

24th International Symposium on  
**P**Articles, **S**trings & **c****o****s**mology

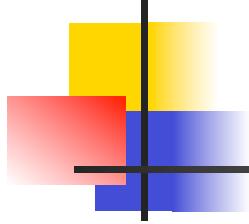


# Neutrino mass and Grand unification

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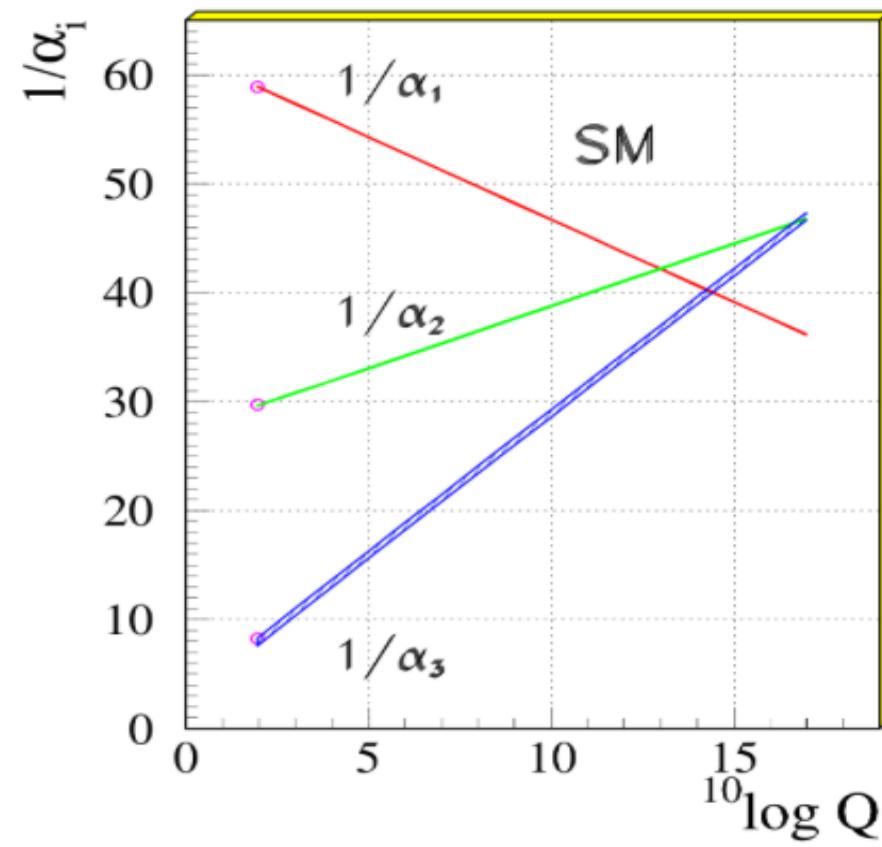
# Grand Unified Theories (GUTs): Elegant and ambitious

- Unifies all matter and forces
- Makes theory more predictive
  - e.g. can predict  $\sin^2 \theta_W$  + more
- Also quantizes electric charges

(Pati, Salam; Georgi, Glashow)

# Key ingredient: coupling unification ( $g_1=g_2=g_3$ )

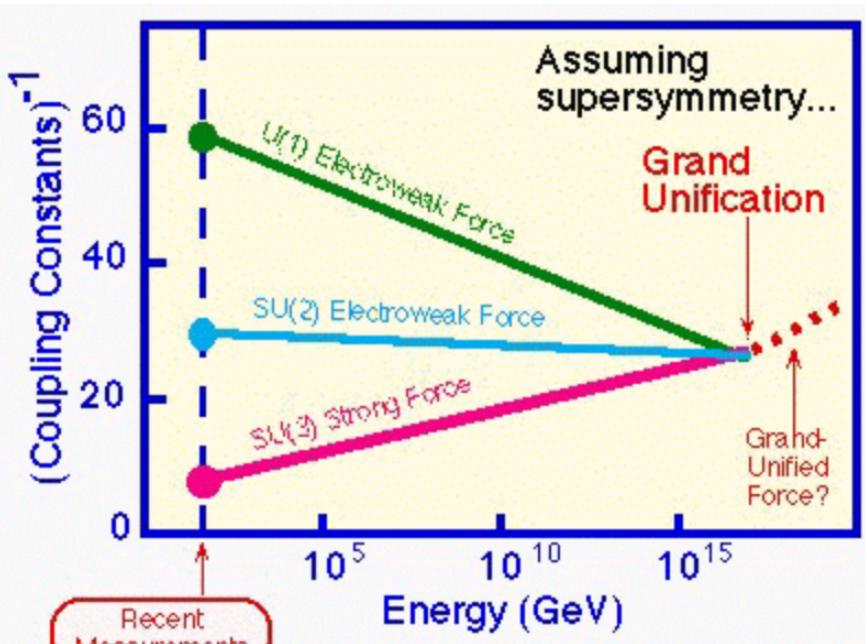
- Prospects in the standard model: g's run towards each other at shorter distances (Georgi, Quinn, Weinberg'74)



→ *But don't quite unify*  
suggesting possibly new  
physics below GUT scale

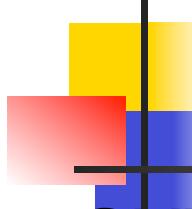
# SUSY GUTs

- Couplings unify at scale  $\sim 10^{16}$  GeV with susy breaking at TeV required to solve gauge hierarchy problem



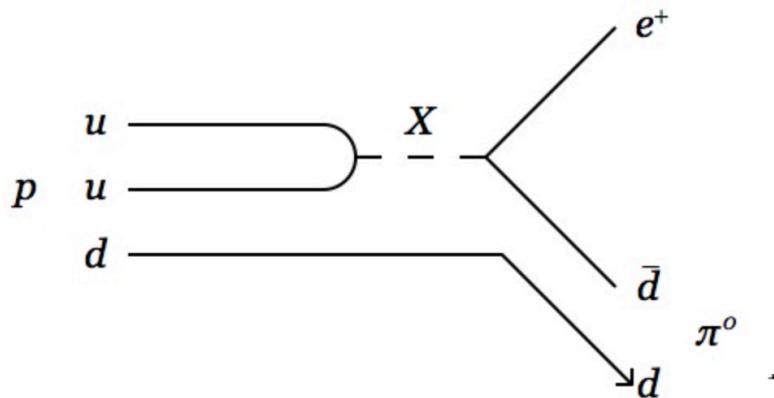
Dimopoulos, Raby, Wilczek'81; Ibanez, Ross'81

Where is SUSY?



