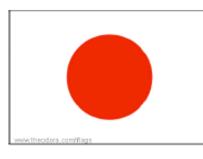
# **GRAPES-3** Detector System

Atul Jain on behalf of GRAPES-3 collaboration Cosmic Ray Laboratory (CRL) Tata Institute of Fundamental Research (TIFR) Ooty, India

## 21 FEB 2019 15<sup>th</sup> Vienna Conference on Instrumentation



#### Current status of GRAPES-3 (Gamma Ray Astronomy at Pev EnergieS (An India-Japan collaboration)



S.K. Gupta, H.M. Antia, S.R. Dugad, I. Mazumdar, P.K. Mohanty, P.K. Nayak, P. Jagadeesan, A. Jain, S.D. Morris, B.S. Rao, Y. Hayashi, S. Kawakami, S. Ogio, H. Kojima, R. Koul, V.K. Dhar, K. Venugopal, S. Das, S.K. Ghosh, S, Raha, P. Joarder, P. Subramanian, P. Jain, A. Oshima, H. Tanaka, S. Shibata, U.D. Goswami, S. Ahmad Badruddin, R. Hasan, A. Bhadra, R.K. Dey, S.K. Sarkar, C.S. Garde

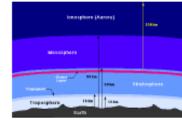
Tata Institute of Fundamental Research, Mumbai, India
 Osaka City University, Osaka, Japan
 Aichi Institute of Technology, Toyota, Japan
 Bhabha Atomic Research Centre, Mumbai, India
 J.C. Bose Institute, Kolkata, India
 Indian Institute of Science and Engineering Research, Pune, India
 Indian Institute of Technology, Kanpur, India
 National Astronomical Observatory of Japan, Tokyo, Japan
 IPMU, University of Tokyo, Tokyo, Japan
 Chubu University, Kasugai, Japan
 University of Dibrugarh, Dibrugarh, India
 Aligarh Muslim University, Aligarh, India
 North Bengal University, Siliguri, India

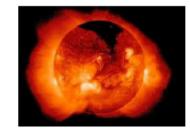
14. Vishwakarma Institute of Information Technology, Pune, India

## **Objective: Universe at high energies**

Acceleration, propagation of high energy particles, Extreme conditions may require new physics ...

- 1. Acceleration in atmospheric electric field Energy  $\sim 100 \text{ MeV}$
- 2. Solar flares, Coronal Mass Ejections Energy ~10 GeV
- 3. Galactic Cosmic Rays at "Knee" Energy ~1 PeV
- 4. Diffuse multi-TeVγ-rays Energy ~100 EeV









#### All of the above science requires the following

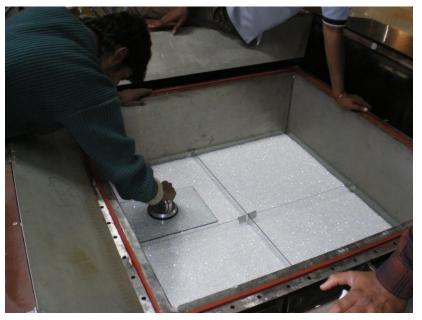
- High performance detectors
- High speed signal processing
- ✓ Large data acquisation systems (DAQ)

Commercial equipment is expensive and does not always meet our exacting requirements
 Indigenous development of detectors and electronics has become a necessity for us

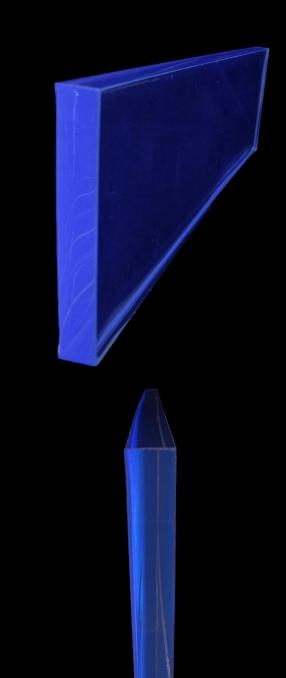
#### **Plastic Scintillator Casting @ GRAPES-3 Ooty**



