

PULSAR TIMING ARRAYS:

THE NEXT WINDOW ON THE GRAVITATIONAL-WAVE UNIVERSE

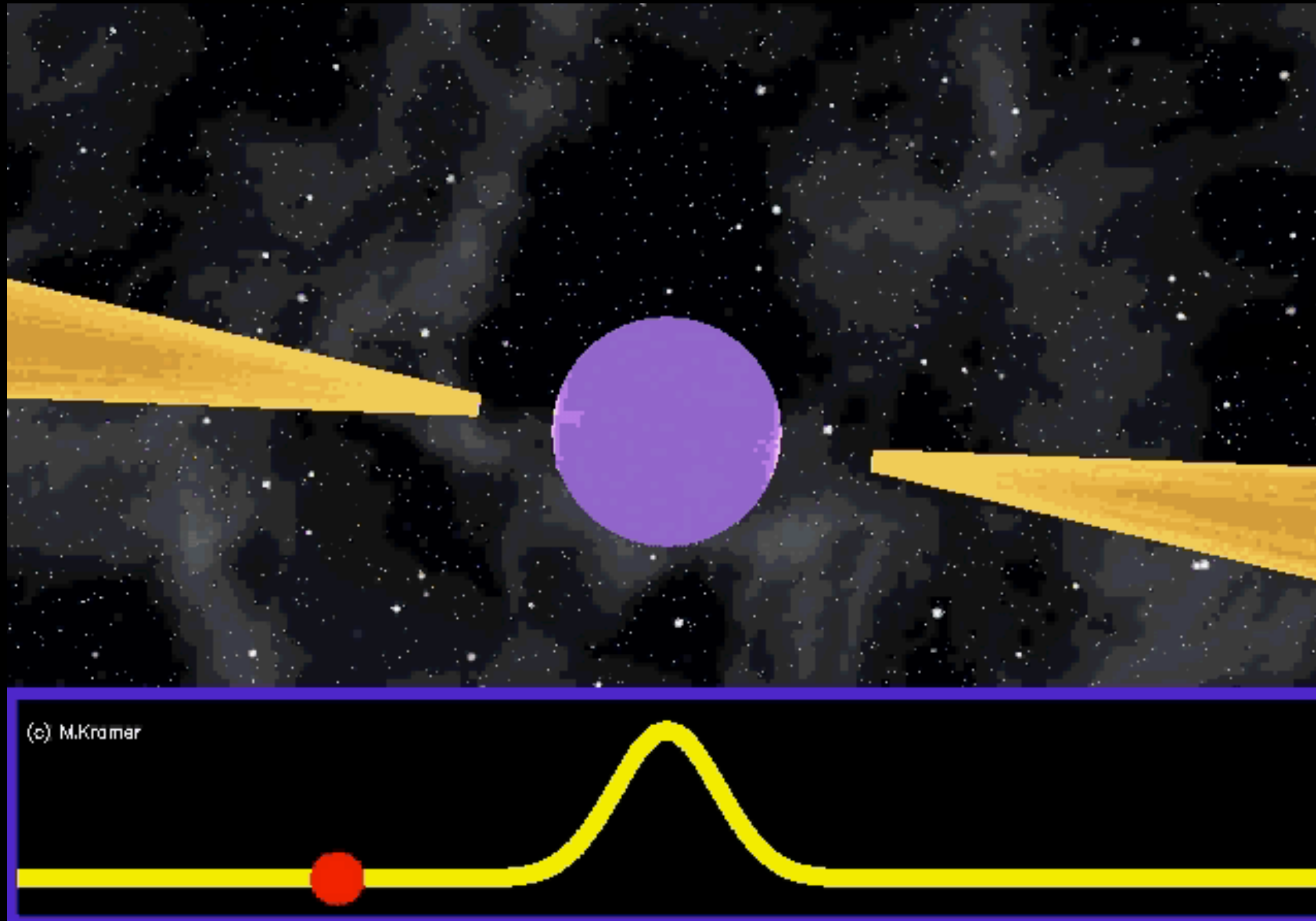
CHIARA MINGARELLI

Flatiron Research Fellow,
Co-Chair, International Pulsar Timing Array Gravitational-Wave Working Group

Think Big. Think the Size of the Milky Way.



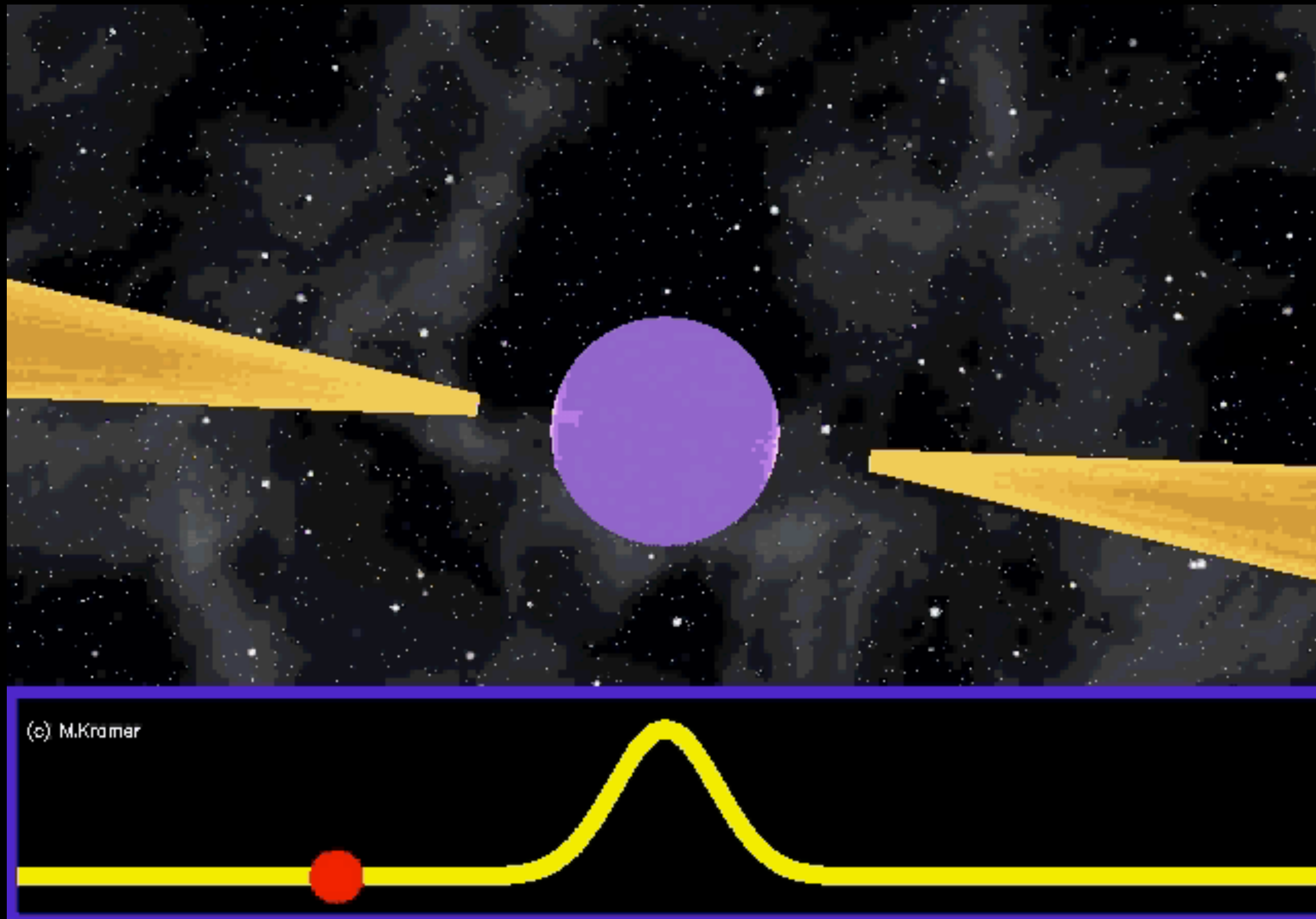
- Pulsars are excellent clocks
- Gravitational waves change the distance between Earth and pulsars
- Supermassive black hole binaries emit gravitational waves with periods of decades
- Need to time pulsars for decades
- This is a galactic-scale gravitational-wave detector



