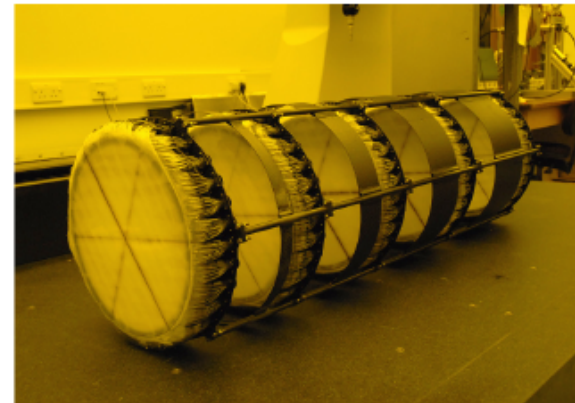
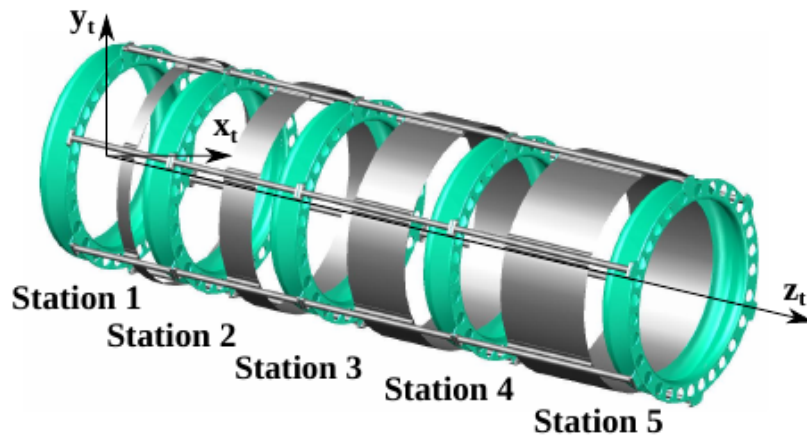
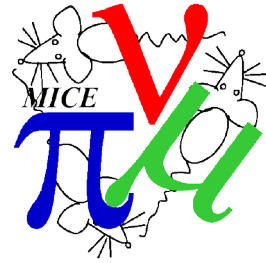


Emittance Evolution Update



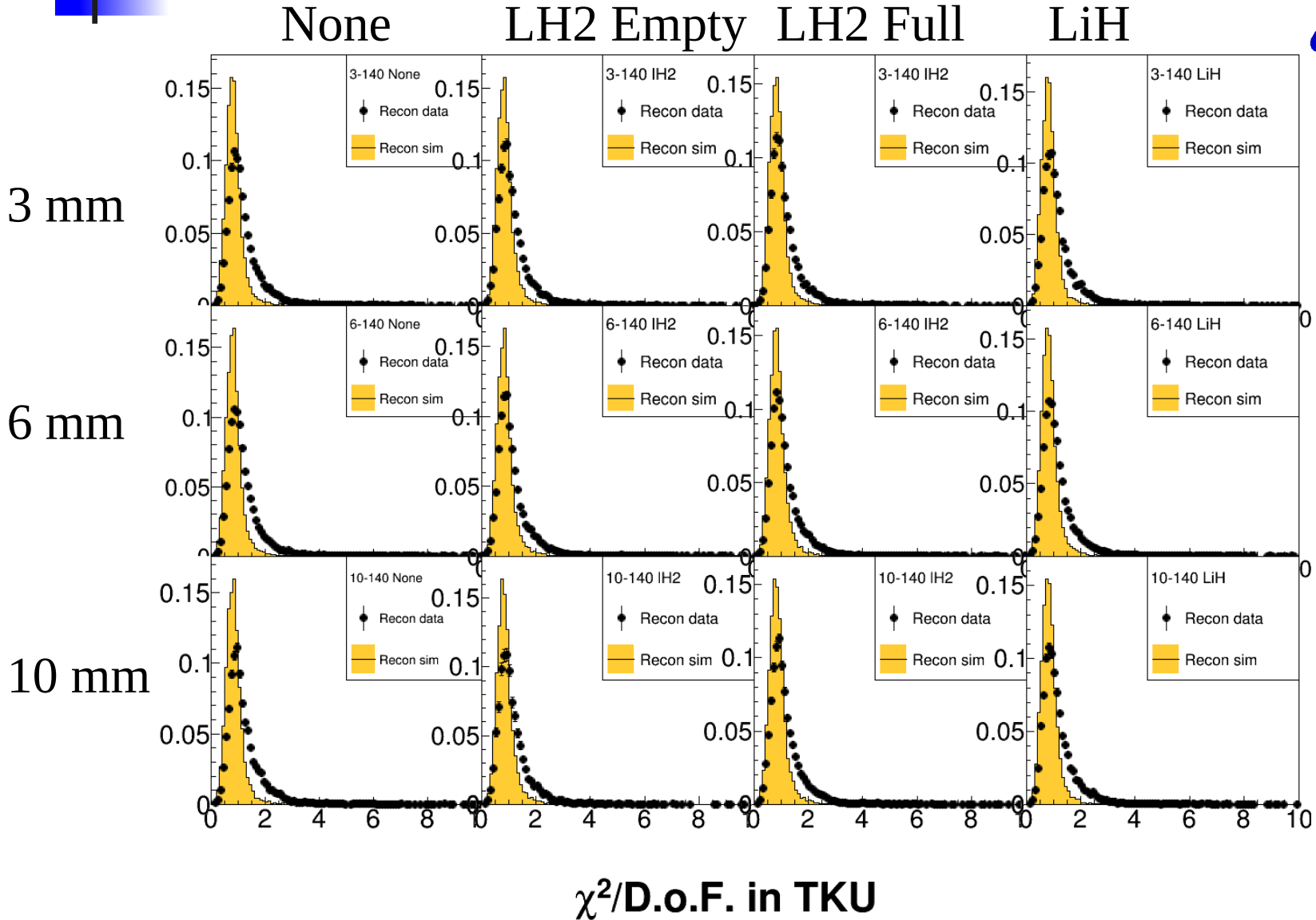
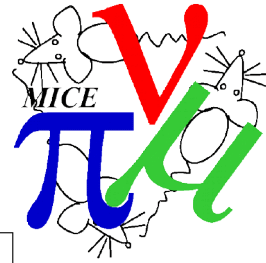
C. Rogers, ISIS Intense Beams Group
Rutherford Appleton Laboratory

Status

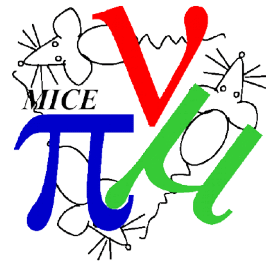


- Since CM50
 - Characterised tracker noise; optimised NPE cut and added dark current model
 - Fixed and tuned tracker density issue
 - Implemented routine to enhance statistical sample for MC
 - Systematic uncertainties due to tracker misalignment, field non-uniformity

Chi2/dof (from CM50)

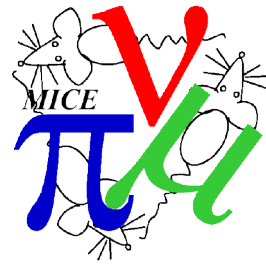


Sources of uncertainty

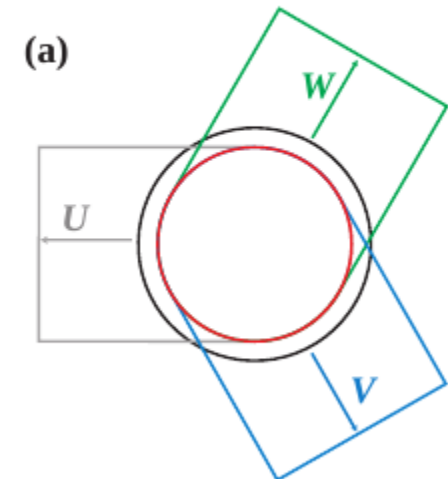


- Oh no!
 - E.g. 3-140 IH2 empty setting
 - MC has peak at 0.8; mean 0.9
 - Data has peak at 0.9; mean 1.2 → longer tail
- What are sources of uncertainty?
 - (1) Scattering and energy straggling in reconstruction
 - (2) Field non-uniformity
 - (3) Tracker rotation compared to solenoid field
 - (4) Detector noise
- (1) and (2) have reasonable model in MAUS
- (3) – I believe Chris Hunt has studied and found no effect
- What about detector noise?

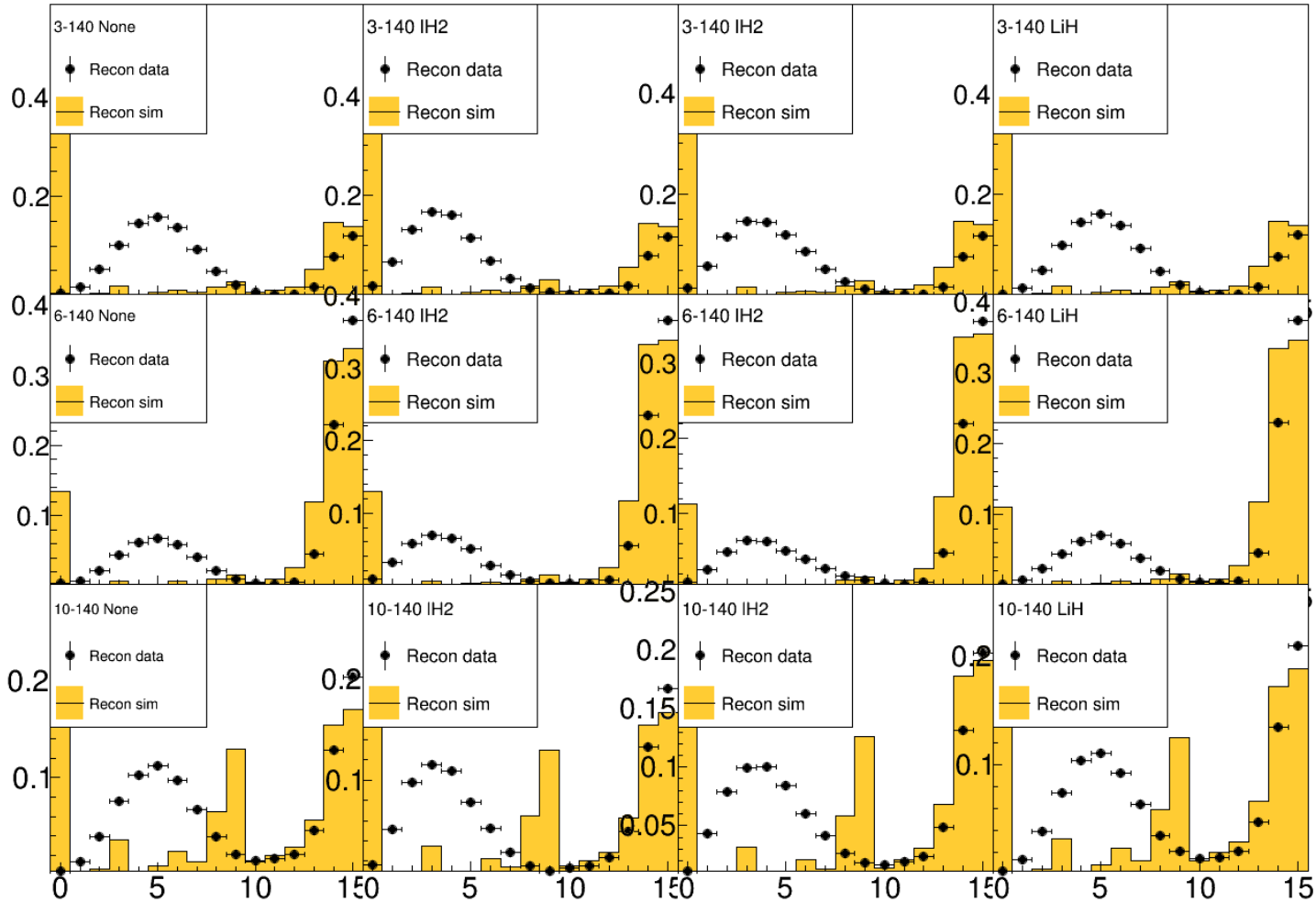
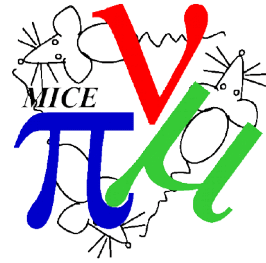
Reminder of tracker recon



- Electronics signal makes a “digit”
- Neighbouring “digits” in the same plane make a “cluster”
 - Channel number \rightarrow distance across the plane
- Crossings of “clusters” in a plane make “space points”
 - Two of three “clusters” makes a “doublet”
 - Three of three “clusters” makes a “triplet”
 - Space point has x-y-z coordinate
- Reject spurious space points using naive helix fit
 - “Pattern Recognition”
- Make a full fit to get the “best” fitted track

