

Global analysis of neutrino oscillation parameters



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Open questions in neutrino Physics

- Dirac/Majorana nature of neutrinos
- Absolute masses
- CP violation
- Mass Ordering
- Sterile neutrinos
- Nonstandard interactions

The knowledge of mass-mixing oscillation parameters can help to answer these questions

In this talk → status of parameter determination from global analyses with focus on θ_{23} , δ and mass ordering

Precision era in neutrino oscillation phenomenology

Standard 3ν mass-mixing framework parameters

What we know

$$\delta m^2 \quad 2.2\%$$

$$\Delta m^2 \quad 1.4\%$$

$$\sin^2 \theta_{12} \quad 4.4\%$$

$$\sin^2 \theta_{13} \quad 3.8\%$$

$$\sin^2 \theta_{23} \quad 5.2\%$$

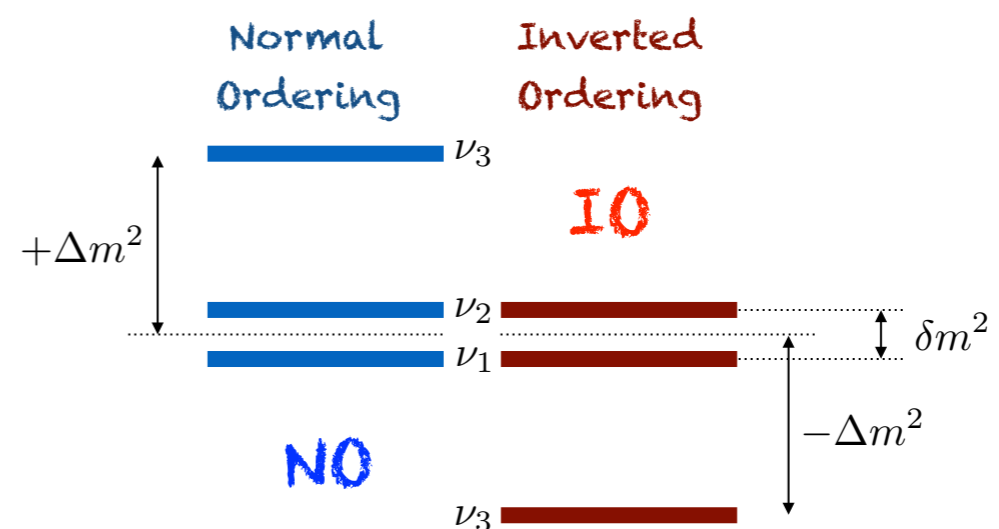
What we still do not know

CP-violating phase δ

Octant of θ_{23}

Mass Ordering $\rightarrow \text{sign}(\Delta m^2)$

$$\Delta m^2 = (\Delta m_{13}^2 + \Delta m_{23}^2)/2$$



To understand how bounds on the oscillation parameter arise it is useful to look at their correlations and to consider the progressive contribution of different data sets:

(1) LBL acc + Solar + KamLAND

Solar + KL data provide the necessary input for $(\delta m^2, \theta_{12})$, but also independent -although weak- constraints on θ_{13} . The data set (1) provides, by itself, a measurement of θ_{13} .

(2) LBL acc + Solar + KamLAND + SBL Reactors

SBL reactors not only provide the most accurate determination of θ_{13} but also an independent determination of Δm^2

(3) LBL acc + Solar + KamLAND + SBL Reactors + Atmospheric

Atmospheric neutrino data (SK + DeepCore) sensitive in different ways to all the oscillation parameters via disappearance and appearance channels. Because of matter effects they depends on all parameters in the 3v framework, but dominantly on $(\Delta m^2, \theta_{23})$

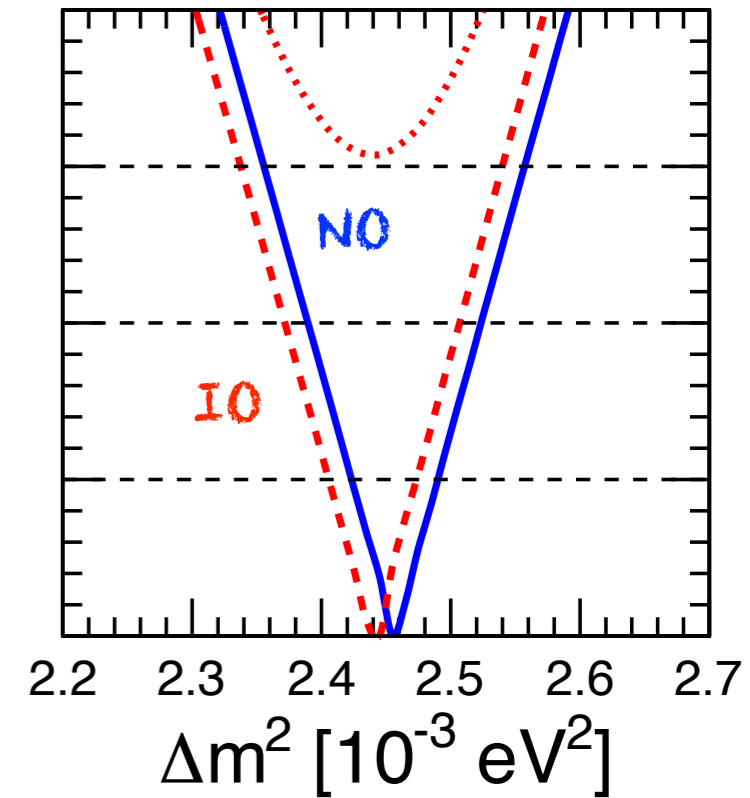
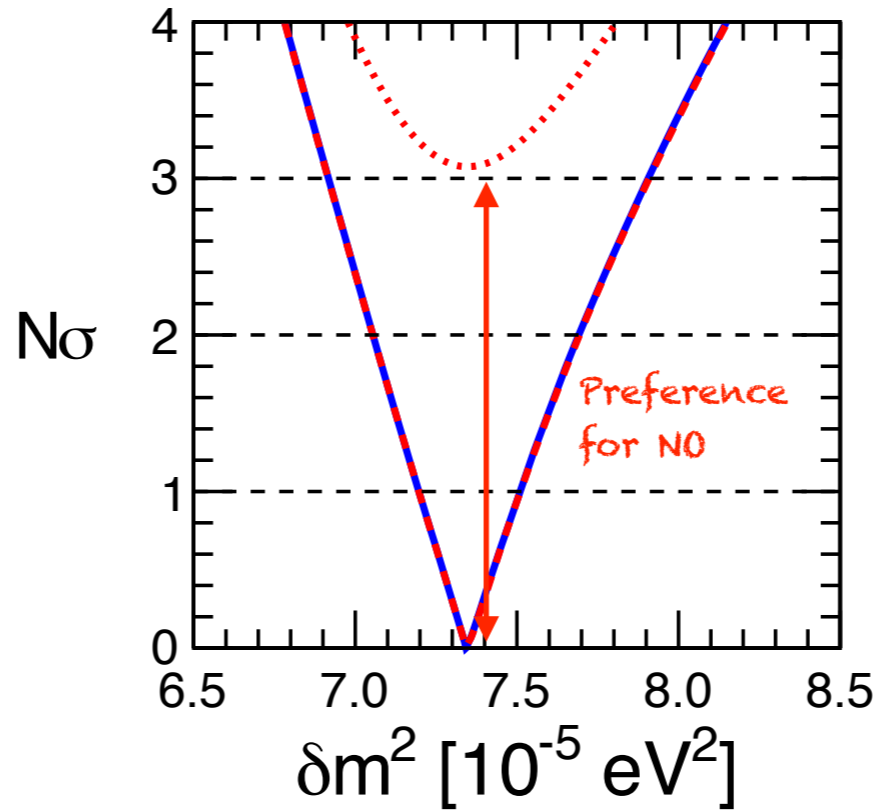
Mass Differences

$$\Delta m^2 = (\Delta m_{13}^2 + \Delta m_{23}^2)/2$$

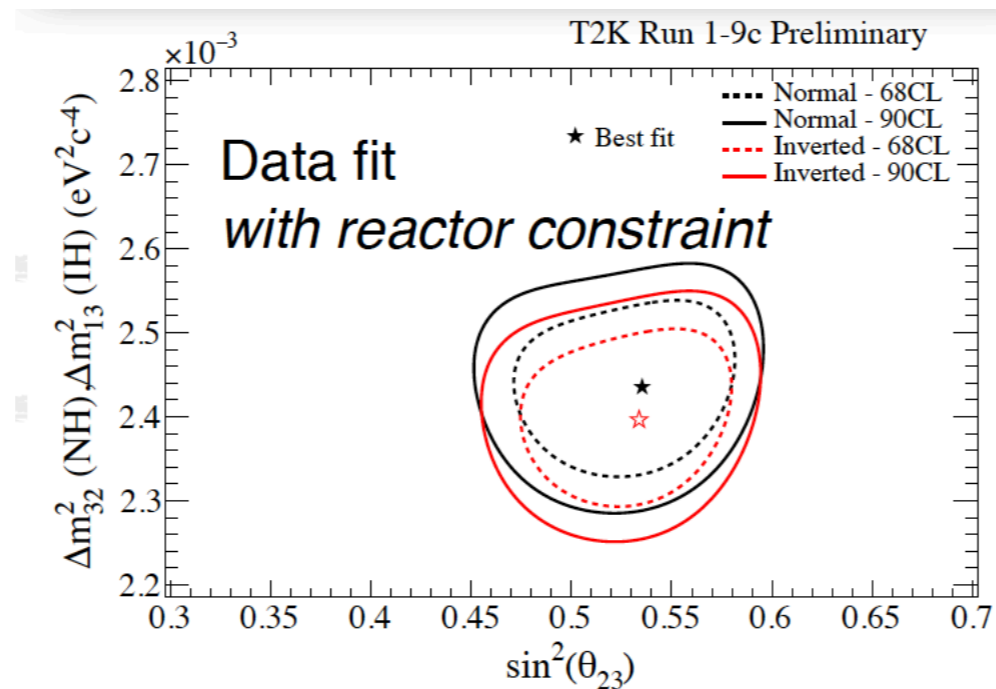
Mass Ordering = sign of Δm^2

Squared mass differences have both lower and upper bounds at more than 3σ

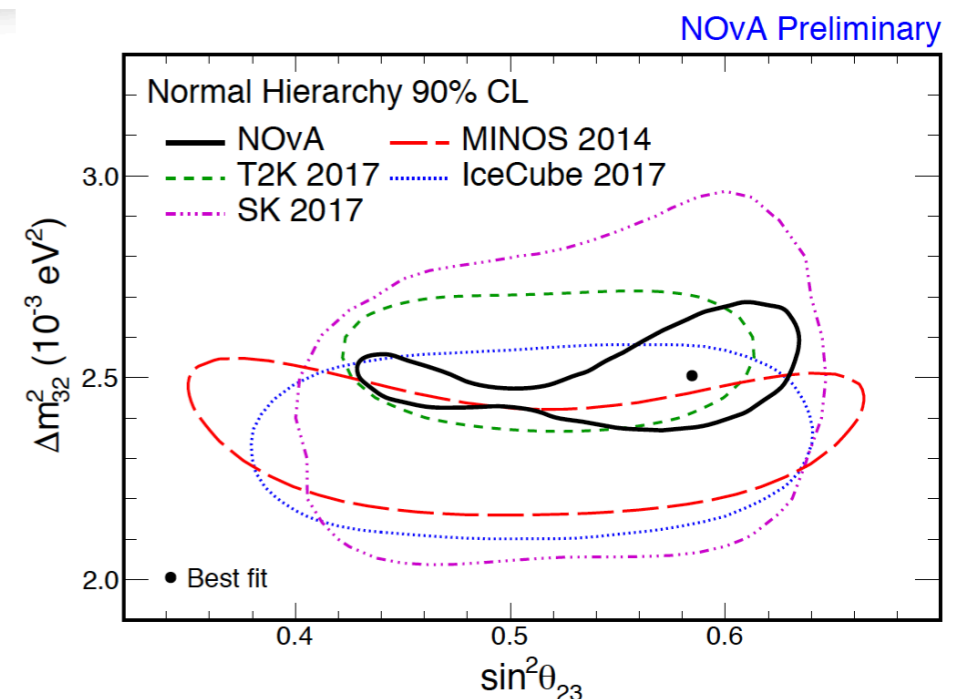
Nearly Gaussian uncertainties for Δm^2 and to a lesser extent for δm^2



Neutrino 2018 updates (still not included)

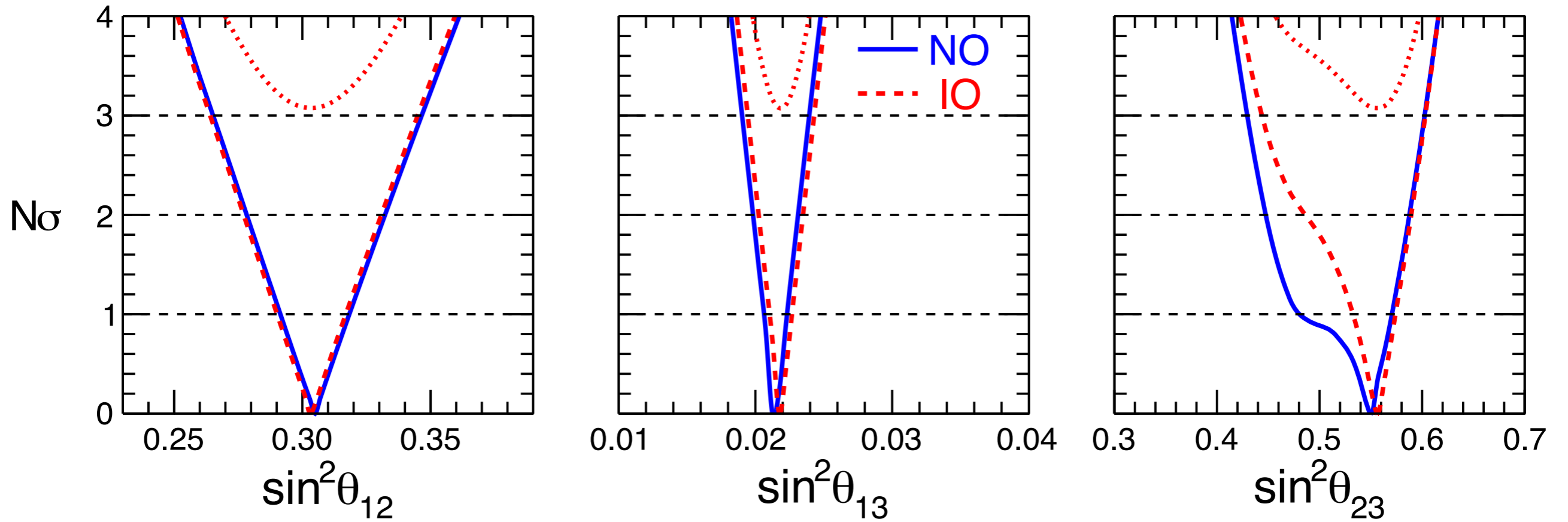


Wascko - Neutrino 2018



Sanchez - Neutrino 2018

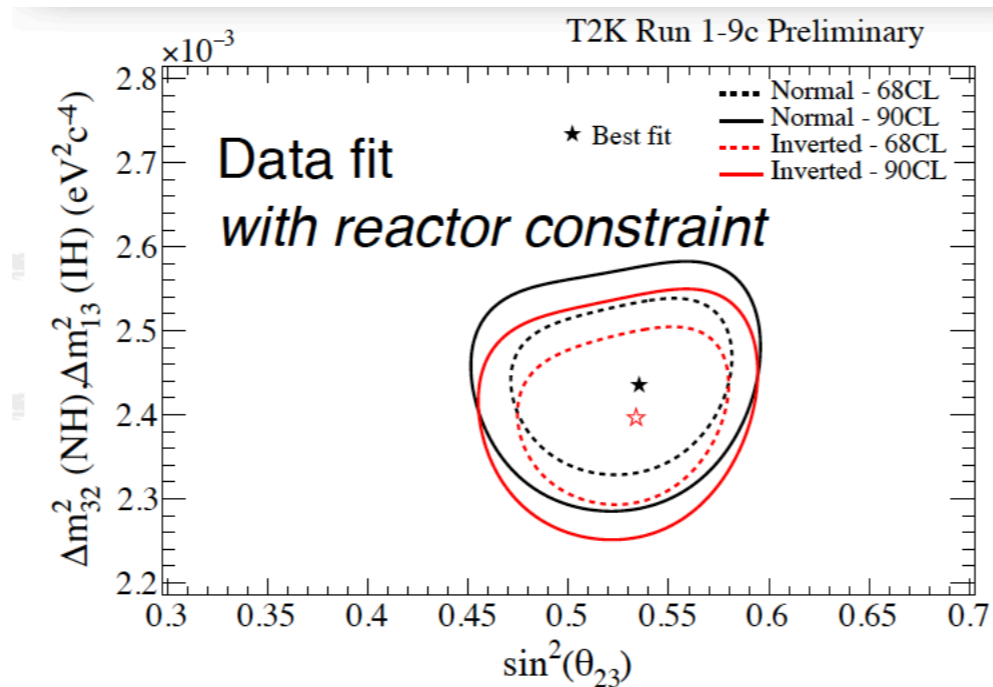
Mixing Angles



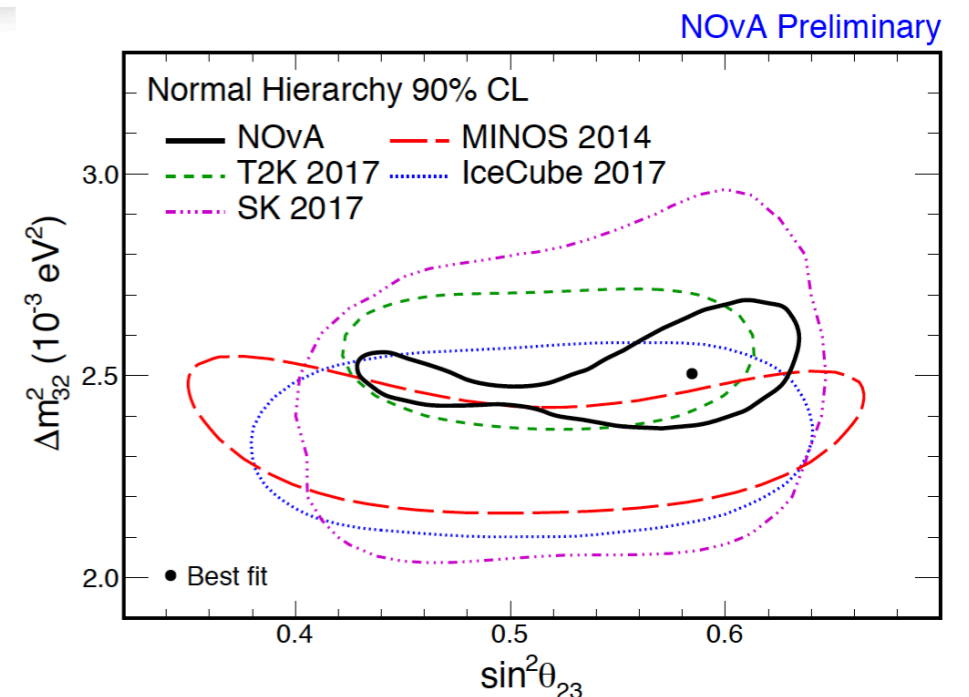
Mixing angles (θ_{23}, θ_{12}) have both lower and upper bounds at more than 3σ

