

30 May 2018

Kimio Niwa

丹羽 公雄

befor OPERA after OPERA

PHOTOGRAPHIC EMULSION in the physics research

- 1896 A Becquerel found **radioactivity**
 - by photographic plate , naked eye.
- 1910 S Kinoshita 木下季吉 studied
 - the darkness by micro scope.
 - darkness made by tracks
 - **“discovery of alpha particle”**

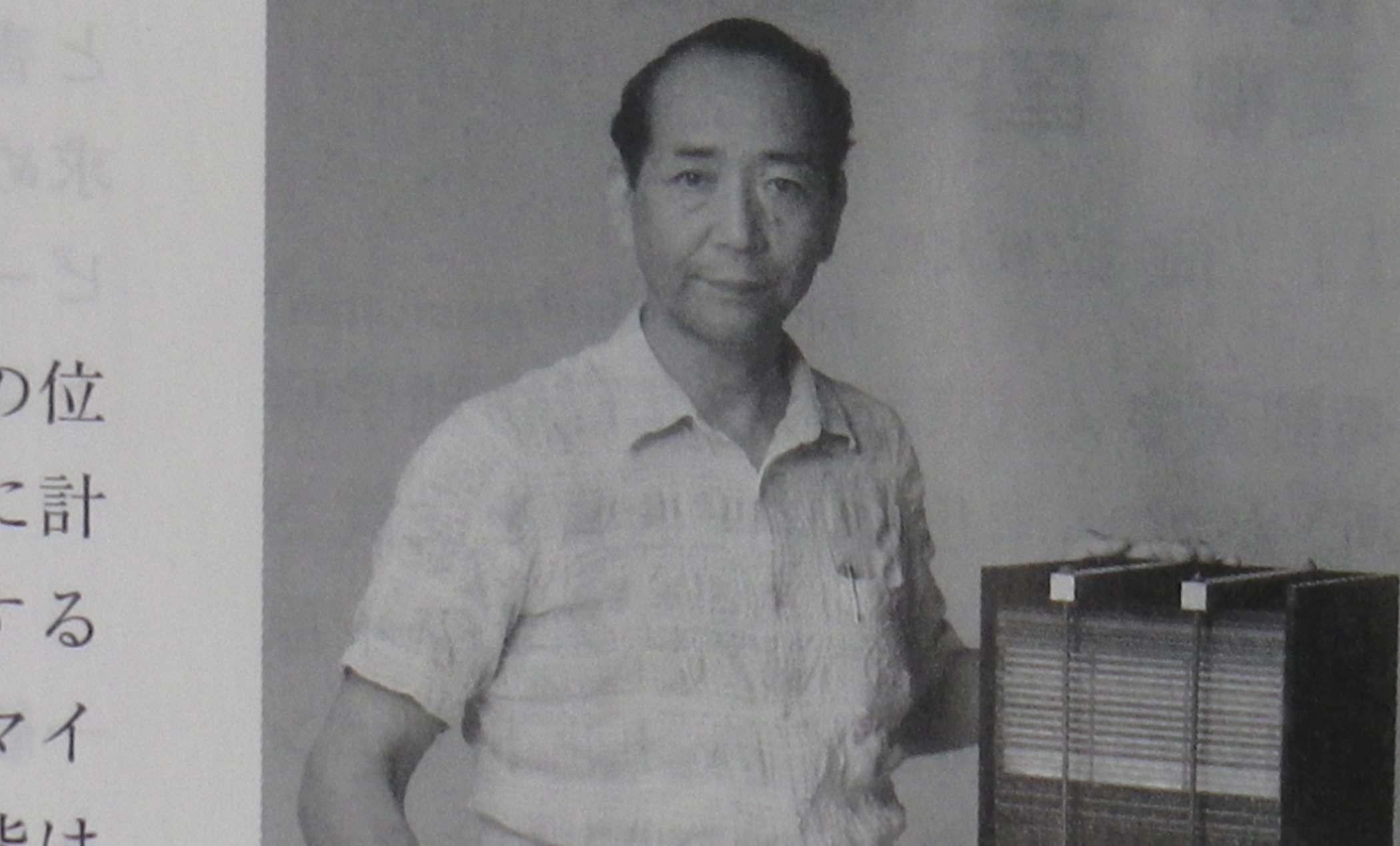
S KINOSHITA



1935 Powell developed electron sensitive gel G5 with Ilford Co. Kodak Co. 1947 μ/e detected in cosmic ray, thick gel [pericle]

1970 both side coated emulsion plate introduced by Japanese physicists Niu, Koshiba, Fuji Co. Both side coated plate is 3D tracking device (x, y, t_x, t_y and D) with high resolution.

