

Connection to Cosmology

The electroweak phase transition and the flavour puzzle

Iason Baldes



Higgs tasting workshop - Benasque, 19 May 2016

Electroweak baryogenesis - an exciting link between the Higgs and cosmology.



- Small review of electroweak baryogenesis
- Some new ideas on the flavour puzzle and its possible role in electroweak baryogenesis

Electroweak baryogenesis - basic picture

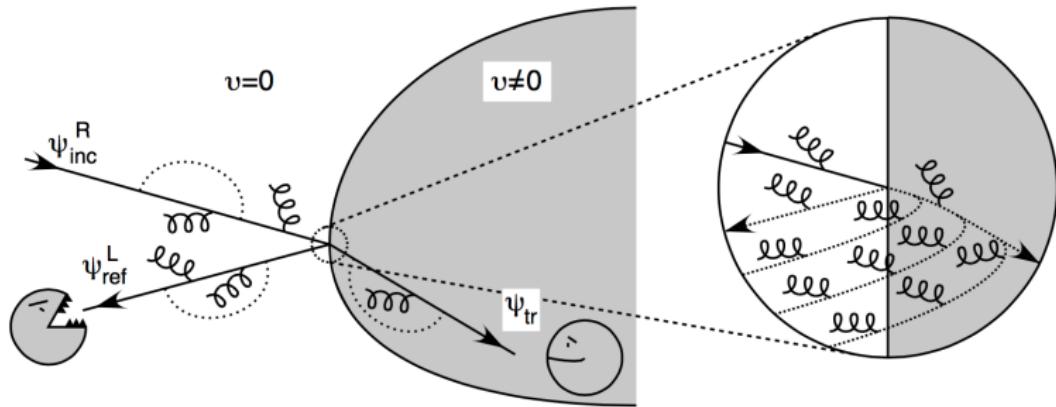
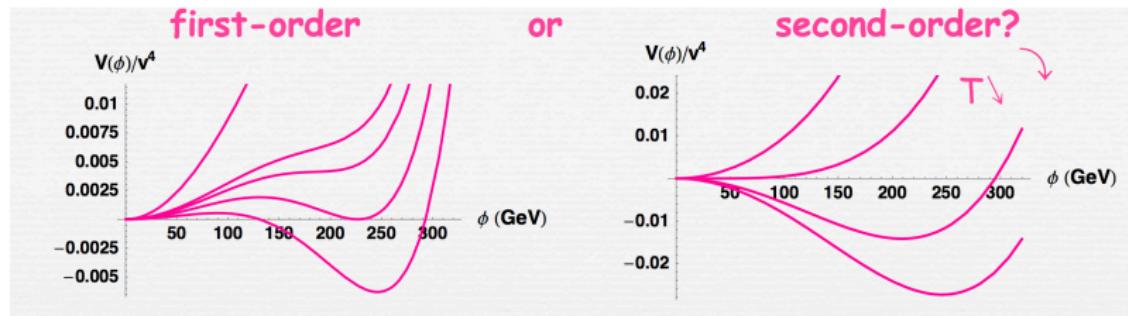


Image from - Gavela, Hernandez, Orloff, Pène, Quimbay [hep-ph/9406289]

- CP violating collisions with the bubble walls lead to a chiral asymmetry.
- Sphalerons convert this to a Baryon Asymmetry.
- This is swept into the expanding bubble where sphalerons are suppressed.

Electroweak baryogenesis - Requirements



Electroweak baryogenesis requires:

- A strong first order phase transition
- Sufficient CP violation

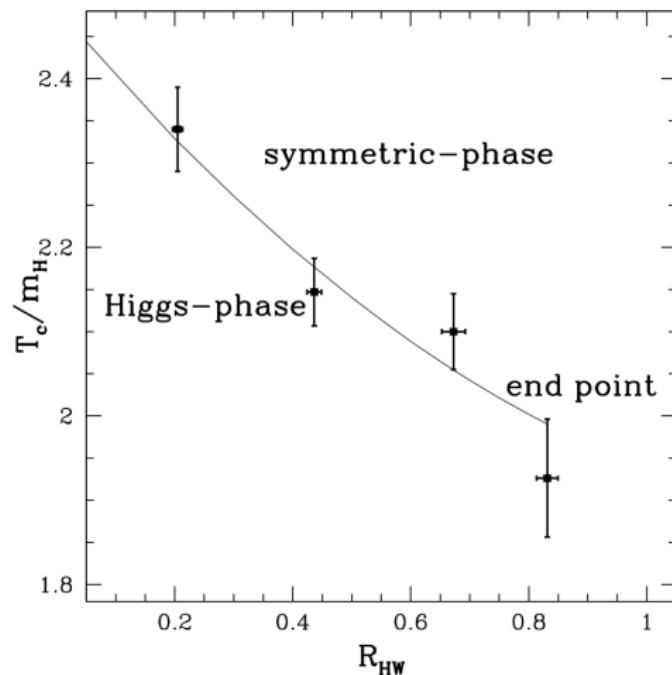
However in the SM:

- The Higgs mass is too large
- Quark masses are too small

We require new (EW-scale) physics!

Electroweak phase transition

Lattice calculations show the SM Higgs mass is too large.