DETWEILER 1979; HELINGS & DOWNS 1983

REVIEWS: LOMMEN 2015, BURKE-SPOLAOR 2015

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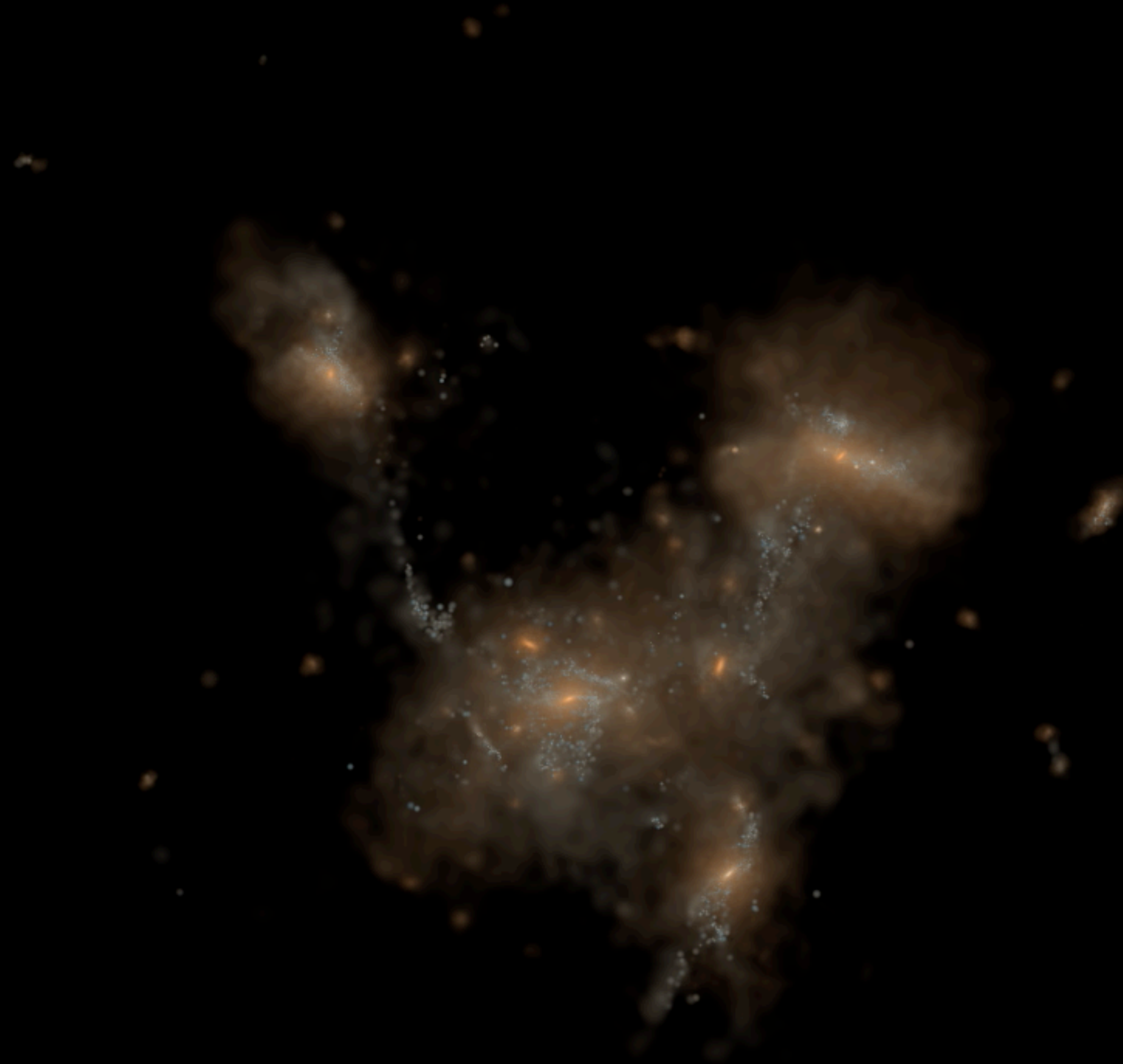


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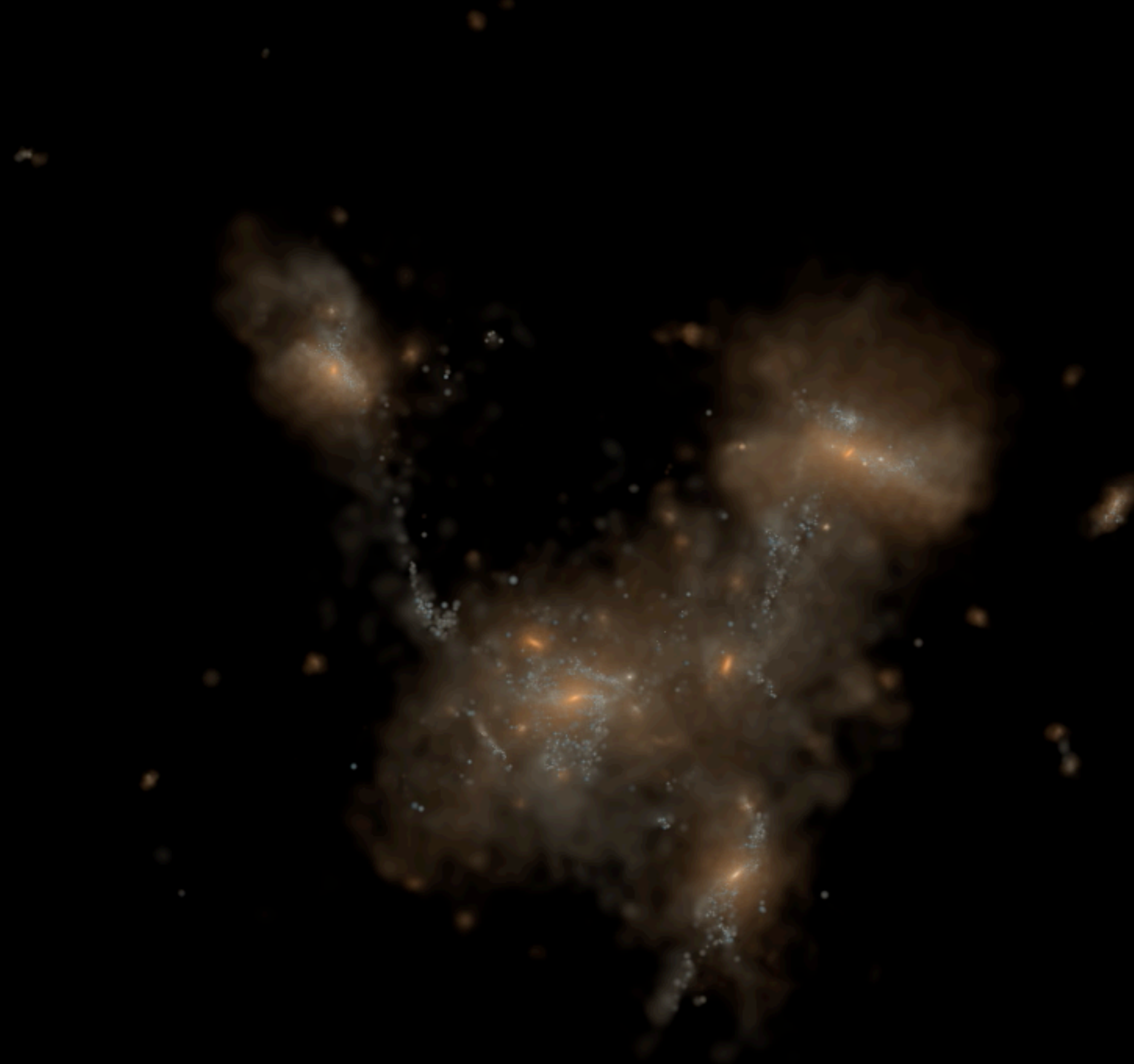
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ILLUSTRIS

