

# Combined analysis status

M. Tenti

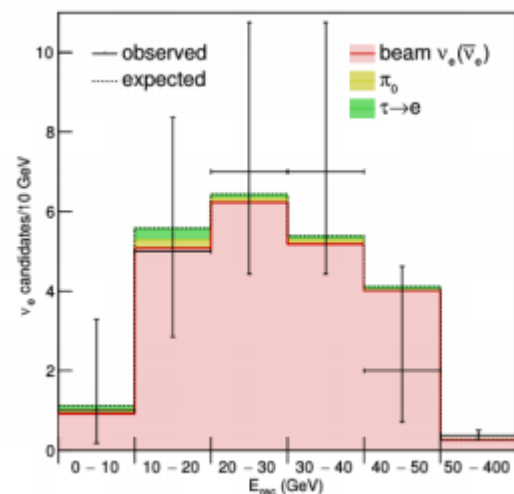
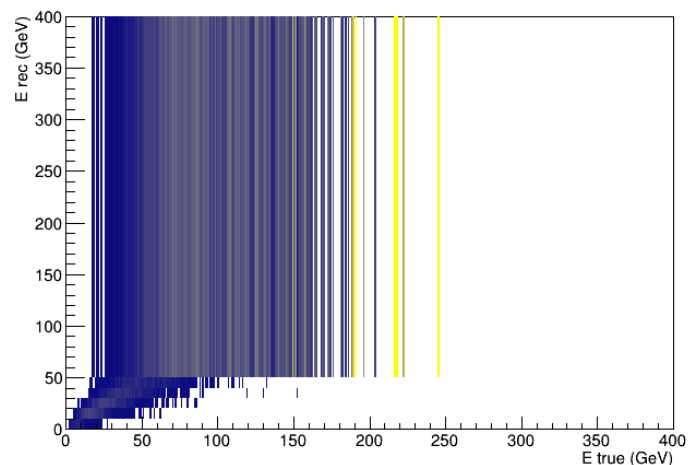
05/31/2018

Villa Orlandi, Anacapri (NA)

# $\nu_e$ channel (1/2)

- Selection:  $\nu_e$  candidate
- Observable:
  - reconstructed energy
  - binnig:  
0,10,20,30,40,50,400 (GeV)
- Energy sampling:
  - 400 points [0 - 400 GeV]
- Observation ( $n_i$ ):
  - 1,5,7,7,2,13
- Normalization:
  - 30.7 w/o  $\nu$  oscillations

Smearing Matrix



# $\nu_e$ channel (2/2)

- Signal ( $s_i$ ):
  - $\nu_\mu \rightarrow \nu_e, \bar{\nu}_\mu \rightarrow \bar{\nu}_e$
  - $\nu_e \rightarrow \nu_e, \bar{\nu}_e \rightarrow \bar{\nu}_e$
- Background (not osc) ( $b_i$ ):
  - $NC, \tau \rightarrow e$
  - 0.25, 0.66, 0.19, 0.11, 0.07, 0.1
- Syst. uncertainties ( $\sigma_j$ ):
  - 20%  $E_\nu < 10$  GeV; 10%  $E_\nu > 10$  GeV
  - 45% on not. osc. background

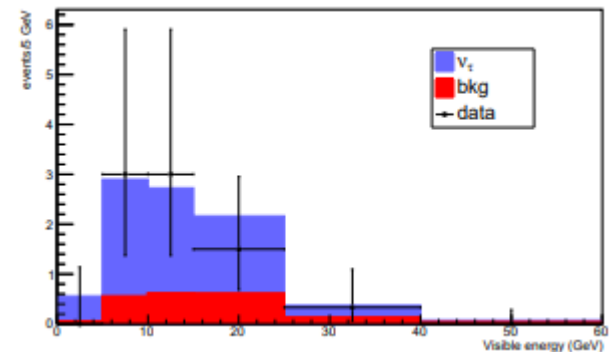
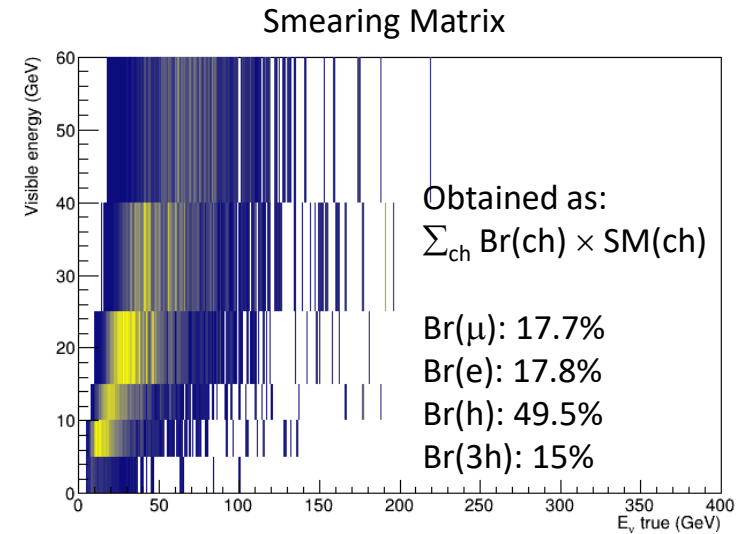
Expected background	Expected $\nu_e$
$1.2 \pm 0.54$	$30.7 \pm 3.1$

- Likelihood:
  - $s_i$  = signal
  - $b_i$  = not osc. bkg.
  - $\mu_i$  = expectation
    - $\mu_i = s_i(1 + k_j) + b_i(1 + k_3)$
    - $j = 1$ , if  $i = 1$
    - $j = 2$ , otherwise
  - $n_i$  = observation
  - $\sigma_j$  = syst. uncertainties

$$-2 \ln L_e = -2 \sum_{i=1}^N (n_i \ln \mu_i - N \mu_i) + \sum_{j=1}^3 \frac{k_j^2}{\sigma_j^2}$$

# $\nu_\tau$ channel (1/2)

- Selection:  $\nu_\tau$  candidate [looser selection]
- Observable:
  - visible energy
  - binnig: 0,5,10,15,25,40,60 (GeV)
- Energy sampling:
  - 400 points [0 - 400 GeV]
- Observation ( $n_i$ ):
  - 0,3,3,3,1,0
- Normalization:
  - 6.8
    - 2 flavours
    - maximal mixing
    - $\Delta m^2 = 2.5 \times 10^{-3} \text{ eV}^2$



# $\nu_\tau$ channel (2/2)

- Signal ( $s_i$ ):

- $\nu_\mu \rightarrow \nu_\tau$

- Background (not osc) ( $b_i$ ):

- *charm, LAS, had. reint.*

- 0.0633, 0.552, 0.622, 0.608, 0.131, 0.033

- Syst. uncertainties ( $\sigma_j$ ):

- 11% on signal

- 20% on not. osc. background

- Likelihood:

- $s_i$  = signal

- $b_i$  = not osc. bkg.

