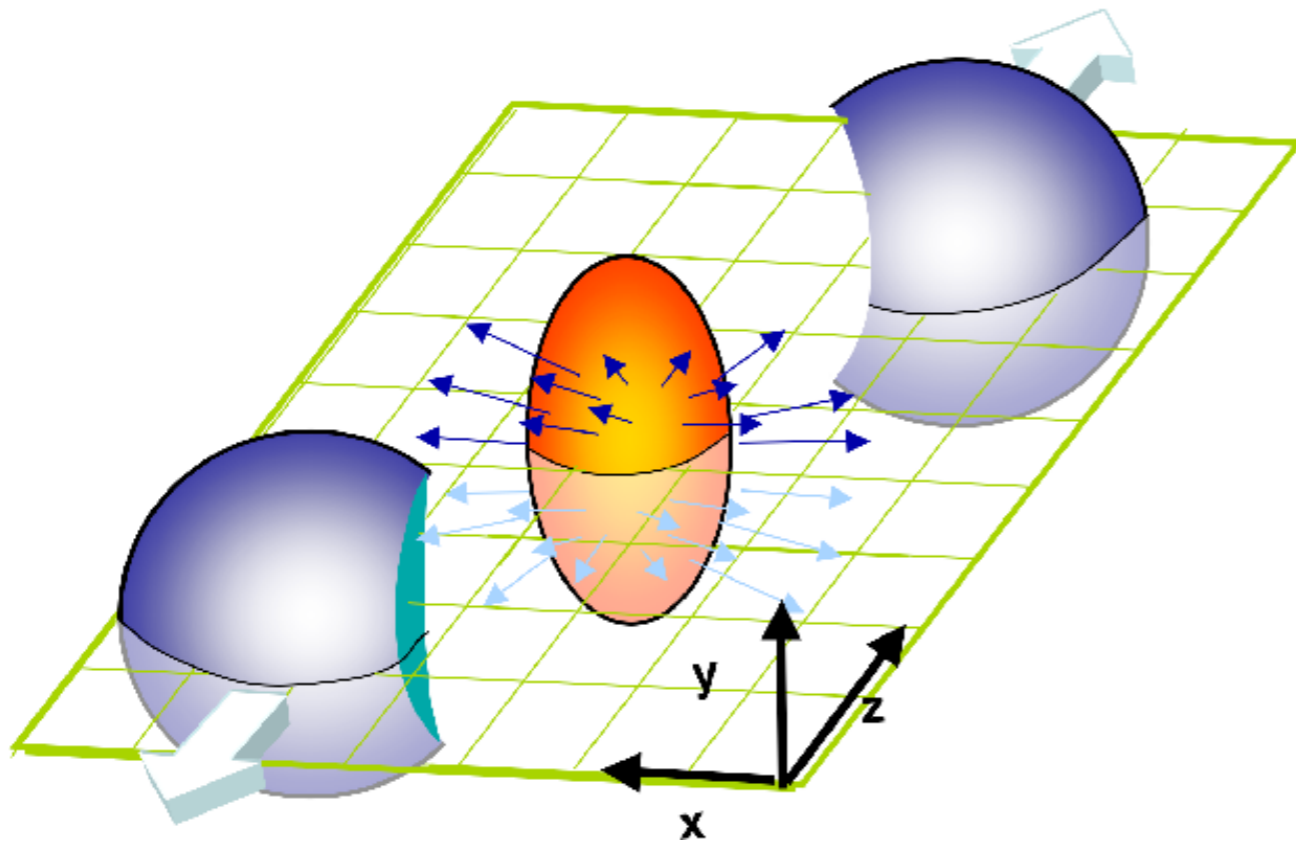
The QCD phase space



- QCD behaves differently depending on conditions of temperature and baryon density
- Low temperature and densities: **hadronic phase (confinement and spontaneously broken chiral symmetry)**
- Lattice simulations indicate a transition at high temperature to a **deconfined, chiral-symmetric phase: The QUARK-GLUON PLASMA**

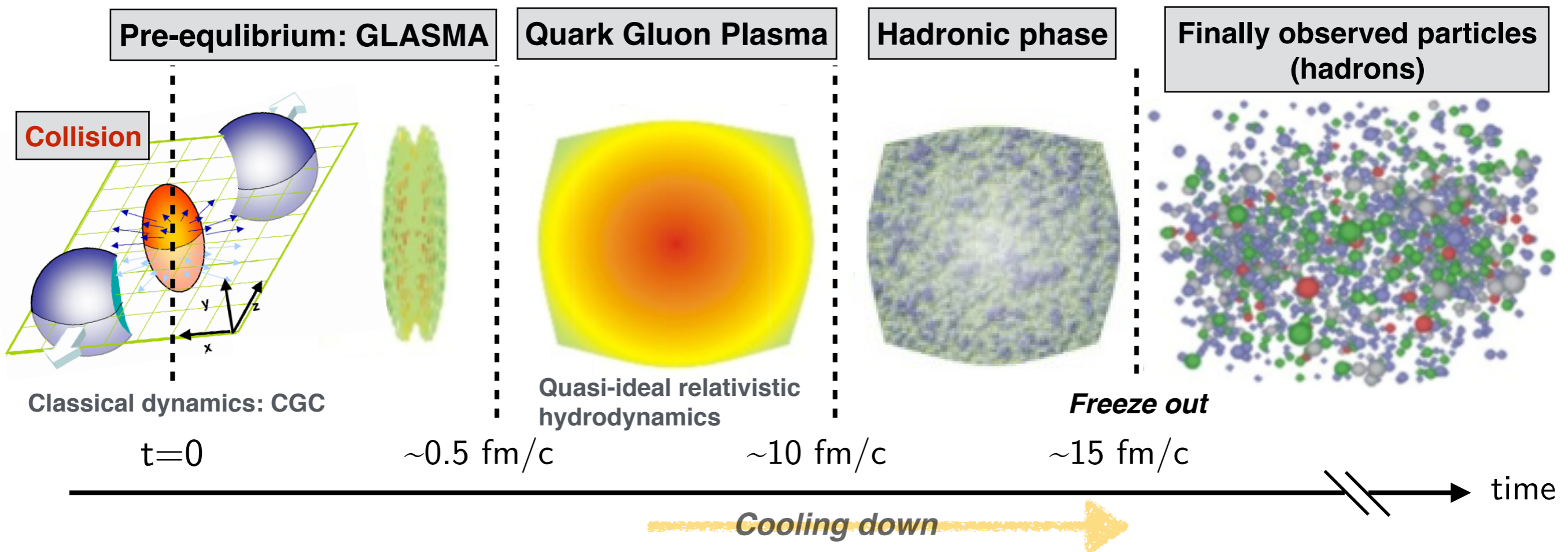
The QCD phase space

- This state of matter can be accessed in particle colliders through **Heavy Ion Collision** experiments



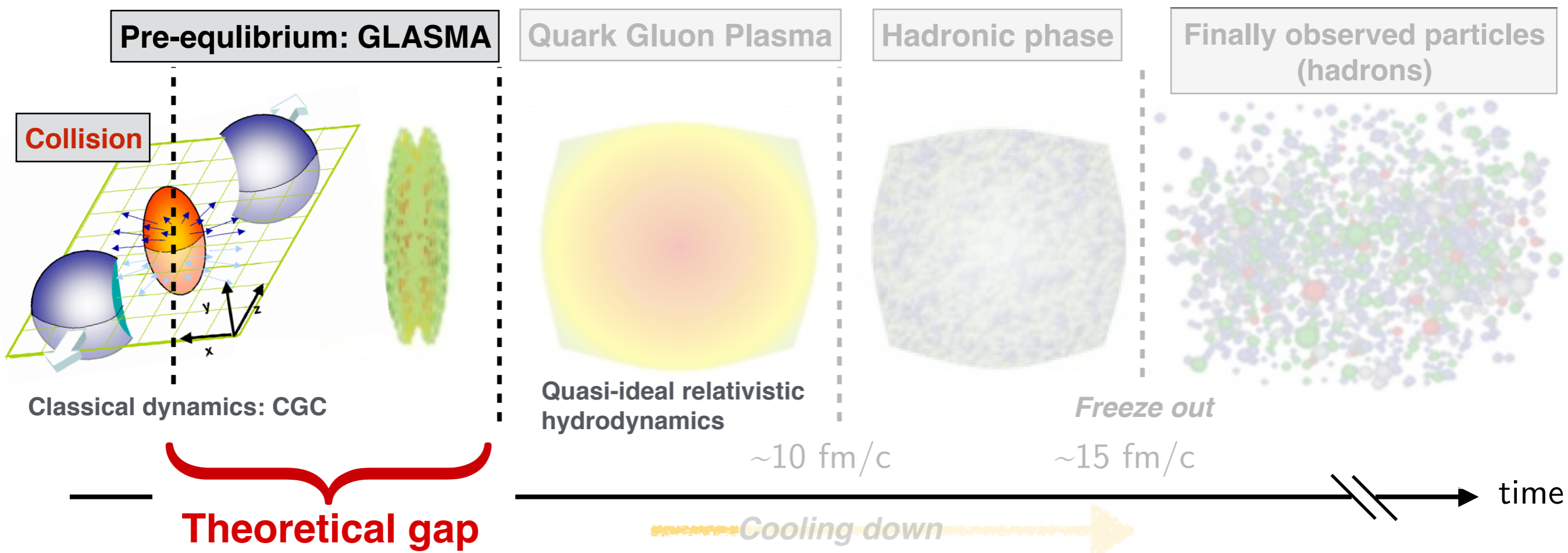
- Performed at Brookhaven National Laboratory's Relativistic Heavy Ion Collider (**RHIC**) and CERN's Large Hadron Collider (**ALICE** experiment)

Stages of a heavy ion collision



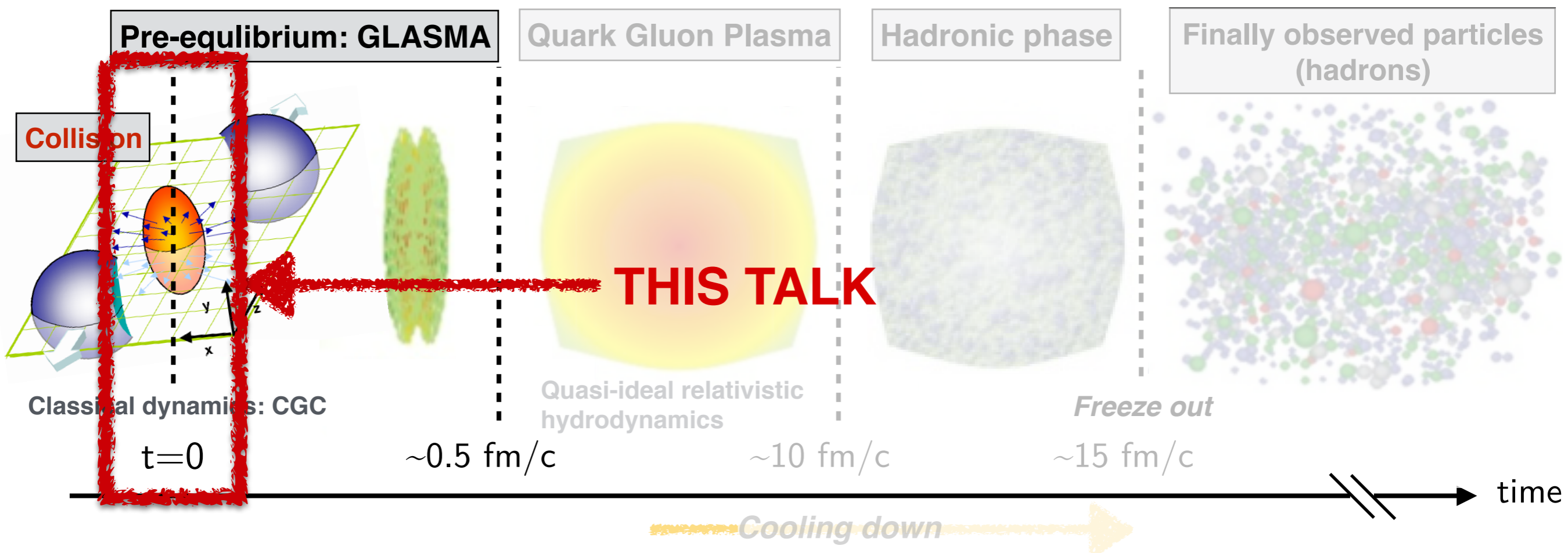
- After the collision, matter goes through different phases as it cools down
- In the last part, it reaches the hadronic phase, and this is how it appears in the detectors

