

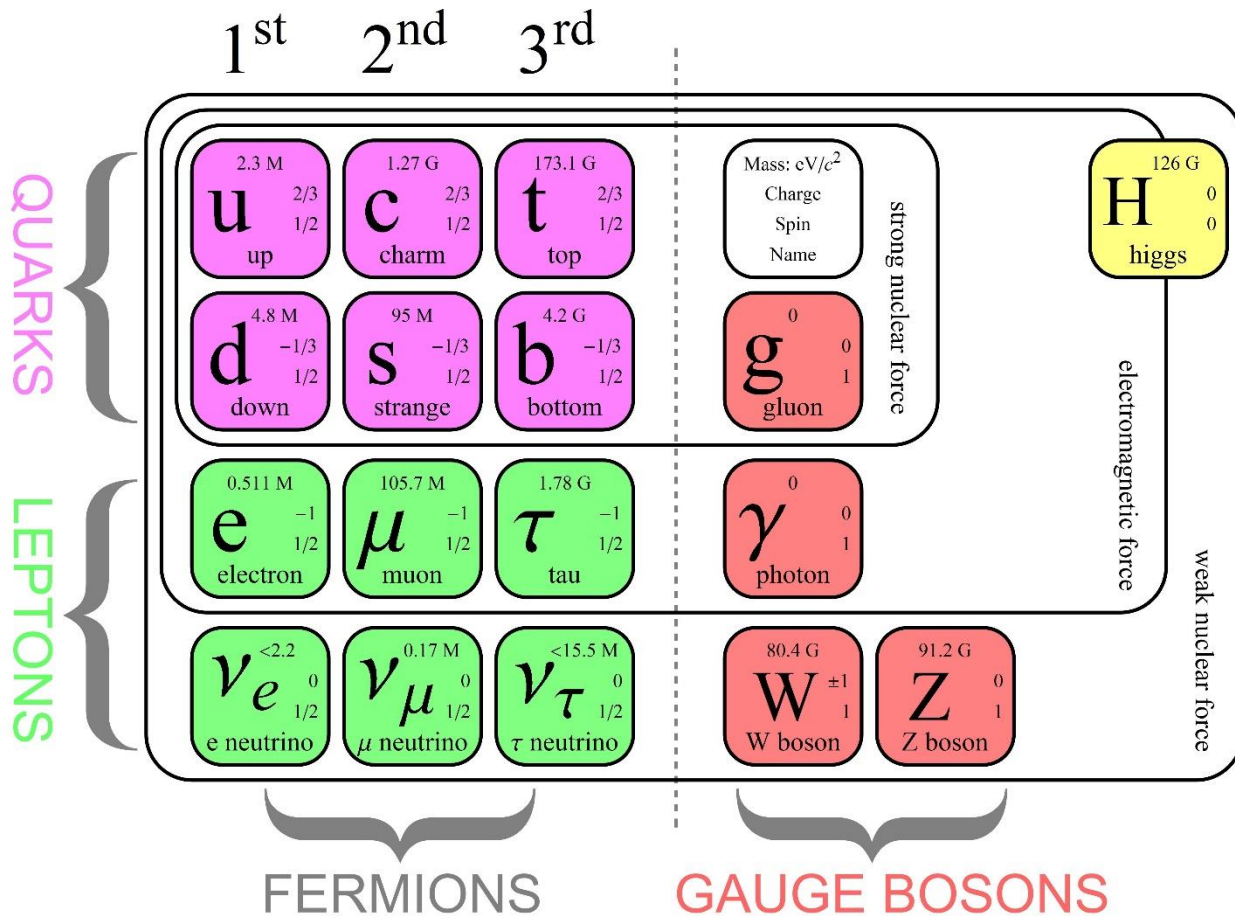
NOvA long-baseline neutrino experiment: recent results

Kirk Bays

California Institute of Technology

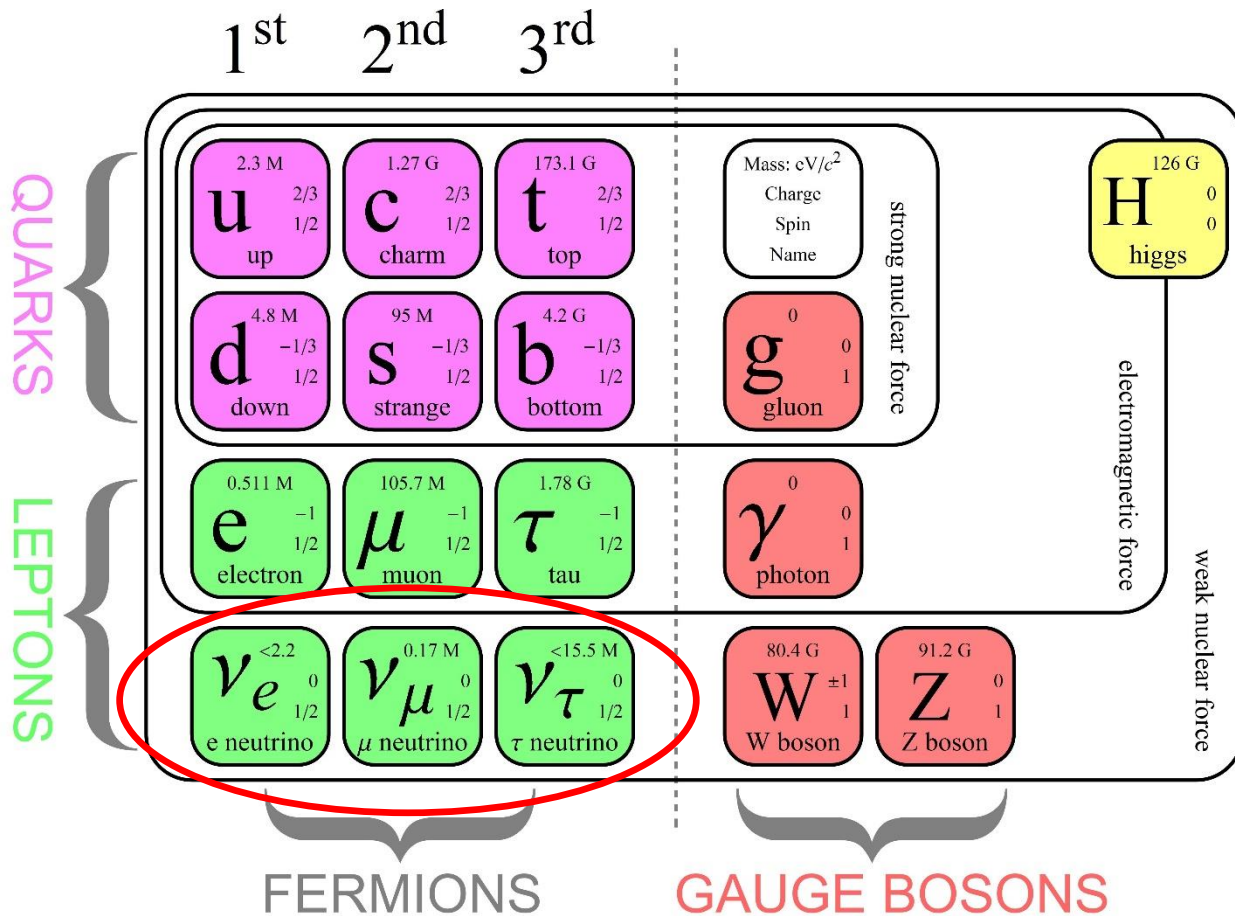
April 3, 2018





Standard model of physics:

12 fundamental particles
 3 forces / 4 gauge bosons
 Higgs

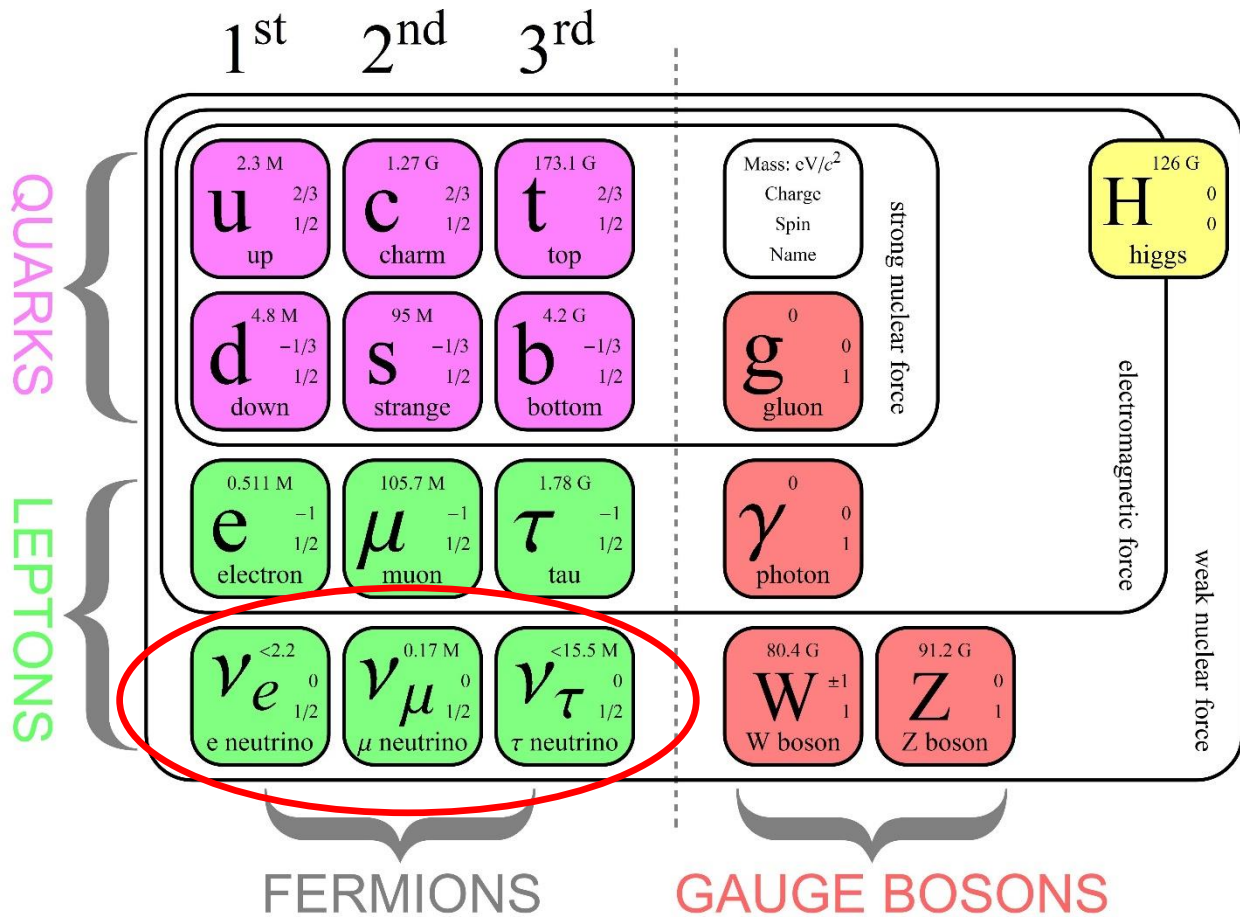


Standard model of physics:

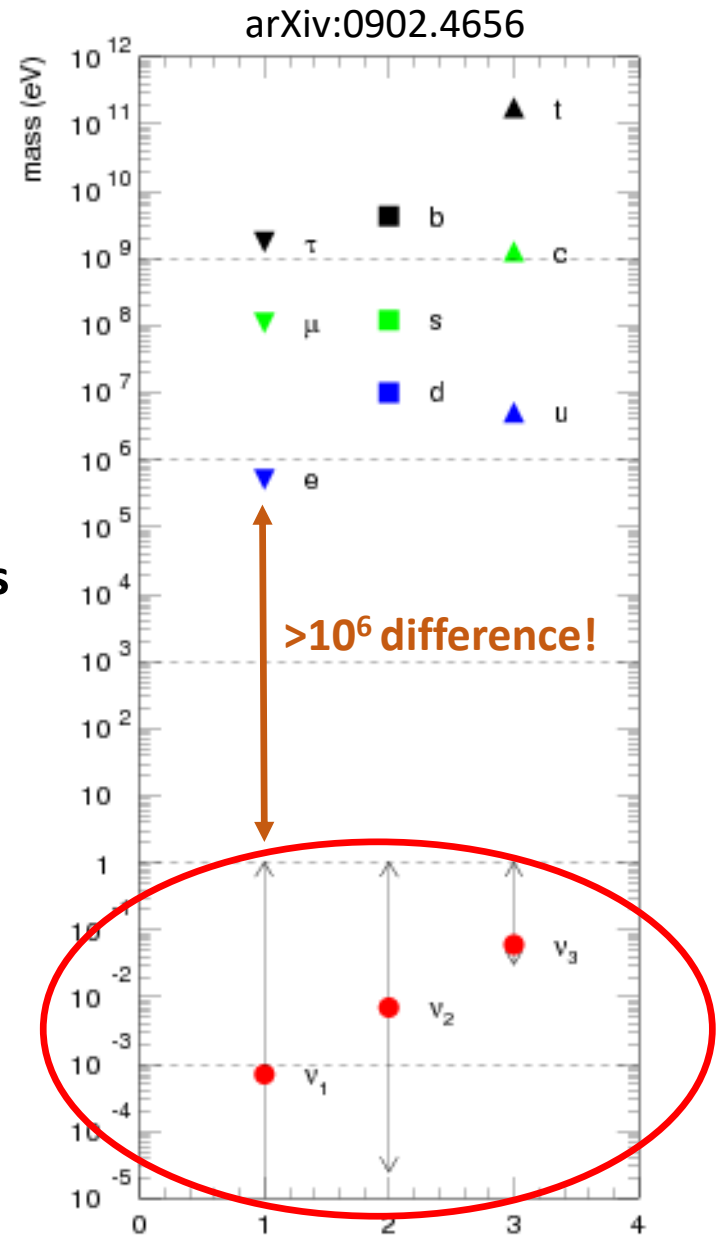
12 fundamental particles
3 forces / 4 gauge bosons
Higgs

3 are neutrinos
Neutrinos are interesting! Why?

original theory: neutrino mass = 0
but we know this is wrong!



Neutrino masses are very small



Neutrinos oscillate - their flavor states (e, μ, τ) are different than their mass states (1, 2, 3)

Connected by mixing matrix:
$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = U^* \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

Mixing matrix \rightarrow mixing angles – 2 flavor case:
$$\begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix} = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}$$

3 flavor case:

$$U = \begin{bmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{bmatrix} = \underbrace{\begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta_{23} & \sin \theta_{23} \\ 0 & -\sin \theta_{23} & \cos \theta_{23} \end{pmatrix}}_{\text{Atmospheric}} \underbrace{\begin{pmatrix} \cos \theta_{13} & 0 & \sin \theta_{13} e^{-i\delta_{CP}} \\ 0 & 1 & 0 \\ -\sin \theta_{13} e^{-i\delta_{CP}} & 0 & \cos \theta_{13} \end{pmatrix}}_{\text{Reactors; very recent!}} \underbrace{\begin{pmatrix} \cos \theta_{12} & \sin \theta_{12} & 0 \\ -\sin \theta_{12} & \cos \theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}}_{\text{Solar}}$$

3 angles and a CP-violation term determine the matrix: $\theta_{12}, \theta_{13}, \theta_{23}, \delta_{CP}$

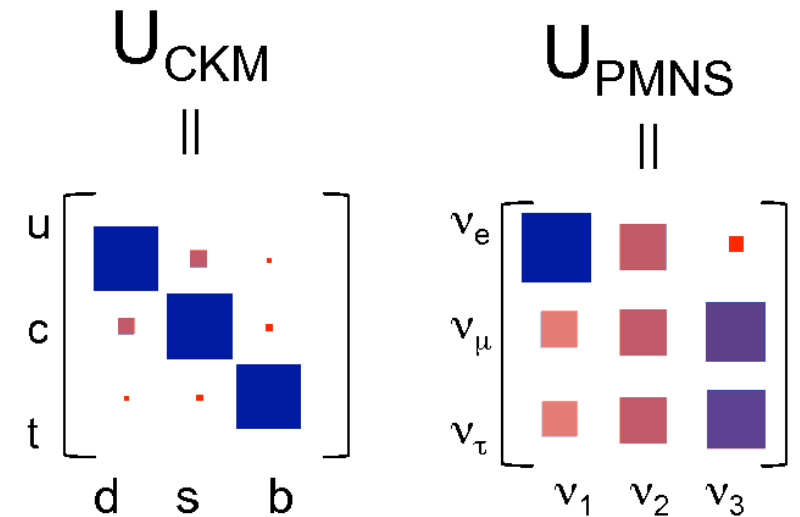
How well are angles measured so far? (PDG)

$$\sin^2(\theta_{12}) = 0.307 \pm 0.013$$

$$\sin^2(\theta_{23}) = 0.51 \pm 0.04 \text{ ('maximal'? or which octant?)}$$

$$\sin^2(\theta_{13}) = 0.021 \pm .0011$$

Neutrino mixing (U_{PMNS}) analogous to quark mixing (U_{CKM}) but much less diagonal



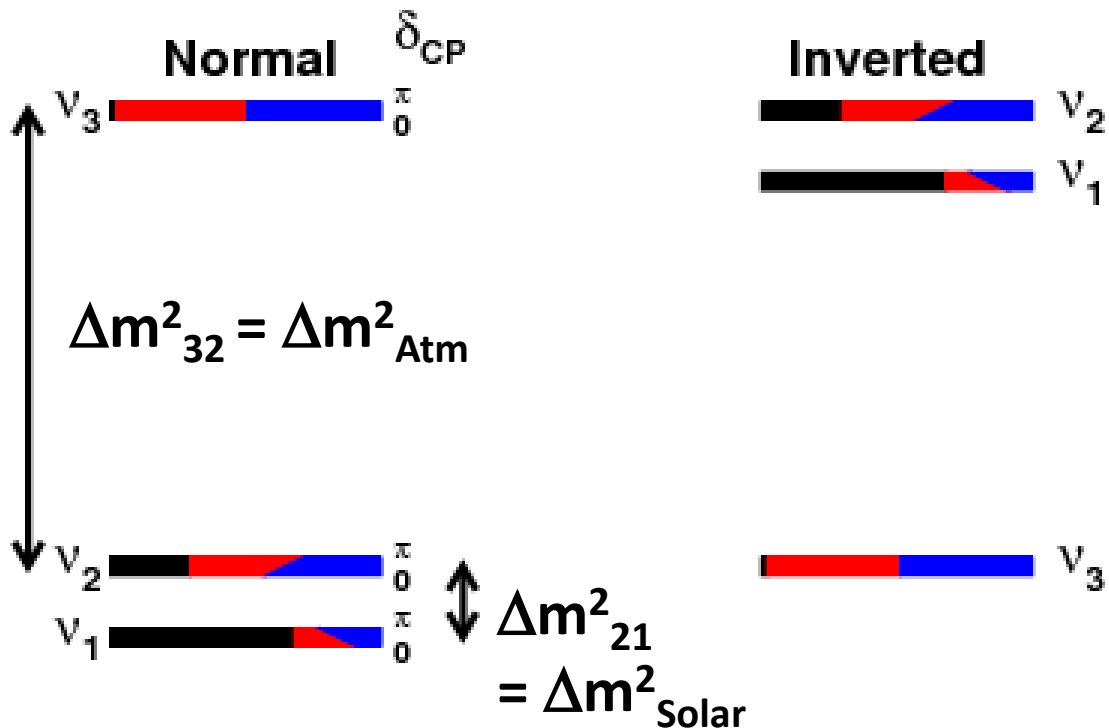
$$P_{\alpha \rightarrow \beta} = |\langle \nu_{\beta}(t) | \nu_{\alpha} \rangle|^2 = \left| \sum_i U_{\alpha i}^* U_{\beta i} e^{-im_i^2 L/2E} \right|^2$$

$$P(\nu_{\mu} \rightarrow \nu_{\mu}) \approx 1 - \sin^2(2\theta_{23}) \sin^2(1.27 \Delta m_{32}^2 L/E)$$

L = baseline
 E of neutrino } experiment setup

$\Delta m_{32}^2 = m_3^2 - m_2^2$ (to be measured)
 also Δm_{21}^2

arXiv:1505.01891



$$\Delta m^2_{21} = (7.53 \pm 0.18) \times 10^{-5} \text{ eV}^2$$

$$\Delta m^2_{32} = (2.45 \pm 0.05) \times 10^{-3} \text{ eV}^2$$

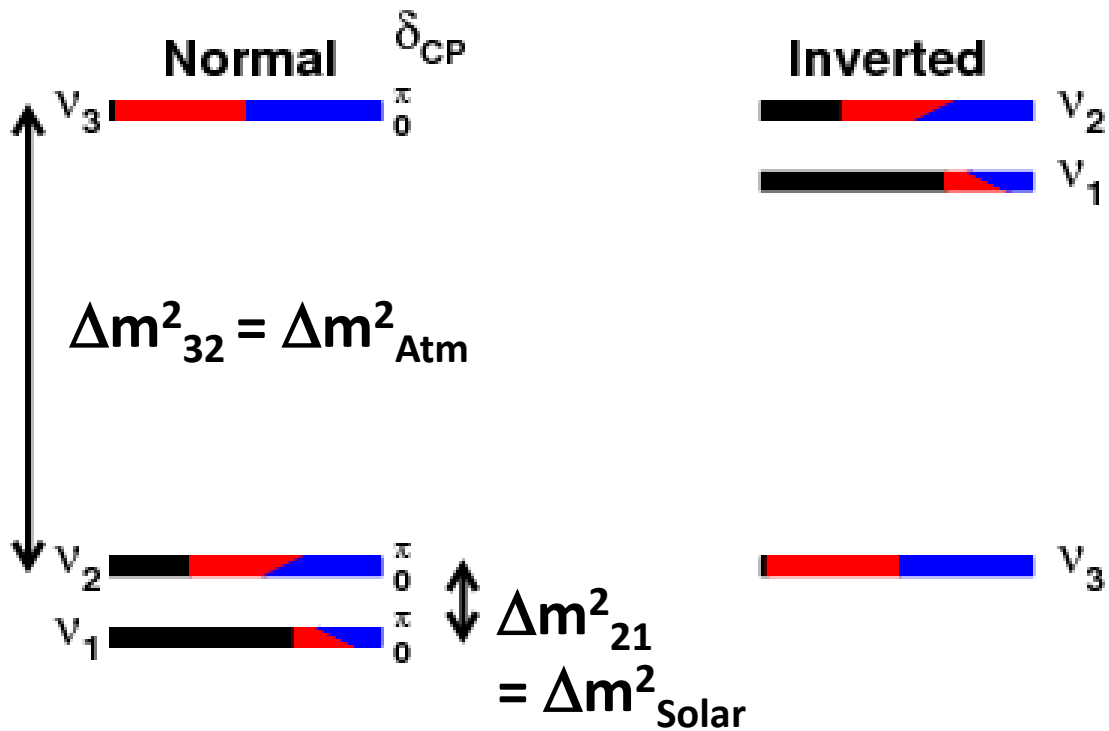
Two possible mass orderings ('hierarchies'): 'Normal' and 'Inverted'. Can't tell which yet!