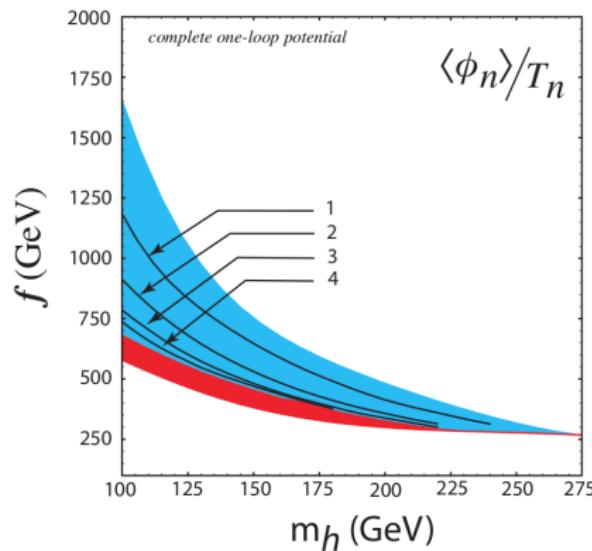


$$R_{HW} \equiv m_H/m_W$$

Endpoint at:
 $m_H \approx 67$ GeV

- Csikor, Fodor, Heitger, Phys. Rev. Lett. 82, 21 (1999)

Require a modification of the Higgs potential



$$V(H) = m^2 |\Phi|^2 + \lambda |\Phi|^4 + \frac{1}{f^2} |\Phi|^6$$

Other options:

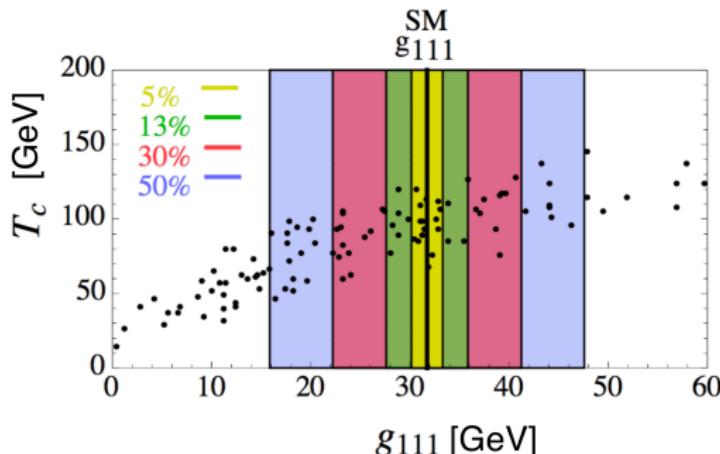
- Singlet models/tree level barriers
- Thermal barriers from bosonic loops
- Multi-step transitions

- Delaunay, Grojean, Wells [0711.2511]

Successful electroweak baryogenesis requires:

$$\Gamma_{\text{sph}} \sim 10^{1/4} \left(\frac{\alpha_W T}{4\pi} \right)^4 \left(\frac{2M_W(\phi)}{\alpha_W T} \right)^7 \text{Exp} \left[-\frac{3.2M_W(\phi)}{\alpha_W T} \right] \lesssim H \Rightarrow \frac{\phi_c}{T_c} \gtrsim 1$$

Collider signatures - example



Correlation between T_c and triple Higgs couplings $g_{111} h^3$ in a singlet model. - Profumo, Ramsey-Musolf, Wainwright, Winslow [1407.5342]

- Example of how the Higgs potential can be probed by experiment.
- This would also constrain the parameter f in the previous example.
- Other signals can also be found in the literature e.g. $h \rightarrow \gamma\gamma$ in inert 2HDM models. - Blinov, Profumo, Stefaniak [1504.05949]

Baryogenesis from charge transport with SM CP violation

$$R_{LR} = \text{Diagram 1} + \text{Diagram 2} + \dots$$

The equation $R_{LR} =$ is followed by two Feynman diagrams separated by a plus sign, indicating a sum of contributions. The first diagram shows a loop with two external lines labeled L and R entering a central vertex labeled \mathcal{M} . The second diagram shows a more complex loop structure with internal lines labeled M^\dagger , R , L , and \mathcal{M} .

$$\epsilon_{\text{CP}} \sim \frac{1}{M_W^6 T_c^6} \prod_{\substack{i>j \\ u,c,t}} (m_i^2 - m_j^2) \prod_{\substack{i>j \\ d,s,b}} (m_i^2 - m_j^2) J_{\text{CP}}$$

- Gavela, Hernandez, Orloff, Pène, Quimbay [hep-ph/9406289],
- Huet, Sather [hep-ph/9404302].

SM quark masses are too small!

EDMs

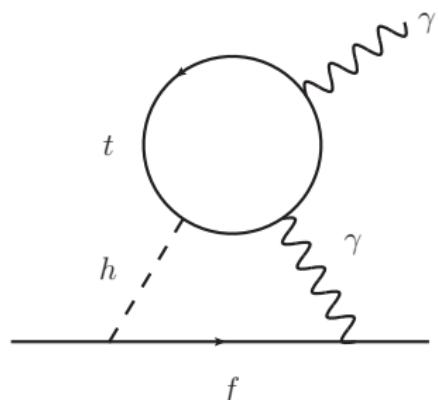
Additional CP violation

$$\mathcal{L} \supset y_{ij} \overline{Q_i} \Phi u_j + x_{ij} \frac{(\Phi^\dagger \Phi)}{\Lambda_{\text{CP}}^2} \overline{Q_i} \Phi u_j$$

- Huber, Pospelov, Ritz [hep-ph/0610003], Konstandin [1302.6713]