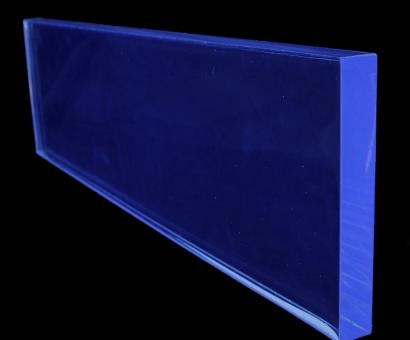


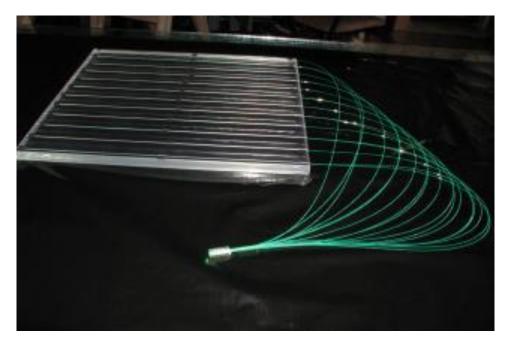






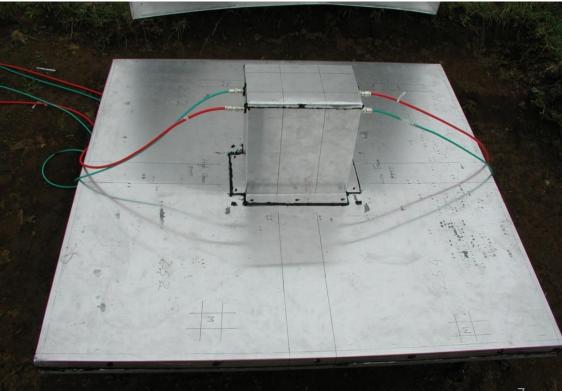
#### **GRAPES-3** Plastic Scintillator











## **GRAPES-3 Plastic Scintillator**

- These state-of-the-art scintillators have a high photon output (55% of anthracene)
- Long attenuation length (>100 cm)
- Short decay time (1.5 nanosecond)
- With the aid of wavelength shifter (WLS) fibre readout, the GRAPES-3 scintillator detector provides a uniform response (r.m.s. non-uniformity <2.5% over 1m<sup>2</sup> area)
- Large signal from minimum ionizing particles (18 photo-electrons)
- Large dynamic range (1-500 particles with single-PMT, >10000 particles with dual-PMT readout) P.K. Mohanty et al. Astropart. Phys. 31 24 (2009)
- In addition, a Monte Carlo code G3sim developed by the GRAPES-3 team allows precision simulation of the response of assembled scintillator detector with WLS readout P.K. Mohanty et al. Rev. Sci. Instrum. 83, 043301 (2012)
- G3sim was used to optimize the design of 200 new scintillator detectors used during expansion of the GRAPES-3 experiment as well as to design detectors for other collaborating institutions