When travelling through matter, there are additional oscillation effects as ν_e feels extra 'drag'

The farther the path through matter the better we measure the hierarchy

arXiv:1505.01891

 ν_e ν_μ ν_τ



Things to be measured with
neutrino oscillation experiments:

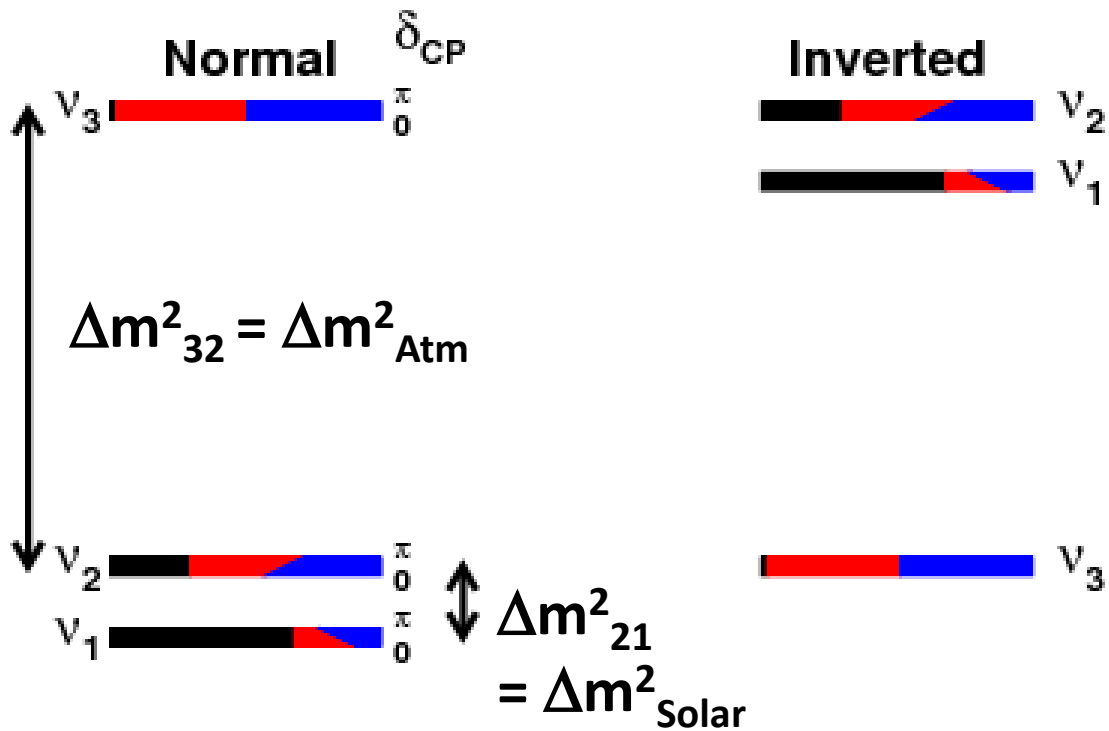
$\theta_{12}, \theta_{13}, \theta_{23}$

$\Delta m^2_{32}, \Delta m^2_{21}$, mass hierarchy

δ_{CP} (is CP violated by neutrinos?)

arXiv:1505.01891

ν_e
 ν_μ
 ν_τ



Long-baseline ν oscillation experiments can measure these especially well

$\theta_{12}, \theta_{13}, \theta_{23}$
 $\Delta m^2_{32}, \Delta m^2_{21}$, mass hierarchy
 δ_{CP} (is CP violated by neutrinos?)

Why does it matter?

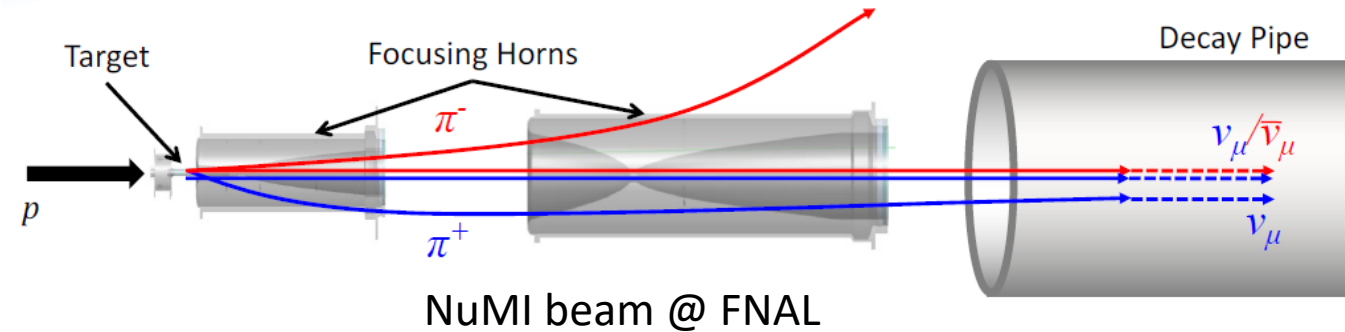
- Fundamental properties of neutrinos affect lots of other things:
 - Cosmology, astrophysics
Why is universe matter and not anti-matter? $\sin(\delta_{CP}) \neq 0 \rightarrow$ leptogenesis?
 - Phenomenology, GUTs
Are neutrinos their own anti-particle (Majorana)? $0\nu\beta\beta$, see-saw mechanism
- Can also measure neutrinos to learn about interesting sources
 - supernova neutrinos
 - sterile neutrinos
 - solar, cosmic ray neutrinos



Detectors used in long-baseline experiments can do this too

A long-baseline experiment: NOvA

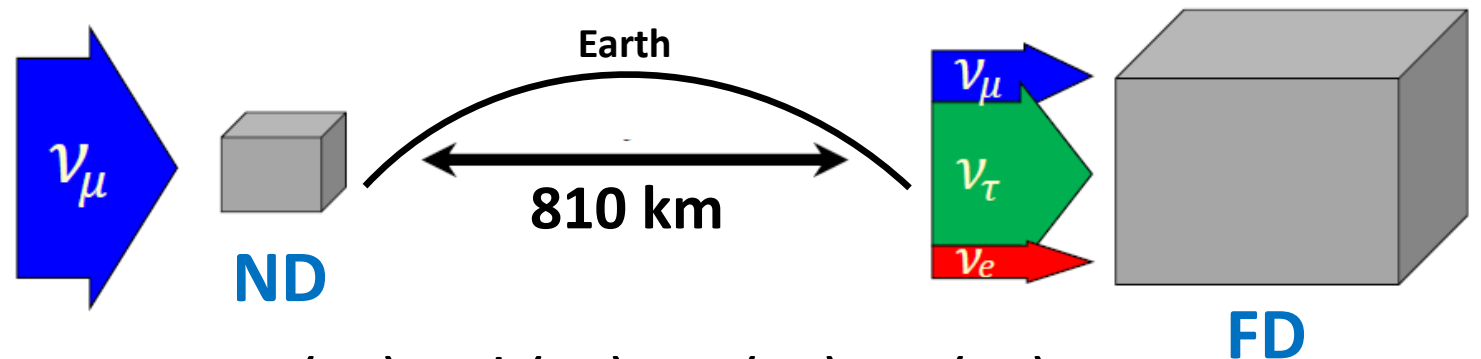
First, make a neutrino beam:
(almost pure ν_μ)



Then shoot it at your detectors:

A Near Detector (**ND**) near the beam, before oscillations

And a Far Detector (**FD**) far away

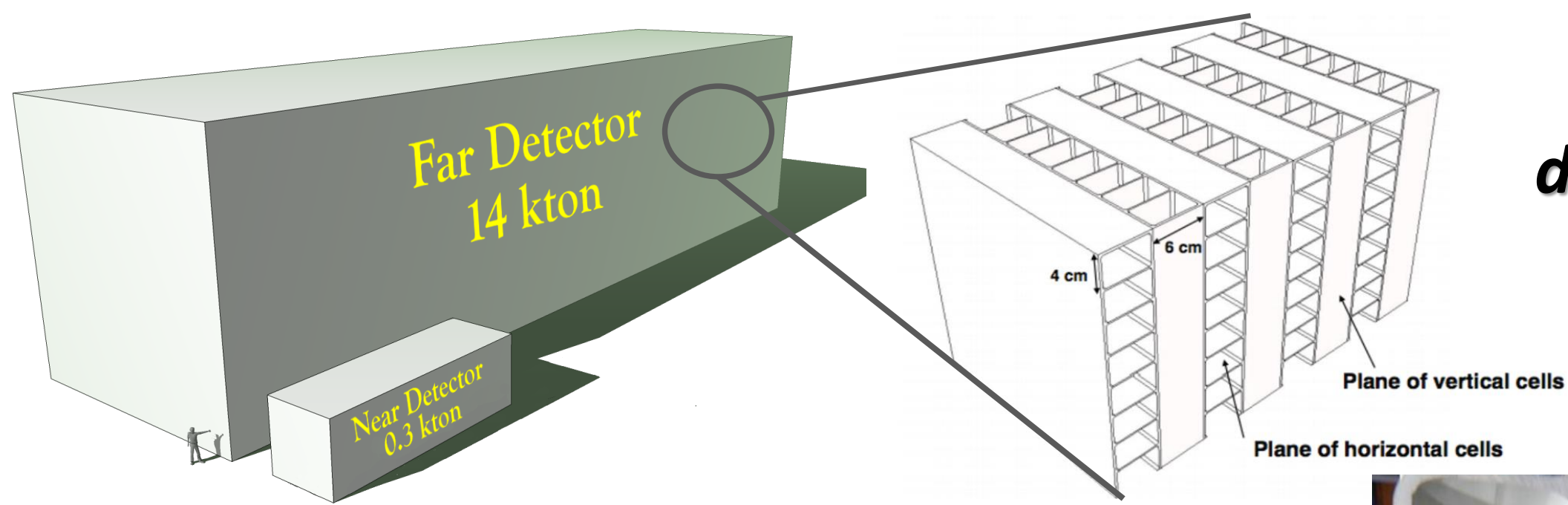


$$N_{\text{ND}}(E_\nu) = \phi(E_\nu) \times \sigma(E_\nu) \times \varepsilon(E_\nu)$$

$$N_{\text{FD}}(E_\nu) = \phi(E_\nu) \times \sigma(E_\nu) \times \varepsilon(E_\nu) \times P(E_\nu) \dots$$

ND helps constrain FD uncertainties

NOvA detectors

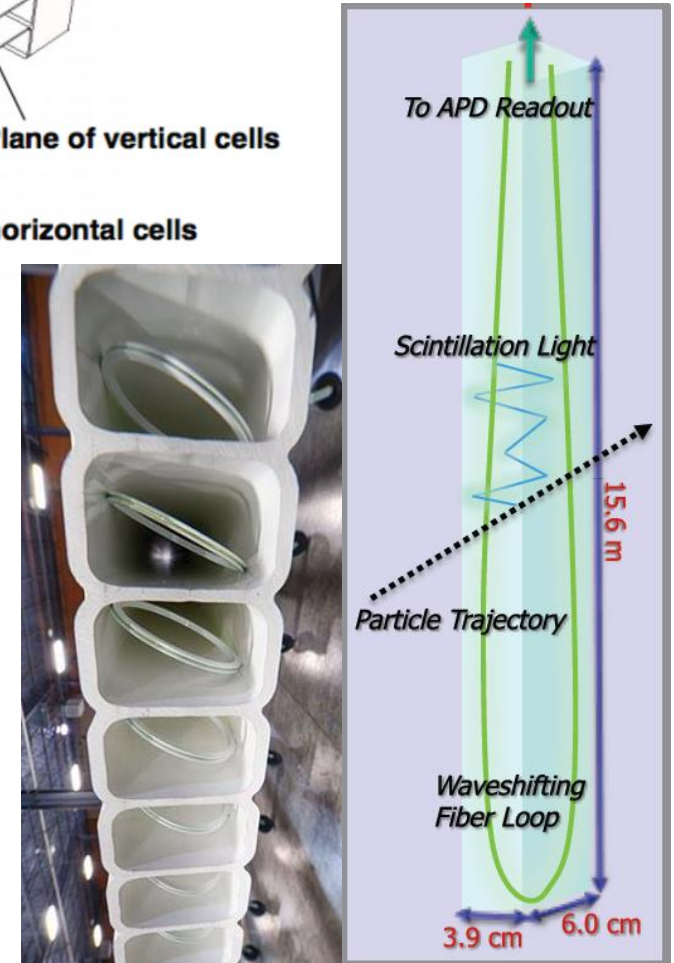


FD (at Ash River, MN, 810 km baseline):

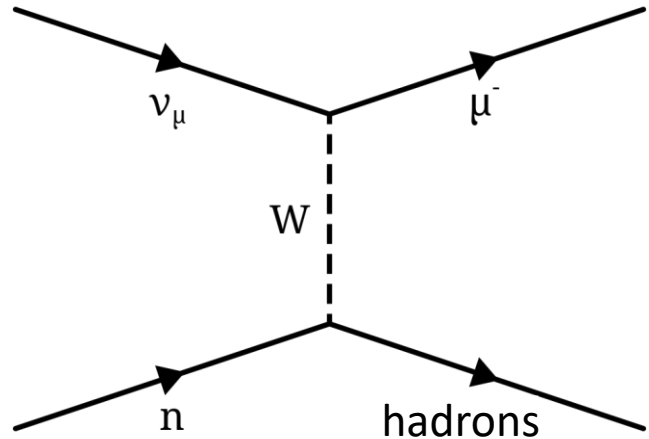
16m x 16m x 60m, 14kton, on surface (some barite overburden)
~2/3 liquid scintillator by mass, ~344,000 cells, 896 planes
low-Z, finely-segmented, 62% active
1 radiation length ~ 6-10 cells

ND (@ FNAL, 1km from NuMI target):

4m x 4m x 16m, 0.3kton, underground
~20,000 cells, design similar to FD



Neutrino interactions



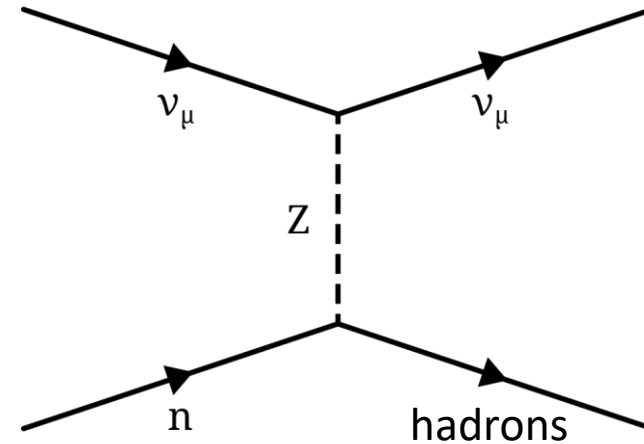
Charged current:

$\nu_x + \text{nucleon} \rightarrow x + \text{hadrons}$

$x = e, \mu, \tau$

hadrons: can be single p (QE)

can be shower (p, π, \dots)



Neutral current:

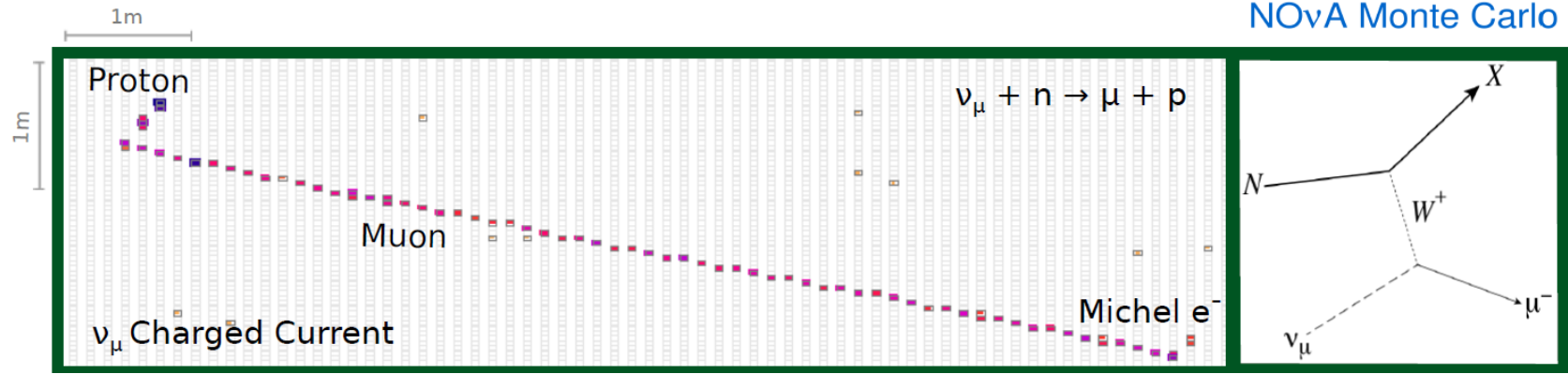
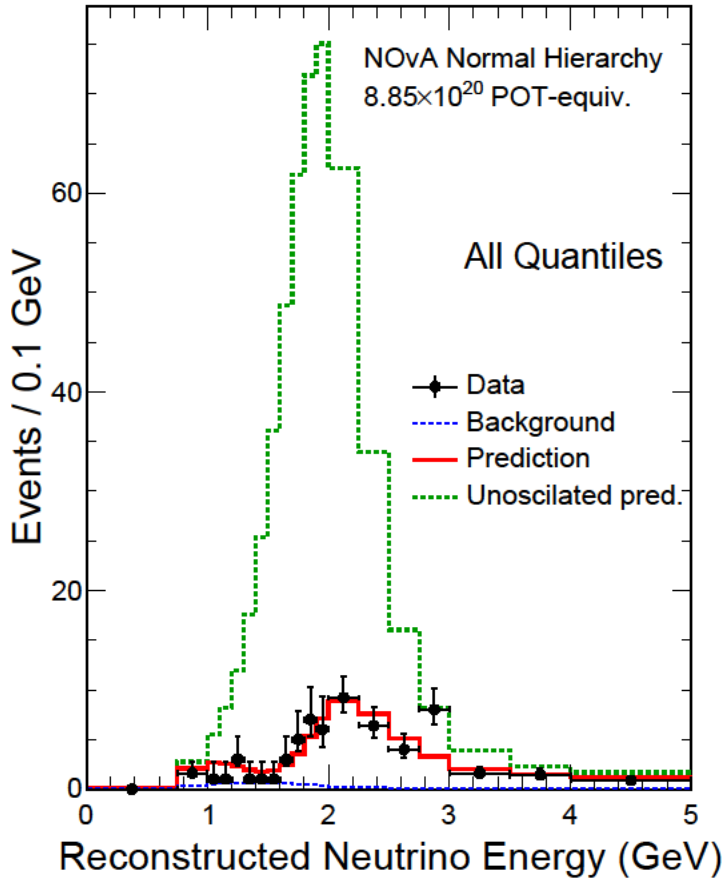
$\nu + \text{nucleon} \rightarrow \nu + \text{hadrons}$

flavor blind

no lepton

ν_μ disappearance

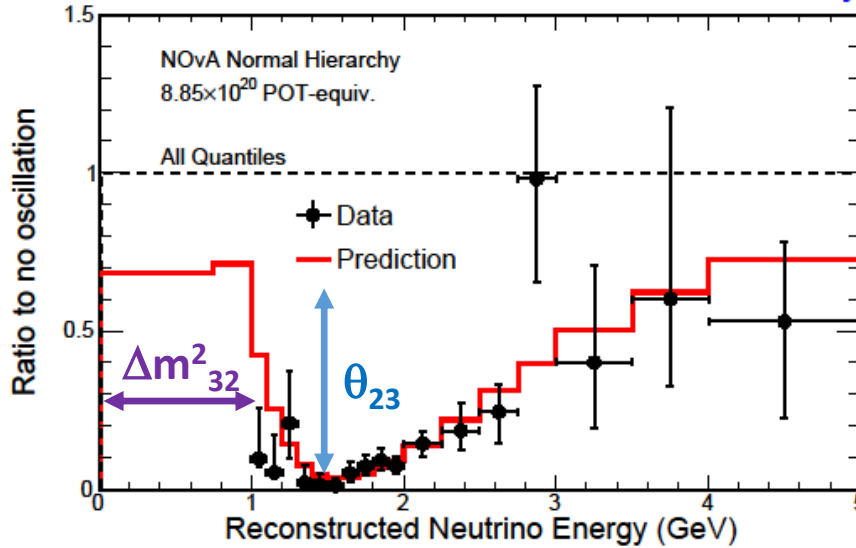
NOvA Preliminary



$$P(\nu_\mu \rightarrow \nu_\mu) \approx 1 - \sin^2(2\theta_{23})\sin^2(1.27\Delta m_{32}^2 L/E_\nu)$$

