# High performance computing in nuclear structure studies

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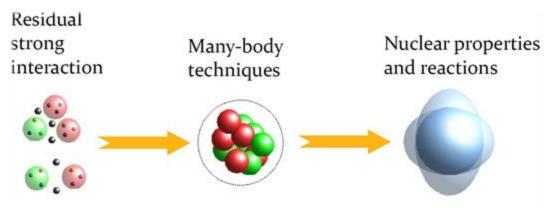
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### Ab-initio theory from nuclear physics perspective

 Ab-initio methods: solution of the nuclear many-body problem starting from *"realistic"* inter -nucleon (NN+NNN) (and? NNNN .... ) force.



- Exact solution for  $A \leq 4$ ? (Fadeev, Fadeev-Yakubowski)
- A>4: development of controlled and improvable *ab-initio* many-body computational methods: No-Core Shell Model (and extensions), Coupled-Cluster method, Green Functions many-body theory...

#### No-Core Shell Model (NCSM) and No-Core Full Configuration(NCFC)

simple, most versatile, access to excited states and transitions, even and odd systems. NCSM review: *Barrett et al., Progress in Particle and Nuclear Physics 69 (2013)*.

### **NN+NNN** interactions

#### Meson exchange potentials

- Bryan Scott (1967-69) •
- Paris (Lacombe et al. 1980) ٠
- Nijmegen (1994) ٠
- Bonn (Bonn A,CD-Bonn) ٠ (*Machleidt et al*.1987-2001)
- Argonne potentials (1995)

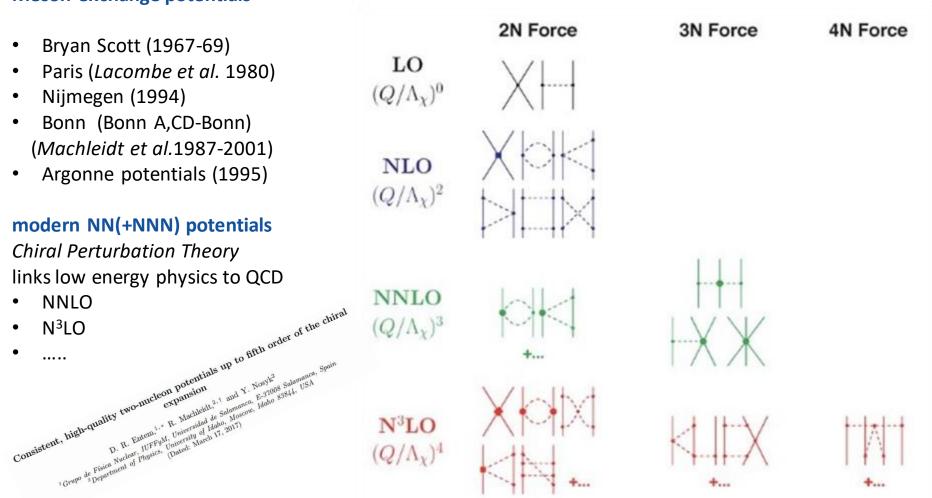
#### modern NN(+NNN) potentials

Chiral Perturbation Theory links low energy physics to QCD

Fisica Nuclear n. Fisica Nuclear

- **NNLO**

1 Grupo



#### Machleidt et al. 2011

short distance physics not resolved, but captured in short range couplings, fitted to NN scattering data, A=3,4 nuclei (recently, heavier systems up to A=16 included)

hierarchy of diagrams, effective 3-, 4- ... body forces

#### **NCSM** essentials

• Solution of many-body Schrodinger equation for bound states

$$H\Psi(\vec{r}_1,\vec{r}_2,\ldots,\vec{r}_A)=E\Psi(\vec{r}_1,\vec{r}_2,\ldots,\vec{r}_A)$$

for A, (or N,Z) point-like nucleons

NCSM (NCFC) assumes intrinsic non-relativistic Hamiltonian with "realistic" NN+NNN interaction

$$H_A = \frac{1}{A} \sum_{i < j} \frac{(\vec{p}_i - \vec{p}_j)^2}{2m} + \sum_{i < j}^A V_{\text{NN},ij} + \sum_{i < j < k}^A V_{\text{NNN},ijk}$$

- All nucleons active (no-core)
- Solution: expansion in 3D spherical harmonic oscillator many-body basis states  $\rightarrow$  Slater determinants constructed from HO s.p. states (with HO length b)

$$\varphi_{nljmm_t}(\vec{r},\sigma,\tau;b) = R_{nl}(r;b)(Y_l(\hat{r})\chi(\sigma))_m^{(j)}\chi(\tau)_{m_t}$$

• Convergence of observables due to the finite basis expansion is the only source of uncertainty

### **NCSM** essentials

• Slater determinants expansion depends on  $\mathbf{\Omega}$  (or eq. HO length *b*)

$$\Psi(\vec{r}_1, \vec{r}_2, ..., \vec{r}_A) = \sum_k c_k \psi_k(\vec{r}_1, \vec{r}_2, ..., \vec{r}_A)$$

determinants  $\Psi_k$  are constructed from single particle states of type

$$\varphi_{nljmm_t}(\vec{r},\sigma,\tau;b) = R_{nl}(r;b)(Y_l(\hat{r})\chi(\sigma))_m^{(j)}\chi(\tau)_{m_t}$$

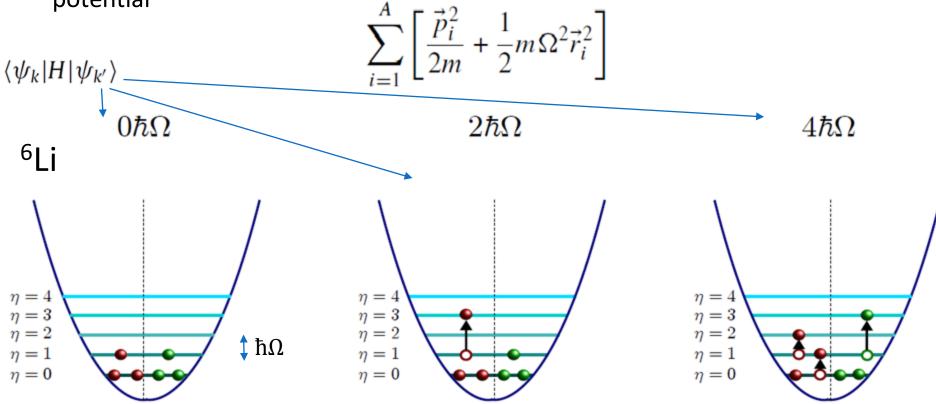
• Many-body problem transformed to symmetric eigenvalue problem

$$\sum_{k'} H_{kk'} c_{k'} = E c_k$$
$$H_{kk'} = \langle \psi_k | H | \psi_{k'} \rangle$$

- Easy task → find (lowest) eigenvalues and eigenvectors of large sparse matrix H
   → HPC (High Performance Computing)
- 1. Contruct **H**<sub>kk</sub>,
- 2. Diagonalize **H**
- 3. Calculate observables
- 4. Check convergence

### **NCSM** essentials

basis construction → A nucleons moving independently in spherical 3D HO potential



• Configuration mixing via "residual" interaction

$$\sum_{i$$

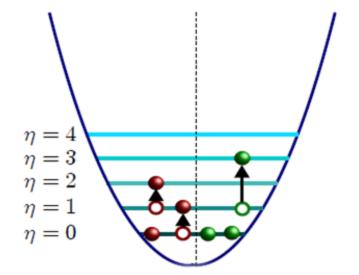
#### **NCSM dimensions**

- dimension of space grows rapidly for heavier systems
- 3D HO with completely filled major shells up to a principal HO quantum number  $\eta$  contains  $n = (\eta+1)(\eta+2)(\eta+3)/3$  single-particle states

 $\eta=0$ n=2 $\eta=1$ n=8 $\eta=2$ n=20 $\eta=8$ n=330 $\eta=20$ n=3542

How to distribute *N*,*Z* nucleons over *n* states? dim= $\binom{n}{N} \times \binom{n}{Z}$ 

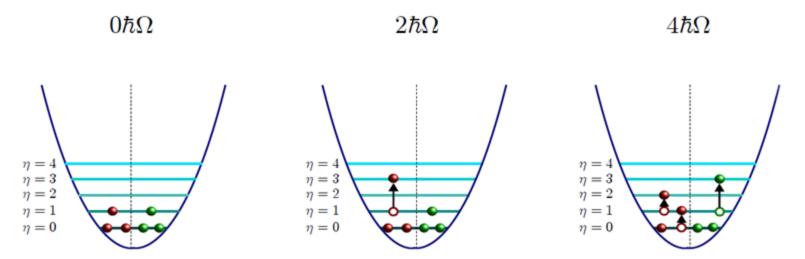
<sup>12</sup> C		<sup>16</sup> O	
η <b>=1</b>	1120	1	
<b>η=6</b>	7. 10 <sup>5</sup> x 7. 10 <sup>5</sup>	1. 10 <sup>12</sup> x 1. 10 <sup>12</sup>	
<b>η=8</b>	3. 10 <sup>11</sup> x 3. 10 <sup>11</sup>	3. 10 <sup>15</sup> x 3. 10 <sup>15</sup>	



### **NCSM dimensions**

• In practice, we truncate basis

In the NCSM states with N=0 HO quanta up to  $N_{max}$  above lowest configuration are included (preserves centre-of-mass factorization)



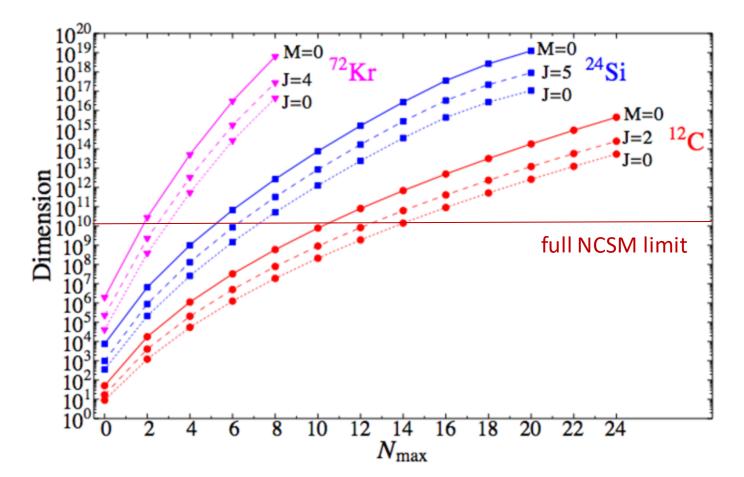
Moreover, one can assume symmetries of the Hamiltonian and construct basis states with required symmetry

 $[\boldsymbol{H}, \boldsymbol{J}_{\boldsymbol{z}}] = 0 \qquad [\boldsymbol{H}, \boldsymbol{J}^2] = 0$ 

projection of total angular momentum  $\rightarrow$  M - scheme magnitude of total angular momentum  $\rightarrow$  J - scheme

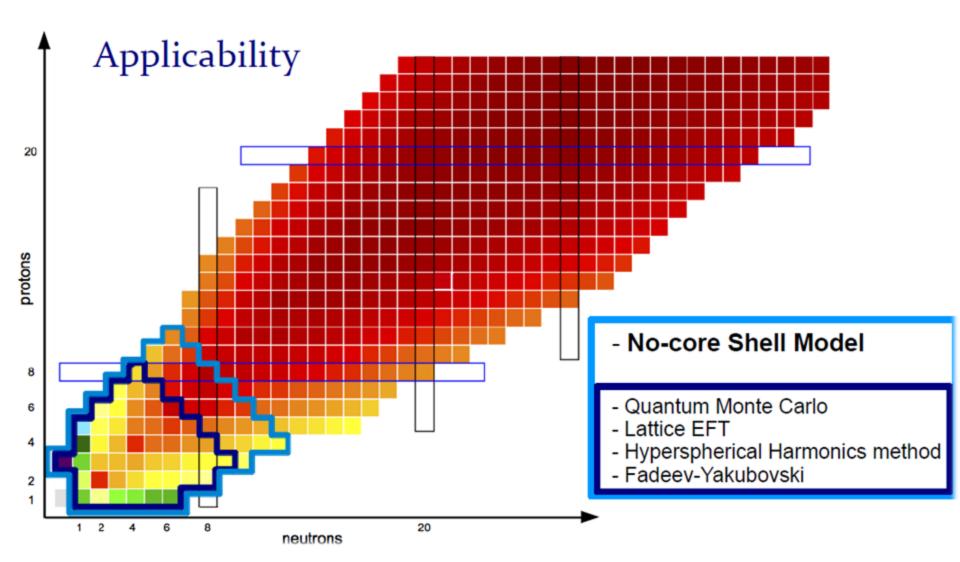
### **NCSM dimensions**

- M-scheme + trivial construction of basis states (Fock space |110010011010....>)
  - + simple calculation of m.e.
  - larger dimension of matrices
- J-scheme + few orders od magnitude reduction
  - involved calculation of m.e., more dense matrix



### Ab-initio theory from nuclear physics perspective

- complete NCSM calculations manageable up to  $N_{max} = 8$  in <sup>16</sup>O
- ground state energy converged for *"soft"* interactions, not for excited states



### How to diagonalize large (huge) matrix? Lanczos algorithm

Lánczos Kornél (1893-1974)

Idea: construct orthogonal basis which turns **H** to tridiagonal form

#### Algorithm :

i=1,  $\beta_1$ =0,  $p_1$  with  $||p_1||$ =1 (pivot vector)

#### do while convergence

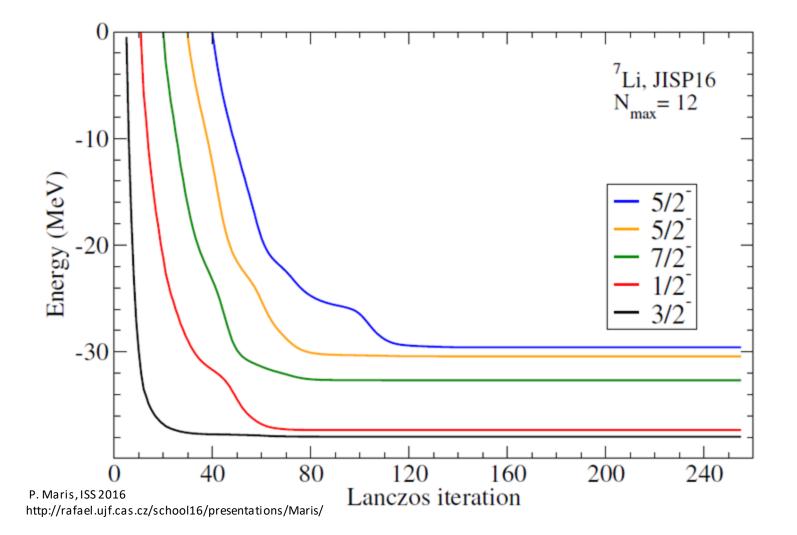
 $p = Hp_i$   $\alpha_i = p_i^T Hp_i$   $k = p - \alpha_i p_i - \beta_i p_{i-1} \text{ (ortog k w.r.t. } p_i)$   $\beta_{i+1} = ||k||$   $p_i = k/\beta_{i+1}$  i = i+1diagonalize tridiagonal matrix T check convergence of lowest eigenvalues end do



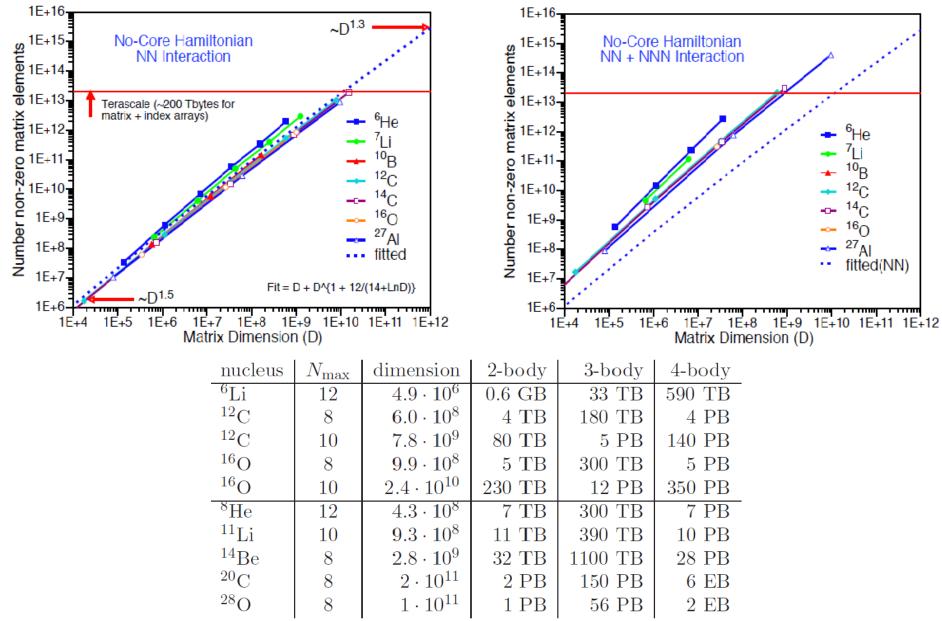
$$\mathsf{T} = \begin{bmatrix} \alpha_1 & \beta_2 & & & \\ \beta_2 & \alpha_2 & \beta_3 & & \\ & \beta_3 & \alpha_3 & \beta_4 & \\ & & \beta_4 & \alpha_4 & \beta_5 & \\ & & & & & & & \\ & & & & & & & & \\ & & & & & & & & & \\ & & & & & & & & & \\ & & & & & & & & & & \\ & & & & & & & & & & \\ & & & & & & & & & & \\ & & & & & & & & & & \\ & & & & & & & & & & \\ & & & & & & & & & & \\ & & & & & & & & & & \\ & & & & & & & & & & \\ & & & & & & & & & & \\ & & & & & & & & & & \\ & & & & & & & & & \\ & & & & & & & & & \\ & & & & & & & & \\ & & & & & & & & & \\ & & & & & & & & & \\ & & & & & & & & & \\ & & & & & & & & \\ & & & & & & & & \\ \end{array}$$

#### Lanczos algorithm - convergence

- matrix dimension 252.10<sup>6</sup> with 4.10<sup>11</sup> n.n.z. matrix elements
- highly optimized parallel code *MFDn*
- 124 nodes using 496 MPI ranks with 6 OpenMP threads/MPI
- total runtime less than 10 minutes (scales up to 100 000 cores!)