Stages of a heavy ion collision



- There is a theoretical gap between the description of the early phase and the simulations of the expansion of the QGP
- Solid theoretical results are needed to mediate between both frameworks

Stages of a heavy ion collision



- There is a theoretical gap between the description of the early phase and the simulations of the expansion of the QGP
- Solid theoretical results are needed to mediate between both frameworks
- We provide a first-principles analytical calculation of:

$$\langle T^{\mu\nu}(x_{\perp}) \rangle$$

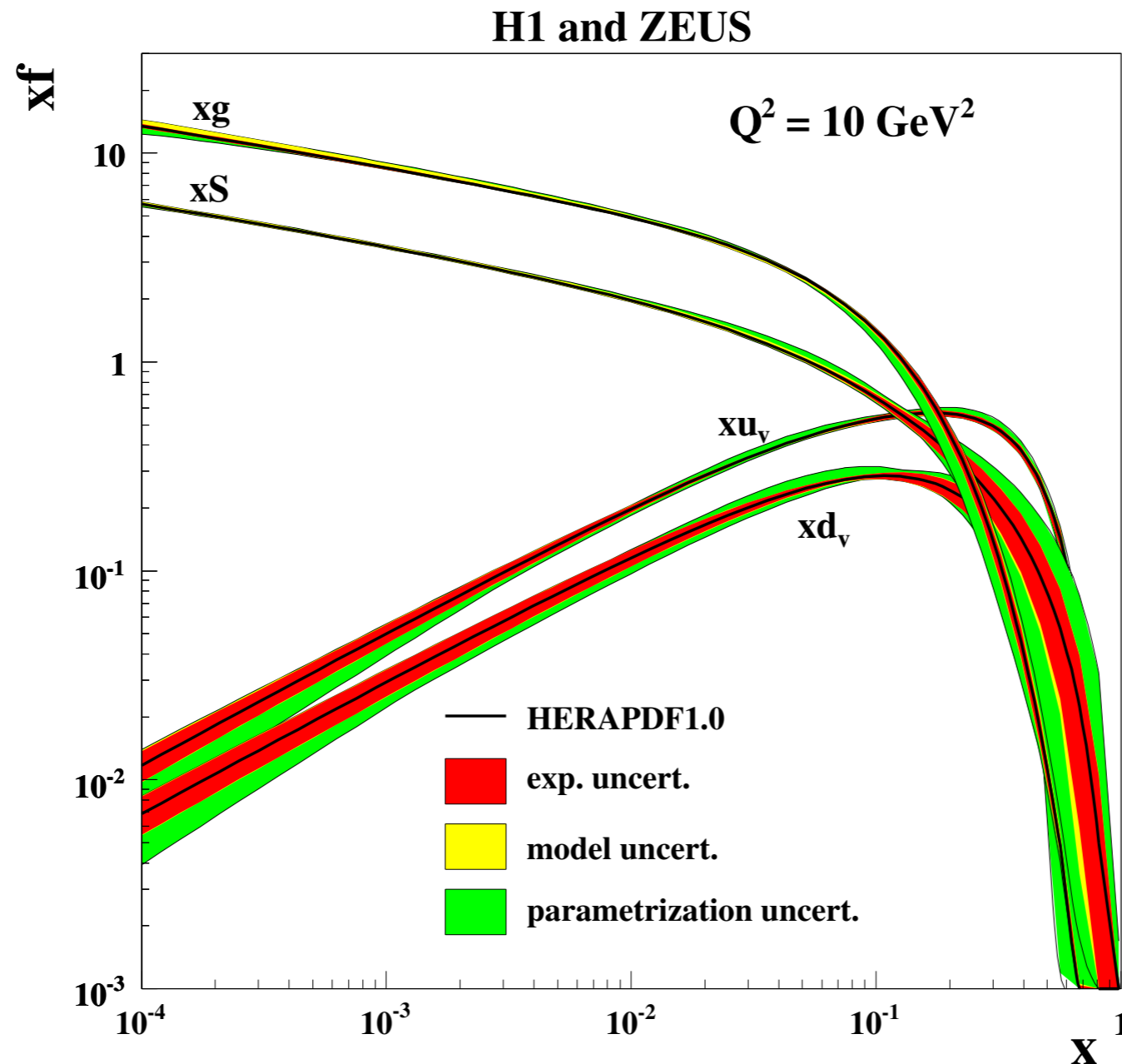
$$\langle T^{\mu\nu}(x_{\perp}) T^{\mu\nu}(y_{\perp}) \rangle$$

In the classical approximation (MV model)

Initial conditions: the Color-Glass Condensate

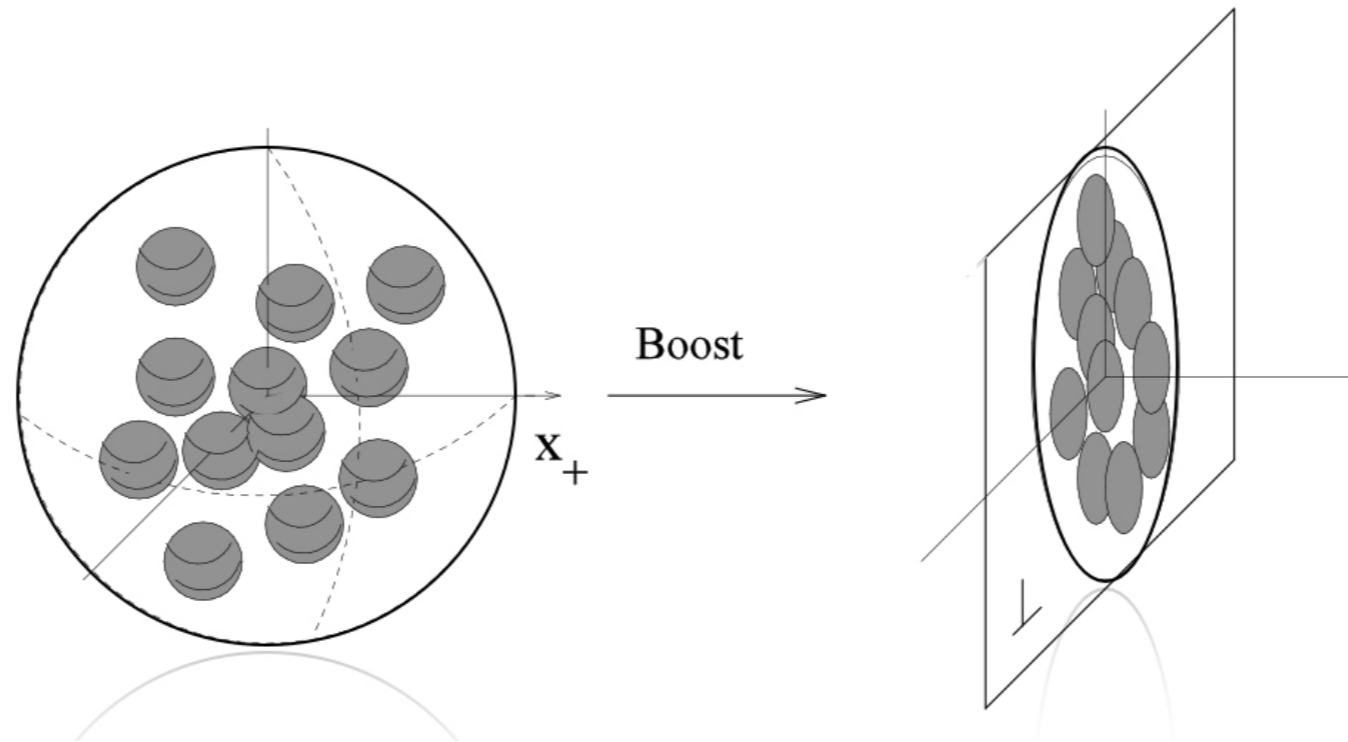
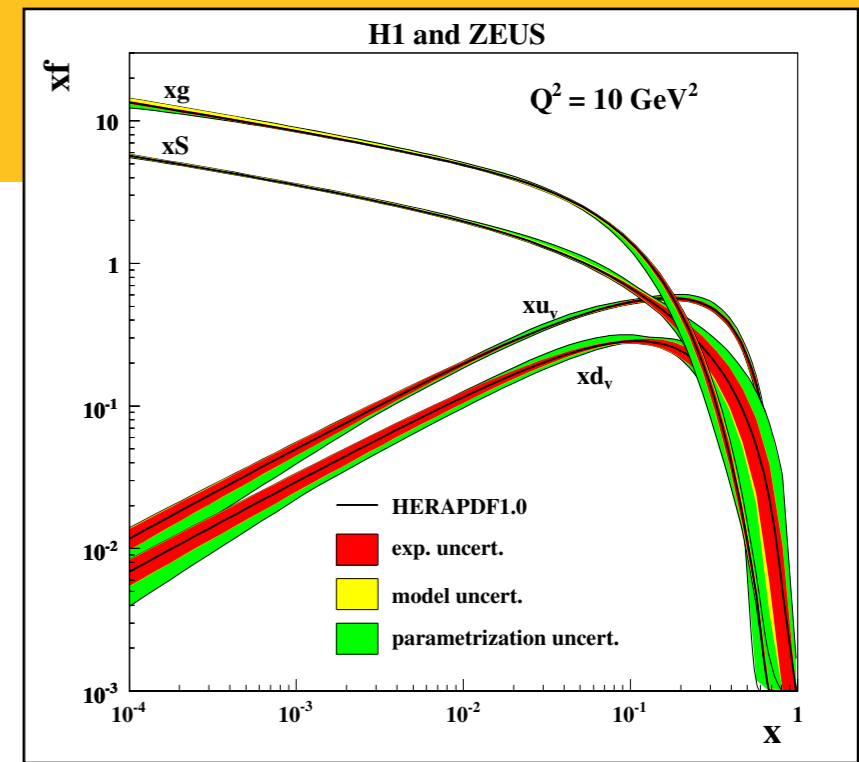
Highly Energetic Heavy Ion Collisions

- At high energies (or equivalently, low x) the partonic content of protons and neutrons is **vastly dominated by a high density of gluons**