# Plastic scintillation detectors - 400 Spread over an area - 25,000 m<sup>2</sup>

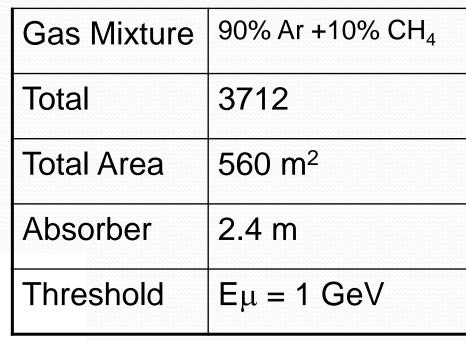
# GRAPES-3 CRL, TIFR, Ooty, India

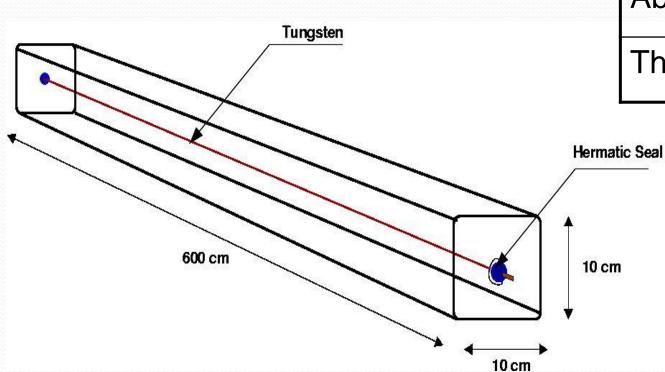
A new A see 200 A 100 A 100 A 100 A 200 A 200 A 200 A 200 A 200 A 100 A 200 A A 1995 A 1976 A 1997 A 199 A 199 A 199 A 199 A 199 A 197 A 199 A 195 A 194 A 195 A 199 A 201 \* 307 \* 500 \* 455 \* 424 \* 355 \* 320 \* 354 \* 320 \* 326 \* 326 \* 327 \* 326 \* 326 \* 326 \* 326 \* 326 \* 320 \* 340 \* 340 \* 340 \* 340 \* and A 2011 A 2011 A 2011 A 2021 A 202 \* 654 \* 557 \* 558 \* 459 \* 559 \* \*\*\*\* ★ and ★ see ★ see ★ and ★ and ★ see k see A and A sea A dea A dea A dea A sea × and have & see & \* 201 \* A case & 507 & 500 & 401 & 572 & 500 & 502 & 501 & 155 & 607 & 155 & 605 & 160 & 605 & 160 & 601 & 165 & 605 & 505 A 600 A 644 A 518 A 445 A 574 A 501 A 524 A 555 A 256 A 557 A 258 A 557 A 268 A 551 A 200 A 551 A 200 A 555 A 555 A 550 A 651 A 650 A case & soy & soo & see & soo nor & now & nor & EAS Landing at GRAPES-3 \* age \* and A see A res A 662 A 664 A 664 A 668 A 669 A 660 A 660 A 700 A 700 A 702 A 703 A 704 A 705 A 705 A 705 A 705 A 705 A 705

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# Proportional CounterTotal area

3712 nos
560 m<sup>2</sup>

## **Expansion of Muon Detectors**

• Double the area of Muon Detectors

560m <sup>2</sup>	$\rightarrow$	1130m <sup>2</sup>
3712nos	$\rightarrow$	+3776 = 7488nos

#### Fabrication and Installation of PRCs

- ✤ Target 3776 nos
- Status as on 14 Feb 2019
- Fabrication 3803 nos
- Installed in Field 3009 nos
- Final Test Bench 645 nos
- Ready for Final Test 149 nos