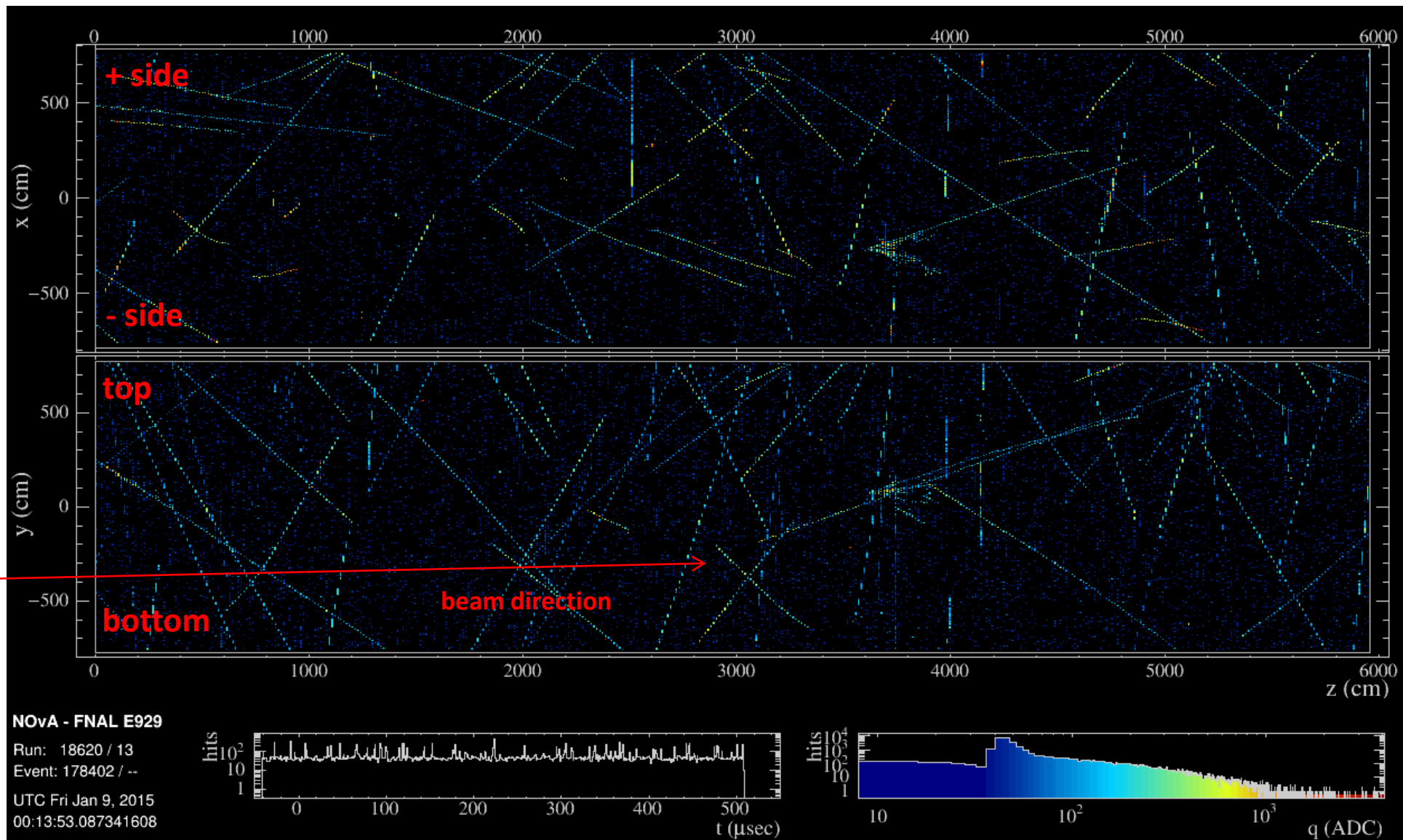
directly measure θ_{23} and Δm_{32}^2

NOvA Preliminary

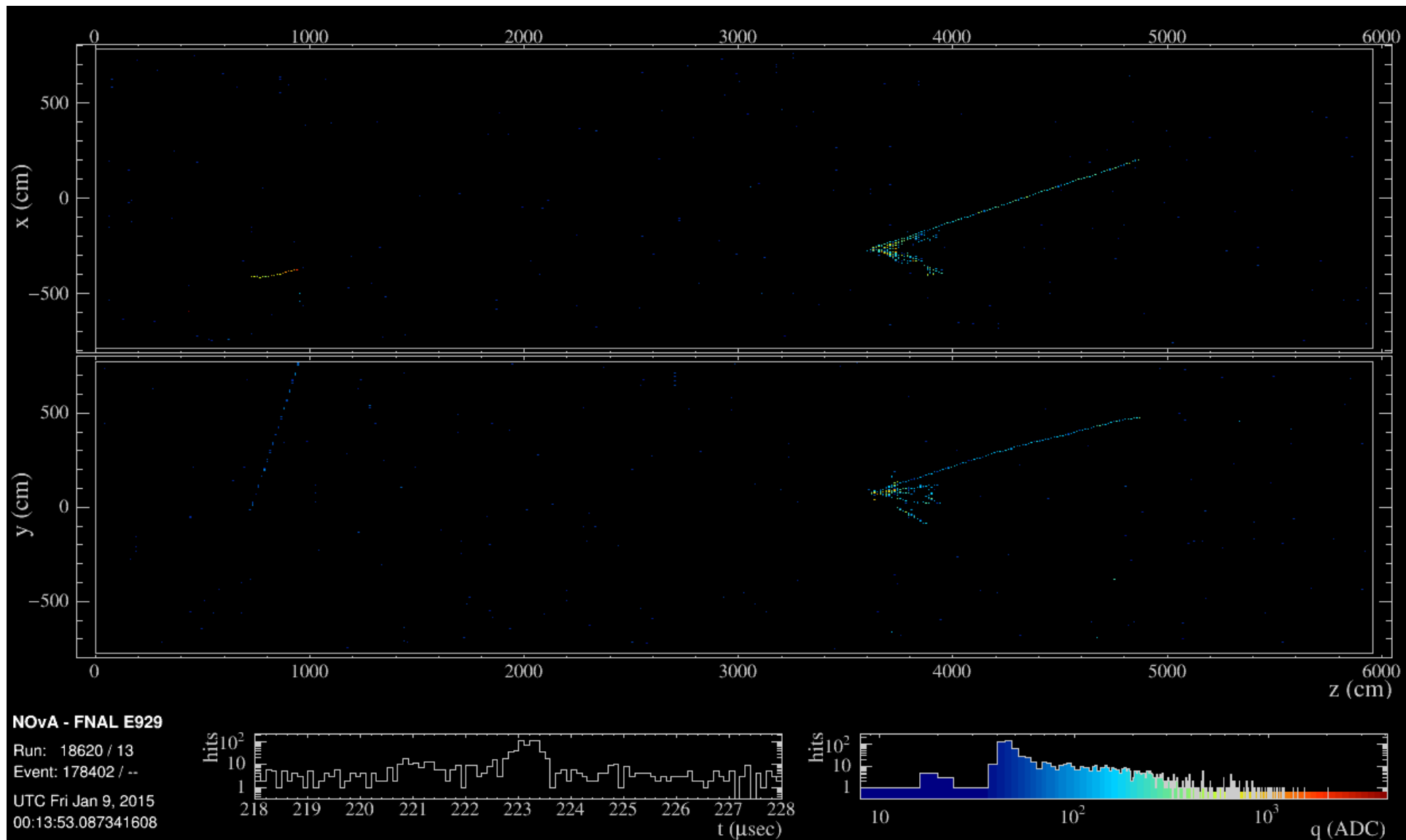


- signature: muon**
- backgrounds:**
 - neutral current**
 - π , p can look like muons**
 - cosmic rays**
 - surface detector, real muons**

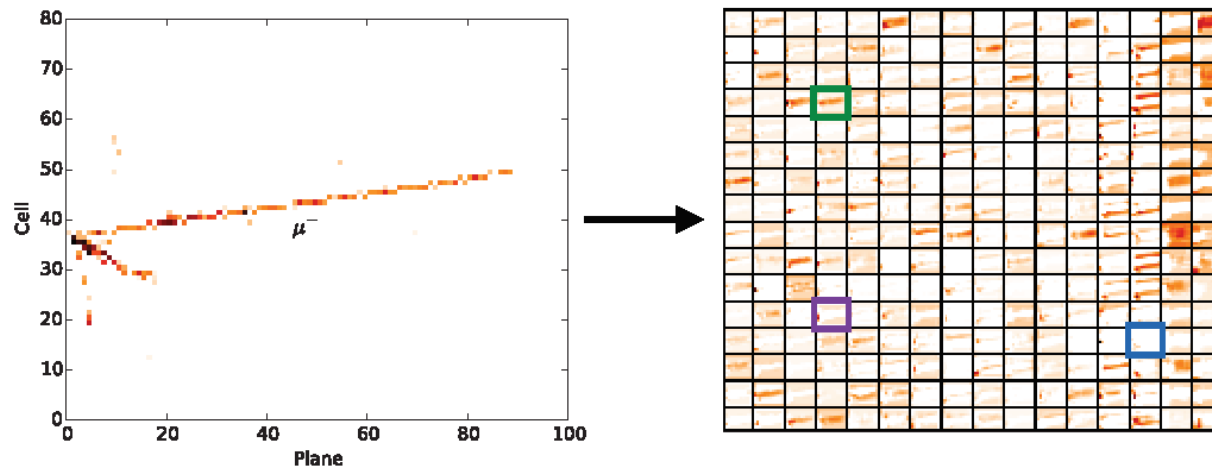
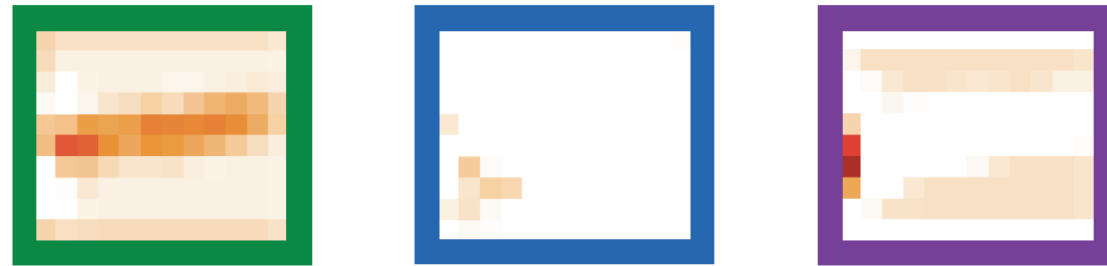
550 μs exposure of the Far Detector



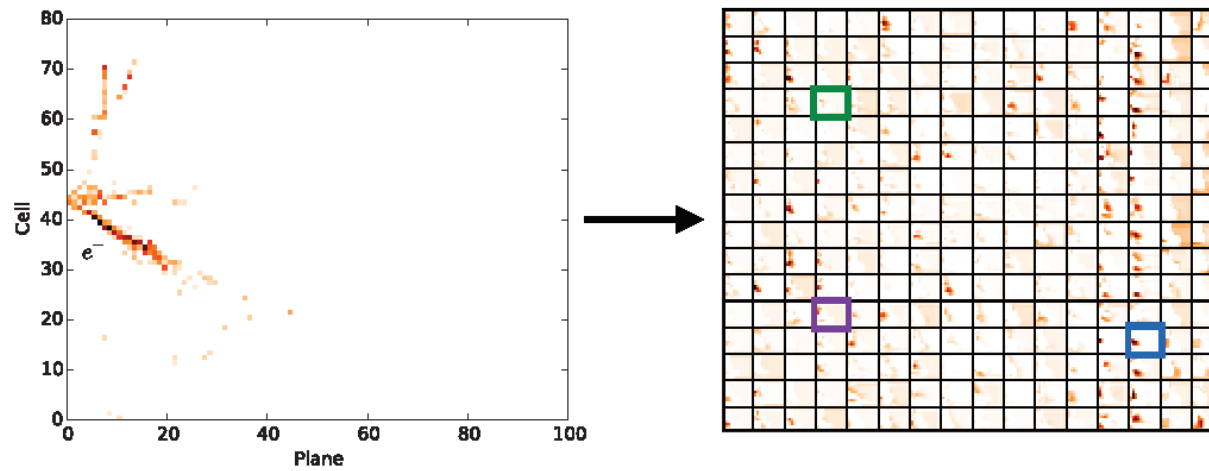
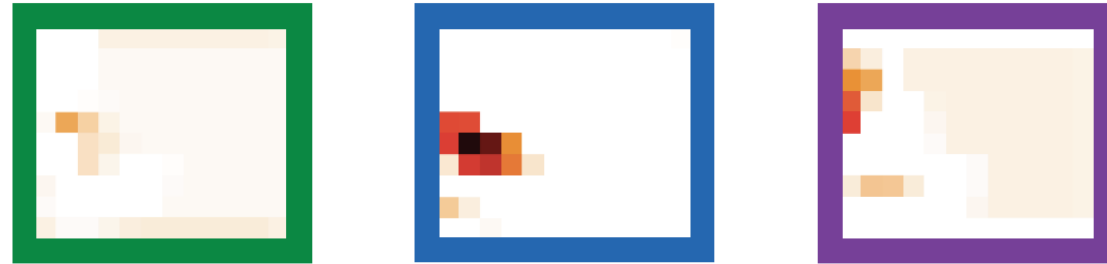
Time-zoom on 10 μ s interval during NuMI beam pulse

