

R a PY from Candidate Materials for the FCC-hh Vacuum System.

R. Cimino¹, E. La Francesca,^{1, 2} M. Angelucci,¹ A. Liedl,¹ L. A. Gonzalez,^{1, 3} I. Bellafont,^{3, 4} F. Siewert,⁵ M.G. Sertsu,⁵ A. Sokolov,⁵ and F. Schäfers⁵

¹ LNF-INFN, Via E. Fermi 40, Frascati (Rome) Italy.

² Università di Roma "La Sapienza", Rome, Italy

³ CERN, Geneva, Switzerland.

⁴ ALBA Synchrotron Light Source, Barcelona, Spain

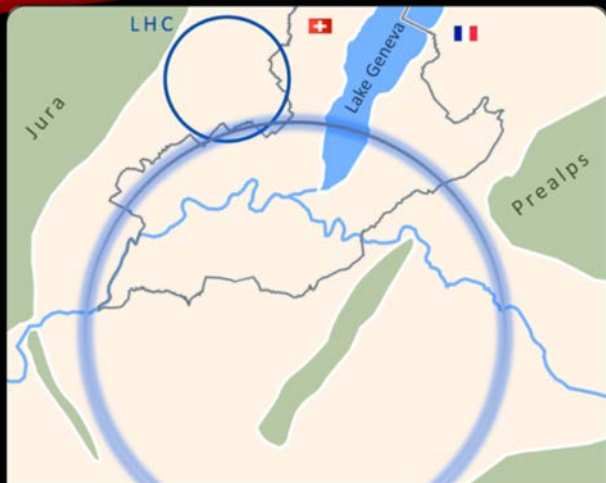
⁵ Helmholtz-Zentrum-Berlin, Berlin, Germany

outline:

- The problem
- The experimental approach
- Selected results
- Conclusion

outline:

- The problem
- The experimental approach
- Selected results
- Conclusion



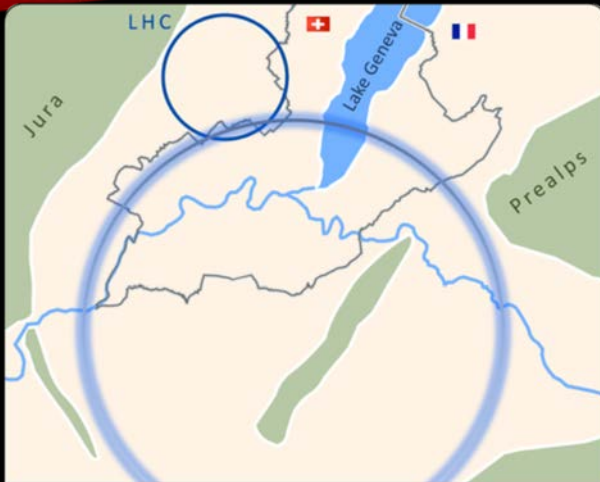
Schematic of an 80 – 100 km long tunnel

The case of FCC-hh

FCC-hh Key Parameters

<i>Version 1.0 (2014-02-11)</i>	LHC	HL-LHC	FHC-hh
c.m. Energy [TeV]	14		100
Circumference C [km]	26.7		100 (83)
Dipole field [T]	8.33		16 (20)
Injection energy [TeV]	0.45		3.3
Peak luminosity [$10^{34} \text{ cm}^{-2}\text{s}^{-1}$]	1.0	5.0	5.0
Stored beam energy [GJ]	0.392	0.694	8.4 (7.0)
SR power per ring [MW]	0.0036	0.0073	2.4 (2.9)
Arc SR heat load [W/m/aperture]	0.17	0.33	28.4 (44.3)
Critical photon energy [keV]	0.044		4.3 (5.5)



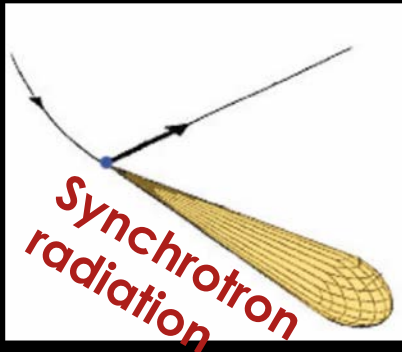


Schematic of an 80 – 100 km long tunnel

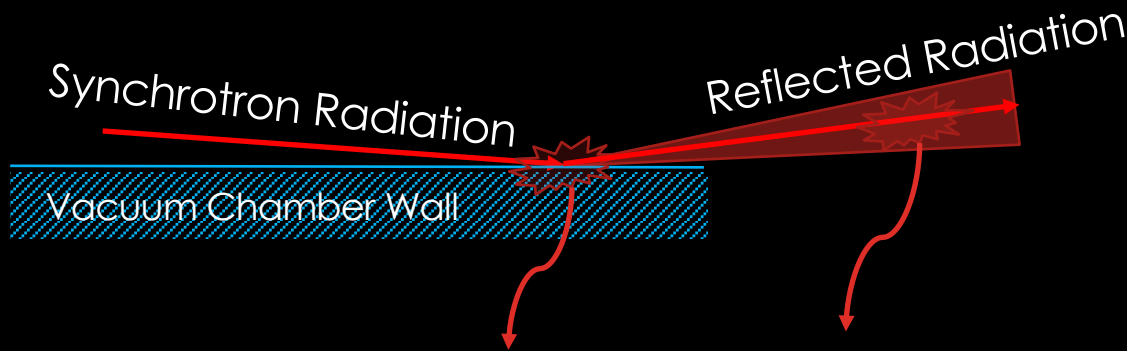
FCC-hh Key Parameters

Cryogenic Operating Temperature (1.9 K at the dipole walls)

Version 1.0 (2014-02-11)	LHC	HL-LHC	FCC-hh
c.m. Energy [TeV]	13	14	100
Circumference C [km]	27	27	100 (83)
Dipole field [T]	8.33	8.33	16 (20)
Injection energy [TeV]	0.45	0.45	4.5
Peak luminosity [$10^{34} \text{ cm}^{-2} \text{ s}^{-1}$]	1.0	5.0	5.0
Stored beam energy [GJ]	0.36	0.694	8.4 (7.0)
SR power per ring [MW]	0.0036	0.0073	2.4 (2.9)
Arc SR heat load [W/m/aperture]	0.17	0.33	28.4 (44.3)
Critical photon energy [keV]	0.044	0.044	4.3 (5.5)



$$\text{Incident Radiation} = \text{Transmitted} + \text{Reflected} + \text{Absorbed}$$

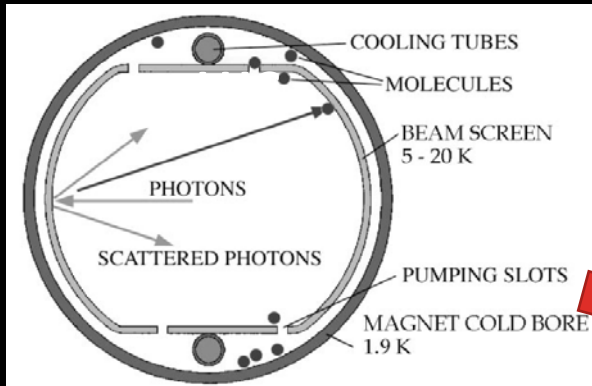


- Absorbed Radiation
- Heat Load
 - Photo Electrons
 - Photo Stimulated Desorption
- On the orbit plane

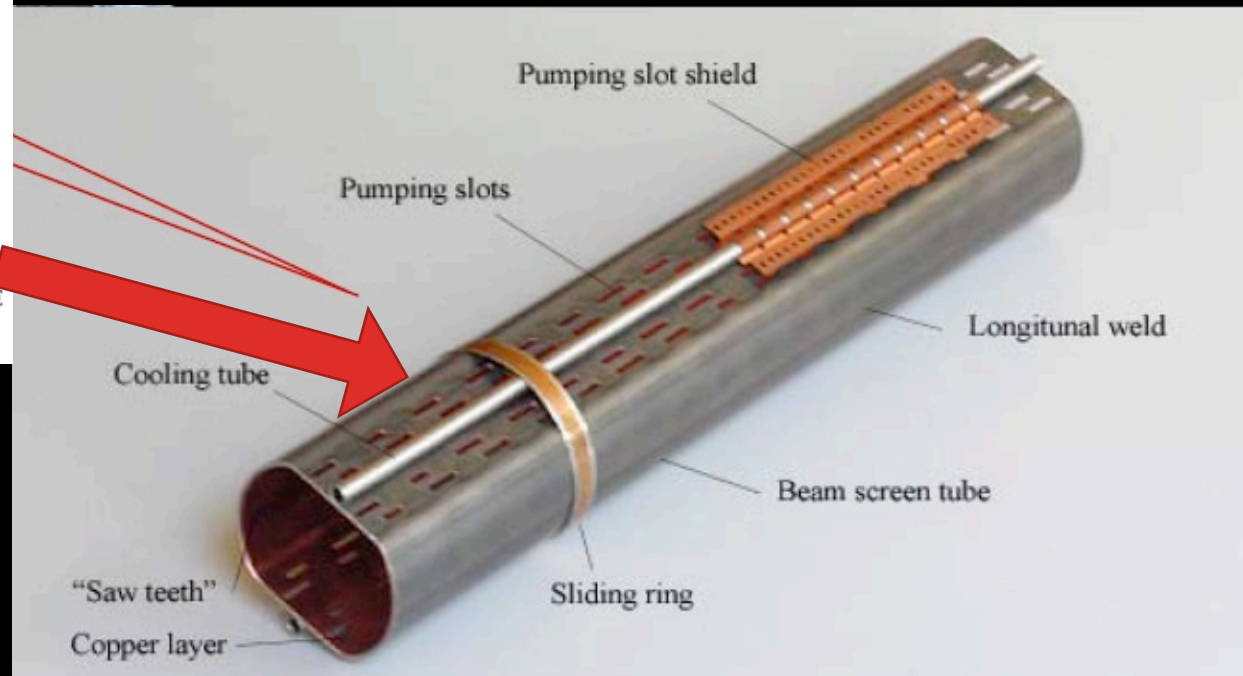
- Reflected Radiation
- Heat Load
 - Photo Electrons
 - Photo Stimulated Desorption
- ALL THIS IN OTHER PARTS OF THE MACHINE

From LHC Beam screen concept (5 K < T < 20 K)

F. Zimmermann et al., HF2014, Beijing, China (THP3H1)



43.6 mm



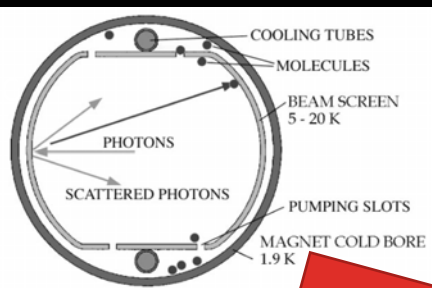
O. Gröbner, Vacuum 60 (2001) 25-34

SR = 0.17 W/m/ap
Tot. Power = 0.007 MW

LHC Beam screen concept (5 K < T < 20 K)



To FCC-hh Beam screen optimization (40 K < T < 60 K)

