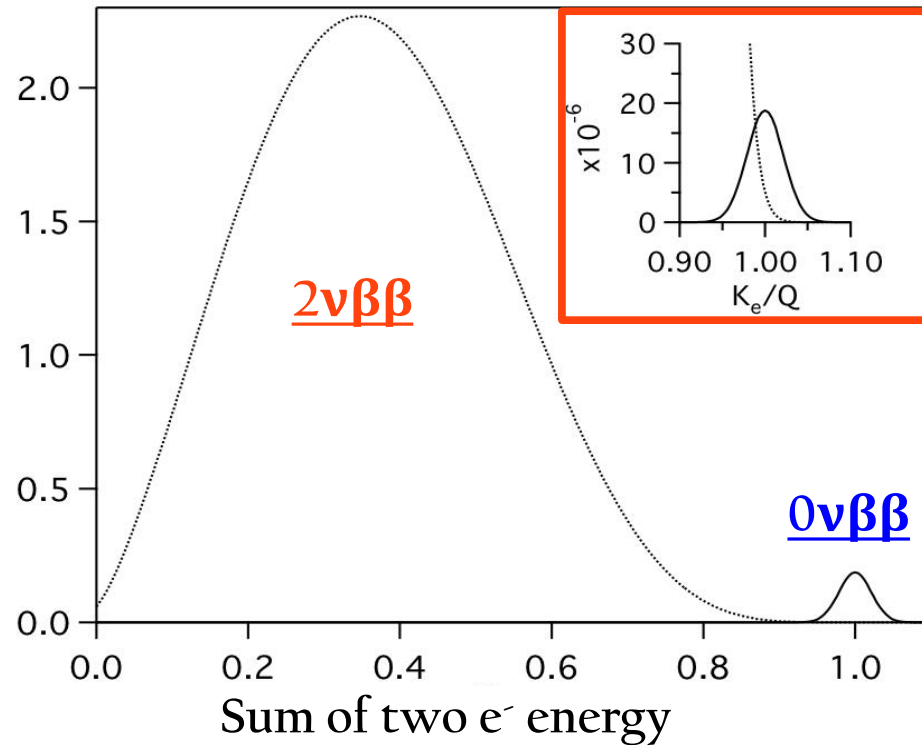

Final results of the CUPID-0 Phase I experiment

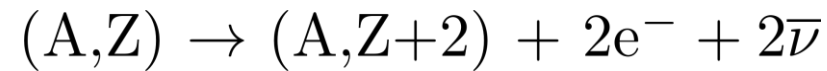
Mattia Beretta

On behalf of the CUPID-0 collaboration

The first enriched scintillating bolometer $\beta\beta$ experiment



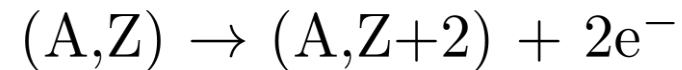
$2\nu\beta\beta$



^{82}Se

$$Q_{\beta\beta} = (2997 \pm 0.3) \text{ keV}$$

$0\nu\beta\beta$:



Performing resolution
At the Q value

Low Background
Few counts expected

CUPID-0 Detector



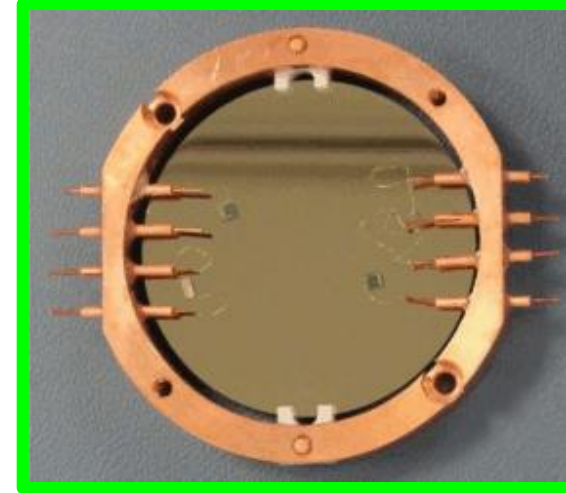
Heat signal:
bolometric high
resolved output

- 26 ZnSe crystals
 - 24 enriched in ^{82}Se (95%)
+ 2 naturals
- Total mass= 10.5 kg
 - ^{82}Se mass =5.17 kg ($3.8 \cdot 10^{25}$
 $\beta\beta$ emitters)

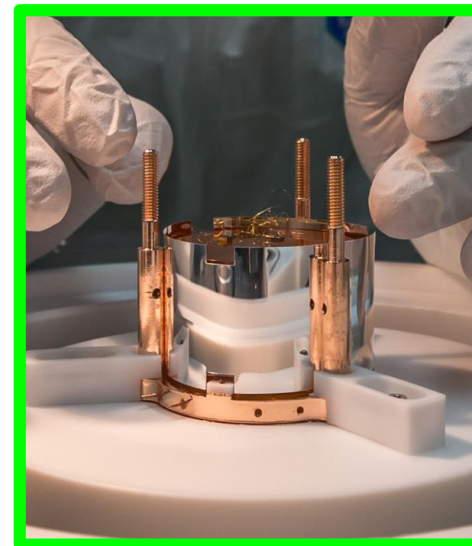
Scintillating
 Zn^{82}Se crystals.



Bolometric Ge Light
detectors



Light signal:
particle
identification



Vikuiti Reflector
More collected light

NOSV Copper
Surface cleaned

CUPID-0 Time-line



Commissioning

Phase I

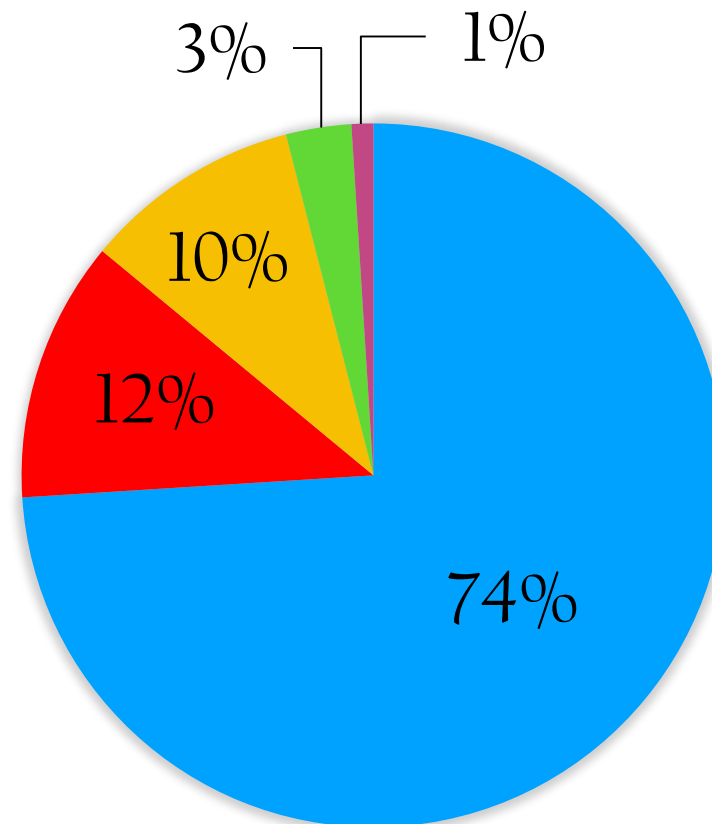
560 d with 74% of livetime

Preparation of
Phase II

^{56}Co Energy Calibration

^{232}Th Energy Calibration

System maintenance



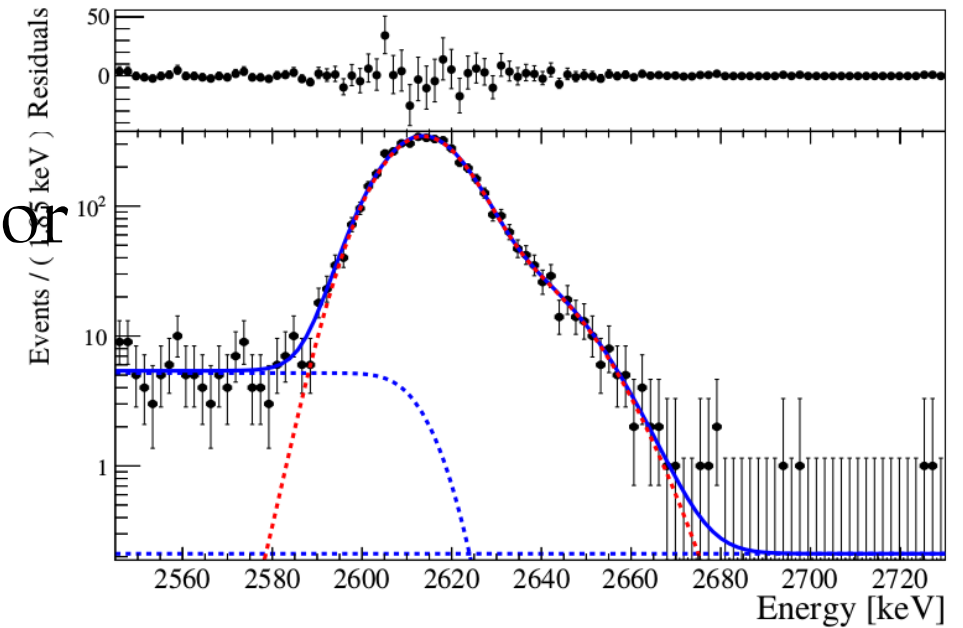
AmBe source
 $\beta\gamma$ Shape Characterization in
the ROI

$\beta\beta$ physics
Exposure: 9.95 kg·y

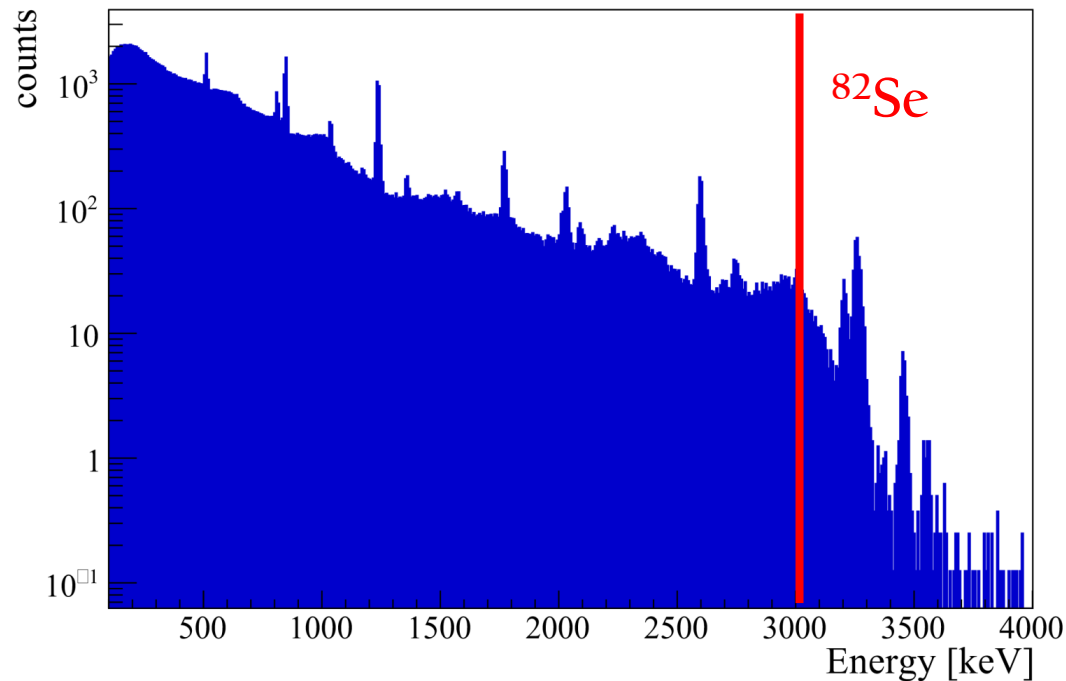
Calibrations

^{232}Th Energy Calibration

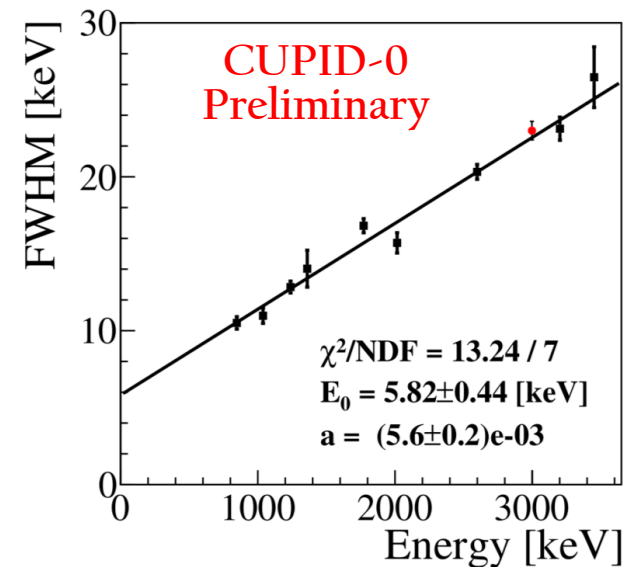
- Periodic Bolometer calibration and light detector intercalibration
- Response function: Double Gaussian
 - Also observed in other bolometers



^{56}Co Energy Calibration



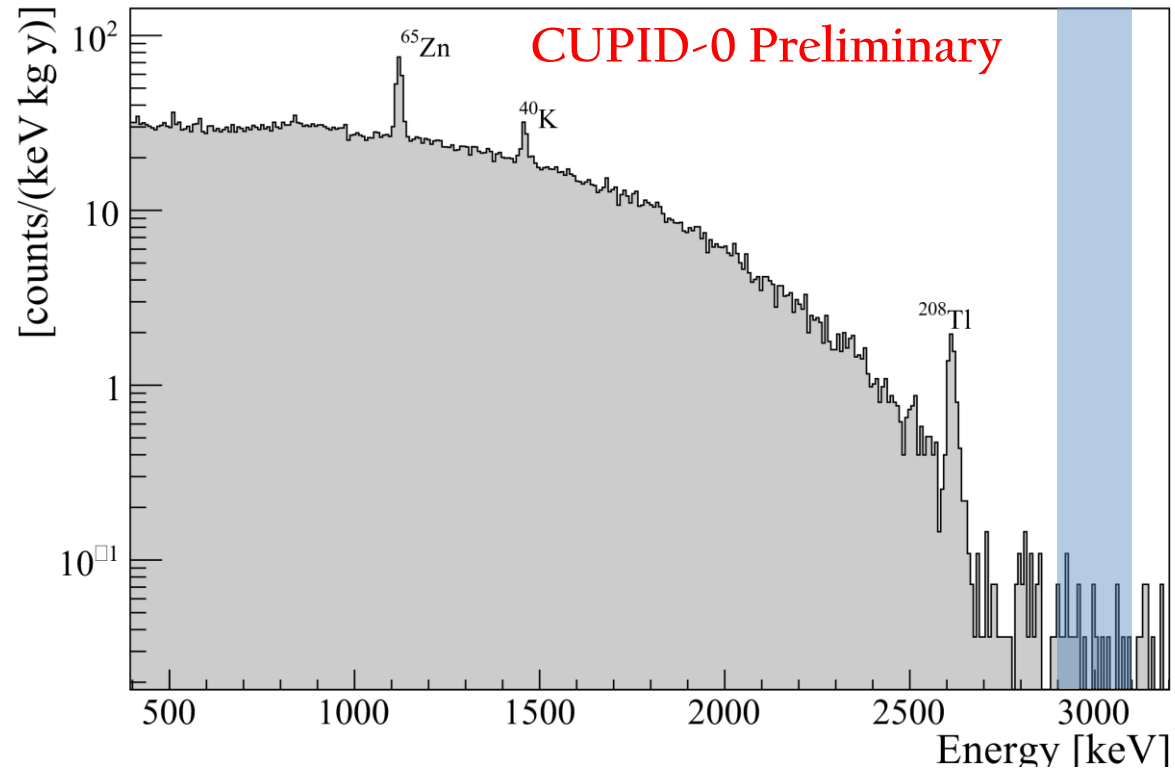
- Check of the energy reconstruction
- Evaluation of FWHM energy resolution @ ^{82}Se Q



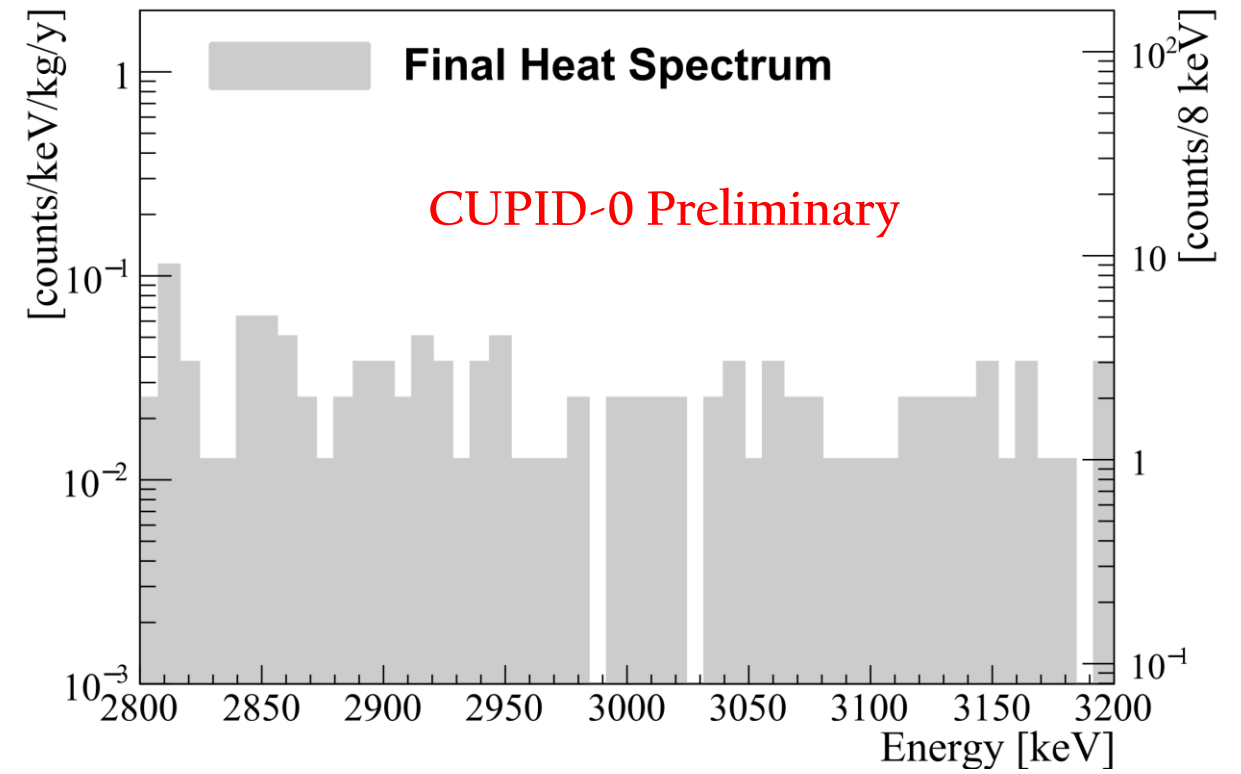
FWHM @ $Q_{\beta\beta}$
(20.0 ± 0.6) keV

