# SBN physics case (A short review)

Sandro Palestini EP-NU meeting – 19/04/18

## Situation (to be confirmed): more than 3 neutrinos

- LSND oscillations (anti- $\nu_{\mu}$  -> anti- $\nu_{e}$  from muons at rest, Los Alamos, at ≈0.25% fraction ) 3.8 σ evidence of appearance
- Confirmed by MiniBooNE (BNB beam at Fermilab) with 3.4 (2.8)  $\sigma$  for (anti)neutrinos
- Low energy anomalies: v<sub>e</sub> flux deficit ≈7% observed at ≈3 σ with nearby detectors at reactors, and with calibration sources at solar neutrino experiments SAGE and Gallex.
- − Phenomena at *L/E*≈1 m/MeV, or  $\Delta m^2 \approx 1 \text{ eV}^2$ , far from the scale of oscillations established at  $\Delta m^2 = 2.5 \times 10^{-3}$ , 0.76×10<sup>-4</sup> eV<sup>2</sup> for atmospheric and solar neutrinos.
- An additional mass difference implies at least one additional neutrino, which does not take part in SM weak interactions.



LSND and MiniBooNE have observed an excess of  $v_e$  equal to  $\approx 0.25\%$  of  $v_{\mu}$ , after subtraction of other sources and bkg.

Band of solution for initial oscillation with (Δm<sup>2</sup>)<sup>2</sup>×sin<sup>2</sup>(2θ)≈constant



# Phenomenology

- New mass difference: at least one new sterile u
- Schemes:
  - 3+1: 3 u's in the usual mass pattern plus a fourth one (which does not interact with leptons, Z ...) well separated from the others.
  - 2+2 : doublet internally separated by dm<sup>2</sup><sub>solar</sub>, dm<sup>2</sup><sub>atmospheric</sub>, and globally separated by ≈1eV <sup>2</sup>
  - >1 sterile neutrinos



# Phenomenology of 3+1 v's

- 3+1 is more frequently taken as paradigm:
  - In the regime of oscillations driven by  $\Delta m_{4i}^2$ :
    - $Prob^{3+1}(\upsilon_{\mu} \rightarrow \upsilon_{e}) = \sin^{2}(2\theta_{\mu e}) \sin^{2}(\Delta m^{2}_{4i}L/4E)$
    - $Prob^{3+1}(\upsilon_{\mu} > \upsilon_{\mu}) = 1 \sin^2(2\theta_{\mu\mu}) \sin^2(\Delta m_{4i}^2 L/4E)$
    - $\sin^2(2\theta_{\mu e}) = 4 |U_{\mu 4}U_{e 4}|^2$
    - $\sin^2(2\theta_{\mu e}) = 4 |U_{\mu 4}|^2 (1 |U_{\mu 4}|^2)$
  - Similar equations can be written in the regimes of oscillation driven by  $\Delta m_{21}^2$  and  $\Delta m_{32}^2 \cong \Delta m_{31}^2$ .
  - Observed amplitude of the oscillations in the 2-neutrinos scheme implies  $|U_{\mu4}|^2 \approx |U_{e4}|^2 << 1$ .

## SBN concept in few words

- Improve on MiniBooNE with better detector(s) (better ID, less background)
- Expand the L/E range using three similar detectors
  - SBND, MicroBooNE, Icarus (T600) at 150m, 470m, 600m and 80t, 90t, 480 t fiducial volume respectively
  - Similar (LArTPC) technology to improve performance and systematics
  - Reduction of beam systematics due to knowledge of BNB and flux scaling properties (in particular between 470 and 600 m)
- Experiments at surface, cosmic-ray induced background is relevant.

## Measurement main target: u<sub>e</sub> appearance

To be obtained measuring the *E*, *L* dependence of Charged-Current events in excess of non-oscillating  $v_{\rho}$  in the beam (from K,  $\mu$  decays) and of backgrounds



### Measurement target: mainly u<sub>e</sub> appearance

#### • Detectors:

- Charged particle reconstruction in LAr TPC , with collection time 1.3 ms, 1.6 ms , 1.0 ms respectively, with 3 layers read-out wires, 3 mm spacing (read out in Lar).
- $E_{v}$  from energy of electron (CC interaction)
- Scintillation light signal to enforce coincidence with beam spill
- Beam: Fermilab Booster Neutrino Beam (8 GeV p)
  5E12 p per spill, 1.6 micros spill, 2ns/19ns substructure
  5 Hz cycle, 6.6E20 p on target, 3 years run
  (211 s total *beam time*)

# Backgrounds

- Beam interaction in detectors:
  - NC gamma misidentified as electron events (reject with local energy deposition) (E > 200 MeV taken as selection criteria, with 80% reconstruction efficiency)
  - $\upsilon_{\mu}$  CC with  $\pi^{0}$ ->  $\gamma$ 's mistaken as e<sup>±</sup> and  $\mu$  misidentified
  - $v_e$  e- elastic scattering, but the cross section is small
- Beam interactions near detector (*dirt events*):
  - Fake e<sup>±</sup> mostly near detector front and side walls, mitigated with fiducial volume < LAr volume</li>
- Background related to cosmic rays (cosmogenic bkg.):
  - gamma conversion taken as  $\upsilon_e$  CC

### Cosmogenic bkg. continued

- Needs to be in time with beam spill
- Better, in time with spill substructure (2 ns  $\approx \sigma_{\text{time}} \ll$  19 ns)
- Tends to come together with a muon entering/crossing the detector
  - Reject if  $v_e$  CC candidate near a muon (or near wall)
- Can be present also if in time only with the detector collection time (≅1 ms)
  - if a second cosmogenic background event occurs in a triple coincidence with the spill (1.4  $\mu s$ ) (triple coincidence)
  - or if a beam interaction (e.g.  $\upsilon_{\mu}$  CC) takes place in the same spill.
- Mitigate with:
  - association of  $\upsilon_e$  CC candidate with scintillating signal to enforce timing to 1.4  $\mu s$
  - External veto tagging (position and time for 1.4  $\mu s$  coincidence)



# Example of expected dN<sub>ue</sub>/dE

Note the large value of the oscillation amplitude chosen here, and the differences in event yield, related to position and volume. The muon tagger reduces the cosmogenic bkg. by a factor > 100.



## Beam related systematics

- Normalization uncertainties in the total flux are at the level of 10-15% in both neutrino flux and the neutrino interaction uncertainties.
- However the effect of these uncertainties cancel to large extent:
  - Relative rates in the three detectors are used in the measurement
  - Very similar detector response, and use of the same beam result in largely coherent systematic errors (hence favorable) in reconstructed energy spectra among the different detectors and the between  $v_e$  and  $v_{\mu}$  CC events.

## Coming to the conclusion:



