

EFT for Non-Standard neutrino Interactions

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October 12, 2018

Lot of interest and efforts (experiment and theory) in Neutrino physics.

New Physics can generate new Non-Standard Interactions (NSI) between neutrino and matter through new mediators.

NSI are not a new idea [[Wolfenstein '78](#)]:

$$\mathcal{L}_{NSI} = \frac{G_F}{\sqrt{2}} \sum_q (\bar{\nu}_\beta \gamma_\mu P_L \nu_\alpha) \left(\varepsilon_{\alpha\beta}^{qV} \bar{q} \gamma^\mu q + \varepsilon_{\alpha\beta}^{qA} \bar{q} \gamma^\mu \gamma_5 q \right)$$

Oscillations exp can only access these.

Measurements of Elastic neutrino-nucleus scattering [[Akimov et al.: 1708.01294](#)] can probe different NSI. (see also [K. Scholberg's talk from Tuesday](#))

Framework of NSI can be systematically described by an Effective Field Theory. [[W. Altmannshofer, MT, J. Zupan: 1810.xxxx](#)].

New Physics at some scale $\sim \Lambda$, typical energy at experiment $p \ll \Lambda \rightarrow$ organize the Lagrangian as a series in p/Λ .

$$\mathcal{L}_{eff} = \sum_{i,d} \hat{\mathcal{C}}_i^{(d)} \mathcal{O}_i^{(d)} \quad \text{where} \quad \hat{\mathcal{C}}_i^{(d)} = \frac{\mathcal{C}_i^{(d)}}{\Lambda^{d-4}}$$

The "full theory" (at Λ scale) is unknown \rightarrow fit Wilson coefficients to experiments.

- This Lagrangian gives a systematic expansion in p/Λ of the full theory;
- it predicts low energy phenomenology in terms of "few" parameters with (in principle) an arbitrary small theoretical uncertainty of $\mathcal{O}(p/\Lambda)^{n+1}$;
- Measurements of Wilson coefficients give info on NP scale Λ .

[[Les Houches 2017 lectures, Pich: 1804.05664 and Manohar: 1804.05863](#)]

NSI operator basis

3 - Flavors basis: $f = e, \mu, u, d, s$

Dimension five:

$$Q_1^{(5)} = \frac{e}{8\pi^2} (\bar{\nu}_\beta \sigma^{\mu\nu} P_L \nu_\alpha) F_{\mu\nu},$$

Dimension six:

$$Q_{1,f}^{(6)} = (\bar{\nu}_\beta \gamma_\mu P_L \nu_\alpha) (\bar{f} \gamma^\mu f), \quad Q_{2,f}^{(6)} = (\bar{\nu}_\beta \gamma_\mu P_L \nu_\alpha) (\bar{f} \gamma^\mu \gamma_5 f),$$

Dimension seven:

$$\begin{aligned} Q_1^{(7)} &= \frac{\alpha}{12\pi} (\bar{\nu}_\beta P_L \nu_\alpha) F^{\mu\nu} F_{\mu\nu}, & Q_2^{(7)} &= \frac{\alpha}{8\pi} (\bar{\nu}_\beta P_L \nu_\alpha) F^{a\mu\nu} \tilde{F}_{\mu\nu}, \\ Q_3^{(7)} &= \frac{\alpha_s}{12\pi} (\bar{\nu}_\beta P_L \nu_\alpha) G^{a\mu\nu} G_{\mu\nu}^a, & Q_4^{(7)} &= \frac{\alpha_s}{8\pi} (\bar{\nu}_\beta P_L \nu_\alpha) G^{a\mu\nu} \tilde{G}_{\mu\nu}^a, \\ Q_{5,f}^{(7)} &= m_f (\bar{\nu}_\beta P_L \nu_\alpha) (\bar{f} f), & Q_{6,f}^{(7)} &= m_f (\bar{\nu}_\beta P_L \nu_\alpha) (\bar{f} i \gamma_5 f), \\ Q_{7,f}^{(7)} &= m_f (\bar{\nu}_\beta \sigma^{\mu\nu} P_L \nu_\alpha) (\bar{f} \sigma_{\mu\nu} f), & Q_{8,f}^{(7)} &= (\bar{\nu}_\beta i \overleftrightarrow{\partial}_\mu P_L \nu_\alpha) (\bar{f} \gamma^\mu f), \\ Q_{9,f}^{(7)} &= (\bar{\nu}_\beta i \overleftrightarrow{\partial}_\mu P_L \nu_\alpha) (\bar{f} \gamma^\mu \gamma_5 f), & Q_{10,f}^{(7)} &= \partial_\mu (\bar{\nu}_\beta \sigma^{\mu\nu} P_L \nu_\alpha) (\bar{f} \gamma_\nu f), \\ Q_{11,f}^{(7)} &= \partial_\mu (\bar{\nu}_\beta \sigma^{\mu\nu} P_L \nu_\alpha) (\bar{f} \gamma_\nu \gamma_5 f). \end{aligned}$$

