# Precision tests with trapped antimatter: A glimpse of the 1S - 2S transition in antihydrogen



#### Dr. Will Bertsche

The University of Manchester The Cockcroft Institute







• Precision measurements on antimatter using Antihydrogen atoms







• Particles have twins with same mass, opposite charge







The University of Manchester

The Cockcroft Institute

• Atoms and antimatter atoms?





• Watch out when they meet their twin!









- <span id="page-5-0"></span>• Annihilation!
- Conversion to (lots of) Energy $E = mc^2$





## Annihilations

- Positron / Electron: photons (511 keV)
- Antiproton / Proton: Many possibilities Pions, etc.







#### First Observation: Positrons

• 1932: Carl Anderson follows up theory quickly: Positrons in Cosmic Rays









#### First Observation: Antiprotons

• 1955: Owen Chamberlain and Emilio Segrè Antiprotons from 1 GeV Protons on Cu Target

**BEVATRON** 





Chamberlain Segrè





CERN Visiting Schools, June 2017

FIG. 1. Diagram of experimental arrangement.

For details see Table I.

Bevatron, Lawrence Berkeley National Lab

**HIFLOING** 

IO FEET



100

## What's the matter with Antimatter?

• Should be equal amounts produced at Big Bang...





CERN Visiting Schools, June 2017



# Possible Explanations: Fundamental Flaw?

• C. P. T. Symmetry: Fundamental Feature of Universe

- 1. Take any experiment
- 2. Swap Charge, Parity, and run Time backwards "CPT Transformation"
- 3. Outcome should be the same

- CPT violation has never been observed
- It is an assumption in essentially all Physic
- Replacing matter with antimatter: a CPT Transformation
- CPT Test: Compare properties of Matter and Antimatter





# Possible Explanations: Gravity?

• Gravity?



#### Where do Positrons come from?

- Easy: Some radioactive isotopes
	- Naturally occurring Potassium-40 (in Bananas: ~ 15 Positrons / sec)
	- 'Manufactured' Sodium-22



"I am a banana!" Don Hertzfeld



#### $^{22}Na \rightarrow ^{22}Ne + e^+ + \nu_e + \gamma$





# Where do Antiprotons come from?

1. Energetic proton creates Proton/Antiproton pair 2. Charge/Mass selected



Cern Proton Synchrotron







 $~\sim$ 3 GeV

(and other stuff)







CERN Visiting Schools, June 2017



# Antimatter: What's it good for?

- Material Characterization (positrons)
- Medical imaging: PET scans
	- Positron Emission Spectroscopy
- High energy physics (antiprotons)









CERN Visiting Schools, June 2017



# Antimatter: What's it good for?

Hollywood has some suggestions...



a surprisingly timely movie...





# Wouldn't that make a good bomb?

Dan Brown writes a novel: "Angels and Demons"

- About researchers from CERN ...
- ... trapping 1/4 gram of antimatter ...
- ... that is stolen by the Illuminati ...
- ... who threaten to blow up the Vatican.









...amidst scandal on the eve of a new Pope





# Antimatter Bombs

- Start with antiprotons: (more bang for the buck)
- Mass conversion efficiency is a fundamental limit.

$$
\rm P \to \sim 10^{-6} \bar P ~at \sim 20 ~GeV/c
$$

- Ignoring practicalities of storage...
- CERN could produce this 1/4 gram in ... ... about 1 Million years

# No bombs. Economics don't make sense.





#### But surely...



Original: Zach Weiner, SMBC-Comics.com



CERN Visiting Schools, June 2017



## History: Forming Antihydrogen

- 1980's Antiprotons at CERN / Fermilab
- 'Fixed Target' smack antiprotons into Xenon
- 9 atoms at CERN (1995), 99 at Fermilab (1996)
- Momenta way too high!







PS210: Baur *et al.* Phys. Lett. B 368, 251 (1996) E862: Blanford *et al.* Phys. Rev. Lett. 80, 3037 (1998)





# History: Forming **COLD** Antihydrogen

- CERN Antiproton Decelerator (1999)
- AD: 10,000,000 antiprotons / 2 minutes
	- "Low Energy" = 5 Million Volts!
- First 'Cold' antihydrogen: ATHENA and ATRAP (2002)
- 100's of millions produced since



CERN Visiting Schools, June 2017





**MANCHESTER** 

## Recipe for Cold Antihydrogen

- 1. Trap ~10 Thousand antiprotons
- 2. Trap ~10 Million positrons
- 3. Chill ingredients to 10's of Kelvin
- 4. Mix, while keeping species cold and confined

5. Bam!

$$
\bar{H}! \stackrel{\scriptscriptstyle e^+}{\longrightarrow} \overline{\phantom{a}^{\!\!\!\!\!-1}}
$$





# ALPHA Apparatus







## ALPHA Apparatus







# Antimatter: Confinement

- Non-neutral plasmas: gas of single-charged particles
	- Pure ensembles of electrons, positrons, antiprotons, etc.







