Charged Lepton Flavor Violation:

$$\mu^- N \rightarrow e^- N$$

R. Bernstein
Fermilab

PASCOS 2018
CWRU, Cleveland IL

What is CLFV?

- a transition among μ, e, τ that doesn't conserve lepton family number

  - muon decay: $\mu^- \rightarrow e^- \bar{\nu}_e \nu_\mu$ has two neutrinos

  - CLFV is seen in (for example) $\mu \rightarrow e\gamma$ or $\mu \rightarrow 3e$ with NO neutrinos; both being searched for (MEG, Mu3e)

  - similar τ decays: $\tau \rightarrow \mu, e + X$ (and no neutrinos)

  - in K system, $K_L \rightarrow \mu e$ and charged B, K to dileptons

  - $H \rightarrow \mu, e, \tau + X$

Two Ways to Look at CLFV

$$\mathcal{L} = \sum_n \frac{\text{couplings}}{\Lambda^n}$$

1. What is the mass scale of new physics?

   - Muon Lifetime and Size of G_F told us about a mass scale M_W ;

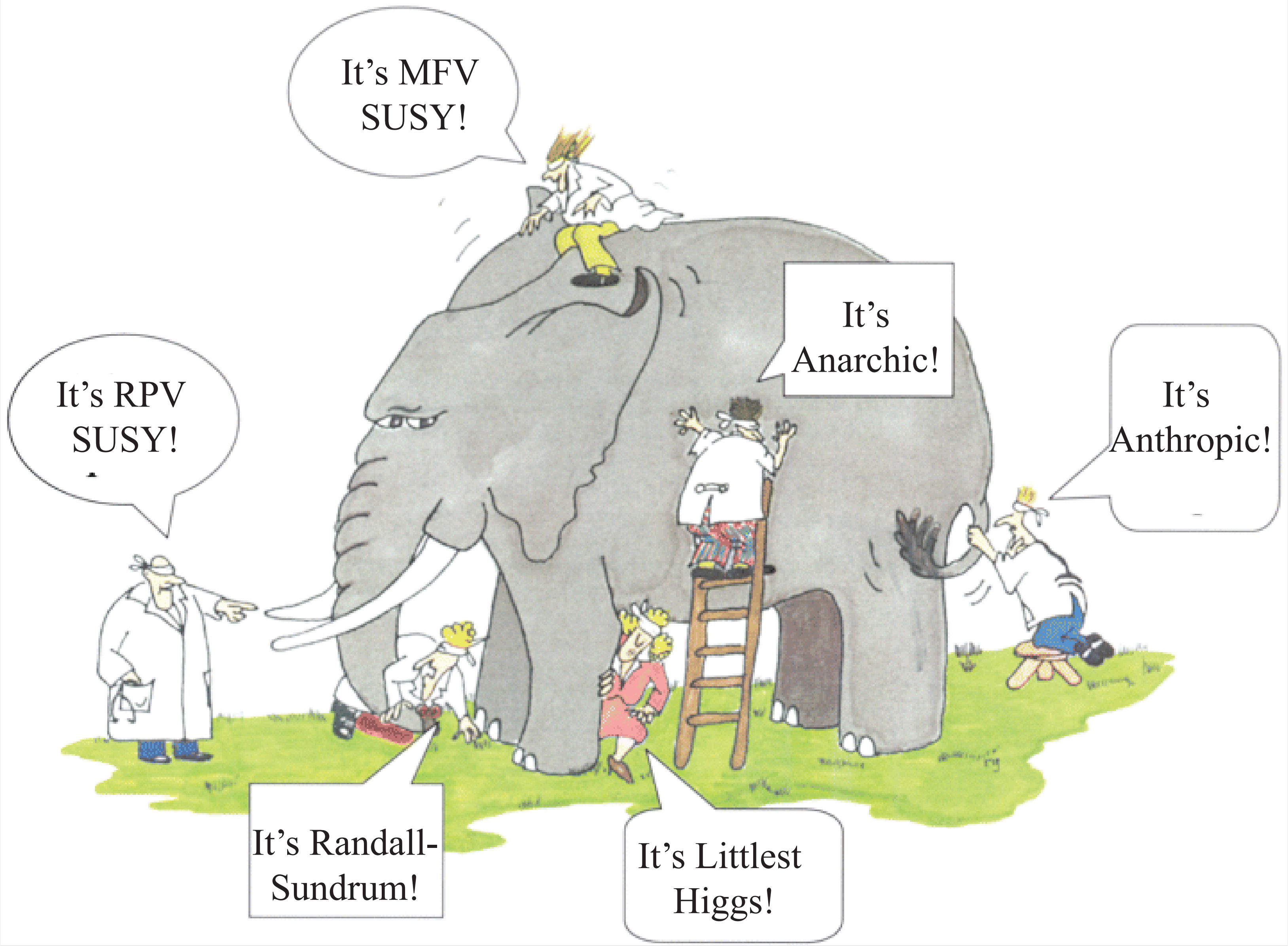
2. What are the symmetries and flavor structure of new physics?

   - the mass scale is not everything; you don’t know the couplings or the sizes of mixing angles

Family number: not a fundamental symmetry

- family number is not a symmetry of Lagrangian like charge
  - quark family number is violated in weak decays in the CKM matrix
  - we know it's violated in neutral leptons: neutrino oscillations (PMNS matrix)
- but we've never seen it in charged leptons. Most "natural" new physics models predict we should have seen it already, even if small. Why haven't we?

How It Seems With Single Measurements



Muon Processes

- Decay Experiments:
  - $\mu \to e\gamma$
    - oldest studied, most powerful limits, and the best experiment so far: MEG at PSI. Back-to-back, limited by resolution and accidentals
  - $\mu \to eee$
    - ambitious and unique, in planning at PSI
- Conversion Experiment:
  - $\mu N \to eN$
    - three experiments upcoming at FNAL and JPARC

What is Muon to Electron Conversion?

- Muon converts to an electron in the field of a nucleus:

$$R_{\mu e} = \frac{\Gamma(\mu^- + N(A,Z) \rightarrow e^- + N(A,Z))}{\Gamma(\mu^- + N(A,Z) \rightarrow \text{all muon captures})}$$

no neutrinos

- Effectively zero Standard Model background
  - a signal is new physics
- Experiments to be discussed can measure a signal with SES of $\sim 3 \times 10^{-17}$, discovery at 2×10^{-16}
  - typical predictions for EWSB $\sim 10^{-15}$

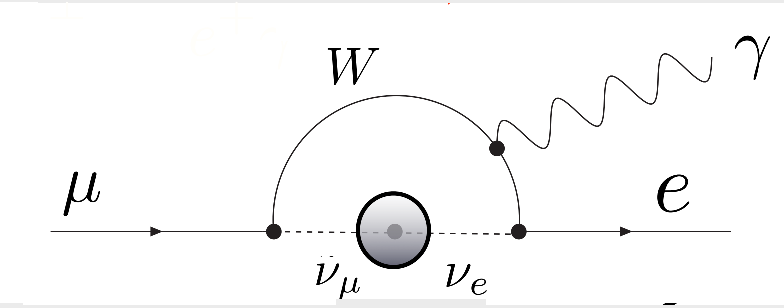
How Small is Our SES?

- Pretty Rare: let us know if this happens to you! You can immediately join the experiment

Probability of...	
rolling a 7 with two dice	1.67E-01
rolling a 12 with two dice	2.78E-02
getting 10 heads in a row flipping a coin	9.77E-04
drawing a royal flush (no wild cards)	1.54E-06
getting struck by lightning in one year in the US	2.00E-06
winning Pick-5	5.41E-08
winning MEGA-millions lottery (5 numbers+megaball)	3.86E-09
your house getting hit by a meteorite this year	2.28E-10
drawing two royal flushes in a row (fresh decks)	2.37E-12
your house getting hit by a meteorite today	6.24E-13
getting 53 heads in a row flipping a coin	1.11E-16
your house getting hit by a meteorite AND you being struck by lightning both within the next six months	1.14E-16
your house getting hit by a meteorite AND you being struck by lightning both within the next three months	2.85E-17

Neutrinos and Mu2e

- Neutrino Oscillations are the only Standard Model background (except neutrino oscillations are not in the Standard Model)



$$\mathrm{BR}(\mu \to e\gamma) \quad = \quad \frac{3\alpha}{32\pi}\left|\sum_{i=2,3} U^*_{\mu i} U_{ei} \frac{\Delta m^2_{1i}}{M^2_W}\right|^2 < 10^{-54}$$

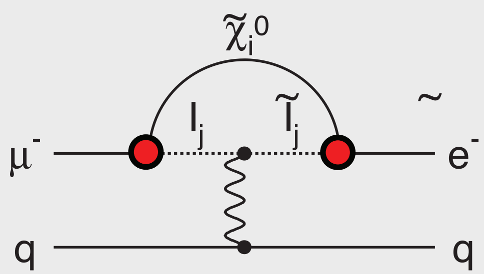
- nobody understood why $\mu \to e\,\gamma$ wasn't 10^{-4} until we hypothesized two neutrinos ($\nu_\mu \neq \nu_e$)

- linkage of CLFV to oscillations, see-saw, Majorana vs. Dirac are getting more interest

Contributions to μe Conversion

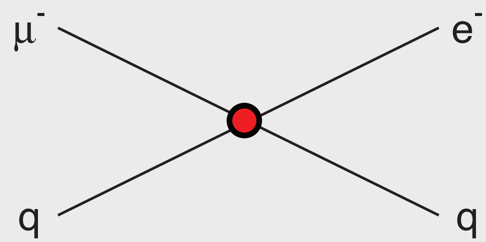
Supersymmetry

rate ~ 10^{-15}



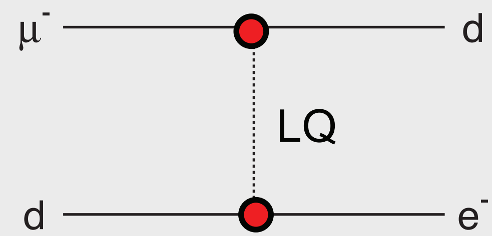
Compositeness

Λ_c ~ 3000 TeV



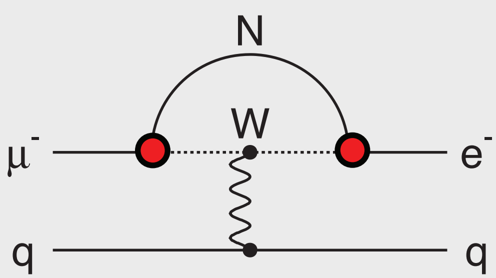
Leptoquark

$M_{LQ} =$
$3000\ (\lambda_{\mu d}\lambda_{ed})^{1/2}$ TeV/c^2



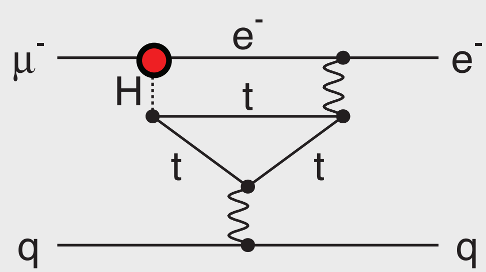
Heavy Neutrinos

$|U_{\mu N}U_{eN}|^2$ ~ 8×10^{-13}



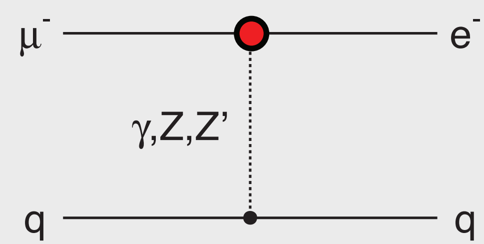
Second Higgs Doublet

$g(H_{\mu e})$ ~ $10^{-4} g(H_{\mu\mu})$



Heavy Z' Anomal. Z Coupling

$M_{Z'} = 3000$ TeV/c^2



also see Flavour physics of leptons and dipole moments, arXiv:0801.1826 ;
Marciano, Mori, and Roney, Ann. Rev. Nucl. Sci. 58, doi:10.1146/annurev.nucl.58.110707.171126 ;
de Gouvea and Vogel, arXiv:1303.4097v2

Effective Lagrangian

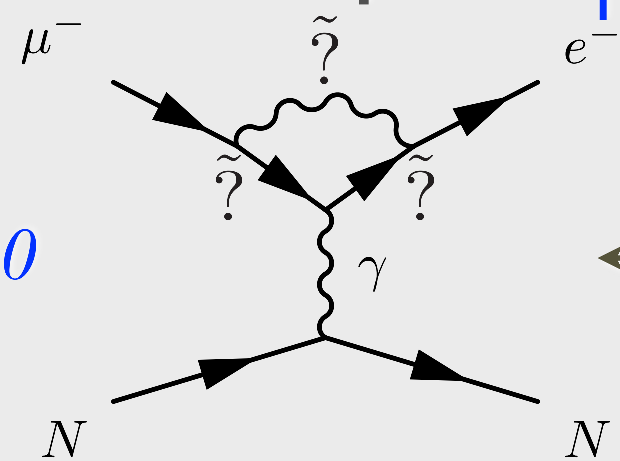
$$\mathcal{L}_{\rm CLFV} = \frac{m_\mu}{(\kappa+1)\Lambda^2}\bar{\mu}_R\sigma_{\mu\nu}e_L F^{\mu\nu} + \frac{\kappa}{(1+\kappa)\Lambda^2}\bar{\mu}_L\gamma_\mu e_L(\bar{u}_L\gamma_\mu u_L + \bar{d}_L\gamma_\mu d_L)$$

"Loops"

mass scale Λ

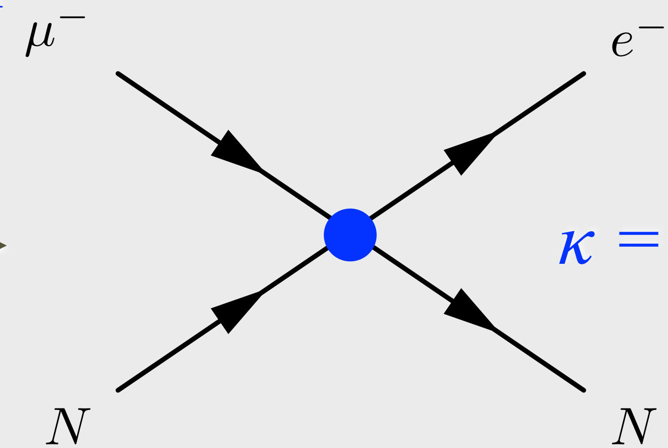
"Contact Terms"

$\kappa = 0$



κ

$\longleftrightarrow$



$\kappa = \infty$

Supersymmetry and Heavy Neutrinos

Contributes to $\mu \rightarrow e\gamma$

(just imagine the photon is real)

New Particles at High Mass Scale (leptoquarks, heavy Z,...)

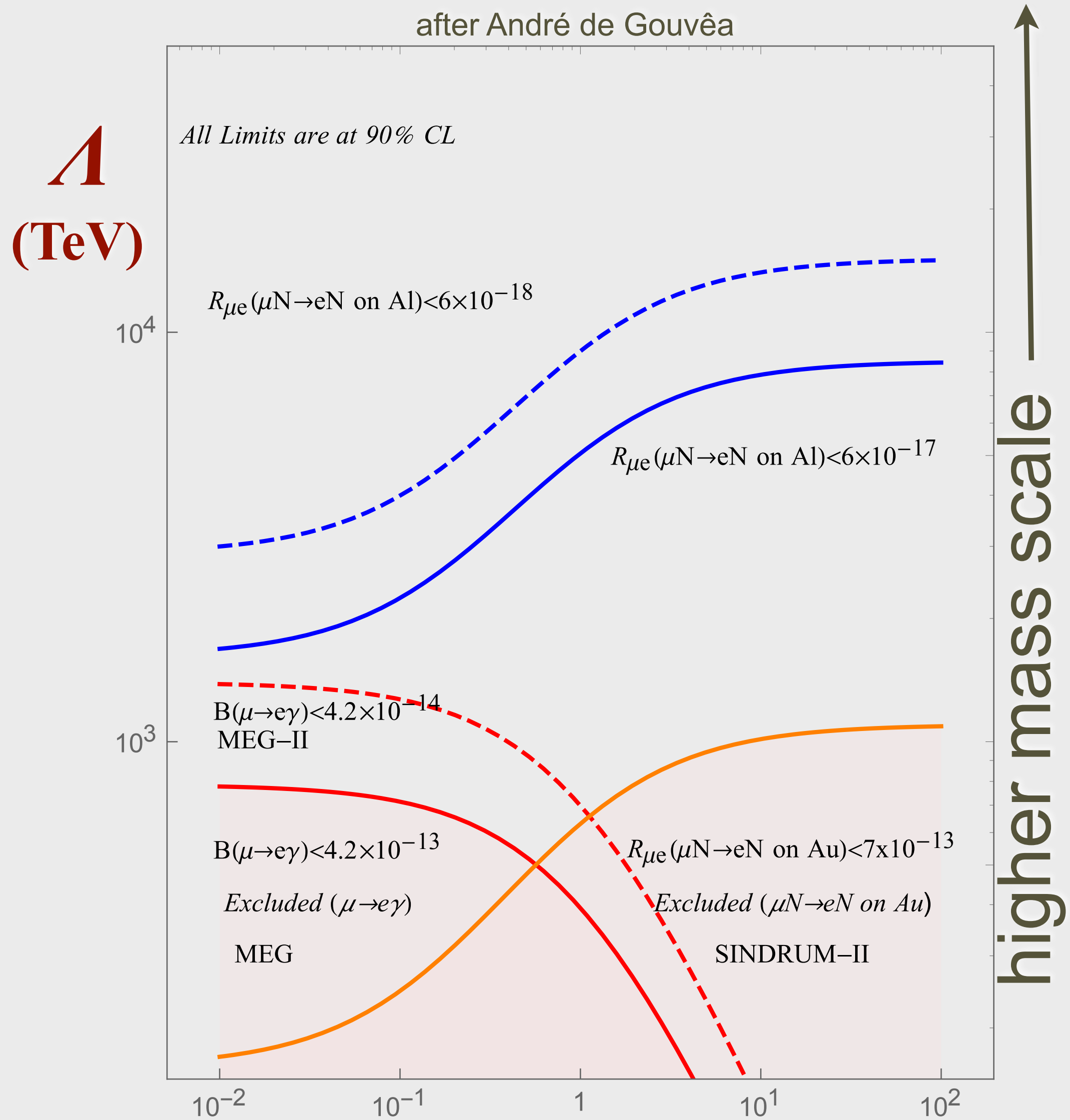
Does not produce $\mu \rightarrow e\gamma$

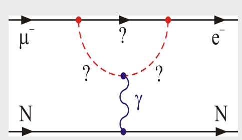
μe Conversion and $\mu \rightarrow e\gamma$

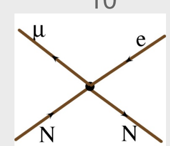
1) *Mass Reach to ~10^4 TeV for unit coupling, x10 in mass scale for dim-6*

2) *Mu2e and MEG-II are complementary in loop-dominated physics.*

3) These are discovery experiments



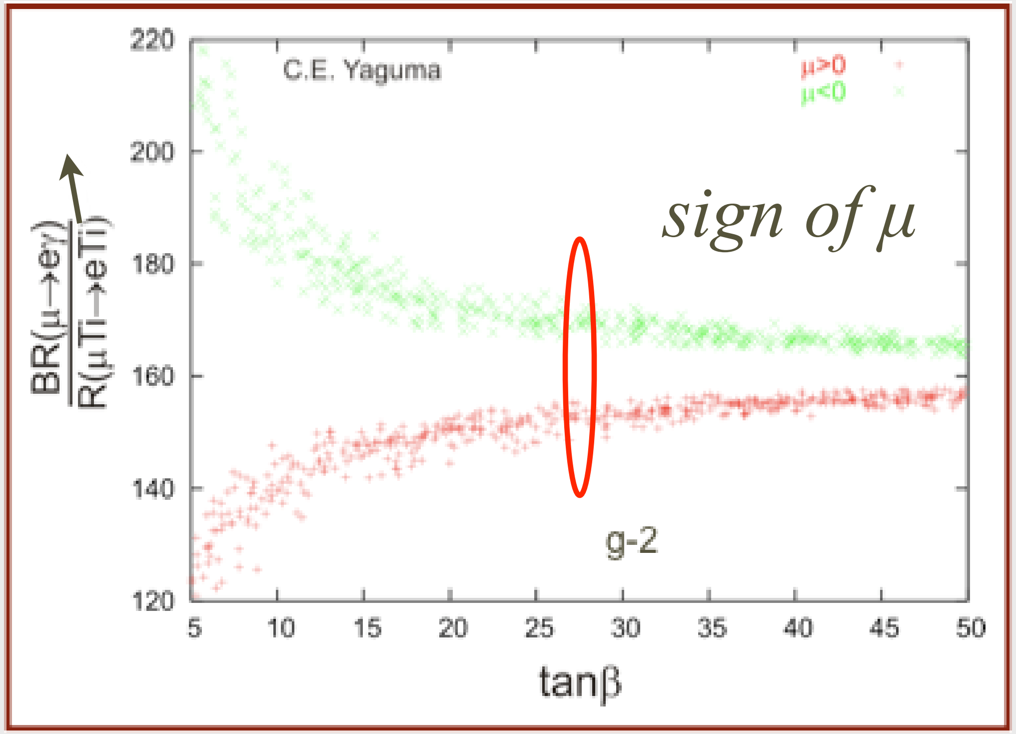




Mu2e, g-2, and $\mu \to e\gamma$

Yaguna, hep-ph/0502014v2

MSSM w mSUGRA



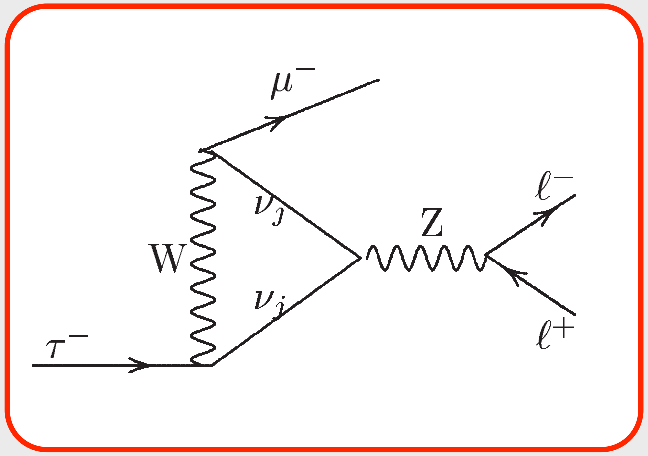
$$M_{\mathrm{H}} = \sqrt{2}\,|\mu|$$

- Need:
  - observation of CLFV in more than one channel, and/or
  - evidence from LHC, g-2, or elsewhere

to allow discrimination among different models

CLFV and Tau Processes

- Advantage:
  - Beyond SM rates can be orders of magnitude higher than in the corresponding muon channel
- Disadvantage
  - τ's hard to produce:
    - ~10^{11} τ/*yr* vs ~10^{11} μ/*sec* in muon experiments



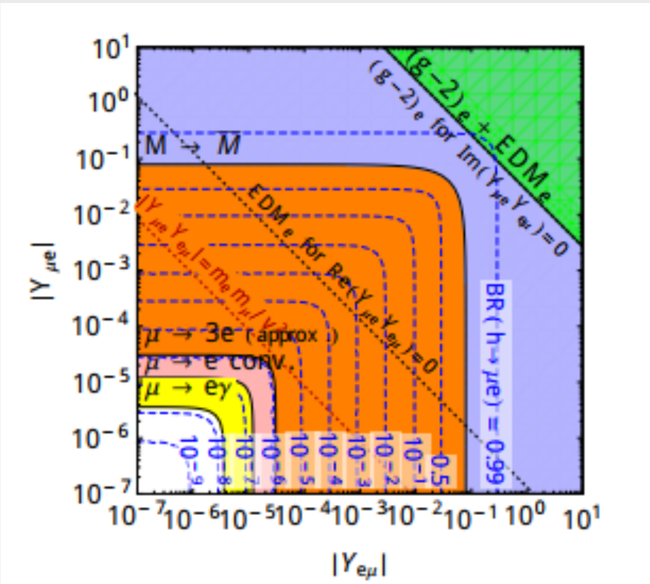
SM ~ 10^{-14} ?