Nearly Gaussian uncertainties for θ_{23} and to a lesser extent for θ_{12}

θ_{23} maximal mixing allowed at about 2σ in both NO and IO

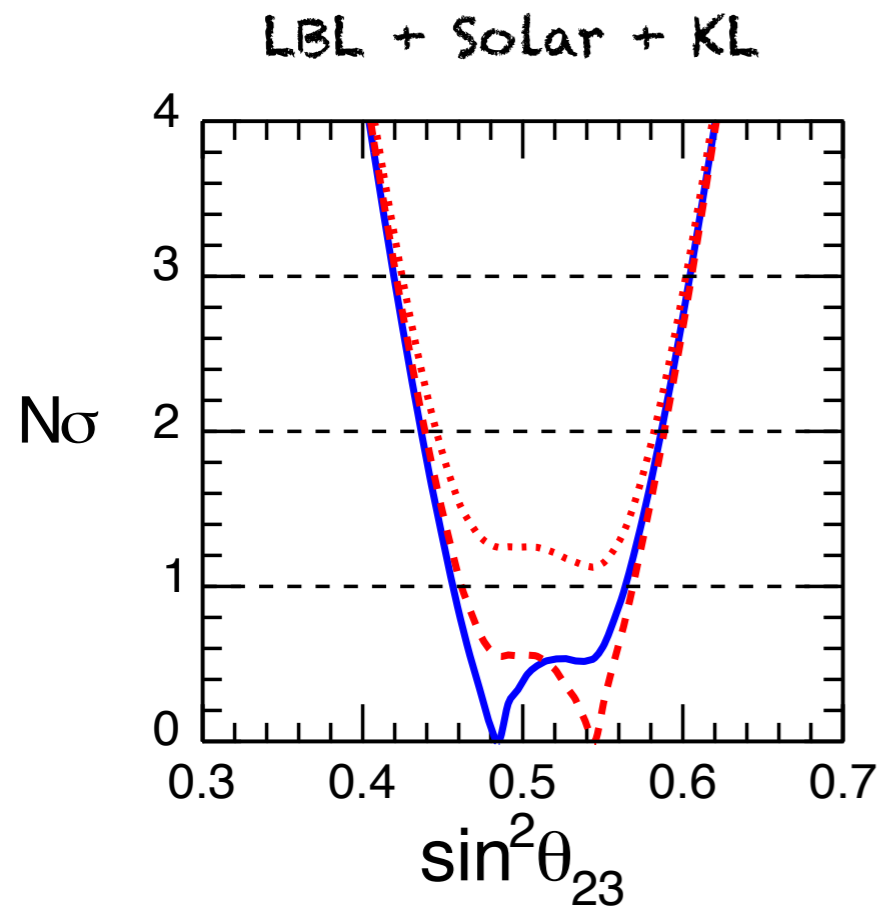


NOvA and MINOS prefer nonmaximal mixing



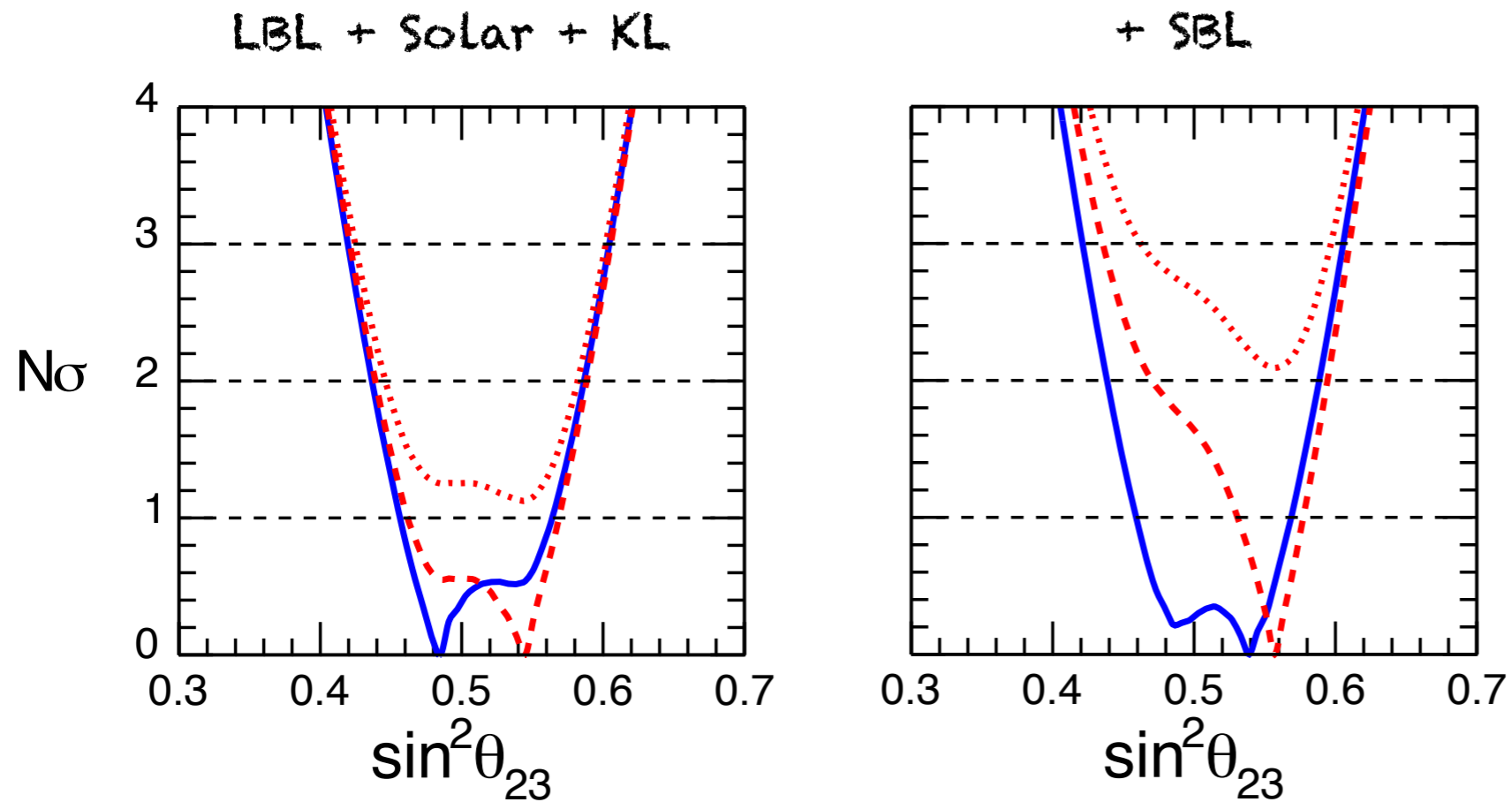
θ_{23} octant ambiguity

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Slight preference for $\theta_{23} < \pi/4$ in NO and $\theta_{23} > \pi/4$ in IO, but both octants are allowed at 1σ

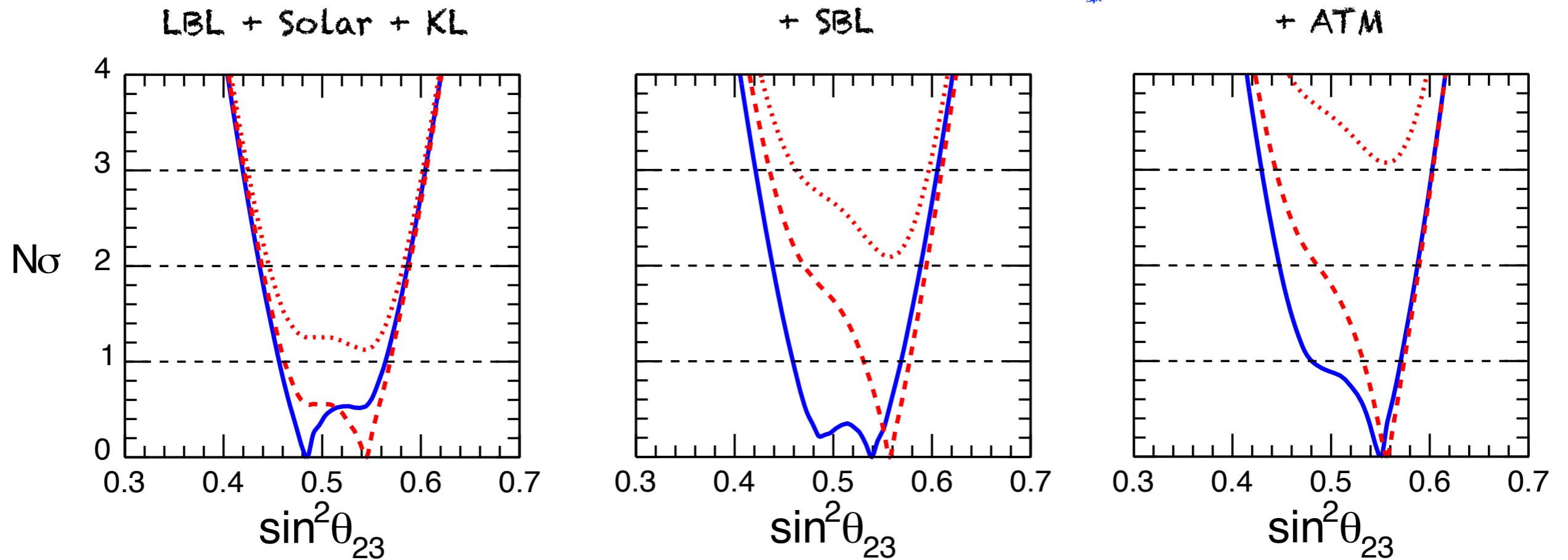
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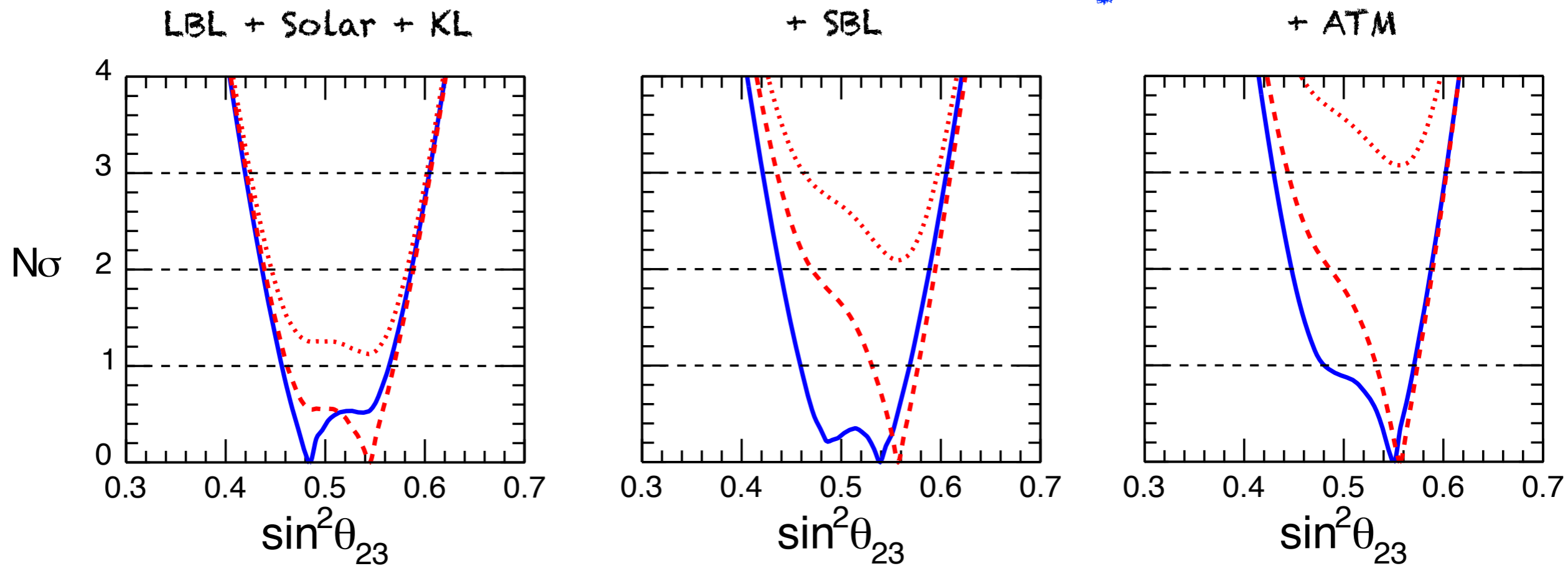


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More pronounced preference for $\theta_{23} > \pi/4$ but both octants allowed at $< 2\sigma$.

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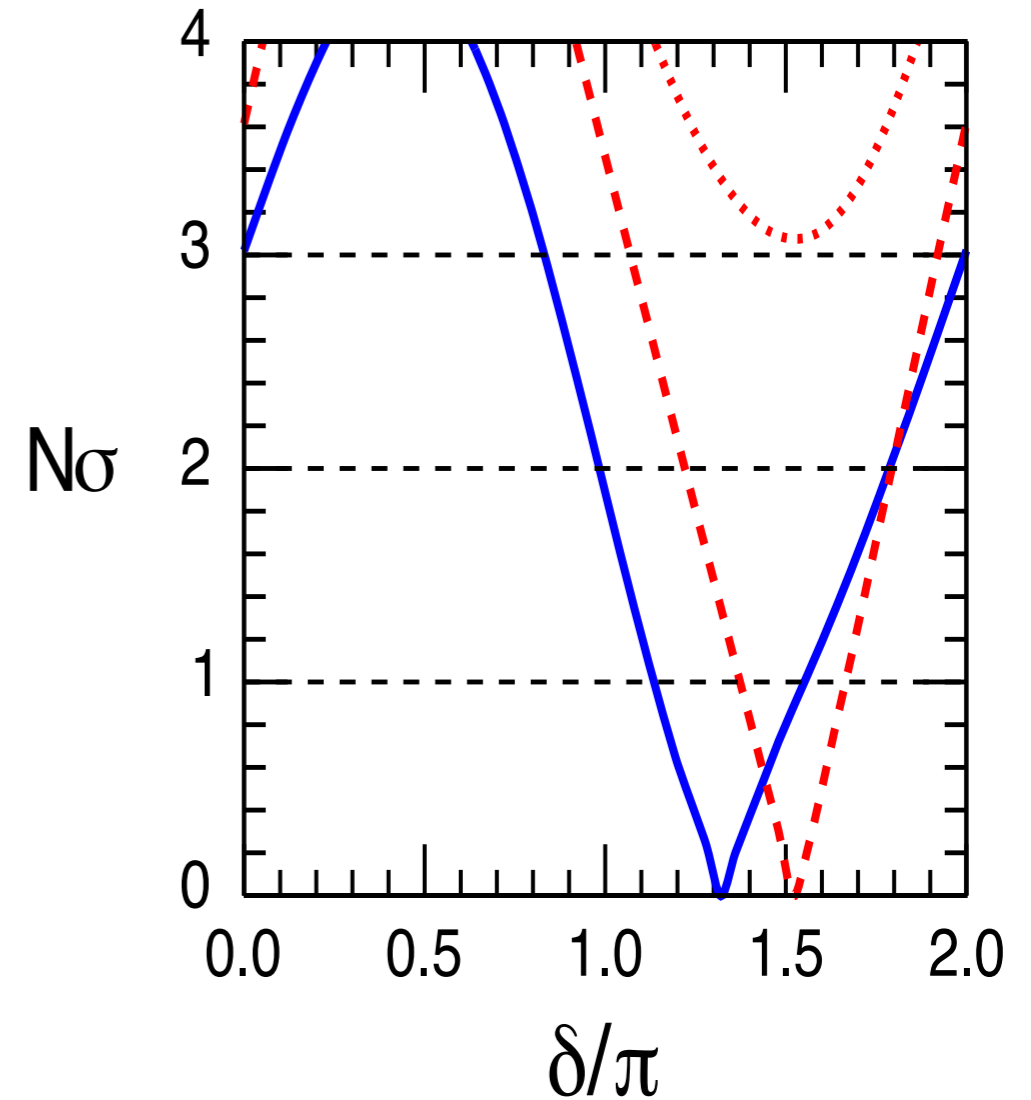
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More pronounced preference for $\theta_{23} > \pi/4$ but both octants allowed at $< 2\sigma$.

