

A visualization of particle tracks, likely from a detector, showing several intersecting lines of blue and green points against a dark blue background.

Machine Learning for Particle Image Neutrino Detectors *at the HEP Intensity Frontier*

*Kazuhiro Terao & Laura Domine
SLAC National Accelerator Lab.*

ACAT @ Saas-Fee (11-15 Mar. 2019)

A visualization of particle tracks on a blue background. Several lines, primarily blue and green, originate from a point at the top left and extend across the frame. Some tracks are straight, while others show slight curves or branching. Small colored dots (yellow, green, red) are scattered along the tracks.

Outline:

- Neutrino detectors
- Machine learning applications
- Toward 3D ML-based data reconstruction

Me: Neutrino Physicist

- **Neutrinos?**
 - One of least understood elementary particles

Me: Neutrino Physicist

- **Neutrinos?**

- One of least understood elementary particles

- **They are everywhere**

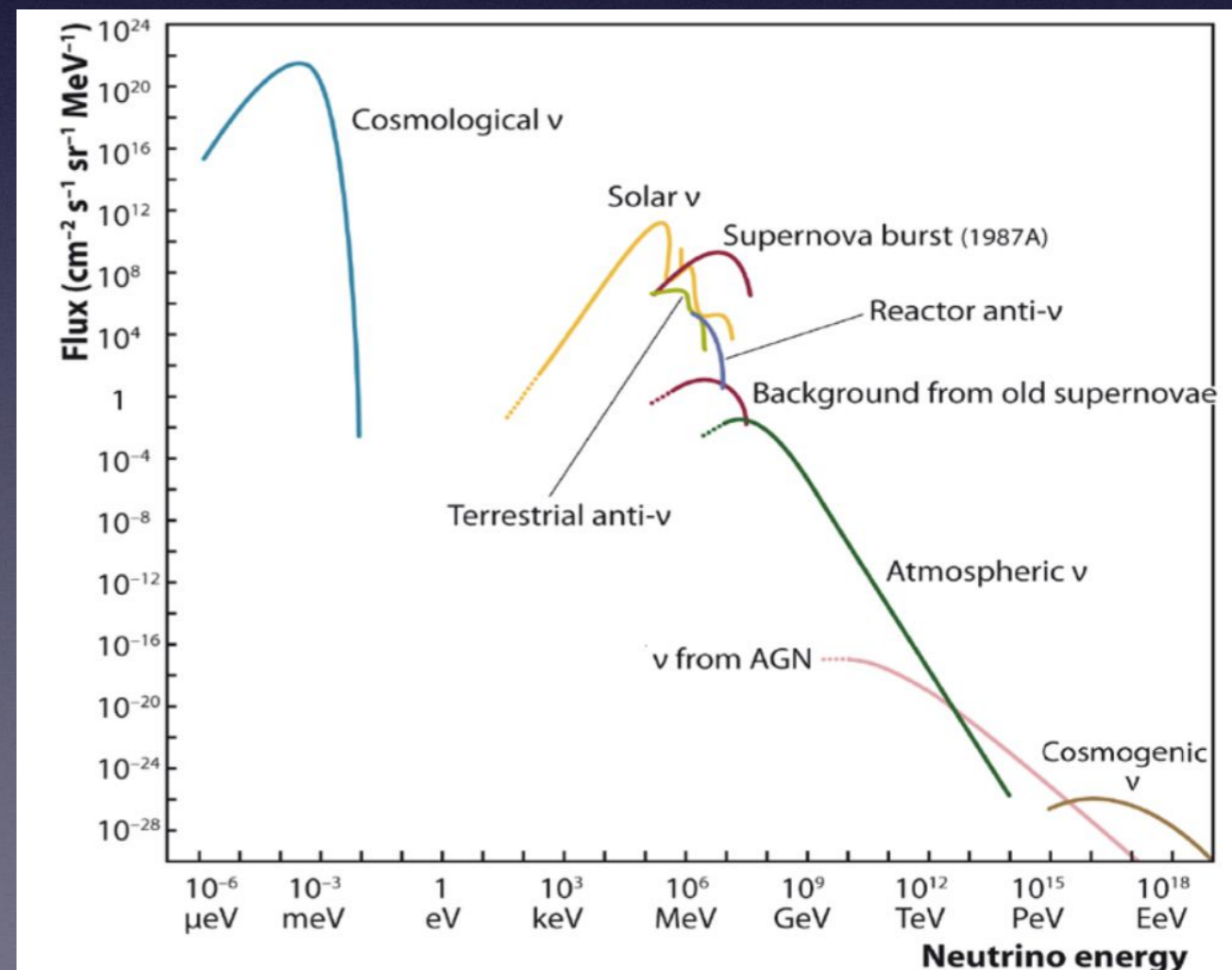
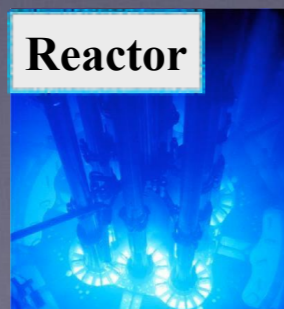
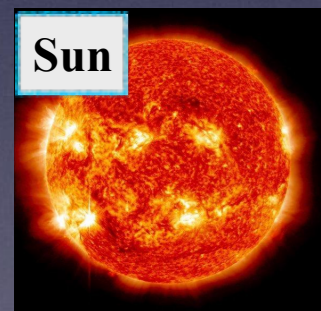
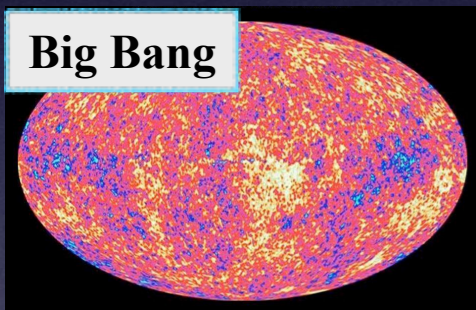
- 400 trillion neutrinos pass your body every second

- Your body generates ~340 million neutrinos a day

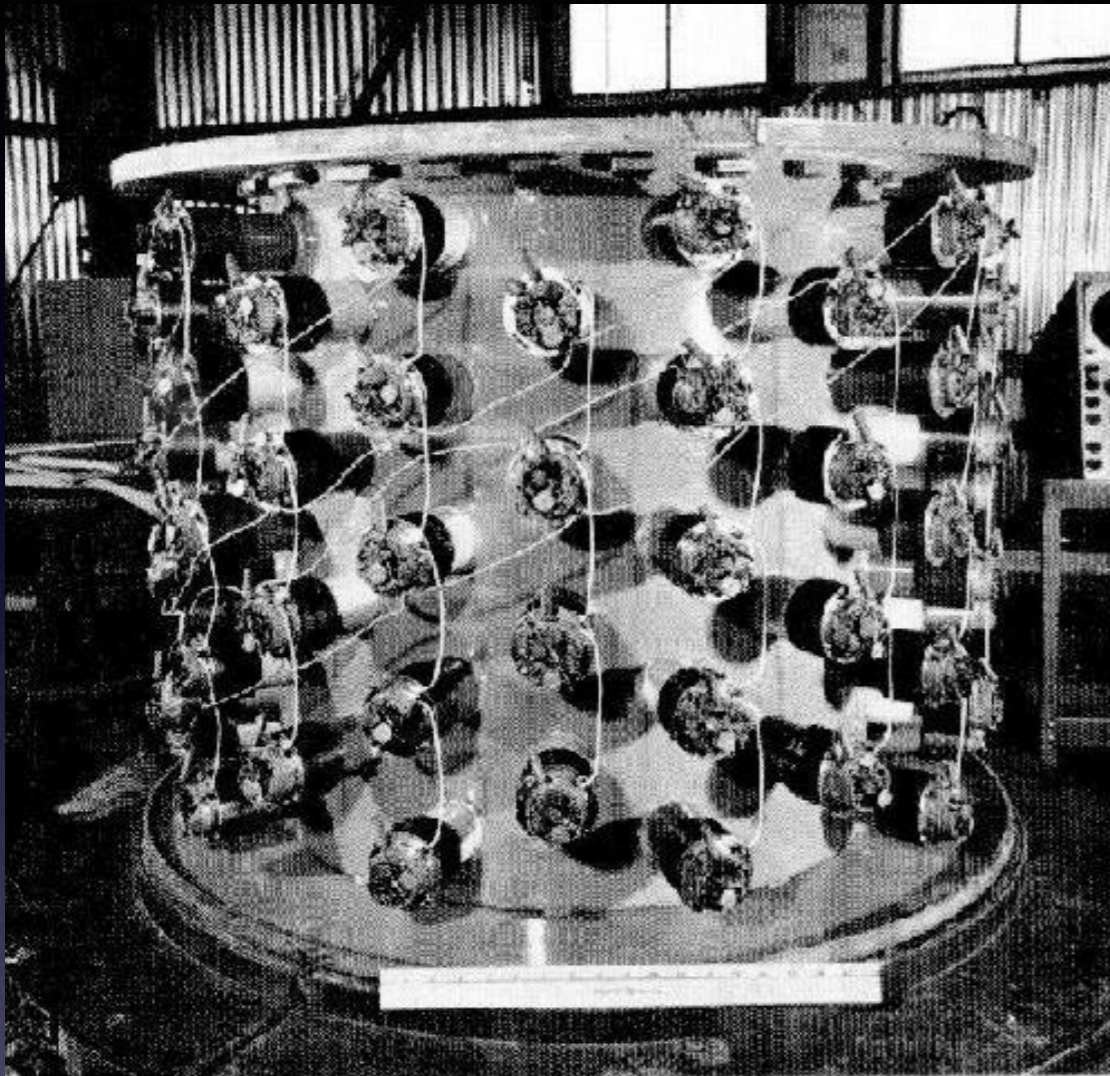
- **They come from everywhere**

100s
neutrinos/second
from our Sun

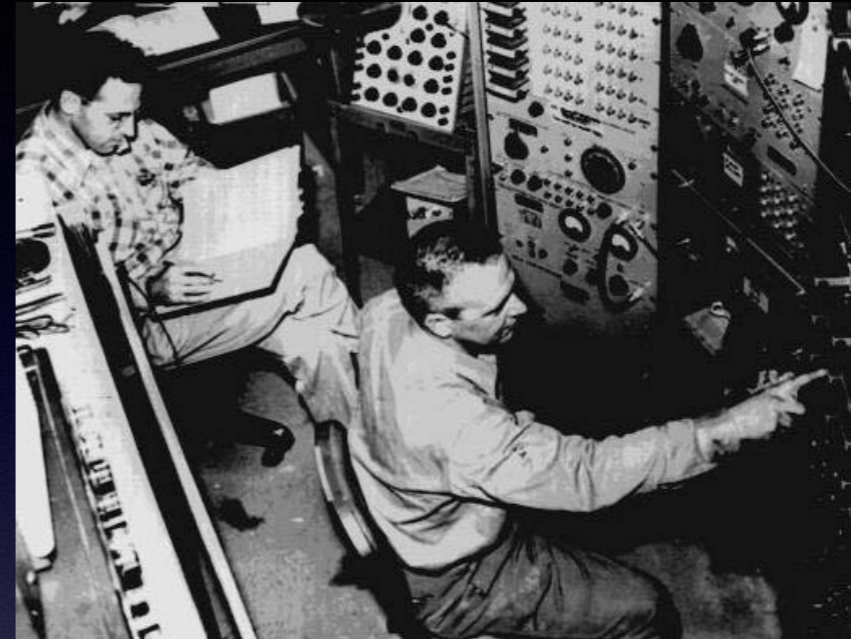
EPJ H37 (2012) 3:515-565



Early days neutrino detection



Cd-doped water
0.4 ton, 100 PMTs
(1956)



Inverse Beta Decay (IBD)



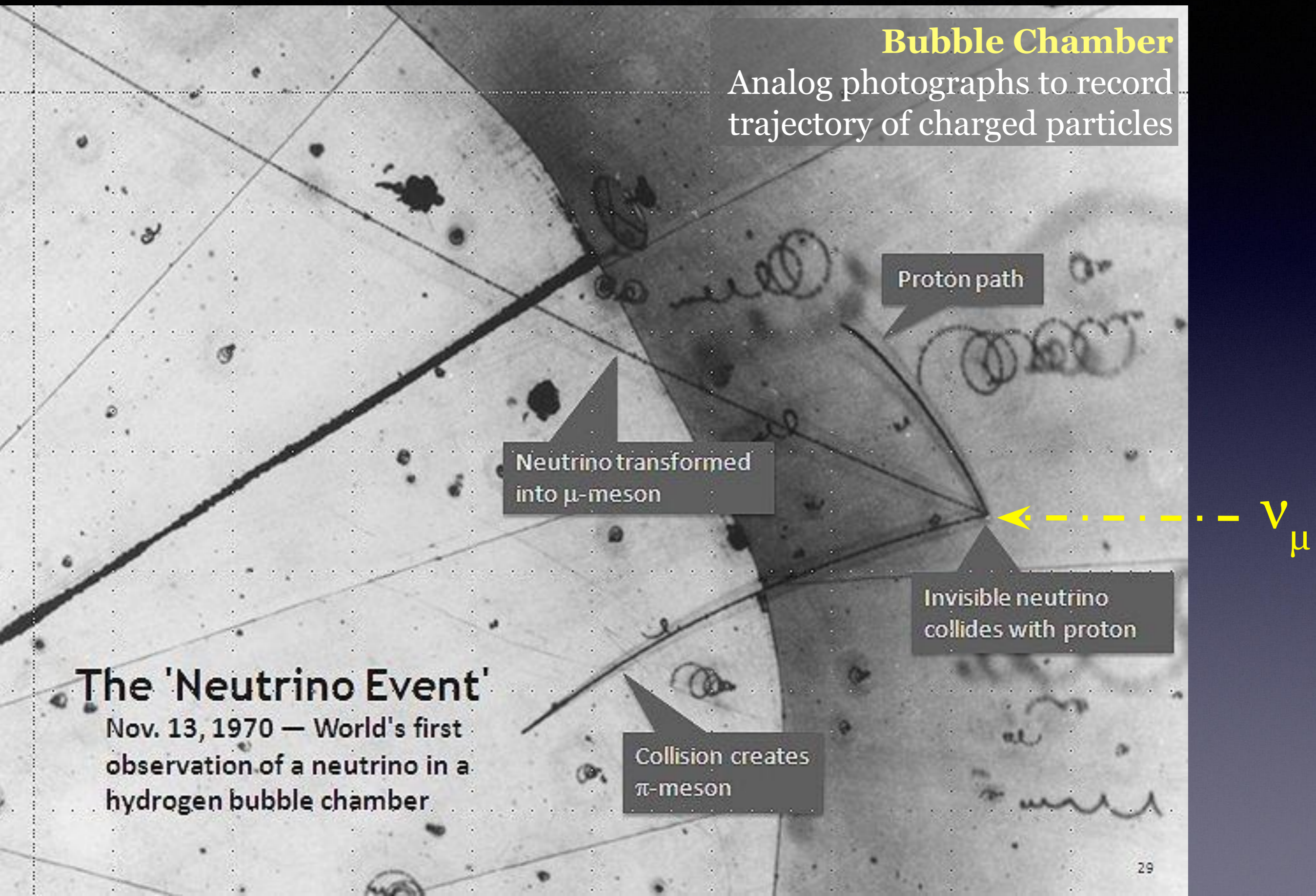
by Reines & Cowan (Nobel Prize 1995)

**First neutrino
detection**

Early days particle imaging

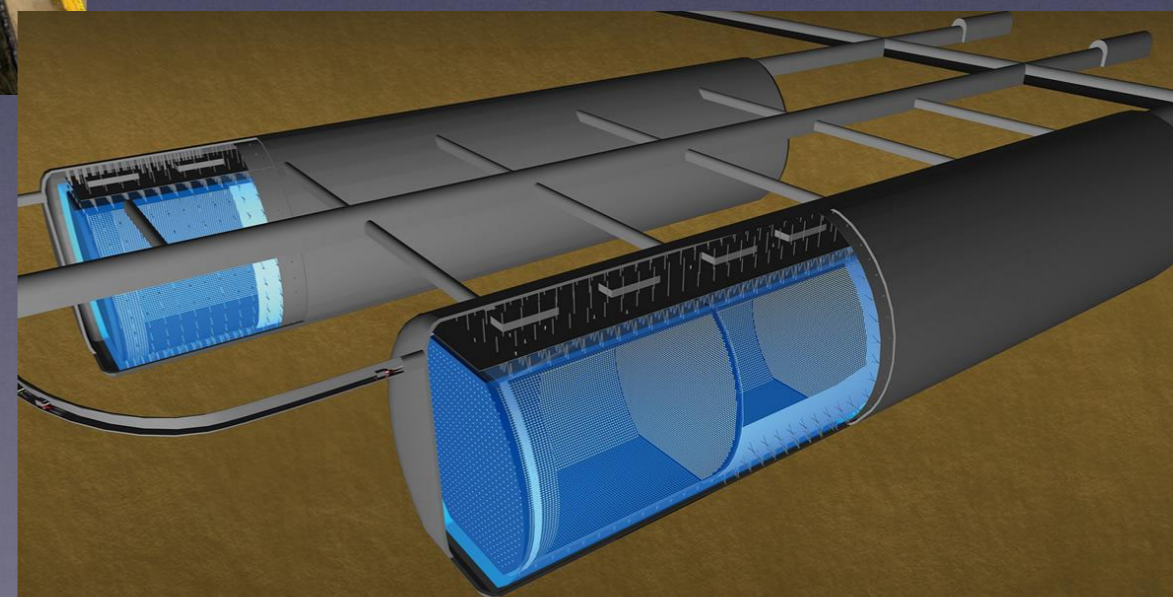
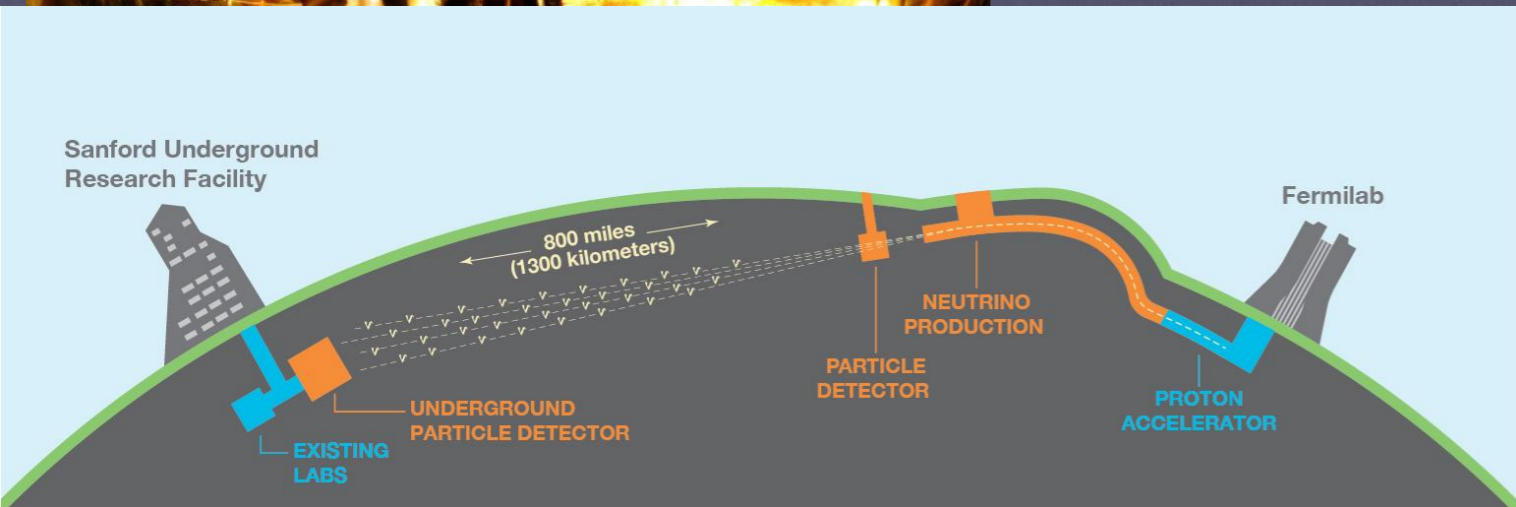
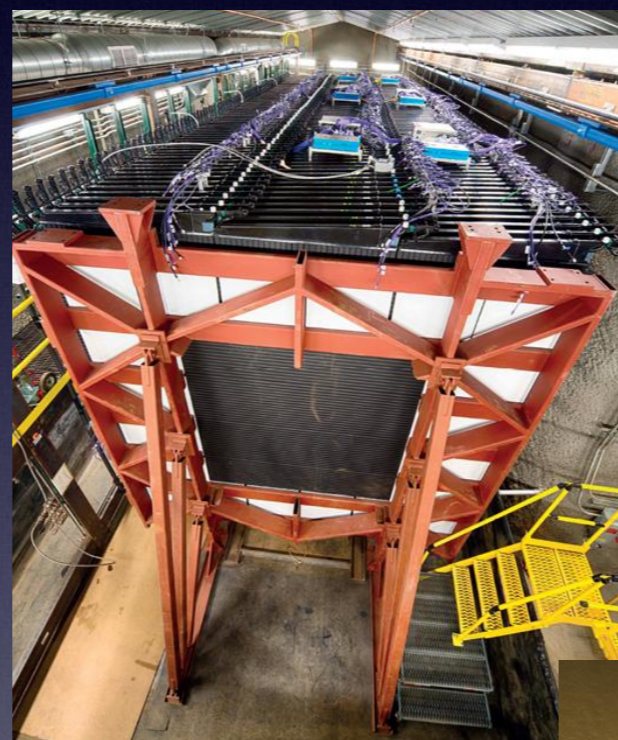
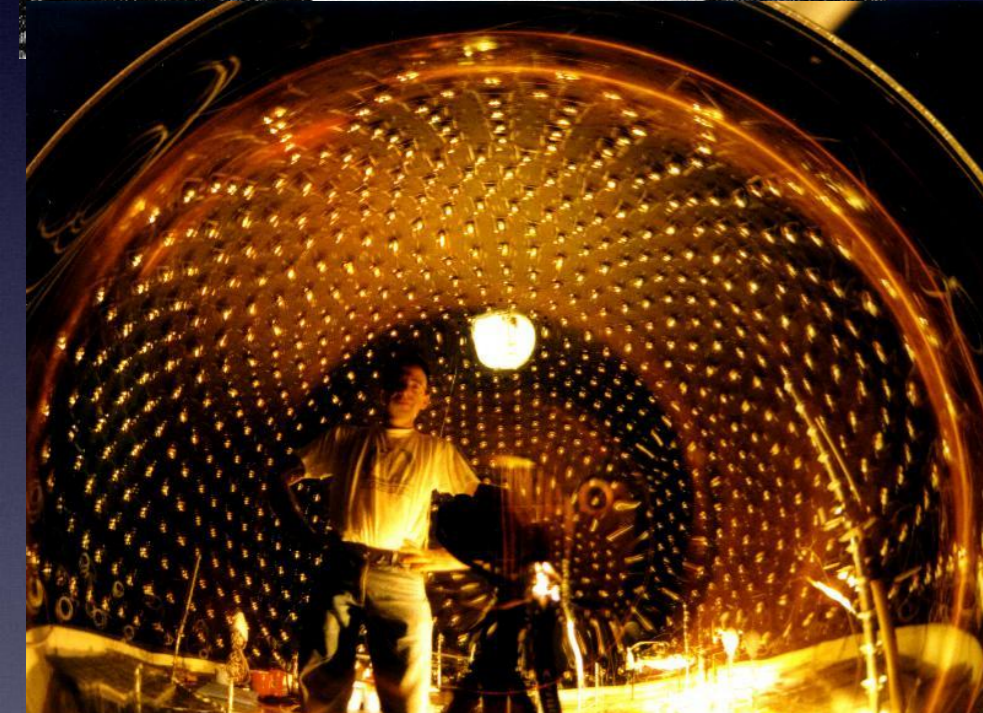
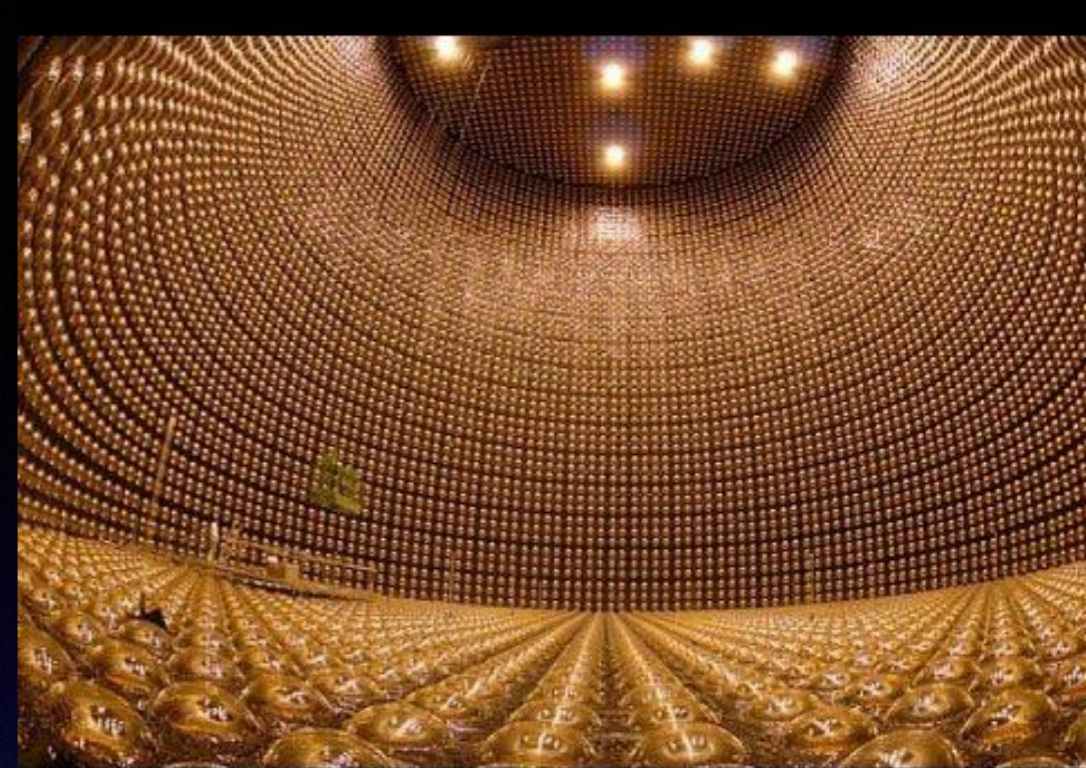
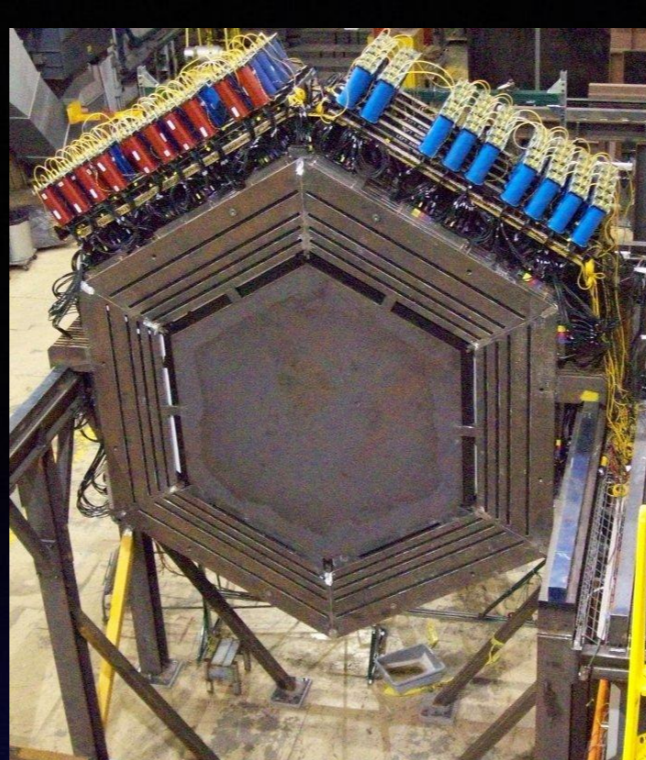
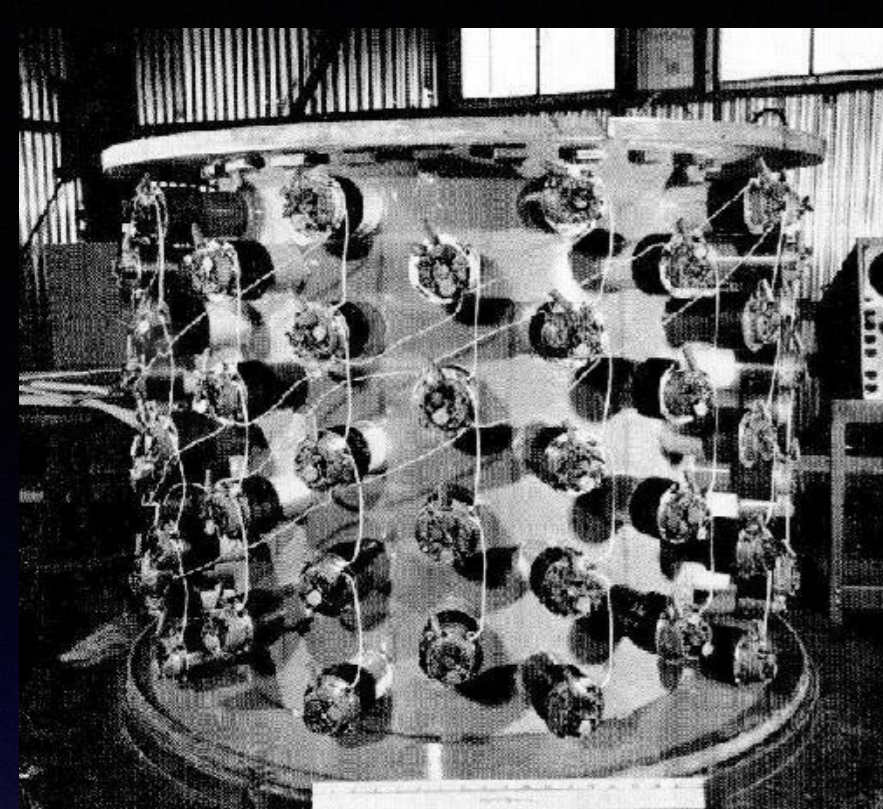
Bubble Chamber

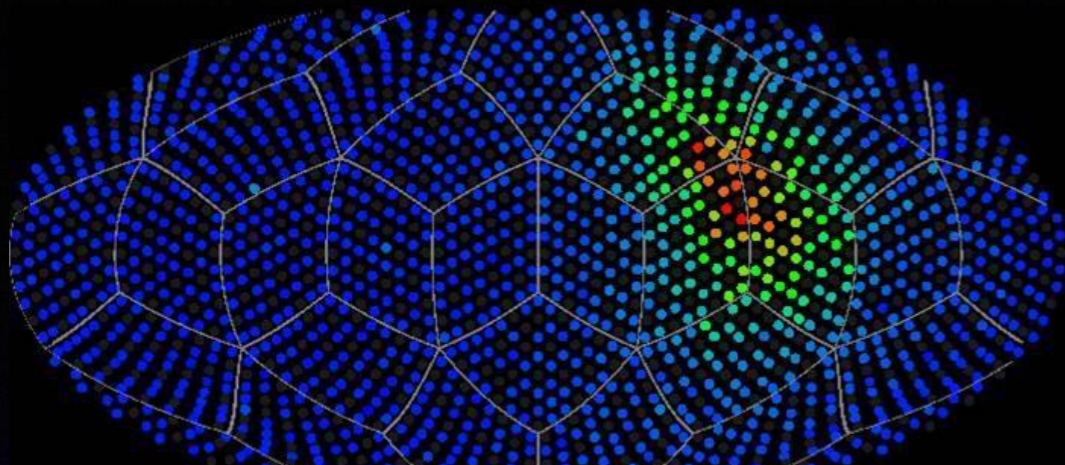
Analog photographs to record trajectory of charged particles