# Key prediction: proton decay

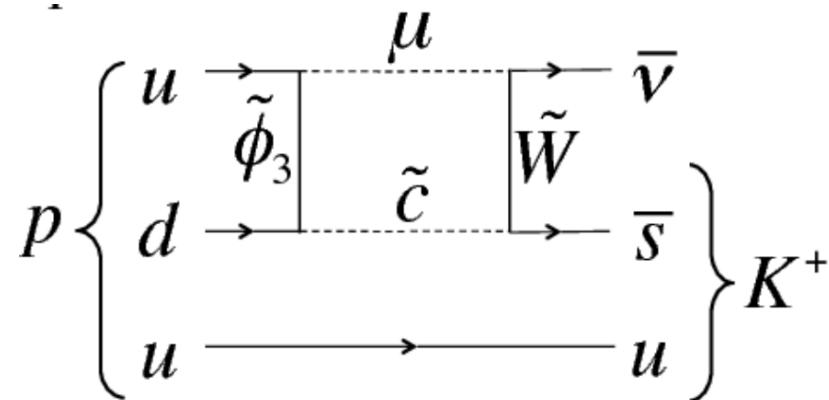
- GUT group  $\rightarrow$  Q-L unification  $\rightarrow$  proton decay
- Non-SUSY GUTs  $p \rightarrow e^+ \pi^0$  (Reviews: Nath Perez)



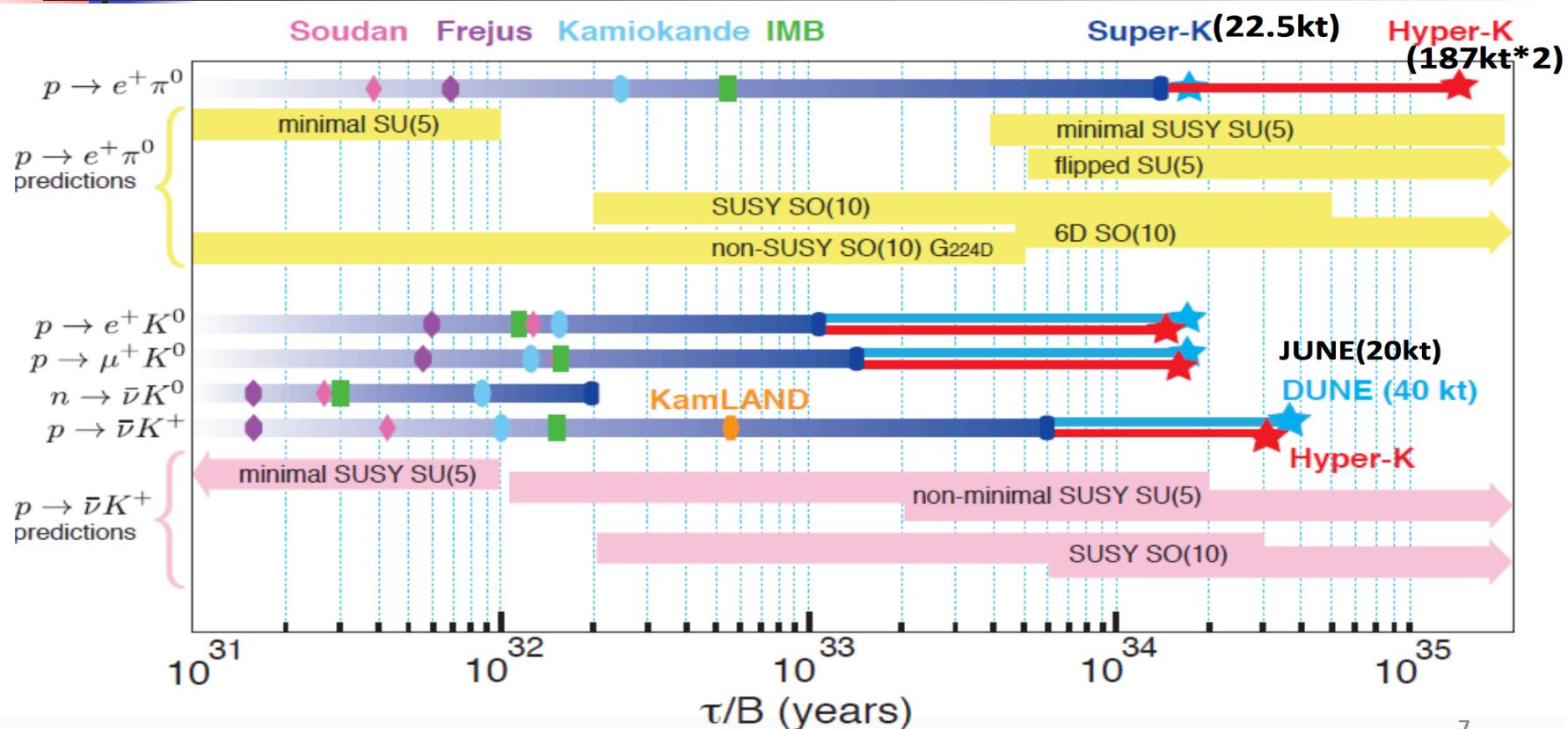
(True test of GUTs)

- SUSY GUTs

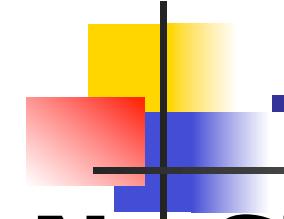
$$p \rightarrow K^+ + \bar{\nu}$$



# Current limits on proton life time



No sign of proton decay!

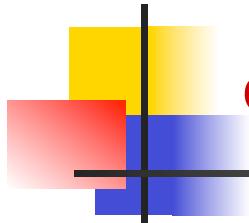


**No SUSY, No proton decay; don't  
need it to understand electric  
charge quantization;**

**Are GUT theories dead?**

**Neutrino masses may have  
provided new life to GUTs**

# Neutrino masses: 20 years ago (<1998)



# Neutrino masses and mixings now

NuFIT 3.2 (2018)

	Normal Ordering (best fit)		Inverted Ordering ( $\Delta\chi^2 = 4.14$ )		Any Ordering
	bfp $\pm 1\sigma$	$3\sigma$ range	bfp $\pm 1\sigma$	$3\sigma$ range	$3\sigma$ range
$\sin^2 \theta_{12}$	$0.307^{+0.013}_{-0.012}$	$0.272 \rightarrow 0.346$	$0.307^{+0.013}_{-0.012}$	$0.272 \rightarrow 0.346$	$0.272 \rightarrow 0.346$
$\theta_{12}/^\circ$	$33.62^{+0.78}_{-0.76}$	$31.42 \rightarrow 36.05$	$33.62^{+0.78}_{-0.76}$	$31.43 \rightarrow 36.06$	$31.42 \rightarrow 36.05$
$\sin^2 \theta_{23}$	$0.538^{+0.033}_{-0.069}$	$0.418 \rightarrow 0.613$	$0.554^{+0.023}_{-0.033}$	$0.435 \rightarrow 0.616$	$0.418 \rightarrow 0.613$
$\theta_{23}/^\circ$	$47.2^{+1.9}_{-3.9}$	$40.3 \rightarrow 51.5$	$48.1^{+1.4}_{-1.9}$	$41.3 \rightarrow 51.7$	$40.3 \rightarrow 51.5$
$\sin^2 \theta_{13}$	$0.02206^{+0.00075}_{-0.00075}$	$0.01981 \rightarrow 0.02436$	$0.02227^{+0.00074}_{-0.00074}$	$0.02006 \rightarrow 0.02452$	$0.01981 \rightarrow 0.02436$
$\theta_{13}/^\circ$	$8.54^{+0.15}_{-0.15}$	$8.09 \rightarrow 8.98$	$8.58^{+0.14}_{-0.14}$	$8.14 \rightarrow 9.01$	$8.09 \rightarrow 8.98$
$\delta_{CP}/^\circ$	$234^{+43}_{-31}$	$144 \rightarrow 374$	$278^{+26}_{-29}$	$192 \rightarrow 354$	$144 \rightarrow 374$
$\frac{\Delta m_{21}^2}{10^{-5} \text{ eV}^2}$	$7.40^{+0.21}_{-0.20}$	$6.80 \rightarrow 8.02$	$7.40^{+0.21}_{-0.20}$	$6.80 \rightarrow 8.02$	$6.80 \rightarrow 8.02$
$\frac{\Delta m_{3e}^2}{10^{-3} \text{ eV}^2}$	$+2.494^{+0.033}_{-0.031}$	$+2.399 \rightarrow +2.593$	$-2.465^{+0.032}_{-0.031}$	$-2.562 \rightarrow -2.369$	$\begin{bmatrix} +2.399 \rightarrow +2.593 \\ -2.536 \rightarrow -2.395 \end{bmatrix}$