K. 1971 Niu discovered charm pti.
in cosmic ray by ECC



ECC

.Sandwich structure by

- emulsion plate and absorber
- Absorber: iron and or lead etc.
- Emulsion plate: double side gel coated
- **E**mulsion **C**loud **C**hamber

.Multiple layer of Spark chamber was used in Cosmic ray study. In 1950-1970 and U_e U_μ identification EXP.

After J/ Ψ discovery 1974

- Beauty particle found
- Tau lepton found
- Interest changed!
- From quark to neutrino

- 1990 KAMIOKANDE obtained atmospheric neutrino. U_e/U_μ separated
- up going neutrino number is missing compare to down going neutrino!
- Neutrino oscillate ? or sterile ?
- Neutrino oscillation was introduced 1962 by MINKAGAWA et al.
- If oscillate, Neutrino is not massless !!
- tau neutrino is dark matter!! **Harari 1998**
- Our interest changed from quark to **neutrino world!**
- CHORUS 1990 started at CERN
- DONUT at FNAL tau neutrino search

DONUT

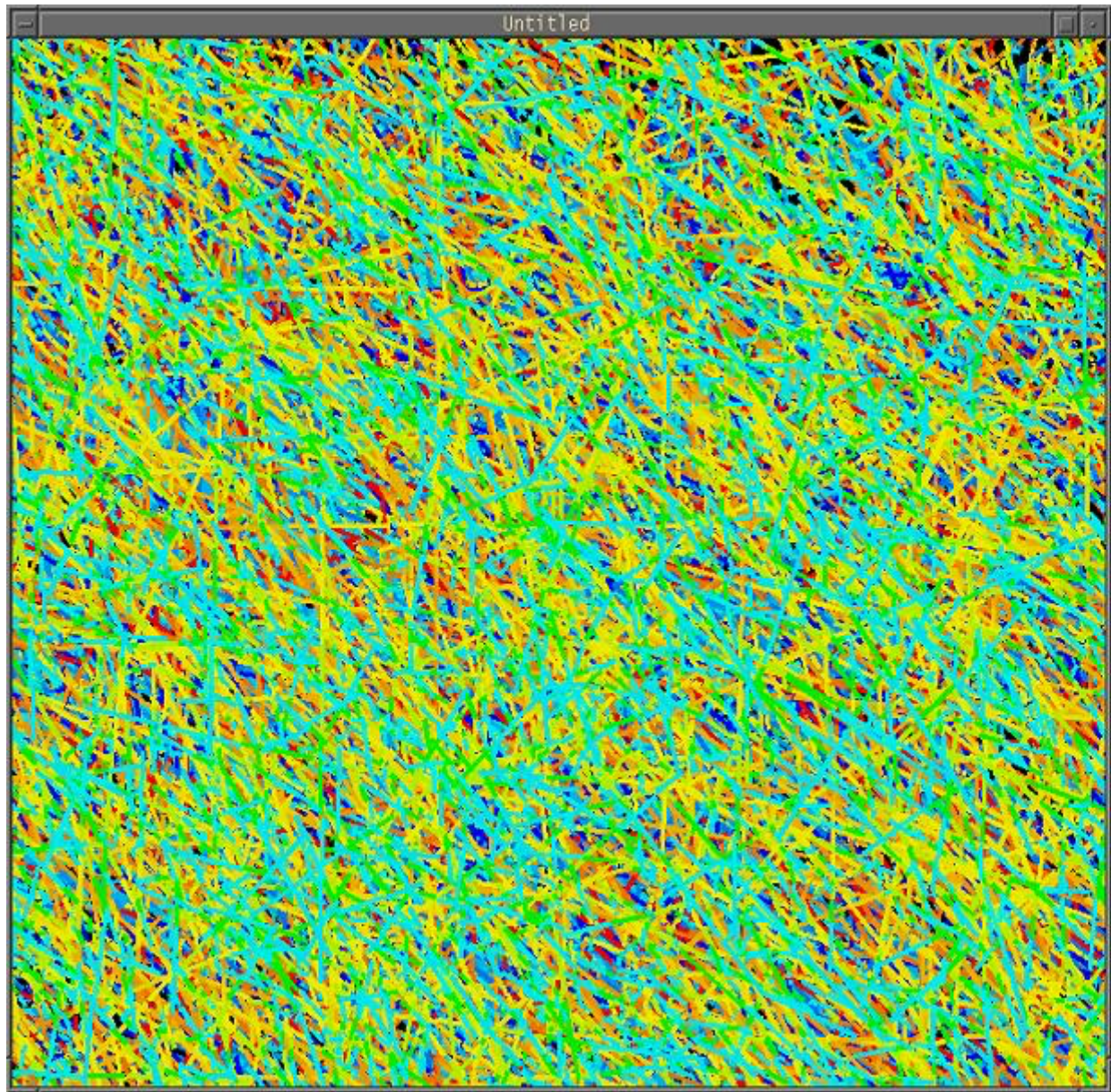
- Beam dump of FNAL TEVATRON
 - 800GeV proton
 - Detector is ECC+muon spectrometer
 - Hand made both side em coated plate
 - Absorber is iron with thickness 1mm
 - Size 50cm by 50 cm
 - Automatic scanning and NET-SCAN was applied
 - Track density $\sim 10^7/\text{cm}^2$