$0\nu\beta\beta$ search

Total Background spectrum



$0\nu\beta\beta$ ROI



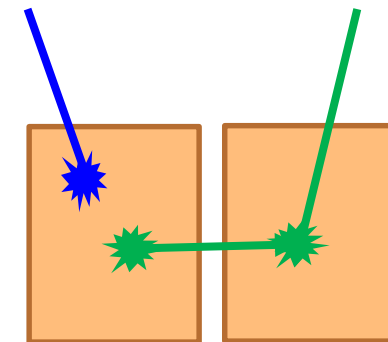
$$\text{BKG} = (3.2 \pm 0.4) \cdot 10^{-2} \text{ cnts}/(\text{keV} \cdot \text{kg} \cdot \text{y})$$

Basic Selections

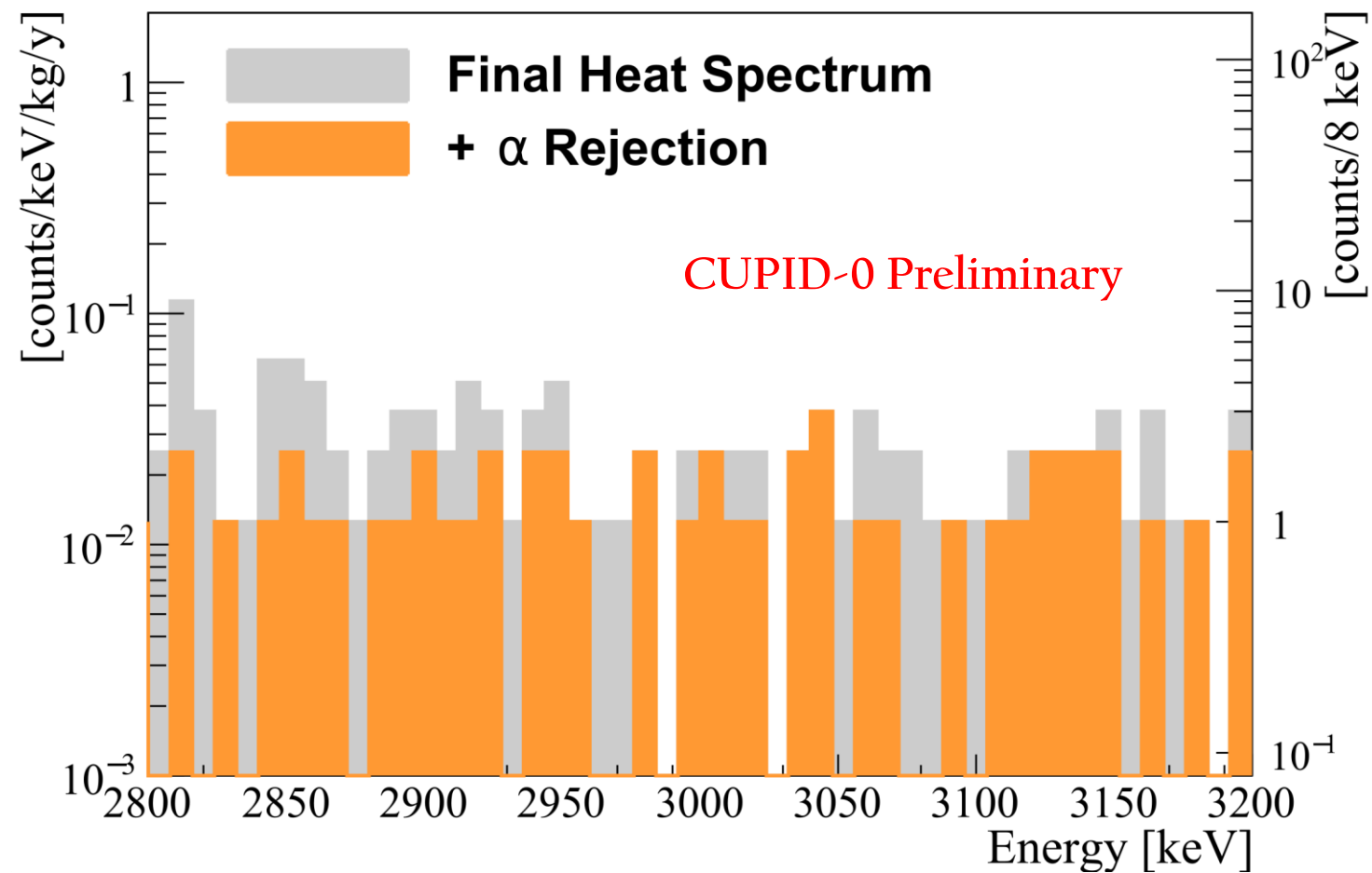
Rejection of “non-particle-like” events through pulse shape on thermal pulses

Multiplicity (M)

Anti-coincidence between crystals ($\Delta T=20\text{ms}$)



Background – Alpha Rejection

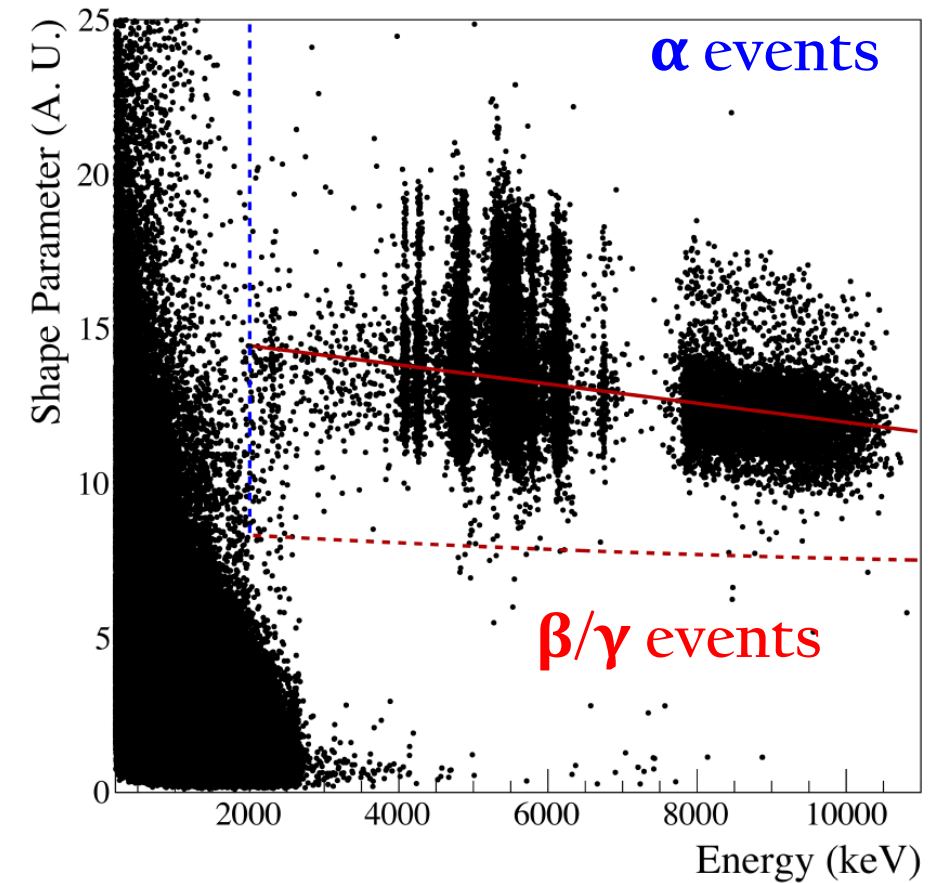


$$\text{BKG} = (3.2 \pm 0.4) \cdot 10^{-2} \text{ cts}/(\text{keV} \cdot \text{kg} \cdot \text{y})$$

$$\text{BKG} = (1.3 \pm 0.2) \cdot 10^{-2} \text{ cts}/(\text{keV} \cdot \text{kg} \cdot \text{y})$$

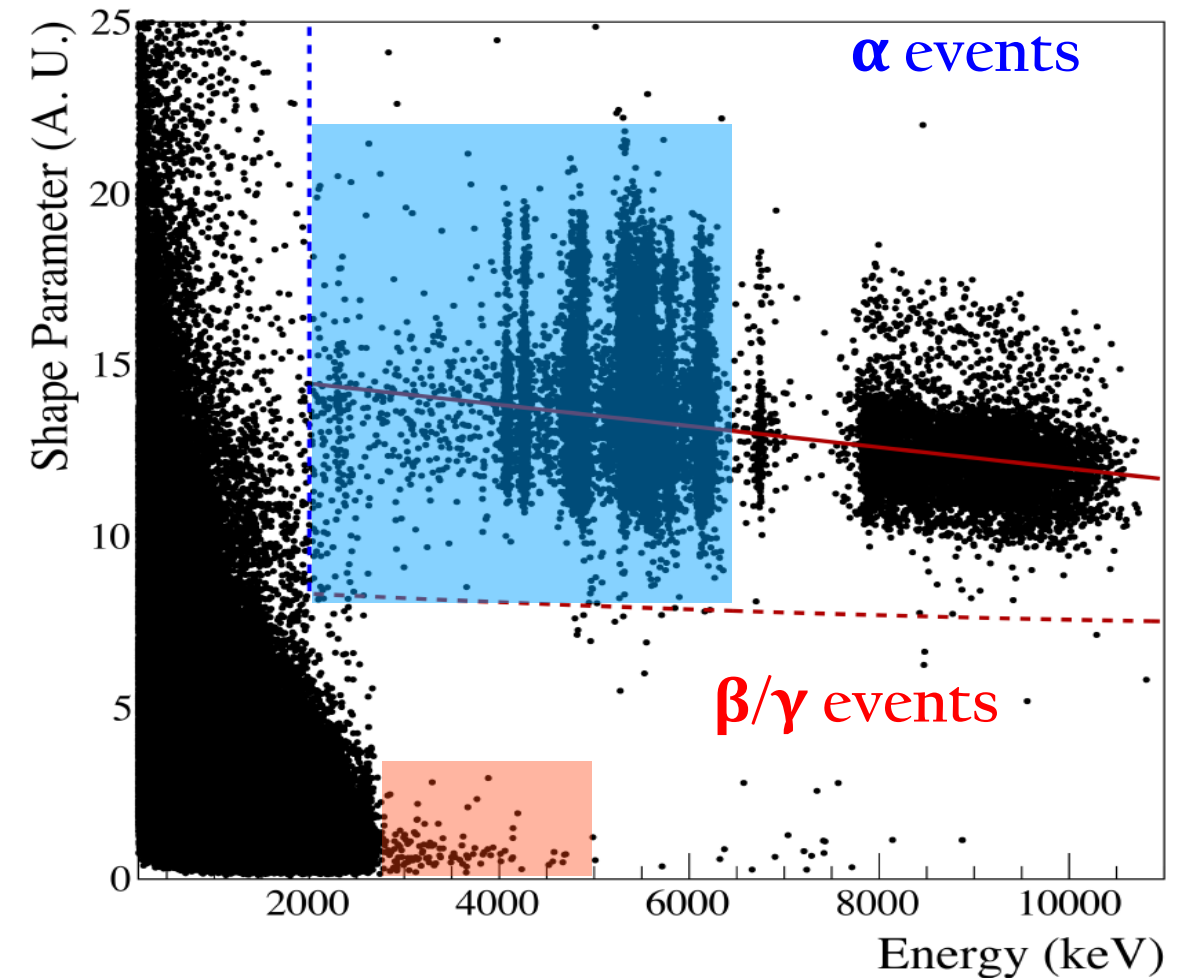
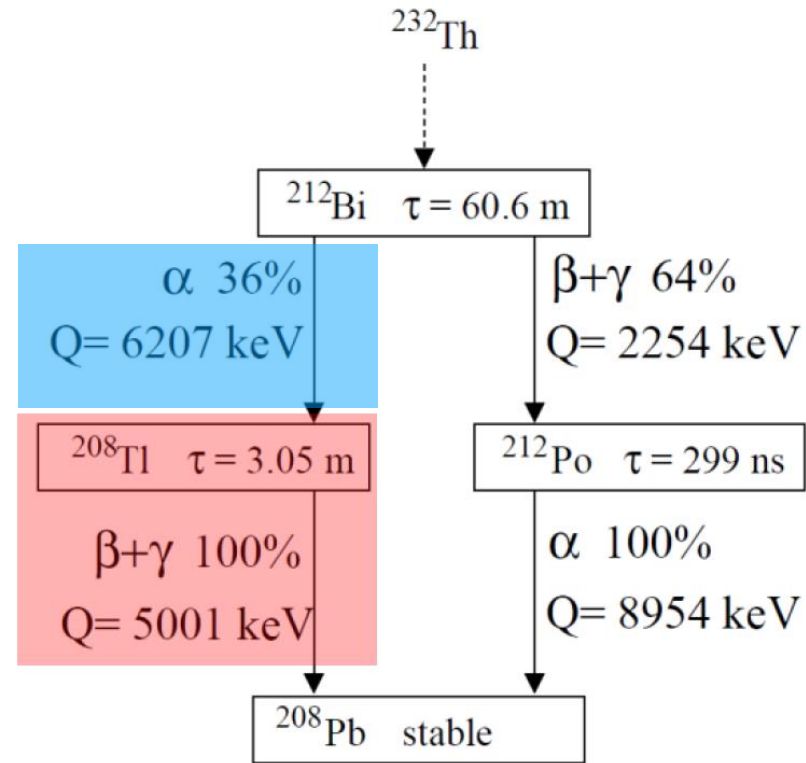
Light Signal depends on particle type

Selection based on light shape parameter



Background – Delayed coincidences rejection

Delayed ^{212}Bi - ^{208}Tl (α/β) coincidences

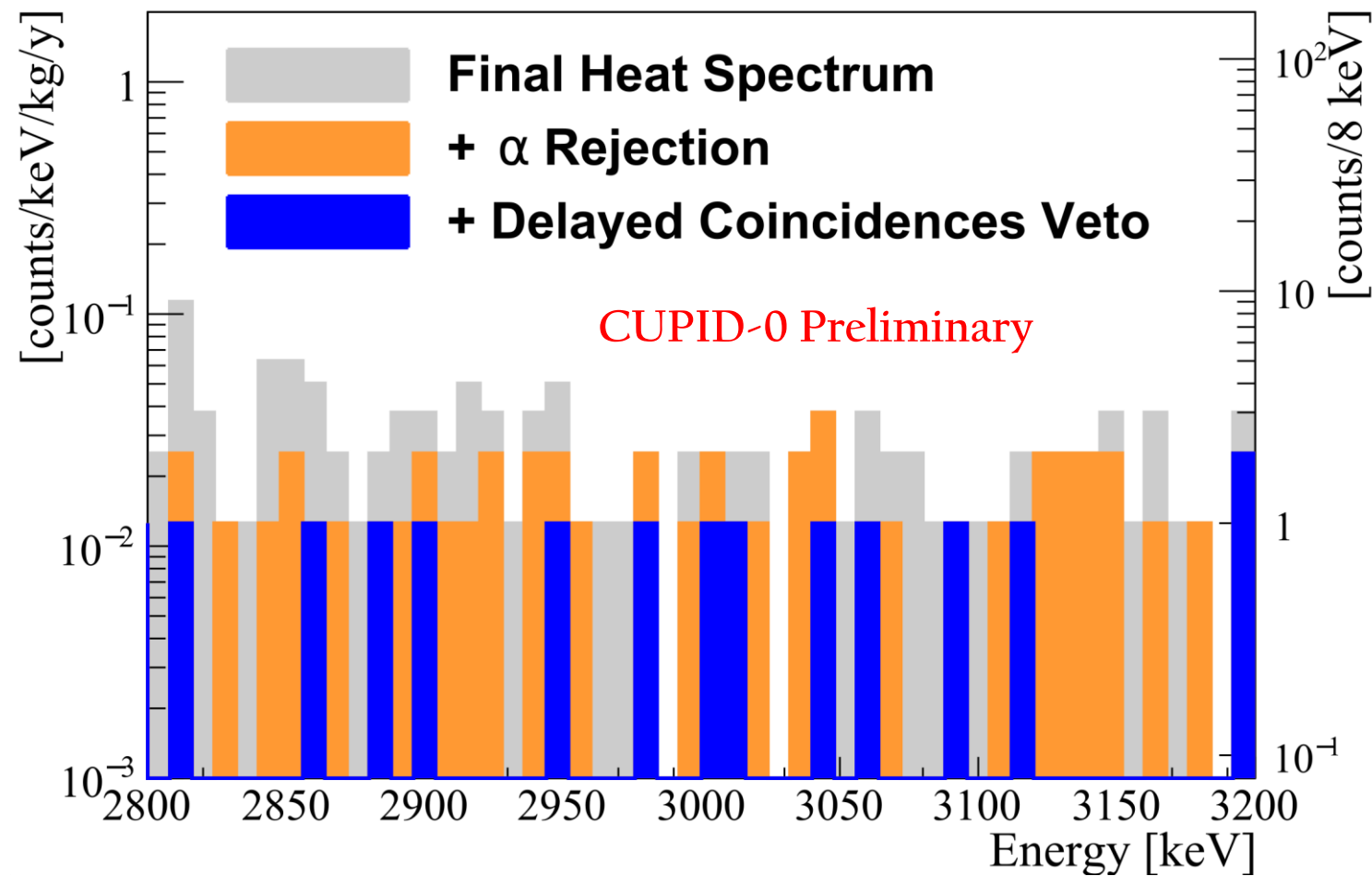


Selection of ^{212}Bi α events

- α pulse shape
- $2.0\text{ MeV} < \text{Energy} < 6.5\text{ MeV}$
 - Degraded tag

————→ Veto for 7 half-life

Background – Delayed coincidences rejection



$$(3.5^{+1}_{-0.9}) \cdot 10^{-3} \text{ cnts}/(\text{keV} \cdot \text{kg} \cdot \text{y})$$

Lowest background ever
measured with a calorimeter

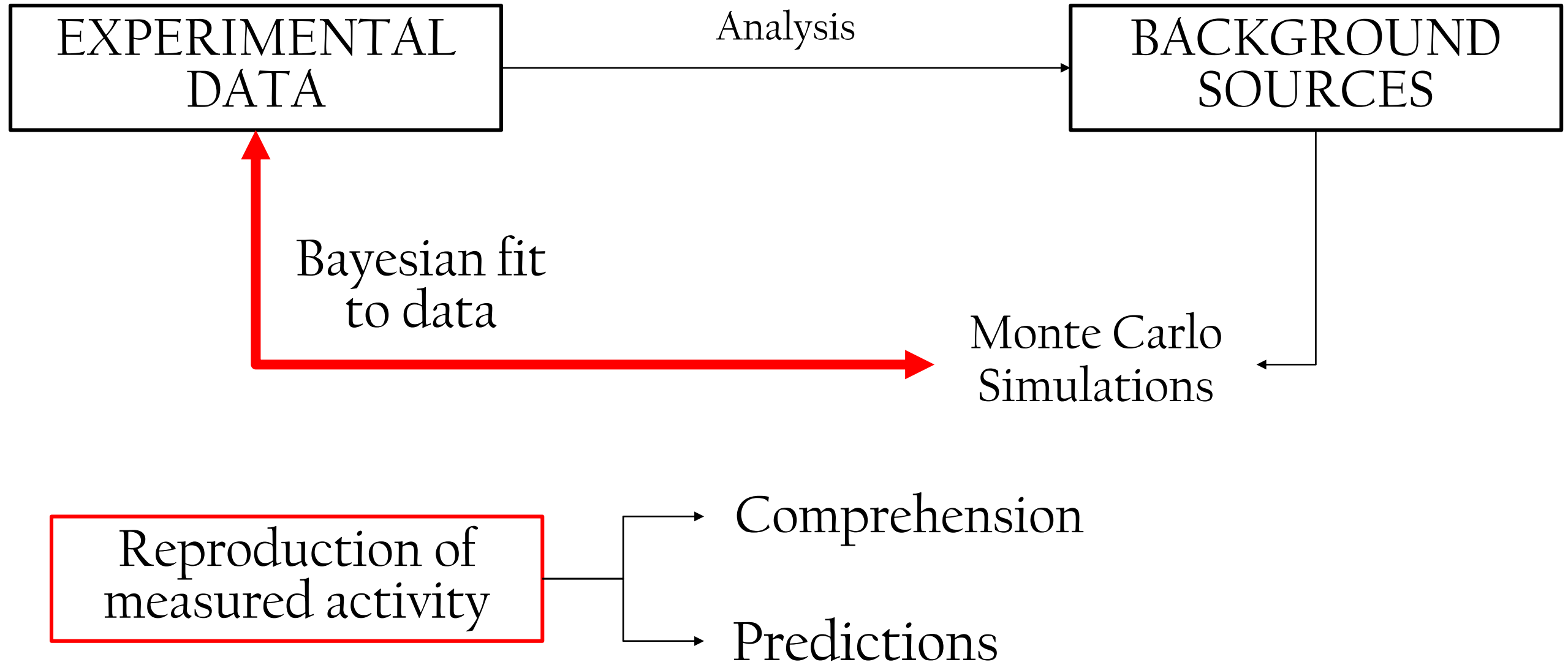
Total cut efficiency:
(86±1)%

No evidence of $0\nu\beta\beta$
signal

$$T_{1/2}^{0\nu} > 3.5 \cdot 10^{24} \text{ yr} \quad @90\% \text{ C.I.}$$

Background Model

A complete model of the background sources

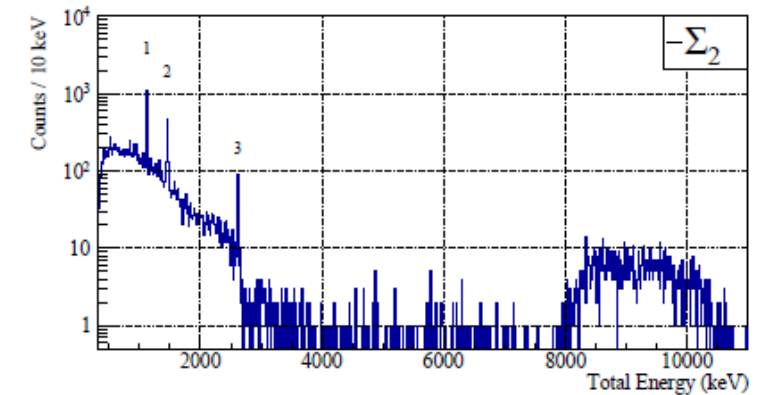
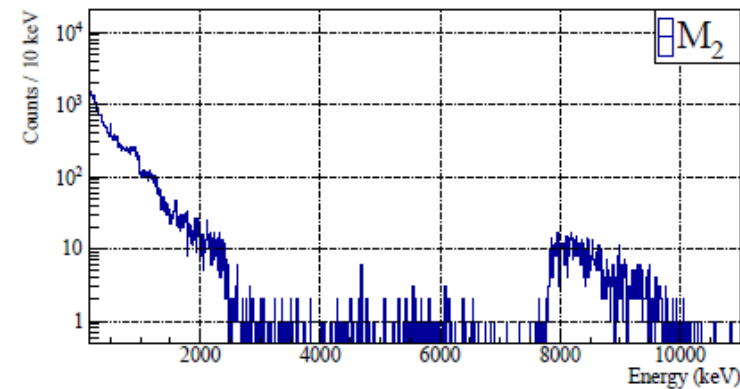
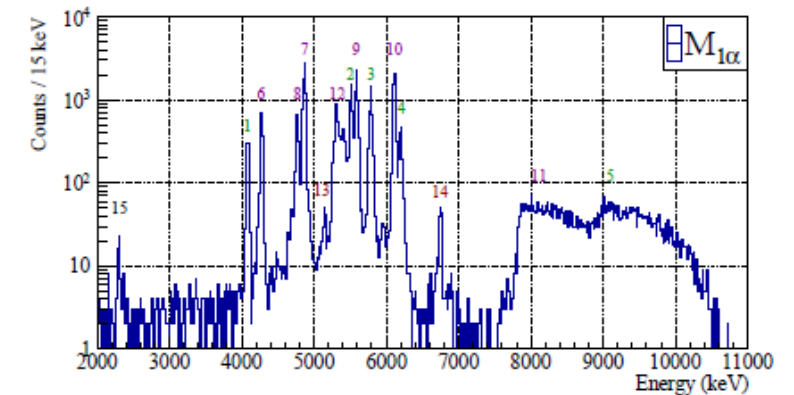
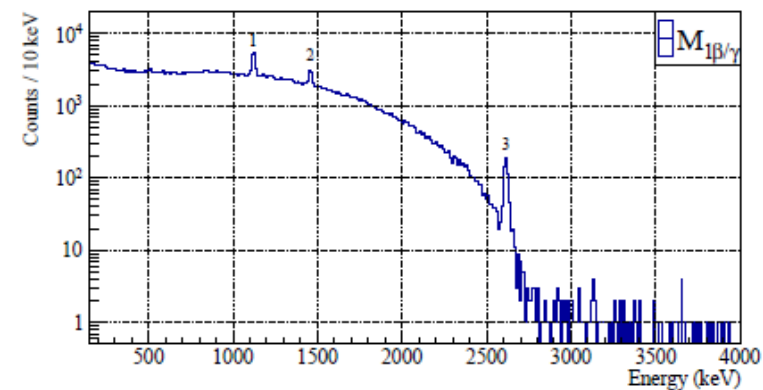
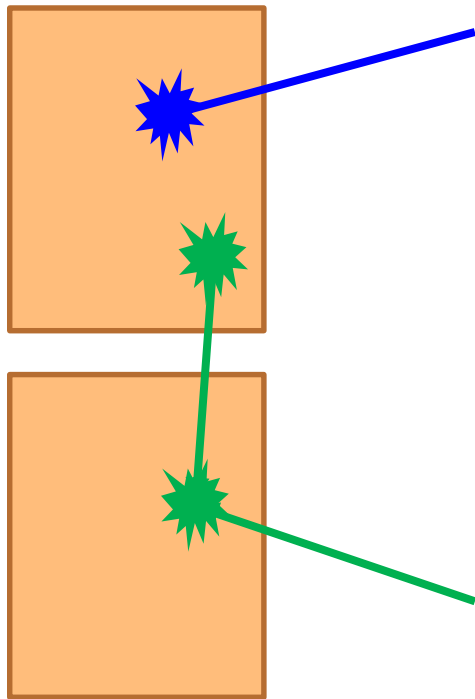


Experimental data for the model

Divided according to
multiplicity and particle type

- M1 $\alpha - \beta/\gamma$
- M2 / M2 sum ($\Sigma 2$)
- M>3 (to constraint Muons)

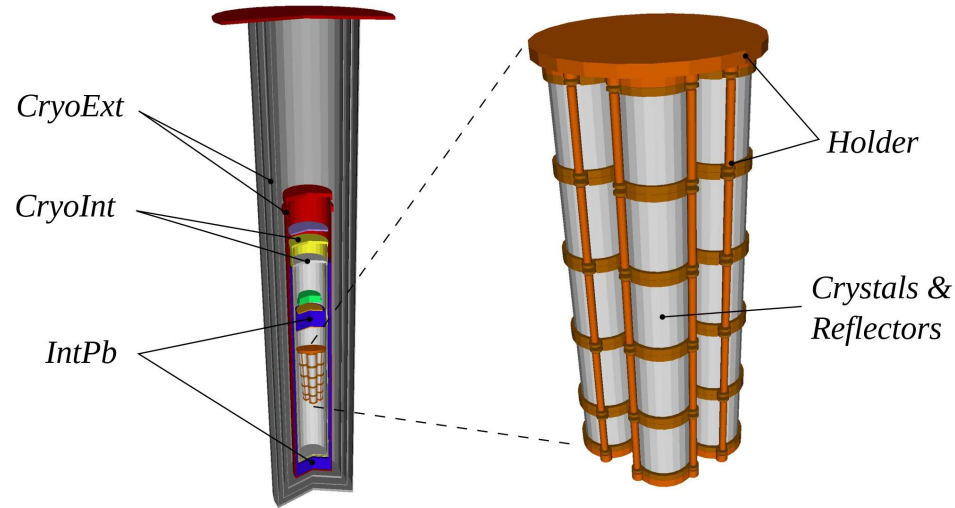
Multiplicity (M)
Crystals triggered in 20ms



Background source identification

Background sources

- Localization in the detector



- Depth of contamination



Surface

Exponential
profile

- Radiation type

Natural Chains

- Fathers + saecular equilibrium breaking points

Single isotopes

- ^{40}K , ^{54}Mn , ^{65}Zn , ^{60}Co , ...

Muons

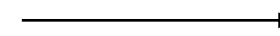
Monte Carlo simulations

Generation

Detection

MODEL

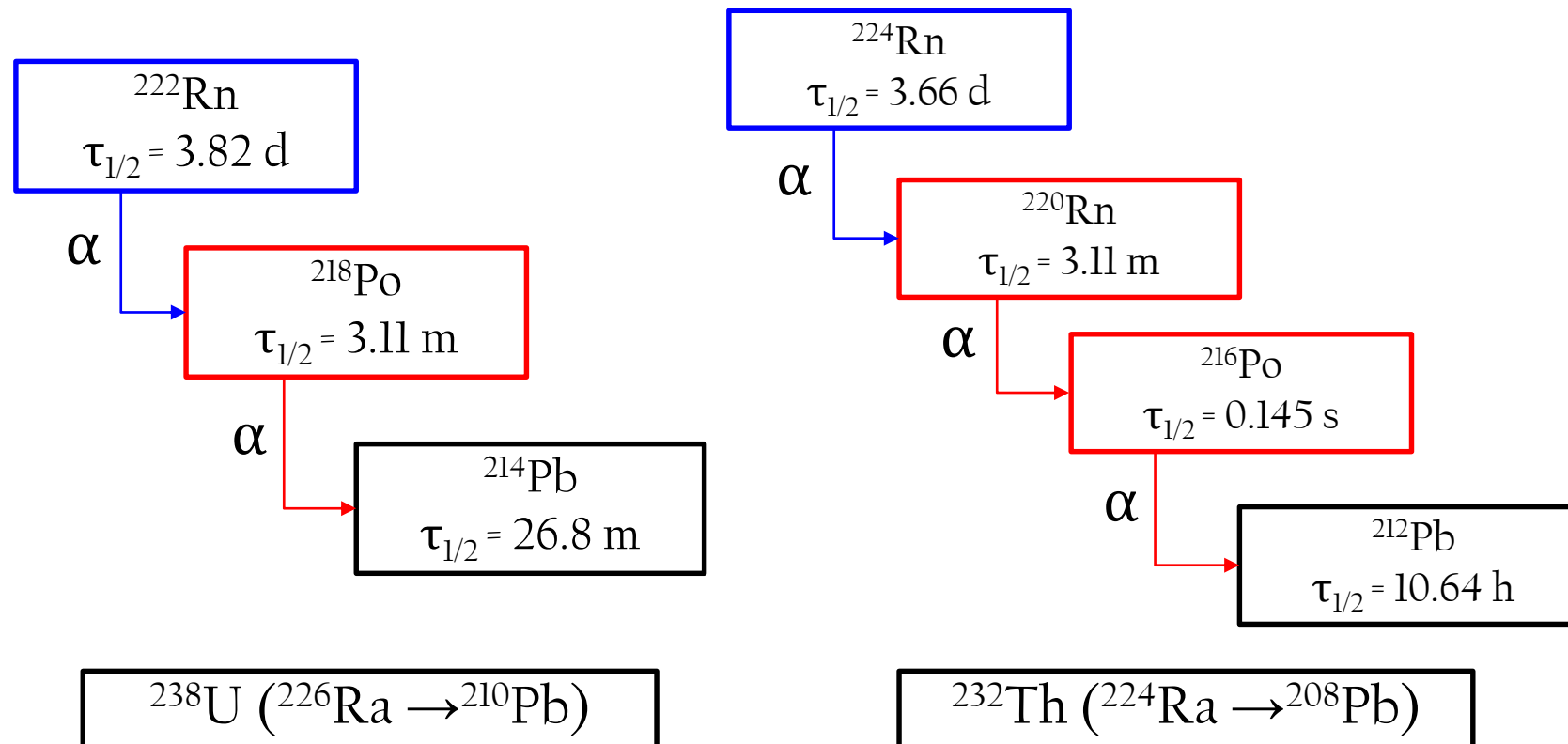
- 33 background sources



Linear combination
Coefficients = Activities

PRIORS

- Experimental signatures
 - α/α coincidences
- Previous contamination measurements
 - Reflective foil

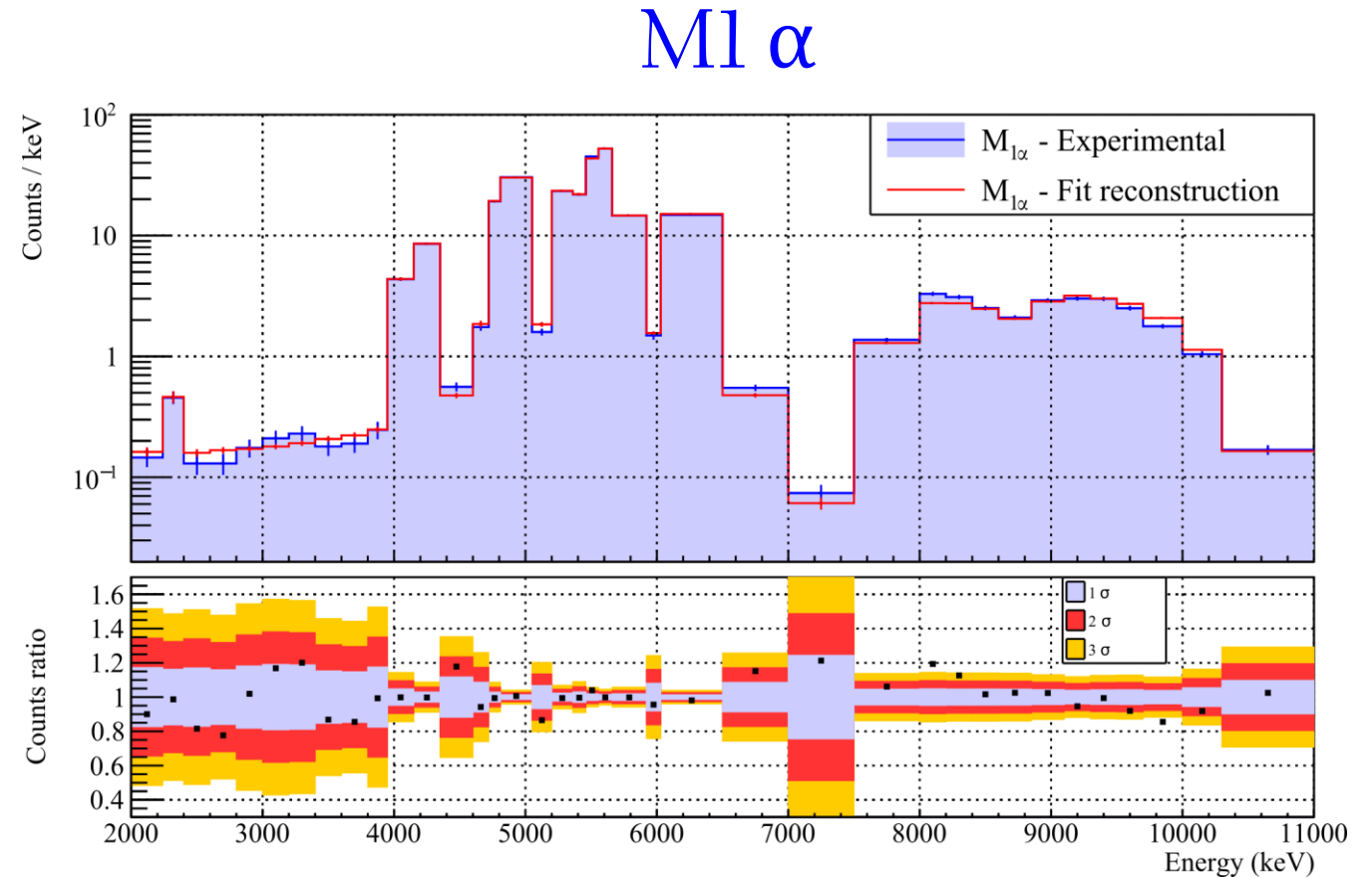
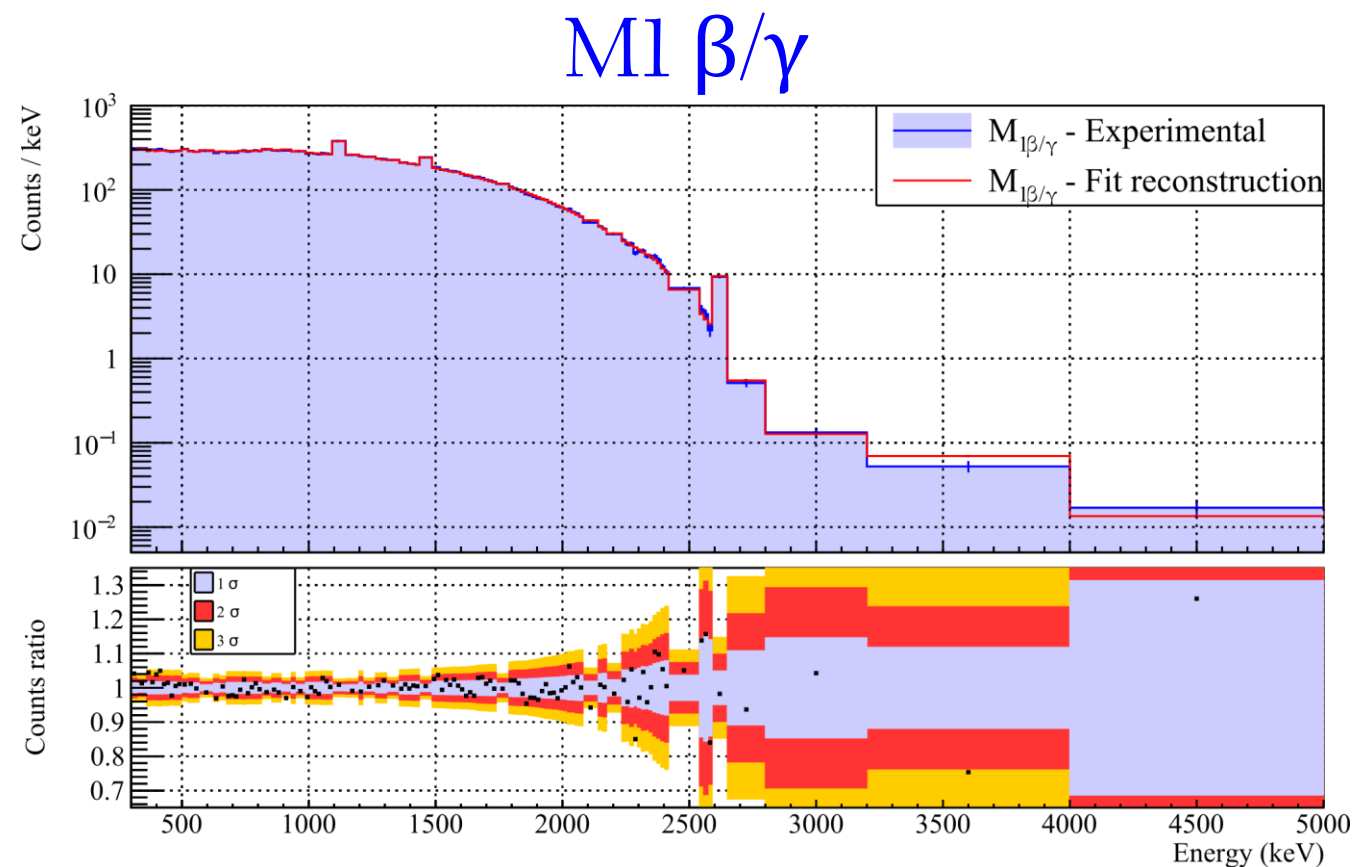


$P(\text{N}|\text{N}) Q_{\text{value}}$
Depends on source
localization

Daughter/parent gives a
prior on surface vs bulk
contaminations

Reconstruction results: M1 Spectra

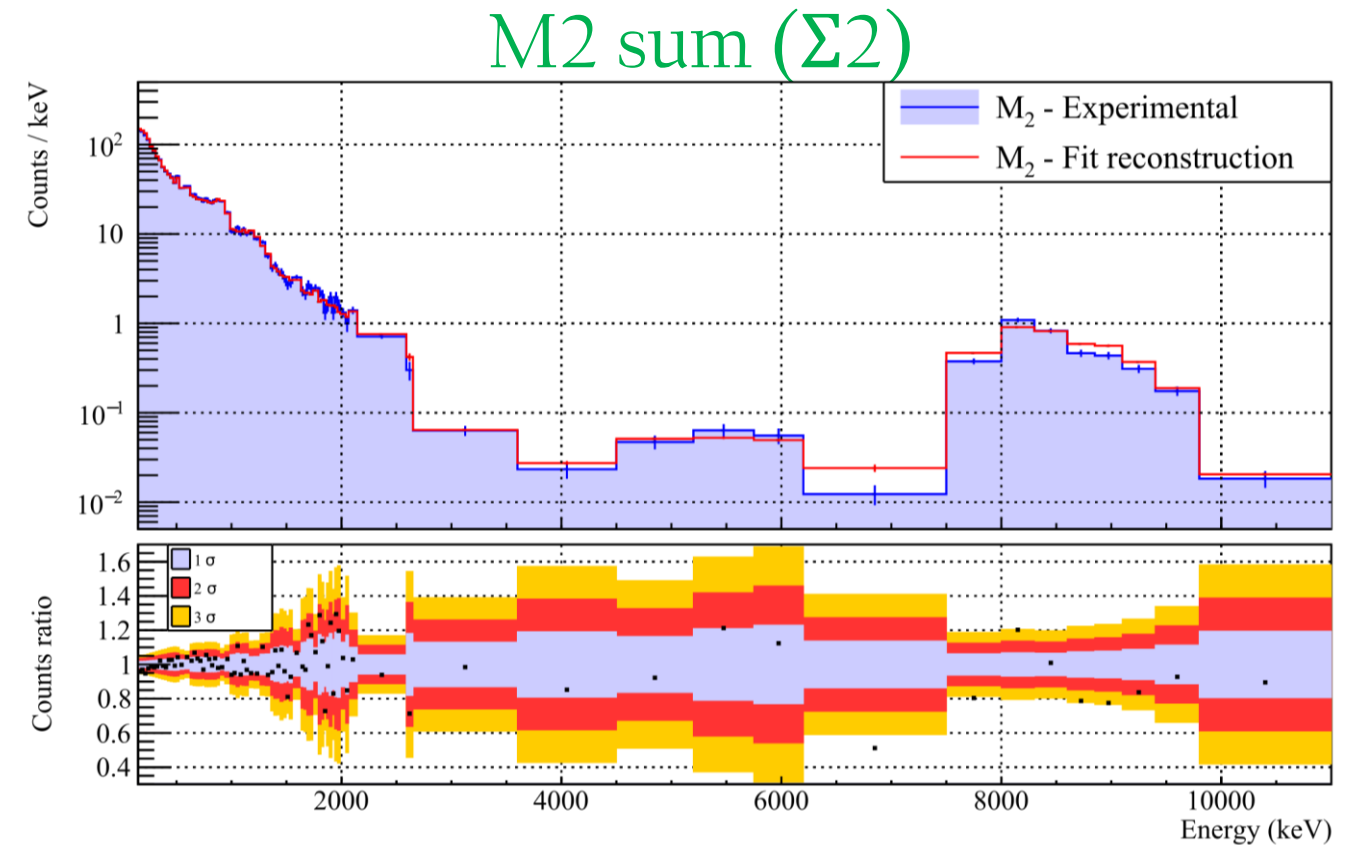
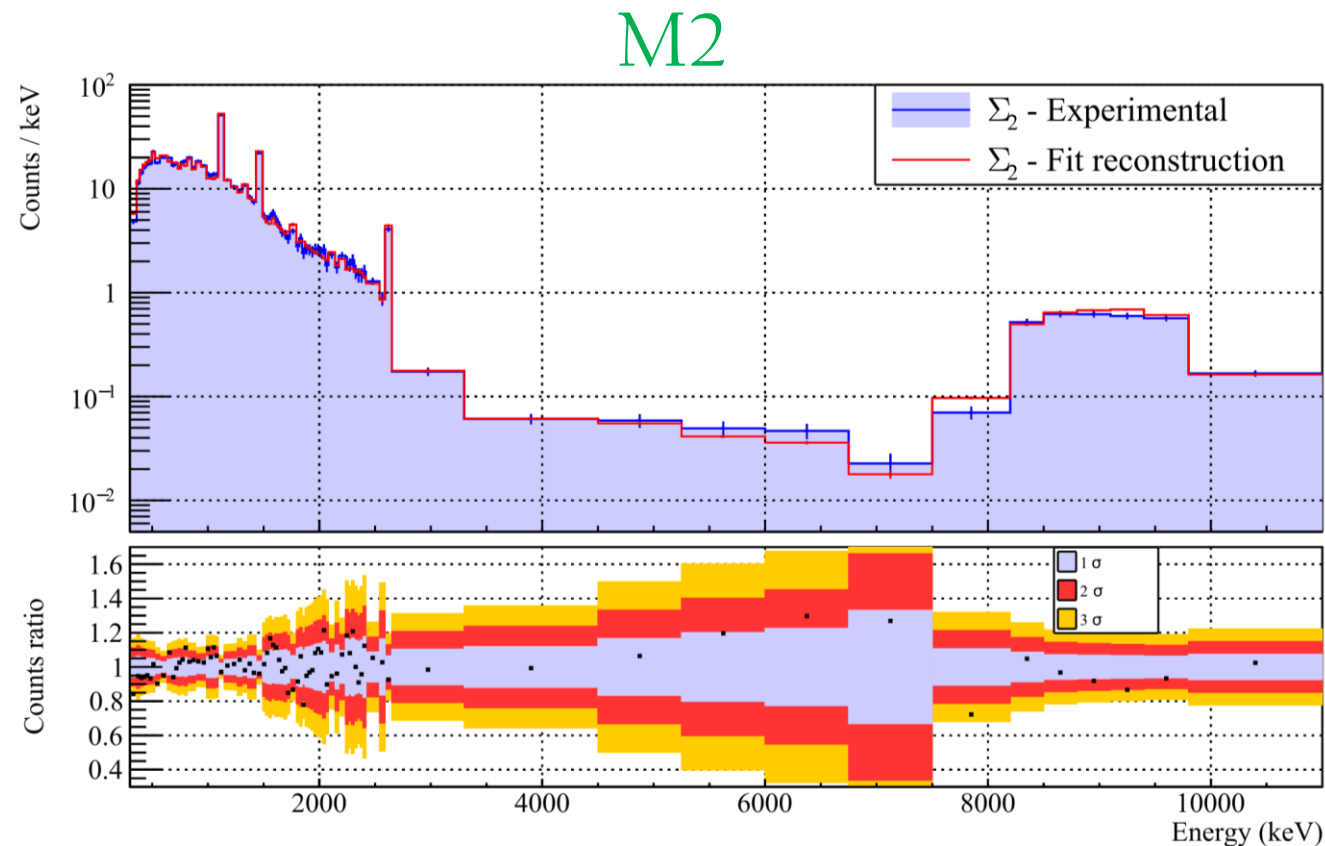
- Full spectrum reconstruction
- Peaks and continuum are well modelled



The $\alpha - \beta/\gamma$ separation allows to disentangle the different contributions

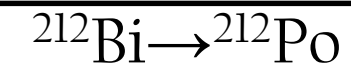
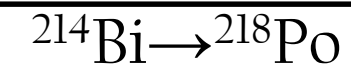
Reconstruction results: M2 spectra

- Both α and β/γ regions are well modelled in peaks and continuum
 - The surface/bulk prior is a key ingredient

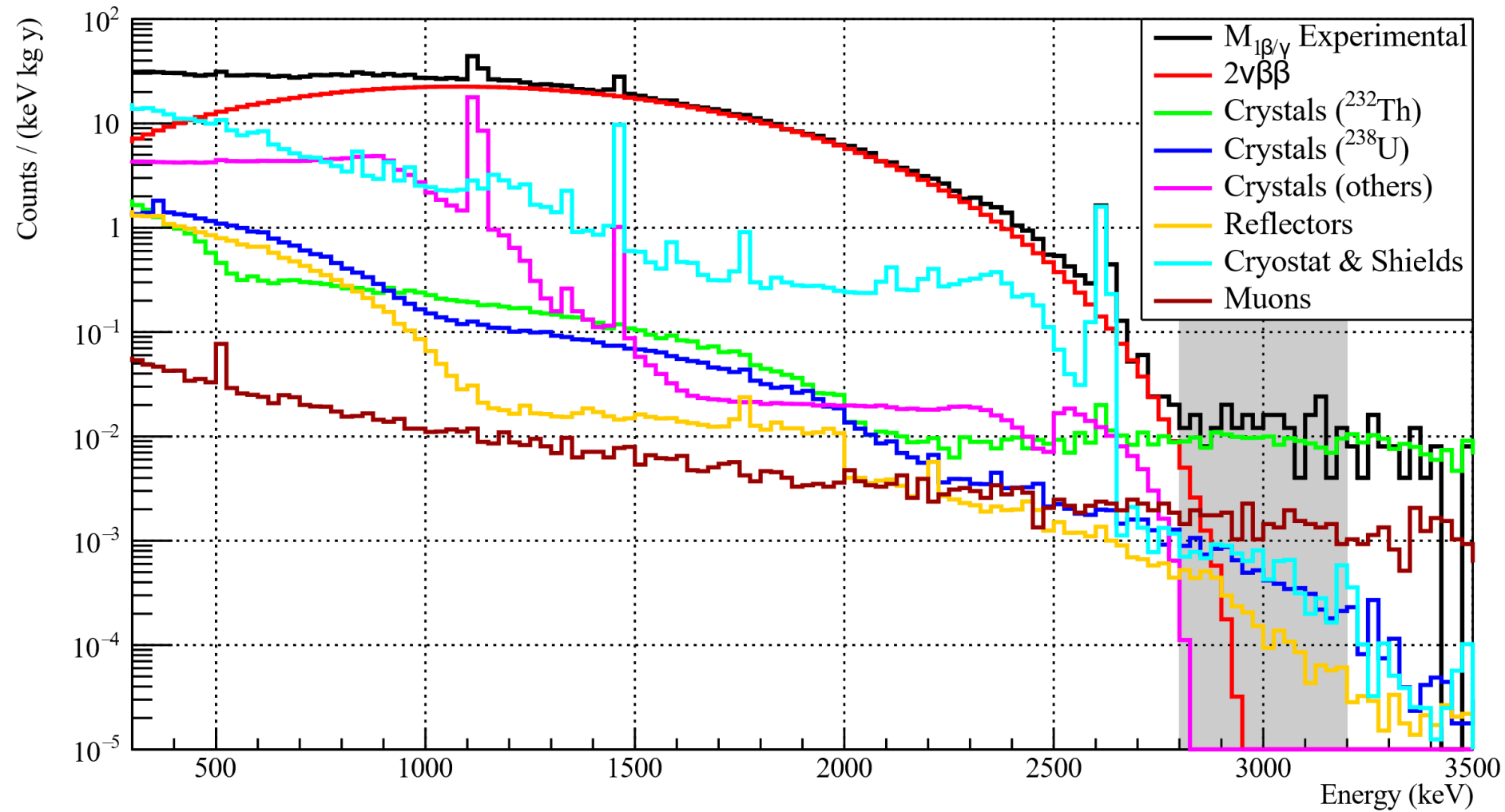


Some differences on the Bi-Po pileup

- Imperfect reconstruction of the deposited energy



Result: Beta/Gamma spectrum



In the ROI

^{232}Th is dominant,
because no delayed cut
is applied

Delayed
coincidences

Muons give 44% of
residual background

$2\nu\beta\beta$ is a dominant
contribution

Possibility to perform detailed
study on this decay

Phase II upgrade

- μ are the main residual background
 - Installation of μ -veto



No reflective foil

- Sensitivity to M2 α events

New clear Cu Shield

- Thermalization
- Additional shielding



NOW COOLING

CUPID 0: current results and future perspectives

- CUPID-0 is the first large array of enriched scintillating bolometers
- We reached the lowest background level achieved with bolometric experiments:

$$(3.5^{+1}_{-0.9}) \cdot 10^{-3} \text{ cnts}/(\text{keV} \cdot \text{kg} \cdot \text{y})$$

- A complete background model has been developed
 - Major ROI background (^{208}Tl β events) is reduced with delayed cut
 - Muons give 44% of residual counts
- Phase II upgrade focused on background improvement
 - Muon veto installed
 - No Reflective foil: M2 alpha events direct tagging
 - Additional shielding

BackUp Slides

Double beta decay ($\beta\beta$)