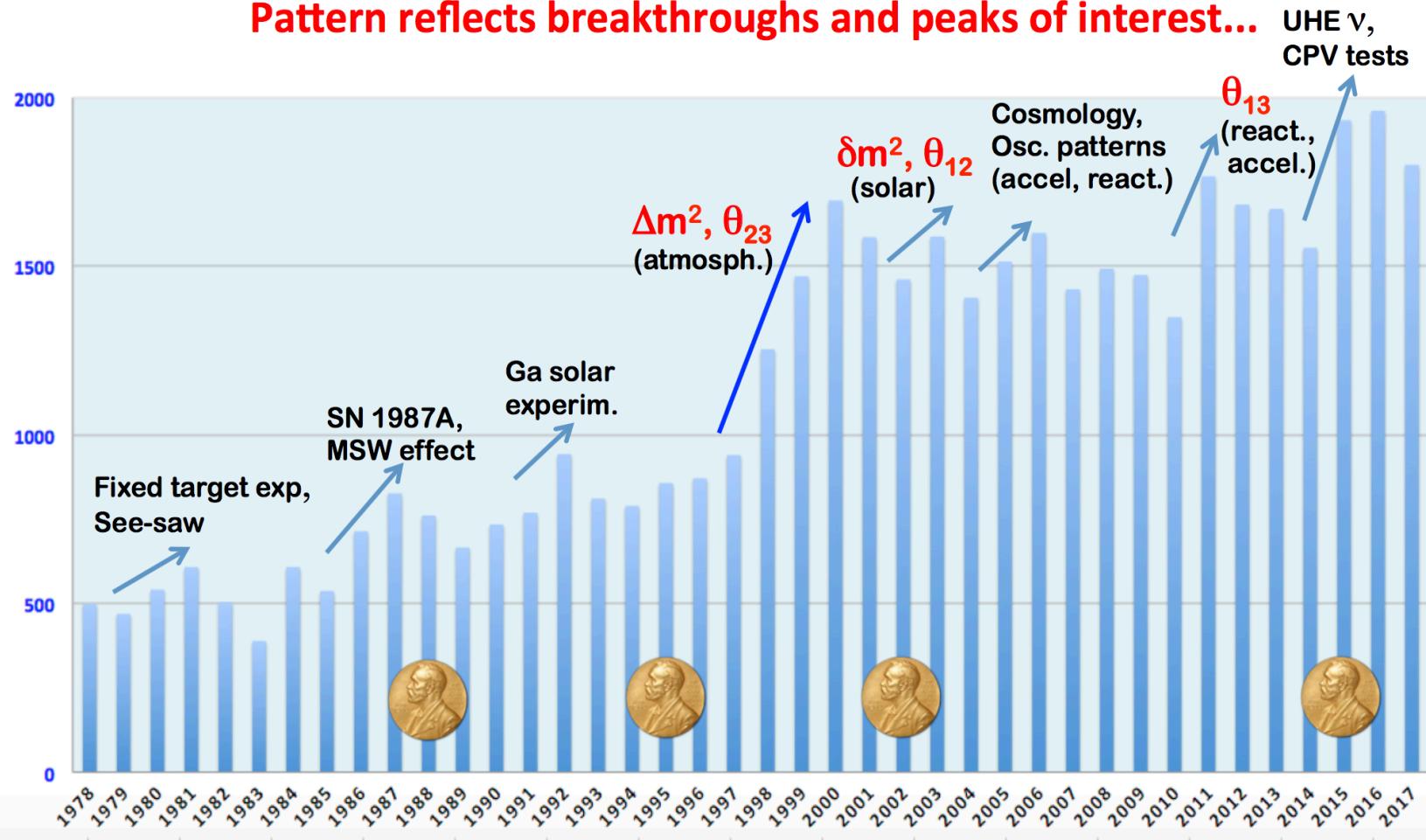
Cosmology  $\sum m_\nu [\text{eV}] < 0.3$  (EUCLID, LSST)  $\rightarrow .02 \text{ eV}$

# Neutrino activity

N. of #neutrino# preprints per year (1978-2018) from [INSPIRE](#)

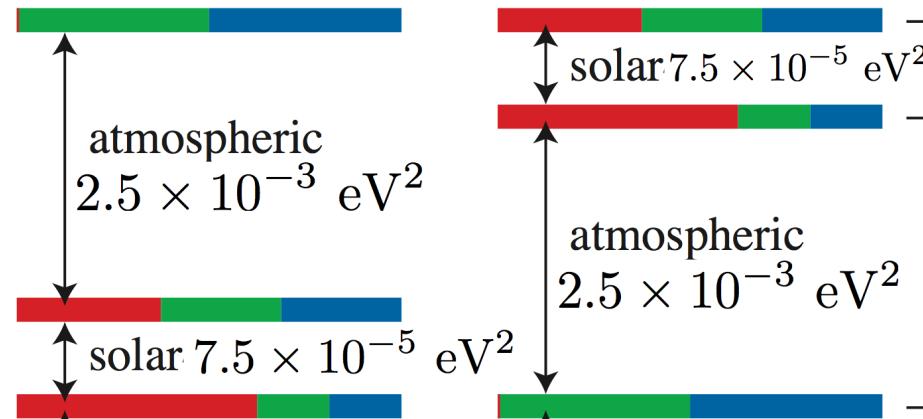
E. Lisi talk,  
Nu2018.

Pattern reflects breakthroughs and peaks of interest...

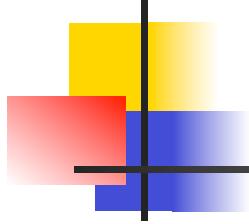


# Things we do not know

- Mass ordering



- CP phase  $\delta_{CP}$  (DUNE, T2K,...)
- Are there sterile neutrinos?
- Absolute scale  $\sum m_\nu$  (EUCLID, LSST)
- Neutrinos Majorana or Dirac?  $\beta\beta_{0\nu}$  Searches



# Neutrino mass and revival of GUTs

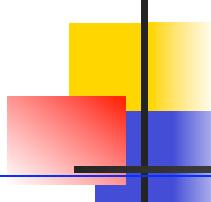
- Two new puzzles from neutrino mass discovery
  - (i)  $m_\nu \ll m_q, m_\ell$  why?
  - (ii) Lepton mixings very different from quarks:  
why?