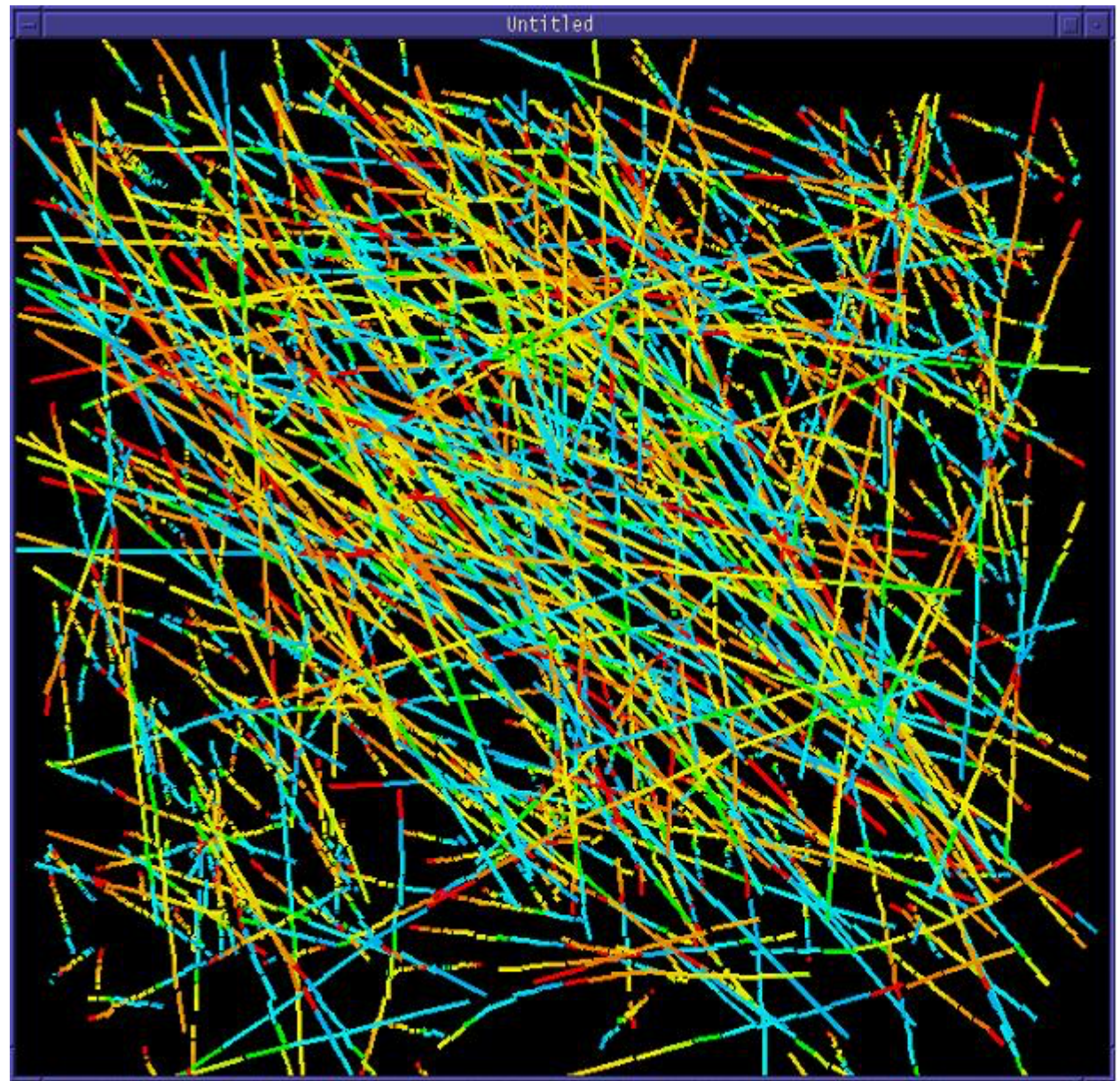
DONUT detector