Globally, relatively low uncertainty of $\sim 5\%$ on $\sin^2\theta_{23}$ (NOvA data in better agreement with quasi-maximal mixing). Maximal mixing allowed at less than 2σ in both NO and IO (but Neutrino 2018 NOvA analysis prefers again nonmaximal mixing)

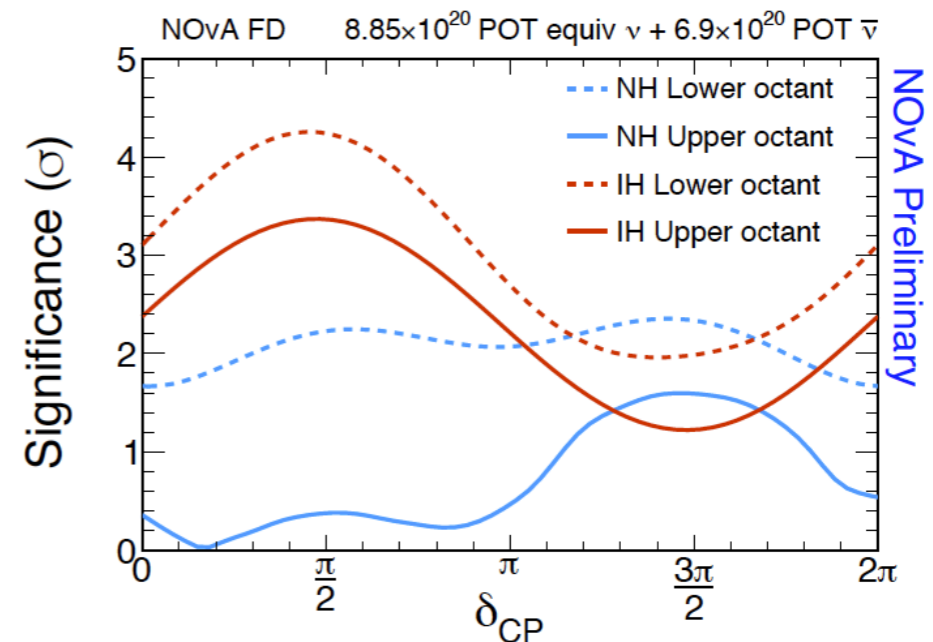
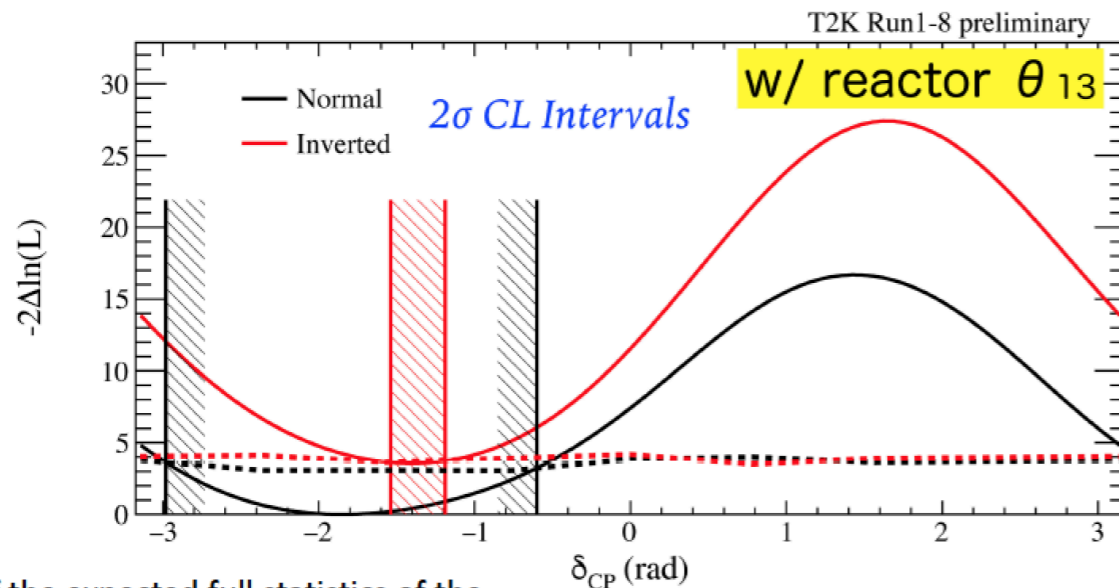
CP phase δ

CP phase: $\delta \sim 1.3\pi$ (1.5π) at best fit
 CP-conserving cases ($\delta = 0, \pi$)
 disfavoured at $\sim 2\sigma$ level or more
 Significant fraction of the $[0, \pi]$
 range disfavoured at $>3\sigma$ in NO,
 at $>4\sigma$ in IO



New T2K results: KEK seminar on 4 August 2017

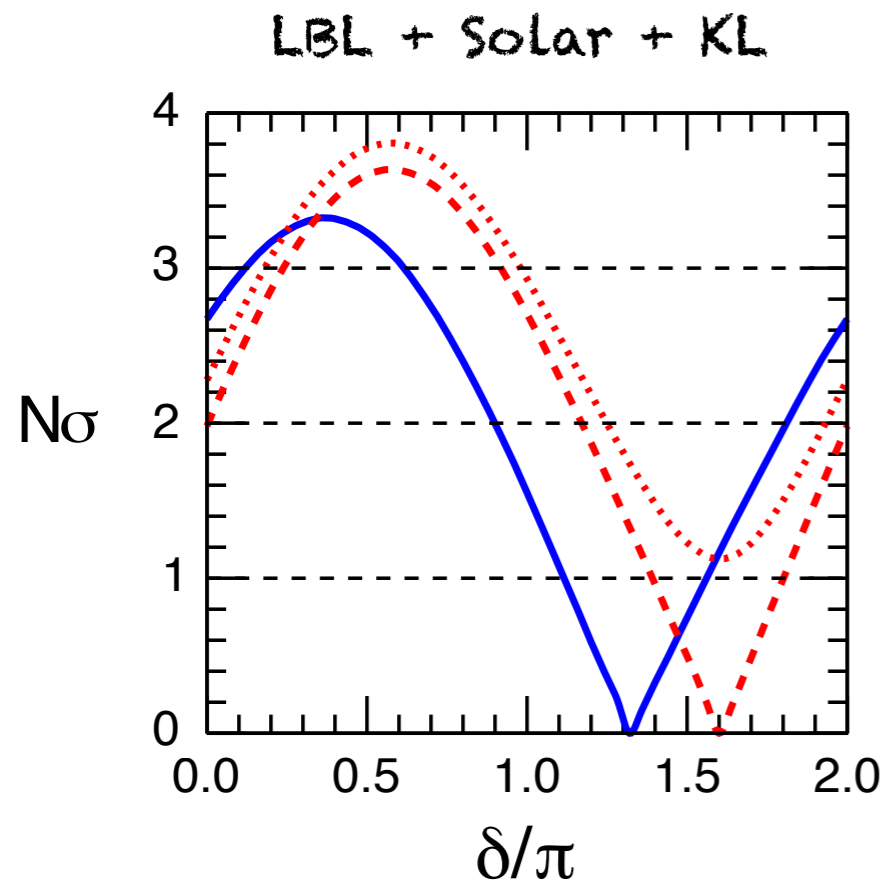
Based on 89 ν_e and 7 $\bar{\nu}_e$ events



- 30% of the expected full statistics of the experiment
- 30% improvement in efficiency x acceptance
- Important improvements in neutrino interactions modelling
- δ_{CP} determination is very important for future searches of MH in long baseline experiments

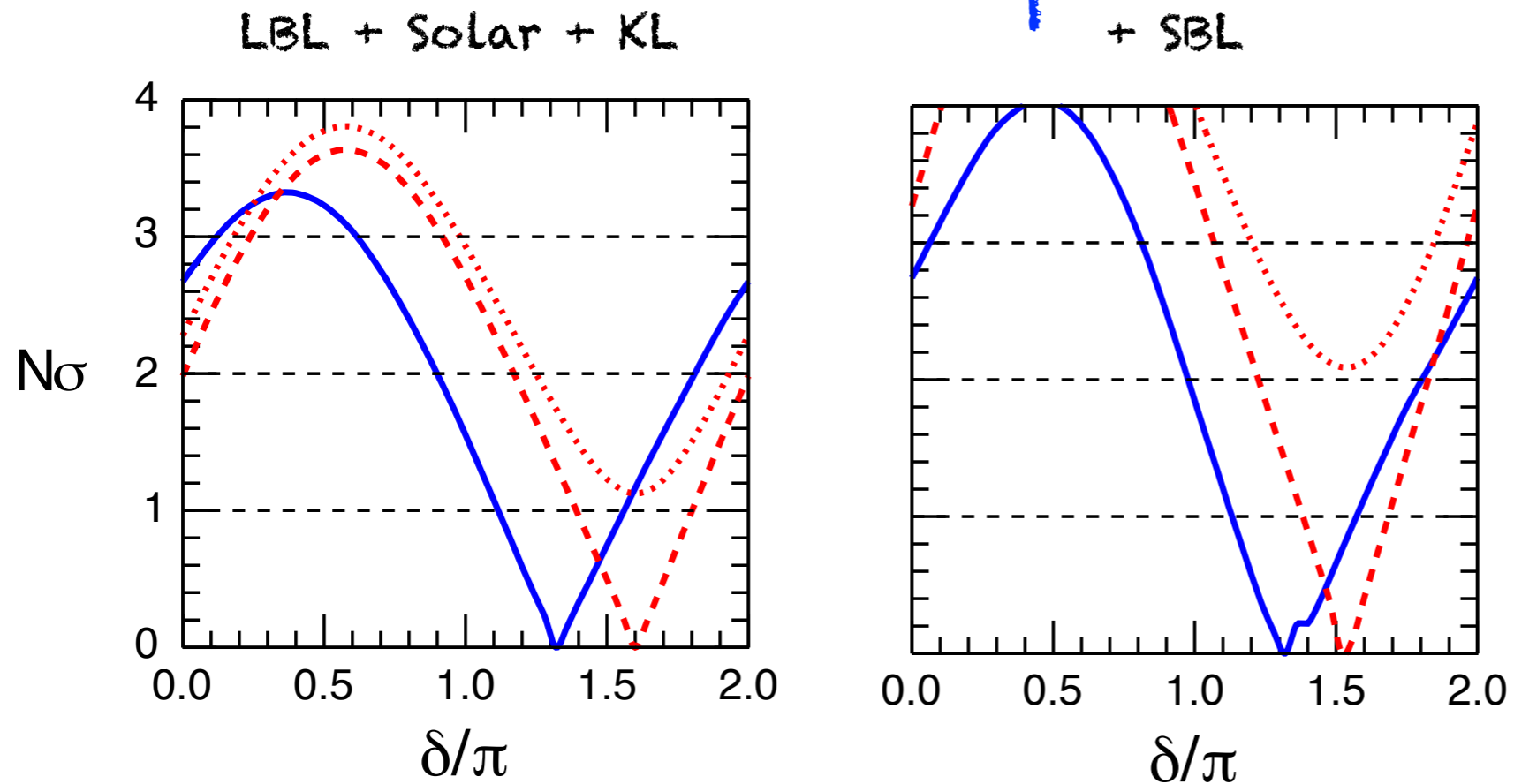
CP phase δ

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CP-conserving values $\delta=\{0,\pi\}$ allowed at $\sim 2\sigma$ (3σ) in NO (IO). Clear preference $\delta\sim 3\pi/2$. $\delta\sim\pi/2$ disfavoured at more than 3σ

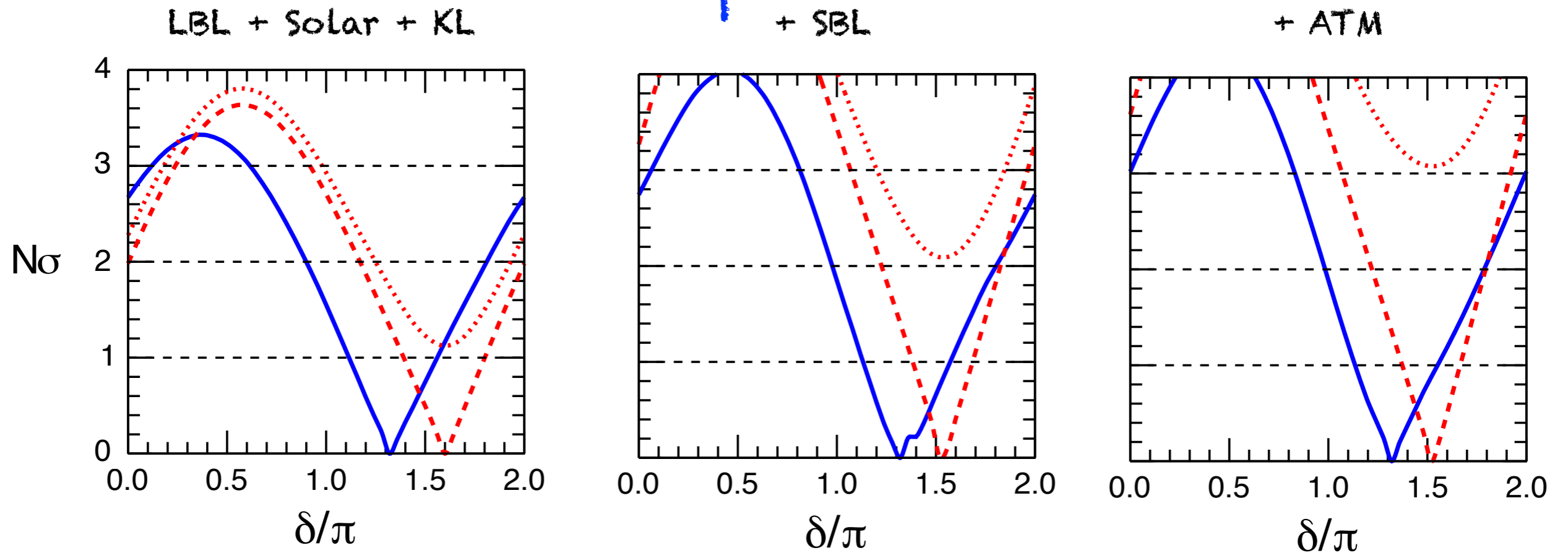
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Indications on $\delta\sim 3\pi/2$ strengthened. CP-conserving values disfavoured at $>1.8\sigma$ in NO and $>3\sigma$ in IO. Significant ranges excluded at $>3\sigma$ in both NO and IO.

CP phase δ

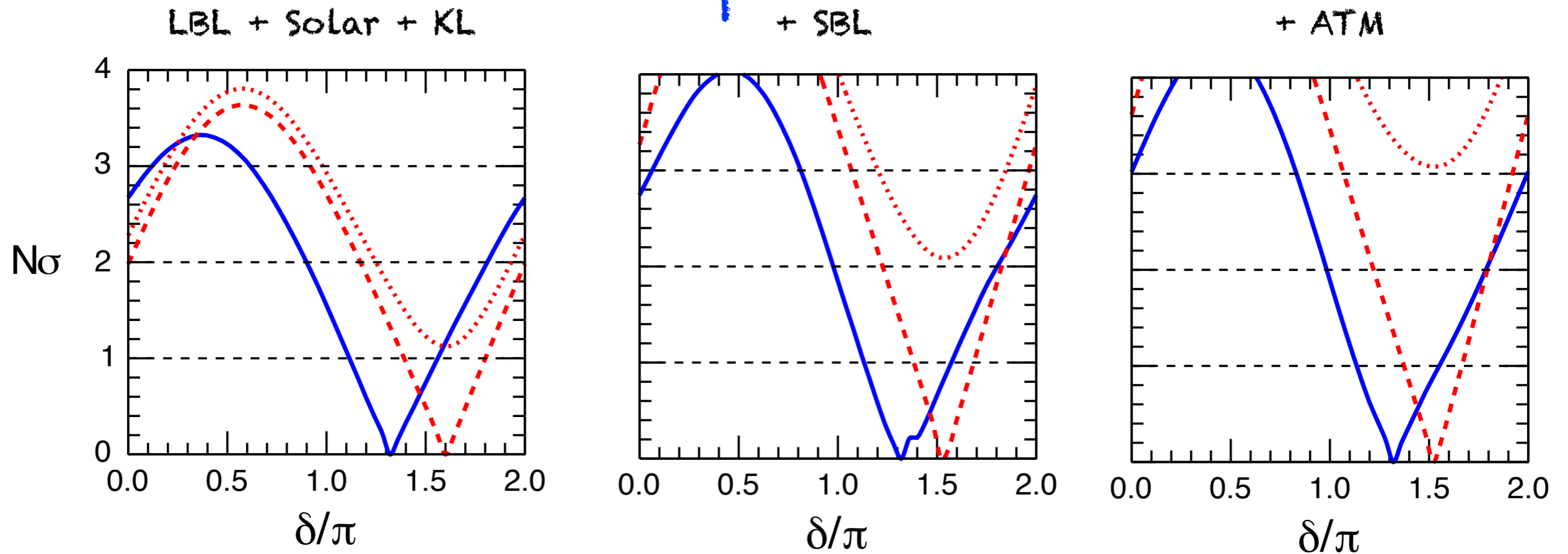


CP-conserving values $\delta = \{0, \pi\}$ allowed at $\sim 2\sigma$ (3σ) in NO (IO). Clear preference $\delta \sim 3\pi/2$. $\delta \sim \pi/2$ disfavoured at more than 3σ

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Preference for CP violation with $\sin\delta < 0$ confirmed, while CP conservation is disfavoured at $> 1.9\sigma$ for NO and $> 3.5\sigma$ for IO.