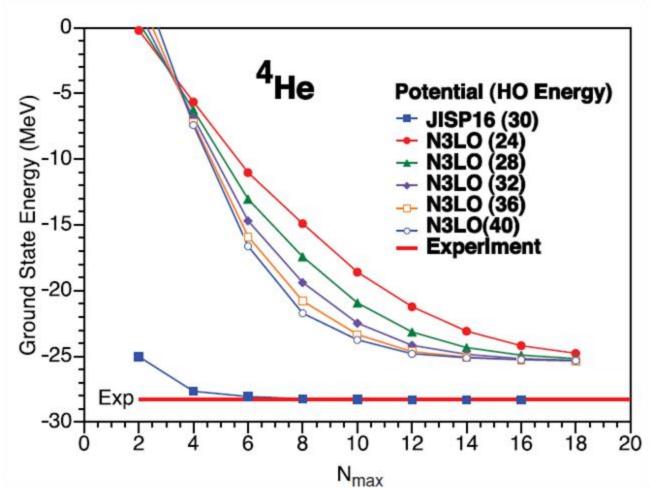


#### No-core shell model – storage



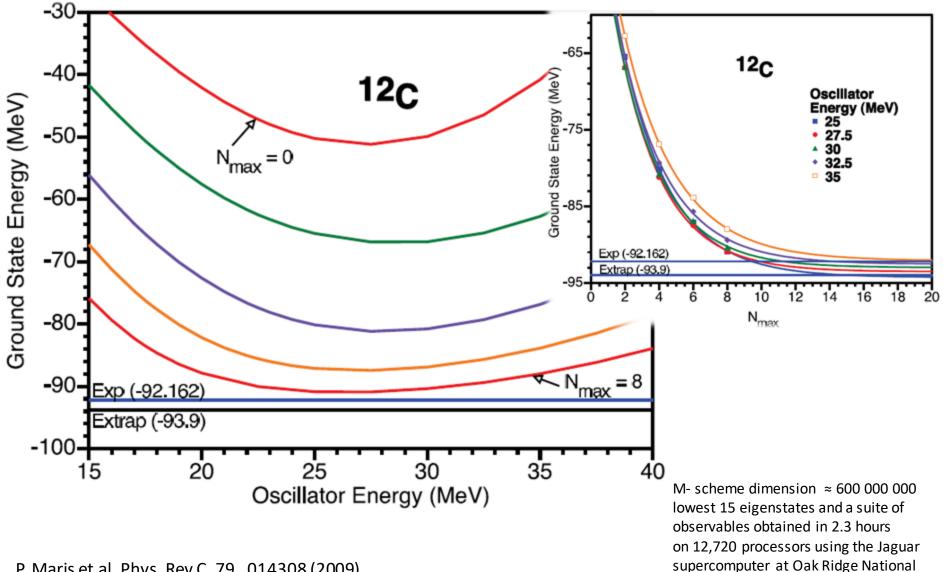
J. Vary et al, Journal of Physics: Conference Series 180 (2009) 012083

- interaction determines convergence rate!
- for "soft" interactions smaller space is needed "hard" interactions → "renormalization" techniques (transformation of "bare" interaction → many body forces)



P. Maris et al, Phys. Rev C. 79, 014308 (2009)

for heavier nuclei fully converged results are not available  $\rightarrow$  extrapolation •



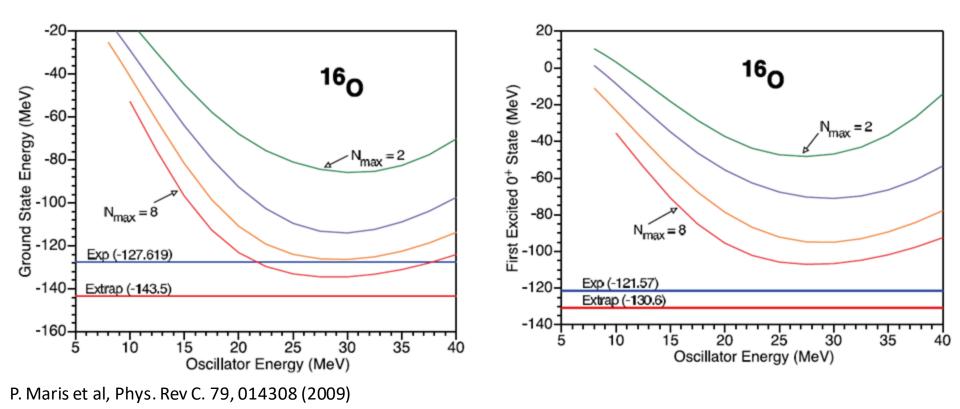
Laboratory (ORNL).

P. Maris et al, Phys. Rev C. 79, 014308 (2009)

- some excited states converge very slowly in HO S.D. basis
- typical examples are cluster states in <sup>16</sup>O and <sup>12</sup>C

```
N<sub>max</sub> =8
matrix dim. 996,878,170
n.n.z 805,811,591,748 (assuming symmetric matrix)
storage of one vector 4.0 GB
```

lowest 8 eigenstates and a suite of observables obtained in 4.5 hours on 12,090 processors @Franklin supercomputer at the National Energy Research Supercomputer Center (NERSC).