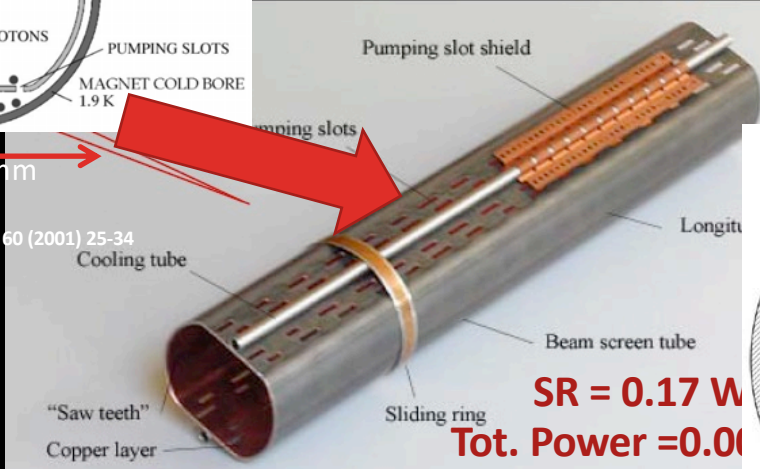


F. Zimmermann et al., HF2014, Beijing, China (THP3H1)

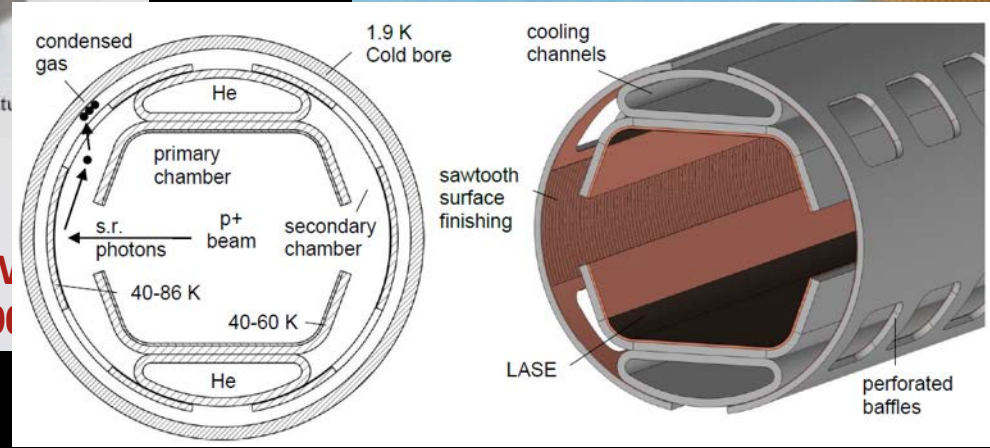


43.6 mm

O. Gröbner, Vacuum 60 (2001) 25-34



SR = 0.17 W
Tot. Power = 0.00



See: EuroCirCol-P2-WP4-D4.4_Analysis of beam-induced vacuum effects



Two main properties of material needs to be carefully investigated:

Reflectivity (R)

The part of incident radiation which is reflected by the wall chambers

Photo Yield (PY)

How many electrons are produced per incident photon

These properties depend on:

- Incident Angle: θ_i
- Energy of Radiation
- Chemical Composition of the Surface
- Surface Treatment
- Coating

These properties depend on:

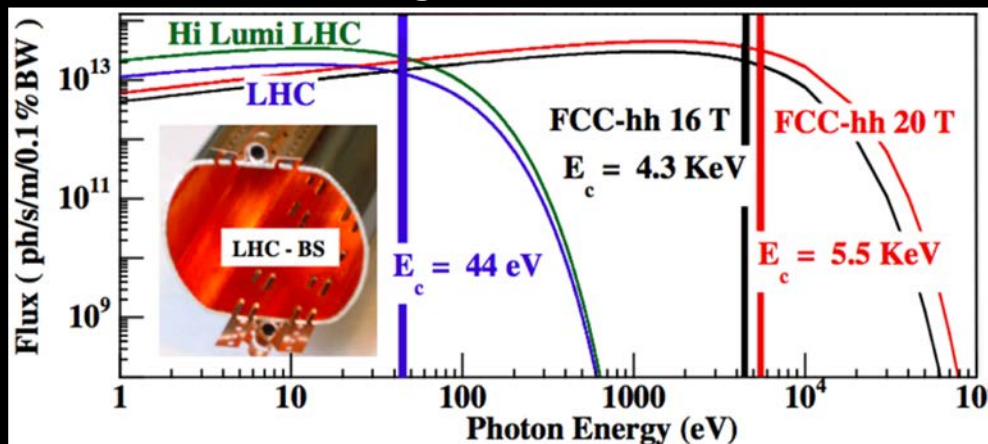
- Incident Angle: θ_i
- Energy of Radiation
- Chemical Composition of the Surface
- Surface Treatment
- Coating

Very Grazing angles:

LHC $\sim 0.28^\circ$

FCC-hh $\sim 0.08^\circ$

Radiation Energy Spectrum: from eV to several keV



R. Cimino, V. Baglin and F. Schäfers, PRL. 115 (2015) 264804

These properties depend on:

- Incident Angle: θ_i
- Energy of Radiation
- Chemical Composition of the Surface
- Surface Treatment
- Coating

Very Grazing angles:

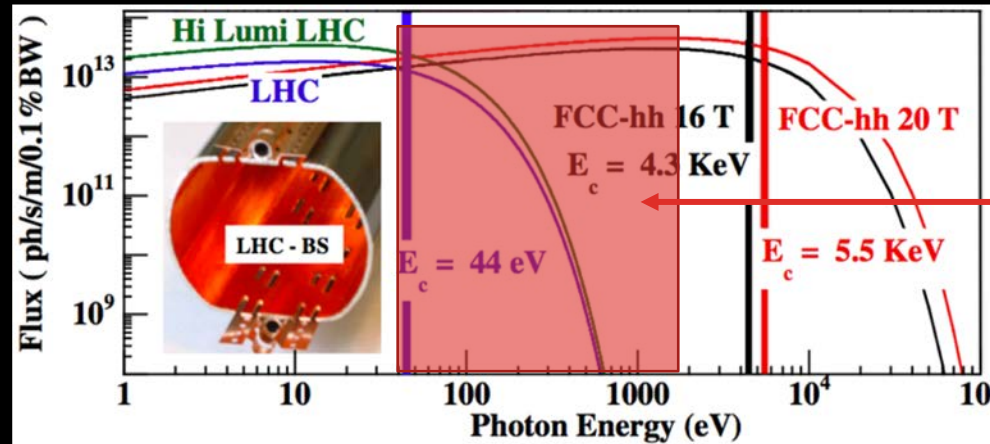
LHC $\sim 0.28^\circ$

FCC-hh $\sim 0.08^\circ$

There are indications suggesting that **R**, and **Photo induced desorption** are more significant in the lower energy part of the spectrum.

PY should slowly increase with photon energy.

Radiation Energy Spectrum: from eV to several keV

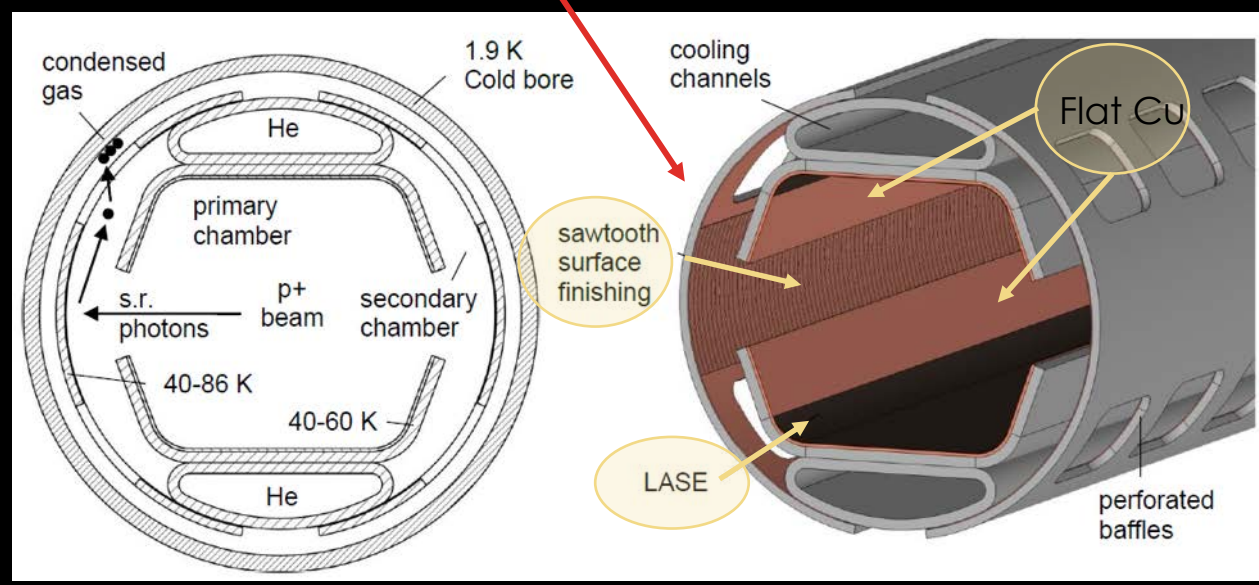


We first investigate the 35-1850 eV region

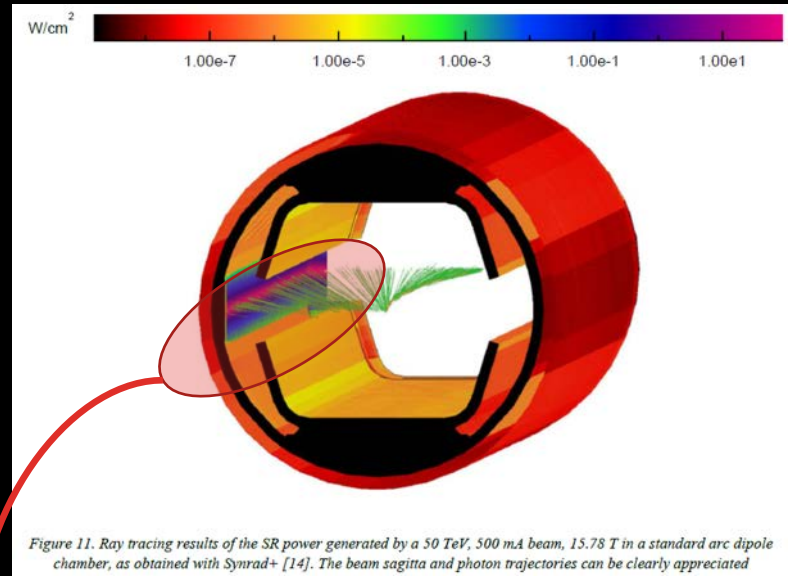
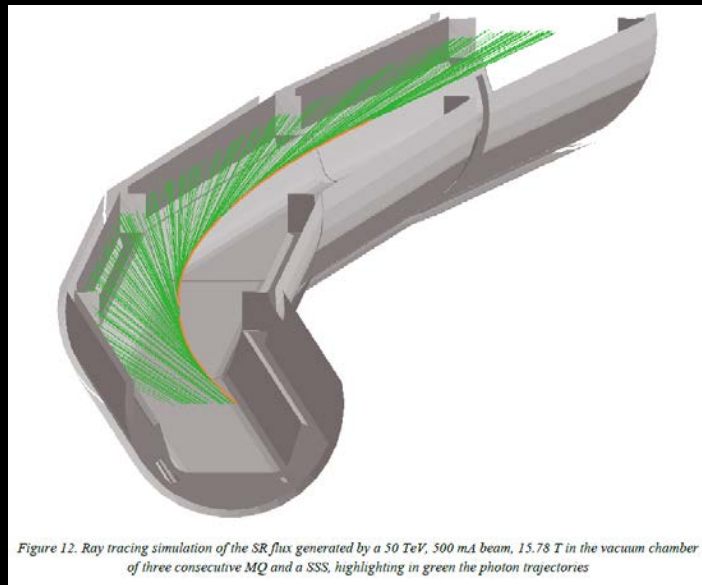
R. Cimino, V. Baglin and F. Schäfers, PRL. 115 (2015) 264804

These properties depend on:

- Incident Angle: θ_i
- Energy of Radiation
- Chemical Composition of the Surface
- Surface Treatment
- Coating



“Ray tracing” programs are at the base of the design phase and generally use R simulated values (from Reflectivity programs)



Thanks to:
I. Bellafont

- ❖ Are simulated R data good enough?
- ❖ Or should we measure R and PY for all proposed materials in realistic geometrical conditions to validate all simulations?

