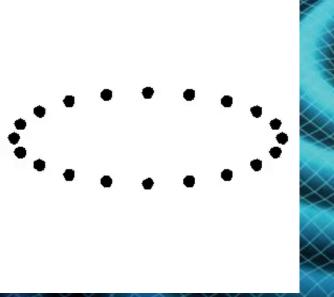




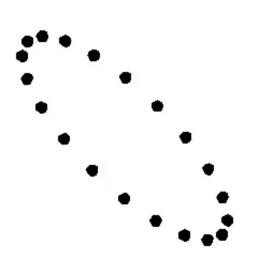
## Advanced LIGO computing requirements

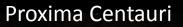
Stephen Fairhurst & Paul Hopkins

# Accelerating Mass produces Gravitational Waves





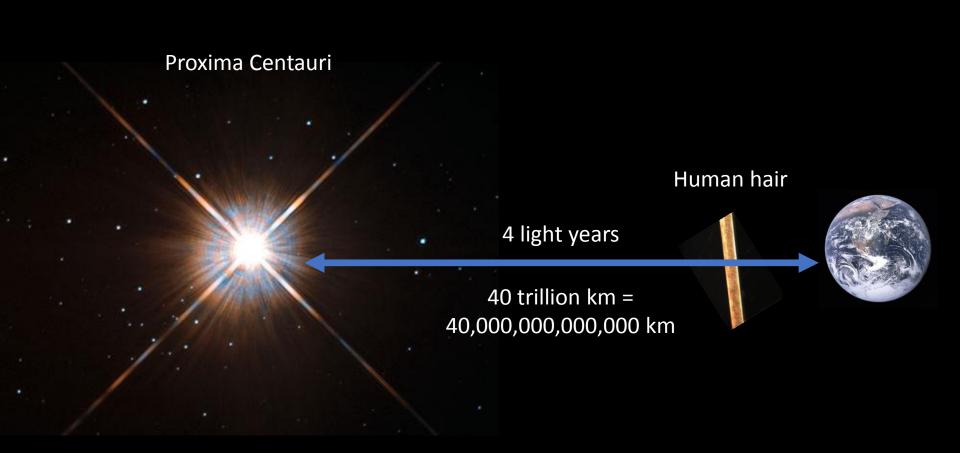




4 light years

40 trillion km = 40,000,000,000,000 km





WARNING: Not to scale!

#### LIGO Hanford Observatory

#### LIGO Hanford

13-

LIGO Livingston

#### LIGO Hanford

LIGO Livingston

Operational Under Construction Planned

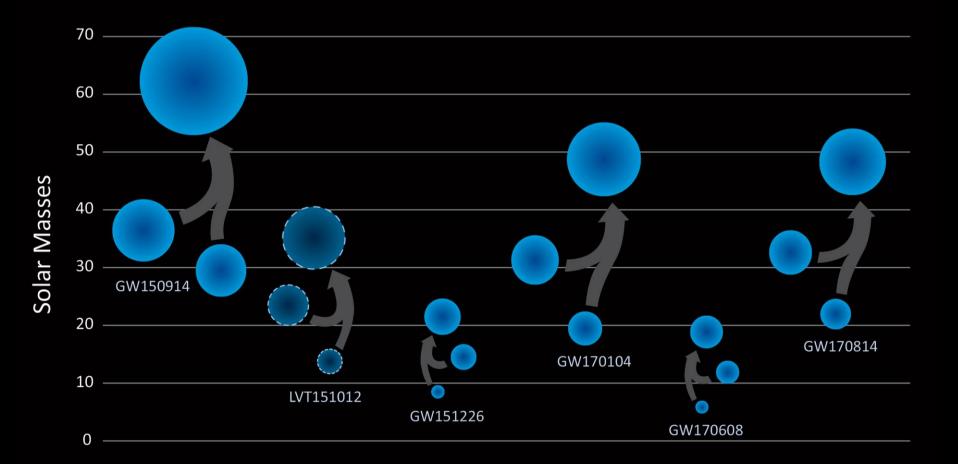
### **Gravitational Wave Observatories**