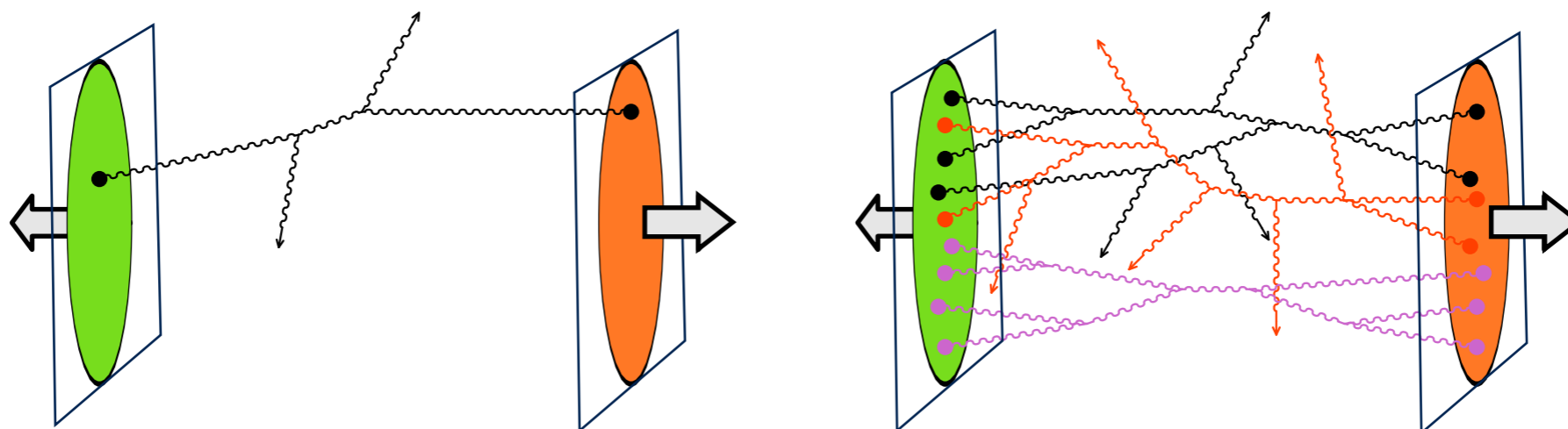
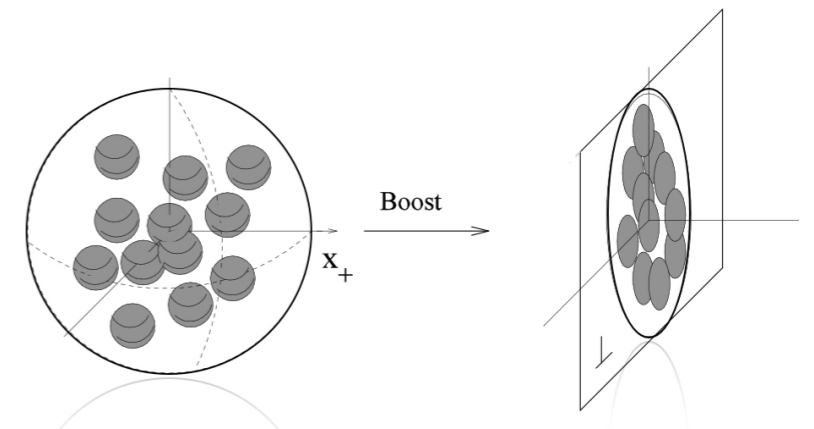
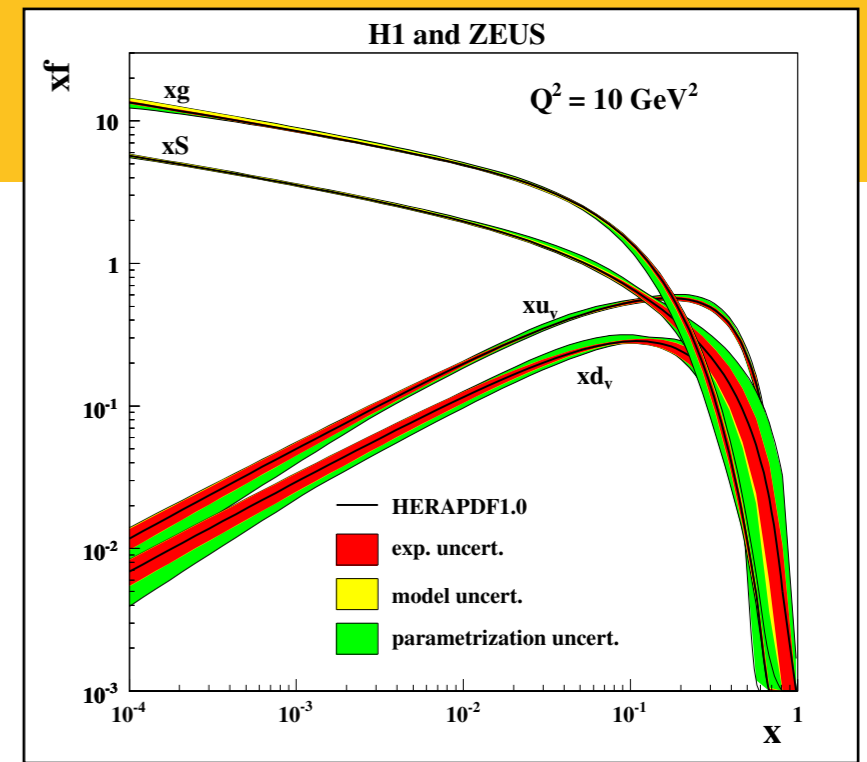
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- Relativistic kinematics: at high energies, the nuclei appear almost two-dimensional in the laboratory frame due to **Lorentz contraction**



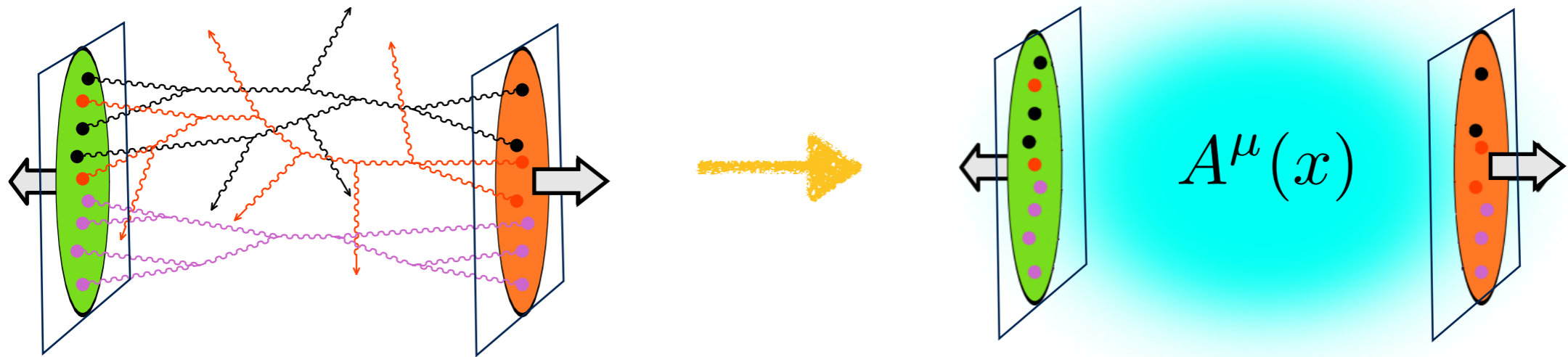
Highly Energetic Heavy Ion Collisions

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- Relativistic kinematics: at high energies, the nuclei appear almost two-dimensional in the laboratory frame due to **Lorentz contraction**
- QCD becomes **non-linear** and **non-perturbative!**



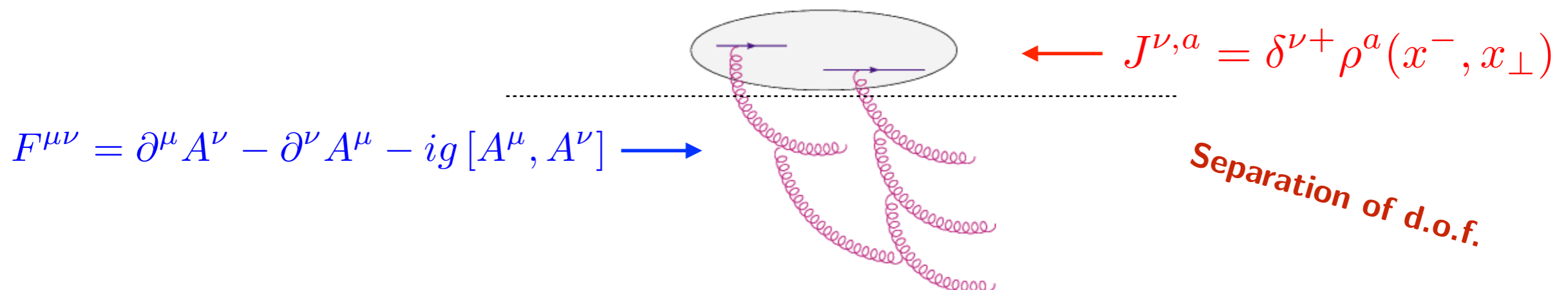
Color Glass Condensate: McLerran-Venugopalan model

- We use an approximation of QCD for high gluon densities where we replace the gluons with a **classical field** generated by the valence quarks



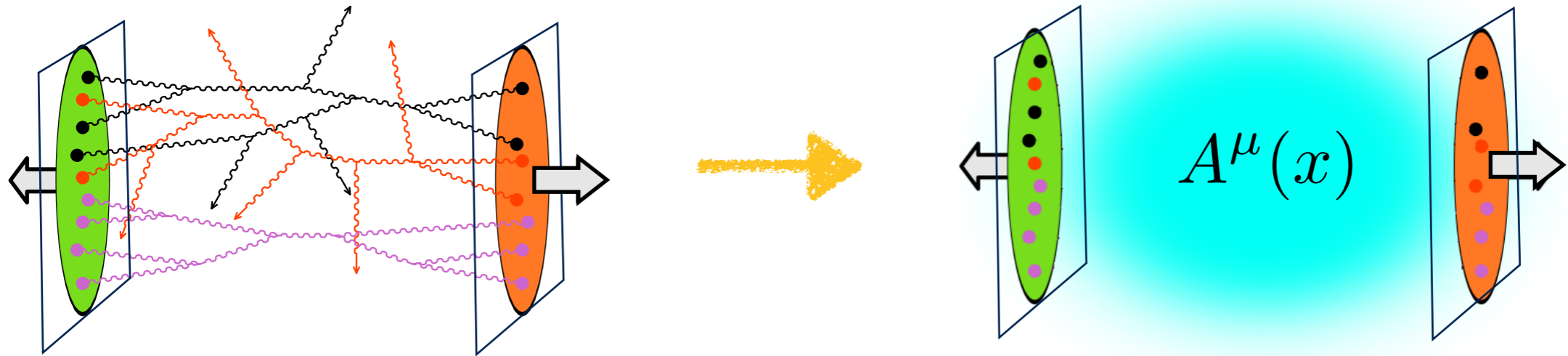
- Dynamics of the field described by **Yang-Mills** classical equations:

$$[D_\mu, F^{\mu\nu}] = J^\nu \propto \rho(x)$$



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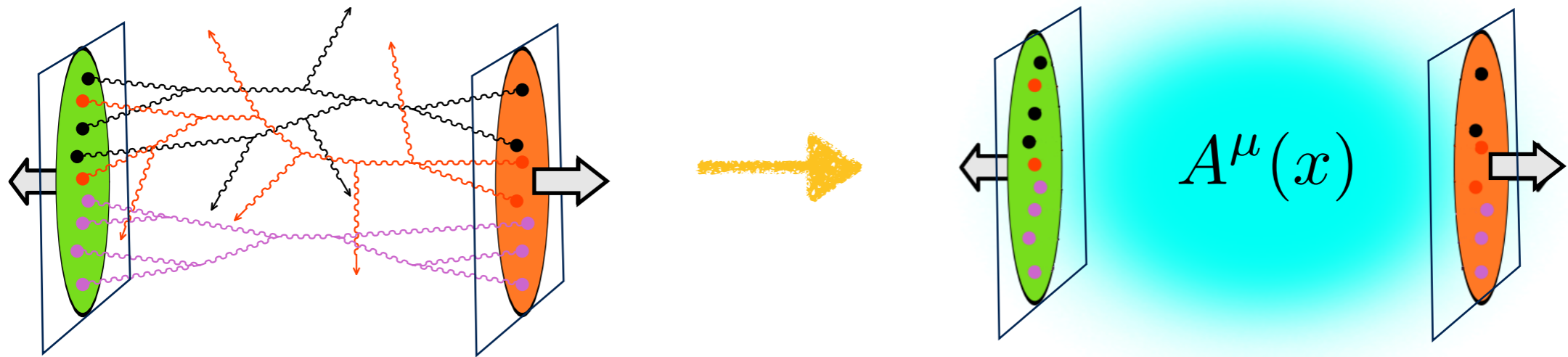
$$[D_\mu, F^{\mu\nu}] = J^\nu \propto \rho(x)$$