- $\mu_i$  = expectation

- $\mu_i = s_i(1 + k_1) + b_i(1 + k_2)$

- $n_i$  = observation

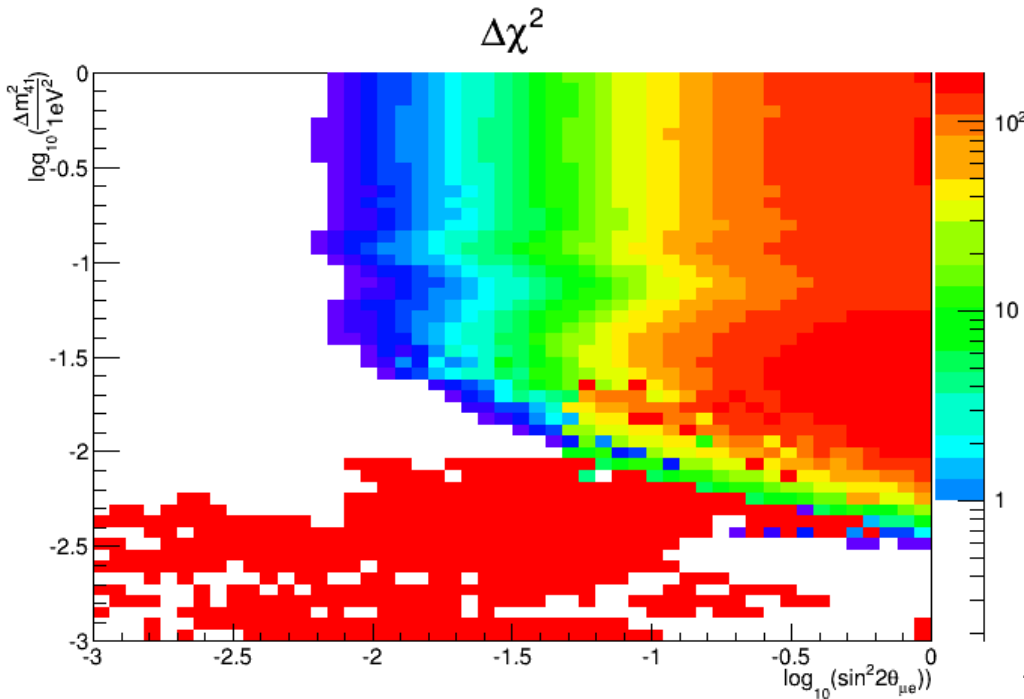
- $\sigma_j$  = syst. uncertainties

$$-2 \ln L_\tau = -2 \sum_{i=1}^N (n_i \ln \mu_i - N\mu_i) + \sum_{j=1}^2 \frac{k_j^2}{\sigma_j^2}$$

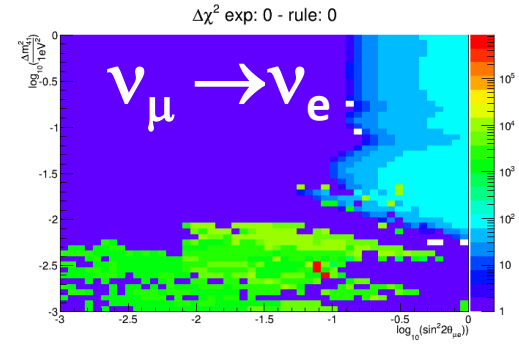
Expected background	Expected $\nu_\tau$
$2.0 \pm 0.4$	$6.8 \pm 0.75$

# Bug

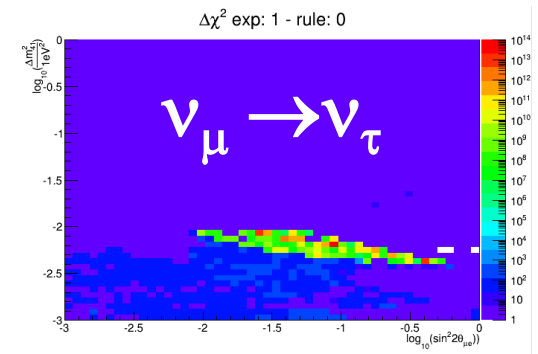
These are the results of the last run



$$L = L_e + L_\tau + \text{Prior}$$



$$-2 \ln L_e = -2 \sum_{i=1}^N (n_i \ln \mu_i - N \mu_i) + \sum_{j=1}^3 \frac{k_j^2}{\sigma_j^2}$$



$$-2 \ln L_\tau = -2 \sum_{i=1}^N (n_i \ln \mu_i - N \mu_i) + \sum_{j=1}^2 \frac{k_j^2}{\sigma_j^2}$$

Bug: penalty term for systematics was not present

# Statistical analysis

- Priors

- Gaussian on  $\Delta m_{31}^2$   
 $(2.56 \pm 0.05) \times 10^{-3} \text{ eV}^2$   
[\[http://pdg.lbl.gov/2017/reviews/rpp2017-rev-neutrino-mixing.pdf\]](http://pdg.lbl.gov/2017/reviews/rpp2017-rev-neutrino-mixing.pdf)

- Likelihood:

- $L = L_e + L_\tau + \text{Prior}$

- Test statistic:  
 profile likelihood ratio

- Wilks' theorem

- To extract limits and confidence/exclusion regions

Parameter	best-fit
$\Delta m_{21}^2 [10^{-5} \text{ eV}^2]$	7.37
$\Delta m_{31(23)}^2 [10^{-3} \text{ eV}^2]$	2.56 (2.54)
$\sin^2 \theta_{12}$	0.297
$\sin^2 \theta_{23}, \Delta m_{31(32)}^2 > 0$	0.425
$\sin^2 \theta_{23}, \Delta m_{32(31)}^2 < 0$	0.589
$\sin^2 \theta_{13}, \Delta m_{31(32)}^2 > 0$	0.0215
$\sin^2 \theta_{13}, \Delta m_{32(31)}^2 < 0$	0.0216
$\delta/\pi$	1.38 (1.31)

# Oscillation model & Mixing matrix representation

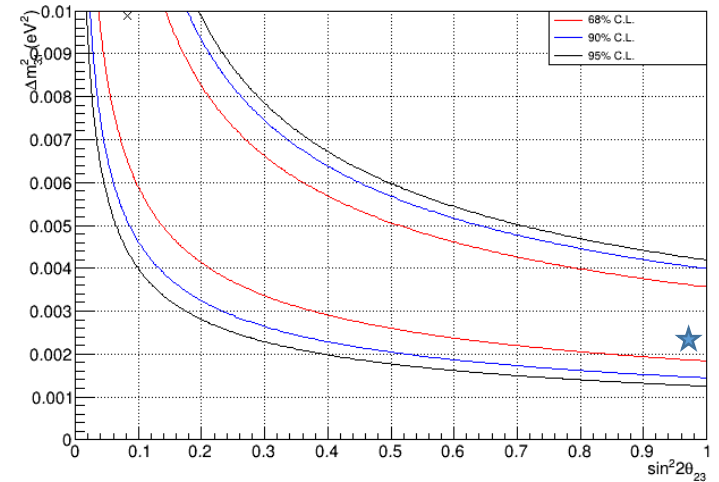
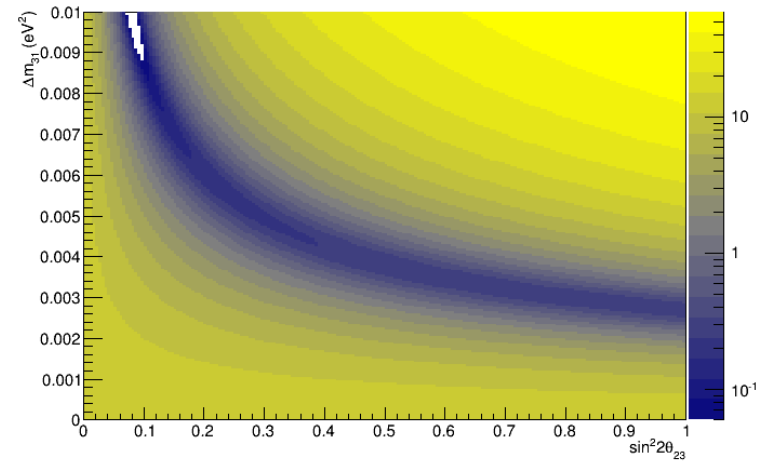
- 3 flavours model
  - $U = R(\theta_{23}) R(\theta_{13}, \delta_{CP}) R(\theta_{12})$
- Matter effects (constant density)
- NH