GEO600

Arge

KAGRA

**LIGO India** 



#### LIGO/VIRGO

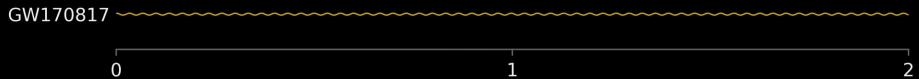


#### LVT151012 ~~~~~

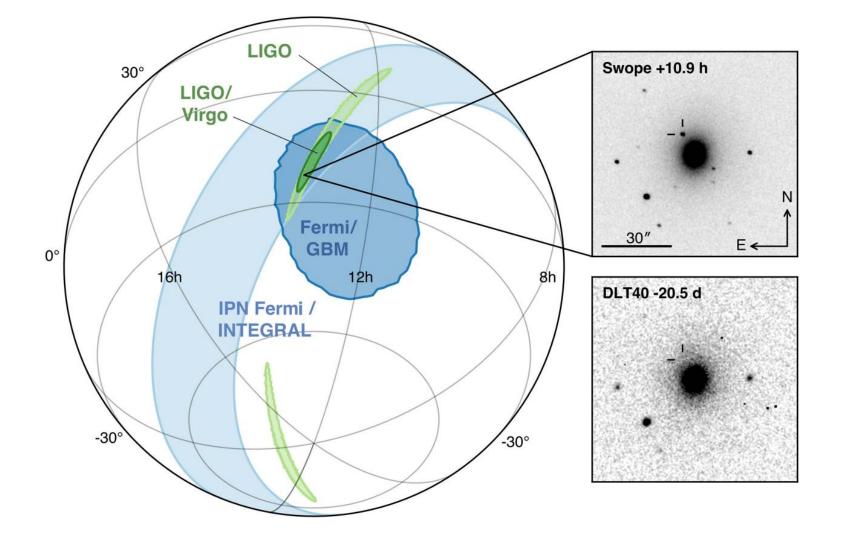


#### 

GW170814 /////////

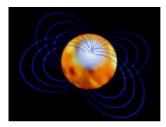


#### time observable (seconds)



### **Computational Challenges**

### Gravitational Wave Sources/Searches



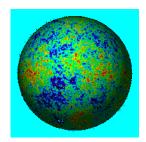
**Continuous Waves:** Spinning Neutron Stars