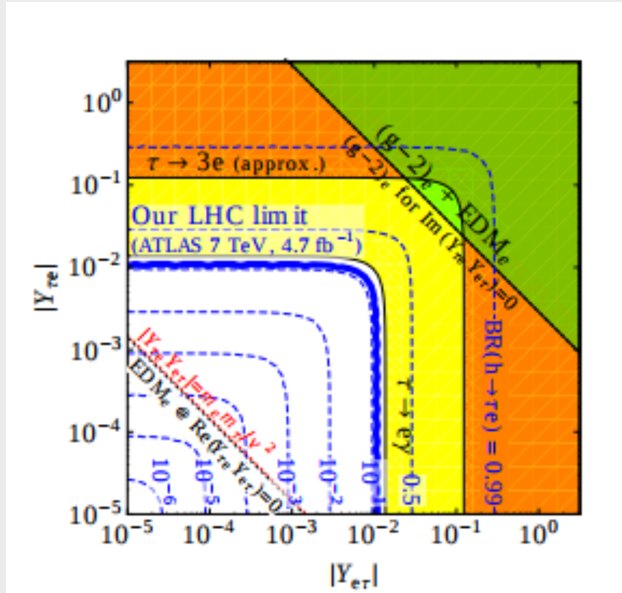
τ's help pin down models; e.g. Cheng-Sher Ansatz $\sqrt{m_i m_j}$

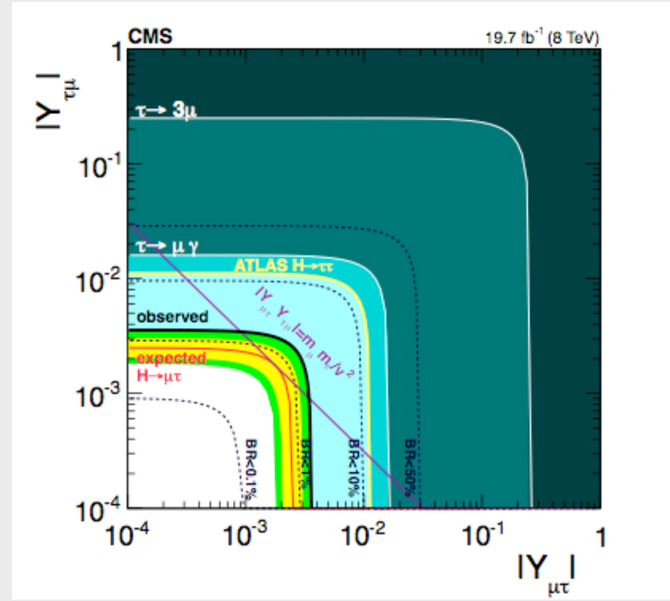
newer: CMS 1502.07400

Combine LHC/Direct

$Y_{l_i l_j}^2 + Y_{l_j l_i}^2 = \text{constant}$ ($\not{P}$ can make mixing asymmetric)







Lepton Flavor Violating Higgs Decays

Harnik, Kopp, Zupan, arXiv:1209.1397

- all $H \to l_i l_j$ consistent with zero
- muon experiments much better in $e\mu$ modes
- colliders unique in τ, can't make in low-energy μ beams

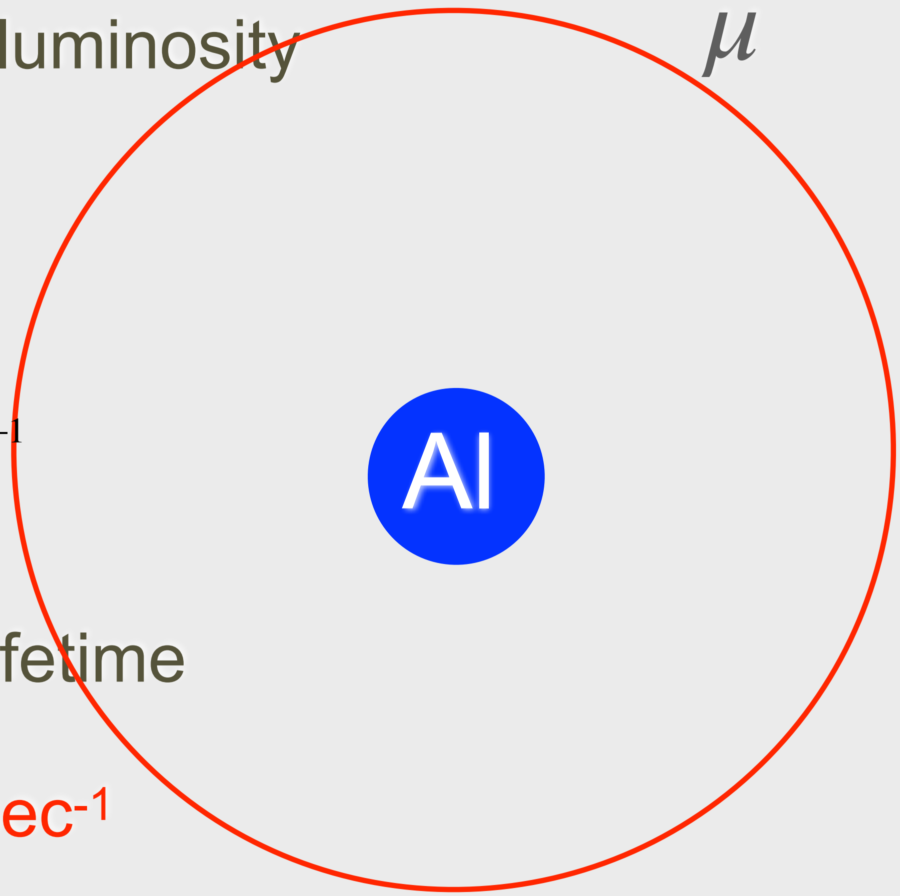
Collider or Muon Beam?

- 1st/2nd generation processes (“solar”)
    - ***Muon experiments orders of magnitude better than colliders in muon-electron processes because we can make beams***
- 2nd/3rd,1st/3rd (“atmospheric and θ_{13}”)
    - ***can’t make τ beams, so colliders best there***
    - ***B*, K*, b→sll, b→ll and b→sγ puzzles: lepton universality***
    - *pp* or e+e- both have contributions depending on channel; e.g. CLFV Higgs decays or τ CLFV (ATLAS/CMS, LHCb, and BELLE-II or BES-III; don’t overlook e+e- and B mesons)

Measuring 10^{-17} in Collider Units

- The captured muon is in a 1s state and the wave function overlaps the nucleus (*picture ~ to scale*)

- We can turn this into an effective luminosity



- Luminosity = density x velocity

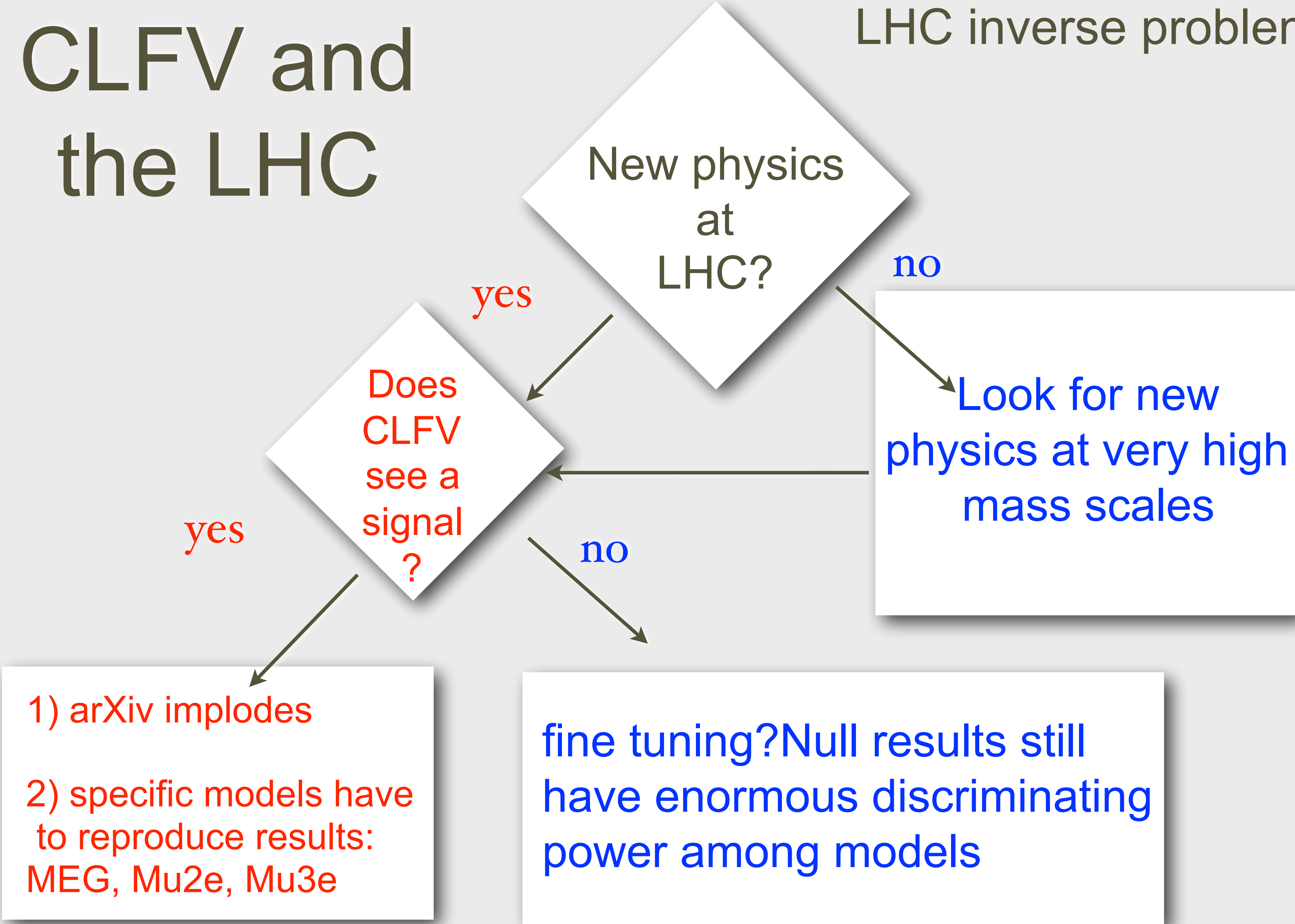
$$|\psi(0)|^2 \times \alpha Z = \frac{m_\mu^3 Z^4 \alpha^4}{\pi} = 8 \times 10^{43}\,\mathrm{cm}^{-2}\,\mathrm{sec}^{-1}$$

- Times 10^{10} muons/sec X 2 μsec lifetime

- Effective Luminosity of 10^{48} cm^{-2}sec^{-1}
  - vs HL-LHC at ~ 10^{35} cm^{-2}sec^{-1} without pile-up

CLFV and the LHC

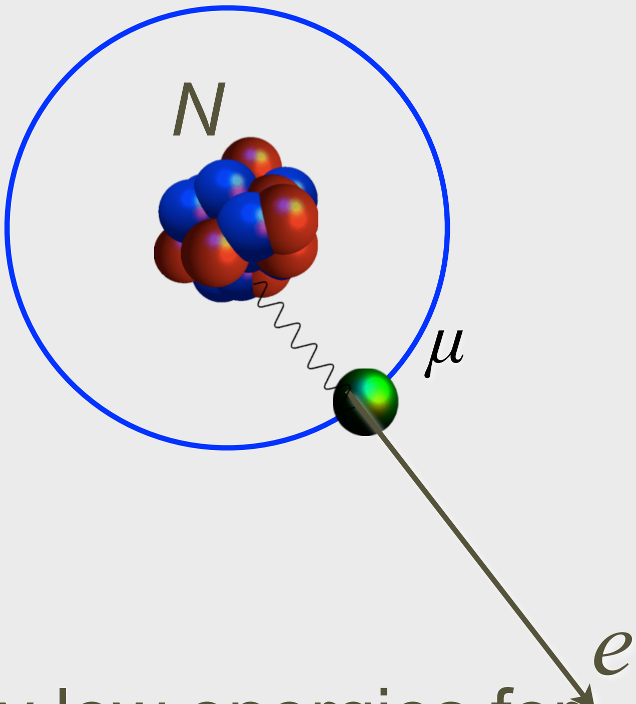
LHC inverse problem



Experimental Signal

$$\mu^- N \rightarrow e^- N$$

- A Single Mono-energetic Electron
  - muon falls into 1s state in "stopping target"
- Nucleus coherently recoils off outgoing electron, no breakup
- If N = Al, E_e ~ 105. MeV ~m_μ
  - electron energy depends on Z
  - physics in Z-dependence of rate



very low energies for HEP!

Mu2e: x10,000 improvement

- World's most intense muon beam:
  - 10^{10} μ/sec at detector under construction
  - Superconducting solenoid system to form muon beam
    - Efficient muon collection; see ~10^{18} muons
- Have to reduce backgrounds to take advantage of statistics
  - Pulsed Beam reduces a background
  - And a high-resolution spectrometer to make muon-related backgrounds small

Backgrounds and Design

	source	*scales with*	*solution*
Intrinsic Backgrounds	decay-in-orbit	# of muons	detector resolution
Beam	radiative pion capture	closeness to beam pulse	pulsed beam
Running Time	cosmic ray	live time	veto system

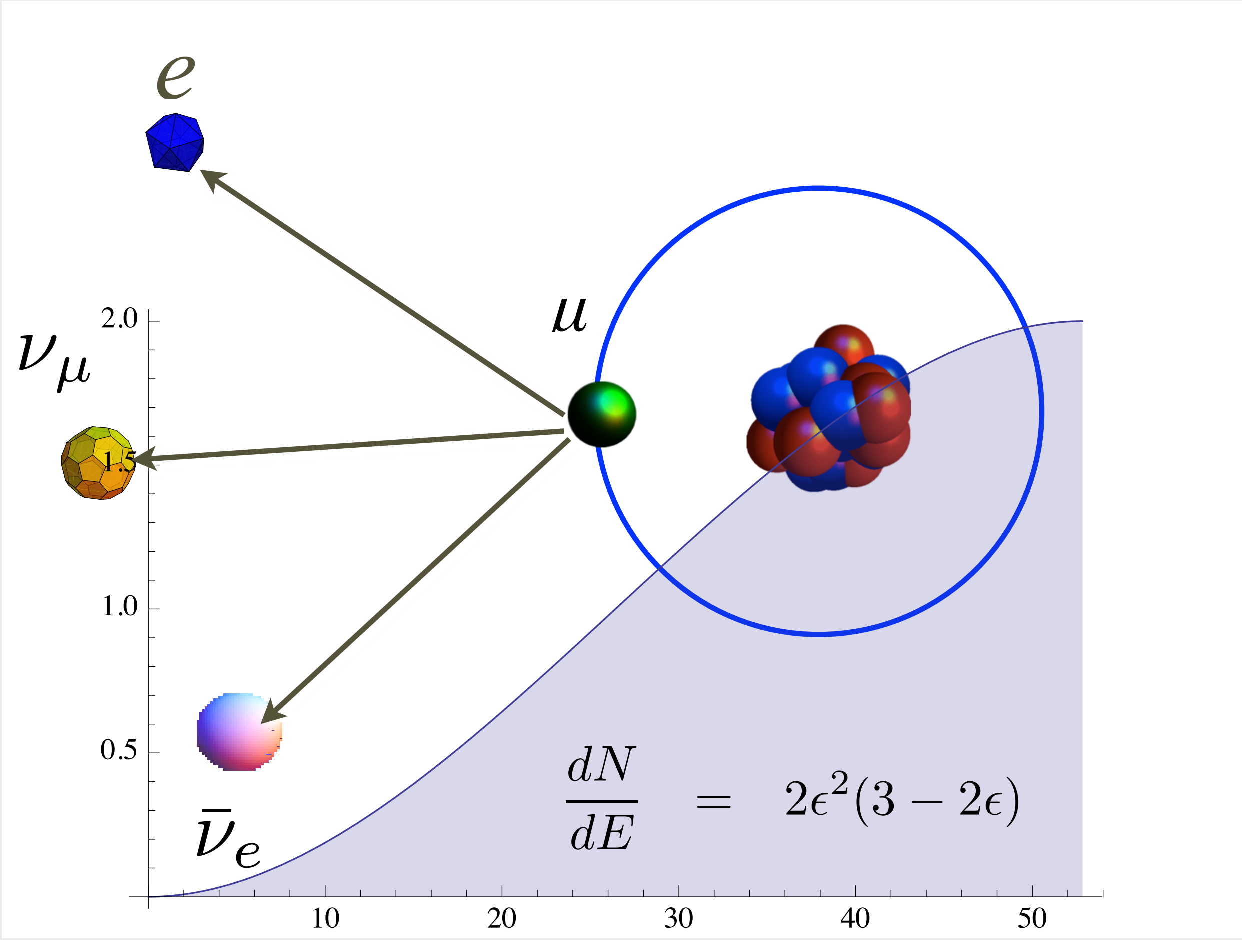
Decay-In-Orbit: usually not background

- Peak and Endpoint of free muon decay spectrum is at

$$E_{\max} = \frac{m_\mu^2 + m_e^2}{2m_\mu} \approx 52.8 \text{ MeV}$$

- Detector will be insensitive to electrons at this energy

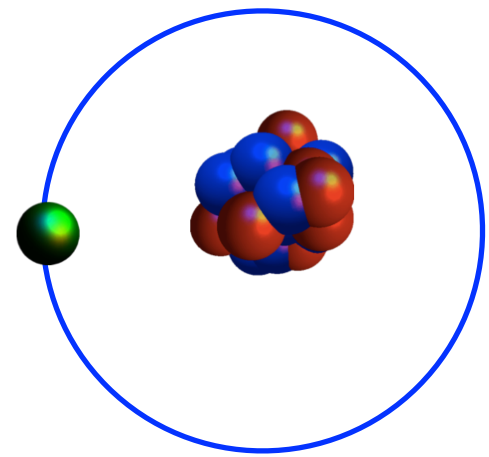
- Recall *signal* at 105 MeV>>52.8 MeV



Michel=free muon decay

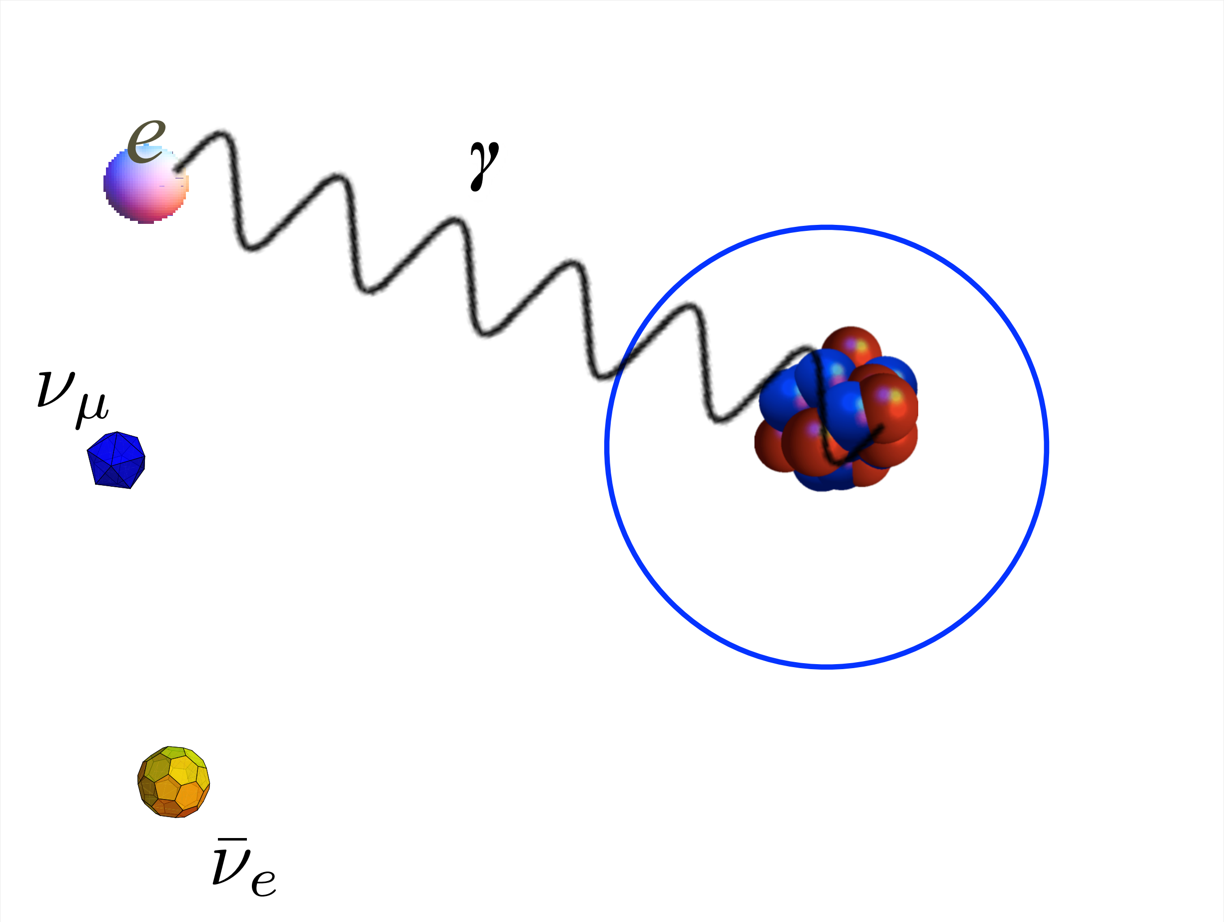
Decay-In-Orbit Background

- Same process as before
- But this time, include electron recoil off nucleus
- If neutrinos are at rest, the DIO electron can be exactly at conversion energy (up to neutrino mass)