Usually, analytical programs are used for optical calculation
 To calculate **SPECULAR REFLECTIVITY**

R (specular reflectivity) is the normalized number of photons emerging from the sample with the same divergence of the incoming beam and at the geometrical reflection

R_t (total reflectivity) is the normalized (to the incoming flux) number of photons emerging from the sample in all directions.

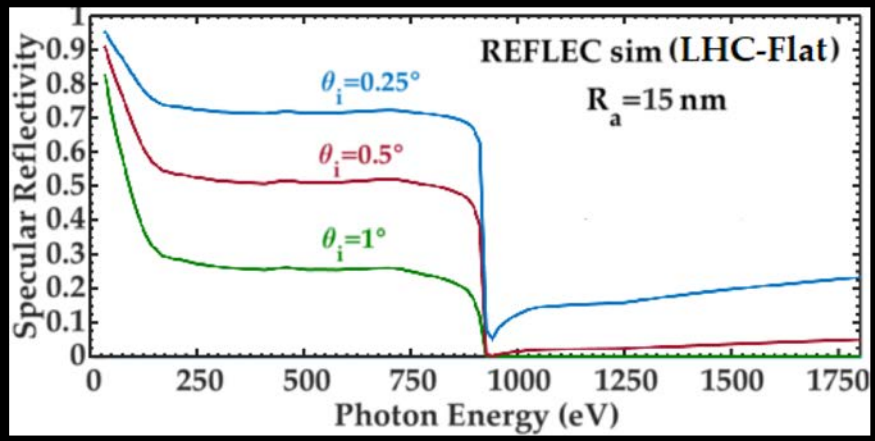
In machine simulations both R and R_t are important.

In optical simulation codes R_q is a considered as a (Debye-Waller) perturbation. It does not work for:

- high roughness technical surfaces.
- high photon energies
- grazing angles
- also, at very grazing angles ->

contaminants should play a major role (difficult to predict)

- Different Surface Geometry -> Saw Tooth profile (ST)
- Treated Surface -> Laser Treated Samples (LASE)



REFLEC simulations of Specular reflectivity VS Photon Energy at three incidence angle for a LHC-Flat Cu sample vs. photon energy

Experimental data are needed!

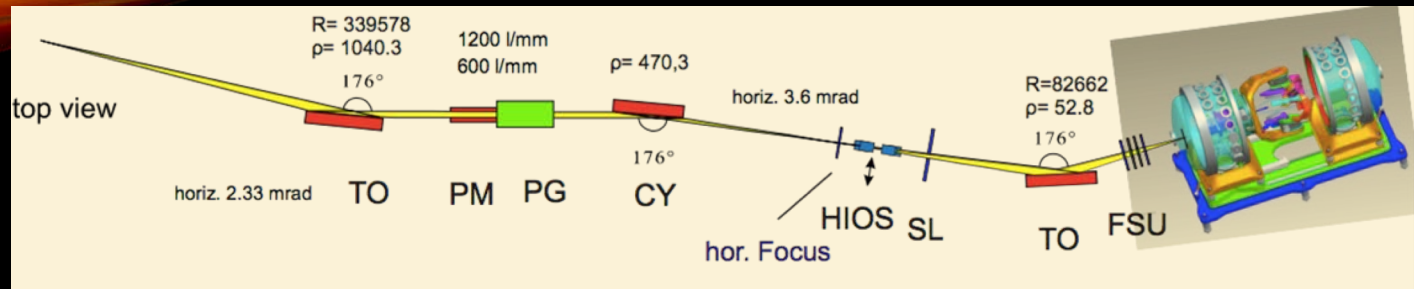


outline:

- The problem
- The experimental approach
- Selected results
- Conclusion

Strategy:

- Identified a close to optimal experimental set-up to perform Reflectivity and Photo Yield studies on technical materials of interest to FCC-hh at very grazing angle and in the wider energy range available.
- LNF launched a long term proposal (MICA) and received support by INFN and beamtime by the project **CALIPSOplus**. (under the Grant Agreement 730872 from the EU Framework Programme for Research and Innovation HORIZON 2020)
- In 2015 – 2018 an intense experimental campaign at the Synchrotron radiation Facility BESSY-2 using the Optics beamline and reflectometer (with guest experimentalists).



A. Sokolov, et al., Journal of Synchrotron Radiation 25, 100 (2018).

Experimental Parameters

- Phot. Energy range 35 ÷ 1800 eV
- Beam height $h=0.3$ mm
- Incident Beam measurement
- GaAsP Photodiodes (4x4mm) (0.1*4mm)
- Low Incident angles down to 0.25°
- High second order suppression

The Beamline

The Reflectometer

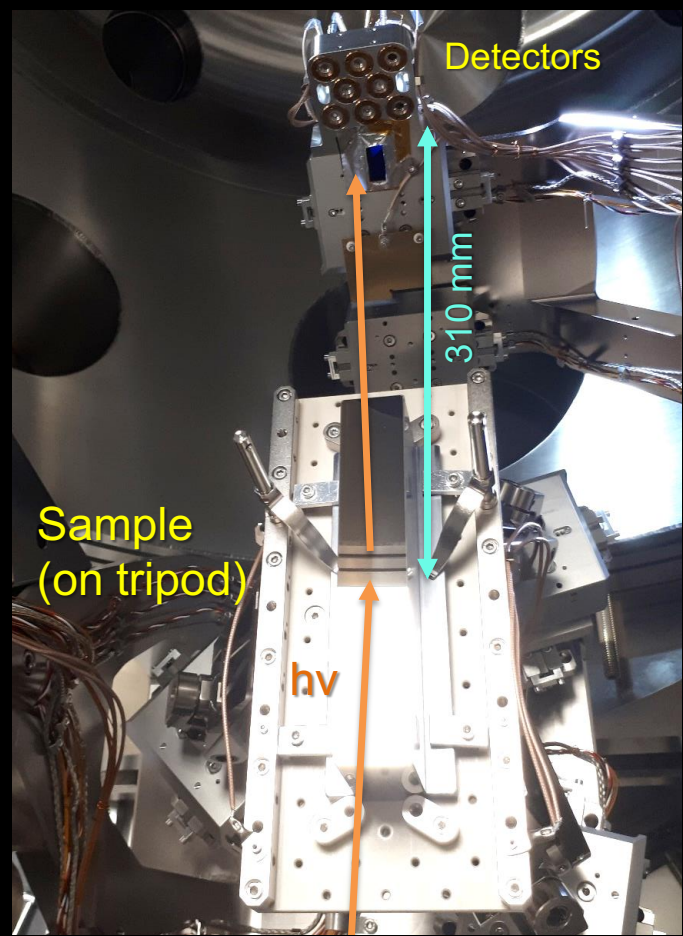
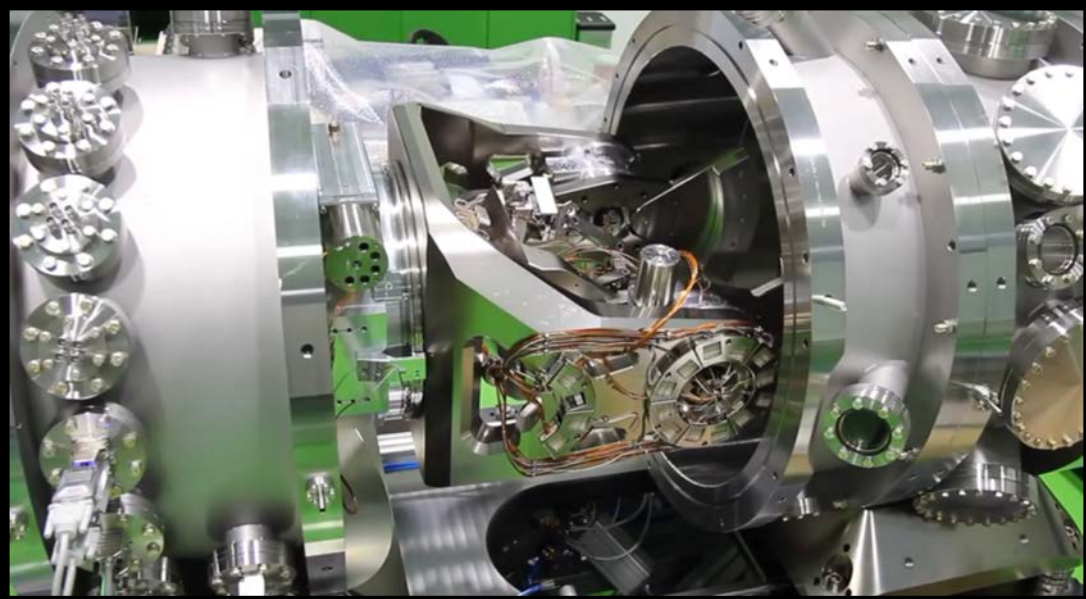
Axis	Hardware	Range	Pos. accuracy
Azimuth angle β	HUBER 430	$-180^\circ - 180^\circ$	3.6"
Sample angle θ	HUBER 411	$-90^\circ - 270^\circ$	3.6"
Detector angle 2θ	HUBER 411	$-180^\circ - 180^\circ$	3.6"
Detector off-plane (2 axes)	Ceramic motors	-25 mm – 25 mm ($-4^\circ - 4^\circ$)	50 nm
Sample Adjustment T_x, T_y, T_z	Ceramic motors	-20 mm – 20 mm (not simul.)	500 nm
Sample Adjustment R_x, R_y, R_z	Ceramic motors	$-10^\circ - 10^\circ$ (not simul.)	1"

A.A.Sokolov, et al, Proc. of SPIE92060J-1-13(2014)



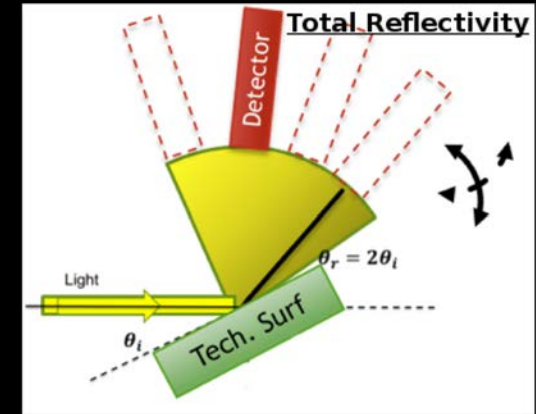
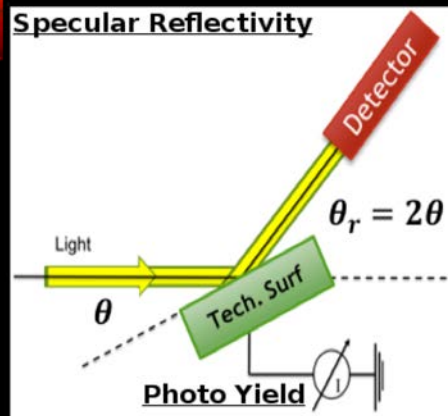
BESSYII – OPTICS Beamline end station