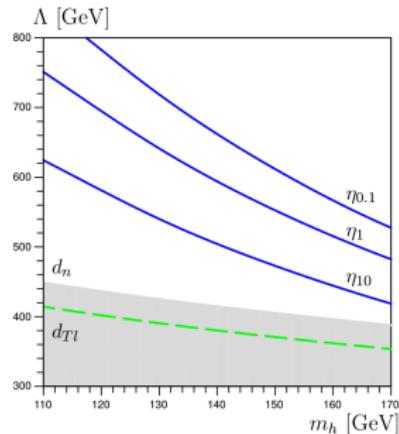
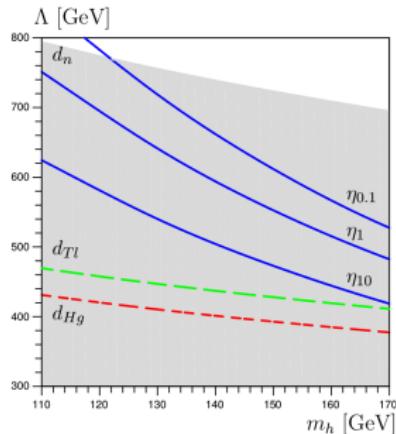
Neutron EDM: $|d_n| < 3 \times 10^{-26} e \text{ cm}$

- Such operators are constrained from EDMs and FCNCs.
- Constraint from neutron EDM:
 $\Lambda_{\text{CP}} \gtrsim \sqrt{\text{Im}[x_{33}]} \times 750 \text{ GeV}.$
- Small Λ_{CP} possible with $x_{ij} \sim y_{ij}$.



Additional CP violation

$$\mathcal{L} \supset y_{ij} \overline{Q_i} \Phi u_j + x_{ij} \frac{(\Phi^\dagger \Phi)}{\Lambda_{\text{CP}}^2} \overline{Q_i} \Phi u_j$$



Plots for $\Lambda \equiv f = \Lambda_{\text{CP}}$. Left: top only (x_{33}). Right: MFV.

- Huber, Pospelov, Ritz [hep-ph/0610003]

Solutions to the flavour puzzle

Yukawa interactions

$$y_{ij} \bar{f}_L^i \Phi^{(c)} f_R^j$$

Possible solutions

- Froggatt-Nielsen
- Composite Higgs
- Randall-Sundrum Scenario

Froggatt-Nielsen Yukawas: $y_{ij} \sim \left(\frac{\langle\chi\rangle}{\Lambda}\right)^{-q_i+q_j+q_H}$

Previous work on flavour cosmology and EWBG

① Baryogenesis from the Kobayashi-Maskawa phase

- Berkooz, Nir, Volansky - Phys. Rev. Lett. 93 (2004) 051301

- Froggatt-Nielsen scenario.
- Pointed out that CP violation could be unsuppressed before EWPT.
- No EDM signal.
- Ignored phase transition strength.

② Split fermions baryogenesis from the Kobayashi-Maskawa phase

- Perez, Volansky - Phys. Rev. D 72 (2005) 103522

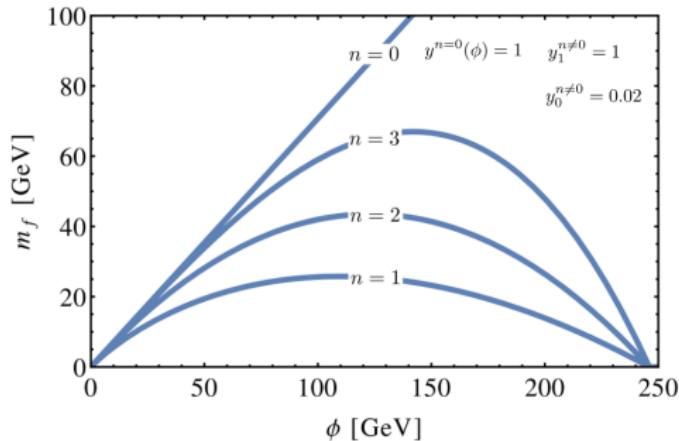
- Baryogenesis from the localiser phase transition.
- CP violation unsuppressed before localiser phase transition.
- Non-standard EWBG: unsuppressed $B - L$ violating operators play the role of sphalerons.

However:

Varying Yukawas can themselves change the nature of the EWPT.

Varying Yukawas

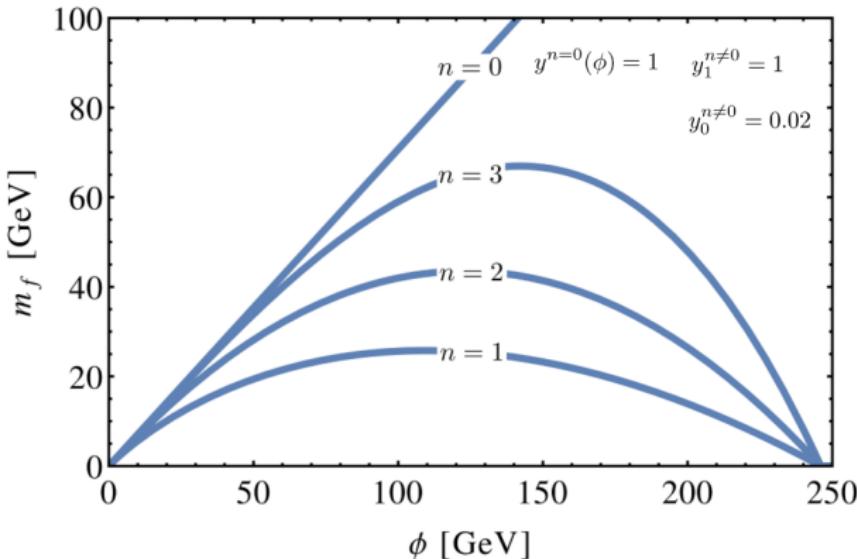
Study the strength of the EWPT with varying Yukawas in a model independent way. - IB, Konstandin, Servant (1604.04526)



Ansatz

$$y(\phi) = \begin{cases} y_1 \left(1 - \left[\frac{\phi}{v}\right]^n\right) + y_0 & \text{for } \phi \leq v, \\ y_0 & \text{for } \phi \geq v. \end{cases}$$