CP phase δ



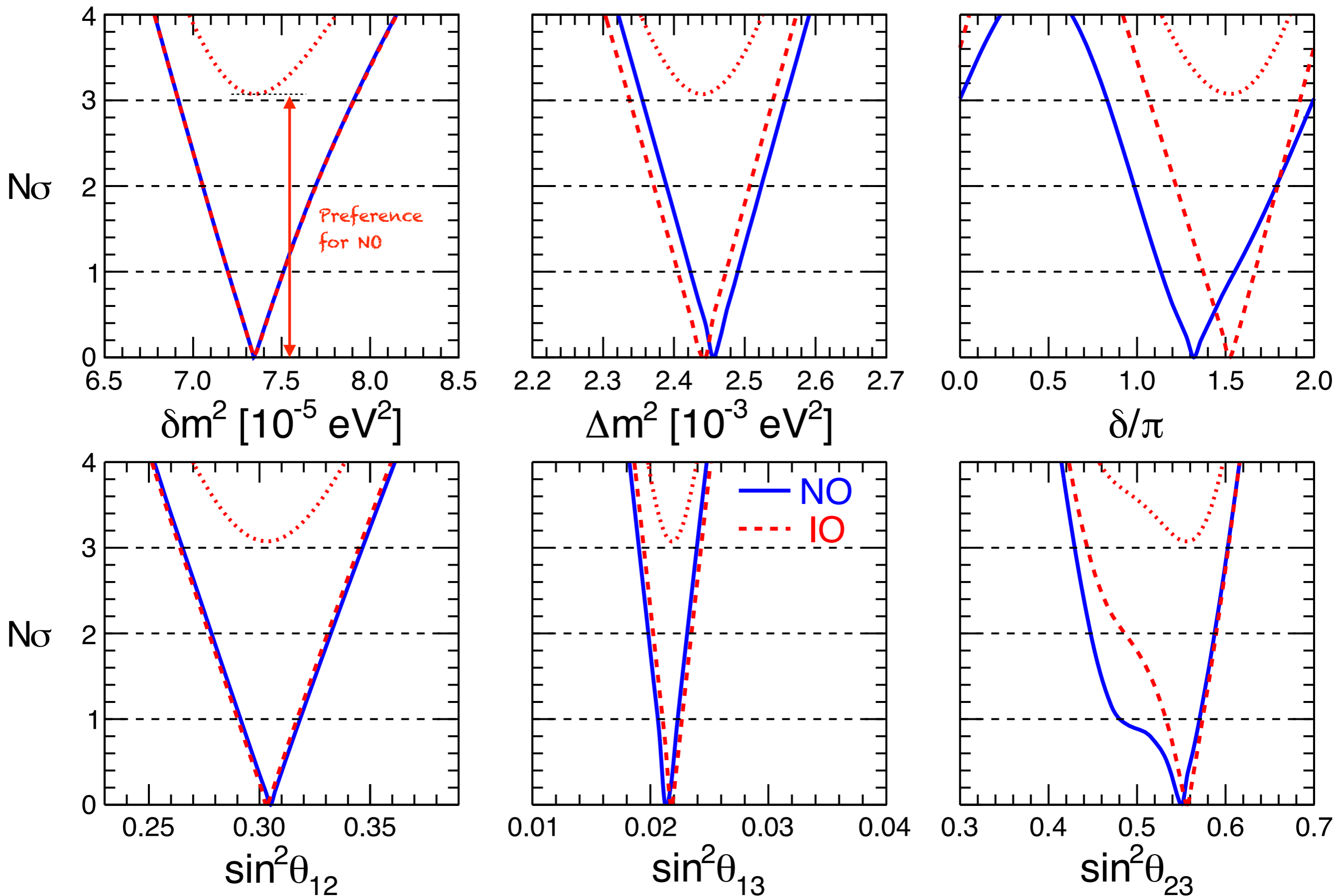
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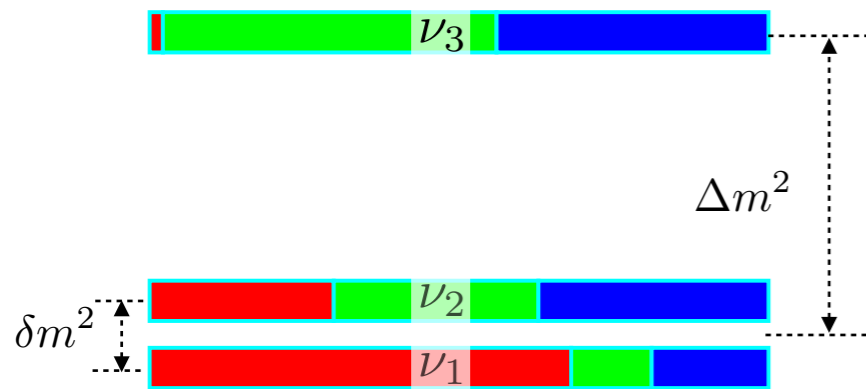
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"Effective" 1σ accuracy of $\sim 15\%$ in NO and $\sim 9\%$ in IO. Rejection of the CP-conserving case $\delta = 0$ at 3σ in NO, but not enough to exclude $\delta = \pi$ at 2σ . Both cases excluded at 3σ in IO.

LBL Acc + Solar + KamLAND + SBL Reactors + Atmos



Comparison of global analyses



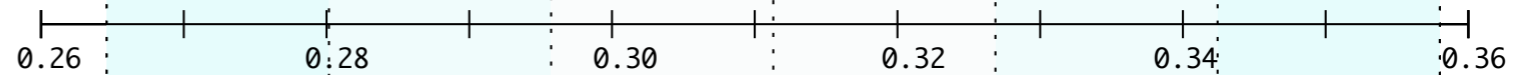
- ν_e
- ν_μ
- ν_τ

- Bari
- NuFit
- Valencia

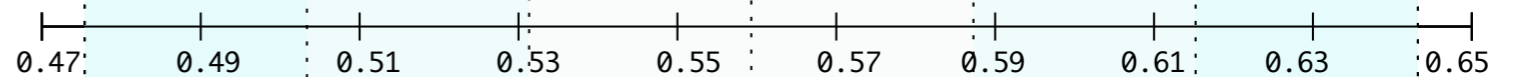
The three global analyses are in good agreement with some difference about θ_{23} ranges & octant (differences on θ_{13} mostly due to Oct 2018 Day Bay update, included in NuFit)

Normal Ordering

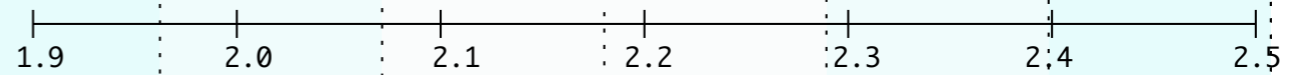
$\sin^2 \theta_{12}$



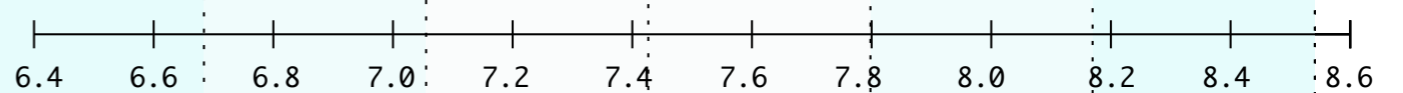
$\sin^2 \theta_{23}$



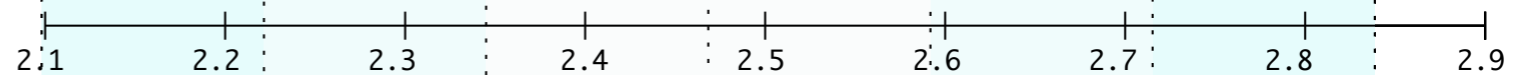
$\sin^2 \theta_{13}$



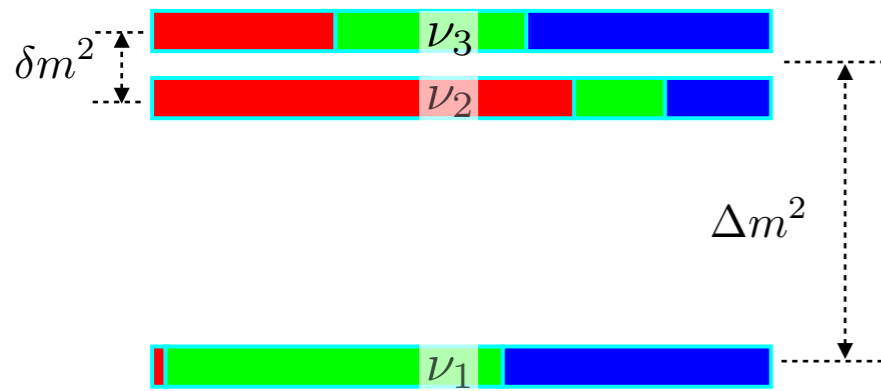
δm^2



Δm^2



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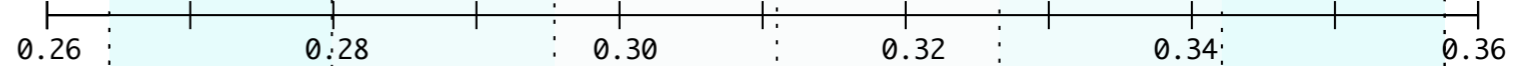
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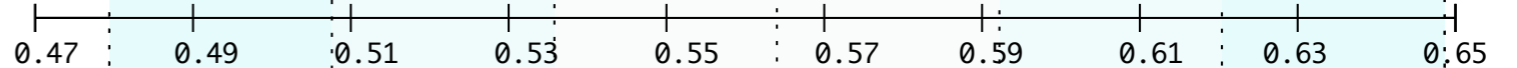
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Inverted Ordering

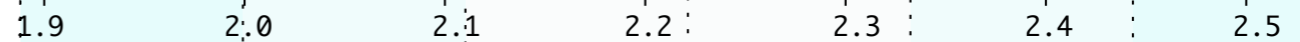
$\sin^2 \theta_{12}$



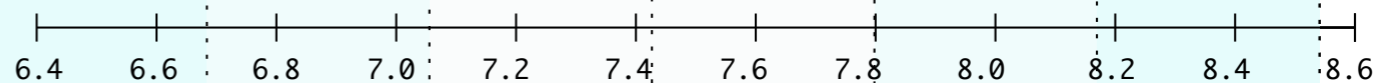
$\sin^2 \theta_{23}$



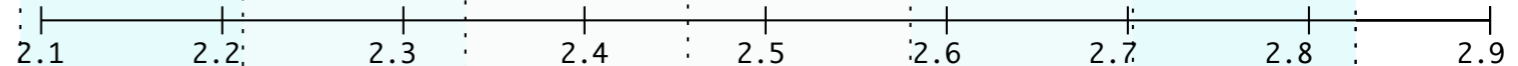
$\sin^2 \theta_{13}$



δm^2

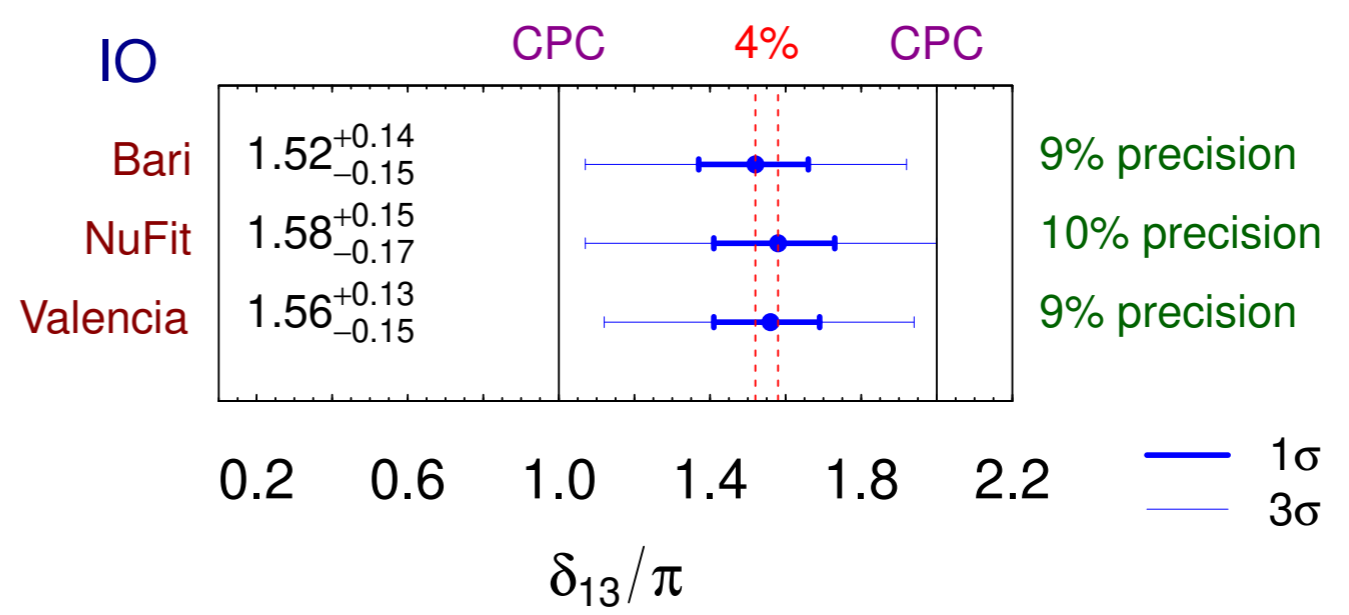
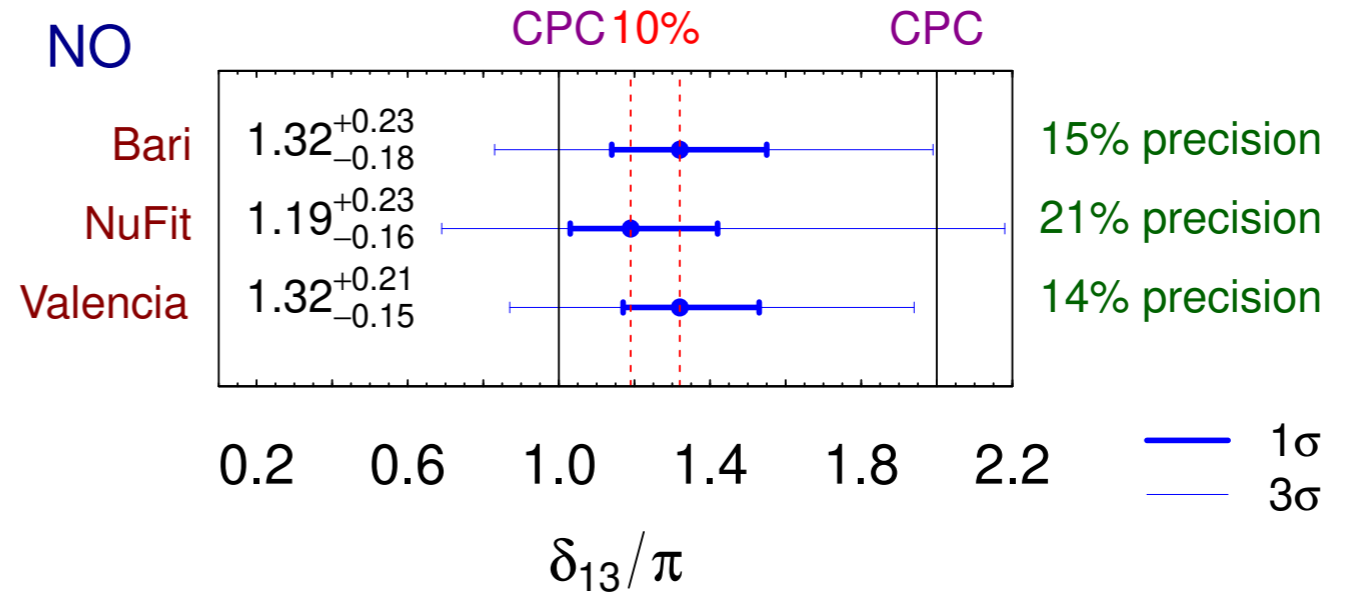
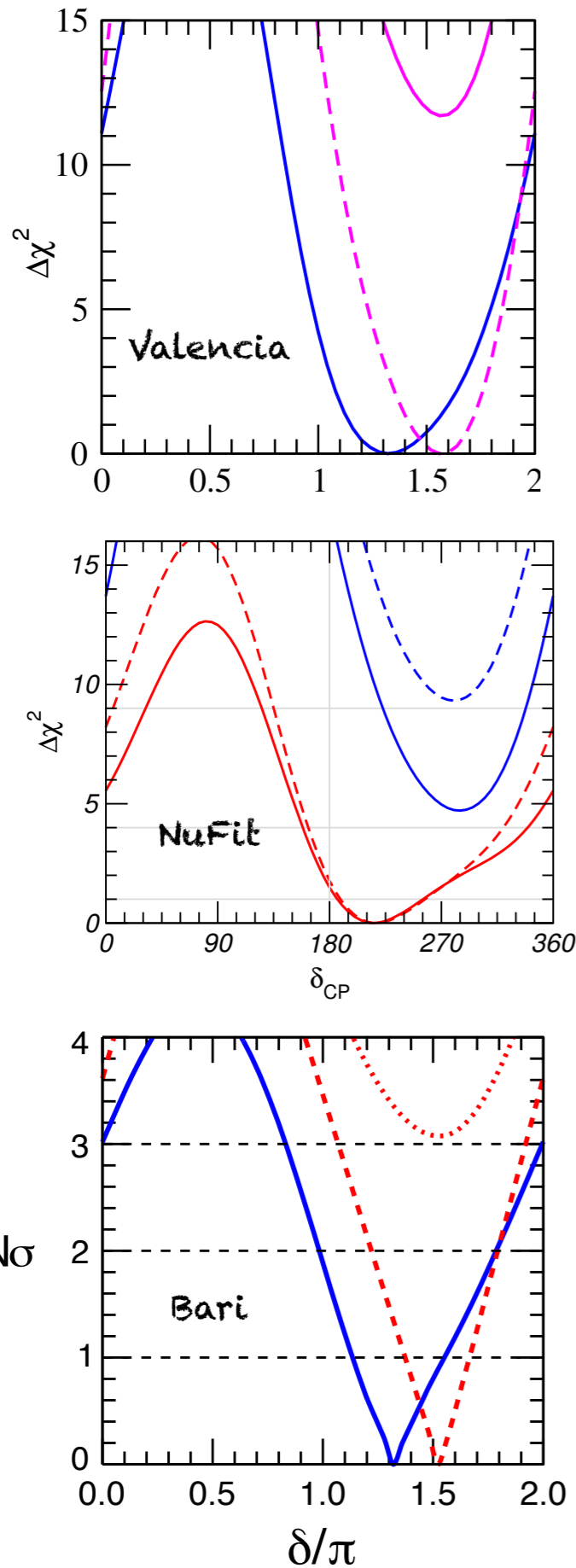


Δm^2



10-15% "Determination" of δ

CP Violation?