# Where does neutrino mass come from ?

- Charged fermion masses come from the Higgs vev:

$$m_f = h_f v_{wk} \quad v_{wk} = \langle h^0 \rangle$$

★ Discovery of the 125 GeV Higgs  $h^0$  confirms this.

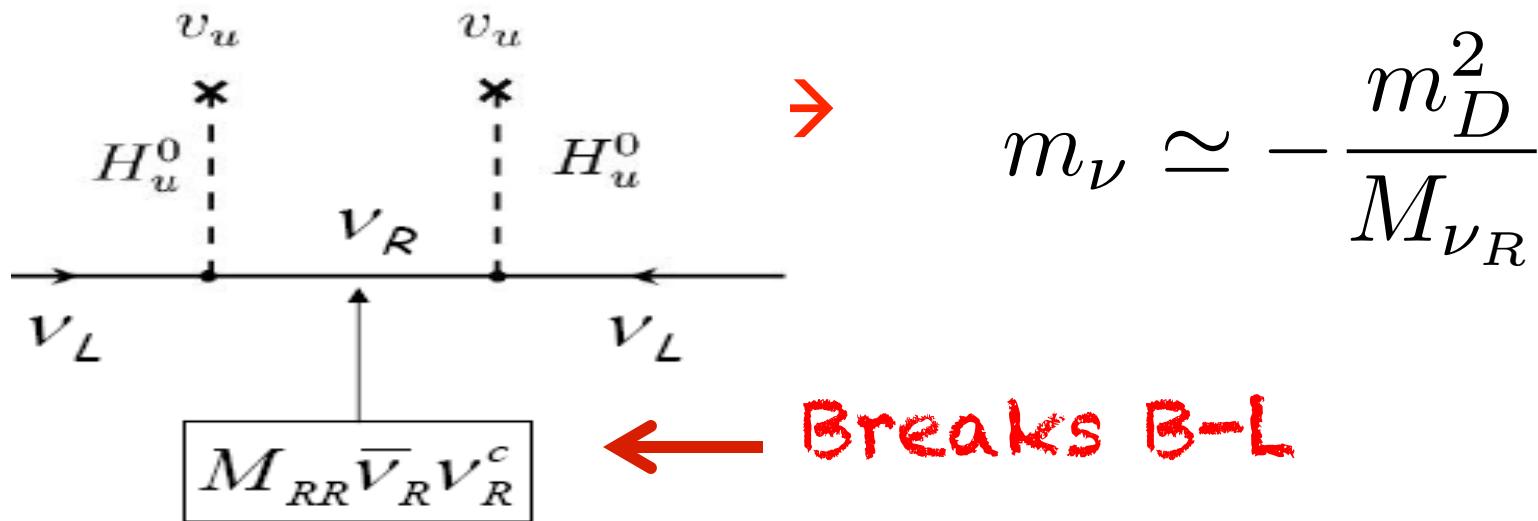


# Where does neutrino mass come from ?

- To get neutrino mass, we can simply add the right handed neutrino  $\nu_R$  to SM and use the same mechanism as for other fermions to get a Dirac mass.
- However, we get too large mass unless  $h_\nu \leq 10^{-12}$
- **We need to go beyond!**

# Seesaw paradigm for neutrino mass and GUTs

- SM+ RH neutrinos  $\nu_R$  but with heavy Majorana mass



(Minkowski'77; Gell-Mann,Ramond, Slansky; Yanagida; Glashow; Mohapatra,Senjanovic'79)

- $\nu_R$  and B-L two essential ingredients of seesaw

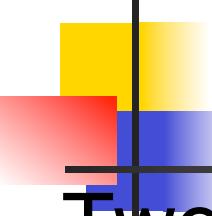


# A Major plus for seesaw

- It has built in mechanism for explaining the observed matter-anti-matter asymmetry of the universe:

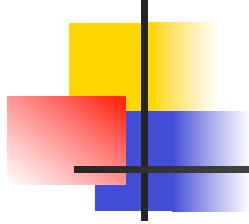
$$\epsilon_\ell = \frac{\Gamma(N \rightarrow \ell + H) - \Gamma(N \rightarrow \bar{\ell} + \bar{H})}{\Gamma(N \rightarrow \ell + H) + \Gamma(N \rightarrow \bar{\ell} + \bar{H})}$$

- Lepton asymmetry gets converted to baryons by the SM sphalerons; (Fukugita and Yanagida'86)



# What is the seesaw scale?

- Two classes of models where the key ingredients of seesaw i.e. Right handed neutrino and B-L symmetry are automatic.
- (i) Left-right symmetric models: the scale can be anywhere from a few TeV to very high.
- (ii) SO(10)- seesaw scale near GUT scale.
- **Leptogenesis works in both cases!!**



# Seesaw and GUTs

- SO(10) theories generally predicts  
→  $m_{D33} \sim m_t$
- + Seesaw formula  $m_\nu \simeq -\frac{m_D^2}{M_{\nu_R}} \rightarrow M_{\nu_R} \sim 10^{14} \text{ GeV}$   
 $M_U \sim 10^{16} \text{ GeV}$
- Fits naturally into GUT framework :  $M_{\nu_R} \sim M_U$

# GUT symmetry just right for seesaw: SO(10)

- {16}- spinor of SO(10)

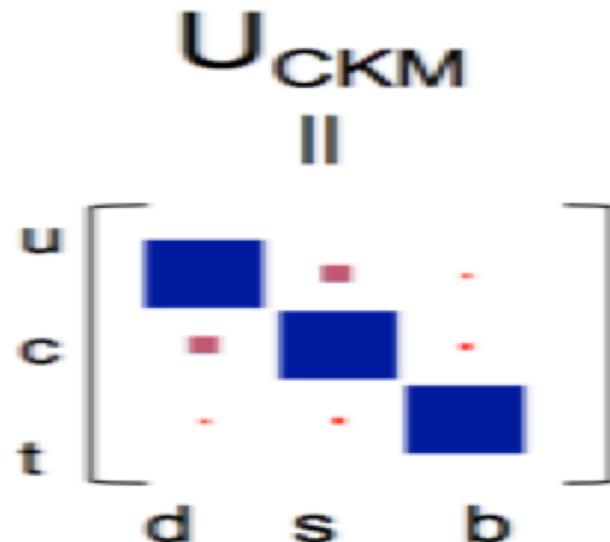
{16}=

$$\begin{pmatrix} u & c & t & \nu \\ d & s & b & e \end{pmatrix}_{L,R}$$

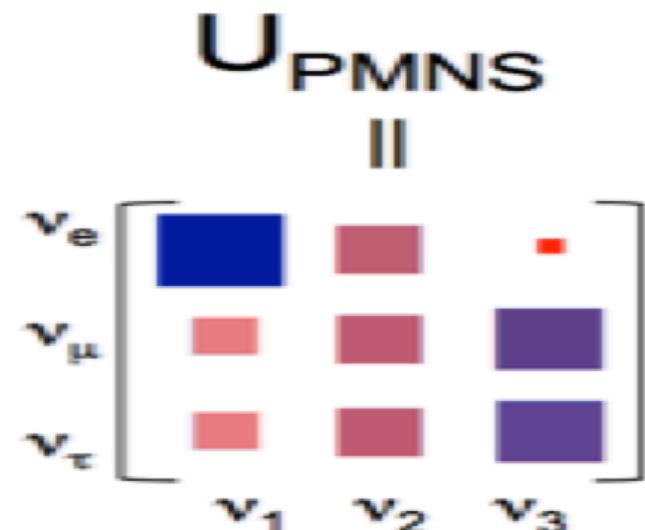
⊃ SM fermions + N\_R

- SO(10) contains B-L as a subgroup

# Flavor: a challenge for GUTs



$$V_{CKM} \sim \begin{bmatrix} 0.976 & 0.22 & 0.004 \\ -0.22 & 0.98 & 0.04 \\ 0.007 & -0.04 & 1 \end{bmatrix}$$