- Calculation of observables: **average** over background classical fields

$$\langle \mathcal{O}[\rho] \rangle = \int [d\rho] \exp \left\{ - \int dx \text{Tr} [\rho^2] \right\} \mathcal{O}[\rho]$$

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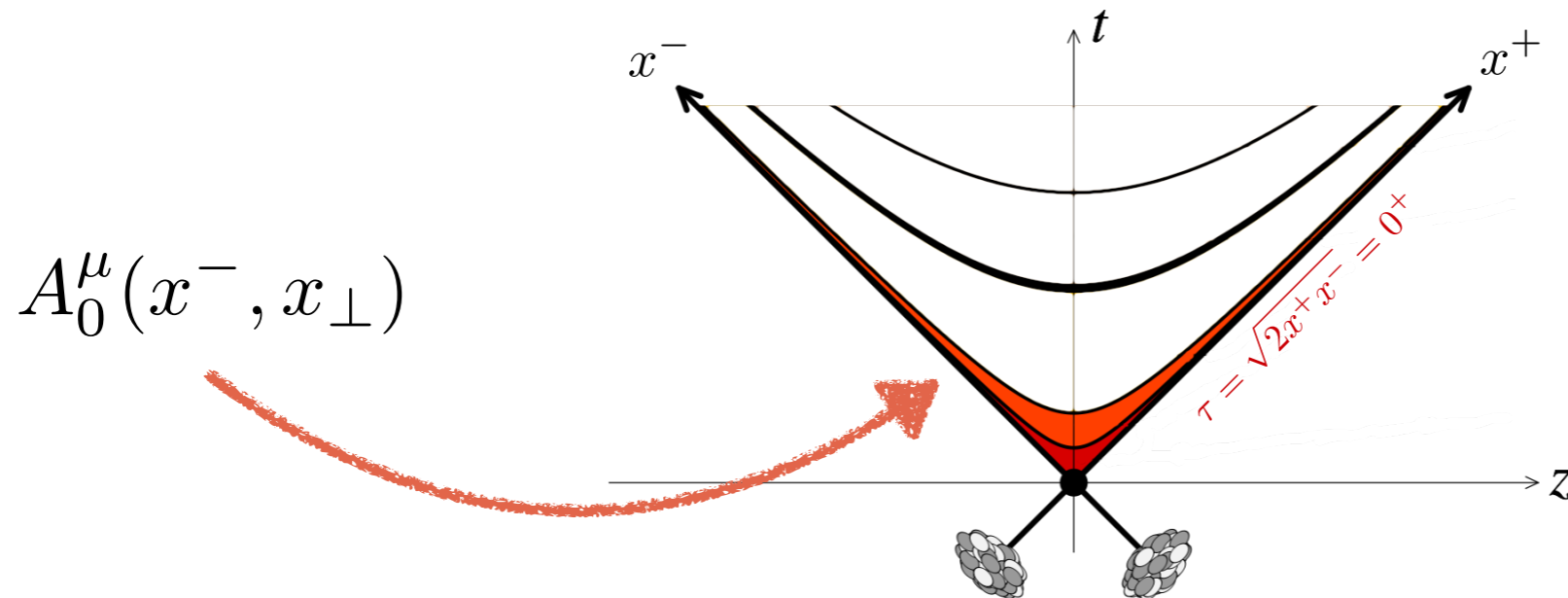
$$[D_\mu, F^{\mu\nu}] = J^\nu \propto \rho(x)$$

- Calculation of observables: **average** over background classical fields
- Basic building block: **2-point correlator (McLerran-Venugopalan)**

$$\langle \rho^a(x^-, x_\perp) \rho^b(y^-, y_\perp) \rangle = \mu^2(x^-) \delta^{ab} \delta(x^- - y^-) \delta^{(2)}(x_\perp - y_\perp)$$

Steps for the calculation

1) Calculate the gluon fields at early times in a HIC



2) Build the energy-momentum tensor

$$T_0^{\mu\nu}(x_\perp) = 2 \text{Tr} \left\{ \frac{1}{4} g^{\mu\nu} F^{\alpha\beta} F_{\alpha\beta} - F^{\mu\alpha} F^\nu{}_\alpha \right\}_0$$

3) Average over the color source distributions

$$\langle T_0^{\mu\nu}(x_\perp) \rangle = \int [d\rho_1] W_1[\rho_1] [d\rho_2] W_2[\rho_2] T_0^{\mu\nu}(x_\perp) [\rho_1, \rho_2]$$

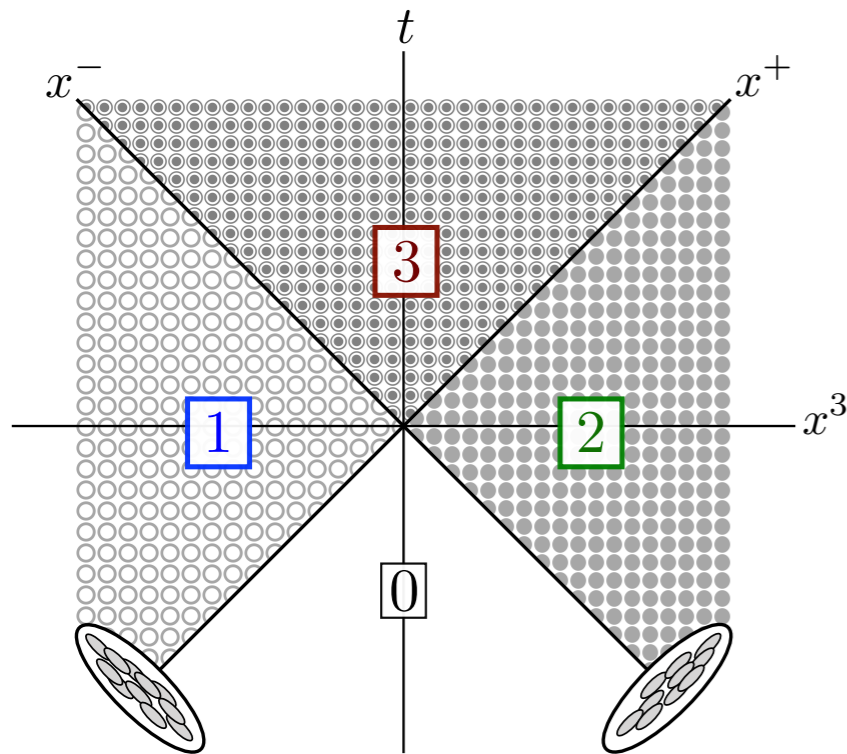
$$\langle T_0^{\mu\nu}(x_\perp) T_0^{\sigma\gamma}(y_\perp) \rangle = \int [d\rho_1] W_1[\rho_1] [d\rho_2] W_2[\rho_2] T_0^{\mu\nu}(x_\perp) T_0^{\sigma\gamma}(y_\perp) [\rho_1, \rho_2]$$

Calculation of the gluon fields

$$[D_\mu, F^{\mu\nu}] = J_1^\nu + J_2^\nu$$

$$J_1^\nu = \rho_1(\mathbf{x}_\perp) \delta(\mathbf{x}^-) \delta^{\nu+}$$

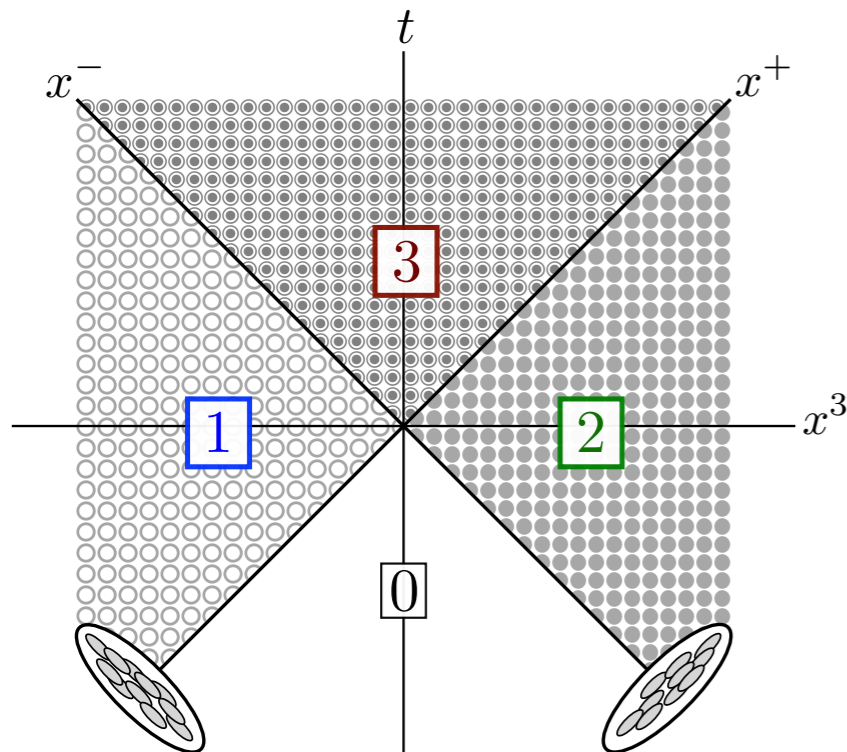
$$J_2^\nu = \rho_2(\mathbf{x}_\perp) \delta(\mathbf{x}^+) \delta^{\nu-}$$



$$[D_\mu, F^{\mu\nu}] = J_1^\nu + J_2^\nu$$

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[1, 2] Single nucleus solution

$$A_1^\pm = 0$$

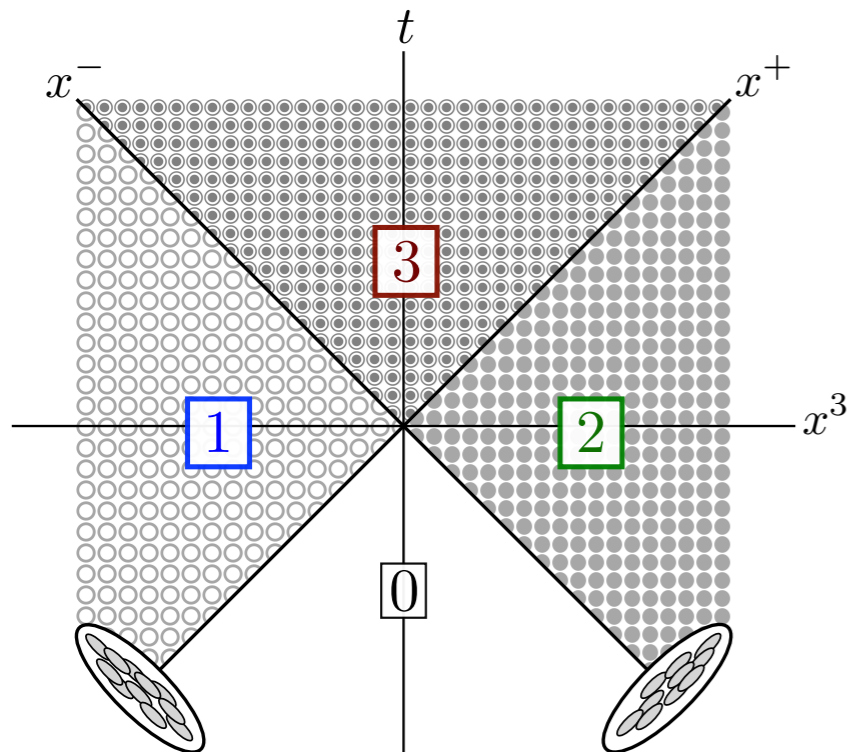
$$A_1^i = \theta(x^-) \int_{-\infty}^{\infty} dz^- U_1^\dagger(z^-, x_\perp) \frac{\partial^i \tilde{\rho}_1(z^-, x_\perp)}{\nabla^2} U_1(z^-, x_\perp) \equiv \theta(x^-) \alpha_1^i(x_\perp)$$

$$U_1(x^-, x_\perp) = \text{P}^- \exp \left\{ -ig \int_{x_0^-}^{x^-} dz^- \frac{1}{\nabla^2} \tilde{\rho}_1(z^-, x_\perp) \right\}$$

$$[D_\mu, F^{\mu\nu}] = J_1^\nu + J_2^\nu$$

$$J_1^\nu = \rho_1(\mathbf{x}_\perp) \delta(\mathbf{x}^-) \delta^{\nu+}$$

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[1, 2] Single nucleus solution