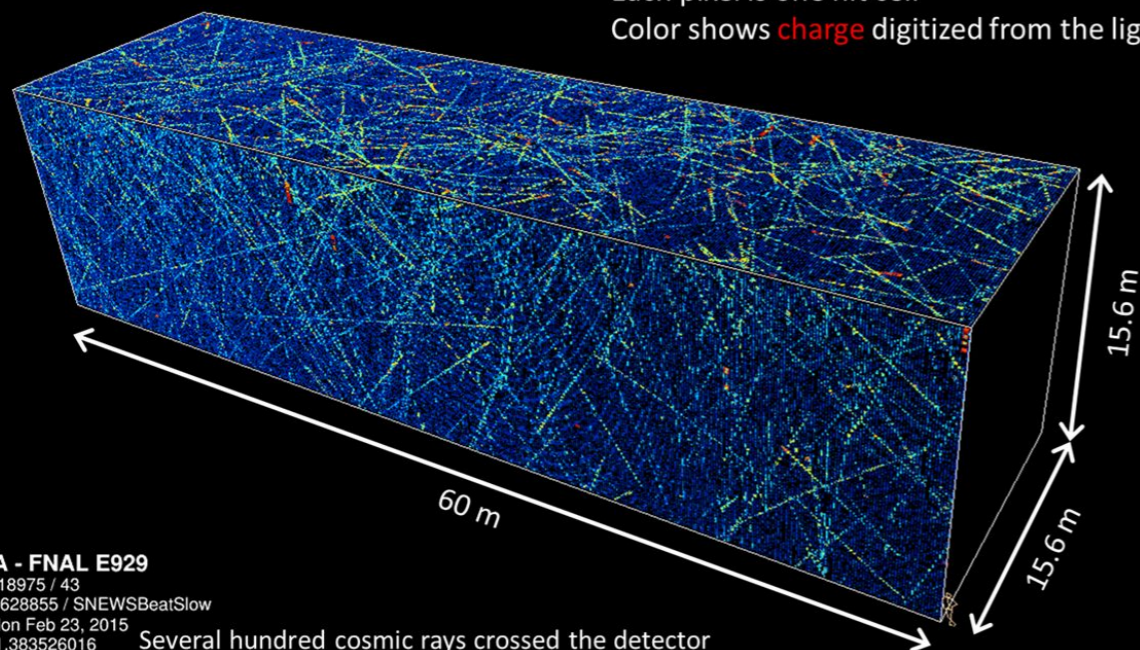
The 'Neutrino Event'

Nov. 13, 1970 — World's first observation of a neutrino in a hydrogen bubble chamber



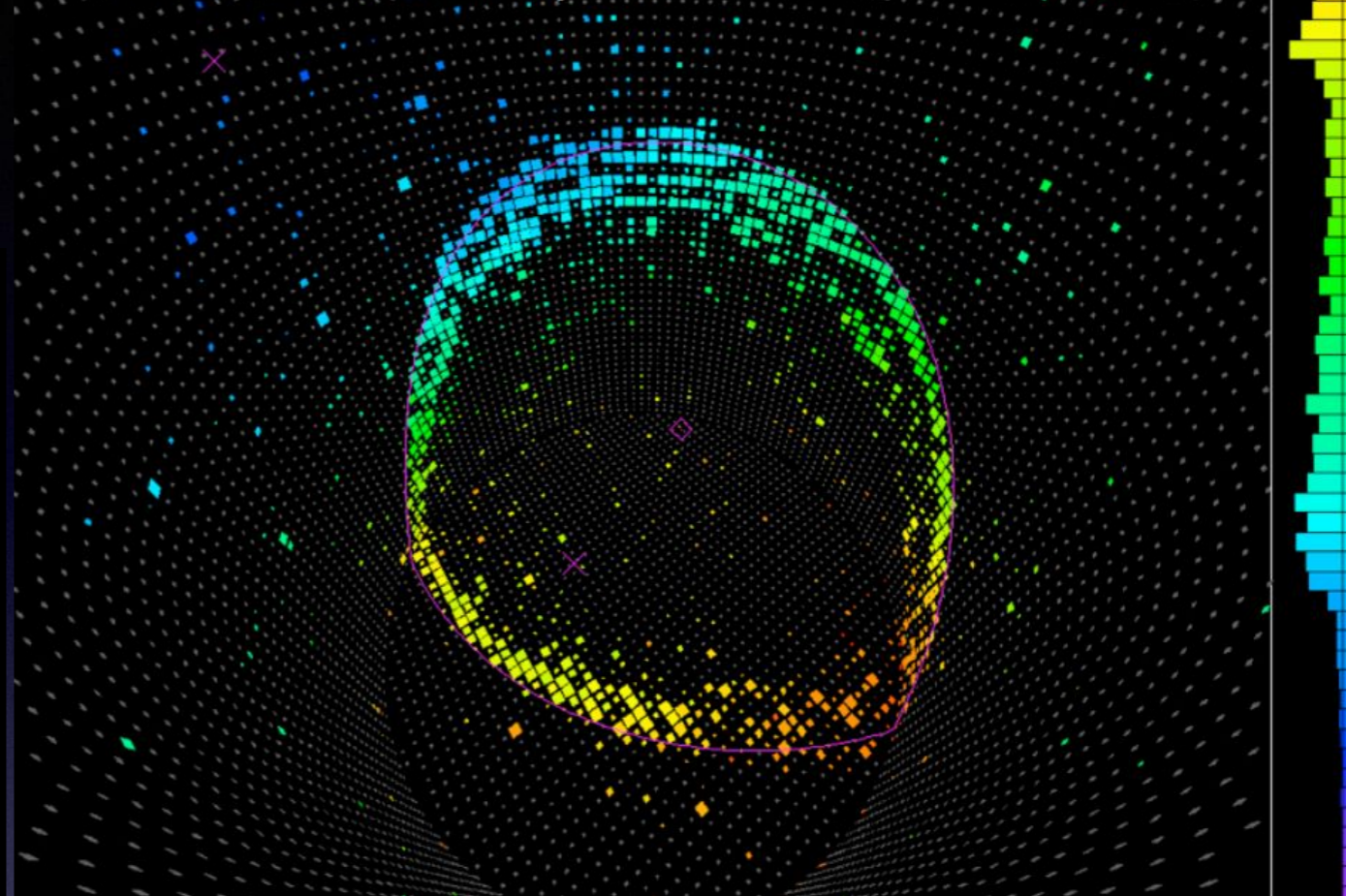


5ms of data at the NOvA Far Detector
 Each pixel is one hit cell
 Color shows **charge** digitized from the light

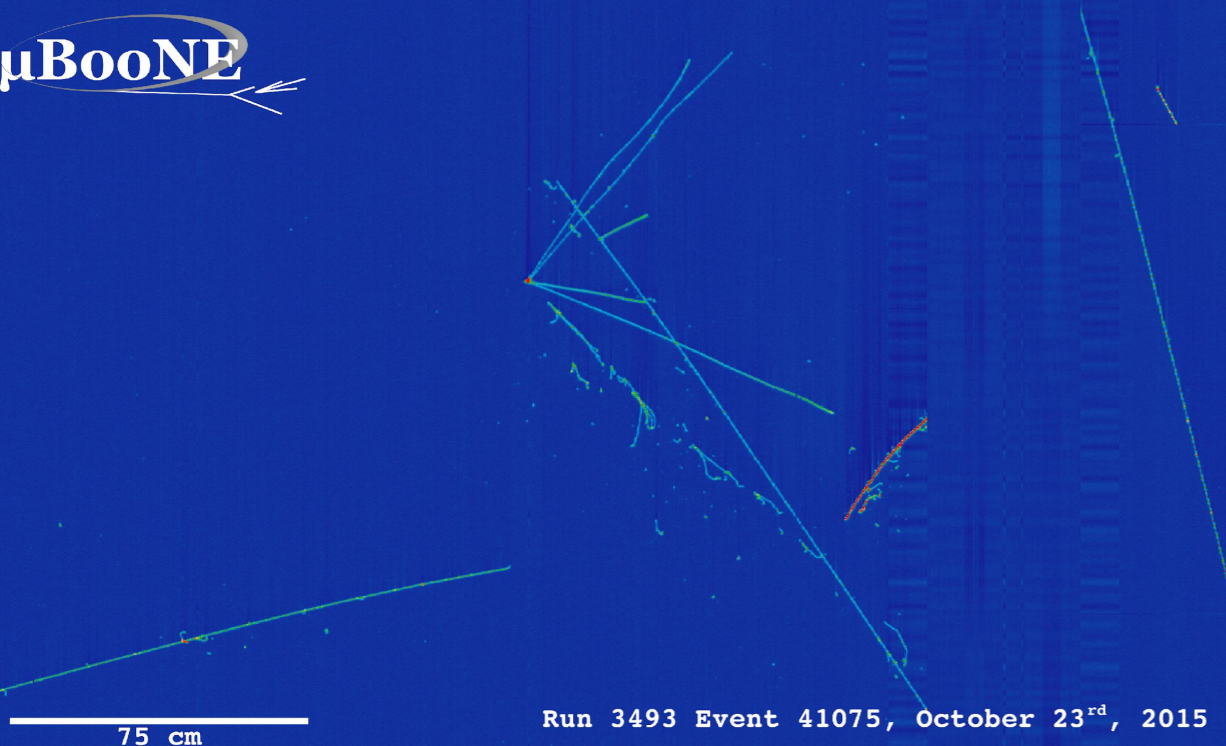


NOvA - FNAL E929
 Run: 18975 / 43
 Event: 628855 / SNEWSBeatSlow
 UTC Mon Feb 23, 2015
 14:30:1.383526016 Several hundred cosmic rays crossed the detector

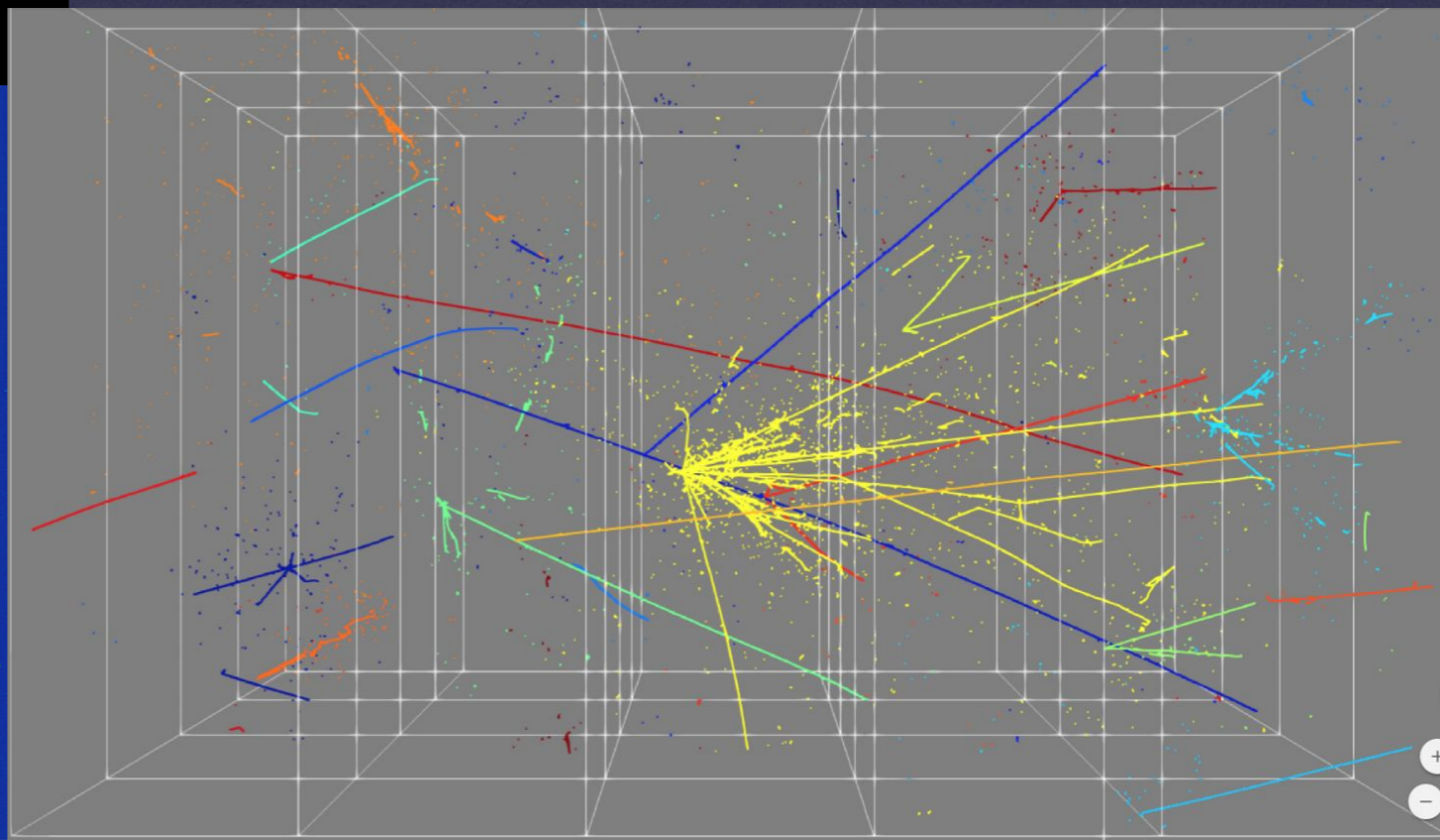
A 603MeV muon in Super-K.

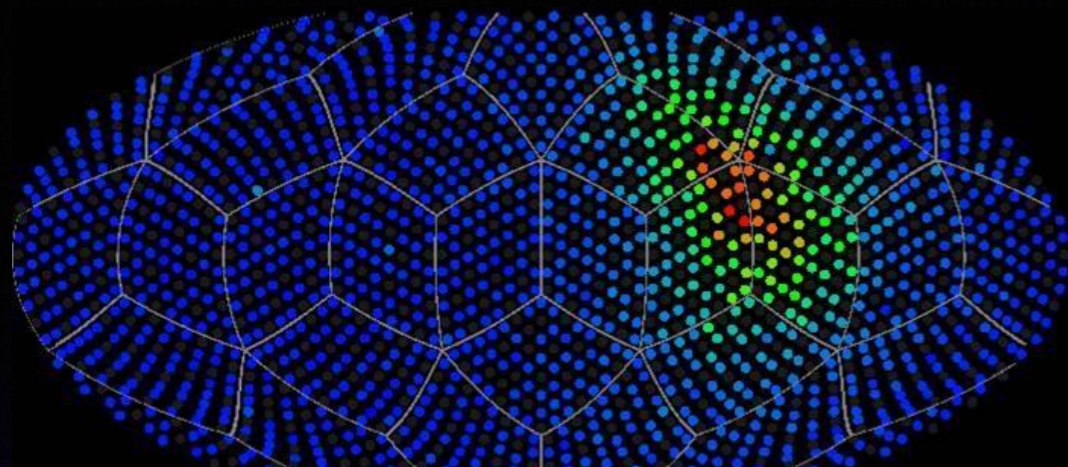


μBooNE

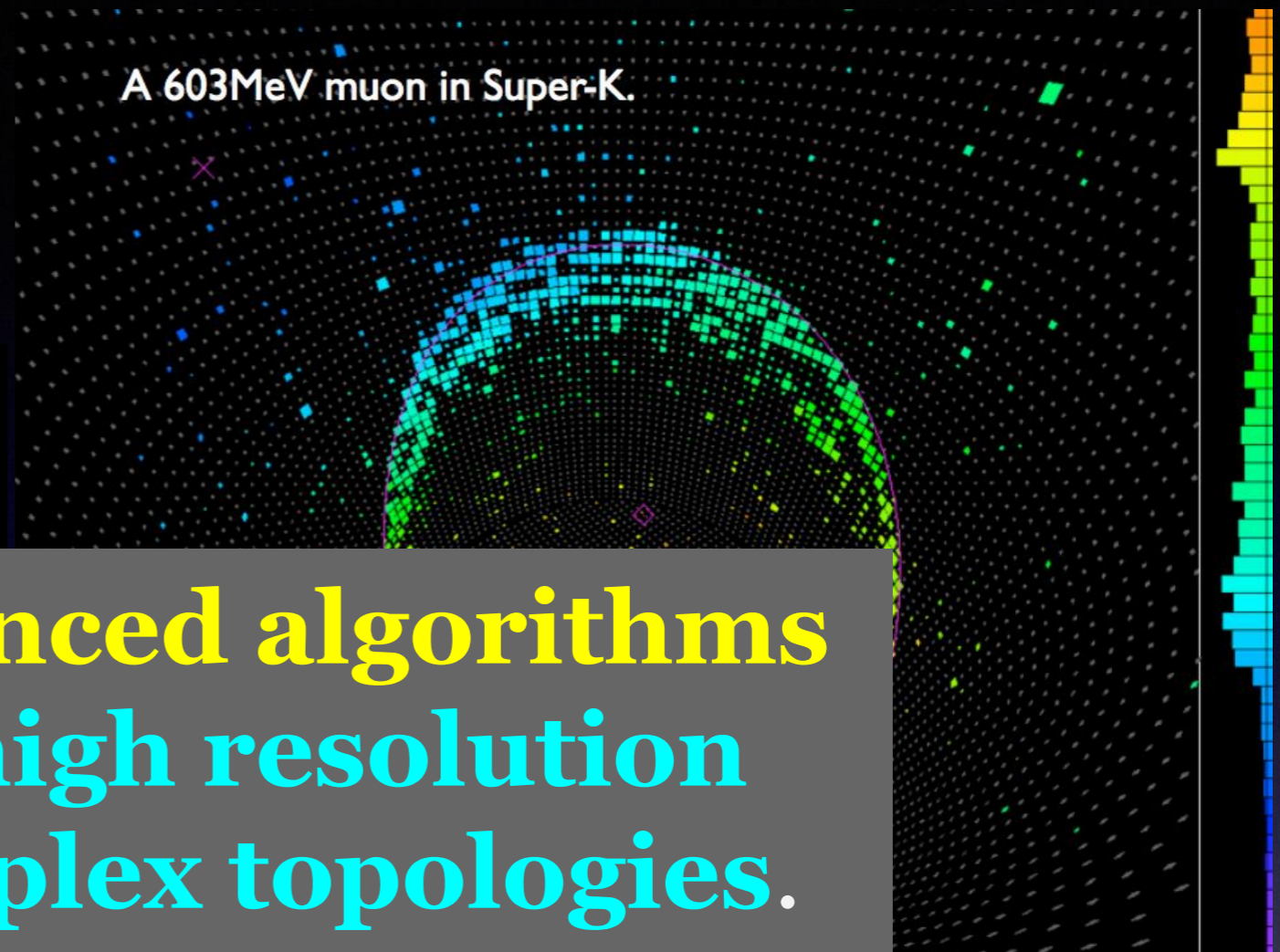


Run 3493 Event 41075, October 23rd, 2015





5ms of data at the NOvA Far Detector
Each pixel is one hit cell
Color shows **charge** digitized from the light

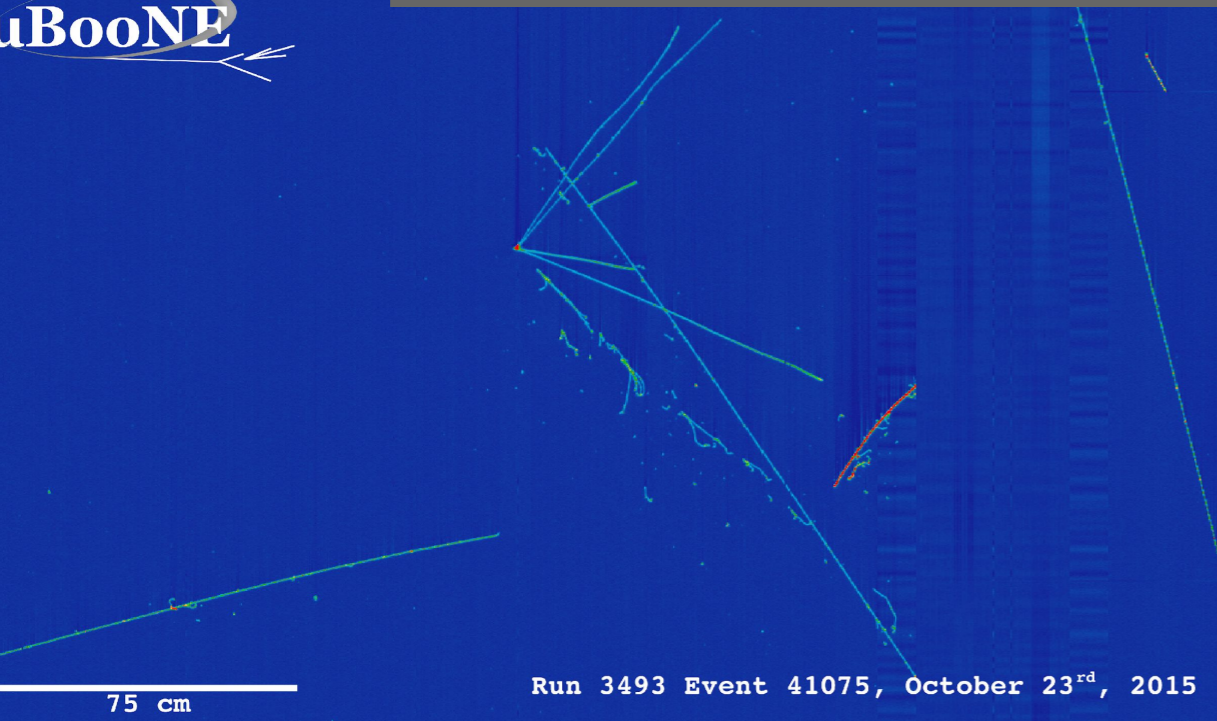


A 603MeV muon in Super-K.

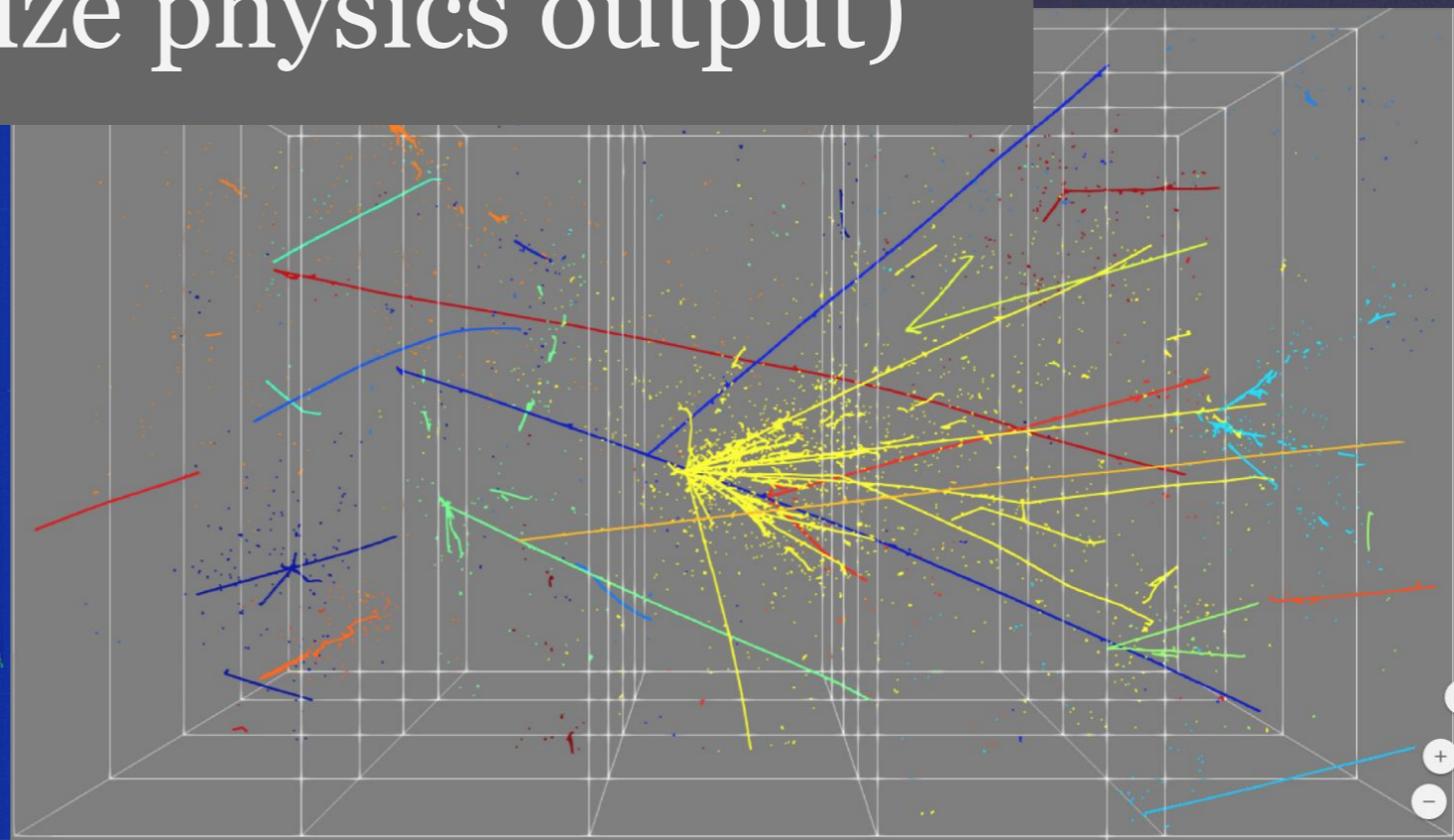
Need for advanced algorithms
for analyzing **high resolution**
data with **complex topologies**.
(**goal**: maximize physics output)

NOvA - FNAL E929
Run: 18975 / 43
Event: 628855 / SNEWSBeatSlow
UTC Mon Feb 23, 2015
14:30:1.383526016 Several hundred

μBooNE



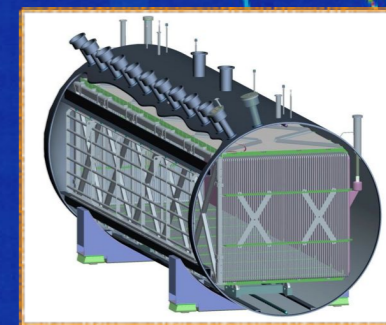
Run 3493 Event 41075, October 23rd, 2015



Hi-Res Particle Imaging

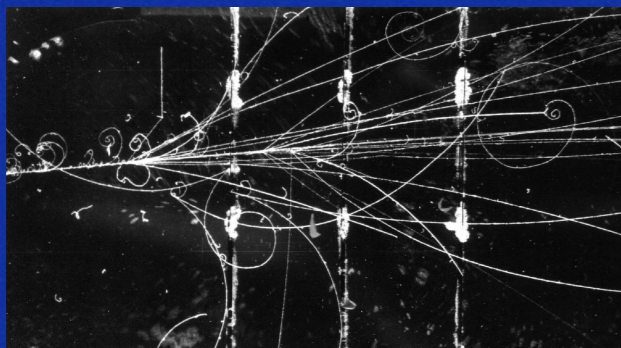
μ BooNE

**~mm/pixel spatial resolution
~MeV level sensitivity**



MicroBooNE
~87 ton (school bus size)

ν_μ



Bubble Chamber

Liquid Argon Time Projection Chamber

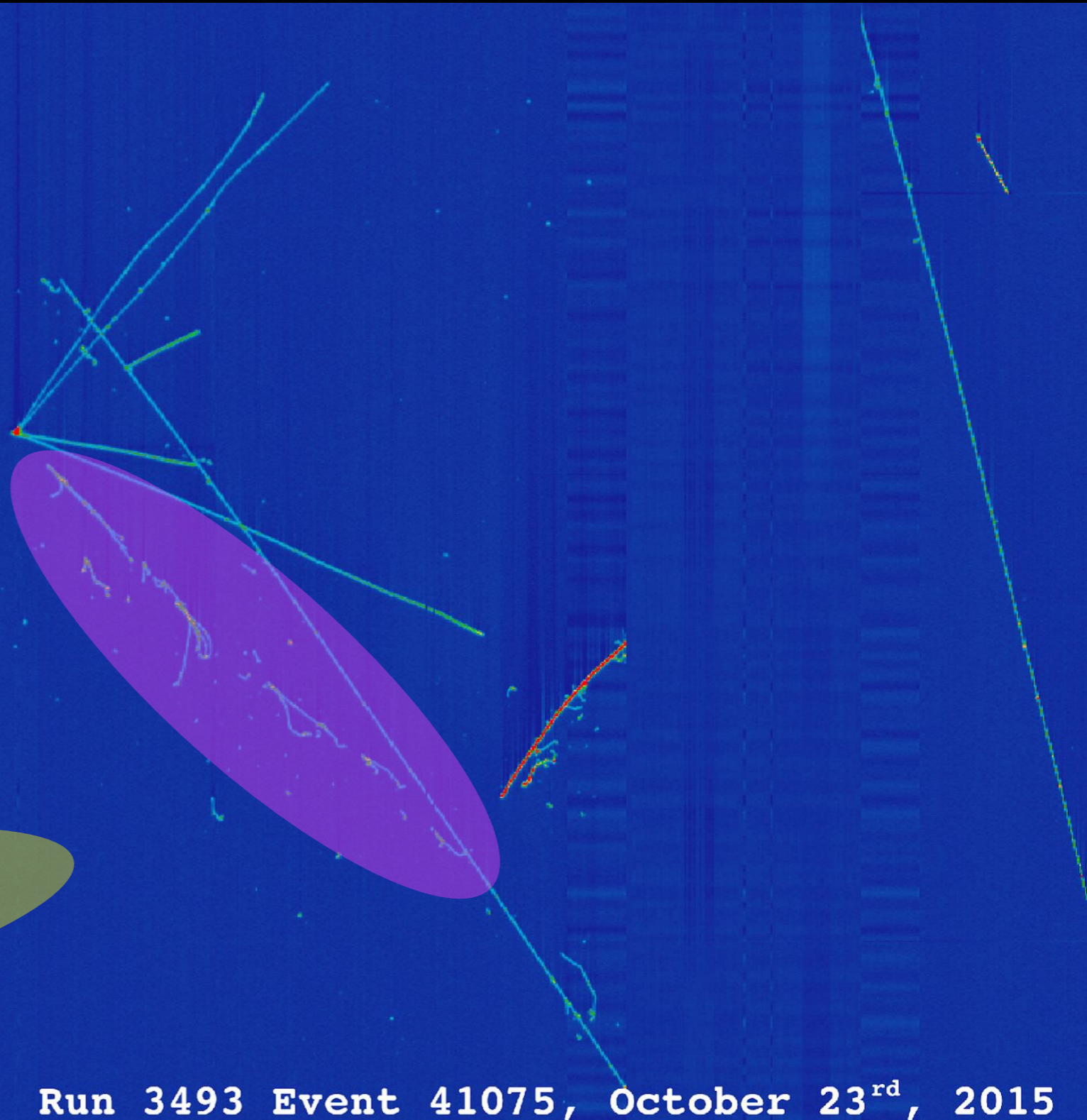
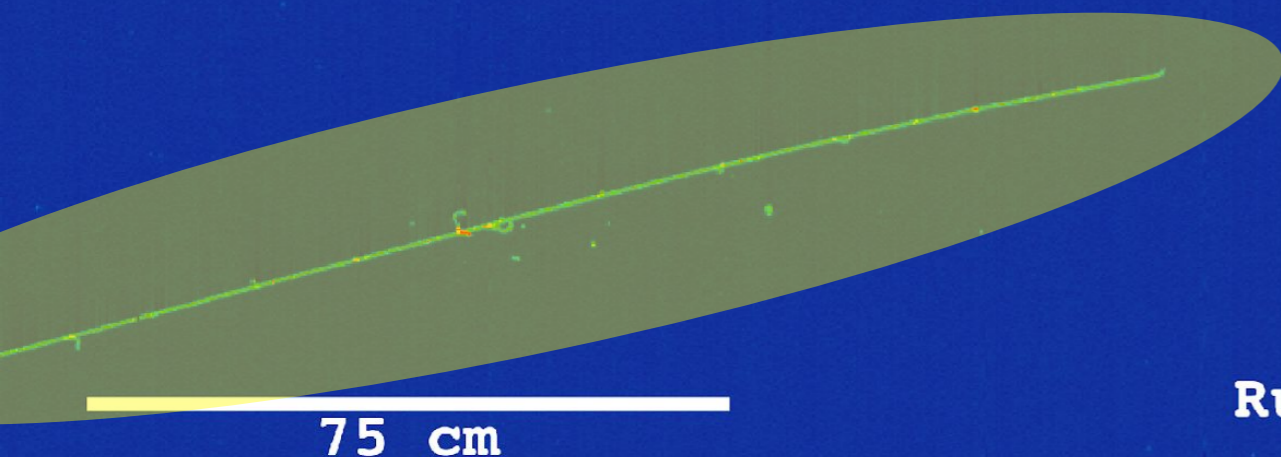
- High resolution photograph of charged particle trajectories
- Calorimetric measurement + scalability to a large mass