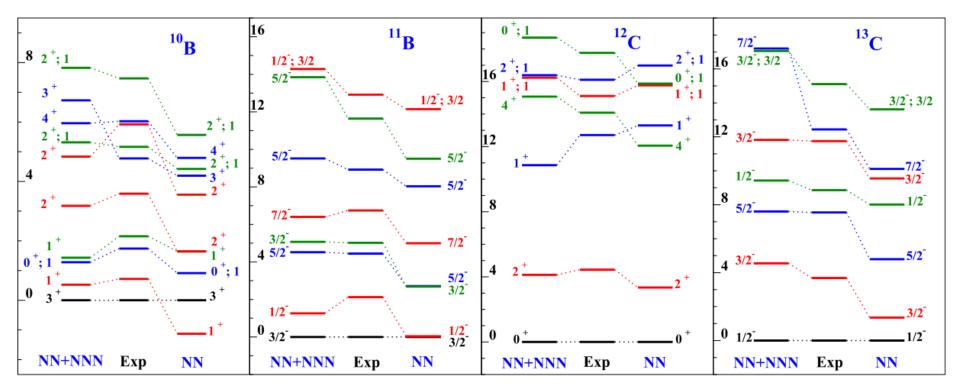


Matrix 6.5 TB

• Nevertheless NCSM can test different NN (+NNN) potentials.

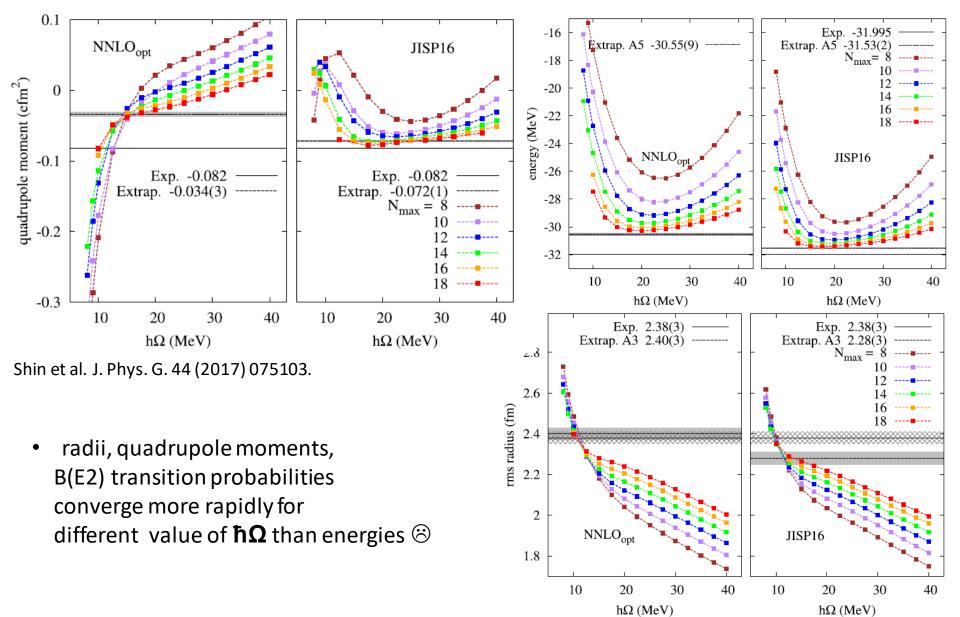
Do three-body forces enter the description? How large are three-body forces?

\* NCSM calculations with chiral NN and NNN potential  $\rightarrow$  role of NNN forces in inversion of 3+ and 1+ in  $^{10}\text{B}$ 



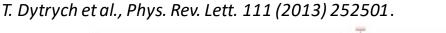
P. Navrátil et al., Phys. Rev. Lett. 99 (2007) 042501

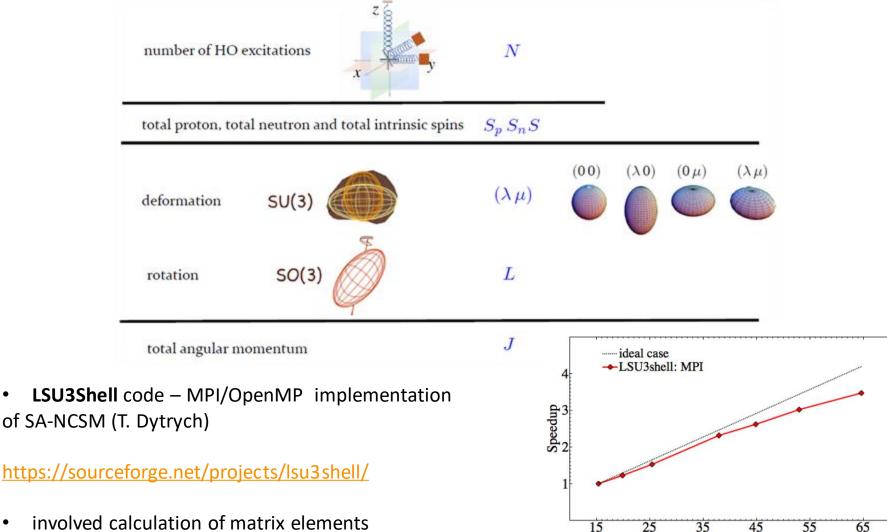
But, what about other observables? Well, it depends .....



#### Symmetry-adapted NCSM (SA-NCSM)-combines algebraic techniques with the NCSM

• multishell extension of Elliot SU(3) model





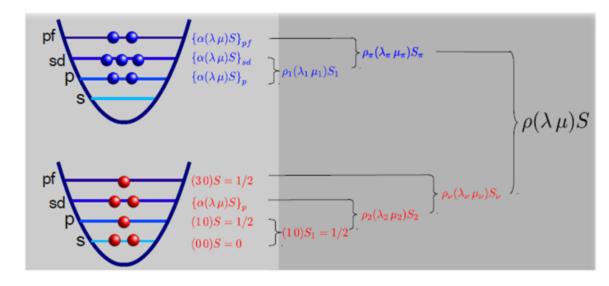
Number of MPI processes [10<sup>3</sup>]

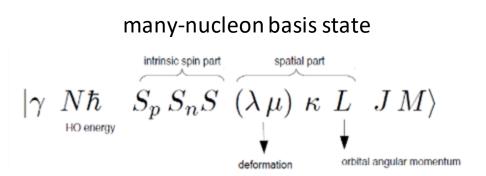
 involved calculation of matrix elements 95% of runtime!