Oscillations

Matter effect on neutrino oscillations (MSW effect). Described by an effective potential due to neutrino interactions with electrons and nuclei (for non-relativistic, electrically neutral and unpolarized medium)

Vector interactions: $\bar{\nu}\gamma_\mu P_L\nu$ (in the $q \rightarrow 0$ limit)

$$\mathcal{V}_{\text{eff}}^{(-)} \Big|_{\text{NSI}} \simeq G_F n_f \left(1 + \varepsilon_{\alpha\beta}^{fV}\right), \quad \mathcal{V}_{\text{eff}}^{(+)} \sim \mathcal{O}\left(\frac{m_\nu^2}{E_\nu}\right)$$

New interactions:

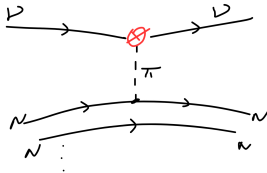
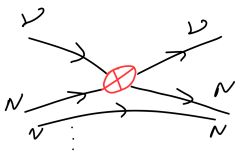
$$S : \bar{\nu} P_L \nu, \quad \bar{\nu}_\beta \overset{\leftrightarrow}{i}\partial_\mu P_L \nu_\alpha \rightarrow \mathcal{V}_{\text{eff}}^S \propto \frac{m_\nu}{E_\nu},$$

$$T : \bar{\nu} \sigma_{\mu\nu} P_L \nu, \quad \partial_\mu (\bar{\nu}_\beta \sigma^{\mu\nu} P_L \nu_\alpha) \rightarrow \mathcal{V}_{\text{eff}}^T = 0,$$

N.B.: tensor operator can contribute in polarized mediums [[Bergmann, Grossmann and Nardi: 9903517](#)]

Elastic neutrino-nucleus scattering

Very low-energy neutrinos ($E_\nu \sim q \sim \mathcal{O}(10)$ MeV) \rightarrow elastic interaction with non-relativistic nuclei in the detector.



At LO single nucleon interaction \rightarrow hadronization described with form factors for single-nucleon currents [Bishara et al.: 1611.00368, 1707.06998]

$$\langle N' | \bar{q} \gamma^\mu q | N \rangle = \bar{u}'_N \left[F_1^{q/N}(q^2) \gamma^\mu + \frac{i}{2m_N} F_2^{q/N}(q^2) \sigma^{\mu\nu} q_\nu \right] u_N,$$

$$\langle N' | \frac{\alpha_s}{8\pi} G^{a\mu\nu} \tilde{G}_{\mu\nu}^a | N \rangle = F_G^N(q^2) \bar{u}'_N i \gamma_5 u_N,$$

At small q^2

$$F_i(q^2) = F_i(0) + F'_i(0)q^2 + \dots, \quad F_G^N(q^2) = \frac{q^2}{m_\pi^2 - q^2} a_{\tilde{G},\pi}^N + b_{\tilde{G}}^N + \dots,$$

Complex structure of nuclei is described by nuclear response functions
[Fitzpatrick et al.: 1203.3542]

- *Spin-Independent*: $\sigma_{tot} \propto \sigma_N A^2 \rightarrow$ prefer high A materials for detectors (for CsI $A \simeq 130 \rightarrow 10^4$ enhancement);

Induced by vector and scalar operators.

- *Spin-Dependent*: $\sigma_{tot} \propto \sigma_N S_A$ (nuclear spin $S_A \sim \mathcal{O}(1)$);

Induced by axial, pseudoscalars and tensor operators.