See movie at: https://www.helmholtz-berlin.de/pubbin/igama_output?modus=datei&did=887



REFLECTOMETER

Station dedicated to the at wavelength characterization of precision gratings and nano-optical devices



$\theta/2\theta$ geometry allows to detect the specular reflected signal.
Two possible way

Simultaneously measure the current generated on the sample is detected for **PY**

Fixed $\theta/2\theta$
 Beam Energy Scan

Fixed Beam Energy
 $\theta/2\theta$ Scan

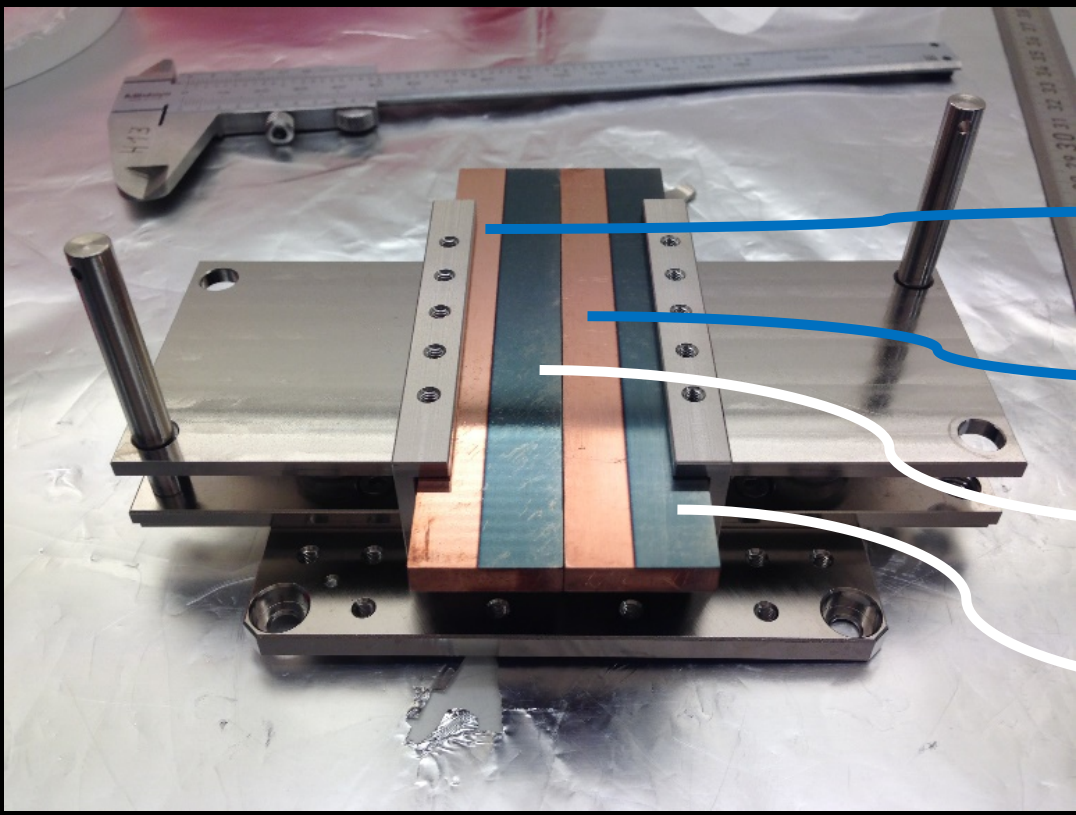
$$PY = \frac{N_e}{N_\gamma}$$

$$R = \frac{I}{I_0}$$

Fixed photon energy and incident angle
 Moving the detector in θ_r and Φ , the **angular distribution** of the Reflected signal is mapped.

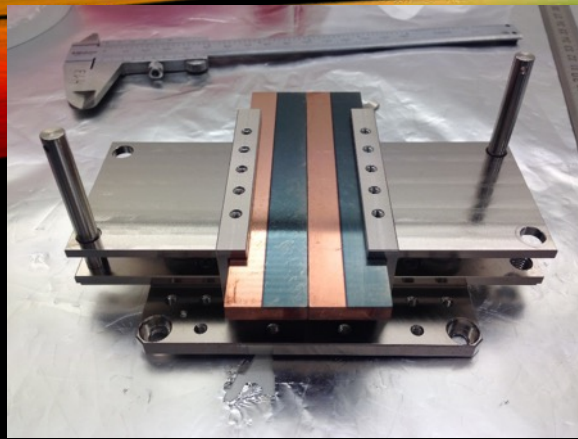
To study the impact of Ra on reflectivity

AFM data



Sample	RMS Roughness (R_a)
Cu 1A	10 nm
Cu 2A	27 nm
Cu 1B	25 nm
Cu - LHC	15 nm
Cu 1A CC	13 nm
Cu 2A CC	28 nm
Cu 1B CC	32 nm
Cu - LHC CC	20 nm

Cu (different Roughness)
Cu+ α -Carbon Thin Film (50 nm)



The sample studied:
Cu (different Roughness)

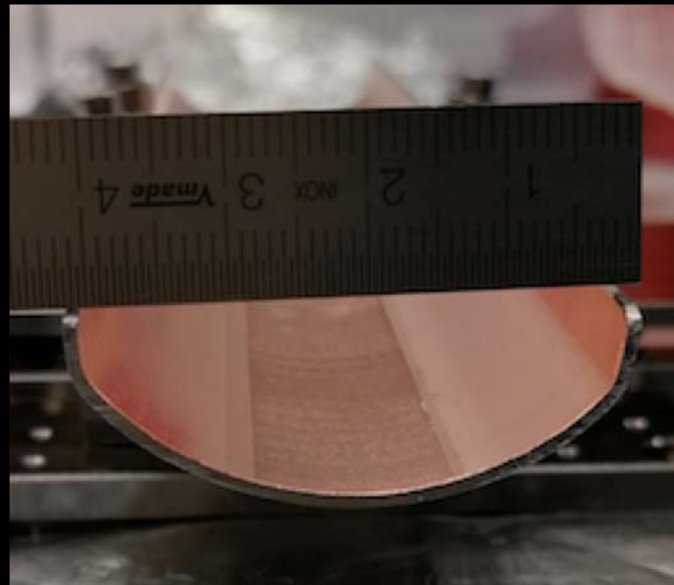
LHC - Saw Tooth

LASE

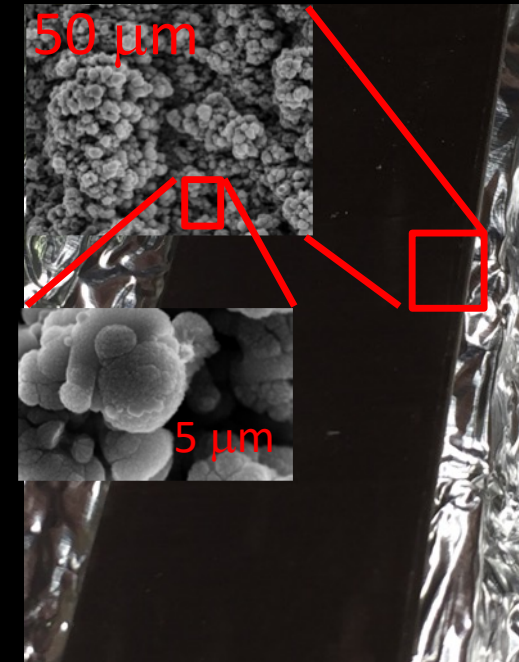
Cu+ amorphous Carbon Thin Film
and: NEG, Stainless Steel, ...



Cu LHC



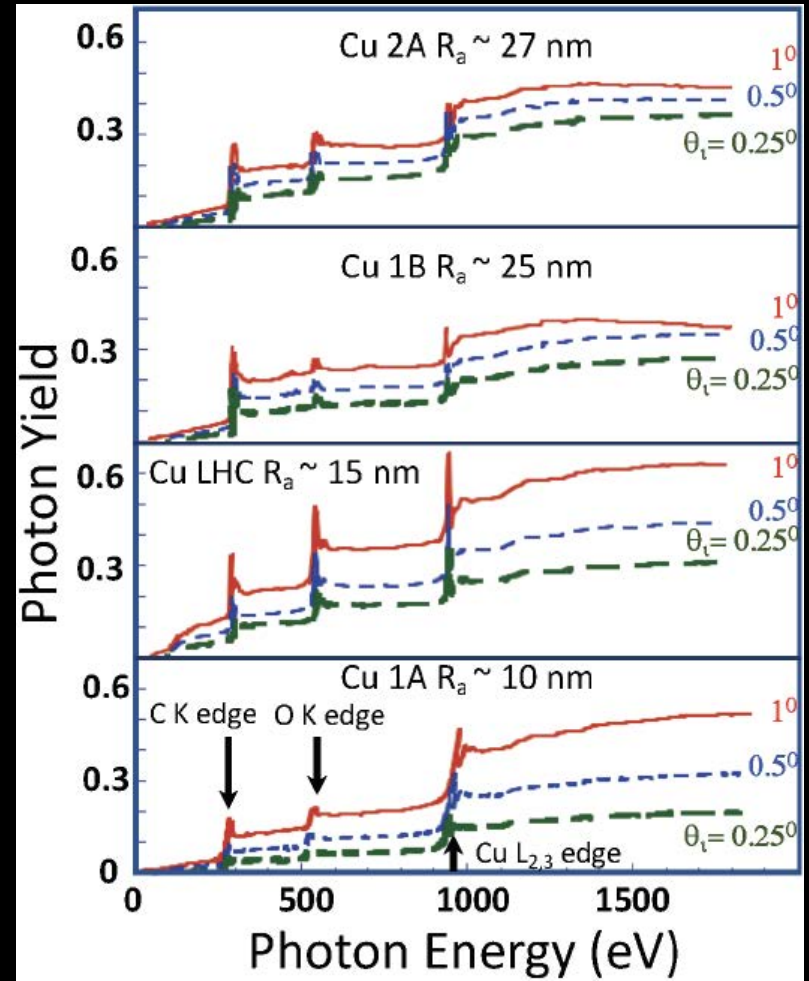
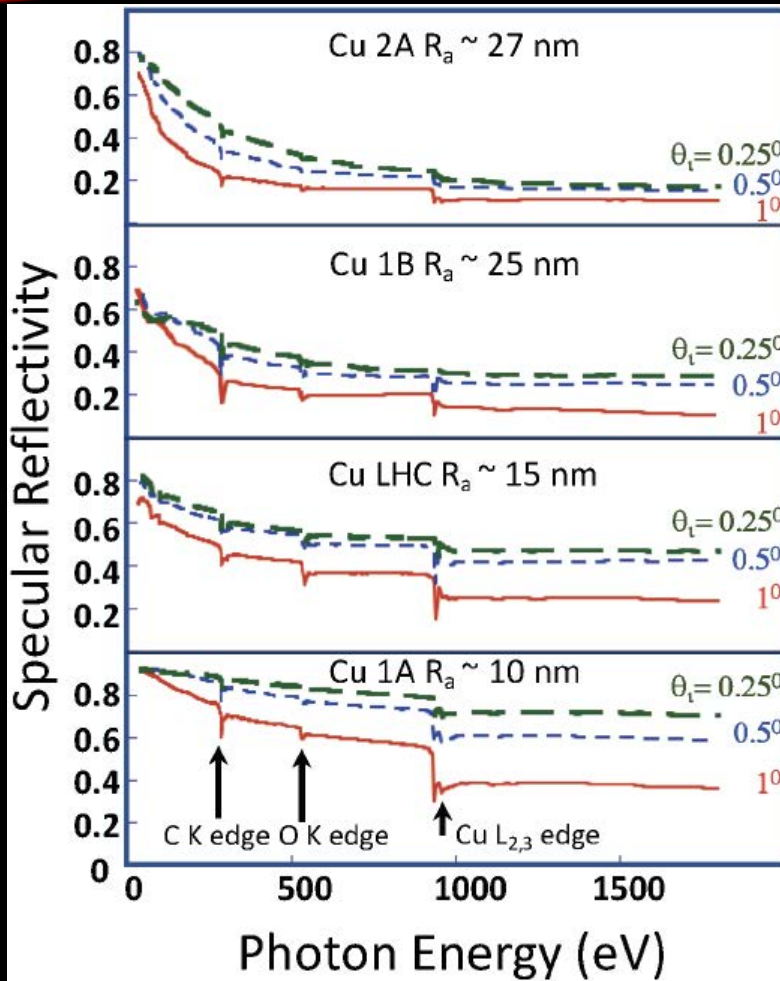
LHC Saw Tooth