2015

Hi-Res Particle Imaging

μ BooNE

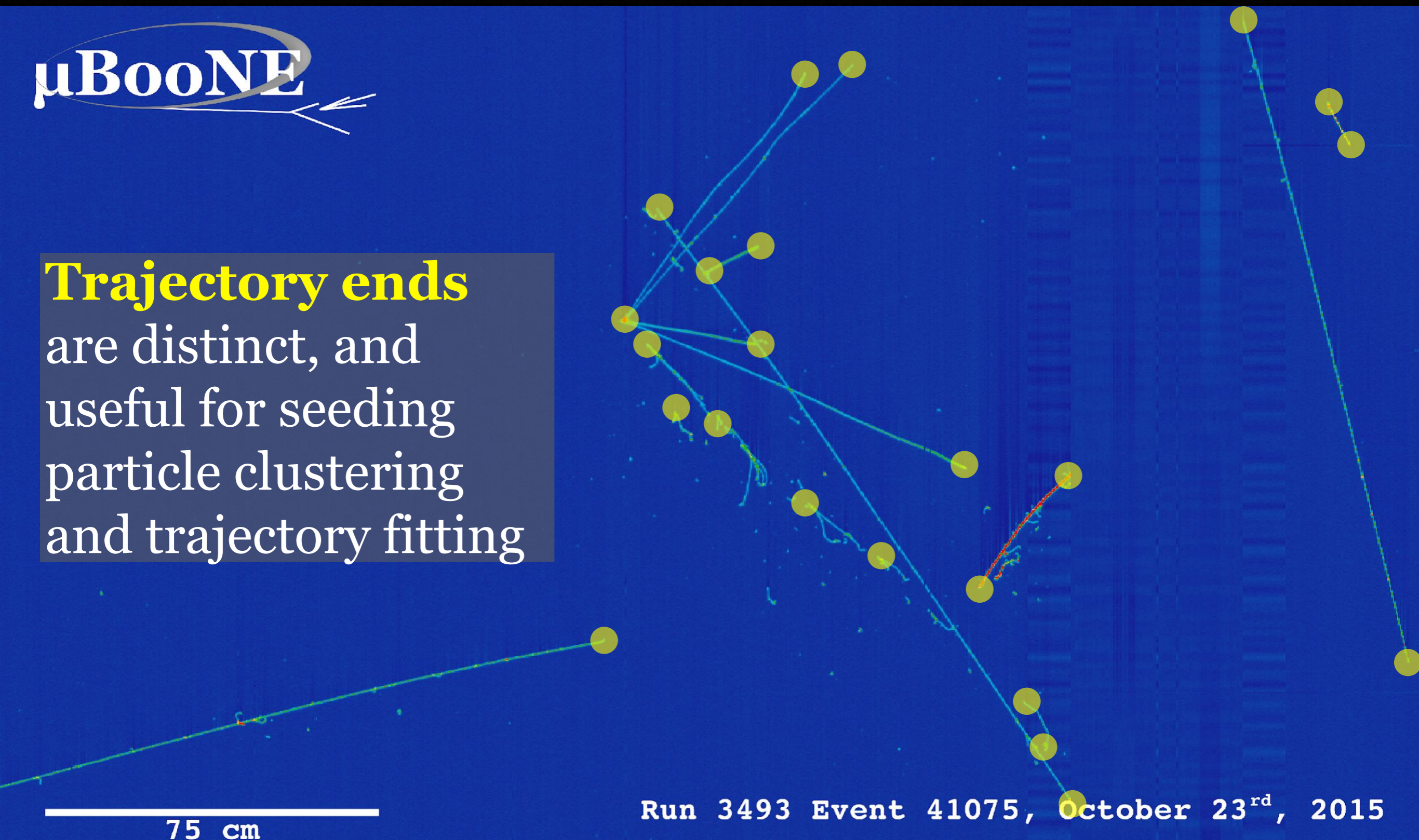
Topological shape
difference is a major
distinction for
“shower” particles



Hi-Res Particle Imaging



Trajectory ends
are distinct, and
useful for seeding
particle clustering
and trajectory fitting



Hi-Res Particle Imaging



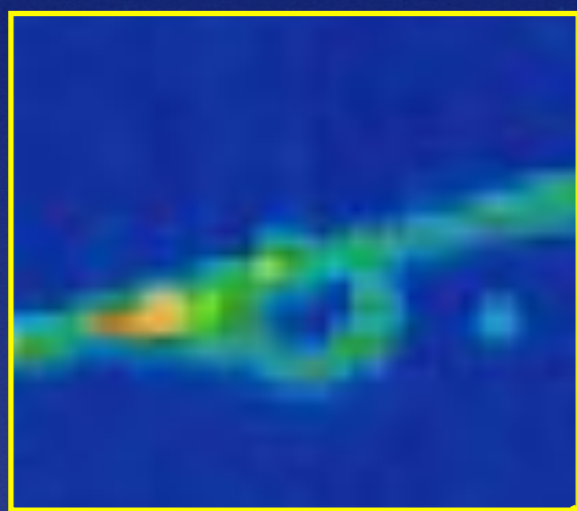
Many, local kinks
caused by Multiple
Coulomb Scattering
process can be used for
momentum estimation

75 cm

Run 3493 Event 41075, October 23rd, 2015

Hi-Res Particle Imaging

μ BooNE



75 cm

Small branches on muon-like trajectories are knocked-off electrons, useful key for the direction

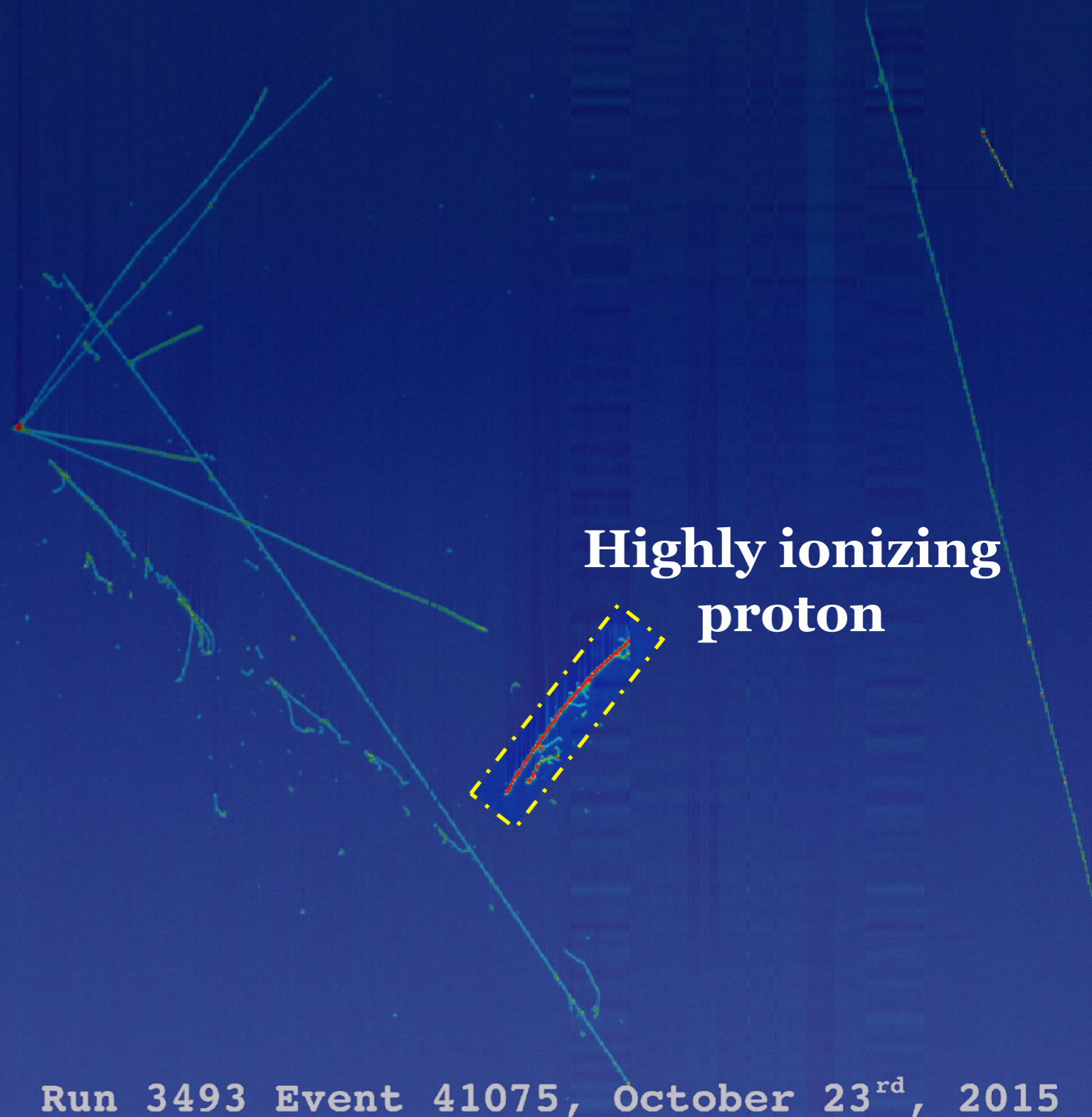
Run 3493 Event 41075, October 23rd, 2015

Hi-Res Particle Imaging



Energy deposition patterns (dE/dX)

vary with particle mass & momentum, useful for analysis

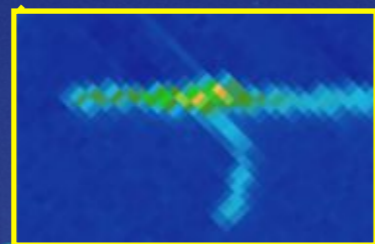


Hi-Res Particle Imaging

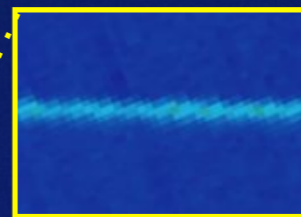


Energy deposition patterns (dE/dX)

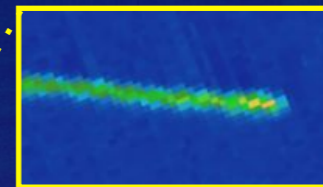
vary with particle mass & momentum, useful for analysis



e- vs. γ
using dE/dX



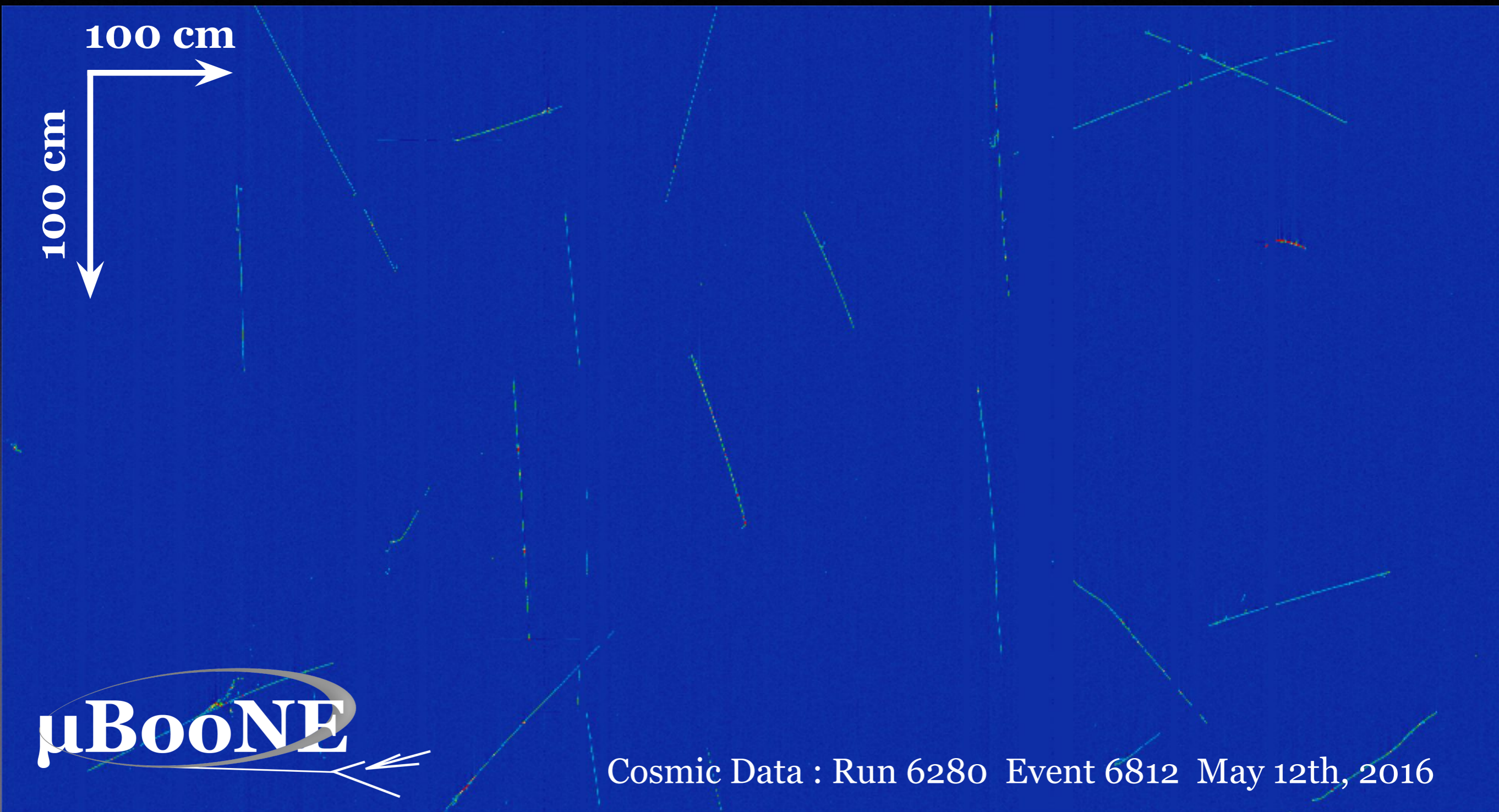
Stopping
particle



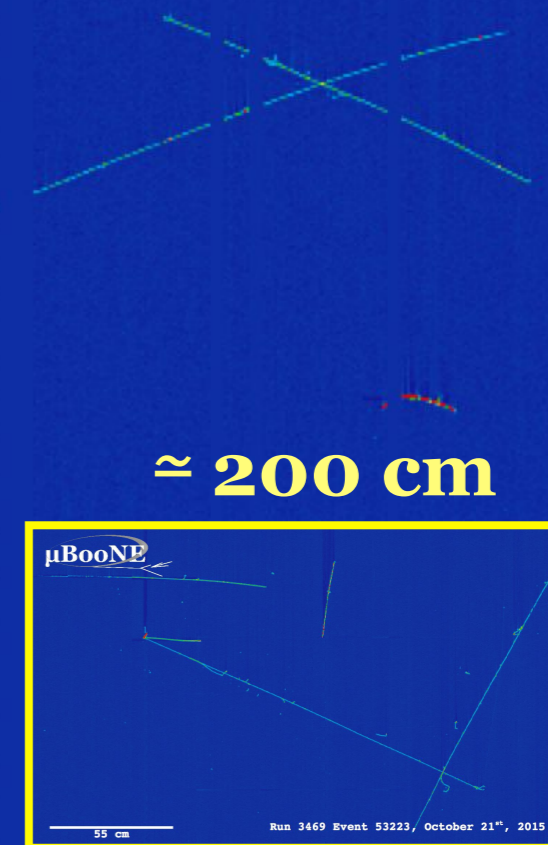
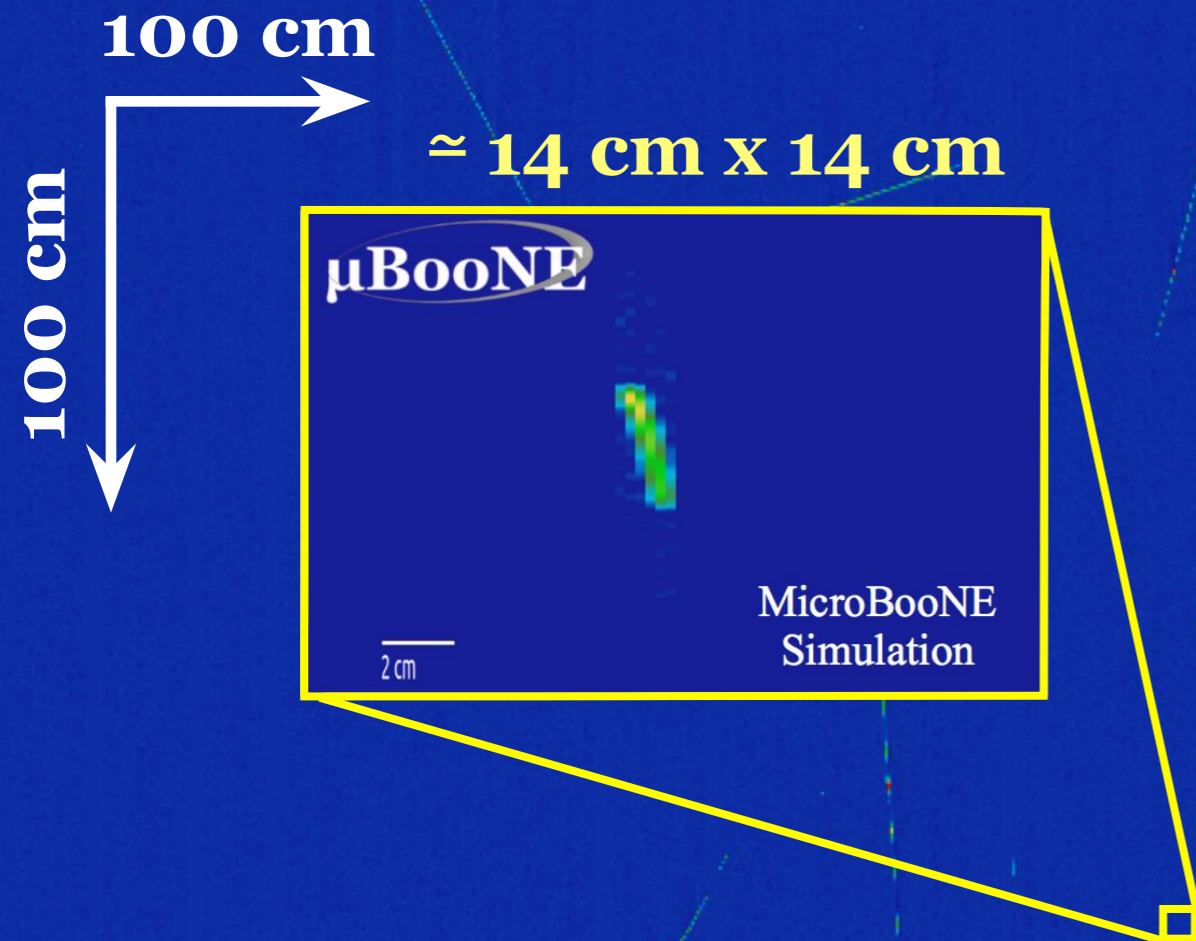
75 cm

Run 3493 Event 41075, October 23rd, 2015

Hi-Res Particle Imaging



Hi-Res Particle Imaging



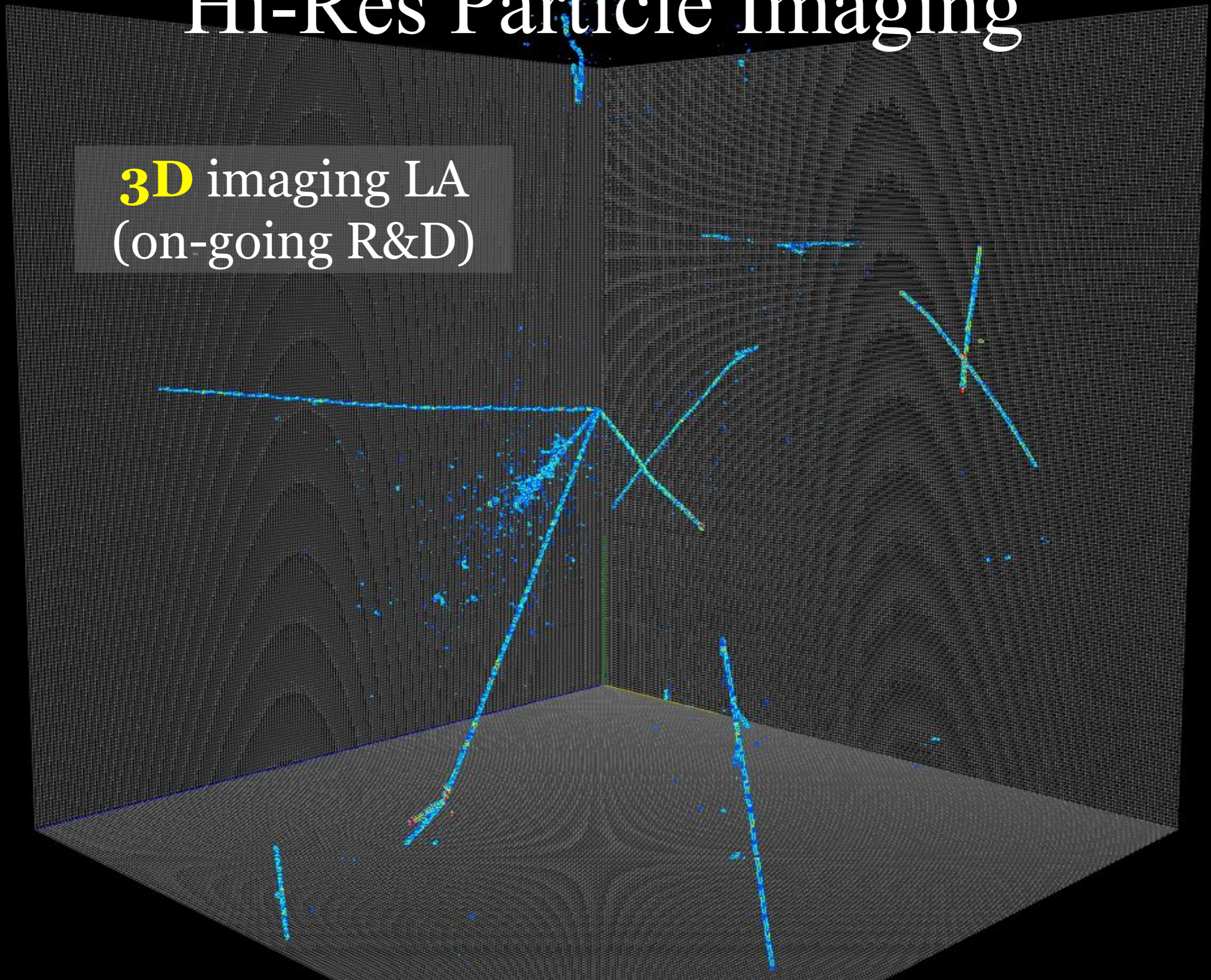
**Interaction vertex can be anywhere
in LAr, varying in size (cm ~ meters)**



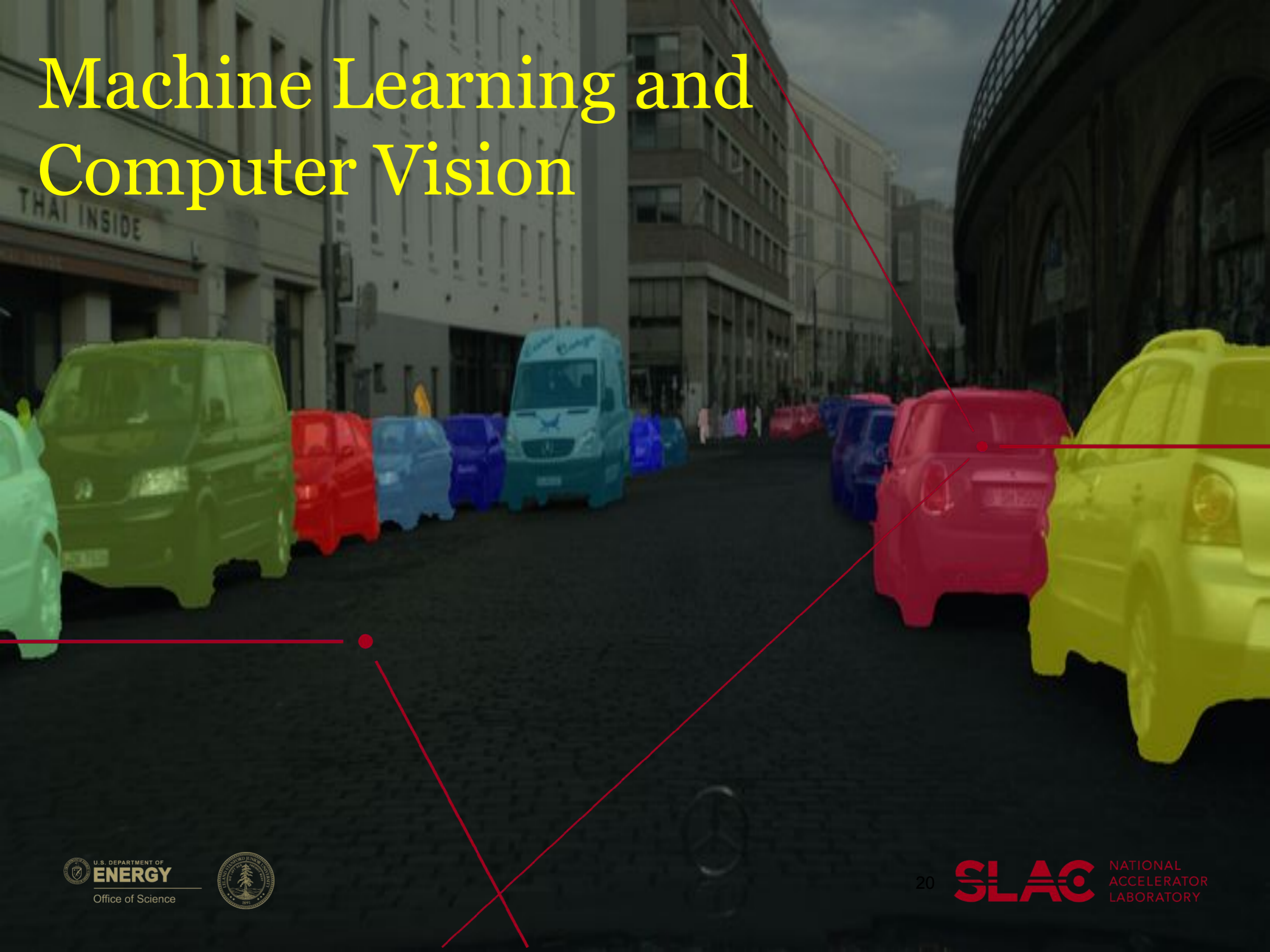
Cosmic Data : Run 6280 Event 6812 May 12th, 2016

Hi-Res Particle Imaging

3D imaging LA
(on-going R&D)

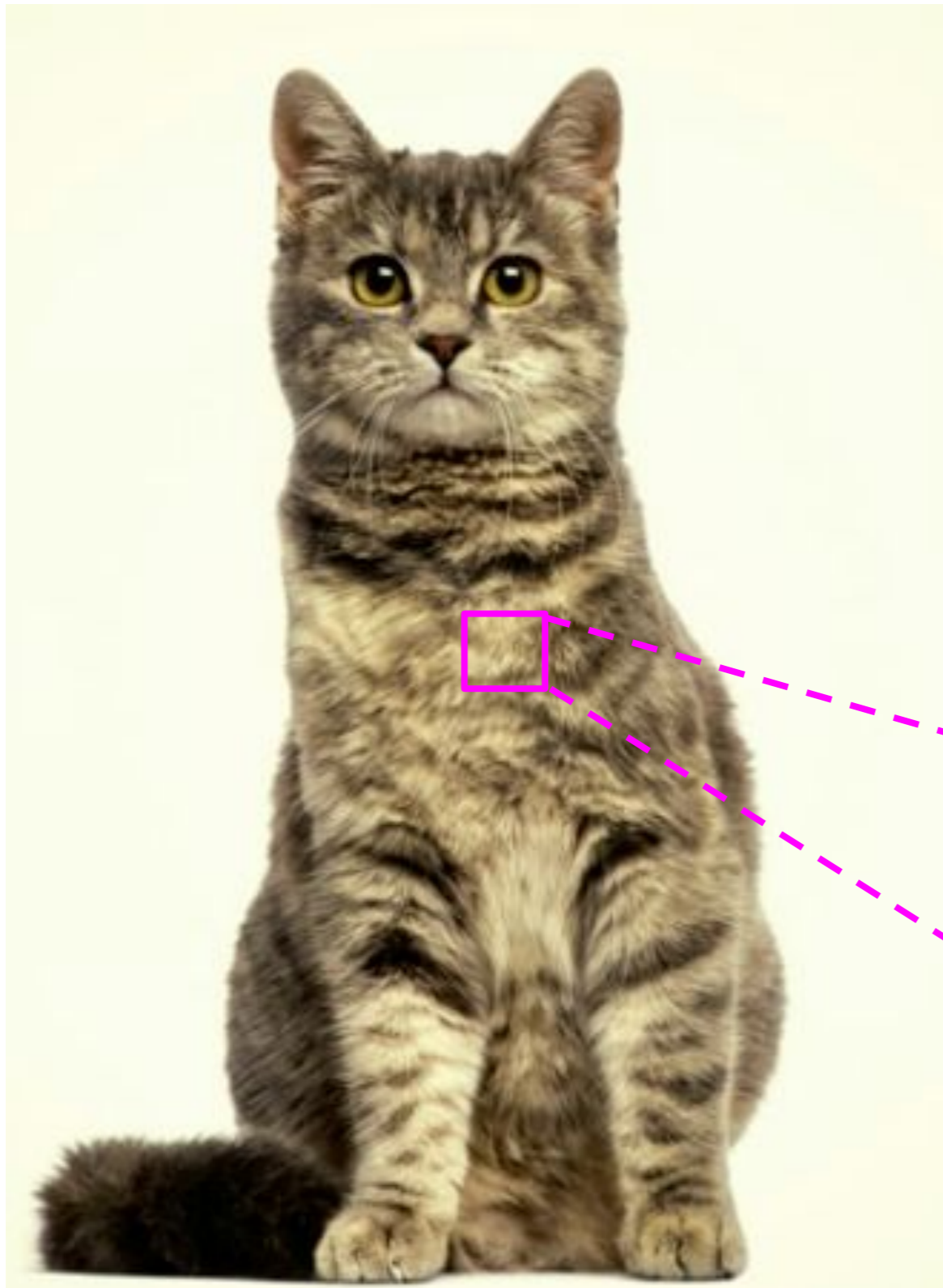


Machine Learning and Computer Vision



Find A Cat

Machine Learning in Neutrino Physics



How to write an algorithm
to identify a cat?

... very hard task ...

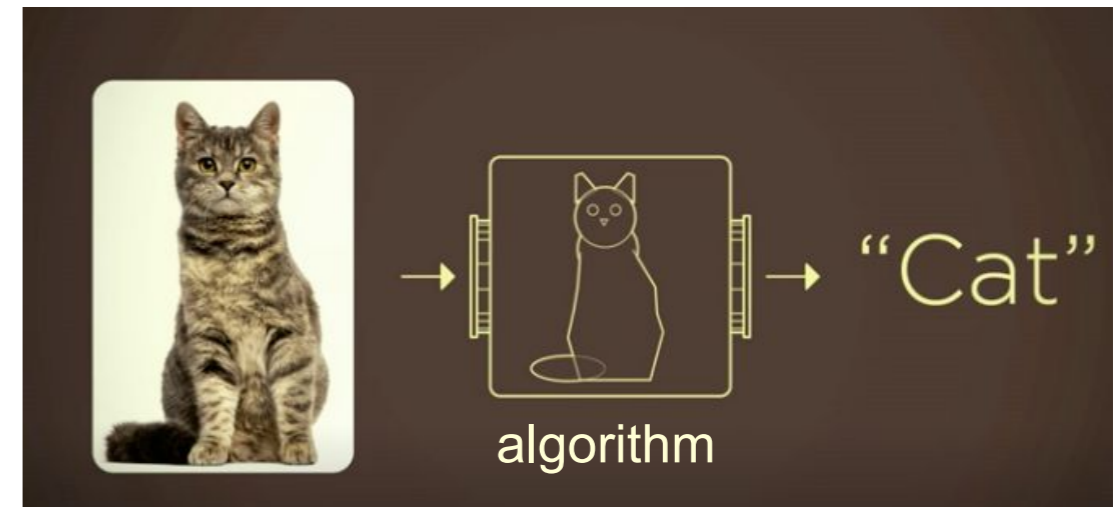
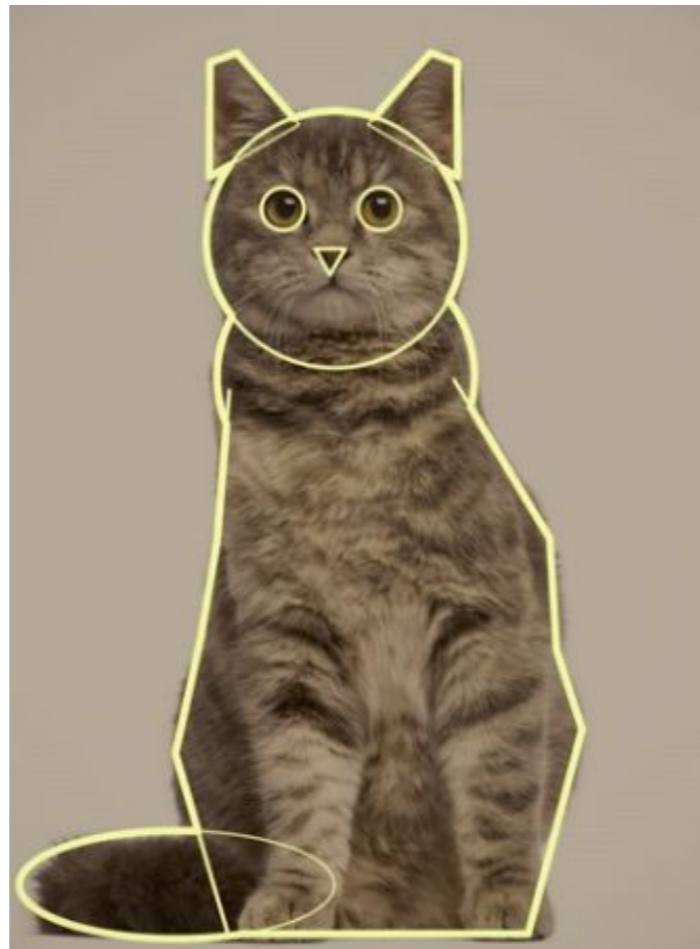
| | | | | | |
|----|----|----|----|----|----|
| 16 | 08 | 67 | 15 | 83 | 09 |
| 37 | 52 | 77 | 23 | 22 | 74 |
| 35 | 42 | 48 | 72 | 85 | 27 |
| 68 | 36 | 43 | 54 | 21 | 33 |
| 79 | 60 | 10 | 25 | 54 | 71 |
| 18 | 55 | 38 | 73 | 50 | 47 |

Find A Cat

Machine Learning in Neutrino Physics

Development Workflow for non-ML reconstruction

1. Write an algorithm based on physics principles



A cat = collection of
(or, a neutrino) certain shapes

Find A Cat

Machine Learning in Neutrino Physics

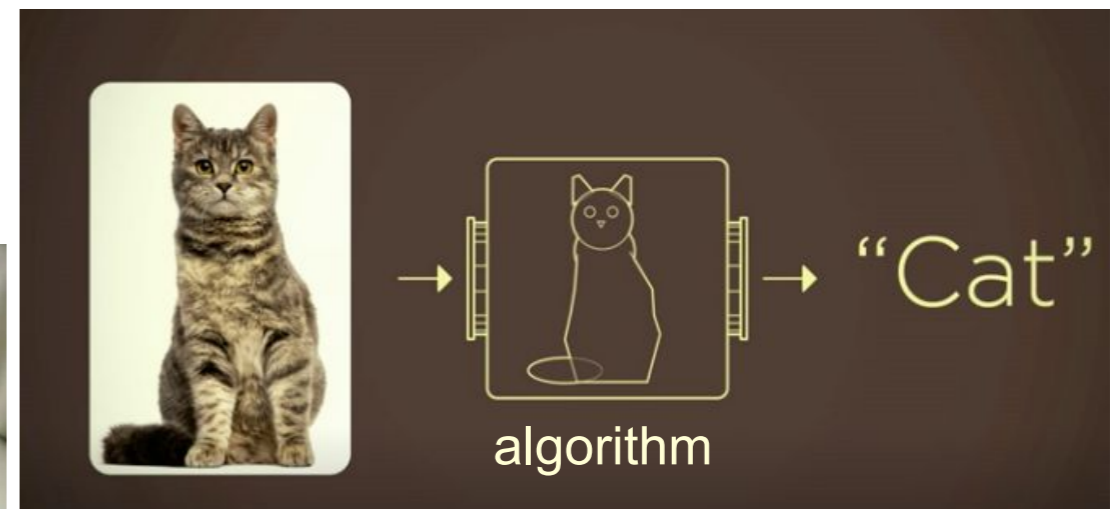
Development Workflow for non-ML reconstruction

1. Write an algorithm based on physics principles
2. Run on simulation and data samples
3. Observe failure cases, implement fixes/heuristics
4. Iterate over 2 & 3 till a satisfactory level is achieved
5. Chain multiple algorithms as one algorithm, repeat 2, 3, and 4.