# Capturing Antiprotons

- Degrade antiprotons 5 Million Volts is still a lot...
- Antiprotons equilibrate with electrons







# Antiproton / Electron Plasma

• Image the equilibrium Antiproton / Electron plasma







# Antihydrogen Formation: Mixing Antiprotons and Positrons

- 1. Antiprotons injected into 'Nested Potential'
- 2. Antiprotons lose energy by collisions with Positrons
- 3. Form Antihydrogen, leaves electrostatic trap $\Phi(z)$







# Antihydrogen Detection

- Silicon-strip detector
- 3D 'Digital Camera' - Particle tracks point to vertex
- Vertex resolution ~ 3mm
- > 50% efficiency for annihilations









CERN Visiting Schools, June 2017



## Antihydrogen Detection During Mixing





CERN Visiting Schools, June 2017



# Trapping Antihydrogen

- Atoms are neutral: Not confined by penning traps
- Antihydrogen has a small magnetic moment - Like a little refrigerator magnet
- Can use a magnetic minimum trap (superconducting)
- Orientation matters (solenoid keeps alignment)
- Makes a shallow 'Bathtub' for T < 0.5 K (-272.65 C)







## Trapping Antihydrogen: Search

- 1. Turn on magnetic trap
- 2. Mix and Form Antihydrogen
- 3. Eject remaining charged particles
- 4. Rapidly (< 30 ms) shut off trap ("Quench")
- 5. Detect annihilations







CERN Visiting Schools, June 2017



#### Antihydrogen Search with Bias Fields

• No spatial bias in signal



# Trapped Antihydrogen!

- Antihydrogen trapped. - 1 atom / 15 minutes.
- 100's of atoms for 100's of seconds

The University of Manchester

#### LETTER

#### **Trapped antihydrogen**

G. B. Andresen<sup>1</sup>, M. D. Ashkezari<sup>2</sup>, M. Baquero-Ruiz<sup>3</sup>, W. Bertsche<sup>4</sup>, P. D. Bowe<sup>1</sup>, E. Butler<sup>4</sup>, C. L. Cesar<sup>5</sup>, S. Chapman<sup>3</sup>, M. Charlton<sup>4</sup>, A. Deller<sup>4</sup>, S. Eriksson<sup>4</sup>, J. Fajans<sup>3,6</sup>, T. Friesen<sup>7</sup>, M. C. Fuji N. Madsen<sup>4</sup>, S. Menary<sup>11</sup>, P. Nolan<sup>12</sup>, K. Olchanski<sup>8</sup>, A. Olin<sup>8</sup>, A. Povilus<sup>3</sup>, P. Pusa<sup>12</sup>, F. Robicheaux<sup>13</sup>, E. Sarid<sup>14</sup>, S. Seif el Nasr<sup>9</sup>, D. M. Silveira<sup>15</sup>, C. So<sup>3</sup>, J. W. Storey<sup>8</sup>†, R. I. Thompson<sup>7</sup>, D.

doi:10.1038/nature09610



# Comparing Matter and Antimatter: Color

- Color is a property of Light
- Light is composed of 'Photons'
	- Electromagnetic waves color from wavelength and frequency
- White light from the sun is composed of many colors
	- Use a prism to separate the different types of photons
- What color is Hydrogen?







#### Hydrogen Spectroscopy

- Excited atoms emit discrete wavelengths
- Spectroscopy is the measurement of these photons
- Hydrogen spectrum: well- measured and predicted
	- Ground-state (1S) to first excited state (2S)



#### Atomic States

- Atoms can exist in many discrete (quantum) states
- Different states have different energies<br>- Lowest energy: 'Ground state'
	-
	- Highest energy: Unbound (no longer an atom)
- Atoms transition between states by absorbing or emitting photons
- 'Color' of the photon relates to the energy differences







#### Atomic Spectra and CPT

- Atomic spectra should transform directly by CPT
- Accomplished by swapping in Antihydrogen



#### ALPHA-2: Laser Access Required!

- Modularity for interfacing with CERN/ELENA upgrade - More antiprotons
- Increase antihydrogen trapping rate
- Lasers for Spectroscopy and Cooling
	- 243 nm 2-photon spectroscopy, 121 nm Lyman-alpha laser cooling
- Built from 2012 today

The University of Manchester





#### ALPHA-2: After LS1 (September 2014)







#### 1S - 2S Laser System

- GPS-reference Menlo Systems Frequency Comb locked to ULE cavity
- 243 nm Toptica laser (~ 100 mW) locked to ULE
- *In-situ* PDH-locked cryogen build-up cavity (~ 1 W)