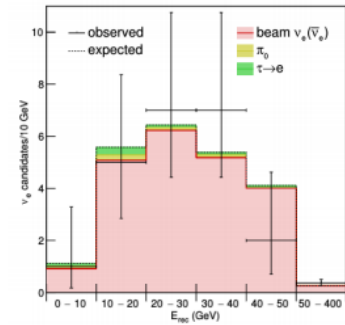
$\Delta m^2_{31}$  VS  $\theta_{23}$

# POI and nuisance params

- Parameters of interest:
  - $\theta_{23}, \Delta m^2_{31}$  (prior removed)
- Nuisance parameters:
  - $\theta_{12}, \theta_{13}, \delta_{CP}$
- Constants
  - $\Delta m^2_{21} = 7.37 \times 10^{-5} \text{ eV}^2$

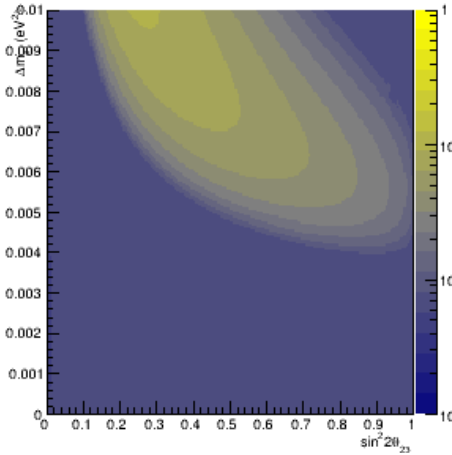


# Bin contribution ( $\nu_e$ channel)

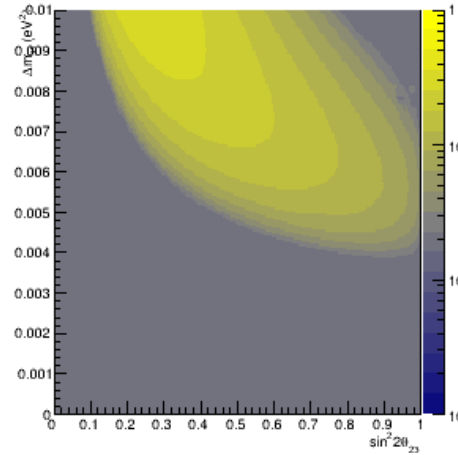


Z axis range: [0 - 1]

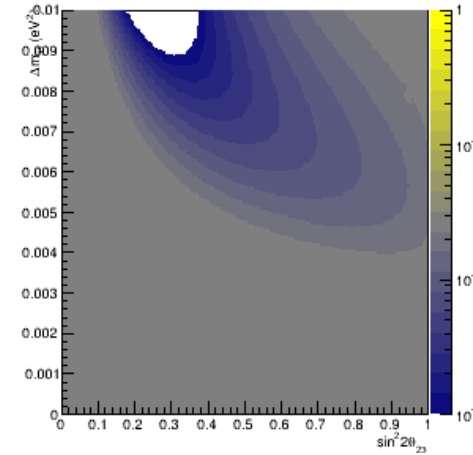
Bin: 1



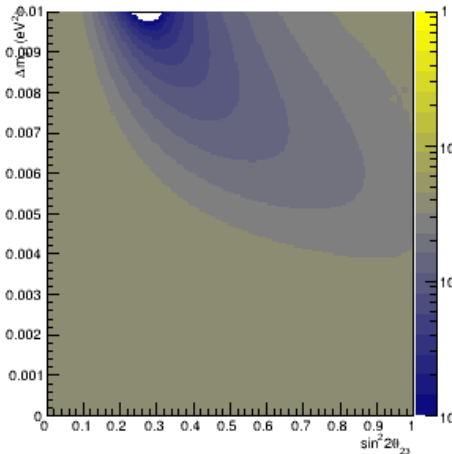
Bin: 2



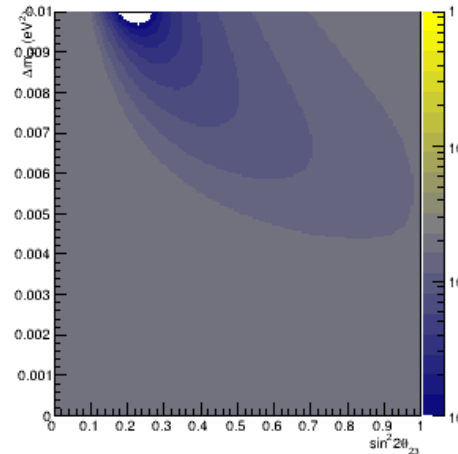
Bin: 3



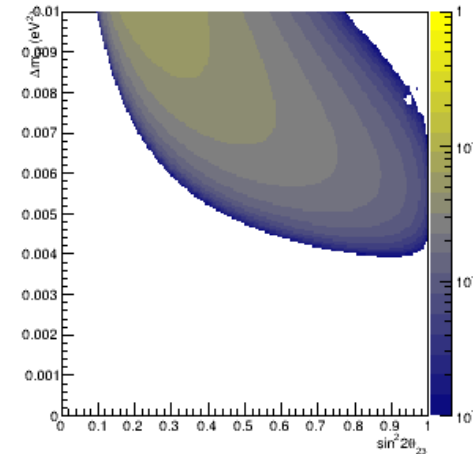
Bin: 4



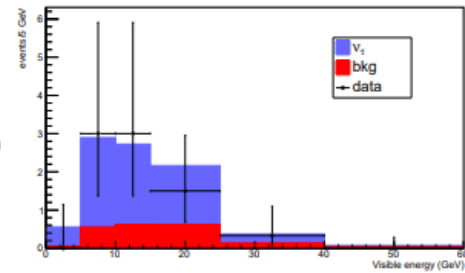
Bin: 5



Bin: 6

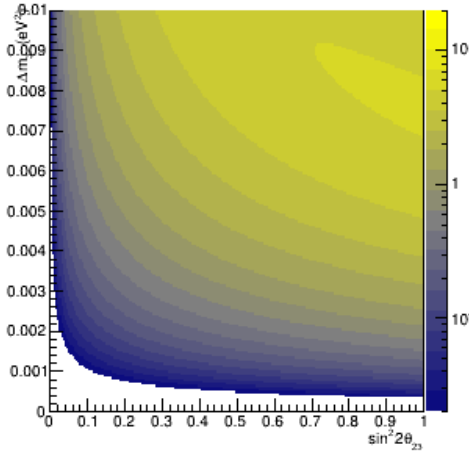


# Bin contribution ( $\nu_\tau$ channel)

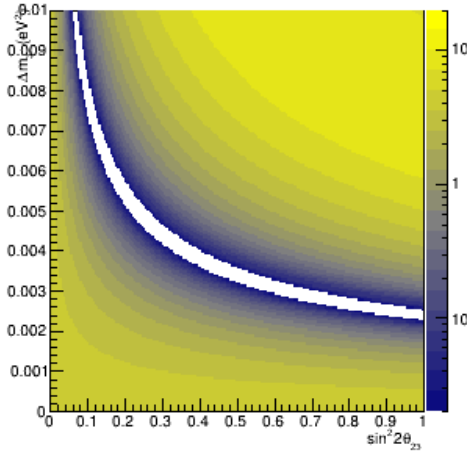


Z axis range: [0 - 20]