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[3] Forward light cone $\tau = 0^+$

$$A^\pm = \pm x^\pm \alpha(\tau = 0^+, x_\perp)$$

$$A^i = \alpha^i(\tau = 0^+, x_\perp)$$

$$\alpha^i(\tau = 0^+, x_\perp) = \alpha_1^i(x_\perp) + \alpha_2^i(x_\perp)$$

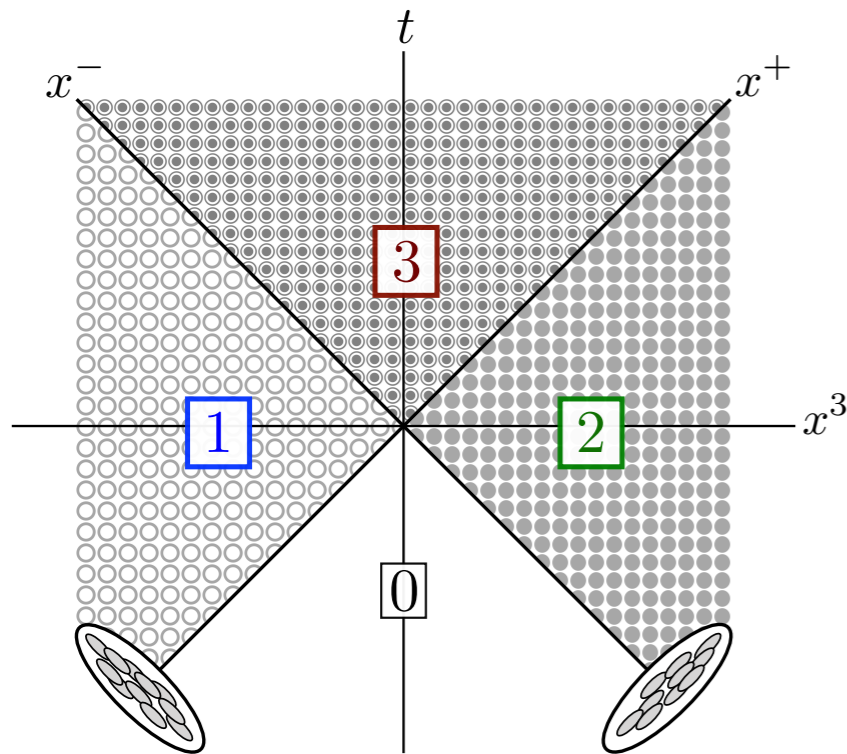
$$\alpha(\tau = 0^+, x_\perp) = \frac{ig}{2} [\alpha_1^i(x_\perp), \alpha_2^i(x_\perp)]$$

Calculation of the energy-momentum tensor $T^{\mu\nu}(\tau = 0^+)$

$$[D_\mu, F^{\mu\nu}] = J_1^\nu + J_2^\nu$$

$$J_1^\nu = \rho_1(\mathbf{x}_\perp) \delta(\mathbf{x}^-) \delta^{\nu+}$$

$$J_2^\nu = \rho_2(\mathbf{x}_\perp) \delta(\mathbf{x}^+) \delta^{\nu-}$$



- We can obtain the early-time energy-momentum tensor as:

[1, 2] Single nucleus solution

$$A_1^\pm = 0$$

$$A_1^i = \theta(x^-) \int_{-\infty}^{\infty} dz^- U_1^\dagger(z^-, x_\perp) \frac{\partial^i \tilde{\rho}_1(z^-, x_\perp)}{\nabla^2} U_1(z^-, x_\perp) \equiv \theta(x^-) \alpha_1^i(x_\perp)$$

$$U_1(x^-, x_\perp) = P^- \exp \left\{ -ig \int_{x_0^-}^{x^-} dz^- \frac{1}{\nabla^2} \tilde{\rho}_1(z^-, x_\perp) \right\}$$

[3] Forward light cone $\tau = 0^+$

$$A^\pm = \pm x^\pm \alpha(\tau = 0^+, x_\perp) \quad \alpha^i(\tau = 0^+, x_\perp) = \alpha_1^i(x_\perp) + \alpha_2^i(x_\perp)$$

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$$\begin{aligned} T_0^{\mu\nu} &= \frac{1}{4} g^{\mu\nu} F^{\alpha\beta,a} F_{\alpha\beta}^a - F^{\mu\alpha,a} F_{\alpha}^{\nu,a} \\ &= -\frac{g^2}{2} (\delta^{ij} \delta^{kl} + \epsilon^{ij} \epsilon^{kl}) \left([\alpha_1^i, \alpha_2^j] [\alpha_1^k, \alpha_2^l] \right) \times \text{diag}(1, 1, 1, -1) \\ &\equiv \epsilon_0 \times \text{diag}(1, 1, 1, -1) \equiv \epsilon_0 \times t^{\mu\nu} \end{aligned}$$

Correlators of the energy-momentum tensor at $\tau = 0^+$

$$\langle T^{\mu\nu}(x_\perp) \rangle = \langle \epsilon_0 \rangle t^{\mu\nu}$$

- For the 1-point correlator of $T^{\mu\nu}$:

$$\begin{aligned} \langle \epsilon_0 \rangle &= -g^2 (\delta^{ij} \delta^{kl} + \epsilon^{ij} \epsilon^{kl}) \left\langle \text{Tr} \left\{ [\alpha_1^i, \alpha_2^j] [\alpha_1^k, \alpha_2^l] \right\} \right\rangle \\ &= -g^2 (\delta^{ij} \delta^{kl} + \epsilon^{ij} \epsilon^{kl}) \left\langle \alpha_1^{i,a} \alpha_2^{j,b} \alpha_1^{k,c} \alpha_2^{l,d} \right\rangle \text{Tr} \left\{ [t^a, t^b] [t^c, t^d] \right\} \\ &= \frac{g^2}{2} (\delta^{ij} \delta^{kl} + \epsilon^{ij} \epsilon^{kl}) f^{abm} f^{cdm} \left\langle \alpha_1^{i,a}(x_\perp) \alpha_1^{k,c}(x_\perp) \right\rangle_1 \left\langle \alpha_2^{j,b}(x_\perp) \alpha_2^{l,d}(x_\perp) \right\rangle_2 \end{aligned}$$

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- We momentarily take two different transverse coordinates:

$$\langle \alpha^{i,a}(x_\perp) \alpha^{j,b}(y_\perp) \rangle = \int_{-\infty}^{\infty} dz^- dz'^- \left\langle \frac{\partial^i \tilde{\rho}^{a'}(z^-, x_\perp)}{\nabla^2} \underbrace{U^{a'a}(z^-, x_\perp)}_{\sim e^{i\rho}} \frac{\partial^j \tilde{\rho}^{b'}(z'^-, y_\perp)}{\nabla^2} \underbrace{U^{b'b}(z'^-, y_\perp)}_{\sim e^{i\rho}} \right\rangle$$

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Luckily, Wilson lines and (external) color source densities factorize

$$\langle T^{\mu\nu}(x_\perp) \rangle = \langle \epsilon_0 \rangle t^{\mu\nu}$$

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Where:

$$L(x_\perp - y_\perp) = \int d^2 z_\perp G(x_\perp - z_\perp) G(y_\perp - z_\perp).$$

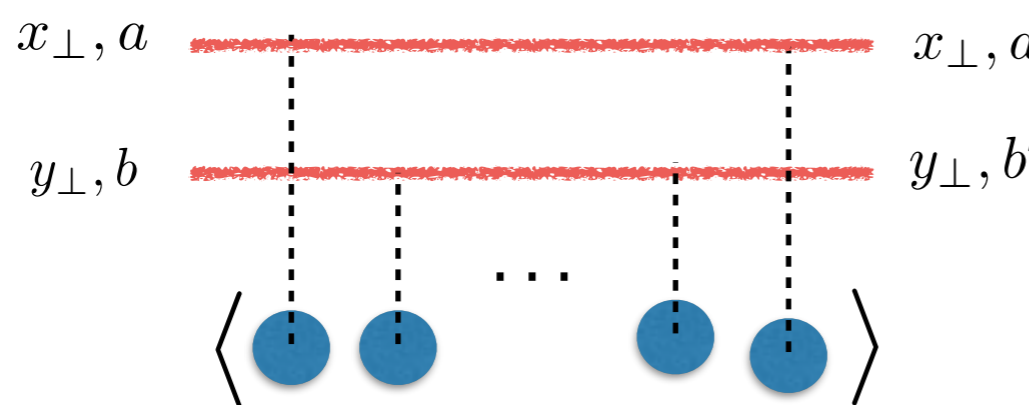
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- For the 1-point correlator of $T^{\mu\nu}$:

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$$\frac{\delta^{ab} \delta^{a'b'}}{N} \exp \left[-g^2 \frac{N}{2} \Gamma(x_\perp, y_\perp) \bar{\mu}^2(x^-) \right]$$

$$\langle T^{\mu\nu}(x_\perp) \rangle = \langle \epsilon_0 \rangle t^{\mu\nu}$$

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Notation:

$$\bar{\mu}^2 = \int_{-\infty}^{\infty} dz^- \mu^2(z^-)$$

- Here we have introduced a **momentum scale** characterizing each nucleus:

$$\bar{Q}_s^2 = \alpha_s N_c \bar{\mu}^2(x_\perp)$$

- In the MV model the factor $\partial^2 L(0_\perp)$ yields a **logarithmic IR divergence**.

$$\langle T^{\mu\nu}(x_{\perp})T^{\sigma\rho}(y_{\perp}) \rangle = \langle \epsilon(x_{\perp})\epsilon(y_{\perp}) \rangle t^{\mu\nu} t^{\sigma\rho}$$

- For the 2-point correlator of $T^{\mu\nu}$: prepare for trouble and make it double

$$\begin{aligned} \langle \epsilon(x_{\perp})\epsilon(y_{\perp}) \rangle = & \frac{g^4}{4} (\delta^{ij}\delta^{kl} + \epsilon^{ij}\epsilon^{kl}) (\delta^{i'j'}\delta^{k'l'} + \epsilon^{i'j'}\epsilon^{k'l'}) f^{abn} f^{cdn} f^{a'b'm} f^{c'd'm} \\ & \times \langle \alpha_{1x}^{ia} \alpha_{1x}^{kc} \alpha_{1y}^{i'a'} \alpha_{1y}^{k'c'} \rangle \langle \alpha_{2x}^{jb} \alpha_{2x}^{ld} \alpha_{2y}^{j'b'} \alpha_{2y}^{l'd'} \rangle \end{aligned}$$