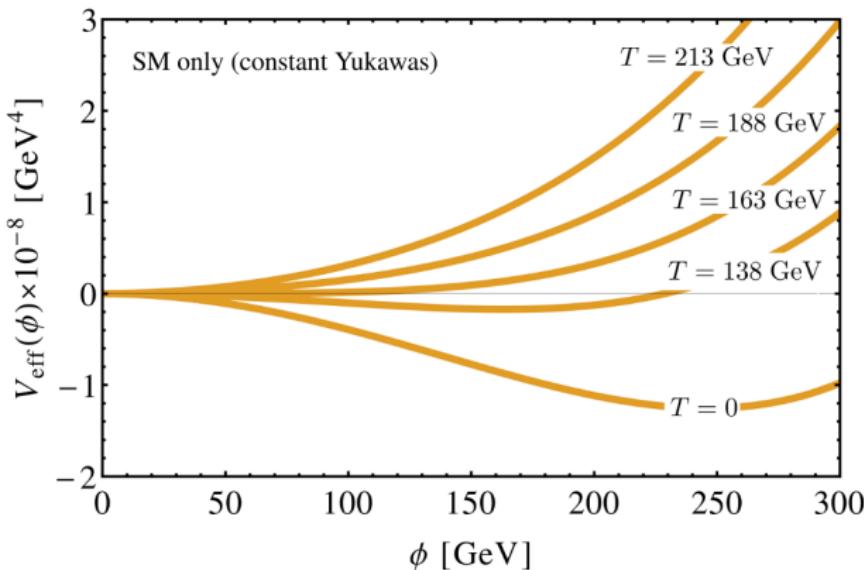
Effective Potential



Thermal correction

$$V_{\text{eff}} \supset -\frac{g_* \pi^2}{90} T^4$$

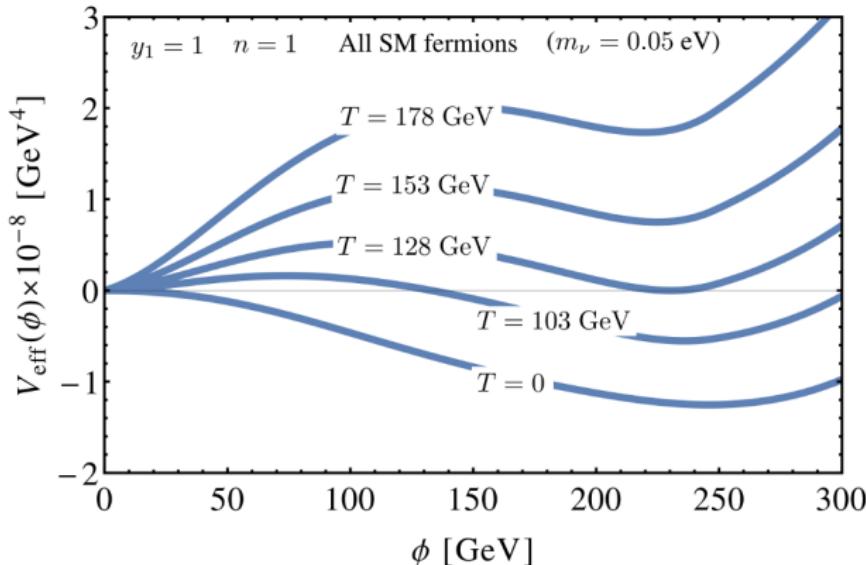
Effective Potential - SM case



Second order phase transition $T_c = 163 \text{ GeV}$.

$$V_{\text{eff}} = V_{\text{tree}}(\phi) + V_1^0(\phi) + V_1^T(\phi, T) + V_{\text{Daisy}}(\phi, T)$$

Effective Potential - Varying Yukawas



Strong first order phase transition

$$\phi_c = 230 \text{ GeV}, \quad T_c = 128 \text{ GeV}, \quad \phi_c/T_c = 1.8$$

Effective Potential - $T = 0$ terms

$$V_{\text{eff}} = V_{\text{tree}}(\phi) + V_1^0(\phi) + V_1^T(\phi, T) + V_{\text{Daisy}}(\phi, T)$$

$$V_{\text{tree}}(\phi) = -\frac{\mu_\phi^2}{2}\phi^2 + \frac{\lambda_\phi}{4}\phi^4$$

$$V_1^0(\phi) = \sum_i \frac{g_i(-1)^F}{64\pi^2} \left\{ m_i^4(\phi) \left(\text{Log} \left[\frac{m_i^2(\phi)}{m_i^2(v)} \right] - \frac{3}{2} \right) + 2m_i^2(\phi)m_i^2(v) \right\}$$

Gives a large negative contribution to the ϕ^4 term.

- Can lead to a new minimum between $\phi = 0$ and $\phi = 246$ GeV.
- Not an issue for previous $y_1 = 1, n = 1$ example.
- Can make phase transition weaker.