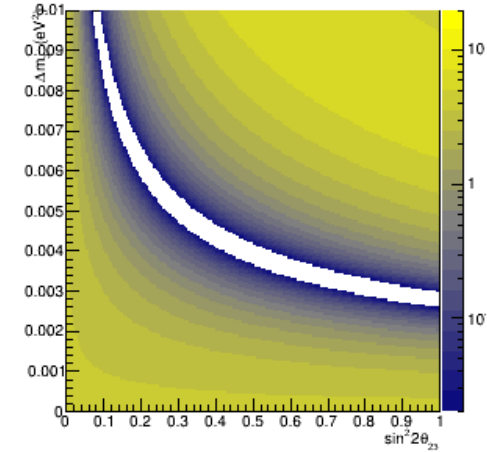
Bin: 1



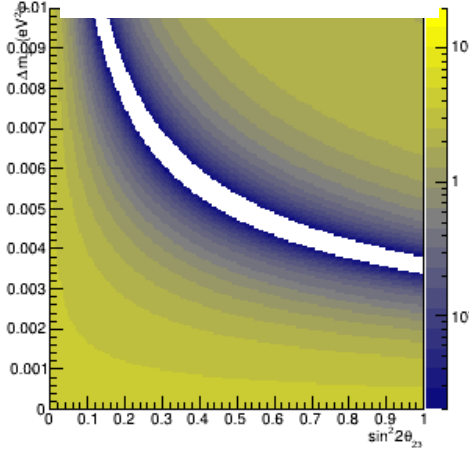
Bin: 2



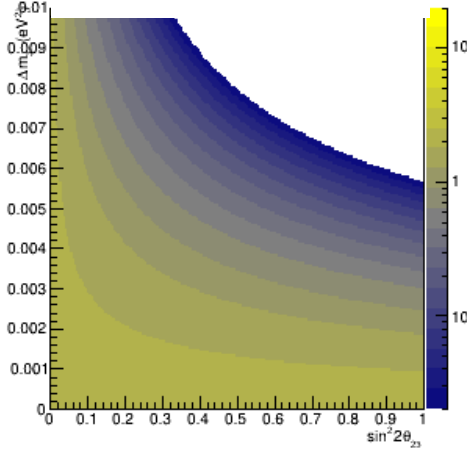
Bin: 3



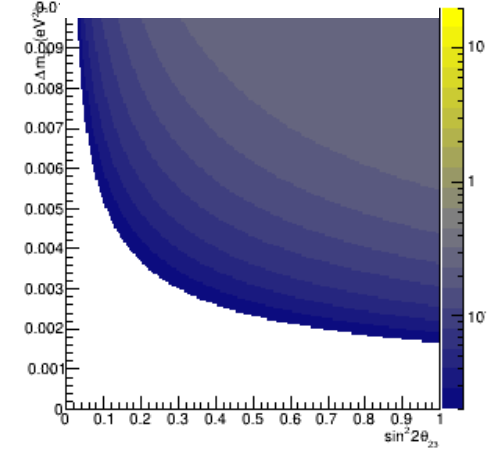
Bin: 4



Bin: 5



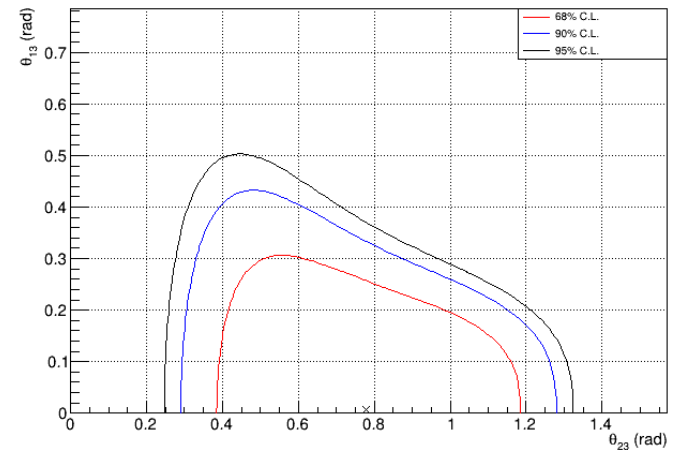
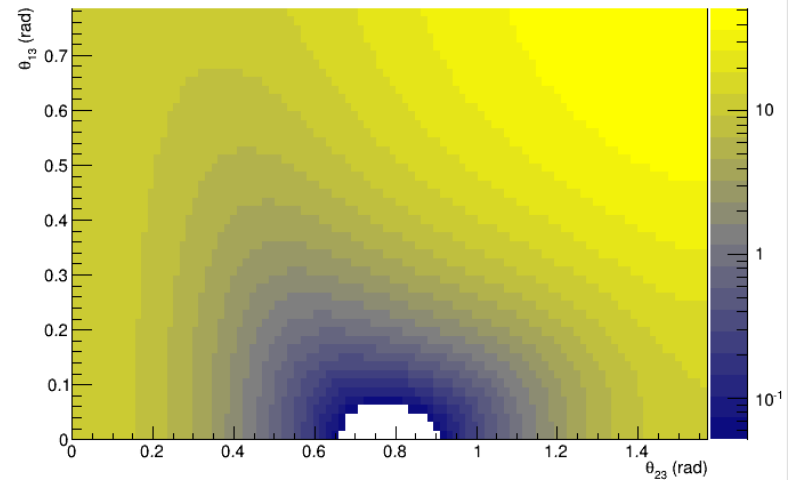
Bin: 6