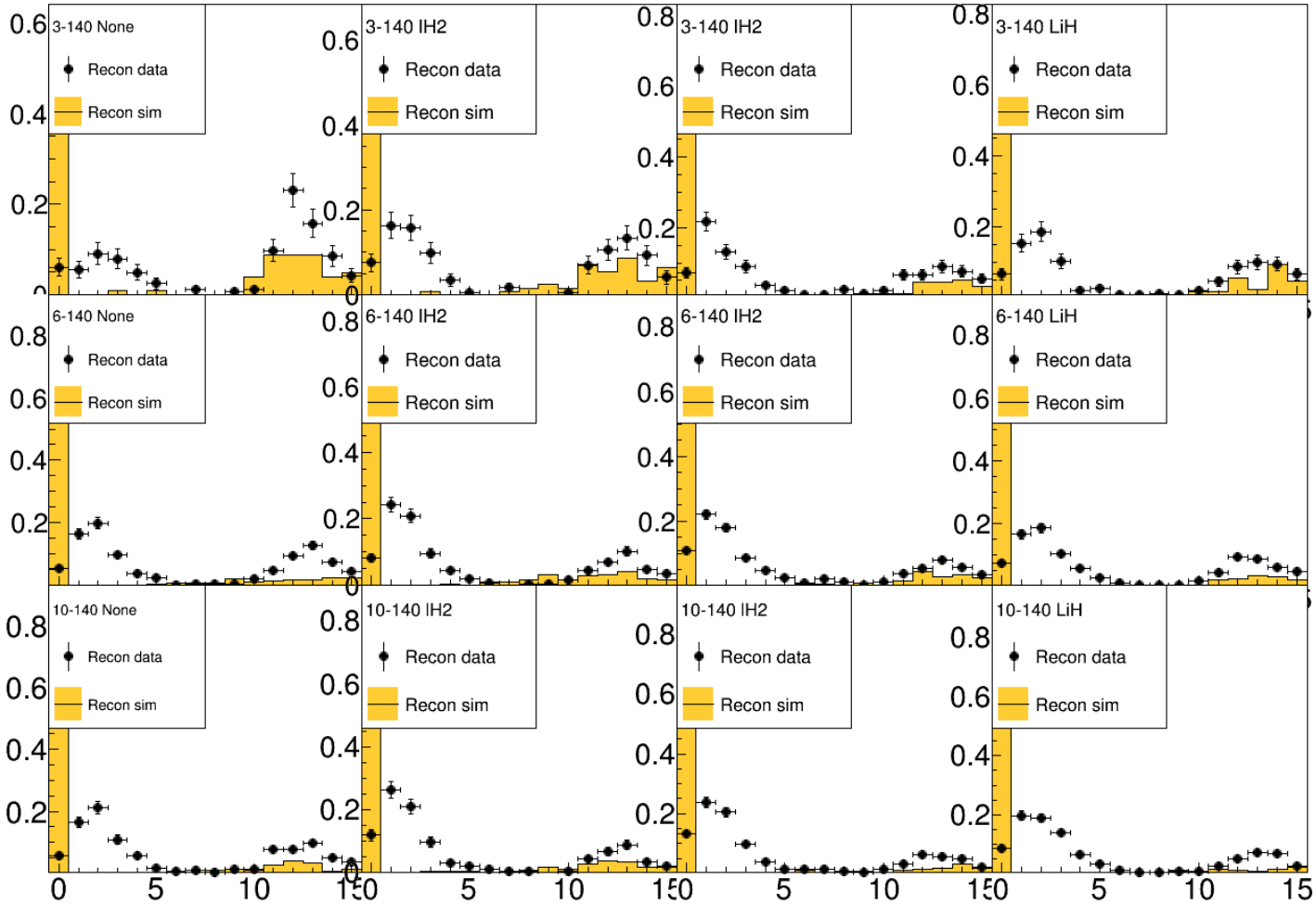
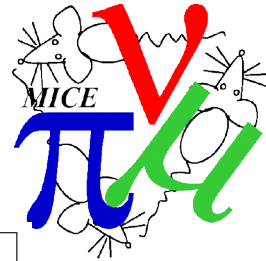


TKU clusters (from CM50)



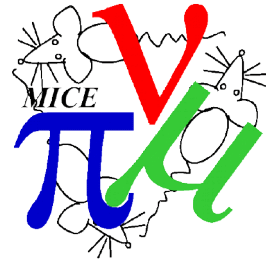
Number of planes with clusters in TKU
For events that DO NOT form a track

TKD clusters (from CM50)



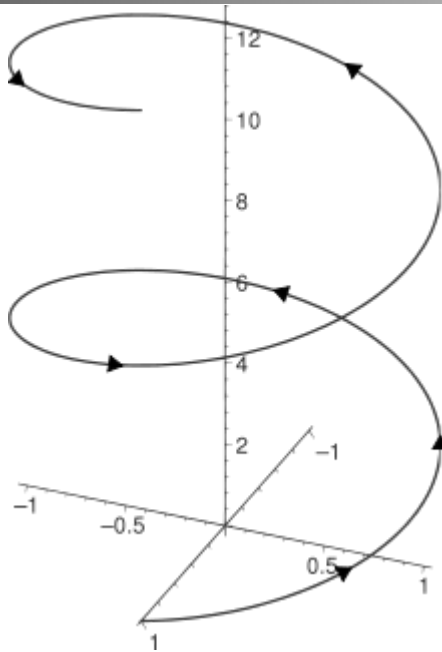
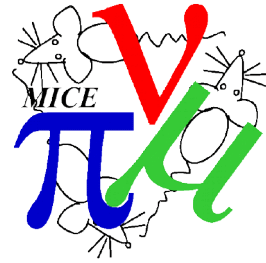
Number of planes with clusters in TKD
For events that DO NOT form a track

Detector noise



- There is no tracker noise model active in MAUS by default
 - But is noise an issue?
 - How much noise is there anyway?
 - If we switch on MC noise, is it right?
- Seek to decide if noise is significant contributor to resolution
 - Look at pattern recognition (PR)
 - Is PR rejecting space points as noise that could have made good tracks?
- Seek to characterise the noise in the data, and then ensure MAUS correctly reflects this effect

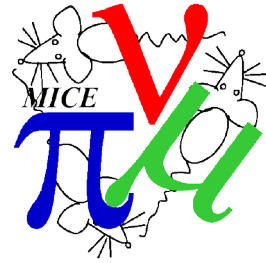
PR Routine



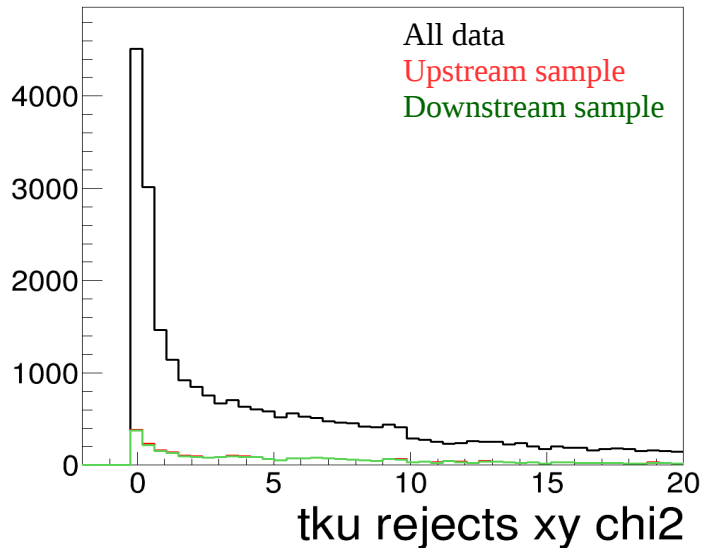
■ PR Routine

- Attempts to fit a circle to the projected space points in xy and returns xy_chi2
- Attempts to fit a line to the space points in sz and returns sz_chi2
 - S is the distance between points around the circumference
- Tries the fit over every possible combination of 5 and 4 space point tracks
- Picks the lowest $(sz_chi2)+(xy_chi2)$
 - All “unused” space points are assumed to be “noise”

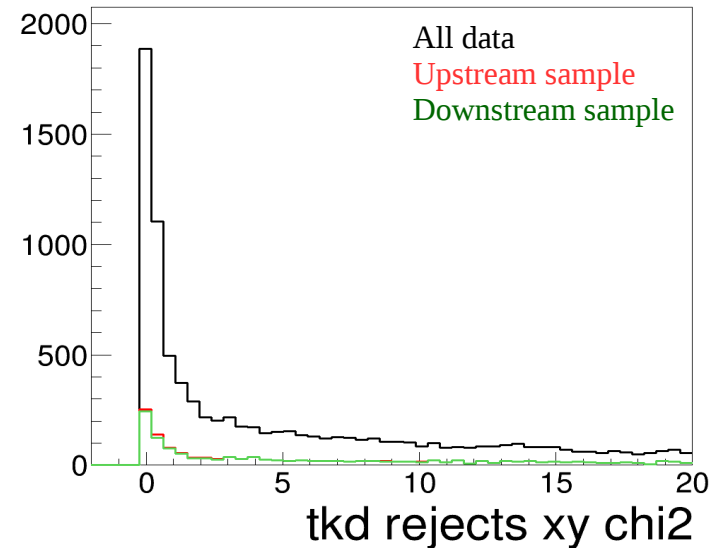
Pattern recognition



2017-2.7 3-140 IH2 empty

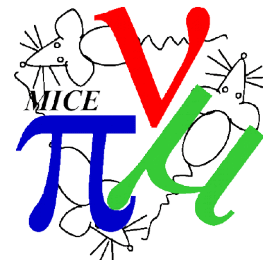


2017-2.7 3-140 IH2 empty

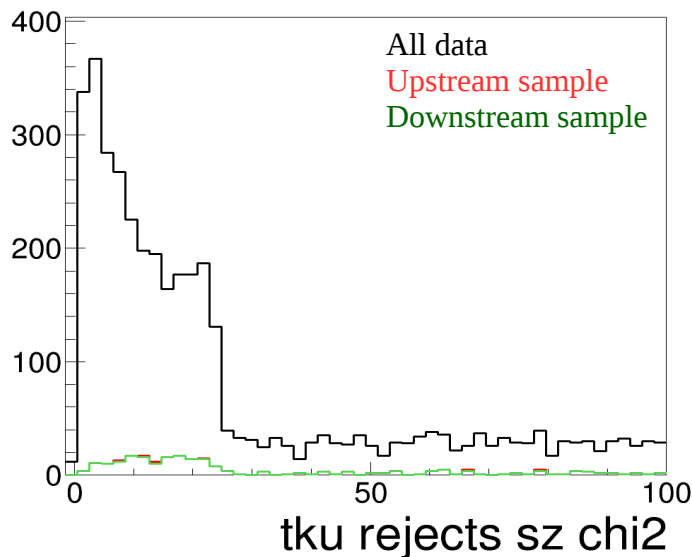


- Circle fit returns a chi2 value
 - Chi2 < 10 => accept track
- Plot Chi2 distribution of the second best 5 space point PR track candidate
 - This is the best candidate that did not make a track
 - “5 space point candidates only” to avoid double counting

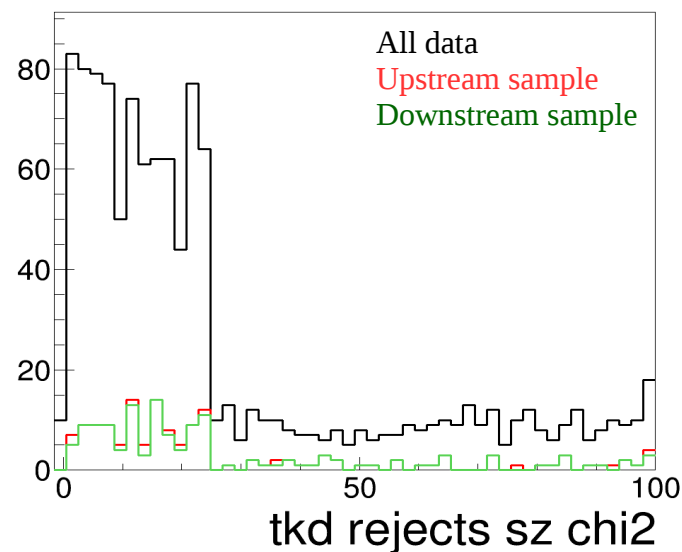
Pattern recognition



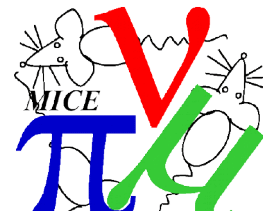
2017-2.7 3-140 IH2 empty



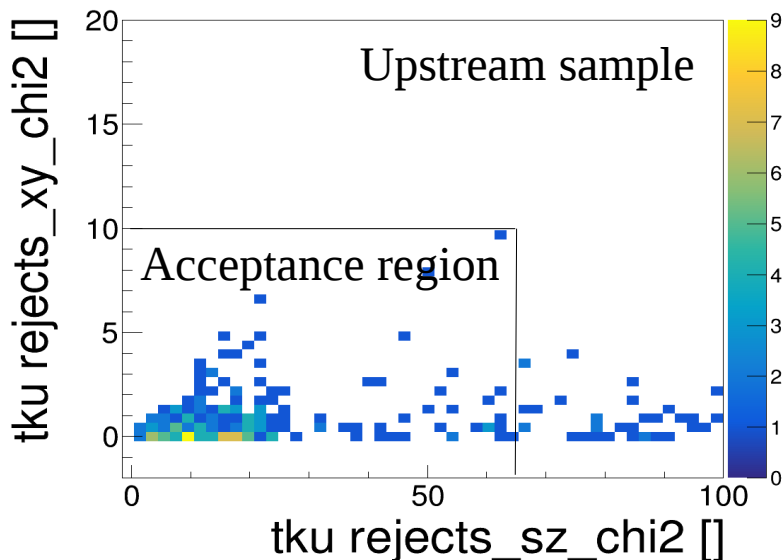
2017-2.7 3-140 IH2 empty



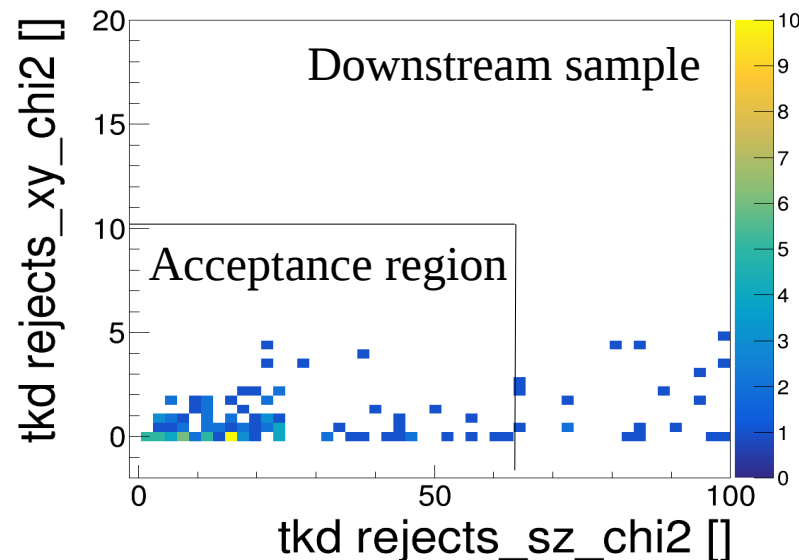
- Helix fit returns a χ^2
 - $\chi^2 < 65 \Rightarrow$ accept track
- χ^2 distribution of the second best 5 space point PR track candidate
 - This is the best candidate that did not make a track
 - “5 space point candidates only” to avoid double counting