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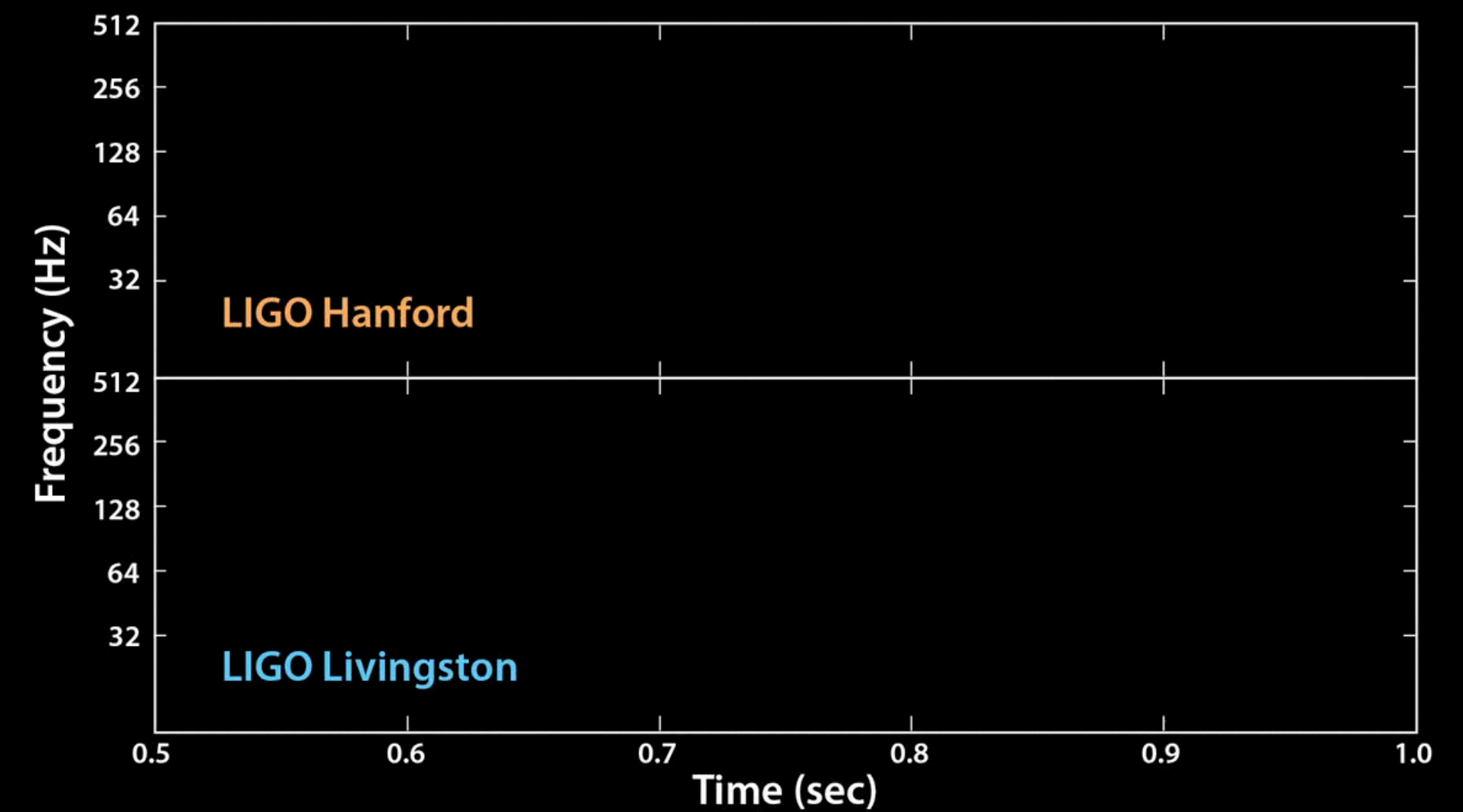
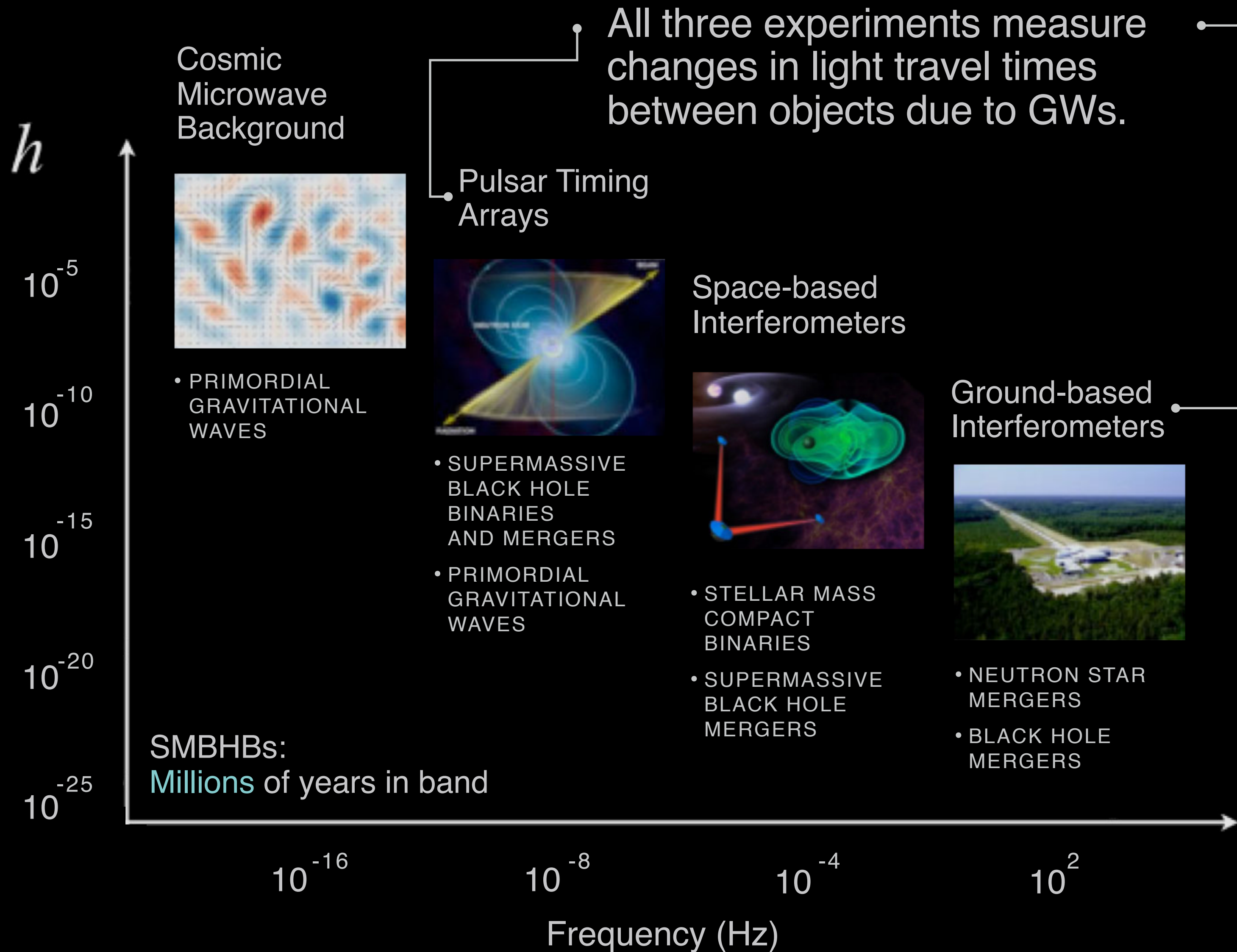