• SU(3) clasification scheme for spatial part  $\rightarrow$  LS coupling  $\rightarrow$  J=L+S

Step 1: SU(3) coupling of succesive shells

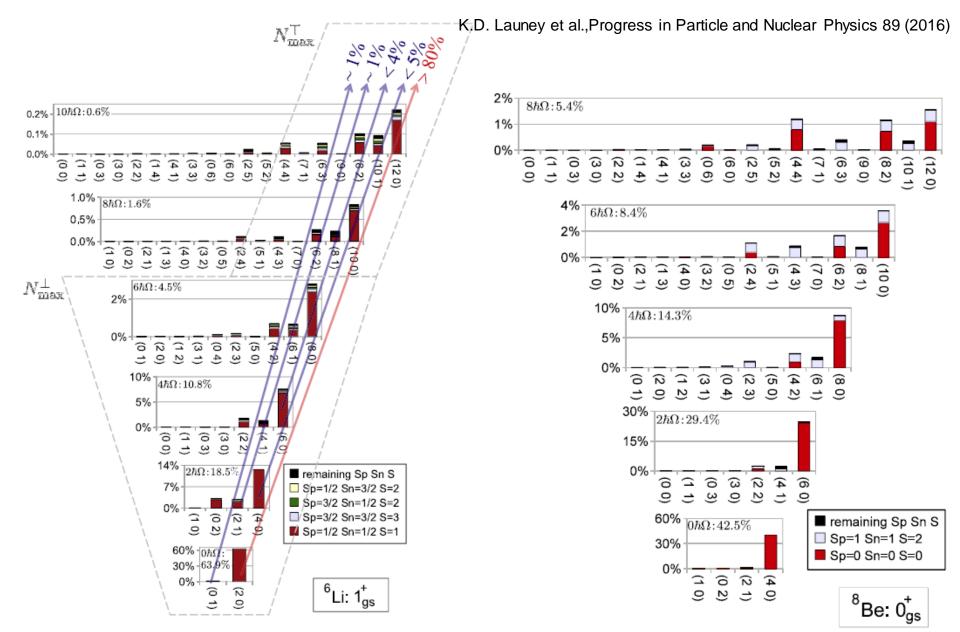
Step 2: SU(3) coupling of protons and neutrons





Why SU(3) coupling scheme? truncation of the model space

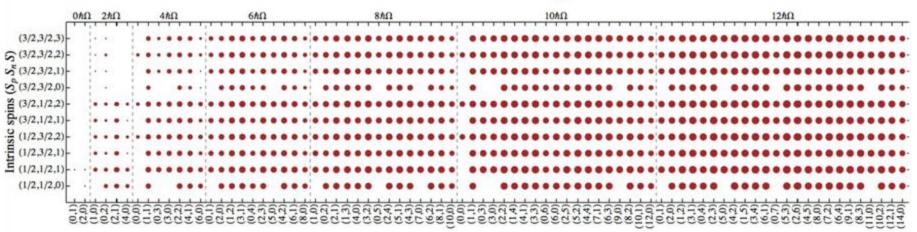
• decomposition of NCSM model space  $\rightarrow$  dominant components in the w.f.  $\rightarrow$  truncation



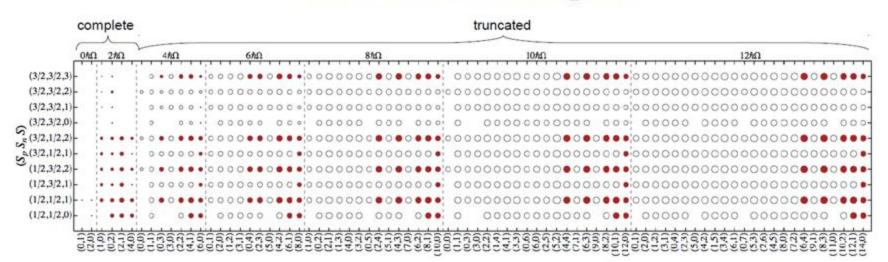
<sup>6</sup>Li model space

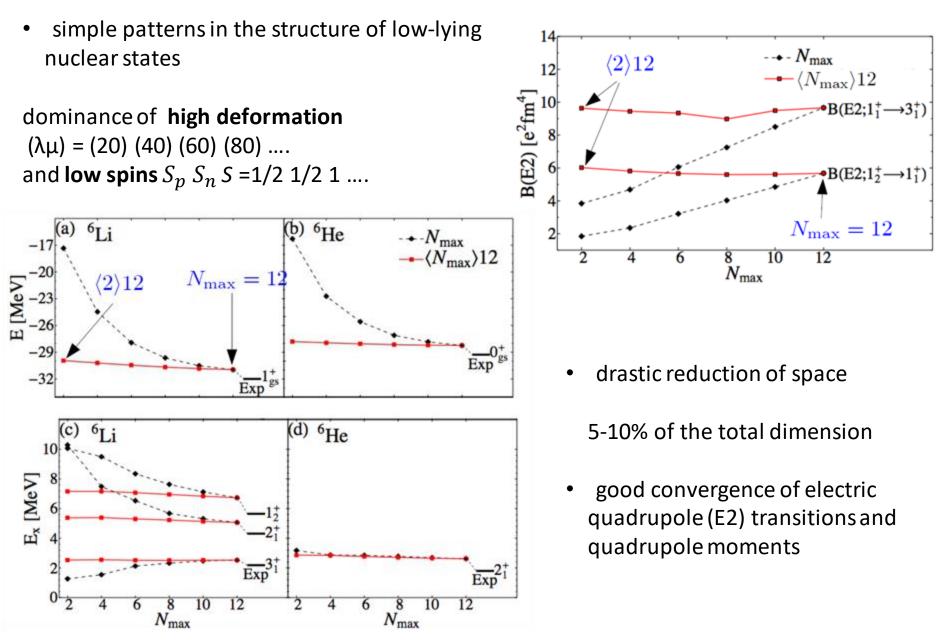
K.D. Launey et al., Progress in Particle and Nuclear Physics 89 (2016)

#### NCSM model space



#### SA-NCSM model space





K.D. Launey et al., Progress in Particle and Nuclear Physics 89 (2016)

#### **Computer resources**

## **BLUE WATERS**

 Fastest supercomputer at university campus total performance ≈ 1 Pflop/s (on a sustained basis) total peak performance 13.24 Pflop/s total system memory 1.634 PB



- 22,640 Cray XE6 nodes each 64 GB RAM, 16 cores
- 4,228 Cray XK7 nodes 32 GB RAM, 16 cores +2688 CUDA
   (≈ 400 000 cores)
- 26 PB storage quick access
- 380 PB long term data storage

Computing time: US National Science Foundation (NFS) grant.