Effective Potential - one-loop $T \neq 0$ correction

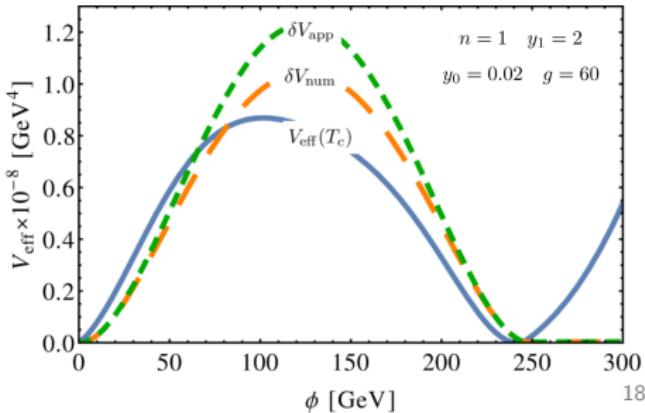
$$V_1^T(\phi, T) = \sum_i \frac{g_i(-1)^F T^4}{2\pi^2} \times \int_0^\infty y^2 \text{Log} \left(1 - (-1)^F e^{-\sqrt{y^2 + m_i^2(\phi)/T^2}} \right) dy$$

$$V_f^T(\phi, T) = -\frac{g T^4}{2\pi^2} J_f \left(\frac{m_f(\phi)^2}{T^2} \right)$$

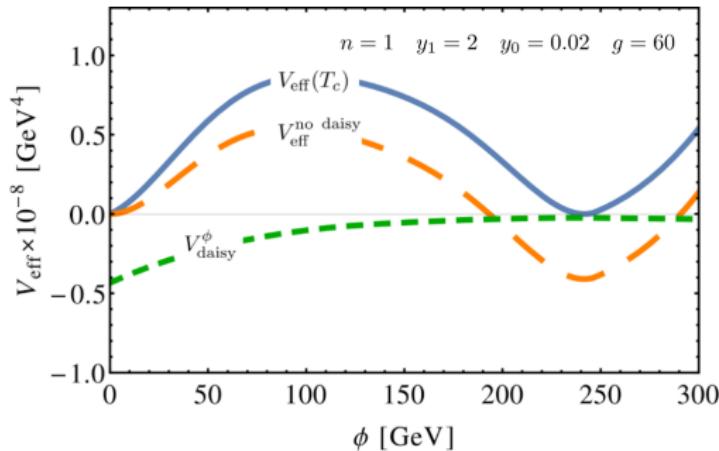
$$J_f \left(\frac{m_f(\phi)^2}{T^2} \right) \approx \frac{7\pi^4}{360} - \frac{\pi^2}{24} \left(\frac{m}{T} \right)^2 - \frac{1}{32} \left(\frac{m}{T} \right)^4 \text{Log} \left[\frac{m^2}{13.9 T^2} \right], \quad \text{for } \frac{m^2}{T^2} \ll 1,$$

$$\delta V \equiv V_f^T(\phi, T) - V_f^T(0, T)$$

$$\approx \frac{g T^2 \phi^2 [y(\phi)]^2}{96}$$



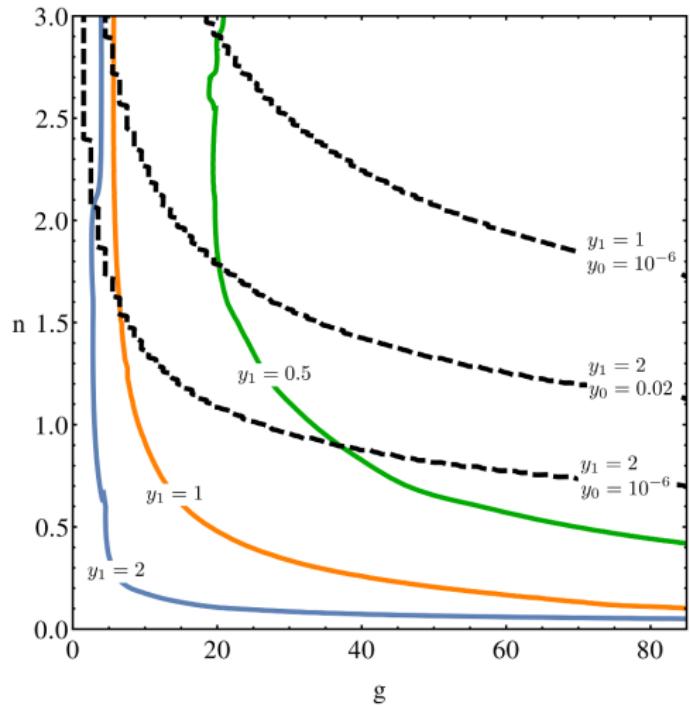
Effective Potential - daisy correction



$$V_{\text{Daisy}}^\phi(\phi, T) = \frac{T}{12\pi} \left\{ m_\phi^3(\phi) - [m_\phi^2(\phi) + \Pi_\phi(\phi, T)]^{3/2} \right\}$$

$$\Pi_\phi(\phi, T) = \left(\frac{3}{16}g_2^2 + \frac{1}{16}g_Y^2 + \frac{\lambda}{2} + \frac{y_t^2}{4} + \frac{gy(\phi)^2}{48} \right) T^2$$

Overall strength



$$y(\phi) = y_1 \left(1 - \left[\frac{\phi}{v} \right]^n \right) + y_0 \quad \text{for } \phi \leq v$$

Introducing the Flavon — An illustrative example

We now introduce the flavon dof - IB, Konstandin, Servant [16mm.xxxxx]

Froggatt-Nielsen

$$\begin{aligned}\mathcal{L} = & \tilde{y}_{ij} \left(\frac{S}{\Lambda_s} \right)^{n_{ij}} \overline{U}_i \tilde{\Phi} Q_j + y_{ij} \left(\frac{S}{\Lambda_s} \right)^{m_{ij}} \overline{D}_i \Phi Q_j \\ & + \tilde{f}_{ij} \left(\frac{X}{\Lambda_\chi} \right)^{2n_{ij}} \overline{U}_i \tilde{\Phi} Q_j + f_{ij} \left(\frac{X}{\Lambda_\chi} \right)^{2m_{ij}} \overline{D}_i \Phi Q_j + H.c.\end{aligned}$$

Under $U(1)_{\text{FN}}$: S (-1) and X (-1/2). Define $\epsilon_s \equiv \langle S \rangle / \Lambda_s$, $\epsilon_\chi \equiv \langle X \rangle / \Lambda_\chi$.