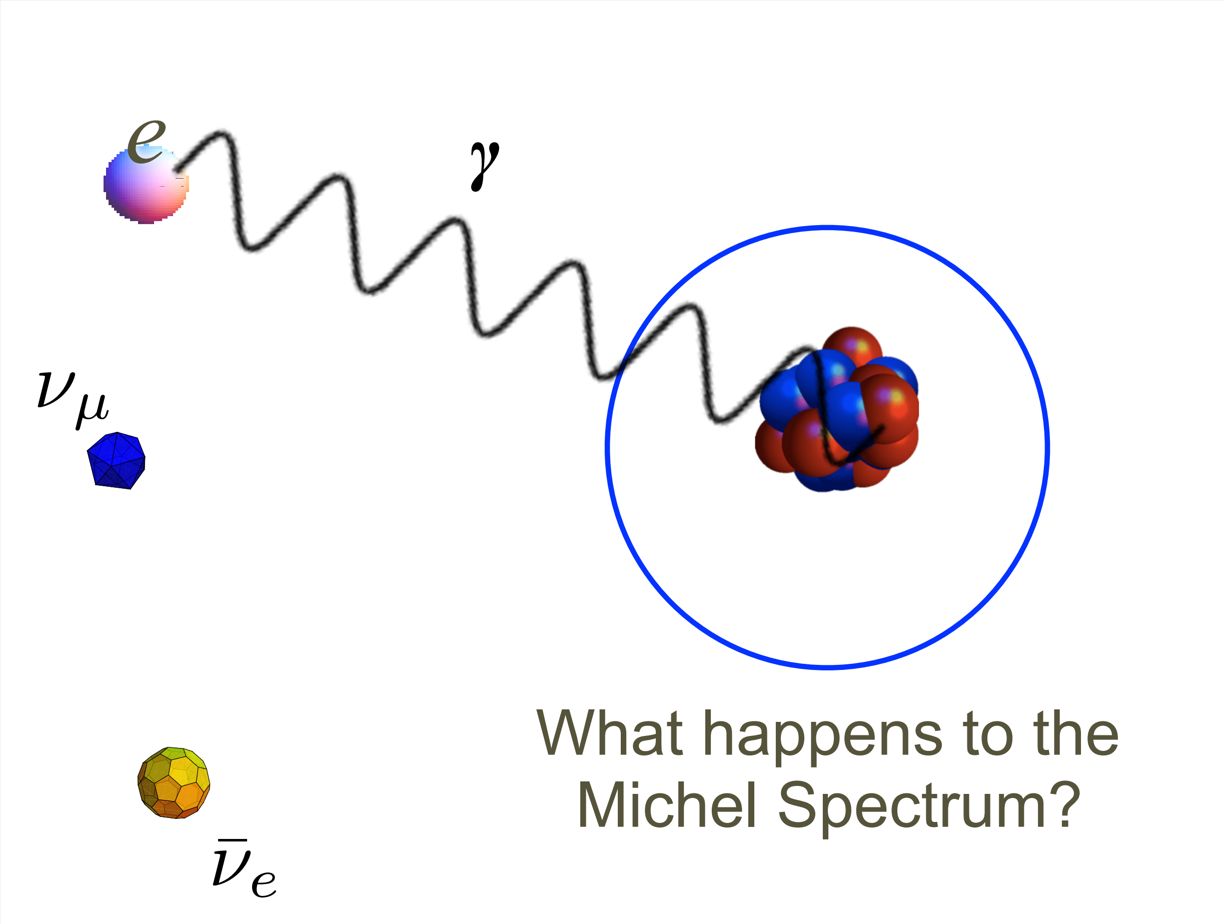
Decay-In-Orbit Background

- Same process as before
- But this time, include electron recoil off nucleus
- If neutrinos are at rest, the DIO electron can be exactly at conversion energy (up to neutrino mass)



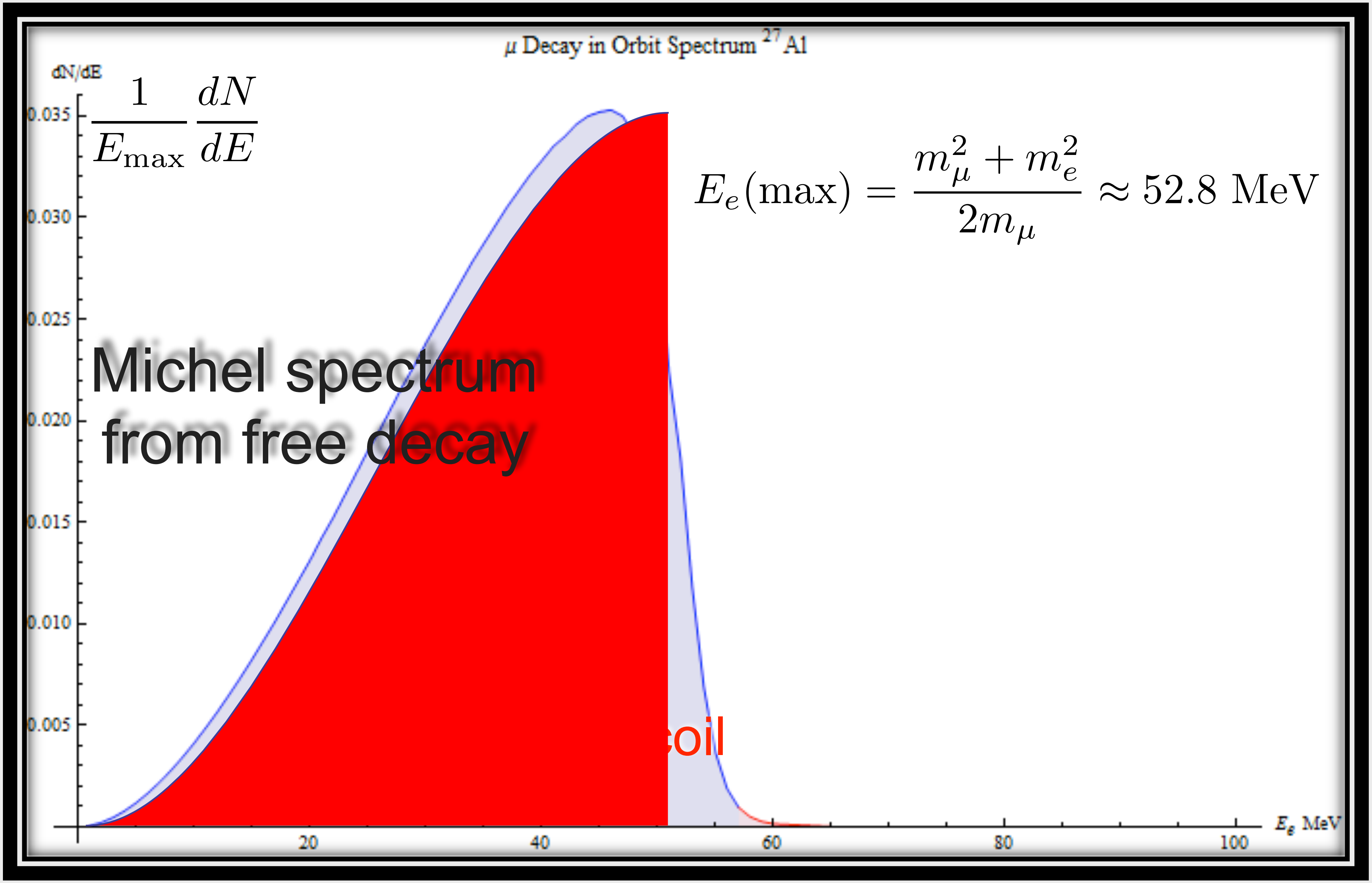
Decay-In-Orbit Background

- Same process as before
- But this time, include electron recoil off nucleus
- If neutrinos are at rest, the DIO electron can be exactly at conversion energy (up to neutrino mass)



Decay-in-Orbit Shape

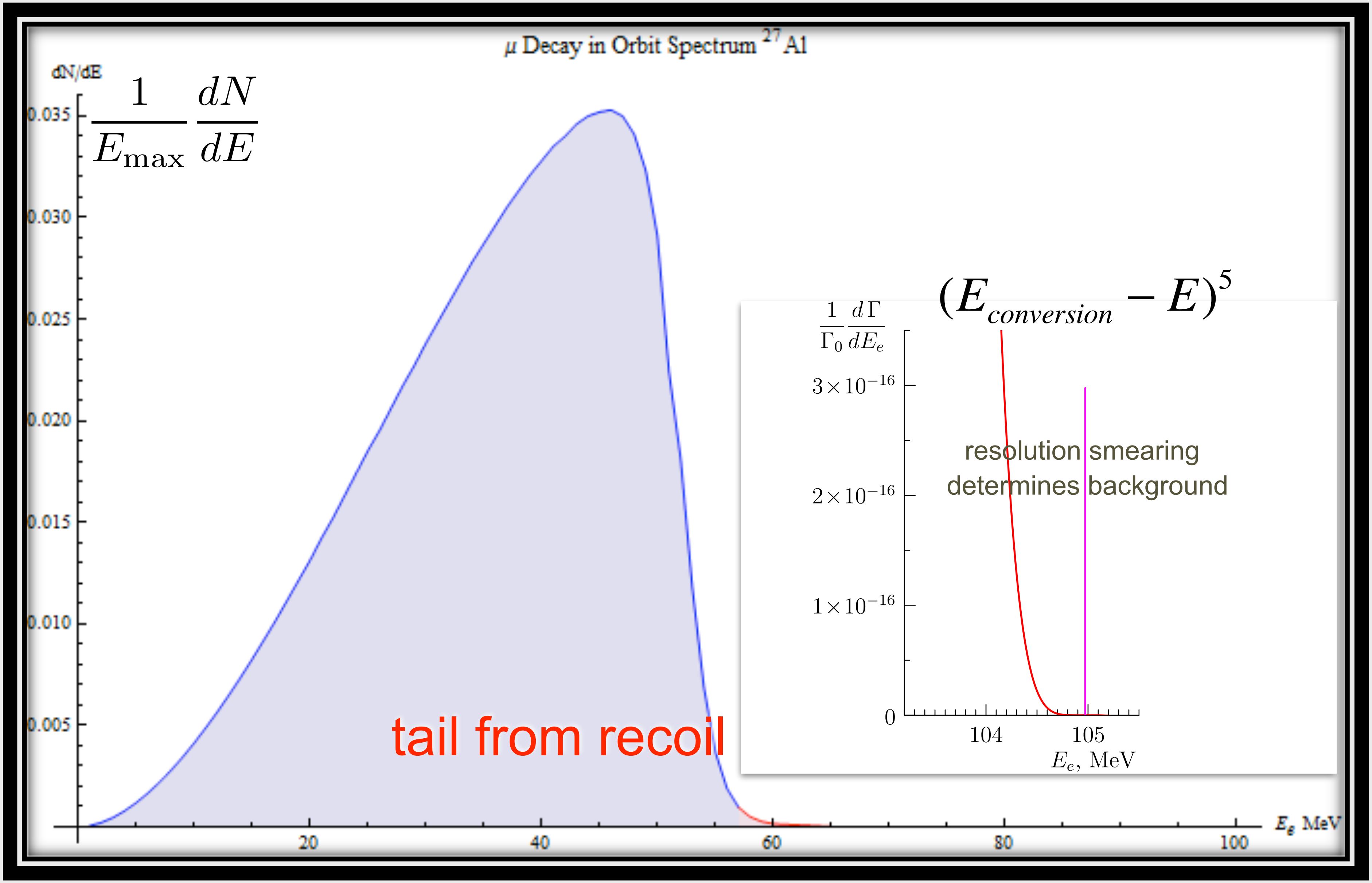
Szafron 10.5506/APhysPolB.46.2279: Radiative Corrections



Czarnecki: 10.1016/j.physletb.2015.12.008 many other papers

Decay-in-Orbit Shape

Szafron 10.5506/APhysPolB.46.2279: Radiative Corrections



Czarnecki: 10.1016/j.physletb.2015.12.008 many other papers

Backgrounds and Design

	source	*scales with*	*solution*
Intrinsic Backgrounds	decay-in-orbit	# of muons	detector resolution
Beam	radiative pion capture	closeness to beam pulse	pulsed beam
Running Time	cosmic ray	live time	veto system

Radiative Pion Capture

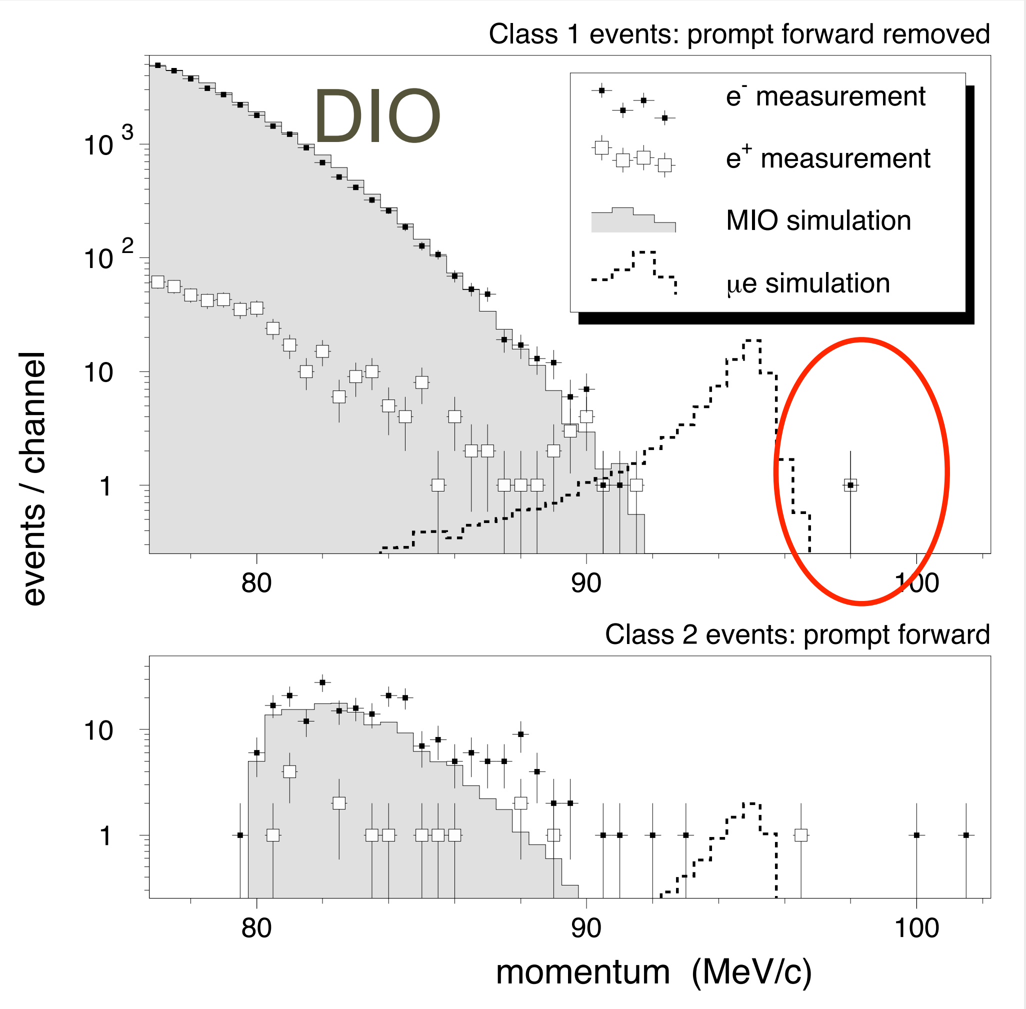
- Need π's to make μ's
- Sometimes those π's live long enough to make it to the stopping target
- They can then undergo radiative pion capture (RPC):
  - $\pi N \rightarrow \gamma N'$ and the γ converts to e^+e^- in the Al.
  - $\pi N \rightarrow \gamma N'$ and the γ internally converts
- Sometimes the e^- is at the signal energy

SINDRUM-II Results

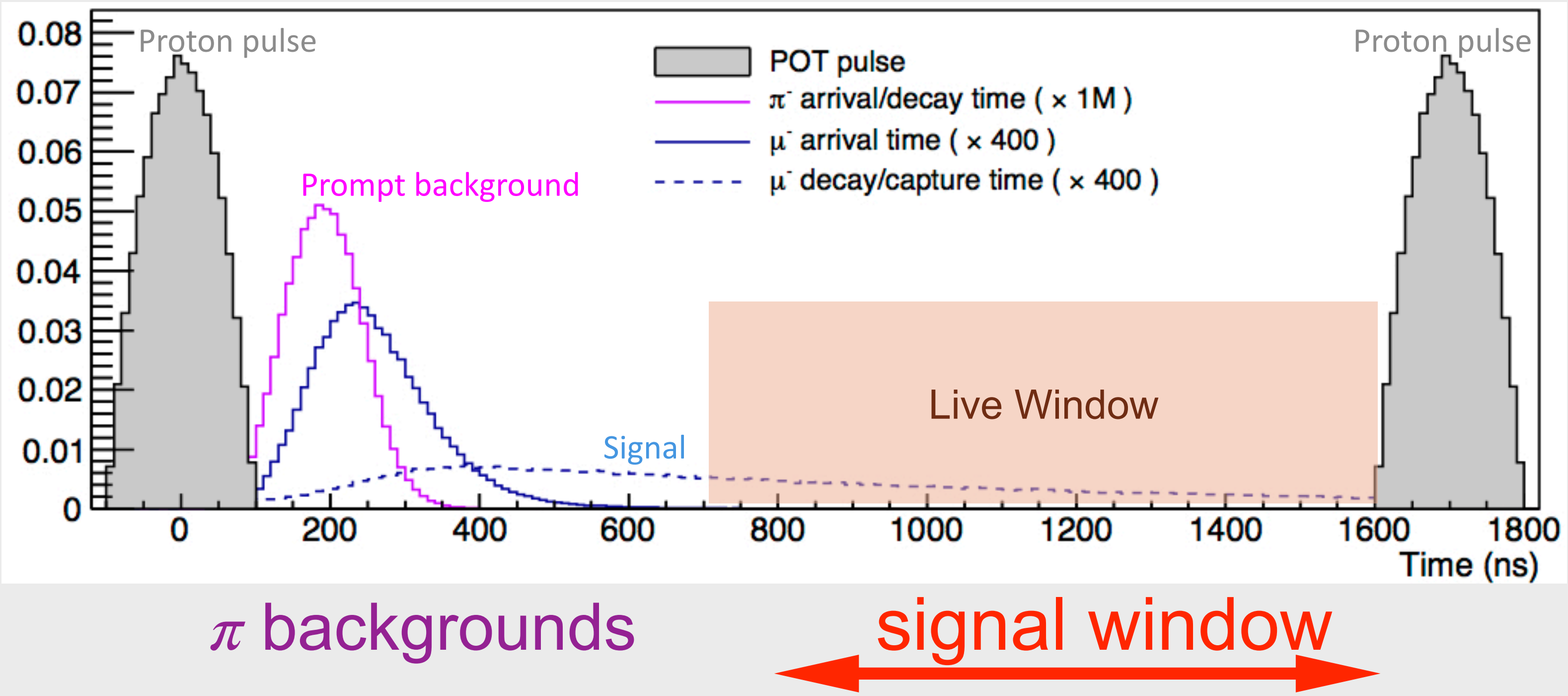
W. Bertl et al., Eur. Phys. J. C 47, 337–346 (2006)

- Effectively constant beam
    - 51 MHz (~20 nsec) repetition rate, ~0.3 nsec pulse
- Small time separation between signal and prompt pion backgrounds
    - bottom plot is first half of 20 nsec, top plot is 2nd half
    - delay lowers background

$$B_{\mu e}^{\mathrm{Au}} \quad < \quad 7 \times 10^{-13} \text{ @ } 90\% \text{ CL}$$

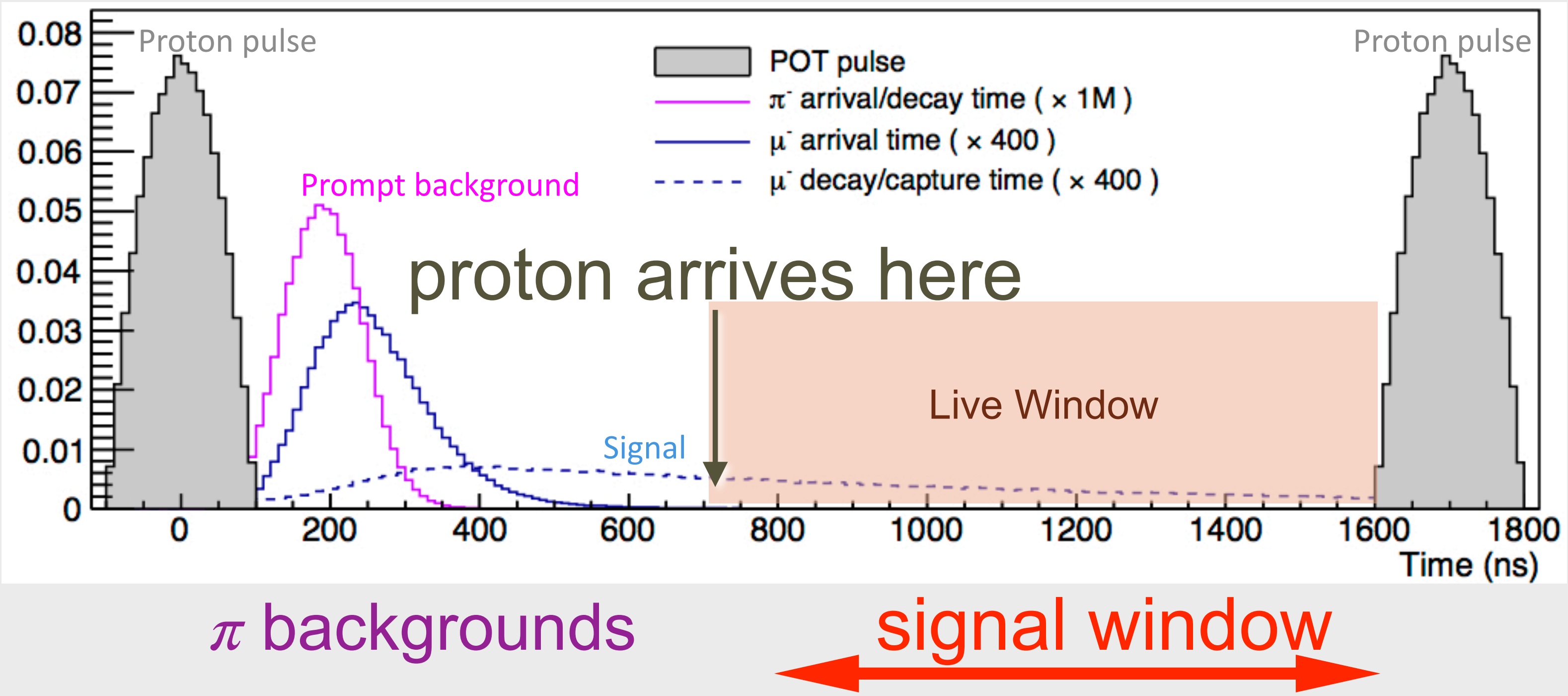


Pulsed Beam Structure



- Wait out pion backgrounds, then open signal window
- Can't "restart the clock" with out-of-pulse protons
- Need 10^{-10} in-pulse/out-of-pulse ratio (**extinction**)
  - will measure and monitor during experiment