## **PRC Cleaning**





# **PRC end plate welding**



### Needle valve assembly and brazing





Needle valve



#### Counter fixed with needle valve

#### Tungsten wire assembly & hermetic seal fixing





**Tungsten wire** 



Hermetic seal

### Leak detection, evacuation and gas filling











### Placement of PRCs in the field





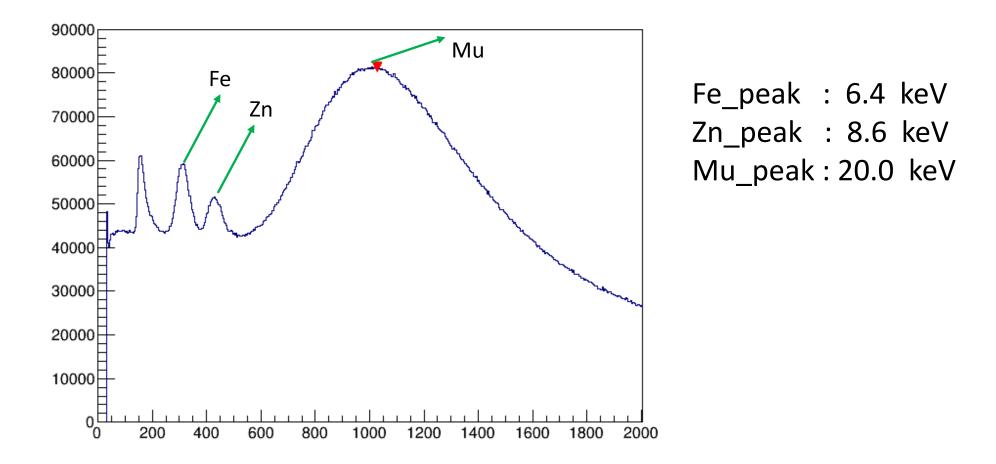


PRC Mobile Test Bench

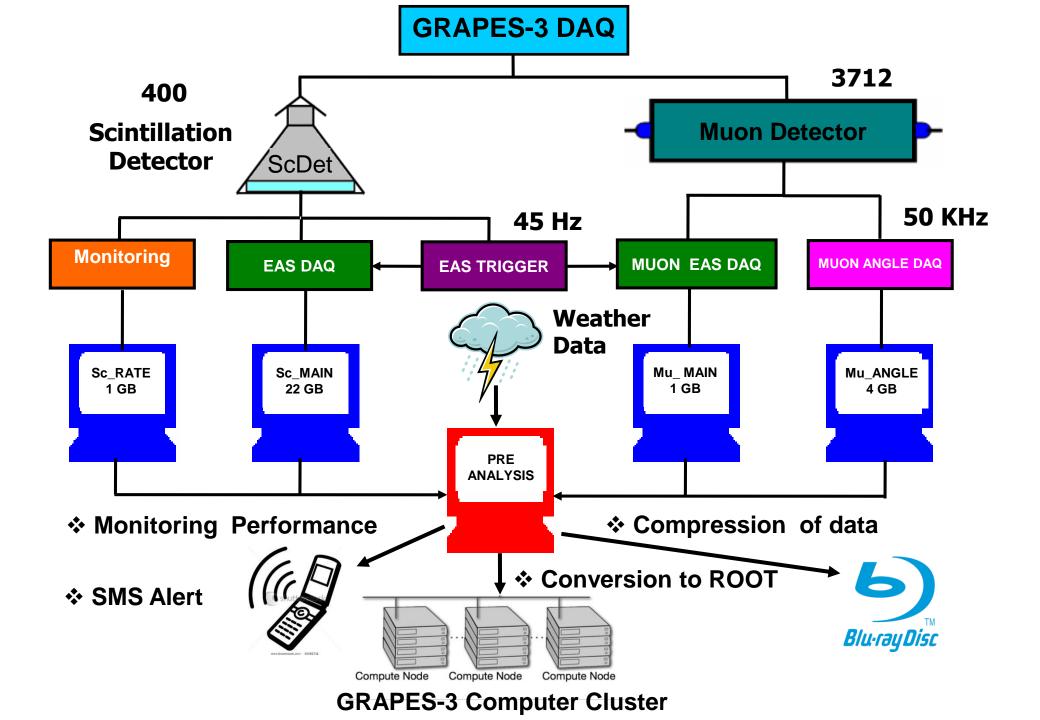




### **Proportional Counter Performance**

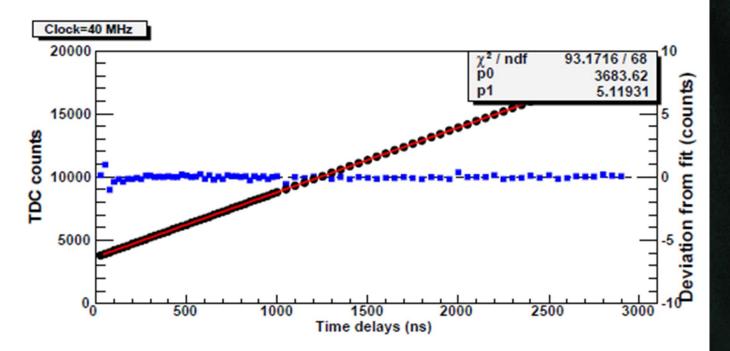


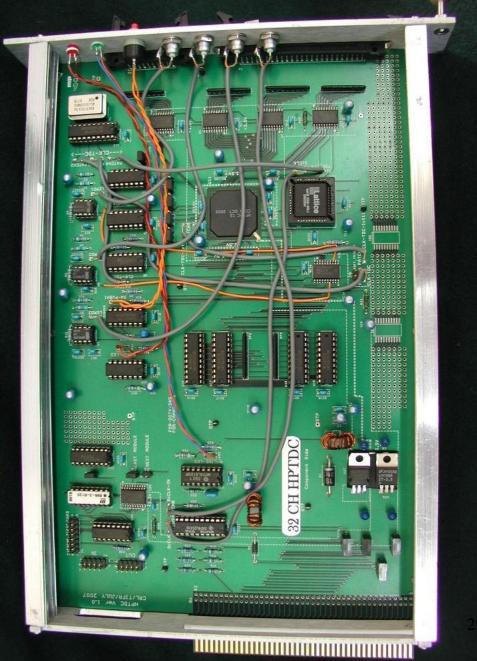
The muon tracking telescope provides a high statistics, directional measurement of the muons, and permits small changes of ~0.1% in the muon flux to be accurately measured on a time scale of ~5 min