Partial cat
(escaping the detector)

Stretching cat
(Nuclear FSI)



A cat = collection of
(or, a neutrino) certain shapes

Development Workflow for non-ML reconstruction

1. Write an algorithm based on physics principles
2. Run on simulation and data samples
3. Observe failure cases, implement fixes/heuristics
4. Iterate over 2 & 3 till a satisfactory level is achieved
5. Chain multiple algorithms as one algorithm, repeat 2, 3, and 4.

Machine Learning

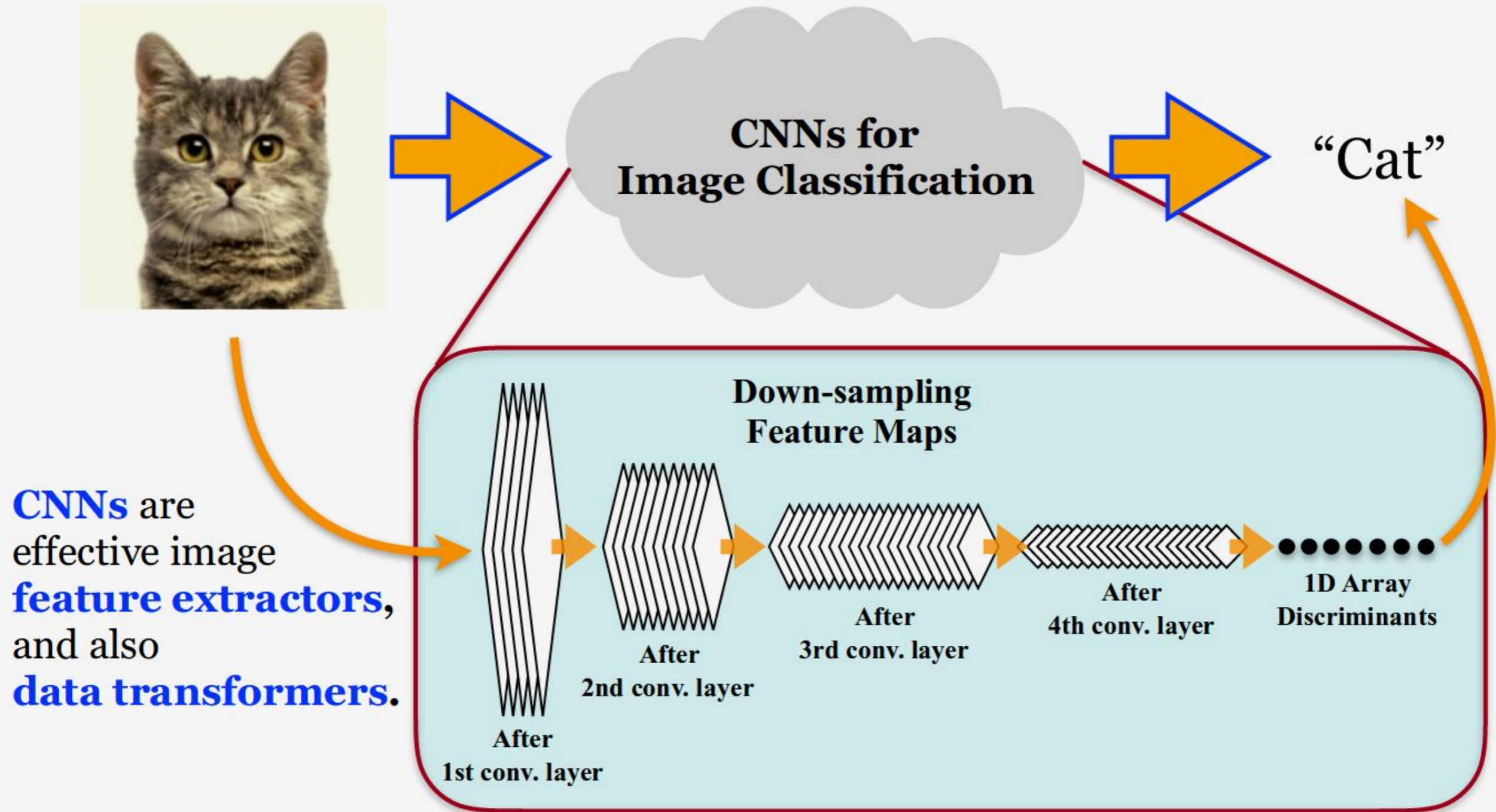
- **“Learn patterns from data”**
 - automation of steps 2, 3, and 4
- **“Chain algorithms & optimize”**
 - step 5 addressed by design



CNN for “Image Classification”

Machine Learning in Neutrino Physics

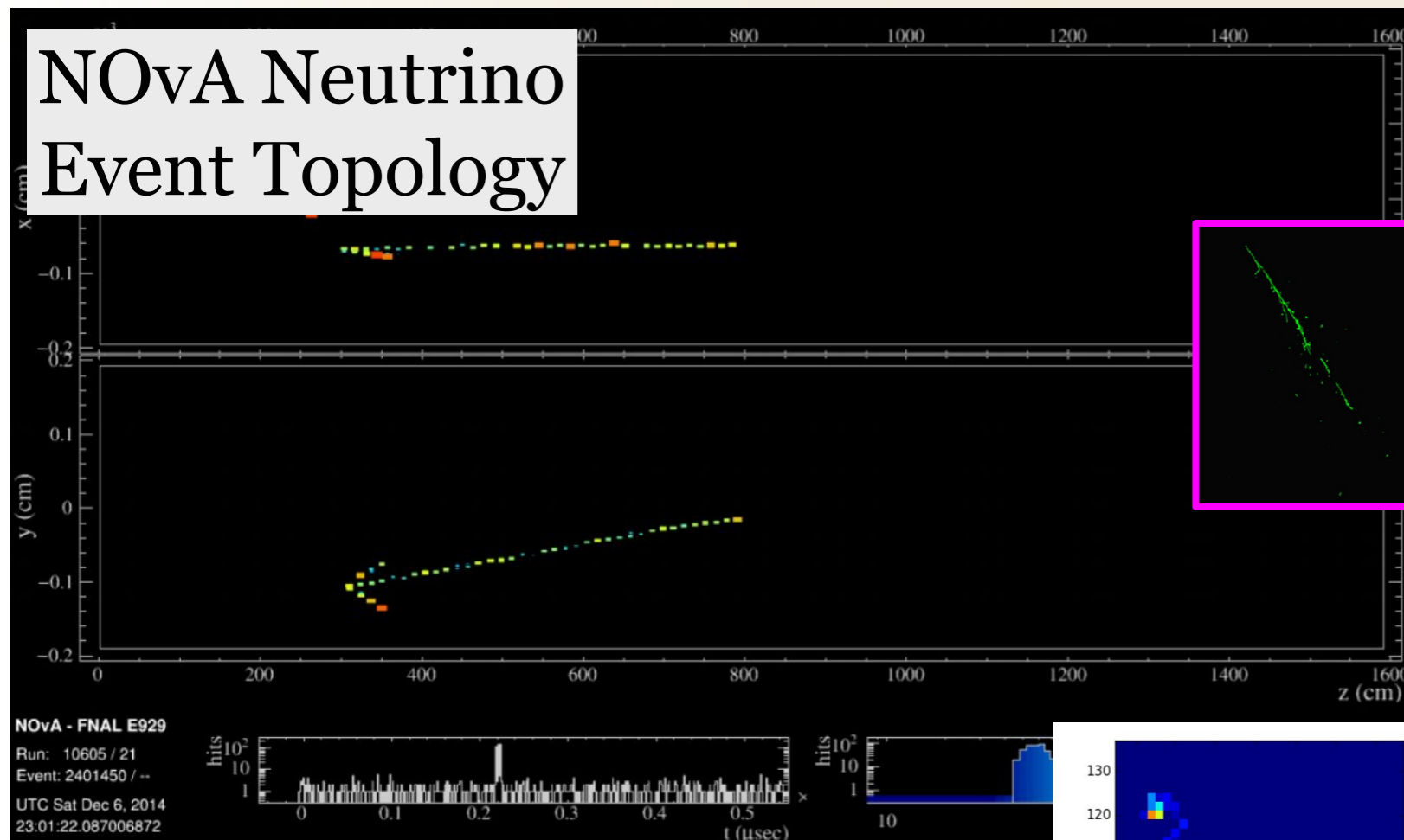
Convolutional Neural Networks (CNNs)



CNN for “Image Classification”

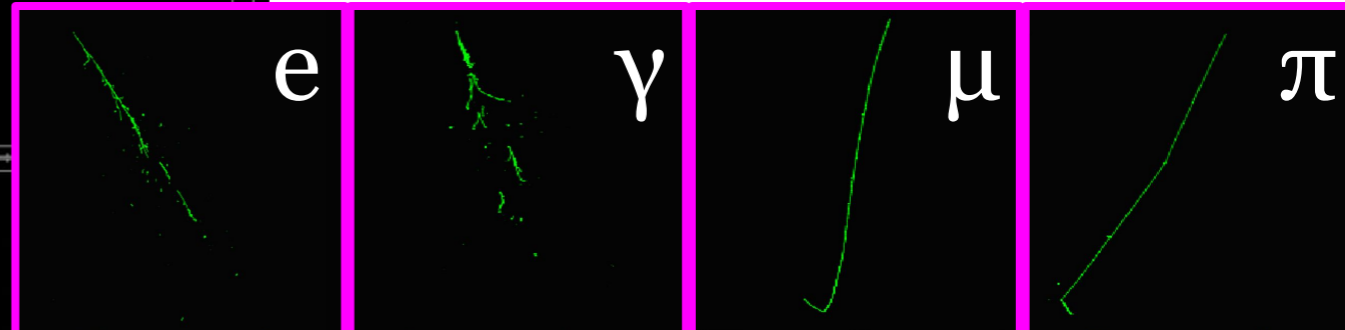
Machine Learning in Neutrino Physics

SLAC



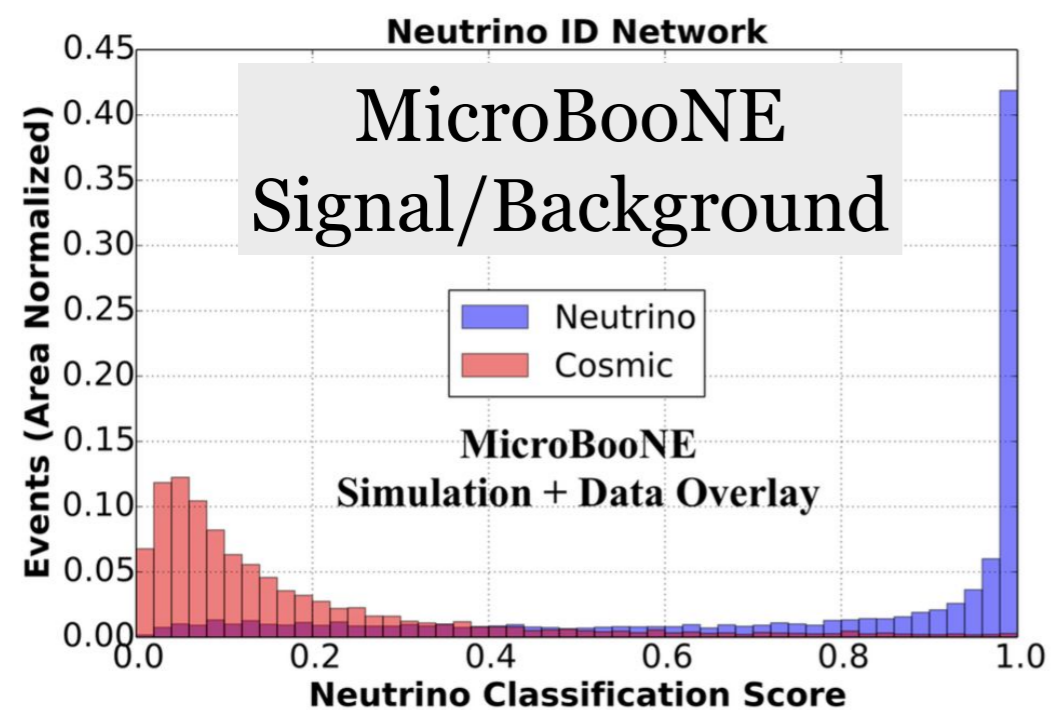
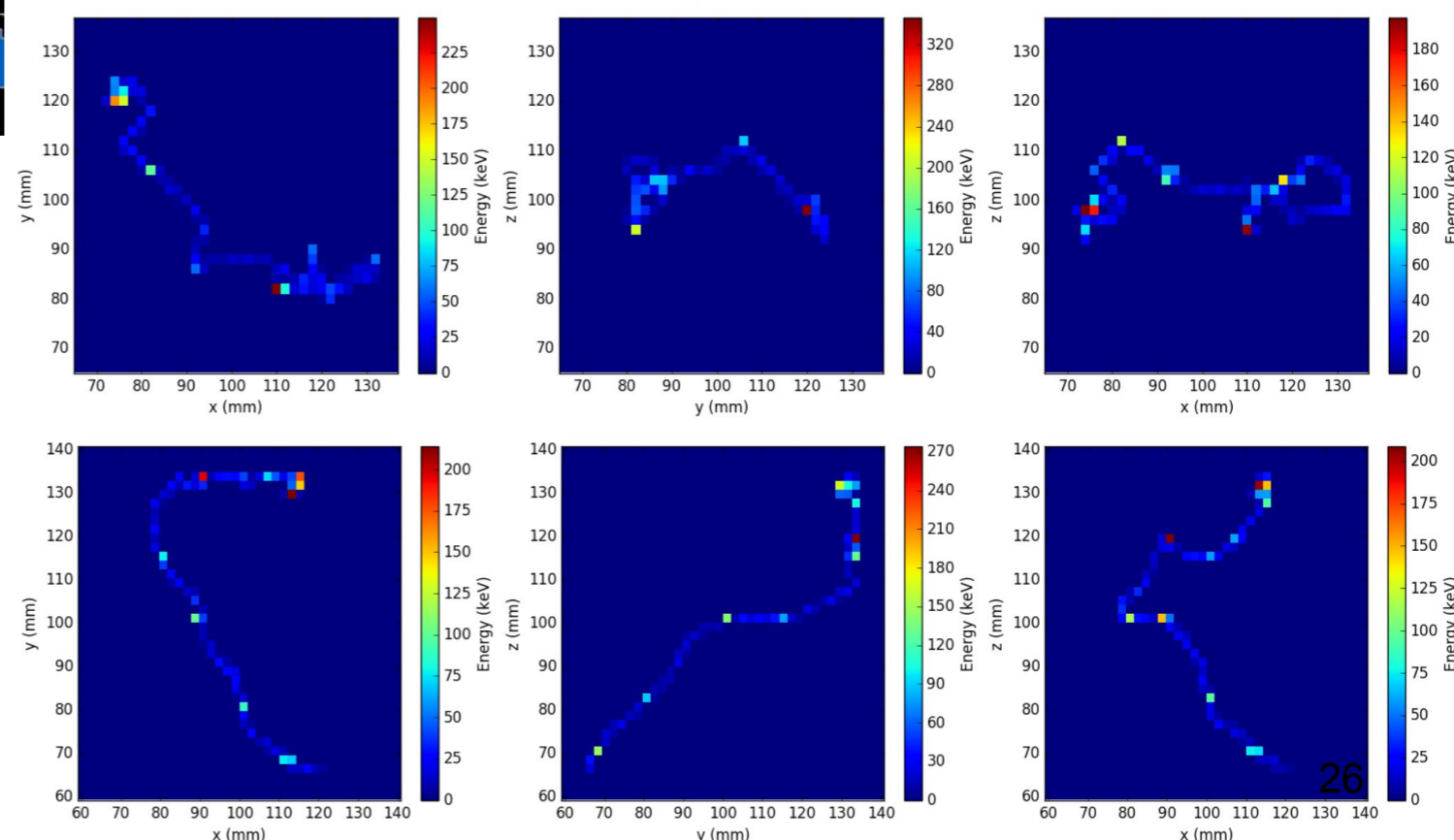
LArLIAT

Signal vs. Background



NEXT

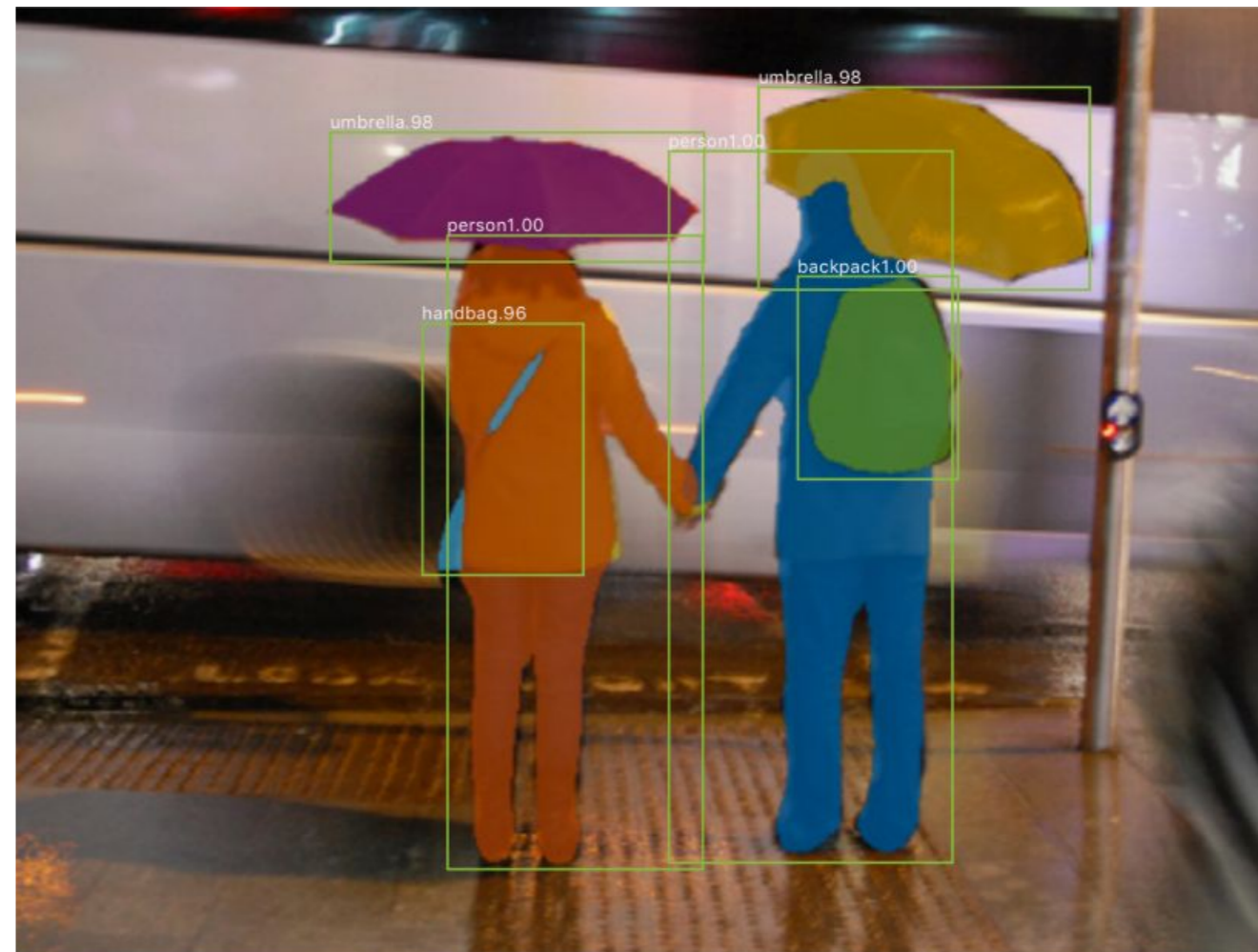
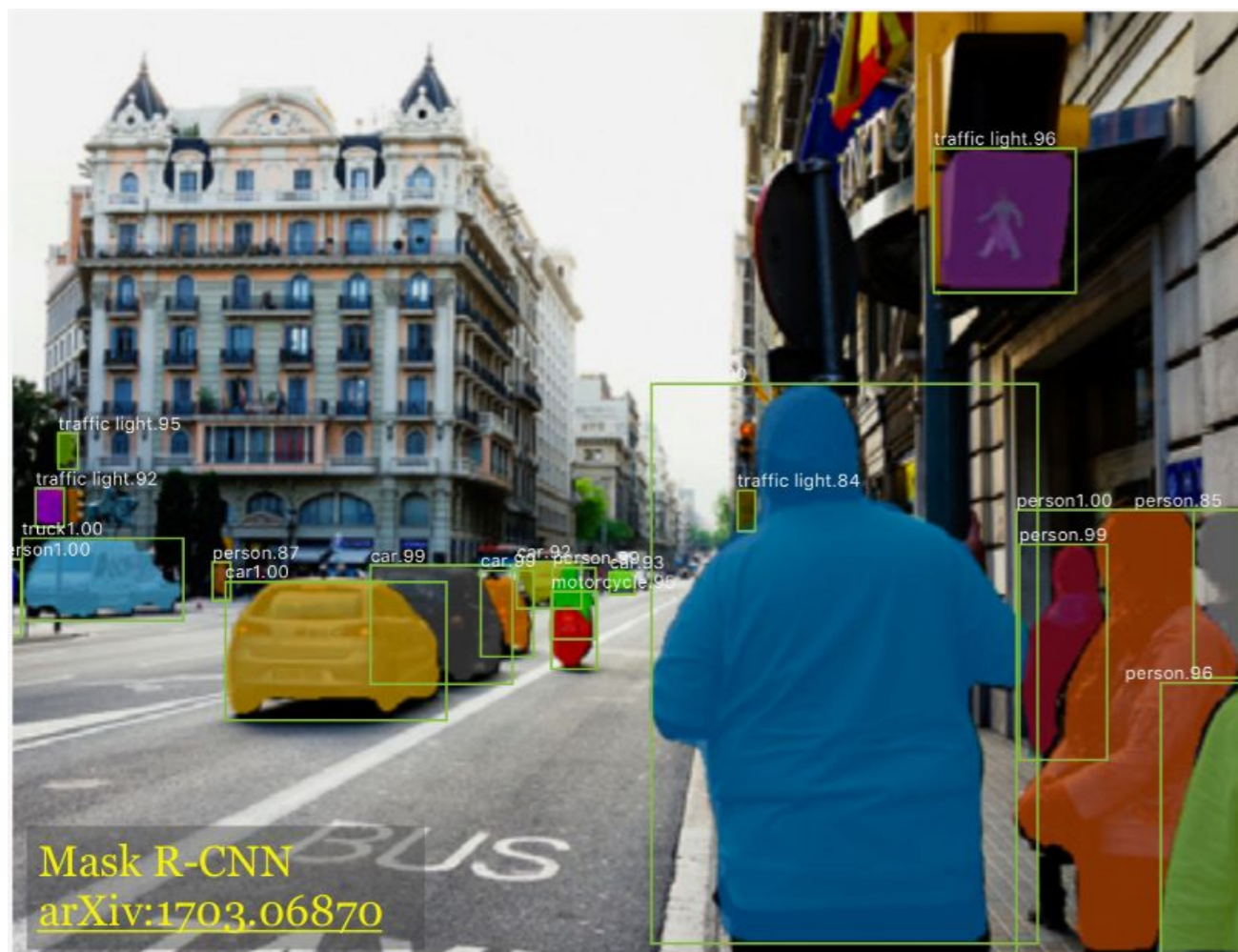
Signal vs. Background



WHAT is WHERE in an image?

Machine Learning in Neutrino Physics

Image Context Detection



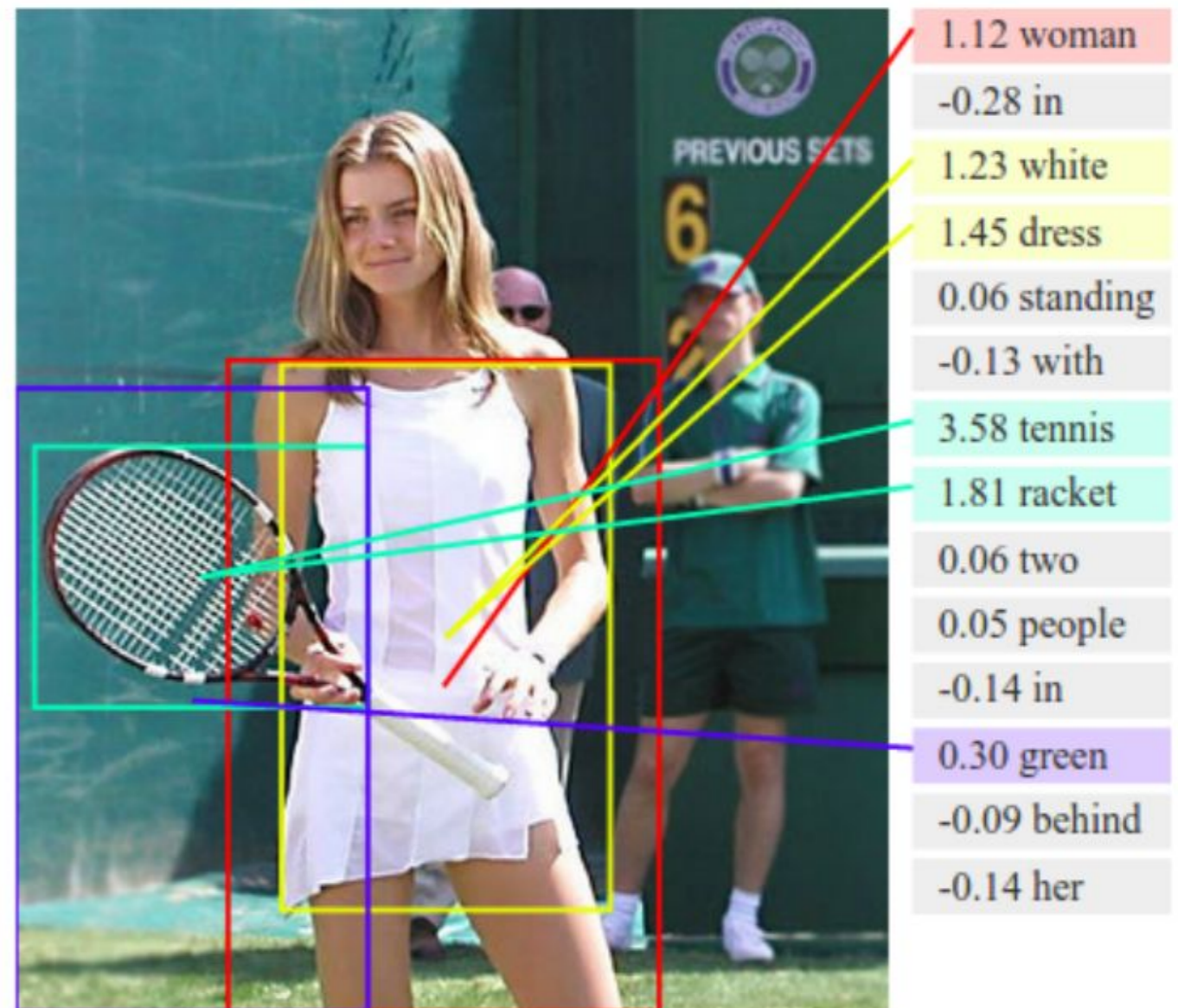
WHAT is WHERE and HOW in an image?

Machine Learning in Neutrino Physics

Interpretation of Contexts' Correlation



"girl in pink dress is jumping in air."