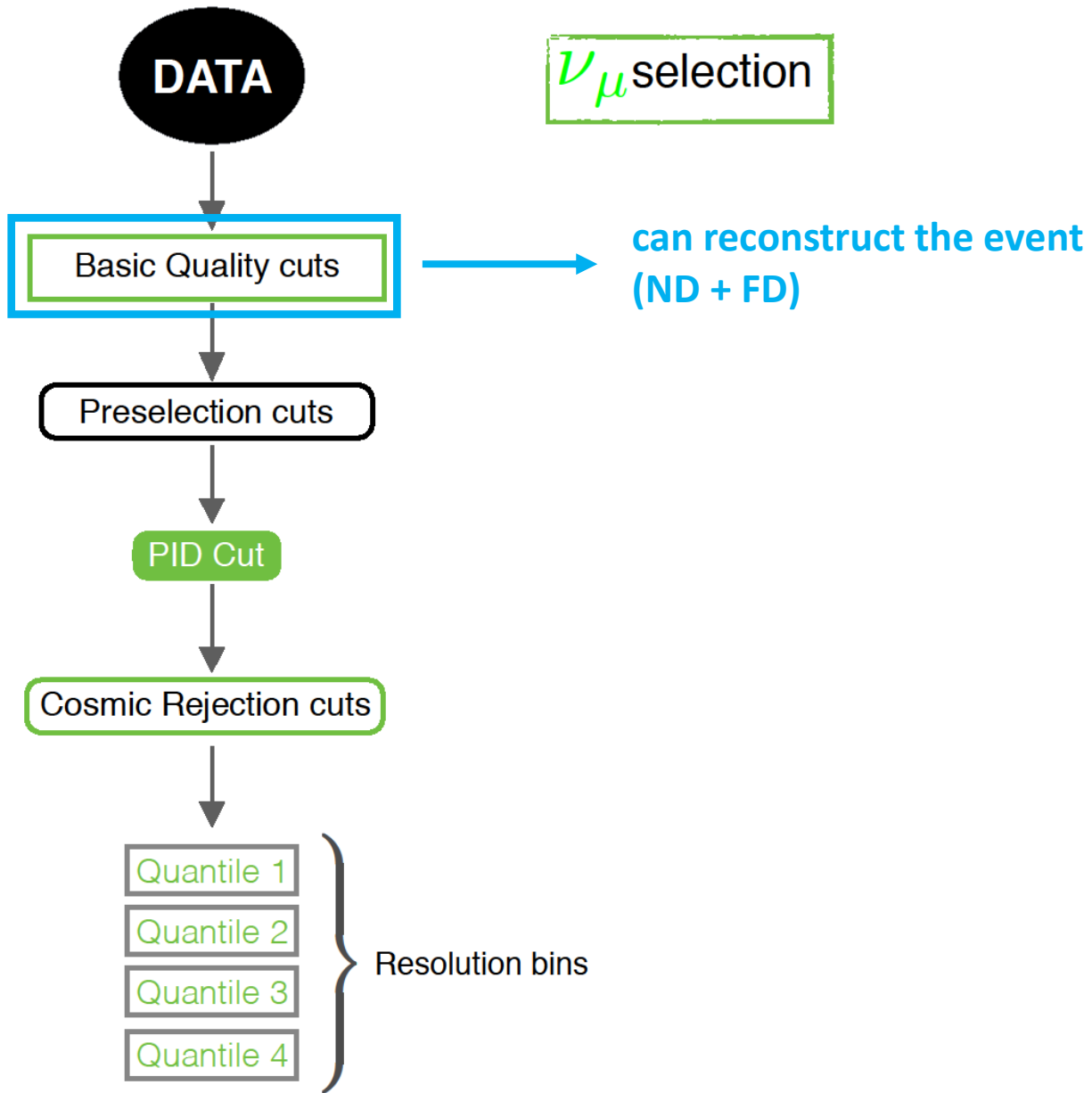


CVN - disappearance



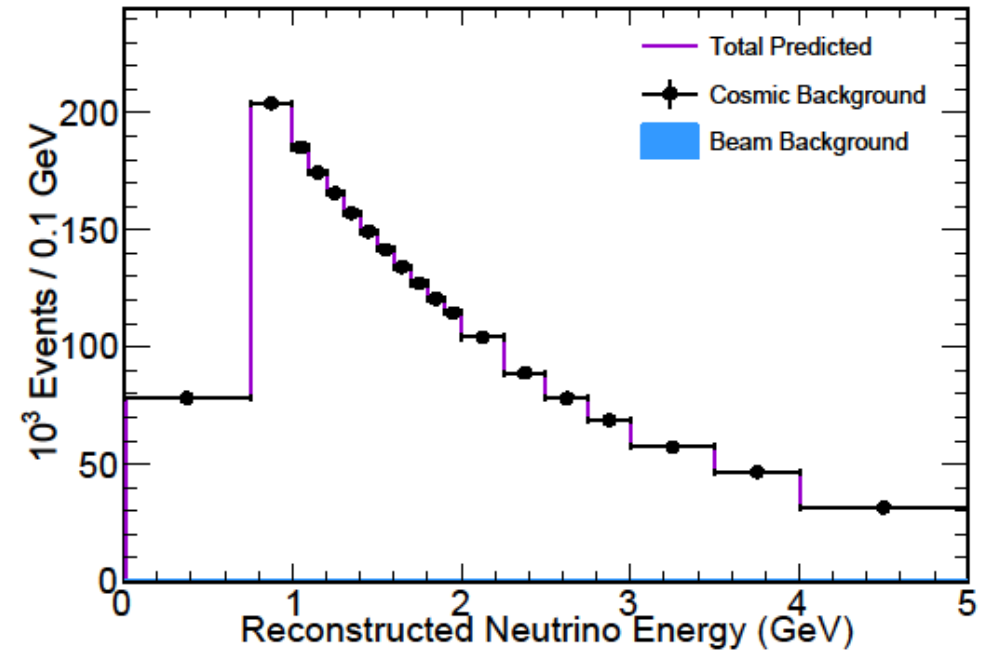
CVN - appearance

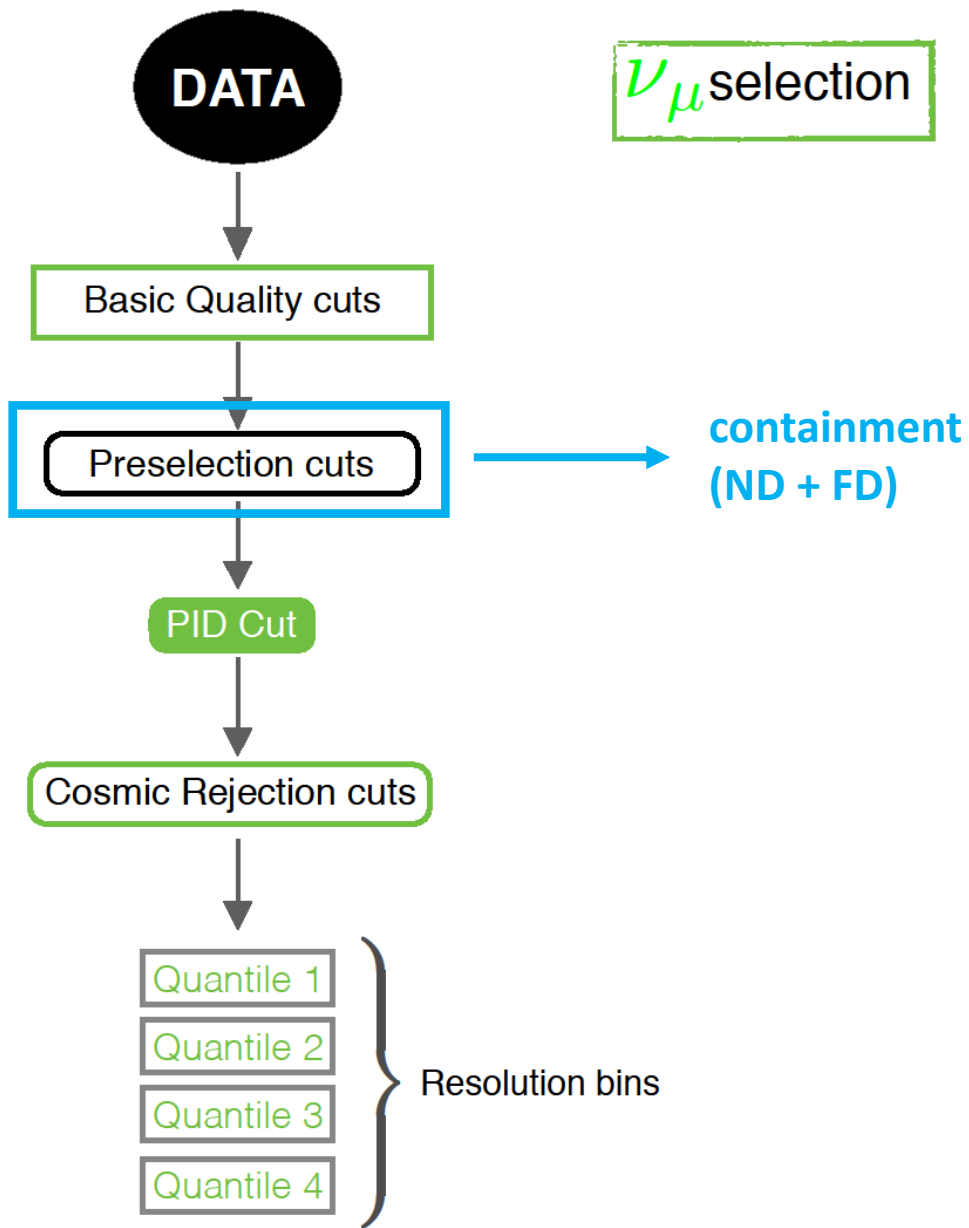




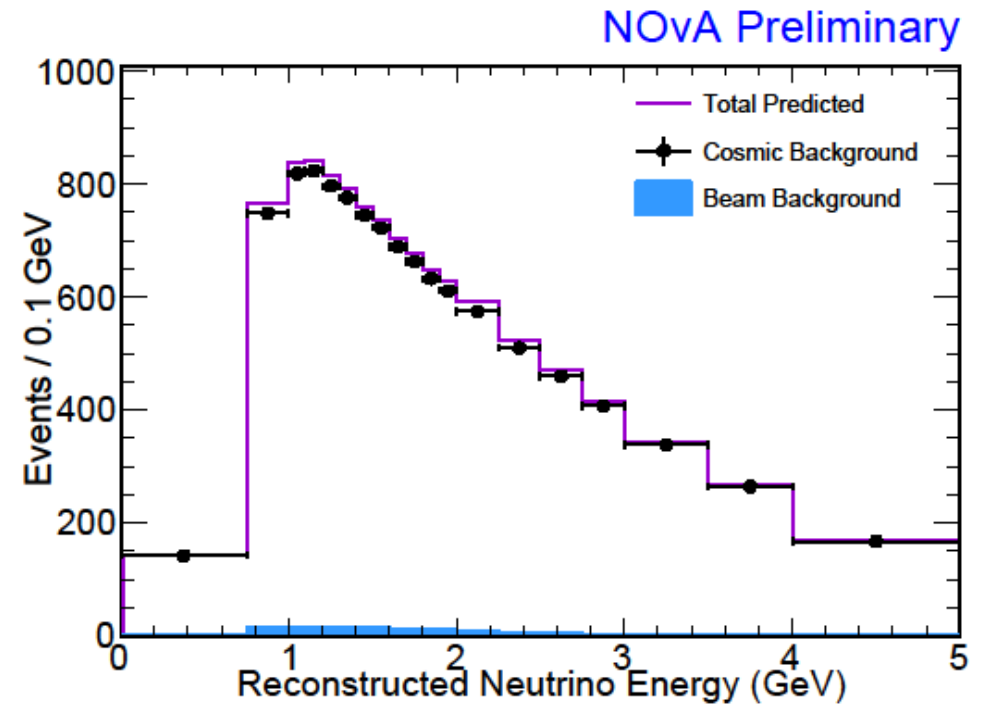
FD: $\sim 10^6$ events

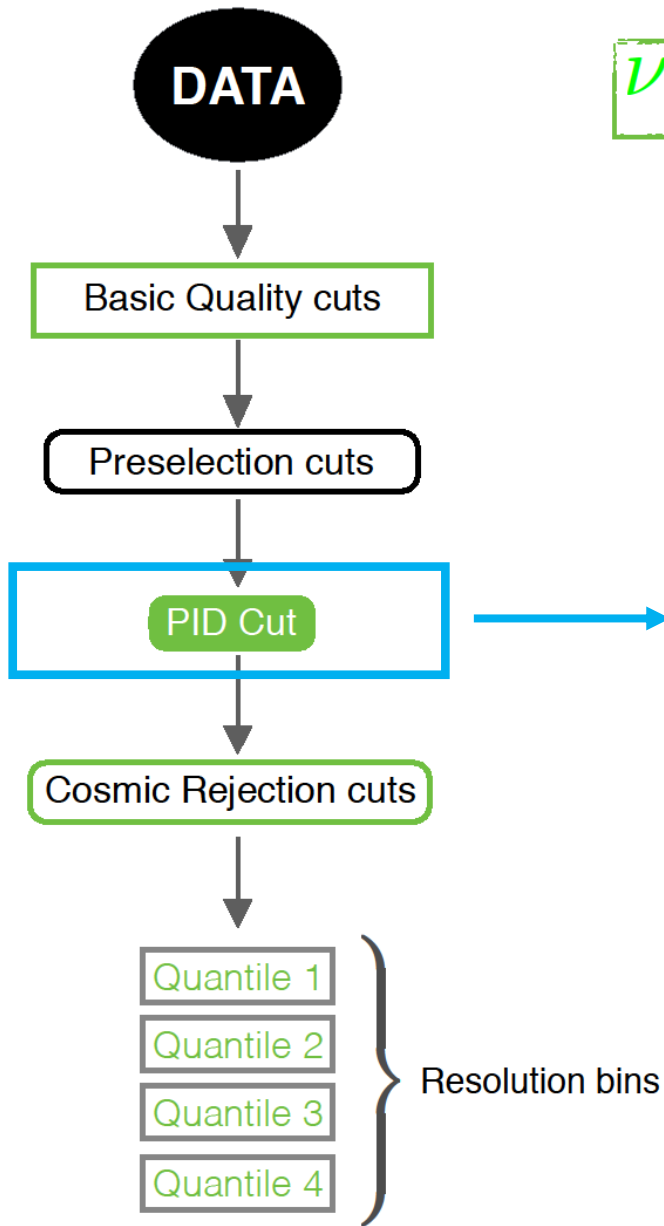
NOvA Preliminary



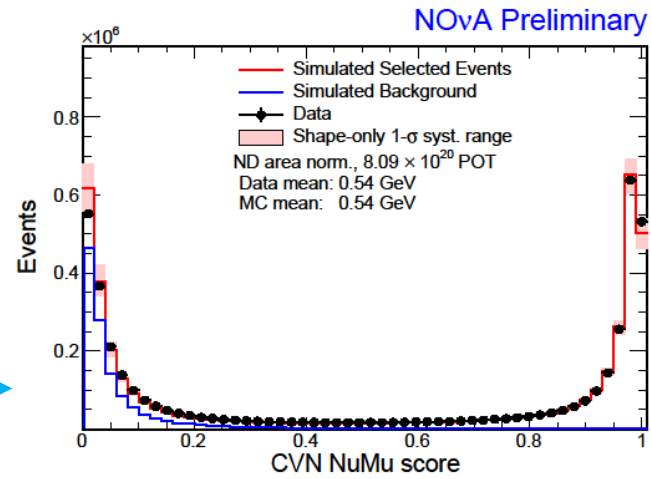


FD: $\sim 10^3$ events



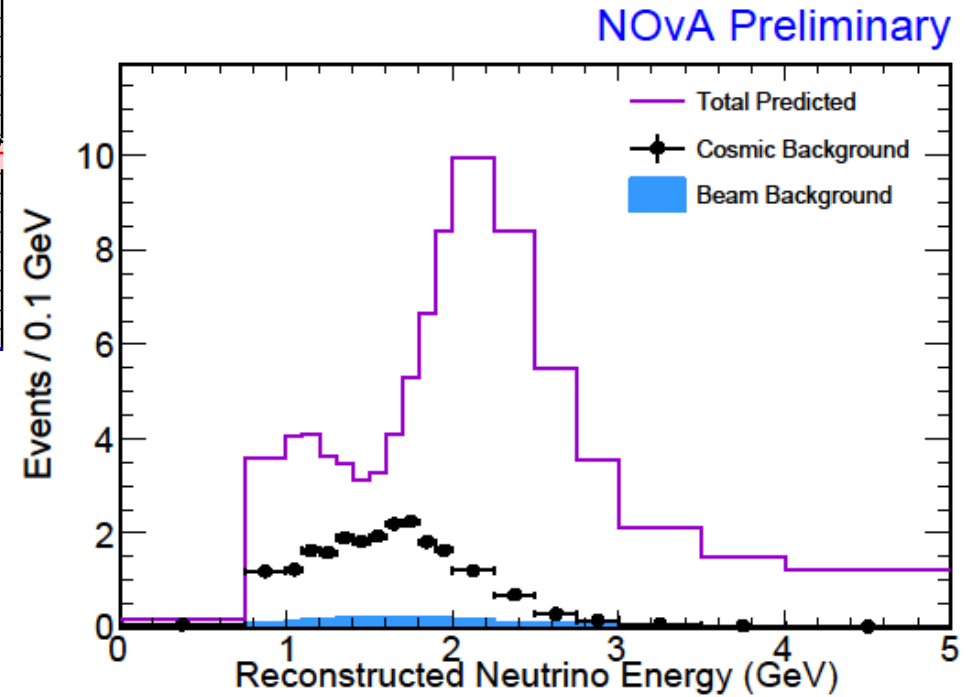


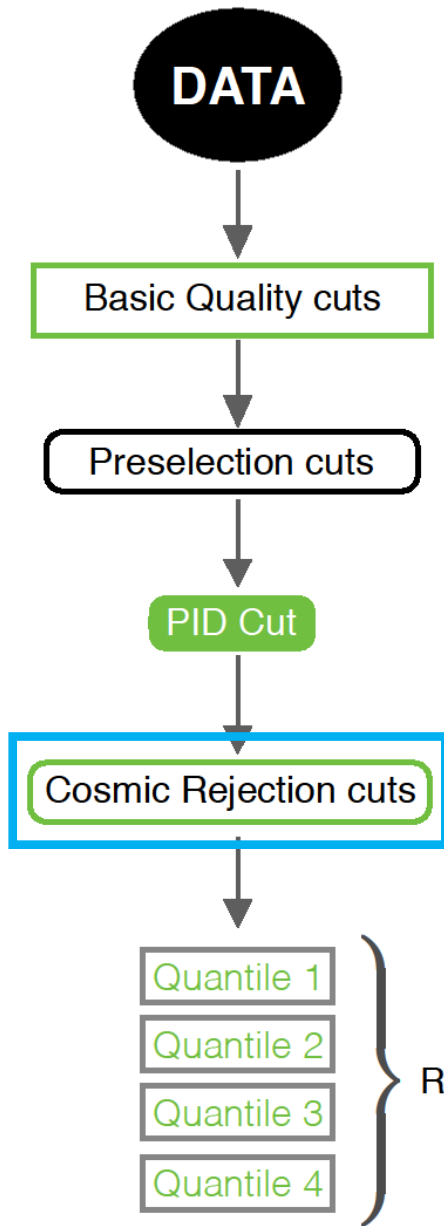
ν_μ selection



is it a ν_μ CC?
 deep learning: CVN
 removes NC, cosmics
 (ND + FD)

FD: ~ 100 events

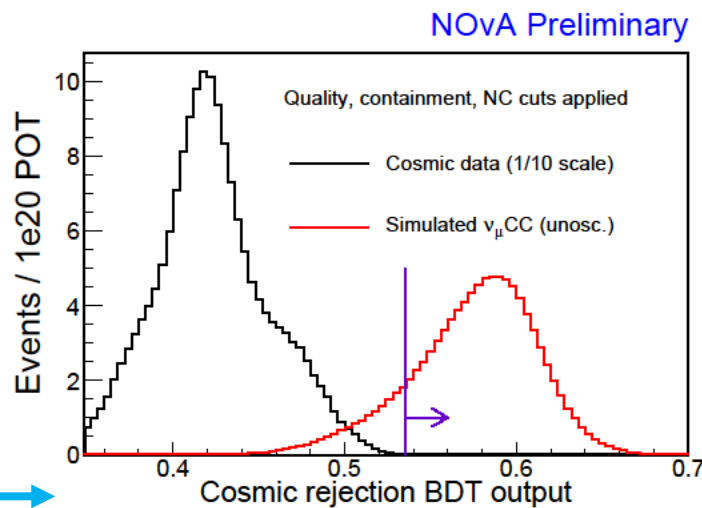




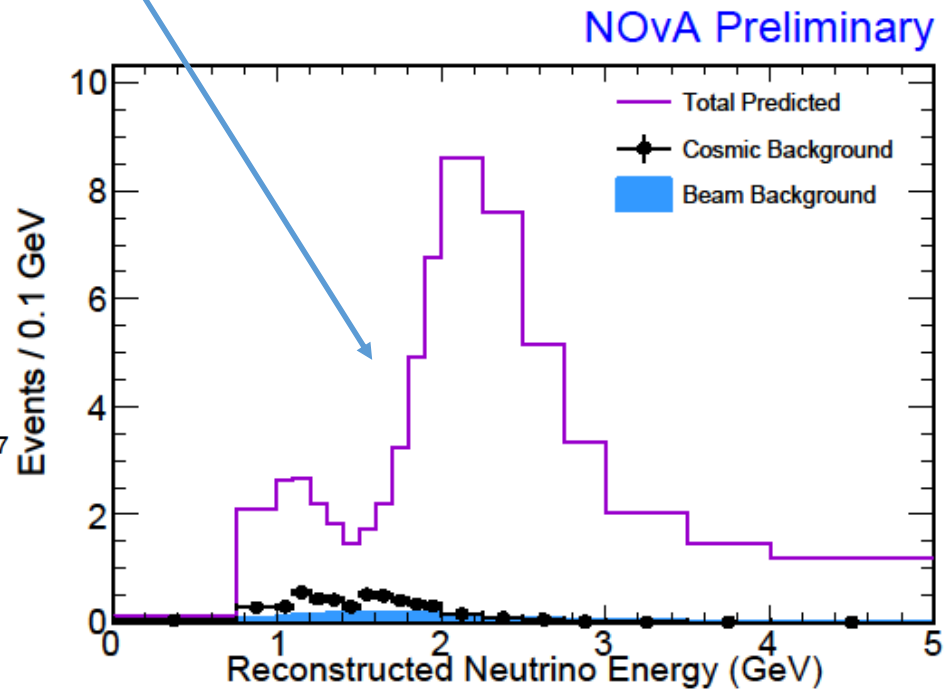
ν_μ selection

shape depends on
oscillation parameters
 θ_{23} and Δm^2_{32}

FD: ~100 events



BDT to remove
remaining cosmics
(FD only)



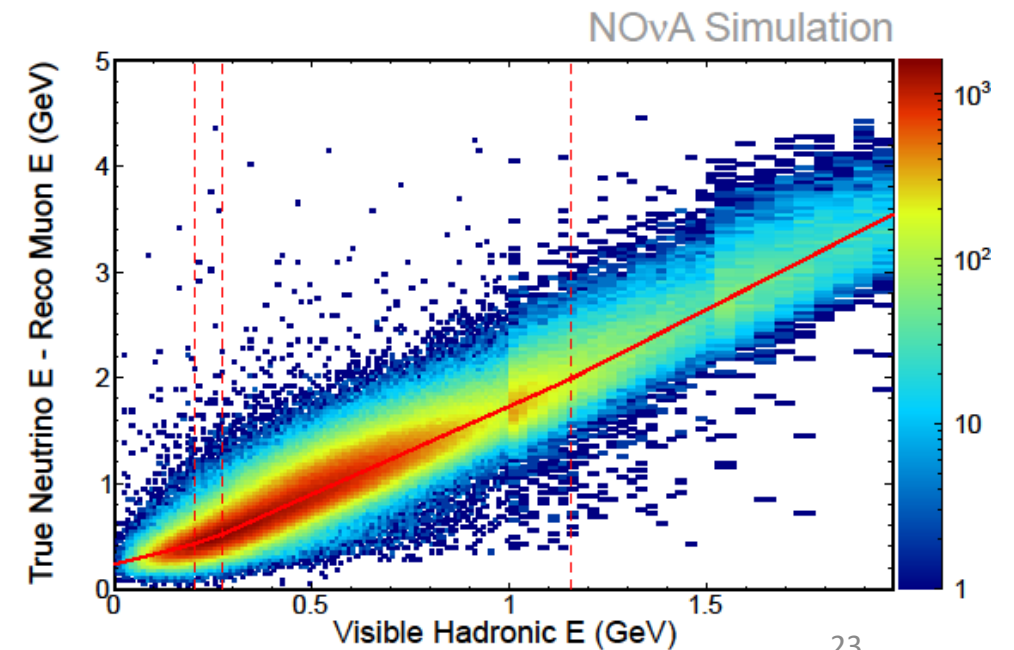
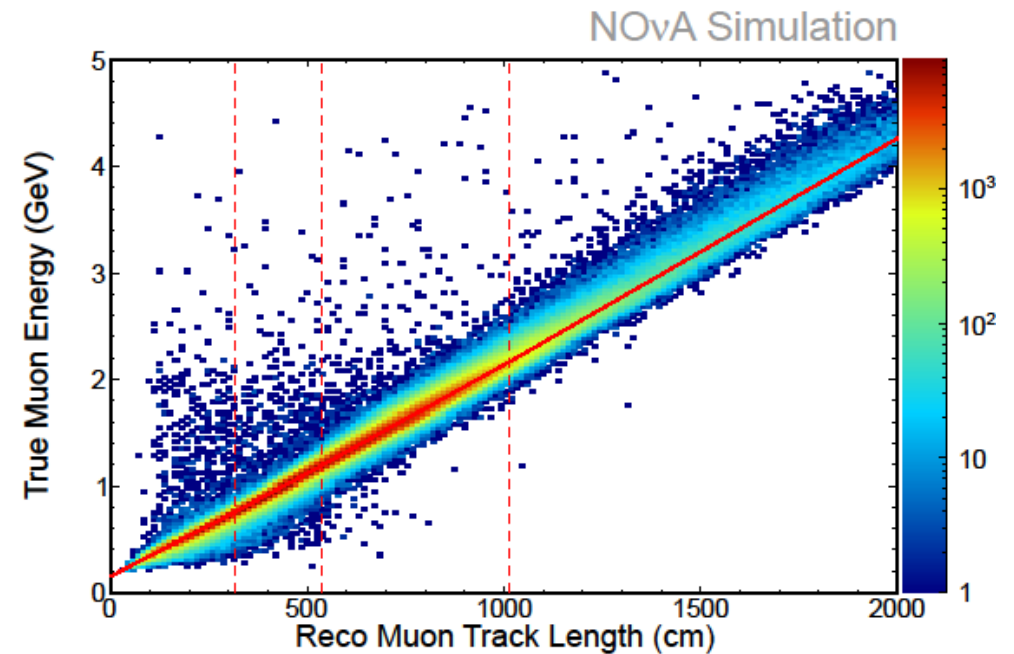
ν_μ energy reconstruction

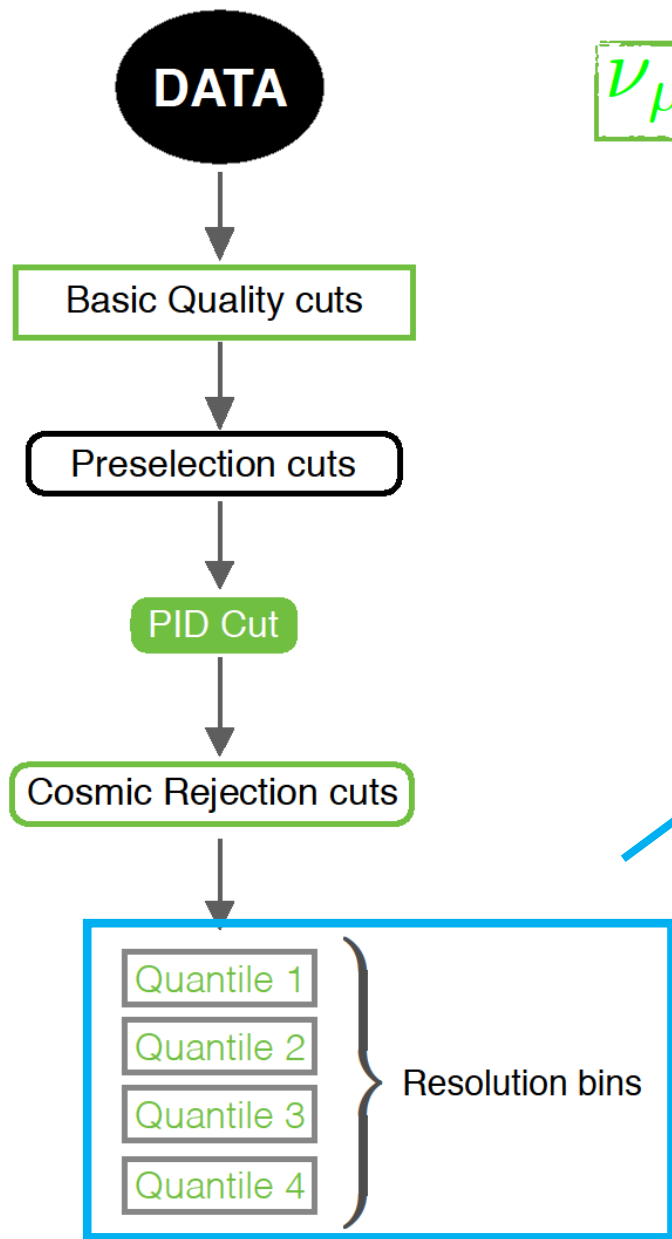
Reconstructed neutrino E:
based on simulation

lepton part ($E_{lep}^{res} = 3\%$)

hadronic part ($E_{had}^{res} = 30\%$)

$E_\nu = E_{lep} + E_{had}$ ($E_\nu^{res} = 9\%$)

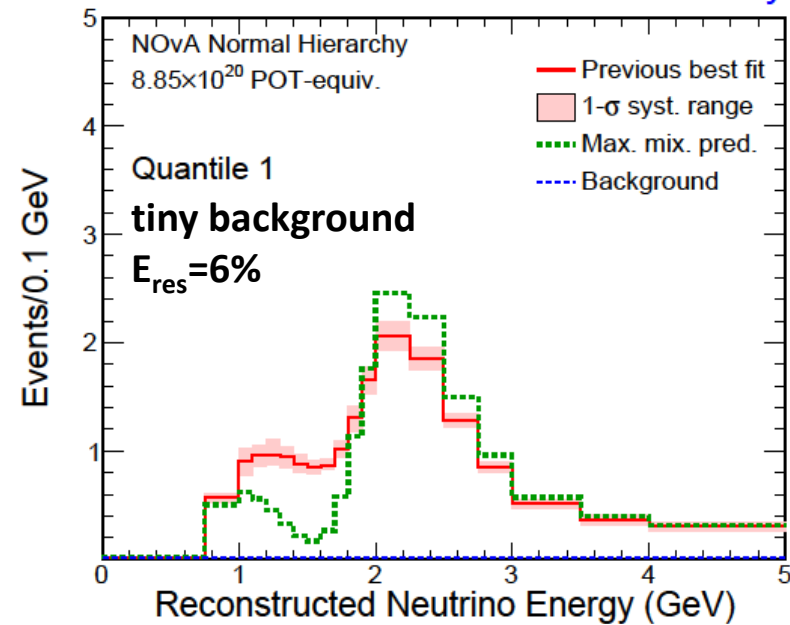




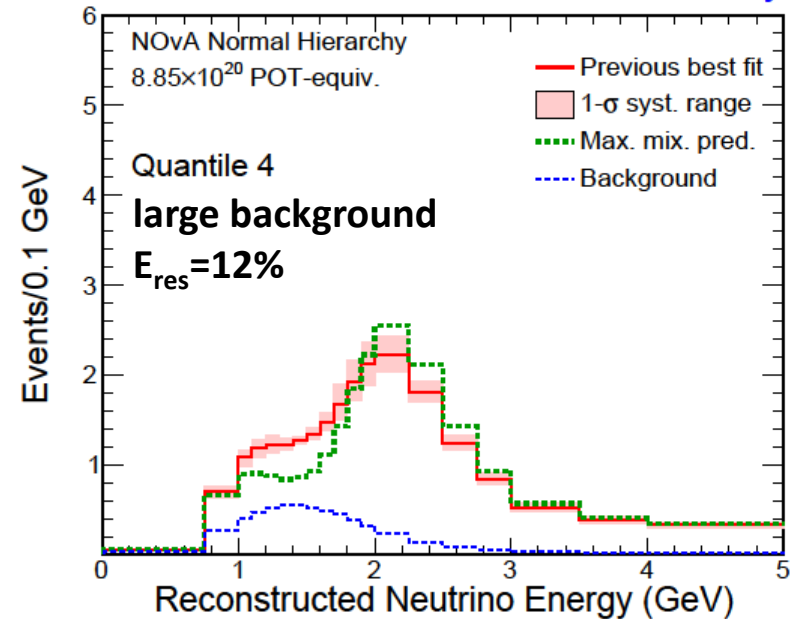
ν_μ selection

the less hadronic E the better the E resolution
 divide into 4 E_{had}/E_ν quantiles