2017-2.7 3-140 IH2 empty

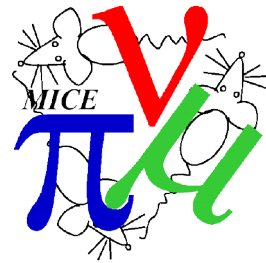


2017-2.7 3-140 IH2 empty



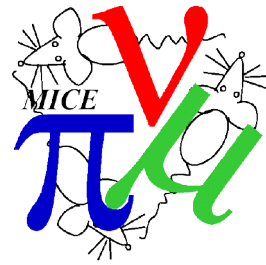
- Chi2 distribution of the second best 5 space point PR track candidate
 - This is the best candidate that did not make a track
 - 5 space point candidates only to avoid double counting
- Conclusion: there are lots of “unused” space points that could have made a track
 - Are we sure these are really noise?
 - What if some of the “data” is really noise and vice versa?

Detector noise



- Is noise a significant contributor to resolution
 - There are a significant amount of hits that are considered as noise but could have made a good track
 - Guess there are noise events that have been included in the track reconstruction
 - → quite possibly noise is a significant contributor
- Seek to characterise the noise in the data
 - Look at amount of light produced by “noise events” (NPE)
 - Look at physical location of noise events versus real events

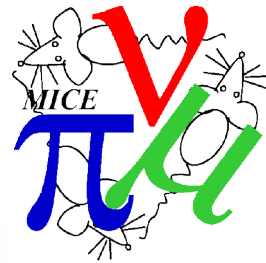
Number of Photoelectrons



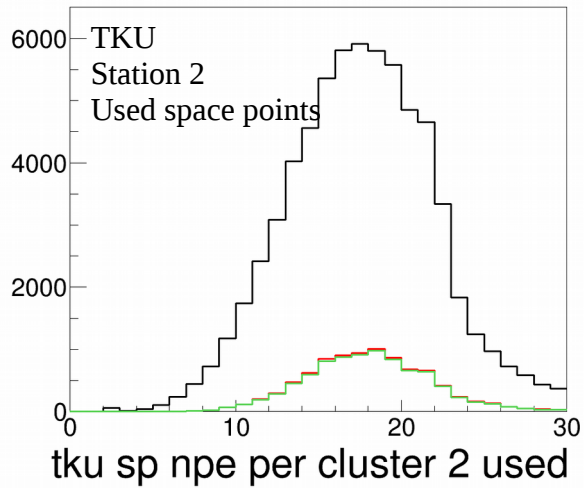
- Light appears in scintillator when particles go through
- Light travels down waveguides to photon counter
- Photon counter (VLPC) makes light into photoelectrons
- EM shower \rightarrow electron multiplication
- \rightarrow Number of Photoelectrons (NPE)

- What causes noise
 - Thermal noise in the VLPC makes some electrons appear
 - Light leaks into the system from elsewhere (we hope not!)
 - Non-beam particles make it into the detector
 - e.g. muon “knocks-on” an electron from surface of the detector into another detector element

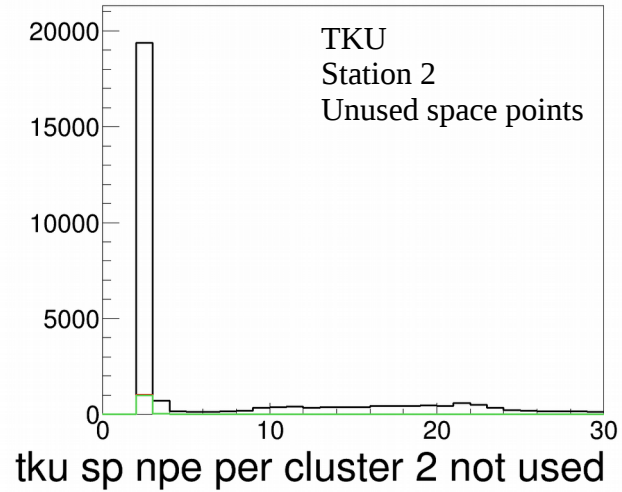
Number of Photoelectrons



2017-2.7 3-140 IH2 empty

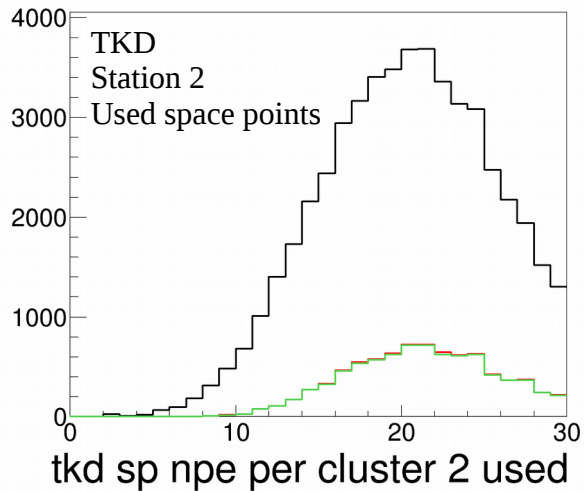


2017-2.7 3-140 IH2 empty

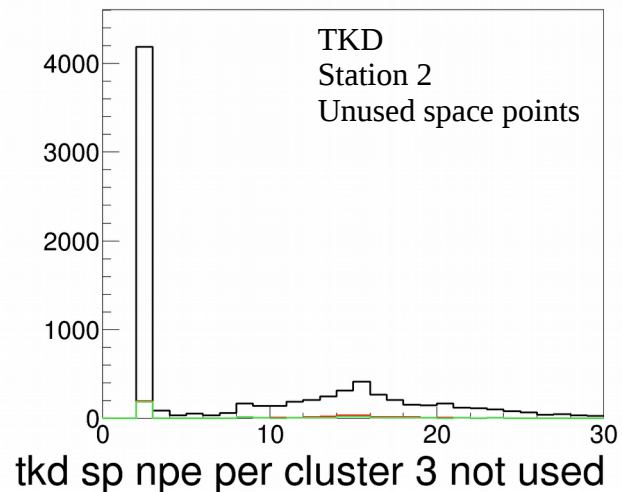


All data
Upstream sample
Downstream sample

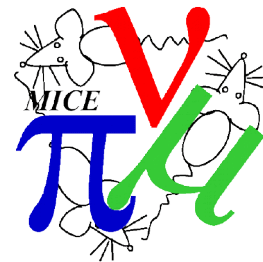
2017-2.7 3-140 IH2 empty



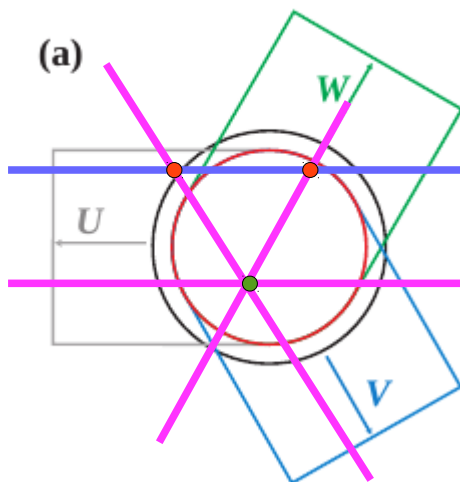
2017-2.7 3-140 IH2 empty



NPE - Conclusions



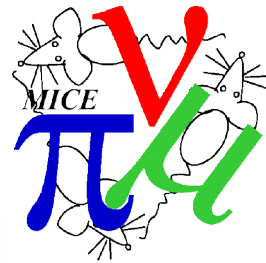
- Lots of “unused” spacepoints have 3-4 npe/cluster
- Almost no “used” spacepoints have 3-4 npe/cluster
 - Those events that do have 3-4 npe/cluster are rejected by some other cut
- Assume this is thermal noise
 - → propose 4 npe cut on data
- Note these space points are rejected anyway
 - → this might not be the source of our chi2 anomaly...
- Reminder:



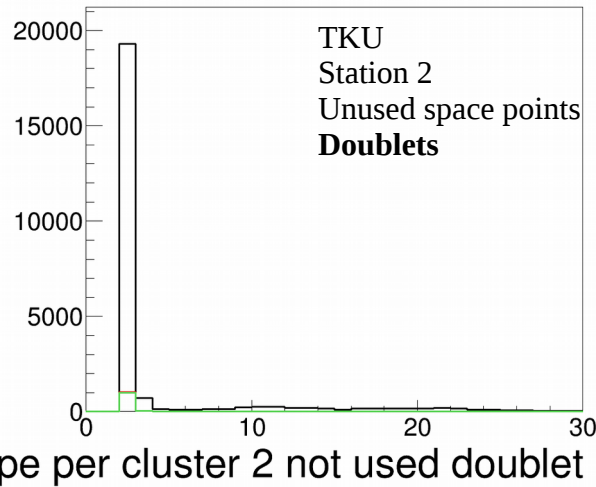
Two planes hit → doublet space point

Three planes hit → triplet space point

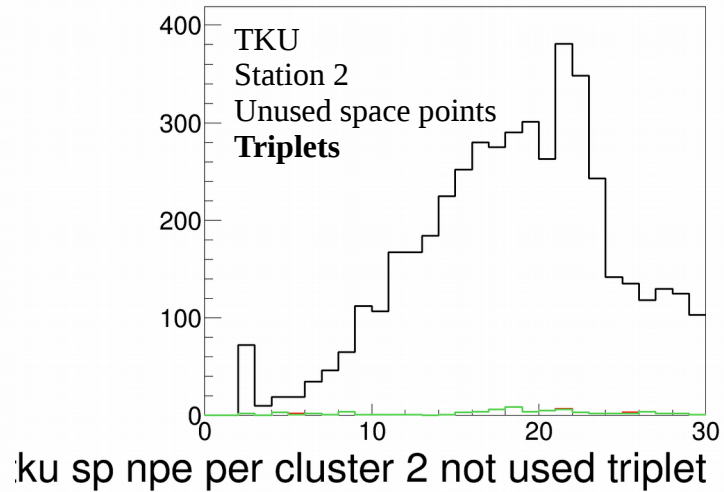
Number of Photoelectrons



2017-2.7 3-140 IH2 empty

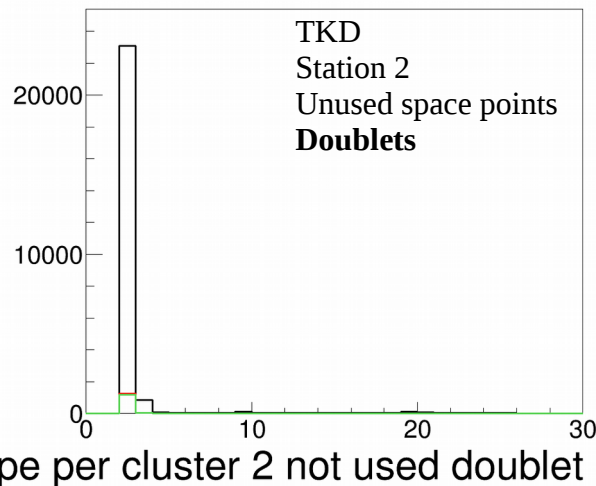


2017-2.7 3-140 IH2 empty

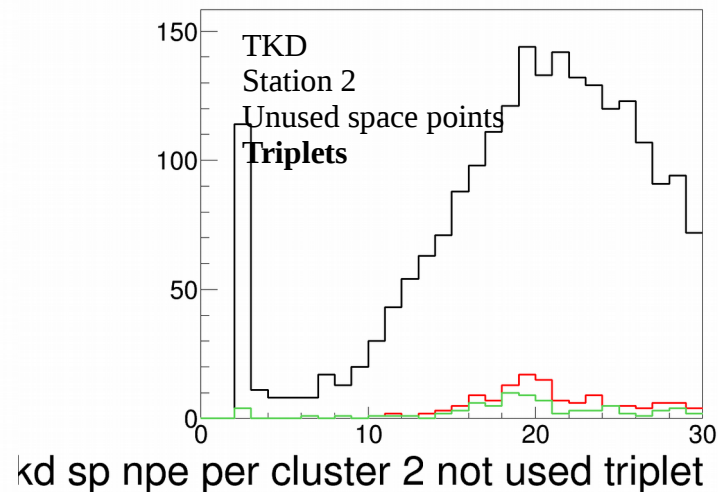


All data
Upstream sample
Downstream sample

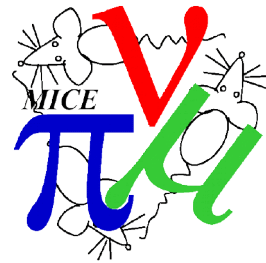
2017-2.7 3-140 IH2 empty



2017-2.7 3-140 IH2 empty

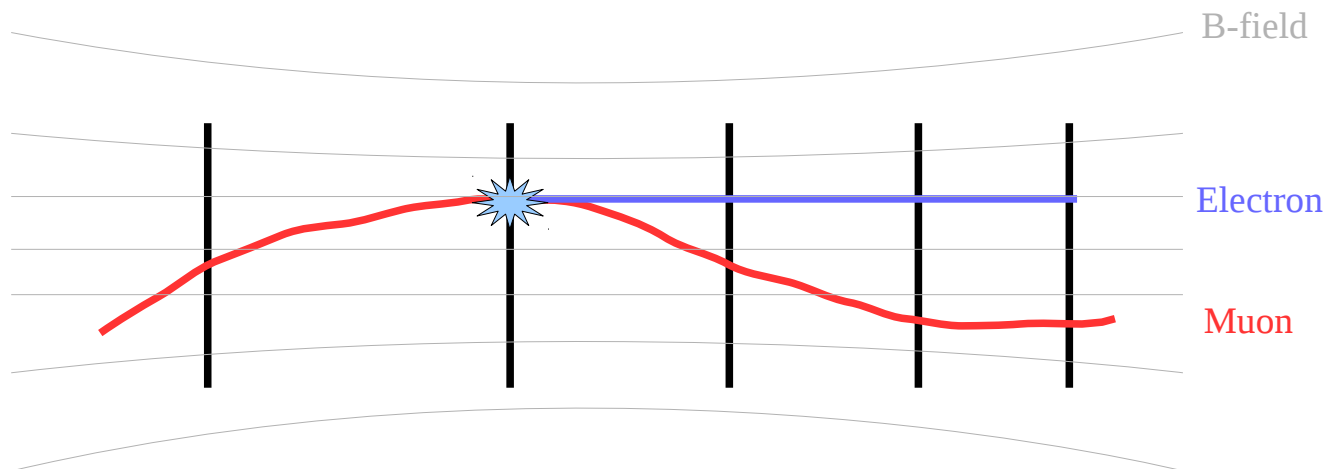
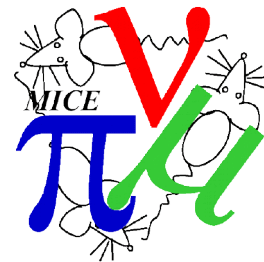


Triplet space points



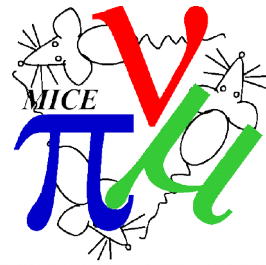
- Triplets are unlikely to be due to thermal noise
 - Unlikely to get coincidence of three planes unless the trackers are really noisy
 - → they are not
- What can cause triplets?
 - Something physical happening in the plane
 - Knock-on electrons?
 - Particle traversing the detector knocks an electron off one station and it hits the next station
 - Cross-talk in the scintillating fibre?
 - Light from a “real” hit goes into a neighbouring fibre
 - Cluster finding doesn't pick it up and we make extra space points

Knock-on electrons

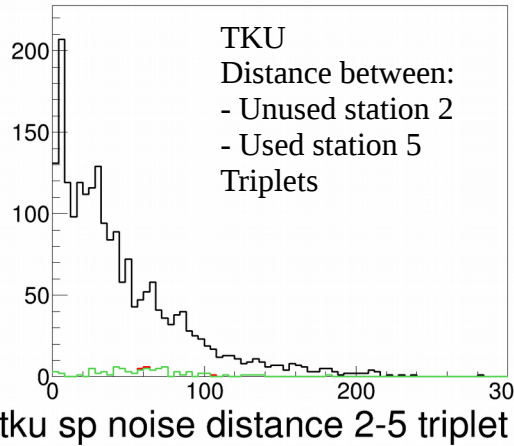


- Knock-on electron gets excited from station
- Traverses through several tracker planes
 - Travels along field lines! i.e. always pretty much straight
- Position of triplets should be correlated with space points in neighbouring stations
- As-opposed to cross-talk, which correlates space points in the same station
- Look at (x, y) distance, dr , in between “unused” space point and “used” space point in adjacent stations

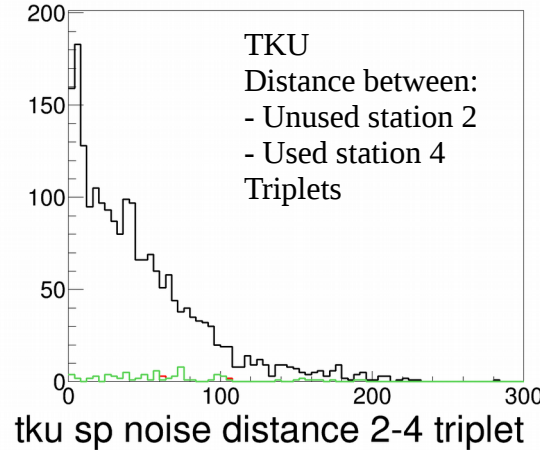
E.g. Noise in station 2 TKU



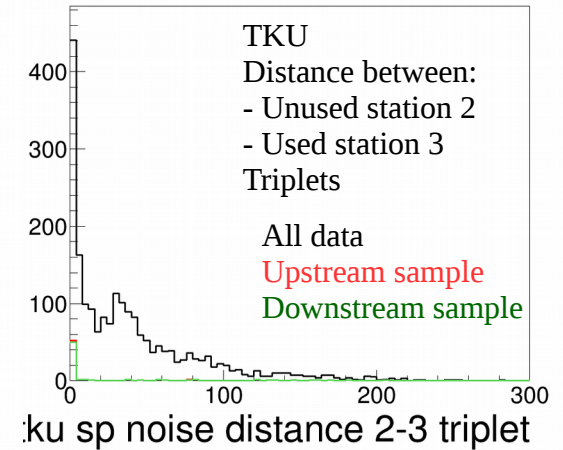
2017-2.7 3-140 IH2 empty



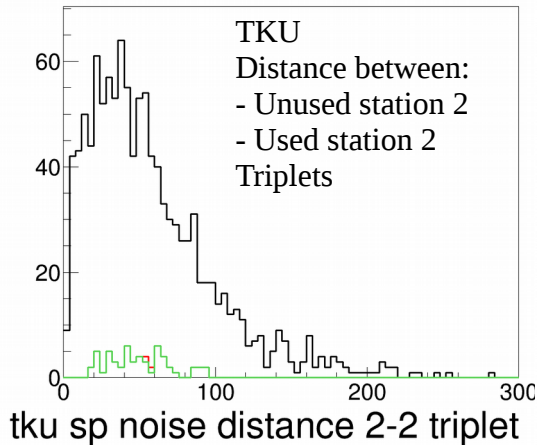
2017-2.7 3-140 IH2 empty



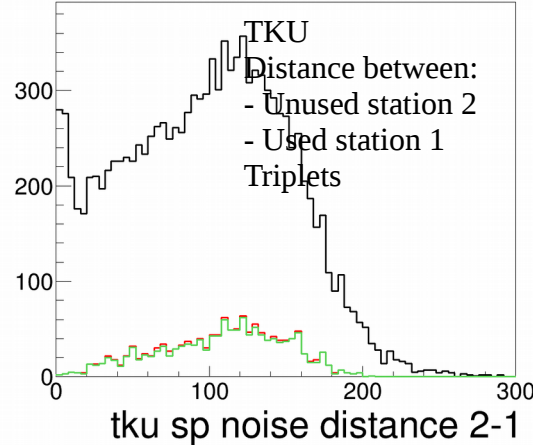
2017-2.7 3-140 IH2 empty



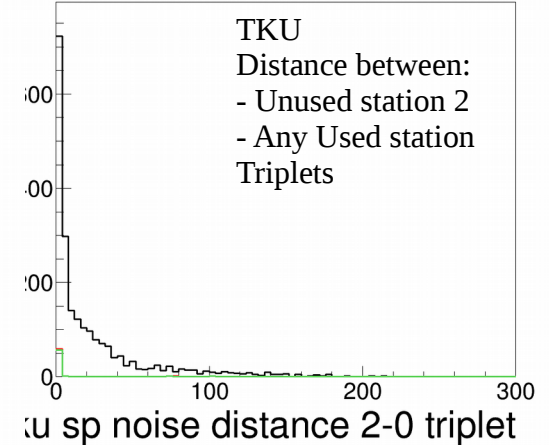
2017-2.7 3-140 IH2 empty



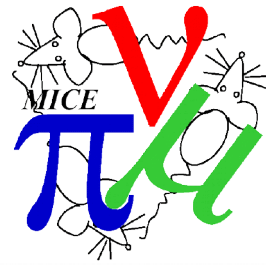
2017-2.7 3-140 IH2 empty



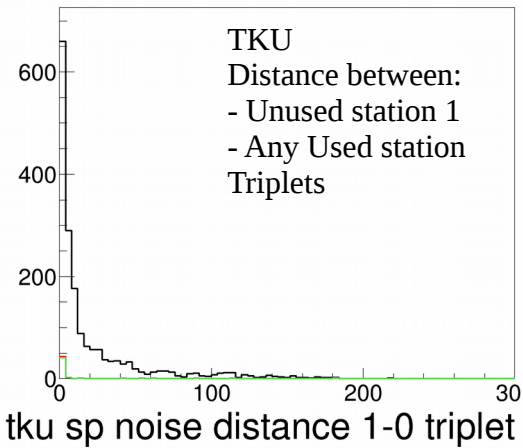
2017-2.7 3-140 IH2 empty



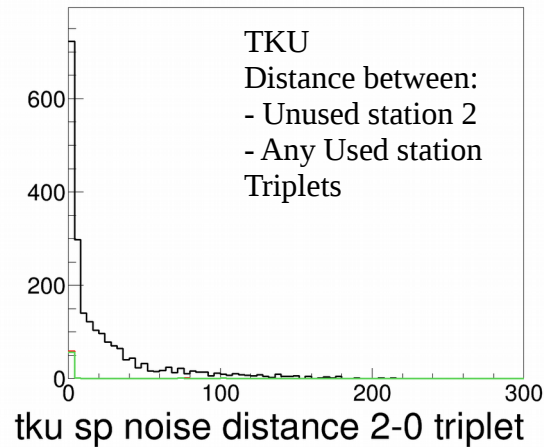
E.g. noise in all stations (TKU)