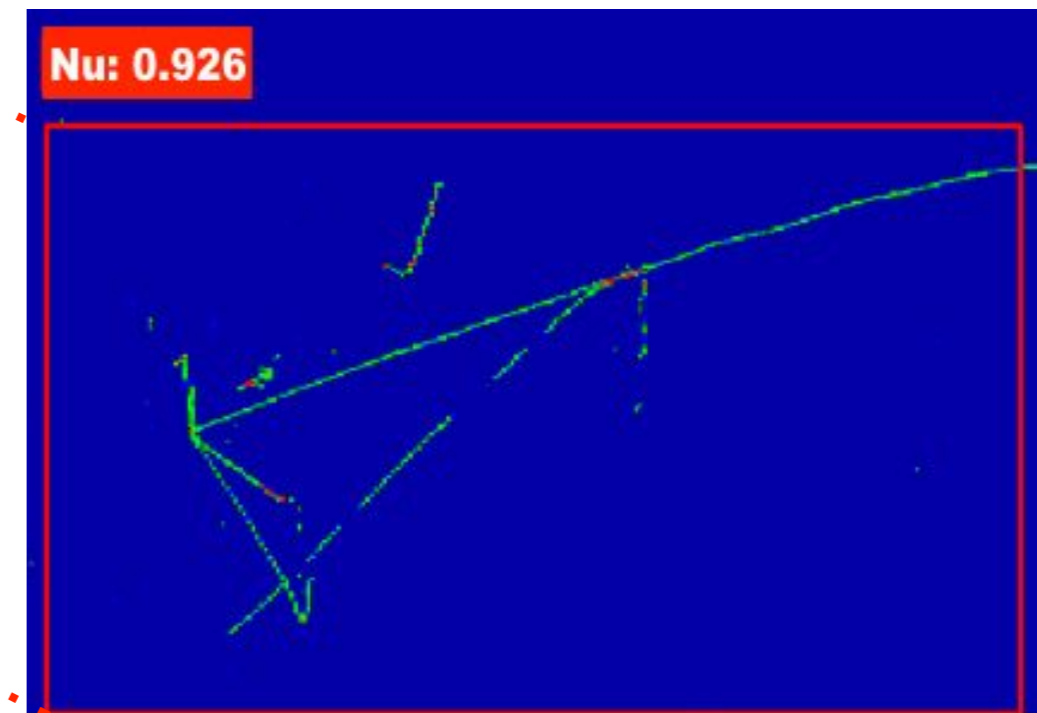
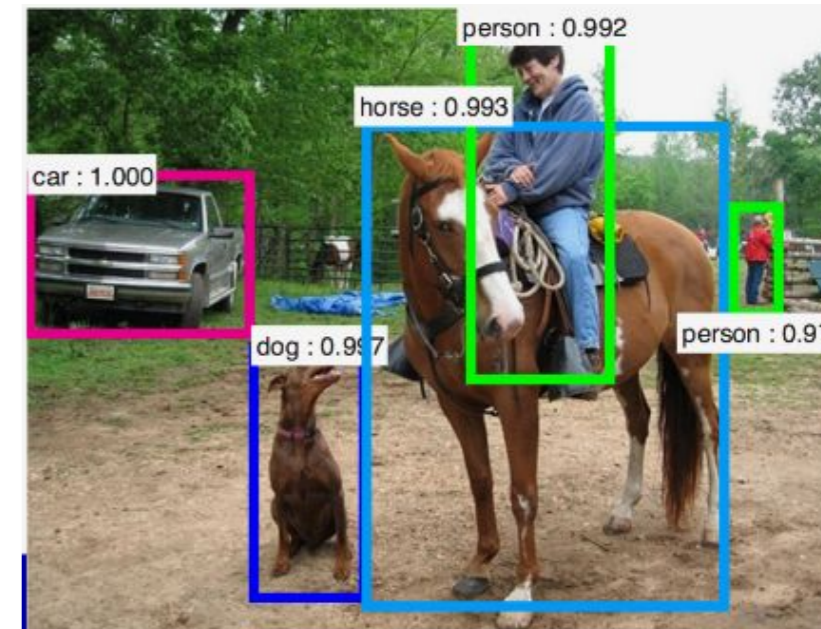
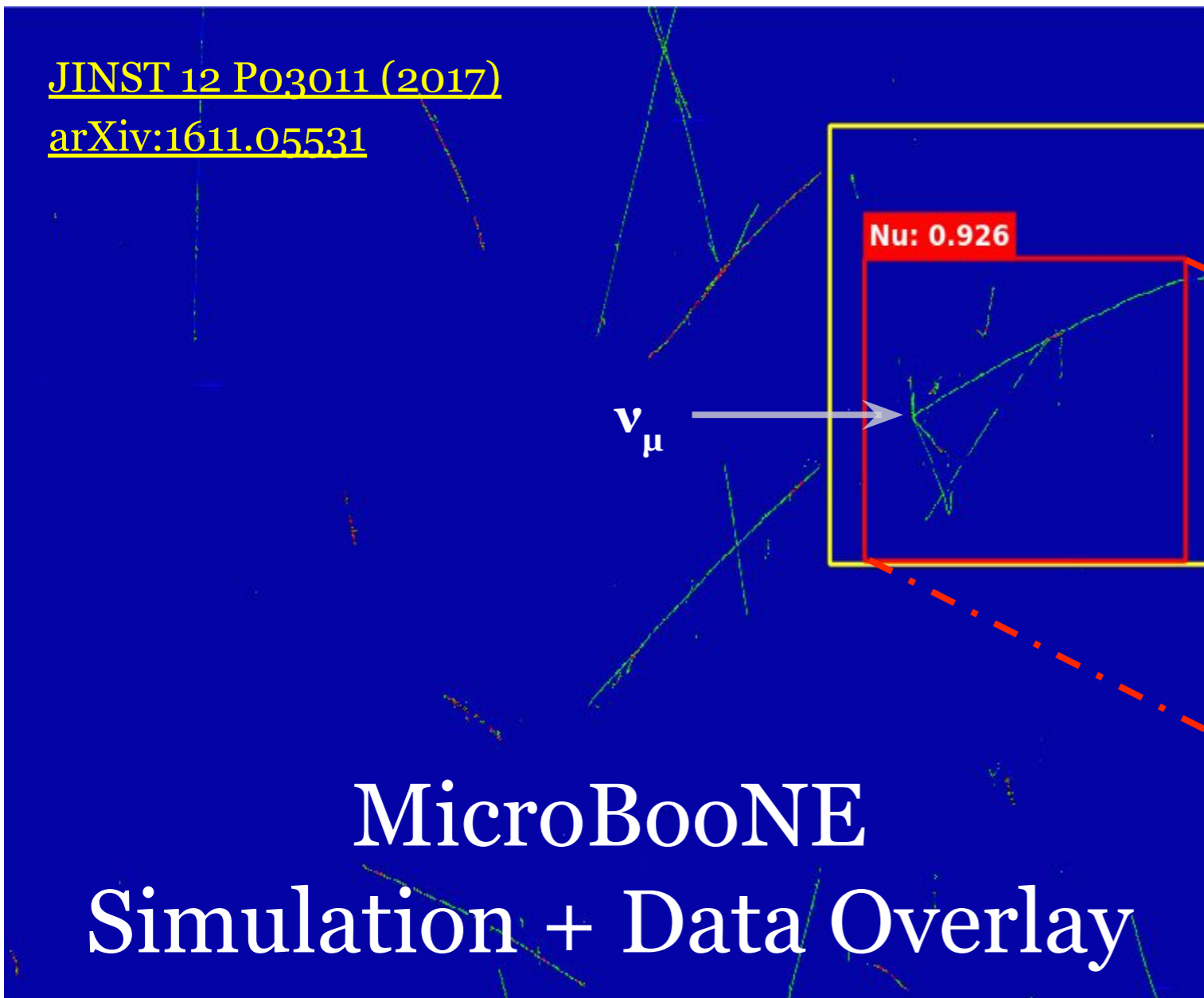


Beyond Image Classification

Machine Learning in Neutrino Physics

Object Detection for Neutrino Finding (MicroBooNE LArTPC)

[JINST 12 P03011 \(2017\)](#)
[arXiv:1611.05531](#)



Task: propose a rectangular box (location & size) that contains neutrino interaction²⁹

Beyond Image Classification

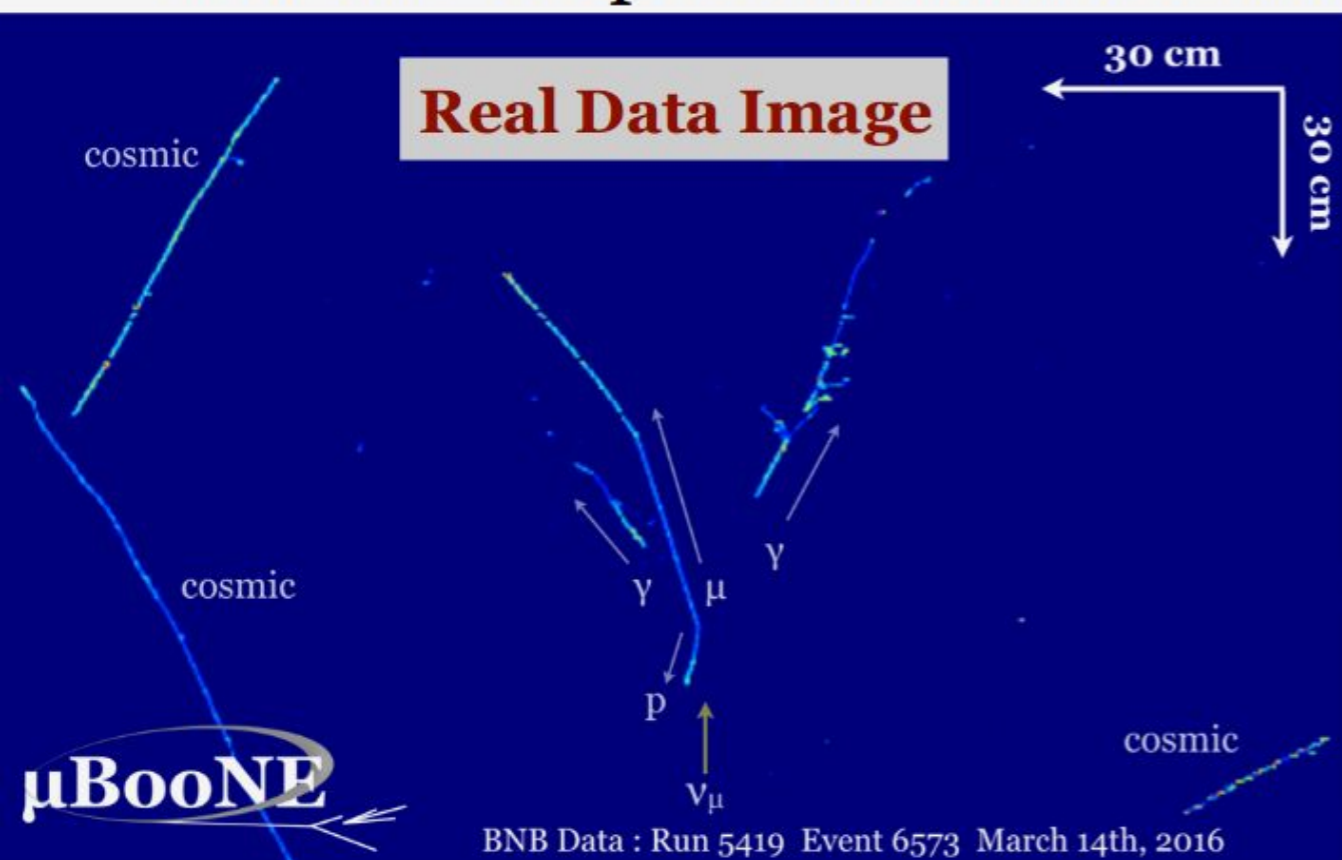
Machine Learning in Neutrino Physics

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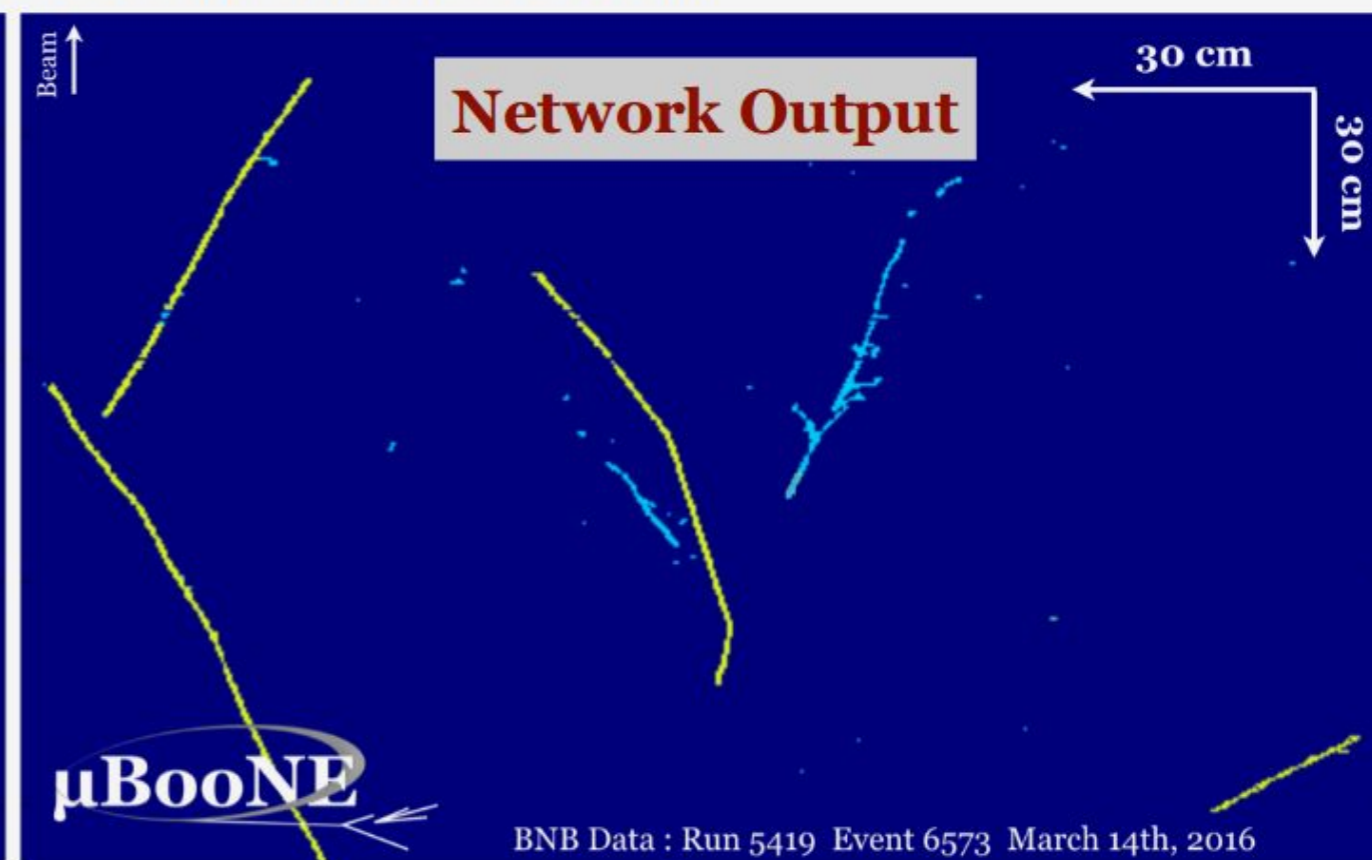
Semantic Segmentation

- Recently published ... [arXiv:1808.07269](https://arxiv.org/abs/1808.07269)
- Pixel-level object classification
 - Separation of EM-particle from other types
 - Key input information for particle clustering
- First time deep neural network validated on LArTPC data

ML Technique @ MicroBooNE LArTPC Detector



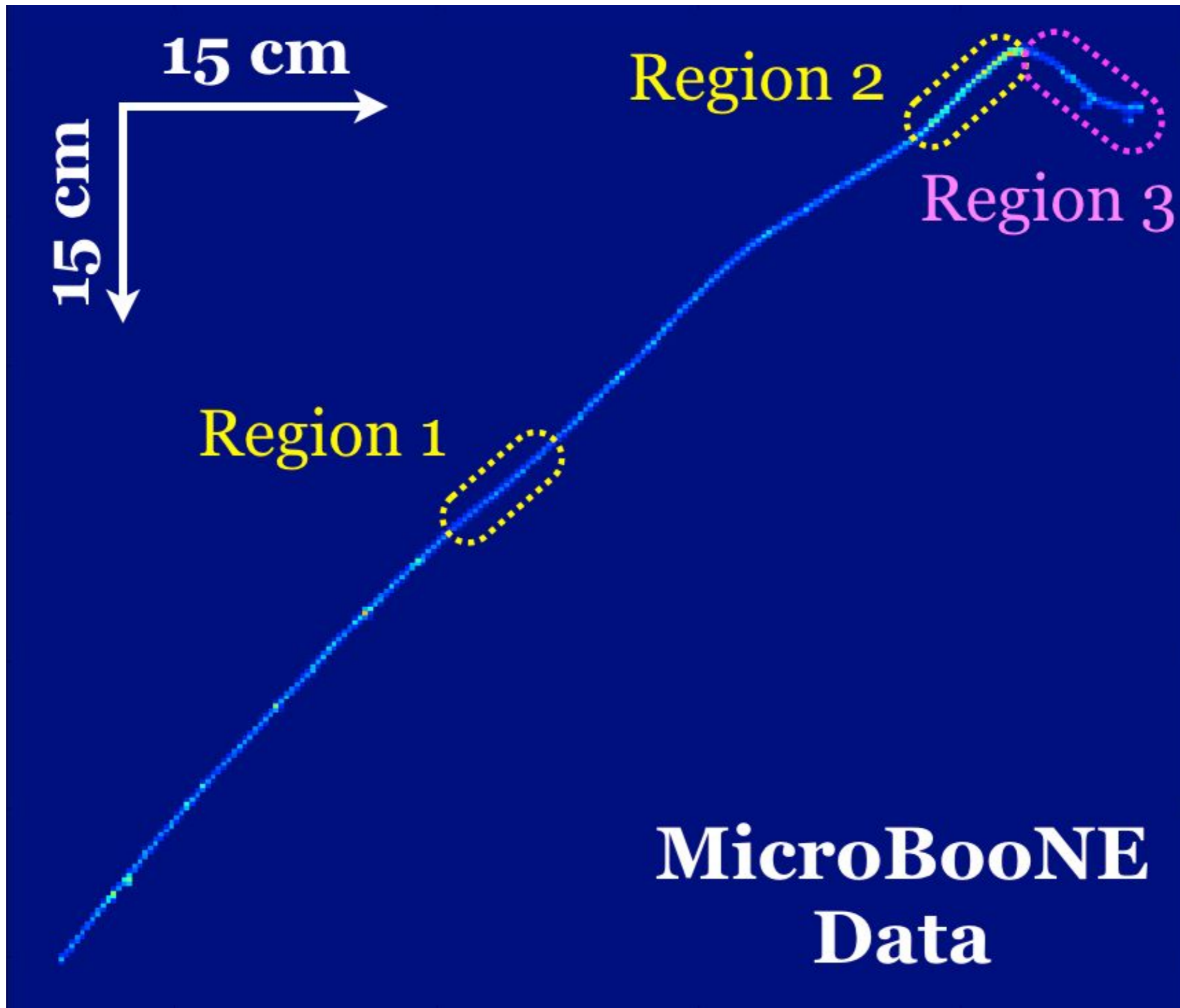
Network Input



Network Output

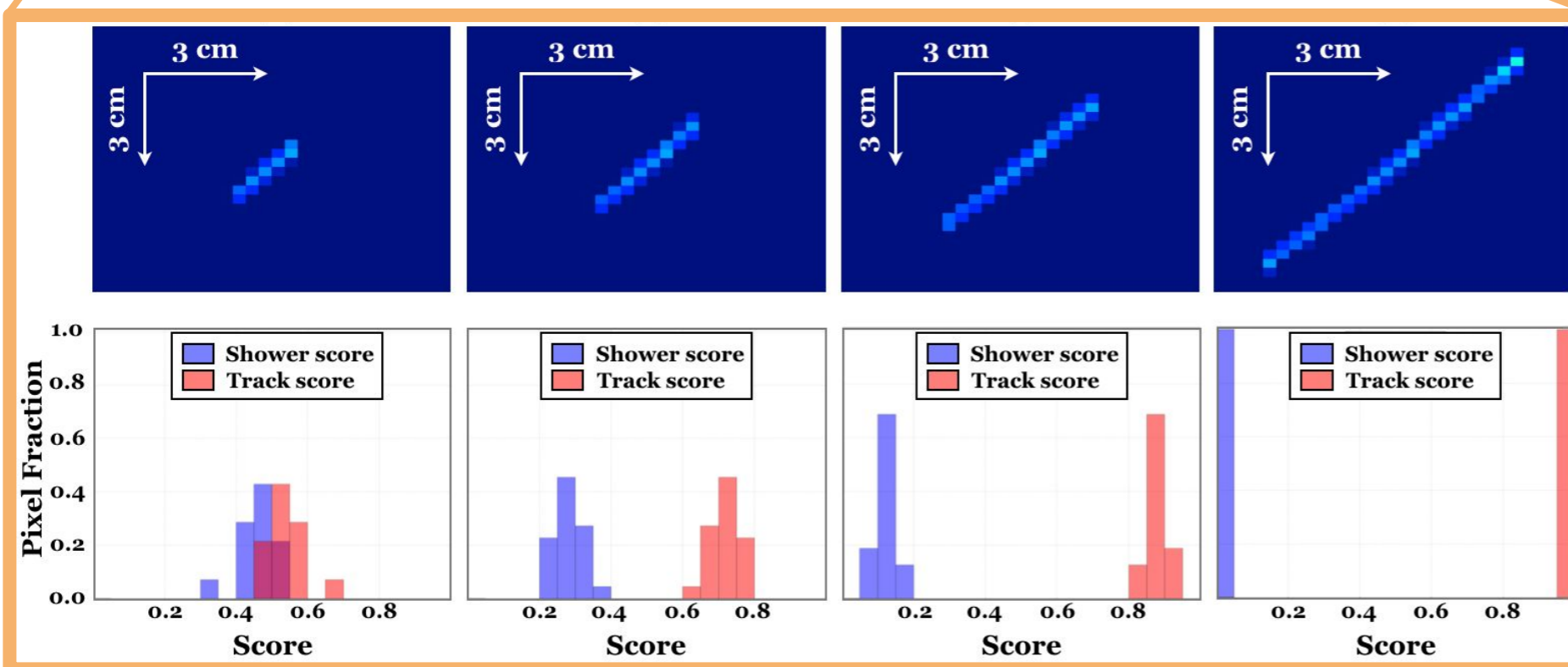
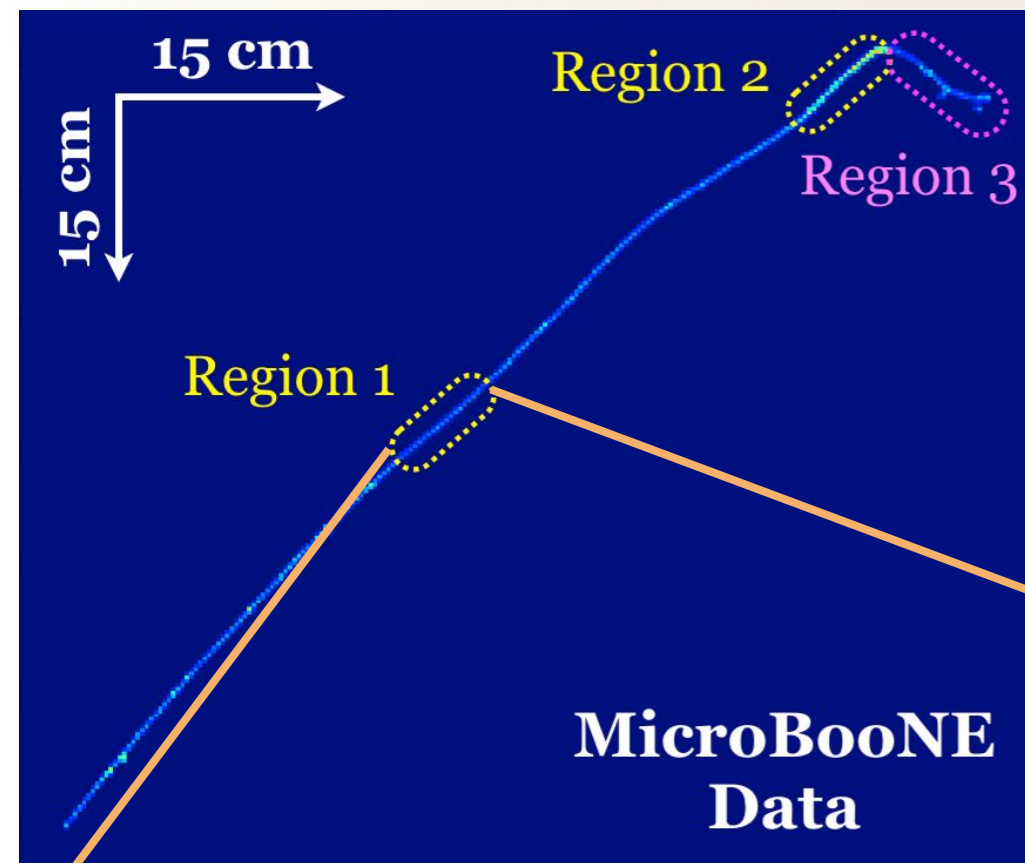
Pixel-level Feature Information

Machine Learning in Neutrino Physics



Pixel-level Feature Information

Machine Learning in Neutrino Physics

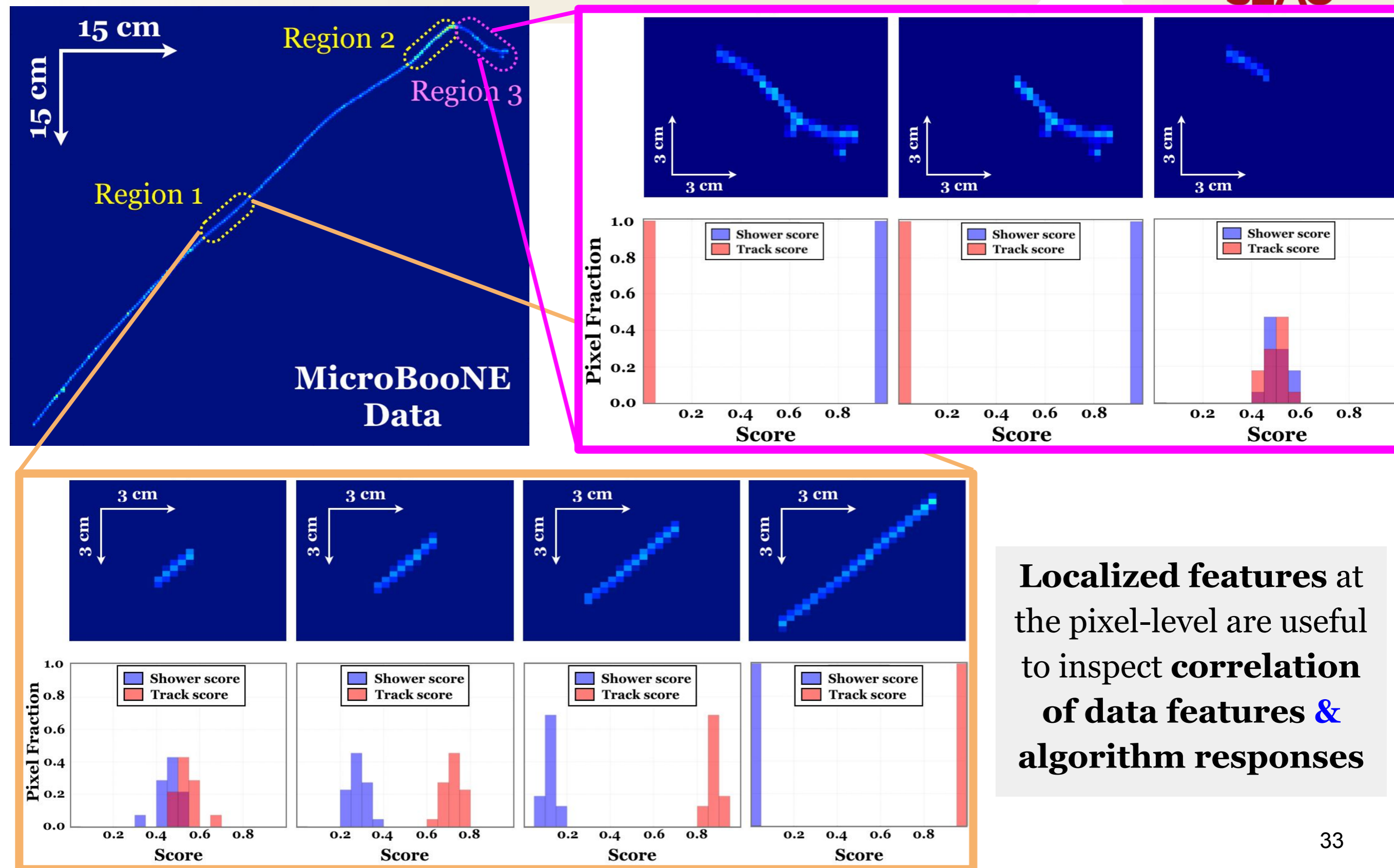


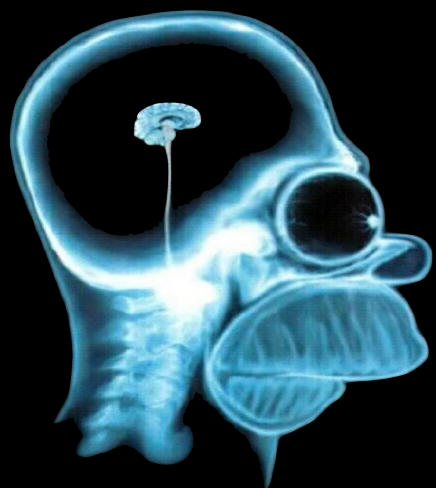
Localized features at the pixel-level are useful to inspect **correlation of data features & algorithm responses**

Pixel-level Feature Information

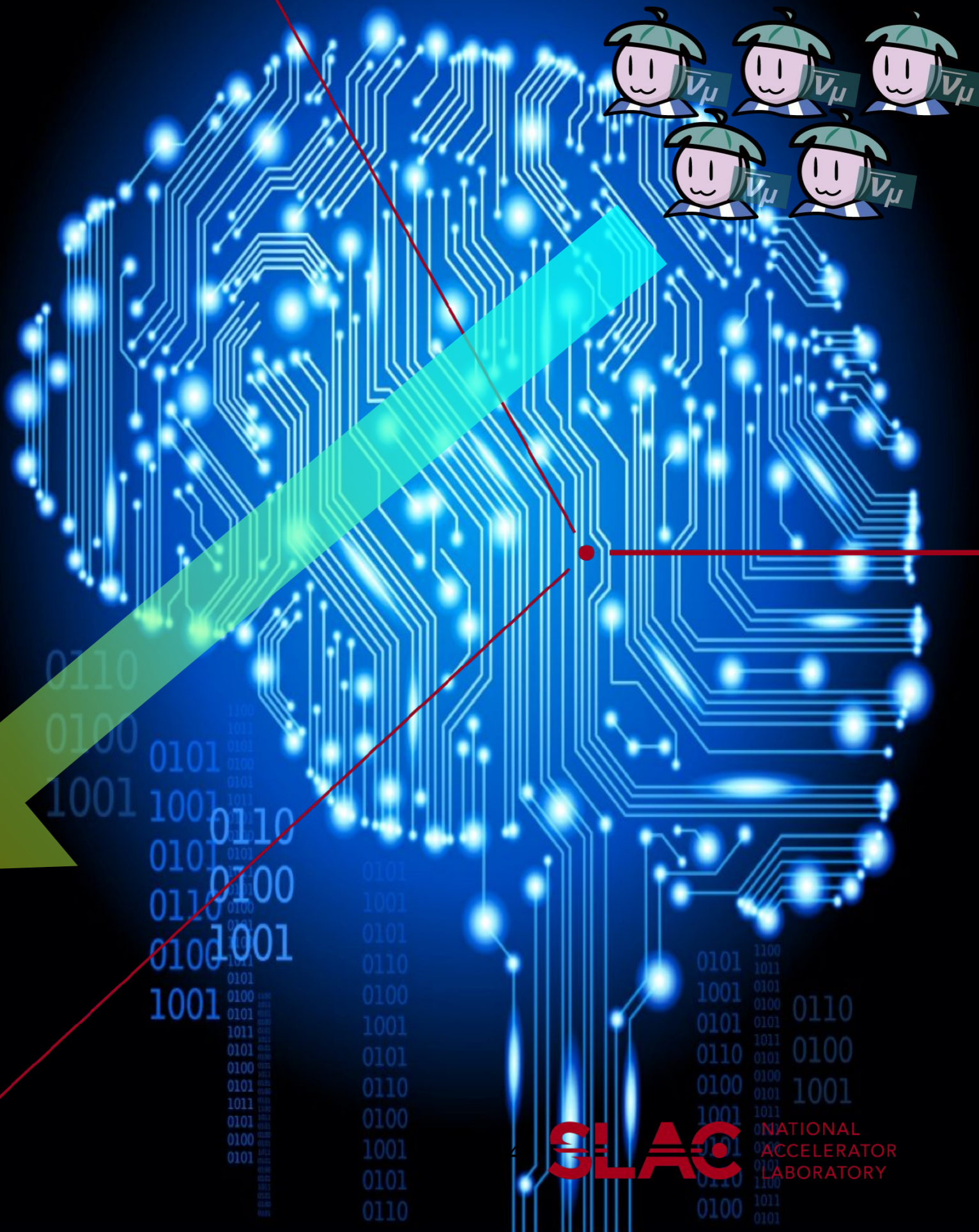
Machine Learning in Neutrino Physics

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ML-based 3D Data Reconstruction