$$U_{PMNS} \sim \begin{bmatrix} 0.85 & -0.54 & 0.16 \\ 0.33 & 0.62 & -0.72 \\ -0.40 & -0.59 & -0.70 \end{bmatrix}$$

- *If quarks and leptons are similar, how to explain this diversity?*

# A Simple SO(10) set-up to address the flavor issue

$$16_f \times 16_f = 10_H + 126_H + 120_H$$

- Minimal models 10+126-Higgs (Babu, RNM'92)

$$W = h\psi\psi H + f\psi\psi\bar{\Delta}$$

- Slightly generalized:

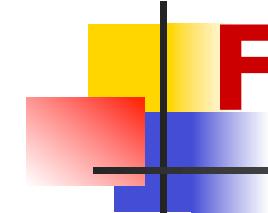
$$10+126+120+\text{CP} \ (\# \text{ param}=17)$$

(Dutta, Mimura, RNM'05; Bertolini, Frigerio, Malinsky'05)

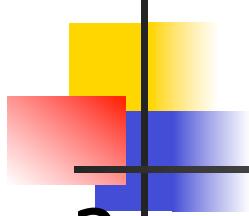
$$W = h\psi\psi H + f\psi\psi\bar{\Delta} + g\psi\psi\Sigma$$

- Lead naturally to DM without extra assumptions

# Explaining lepton-quark Flavor diversity: 10+126



- $$\mathcal{M}_\nu = f v_L - M_d \frac{1}{f v_R} M_d$$
- $$; \quad f = \frac{1}{4\kappa_d} (M_d - M_\ell)$$
- in GUTs  $m_b(M_U) \simeq m_\tau(M_U)$  endows  $\mathcal{M}_\nu$  with different flavor structure compared to  $M_{u,e,d}$  and provides a way to understand diverse quark-leptons flavor patterns: (Bajc, Senjanovic, Vissani'2003)



# How does it work?

- 2 gen; if f- dominates  $\mathcal{M}_\nu = c(M_d - M_\ell)$

$$M_d = m_b \begin{pmatrix} \sim \lambda^2 & \sim \lambda^2 \\ \sim \lambda^2 & 1 \end{pmatrix} \quad M_\ell = m_\tau \begin{pmatrix} a_1 \lambda^2 & a_2 \lambda^2 \\ a_2 \lambda^2 & 1 \end{pmatrix}$$

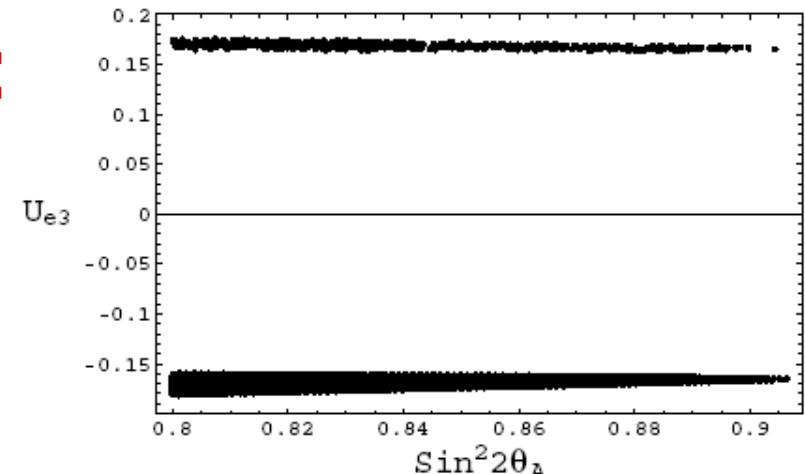
- At GUT scale  $m_b \simeq m_\tau (1 + \lambda^2)$
- Then  $\mathcal{M}_\nu = \begin{pmatrix} a' \lambda^2 & b' \lambda^2 \\ b' \lambda^2 & \lambda^2 \end{pmatrix}$
- $a_1, a_2, a', b' \sim 1 \rightarrow$  large  $\theta_{23}$  natural; easily extends to 3 generation case

# Three gen. extension

- Works quantitatively: (10+126)
- Predicts normal hierarchy:

-   $\theta_{12}, \theta_{23}$  large

-   $\theta_{13} \approx \lambda$  “large”



(Goh, RNM, Ng, 03 ; Babu, Macesanu'05)