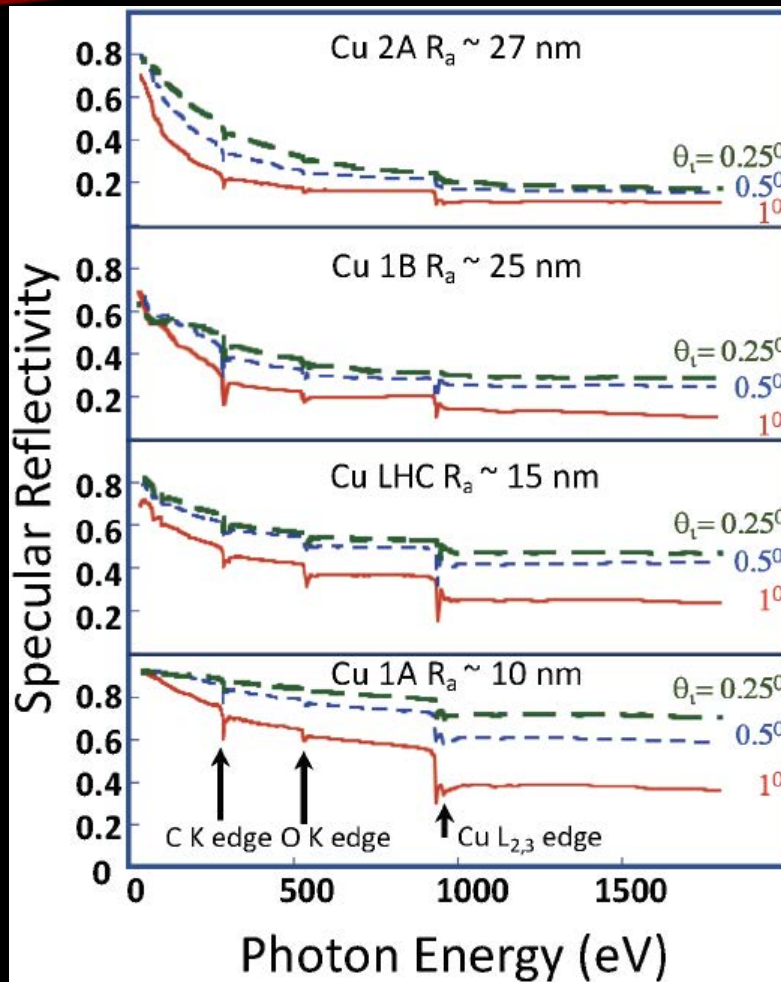
LASE Cu

outline:

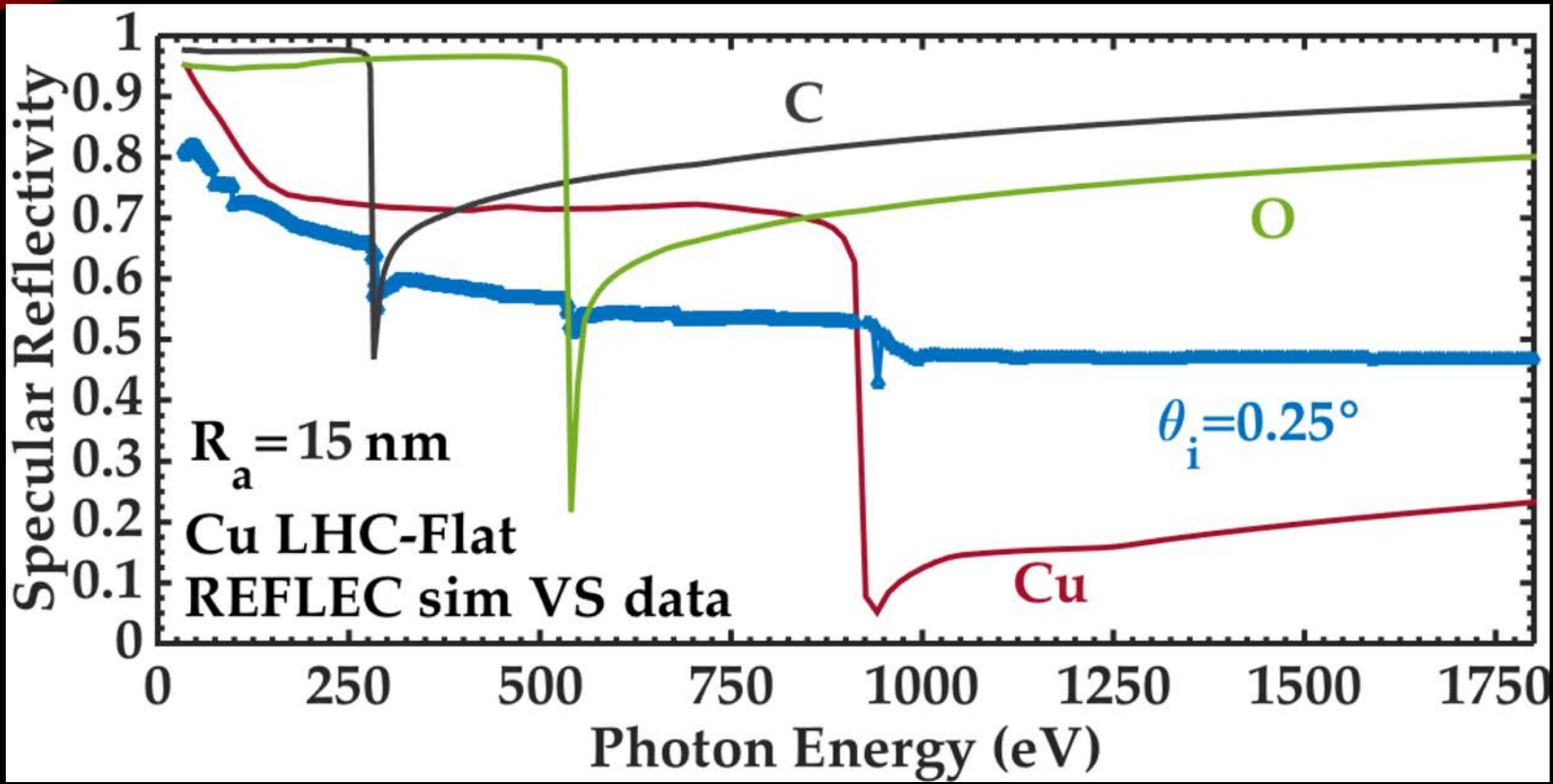
- The problem
- The experimental approach
- **Selected results**
- Conclusion



Specular Reflectivity and Photo Yield



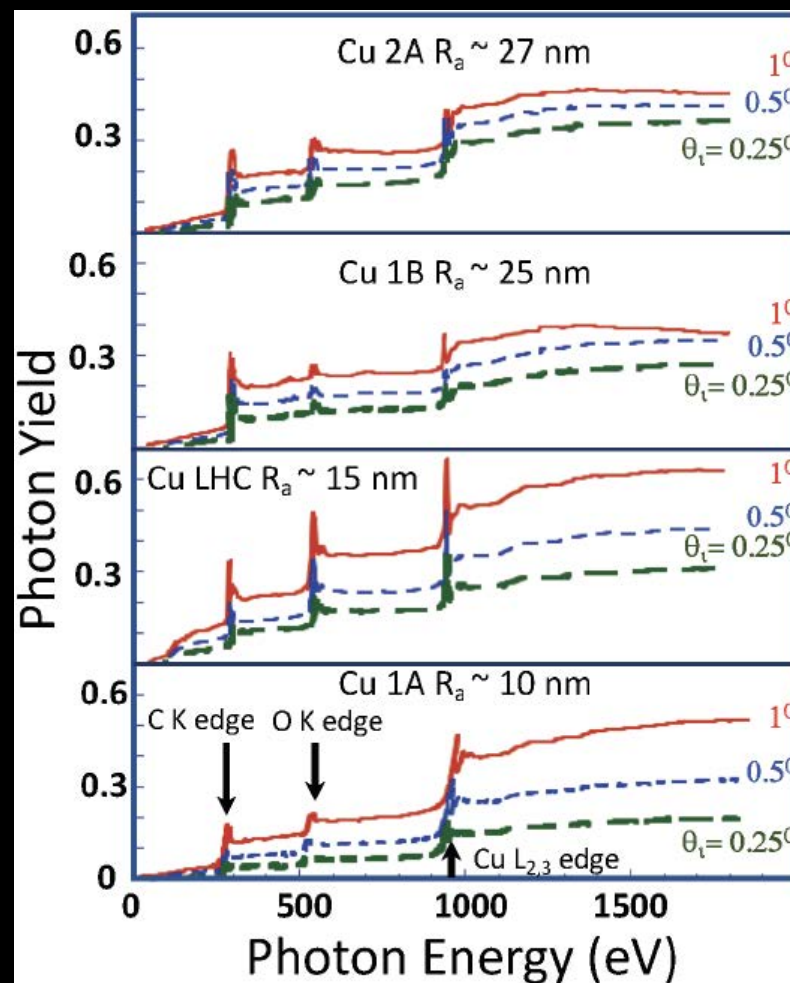
- In all cases, R is higher at lower photon energies.
- In all cases, R is higher at lower angle of incidence θ_i .
- Reflectivity, after the observed Cu- L_{2-3} absorption edge at 930-950 eV is much higher than in simulation.
- In all spectra we measure a significant effect due to the absorption edges of C K-edge at 280 eV and O K edge at 530 eV (surf. cont.)
- C and O K absorption edges are more effective at lower incidence angles.
- Roughness, as expected, plays a major role in determining the ability of a surface to specularly reflect impinging photons.



Comparison of REFLEC simulations and experimental data of Specular Reflectivity VS Photon Energy of Cu LHC sample ($\theta_i = 0.25$). To understand the role of air contaminants simulations for Carbon and Oxygen are also reported.