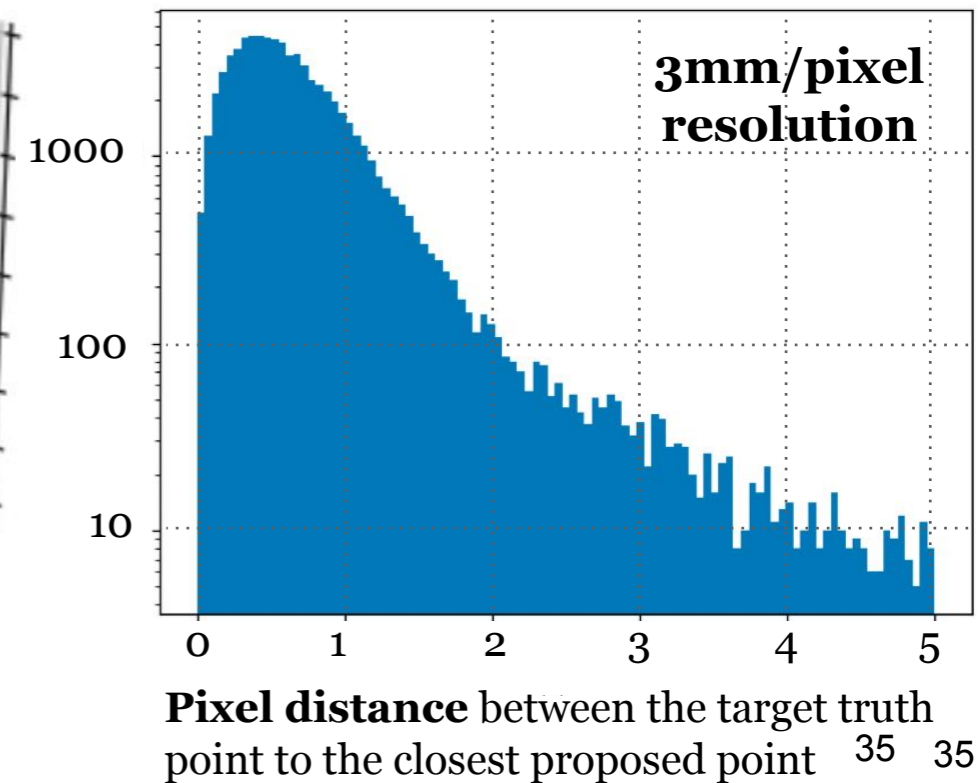
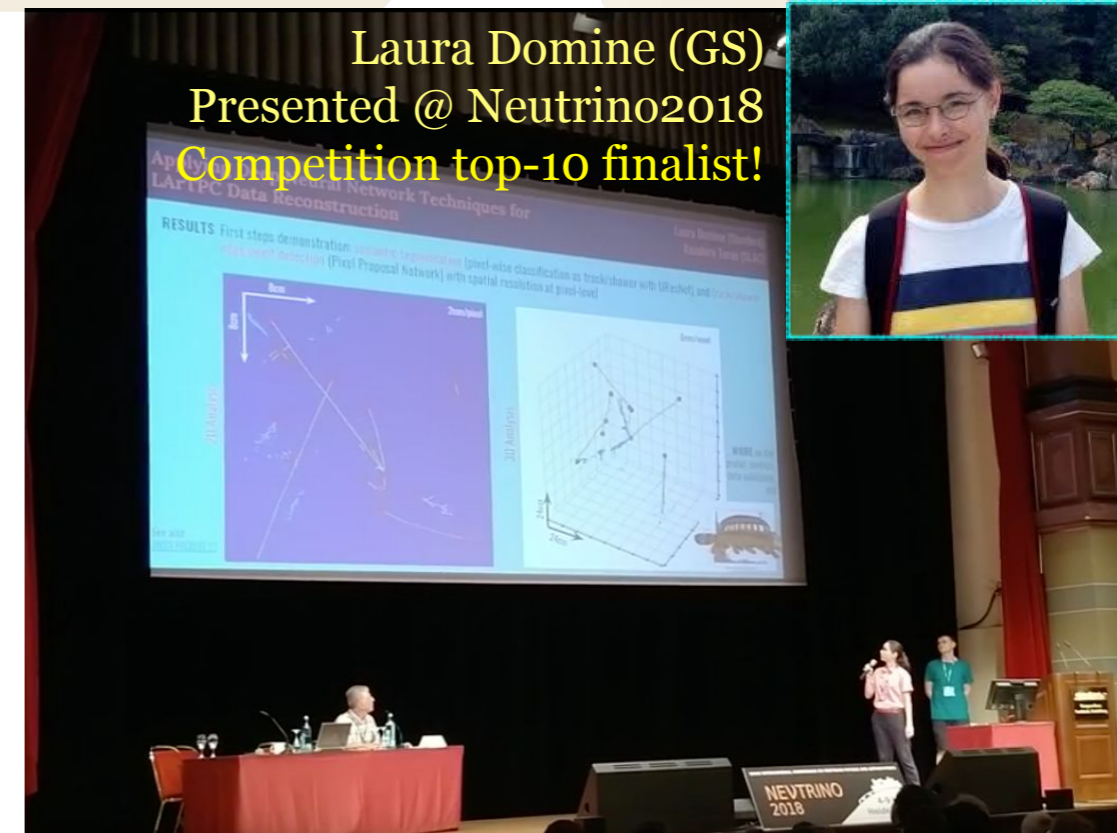
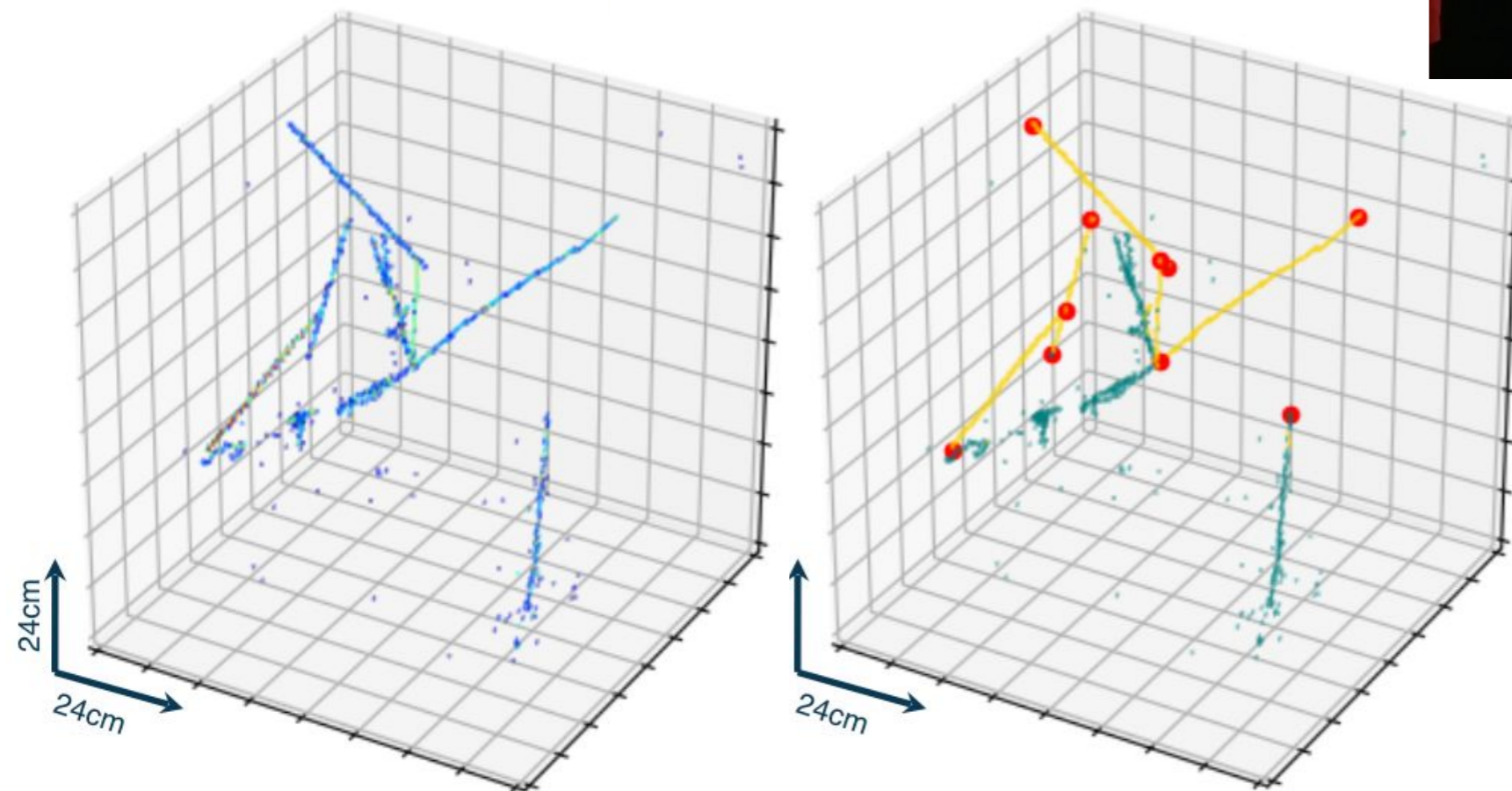


Toward “Reconstruction Chain” Machine Learning in Neutrino Physics

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Multi-Task Network Cascade

- **Chain of Segmentation + Detection**
 - Feature points: “shower start” and “track edges”
 - Classify each pixel into “shower” vs. “track”
- **Extension to 3D data**
 - Change in tensor dimensions, identical algorithms

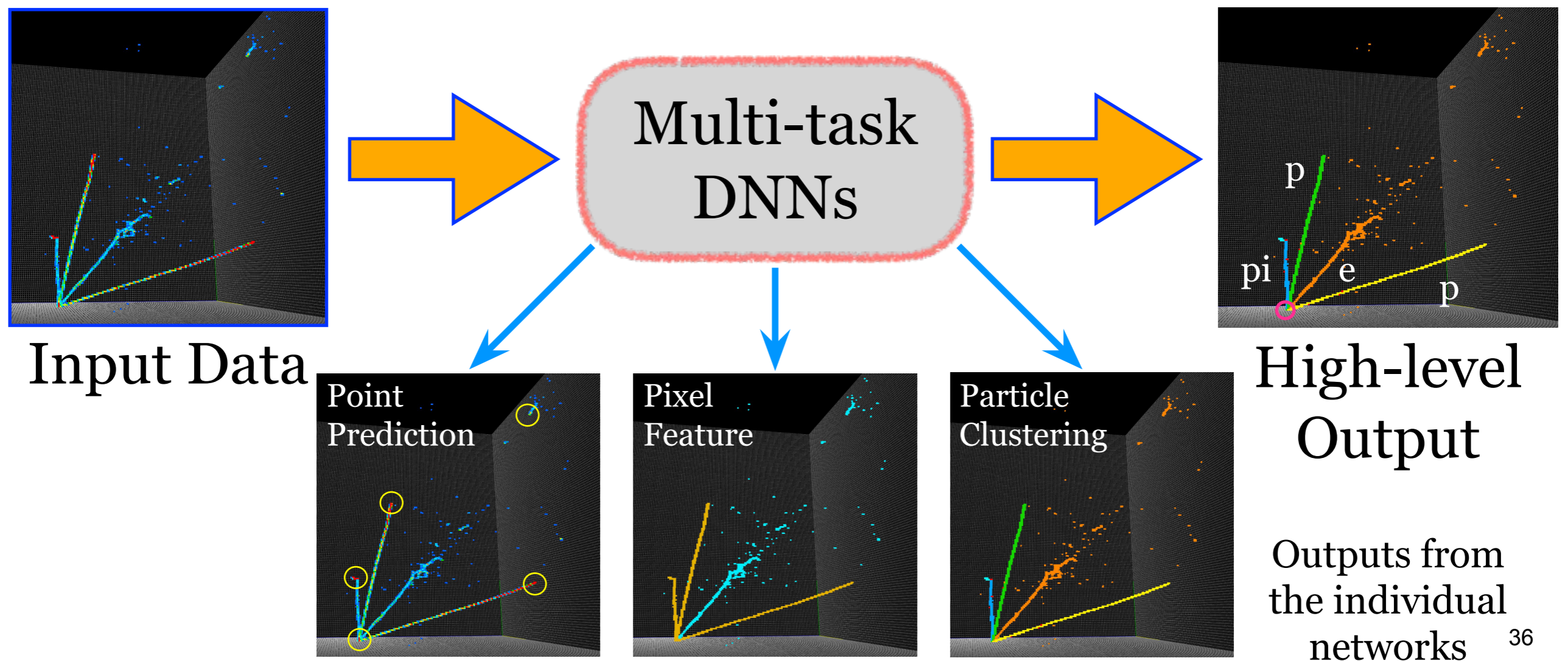


Toward “Reconstruction Chain”

Machine Learning in Neutrino Physics

Multi-task DNN for Physics Reconstruction

Introduce physical feature extraction tasks (auxiliary targets) to bias the data transformation path to support producing a logical conclusion. Optimize the whole reconstruction chain.



Deep CNN for LARGE Detectors? (scalability)

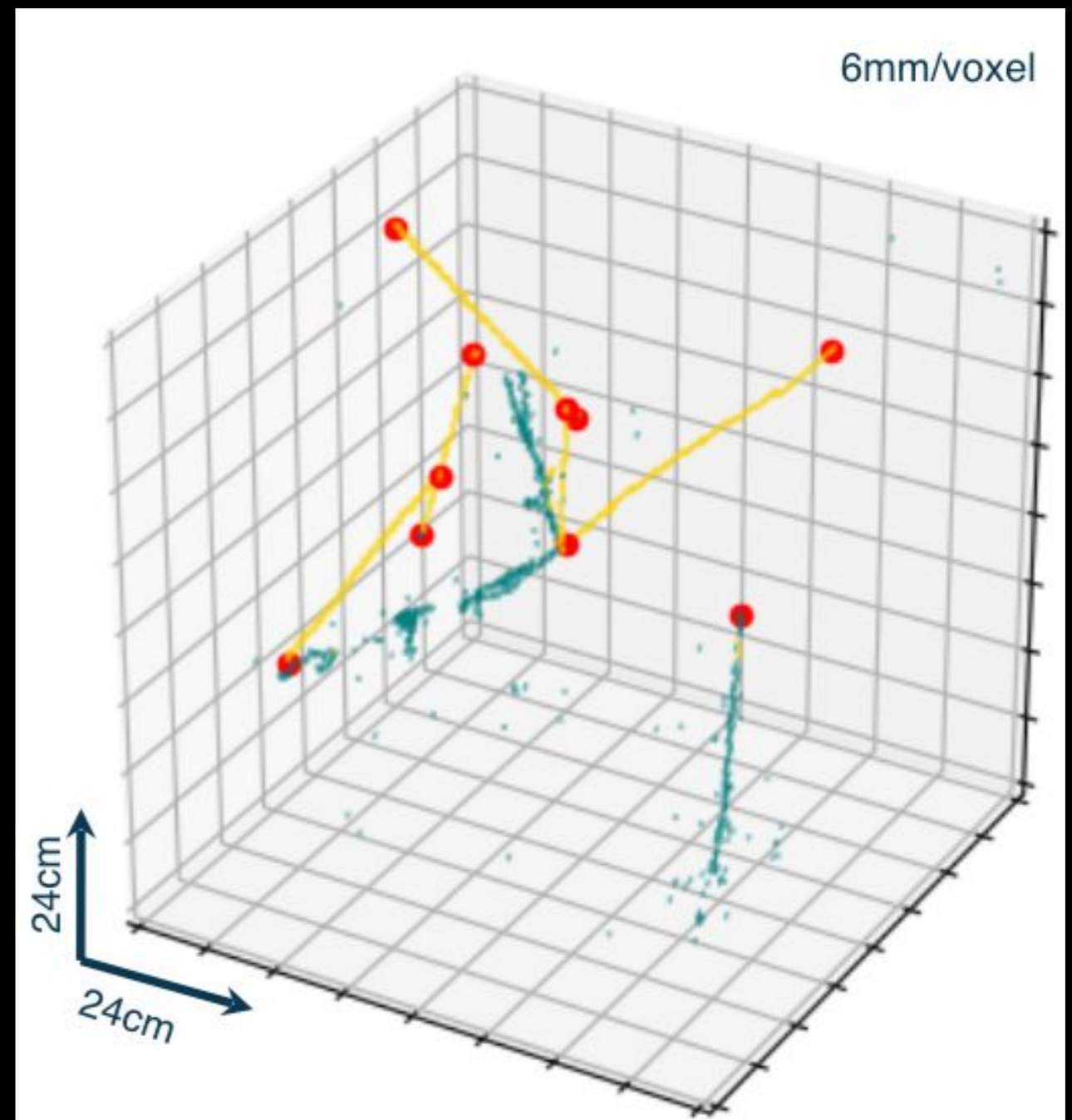
Machine Learning for Particle Image Analysis

Data feature: generally sparse, locally dense image, and very large volume (1 E10-20 pixels)

Issues using standard CNNs

- **Inefficient** calculations (“zero” matrix elements)
- **Prohibitive resource** usage (memory, time)
- **Degraded performance**

... terrible scaling = garbage!



Deep CNN for LARGE Detectors? (scalability)

Machine Learning for Particle Image Analysis

Data feature: generally sparse,
locally dense image, and very large
volume (1 E10-20 pixels)

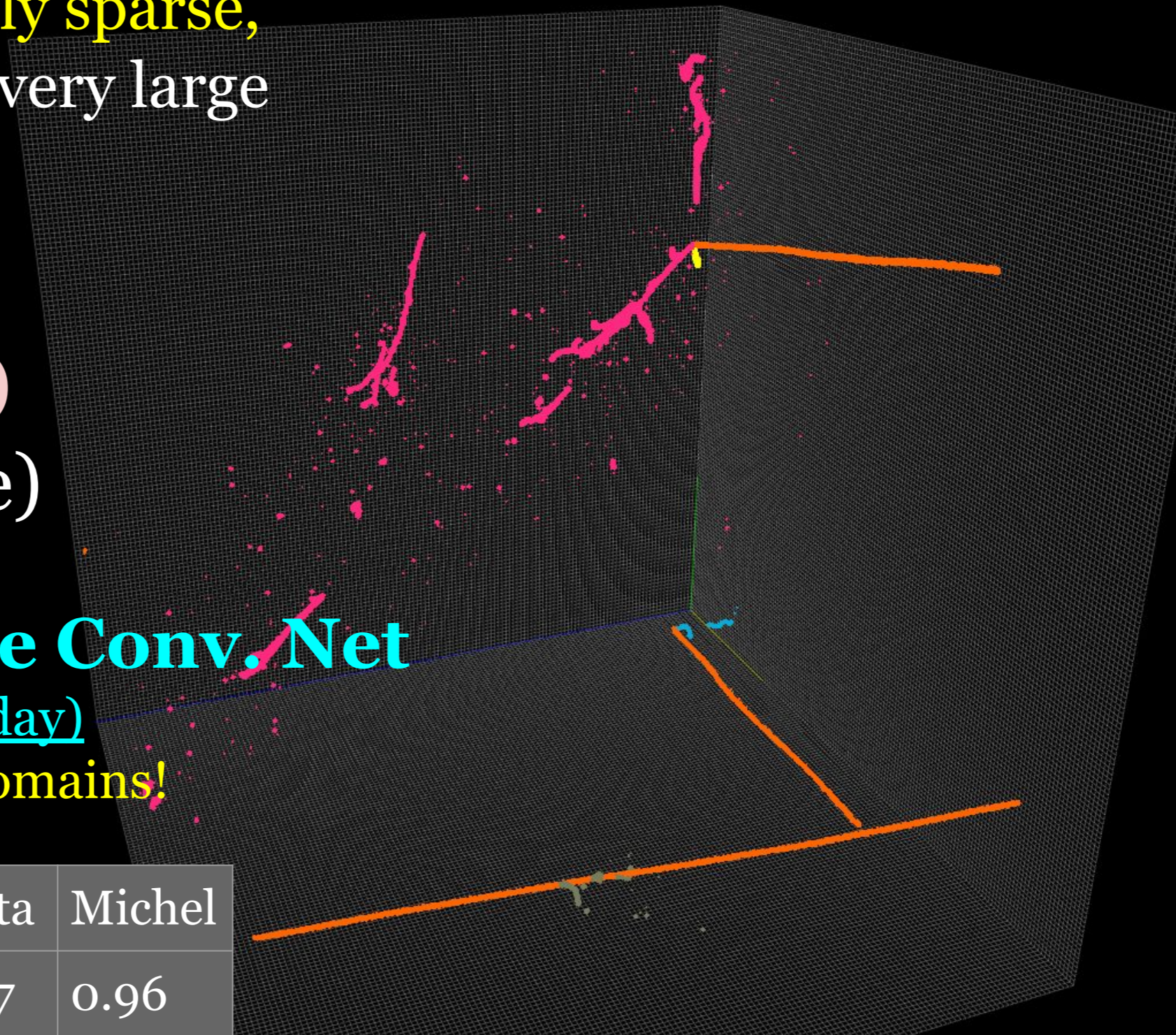
Got a solution :)
(right: 768^3 volume)

Submanifold Sparse Conv. Net

Talk by [Laura Domine \(Thursday\)](#)

Great for LArTPC and other domains!

| Type | HIP | MIP | Shower | Delta | Michel |
|------|------|------|--------|-------|--------|
| Acc. | 0.99 | 0.98 | 0.99 | 0.97 | 0.96 |

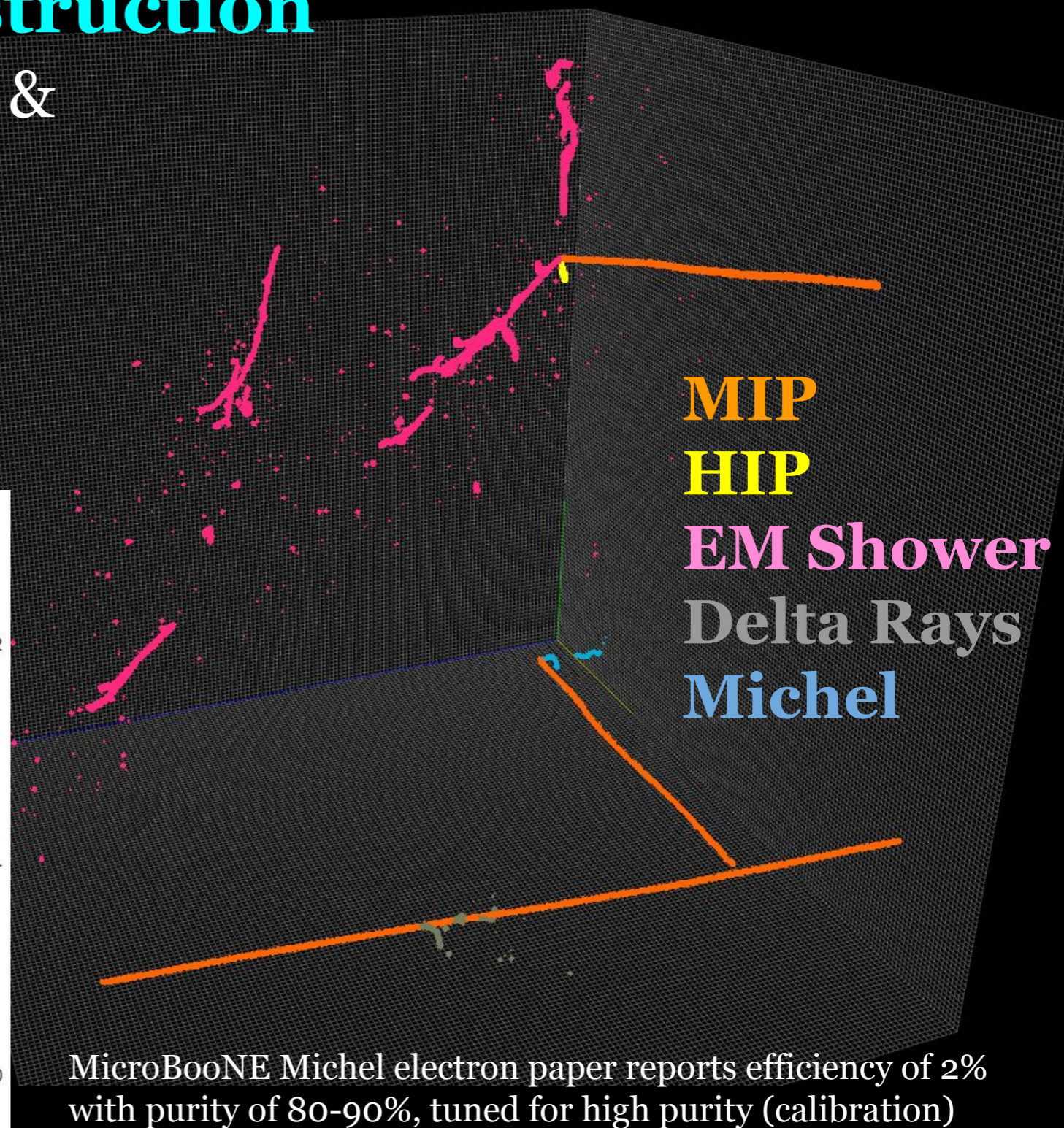
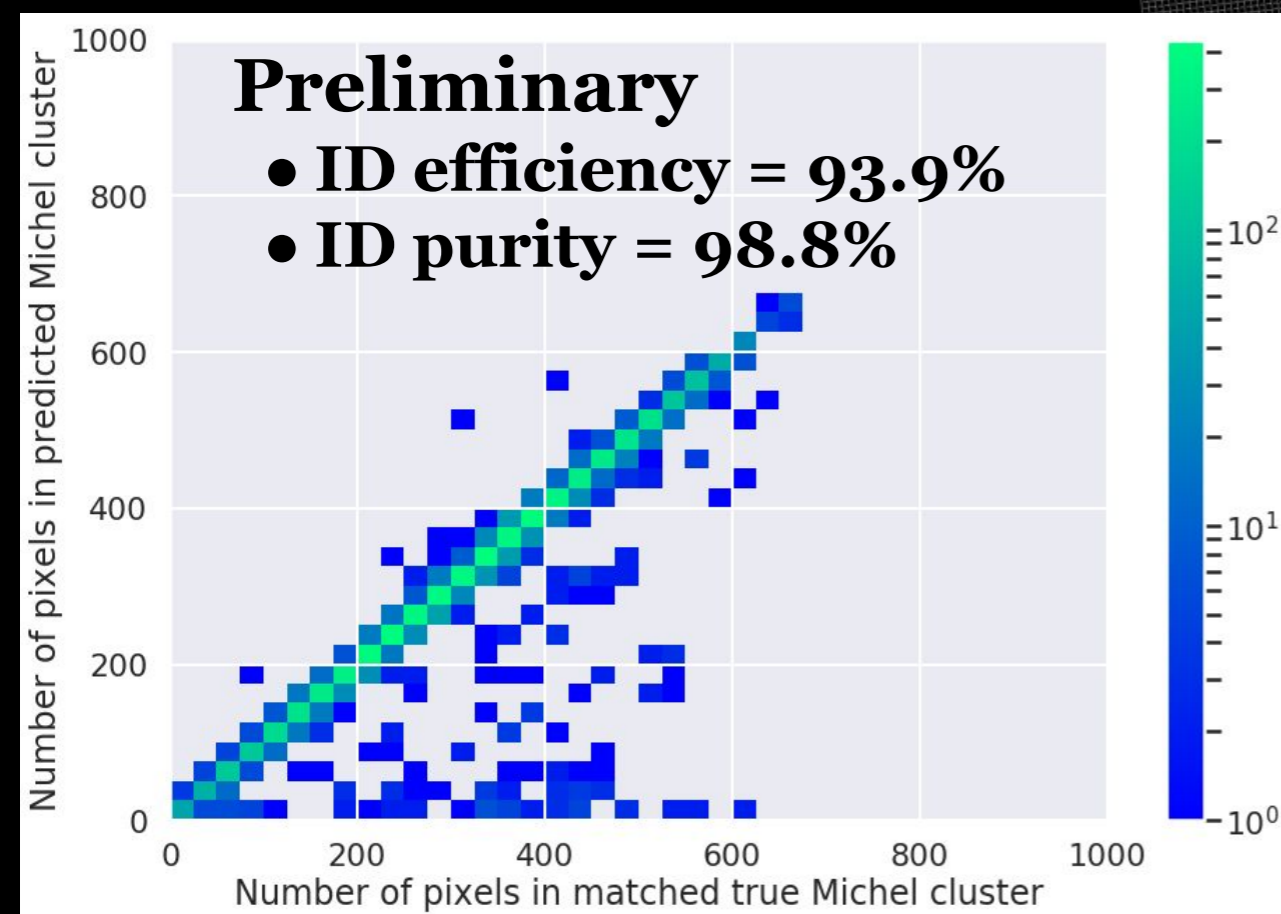


Deep CNN for LARGE Detectors? (scalability)

Machine Learning for Particle Image Analysis

Michel Electron Reconstruction

- Run spatial DBscan for MIP & Michel pixels (separately)
- Keep only Michel clusters which edge touches with an edge of a MIP cluster



Reproducible technique sharing is important...

- Submanifold Sparse Conv. Net for scalability
 - See [Laura's talk](#), and our benchmark ... [arXiv: 1903.05663](#)
 - Open data sample: [DOI 10.17605/OSF.IO/VRUZP](#)
 - Software stuck: [Singularity](#) or [Docker](#) container
 - Implementation: [github repo](#)

Toward HPC: contact them if you want help!

- SSCN + Horovod + custom MPI for production
- [Corey Adams](#) (ANL)
 - KNL/GPU nodes @ ALCF
- [Eric Church](#), [Jan F Strube](#), [Alexander R. Hagen](#) (PNNL)
 - SummitDev Intel Power8, now moving onto Power9

Experimental neutrino physics:

- **Detector trend: particle imaging**
 - LArTPC is the current frontier for imaging
- **Many applications from computer vision**
 - ML-based full data reconstruction being developed
 - Active but not mentioned: data/sim domain adaptation
- **Next few years**
 - Integration of ML-based reconstruction
 - Data/Simulation domain adaptations
 - Software stack development toward HPC

A young girl with dark hair, wearing a bright pink patterned winter jacket, stands on a balcony or walkway. She is pointing her right hand towards a scenic mountain landscape. In the background, a snow-covered mountain peak rises above a small town with wooden buildings. A river flows through the valley below. The sky is blue with scattered white clouds.

Thank you
for listening

and

Thank YOU
for organizing
ACAT2019!

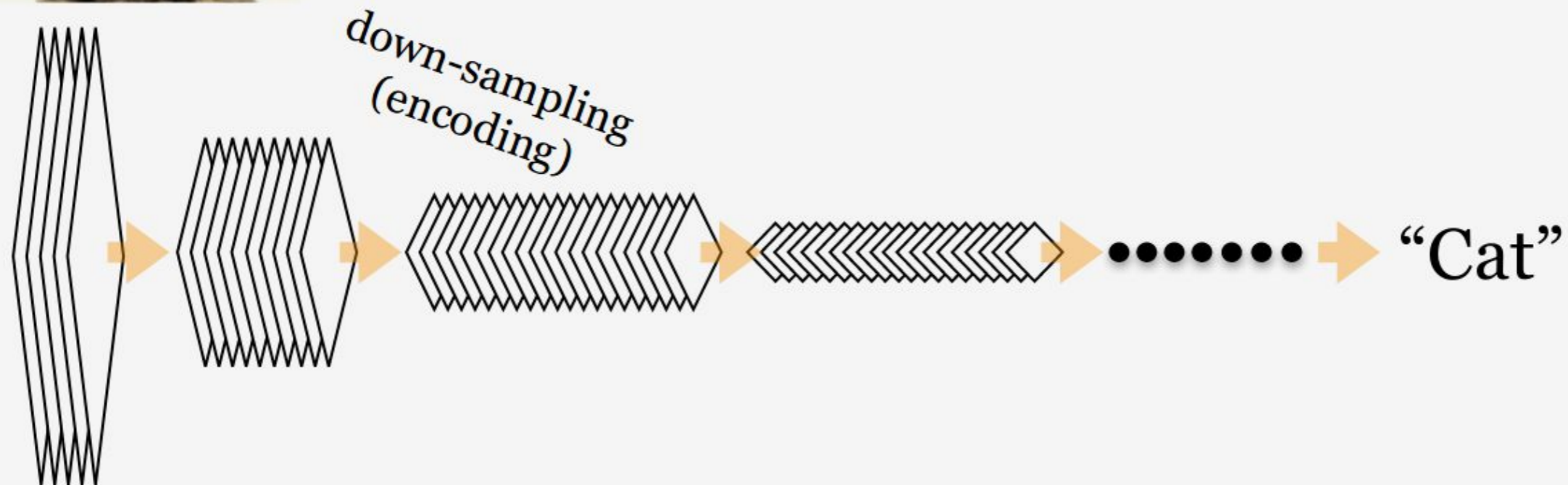
Back Up Slides

HO HO HO

Machine Learning for Particle Image Analysis

SLAC

How image classification works



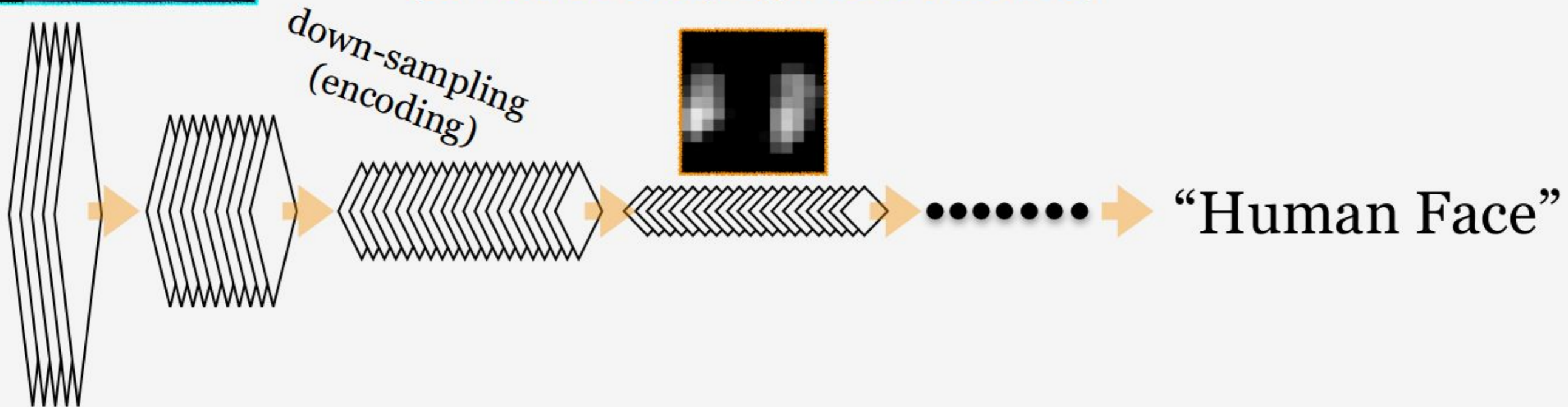
HO HO HO

Machine Learning for Particle Image Analysis

How image classification works



Intermediate Data Tensor
(low-resolution, high-level features)