Pulsed Beam Structure

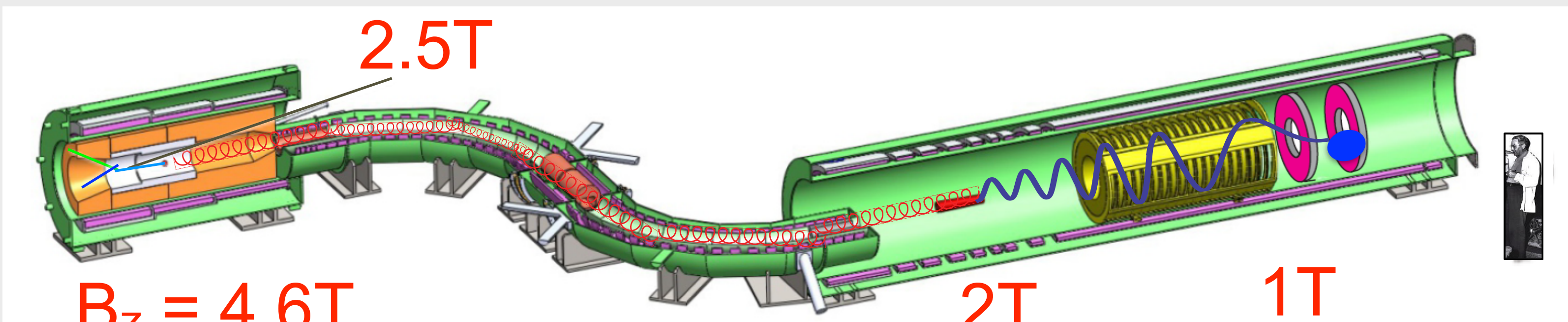


- Wait out pion backgrounds, then open signal window
- Can't "restart the clock" with out-of-pulse protons
- Need 10^{-10} in-pulse/out-of-pulse ratio (**extinction**)
  - will measure and monitor during experiment

Actual Experiments: Mu2e/COMET

mu2e.fnal.gov

4.6T ⟶ B-field gradient ⟶ 1T



Mu2e

Production Transport Detector

- Target protons at 8 GeV inside superconducting solenoid at 8kW: shielding a challenge

- Capture muons and guide through S-shaped region to Al stopping target

Detector Layout

finding electron track and momentum determination: highest precision, needed to reject decay-in-orbit

tracker

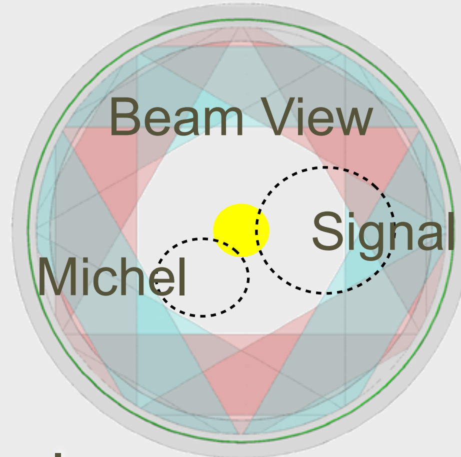
beam stop

stopping target

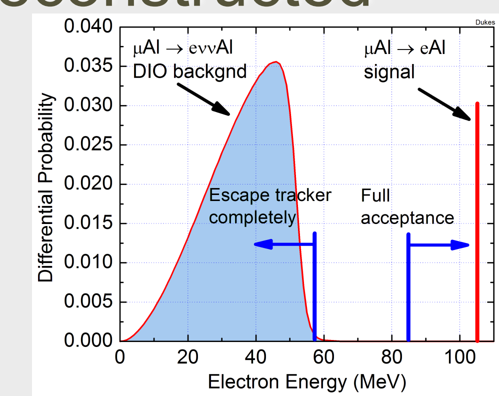
calorimeter

1) PID separation of $\pi/\mu/e$
2) Seed for Track Finder
3) Standalone Trigger

Why Annular?

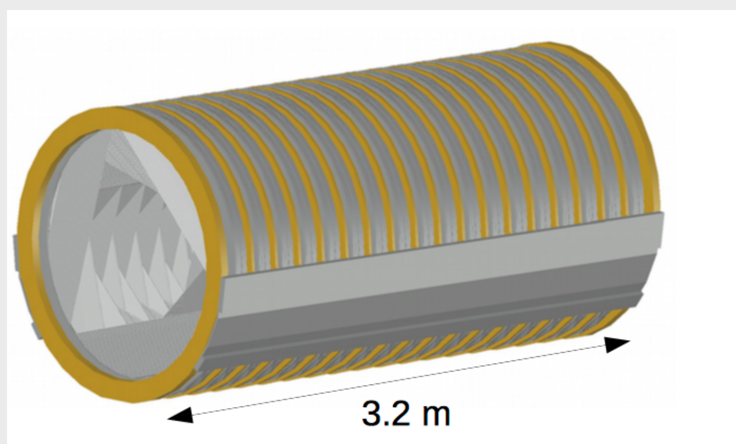


- Reduces beam activity and a background
  - don't want initial "flash" of activity from proton pulse, or remnant muon beam in detector (only ~1/2 of muon beam stops in target)
  - eliminates all but a tiny number of decay-in-orbit background
  - p sinθ ~ qBR in solenoid, Michel e- at small radius
  - so we're not rejecting 10^{18}; only ~10^5 even reconstructed
- COMET uses a different method; not annular
  - advantages and disadvantages



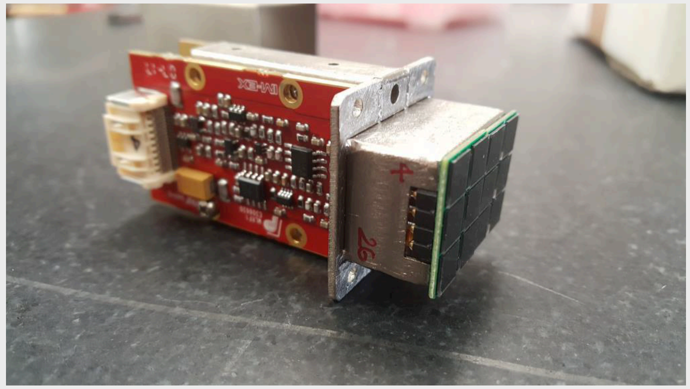
Straw Tube Tracker

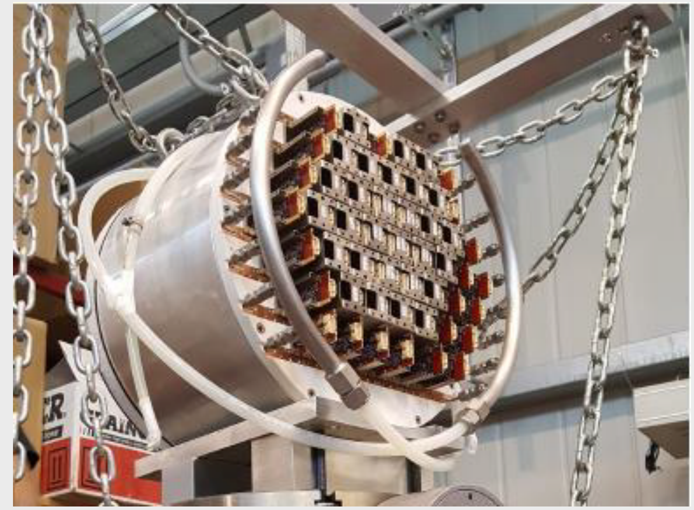
- 15 microns of Kapton (human hair = 10–80 microns)
  - ~1 meter long straws, 5 mm diameter with a central gold-plated tungsten wire
- ~20,000 straws in a vacuum of 10^{-4} Torr



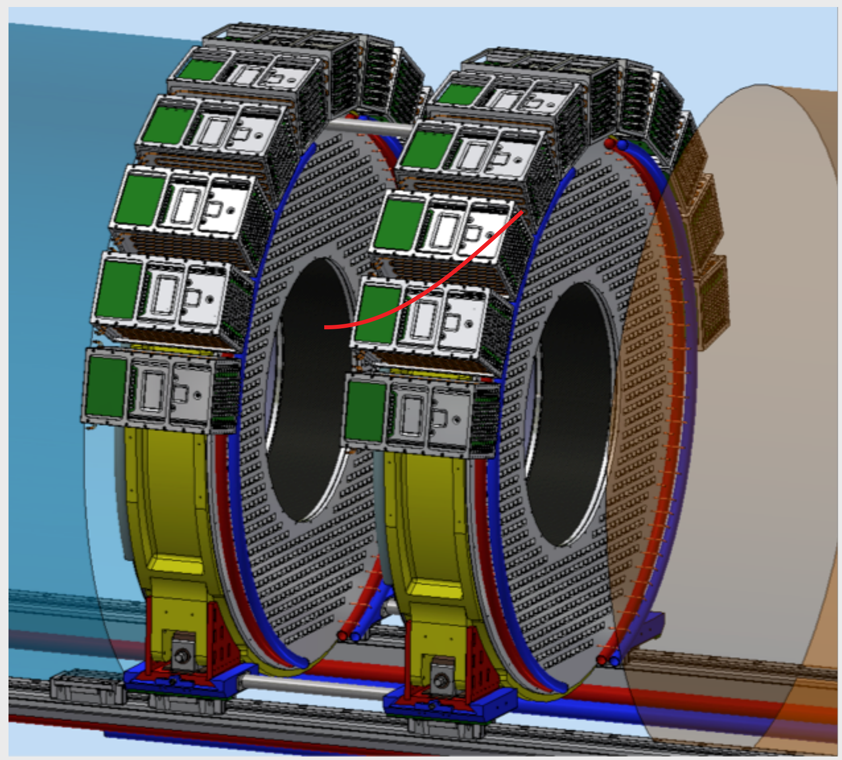
Calorimeter

- Two CsI crystal disks
  - hole, as in tracker, for passage of muon beam
  - distance chosen such that if signal electron goes down hole, hits next disk





“module 0”: 51 crystals

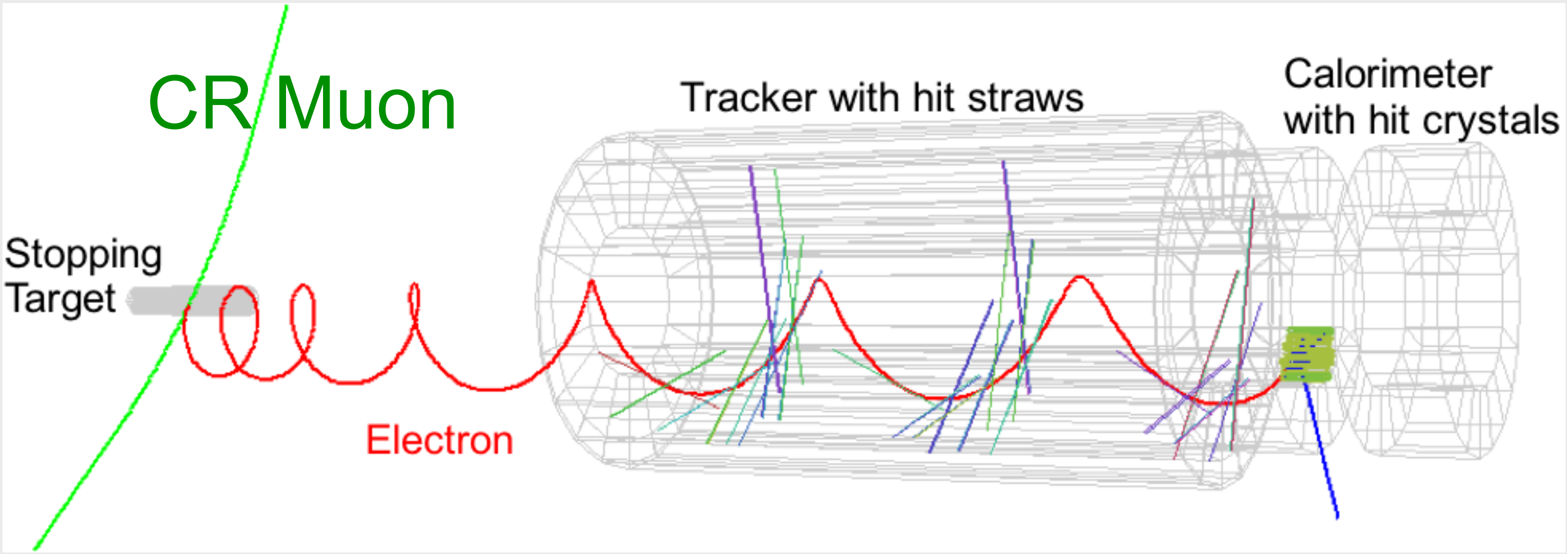


Backgrounds and Design

	source	*scales with*	*solution*
Intrinsic Backgrounds	decay-in-orbit	# of muons	detector resolution
Beam	radiative pion capture	closeness to beam pulse	pulsed beam
Running Time	cosmic ray	live time	veto system

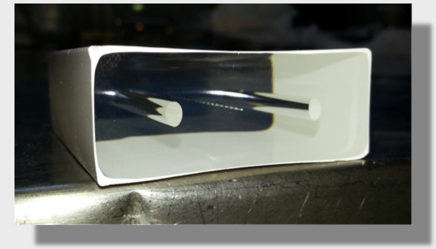
Cosmic Ray Background

- At 10^{-17} there's a lot of rare backgrounds; here's one that is surprising but not too rare: ~1/day



- the experiment needs a 99.99% efficient veto (3/4 layers)

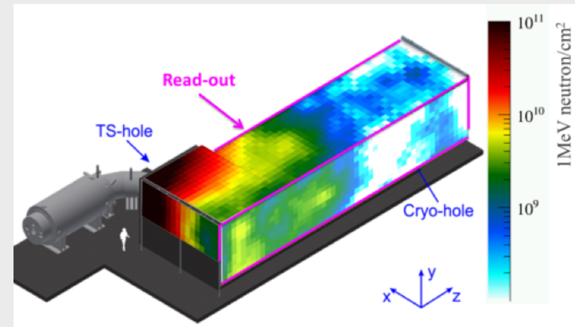
- in an environment with 10^{10} neutrons/cm^2/sec from proton pulse

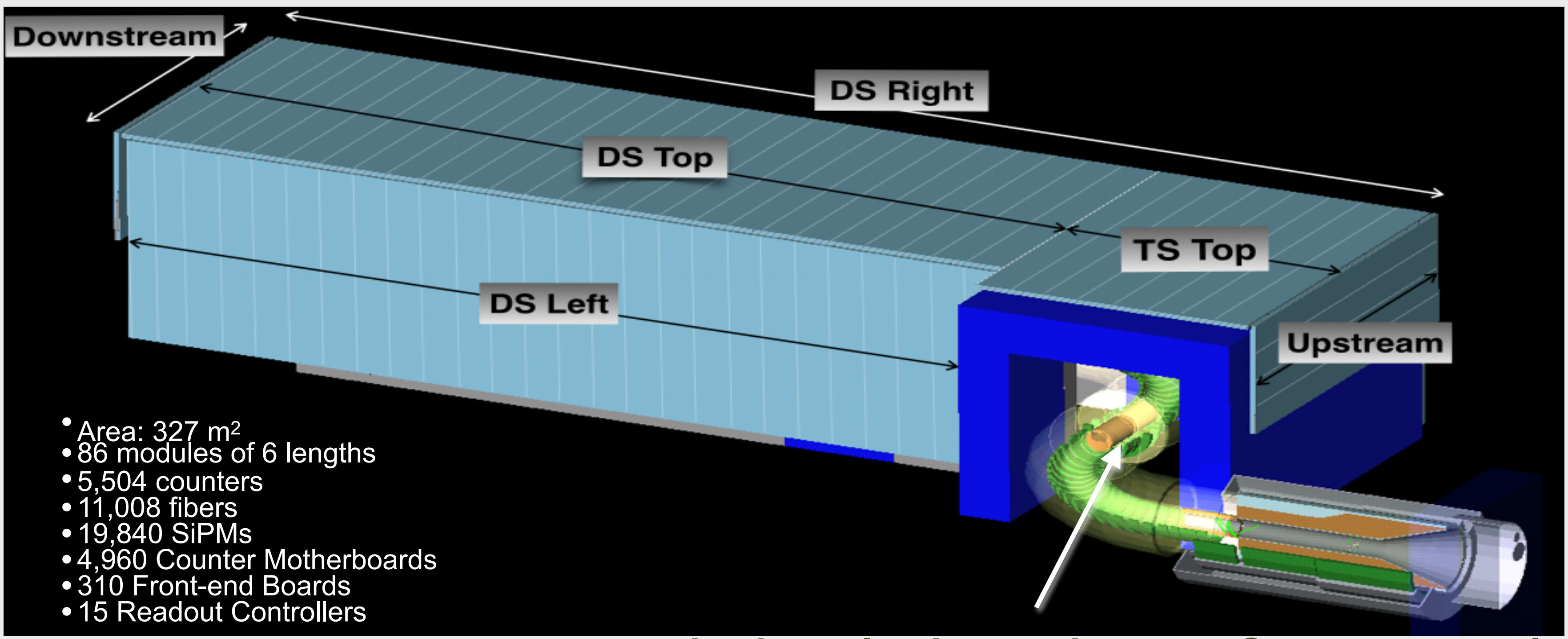


CR Veto Details

- Veto systems are never totally hermetic
  - even with 26 km of veto counters, area of a large house (~3000 ft²) (extrusions completed)
- Simulated 250 experiment live-times in targeted regions

Neutron Rates

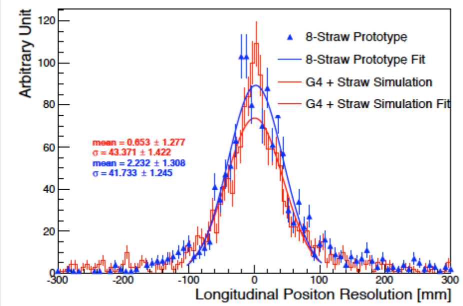




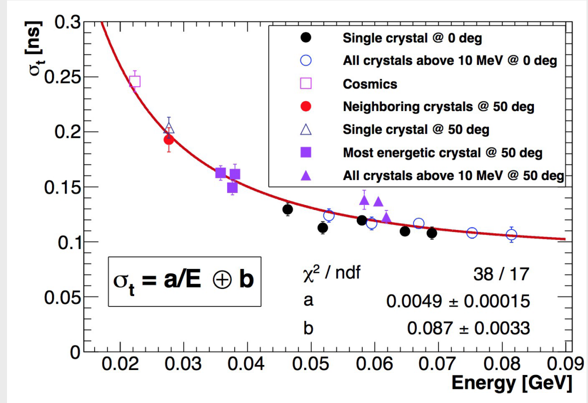
targeted simulations here, for example

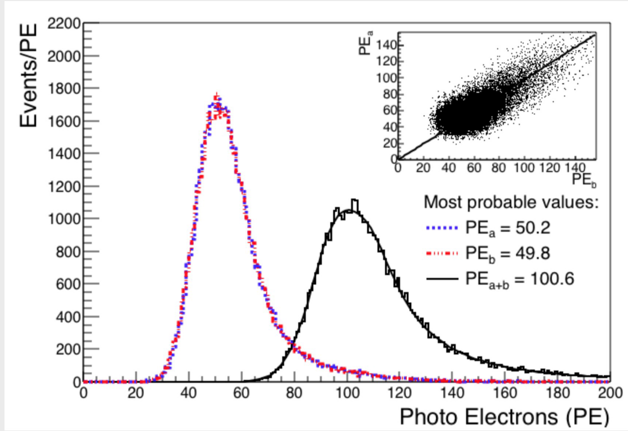
Tracker/Calorimeter/ CRV Test Beam Results

- Energy and time response of 3X3 CsI crystal array
- PE yields of CRV scintillation counters read out by SiPMs
- Tracker Resolution measured and modeled at level of charge cluster formation
- *Measured resolutions and properties in the simulations*



Longitudinal Resolution
(Data vs MC)
$\sigma_{data} = 42 \pm 1$ mm
$\sigma_{MC} = 43 \pm 1$ mm





arXiv: 1709.06587, submitted to NIM

DOI: 10.1088/1748-0221/12/05/P05007

Beam Under Construction

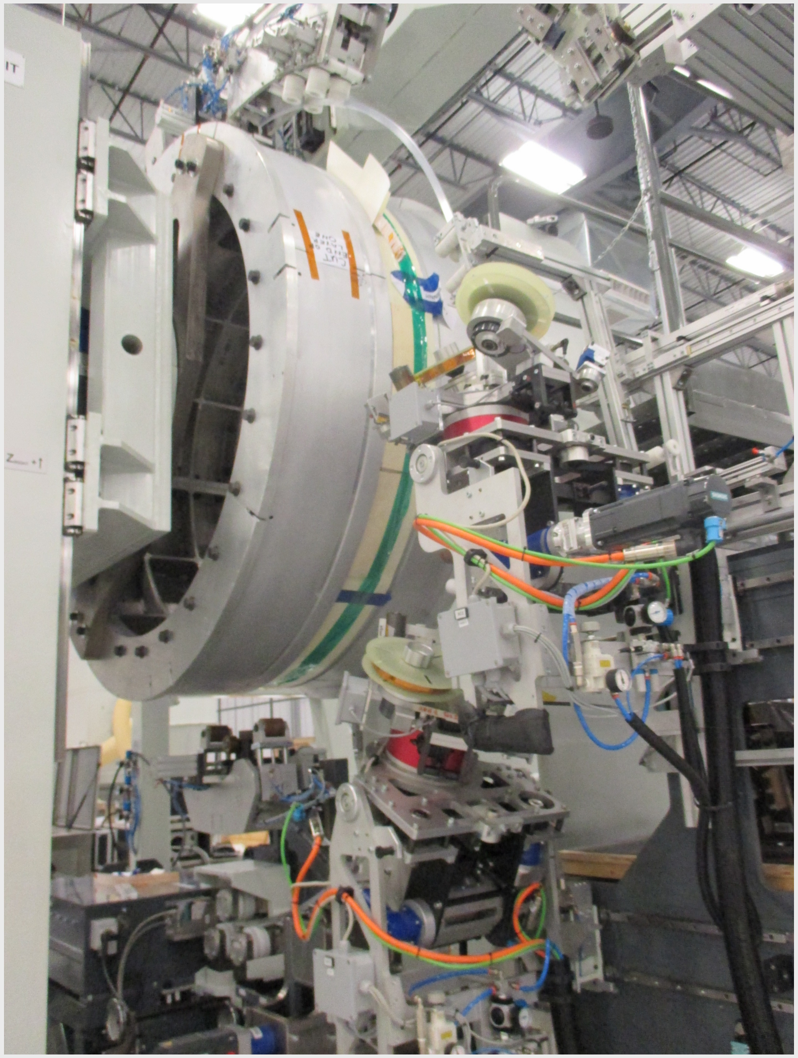
diagnostic absorber near g-2/Mu2e split: is the beam as expected and does the extinction work?

checking polarity now! (before we turn on)

Mu2e Solenoid

Coil modules at ASG

~1 of 75 km of cable and spares (cable complete)

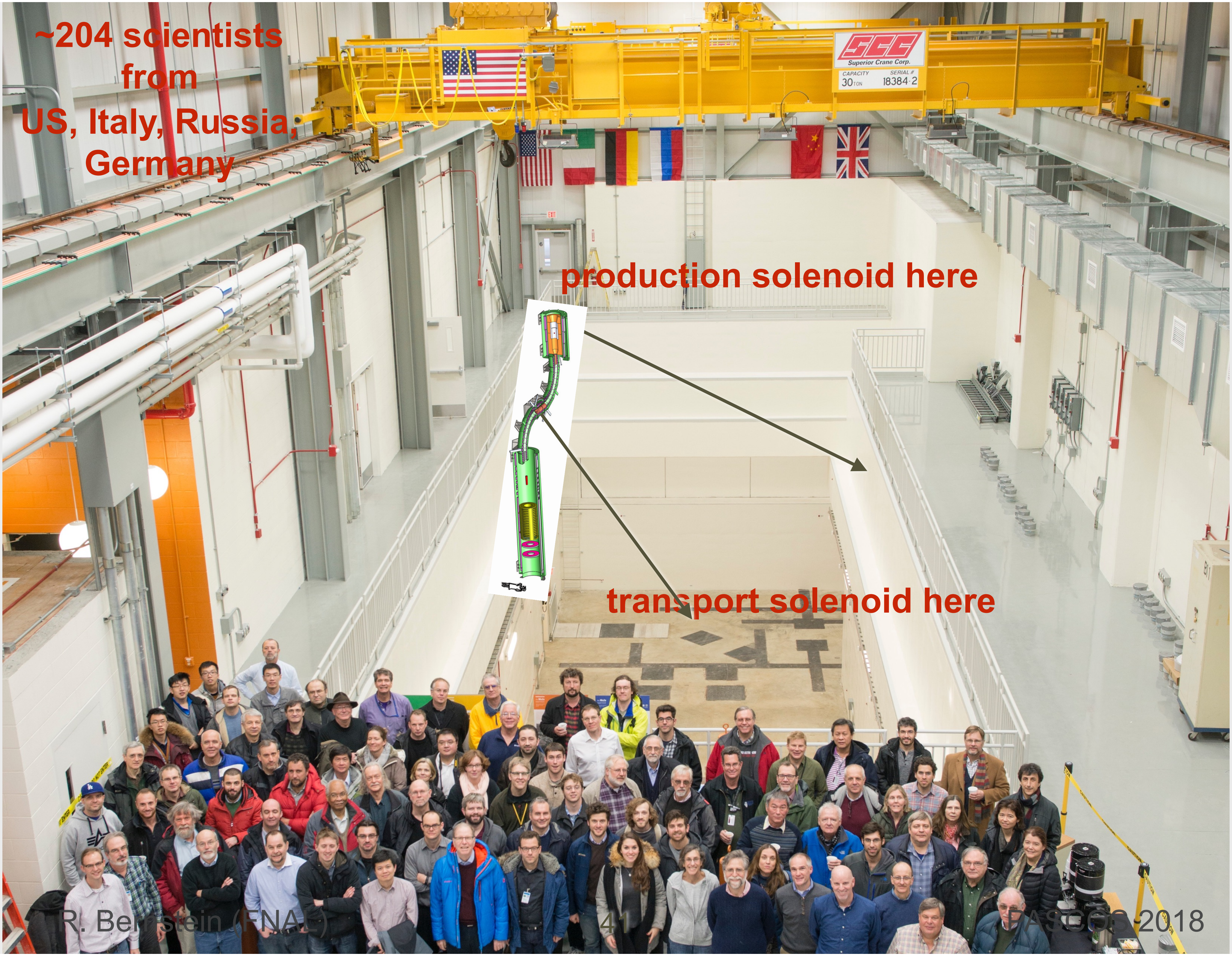


at General Atomics
splice between solenoids

Beginning Production

CDF was here

- Ramping up production and cold tests

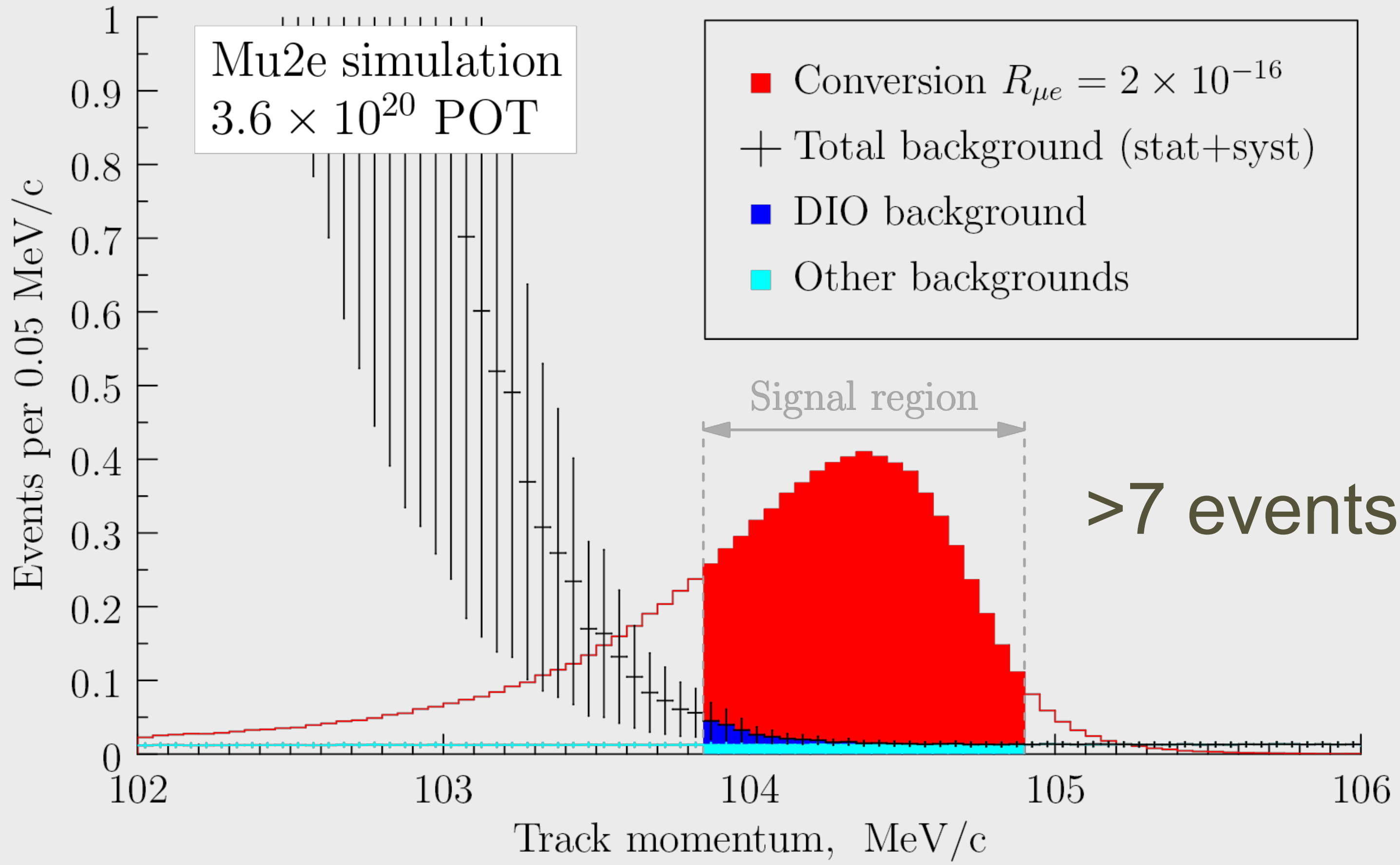

~204 scientists
from
US, Italy, Russia,
Germany
production solenoid here
transport solenoid here
Superior Crane Corp.
CAPACITY
30 TON
SERIAL #
18384-2

Mu2e Expected

90% CL ~ 8 x 10^{-17} 5σ~ 2 x 10^{-16}

x10000 better than SINDRUM-II

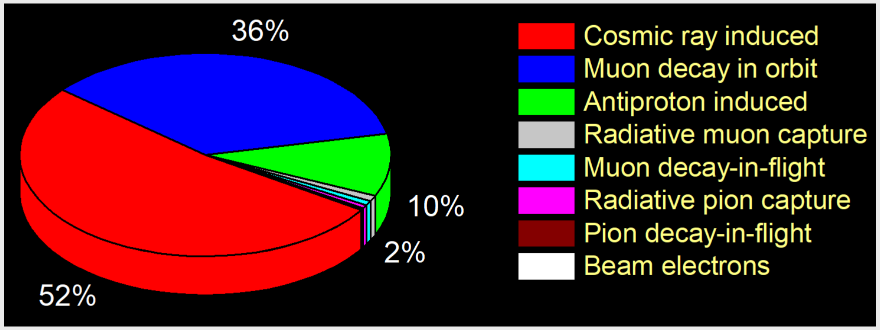
Mu2e



>7 events = 5σ

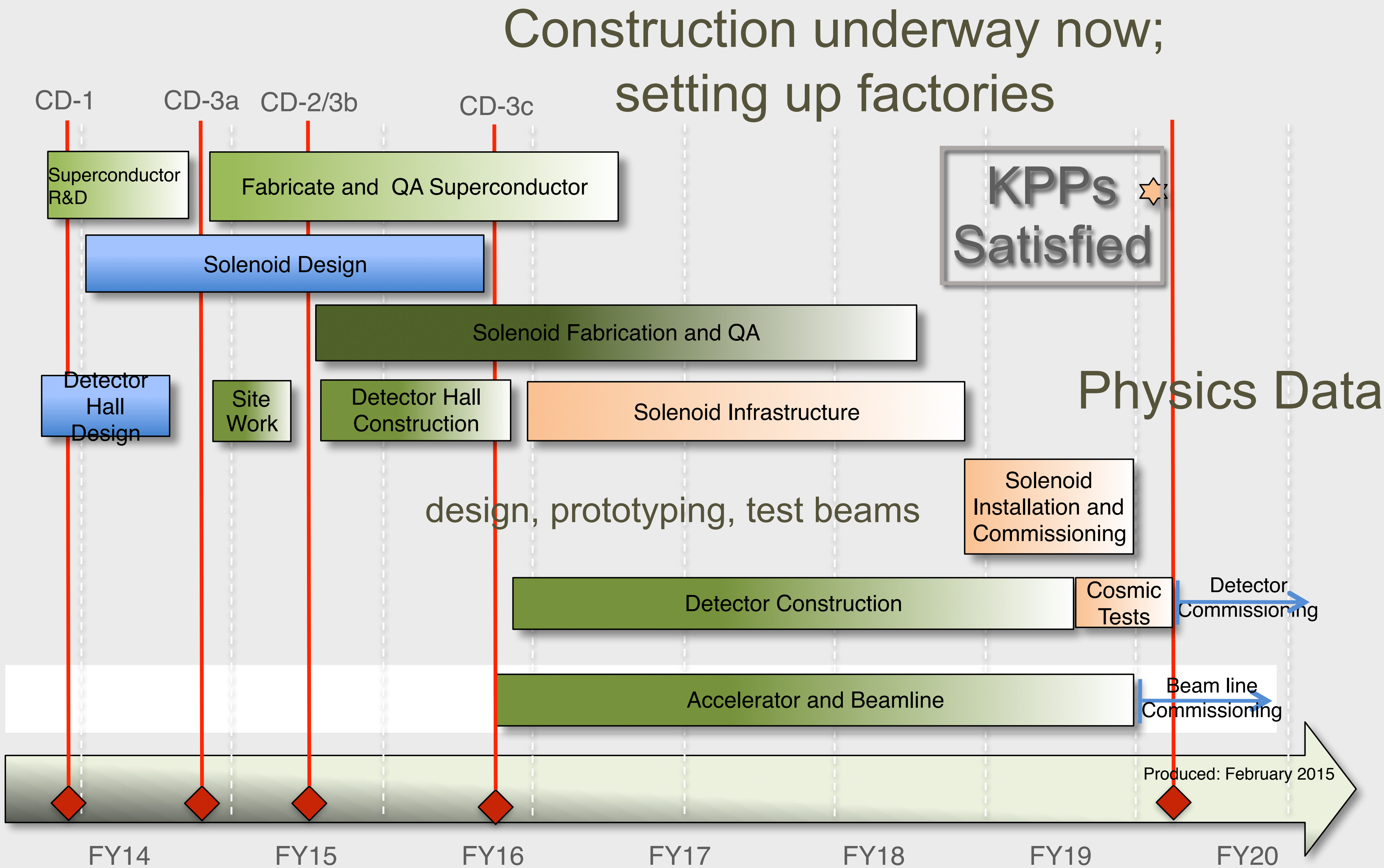
typical SUSY at 10^{-15}: 40 events vs 0.4 bkg

Background Estimates

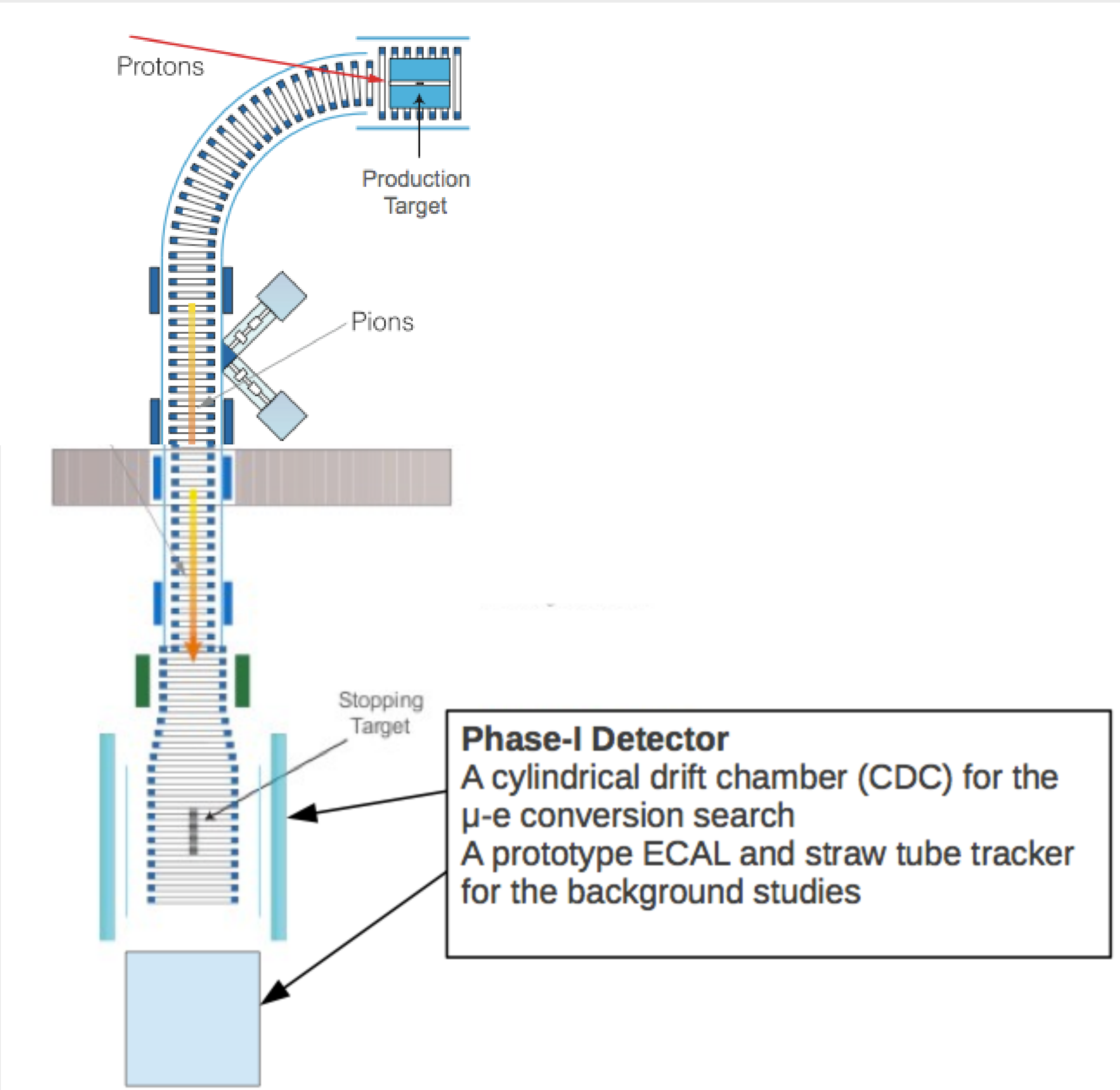


Process	Expected Number
Cosmic Ray Muons	0.25 ± 0.026
DIO	0.14 ± 0.09
RPC	0.025 ± 0.003
Antiprotons	0.047 ± 0.024
Muon DIF	< 0.003
Pion DIF	$0.001 \pm < 0.001$
Beam Electrons	$< 5 \times 10^{-4}$
Total	0.46 ± 0.10

Mu2e Schedule

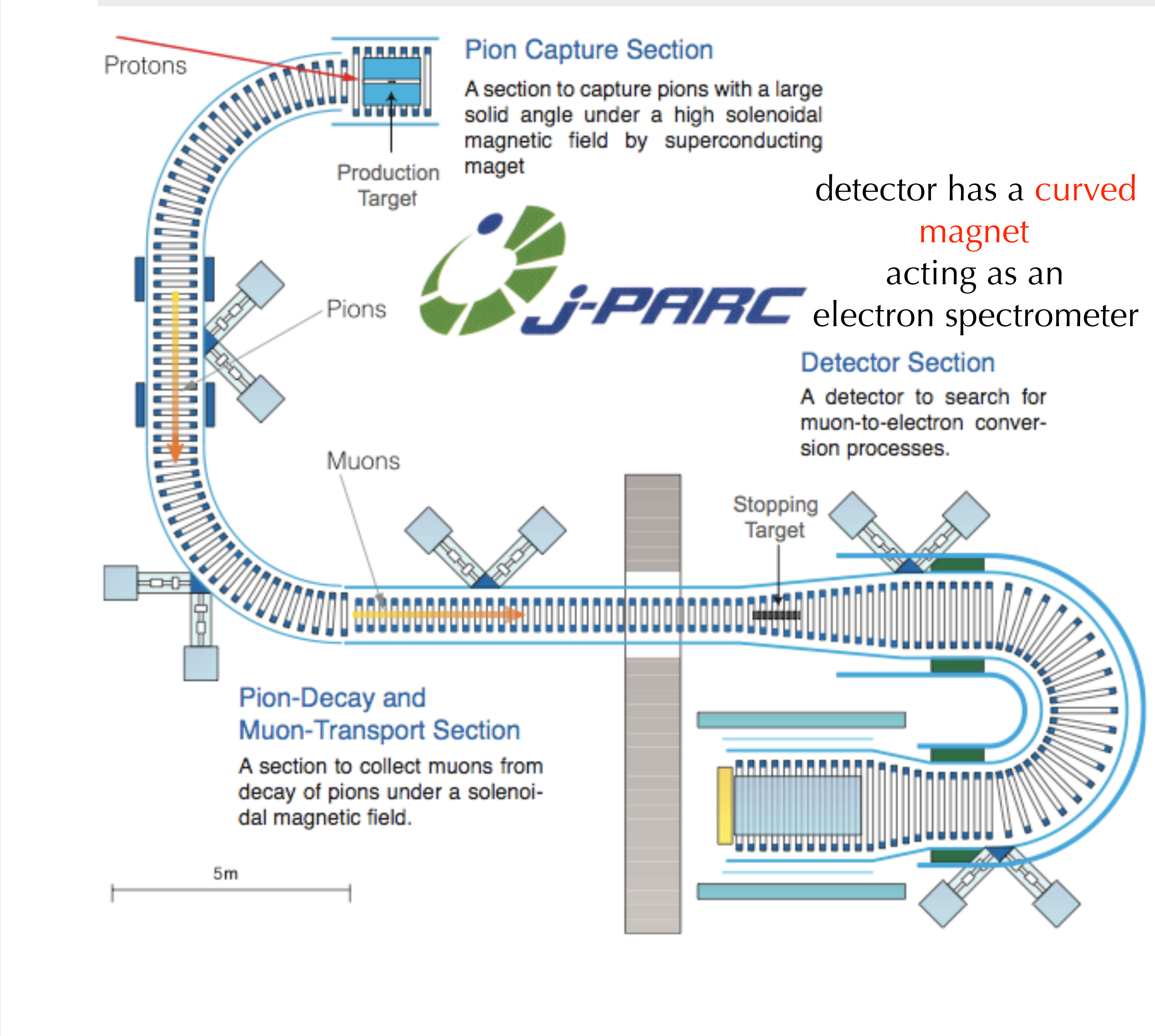


COMET



- Stage I: (2021)
  - x100 better than SINDRUM-II
- Stage II
  - x10,000 better
  - same reach as Mu2e
- Important differences:
  - C- vs S- for transport
  - 2nd bend in detector solenoid: charge/ momentum selects, (save that info for later)