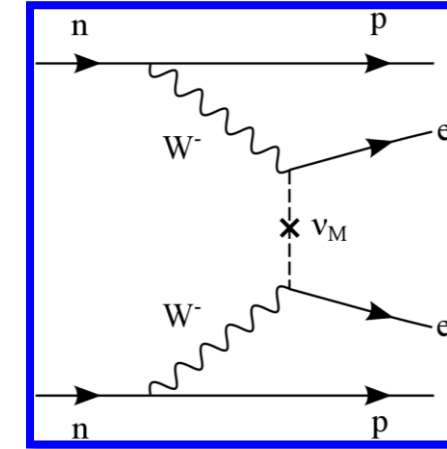
$0\nu\beta\beta$:

$$(A, Z) \rightarrow (A, Z+2) + 2e^-$$

$$m_\nu \neq 0$$

$$\nu \equiv \bar{\nu}$$

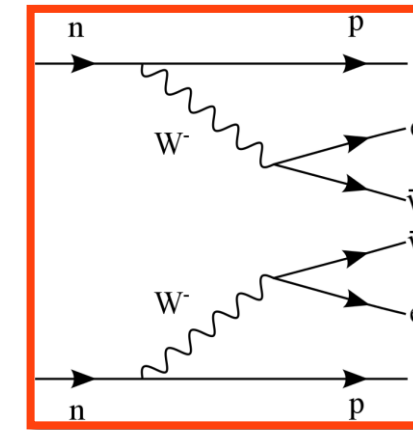
- Prohibited in the Standard Model ($\Delta L=2$)
- Limits: $T_{1/2}^{0\nu} > 10^{24} - 10^{25} \text{ y}$



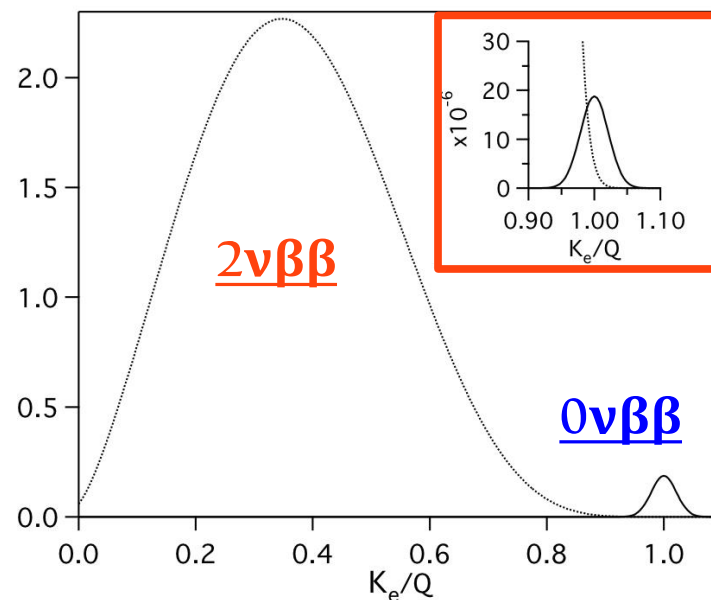
$2\nu\beta\beta$:

$$(A, Z) \rightarrow (A, Z+2) + 2e^- + 2\nu$$

- Predicted and detected



Measuring the two electron energy



Performing resolution

- At 2-3 MeV (Q_{value} of different isotopes)

Low Background

- Observing few counts above background

Experimental search for $0\nu\beta\beta$

Experimental sensitivity

Maximum measurable half-life at a given C.L.

$$S_{0\nu} \propto \sqrt{\frac{M \cdot T}{B \cdot \Delta}}$$

Critical experimental parameters:

- Isotope Mass (M)
- FWHM energy resolution (Δ)
- Background (B)

Mass scalability at low cost and high isotopic abundance

High purity materials
($< \text{ppb}$ radioactive contaminations)
Rejection techniques

Δ of few % at Q_{value} to avoid the $2\nu\beta\beta$ induced background

Ratio $0\nu\beta\beta$ signal/ $2\nu\beta\beta$ background

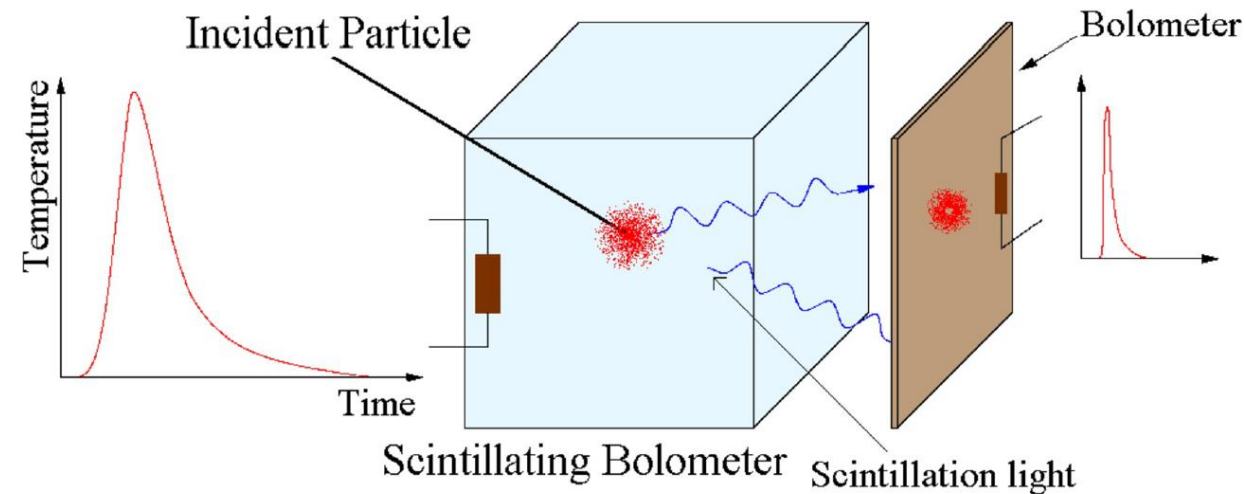
$$\frac{S^{0\nu}}{B^{2\nu}} = \frac{m_e}{7} \frac{T_{1/2}^{2\nu}}{T_{1/2}^{0\nu}} \frac{Q_{\text{value}}^5}{\Delta^6}$$

The first enriched scintillating bolometer $\beta\beta$ experiment

Demonstrating achievable
Background rejection

Precision measurements on
 ^{82}Se $\beta\beta$

$$^{82}\text{Se} - Q_{\beta\beta} = (2997 \pm 0.3) \text{ keV}$$



Heat:

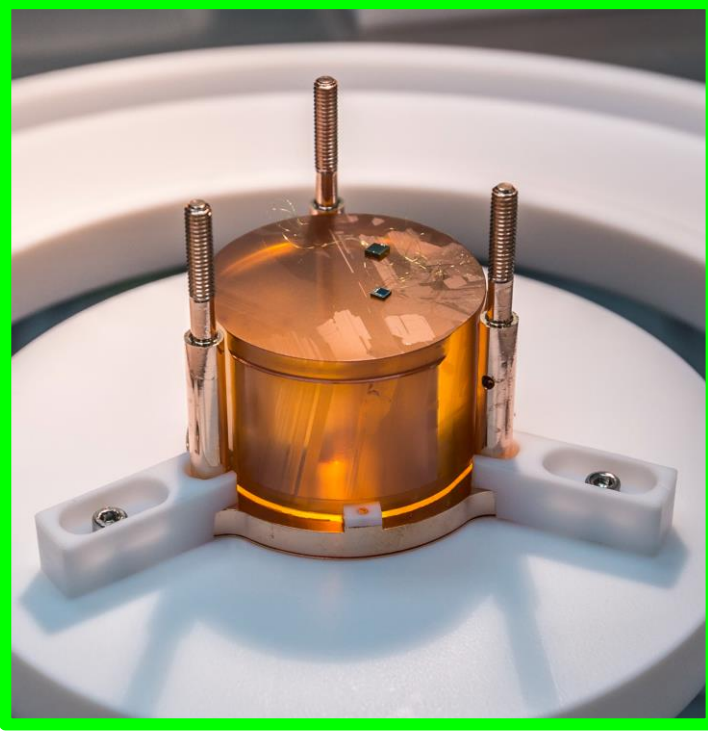
bolometric high resolved output

Light:

particle identification

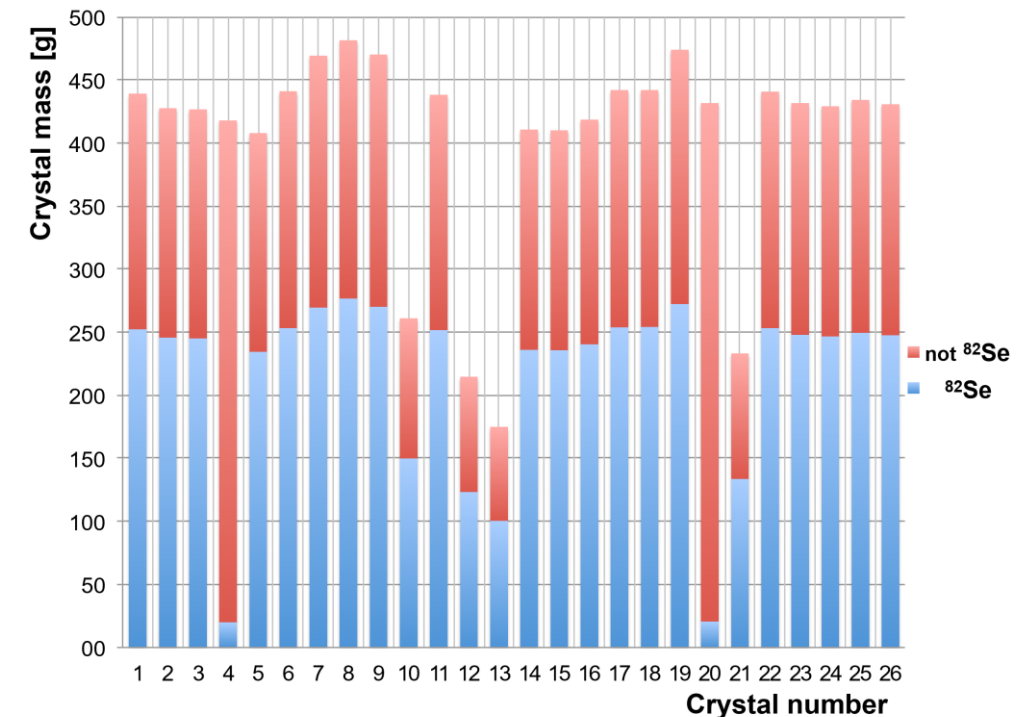


Zn^{82}Se crystals



First large mass Zn^{82}Se enriched crystals ever grown.

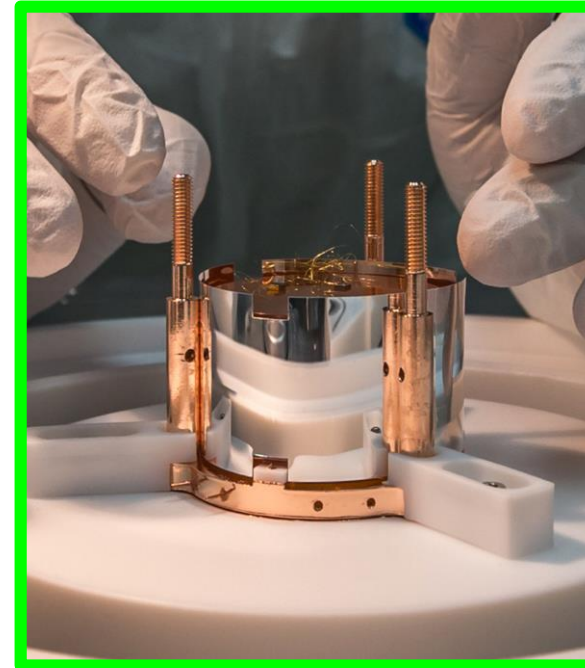
- 26 ZnSe crystals
 - 24 enriched in ^{82}Se (95%) + 2 naturals
- Total mass = 10.5 kg
 - ^{82}Se mass = 5.17 kg
 - $3.8 \cdot 10^{25}$ $\beta\beta$ emitters



Material surrounding the crystal

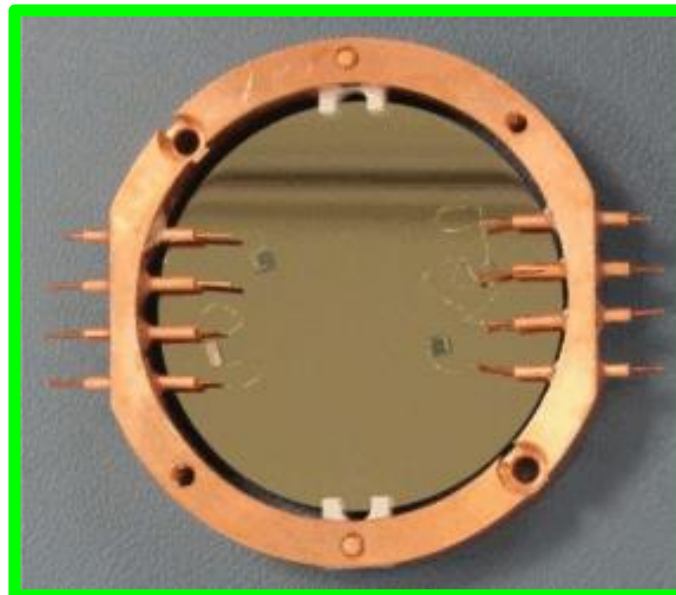
Copper structure

- NOSV ultra-pure copper
- Cleaning procedure vs surface contaminations



Vikuiti Reflector

- Enhances the light output
- Low Th/U contaminations measured

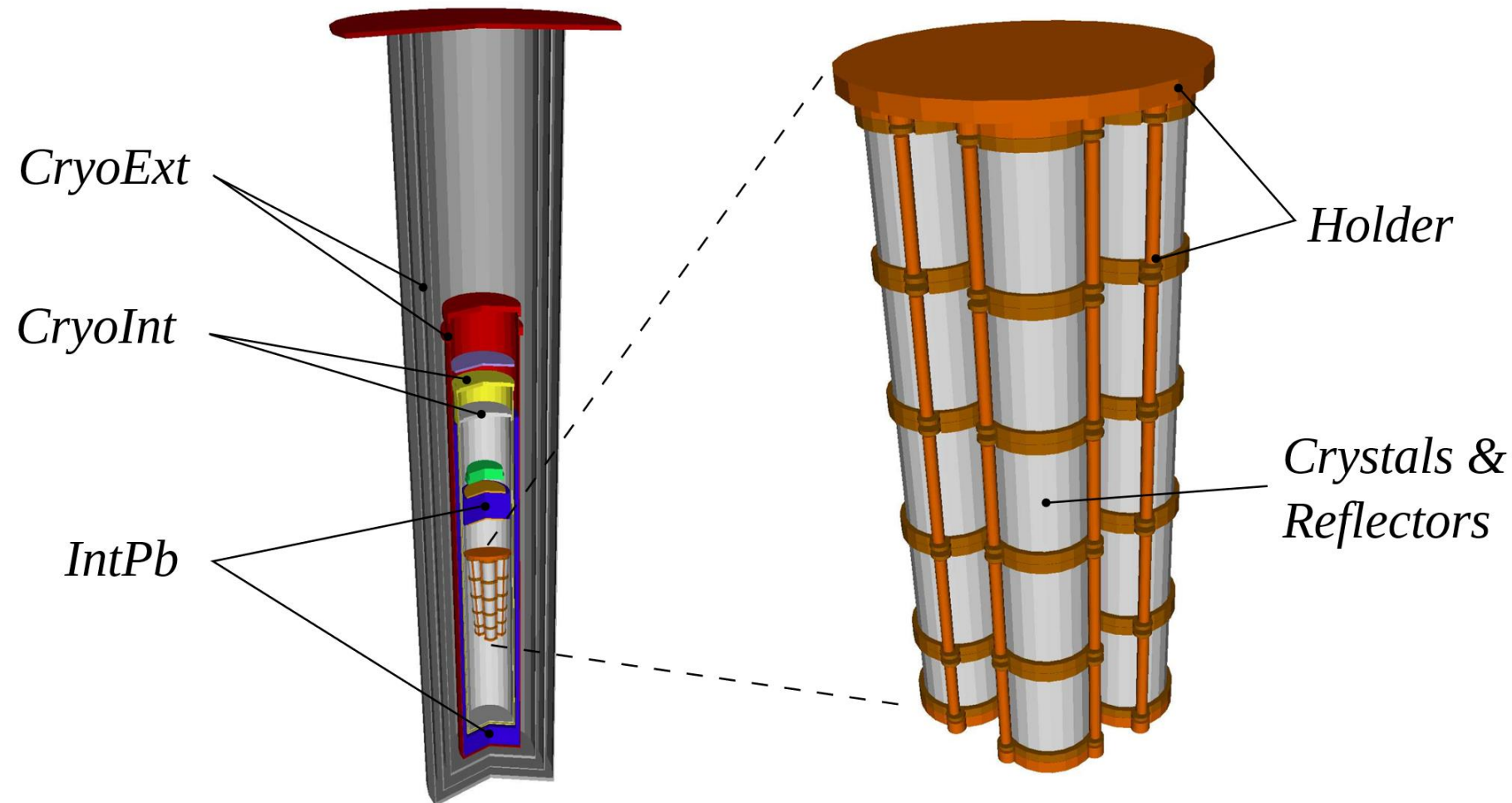


Ge Bolometric Light Detector

- SiO_2 anti-reflective coating
- Sensible to few keV energy deposition

The Cryostat

- Oxford 1000 $^3\text{He}/^4\text{He}$ dilution cryostat (CUORE-0)
- Radially divided by the Roman Lead shield:
 - CryoExt, RomanPb, CryoInt