Mass Ordering: present situation

	LBL + Solar + KL	(+ SBL)	(+Atm)
$\Delta\chi^2_{IO-NO}$	+1.3	+4.4	+9.5

@Neutrino 2018

T2K - preference for NO

$$\Delta\chi^2_{IO-NO} \sim 4$$

NOvA - weak preference for

$$NO, \Delta\chi^2_{IO-NO} \sim 1.3$$

Other groups findings

<http://www.nu-fit.org/> $\Delta\chi^2_{IO-NO} = +9.3$ (4.7, No SK)

M. Tortola @Neutrino 2018 $\Delta\chi^2_{IO-NO} = 11.7$

	$\Delta\chi^2_{IO-NO}$	$N\sigma$
Bari	9.5	3.1
NuFit	9.1	3.0
Valencia	11.7	3.4

NO favoured over IO
at about 3 sigma level

Correlations: (δm^2 , θ_{12})

We include the latest low-energy Borexino and SK-IV data

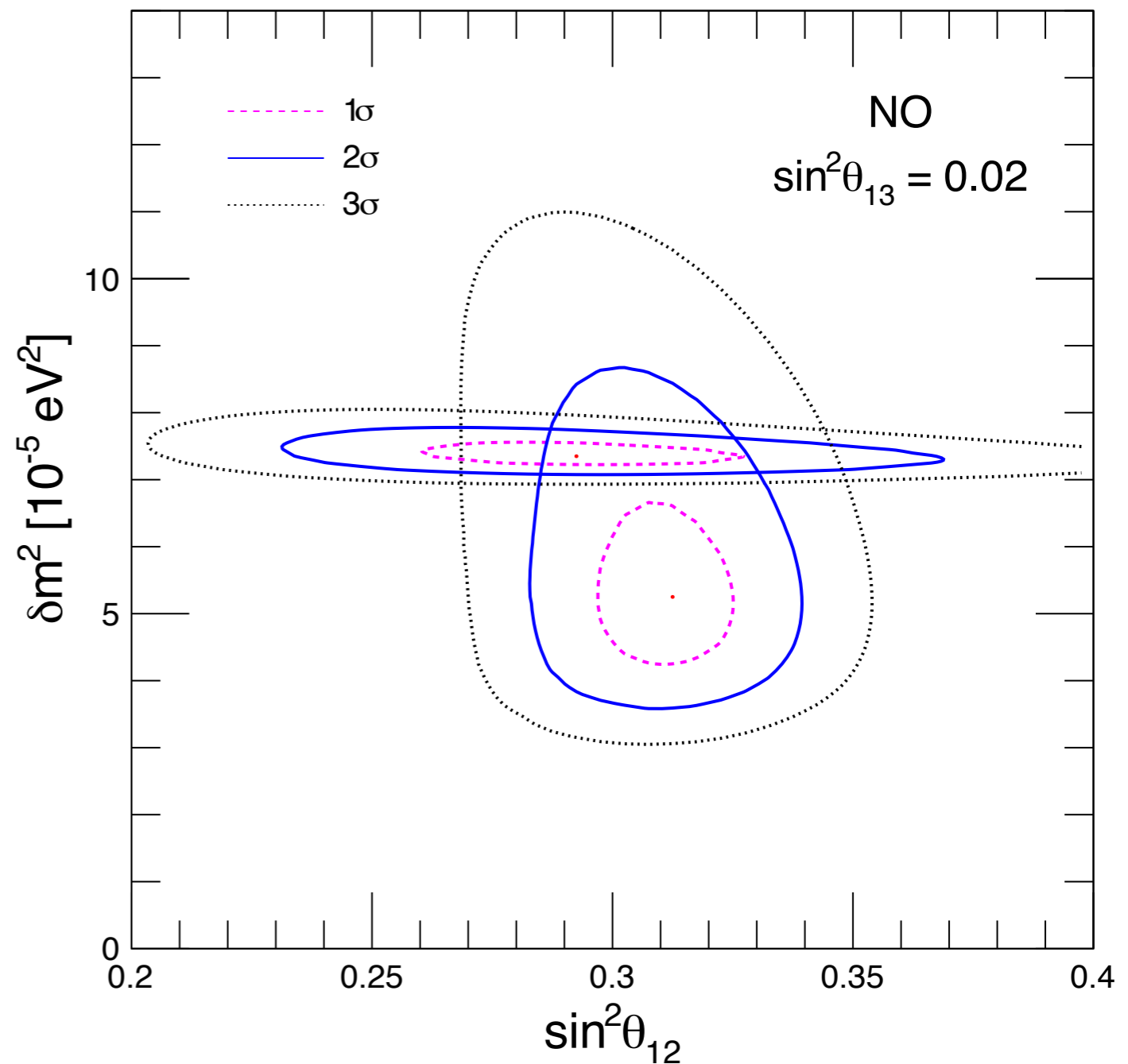
Many input updated. Ga neutrino absorption cross-section leads to a reduction of the unoscillated solar neutrino rate by ~ 6 SNU

Slightly decrease of θ_{12} and increase of δm^2 for nonzero θ_{13}

slight tension between the preferred mass-mixing value

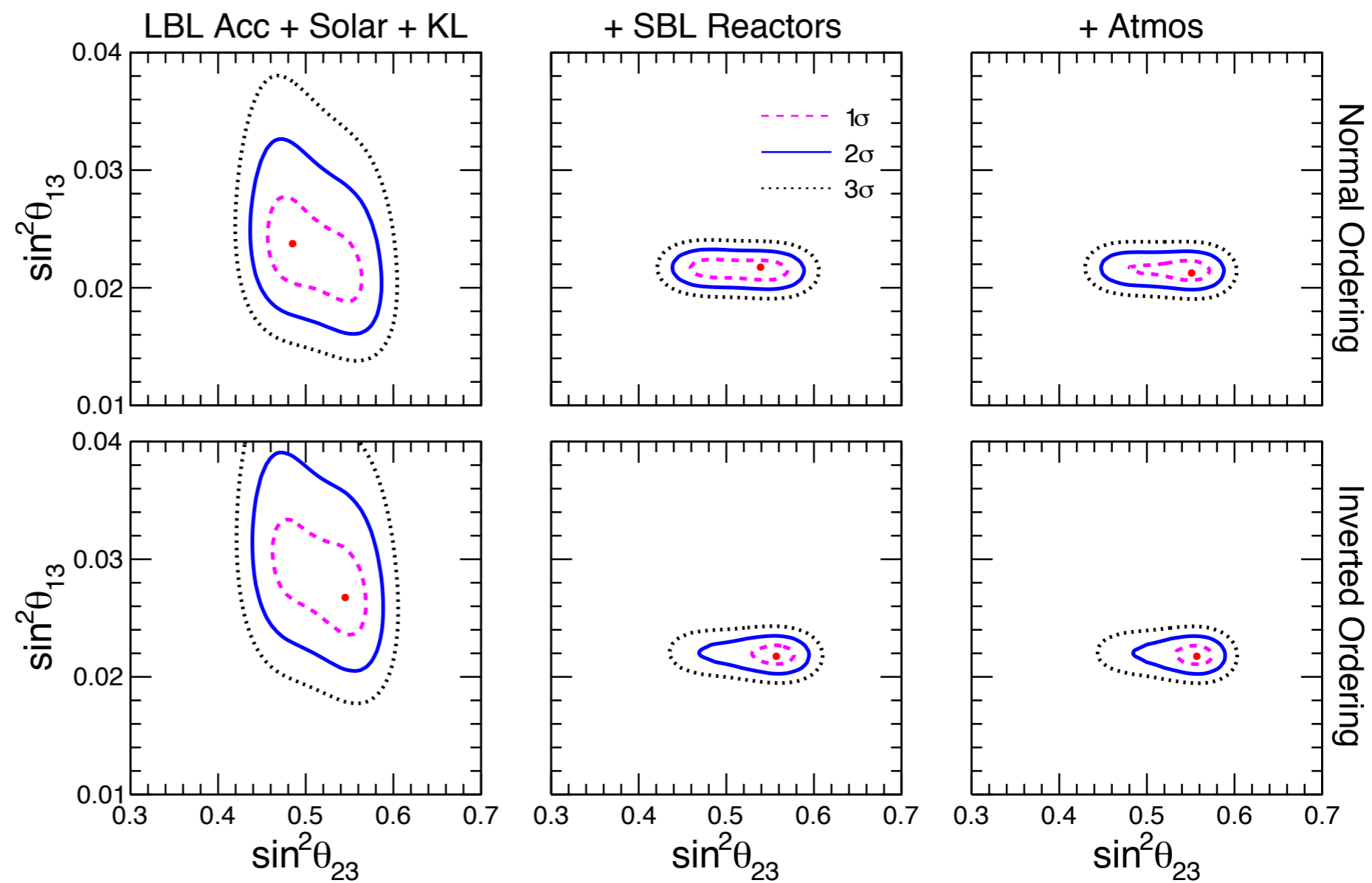
$$\Delta\chi^2 = \chi_{\text{sol+KL}}^2 - (\chi_{\text{sol}}^2 + \chi_{\text{KL}}^2) \sim 2$$

Solar Neutrinos + KamLAND



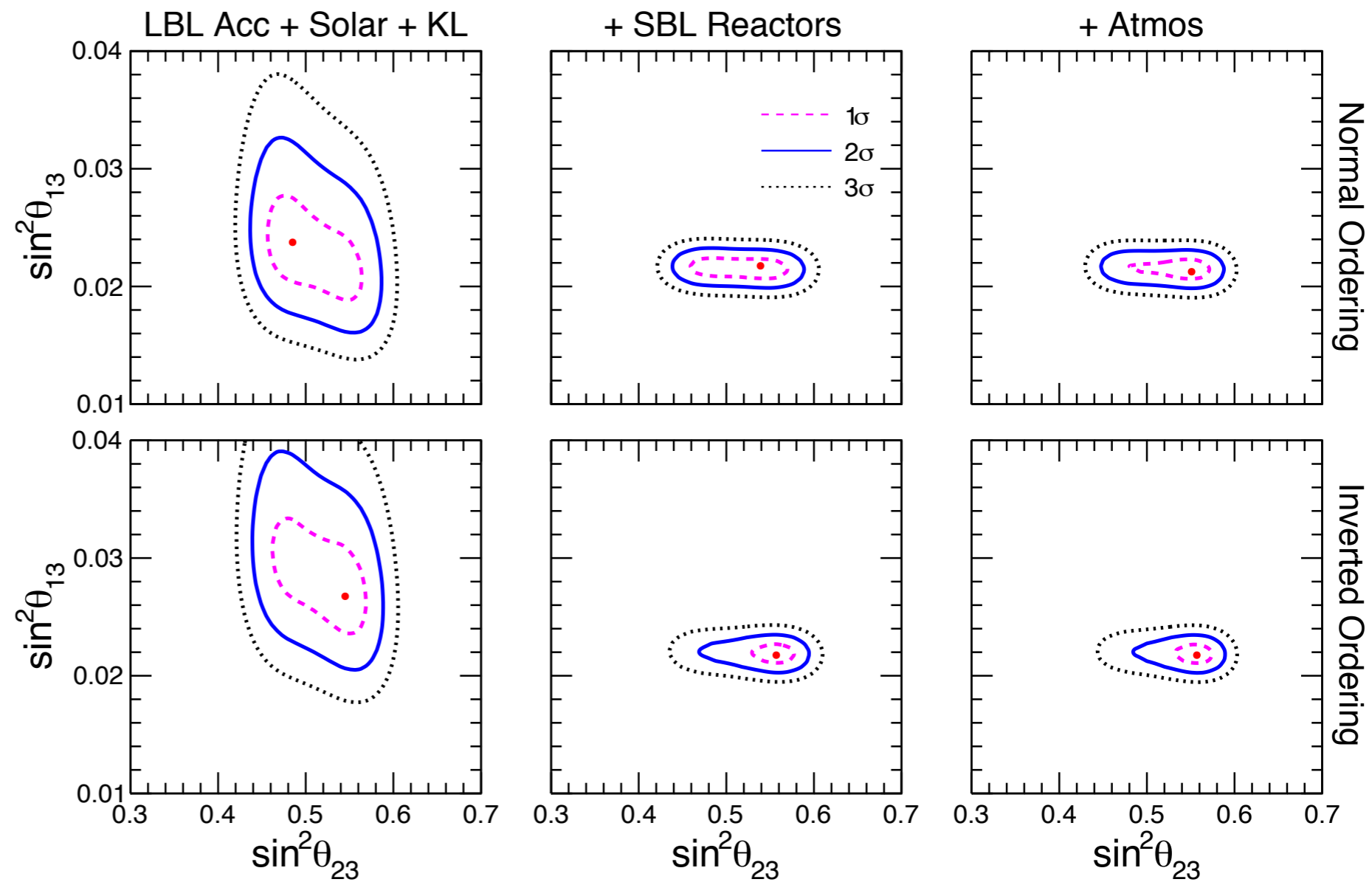
Solar neutrino contours would be slightly different in IO, shifted leftwards with $\delta(\sin^2 \theta_{12}) = -0.02$ to compensate the slightly higher survival probability for IO as compared to NO. In combination with (mass-ordering insensitive) KamLAND data, the overall shift of the best-fit mixing angle amounts to $\delta(\sin^2 \theta_{12}) = -0.01$
 $\Delta\chi^2 = 0.08$ in favor of IO with respect to NO

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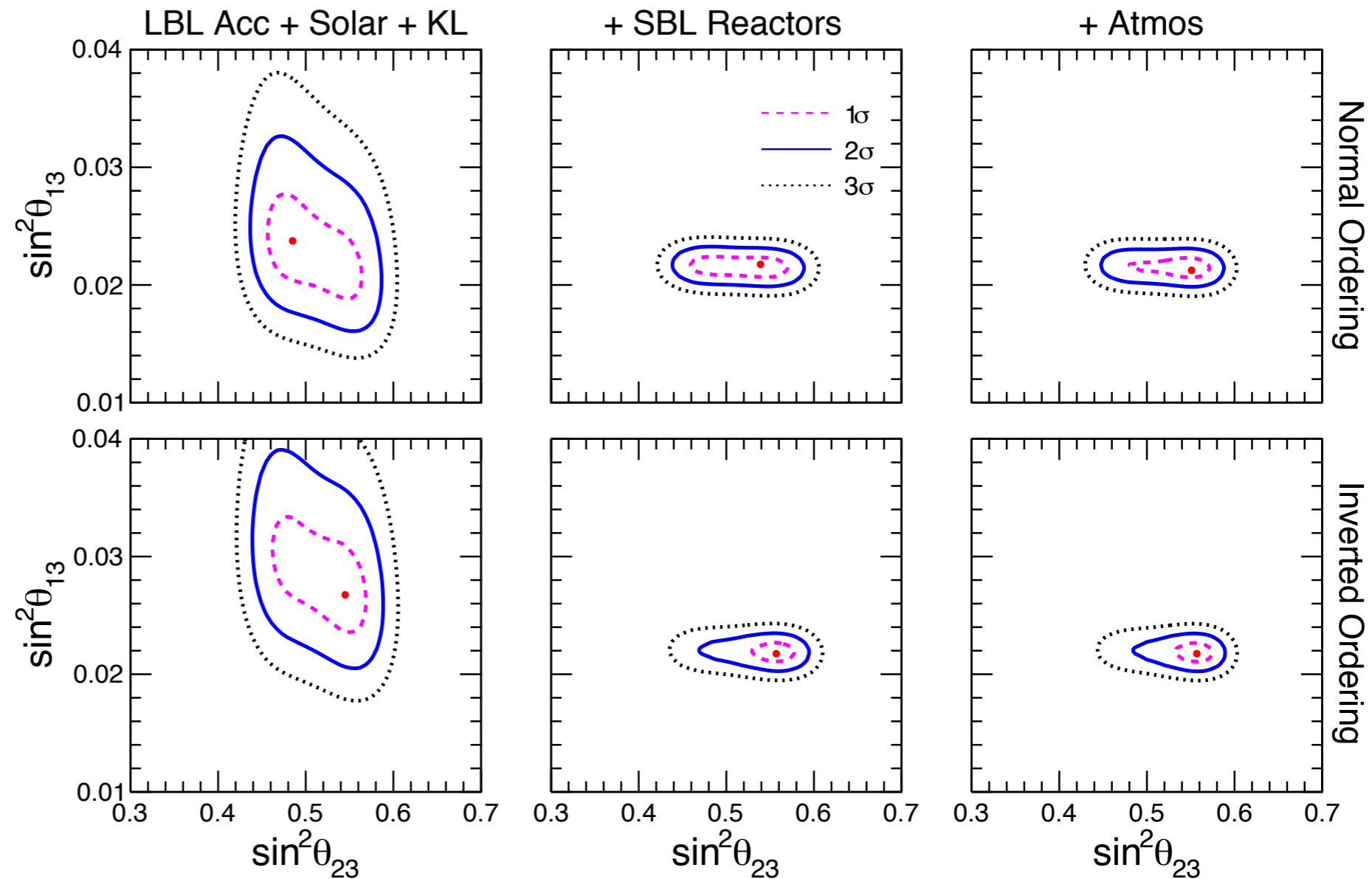
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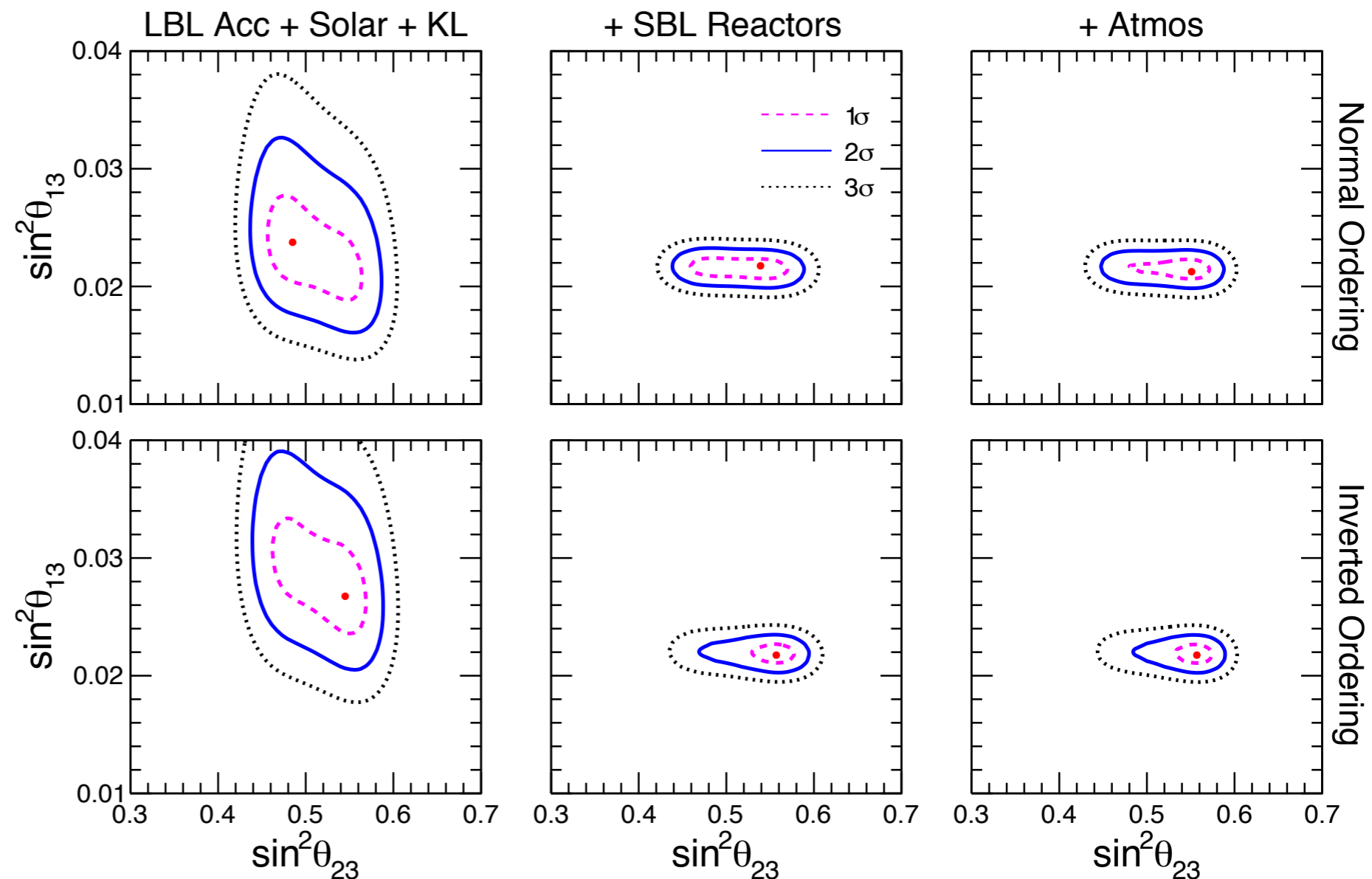


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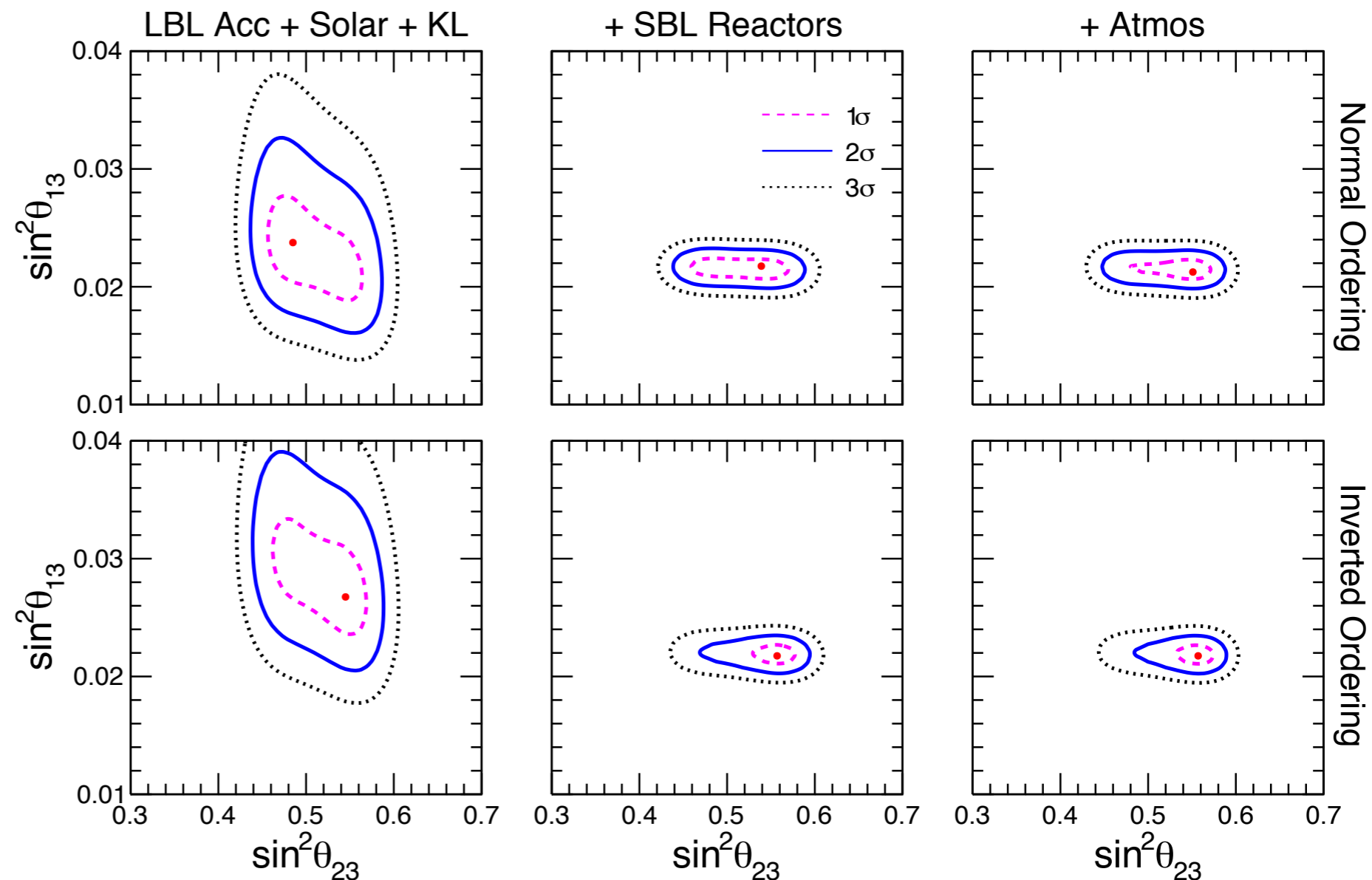
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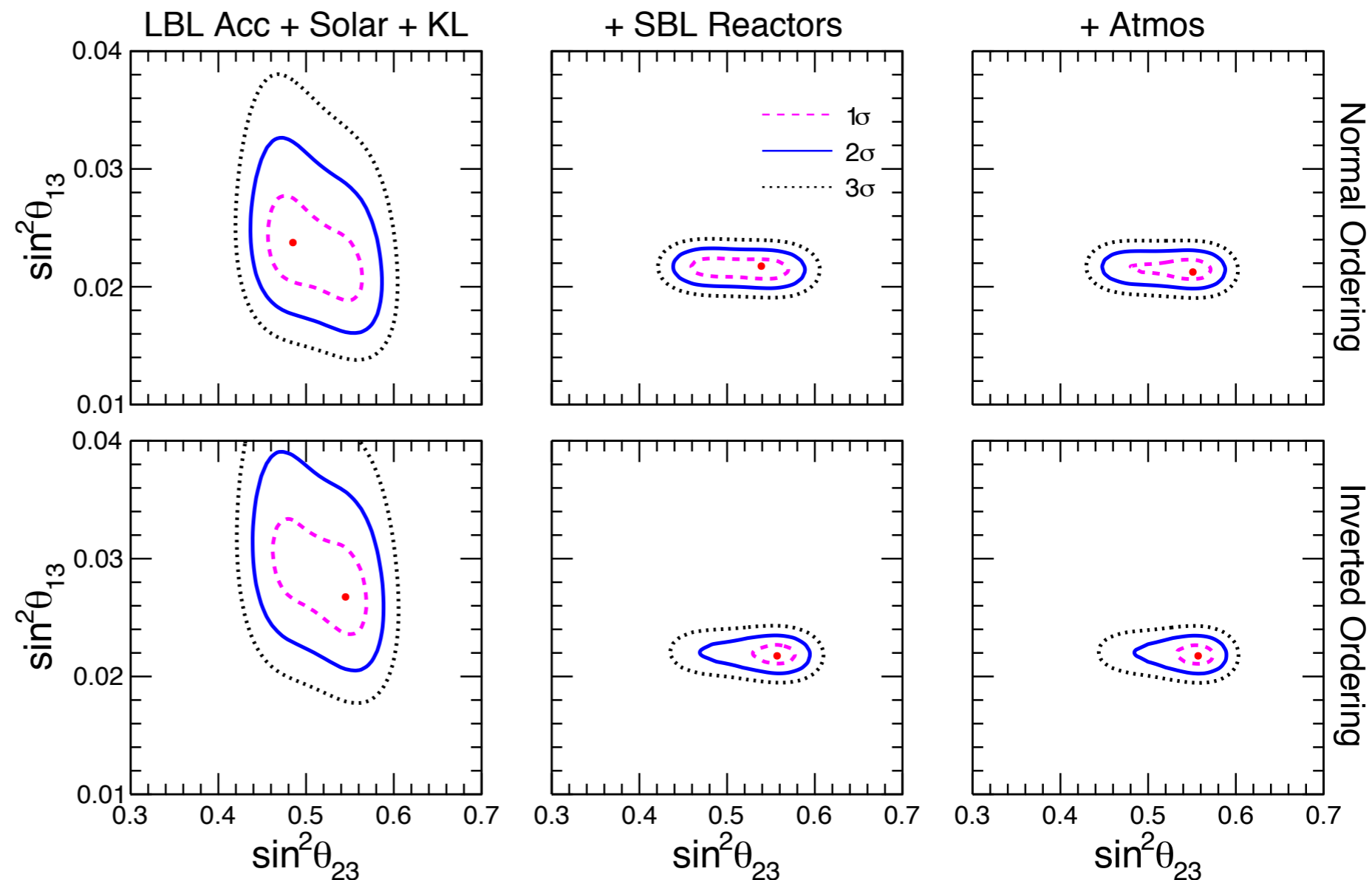
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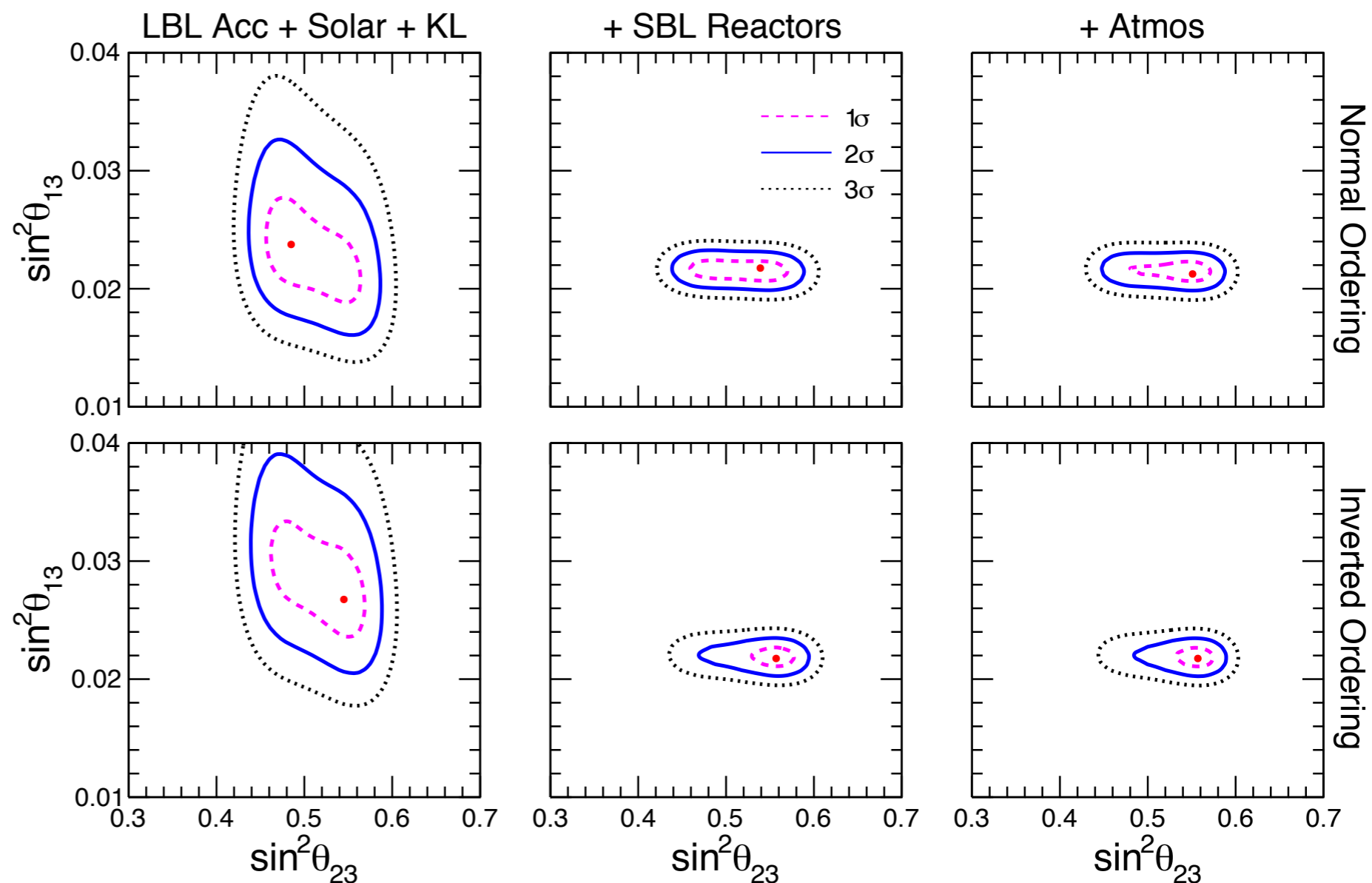
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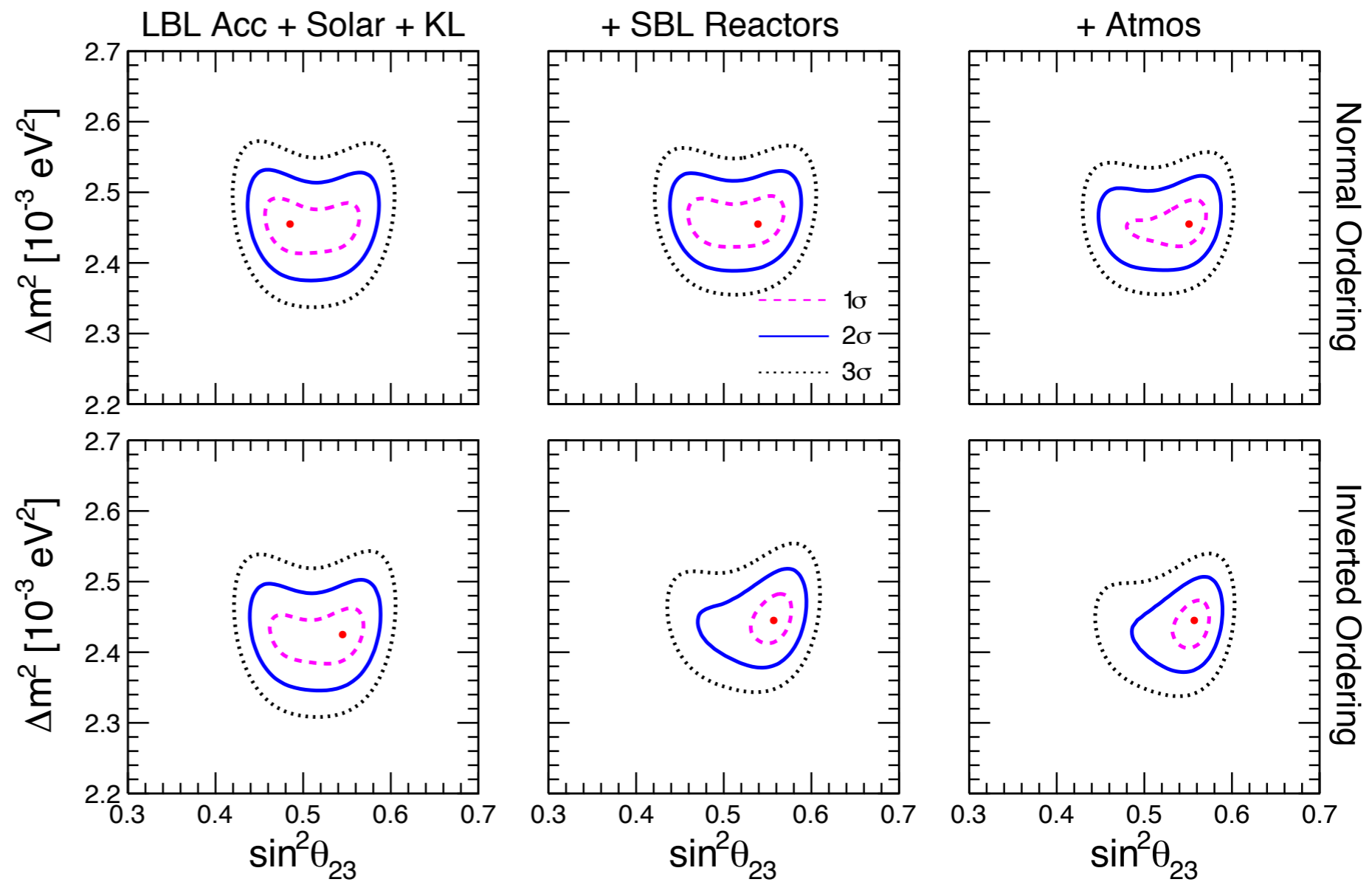
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The octant ambiguity remains unresolved at 2σ level in both NO and IO.