# GRAPES-3 High Performance Time to Digital Converter (HPTDC)

- ✓ 32 Channel CAMAC Interface in Single Module
- ✓ 24 psec Resolution
- ✓ External Clock 40 MHz. (Internally 80,160,320 MHz)
- ✓ Dynamic Range 51 µsec
- ✓ Double Pulse Resolution 5 nsec
- ✓ Trigger Mode of Operation

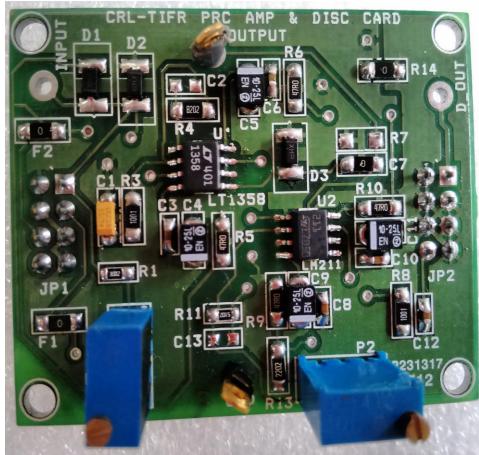




#### GRAPES-3 Integrated Amplifier and Discriminator Card for Proportional Counter

- ✓ Programmable amplification and decay constant adjustment
- ✓ On board protection for transient voltage and short circuit
- $\checkmark\,$  Compact card with plug and play design
- ✓ 7244 / 8000 cards are ready
- ✓ 2060 has been installed