Charges and resulting Yukawas and mixings

$$\begin{array}{llllll}Q_3 (0), & Q_2 (+2), & Q_1 (+3), & y_t \sim 1, & y_c \sim \epsilon_s^3, & y_u \sim \epsilon_s^7, \\ \overline{U}_3 (0), & \overline{U}_2 (+1), & \overline{U}_1 (+4), & y_b \sim \epsilon_s^2, & y_s \sim \epsilon_s^4, & y_d \sim \epsilon_s^6, \\ \overline{D}_3 (+2), & \overline{D}_2 (+2), & \overline{D}_1 (+3). & s_{12} \sim \epsilon_s, & s_{23} \sim \epsilon_s^2, & s_{13} \sim \epsilon_s^3.\end{array}$$

Introducing the Flavon — An illustrative example

Froggatt-Nielsen

$$\begin{aligned}\mathcal{L} = & \tilde{y}_{ij} \left(\frac{S}{\Lambda_s} \right)^{n_{ij}} \overline{U}_i \tilde{\Phi} Q_j + y_{ij} \left(\frac{S}{\Lambda_s} \right)^{m_{ij}} \overline{D}_i \Phi Q_j \\ & + \tilde{f}_{ij} \left(\frac{X}{\Lambda_\chi} \right)^{2n_{ij}} \overline{U}_i \tilde{\Phi} Q_j + f_{ij} \left(\frac{X}{\Lambda_\chi} \right)^{2m_{ij}} \overline{D}_i \Phi Q_j + H.c.\end{aligned}$$

What if Froggatt-Nielsen dynamics takes place close to the EW scale?
This can lead to variation of Yukawa couplings during the EWPT.

Symmetry breaking

$$\Phi \rightarrow \frac{1}{\sqrt{2}} \begin{pmatrix} G_1 + iG_2 \\ v_\phi + \phi + iG_3 \end{pmatrix}, \quad S \rightarrow \frac{v_s + \sigma + i\rho}{\sqrt{2}}, \quad X \rightarrow \frac{v_\chi + \chi + i\eta}{\sqrt{2}}.$$

Long range forces must be suppressed by introducing explicit breaking of $U(1)_{\text{FN}}$.

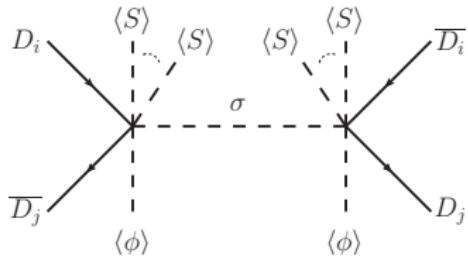
$$V(S) \supset -\mu_s^2 S^\dagger S + \lambda_s (S^\dagger S)^2 - A^2 (SS + S^\dagger S^\dagger)$$

Minimisation of the potential gives the relations

$$m_\sigma^2 = 2\mu_s^2 + 4A^2 = 2\lambda_s v_s^2, \quad m_\rho^2 = 4A^2.$$

We will take $m_\rho > m_\sigma$ below.

Constraints on S



$$\mathcal{H} = C_2^{sd}(\bar{s}Ld)^2 + \tilde{C}_2^{sd}(\bar{s}Rd)^2 + C_4^{sd}(\bar{s}Ld)(\bar{s}Rd) + H.c.$$

$$C_2^{sd} = \left(\frac{5\epsilon_s^4 v_\phi y_{21}}{2\Lambda_s m_\sigma} \right)^2$$

$$\tilde{C}_2^{sd} = \left(\frac{5\epsilon_s^4 v_\phi y_{12}^*}{2\Lambda_s m_\sigma} \right)^2$$

$$C_4^{sd} = y_{12}y_{21}^* \left(\frac{5\epsilon_s^4 v_\phi}{2\Lambda_s m_\sigma} \right)^2$$

$$\sqrt{\Lambda_s m_\sigma} \gtrsim |\text{Im}[y_{12}y_{21}^*]|^{1/4} \times 15 \text{ TeV}$$

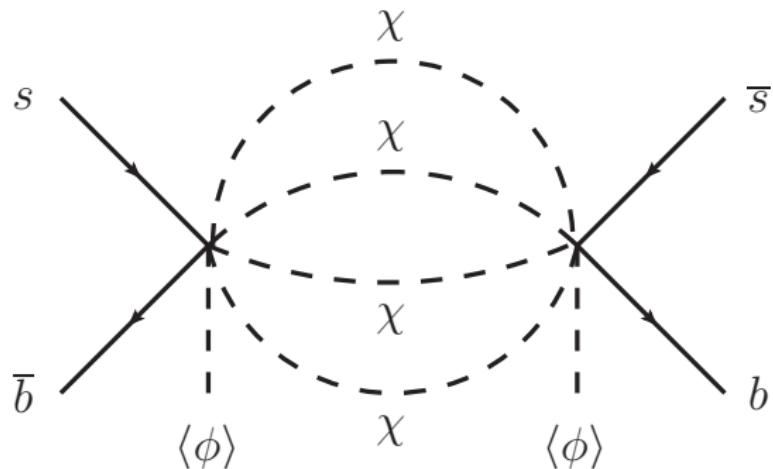
$$\sqrt{\Lambda_s m_\sigma} \gtrsim |\text{Re}[y_{12}y_{21}^*]|^{1/4} \times 4.8 \text{ TeV}$$

$$\sqrt{\Lambda_s m_\sigma} \gtrsim |\text{Im}[y_{21}y_{21}]|^{1/4} \times 10 \text{ TeV}$$

$$\sqrt{\Lambda_s m_\sigma} \gtrsim |\text{Re}[y_{21}y_{21}]|^{1/4} \times 2.7 \text{ TeV}$$