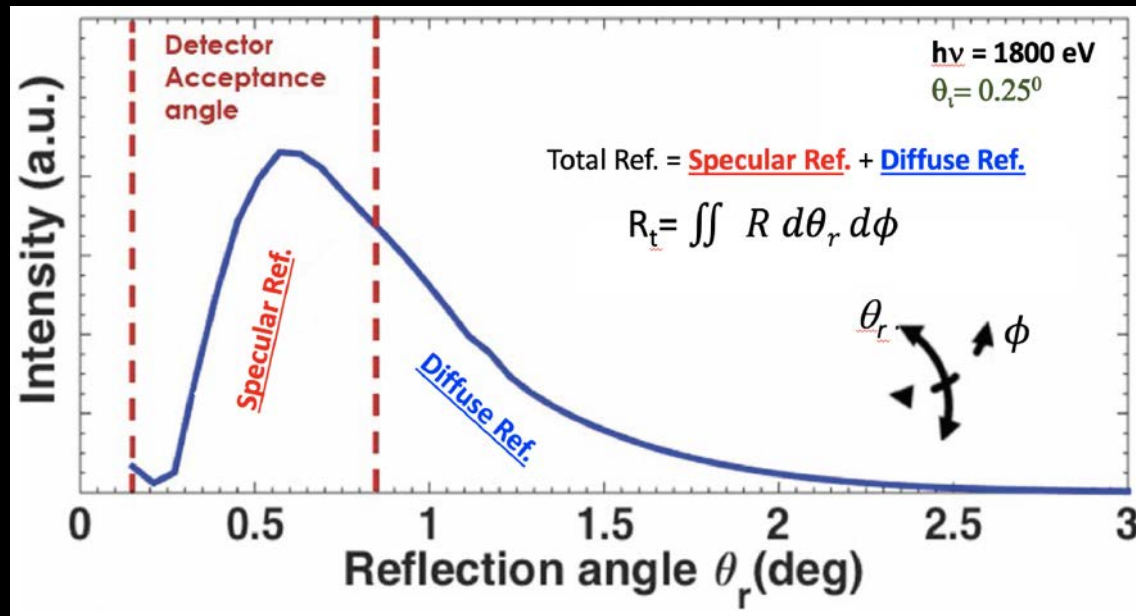
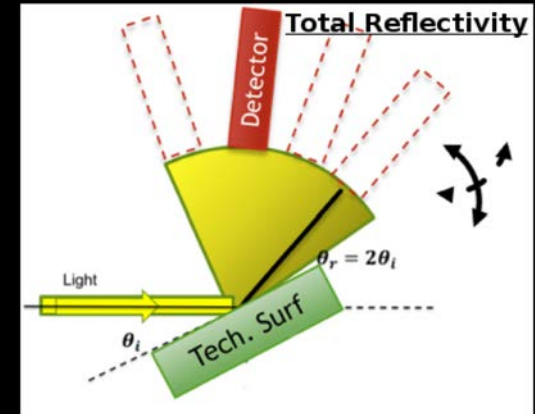


- In all cases, PY is higher at higher photon energies.
- The PY dependence on θ_i is consistently dimmed and finally washed out when surface R_a is increasing.
- In all cases, the Cu-L2-3 absorption edge at 930-950 eV is visible and cause an increase in the measured PY.
- In all spectra we measure a significant effect due to the C K-edge at 280 eV and O K-edge at 530 eV (surf. Contaminants)
- Roughness does influence the PY. The lower is R_a , the highest is the measured PY. Also because Reflected photons do not produce photoelectrons!



Angular Distribution of Reflectivity

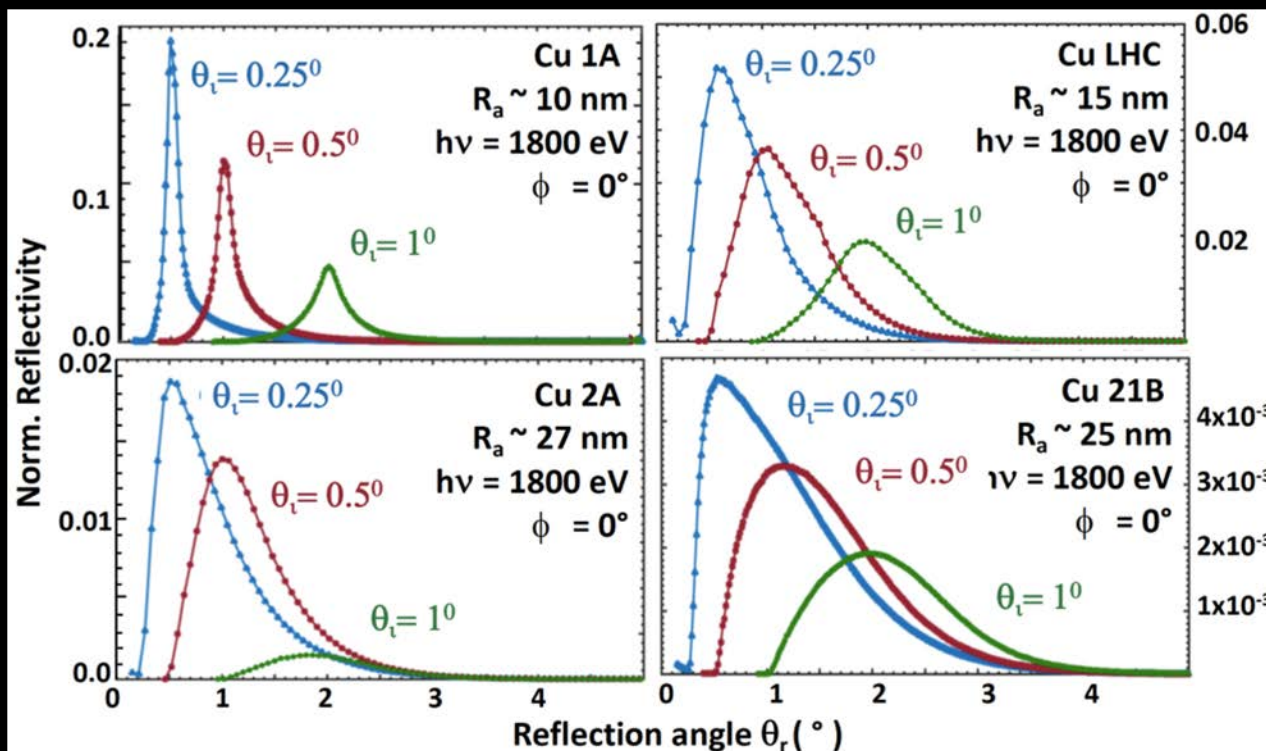
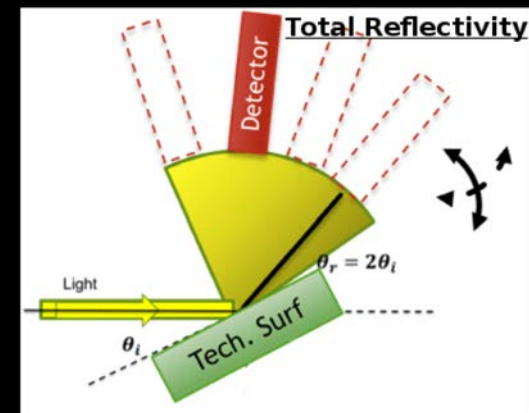
Incident angle } Fixed
 photon Energy }
 The detector moves over the reachable Solid Angle $[\Phi \theta]$



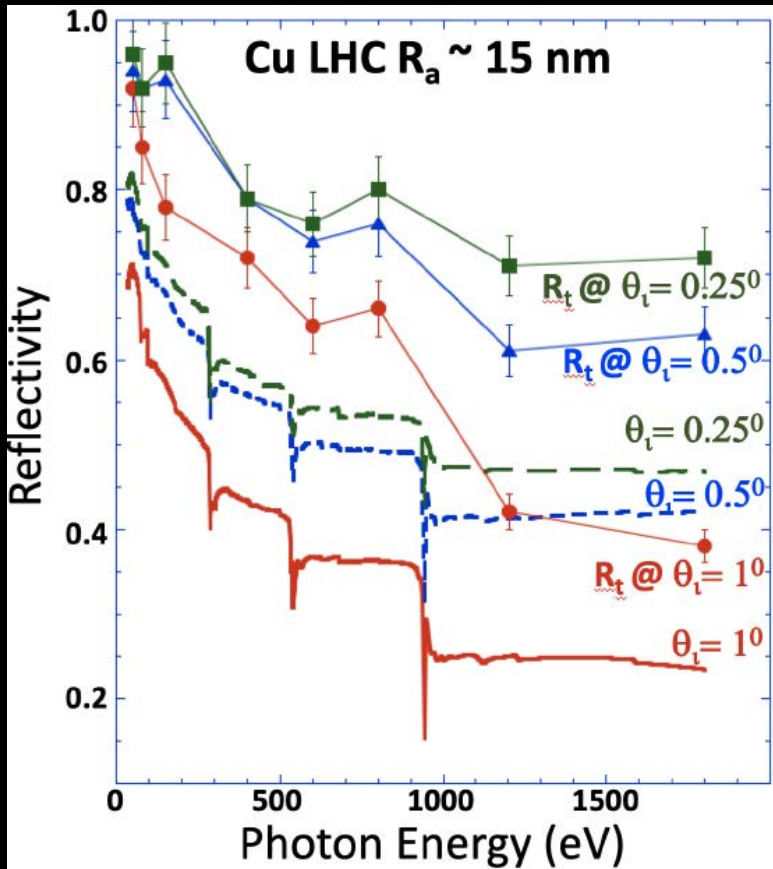
Total Reflectivity is calculated integrating the angular distribution in θ_r and Φ

Angular Distribution of Reflectivity

Reflectivity Angular Distribution of Samples With different Roughness



Specular Reflectivity VS Total Reflectivity

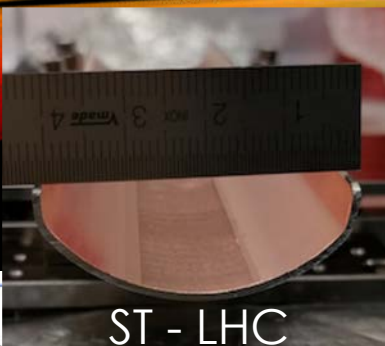


Photon energy $h\nu$ (eV)	Specular Reflec. $\theta_i = 0.25^\circ$ ($\Delta R/R = \pm 2\%$)	Total Reflec. $\theta_i = 0.25^\circ$ ($\Delta R_t/R_t = \pm 5\%$)
1800	0.47	0.72
1200	0.47	0.71
800	0.53	0.80
600	0.54	0.76
400	0.59	0.79
150	0.71	0.95
80	0.75	0.92
50	0.81	0.96