# Building Blocks

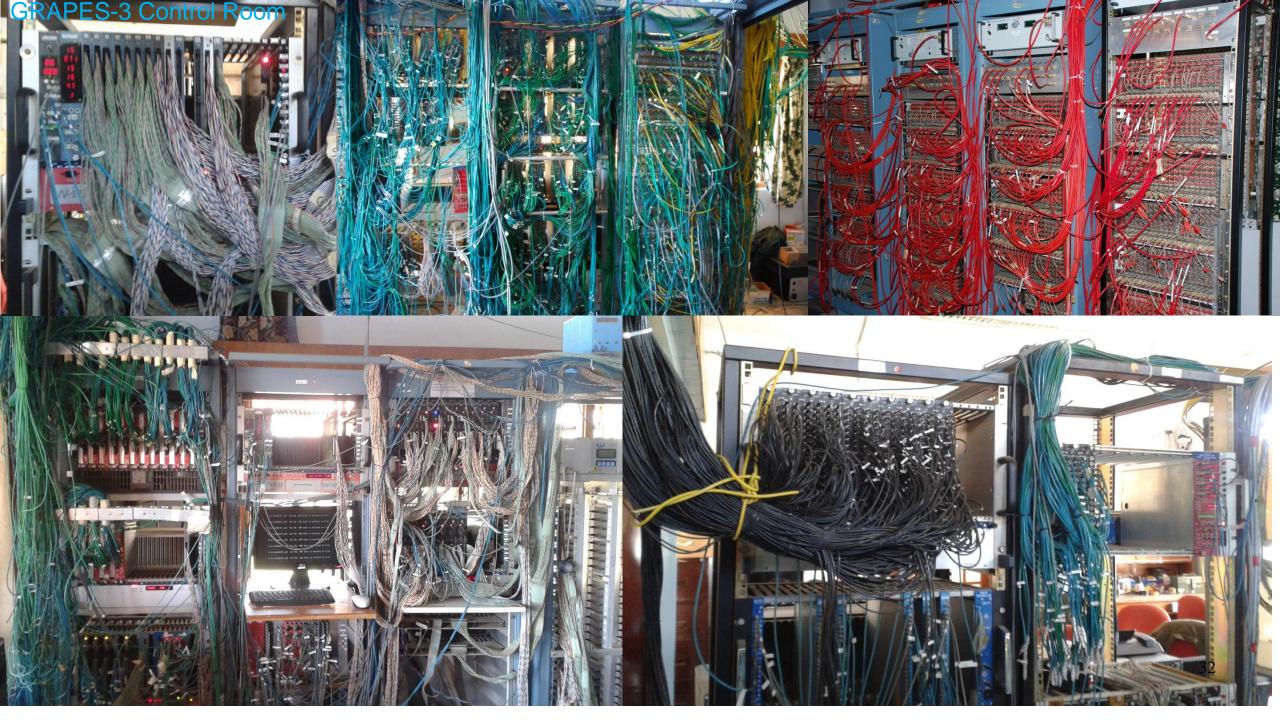


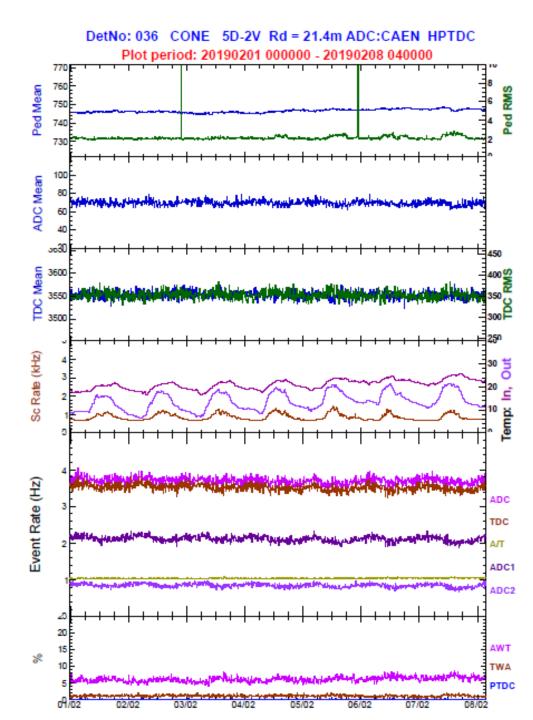






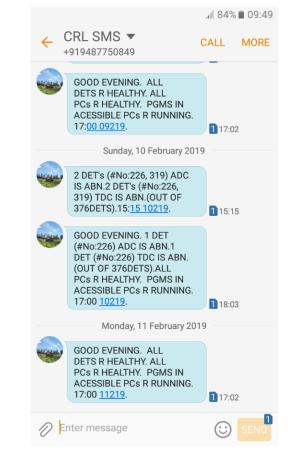
- > 8 channel Integrated Fast Amplifier and Discriminator module
- FPGA based pulse width analyser
- > Multipurpose fast logic unit and scalar module
- TCP/IP interface board for 32 Digital I/O's
- USB interface board for 16 Digital I/O's
- > 32 channel HPTDC
- > High voltage monitoring system
- Temperature Scanner and Controller
- TTL to optical link





## **GRAPES-3 Monitoring Tools**

- ROOT based data analysis tools
- Web based parameter monitoring
- SMS alerts in case of abnormality in DAQ



# **GRAPES-3 Computer Cluster**

- Nodes 40
- Memory 1280GB
- Storage 660GB
- Optical N/W 10Gbps
- Rocks Cluster Ranking 36



week ending 21 OCTOBER 2016

#### Transient Weakening of Earth's Magnetic Shield Probed by a Cosmic Ray Burst

(C) MF Bz (nT) Interplanetry magnetic field (IMF) Data from OMNI Web Measured by wind space craft Located at L1 (1.5 million kms from Earth) -40Muon rate variation ( -1.0 -1.5 -2.0 -2.5 (d) **Muon Flux Variation** Data from GRAPES-3 experiment, Measured by Muon Tracking Telescope Located at Ooty, India 18 h 21-Jun-15 22-Jun-15 23-Jun-\$5



#### **Media Coverage**

- Countries 119

1096 – Reports in 37 Languages

APS Physics -

#### Synopsis: A Crack in Earth's Protective Shield

VOLUME

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ENGLISH

LANGUAGE

UNE 2017

LANGUAGES

UNE 2017

**JUNE 2017** 

October 20, 2016

**EXAMPLE O** 

Observations with India's cosmic-ray teleso weakened during a 2015 geomagnetic storn

#### **GRAPES-3** indicates a crack in Earth's magnetic shield

Phys.org/news/2016-11-grapes-earth-magnetic-shield.html

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**COVERAGE OF TRANSIENT WEAKENING OF EARTH'S MAGNETIC SHIELD PROBED BY GRAPES-3** 



UNE

2017

JUNE 2017

Cosmic Ray Laboratory, Ooty





# Thanks

Web : grapes-3.tifr.res.in Email : atul@crl.tifr.res.in