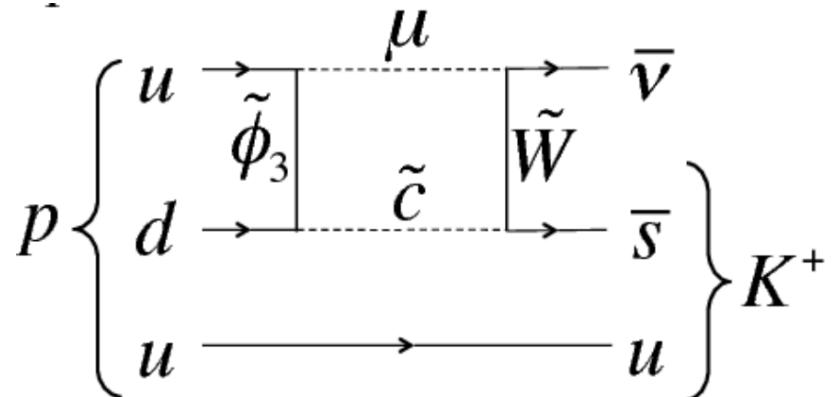
$$\text{--- } \theta_{13} \cong 0.15$$

$$\text{Expt : } \theta_{13} = 0.14 - 0.156$$

$$\text{--- } \frac{m_{solar}}{m_{atmos}} \sim \lambda$$

# Proton decay in SUSY GUTs

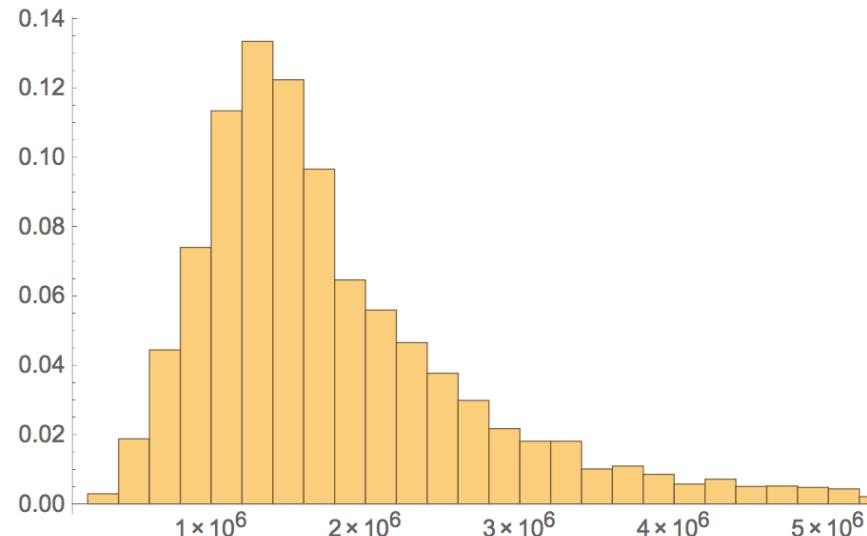
- In Non-SUSY models,  $p \rightarrow e^+ \pi^0$  dominates – does not test seesaw but a key test of GUTs
- In SUSY models, p-decay  $p \rightarrow K^+ \bar{\nu}$  connected to neutrino mixings and hence can test seesaw



# What does proton decay say about these models?

- 10+126 model has tension with p-decay – specially RRRR operators troublesome:

$$p \rightarrow K^+ + \bar{\nu}_\tau \text{ rate} \propto (\tan\beta)^2$$



$M_{\text{susy}} > 232 \text{ TeV}$

(Babu, Bajc, Saad'18)

- What if susy is discovered at lower scale?

# SO(10) model with

## 10+126+120 works better

- In 10+126 models all Yukawas fixed by fermion mass fits, which leads to p-decay tension.
- On the other hand in models with 120, one can choose textures to suppress proton decay: e.g.

$$\tilde{h} = \begin{pmatrix} 0 & & \\ & 0 & \\ & & M \end{pmatrix}, \quad \tilde{f} = \begin{pmatrix} \sim 0 & \sim 0 & f_{13} \\ \sim 0 & f_{22} & f_{23} \\ f_{13} & f_{23} & f_{33} \end{pmatrix}, \quad \tilde{g} = i \begin{pmatrix} 0 & g_{12} & g_{13} \\ -g_{12} & 0 & g_{23} \\ -g_{13} & -g_{23} & 0 \end{pmatrix}.$$

- P-decay bound OK with 3 TeV susy breaking!

(Dutta, Mimura, RNM'05)

# Fermion mass fit and predictions

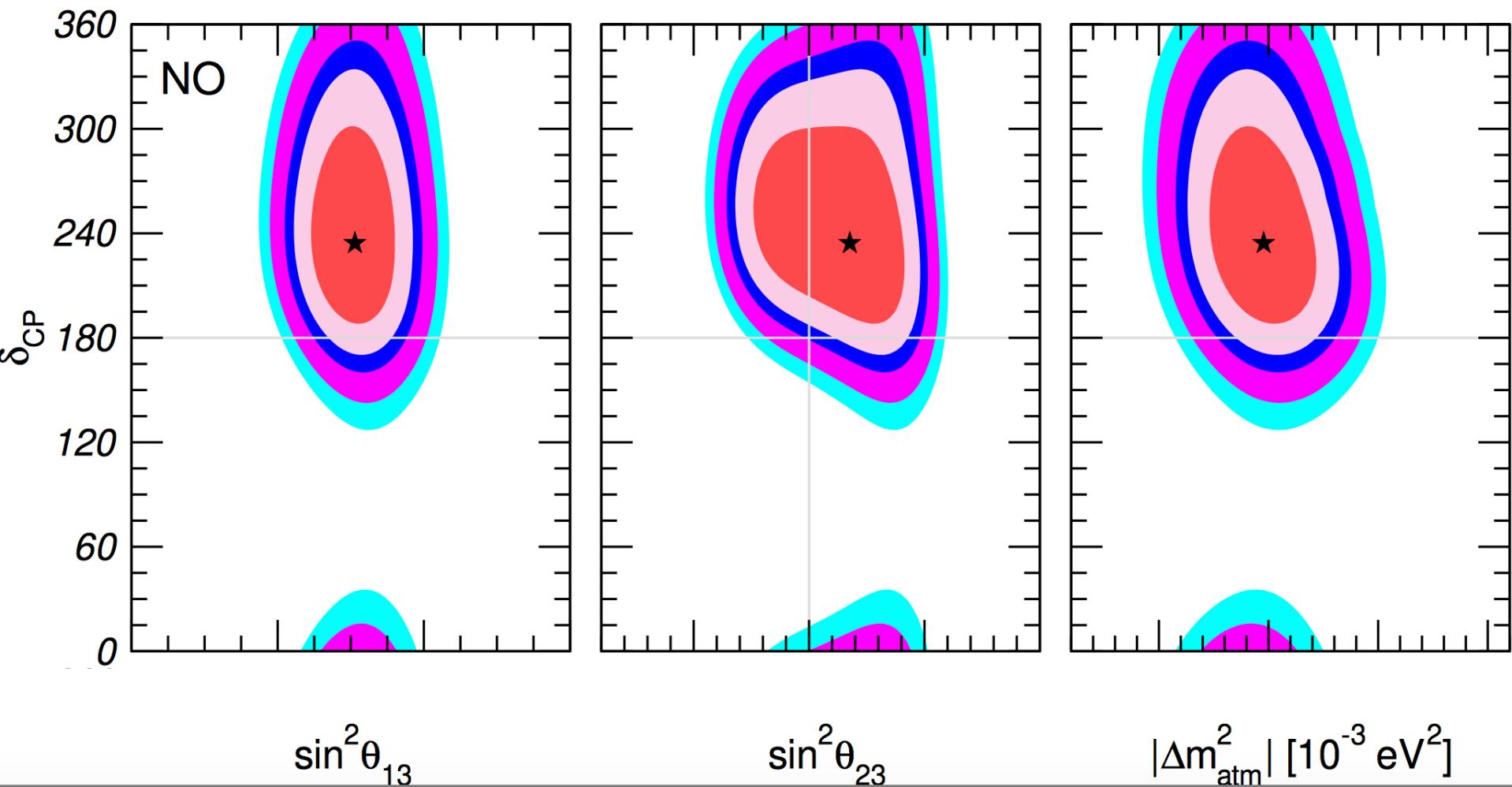
- Best fit values for neutrinos with large CP phase