ANIMATION FROM JOHN ROWE ANIMATION/
AUSTRALIA TELESCOPE NATIONAL FACILITY, CSIRO

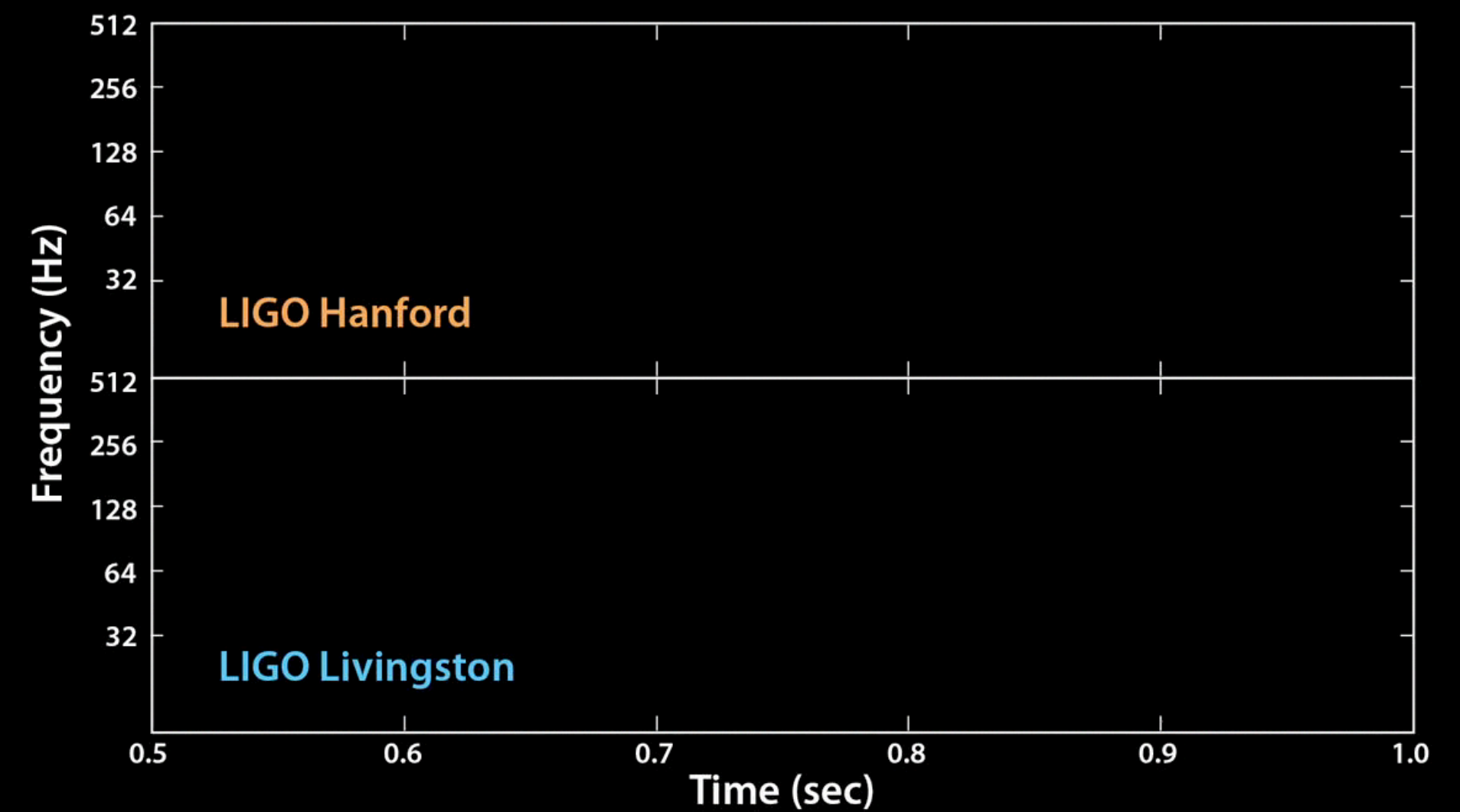
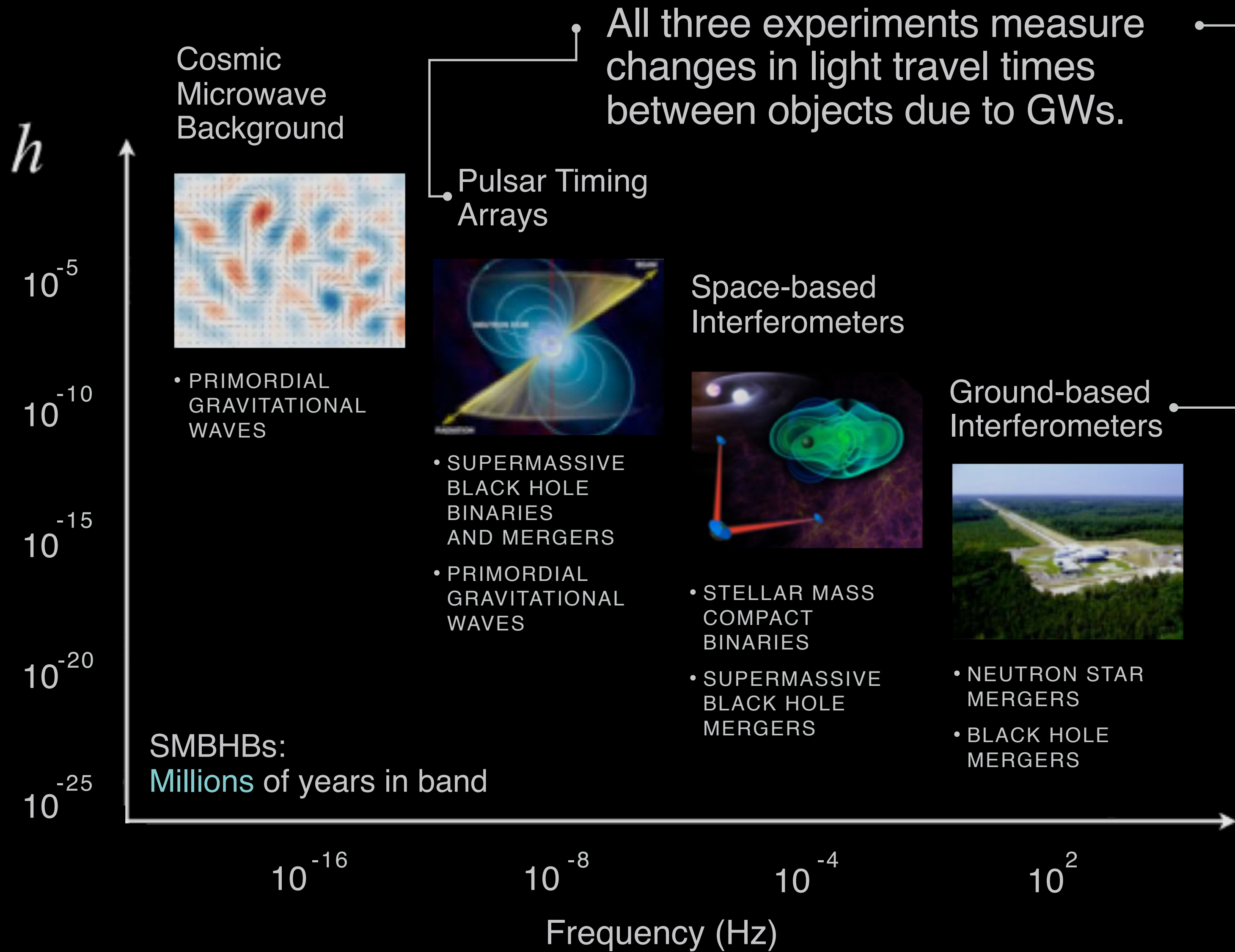


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The Spectrum of Gravitational-Wave Astronomy