HO HO HO

Machine Learning for Particle Image Analysis

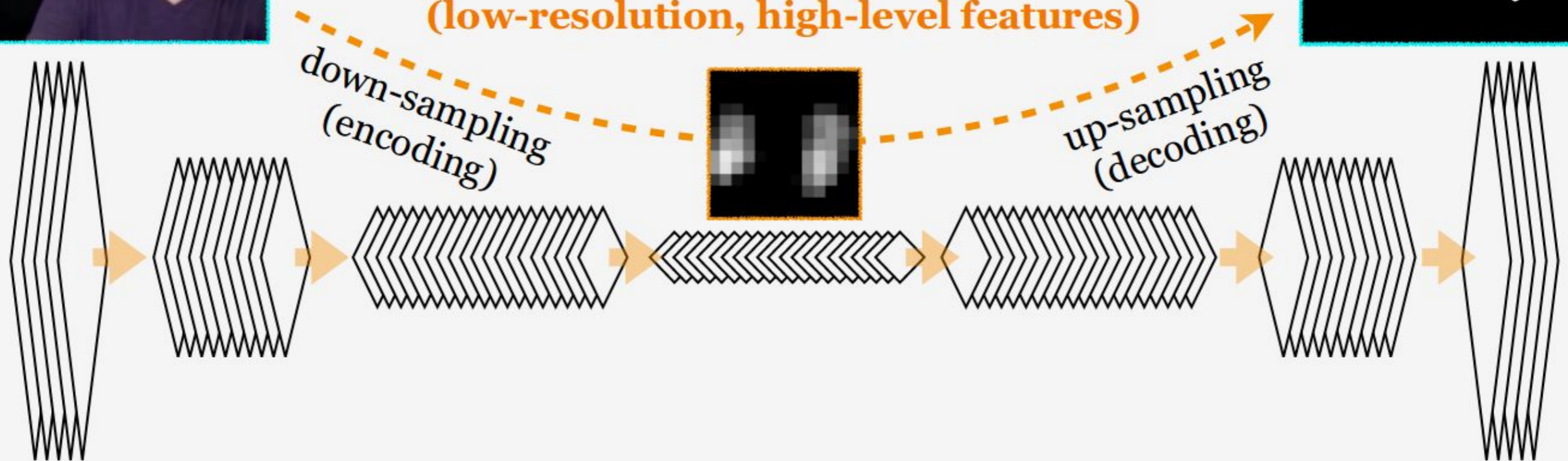
SLAC

How pixel segmentation works

- Combine “up-sampling” + convolutions
- Output: “learnable” interpolation filters



Intermediate Data Tensor
(low-resolution, high-level features)



HO HO HO

Machine Learning for Particle Image Analysis

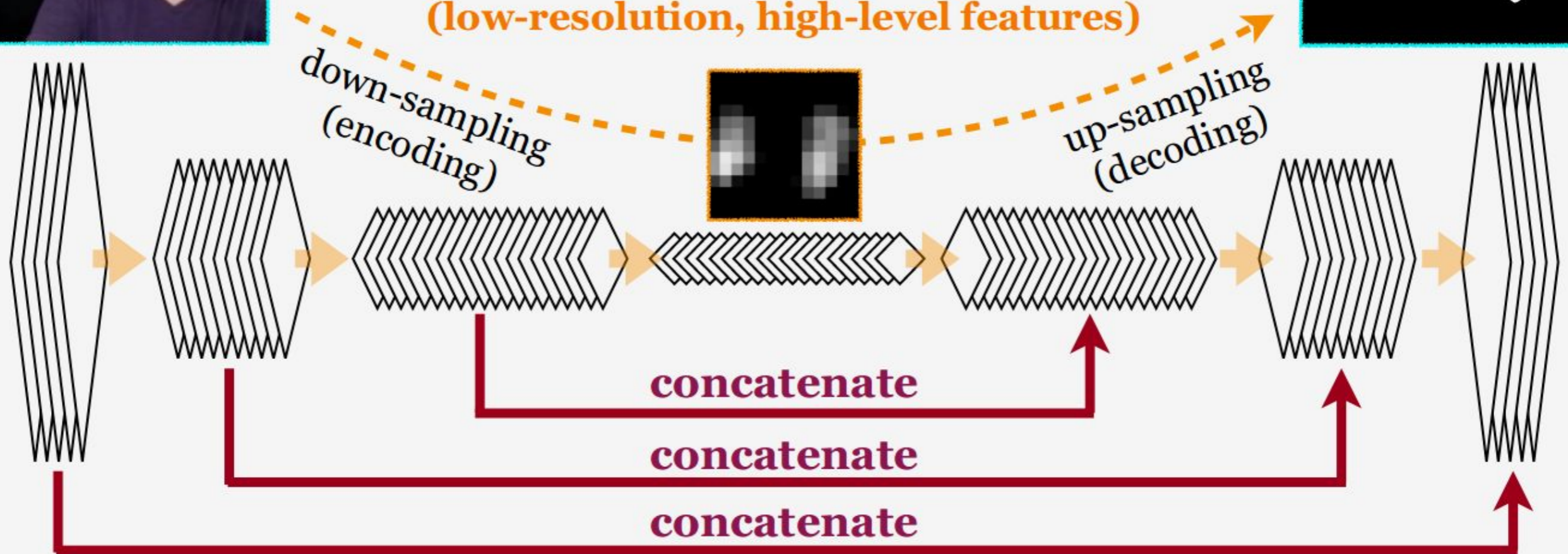
SLAC

How pixel segmentation works

- Combine “up-sampling” + convolutions
- Output: “learnable” interpolation filters



Intermediate Data Tensor
(low-resolution, high-level features)

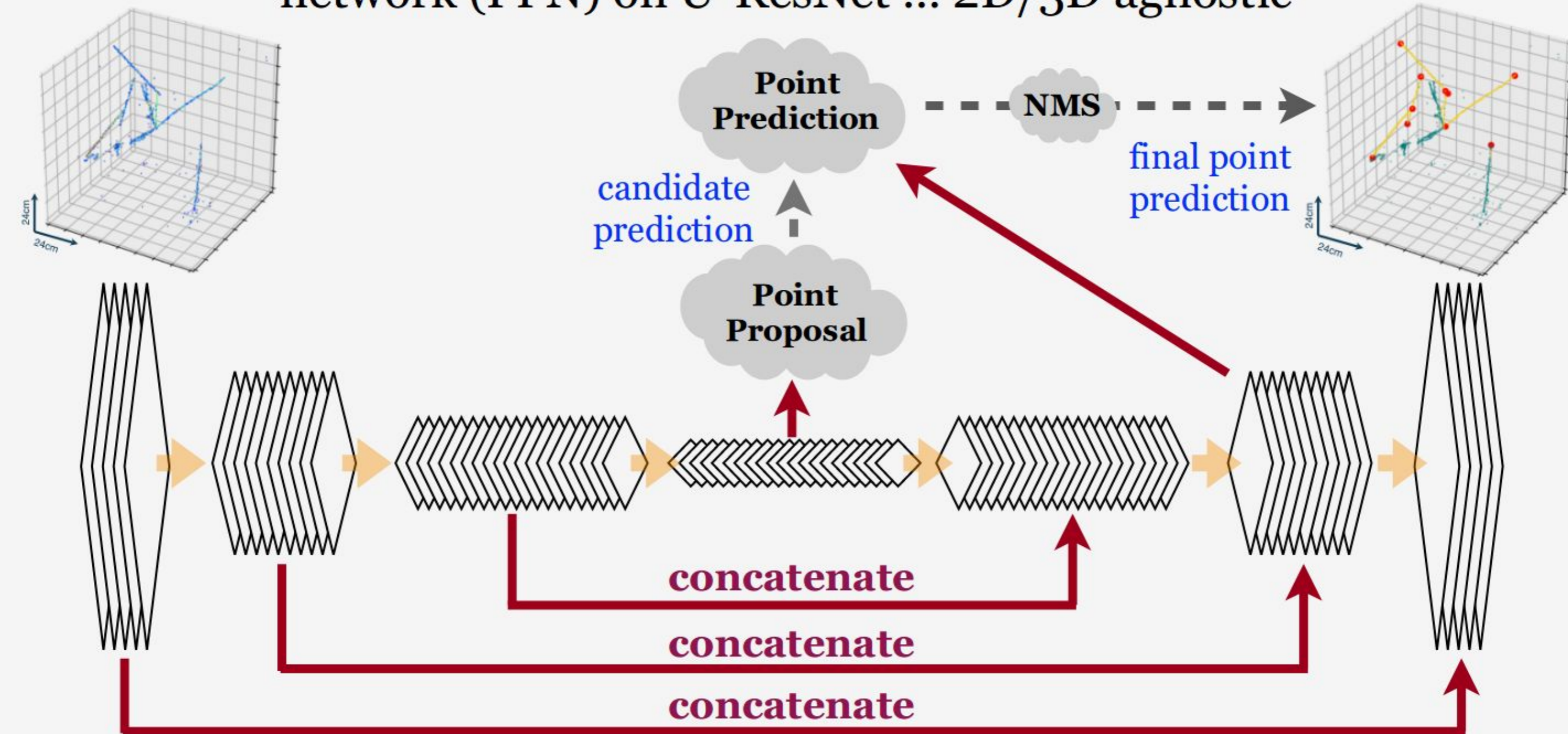


Concatenation recovers spatial resolution information

HO HO HO

Machine Learning for Particle Image Analysis

Parasitic multi-task scheme for point prediction network (PPN) on U-ResNet ... 2D/3D agnostic



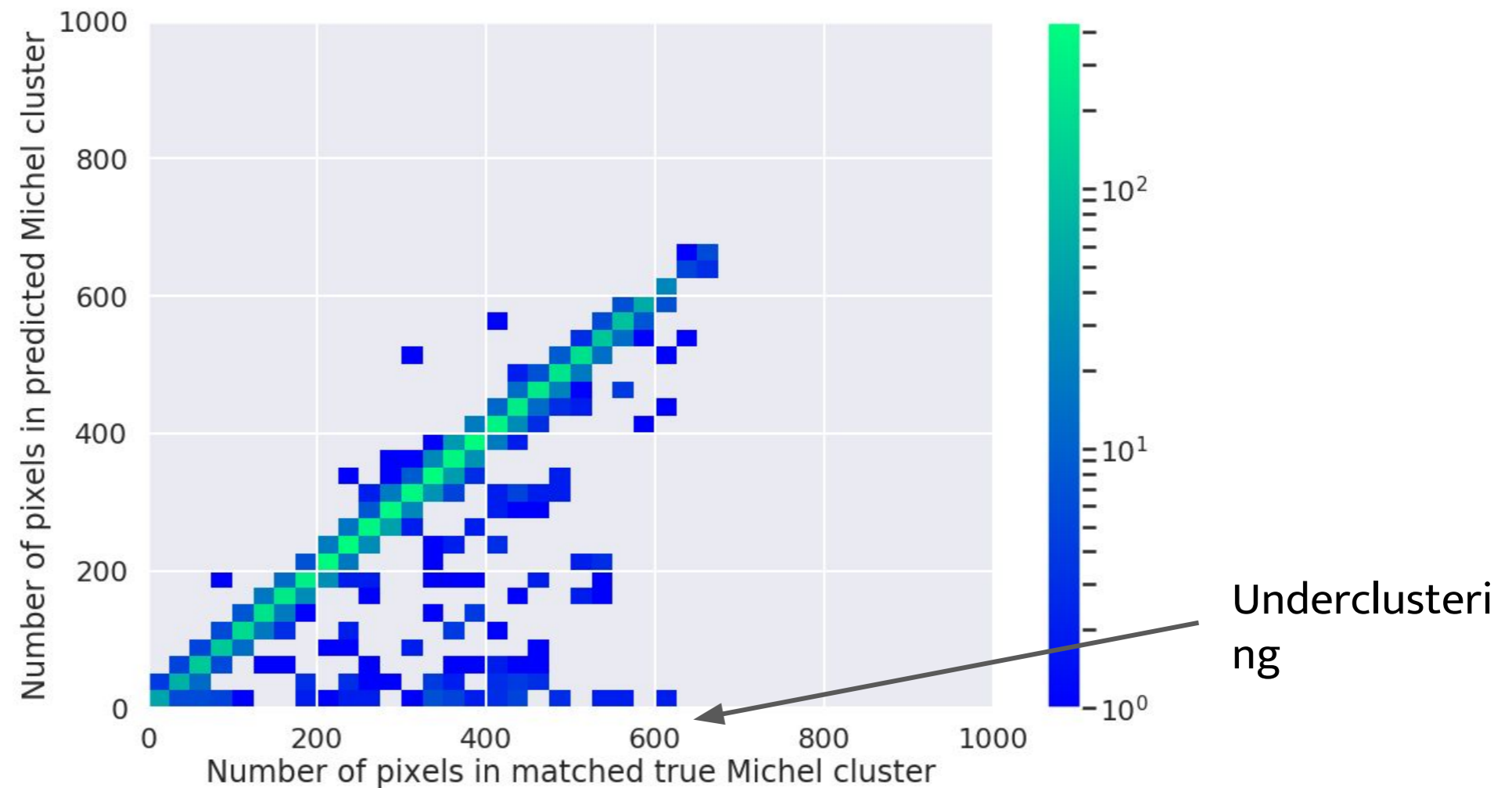
Concatenation recovers spatial resolution information

Relabeling study

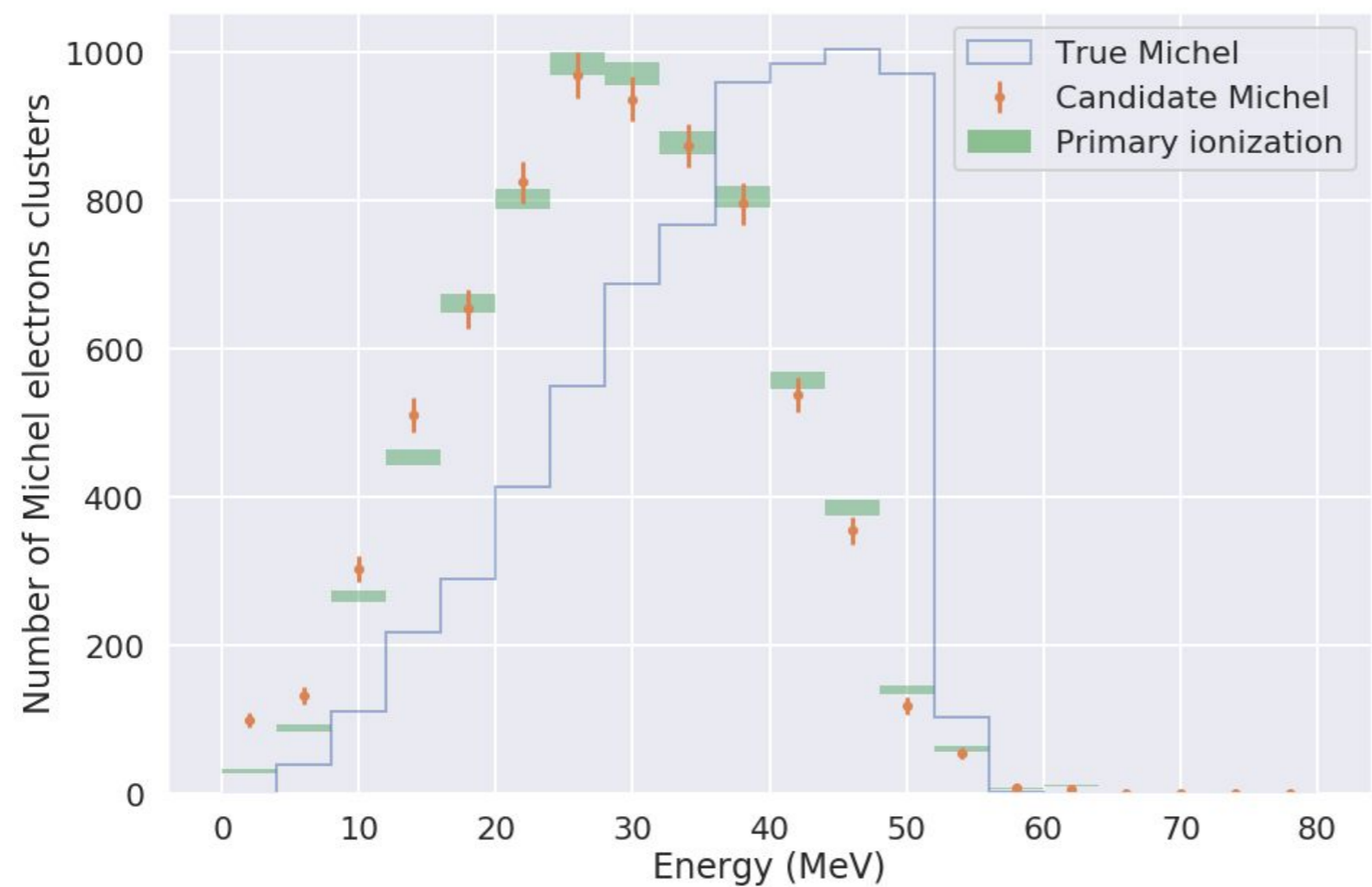
Relabeled dataset = only primary ionization is labelled as Michel electrons

| Train data | Regular | | Relabeled | Relabeled + Weighting |
|---------------------|---------|-----------|-----------|--------------------------|
| Test data | Regular | Relabeled | | |
| HIP | 98.0% | 98.1% | 98.1% | 99.3% |
| MIP | 99.4% | 99.2% | 99.4% | 98.1% |
| EM shower | 99.4% | 97.9% | 99.2% | 99.2% |
| Delta rays | 85.7% | 94.8% | 96.0% | 97.2% |
| Michel electrons | 56.6% | 94.4% | 94.7% | 95.7% |

Number of pixels in candidate vs matched Michel cluster



Michel electrons energy spectrum reconstruction



| | |
|-------------------------------------|-------|
| Sample size | 7105 |
| Identification purity | 98.8% |
| Identification efficiency | 93.9% |
| Cluster efficiency | 96.1% |
| Cluster based purity | 97.8% |
| Simulation only, next step is data! | |

Detectors

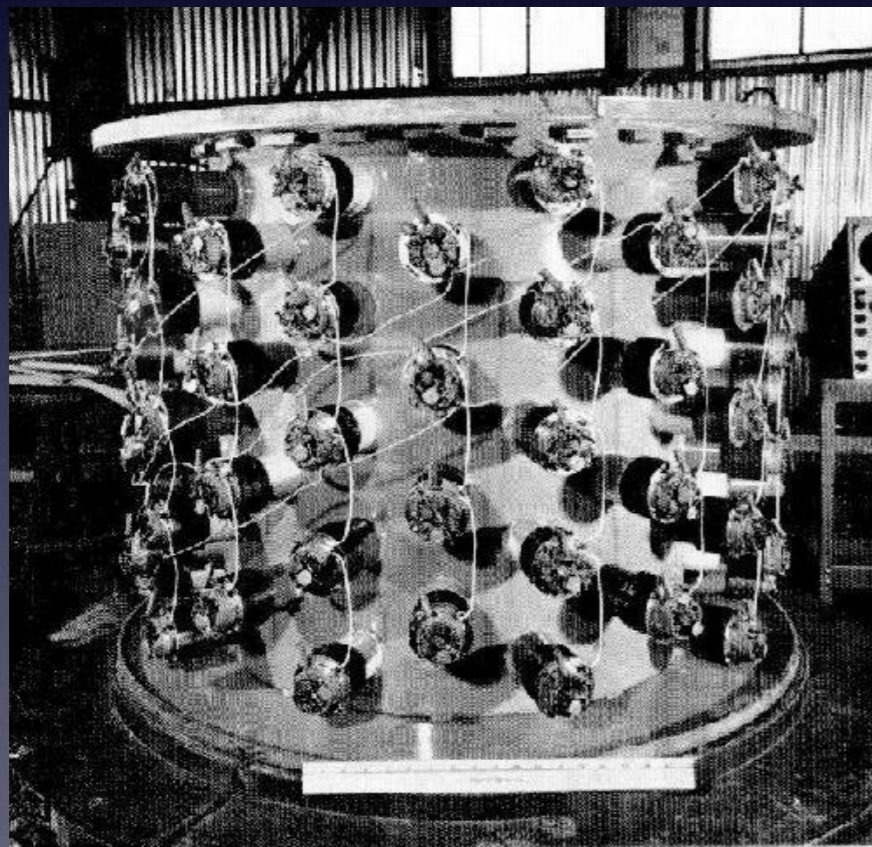
Detecting Neutrinos:

BMB

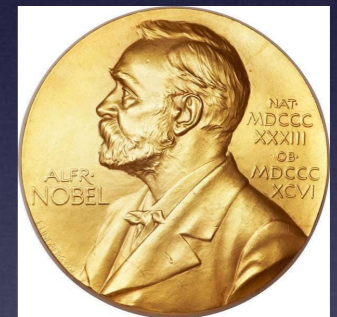
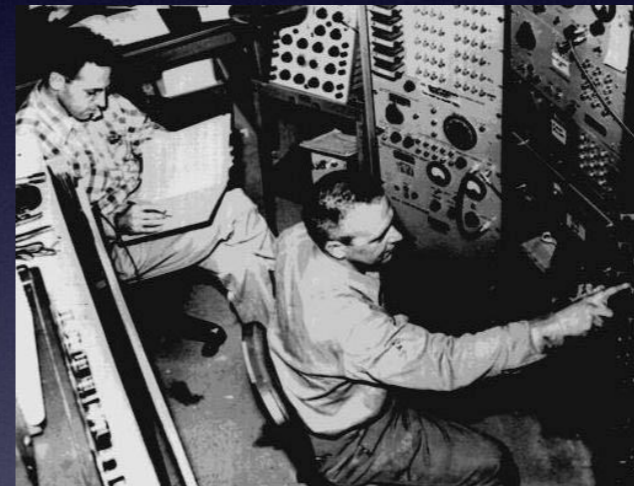
We cannot observe neutrinos, but we can detect particles that come out of a neutrino interaction.

Neutrino Oscillation Experiments (I)

Evolution of Detectors



Cd-doped water
0.4 ton, 100 PMTs



Inverse Beta Decay (IBD)

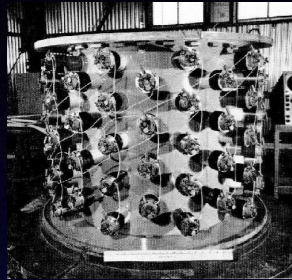


by Reines & Cowan (Nobel Prize 1995)

First neutrino detection

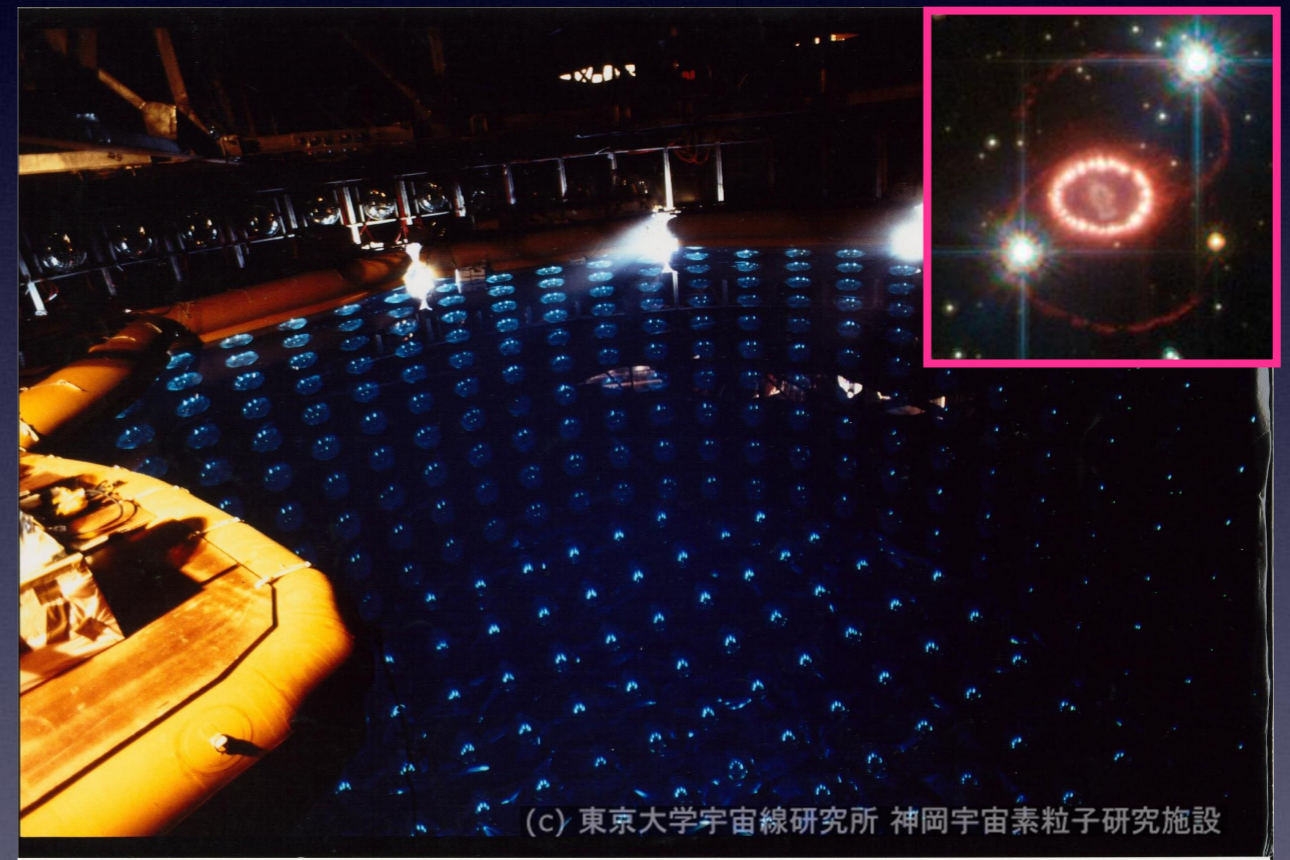
Neutrino Oscillation Experiments (I)

Evolution of Detectors



Cd-doped water
0.4 ton, 100 PMTs
(1956)

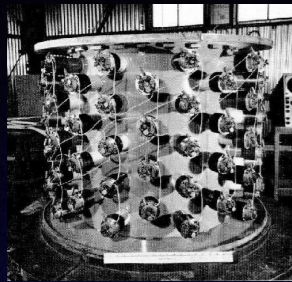
**Birth of neutrino
astrophysics!**



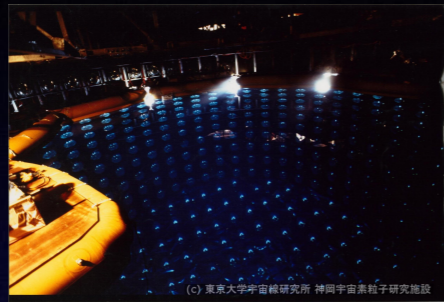
KamiokaNDE Detector
3 kton ultra-pure water, 1000 20" PMTs
(shared Nobel Prize 2002)

Neutrino Oscillation Experiments (I)

Evolution of Detectors



Cd-doped water
0.4 ton, 100 PMTs
(1956)

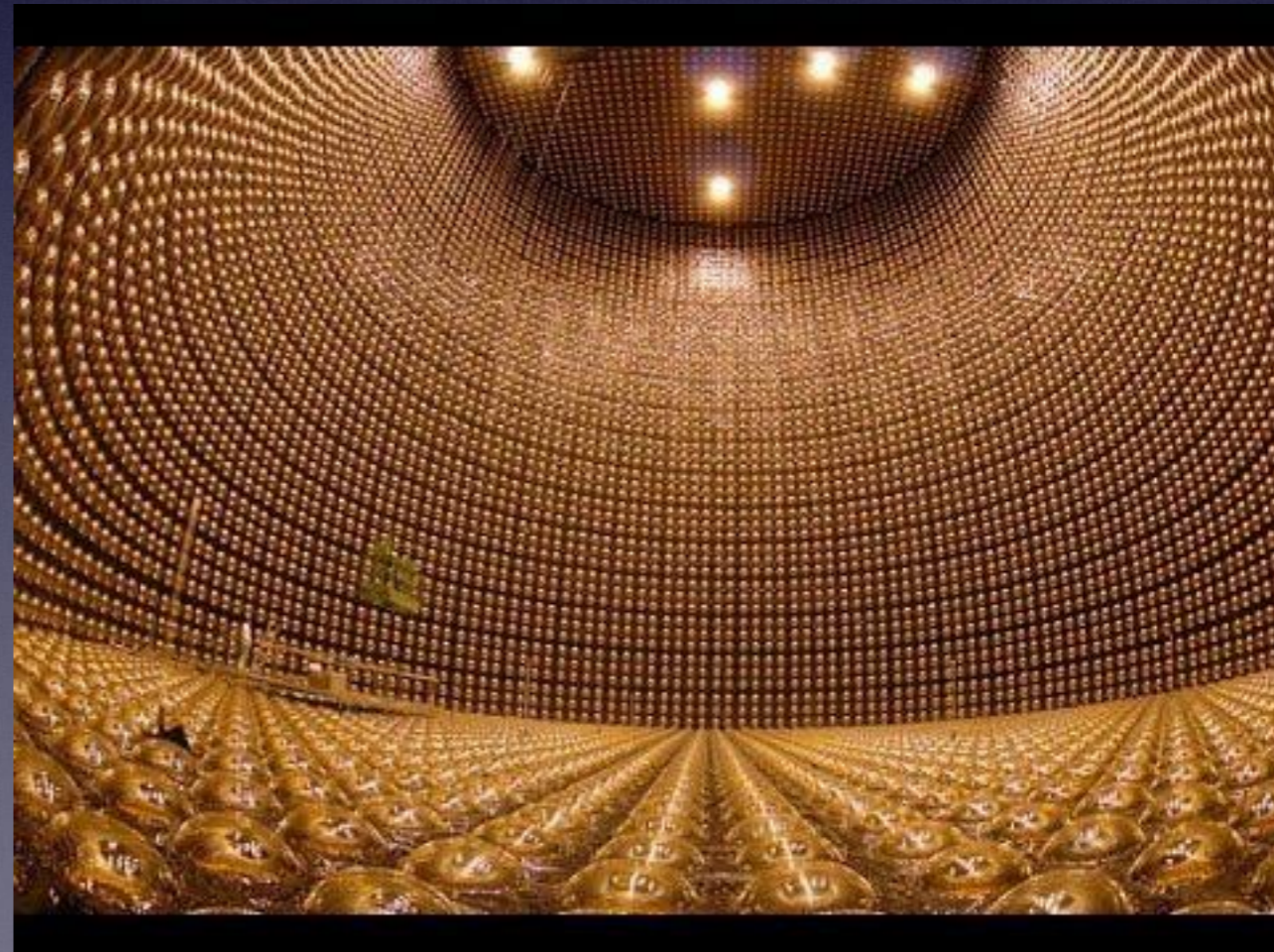


Ultra-pure water
3 kton, 1000 PMTs
(1983)

Discovery of ν_{atmo} oscillation!

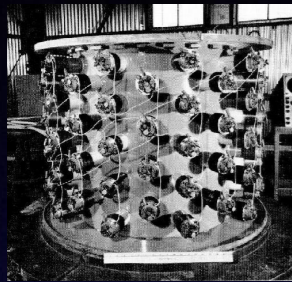


Super-KamiokaNDE
50 kton ultra-pure water,
11000 PMTs
(shared Nobel Prize 2015)

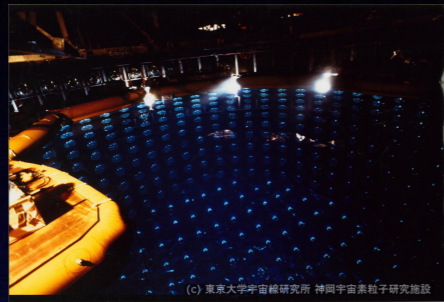


Neutrino Oscillation Experiments (I)

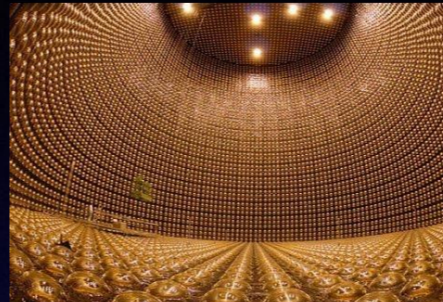
Evolution of Detectors



Cd-doped water
0.4 ton, 100 PMTs
(1956)

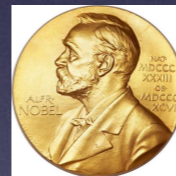


Ultra-pure water
3 kton, 1000 PMTs
(1983)

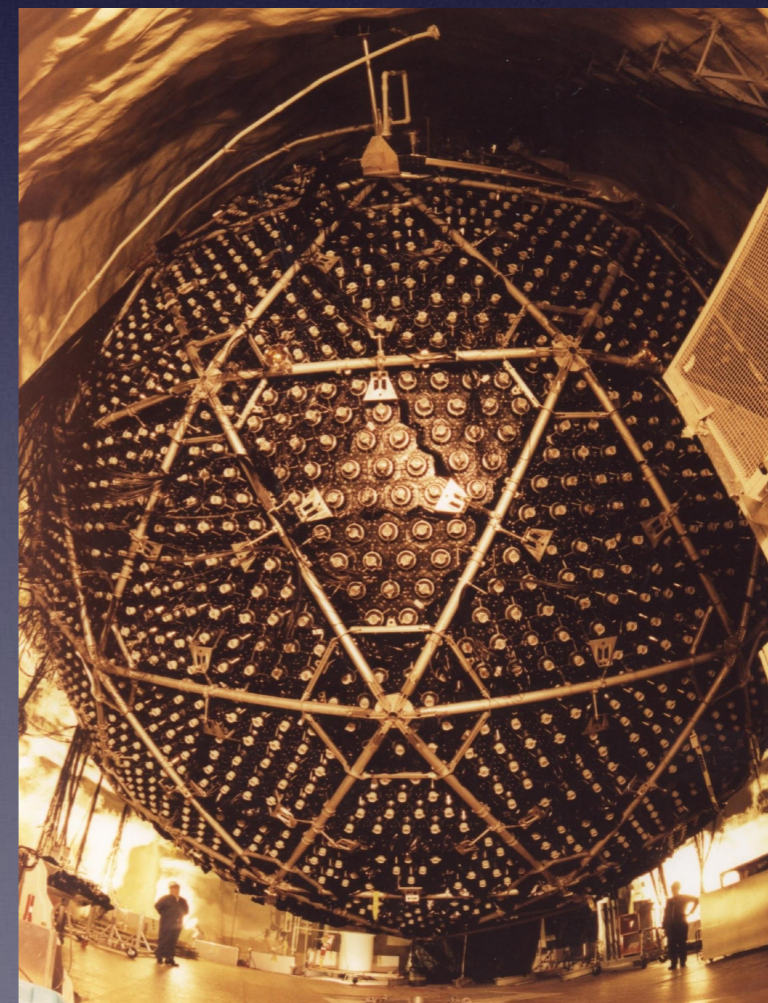


Ultra-pure water
50 kton, 11000 PMTs
(1996)

Discovery of ν_{solar} oscillation!