$\Rightarrow \sqrt{\Lambda_s m_\sigma} \gtrsim \text{few } \times \text{ TeV. For Yukawa variation we need } m_\sigma \sim v_\phi^2 / \Lambda_s.$

Constraints on X



We assume $\langle\chi\rangle = 0$ today.

$$|C_2^{bs}| \approx \left(\frac{|f_{23}| v_\phi}{2 \Lambda_\chi^4} \right)^2 \left\{ \frac{1}{(8\pi)^3} \Lambda_\chi^4 \right\}$$

$$\Lambda_\chi \gtrsim \sqrt{|f_{23}|} \times 360 \text{ GeV}$$

VEV variation to Yukawa variation

We want the VEV of χ to move from:

$$\langle \chi \rangle = v_\chi \sim \Lambda_\chi$$

to:

$$\langle \chi \rangle = 0.$$

during the EWPT.

- Choice of FN charge $-1/2$ for X gives a Z_2 symmetry and allows us to have $\langle \chi \rangle = 0$ in a consistent manner.
- FN charge -1 would allow, e.g. $S^\dagger S^\dagger S X$ term in the potential.

$$y_i = f_i \left(\frac{\chi}{\sqrt{2} \Lambda_\chi} \right)^{2n_i} + y_i^{\text{SM}}$$

Yukawa variation at tree level - renormalisable potential

$$V = \frac{\mu_\phi^2}{2} \phi^2 + \frac{\lambda_\phi}{4} \phi^4 + \frac{\mu_\chi^2}{2} \chi^2 + \frac{\lambda_\chi}{4} \chi^4 + \frac{\lambda_{\phi\chi}}{4} \phi^2 \chi^2.$$

VEV conditions

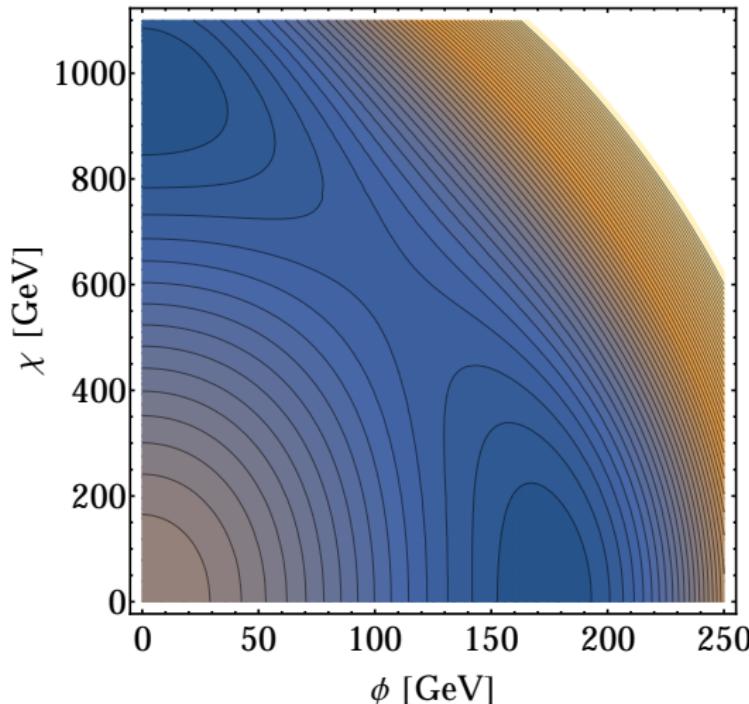
$$\begin{aligned}\mu_\chi^2 + \lambda_\chi v_\chi^2 &= 0, \\ \mu_\phi^2 + \lambda_\phi v_\phi^2 &= 0.\end{aligned}$$

Constraints

$$m_\chi^2 = \mu_\chi^2 + \frac{\lambda_{\phi\chi} v_\phi^2}{2} = -\lambda_\chi v_\chi^2 + \frac{\lambda_{\phi\chi} v_\phi^2}{2} > 0$$

$$\lambda_\chi < \lambda_\phi \left(\frac{v_\phi}{v_\chi} \right)^4 = 4.7 \times 10^{-4} \left(\frac{1 \text{ TeV}}{v_\chi} \right)^4$$

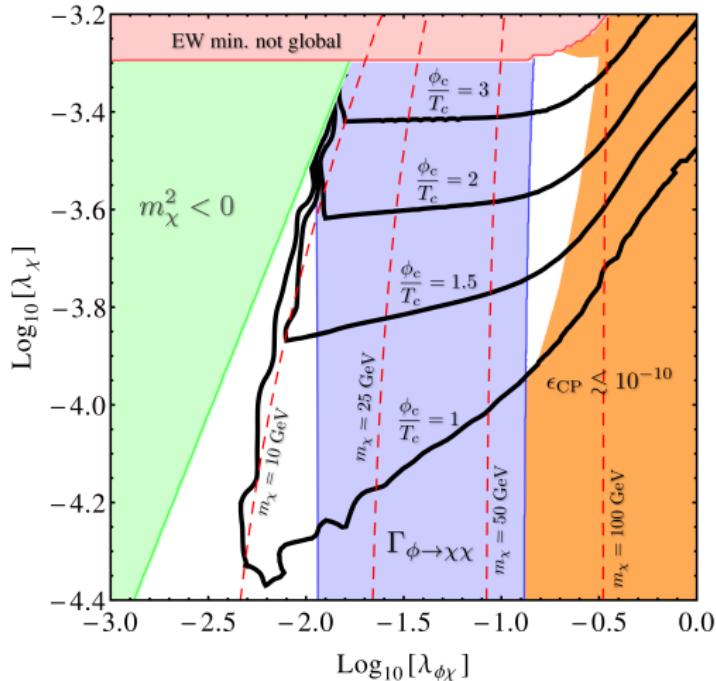
The finite T potential



$v_\chi = 1 \text{ TeV}$
 $\Lambda_\chi = 1 \text{ TeV}$
 $\lambda_\chi = 10^{-4}$
 $\lambda_{\phi\chi} = 10^{-2}$
 $m_\chi = 14 \text{ GeV.}$