The Spectrum of Gravitational-Wave Astronomy



New Avenue to Explore the Universe



Galaxy mergers: how the universe works

Each galaxy hosts a supermassive black hole: merge via gravitational waves

Gravitational waves: important in cosmology, galaxy evolution and fundamental physics

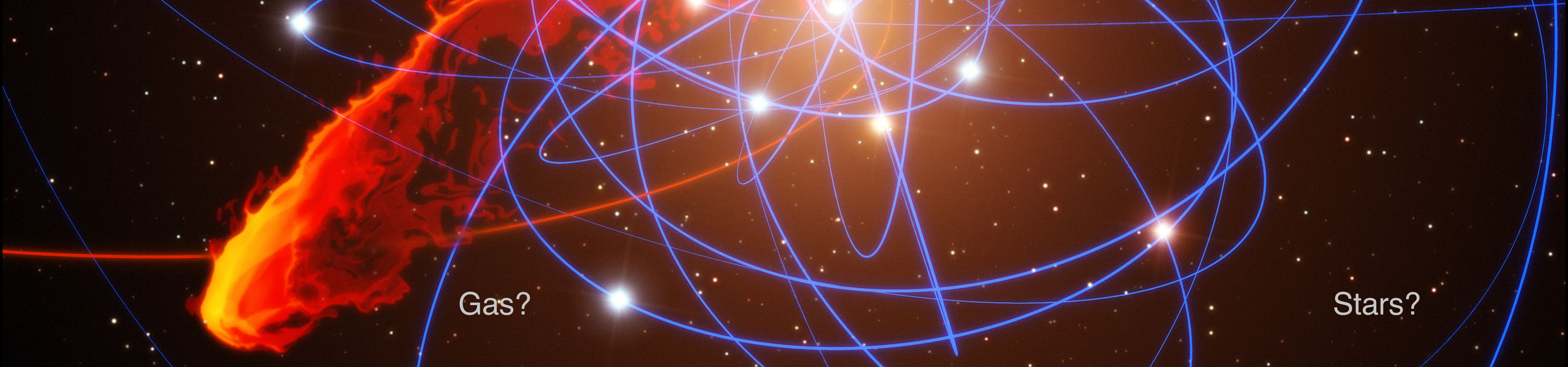
Information **not accessible by any other means**

What's New with PTAs?

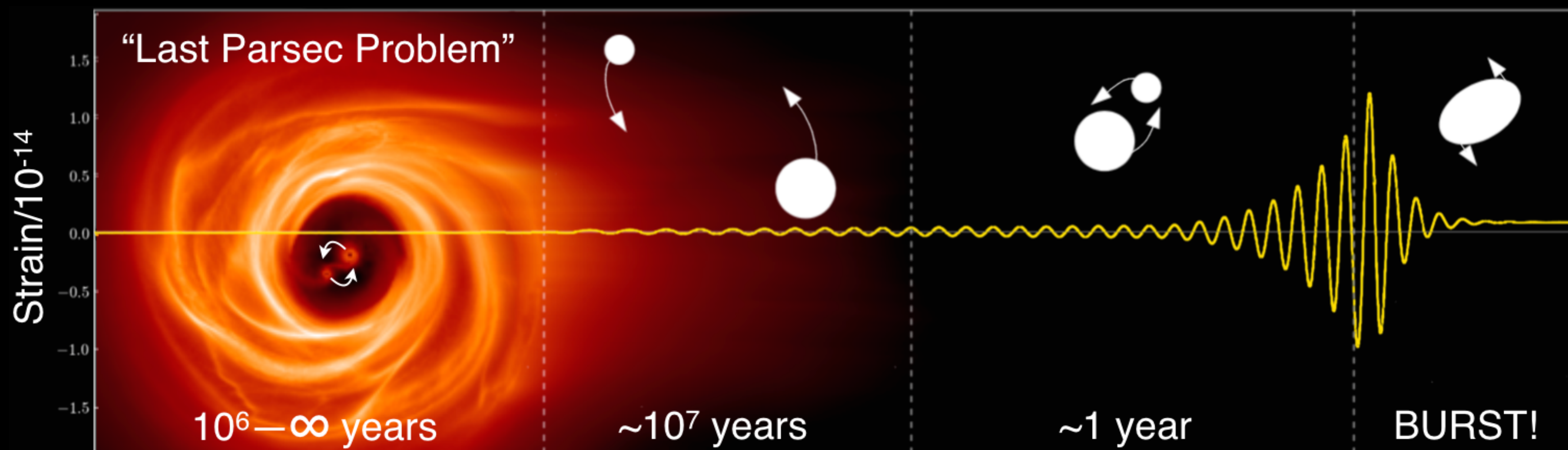
- Datasets are now long enough and of sufficient quality to make a detection.
- At this point, non-detections matter — can tell us about how supermassive black holes merge, galaxy evolution.
- Time to detection is anchored in underlying black hole astrophysics.

Arzoumanian et. al. 2016, 2018; Verbiest et al. 2016;

Lentati et al. 2015; Taylor et al. 2016

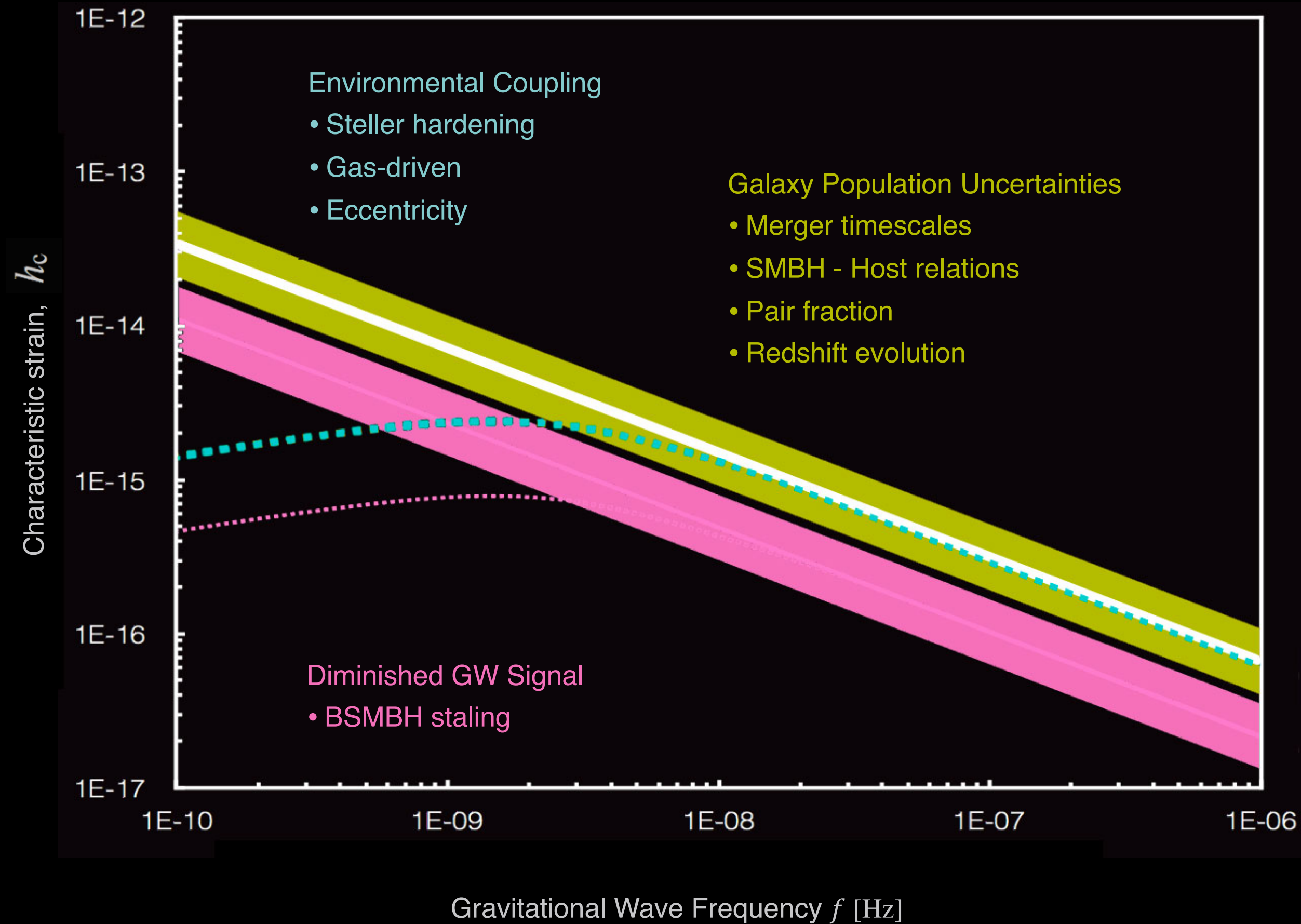


How Do Supermassive Black Holes Merge?