Collaborative Research: Advancing first-principle symmetry-guided nuclear modeling for studies of nucleosynthesis and fundamental symmetries in nature (PI J. Draayer, LSU)

3,430,000 node hrs in 08/2017-08/2018

#### SA-NCSM@BlueWaters

- large calculation: <sup>20</sup>Ne beyond reach of NCSM
- ≈ 22 000 nodes ( ≈ 700 000 cores)
   runtime ≈ 1 hour for calculation of Hamiltonian + diagonalization

- example:  $J^{\pi} = 2^+$
- dim.  $\approx 80 \cdot 10^6$

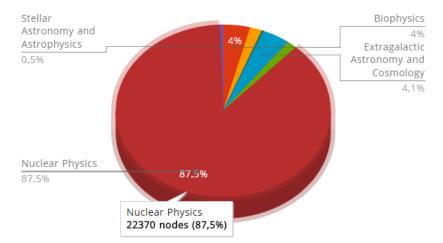
n.n.z. m.e. ≈ 28. 10<sup>12</sup>

matrix storage:

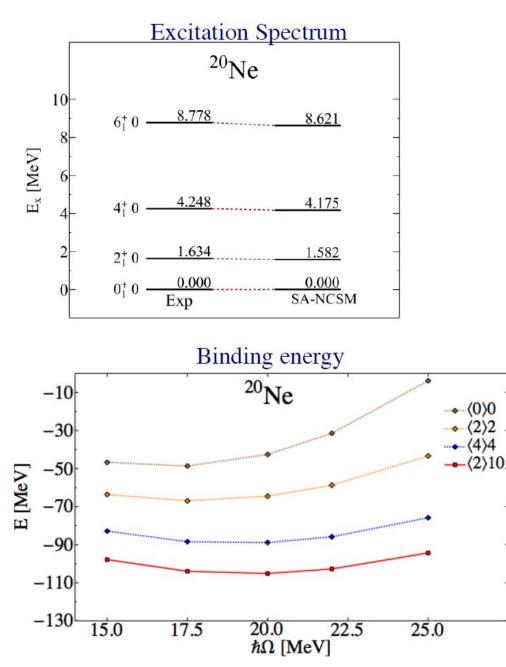
139 TB in VBC 211 TB in CSR

Collaborative Research: Advancing first-principle symmetry-guided nuclear modeling for studies of nucleosynthesis and fundamental symmetries in nature (Jobs:1) PI: Jerry Draayer, Louisiana State University	XE	715712	202,288.04
GPU-enabled General Relativistic Simulations of Jetted Tidal Disruptions of Stars by Supermassive Black Holes (Jobs:4) PI: Alexander Tchekhovskoy, University of California, Berkeley	XK	16880	2,980.98
Unveiling the functions of the HIV-1 and hepatitis B virus capsids through the computational microscope (Jobs:1) PI: Juan Perilla, University of Illinois at Urbana-Champaign	XK	7200	1,269.00
Probing New Physics in Galaxy Formation at Ultra-High Resolution (Jobs:2) PI: Philip Hopkins, California Institute of Technology	XE	2560	446.51
More Power to the Many: Scalable Ensemble-based Simulations and Data Analysis (Jobs:1) PI: Shantenu Jha, Rutgers, the State University of New Jersey	XE	1920	160.27

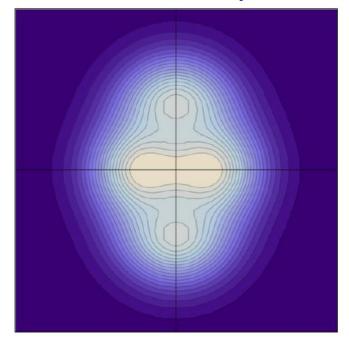
CURRENT RUNNING JOBS BY SCIENCE AREA



#### SA-NCSM@BlueWaters



#### Nuclear density



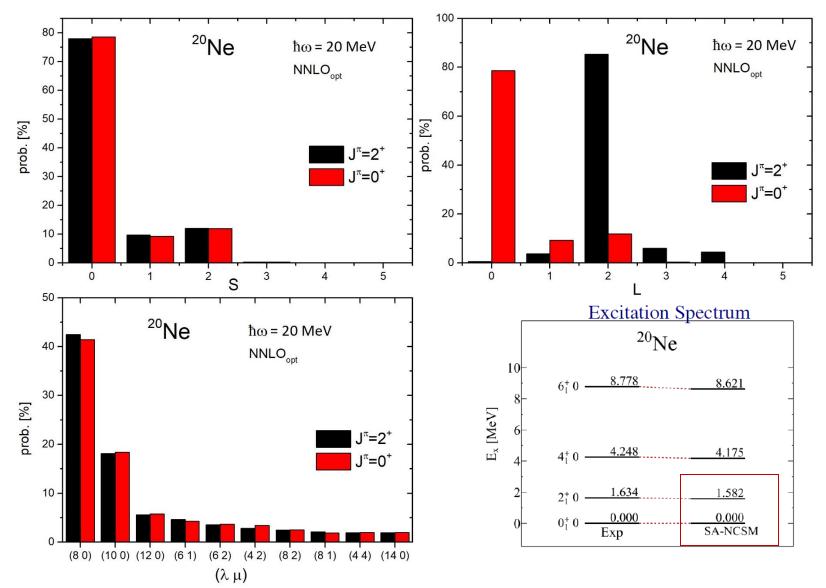
Complete space:  $4 \times 10^{12}$ Symmetry-adapted space:  $1 \times 10^7$ 