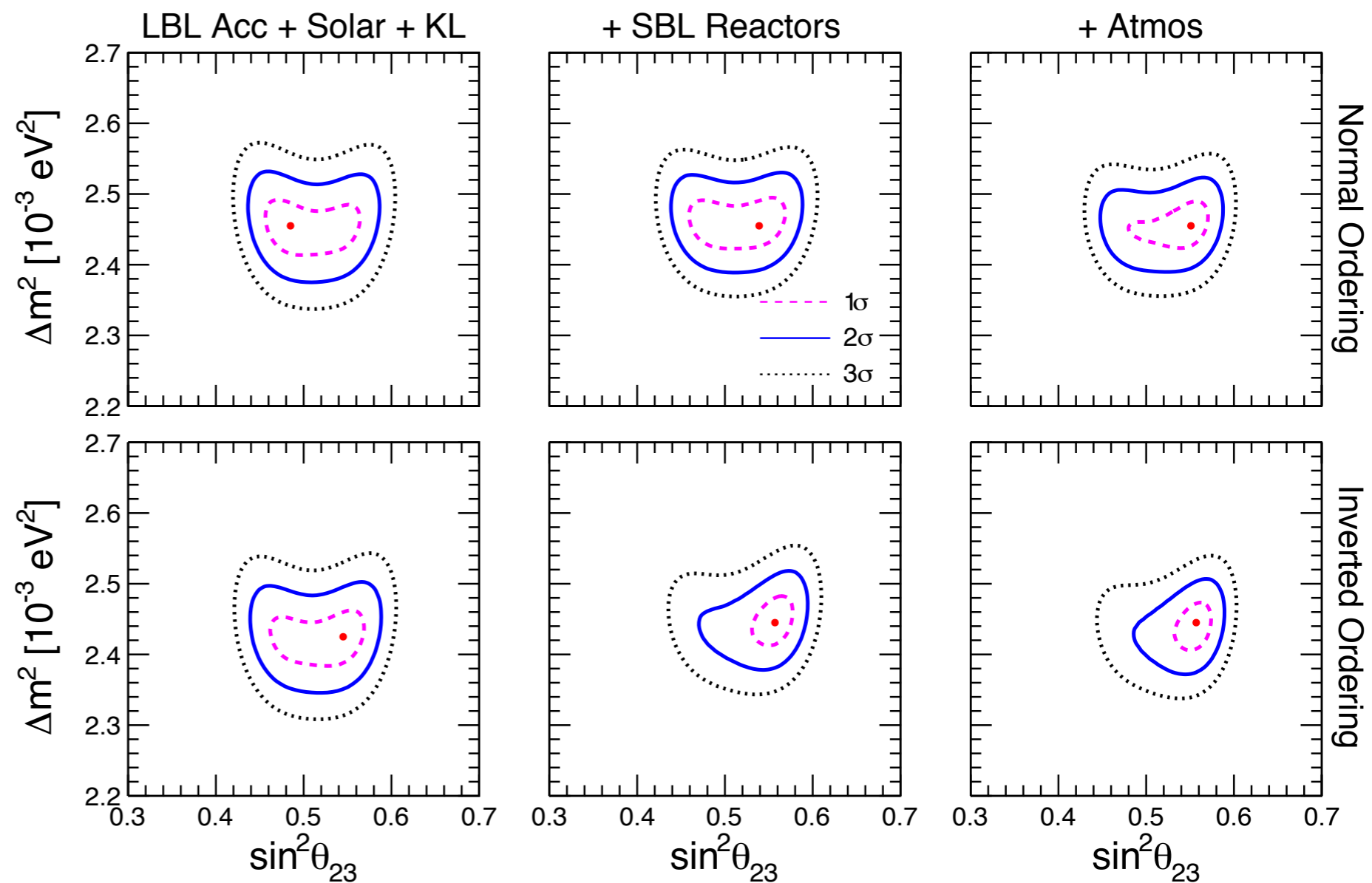
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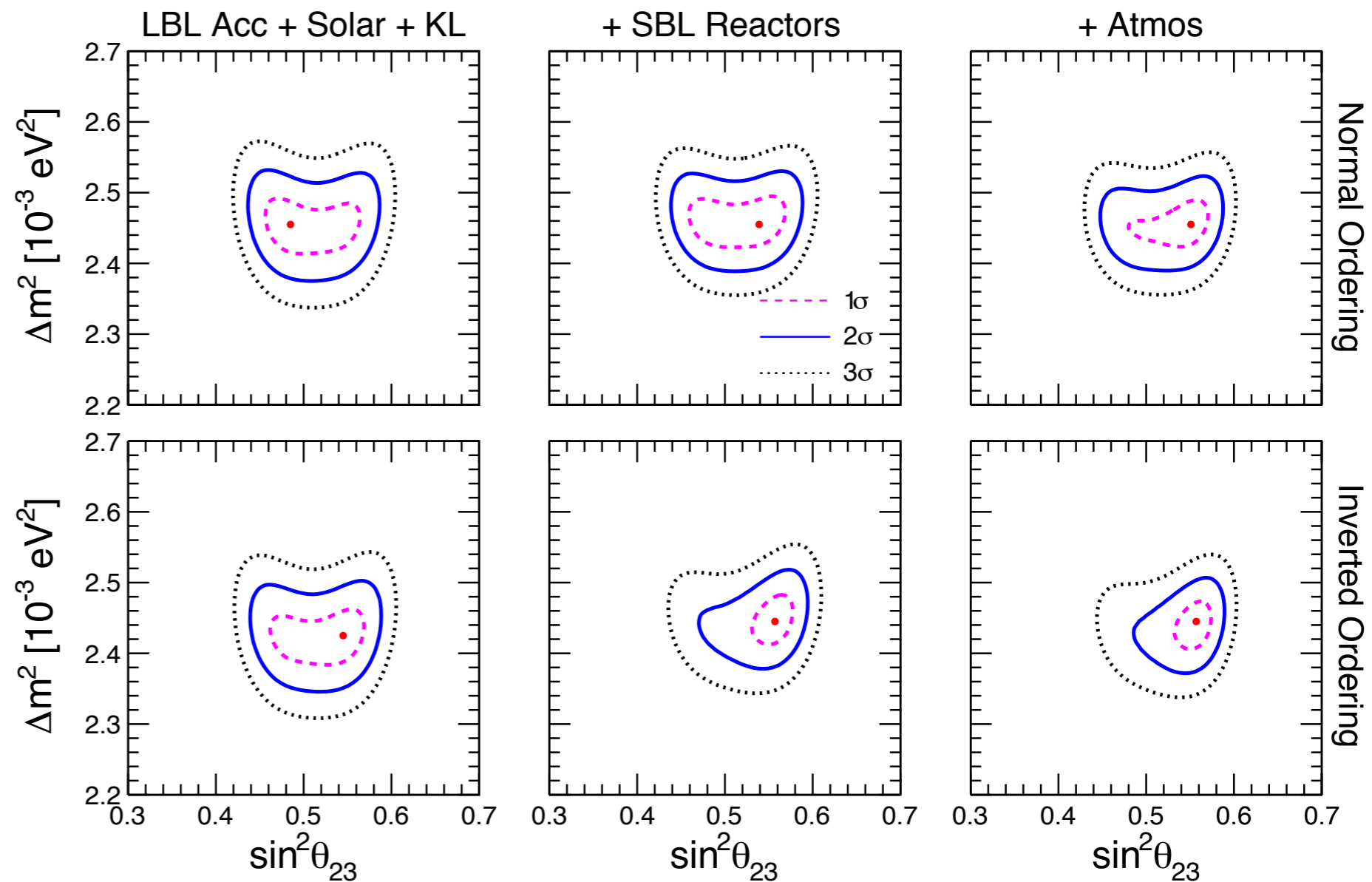
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Δm^2 slightly increases in IO after the addition of SBL reactor data as consequence of the slight increment in $\sin^2\theta_{23}$

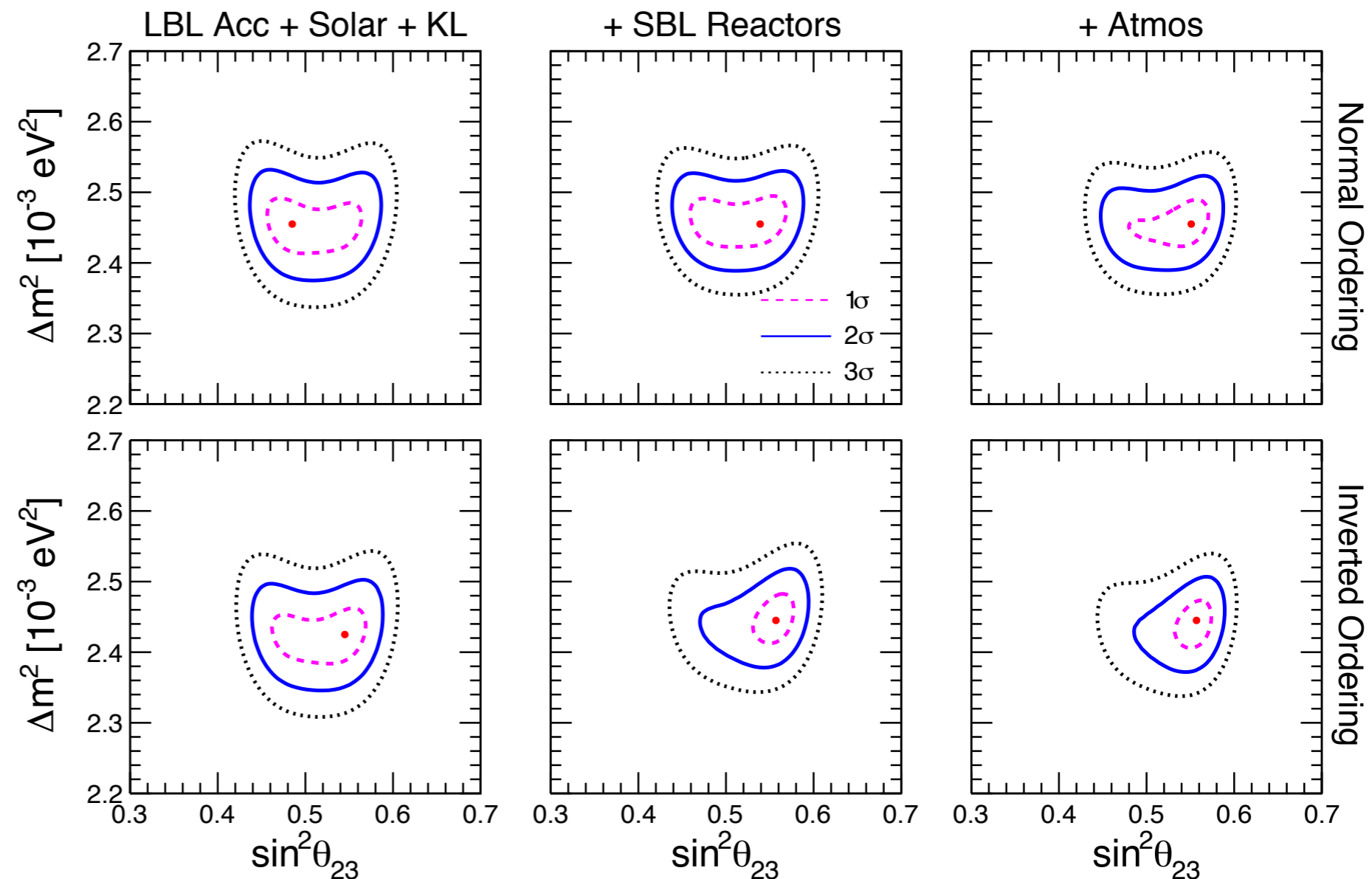


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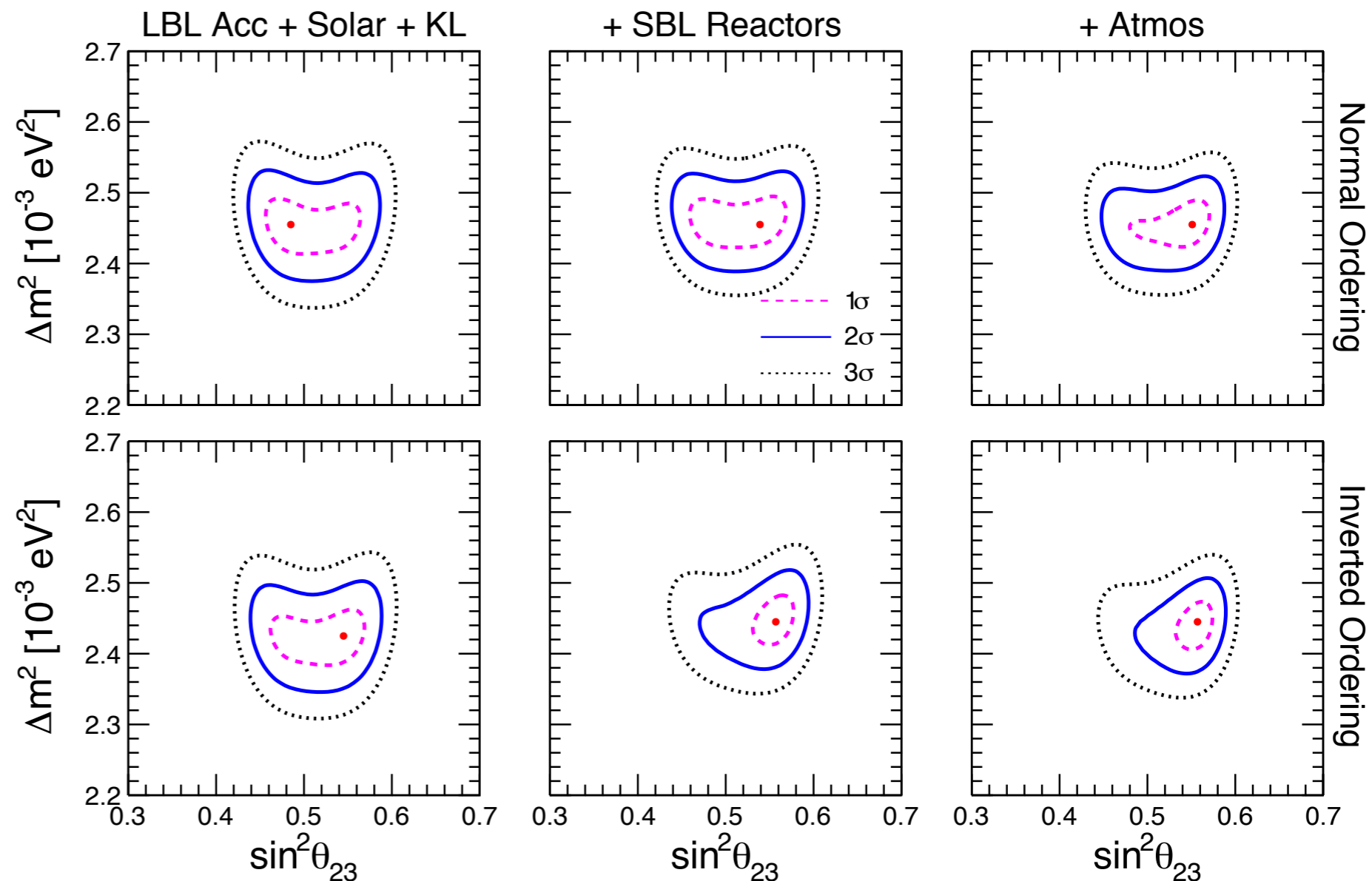
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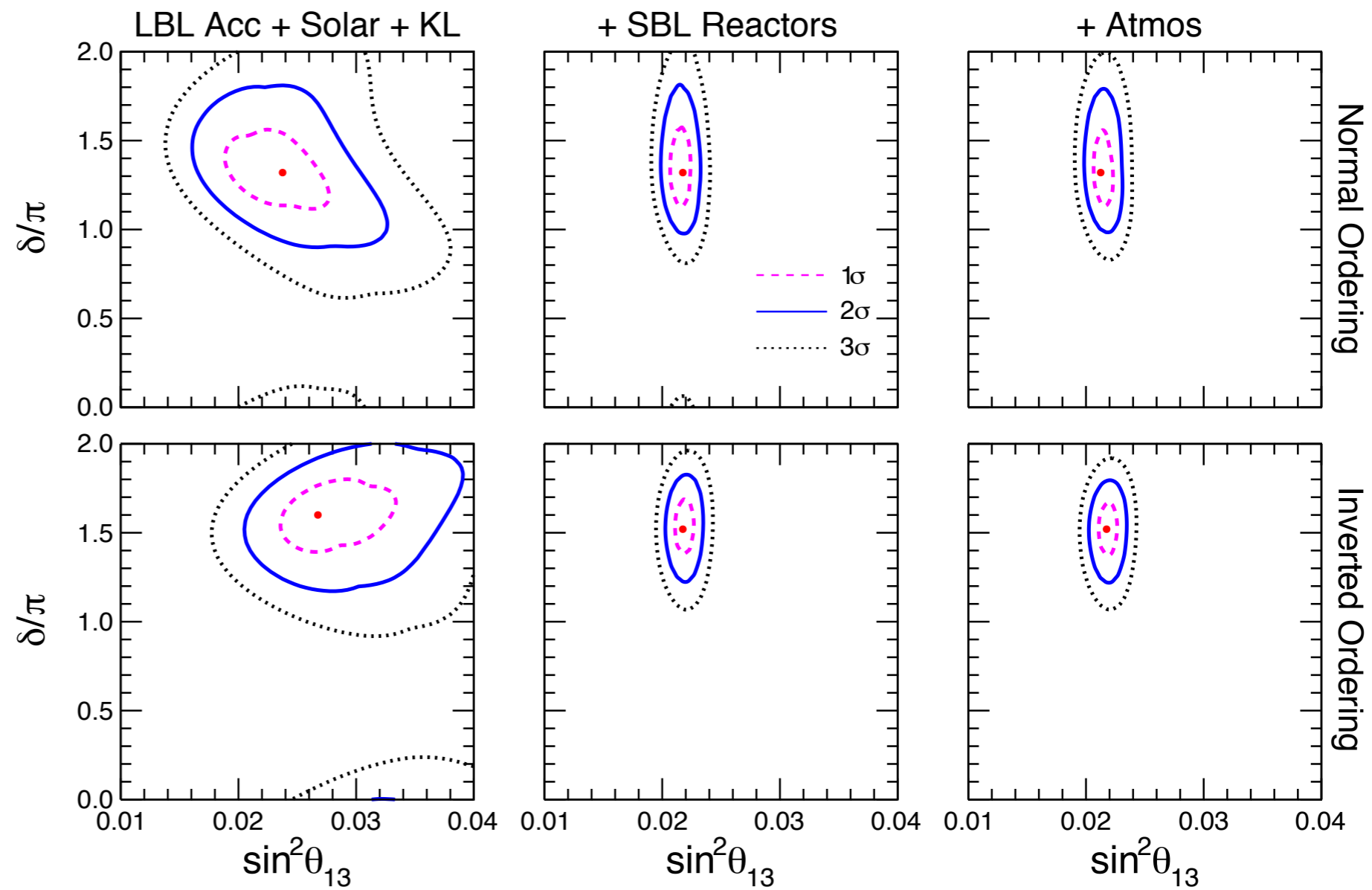
The small increase of Δm^2 slightly worsen the agreement with IC-DC data \rightarrow small contribution to $\Delta\chi^2$ (about one unit) favoring NO