#### **Coalescing Binaries:** Merging black holes and neutron stars

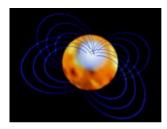


**Short Bursts:** Supernovae, Gamma Ray Bursts



**Stochastic Background:** GW from the Big Bang or Astrophysical sources

### Gravitational Wave Sources/Searches



Known waveform: matched filter Spinning Neutron Stars



...escing Binaries: Merging black holes and neutron stars

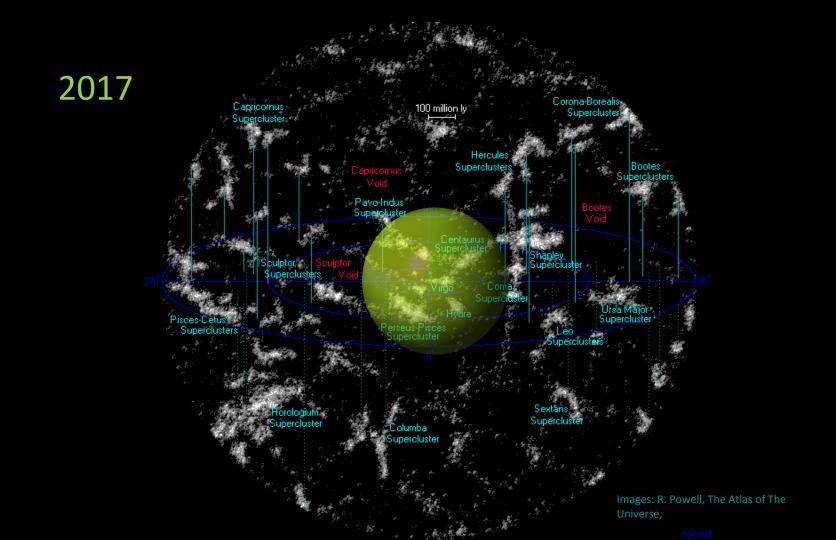


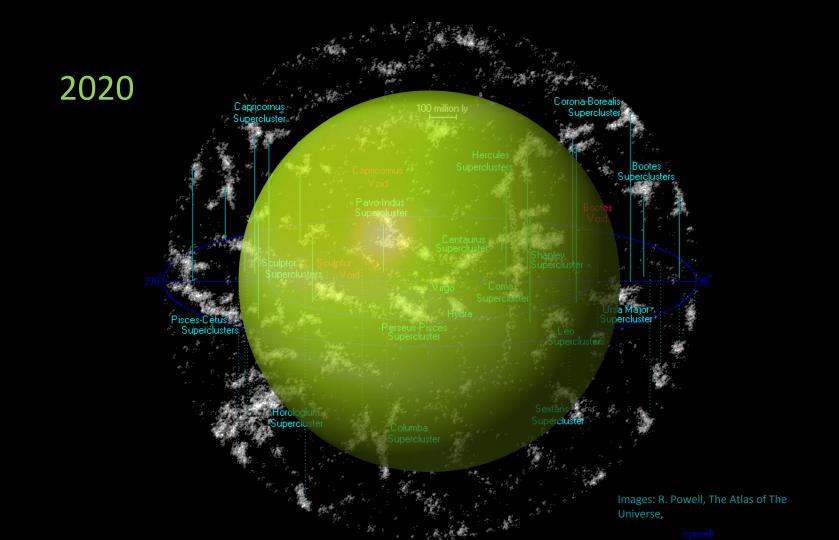
Unknown waveform: cross correlate data **Suc Background:** GW from the Big Bang or Astrophysical sources

## **Computational challenges**

(Mostly) 'embarrassingly parallel' high throughput computing, ~ 1 Petaflop

- Einstein@home project for Continuous Waves all sky search;
  ~ 2.5 Petaflops
- Dominant use on LIGO/Virgo resource:
  - binary search using 300,000 templates. Number and length of templates increases with sensitivity.
  - parameter estimation on events. Scales with number of observed events. Repeat with various waveform models and physical effects incorporated.
  - results in low latency required for electromagnetic followup

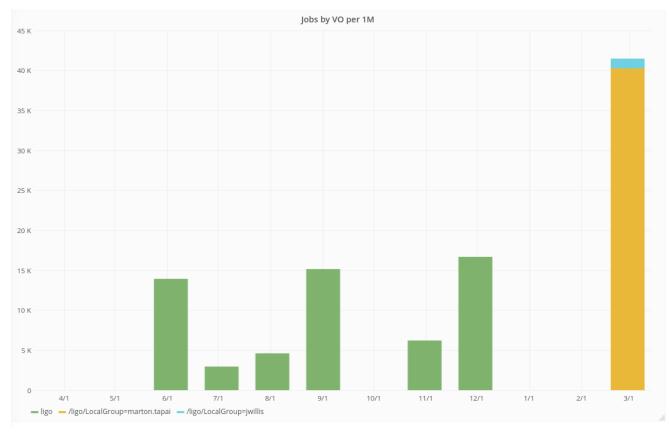




## Production computing environments

- Current Baseline: Dedicated/committed LIGO clusters, and Einstein@Home
- Increasing use of other resources:
  - US Open Science Grid
  - European Virgo clusters
  - UK GridPP Tier 0
- Multiple pipelines adapted to take advantage of GPUs
  - Some LSC clusters have a large number of GPUs, e.g. 2400 at Max Planck Institute
- Containerization light-weight virtualization
- Federated access to resources, including computational clusters, but also e.g. data, web pages, services etc.