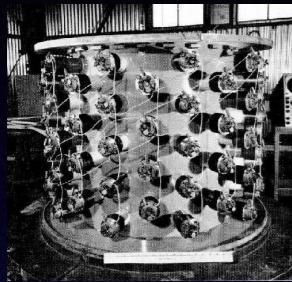


SNO
1 kton heavy-water Cherenkov,
9600 PMTs
(shared Nobel Prize 2015)

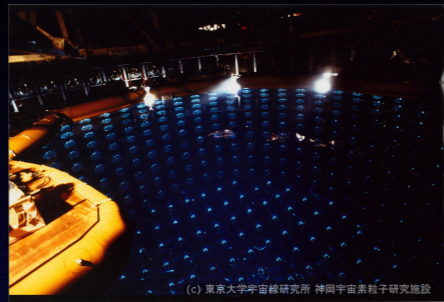


Neutrino Oscillation Experiments (I)

Evolution of Detectors



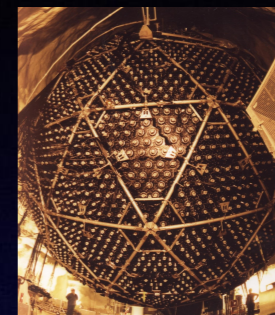
Cd-doped water
0.4 ton, 100 PMTs
(1956)



Ultra-pure water
3 kton, 1000 PMTs
(1983)



Ultra-pure water
50 kton, 11000 PMTs
(1996)



Heavy water
1 kton, 9600 PMTs
(1999)

Reactor neutrino oscillation! (the solar model is right!)

KamLAND

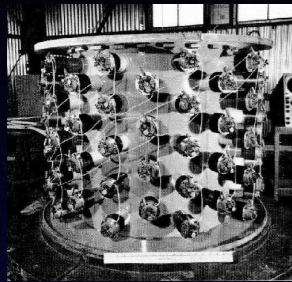
1 kton liquid scintillator, 1900 PMTs

My first neutrino experiment
(undergraduate RA @ UC Berkeley)

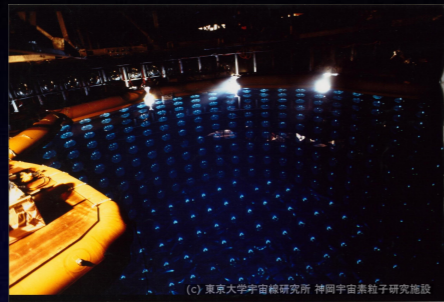


Neutrino Oscillation Experiments (I)

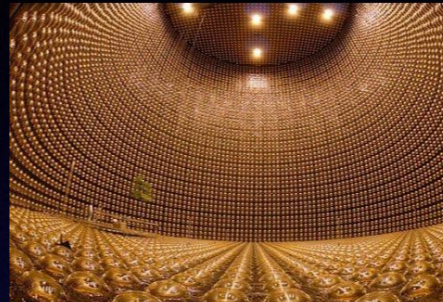
Evolution of Detectors



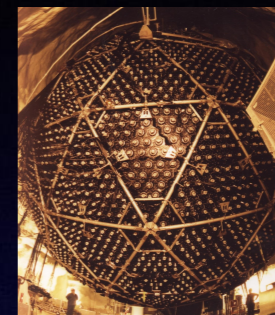
Cd-doped water
0.4 ton, 100 PMTs
(1956)



Ultra-pure water
3 kton, 1000 PMTs
(1983)



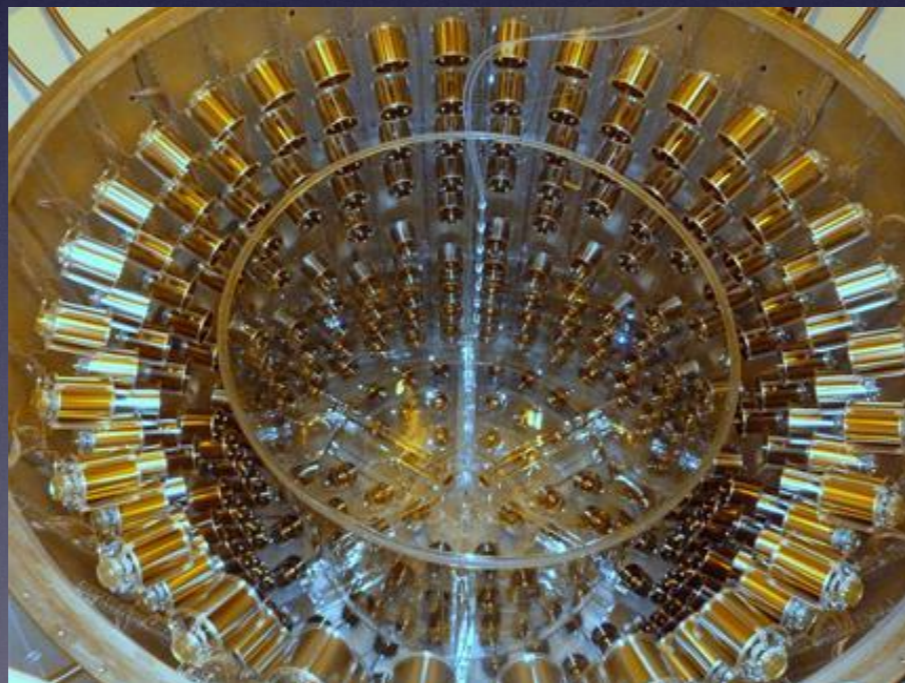
Ultra-pure water
50 kton, 11000 PMTs
(1996)



Heavy water
1 kton, 9600 PMTs
(1999)



Liquid Scintillator
1 kton, 1900 PMTs
(2002)



Gd-doped liquid scintillator
RENO, Daya Bay, Double Chooz

“Near” & “Far” design
2 x 16 ton detectors with 400
PMTs each (Double Chooz)

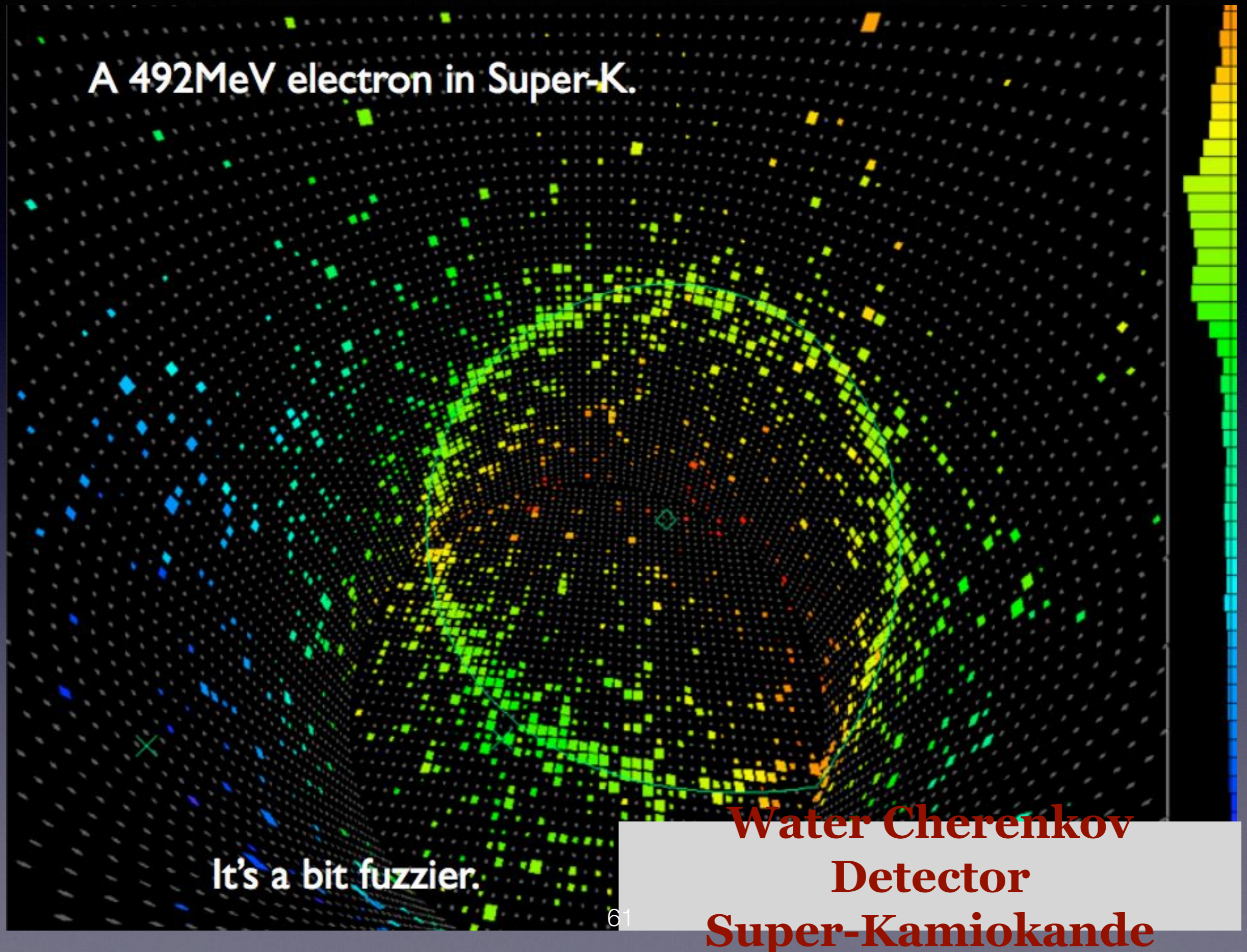
My Ph.D thesis! (MIT)
**“Last mixing
angle”** θ_{13}
Experiments!

Neutrino Oscillation Experiments (I)

A 603MeV muon in Super-K.



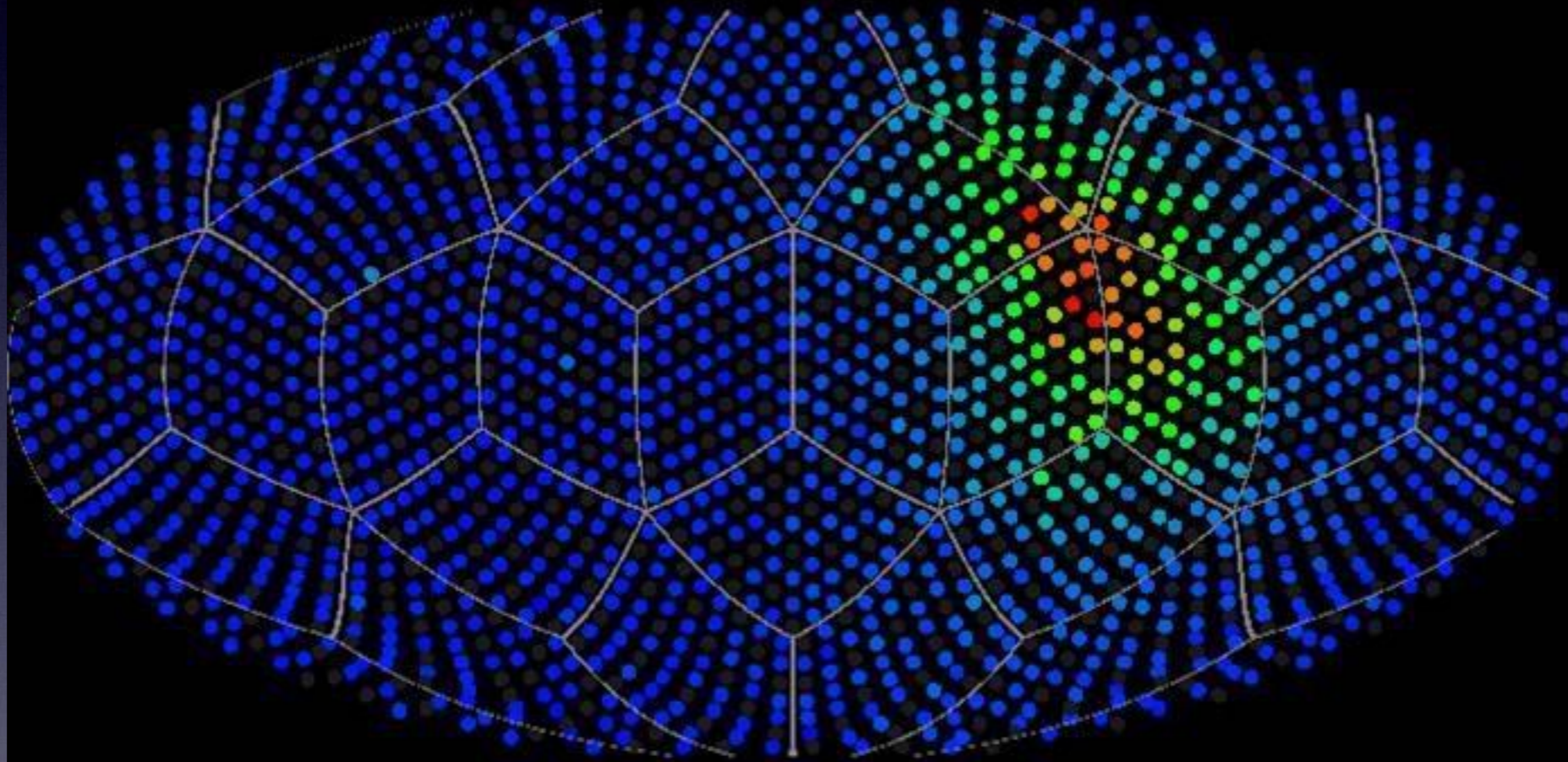
Neutrino Oscillation Experiments (I)



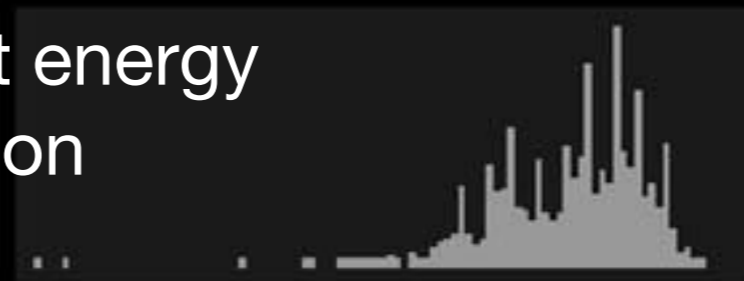
Neutrino Oscillation Experiments (I)

Liquid Scintillator Detector KamLAND

KamLAND Event Display
Run/Subrun/Event : 110/0/192
UT: Sat Feb 23 15:25:11 2002
TimeStamp : 13052924536
TriggerType : 0x3a10 / 0x2
Time Difference 28.3 msec
NumHit/Nsum/Nsum2/NumHitA : 1317/264/1322/46
Total Charge : 3.21e+05 (465)
Max Charge (ch): 2.22e+03 (640)



Less topological information
but excellent energy
resolution



Q : 0.4 222.3 444.1 665.9 887.7 1109.5 1331.3 1553.2 1775 1996.8 2218.6

Neutrino Oscillation Experiments (II)

How can we do better?

Three important detector features for oscillation measurement

$$P(\nu_\mu \rightarrow \nu_e) = \sin^2 2\theta \sin^2 \left(\frac{1.27 \Delta m^2 L}{E_\nu} \right)$$

Good Energy Resolution

Precise E_ν reduce oscillation uncertainty

Large Mass (scalable)

“More” statistics to measure rare physics process

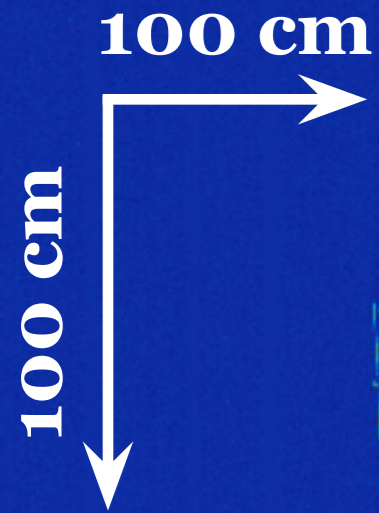
Particle ID Capability

Better ν identification background rejection


Challenge S

Analysis Challenges

100 cm
100 cm



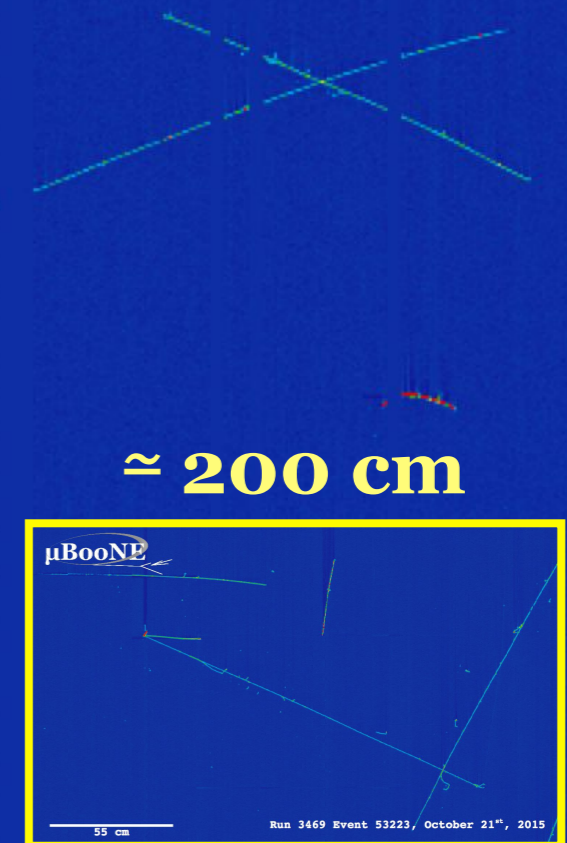
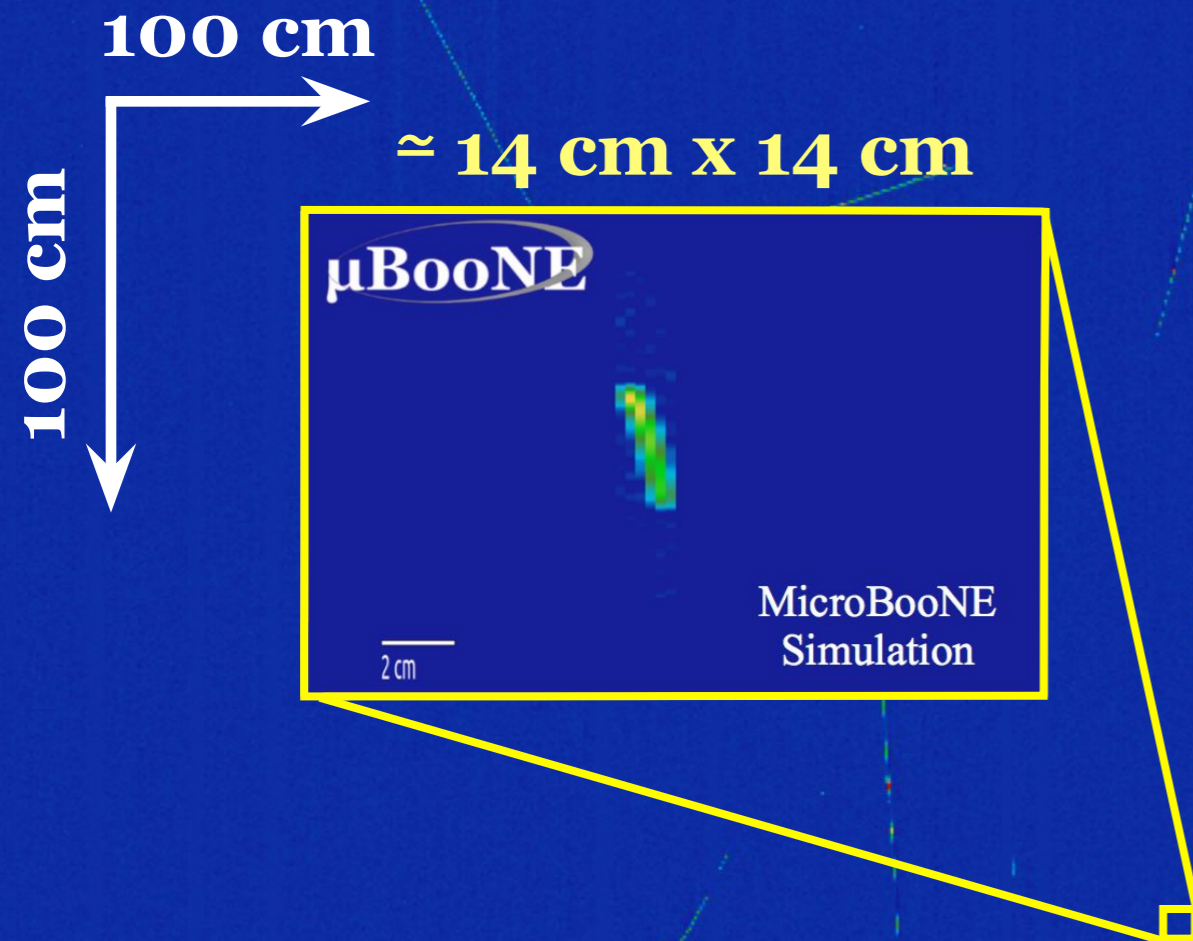
μ BooNE



There may be lots of backgrounds

Cosmic Data : Run 6280 Event 6812 May 12th, 2016

Analysis Challenges



Interaction vertex can be anywhere in LAr, varying in size (cm ~ meters)



Cosmic Data : Run 6280 Event 6812 May 12th, 2016

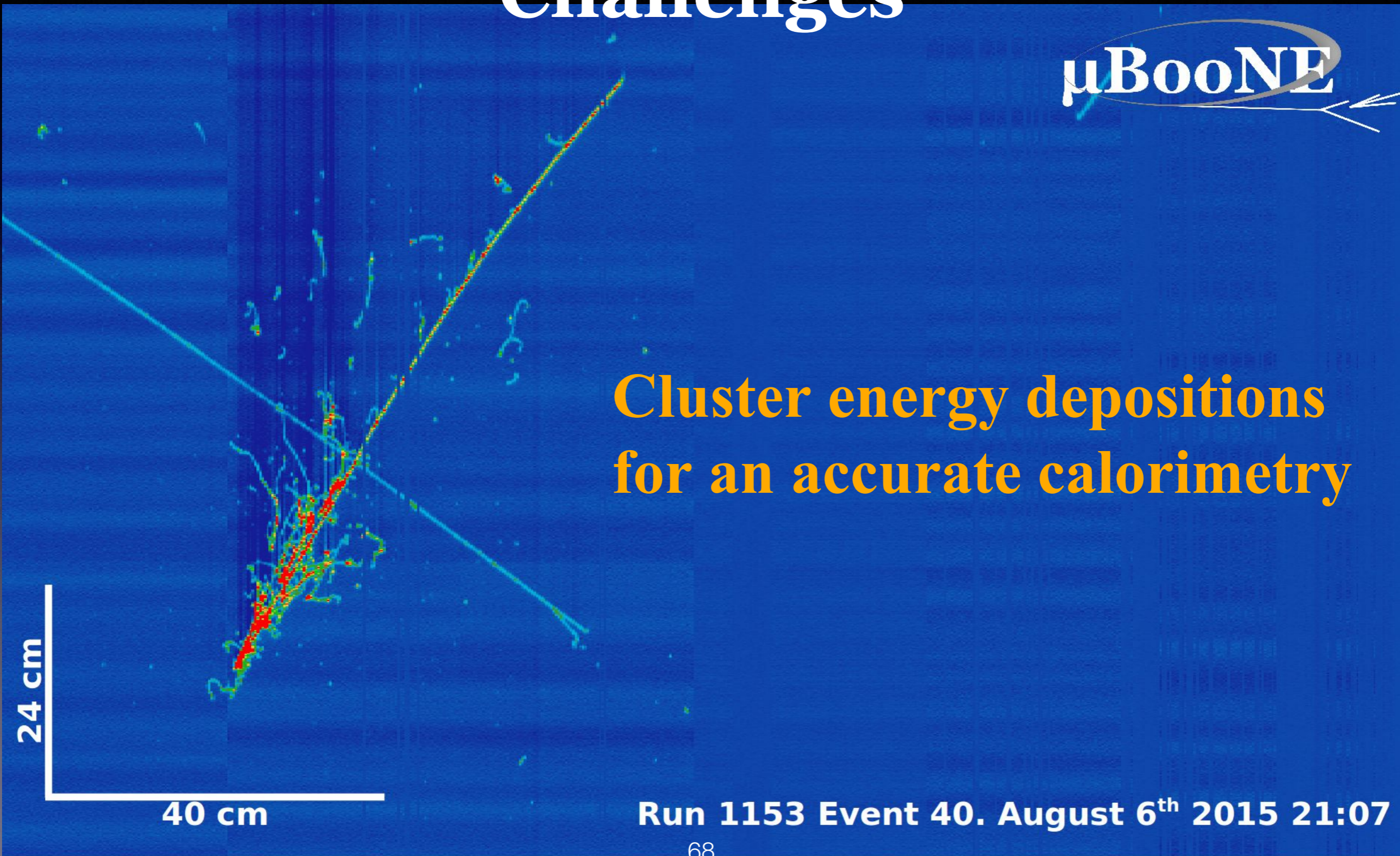
Analysis Challenges



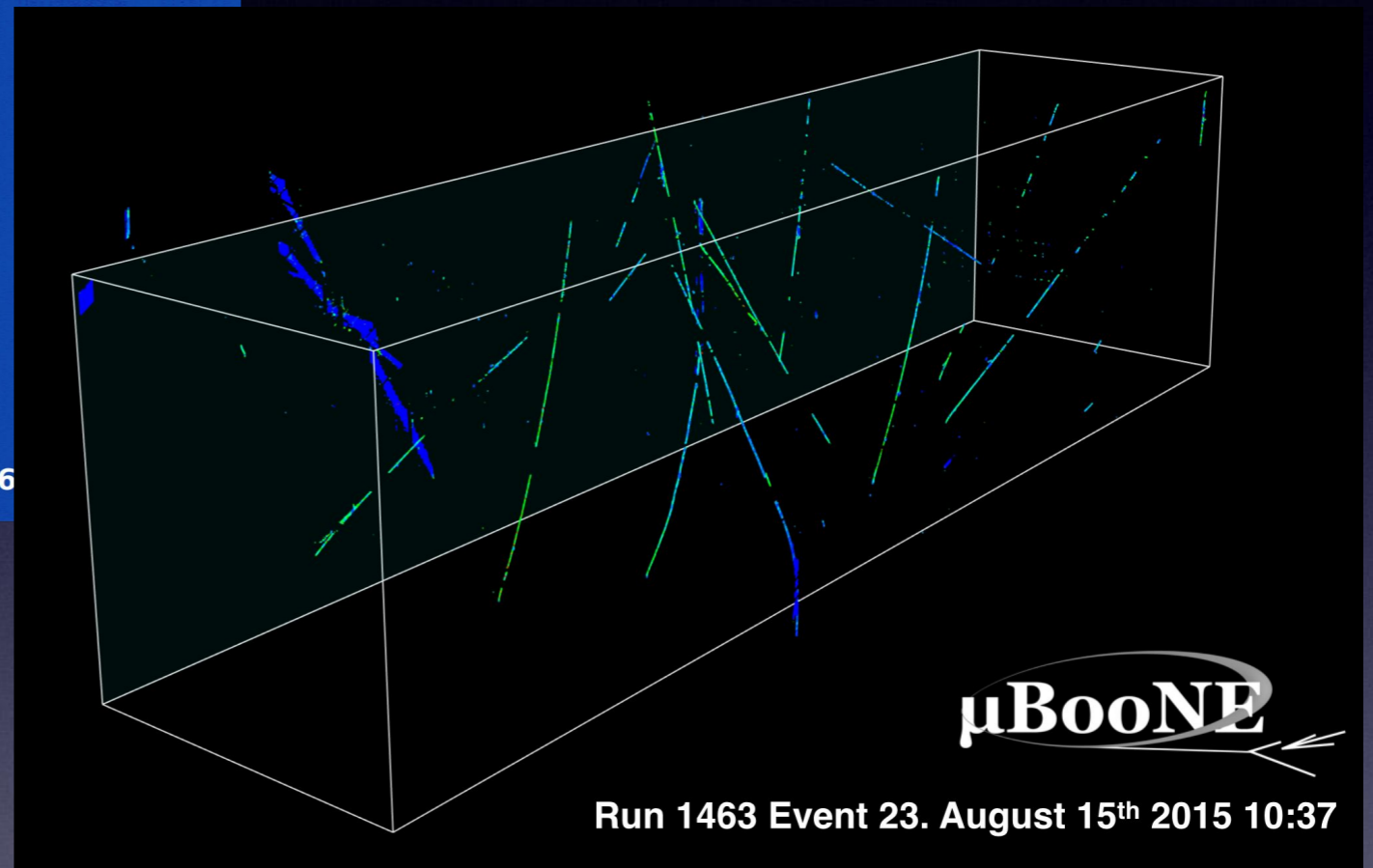
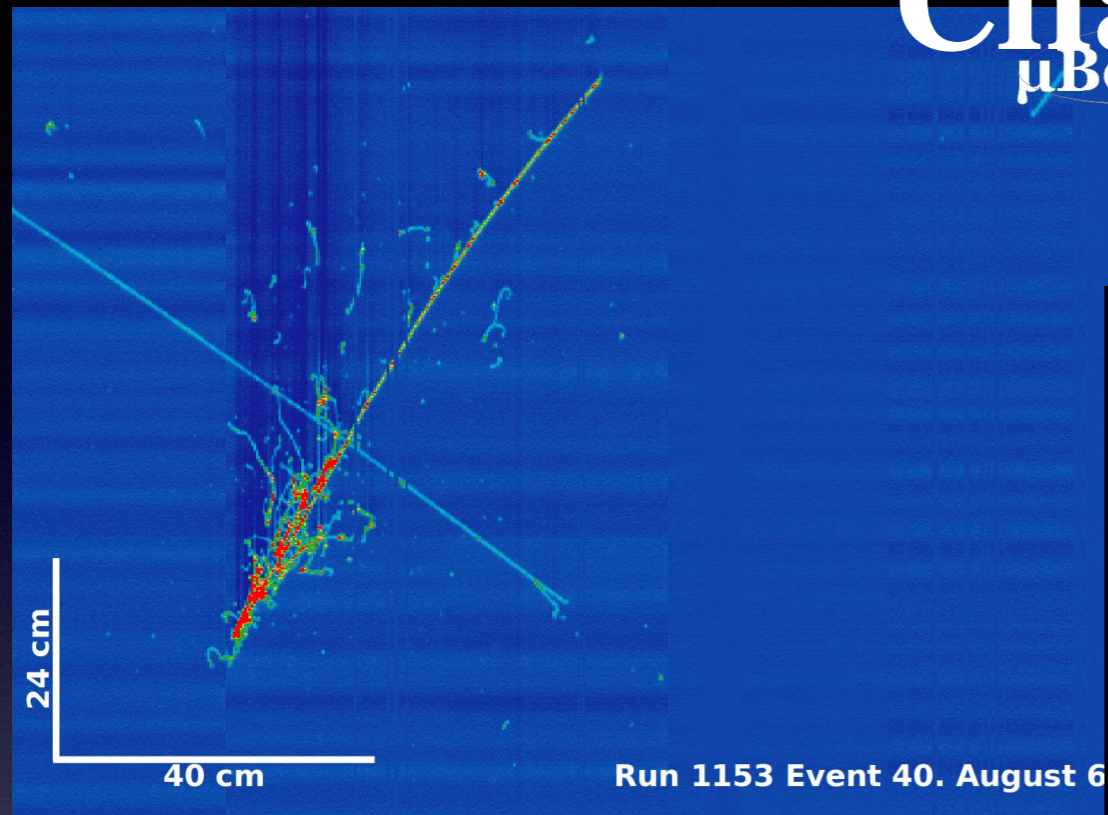
Analysis Challenges

μ BooNE

Cluster energy depositions
for an accurate calorimetry



Analysis Challenges



Deal with optical illusions in 2D projections + 3D pattern recognitions