T. Dytrych

#### SA-NCSM@BlueWaters

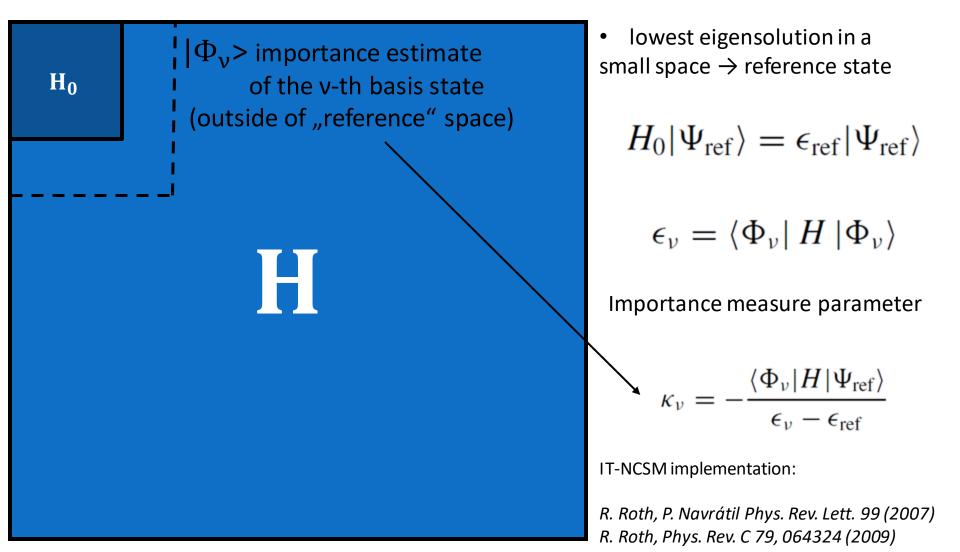


 LS decomposition of w.f. → interpretation of state as rotation of deformed shape nuclear spin J=L+S



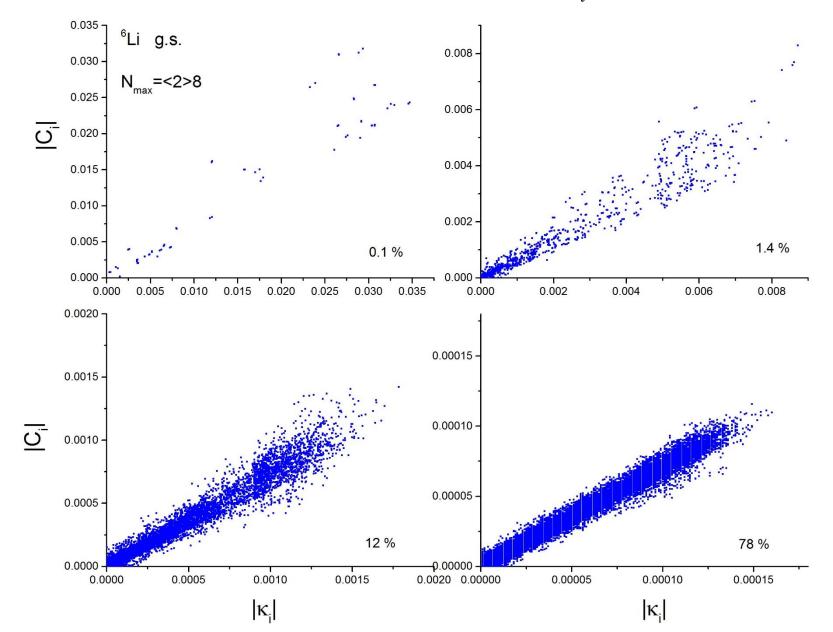
### SA-NCSM with importance truncation

- additional reduction of model space for heavier systems necessary
- quantitative justification of truncation → Importance Truncation (IT) based on 1<sup>st</sup> order many-body perturbation theory



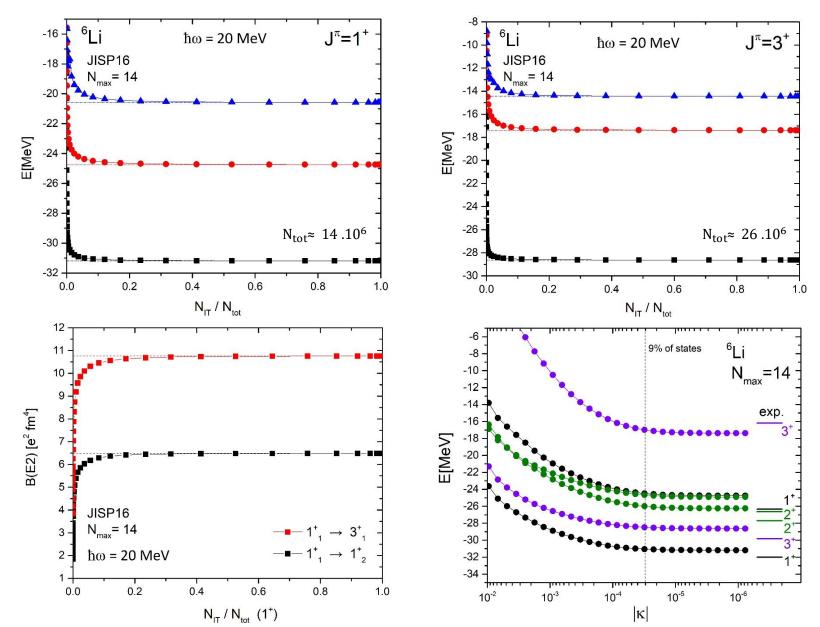
#### SA-NCSM with importance truncation

• correlation between importance measure  $|\kappa_i|$  and amplitudes  $|C_i|$  for i-th component



#### SA-NCSM with importance truncation

• fast convergence of observables → reduction of basis dimensions



Thankyou!