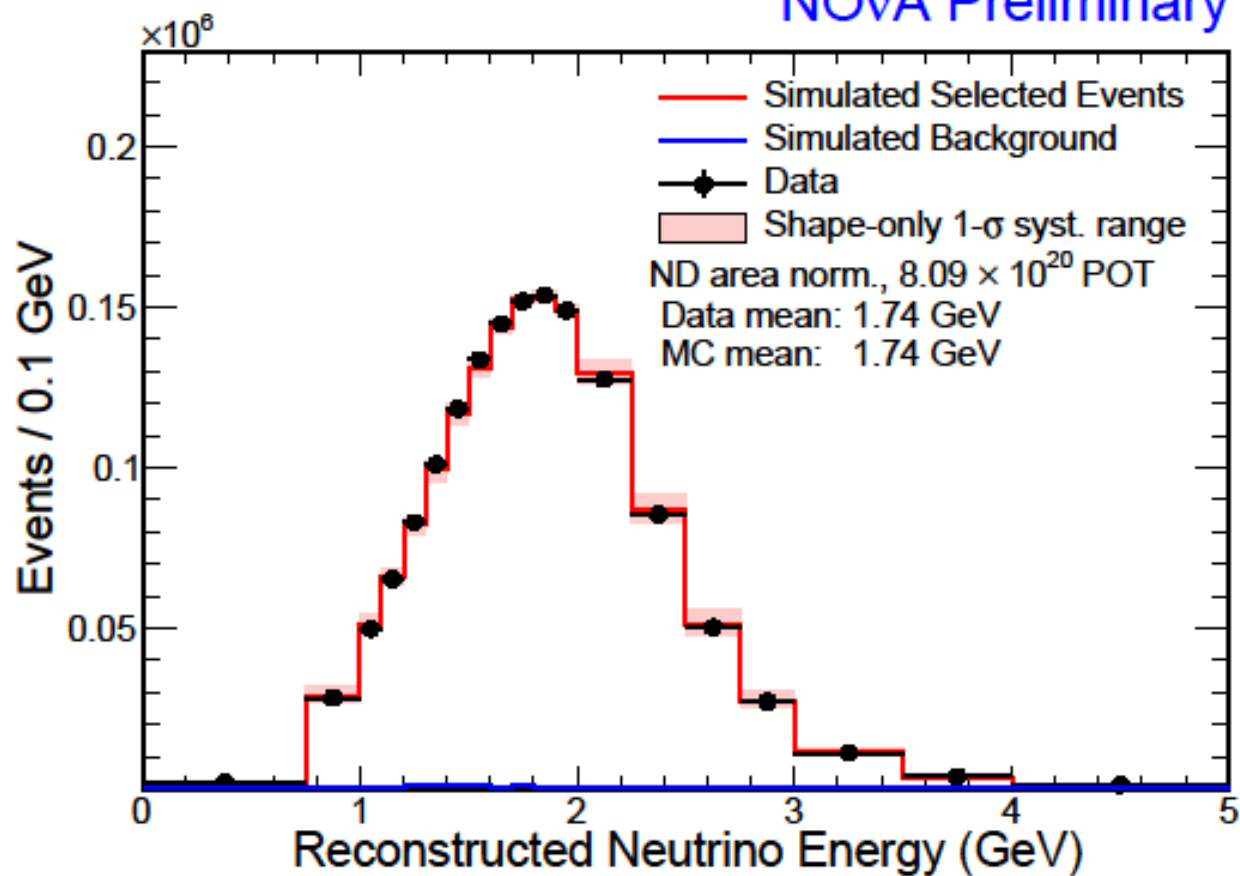
NOvA Preliminary



NOvA Preliminary



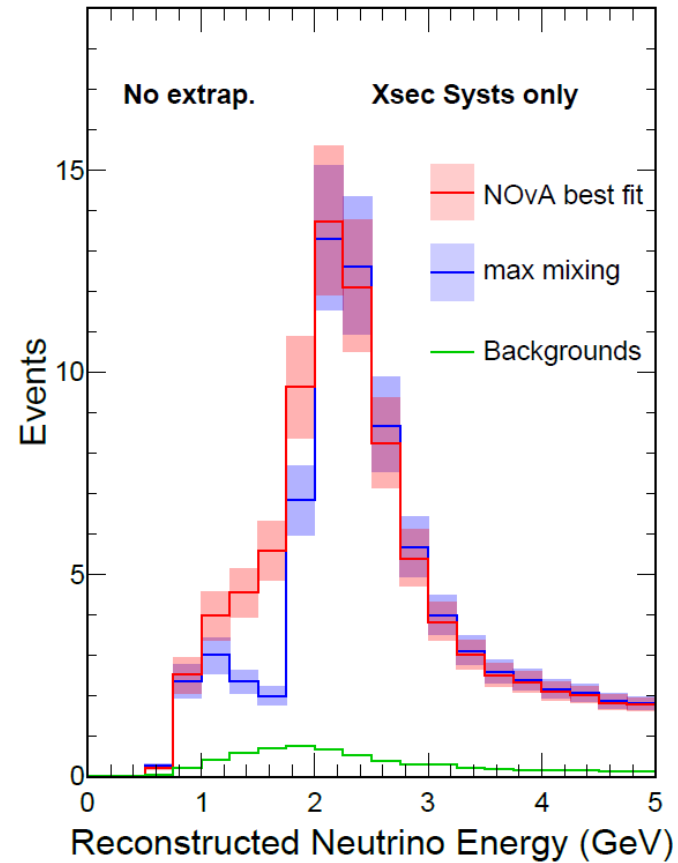
NOvA Preliminary



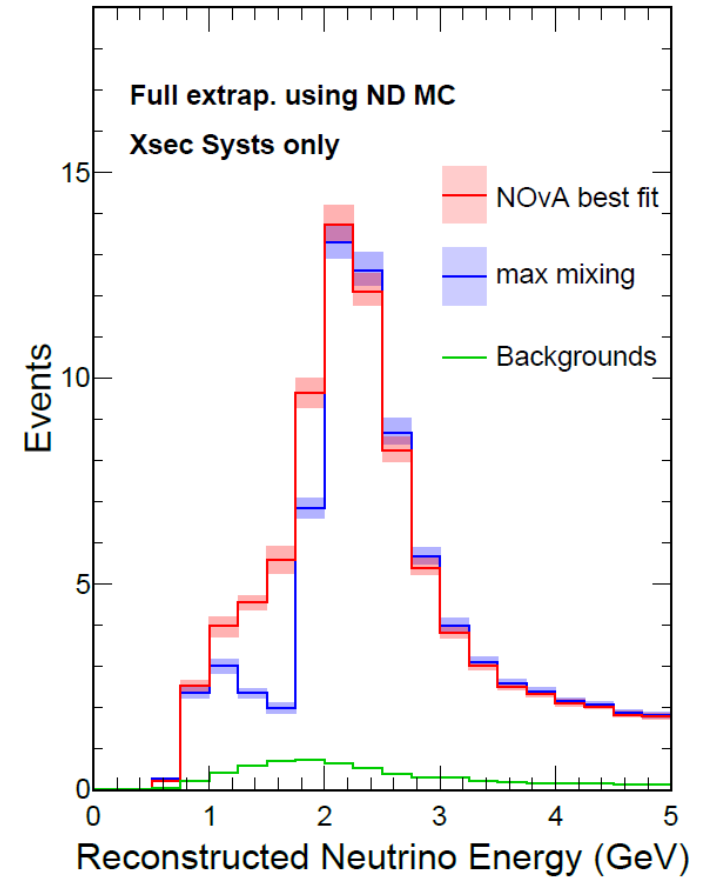
select ν_μ events in ND
data/MC agreement good
tells us simulation not too wrong

Use ND as a measurement to constrain uncertainties in the FD Far/near extrapolation

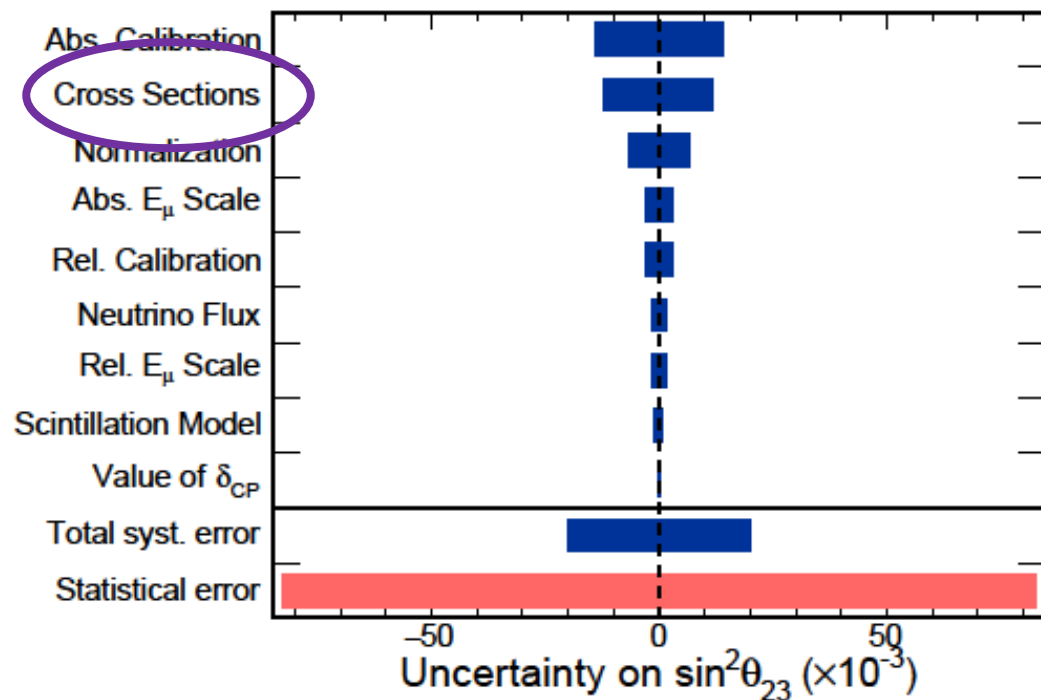
NOvA Simulation



NOvA Simulation

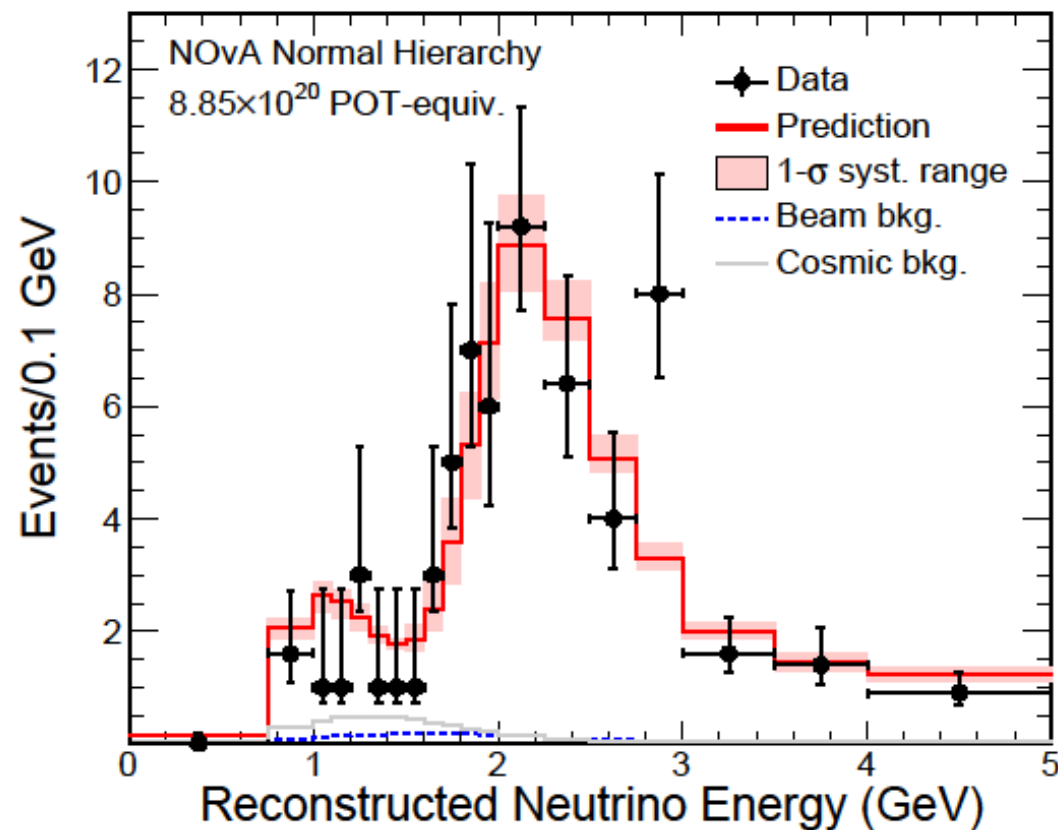


Each systematic is a penalty term in χ^2 fit to determine oscillation parameters



if no oscillations, predict 763 events
observe 126 events

NOvA Preliminary

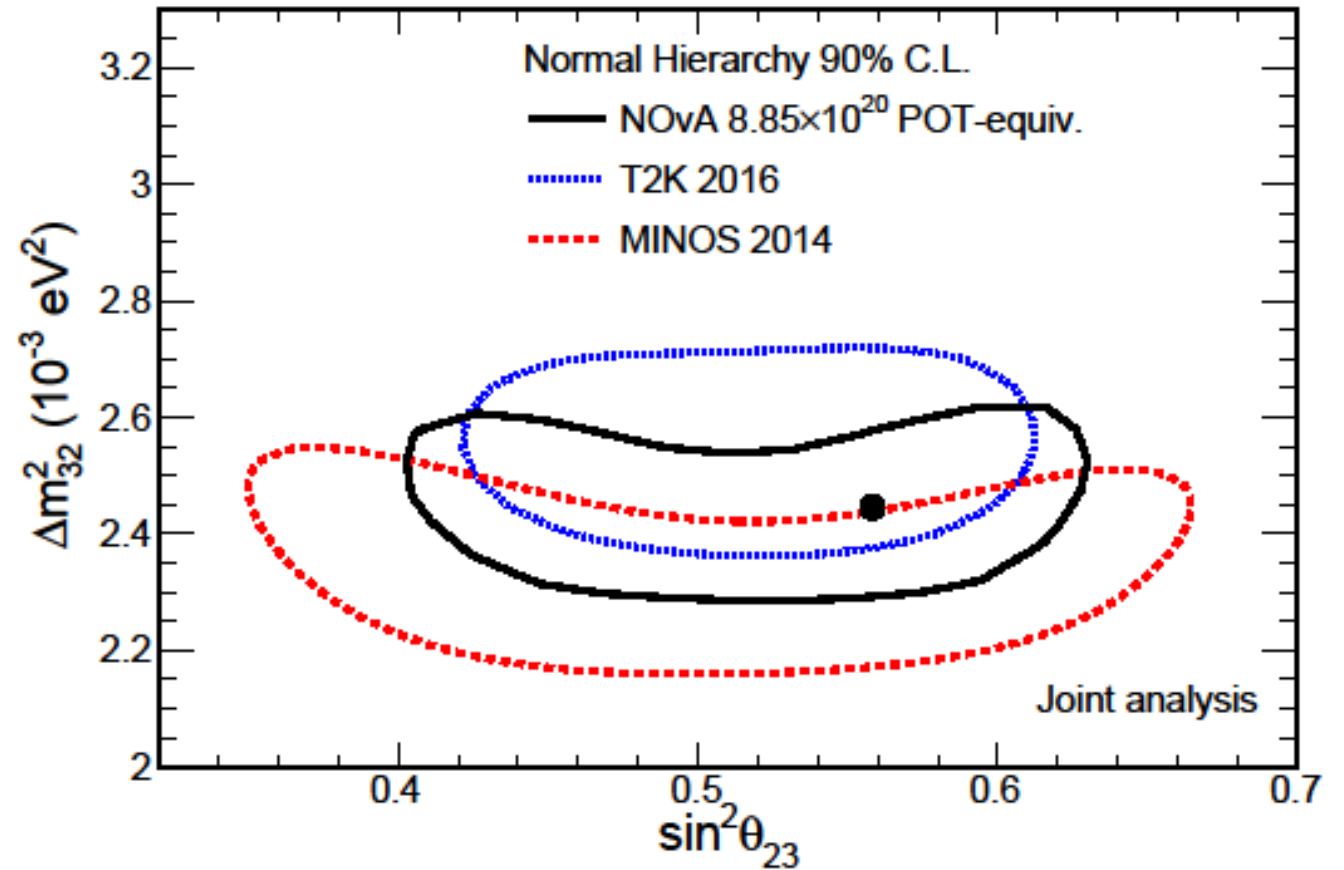


Can plot as 90% confidence contour in θ_{23} , Δm^2_{32} space

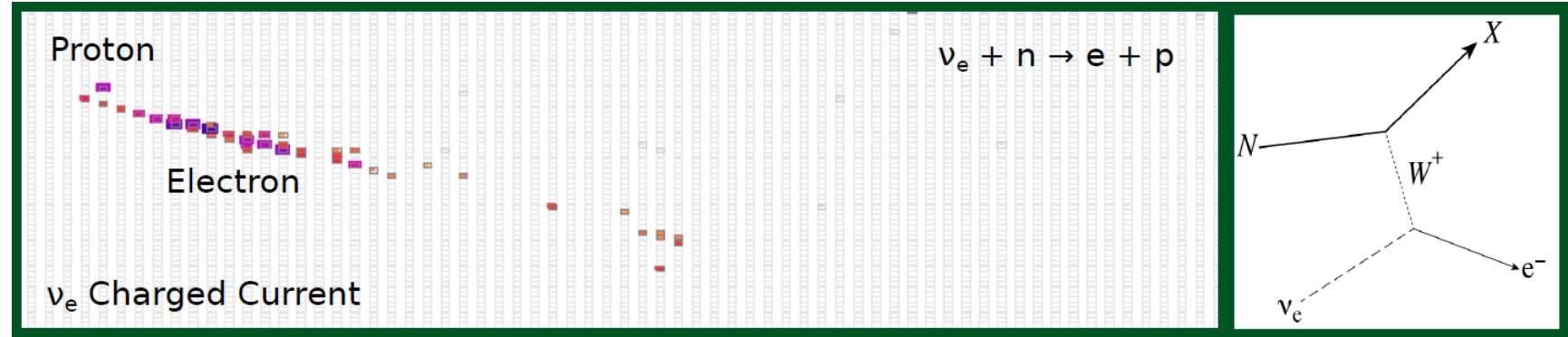
best fits:

$$\sin^2\theta_{23} = 0.558^{+0.041}_{-0.033}$$

$$\Delta m^2_{32} = 2.444^{+0.079}_{-0.077} \times 10^3 \text{ eV}^2$$



ν_e Appearance



$$P(\bar{\nu}_{\mu} \rightarrow \bar{\nu}_{e}) \approx \sin^2 2\theta_{13} \sin^2 \theta_{23} \frac{\sin^2(A-1)\Delta}{(A-1)^2}$$

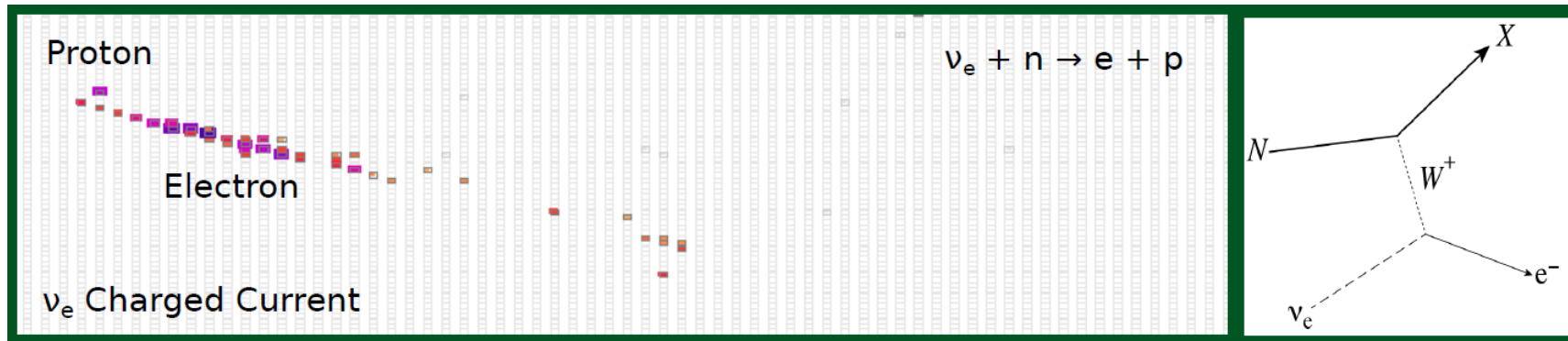
$$- 2\alpha \sin \theta_{13} \sin \delta_{CP} \sin 2\theta_{12} \sin 2\theta_{23} \frac{\sin A\Delta}{A} \frac{\sin(A-1)\Delta}{A-1} \sin \Delta$$

$$+ 2\alpha \sin \theta_{13} \cos \delta_{CP} \sin 2\theta_{12} \sin 2\theta_{23} \frac{\sin A\Delta}{A} \frac{\sin(A-1)\Delta}{A-1} \cos \Delta$$

$$\text{Where: } \alpha = \frac{\Delta m_{21}^2}{\Delta m_{31}^2} \quad \Delta = \Delta m_{31}^2 \frac{L}{4E} \quad A = \frac{(-)}{+} G_f N_e \frac{L}{\sqrt{2}\Delta}$$

directly measure δ_{CP} , the **mass hierarchy**, additional information for θ_{23}

ν_e Appearance

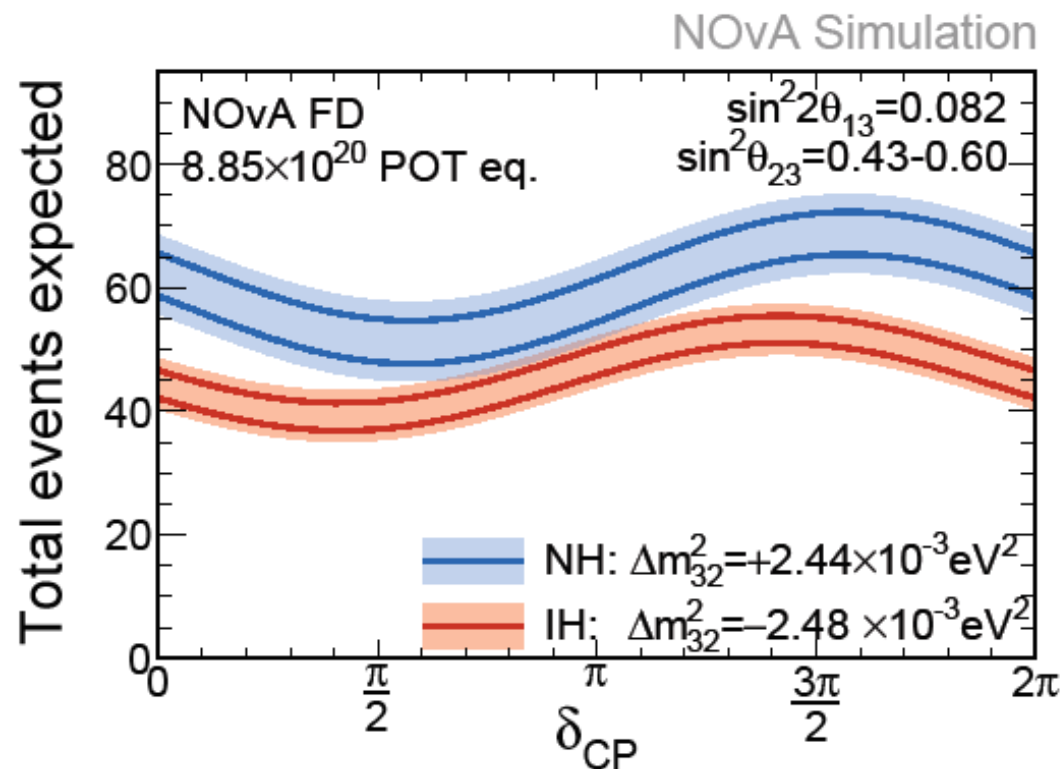


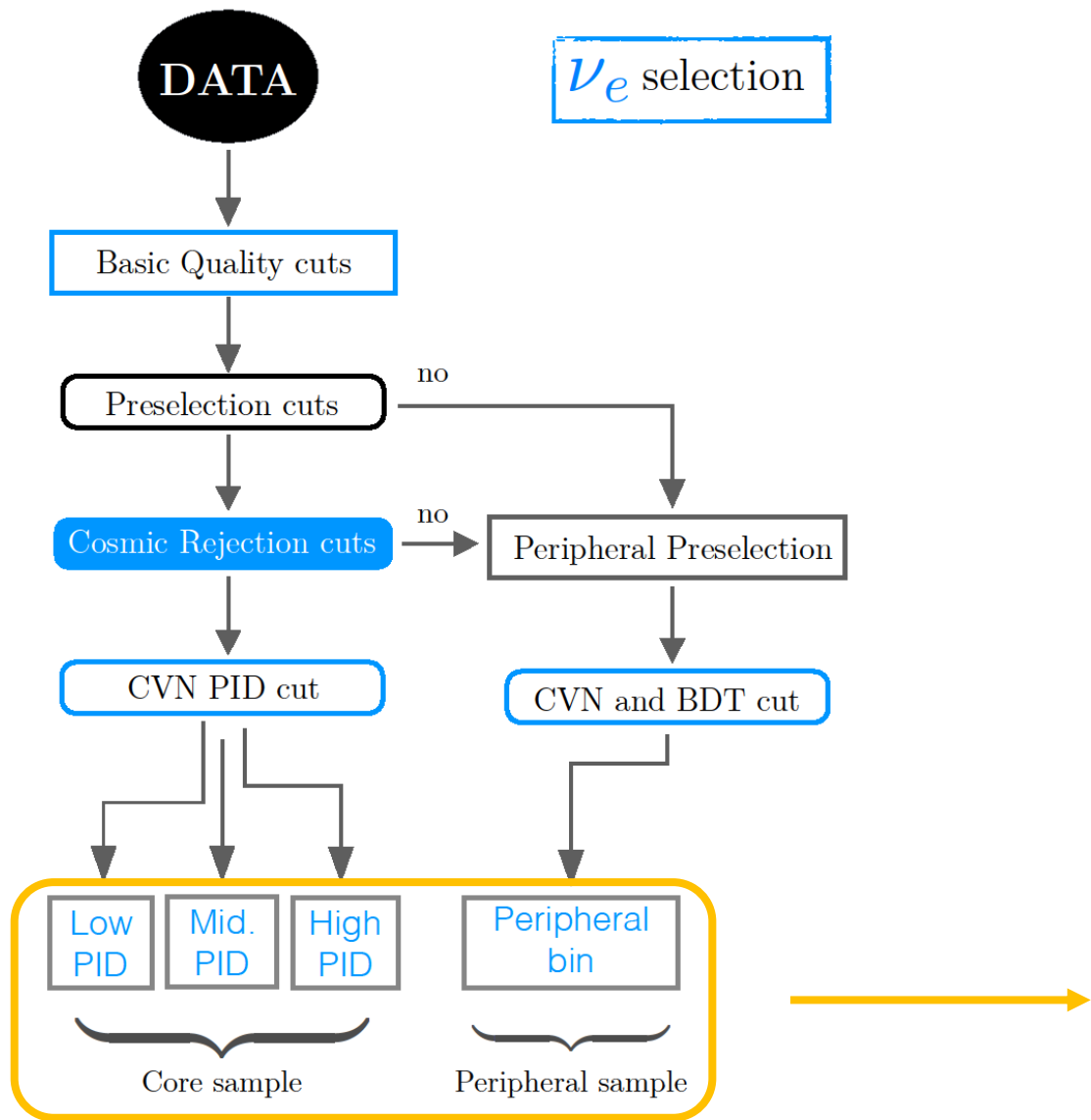
Signature is an electron shower

Backgrounds:

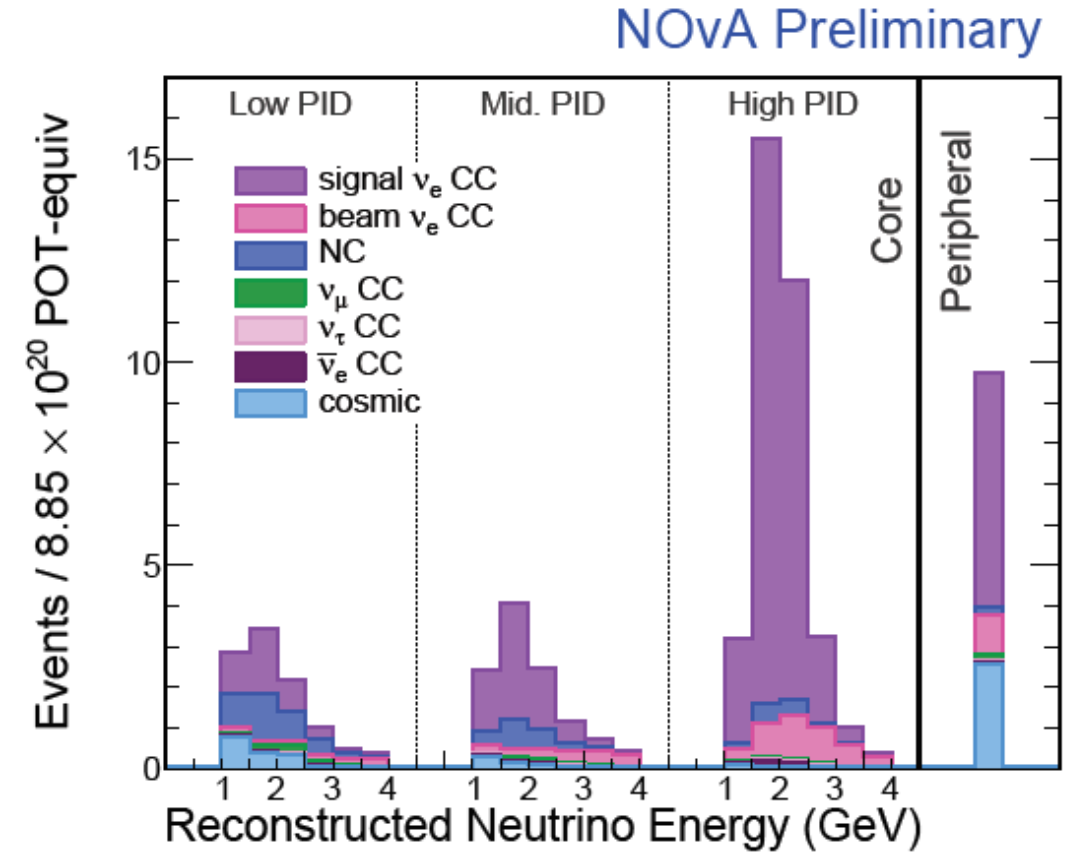
- cosmics
- neutral current
- intrinsic beam ν_e

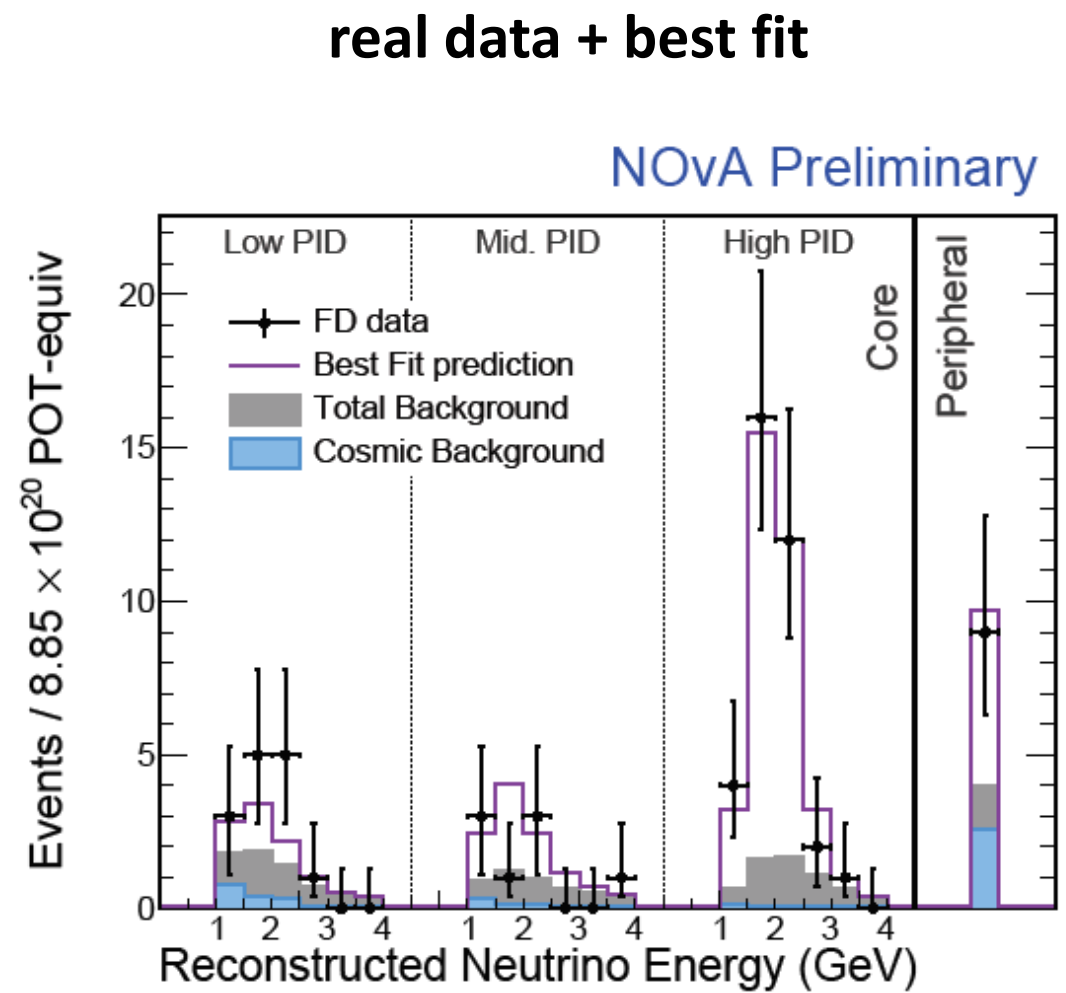
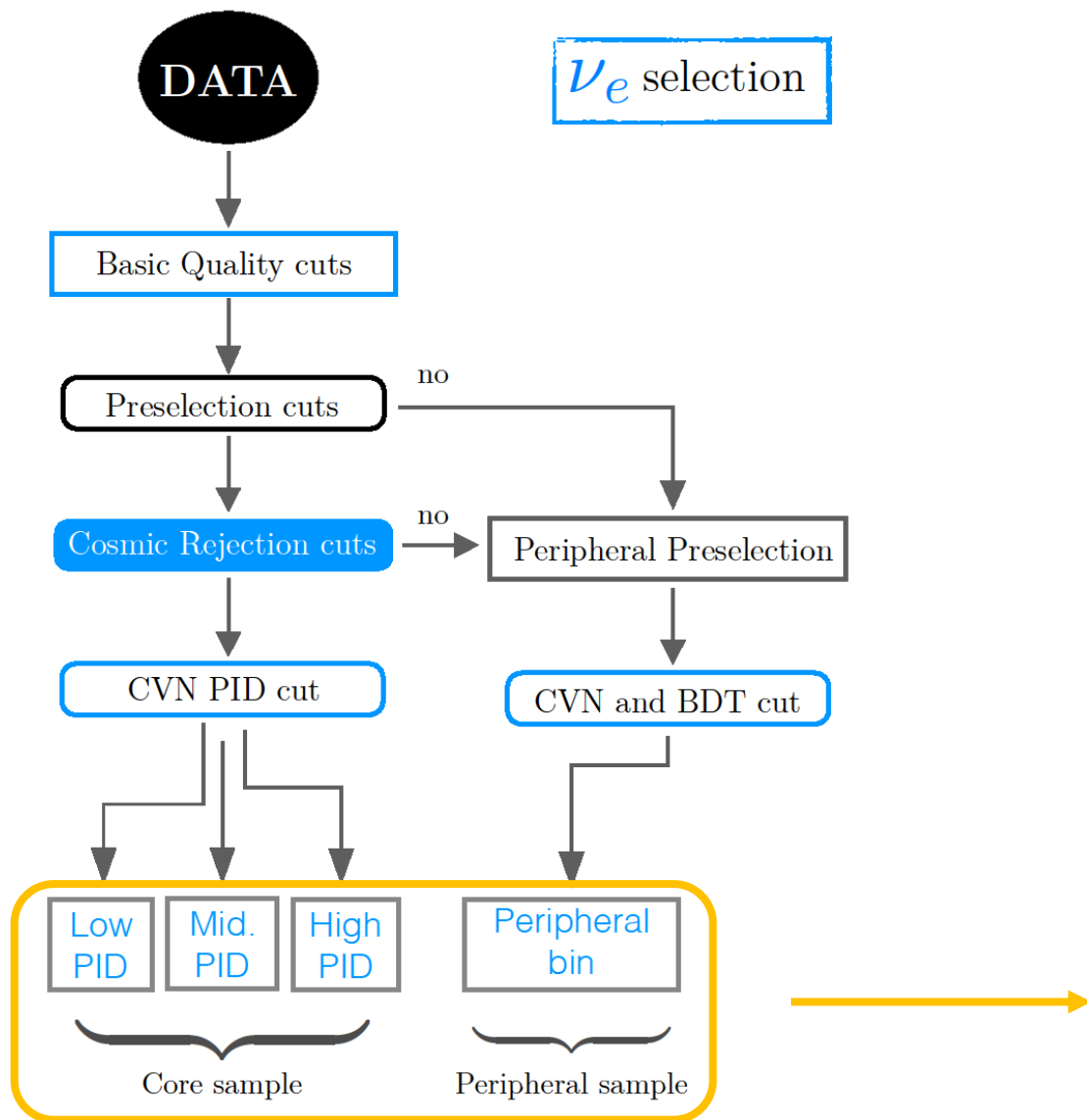
Use ND to measure intrinsic ν_e and neutral current background, extrapolate to FD



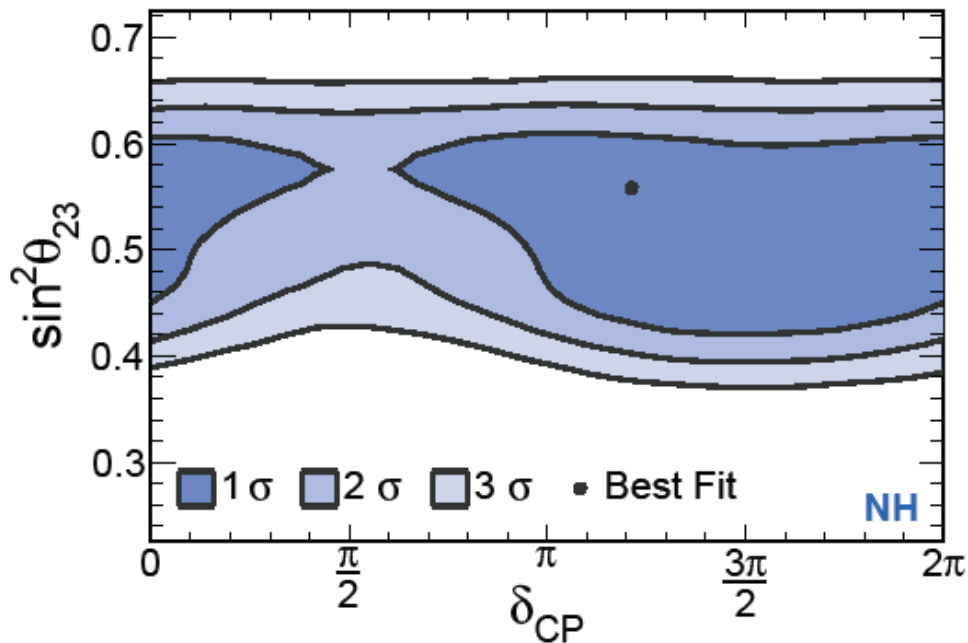


bin in energy for three PID bins + sideband

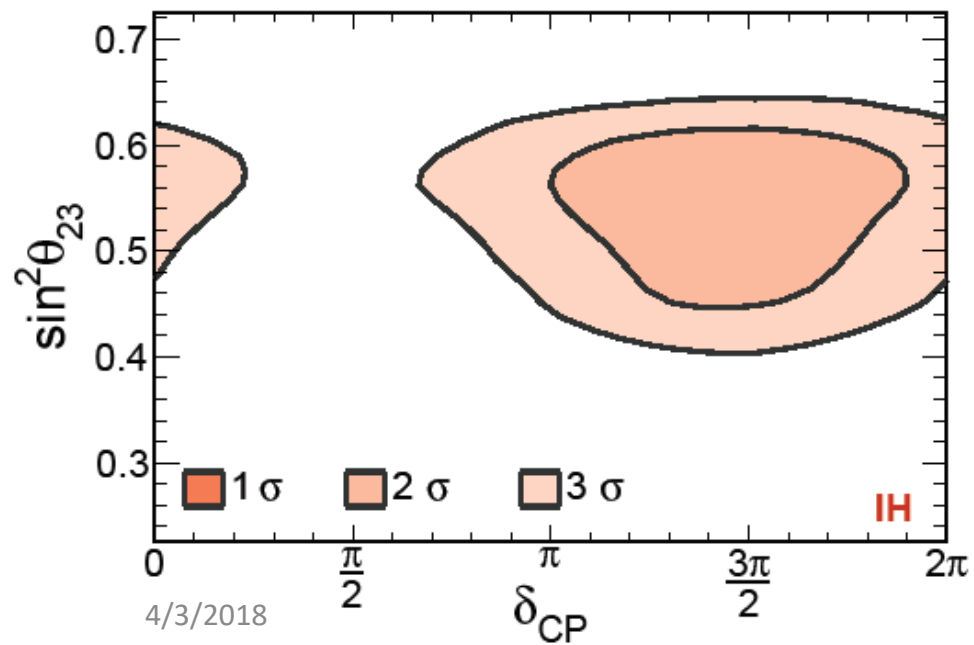




NOvA Preliminary

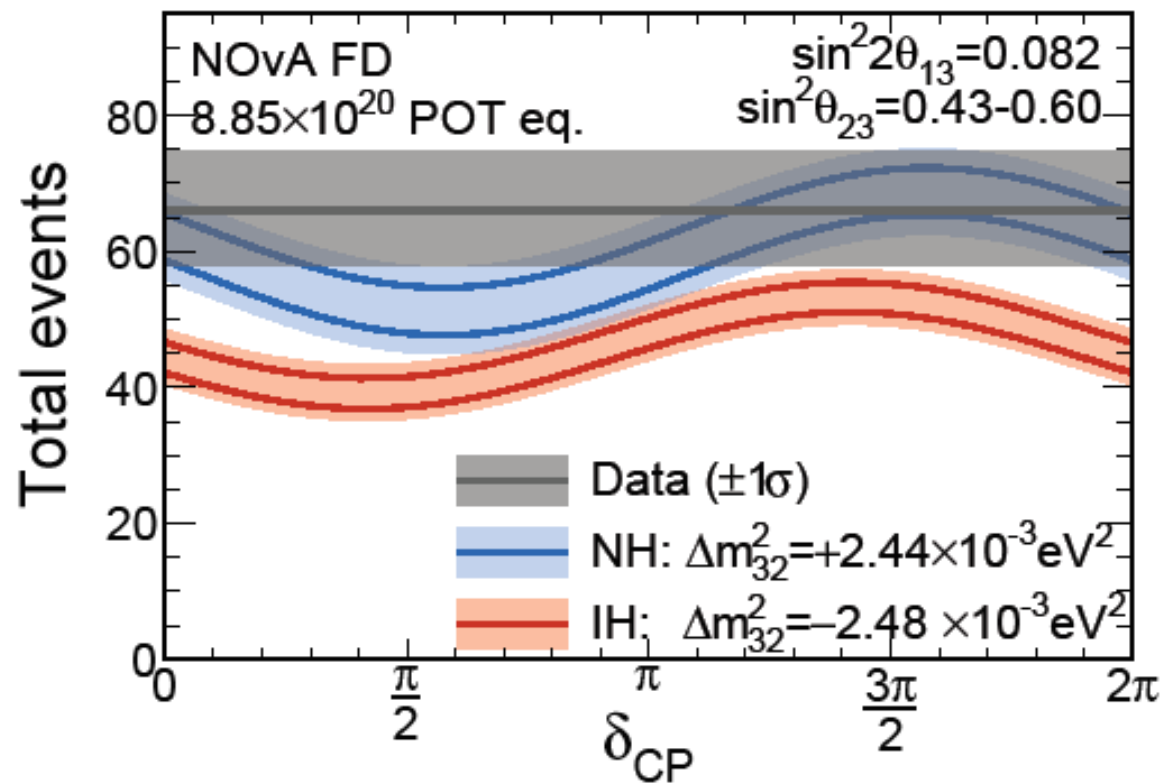


NOvA Preliminary

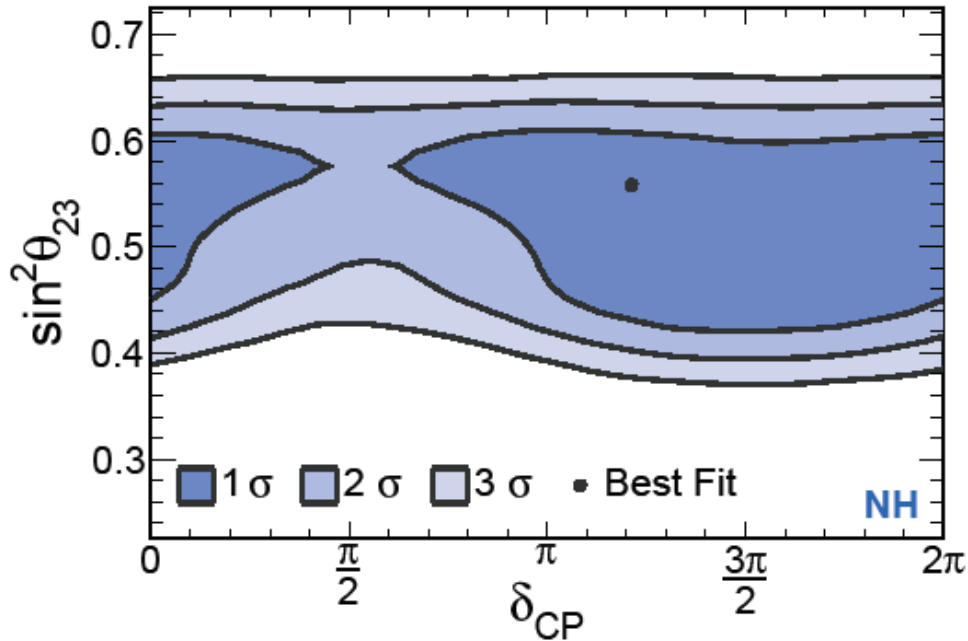


Contours in θ_{23}, δ_{CP} space
Depends a lot on mass hierarchy

NOvA Preliminary



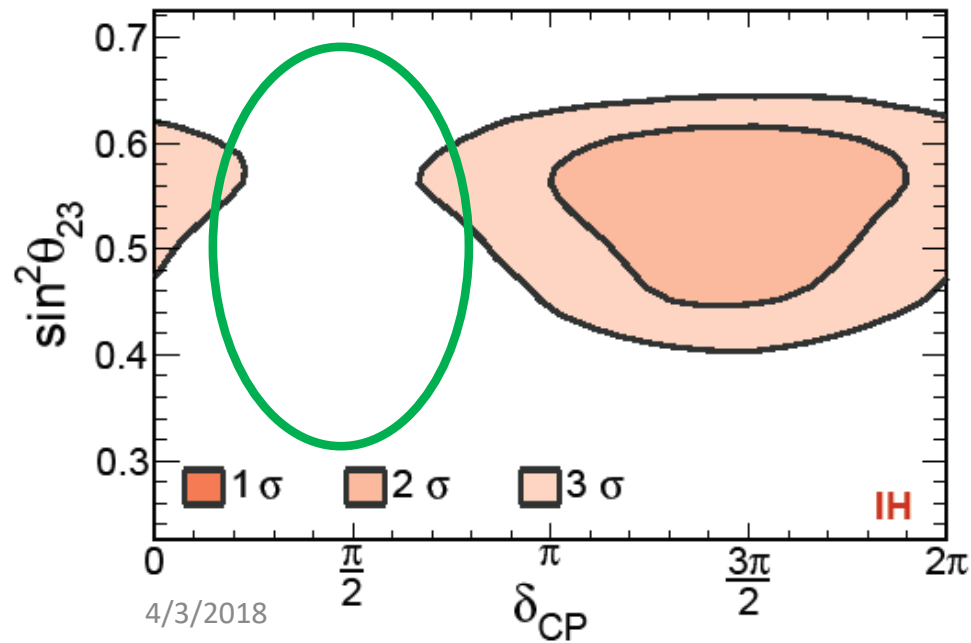
NOvA Preliminary



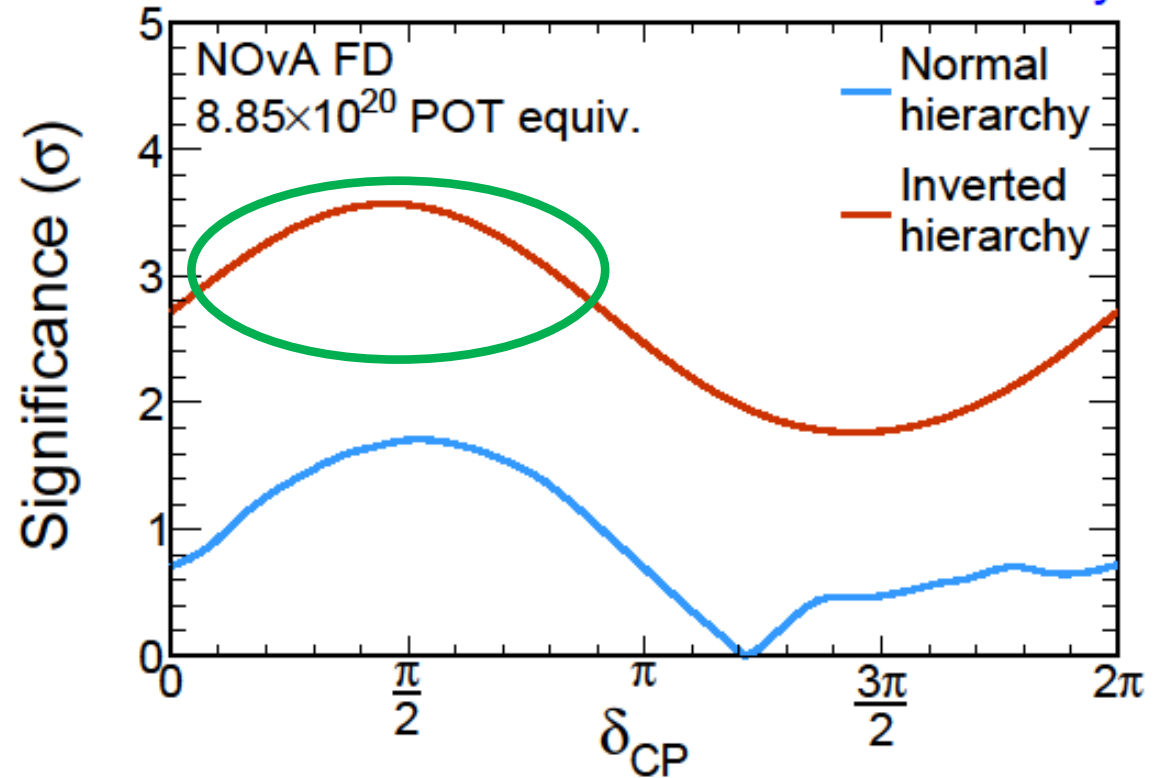
Contours in θ_{23}, δ_{CP} space
Depends a lot on mass hierarchy

For IH, δ_{CP} 0- π is disfavored by $> 3 \sigma$!