	best fit	exp value		best fit	exp value
$m_u$ (MeV)	0.7246	$0.72^{+0.12}_{-0.15}$	$V_{us}$	0.22427	$0.2243 \pm 0.0016$
$m_c$ (MeV)	208.6	$210.5^{+15.1}_{-21.2}$	$V_{ub}$	0.0030	$0.0032 \pm 0.0005$
$m_t$ (GeV)	80.113	$80.45^{+2.9*}_{-2.6}$	$V_{cb}$	0.03497	$0.0351 \pm 0.0013$
$m_d$ (MeV)	1.515	$0.930 \pm 0.38^*$	$J \times 10^{-5}$	2.29	$2.2 \pm 0.6$
$m_s$ (MeV)	24.47	$17.6^{+4.9*}_{-4.7}$	$\Delta m_{21}^2 / \Delta m_{32}^2$	0.0308	$0.0309 \pm 0.0015$
$m_b$ (GeV)	1.311	$1.24 \pm 0.06^*$	$\theta_{13}$ ( $^\circ$ )	9.397	$8.88 \pm 0.385$
$m_e$ (MeV)	0.3565	$0.3565^{+0.0002}_{-0.001}$	$\theta_{12}$ ( $^\circ$ )	33.62	$33.5 \pm 0.8$
$m_\mu$ (MeV)	75.297	$75.29^{+0.05}_{-0.19}$	$\theta_{23}$ ( $^\circ$ )	43.79	$44.1 \pm 3.06$
$m_\tau$ (GeV)	1.61	$1.63^{+0.04}_{-0.03}$	$\delta_{\text{CP}}$ ( $^\circ$ )	-67	
			$\sum \chi^2$	3.16	

$$\sum m_\nu = .074 \text{ eV}$$

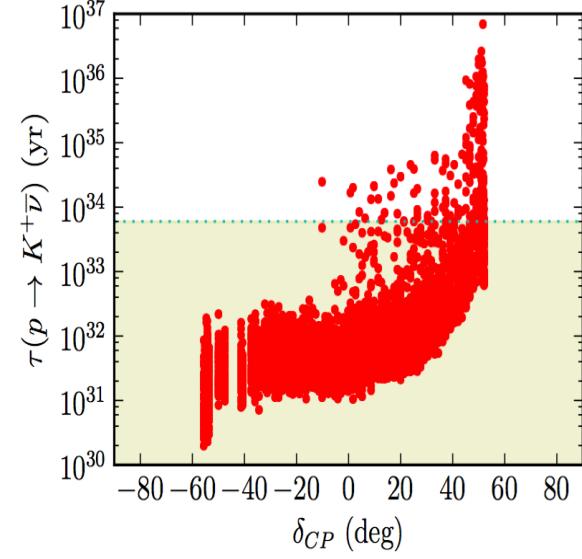
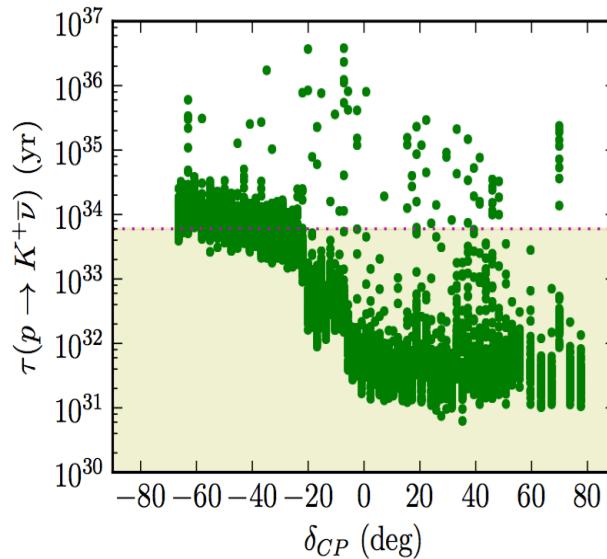
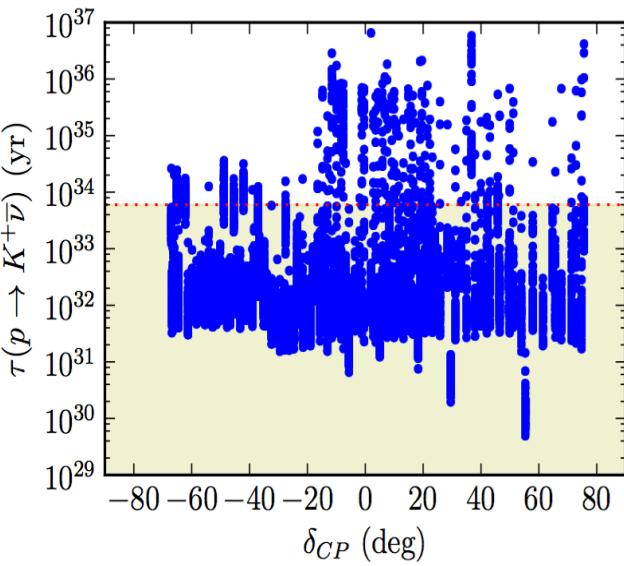
RNM and Severson'18

# Nu-Fit 3.2 result for CP phase



# proton life time and correlation in SO(10)

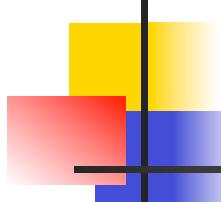
$\delta_{CP}$



$\delta_{CP}$  VS  $\tau_p$        $M_{SUSY} = 5$  TeV

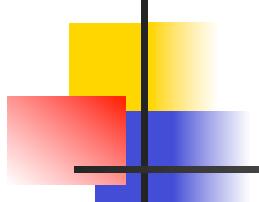
- Can help to test these models

(RNM,Severson'18)



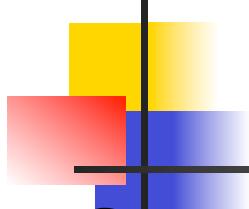
# **SO(10) with 120 solves strong CP without axion**

- Note that in  $10+126+120$  models  $M_q = M_q^\dagger$
- $\rightarrow$  in the effective theory,  $\theta = \text{Arg Det } M_q = 0$
- **Potential** to solve the strong CP problem without the axion.
- Structure natural even in the presence of higher dimensional terms!!



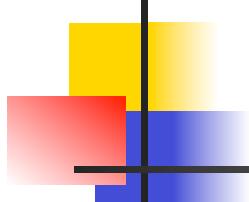
# Beyond Minimal SO(10)

- Add discrete symmetries to understand the Yukawa pattern.
- Embed in orbifold models in 5, 6 D to solve problems of GUTs e.g. doublet triplet splitting, less threshold effect etc
- Possibly eventually embed into string theories?



# Broad message

- Grand unified theories are elegant.
- However, with no evidence for SUSY and proton decay, neutrino masses provide compelling theoretical reason for them.
- Important to search for seesaw scale (e.g.  $N_R$ ,  $W_R$  etc ) in colliders to test TeV vs GUT seesaw.
- Discovery of evidence for TeV scale seesaw will change thinking about GUTs.