$\theta_{13}$  VS  $\theta_{23}$

# POI and nuisance params

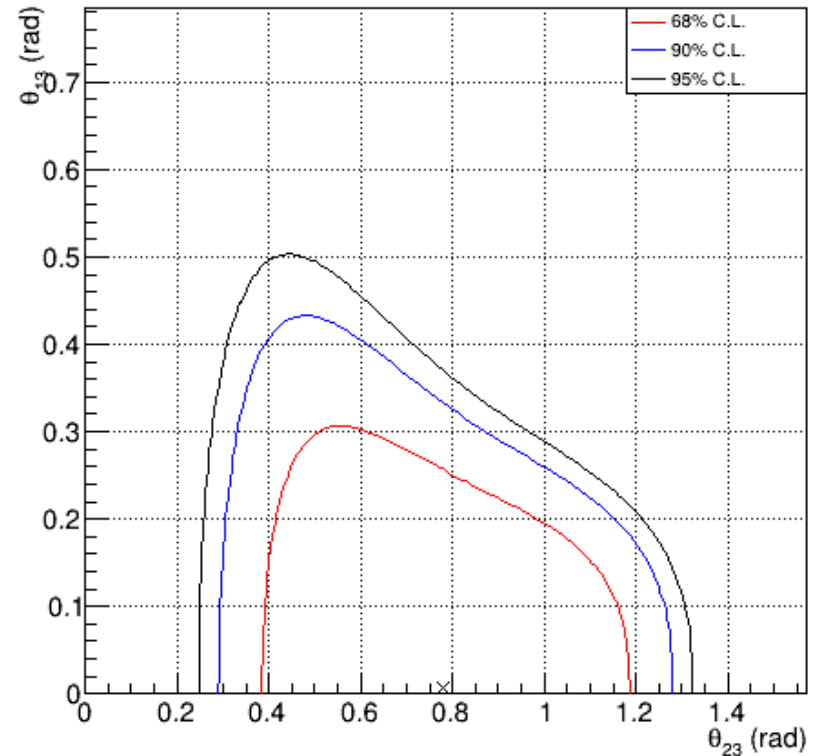
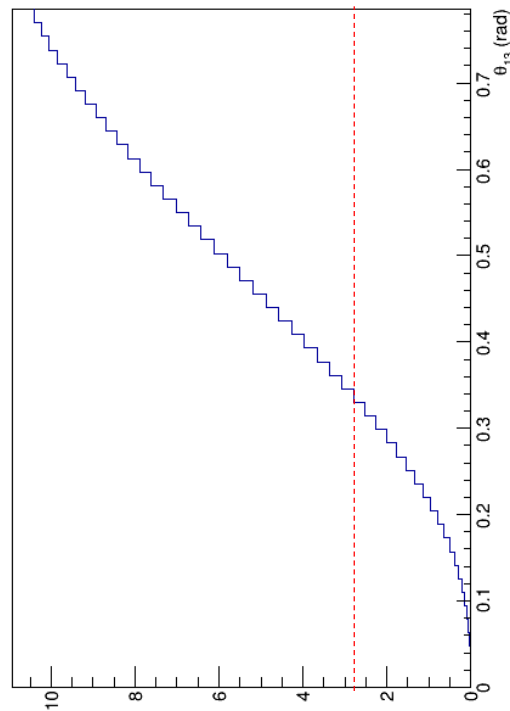
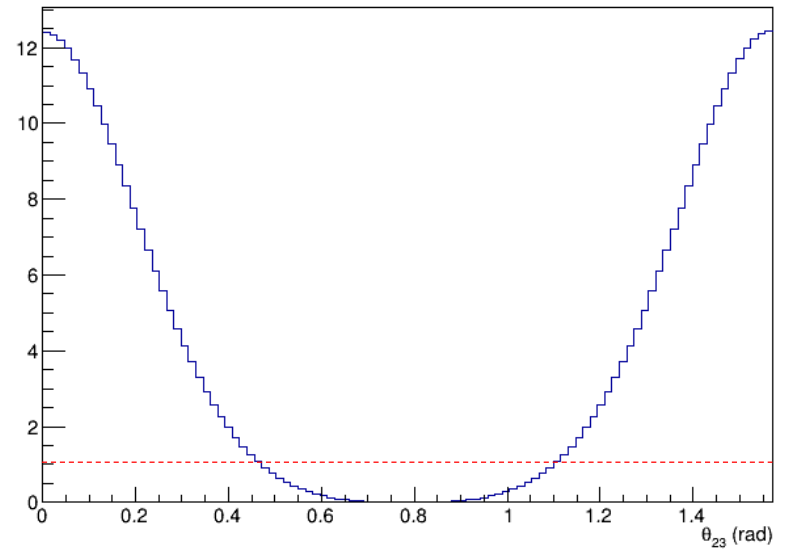
- Parameters of interest:
  - $\theta_{23}, \theta_{13}$
- Nuisance parameters:
  - $\theta_{12}, \delta_{CP}, \Delta m^2_{31}$
- Constants
  - $\Delta m^2_{21} = 7.37 \times 10^{-5} \text{ eV}^2$



# Profiling

$\theta_{23} = [0.47 - 1.10]$  @ 68% C.L.

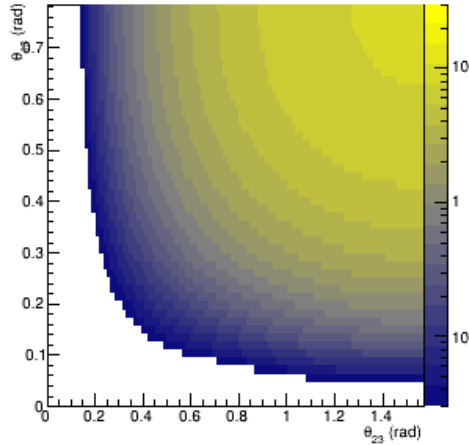
$\theta_{13} < 0.33$  @ 90% C.L.



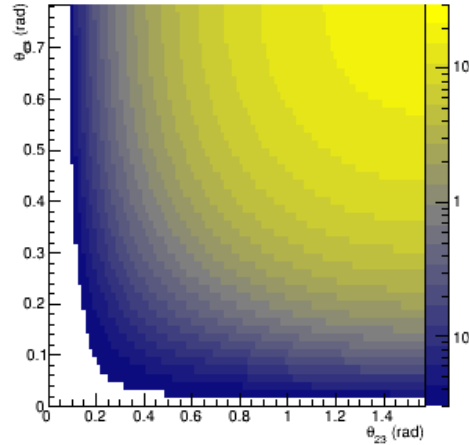
# Bin contribution ( $\nu_e$ channel)

Z axis range: [0 - 30]

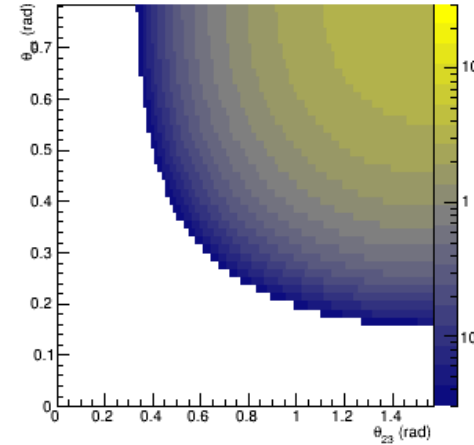
Bin: 1



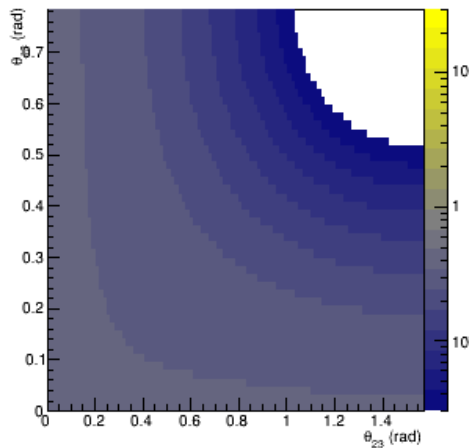
Bin: 2



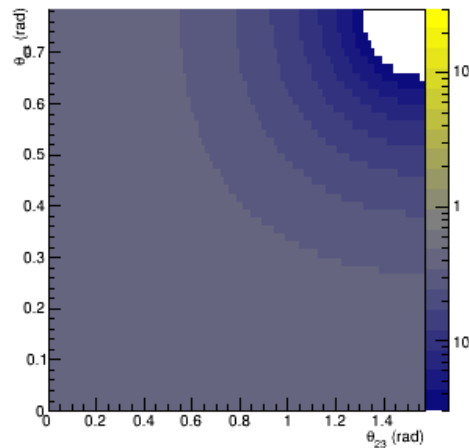
Bin: 3



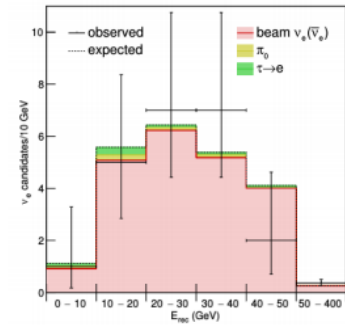
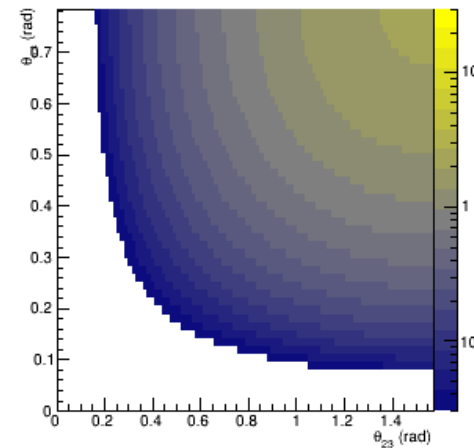
Bin: 4