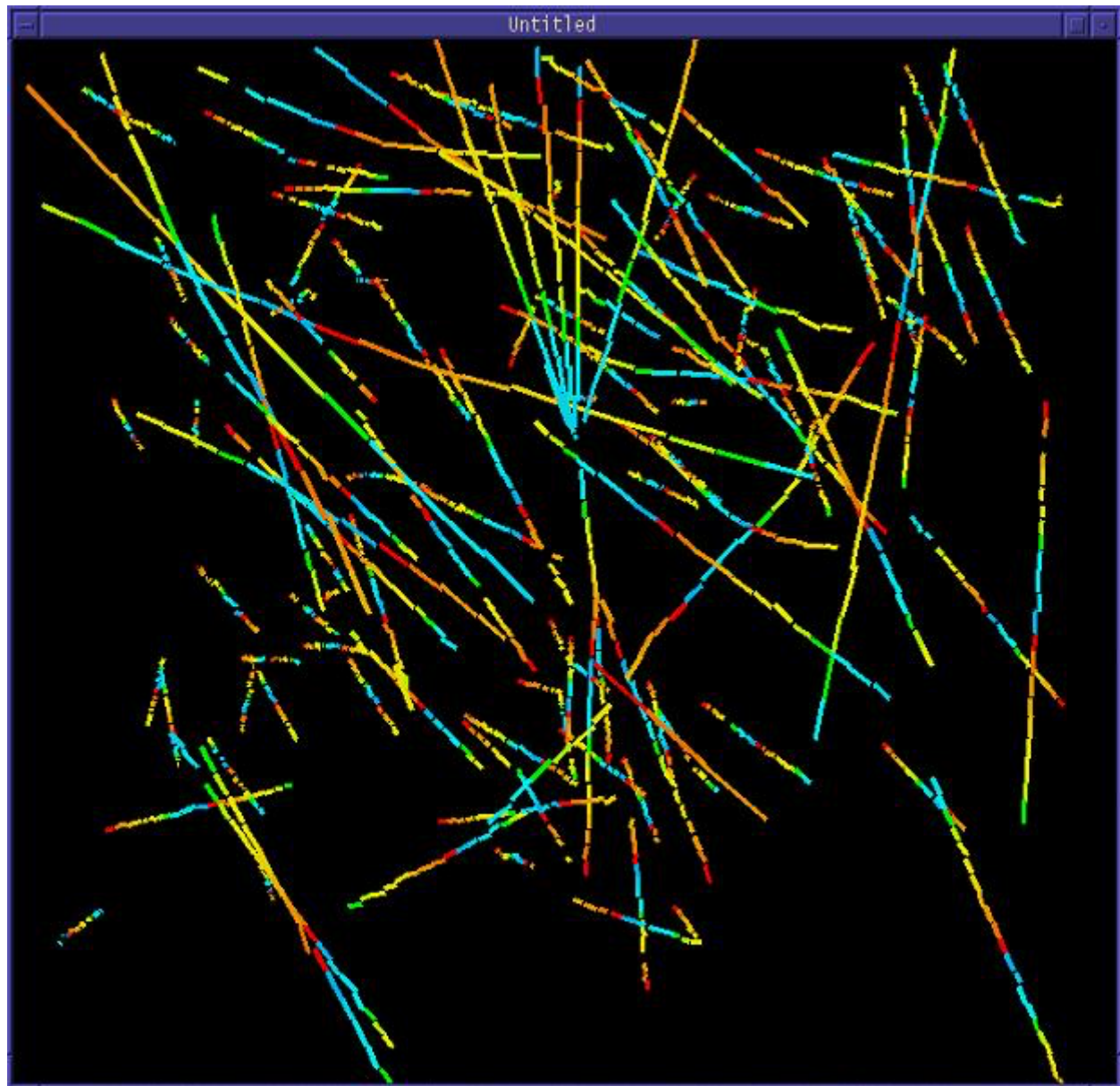
50kg ECC 4 units total 250kg

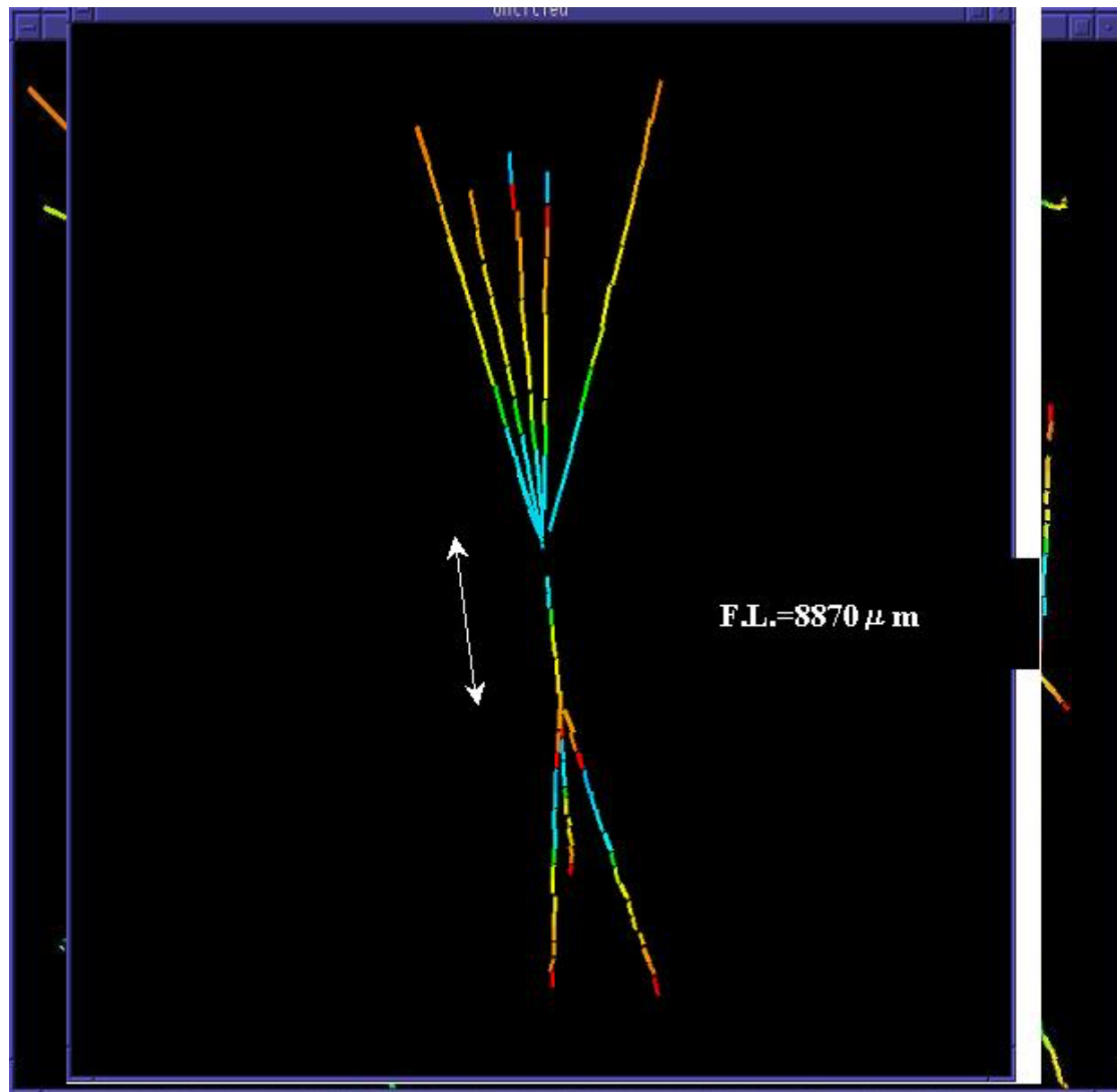
.Sandwich structure, emulsion plate and iron plate