E. La Francesca et al: submitted to PR ST



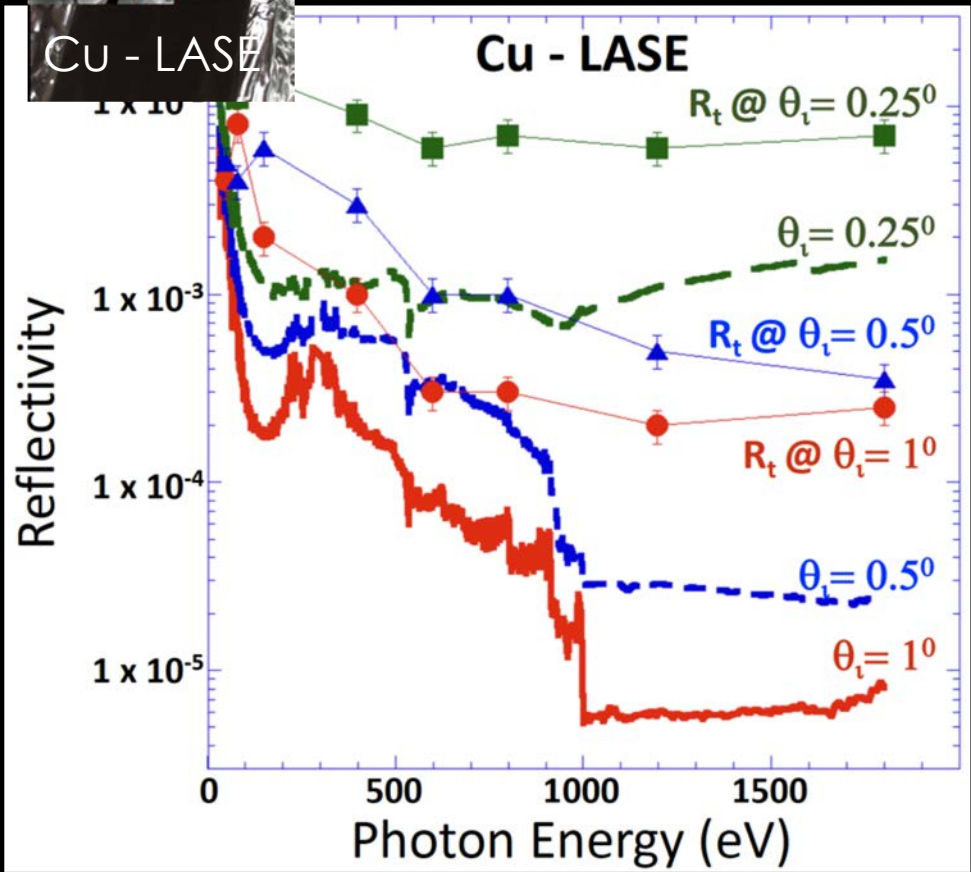
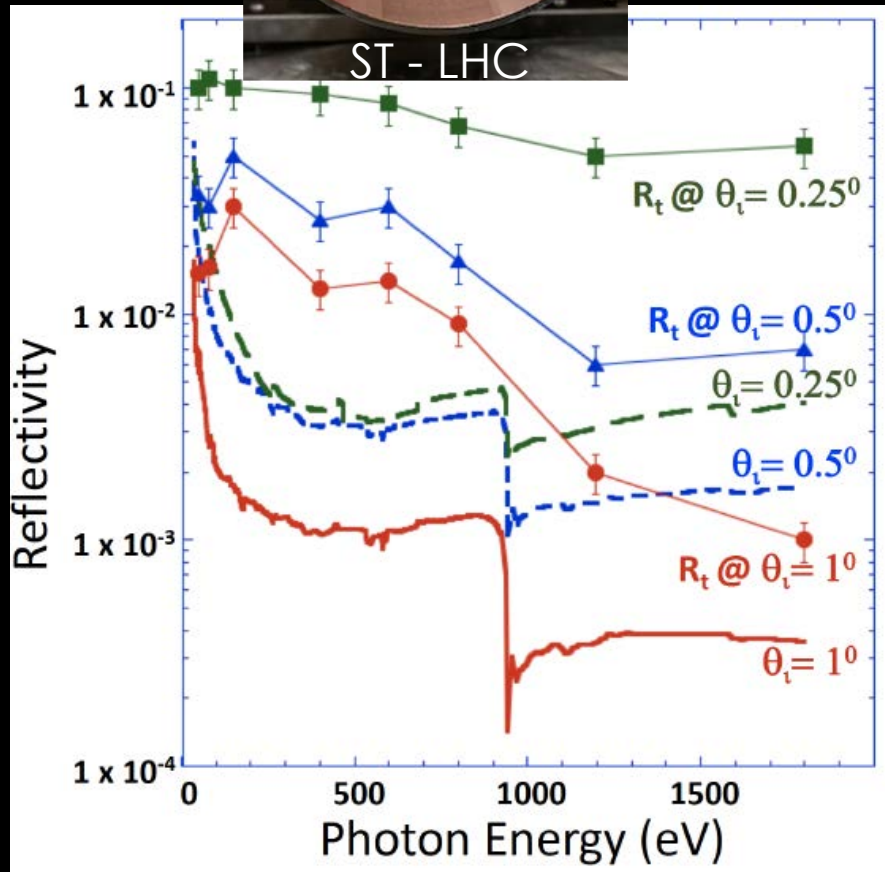


ST - LHC



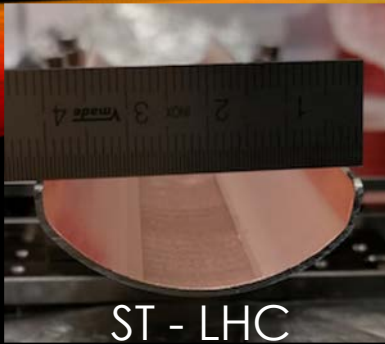
Cu - LASE

Specular Reflectivity VS Total Reflectivity



E. La Francesca et al: submitted to PR ST





Specular Reflectivity VS Total Reflectivity

Photon energy $h\nu$ (eV)	Specular Reflec. $\theta_i = 0.25^\circ$ ($\Delta R/R = \pm 2\%$)	Total Reflec. $\theta_i = 0.25^\circ$ ($\Delta R_t/R_t = \pm 10\%$)
1800	0.004	0.05
1200	0.003	0.05
800	0.0045	0.07
600	0.0035	0.08
400	0.004	0.1
150	0.01	0.10
80	0.02	0.11
50	0.03	0.10

Photon energy $h\nu$ (eV)	Specular Reflec. $\theta_i = 0.25^\circ$ ($\Delta R/R = \pm 2\%$)	Total Reflec. $\theta_i = 0.25^\circ$ ($\Delta R_t/R_t = \pm 20\%$)
1800	0.0015	0.007
1200	0.0011	0.006
800	0.0009	0.007
600	0.0009	0.006
400	0.001	0.009
150	0.001	0.014
80	0.003	0.011
50	0.006	0.012

We can generate experimental values that can be mediated and used in realistic simulations.

E. La Francesca et al: submitted to PR ST

Do alternatives to optimization of HL removal from BM cold masses exist?

PRL 115, 264804 (2015)

PHYSICAL REVIEW LETTERS

week ending
31 DECEMBER 2015

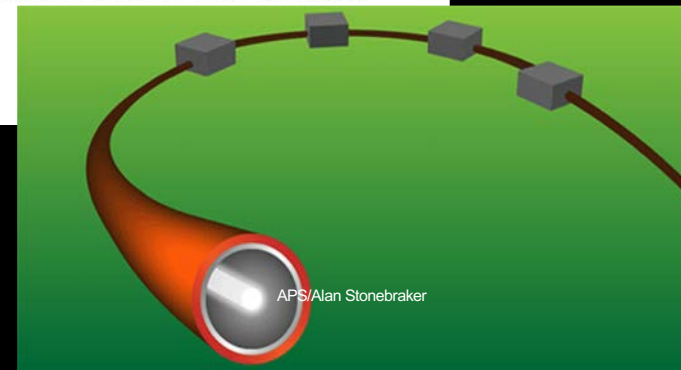
Potential Remedies for the High Synchrotron-Radiation-Induced Heat Load for Future Highest-Energy-Proton Circular Colliders

R. Cimino,^{1,2,*} V. Baglin,² and F. Schäfers³



From:

Synopsis: Cooler Colliders



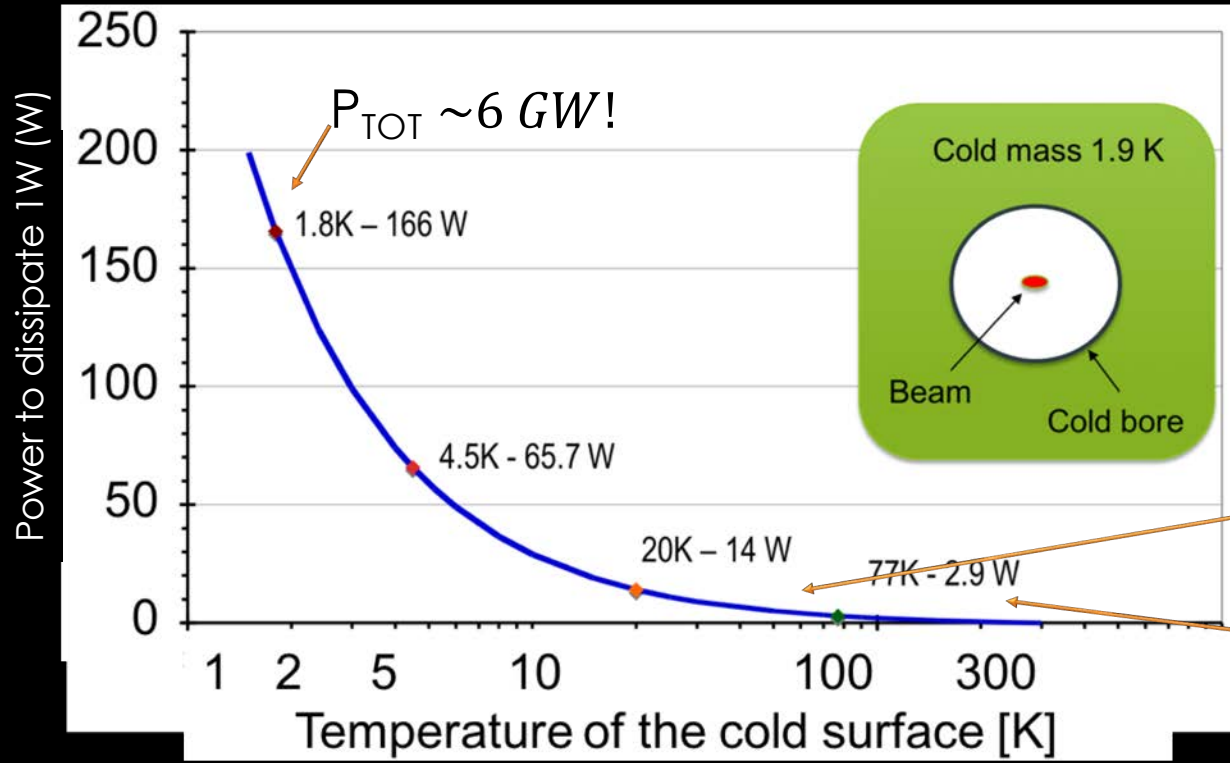
..... The proposal is to coat the interior of the copper tube with a thin layer of carbon that reflects all the incident radiation. the radiation, and the heat it carries, is transported away from colder regions towards periodically placed room-temperature absorbers, which are easier and cheaper to cool than the tube itself. **The authors claim that this design would reduce the power consumption potentially cutting the associated costs in half.**

(– Katherine Wright, APS, Physics)