Bin: 5

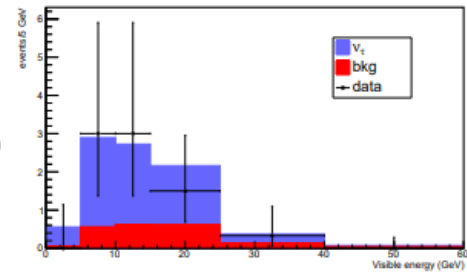


Bin: 6



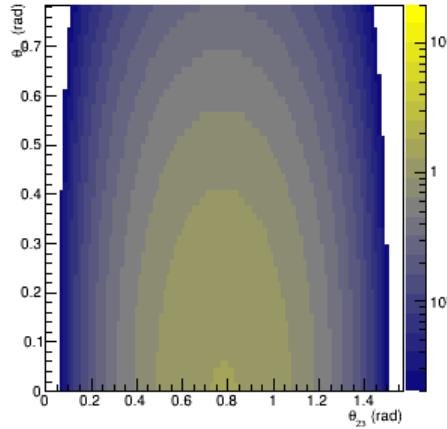


# Bin contribution ( $\nu_\tau$ channel)

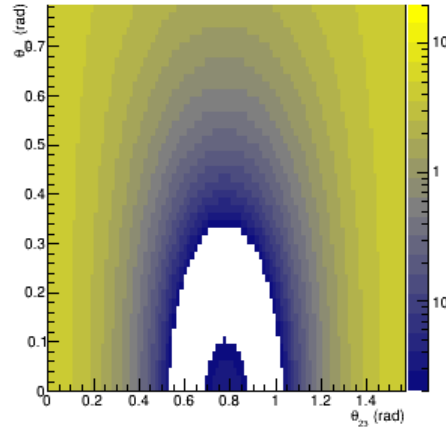


Z axis range: [0 - 20]

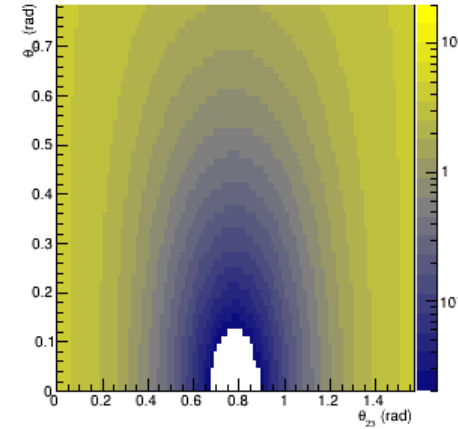
Bin: 1



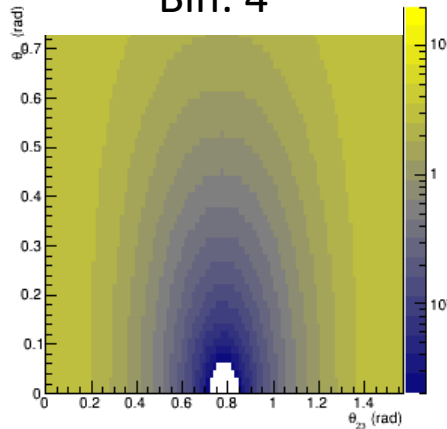
Bin: 2



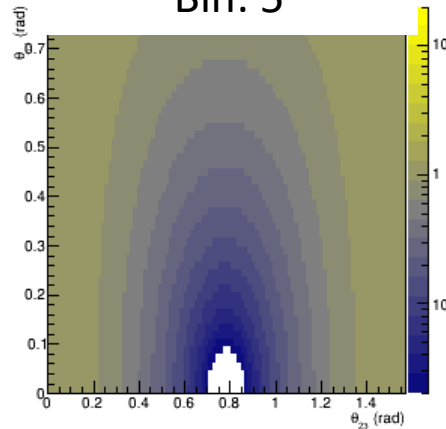
Bin: 3



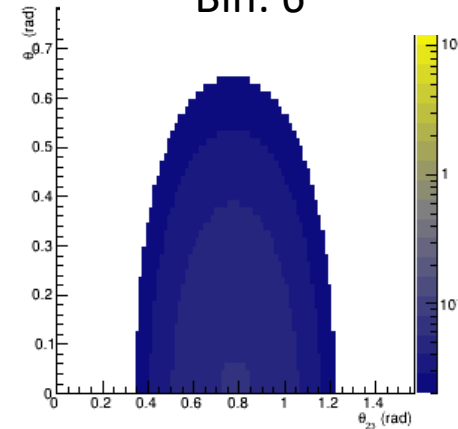
Bin: 4



Bin: 5



Bin: 6



3 + 1 model

# Oscillation model & Mixing matrix representation

- 3+1 Model

- $U = R(\theta_{34}) R(\theta_{24}) R(\theta_{23}, \delta_3) R(\theta_{14}) R(\theta_{13}, \delta_2) R(\theta_{12}, \delta_1)$   
[[arXiv:0704.0388](https://arxiv.org/abs/0704.0388)]

- Last column is  $\begin{pmatrix} s_{14} \\ s_{24}c_{14} \\ s_{34}c_{24}c_{14} \\ c_{14}c_{24}c_{14} \end{pmatrix}$  so  $\begin{cases} \sin^2 2\theta_{\mu e} = 4|U_{\mu 4}|^2|U_{e 4}|^2 = s_{24}^2 \sin^2 2\theta_{14} \\ \sin^2 2\theta_{\mu\tau} = 4|U_{\mu 4}|^2|U_{\tau 4}|^2 = c_{14}^2 s_{34}^2 \sin^2 2\theta_{24} \end{cases}$

- Matter effects (constant density)