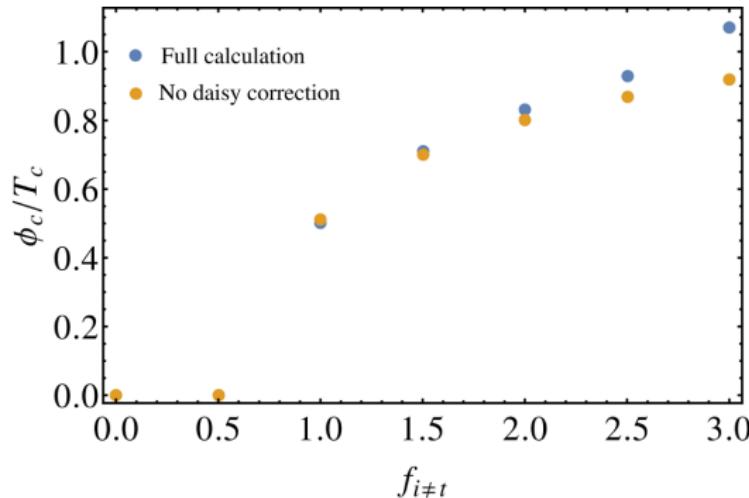
$T_c = 133 \text{ GeV}, \quad \phi_c = 174 \text{ GeV}, \quad \phi_c/T_c = 1.3.$

Parameter scan with $f_i = 1$, $v_\chi = \Lambda_\chi = 1$ TeV



Tree level barrier for: $\lambda_{\phi\chi} \geq -2 \frac{\mu_\phi^2}{v_\chi^2} = 1.56 \times 10^{-2} \left(\frac{\text{TeV}}{v_\chi} \right)^2$

The effect of the Yukawas

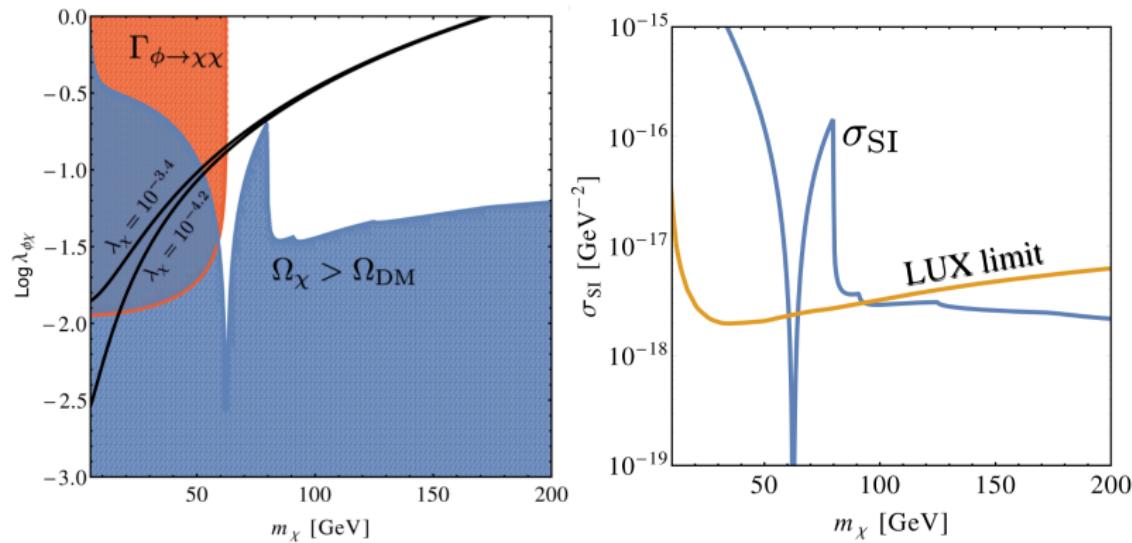


$$\begin{aligned}\nu_\chi &= 1 \text{ TeV} \\ \Lambda_\chi &= 1 \text{ TeV} \\ \lambda_\chi &= 10^{-4} \\ \lambda_{\phi\chi} &= 10^{-2.4} \\ m_\chi &= 14 \text{ GeV.}\end{aligned}$$

We find the varying Yukawas can change the nature of the phase transition near the $\phi_c/T_c = 1$ contour.

$$y_b = f_b \left(\frac{\chi}{\sqrt{2} \Lambda_\chi} \right)^4 + y_b^{\text{SM}}$$

Relic Abundance



The Z_2 symmetry may be softly broken, allowing χ to decay.
(Alternatively new annihilation channels may also be present.)

e.g. $\frac{X^\dagger X \overline{N^c} N}{\Lambda}$

Conclusions

The Higgs may have links to cosmology with experimentally accessible signatures.

Yukawa variation may allow us to address:

- The lack of a strong first order phase transition in the SM,
- The insufficient CP violation for EW baryogenesis,
- The related limits on EDMs (this approach leads to a lack of EDM signals).

This offers additional motivation to consider low scale flavour models and their cosmology.

Other models of flavour are worth looking at too (not just Froggatt-Nielsen).

New experimental signatures should then be accessible as we further probe the Higgs potential!