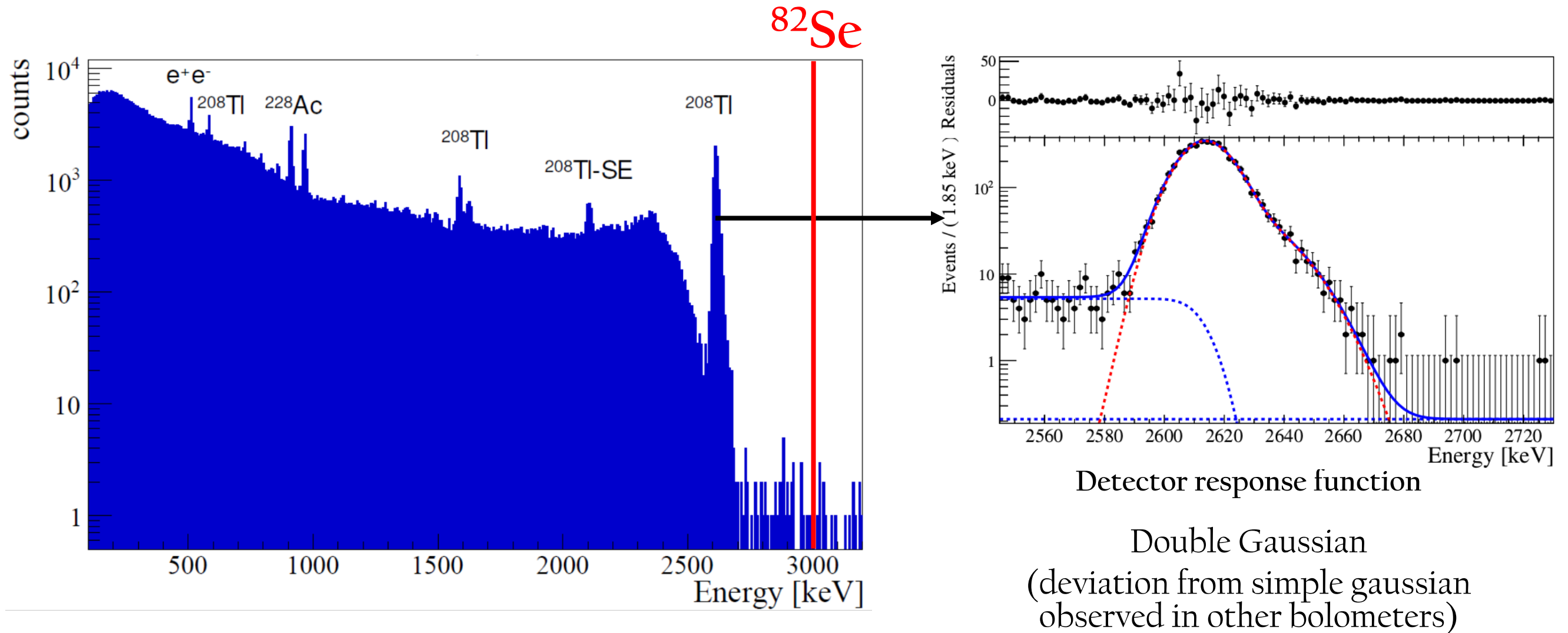


The cryostat contamination have been evaluated from experimental data

Detector Performances

^{232}Th Energy Calibration

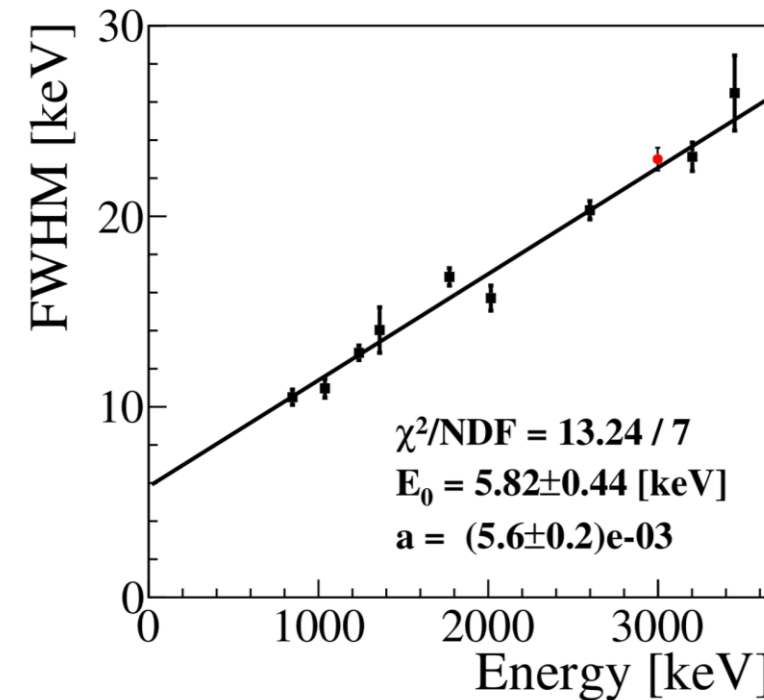
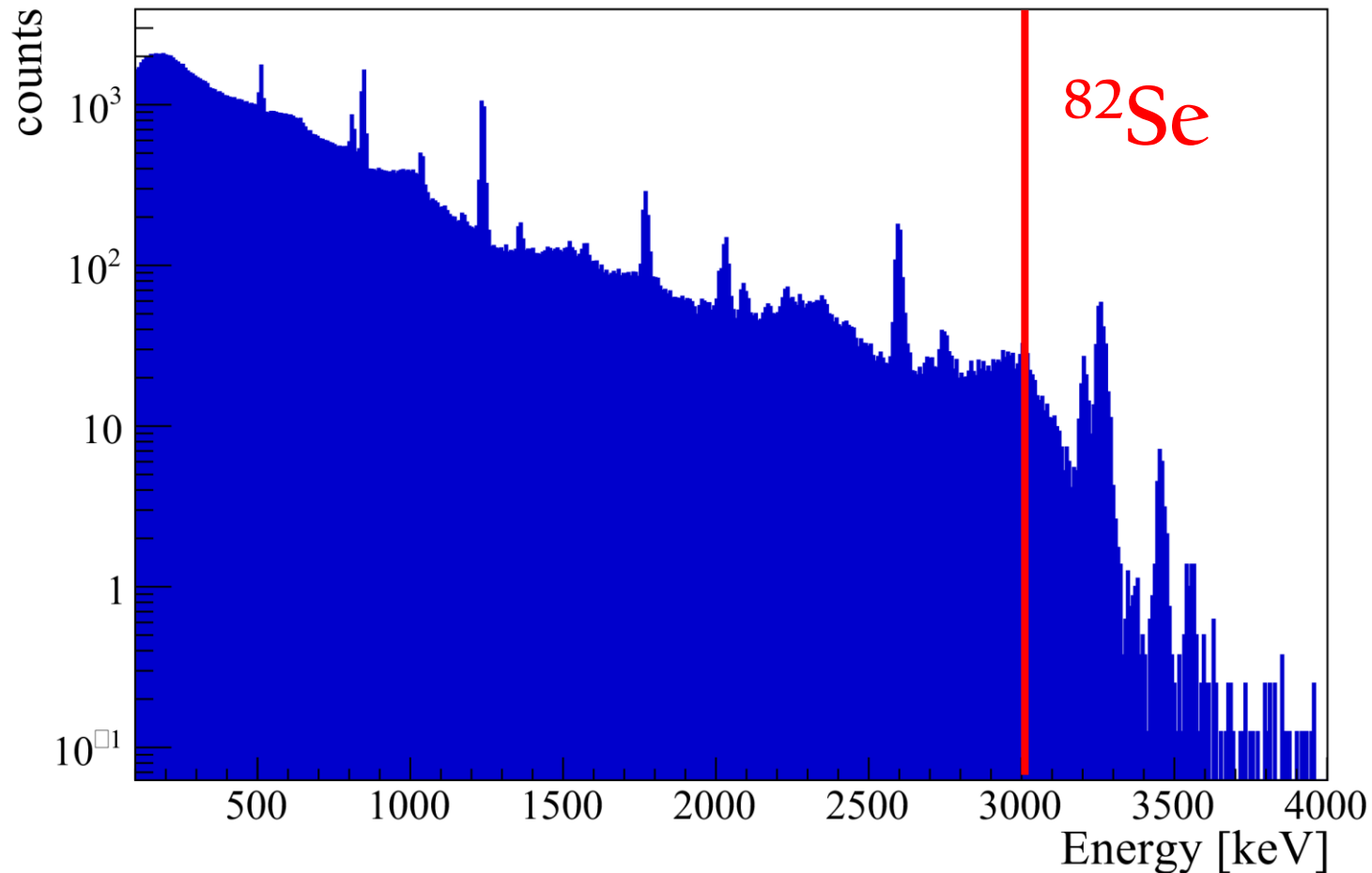
- Bolometer calibration and light detector intercalibration
- Response function evaluation



Detector Performances

^{56}Co Energy Calibration

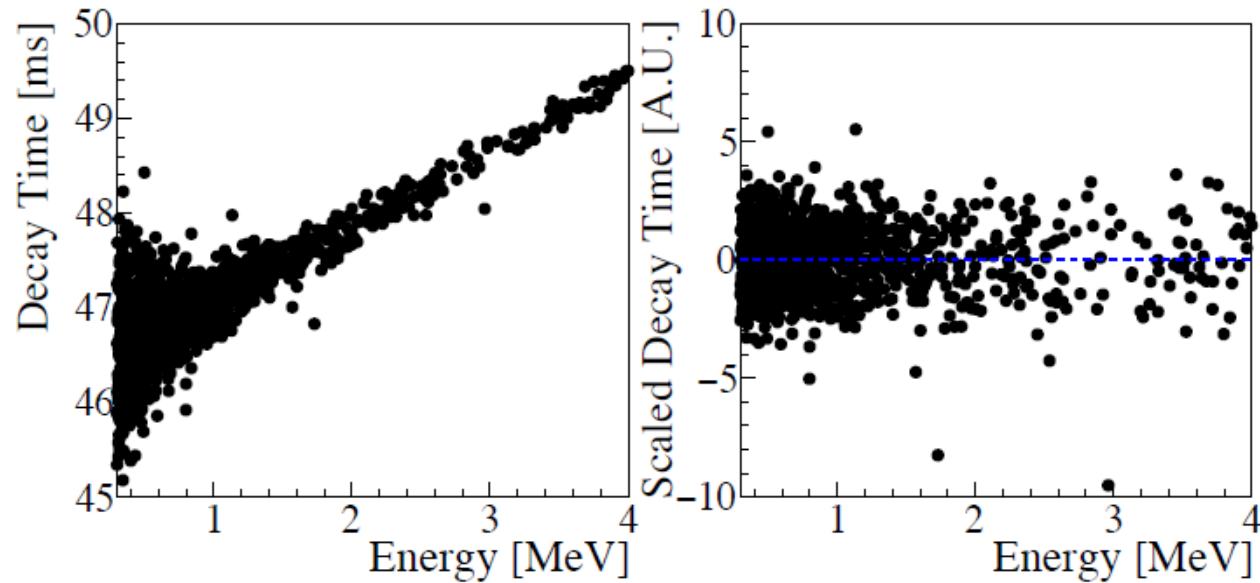
- Check of the energy reconstruction
- Evaluation of FWHM energy resolution @ ^{82}Se Q



FWHM @ $Q_{\beta\beta}$
(20.0 ± 0.6) keV

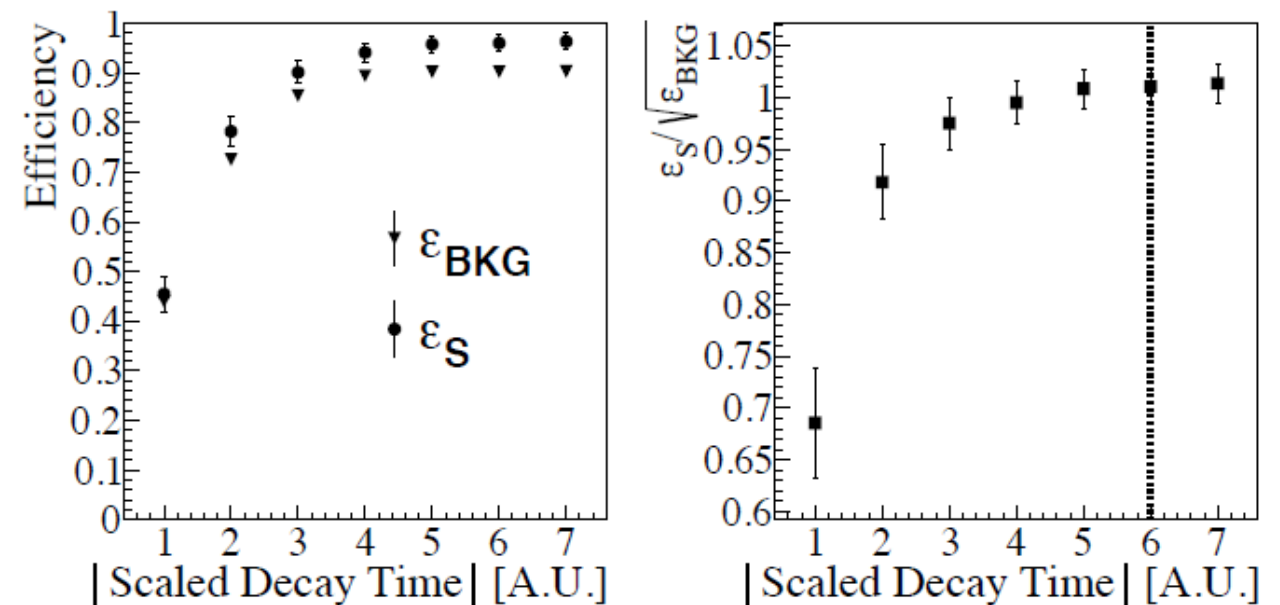
Major contribution is the crystal quality
(average baseline FWHM ~ 5 keV)

Shape parameters cut



Remove energy-dependency of the shape parameters for energy-independent cuts

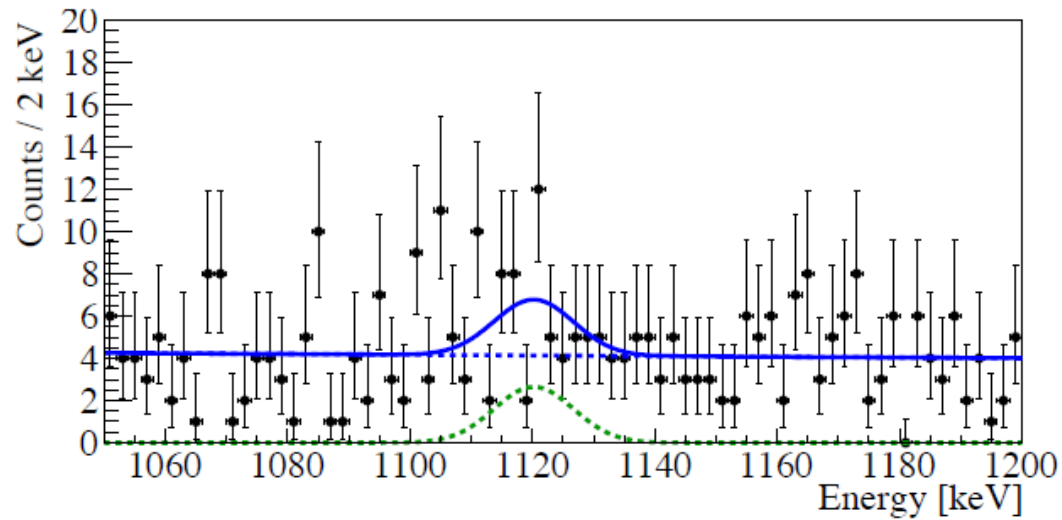
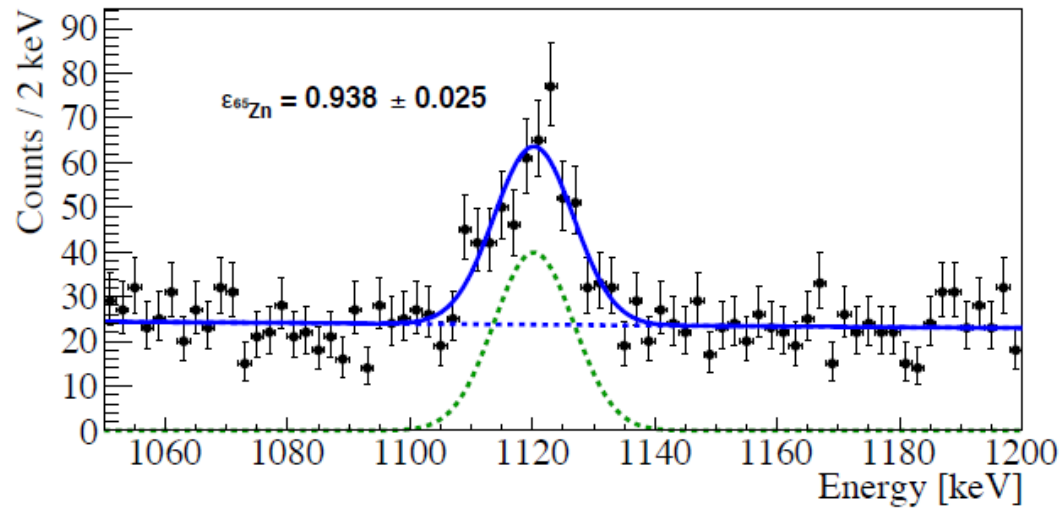
Efficiency and signal/noise as figure of merit to choose cut level



Evaluation of efficiency

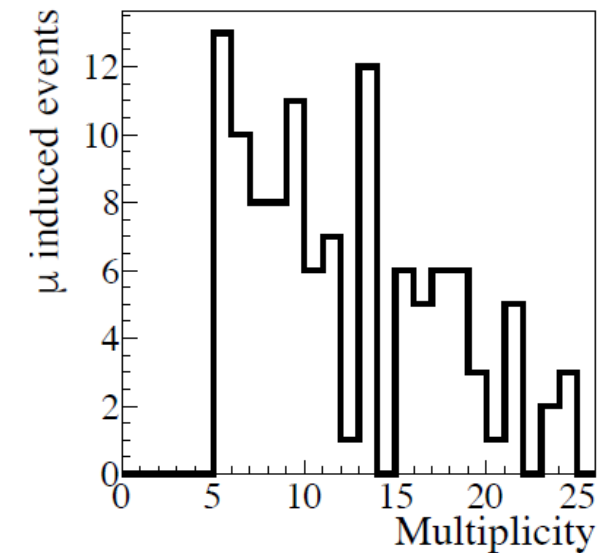
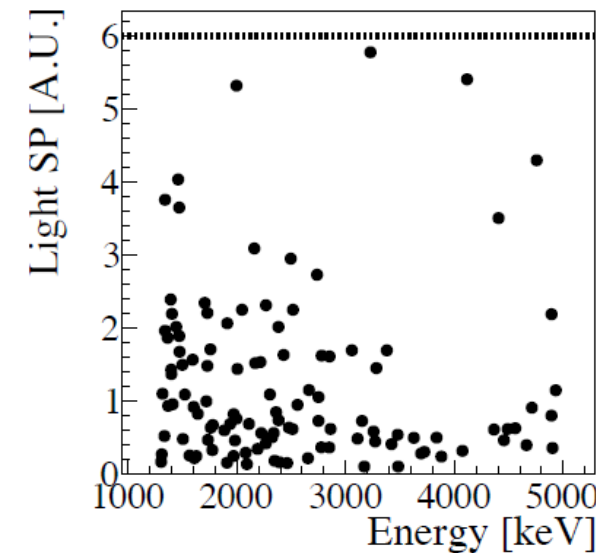
Heat