- NH &  $\Delta m_{41}^2 > 0$

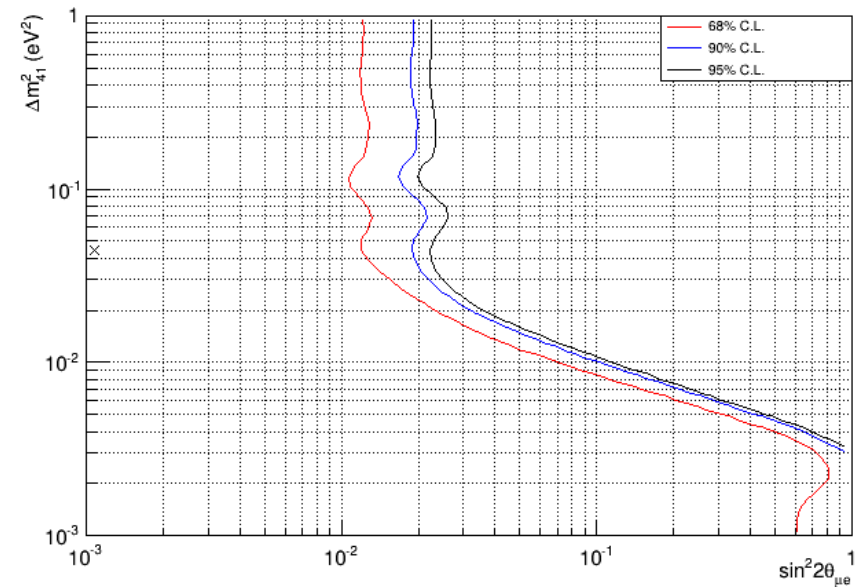
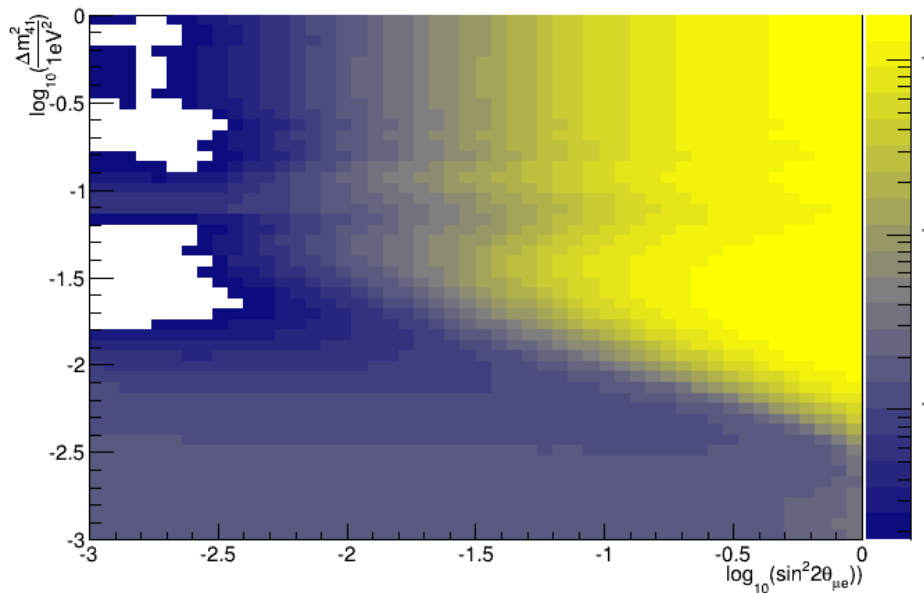
$\Delta m^2_{41}$  VS  $\sin^2 2\theta_{\mu e}$

# POI and nuisance params

- Parameters of interest:
  - $\sin^2 2\theta_{\mu e}, \Delta m^2_{41}$
- Nuisance parameters:
  - $\theta_{12}, \theta_{13}, \theta_{14}, \theta_{23}, \theta_{24}, \delta_1, \delta_2, \delta_3, \Delta m^2_{31}$
  - 10 sampling points in  $\theta_{14}$  for likelihood profiling (to be increased [30])
- Constants
  - $\Delta m^2_{21} = 7.37 \times 10^{-5} \text{ eV}^2$

# $\sin^2 2\theta_{\mu e}$ VS $\Delta m^2_{41}$

Now likelihood is ok

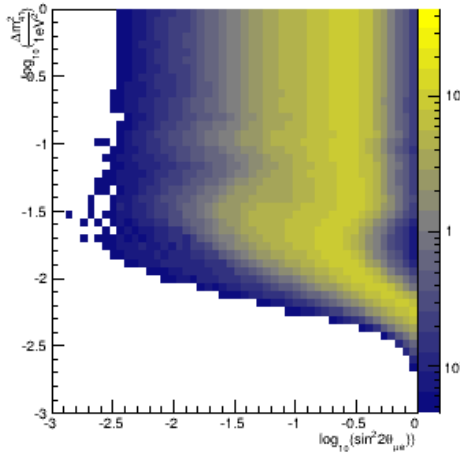


$\sin^2 2\theta_{\mu e} > 0.022$  [90% C.L.]  
@  $\Delta m^2_{41} \sim 1 \text{ eV}^2$

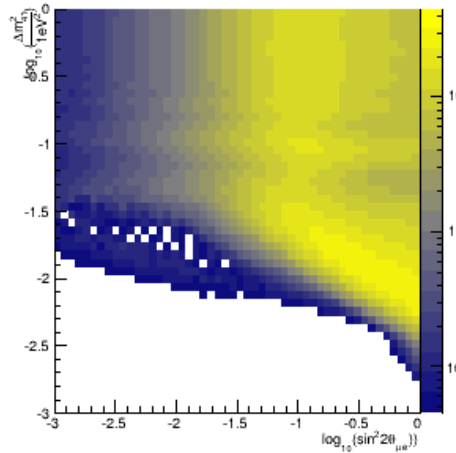
# Bin contribution ( $\nu_e$ channel)

Z axis range: [0 - 45]

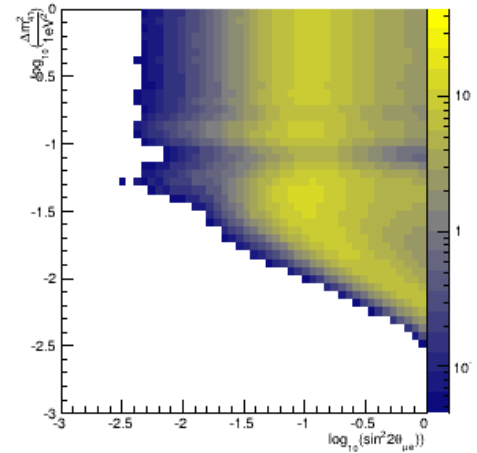
Bin: 1



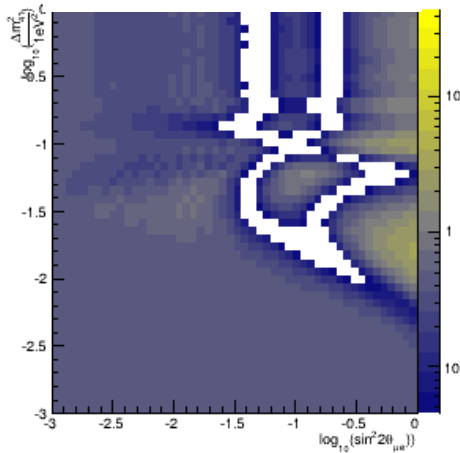
Bin: 2



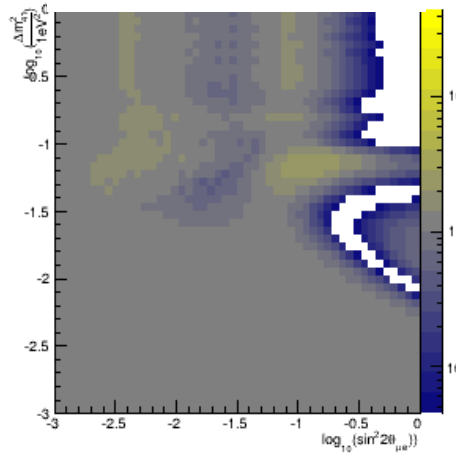
Bin: 3



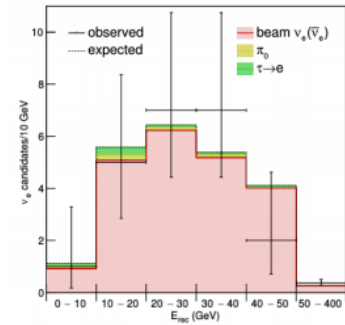
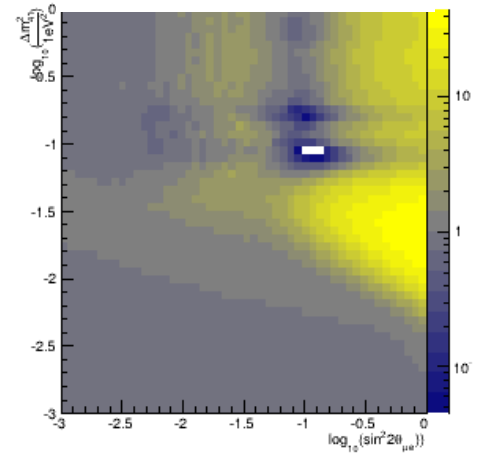
Bin: 4