In general, at nearly maximal mixing one gets the lowest allowed values of Δm^2 , while for nonmaximal mixing (in either octants) the preferred values of Δm^2 increases. Correlation mainly from disappearance data in LBL where a decrease of the leading oscillation amplitude governed by $\sin^2 2\theta_{23}$ can be compensated by an increase of the leading oscillations phase governed by Δm^2



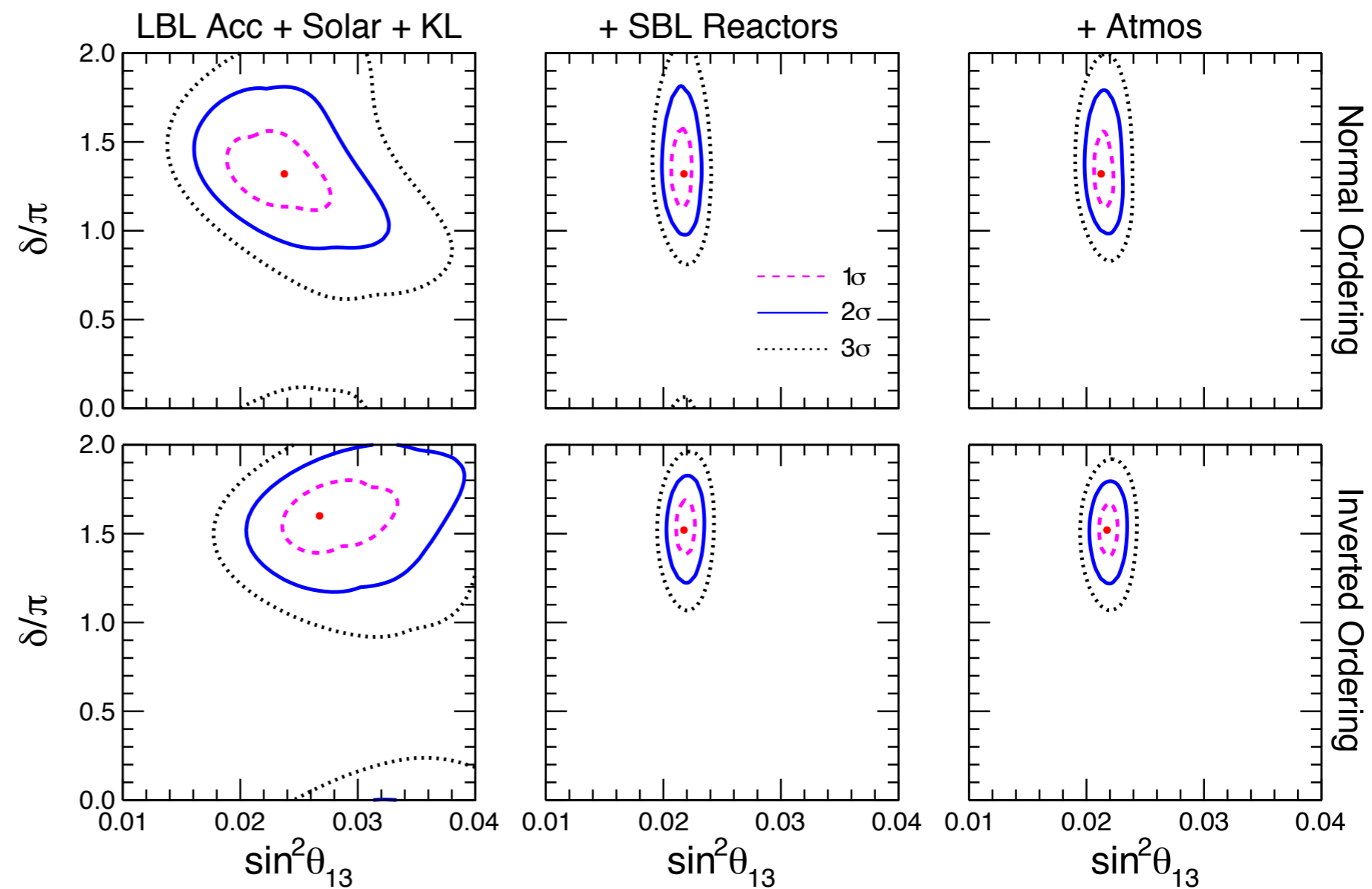
Correlations: (δ, θ_{13})

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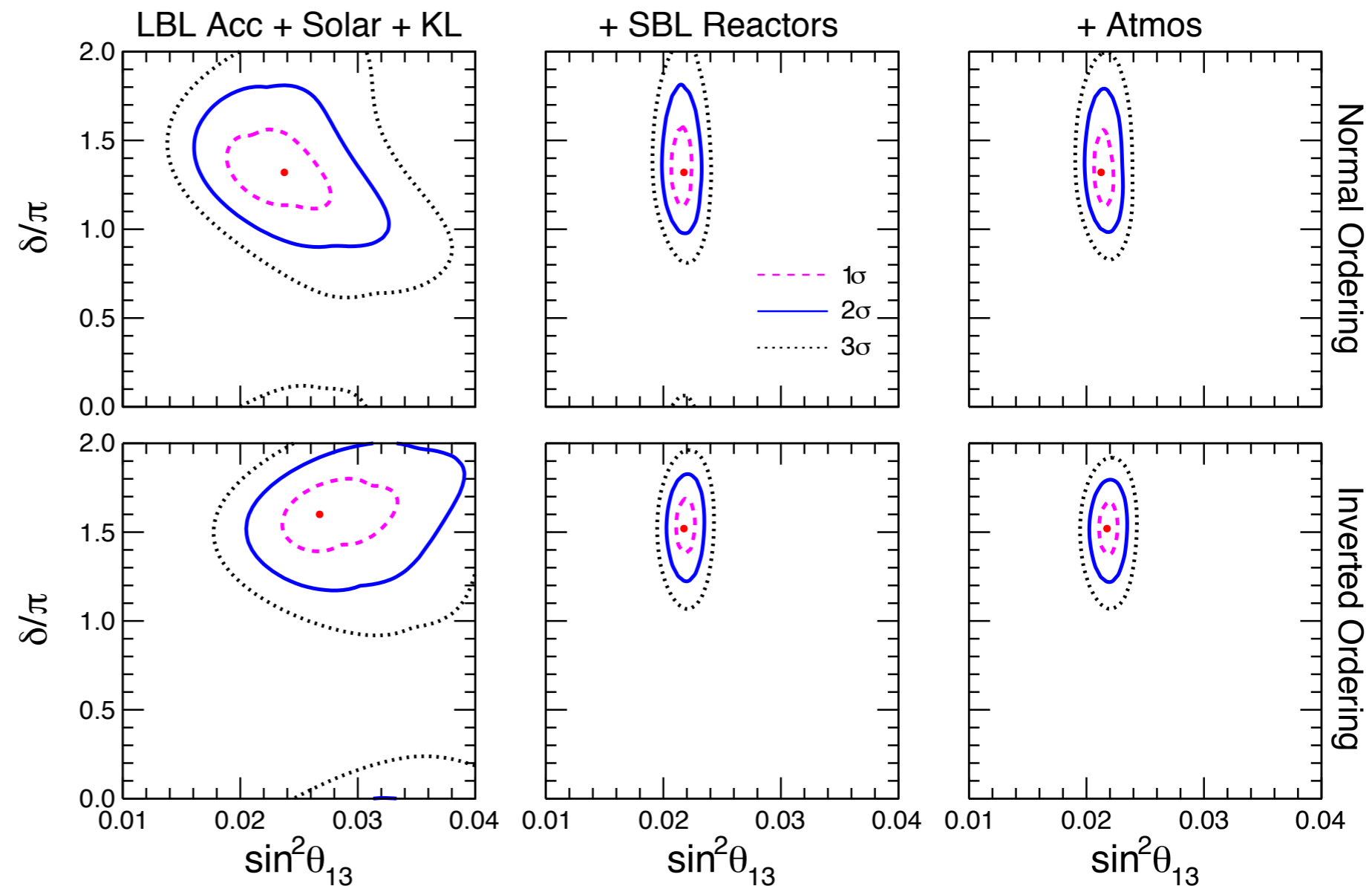
Strong correlations (in the left panels) mainly induced by the interplay between δ and θ_{13} in the subleading terms of the appearance probability for LBL experiments



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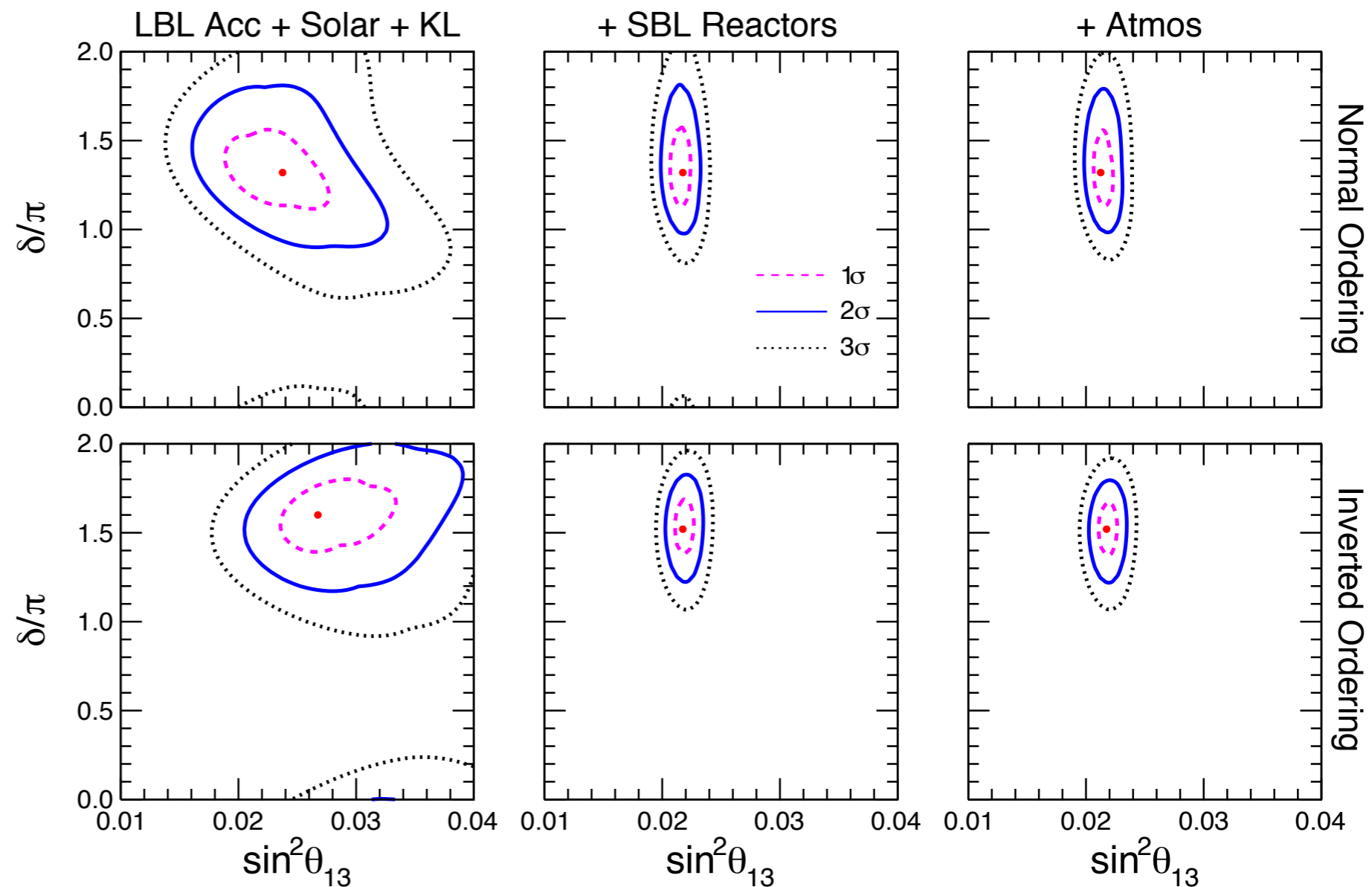


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Consistency of all the datasets towards the same best-fit values of both the Δm^2 , θ_{23} , θ_{13} and δ



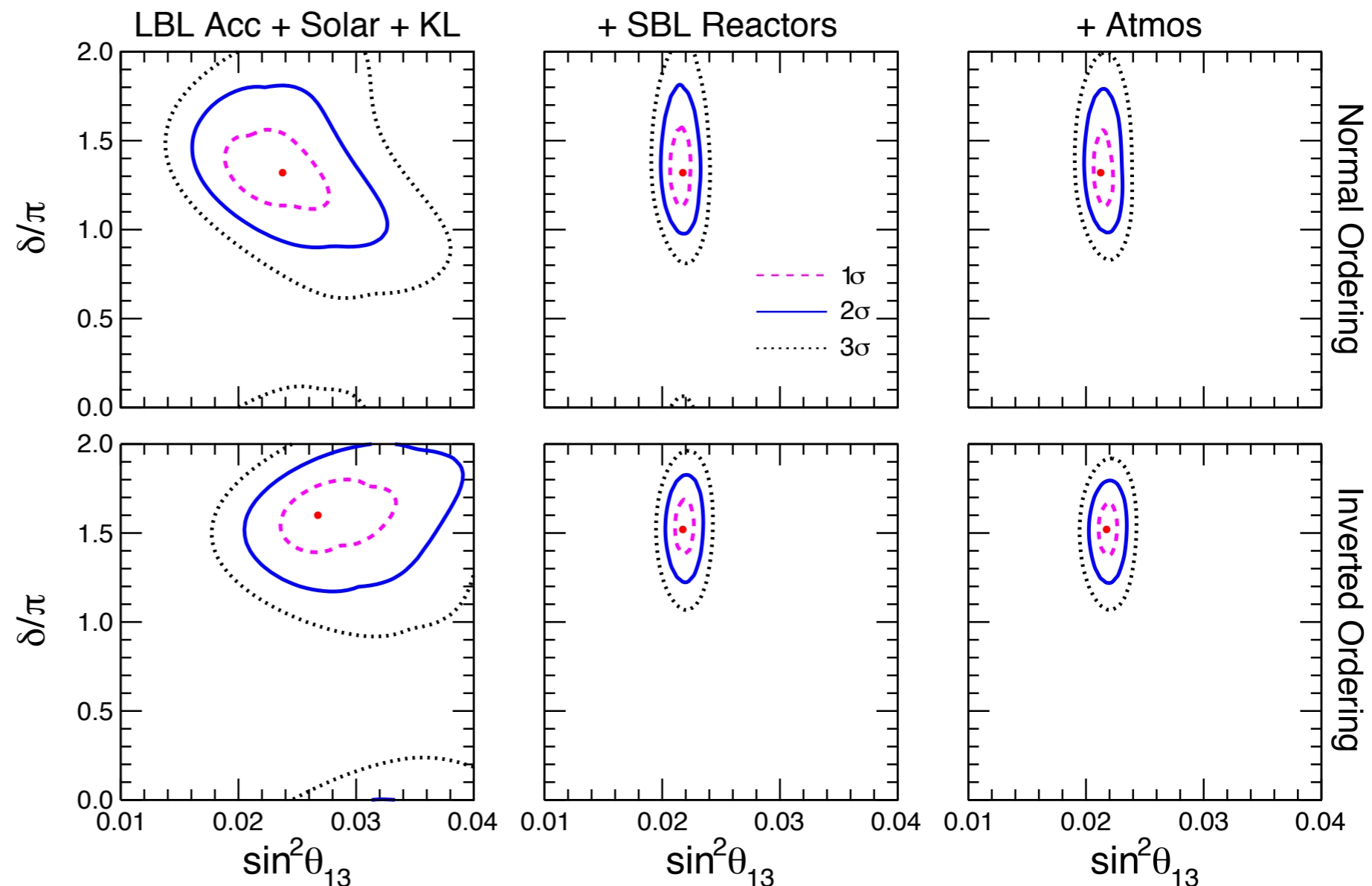
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In IO there is a slight decrease of δ from left to middle panels, correlated to the decrease of θ_{13} .



Conclusions

- Ranges of well-known 3v parameters ($\delta m^2, \theta_{12}$) & ($\Delta m^2, \theta_{13}$) confirmed by v2017-8 data updates

- CPV: $\sin\delta < 0$ preferred

best fit: $\delta/\pi \sim 1.3-1.4 \pm 0.2$ (1σ)

$\delta \sim 0$ (π) disfavoured at 2σ (3σ)

$\sin\delta \sim +1$ disfavoured at $> 4\sigma$

- Octant info: $\theta_{23} > \pi/2$, but still fragile

- Mass Ordering: IO disfavored by oscillation data:

	LBL+Sol+KL	+SBL	+ATM
$\Delta\chi^2(\text{IO-NO})$	1.3	4.4	9.5

(Non oscillation data corroborate NO)

- Info from ongoing - near future experiments