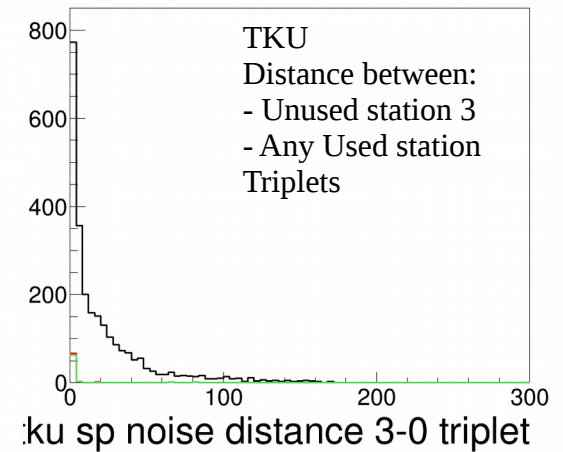
2017-2.7 3-140 IH2 empty



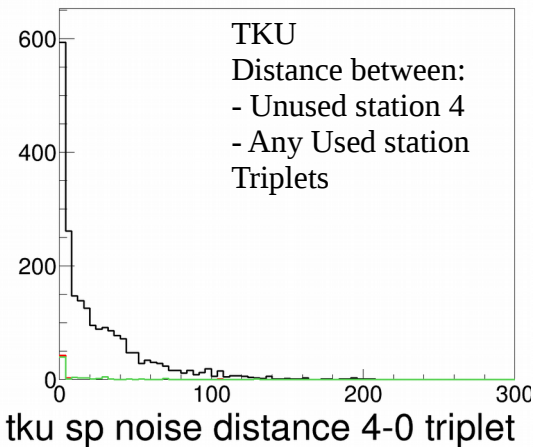
2017-2.7 3-140 IH2 empty



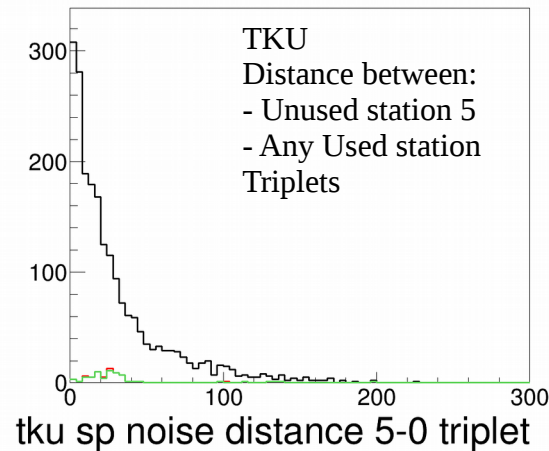
2017-2.7 3-140 IH2 empty



2017-2.7 3-140 IH2 empty

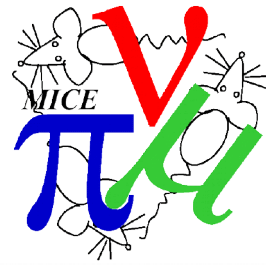


2017-2.7 3-140 IH2 empty

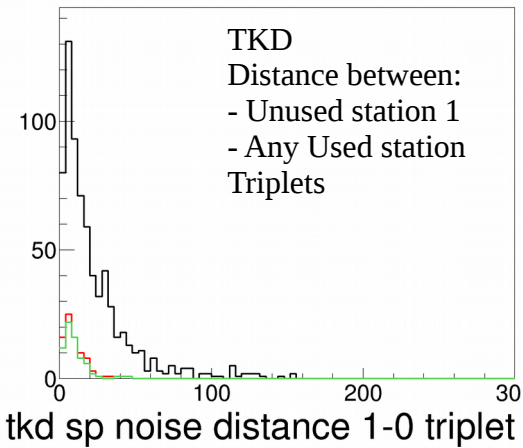


All data
Upstream sample
Downstream sample

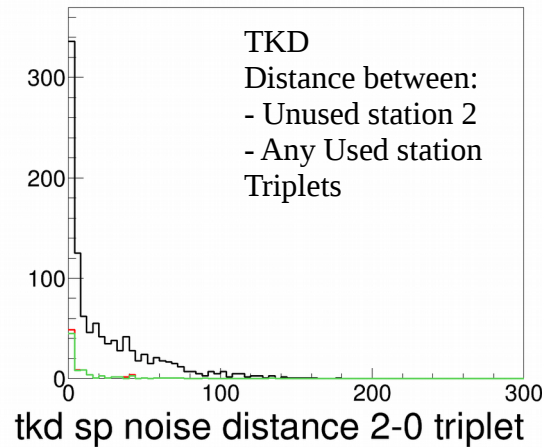
E.g. noise in all stations (TKD)



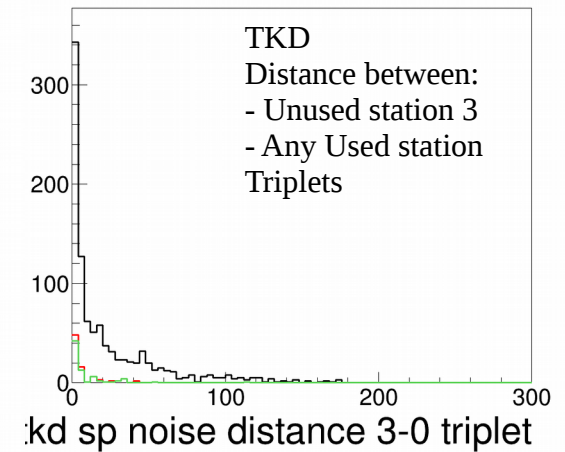
2017-2.7 3-140 IH2 empty



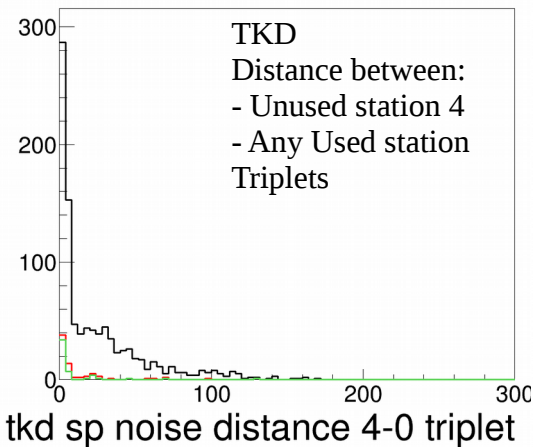
2017-2.7 3-140 IH2 empty



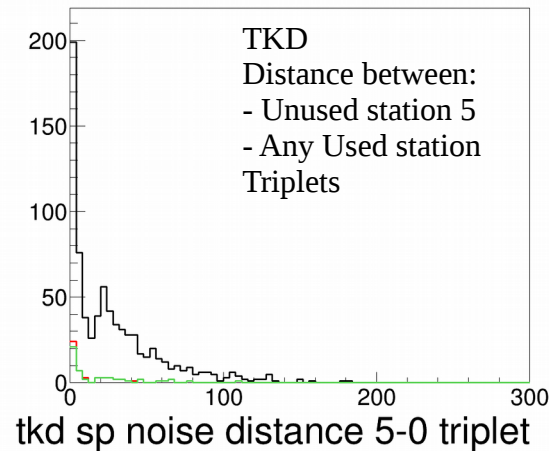
2017-2.7 3-140 IH2 empty



2017-2.7 3-140 IH2 empty

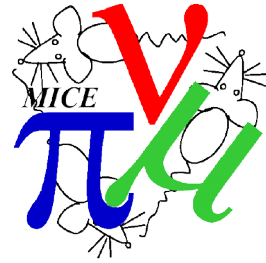


2017-2.7 3-140 IH2 empty



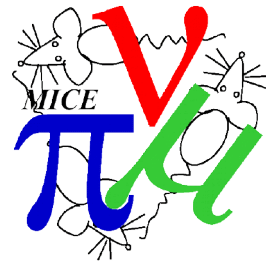
All data
Upstream sample
Downstream sample

Stats



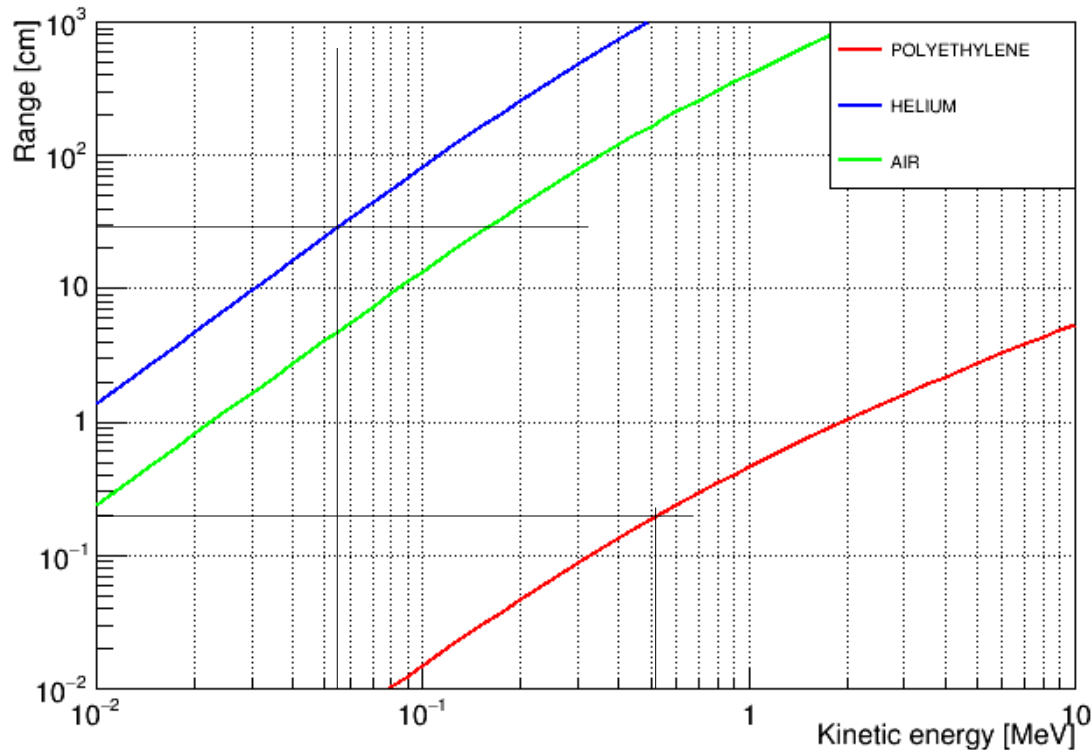
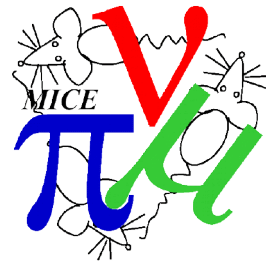
- ~100 “low dr unused space points” in the upstream sample per plane
 - Out of 9055 events
 - → ~500 per tracker in unused sample
 - Unknown number in used sample (enough to smear?)
 - ~80 % from upstream plane
 - ~20 % from upstream plane + 1

Conclusion



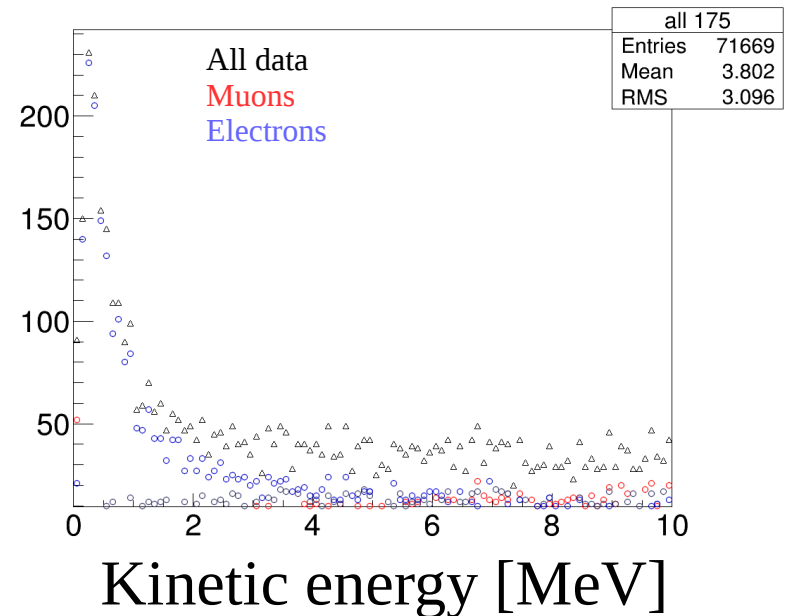
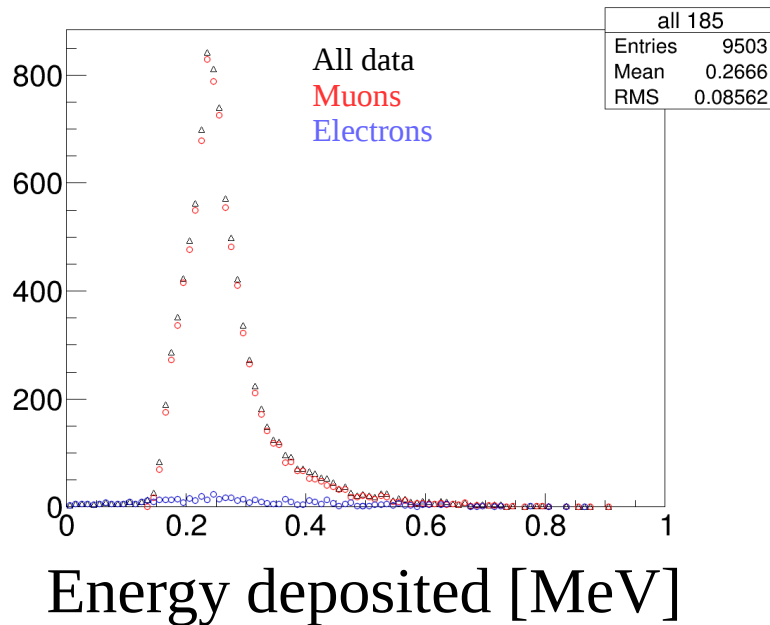
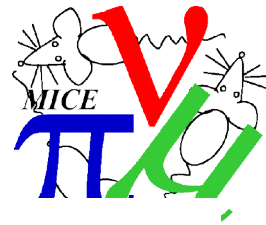
- Nearly all triplet noise can be attributed to knock-on electrons
 - Some of this noise is in the “good muon” sample
- A significant proportion of doublet noise can be attributed to thermal noise
 - None of this noise is in the “good muon” sample
 - Propose increase photoelectron cut to 4 photoelectrons
- Some doublet noise is unaccounted
 - Guess knock-on electrons are at fault
- Some features may still exist
 - But good place to start

Order of magnitude...