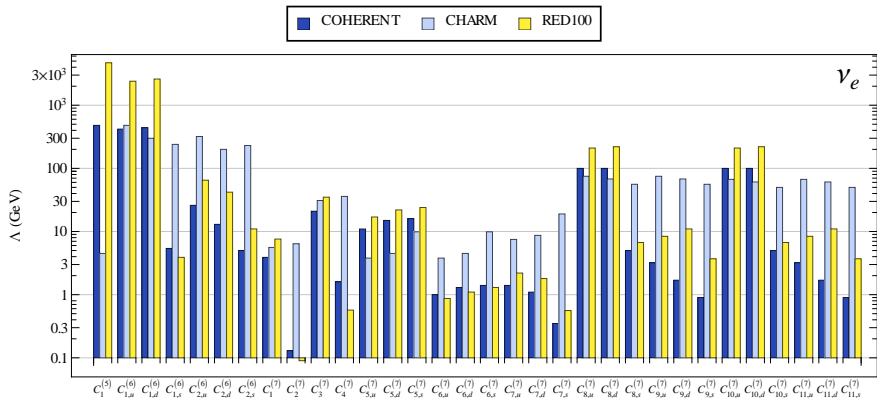
Differential cross section:

$$\frac{d\sigma}{dE_R}(E_\nu) \sim R_\nu W_N, \quad R_\nu \left(\hat{C}_i^{(d)}, E_\nu, q^2 \right) \text{ neutrino kinematics}$$

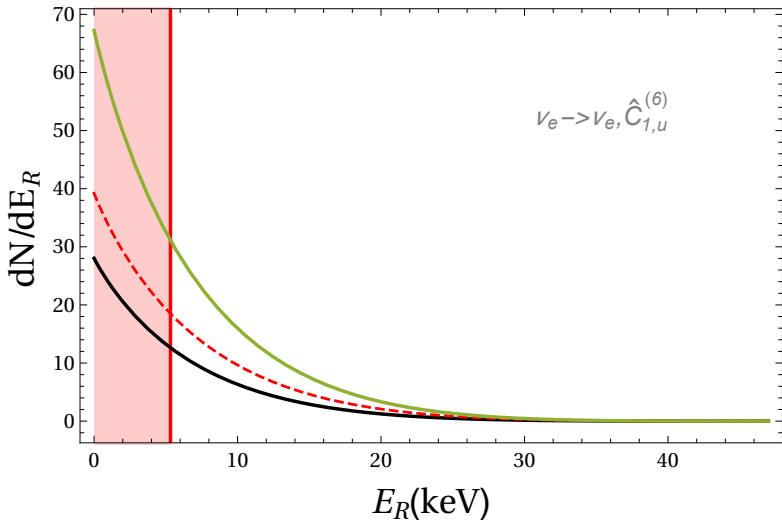
N.B.: form factors carry uncertainty that affect theoretical prediction.

Results (preliminary!)

Assume $C_i^{(d)} = 1 \rightarrow$ put lower limit on Λ

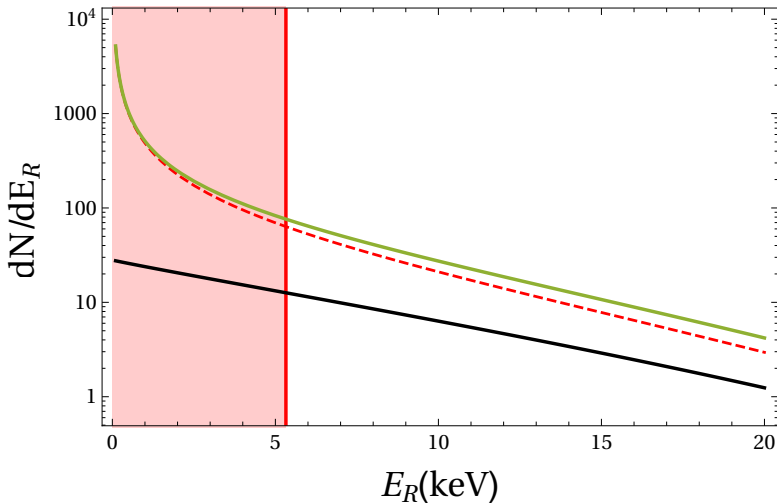


Reduce uncertainties, e.g. $\mathcal{Q}_{1,f}^{(6)} = (\bar{\nu}_\beta \gamma_\mu P_L \nu_\alpha)(\bar{f} \gamma^\mu f)$.