COMET

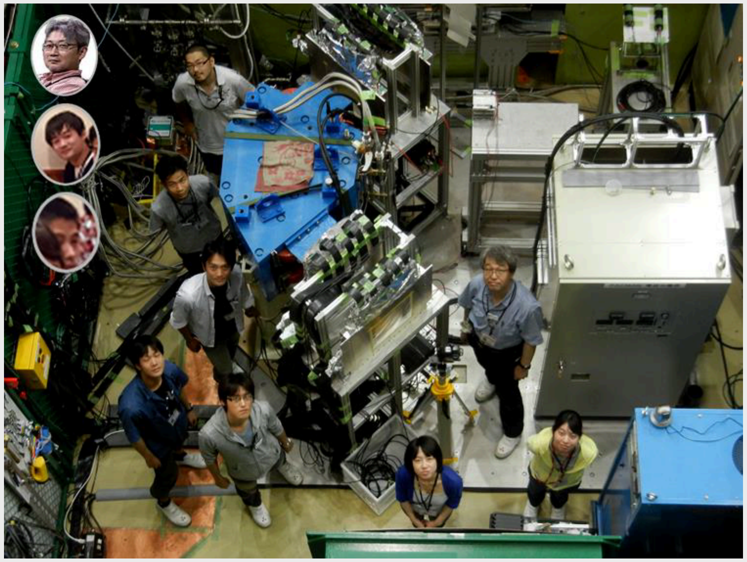


detector has a curved magnet acting as an electron spectrometer

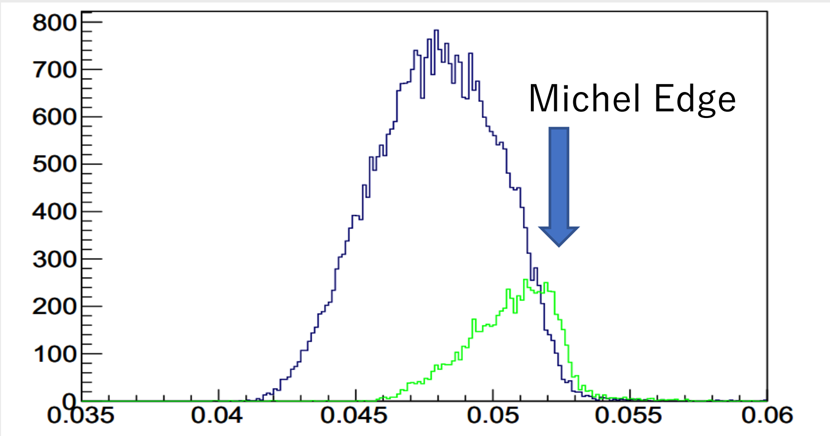
- Stage I: (2021)
  - x100 better than SINDRUM-II
- Stage II
  - x10,000 better
  - same reach as Mu2e
- Important differences:
  - C- vs S- for transport
  - 2nd bend in detector solenoid: charge/momentum selects, (save that info for later)

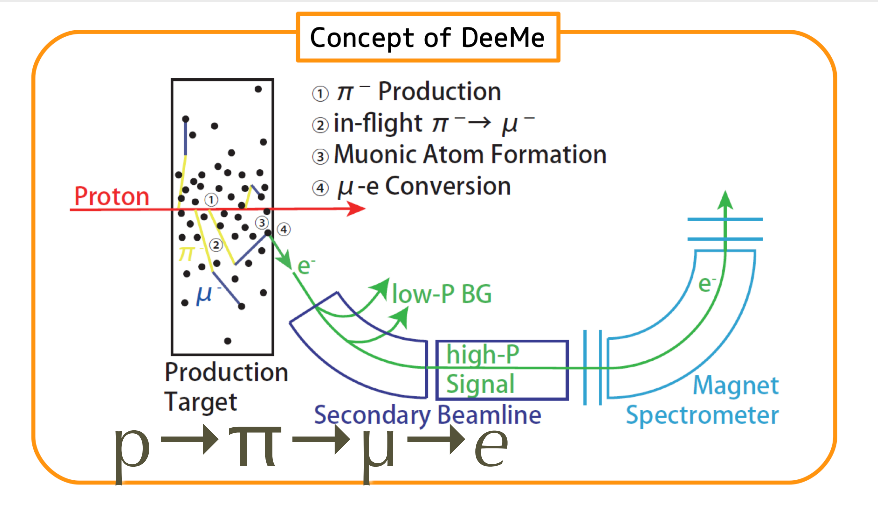
DeeMe

- DeeMe at J-PARC $\mu N \rightarrow eN$ with a 2 x 10^{-14} SES, x10 improvement
- ***production target and conversion target are the same: not like Mu2e or COMET***



Calibration Spectra (μ^+ data)

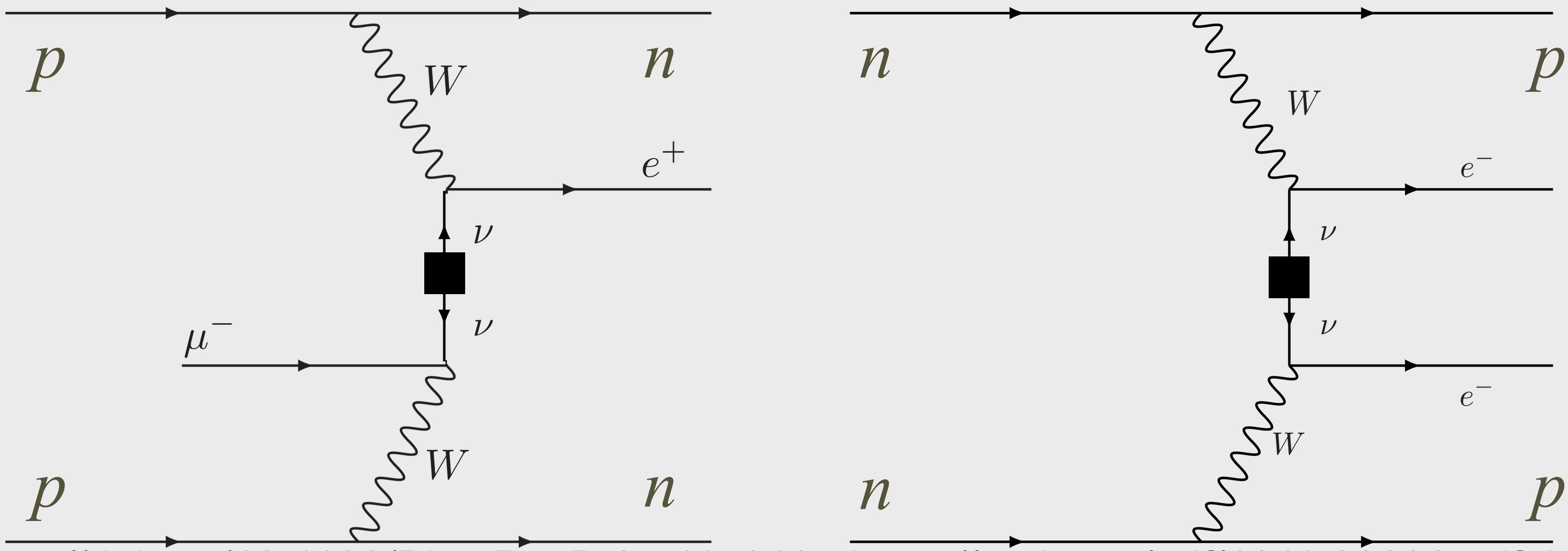




$\mu^- N \to e^+ N^*$

- Closely related to *0ν2β*
    - or a leptoquark (not shown), …

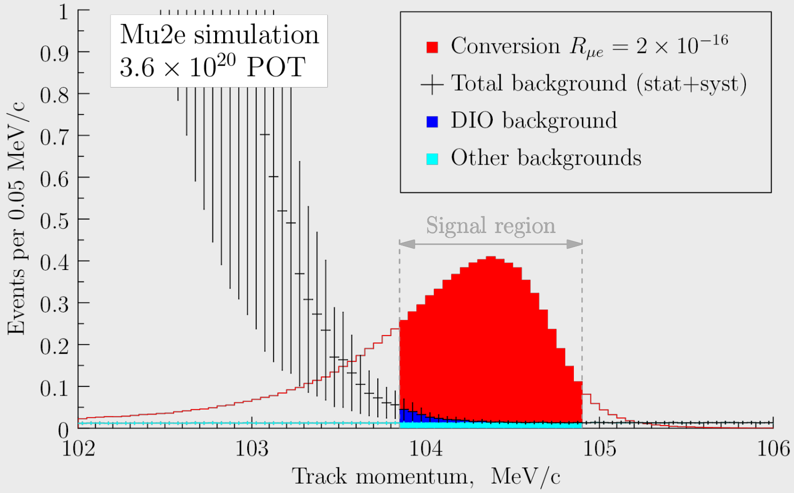
best done at Mu2e



- https://doi.org/10.1103/PhysRevD.95.115010, https://arxiv.org/pdf/1611.00032.pdf, https://arxiv.org/pdf/1705.07464.pdf, Geib et al., Physics Letters B 764 (2017) 157–162 or https://arxiv.org/abs/1609.09088 discusses relation to *0ν2β*

Mu2e Summary

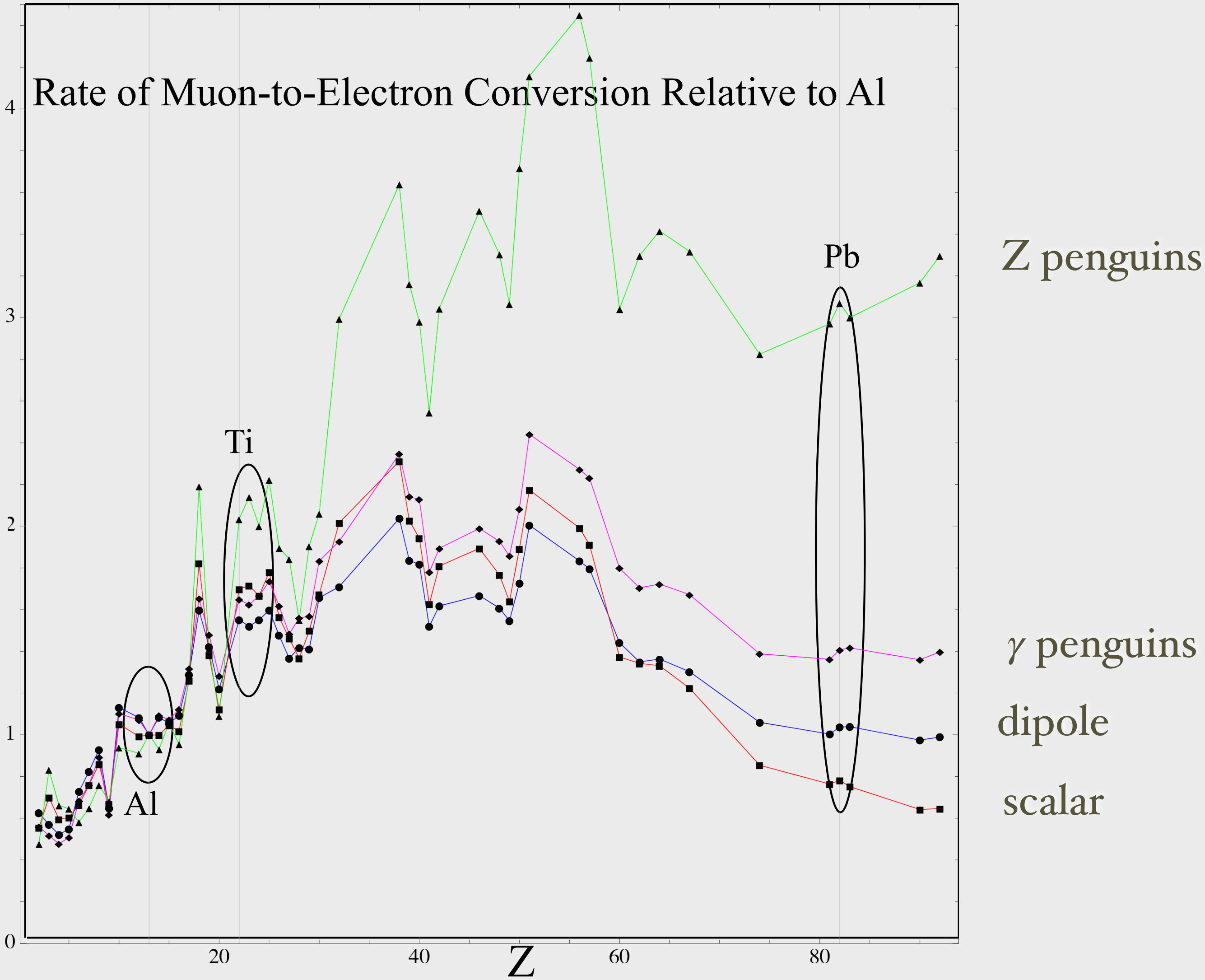
- $\mu^- Al \rightarrow e^- Al$
  - SES ~ 3×10^{-17}
  - 90% CL ~ 8×10^{-17}
  - 5σ discovery ~ 2×10^{-16}
  - Run for 3 years; physics data starting 2022
- Upgrades in intensity and changes in material already under examination



Mu2e Upgrades

- Next Step in cLFV Program:

Upgrades and Z-Dependence



5% measurement on Al/Ti needed to see split

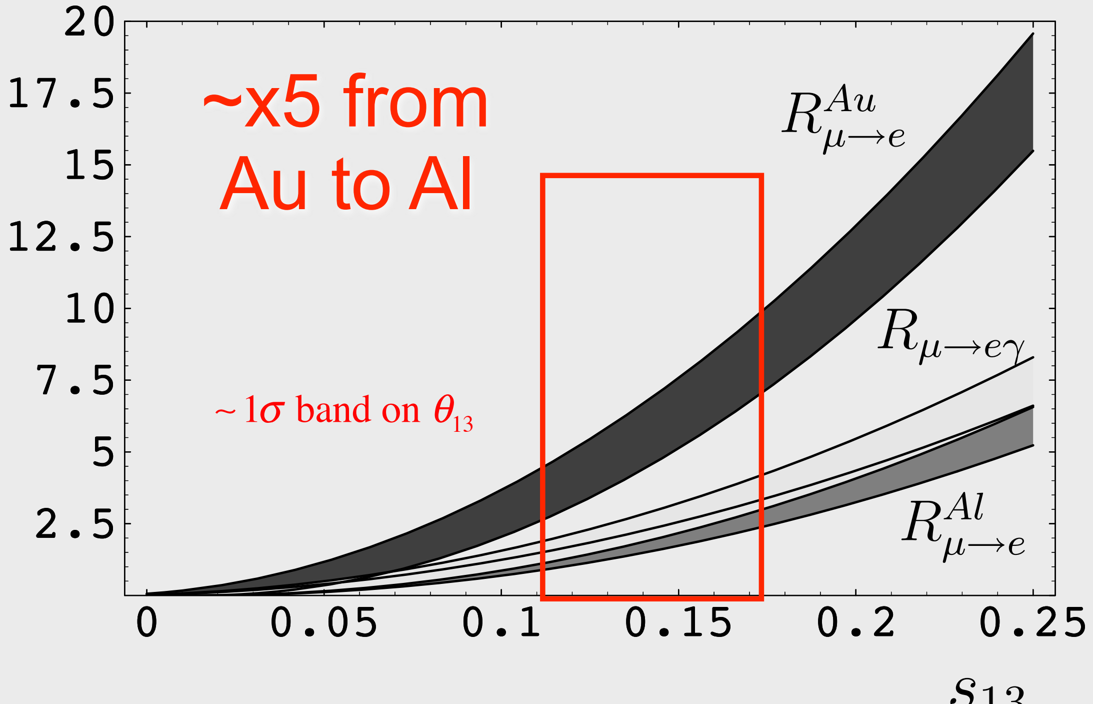
- Different Operators have different *Z*-dependence
- Combine depending on the particular model

Lepton flavor violating mu - e conversion rate for various nuclei
M. Koike et al., J.Phys. G29 (2003) 2051-2054
DOI: 10.1088/0954-3899/29/8/401

Relationship to Neutrino Oscillations (neutral lepton flavor violation)

- want as high a *Z* as possible

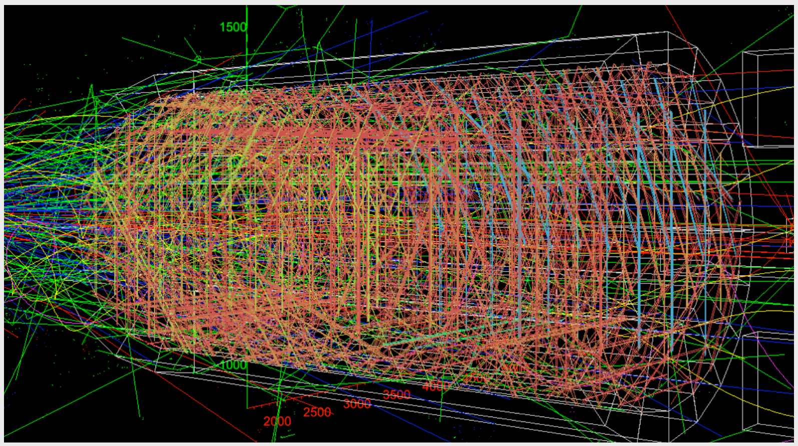
PDG 2017



V. Cirigliano, B. Grinstein, G. Isidori, M. Wise
Nucl.Phys.B728:121-134,2005

Choice of Z for Upgrade

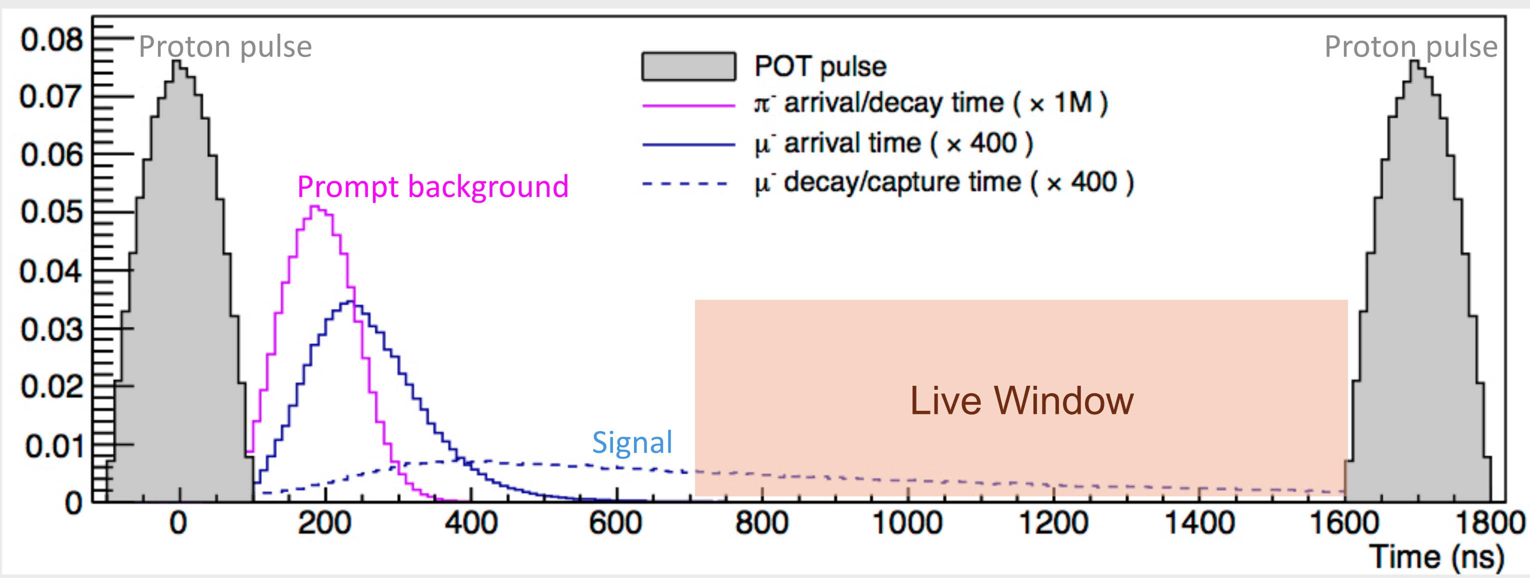
- What Sets Material Choice? Lifetime:



$$\tau_\mu(\mathrm{Al}) = 864 \text{ ns}$$

$$\tau_\mu(\mathrm{Ti}) = 338 \text{ ns}$$

$$\tau_\mu(\mathrm{Au}) = 74 \text{ ns}$$



Mu2e Upgrades

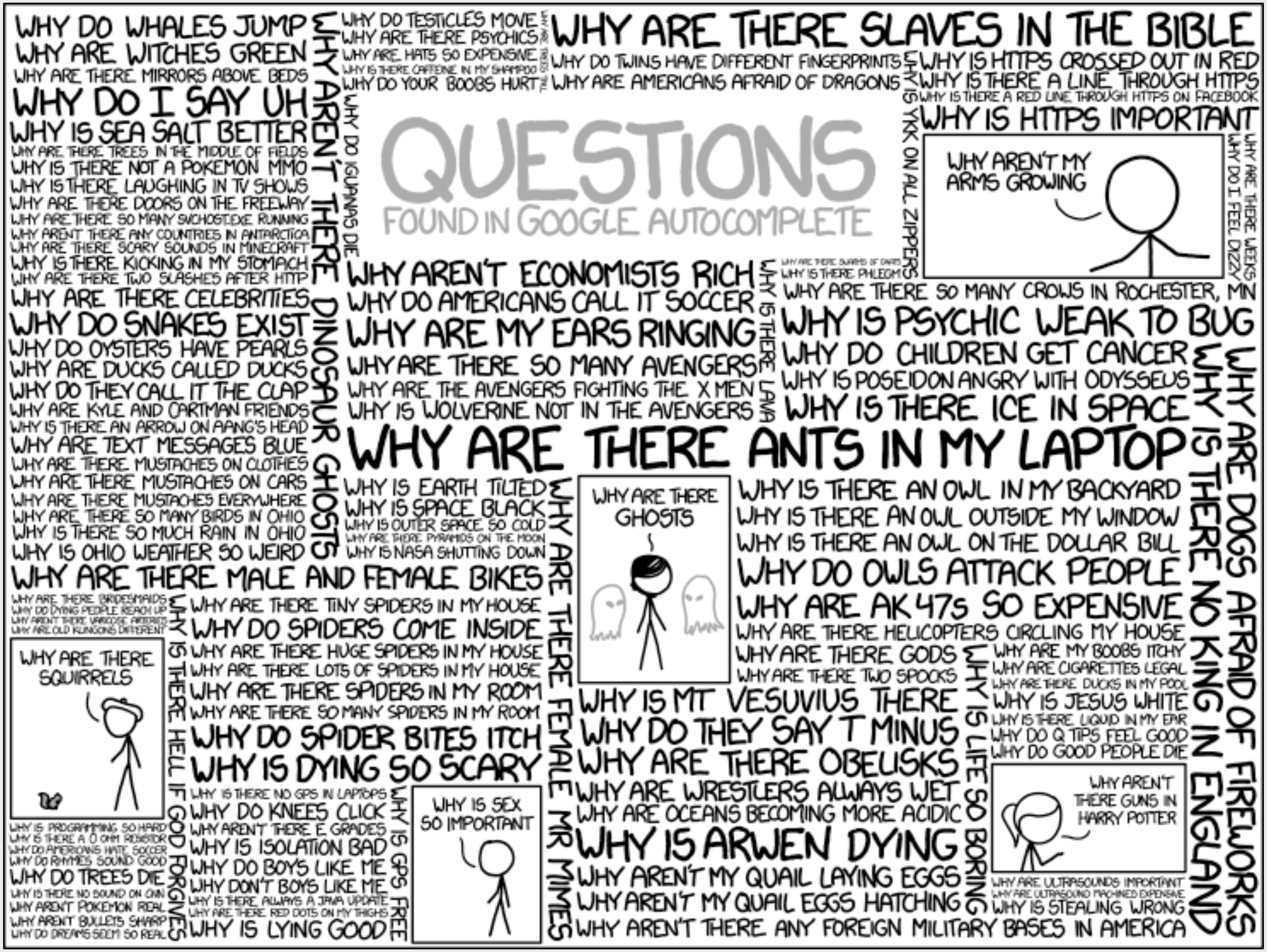
- Studies for x10 improvement with Ti look promising and will be continued; EOI written (1307.1168 and EOI at 1802.02599)
- We need detector and solenoid improvements
  - may need new production solenoid to handle lower energy beam and higher power.
- FNAL PIP-II natural for both pulsed and non-pulsed CLFV, could do $\mu^- N \rightarrow e^\pm N$, $\mu \rightarrow e \gamma$, $\mu \rightarrow 3e$, $\mu^- e^- \rightarrow e^- e^-$ at one facility

What I Hope You'll Remember

- Any signal is unambiguously new physics. We need multiple measurements to understand sort out the nature of the signal (or limits)
  - CLFV, neutrinos, g-2, EDMs, kaon decays, $0\nu 2\beta$, dark matter, collider CLFV are tightly linked and models have to fit *all* the data.
- The experiments are challenging theory and getting better fast, with upgrades on the way.
  - within next 5-10 years: muons improve by 10^4, x10 in mass scale; Run-II data; e^+e^- and LHCb will probe τ , B, and charm sectors. Kaon system continues to be incisive and unique.