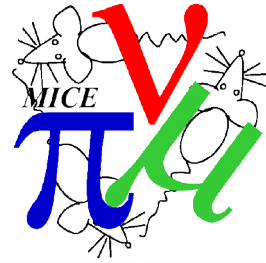
- Electron range in polyethylene, Helium (and air)
 - 30 cm He stops electrons with < 0.05 MeV
 - 0.2 cm polythene stops electrons with < 0.5 MeV

G4 Model for Delta Electrons

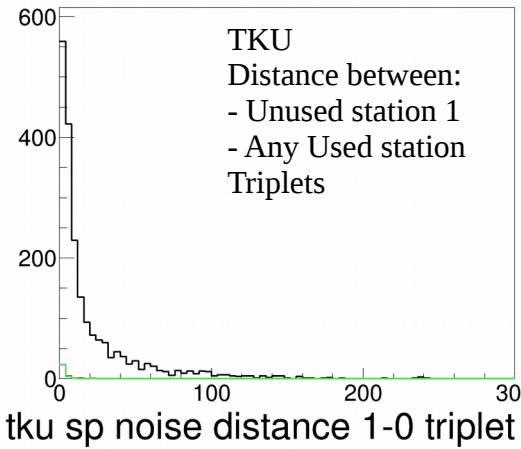


- Following some digging/bug fixing in MAUS
 - We have a few low energy electrons in MAUS
 - (About 1% of muons make an electron)

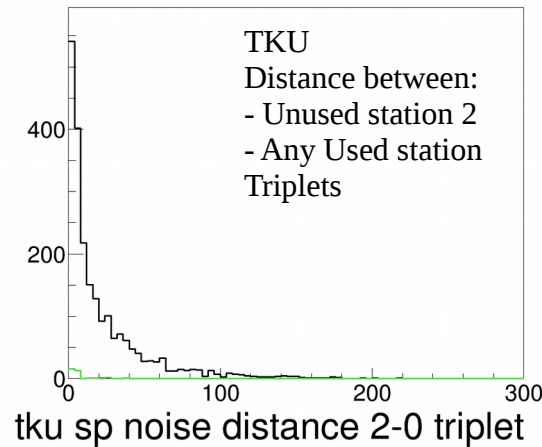
Simulated noise (TKU)



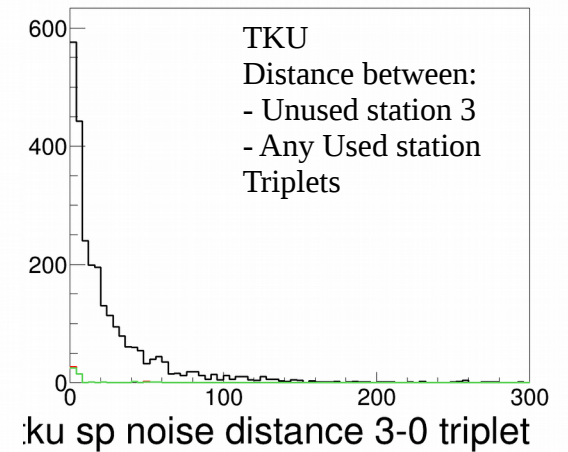
Simulated 2017-2.7 3-140 IH2 empty



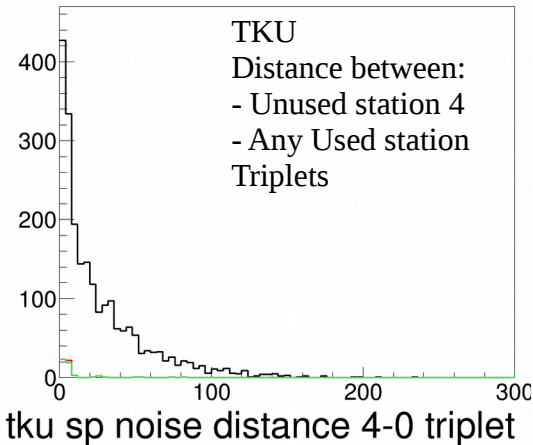
Simulated 2017-2.7 3-140 IH2 empty



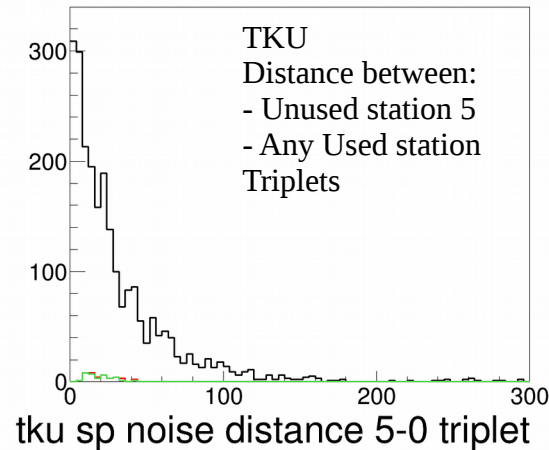
Simulated 2017-2.7 3-140 IH2 empty



Simulated 2017-2.7 3-140 IH2 empty

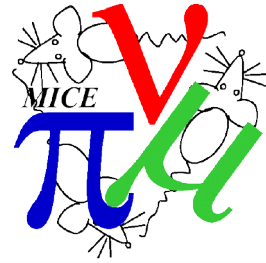


Simulated 2017-2.7 3-140 IH2 empty

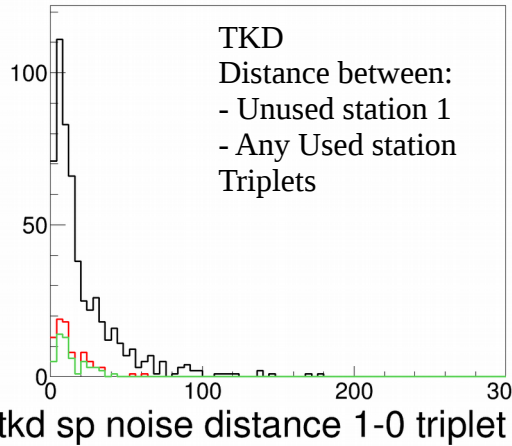


All MC
Upstream sample
Downstream sample

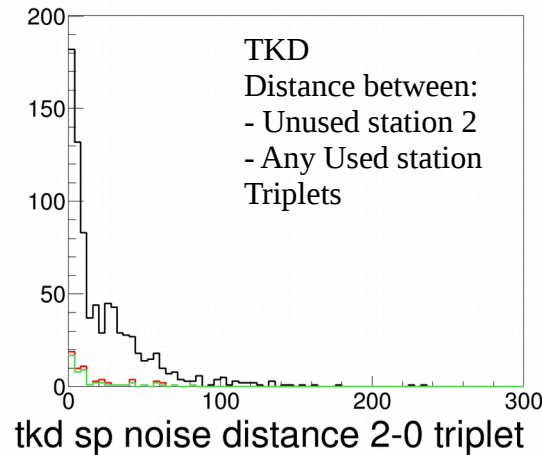
Simulated noise (TKD)



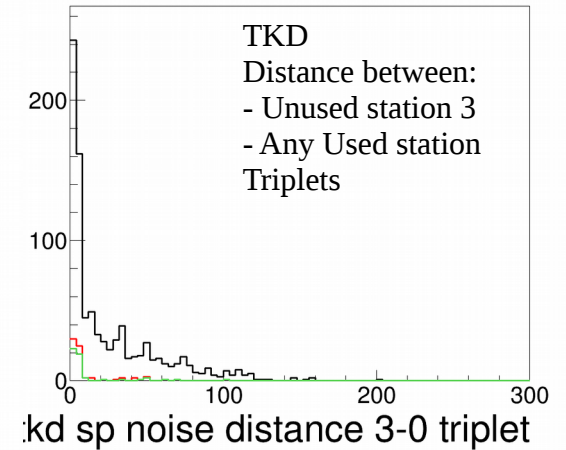
Simulated 2017-2.7 3-140 IH2 empty



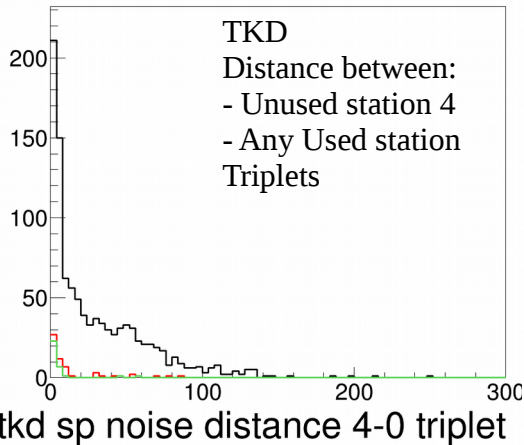
Simulated 2017-2.7 3-140 IH2 empty



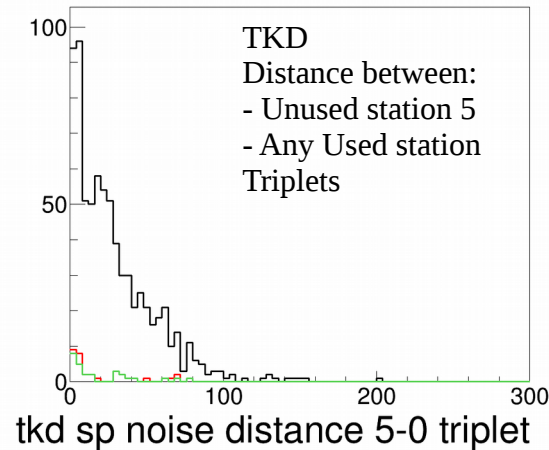
Simulated 2017-2.7 3-140 IH2 empty



Simulated 2017-2.7 3-140 IH2 empty

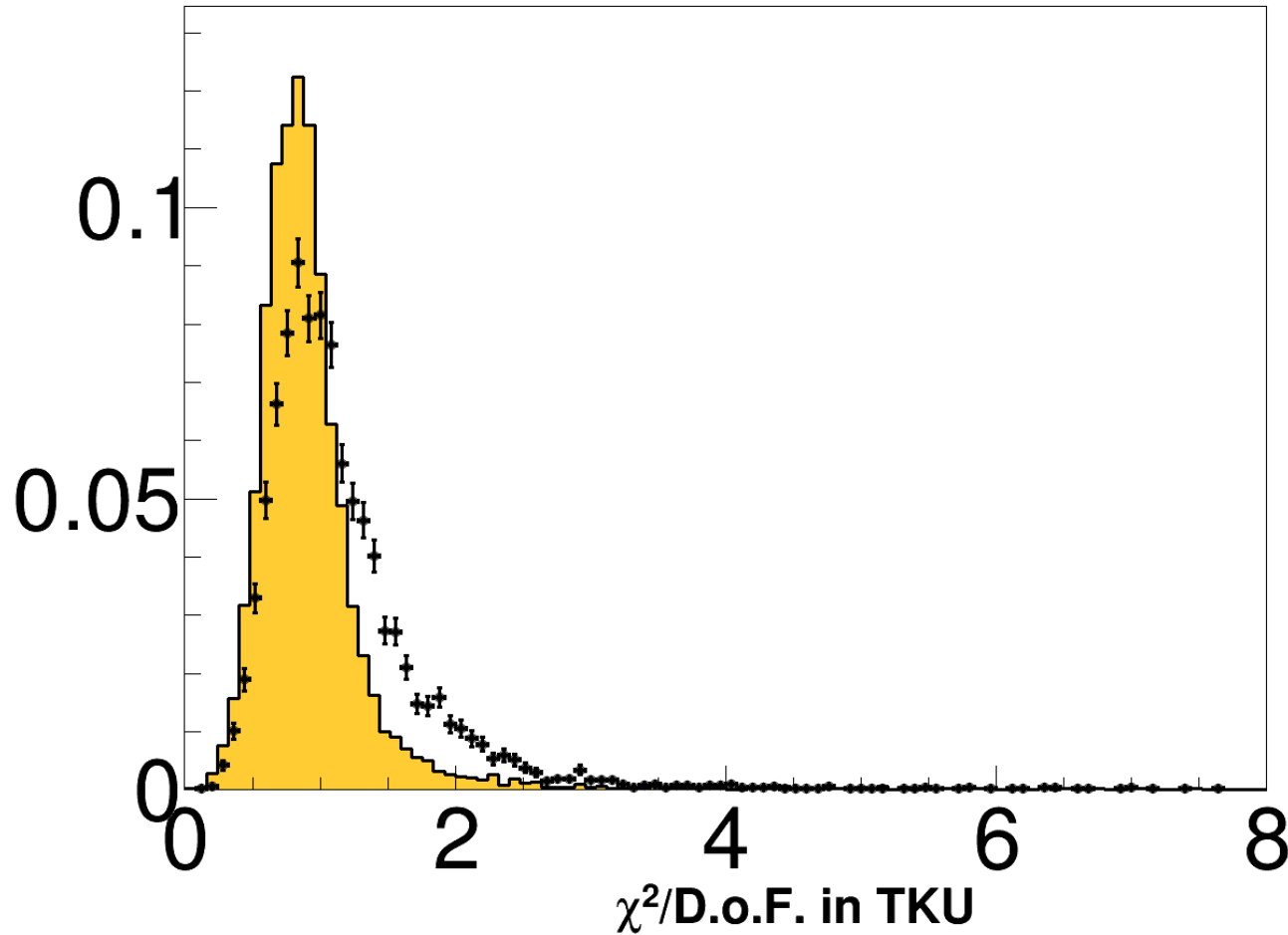
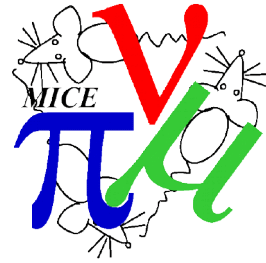