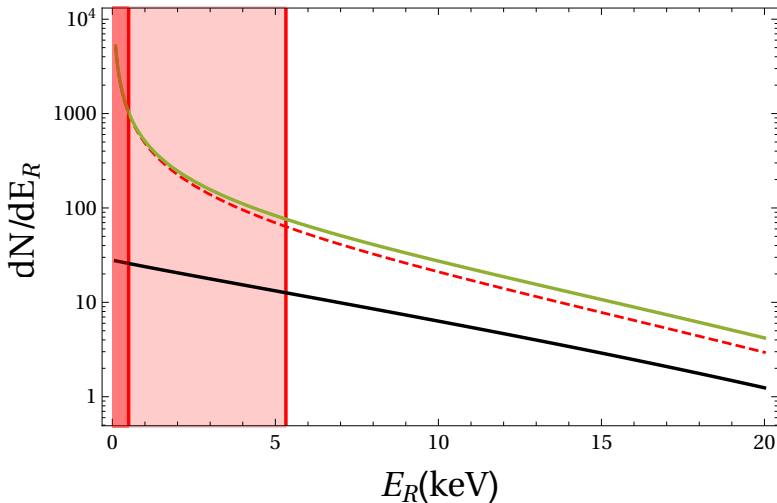
$$R_\nu \propto 4E_\nu^2 - \vec{q}^2 \quad \hat{C}_{1,u}^{(6)} \sim \frac{10^{-6}}{\text{GeV}^2} \rightarrow \Lambda \sim 400 \text{ GeV}.$$

Lower threshold, e.g. $Q_1^{(5)} = \frac{e}{8\pi^2} (\bar{\nu}_\beta \sigma^{\mu\nu} P_L \nu_\alpha) F_{\mu\nu}$.



$$R_\nu \propto 1/\vec{q}^2 \quad \hat{C}_1^{(5)} \sim \frac{10^{-3}}{\text{GeV}} \rightarrow \Lambda \sim 1\text{TeV}.$$

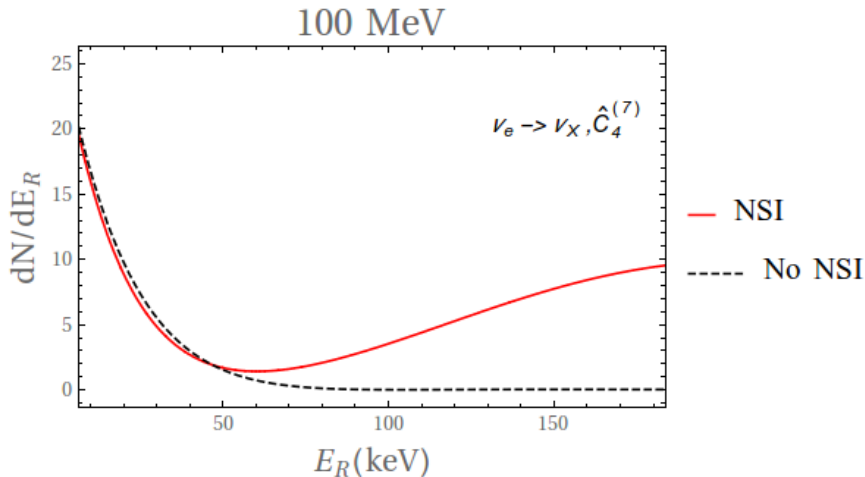
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$$R_\nu \propto 1/\vec{q}^2 \quad \hat{C}_1^{(5)} \sim \frac{10^{-3}}{\text{GeV}} \rightarrow \Lambda \sim 1\text{TeV}.$$

Higher energy, e.g.

$$Q_4^{(7)} = \frac{\alpha_s}{8\pi} (\bar{\nu}_\beta P_L \nu_\alpha) G^{a\mu\nu} \tilde{G}_{\mu\nu}^a \quad \text{or} \quad Q_{6,f}^{(7)} = m_f (\bar{\nu}_\beta P_L \nu_\alpha) (\bar{f} i \gamma_5 f)$$



$$R_\nu \propto \bar{q}^4 \quad \hat{C}_4^{(7)} \sim \frac{10^{-1}}{\text{GeV}^3} \rightarrow \Lambda \sim 2 \text{GeV}.$$

Conclusions

- Elastic neutrino - nucleus scattering provides a possible probe of NSI not accessible to oscillation;
- Need to reduce theoretical and experimental uncertainties;
- Probe different q^2 dependence by lower threshold and higher energy;
- Inelastic scattering;
- Various projects for experiment upgrades... wait for new data.

Thanks!