### LIGO Jobs running at RAL



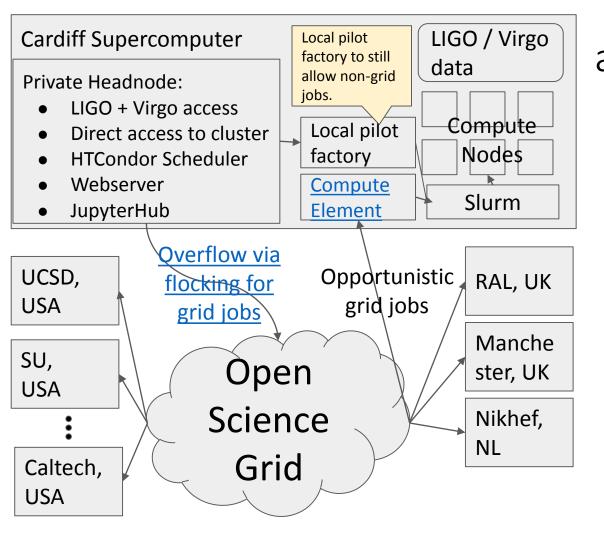
### LIGO Computing Resources

Resources are distributed around the world and inhomogeneous:

- Lab and LSC committed clusters
  - Standard operating system
  - dependencies and software installed on system
  - condor scheduler
- LSC general purpose clusters and Virgo clusters
  - Shared with university or other organization
  - Variety of OS and schedulers
  - Challenges managing software and dependencies
  - Extra hurdles in obtaining accounts and accessing resources

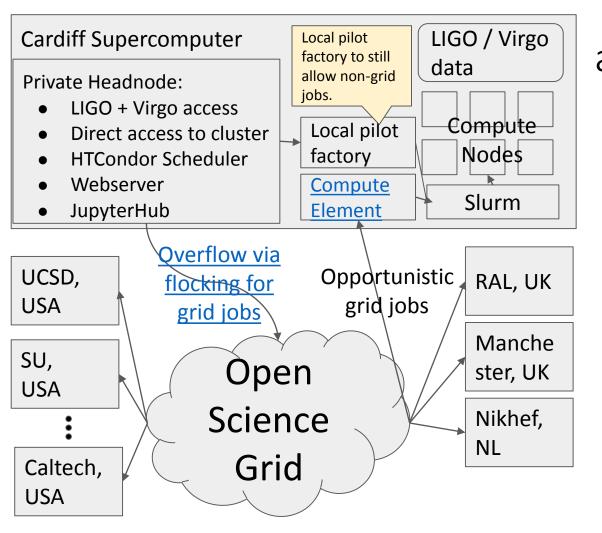
### Data rates

Data Type	File Rate	Data Rate (After Compression)
Raw	492,750 files/yr	10.00 MB/s
Second-Trend	52,560 files/yr	1.10 MB/s
Minute-Trend	8,760 files/yr	0.08 MB/s
RDS	123,188 files/yr	2.00 MB/s
Calibrated Strain	123,188 files/yr	0.12 MB/s
SFTs	17,520 files/yr	0.02 MB/s
Single Interferometer Total:	818000 files/yr	13.32 MB/s
Both LIGO Interferometers:	1.6M files/yr	800 TB/yr



## aLIGO UK upgrade

- Traditional HPC cluster
- Slurm (Currently PBSPro)
- NFS Home, 500GB quota
- Lustre /scratch filesystem
- LIGO/Virgo data on private
  Gluster storage and most
  recent 20TB on /scratch
- Dedicated Headnode with HTCondor and local pilots



## aLIGO UK upgrade

- Retain local pilot factory
- Use cvmfs for data paths utilising locally stored data
- Add a local compute element to accept jobs from OSG
- Overflow grid jobs to OSG via flocking
- Encourage users to grid enable their workflows
- Aim to remove local pilot factory?