# Broad message

- GUT theories have the advantage of being very predictive and therefore testable!
- Renormalizable SO(10) theory with a number of interesting properties, a serious candidate;
- They predicted correct  $\theta_{13}$  prior to discovery;
- Explain matter-anti-matter asymmetry; dark matter of the universe; solve strong CP problem!
- Testable via link of proton decay to neutrino mixings and CP phase ( $\sim 30^\circ$  to  $-70^\circ$ )

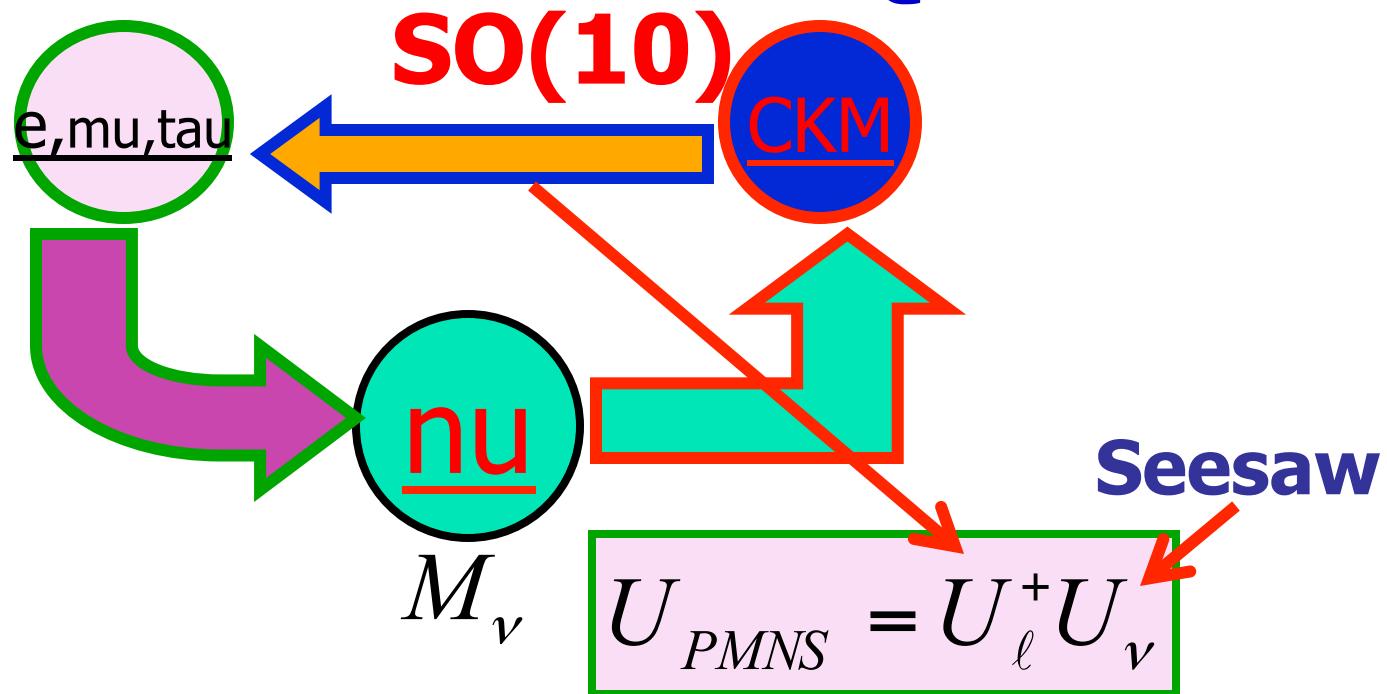


*Thank you !*

# Lepton-quark interplay in SO(10)

Leptons

Quarks



# Moving on to 3 generations

- $M_d$  and  $M_l$  typical:

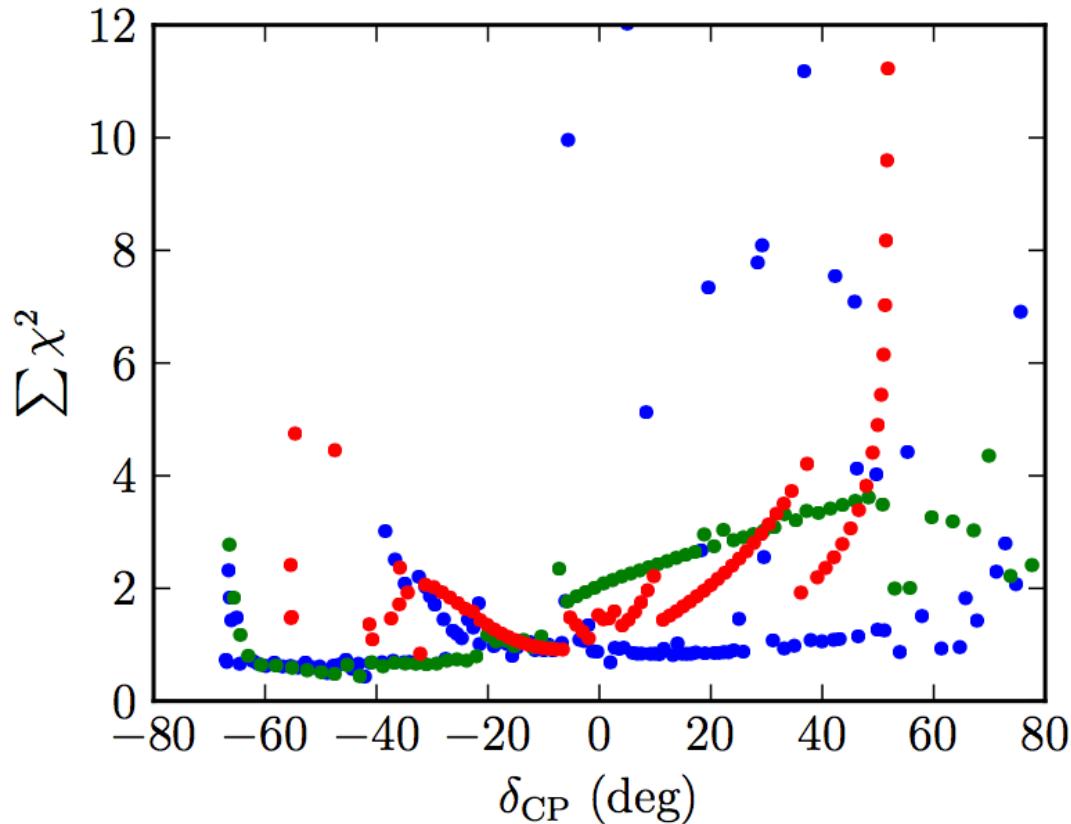
$$M_{d,\ell} \approx m_{b,\tau} \begin{pmatrix} \sim \lambda^4 & \sim \lambda^5 & \sim \lambda^3 \\ \sim \lambda^5 & \sim \lambda^2 & \sim \lambda^2 \\ \sim \lambda^3 & \sim \lambda^2 & 1 \end{pmatrix}$$

- b-tau unif at GUT scale →

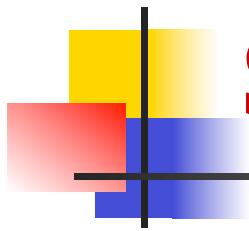
$$M_\nu = c(M_d - M_\ell) \approx m_0 \begin{pmatrix} \lambda^4 & \lambda^5 & \lambda^3 \\ \lambda^5 & \lambda^2 & \lambda^2 \\ \lambda^3 & \lambda^2 & \lambda^2 \end{pmatrix}$$

- Atmospheric, solar, theta13, all large: (Goh, RNM, Ng'03)

# CP phase prediction



$$\delta_{CP} = -67^0 \rightarrow +45^0$$



# **SO(10) to SM: Scale set up**