Conclusions

- Mu2e will:
  - Reduce the limit for $R_{\mu e}$ by more than four orders of magnitude, x10 in mass reach ($R_{\mu e}$ < 8x10^{-17} @ 90% C.L.)
  - Discover unambiguous proof of new physics or
  - Set powerful constraints on a wide variety of models
- Mu2e will therefore both complement LHC results and independently probe up to 10^{4} TeV/c^{2}
  - COMET and DeeMe in progress; MEG upgrade underway
- Upgrades already under consideration with a new muon CLFV program being mapped out


QUESTIONS
FOUND IN GOOGLE AUTOCOMPLETE
WHY DO WHALES JUMP
WHY ARE WITCHES GREEN
WHY ARE THERE MIRRORS ABOVE BEDS
WHY DO I SAY UH
WHY IS SEA SALT BETTER
WHY ARE THERE TREES IN THE MIDDLE OF FIELDS
WHY IS THERE NOT A POKEMON MMO
WHY IS THERE LAUGHING IN TV SHOWS
WHY ARE THERE DOORS ON THE FREEWAY
WHY ARE THERE SO MANY SVCHOST.EXE RUNNING
WHY AREN'T THERE ANY COUNTRIES IN ANTARCTICA
WHY ARE THERE SCARY SOUNDS IN MINECRAFT
WHY IS THERE KICKING IN MY STOMACH
WHY ARE THERE TWO SLASHES AFTER HTTP
WHY ARE THERE CELEBRITIES
WHY DO SNAKES EXIST
WHY DO OYSTERS HAVE PEARLS
WHY ARE DUCKS CALLED DUCKS
WHY DO THEY CALL IT THE CLAP
WHY ARE KYLE AND CARTMAN FRIENDS
WHY IS THERE AN ARROW ON AANG'S HEAD
WHY ARE TEXT MESSAGES BLUE
WHY ARE THERE MUSTACHES ON CLOTHES
WHY ARE THERE MUSTACHES ON CARS
WHY ARE THERE MUSTACHES EVERYWHERE
WHY ARE THERE SO MANY BIRDS IN OHIO
WHY IS THERE SO MUCH RAIN IN OHIO
WHY IS OHIO WEATHER SO WEIRD
WHY AREN'T THERE DINOSAUR GHOSTS
WHY DO IGUANAS DIE
WHY DO TESTICLES MOVE
WHY ARE THERE PSYCHICS
WHY ARE HATS SO EXPENSIVE
WHY IS THERE CAFFEINE IN MY SHAMPOO
WHY DO YOUR BOOBS HURT
WHY ARE THERE SLAVES IN THE BIBLE
WHY DO TWINS HAVE DIFFERENT FINGERPRINTS
WHY ARE AMERICANS AFRAID OF DRAGONS
WHY IS HTTPS CROSSED OUT IN RED
WHY IS THERE A LINE THROUGH HTTPS
WHY IS THERE A RED LINE THROUGH HTTPS ON FACEBOOK
WHY IS HTTPS IMPORTANT
WHY IS YKK ON ALL ZIPPERS
WHY AREN'T MY ARMS GROWING
WHY ARE THERE WEEKS
WHY DO I FEEL DIZZY
WHY AREN'T ECONOMISTS RICH
WHY DO AMERICANS CALL IT SOCCER
WHY ARE MY EARS RINGING
WHY ARE THERE SO MANY AVENGERS
WHY ARE THE AVENGERS FIGHTING THE X MEN
WHY IS WOLVERINE NOT IN THE AVENGERS
WHY IS THERE LAVA
WHY IS THERE PHLEGM
WHY ARE THERE SO MANY CROWS IN ROCHESTER, MN
WHY IS PSYCHIC WEAK TO BUG
WHY DO CHILDREN GET CANCER
WHY IS POSEIDON ANGRY WITH ODYSSEUS
WHY IS THERE ICE IN SPACE
WHY ARE THERE ANTS IN MY LAPTOP
WHY ARE DOGS AFRAID OF FIREWORKS
WHY IS THERE NO KING IN ENGLAND
WHY IS EARTH TILTED
WHY IS SPACE BLACK
WHY IS OUTER SPACE SO COLD
WHY ARE THERE PYRAMIDS ON THE MOON
WHY IS NASA SHUTTING DOWN
WHY ARE THERE FEMALE MR MIMES
WHY ARE THERE GHOSTS
WHY IS THERE AN OWL IN MY BACKYARD
WHY IS THERE AN OWL OUTSIDE MY WINDOW
WHY IS THERE AN OWL ON THE DOLLAR BILL
WHY DO OWLS ATTACK PEOPLE
WHY ARE AK47s SO EXPENSIVE
WHY ARE THERE HELICOPTERS CIRCLING MY HOUSE
WHY ARE THERE GODS
WHY ARE THERE TWO SPOCKS
WHY IS LIFE SO BORING
WHY ARE MY BOOBS ITCHY
WHY ARE CIGARETTES LEGAL
WHY ARE THERE DUCKS IN MY POOL
WHY IS JESUS WHITE
WHY IS THERE LIQUID IN MY EAR
WHY DO Q TIPS FEEL GOOD
WHY DO GOOD PEOPLE DIE
WHY AREN'T THERE GUNS IN HARRY POTTER
WHY ARE ULTRASOUNDS IMPORTANT
WHY ARE ULTRASOUND MACHINES EXPENSIVE
WHY IS STEALING WRONG
WHY ARE THERE MALE AND FEMALE BIKES
WHY ARE THERE BRIDESMAIDS
WHY DO DYING PEOPLE REACH UP
WHY AREN'T THERE VARICOSE ARTERIES
WHY ARE OLD KLINGONS DIFFERENT
WHY ARE THERE SQUIRRELS
WHY IS THERE HELL IF GOD FORGIVES
WHY ARE THERE TINY SPIDERS IN MY HOUSE
WHY DO SPIDERS COME INSIDE
WHY ARE THERE HUGE SPIDERS IN MY HOUSE
WHY ARE THERE LOTS OF SPIDERS IN MY HOUSE
WHY ARE THERE SPIDERS IN MY ROOM
WHY ARE THERE SO MANY SPIDERS IN MY ROOM
WHY DO SPIDER BITES ITCH
WHY IS DYING SO SCARY
WHY IS MT VESUVIUS THERE
WHY DO THEY SAY T MINUS
WHY ARE THERE OBELISKS
WHY ARE WRESTLERS ALWAYS WET
WHY ARE OCEANS BECOMING MORE ACIDIC
WHY IS ARWEN DYING
WHY AREN'T MY QUAIL LAYING EGGS
WHY AREN'T MY QUAIL EGGS HATCHING
WHY AREN'T THERE ANY FOREIGN MILITARY BASES IN AMERICA
WHY IS PROGRAMMING SO HARD
WHY IS THERE A 0 OHM RESISTOR
WHY DO AMERICANS HATE SOCCER
WHY DO RHYMES SOUND GOOD
WHY DO TREES DIE
WHY IS THERE NO SOUND ON CNN
WHY AREN'T POKEMON REAL
WHY AREN'T BULLETS SHARP
WHY DO DREAMS SEEM SO REAL
WHY IS THERE NO GPS IN LAPTOPS
WHY DO KNEES CLICK
WHY AREN'T THERE E GRADES
WHY IS ISOLATION BAD
WHY DO BOYS LIKE ME
WHY DON'T BOYS LIKE ME
WHY IS THERE ALWAYS A JAVA UPDATE
WHY ARE THERE RED DOTS ON MY THIGHS
WHY IS LYING GOOD
WHY IS GPS FREE
WHY IS SEX SO IMPORTANT

NOvA
and
miniBooNE
at Neutrino 2018

(I am a NOvA collaborator but not miniBooNE)

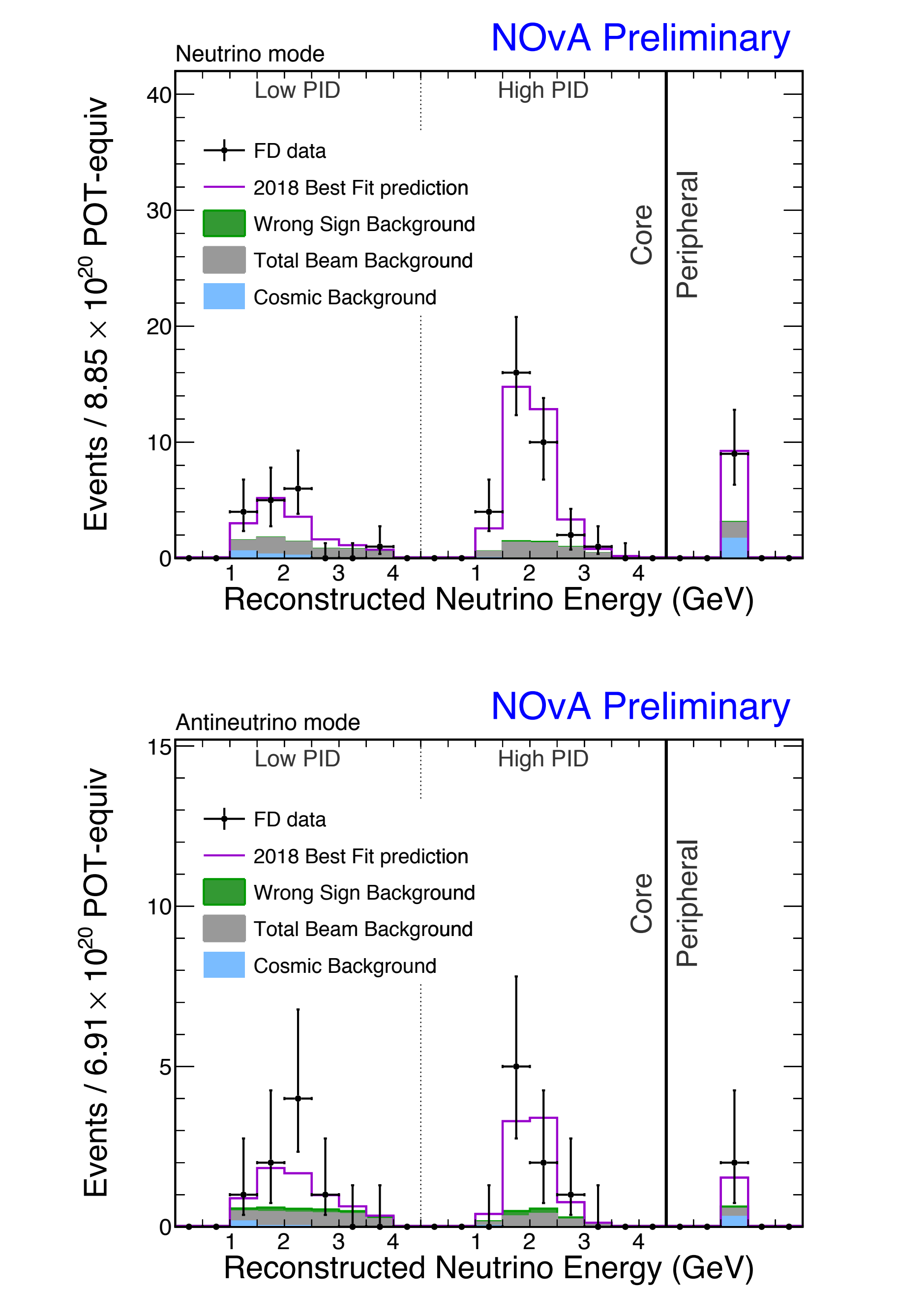
Summary of NOvA Results

- over 4σ ν_e appearance
- analyzed our first $\bar{\nu}$ data and in a joint analysis (ν_μ+ ν_e and $\bar{\nu}_\mu$+ $\bar{\nu}_e$) we see:
  - a preference for the normal hierarchy at 1.8σ, exclusion of IH at the $\delta_{CP} = \pi/2$ point at > 3σ
- reject maximal mixing at 1.8σ and prefer the upper octant solution at similar level

ELECTRON NEUTRINO AND ANTINEUTRINO APPEARANCE

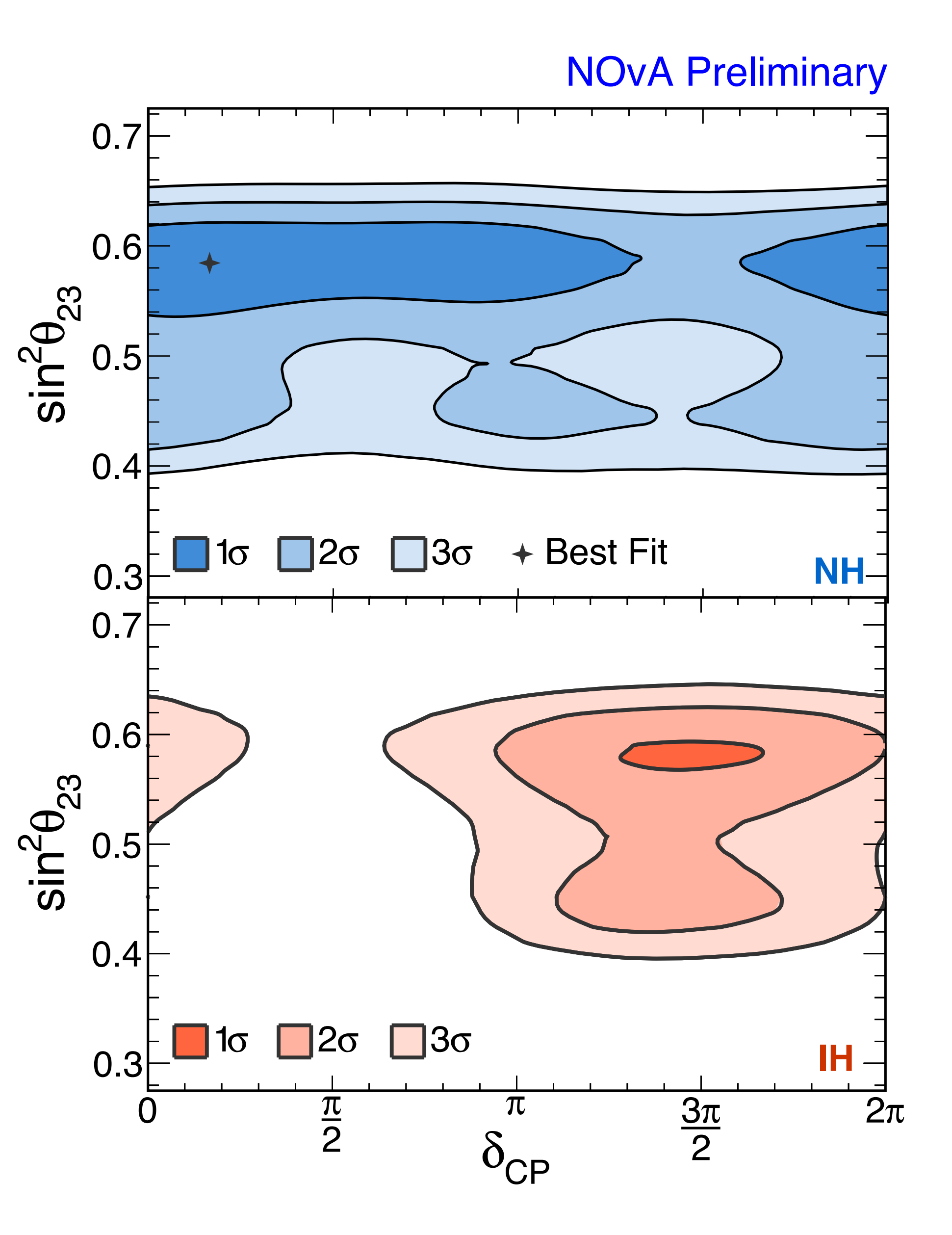
- On the neutrino beam we observe 58 events and expect 15 background interactions:
  - 11 beam, 3 cosmic background and < 1 wrong sign background.
- For the antineutrino beam we observe 18 and expect 5.3 background interactions:
  - 3.5 beam background, < 1 cosmic background and 1 wrong sign background.

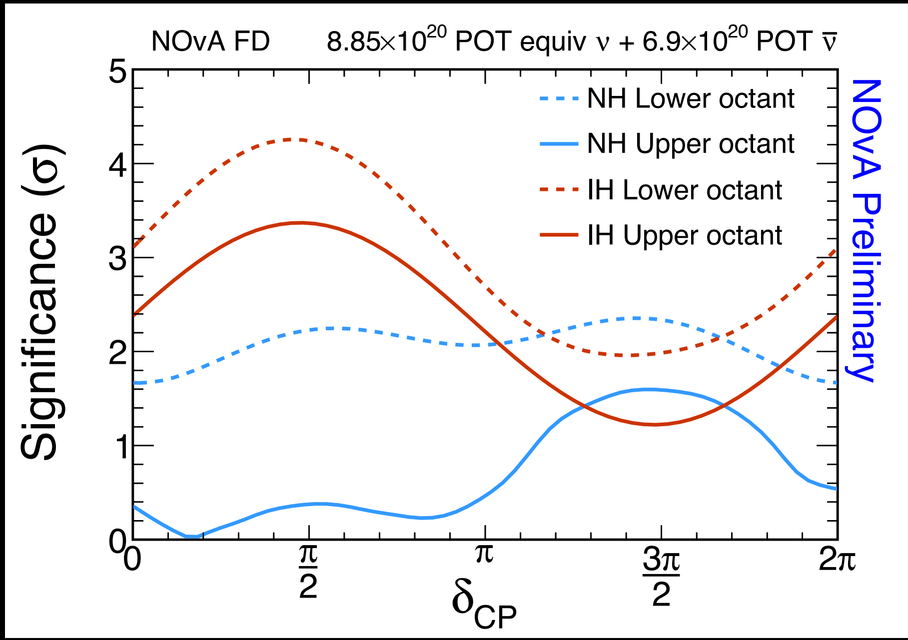
> 4σ evidence of electron antineutrino appearance



see poster #81

ALLOWED OSCILLATION PARAMETERS



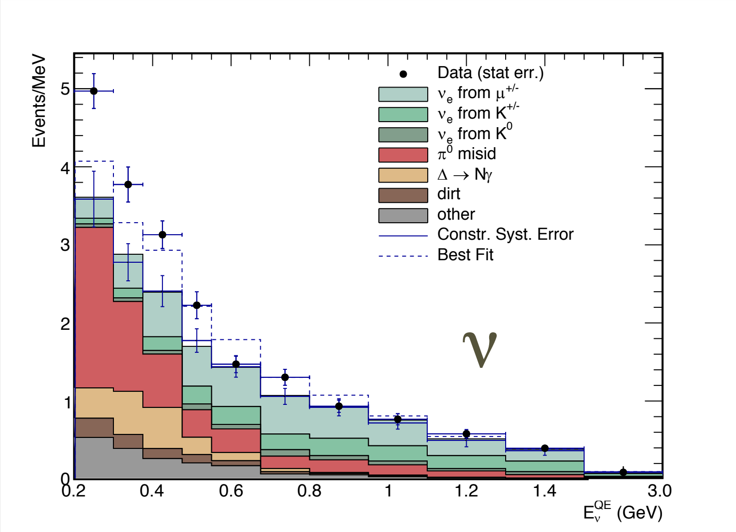


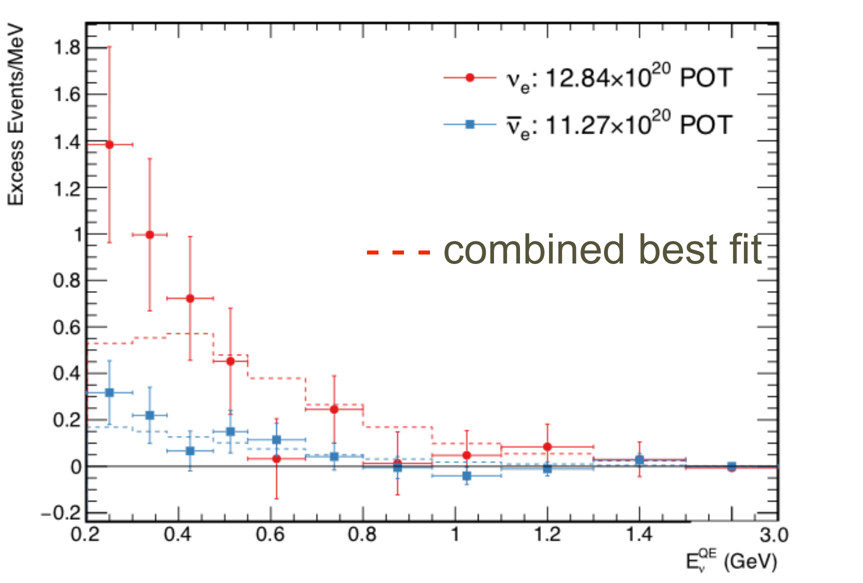
- Best fit: Normal Hierarchy
  δ_{CP}= 0.17π
  $\sin^2\theta_{23}$ = 0.58±0.03 (UO)
  $\Delta m^2_{32} = (2.51^{+0.12}_{-0.08})\cdot 10^{-3}$ eV^2