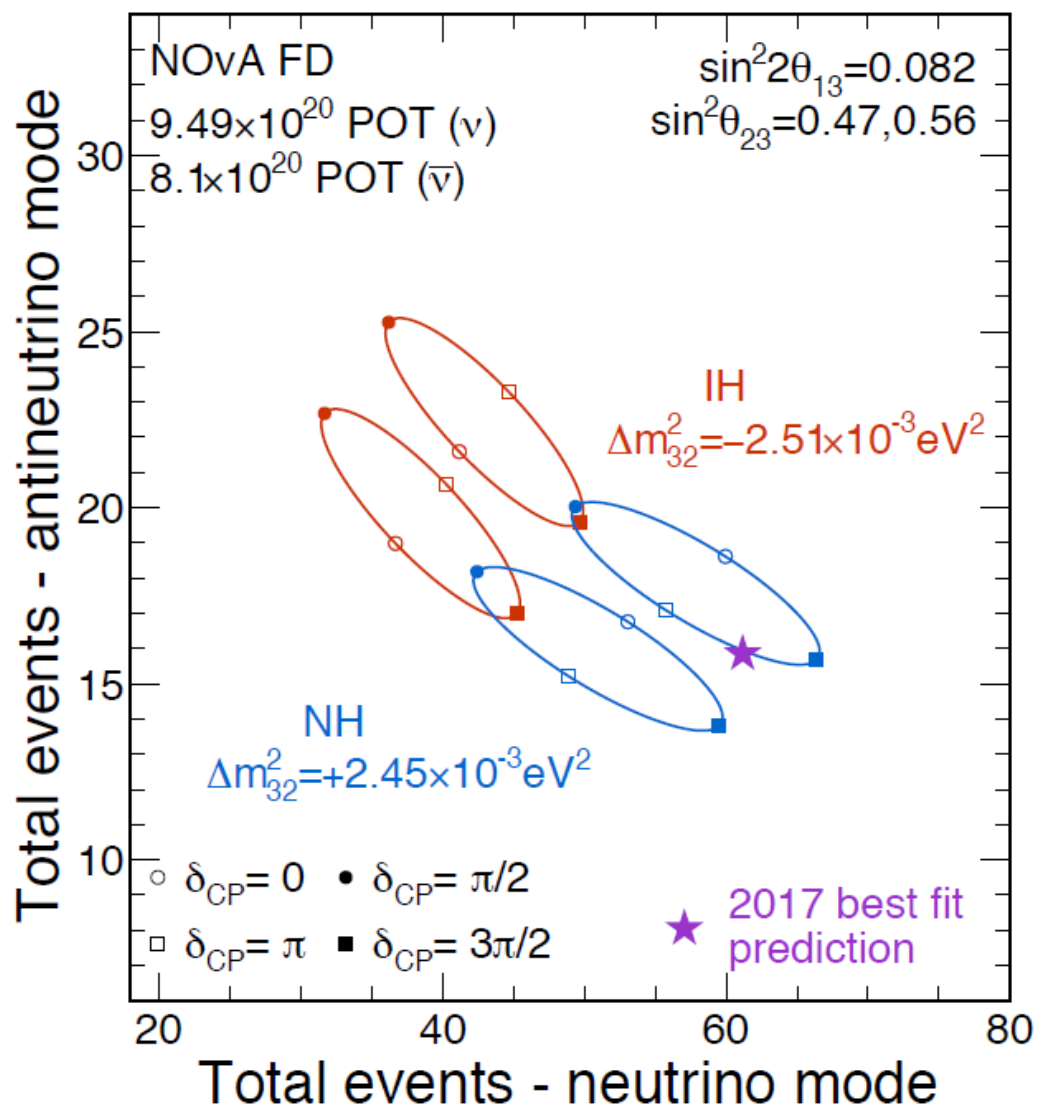
NOvA Preliminary



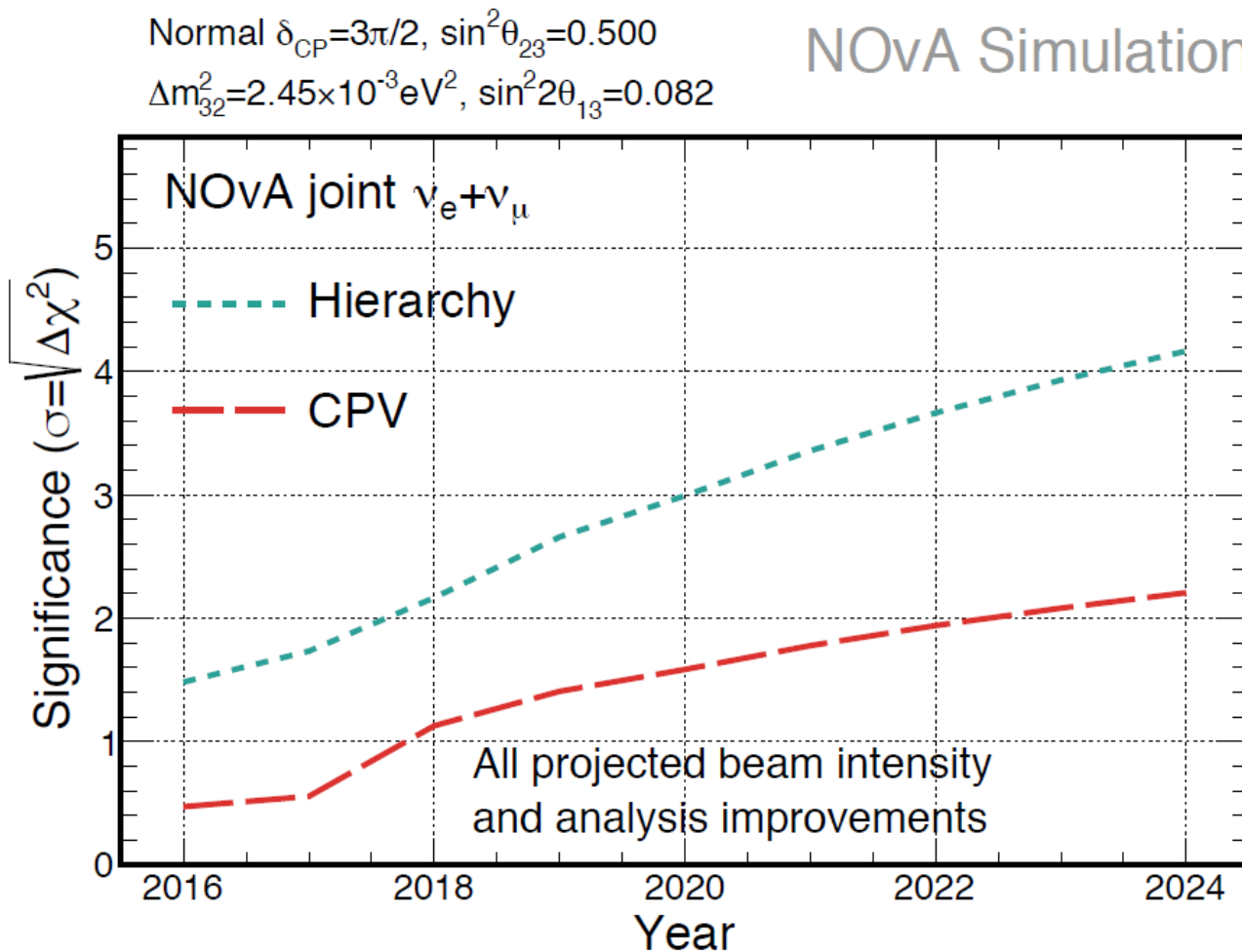
NOvA Preliminary



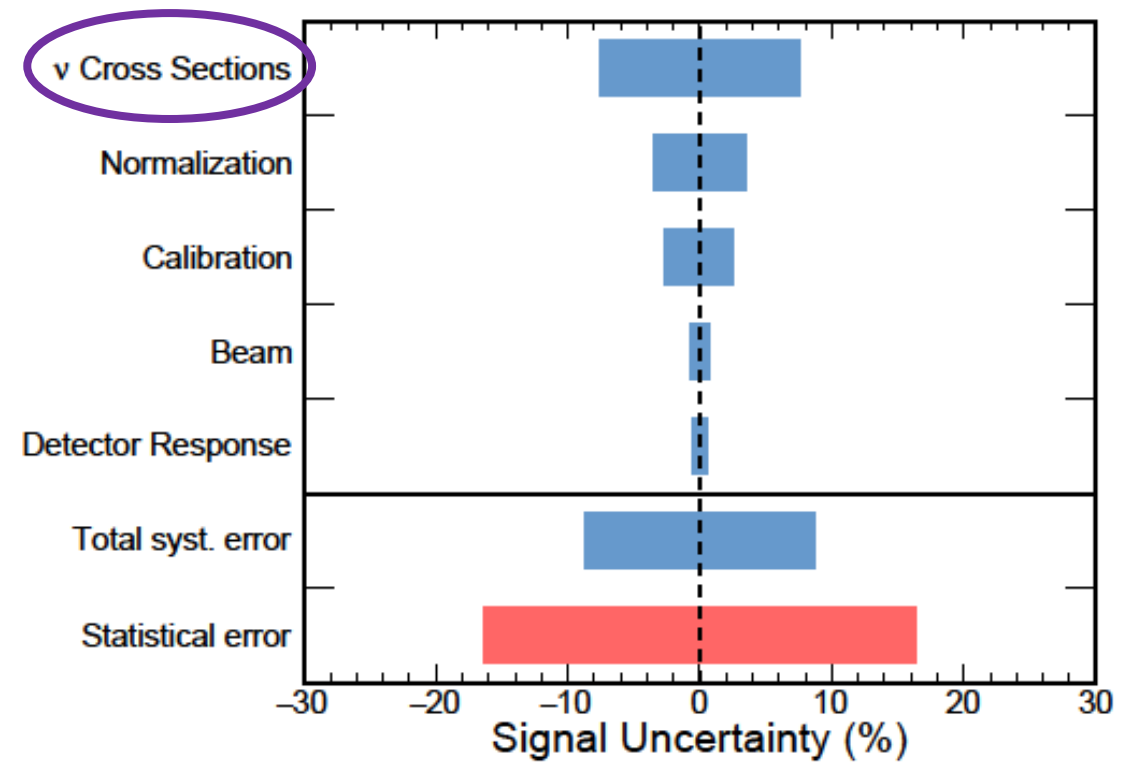
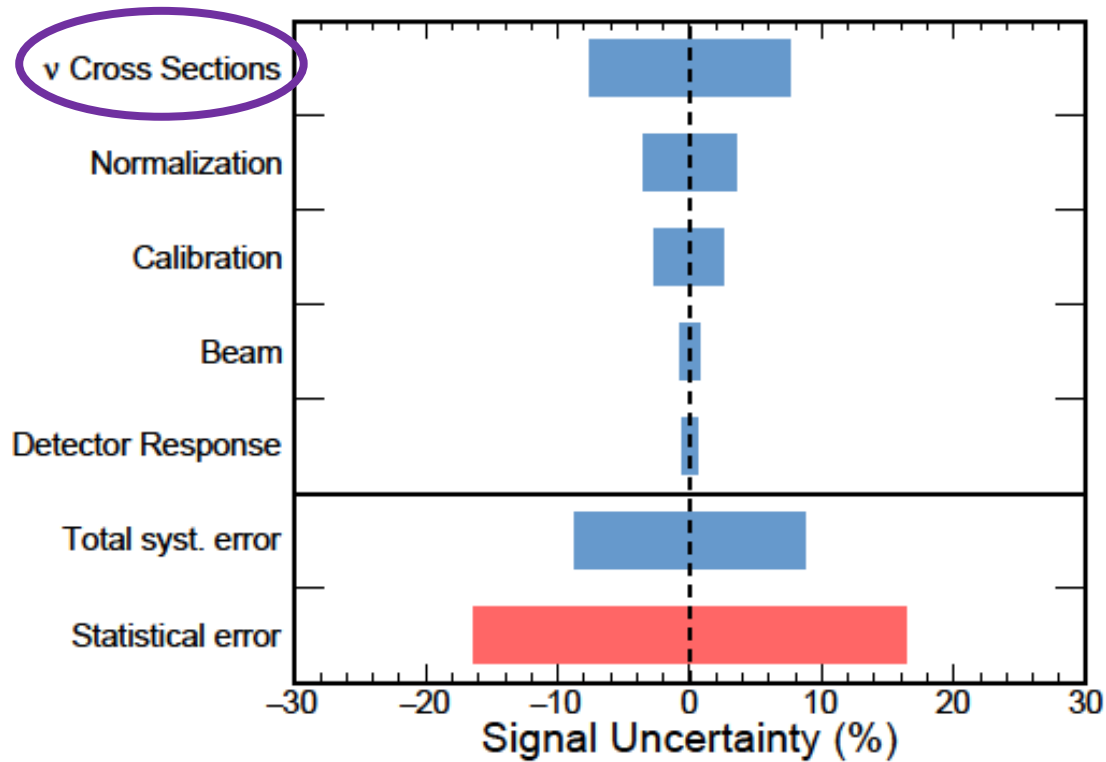
NOvA Simulation



NOvA Simulation



biggest uncertainty for appearance is cross-sections
still statistics limited but getting close
second biggest uncertainty for disappearance



cross-sections are hard! Lots of uncertainty

three standard interaction types:

Quasi-**E**lastic: just lepton and proton

Resonance: hadronic system is resonance (ie delta) which decays

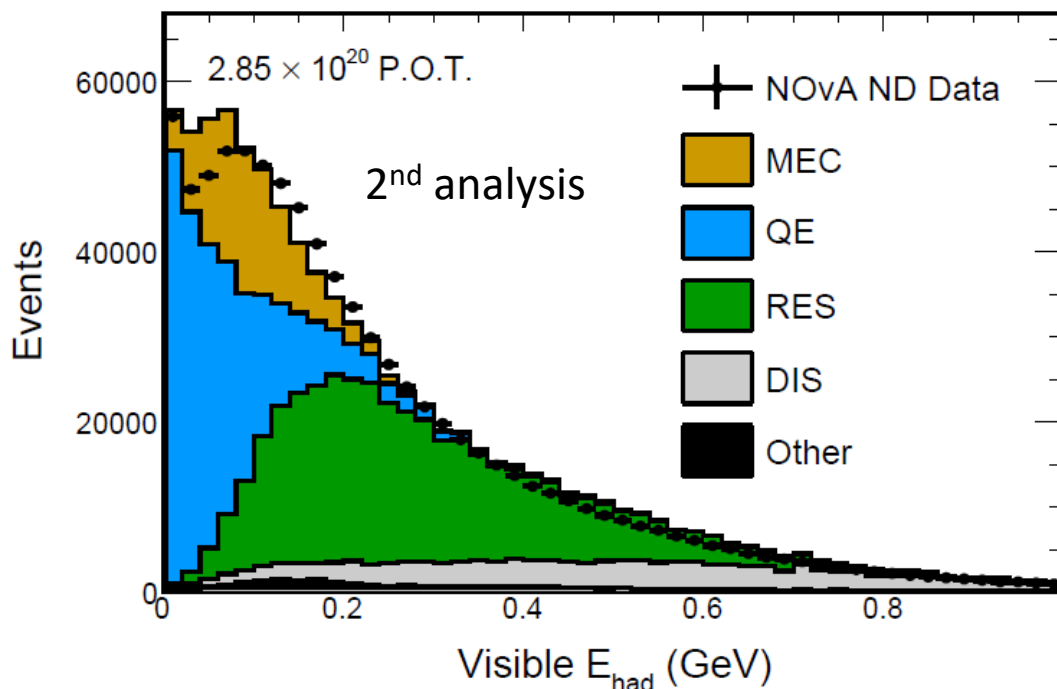
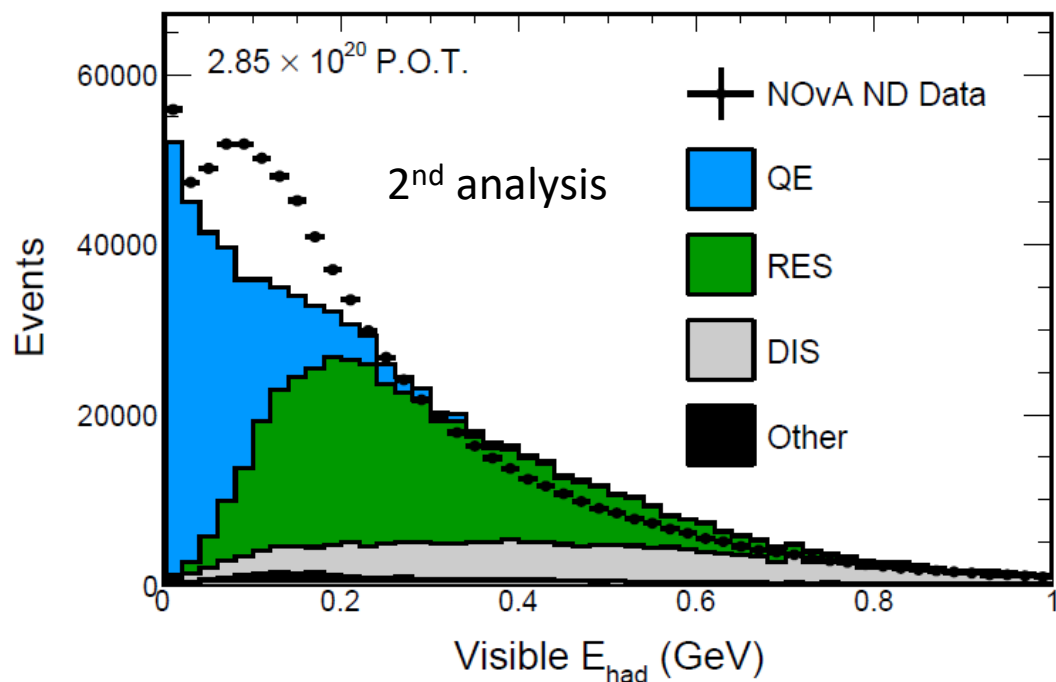
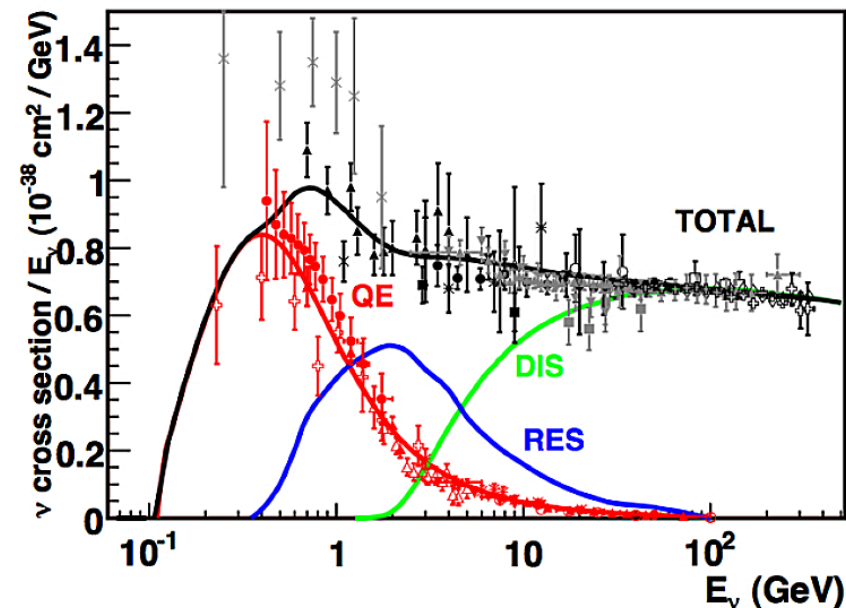
Deep **I**nelastic **S**cattering: neutrino hits quark directly

Meson **E**xchange **C**urrents: only recently discovered

still tons of uncertainty to rate and shape

NOvA is measuring this! Important for not just NOvA!