From Quark to neutrino

- 1996 KAMIOKANDE new data mean
- atmospheric neutrino is misterius !!
- Looks like, neutrino is oscillating...
- Mixing angle is big, mass difference is small
-

1995 at FNAL I_(Niwa) introduced the idea of
long base line ν_μ - ν_τ appearance EXP by
DONUT like 100ton ECC. start of opera
assumed mixing angle 10^{-2}
using KAMIOKANDE (before SK) data

Neutrino mass may be not heavy,

ν_t dose not a candidate of darkmatter.

ν_μ - ν_t dose not osc. in short distance.

CHRUS should be no solution! Long base line
should be done with big target mass

Summary 1947 to 2005

- 1947 C.F.Powell $\pi/\mu/e$ [emulsion thickgel: pericle]
- electron sensitive gel G5
-
- 1971 K Niu Charm ptl by ECC Emulsion cloud chamber
- both side coated film : track position(x,y) and angle(tx,ty)
- manually scanning
- 1998(200) DONUT tau neutrino by ECC
- handmade emulsionfilm
- automatic scanning

1995 I_(丹羽) introduced the idea of long base line

U_μ - U_τ appearance EXP at FNAL by

- DONUT like 100tonECC 700km fromFNAL.
- assumed mixing angle 10^{-2}
- using KAMIOKANDE before SK
-
- Neutrino should not be heavy,
- U_t dose not a candidate of darkmatter.
- μ -neutrino dose not osc. in short distance.
- CHRUS should be no solution!
- Long base line should be done!!

Who interested to Niwa talk ?

- FNAL adam para
- MINOS leader S.Wojiki?
- ECC 100 ton ! Emulsion ?
- At CERN: discussed with chorus p
- K.Winter, Dore, Roberta
- Next supporter!
- L.Fore, Antonio , P strolin
- Seminar at rome univ. etc.
-

まとめ From chorus to opera proposal

- 1998 TAKAYAMA neutrino conference.
 - SK group (Totsuka and Kajita) presented
 - the beautiful result on atmospheric neutrinos
 - Up going neutrino missing compare to down going neutrino. **Mu neutrino disappeared !**
 - **Strongly suggested neutrino oscillation**
 - **Sterile?**
- DONUT presented one neutau event detected
 - In ECC.

LOI : ANTONIO E assembled !

[Appearans] is indispensable to confirm
the neutrino oscillation.

- Tau neutrino detection is possible by ECC
- Big ECC detector construction is possible !
- Fuji Co. can make big volume emulsion film.
- Scanning is possible! $20\text{cm}^2/\text{h}/\text{sys}$

| | | | | |
|--|--|--|--|--|
| | | | | |
| | | | | |

.1999 OPERA proposed

- . DG Maiani supported

.2000 approved **at CERN**

- . 1.5kton ECC sensitive to $3 \cdot 10^{-3}$ mixing

.Maiani pushed OPERA

- . Machine made emulsion film by FUJI Co.
- . BAM, Lead plate, packing, Target tracker,
- . Muon spectrometer, 1.5kton mounting.
- . Brick handling(exchange) system,
- . Automatic develop, CS, CS to Brick, etc.

- .
- . By international collaboration work.

.

Opera progress after approval

- .2007 OPERA detector constructed at Grnsasso
- . CNGS(neutrino beam from CERN)constructed
- .Neutrino exposure 2008-2014
- .2008/5 The first event (first light) detected in ECC
 - . under microscop
- .20010 the first tau neutrino candidate
- .20012 clear tau neutrino (leptonic tau) detected
- .20015 5 tau neutrino events
- .20018 final analysis 6 sigma

Opera summary

- 1995 idea at fnal (DONUT meeting)
- OPERA detector constructed in Gransasso.
- OPERA/CNGS construction 2000-2007
- Exposure 2008-2014
- 2008 dec. first neutrino event is located.
- N_{μ} - N_t 5 events 6-sigma
- Technological out put, useful for future
 - Big volume emulsion film handling
 - High speed scanning machine with $1\text{m}^2/\text{h}$

教訓

- Thanks INFN for money short of “japan”
 - to organize international collaboration
- Trouble: **over light velocity**
 - Good lesson to scientist
- It is important to open OPERA data
 - Big data handling, application soft.
 - for outside researcher and education .

教訓

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Trouble : Neutrino velocity
Money short

friends passed away

G Romano

Giacomelli

Tolun(ankara)

Niu(japan)

Kanazawa Fuji co.

Guy Vanbeek(burussel)

合掌



HTS scanning speed 1m2/h

.HTS の写真

After opera [future]

- .We managed big volume emulsion film
 - by low cost 1m^2 100 ER(big volume only)
 - Other device are more expensive.
- .We have big scanning power
 - $1\text{m}^2/\text{h}$ or more $\sim 10\text{m}^2/\text{h}$ in future

Emulsion future

• Emulsion is the long life detector in physics

• Start 1896 ,

• Used in research and business..

• Very good interaction with industrial technology.

• Emulsion is not behind the times,

• rather ahead of the times.

•

•

OPERA else

- Gamma ray star AOKI
- balloon , 1m², 10 m² E_γ>10MeV,,,1GeV
- Monopol, antiproton ,quark,
- Under ground ,airple,balloon,
- μon radiography
- blast furnace / pyramid /etc
- check the status without distroy
- WIMPS search ?, super fine AgBr crystal
- Double hyper nuclus analyzed ΛΛ force,
important for neutron star study NAKAZAWA

Nuclear reactor

F.Reines ν_e

detector:liquid scintillator+PMTs

BNL 14/28 GeV **beam dump**

Lederman ν_μ

detector:multilayer spark chambers

FNAL 800GeV **beam dump**

DONUT ν_t

detector:FCC

4th Neutrino

e μ τ X₄ ?

V_e V_μ V_τ V₄ ?

rejected Mass (ν₄) ≤ 50 GeV LRP exp
Mass () ≤ 100 GeV LHC

try mass 100 GeV ≤ .. ≤ 7 TeV

- Search for very heavy neutrino ν_4
 - One of dark matter WIMP candidate
 - LEP killed light neutrino ν_4 mass $< 50 \text{ GeV}$
 -
 - LHC no signal on heavy mass neutrino
 - DAMA result is not killed
 - mass $\sim 50 \text{ GeV}??$

LHC

7TeV X 7TeV

- On the ground above CMS/ATLAS
- Collision point to the surface \approx 100m
- Detector area 100m X 100m
- Detect the muon position and impact parameter and momentum

LHC ポンチ絵

Beam dump emulsion exp

| name | Target mass | physics | date year |
|---------------|-------------|--------------------------------|-----------|
| DONUT | 100kg | ν_τ discover | 2000 |
| FNAL tevatron | | | |
| • | | | |
| • OPERA | 1000 ton | $\nu_\mu \rightarrow \nu_\tau$ | 2017 |
| • SPS | | appearance | |
| • New | 1 Mega ton | ν_4 | ? |
| • LHC | | | |

| Experiment | Numbr of proton int | Target mass | Energy cross section |
|-----------------------|--|-------------|------------------------|
| DONUT tau neutrino | 10^{18} | 250kg | 20Gev (800GeV) |
| LHC V_4 search | $100\text{MHz} \cdot 3 \cdot 10^7 \text{s} = 3 \cdot 10^{15} \text{int/y}$ | 1M ton | 14TeV (10^8 GeV) |

present situation LHC is wonderfull !

LHC beam colliding position is
under ground depth 100m

Put the detector on the ground

above ATLAS and/or CMS

detector size $100\text{m} \times 100\text{m}$

ECC muon detector

- Position and angle of muon
- impact parameter analysis
- Momentum measurement
- multiple scattering

- Iron plate and emulsion film ~50 layers
- muon momentum (TeV muon)
- iron plate thickness ~0.5cm 40kg
- Unit ECC weight is ~2 ton

possible ? SCANNING

Opera

new

unit size 10*cm*10cm*0.6cm

100*100*50 layer

mag spectrometer +scat.

Scatterng only

• Unit number 100000 units

10000 units

•scanning power 100cm²/h

10000cm²/h

• Only selected

all

コスト how mach??

- Ecc 10000 units cost
- Ecc cost film 2万/m² 100万円10⁴ euro/unit
- Total detector cost 10⁸ euro 100億円

- 仲間 中国？
- (HP Kamioka:: korea chaina)

- 宇宙線研究 高エネルギーμフラックス