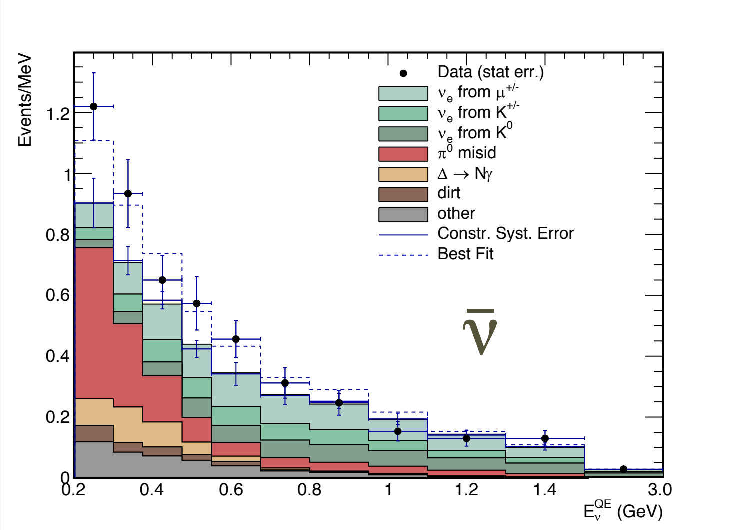
Prefer NH by 1.8σ
Exclude δ=π/2 in the IH at > 3σ

miniBooNE

- Excess of events at low energy
- Combined with LSND (3.8σ) total significance 6.1σ





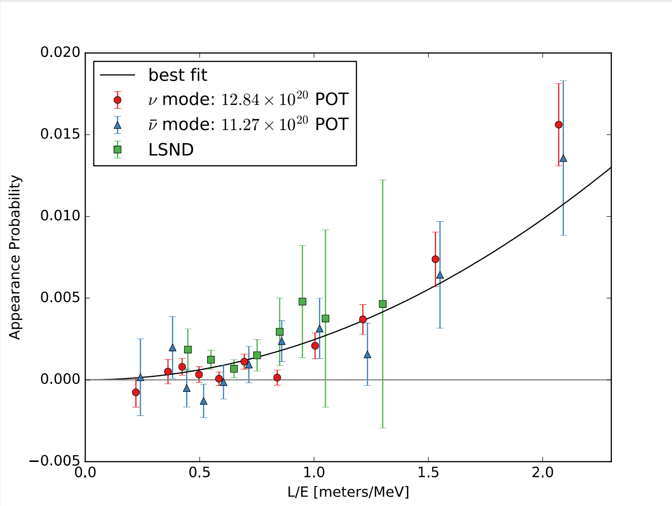


miniBooNE

R. T. Thornton – Recent Results from MiniBooNE

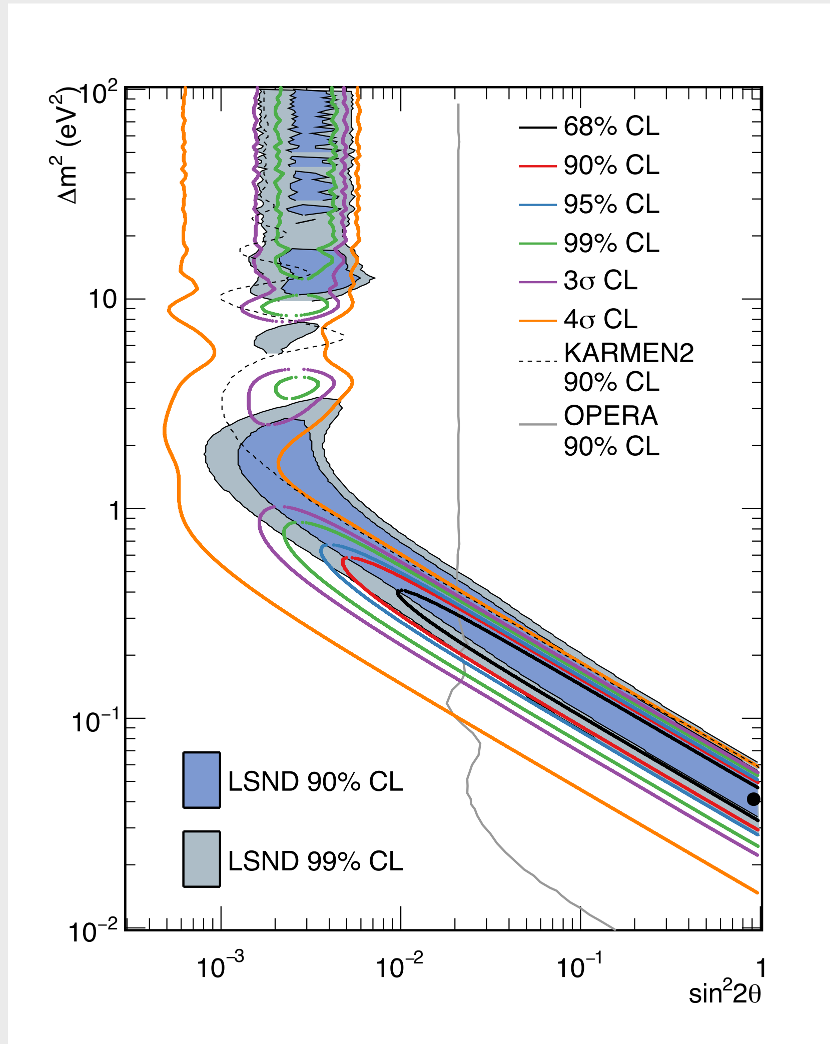
Parameter Space

$\overline{\nu}_\mu \rightarrow \overline{\nu}_e$ or $\nu_\mu \rightarrow \nu_e$



- MiniBooNE ν and $\bar{\nu}$ are **CONSISTENT** with LSND in L/E and appearance probability
- Simple 2ν fit
  - Best fit at maximum mixing
  - But 1σ region is large
  - Hints at more complicated model

- For more information
  - arxiv:1805.12028
  - July 27 W&C Talk

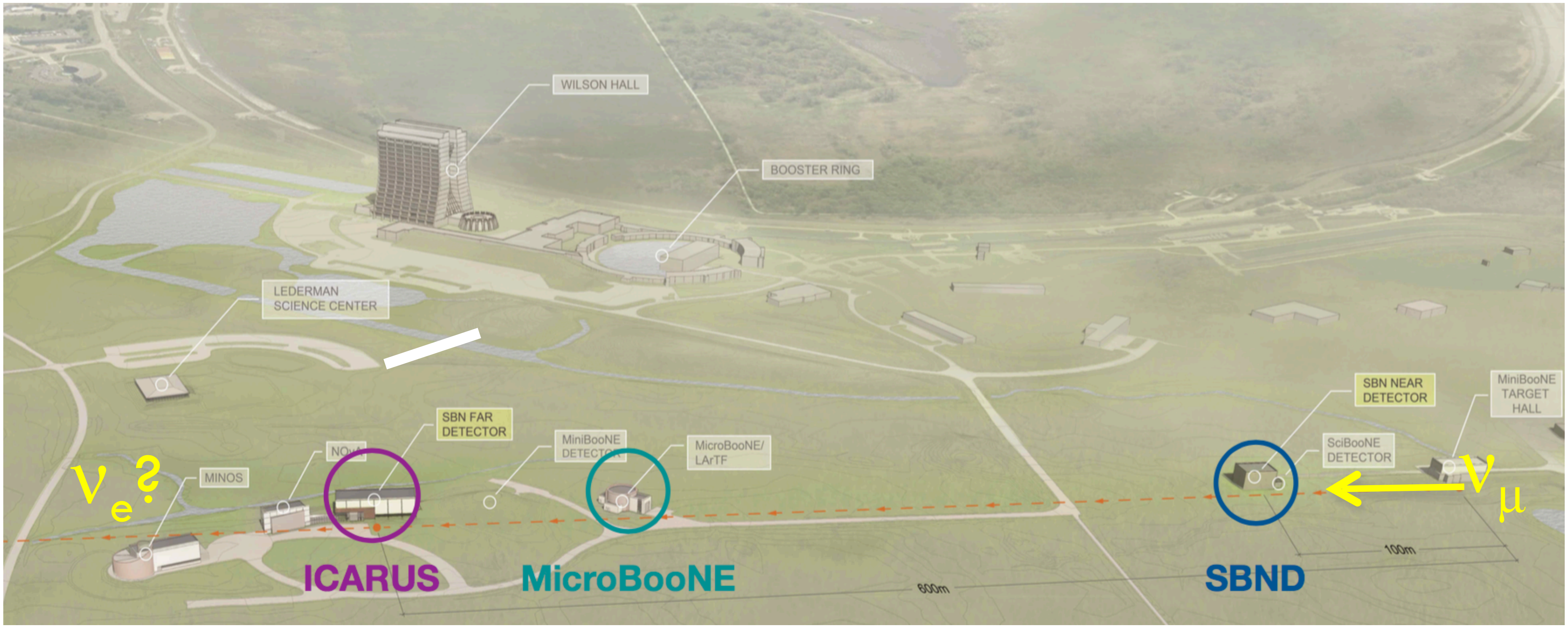


$$(\Delta m^2, \sin^2 2\theta) = (0.041\ \mathrm{eV}^2, 0.958)$$
$$\chi^2/ndf = 19.5/15.4\ (\text{prob.} = 20.1\%)$$

note $\Delta m^2 > .01$: difficult to combine with solar, atmo. at $\Delta m^2 \sim 10^{-3}$ or less, leading to sterile?

R. T. Thornton - Recent Results from MiniBooNE

Short-Baseline Neutrino (SBN) Program

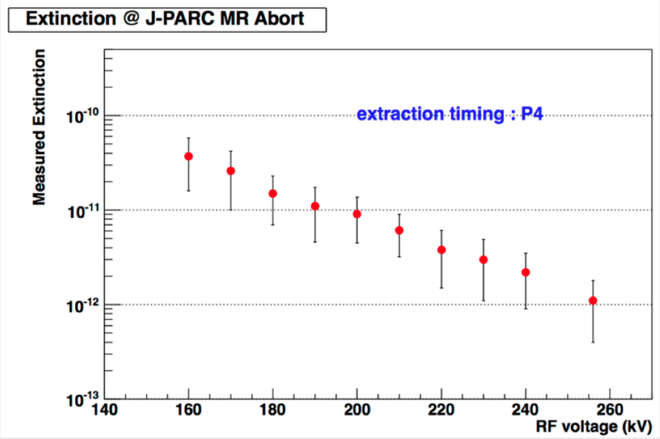


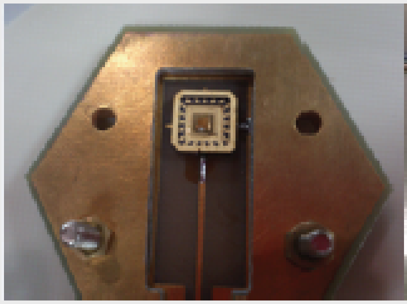
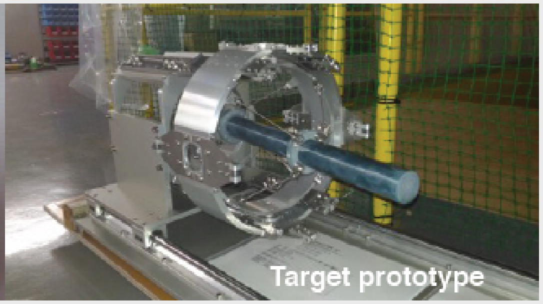
- **MicroBooNE**: what is the MiniBooNE excess? Is it electrons or photons? (Eberly, this conf.)
- **ICARUS + MicroBooNE + SBND**: Is there definitive evidence for sterile neutrinos in the currently allowed parameter space?

COMET Progress

Chen Wu, Tau 2016
and
Y. Kuno, priv. comm

- Extinction measured ~ 10^{-11}
- Prototype Target
- Coil winding for capture solenoid
- Transport Solenoid in Hall
- 14 Coils of Detector Solenoid wound



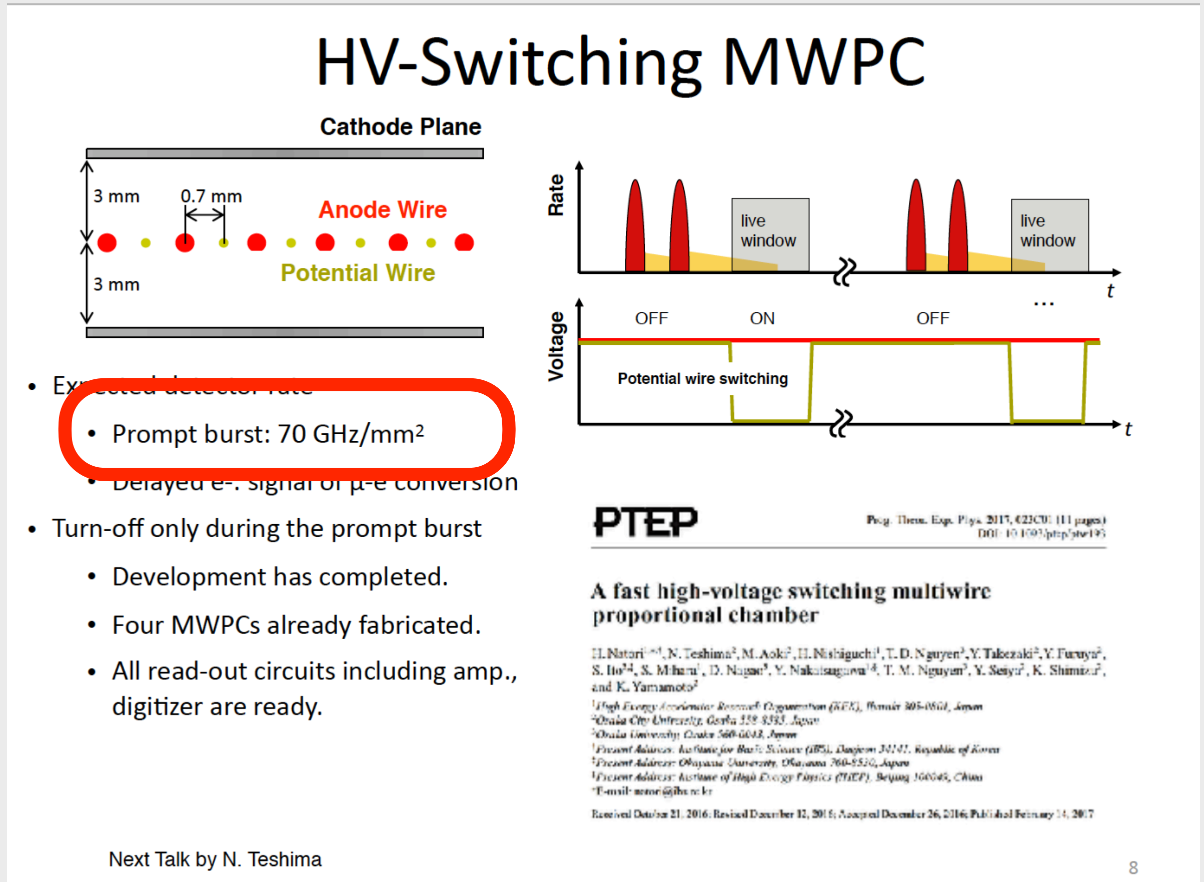
COMET Remaining

- Production Solenoid cable procured, but one has to then design a solenoid shielding system that will withstand heat and radiation
    - dumping up to 56 kW (Stage II) in a superconducting solenoid
    - design still underway
    - then needs construction
- Negotiation over running time with T2K
    - ~150 days live-time needed for Stage 1 (x100 improvement)
- Available schedules out-of-date, but Stage 1 ~ 2021

http://j-parc.jp/researcher/Hadron/en/pac_1707/PAC24thMinutes_final_draft.pdf

DeeMe HV Switching

- Since the detector is so “close” to the production target, the detector has to be turned off near the beam flash
- HV switching on a MWPC



DeeMe Status

- Depends on J-PARC H-line construction (different from COMET, DeeMe in Materials Science section)

Summary

- There is a competitive merit of physics in searching for μ-e conversion at sensitivity of 10^{-14} in timely manner.
- **It is important to maximize the potential of major discovery at J-PARC**.
- DeeMe, yet another mu-e conversion search with totally different method from COMET and Mu2e, creates harmonious diversity for J-PARC.
- DeeMe has already acquired **Stage-2 Approval from muon-PAC of KEK/IMSS**.
- Construction of **detector system has completed** with Grant-in-Aid for Scientific Research of Japan (Basic Science S, 2012–2016).
- An **integration test of the whole detector system** including DAQ had **success**fully performed. A **DIO measurement** had done at the same time, and analysis is ongoing.
- The **Detector system is ready** for the physics data taking.
- We are hoping to start soon after the completion of the beamline construction. No beam-time conflicts with T2K, KOTO or whatever the physics programs with the main ring of J-PARC.
- Meanwhile, we are keeping collaboration activities with detector improvements and by-product research.

Mu2e and COMET: High-Level Differences

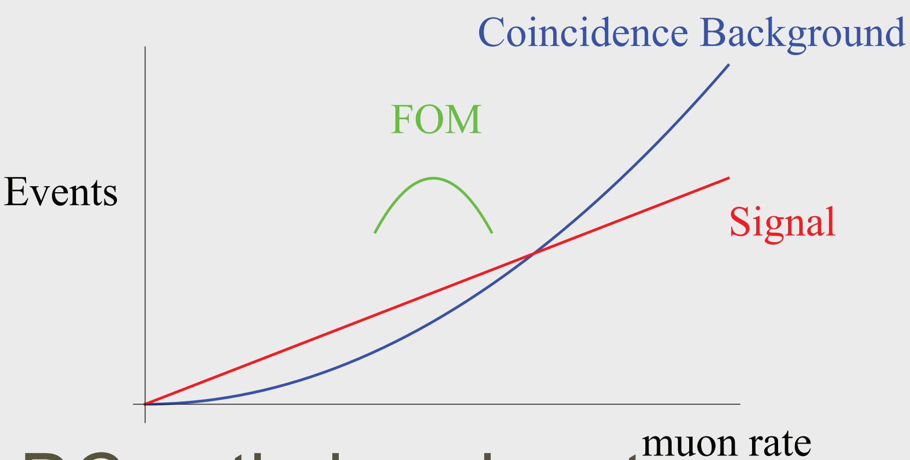
- COMET Staging has obvious advantages for learning as you go
- Charge Symmetry of Detector:
  - Mu2e: charge symmetric, e^+e^- the same
  - 2nd Bend in COMET Stage II momentum selects ~ 105 MeV e^- only
  - Mu2e needs hole, COMET does not

Effect of Charge-Symmetry Choice

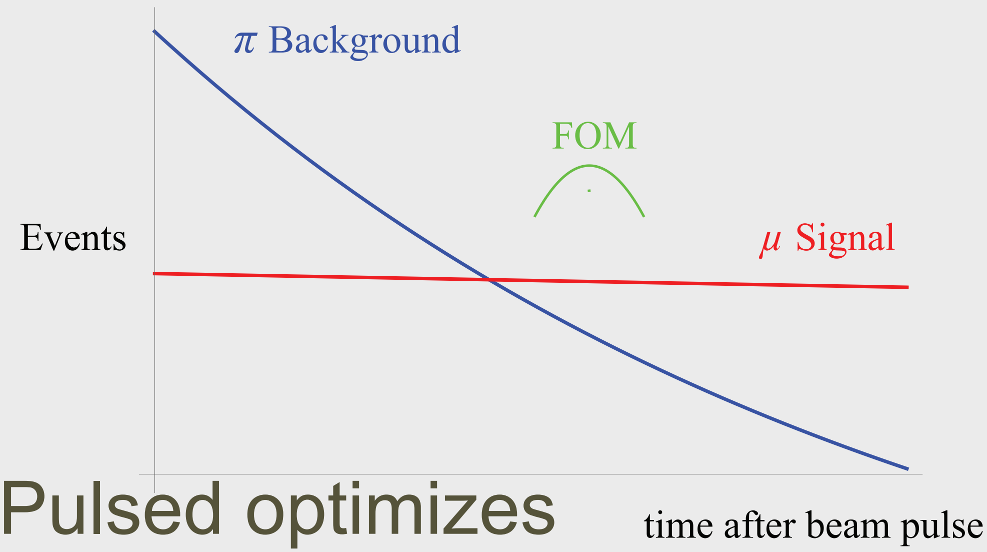
- Backgrounds
  - many background sources of e^- (RPC!) produce equal e^+e^-
  - Mu2e can measure them with e^+ *in situ*
  - *look "inside the box"*, no blinding needed since e^+
- Physics (next topic)
  - Mu2e can simultaneously measure $\mu^- N \rightarrow e^+ N'$, different/unique physics
  - COMET would have to reverse field, not trivial: separate runs

Pulsed vs. "Steady" Beam

- Decay expts, $\mu \to e\gamma$ and $\mu \to eee$ want a steady beam: PSI
  - need to minimize backgrounds from coincidences of two decays
  - $(\text{Rate})^2$ bkg vs Rate(signal)
- muon-electron conversion wants a pulsed beam: FNAL/J-PARC
  - Many pion-induced backgrounds after proton pulse
  - Take advantage of 26 nsec lifetime to "wait it out"



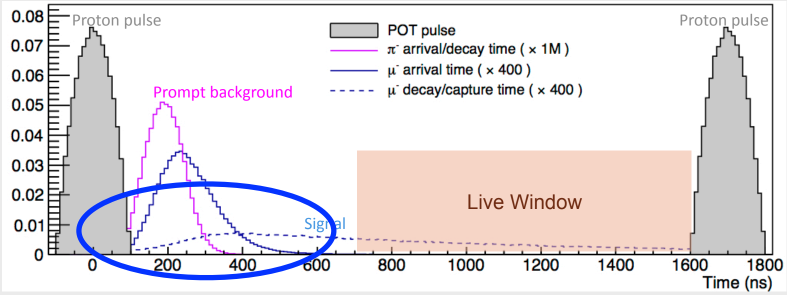
DC optimizes in rate



Pulsed optimizes in time

Beam Flash

- After proton beam strikes target, “flash” of photons and electrons, and many electrons hit stopping target and detector
- Large neutron flux as well that would deaden detectors
- By waiting:



  - activity from beam flash is over
  - And pion backgrounds have decayed by $\sim 10^{11}$
- This early flash has implications for Mu2e and especially for potential upgrades