# Challenges with Antihydrogen Spectroscopy

- Ultimate goal in ALPHA: Measure 1S 2S transition Problem: Few trapped atoms
	- Direct detection of absorbed or radiated photons is presently futile
- Solution:
- Drive antihydrogen from a trapped to untrapped state
- Efficiently detect annihilation





# Untrapping Antihydrogen

• Example: Hyperfine Transition (spin-flip)



 $\lambda_{\rm hf} = 21.1061140541791(13)$  cm

• Example: Ionization with ultraviolet light



M. Niering, et al. Phys. Rev. Lett. 84, 5496 (2000)





# 1S - 2S Transition in (anti) hydrogen

- 2 photon Doppler-free spectroscopy (243 nm)
- Drive between trapped hyperfine states







#### 1S - 2S possible outcomes



# 1S - 2S Experiment

- Produce and trap antihydrogen
- Illuminate experiment (or not) for 600 seconds
	- **- On-Resonance** 
		- Drive  $f_{cc}$  and  $f_{dd}$  (300 seconds each)
	- **- Off-Resonance** 
		- Detune each by 200 kHz
	- **- No-laser**
- Fast magnet ramp-down
	- Look for disappearance
- Also look for appearance
	- Multivariate Analysis…





# 1S - 2S Disappearance

• ON-Resonace de-populates the trap



- ON and OFF resonance trials differ by  $92 \pm 15$  counts (Detector efficiency here is 0.376)
- $(58 \pm 6)\%$  of atoms removed





# 1S - 2S Appearance

#### • Tune MVA for appearance mode



- Difference (ON OFF) resonance totals is  $52 \pm 10$ 
	- (Detector efficiency here is 0.376)

Annihilations in disappearance  $92 / 0.688$  | 134 Annihilations in appearance  $\left(52 / 0.376\right)$  138





#### 1S - 2S Summary



OPF N doi:10.1038/nature21040

#### Observation of the  $1S-2S$  transition in trapped antihydrogen

M. Ahmadi<sup>1</sup>, B. X. R. Alves<sup>2</sup>, C. J. Baker<sup>3</sup>, W. Bertsche<sup>4,5</sup>, E. Butler<sup>6</sup>, A. Capra<sup>7</sup>, C. Carruth<sup>8</sup>, C. L. Cesar<sup>9</sup>, M. Charlton<sup>3</sup>, S. Cohen<sup>10</sup>, R. Collister<sup>7</sup>, S. Eriksson<sup>3</sup>, A. Evans<sup>11</sup>, N. Evetts<sup>12</sup>, J. Fajans<sup>8</sup>, T. Friesen<sup>2</sup>, M. C. Fujiwara<sup>7</sup>, D. R. Gill<sup>7</sup>, A. Gutierrez<sup>13</sup>, J. S. Hangst<sup>2</sup>, W. N. Hardy<sup>12</sup>, M. E. Hayden<sup>14</sup>, C. A. Isaac<sup>3</sup>, A. Ishida<sup>15</sup>, M. A. Johnson<sup>4,5</sup>, S. A. Jones<sup>3</sup>, S. Jonsell<sup>16</sup>, L. Kurchaninov<sup>7</sup>, N. Madsen<sup>3</sup>, M. Mathers<sup>17</sup>, D. Maxwell<sup>3</sup>, J. T. K. McKenna<sup>7</sup>, S. Menary<sup>17</sup>, J.





## 1S - 2S Prospects

- The transition has been found (100's kHz level)
- Measurement of lineshape limited by end of beamtime
- Precision at the 10's kHz level is possible
- $\cdot$  ~10<sup>-10</sup> (Hydrogen)







# Precision gravity?

• Do atoms and anti-atoms gravitate differently?

$$
F_{\rm antimatter} = F \cdot m g
$$

• Antihydrogen will fall out the bottom (or top) of the trap







# Gravitational Deflection: Precision?

• Simulate various F, test exclusion of RCA during quench

 $-65 < F < 110$ 

- Not very precise:
	- Poor statistics, hot population, short distance
- Charge neutrality important





 $COMM$ UNICATIONS | 4:1785 | DOI: 10.1038/nc

Description and first application of a new technique to measure the gravitational mass of antihydrogen



# ALPHA-g: Precision gravitational measurements with antihydrogen

- $\cdot$  ~ 2 m tall antihydrogen trap
- Release + detect falling Hbar
- Measure sign of gbar  $\sim$  1 year
- Measure gbar a ~ 1% - 4 - 5 years







# **Summary**

- Understanding the differences between matter and antimatter is a Grand Challenge of physics
- ALPHA has taken the first steps towards this goal by trapping antihydrogen, performing preliminary measurements on trapped antihydrogen.
- ALPHA-2: Recently demonstrated driving the 1S 2S transition
	- Line shape measurements in the near future!
- •ALPHA-g: Future effort on gravity underway!





#### Thanks!

... Many things you can do with antimatter in a can!





CERN Visiting Schools, June 2017