- Fit of ^{65}Zn Line before/after cuts applied

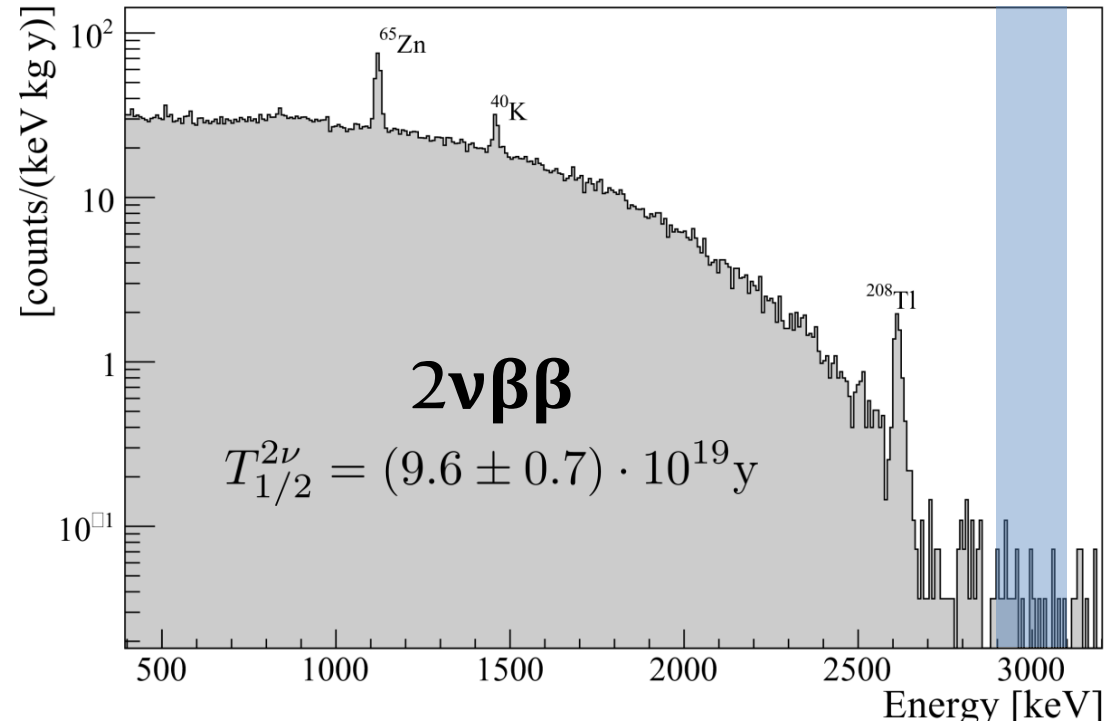


Light

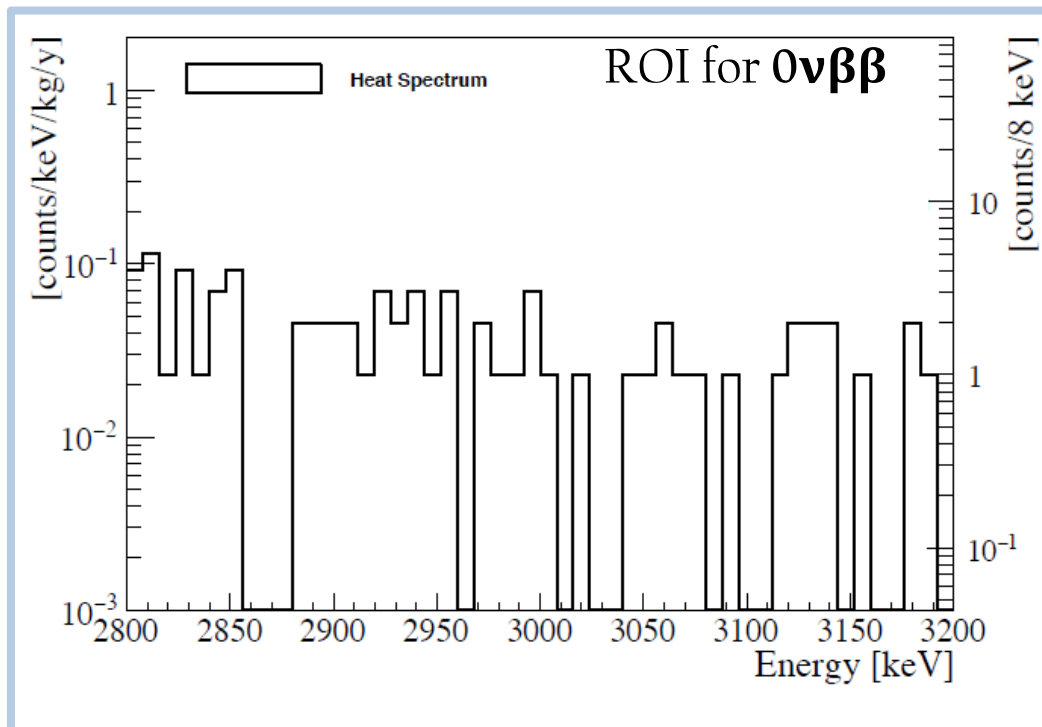
- Cut on events with $M > 6$
 - Muonic showers, almost pure beta/gamma sample



$0\nu\beta\beta$ search



- ^{65}Zn : cosmogenically activated
- ^{40}K and ^{208}Tl : natural radioactivity
- $2\nu\beta\beta$ is the dominant background



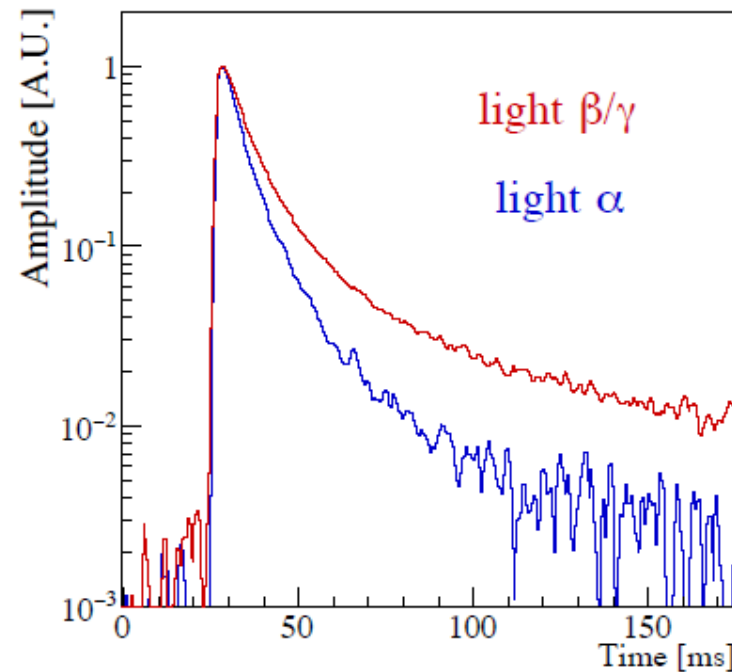
Basic Selections

- Rejection of “non-particle-like” events through pulse shape on thermal pulses
- Anti-coincidence between crystals ($\Delta T=20\text{ms}$)
 - Multiplicity selection

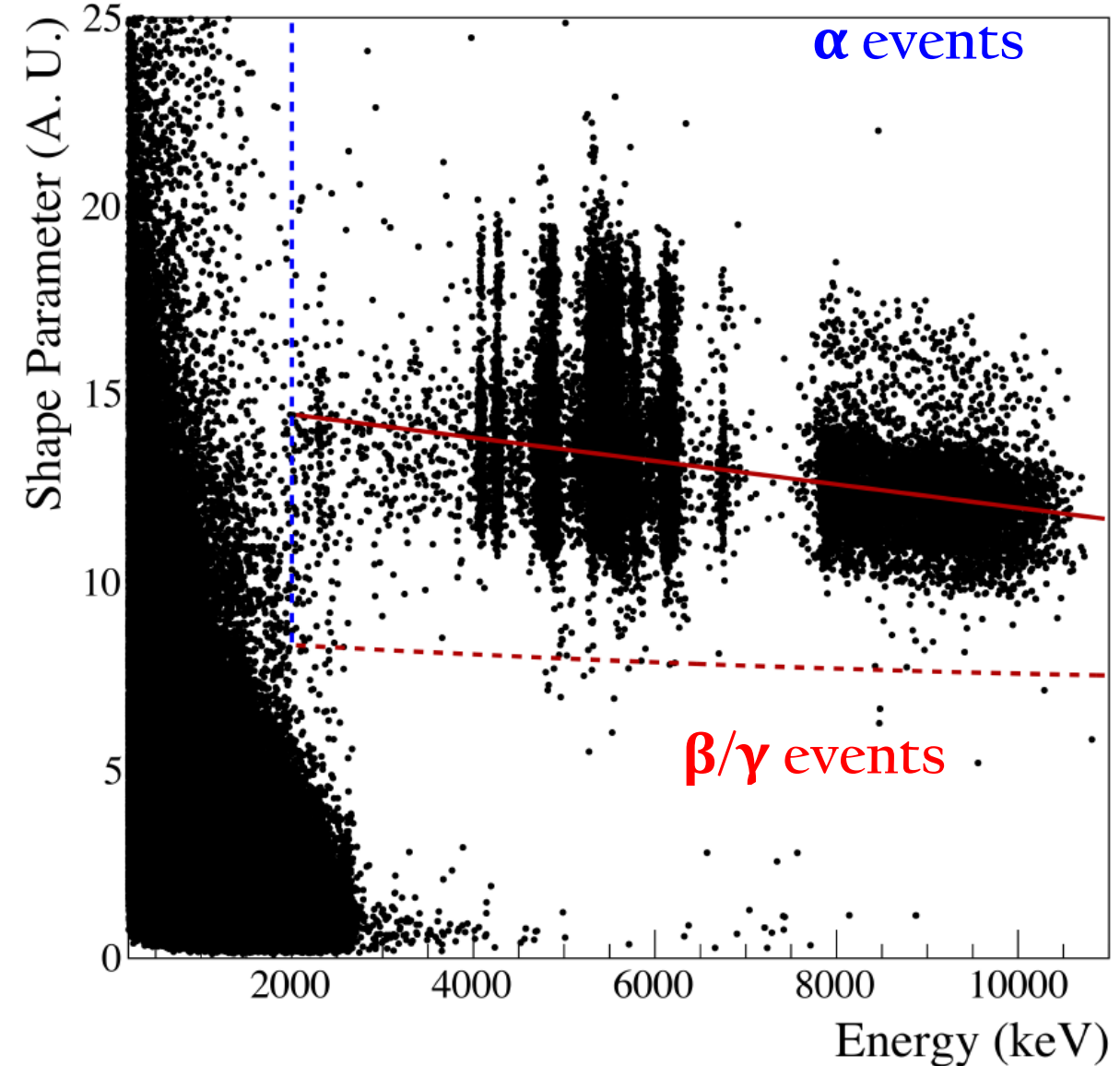
$$\text{BKG} = (3.2 \pm 0.4) \cdot 10^{-2} \text{ cnts}/(\text{keV} \cdot \text{kg} \cdot \text{y})$$

Background – Alpha Rejection

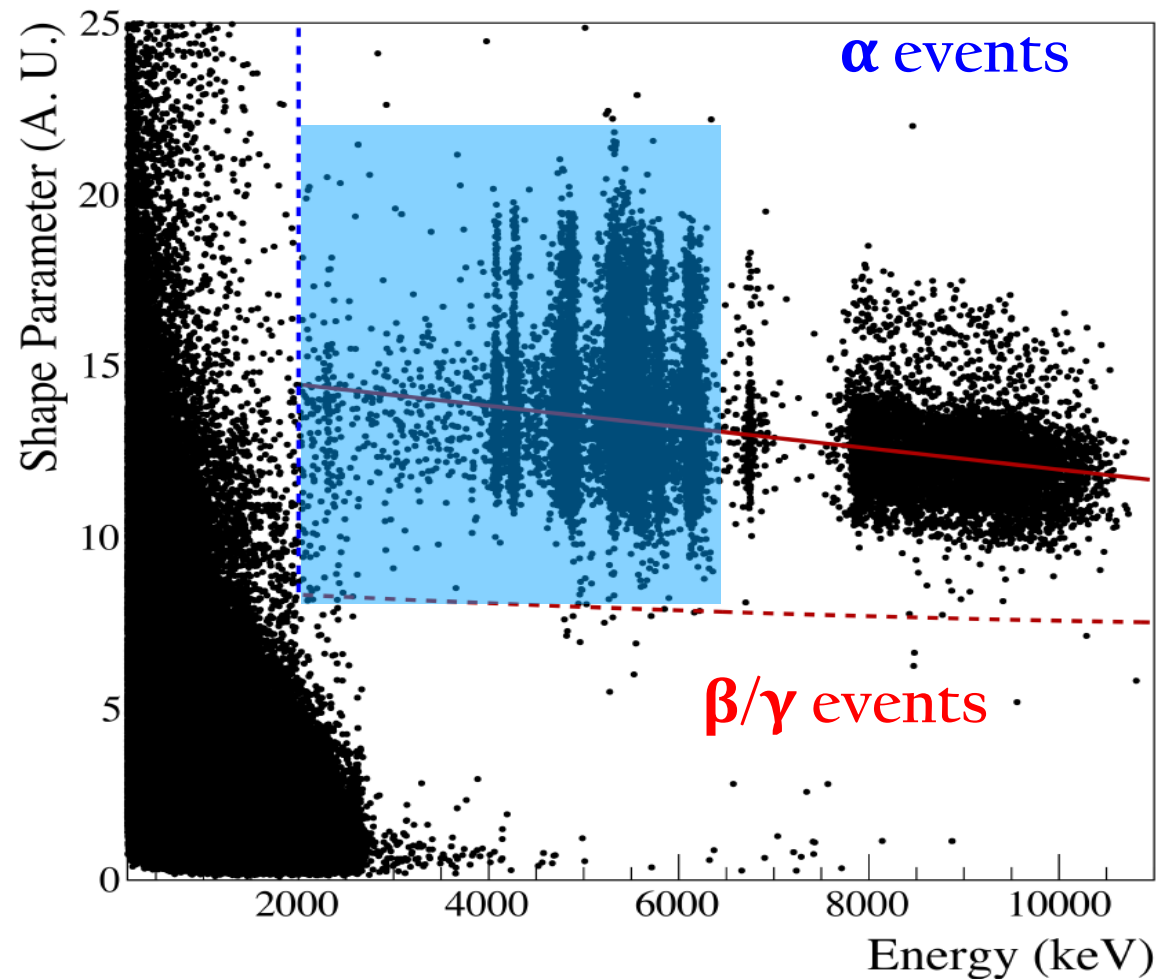
Light Signal depends on particle type



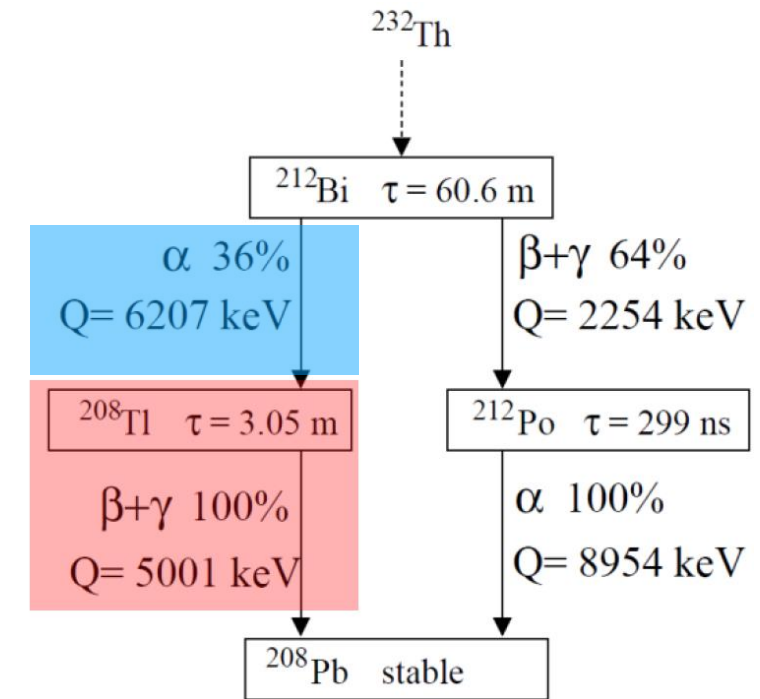
Selection based on light shape parameter