How Do Supermassive Black Holes Merge?

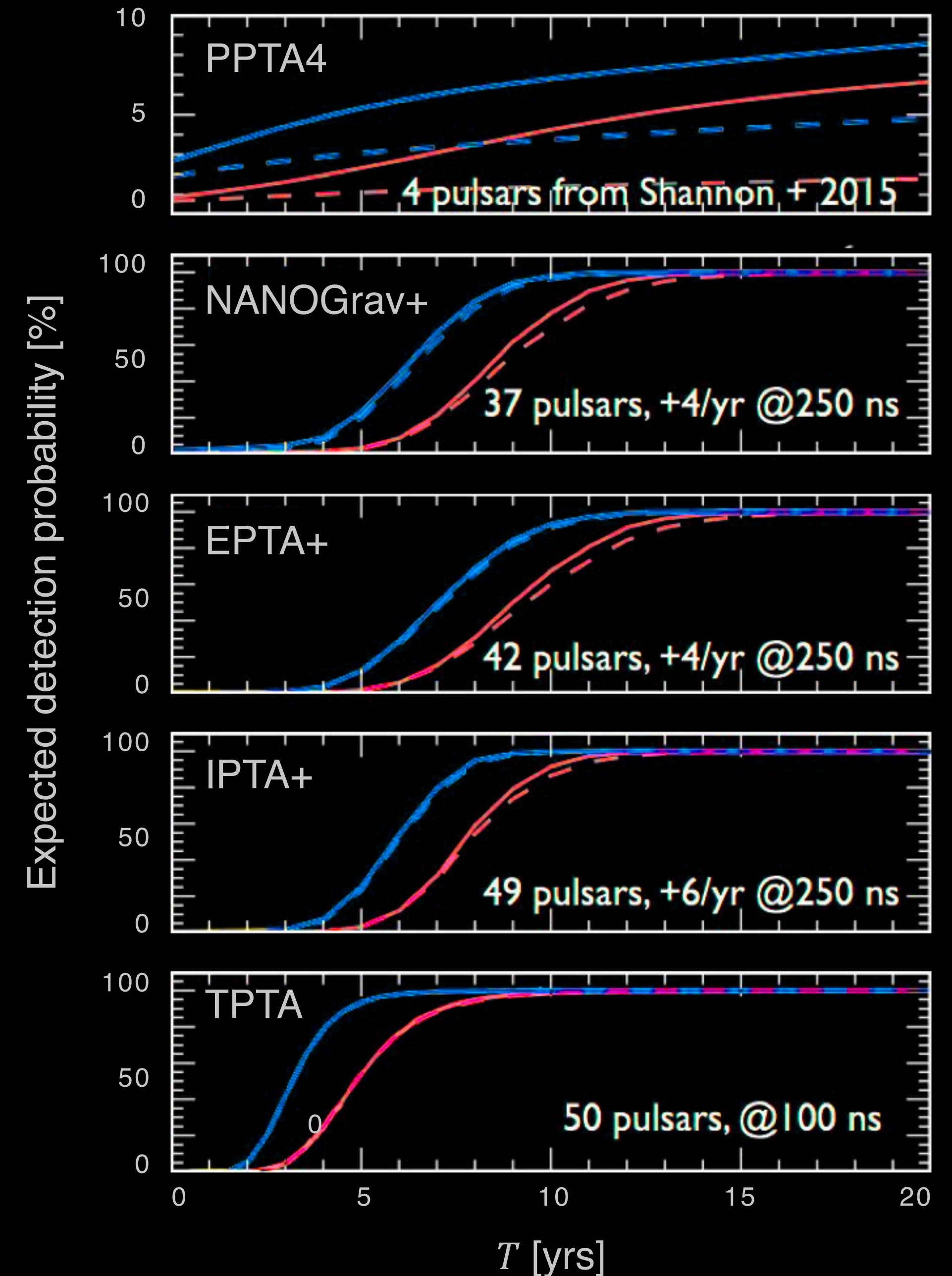
- Cosmic merger history of supermassive black holes forms a gravitational-wave background
- Shape informs us of final parsec solutions
- No gravitational waves = how does the universe work?
- **Constrained for the first time in 2016**



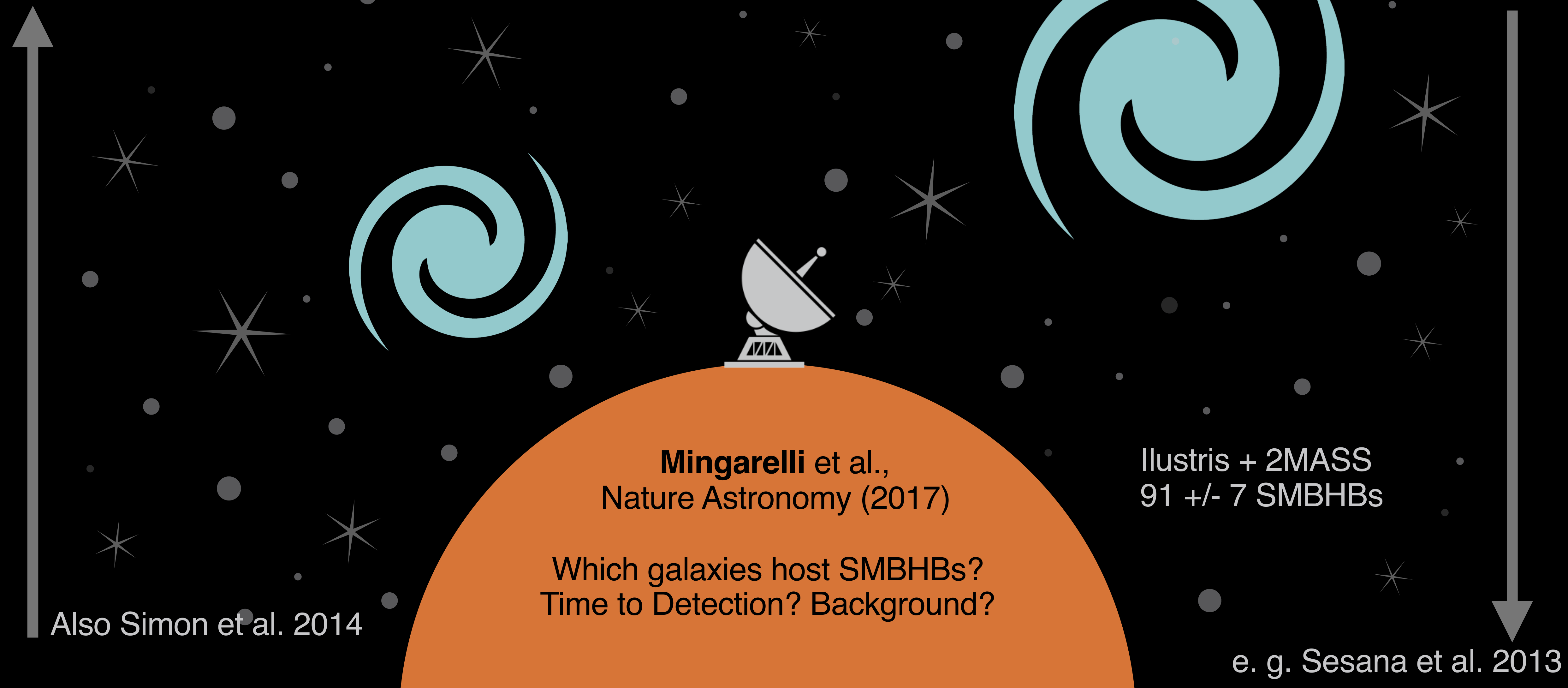
The NANOGrav nine-year data set: Limits on the isotropic stochastic gravitational wave background

Time to Detection of Gravitational-wave Background?