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- The building block:

$$\langle \alpha^{ia}(x_{\perp})\alpha^{kc}(x_{\perp})\alpha^{i'a'}(y_{\perp})\alpha^{k'c'}(y_{\perp}) \rangle = \int_{-\infty}^{\infty} dz^{-} dw^{-} dz^{-'} dw^{-'} \left\langle \frac{\partial^i \tilde{\rho}^e(z^{-}, x_{\perp})}{\nabla^2} U^{ea}(z^{-}, x_{\perp}) \right. \\ \left. \frac{\partial^k \tilde{\rho}^f(w^{-}, x_{\perp})}{\nabla^2} U^{fc}(w^{-}, x_{\perp}) \frac{\partial^{i'} \tilde{\rho}^{e'}(z^{-'}, y_{\perp})}{\nabla^2} U^{e'a'}(z^{-'}, y_{\perp}) \frac{\partial^{k'} \tilde{\rho}^{f'}(w^{-'}, y_{\perp})}{\nabla^2} U^{f'c'}(w^{-'}, y_{\perp}) \right\rangle.$$

$$\langle T^{\mu\nu}(x_{\perp})T^{\sigma\rho}(y_{\perp}) \rangle = \langle \epsilon(x_{\perp})\epsilon(y_{\perp}) \rangle t^{\mu\nu} t^{\sigma\rho}$$

- For the 2-point correlator of $T^{\mu\nu}$: prepare for trouble and make it double

$$\langle \epsilon(x_{\perp})\epsilon(y_{\perp}) \rangle = \frac{g^4}{4} (\delta^{ij}\delta^{kl} + \epsilon^{ij}\epsilon^{kl}) (\delta^{i'j'}\delta^{k'l'} + \epsilon^{i'j'}\epsilon^{k'l'}) f^{abn} f^{cdn} f^{a'b'm} f^{c'd'm}$$

$$\times \langle \alpha_{1x}^{ia} \alpha_{1x}^{kc} \alpha_{1y}^{i'a'} \alpha_{1y}^{k'c'} \rangle \langle \alpha_{2x}^{jb} \alpha_{2x}^{ld} \alpha_{2y}^{j'b'} \alpha_{2y}^{l'd'} \rangle$$

- Technical difficulties:

- The **expansion of the correlator** $\langle \alpha^{ia}(x_{\perp})\alpha^{kc}(x_{\perp})\alpha^{i'a'}(y_{\perp})\alpha^{k'c'}(y_{\perp}) \rangle$ is far more difficult than that of $\langle \alpha^{ia}(x_{\perp})\alpha^{kc}(y_{\perp}) \rangle$. Schematically: *[Fillion-Gourdeau & Jeon '09]*

$$\langle \alpha^{ia}(x_{\perp})\alpha^{kc}(x_{\perp})\alpha^{i'a'}(y_{\perp})\alpha^{k'c'}(y_{\perp}) \rangle = \underbrace{\langle \rho^4 \rangle \langle U^4 \rangle}_{3 \text{ terms}} + \underbrace{\langle \rho^2 \rangle \langle \rho^2 U^4 \rangle}_c_{4 \text{ terms}}$$

(Wick's theorem)

$$\langle T^{\mu\nu}(x_{\perp})T^{\sigma\rho}(y_{\perp}) \rangle = \langle \epsilon(x_{\perp})\epsilon(y_{\perp}) \rangle t^{\mu\nu} t^{\sigma\rho}$$

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$$\langle \alpha^{ia}(x_{\perp})\alpha^{kc}(x_{\perp})\alpha^{i'a'}(y_{\perp})\alpha^{k'c'}(y_{\perp}) \rangle = \langle \rho^4 \rangle \langle U^4 \rangle + \langle \rho^2 \rangle \langle \rho^2 U^4 \rangle_c$$

- Instead of having to calculate the adjoint Wilson line dipole, we need the much more complex adjoint **Wilson line quadrupole** *[Kovner & Wiedemann '01]*

$$\langle U^{ab}(z^-, x_{\perp})U^{cd}(z^-, y_{\perp})U^{ef}(z^-, x'_{\perp})U^{gh}(z^-, y'_{\perp}) \rangle$$

$$\langle T^{\mu\nu}(x_{\perp})T^{\sigma\rho}(y_{\perp}) \rangle = \langle \epsilon(x_{\perp})\epsilon(y_{\perp}) \rangle t^{\mu\nu} t^{\sigma\rho}$$

- For the 2-point correlator of $T^{\mu\nu}$: prepare for trouble and make it double

$$\langle \epsilon(x_{\perp})\epsilon(y_{\perp}) \rangle = \frac{g^4}{4} (\delta^{ij}\delta^{kl} + \epsilon^{ij}\epsilon^{kl}) (\delta^{i'j'}\delta^{k'l'} + \epsilon^{i'j'}\epsilon^{k'l'}) f^{abn} f^{cdn} f^{a'b'm} f^{c'd'm} \\ \times \langle \alpha_{1x}^{ia} \alpha_{1x}^{kc} \alpha_{1y}^{i'a'} \alpha_{1y}^{k'c'} \rangle \langle \alpha_{2x}^{jb} \alpha_{2x}^{ld} \alpha_{2y}^{j'b'} \alpha_{2y}^{l'd'} \rangle$$

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- The **expansion of the correlator** $\langle \alpha^{ia}(x_{\perp})\alpha^{kc}(x_{\perp})\alpha^{i'a'}(y_{\perp})\alpha^{k'c'}(y_{\perp}) \rangle$ is far more difficult than that of $\langle \alpha^{ia}(x_{\perp})\alpha^{kc}(y_{\perp}) \rangle$. Schematically: *[Fillion-Gourdeau & Jeon '09]*

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$$\langle U^{ab}(z^-, x_{\perp})U^{cd}(z^-, y_{\perp})U^{ef}(z^-, x'_{\perp})U^{gh}(z^-, y'_{\perp}) \rangle$$

- The **color structure** of this object is frustratingly complex. Even with all parts analytically calculated, the contraction of the color indices demands a computational treatment (via FeynCalc)

$$\text{Cov}[\epsilon_0](x_\perp, y_\perp) = \langle \epsilon_0(x_\perp) \epsilon_0(y_\perp) \rangle - \langle \epsilon_0(x_\perp) \rangle \langle \epsilon_0(y_\perp) \rangle$$

$$\begin{aligned} \text{Cov}[\epsilon](\tau=0^+; x_\perp, y_\perp) = & -\frac{\partial_x^i \Gamma \partial_y^i \Gamma (N_c^2 - 1) A (4A^2 - B^2)}{16 N_c^2 \Gamma^5 g^4} (f_1 g_2 + f_2 g_1) \\ & + \frac{(N_c^2 - 1)(16A^4 + B^4)}{2 N_c^2 \Gamma^4 g^4} f_1 f_2 + \frac{(\partial_x^i \Gamma \partial_y^i \Gamma)^2 (N_c^2 - 1) A^2}{64 N_c^2 \Gamma^6 g^4} g_1 g_2 \\ & + \frac{(4A^2 + B^2)^2 r^2}{N_c^2 \Gamma^4 g^4} \left(\left[\frac{1}{2} Q_{s1}^2 Q_{s2}^2 r^2 + 4 Q_{s2}^2 e^{-\frac{Q_{s1}^2 r^2}{4}} - 4 Q_{s1}^2 \right] + [1 \leftrightarrow 2] \right) \\ & + \frac{(N_c^2 - 1)(4A^2 + B^2)}{2 N_c^2 \Gamma^2 g^4} (4\pi \partial^2 L(0_\perp))^2 \left(\left[\bar{Q}_{s1}^4 (Q_{s2}^2 r^2 - 4 + 4e^{-\frac{Q_{s2}^2 r^2}{4}}) \right] + [1 \leftrightarrow 2] \right) \\ & + \frac{(4A^2 + B^2)^2}{\Gamma^4 g^4 N_c^2 (N_c^2 - 1)^2 (N_c^2 - 4)^2} \left(\left[-4(N_c^2 - 1)(N_c^2 - 4)(N_c^6 - 3N_c^4 - 26N_c^2 + 16)e^{-\frac{Q_{s1}^2 r^2}{4}} \right. \right. \\ & + (N_c - 3)(N_c + 1)^3 (N_c + 2)^2 N_c^3 \left((N - 2)e^{\frac{Q_{s1}^2 r^2}{4}} - 2(N - 1) \right) e^{-\frac{r^2(N_c Q_{s1}^2 + 2(N_c - 1)Q_{s2}^2)}{4N_c}} \\ & + (N_c + 3)(N_c - 1)^3 (N_c - 2)^2 N_c^3 \left((N + 2)e^{\frac{Q_{s1}^2 r^2}{4}} - 2(N + 1) \right) e^{-\frac{r^2(N_c Q_{s1}^2 + 2(N_c + 1)Q_{s2}^2)}{4N_c}} \\ & \left. \left. + 4(N_c^2 - 8)(N_c^2 - 1)^3 (N_c^2 + 4)e^{-\frac{1}{4}r^2(Q_{s1}^2 + Q_{s2}^2)} \right. \right. \\ & \left. \left. + \frac{1}{2}(N_c - 2)^2 (N_c - 1)^3 (N_c + 3) N_c^4 e^{-\frac{(N_c + 1)r^2(Q_{s1}^2 + Q_{s2}^2)}{2N_c}} \right. \right. \\ & \left. \left. + \frac{1}{2}(N_c - 3)(N_c + 1)^3 (N_c + 2)^2 N_c^4 e^{-\frac{(N_c - 1)r^2(Q_{s1}^2 + Q_{s2}^2)}{2N_c}} \right] + [1 \leftrightarrow 2] \right) \\ & \left. + 2(N_c^2 - 4)^2 (N_c^6 + 2N_c^4 - 19N_c^2 + 8) \right) \end{aligned}$$

with:

$$f_{1,2} = e^{-\frac{Q_{s1,2}^2 r^2}{4}} (Q_{s1,2}^2 r^2 + 4) - 4$$

$$g_{1,2} = e^{-\frac{Q_{s1,2}^2 r^2}{4}} (Q_{s1,2}^4 r^4 + 8Q_{s1,2}^2 r^2 + 32) - 32.$$

Pocket formulae

- Omitting for the moment the issues with the $r \rightarrow 0$ divergencies (GBW-model)

$r \rightarrow 0$

$$\lim_{r \rightarrow 0} \text{Cov}[\epsilon](0^+; x_\perp, y_\perp) = \frac{3C_F}{g^4 2N_c} Q_{s1}^4 Q_{s2}^4$$

$$\lim_{r \rightarrow 0} \frac{\text{Cov}[\epsilon](0^+; x_\perp, y_\perp)}{\langle \epsilon_0(x_\perp) \rangle \langle \epsilon_0(y_\perp) \rangle} = \frac{3}{(N_c^2 - 1)}$$

$r \rightarrow \infty$

$$\lim_{r \rightarrow \infty} \text{Cov}[\epsilon](0^+; x_\perp, y_\perp) = \frac{2(N_c^2 - 1) (Q_{s1}^4 Q_{s2}^2 + Q_{s1}^2 Q_{s2}^4)}{g^4 N_c^2 r^2}$$