Heat Load Dissipation VS Temperature

**Some Numbers:
(Power → Feasibility → Sustainability!)**



@ 60K $P_{TOT} \sim 100 MW$

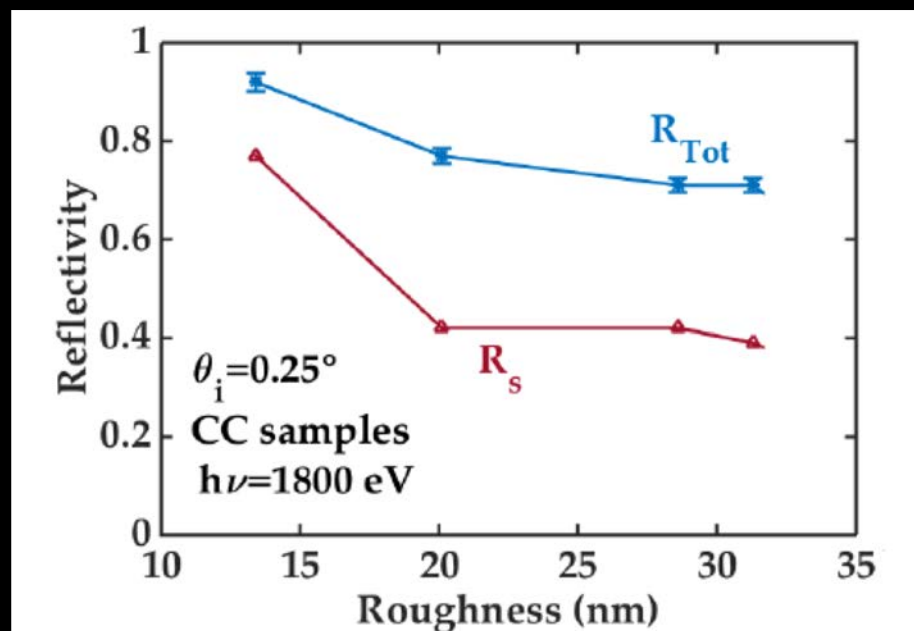
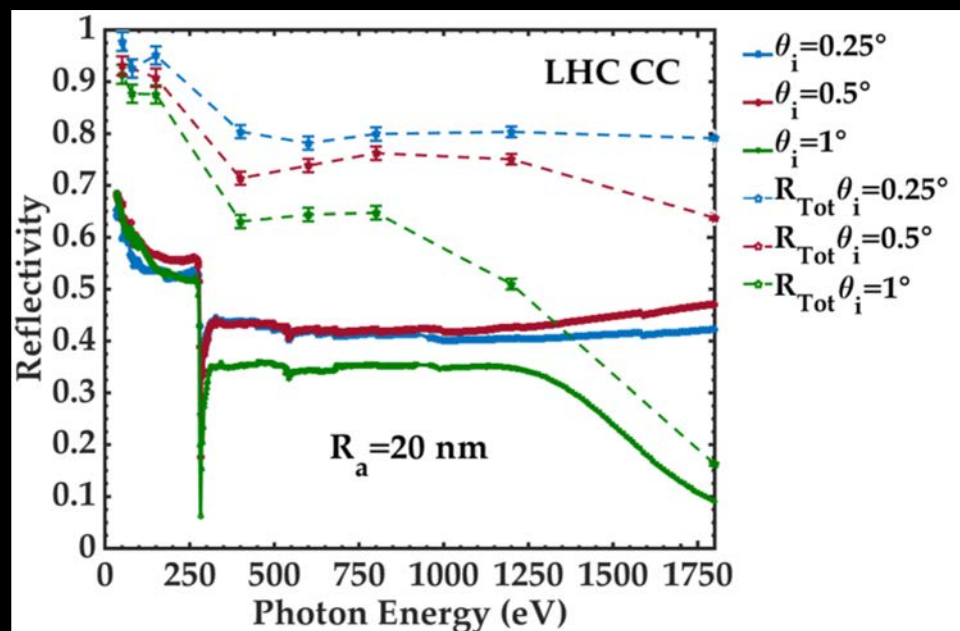
@ RT $P_{TOT} \sim 20 MW$

Credits: R. Kersevan - 2017



Specular and Total Reflectivity VS Photon Energy of LHC-Flat Carbon Coted (CC) sample.

Comparison between Specular and Total Reflectivity VS R_a for Carbon coated samples: at 1800 eV and incidence angle $\theta_i = 0.25^\circ$

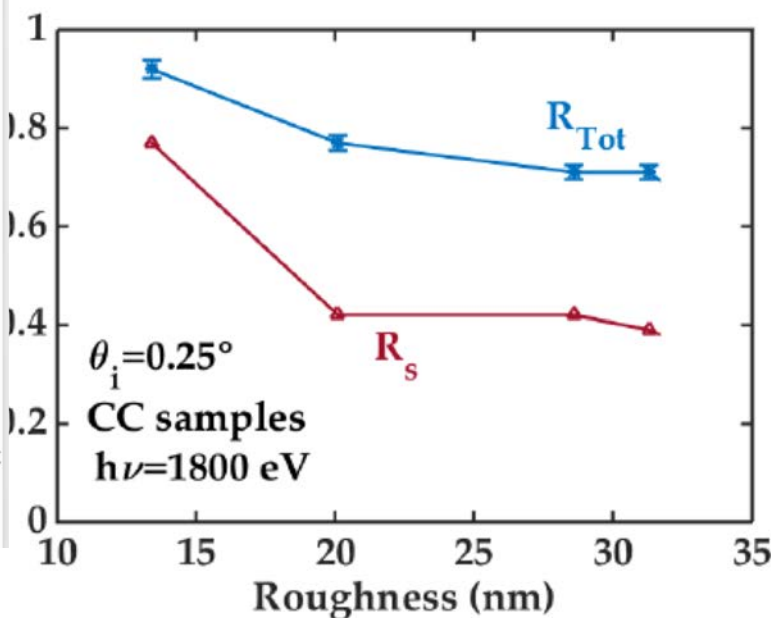
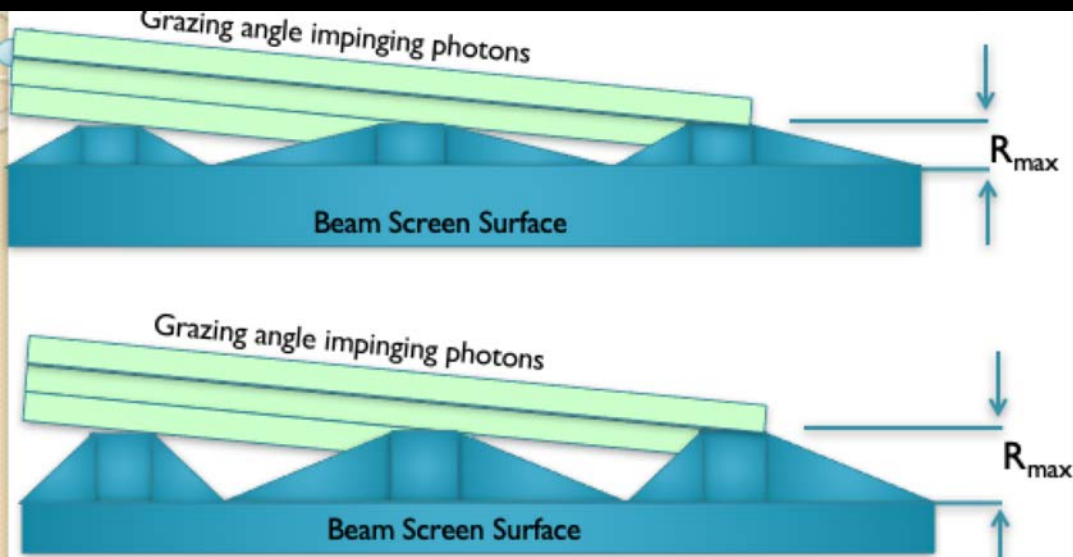


Increases with decreasing θ_i
It stays above 80 %

Does not significantly decreases
with R_a !!

Comparison between Specular and Total Reflectivity VS R_q for Carbon coated samples: at 1800 eV and incidence angle $\theta_i = 0.25^\circ$

R. Cimino, V. Baglin and F. Schäfers, PRL. 115 (2015) 264804



**Increases with decreasing θ_i
It stays above 80 %**

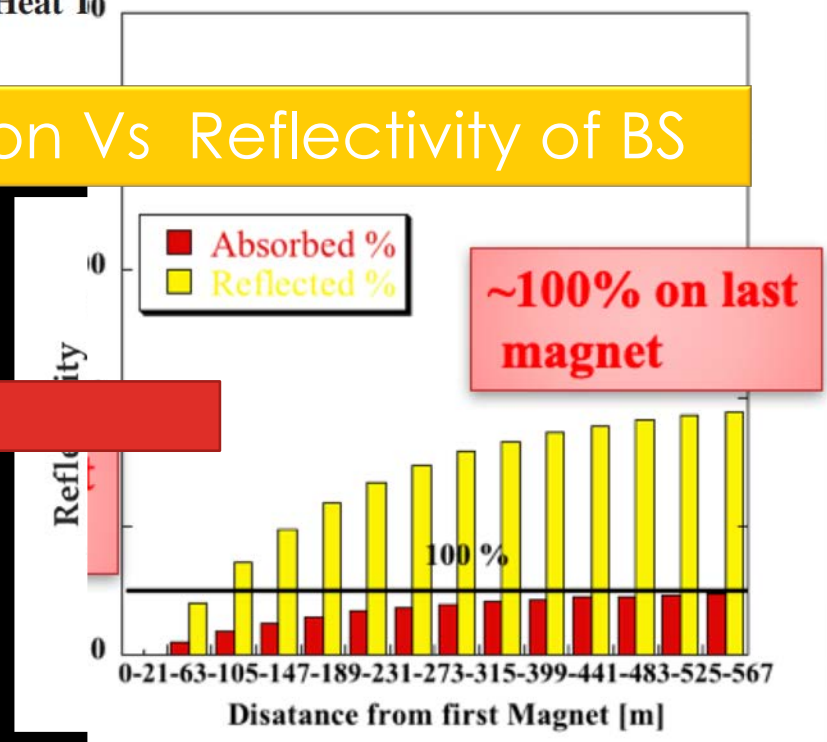
Does not significantly decrease with R_q !!

Potential Remedies for the High Synchrotron-Radiation-Induced Heat In Highest-Energy-Proton Circular Colliders

Heat Load Propagation Vs Reflectivity of BS

Depending on the distance between RT absorbers, one can foresee a significant reduction in the HL to be dissipated at COLD, and reduce constructive and machine running costs.

LHC standard surface finish



~100% on last magnet

Within 500m ~ 70% of expected HL adsorbed @ LT

R. Cimino, V. Baglin and F. Schäfers, PRL. 115 (2015) 264804



outline:

- The problem
- The experimental approach
- Selected results
- Conclusion

- R and PY on technically relevant samples can be experimentally measured!
 - We showed: - the importance of contaminants (LT studies?)
 - the importance of total R
- Simulation for Machine Design study must be supported by:
 - Experimental data of each individual Material Property
 - Dedicated experiments are the only way to produce the necessary inputs close to realistic conditions.
- We experimentally add a positive piece of information to validate the use of low SEY, high Reflectivity Carbon coatings to extract the HL from the cold BS to RT absorbers with a net reduction in building and running costs.

Acknowledgments

This work was supported by **INFN** National committee V through the "**MICA**" project. Research leading to these results has also received funding by the project **CALIPSOplus**, under the Grant Agreement 730872 from the EU Framework Programme for Research and Innovation HORIZON 2020. M.A., I.B., L.S. and L.G.G. acknowledge the support of the WP4 "**EuroCirCol**" project, the European Union's Horizon 2020 research and innovation programme under grant agreement No. 654305. We thank HZB for the allocation of synchrotron radiation beamtime. We thank R. Valizadeh, O. Malyshev for providing us with the LASE sample. We thank N. Kos for helping us with the preparation of some of the samples.