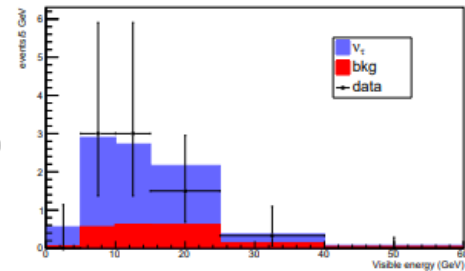
Bin: 5



Bin: 6



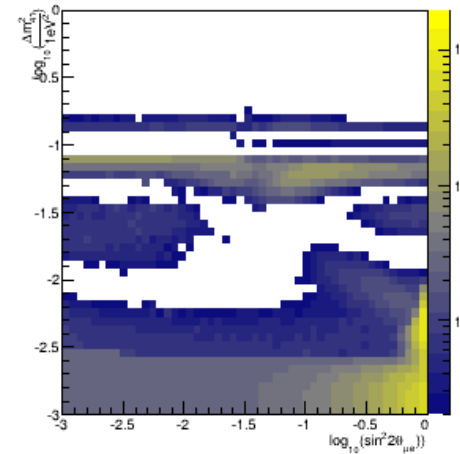
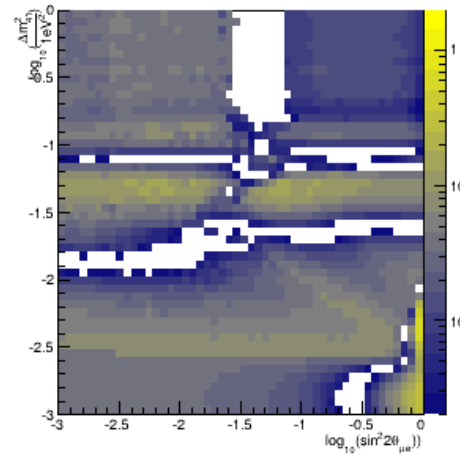
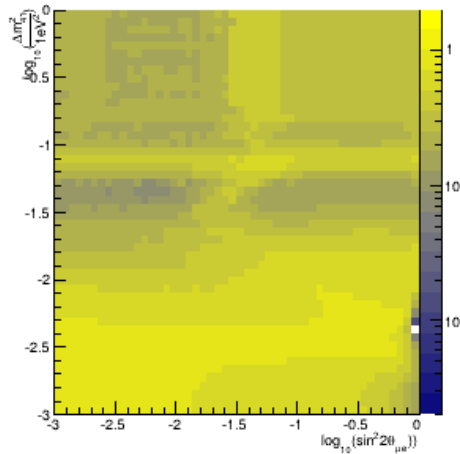
# Bin contribution ( $\nu_\tau$ channel)



Bin: 1

Bin: 2

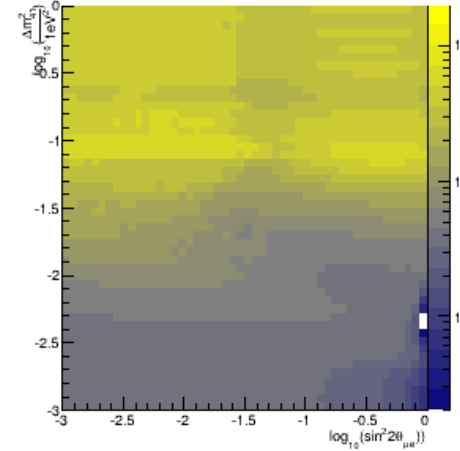
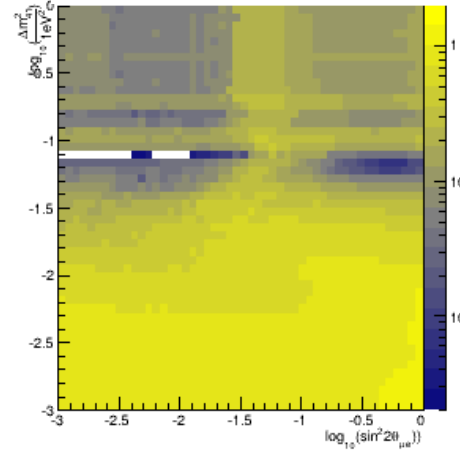
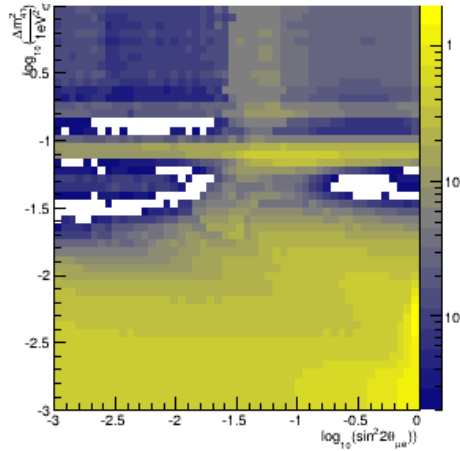
Bin: 3



Bin: 4

Bin: 5

Bin: 6



Z axis range: [0 - 2]



$$\Delta m^2_{41} \text{ VS } \sin^2 2\theta_{\mu\tau}$$

# POI and nuisance params

- Parameters of interest:
  - $\sin^2 2\theta_{\mu\tau}, \Delta m^2_{41}$
- Nuisance parameters:
  - $\theta_{12}, \theta_{13}, \theta_{14}, \theta_{23}, \theta_{24}, \delta_1, \delta_2, \delta_3, \Delta m^2_{31}$
  - 5 x 5 sampling points in  $\theta_{14}, \theta_{24}$  for likelihood profiling (to be increased [10 x 10])
- Constants
  - $\Delta m^2_{21} = 7.37 \times 10^{-5} \text{ eV}^2$

$\sin^2 2\theta_{\mu\tau}$  VS  $\Delta m^2_{41}$

in progress

# Conclusions and outlooks

- Code in
  - [git@baltig.infn.it:opera-combined/opera-combined-globes.git](https://git@baltig.infn.it:opera-combined/opera-combined-globes.git)
  - Readability to be improved
- Machinery seems to work as expected
  - **Additional checks** are on going
- $\sin^2 2\theta_{\mu\tau}$  VS  $\Delta m^2_{41}$  is **in progress**
- Once all is ok:
  - Profiling likelihood sampling will be increased, to improve profiling accuracy
  - IH can be investigated
  - the same for  $\Delta m^2_{41} > 0$
- It will be cool to add also  **$\nu_{\mu}$  disappearance** channel