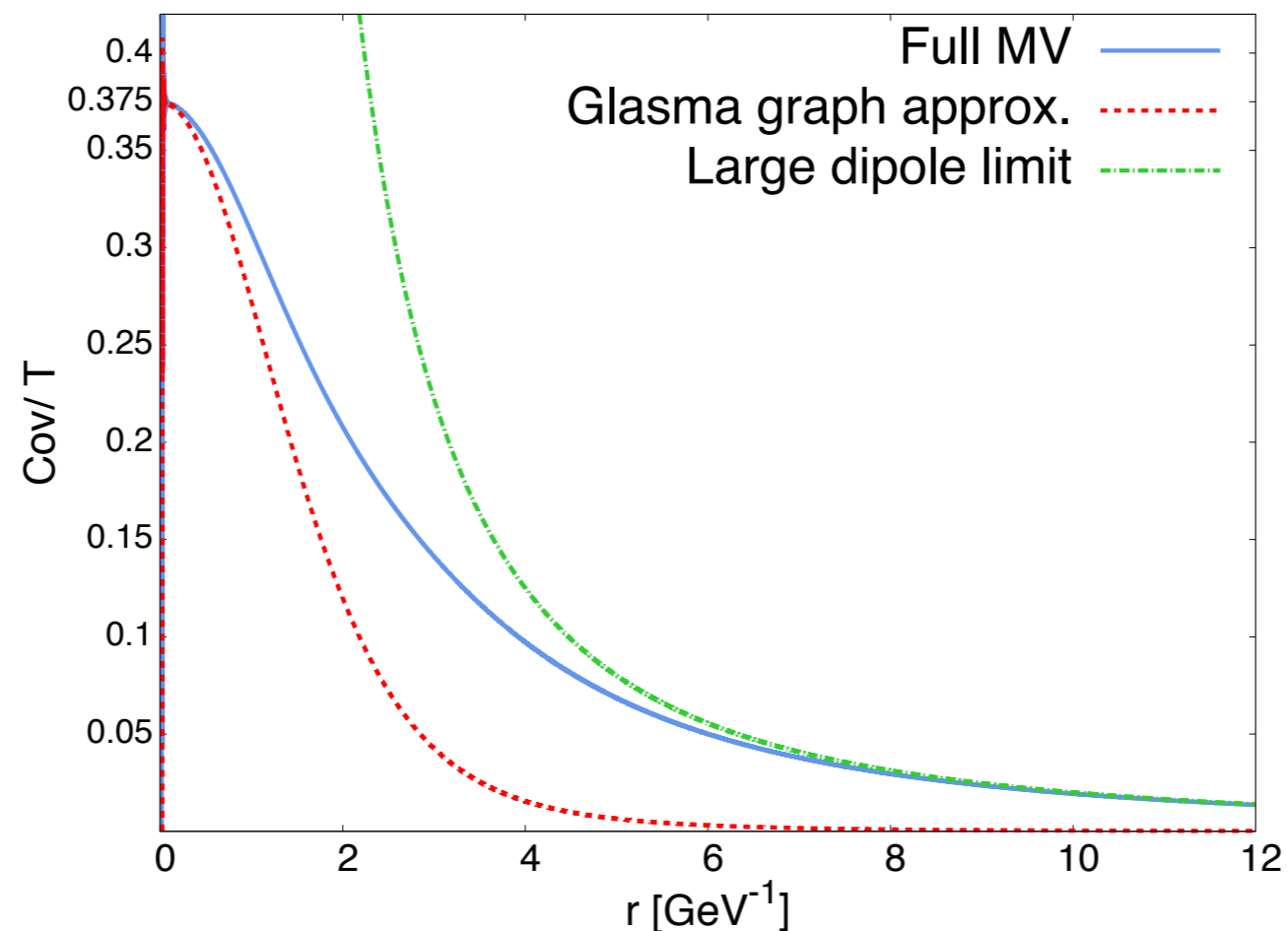
$$\lim_{r \rightarrow \infty} \frac{\text{Cov}[\epsilon](0^+; x_\perp, y_\perp)}{\langle \epsilon_0(x_\perp) \rangle \langle \epsilon_0(y_\perp) \rangle} = \frac{1}{2(N_c^2 - 1)r^2} \left(\frac{1}{Q_{s1}^2} + \frac{1}{Q_{s2}^2} \right)$$

Comparison with the 'Glasma Graph' approximation

- Glasma Graph approximation [*Lappi & Schlichting 2018, Muller & Schaefer 2012*]. Assume Gaussian distribution of the produced gluon fields:

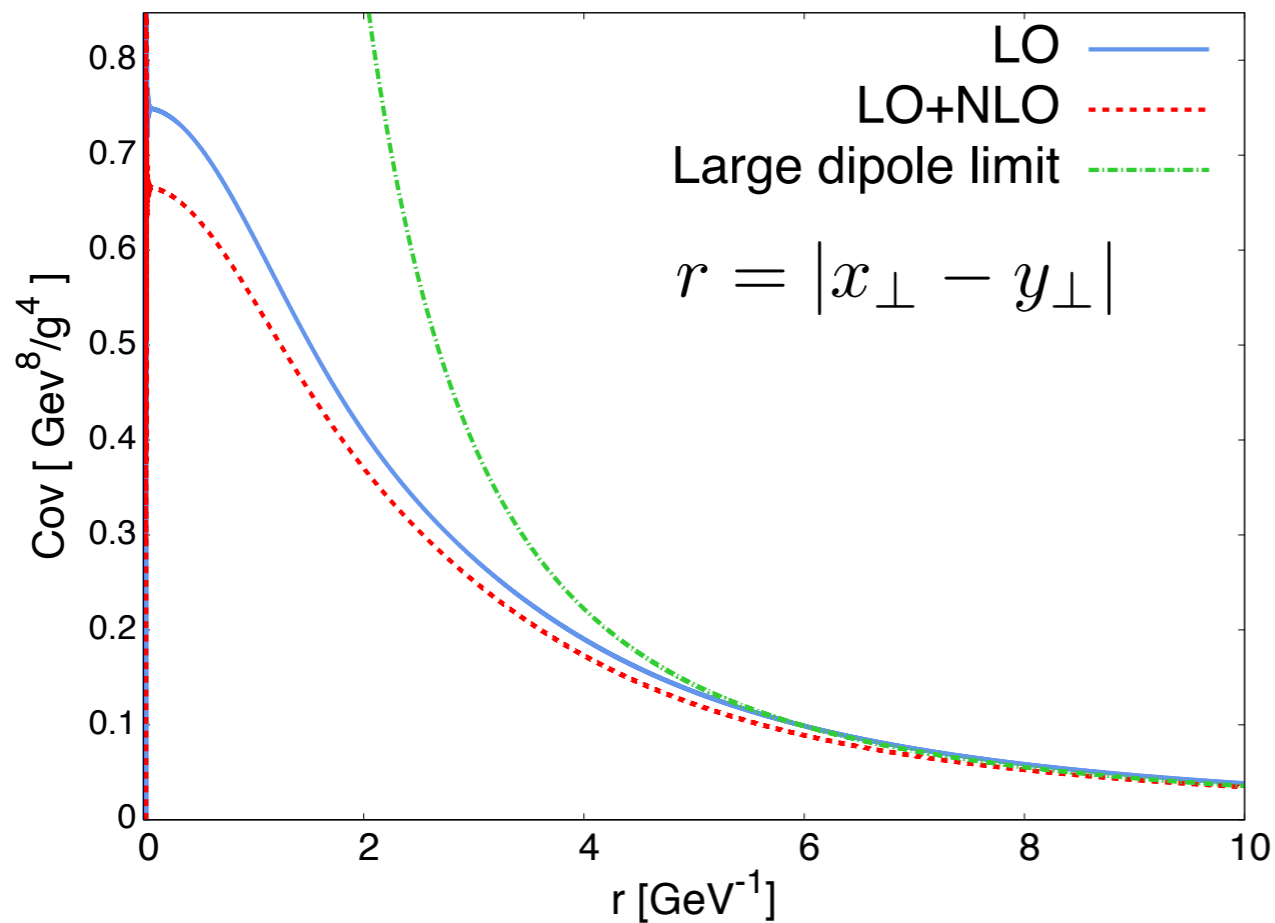
$$\begin{aligned} \langle \alpha^{ia}(x_\perp) \alpha^{kc}(x_\perp) \alpha^{i'a'}(y_\perp) \alpha^{k'c'}(y_\perp) \rangle_{\text{GG}} &= \langle \alpha^{ia}(x_\perp) \alpha^{kc}(x_\perp) \rangle \langle \alpha^{i'a'}(y_\perp) \alpha^{k'c'}(y_\perp) \rangle \\ &+ \langle \alpha^{ia}(x_\perp) \alpha^{i'a'}(y_\perp) \rangle \langle \alpha^{kc}(x_\perp) \alpha^{k'c'}(y_\perp) \rangle \\ &+ \langle \alpha^{ia}(x_\perp) \alpha^{k'c'}(y_\perp) \rangle \langle \alpha^{kc}(x_\perp) \alpha^{i'a'}(y_\perp) \rangle. \end{aligned}$$

- Agreement with full result in the $r \rightarrow 0$ limit. **Strong discrepancies** in the $r \rightarrow \infty$ limit

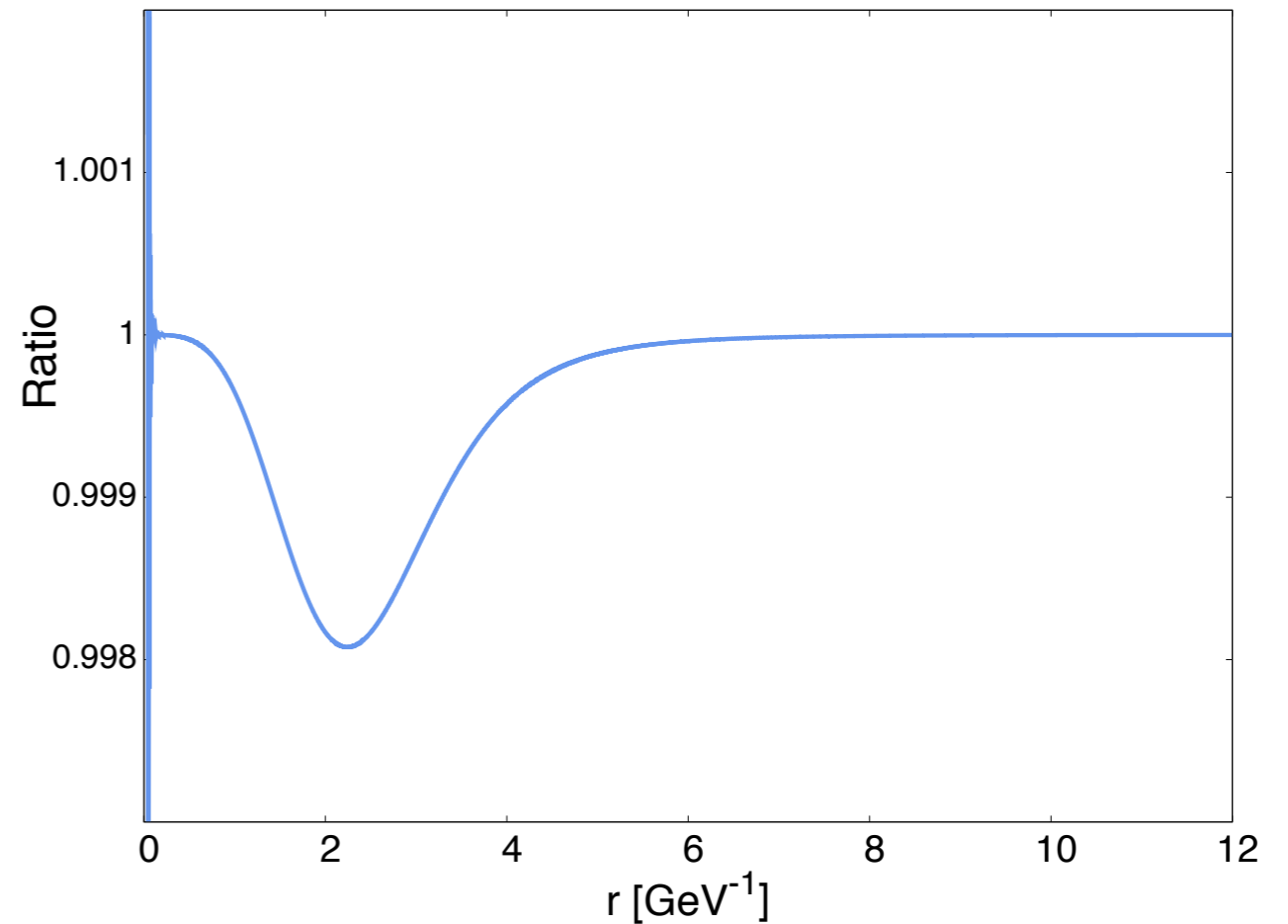


N_c expansion

First orders of the N_c expansion: N_c^0 and N_c^{-2}



Sum of the first two orders of the N_c -expansion of the energy density covariance for $N=3$ in the classical MV model.