- What if Supermassive Black Holes stall?
blue line = no stalling, red line = 90% stalling
- First “hint” likely in next 2-3 years, faster with international collaboration
- No detection? Revisit galaxy evolution models, black hole masses, fundamental physics



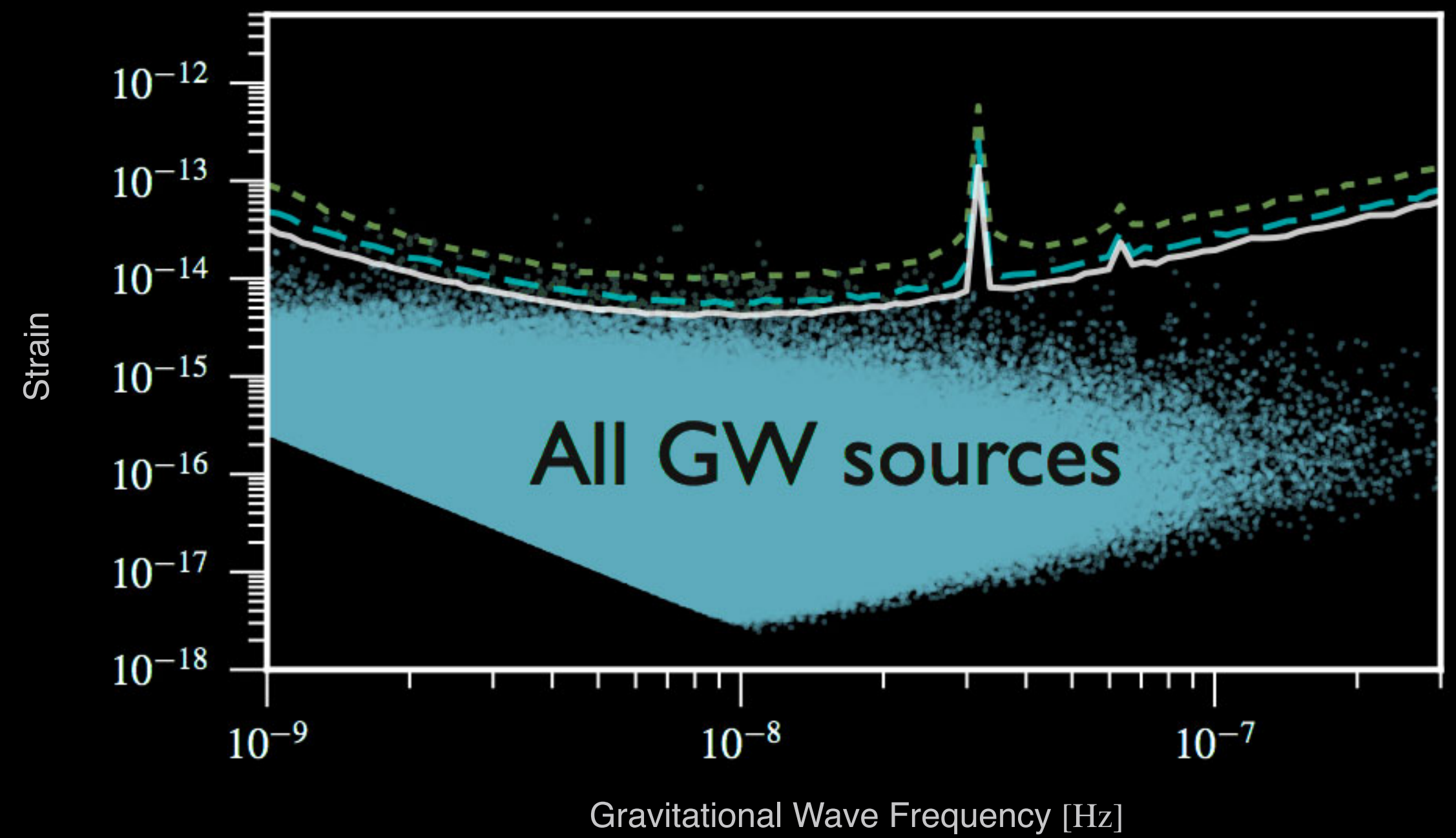
Resolvable Supermassive Black Holes



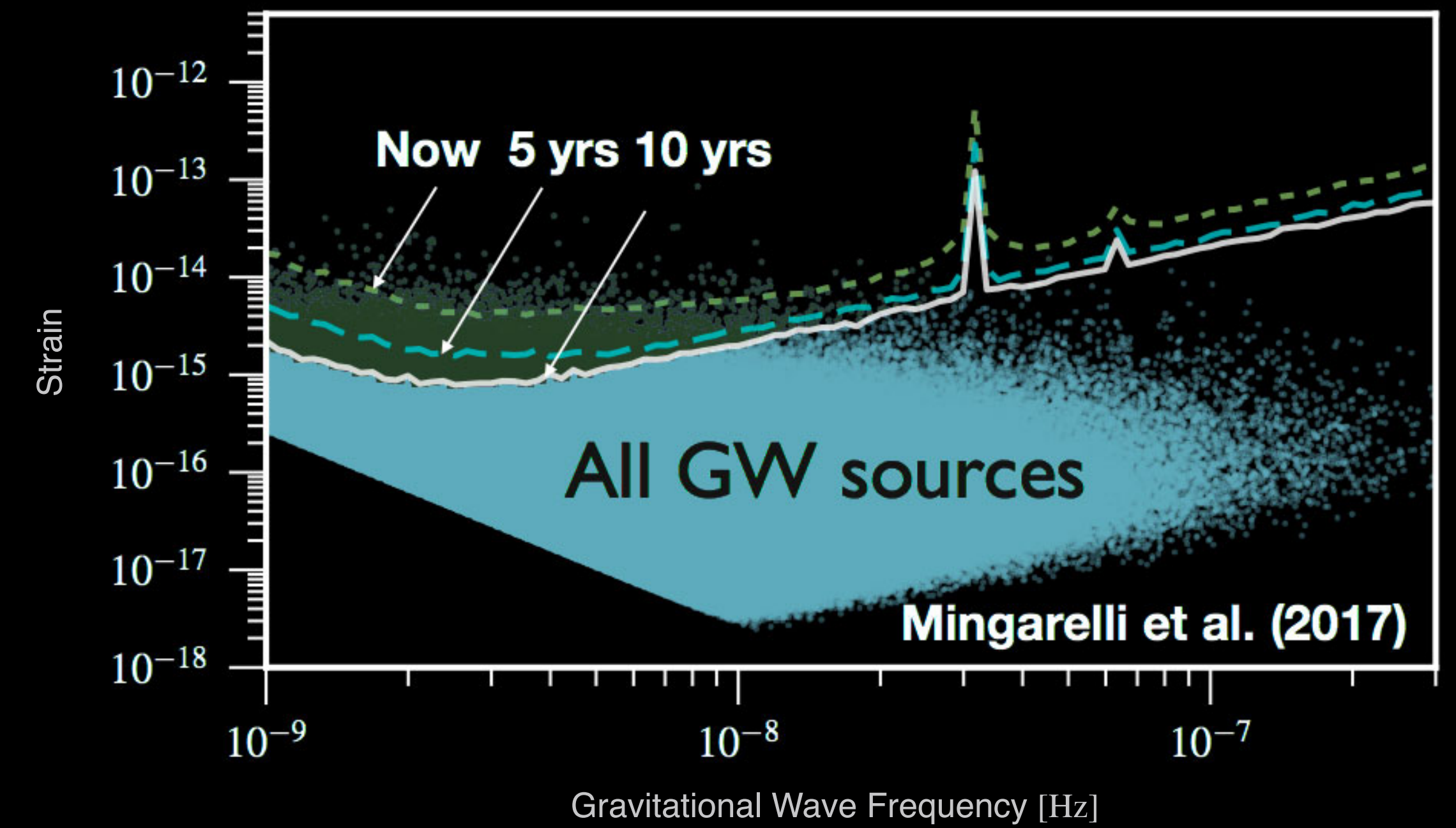
Time to Detection: Single Sources

Detection in < 10 years

Strong Red noise 3σ



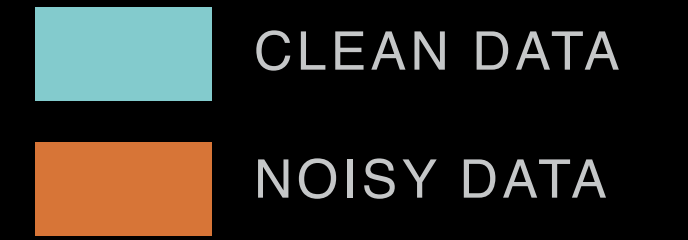
White noise 3σ



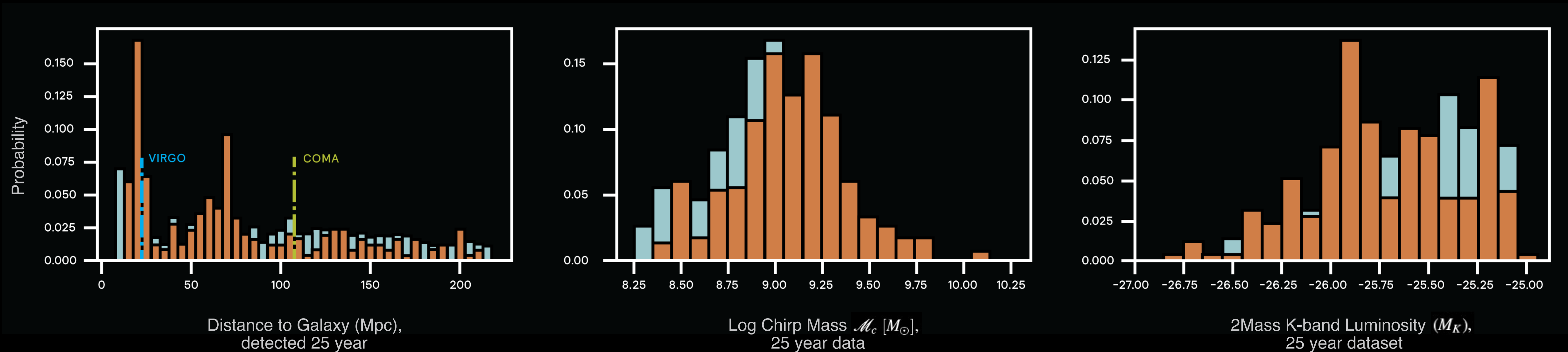
	FAP	Now	5 yrs	10 yrs
Factor of 4!	0.05 (2σ)	8%	96%	100%
sky location	3e-3 (3σ)	2%	36%	100%
	1e-4 (4σ)	1%	16%	100%

