

Goldstone fields with spins higher than $1/2$

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- Typically, higher-spin gauge theories are based on infinite-dimensional symmetries with infinite number of fields of increasing spin. Study of spontaneous symmetry breaking in these theories is highly non-trivial.
- We consider simple 3D models in which massless spin-1 and spin-3/2 fields play the role of the Goldstones of a spontaneously broken higher-spin generalization of supersymmetry algebra.

Higher Spin Algebras

Hietarinta Higher-Spin Algebra - FINITE

Fermionic Superalgebra:

$$\{Q_{\alpha}^{a_1 \dots a_n}, Q_{\beta}^{b_1 \dots b_m}\} = f_{\alpha\beta}^{a_1 \dots a_n, b_1 \dots b_m, c} P_c; [Q_{\alpha}^{a_1 \dots a_n}, P_c] = 0.$$

Bosonic Algebra:

$$[S^{a_1 \dots a_p}, S^{b_1 \dots b_q}] = f^{a_1 \dots a_p, b_1 \dots b_q, c} P_c; [S^{a_1, \dots, a_n}, P_c] = 0.$$



Hietarinta

Conventional Higher-Spin Superalgebra - INFINITE

$$[T_{s_1}, T_{s_2}] = T_{s_1+s_2-2} + T_{s_1+s_2-4} + \dots + T_{|s_1-s_2|+2}$$

Unlike the conventional HS superalgebra, Hietarinta superalgebra does not produce more HS generators of ever-increasing spin.

Volkov-Akulov Lagrangian

Dmitrij Vasilievich Volkov
and **Vladimir P. Akulov**
developed the
Volkov-Akulov Lagrangian
formalism. It realizes
supersymmetry non-linearly.



Volkov



Akulov

Volkov-Akulov formalism enables
the construction of the Lagrangian
of the Goldstone field. It can be
generalised for all higher-spin fields
associated with Hietarinta algebra.

Chern-Simons Goldstone Model

Spin-1

Chern-Simons Goldstone Model

Bosonic algebra: $[S^a, S^b] = 2 \varepsilon^{abc} P_c$, $[S^a, P_b] = 0$,

where S^a is the bosonic vector operator.

$$\begin{aligned} S_1 &= -f^2 \int d^3x (\det E_d^a - 1) \\ &= \int d^3x \left(\underbrace{\varepsilon^{abc} A_a \partial_b A_c}_{\text{Chern-Simons}} - \frac{f^{-2}}{2} \varepsilon^{abc} \varepsilon^{def} A_a A_d \partial_e A_b \partial_f A_c \right) \end{aligned}$$

Stueckelberg Trick

$$A_a \rightarrow \hat{A}_a = A_a - f^{\frac{1}{2}} \partial_a \hat{\varphi}$$

\hat{A}_a is invariant under $\delta A_a = \partial_a \lambda$ and $\delta \hat{\varphi} = f^{-\frac{1}{2}} \lambda$.

In the decoupling limit, the Stueckelberg Lagrangian reduces to

$$\mathcal{L}_{f \rightarrow \infty} = \varepsilon^{abc} A_a \partial_b A_c - \hat{\varphi} ((\square \hat{\varphi})^3 - 3 \square \hat{\varphi} \partial_a \partial^b \hat{\varphi} \partial_b \partial^a \hat{\varphi} + 2 \partial_a \partial^b \hat{\varphi} \partial_b \partial^c \hat{\varphi} \partial_c \partial^a \hat{\varphi}))$$

This is the same as the quartic term in the Galileon Lagrangian in modified gravity.

Rarita Schwinger Goldstino Model

Spin-3/2

Rarita-Schwinger Goldstino Model

Spin 3/2 superalgebra: $\{Q_\alpha^a, Q_\beta^b\} = 2 C_{\alpha\beta} \varepsilon^{abc} P_c, \quad [Q_\alpha^a, P_b] = 0$

The spin-3/2 goldstino action has the following form:

$$S_{3/2} = \int d^3x \left(\underbrace{i \varepsilon^{abc}}_{\text{Rarita-Schwinger}} \chi_a \partial_b \chi_c + \frac{f^{-2}}{2} (\varepsilon^{abc} \varepsilon^{dfg} - \varepsilon^{afc} \varepsilon^{dbg}) (\chi_a \partial_b \chi_c) (\chi_d \partial_f \chi_g) \right. \\ \left. + \frac{i f^{-4}}{6} \varepsilon^{a'b'c'} (\varepsilon^{abc} \varepsilon^{def} - \varepsilon^{abf} \varepsilon^{dec}) (\chi_c \partial_{a'} \chi_f) (\chi_a \partial_{b'} \chi_b) (\chi_d \partial_{c'} \chi_e) \right)$$

Stueckelberg Trick

$$\hat{\chi}_a = \chi_a + f^{\frac{2}{3}} \partial_a \psi$$

$\hat{\chi}_a$ is invariant under $\delta\chi_a = \partial_a \epsilon$ and $\delta\psi = -f^{-\frac{2}{3}} \epsilon$.

In the limit $f \rightarrow \infty$ the Stueckelberg Lagrangian reduces to

$$\mathcal{L}_{f \rightarrow \infty} = i \varepsilon^{abc} \chi_a \partial_b \chi_c + 2 \varepsilon^{abc} \varepsilon^{dfg} (\chi_a \partial_d \partial_c \psi) (\partial_f \psi \partial_b \partial_g \psi) - \frac{1}{3} \text{tr} (M^3)$$

where $M^a_d = i \varepsilon^{abc} \partial_b \psi \partial_d \partial_c \psi$.

Full Spin-3/2 Action as Rarita-Schwinger Action

The action reduces to the free Rarita-Schwinger action

$$S_{RS} = i \int d^3x \varepsilon^{abc} \hat{\chi}_a \partial_b \hat{\chi}_c$$

upon the following non-linear field redefinition:

$$\hat{\chi}_a^\alpha = \chi_a^\alpha + \frac{if^{-2}}{3} \varepsilon^{dfg} \chi_d^\alpha (\chi_f \partial_a \chi_g).$$

In contrast to the vector Goldstone case, this non-linear model retains gauge symmetry which is that of the Rarita-Schwinger action.

Conclusions and Outlook

- It would be interesting to consider the coupling of the vector Goldstone to a 3d gravity theory which is invariant under local symmetry associated with Hietarinta algebra.
- What kind of 3d massive gravity or bi-gravity will we obtain in that case?

Conclusions and Outlook

- One can couple the RS goldstino to conventional 3d (super)gravity and Hypergravity (with spin-2 and spin-5/2 gauge fields) and study the properties of these models.
- One can also consider the RS goldstino model in D=4 associated with the algebra

$$\{Q_\alpha^a, Q_\beta^b\} = 2\varepsilon^{abcd} (\Gamma_5 \Gamma_c)_{\alpha\beta} P_d, \quad (\alpha, \beta = 1, \dots, 4) \quad (a, b, \dots = 0, 1, 2, 3)$$

to see if it reduces to the free Rarita-Schwinger action and has gauge symmetry or not.

- Generalization to yet higher-spin Goldstone fields.

Thank you!