Ratio between the full result and the sum of the first two orders of the N_c -expansion, which turns out to be a very good approximation.

Conclusions

- We have performed an exact analytical calculation of the covariance of the energy momentum tensor of the **Glasma** at $\tau = 0^+$, in the framework of the **Color Glass Condensate**.
- We expect to be able to **generalize this framework** by introducing an impact parameter dependence and relaxing some of the original assumptions, which could potentially open the door to **phenomenological applications**.
- The following steps are computing the **time evolution** of our result towards thermalization time $\tau \sim 1/Q_s$, where it can serve as input for **hydro QGP simulations**.

$$T^{\mu\nu} = T_0^{\mu\nu} + T_1^{\mu\nu} \tau + T_2^{\mu\nu} \tau^2 + \dots$$



Conclusions

- We have performed an exact analytical calculation of the covariance of the energy momentum tensor of the **Glasma** at $\tau = 0^+$, in the framework of the **Color Glass Condensate**.
- We expect to be able to **generalize this framework** by introducing an impact parameter dependence and relaxing some of the original assumptions, which could potentially open the door to **phenomenological applications**.
- The following steps are computing the **time evolution** of our result towards thermalization time $\tau \sim 1/Q_s$, where it can serve as input for **hydro QGP simulations**.

Thanks for your attention

Back-up: Expressions of two first orders of Nc expansion

- Leading order:

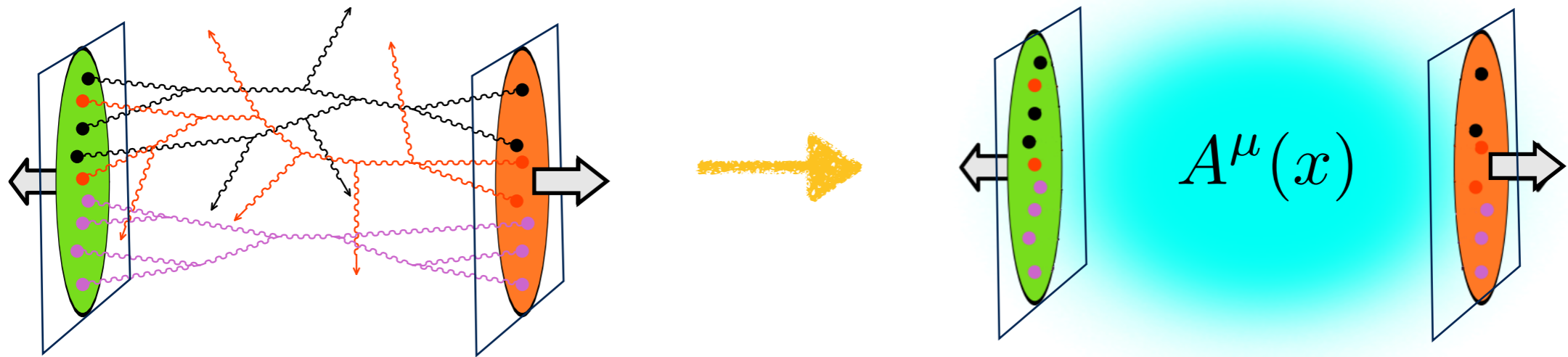
$$\begin{aligned}
 [\text{Cov}[\epsilon_{\text{MV}}](0^+; x_\perp, y_\perp)]_{N_c^0} &= \frac{1}{4g^4 r^8} e^{-\frac{r^2}{2}(Q_{s1}^2 + Q_{s2}^2)} \left(128 + 128 \left(e^{\frac{Q_{s1}^2 r^2}{2}} + e^{\frac{Q_{s2}^2 r^2}{2}} \right) \right. \\
 &\quad \left. - \left[256 e^{\frac{Q_{s1}^2 r^2}{4}} + 16 e^{\frac{r^2}{4}(2Q_{s1}^2 + Q_{s2}^2)} (Q_{s2}^4 r^4 + 8Q_{s2}^2 r^2 - 2Q_{s1}^4 r^4 + 48) \right] - [1 \leftrightarrow 2] \right. \\
 &\quad \left. - e^{\frac{r^2}{4}(Q_{s1}^2 + Q_{s2}^2)} \left[Q_{s1}^4 Q_{s2}^4 r^8 + 4Q_{s1}^2 Q_{s2}^2 r^6 (Q_{s1}^2 + Q_{s2}^2) \right. \right. \\
 &\quad \left. \left. + 128r^2 (Q_{s1}^2 + Q_{s2}^2) + 16r^4 (Q_{s1}^2 + Q_{s2}^2)^2 + 1024 \right] \right. \\
 &\quad \left. + 8e^{\frac{r^2}{2}(Q_{s1}^2 + Q_{s2}^2)} \left[-4r^4 (Q_{s1}^4 + Q_{s2}^4) + Q_{s1}^2 Q_{s2}^2 r^6 (Q_{s1}^2 + Q_{s2}^2) + 80 \right] \right)
 \end{aligned}$$

- First correction:

$$\begin{aligned}
 [\text{Cov}[\epsilon_{\text{MV}}](0^+; x_\perp, y_\perp)]_{N_c^{-2}} &= \frac{1}{4N_c^2 g^4 r^8} e^{-\frac{r^2}{2}(Q_{s1}^2 + Q_{s2}^2)} \left(16 (Q_{s1}^2 r^2 + Q_{s2}^2 r^2 + 8)^2 \right. \\
 &\quad \left. + \left[16Q_{s1}^2 r^2 (8 + Q_{s1}^2 r^2) e^{\frac{Q_{s2}^2 r^2}{2}} - 32(8 + Q_{s1}^2 r^2)(4 + Q_{s1}^2 r^2) e^{\frac{Q_{s2}^2 r^2}{4}} \right] + [1 \leftrightarrow 2] \right. \\
 &\quad \left. + \left[16 r^2 e^{\frac{1}{4}r^2(Q_{s1}^2 + 2Q_{s2}^2)} (r^2 (Q_{s1}^4 - 2Q_{s2}^4) + 8 (Q_{s1}^2 + 2Q_{s2}^2)) \right] + [1 \leftrightarrow 2] \right. \\
 &\quad \left. - 8e^{\frac{1}{2}r^2(Q_{s1}^2 + Q_{s2}^2)} r^2 (Q_{s1}^2 + Q_{s2}^2) (Q_{s1}^2 Q_{s2}^2 r^4 - 4r^2 (Q_{s1}^2 + Q_{s2}^2) + 32) \right. \\
 &\quad \left. + e^{\frac{1}{4}r^2(Q_{s1}^2 + Q_{s2}^2)} \left[Q_{s1}^4 Q_{s2}^4 r^8 + 4Q_{s1}^2 Q_{s2}^2 r^6 (Q_{s1}^2 + Q_{s2}^2) + 128r^2 (Q_{s1}^2 + Q_{s2}^2) \right. \right. \\
 &\quad \left. \left. + 16r^4 (Q_{s1}^2 + Q_{s2}^2)^2 - 1024 \right] \right)
 \end{aligned}$$

Color Glass Condensate: McLerran-Venugopalan model (modified)

- We use an approximation of QCD for high gluon densities where we replace the gluons with a **classical field** generated by the valence quarks



- Dynamics of the field described by **Yang-Mills** classical equations:

$$[D_\mu, F^{\mu\nu}] = J^\nu \propto \rho(x)$$

- Calculation of observables: **average** over background classical fields

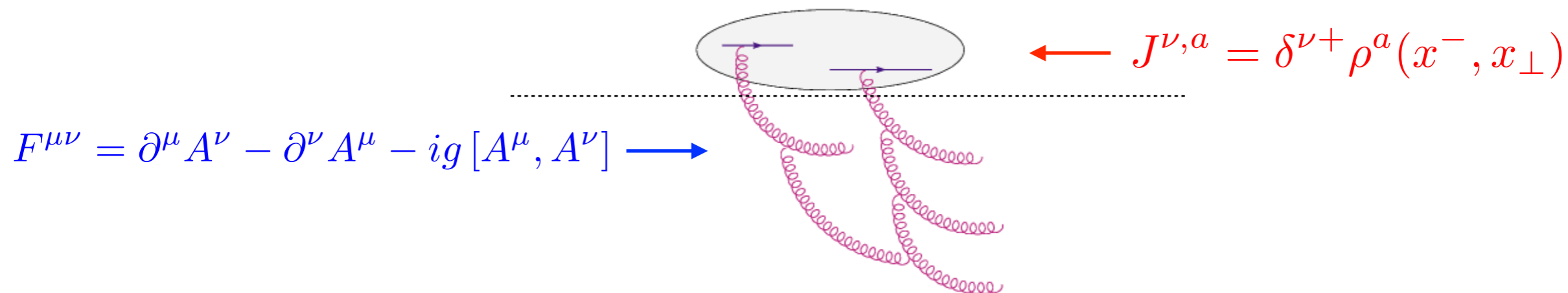
- Basic building block: (**generalized**) **2-point correlator**

$$\langle \rho^a(x^-, x_\perp) \rho^b(y^-, y_\perp) \rangle = \mu^2(x^-) h(b_\perp) \delta^{ab} \delta(x^- - y^-) f(x_\perp - y_\perp)$$

Non-Gaussianities

Back-up: More about the Color Glass Condensate

- Separation of ‘slow’ and ‘fast’ degrees of freedom



- Dynamic relation given by solution to classical Yang-Mills equations:

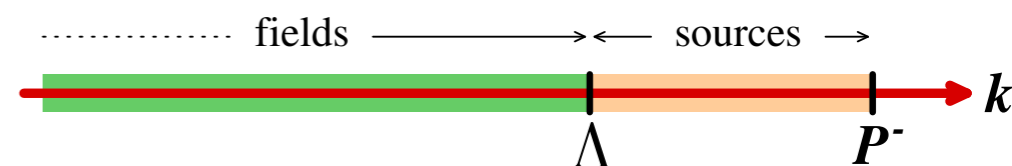
$$[D_{\mu}, F^{\mu\nu,a}] = J^{\nu,a}$$

- Calculation of observable quantities: average over color sources

$$\langle \mathcal{O} \rangle = \int [D\rho] W_{\Lambda}[\rho] \mathcal{O}[\rho]$$

- Scale (in)dependence: JIMWLK equations

$$\frac{\partial W_{\Lambda}}{\partial \log \Lambda} = \mathcal{H} W_{\Lambda}$$



- McLerran-Venugopalan model: W_{Λ} is a Gaussian distribution

$$\langle \rho^a(x^-, x_{\perp}) \rho^b(y^-, y_{\perp}) \rangle = \mu^2(x^-) \delta^{ab} \delta^2(x_{\perp} - y_{\perp}) \delta(x^- - y^-)$$