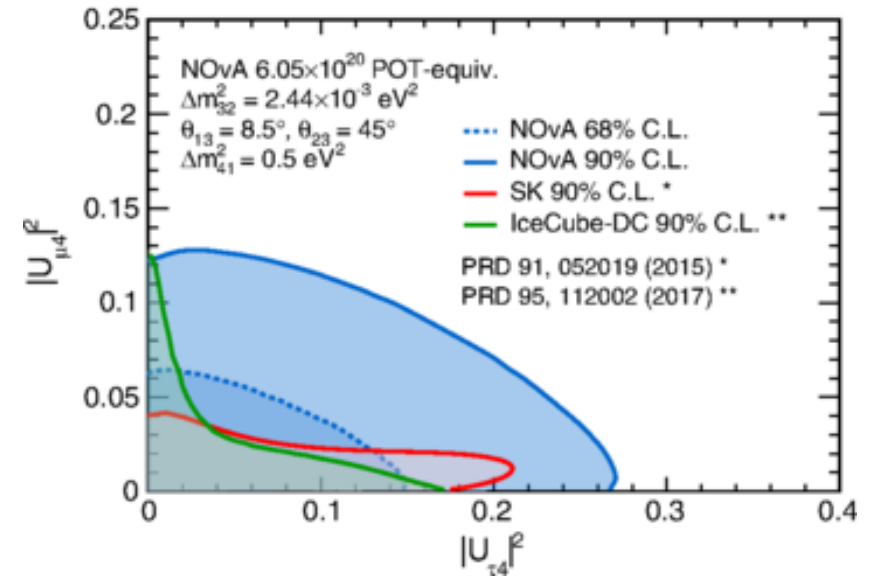
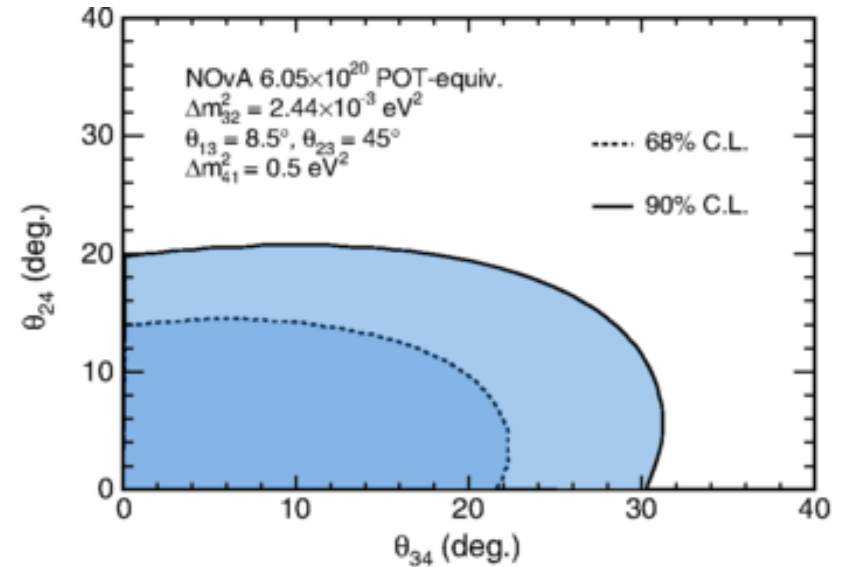
DUNE: current x-sec systs: ~10% needed: ~1%



Other NOvA physics

- Sterile neutrinos: Phys. Rev. D **96**, 072006 →
- CVN classifier: JINST 11 (2016) no. 09, P09001
- Direct x-section measurements
 - NC coherent π^0
 - ν_μ CC π^0 inclusive
- Dark matter searches
- Magnetic monopoles
- Supernova physics
- ... and much more

in paper draft

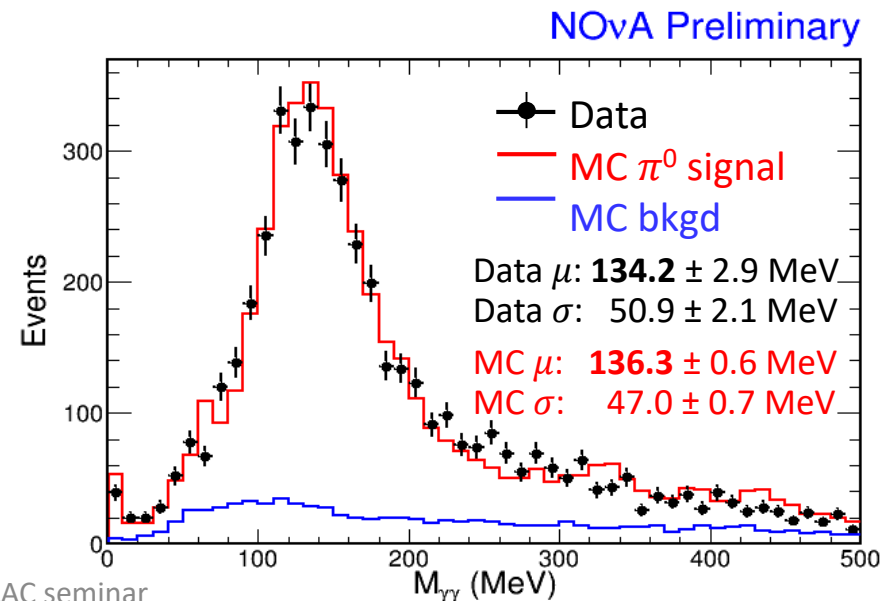
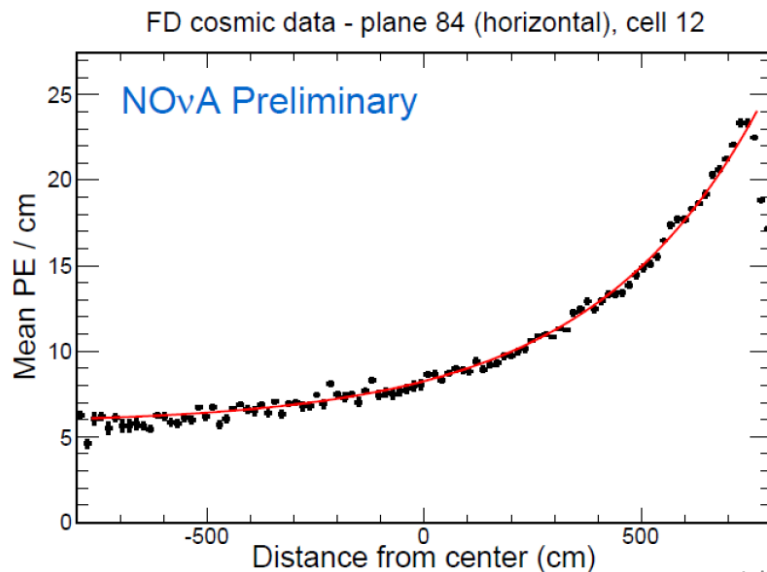




Backups

NOvA Calibration:

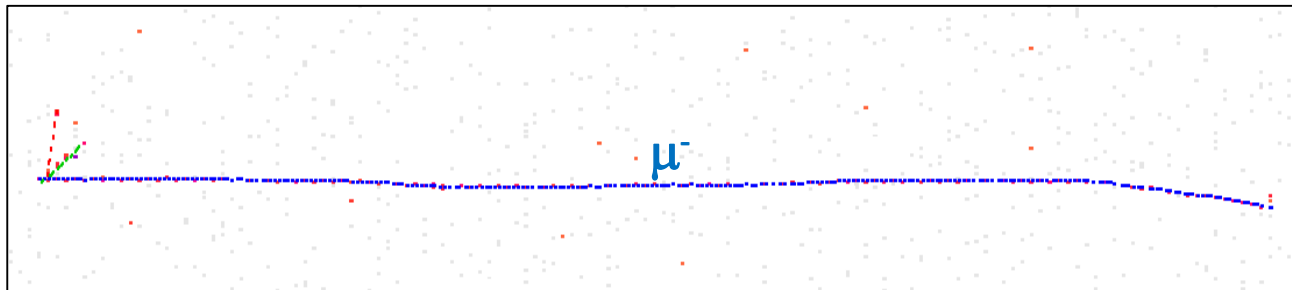
- Critical for any detector. Very briefly:
- Absolute energy scale is calibrated with stopping muons (dE/dx , Bethe-Bloch)
- Biggest cell by cell effect: attenuation in WLS fiber
- Check energy scales with cosmics, beam events, Michels, π^0 mass/hadronic showers in ND data (all agree to $\sim 5\%$)



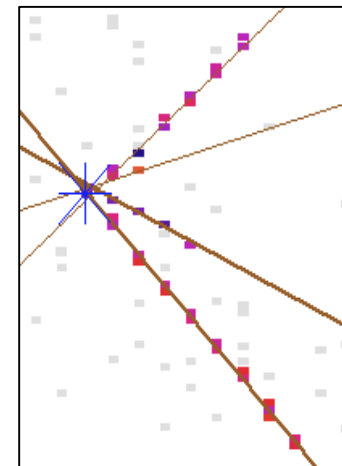
NOvA Reconstruction Basics:

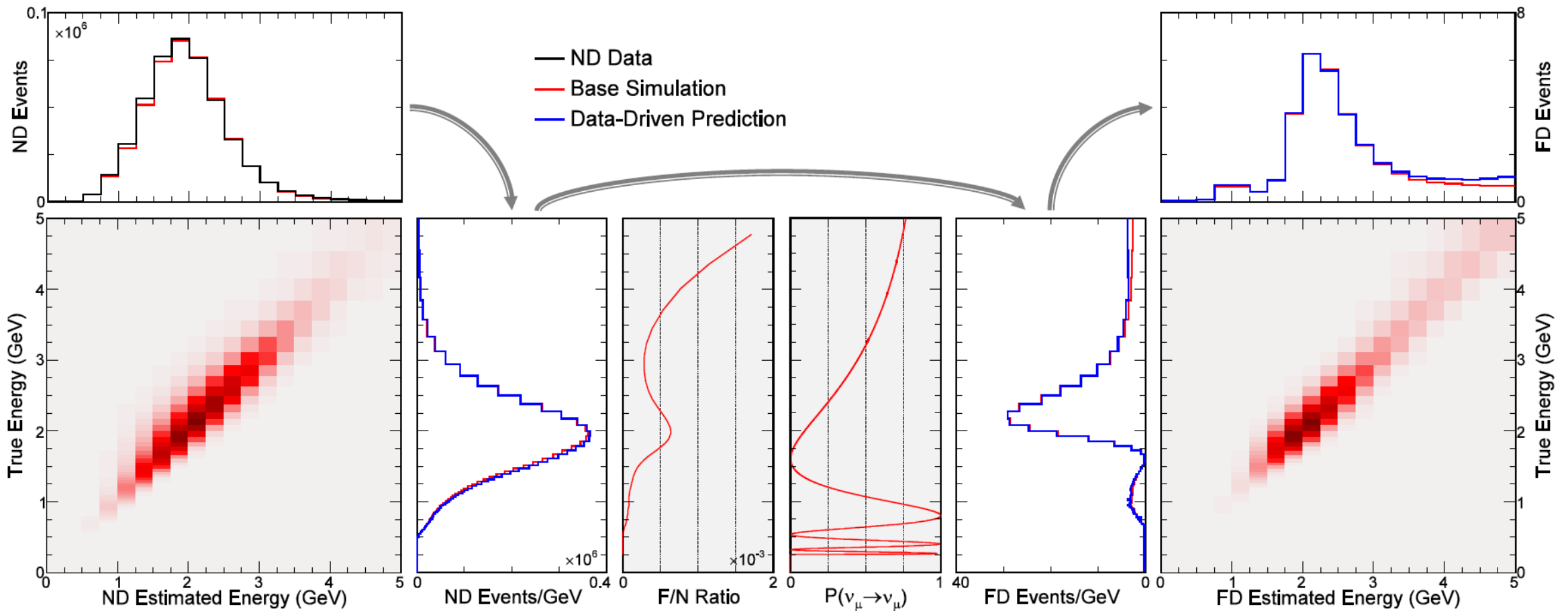
- **Slicing:** cluster hits in space and time to isolate physics interactions; highly accurate, can distinguish between > 50 FD muons in the $550 \mu\text{s}$ spill window with almost no overlap
 - timing resolutions: FD $\sim 150\text{ns}$, ND $\sim 50\text{ns}$
- **Tracking:** for muons especially, also protons and pions (disappearance). Use a Kalman Filter inspired algorithm
- **Vertexing:** for showers, hadronics: track lines of energy deposition back to a single start point (appearance)

tracking example



vertexing example

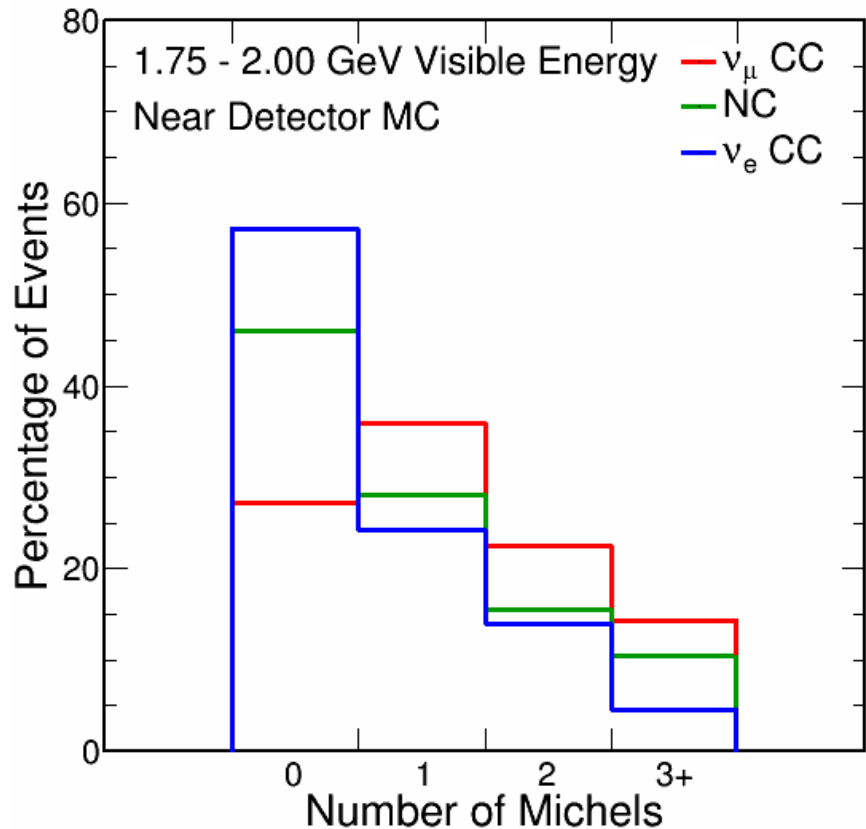




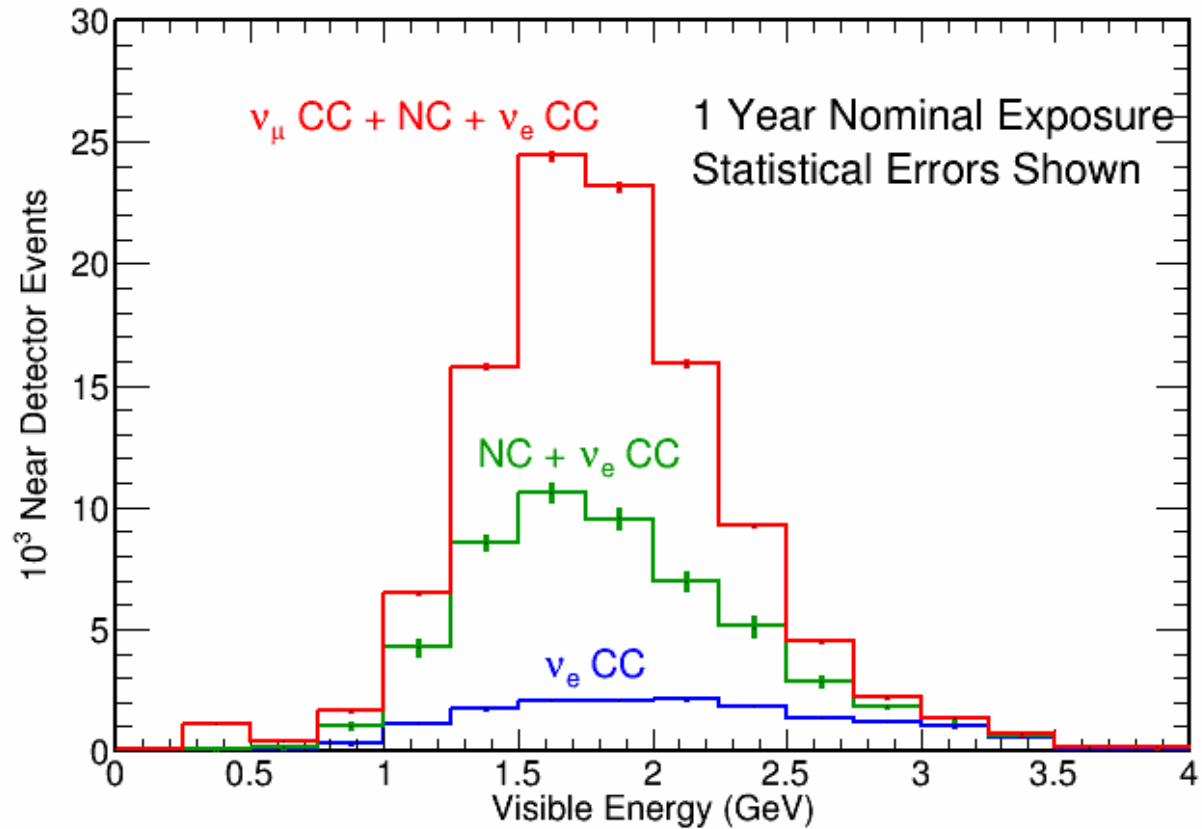
- This entire procedure is re-done beginning to end for each combination of oscillation parameters or systematics being tested
- The extrapolation provides a data-driven approach to help fix any simulation errors and constrain uncertainties

- It is not perfect though – it deals well with normalization effects, but poorly with large energy shifts
- Thus it is also important to make the simulation as accurate as possible

NOvA Simulation



NOvA Simulation



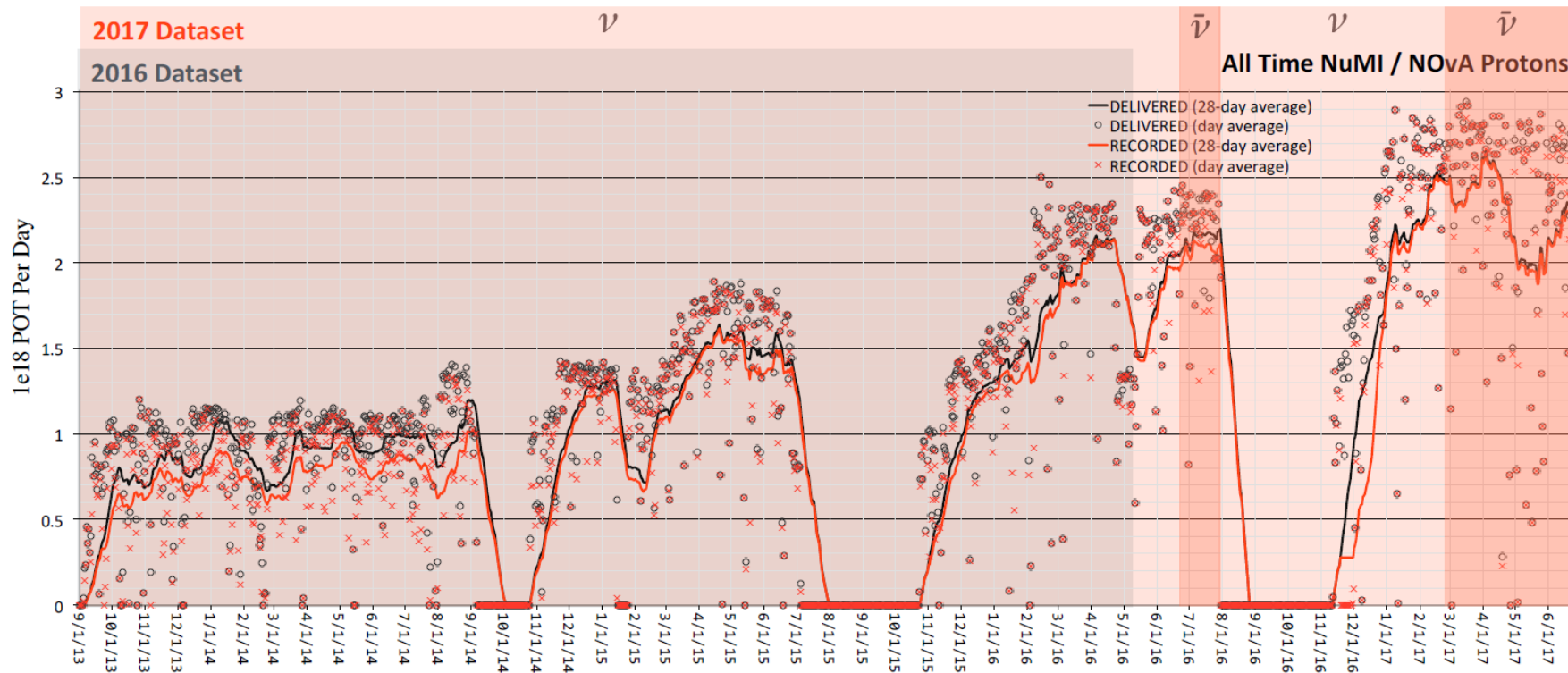
NuMI beam most powerful neutrino beam in the world

Recent upgrades, up to goal of 700 kW

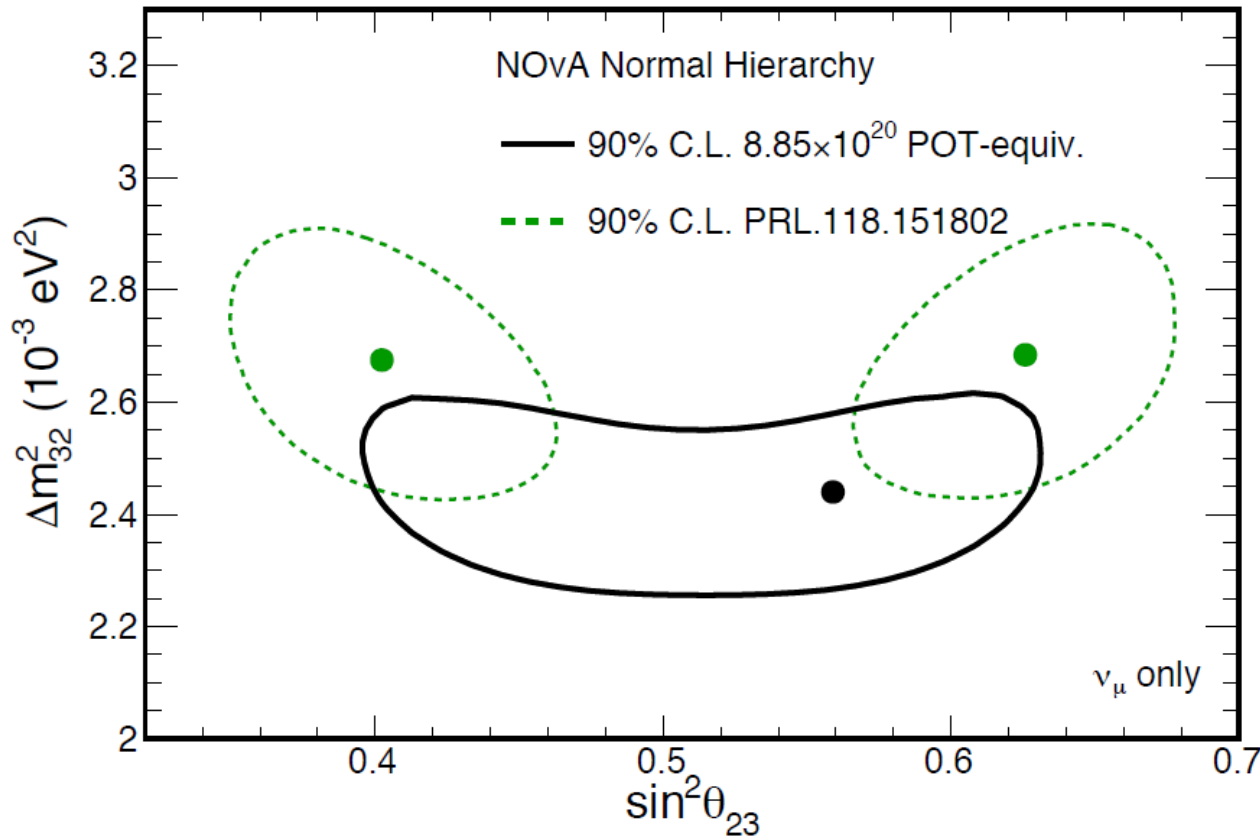
NOvA recently released third set of oscillation results

- based on 8.85×10^{20} POT

All neutrino mode running, anti-neutrino data analyses ongoing



NOvA Preliminary



From $2.6 \rightarrow 0.8 \sigma$ exclusion of max mixing

new light model (include Cherenkov light)
this changes E resolution ($7\% \rightarrow 9\%$) and
shifts hadronic E (~ 70 MeV on average),
which coincidentally pushes 3 events
across bin boundaries (expected: 0.5)

New analysis techniques – energy
resolution binning separates out poor
resolution events that may be background;
removes impact of possible background
fluctuation