Background – Delayed coincidences rejection



Selection made
Delayed α coincidence ^{212}Bi - ^{208}Tl rejection

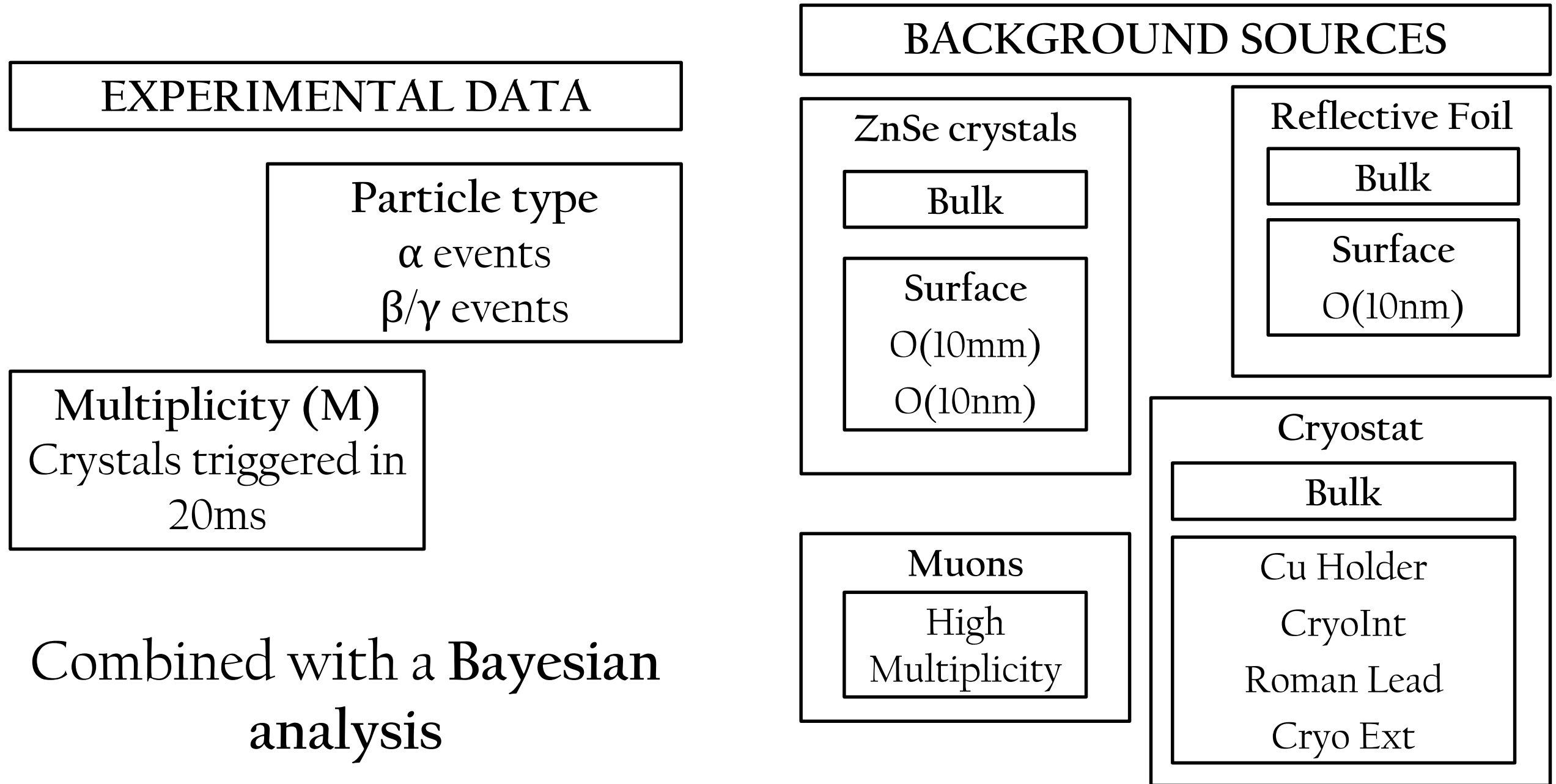


Veto any event succeeding a ^{212}Bi α event in a 7 half-life window

- α pulse shape
- $2.0 \text{ MeV} < \text{Energy} < 6.5 \text{ MeV}$
 - Both peak and surface events

Background model

- A full model is needed to understand the background components



Shape cut for Background Model

$$SP = \frac{1}{\omega_r} \sqrt{\sum_{i=i_M}^{i_M+\omega_r} (y_i - A \cdot S_i)^2}$$

y_i = filtered light pulse

A = maximum amplitude

S_i = filtered average pulse

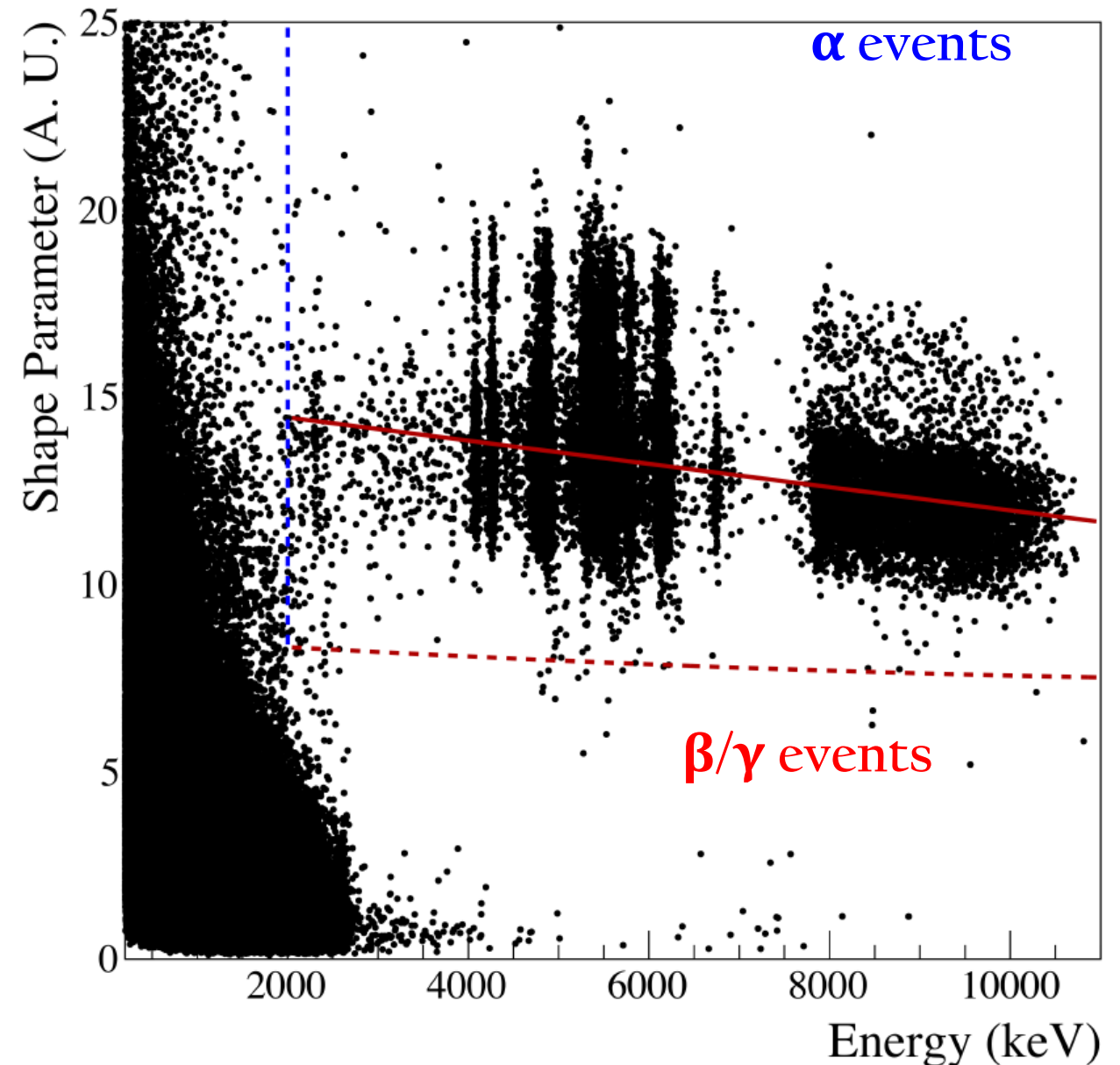
i_M = position of the maximum and

ω_r = right width at half maximum of S_i

Optimized cut:

$$SP = \mu_{\alpha}(E) - 3 \times \sigma_{\alpha}(E)$$

Parameter calculates for $SP > 6$ (pure beta gamma)

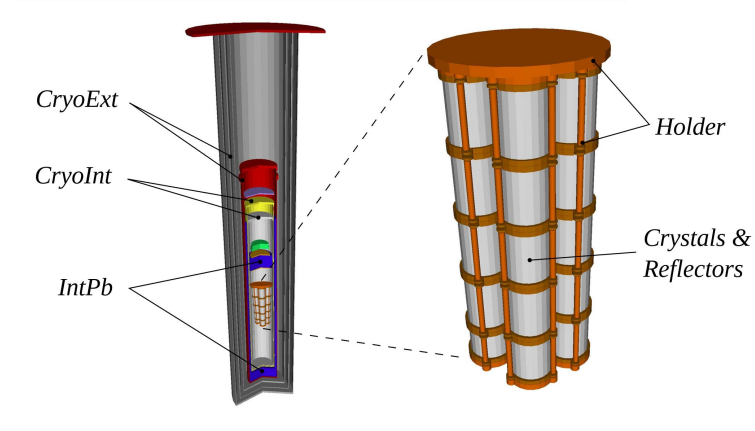


Total Fit Results

Component	Mass (kg)	Source	Index	Activity (Bq/kg)
<i>Crystals</i>	10.5	$2\nu\beta\beta$	1	$(9.96 \pm 0.03) \times 10^{-4}$
		^{65}Zn	2	$(3.52 \pm 0.06) \times 10^{-4}$
		^{40}K	3	$(8.5 \pm 0.4) \times 10^{-5}$
		^{60}Co	4	$(1.4 \pm 0.3) \times 10^{-5}$
		^{147}Sm	5	$(1.6 \pm 0.3) \times 10^{-7}$
		^{238}U - ^{226}Ra	6	$(5.51 \pm 0.10) \times 10^{-6}$
		^{226}Ra - ^{210}Pb	7	$(1.54 \pm 0.02) \times 10^{-5}$
		^{210}Pb - ^{206}Pb	8	$(7.05 \pm 0.16) \times 10^{-6}$
		^{232}Th - ^{228}Ra	9	$(2.74 \pm 0.10) \times 10^{-6}$
		^{228}Ra - ^{208}Pb	10	$(1.20 \pm 0.03) \times 10^{-5}$
		^{235}U - ^{231}Pa	11	$(5.3 \pm 0.7) \times 10^{-7}$
		^{231}Pa - ^{207}Pb	12	$(7.8 \pm 0.4) \times 10^{-7}$
<i>Holder</i>	3.10	^{54}Mn	13	$(2.2 \pm 0.3) \times 10^{-4}$
<i>CryoInt</i> ^(a)	36.9	^{232}Th	14	$< 4.5 \times 10^{-5}$
		^{238}U	15	$(7 \pm 3) \times 10^{-5}$
		^{40}K	16	$(3.0 \pm 0.6) \times 10^{-3}$
		^{60}Co	17	$(6.8 \pm 1.3) \times 10^{-5}$
<i>IntPb</i>	202	^{232}Th	18	$< 6.3 \times 10^{-5}$
		^{238}U	19	$< 7.3 \times 10^{-5}$
<i>CryoExt</i>	832	^{60}Co	20	$(2.6 \pm 0.9) \times 10^{-5}$
<i>ExtPb</i> ^(b)	24694	^{232}Th	21	$(4.3 \pm 0.6) \times 10^{-4}$
		^{238}U	22	$(2.5 \pm 1.2) \times 10^{-4}$
		^{40}K	23	$(2.8 \pm 0.8) \times 10^{-3}$
		^{210}Pb	24	7.8 ± 0.3

^(a) *CryoInt* sources include also a minor contribution from *Holder* bulk contaminations.
^(b) *ExtPb* is used to represent also the *CryoExt* sources, that exhibit degenerate spectra.
^(c) *Reflectors* include also a contribution from light detectors, and from copper surface and other parts directly facing the ZnSe crystals.

Component	Surface (cm ²)	Source	Index	Activity (Bq/cm ²)
<i>Crystals</i>	2574	^{226}Ra - ^{210}Pb - $0.01\mu\text{m}$	25	$(2.63 \pm 0.15) \times 10^{-8}$
		^{228}Ra - ^{208}Pb - $0.01\mu\text{m}$	26	$(6.5 \pm 1.1) \times 10^{-9}$
		^{226}Ra - ^{210}Pb - $10\mu\text{m}$	27	$< 2.3 \times 10^{-9}$
		^{228}Ra - ^{208}Pb - $10\mu\text{m}$	28	$(4.2 \pm 1.6) \times 10^{-9}$
<i>Reflectors</i> ^(c)	2100	^{232}Th - $10\mu\text{m}$	29	$< 7.3 \times 10^{-10}$
		^{226}Ra - ^{210}Pb - $10\mu\text{m}$	30	$(8.7 \pm 1.3) \times 10^{-9}$
		^{210}Pb - ^{206}Pb - $10\mu\text{m}$	31	$(1.0 \pm 0.5) \times 10^{-8}$
		^{210}Pb - ^{206}Pb - $0.01\mu\text{m}$	32	$(1.43 \pm 0.02) \times 10^{-7}$
Muons	Flux in units of $\mu/(\text{cm}^2\text{s})$		33	$(3.7 \pm 0.2) \times 10^{-8}$



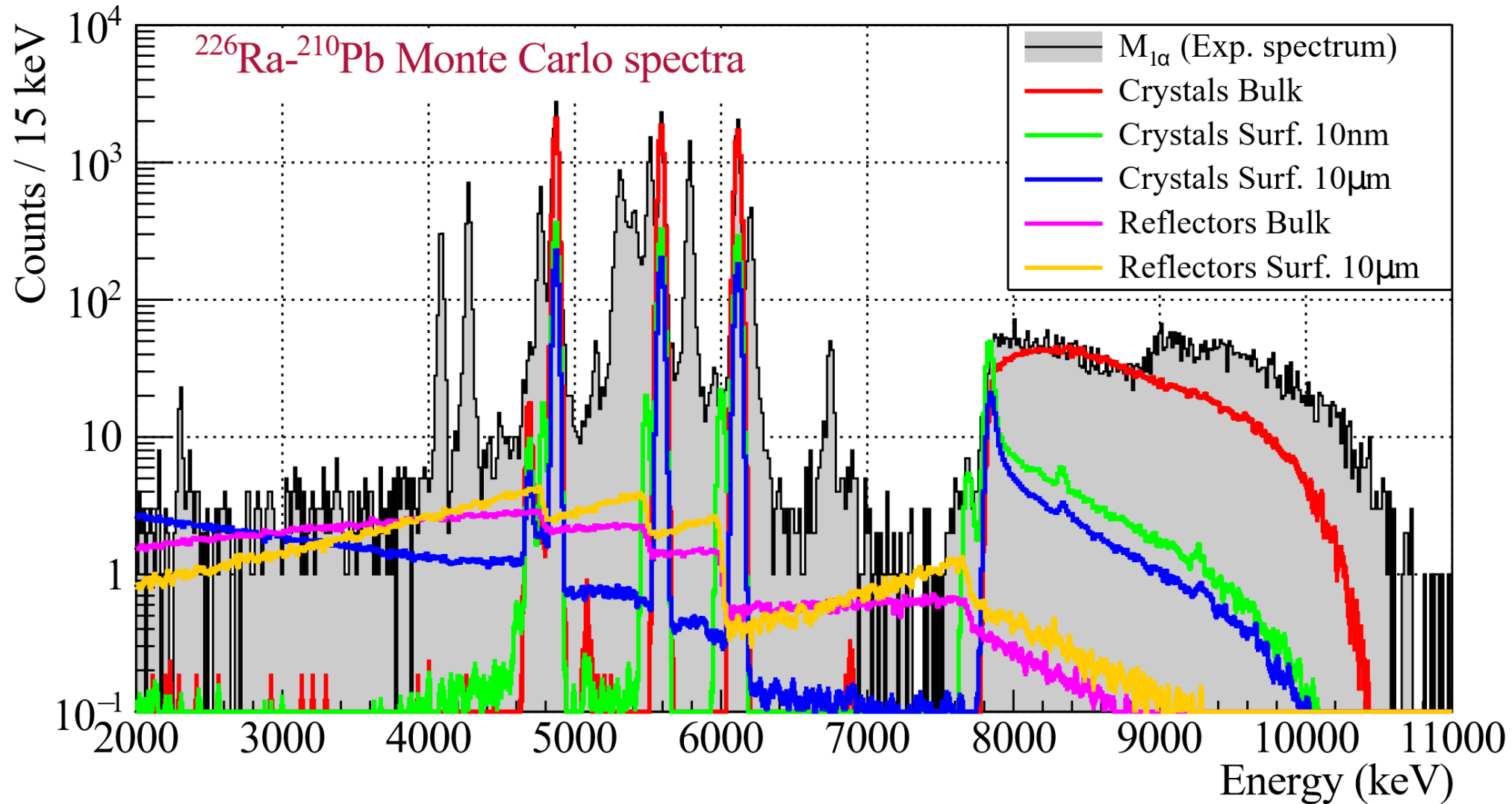
Surface
Exponential
profile

ROI contaminations

After time veto of 7 livetimes

Component	ROI _{bk_g} rate (10 ⁻⁴ counts/(keV kg yr))	Source	ROI _{bk_g} rate (10 ⁻⁴ counts/(keV kg yr))
<i>Crystals</i>	$11.7 \pm 0.6 \begin{smallmatrix} +1.6 \\ -0.8 \end{smallmatrix}$	²³² Th– bulk	$3.4 \pm 0.6 \pm 0.1$
		²³² Th–surf	$3.4 \pm 0.5 \begin{smallmatrix} +1.0 \\ -0.7 \end{smallmatrix}$
		²³⁸ U–surf	$4.9 \pm 0.3 \begin{smallmatrix} +1.3 \\ -0.3 \end{smallmatrix}$
<i>Reflectors & Holder</i>	$2.1 \pm 0.3 \begin{smallmatrix} +2.2 \\ -1.0 \end{smallmatrix}$	²³² Th	< 3.3
		²³⁸ U	$1.8 \pm 0.3 \begin{smallmatrix} +1.4 \\ -0.9 \end{smallmatrix}$
<i>Cryostat & Shields</i>	$5.9 \pm 1.3 \begin{smallmatrix} +7.2 \\ -2.9 \end{smallmatrix}$	²³² Th	$3.5 \pm 1.3 \begin{smallmatrix} +7.4 \\ -3.3 \end{smallmatrix}$
		²³⁸ U	$2.4 \pm 0.4 \begin{smallmatrix} +4.1 \\ -0.7 \end{smallmatrix}$
Subtotal	$19.8 \pm 1.4 \begin{smallmatrix} +6.6 \\ -2.7 \end{smallmatrix}$		
Muons	$15.3 \pm 1.3 \pm 2.5$		
$2\nu\beta\beta$	6.0 ± 0.3 (< 3 × 10 ⁻⁶ counts/(keV kg yr) in [2.95–3.05] MeV range)		
Total	$41 \pm 2 \begin{smallmatrix} +9 \\ -4 \end{smallmatrix}$		
Experimental	$35 \begin{smallmatrix} +10 \\ -9 \end{smallmatrix}$		

Alpha vs Bulk



- Different impact on the continuum