Simulated 2017-2.7 3-140 IH2 empty



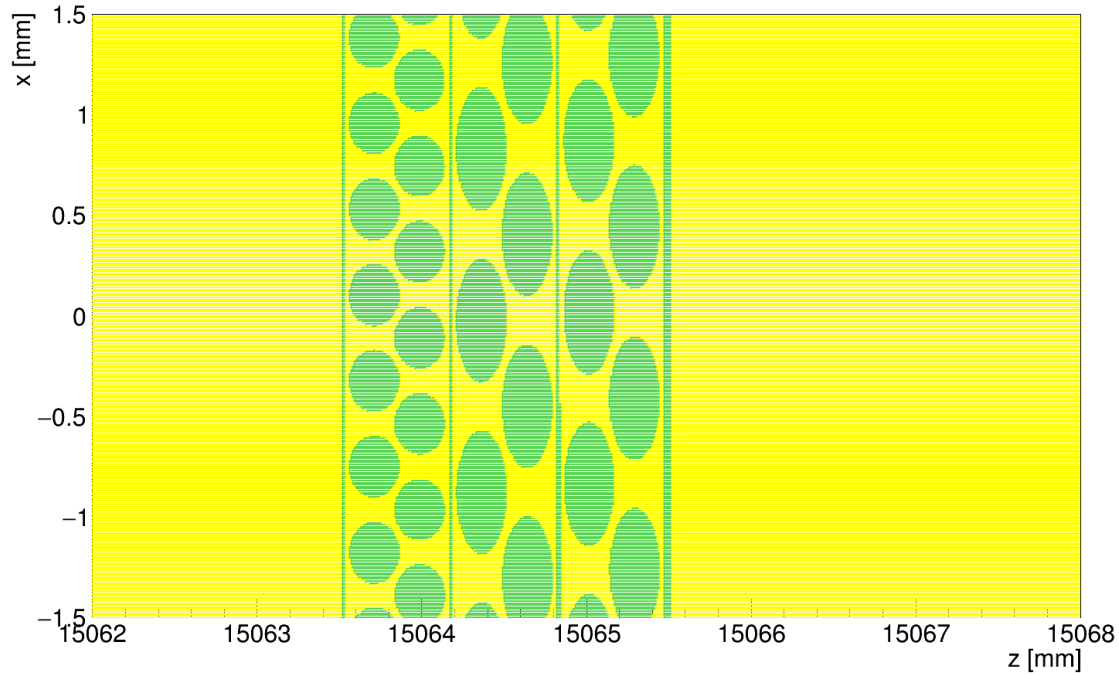
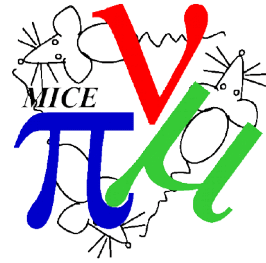
All MC
Upstream sample
Downstream sample

Chi2 - after updates



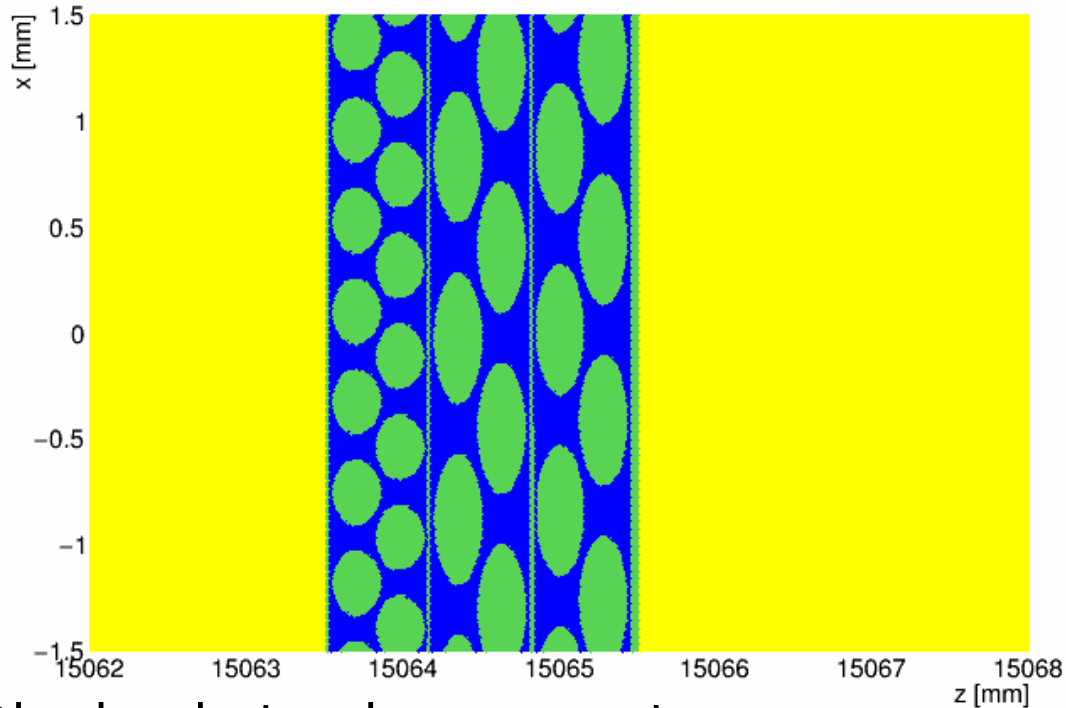
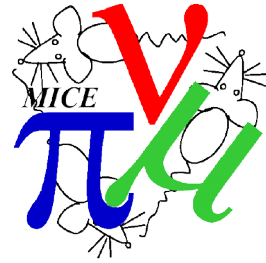
- Little/no improvement in the chi2
 - Keep digging...

Tracker Geometry



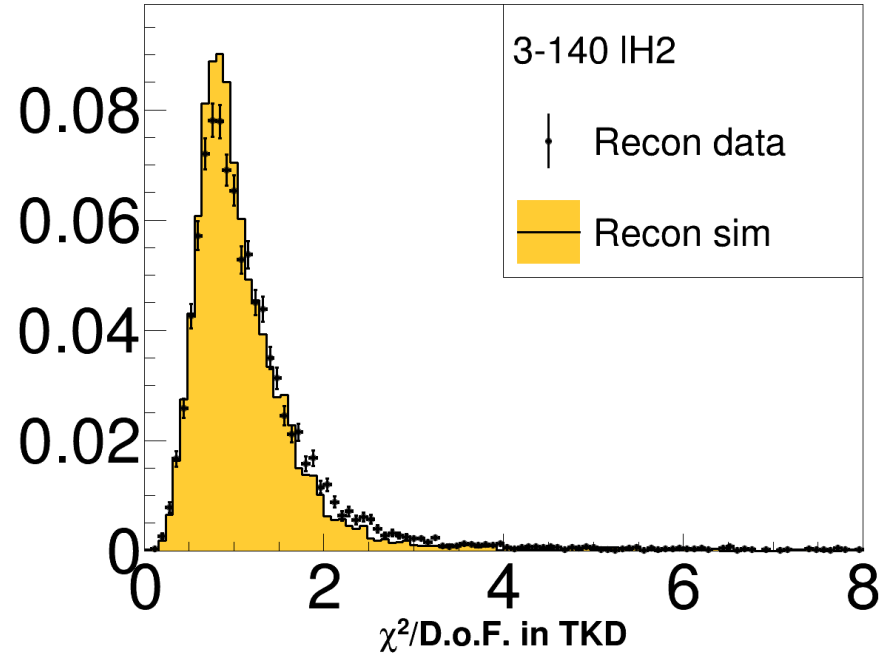
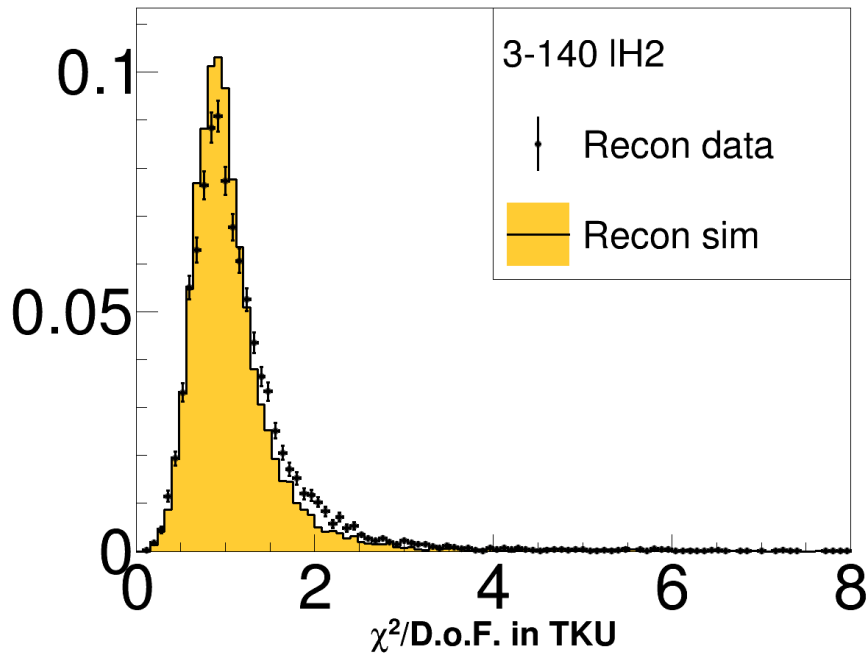
- Digging in tracker geometry
 - Yellow is He
 - Green is polystyrene

Tracker Geometry (Fixed)



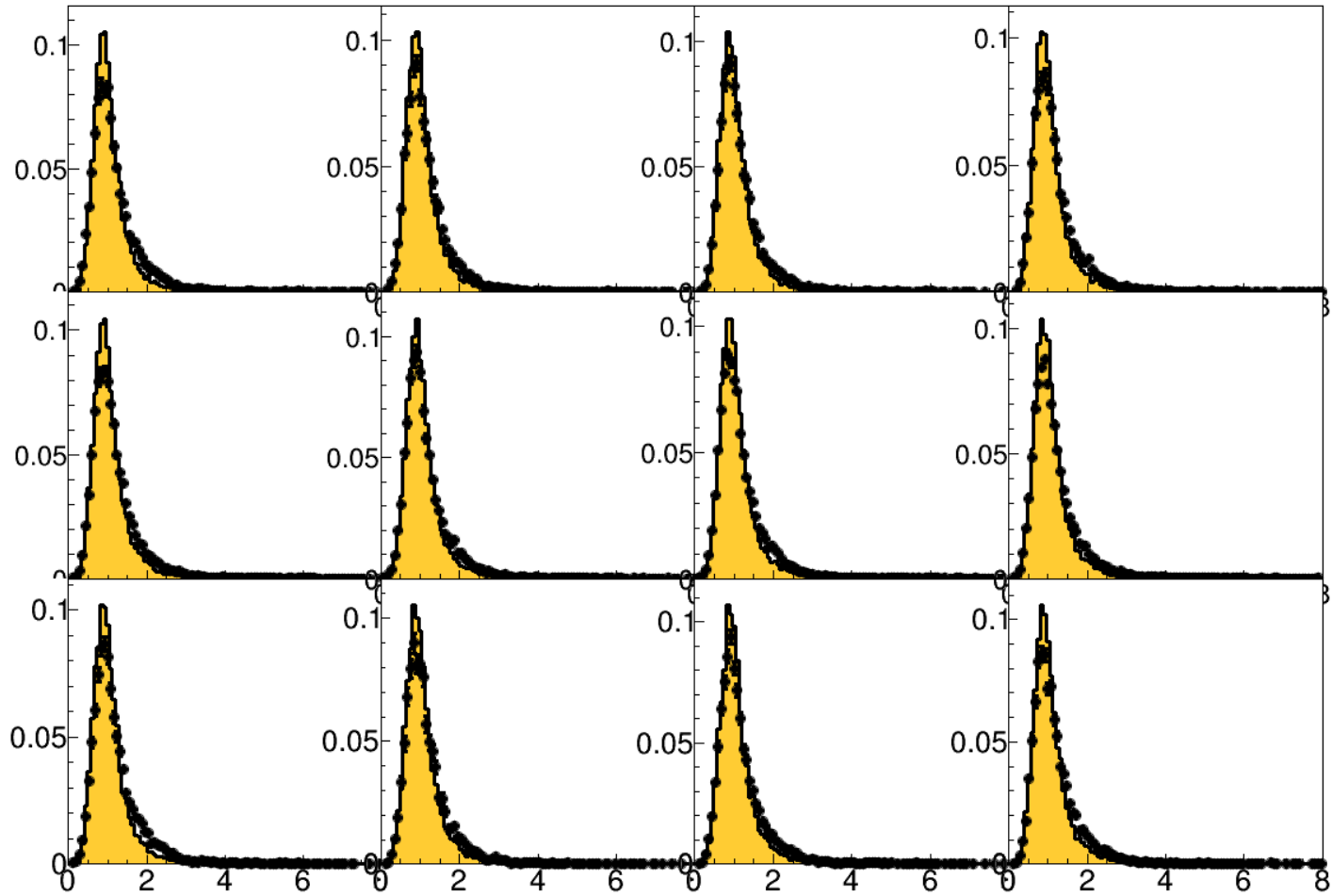
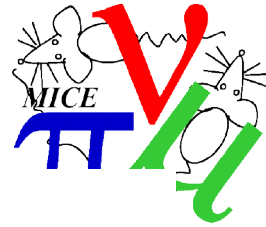
- Digging in tracker geometry
 - Yellow is He
 - Green is polystyrene
 - Blue is glue; optimised density to 2 g/cm^3

Tracker Geometry



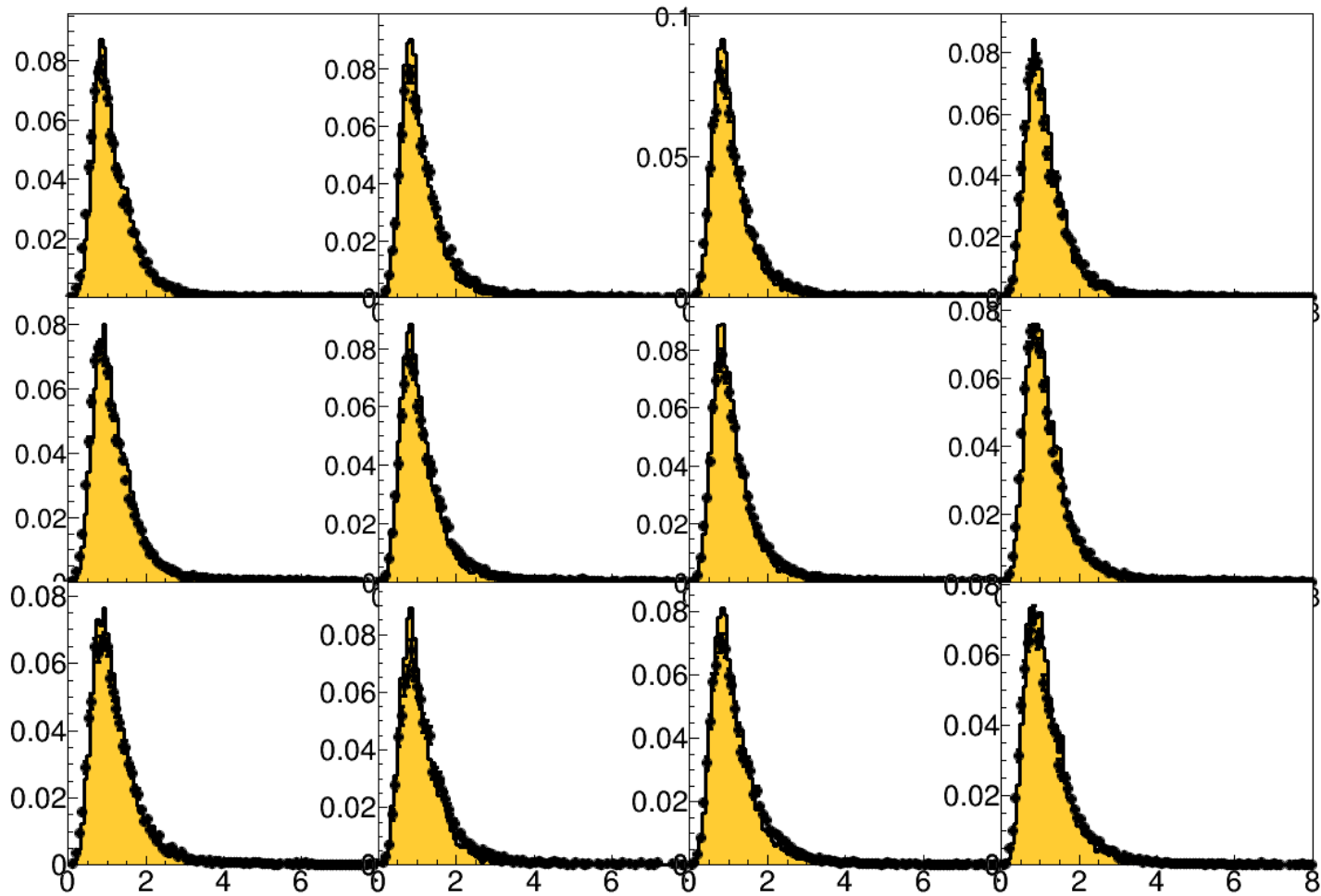
- Quite a bit better
 - Denser glue can improve the comparison
 - Is density $> 2 \text{ g/cm}^3$ physical?

TKU Chi2 following material fix



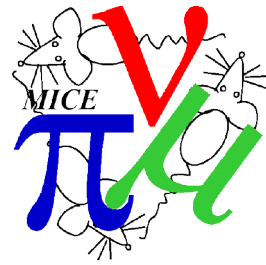
$\chi^2/\text{D.o.F. in TKU}$

TKD Chi2 following material fix



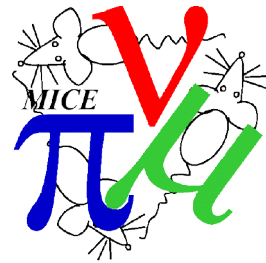
$\chi^2/\text{D.o.F. in TKD}$

Systematic Uncertainties



- Say the real experiment is different to our **nominal** model
 - How does that effect the reconstruction and analysis
- Sample and smear the beam at TKU
- Simulate using a **varied** geometry
 - e.g. Translate the tracker
 - e.g. Tilt the tracker
 - e.g. Vary End field
- Reconstruct using the **nominal** geometry
- Calculate the **varied** correction from MC → Recon
- Systematic uncertainty is the difference between:
 - **nominal** correction (with **nominal** geometry)
 - **Varied** correction and (with procedure above)

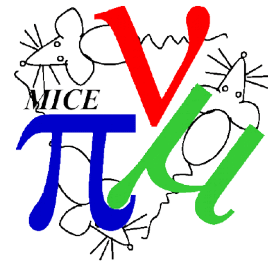
Systematic Uncertainties



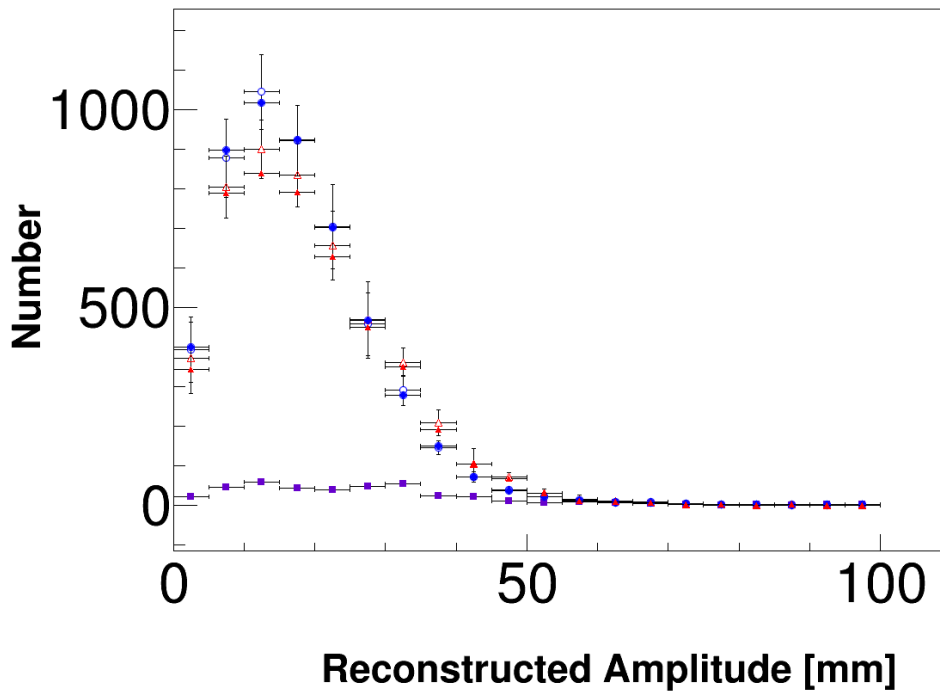
E.g. for 6-140 IH2 empty data

		2.5	7.5	12.5	17.5	22.5	27.5
tku position	1 mm	0.05	0.09	-0.03	-0.01	-0.01	-0.09
tku rotation	1 mrad	0.04	0.04	-0.02	0.01	0.00	-0.07
tku E2 coil	5.00%	0.06	0.04	-0.02	0.01	-0.03	-0.07
tkd position	1 mm	-0.08	0.07	-0.02	0.04	0.00	-0.12
tkd rotation	1 mrad	-0.02	0.03	0.06	0.06	0.05	-0.06
tkd E2 coil	5.00%	-0.04	0.03	0.00	0.05	-0.04	-0.06
		32.5	37.5	42.5	47.5	52.5	57.5
tku position	1 mm	-0.01	0.03	-0.07	-0.08	-0.10	0.08
tku rotation	1 mrad	0.03	0.00	-0.11	0.05	-0.05	0.00
tku E2 coil	5.00%	0.06	-0.02	-0.04	-0.05	-0.05	-0.08
tkd position	1 mm	-0.05	0.03	0.03	-0.02	-0.17	-0.40
tkd rotation	1 mrad	0.01	0.03	0.14	-0.03	0.00	-0.10
tkd E2 coil	5.00%	0.03	-0.05	0.10	0.02	0.10	-0.50

Systematic Uncertainties

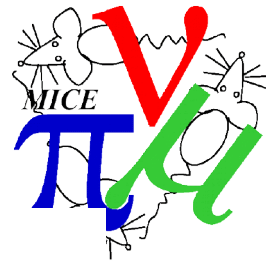


E.g. for 6-140 IH2 empty data

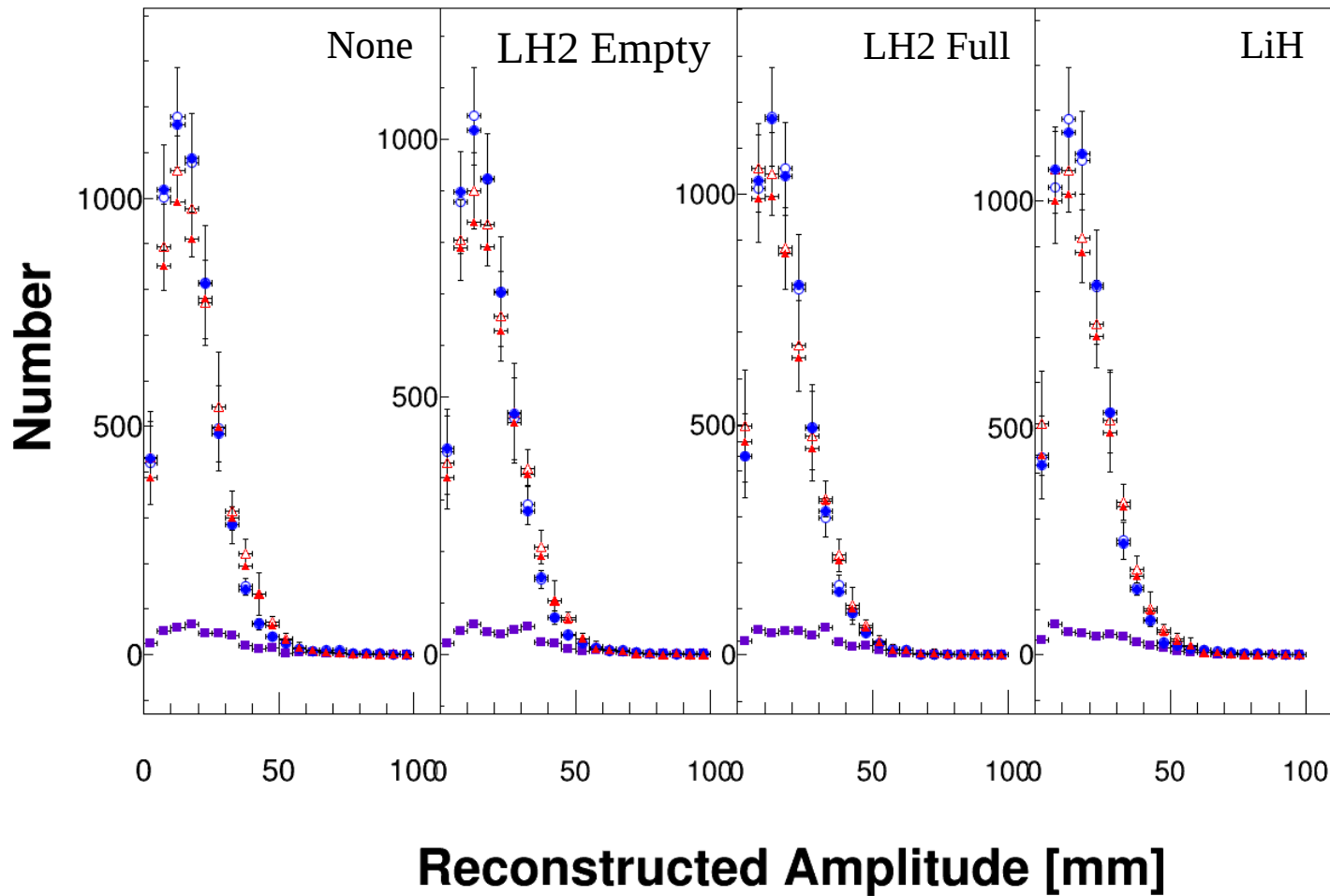


- Statistical and systematic errors added in quadrature
 - Look at the open points
 - Blue is upstream
 - Red is downstream

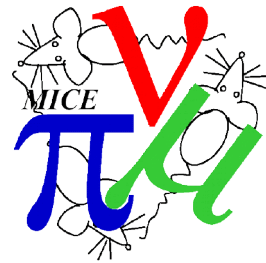
Systematic Uncertainties



6-140 mm data



Conclusions



- Tracker model looks much more robust
 - Model for noise
 - Material model
- Systematic error calculation discussed
 - Needs a second pass on the calculations
 - Plot systematic uncertainty separate to statistical
 - Add in 3-140, 10-140 data
 - More statistics
- Try to beat down the uncertainty
 - Apply Chris Hunt algorithm to rotation
 - Apply improved field map
- Conscious of timeline to IPAC...