Can have > 1

Clues for EM Counterparts



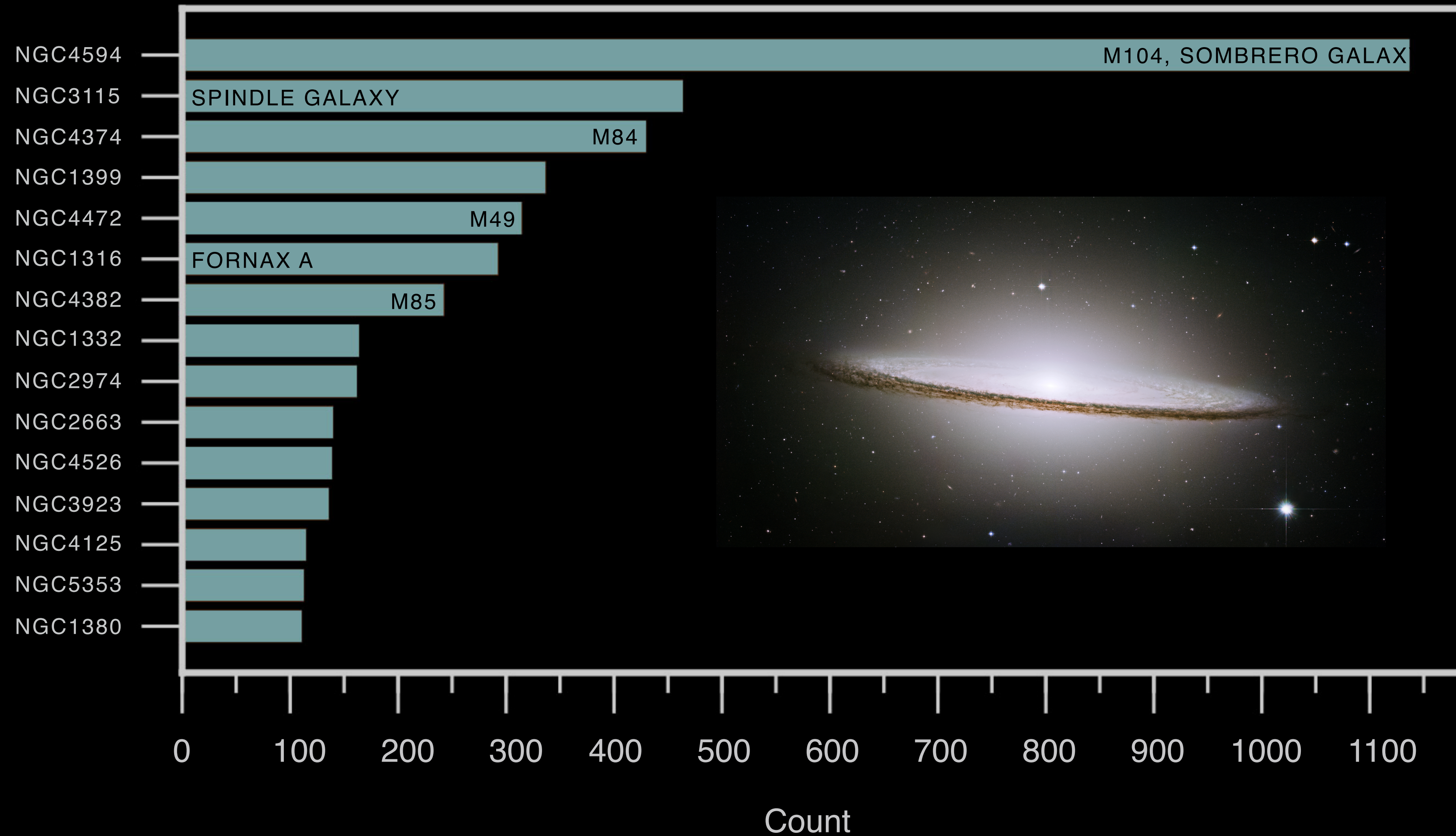
Crucial to identify host galaxies:
Can be many inside error box!



Is there even an EM counterpart?

Tang, Haiman, MacFadyen 2018; D’Orazio & Loeb 2017; many others!.

Hit List



Useful for Event Horizon Telescope

The Future of Pulsar Timing Arrays



- Unlock mysteries in galaxy evolution: how do galaxies grow and evolve?
- Measure supermassive black hole masses and spins over cosmic history
- Does gravity work as we expect it to when dealing with supermassive black holes?
- Most optimistic gravitational-wave background models **ruled out**

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