

# From matrix / tensor models to black holes

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based on arXiv:1707.03431 with T. Azeyanagi and F. Ferrari

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## Desirable features of a black hole model

- A) Macroscopic space-time description [Schwarzschild - 1916; Kerr - 1963; ...]
- Definition of the horizon [Finkelstein - 1958; ...]
  - Description of the interior [Kruskal - 1960; Penrose, Hawking - 1965, 1970; ...]
  - Entropy  $S = A/4G_N$  [Bekenstein - 1972; Bardeen, Carter, Hawking - 1973; ...]

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**Can we study black holes starting from (B) and getting to (A) through holography?**

## Existence of parameter $N$

Loss of time-reversal invariance / Unitarity problems



Thermodynamical irreversibility (limit  $N \rightarrow \infty$ )

## Chaotic dynamics

$$F_\beta(t) = \left\langle \hat{O}(0) \hat{O}(t) \hat{O}(0) \hat{O}(t) \right\rangle_{\beta, \text{con.}} \propto e^{\lambda_L t}$$

Lyapunov exponent saturates bound for black holes

$$\lambda_L \leq \frac{2\pi}{\beta} \quad [ \text{Maldacena, Shenker, Stanford - 2015} ]$$

$N$  Majorana fermions  $\psi_1, \dots, \psi_N$  in  $0 + 1$  dim. with Hamiltonian

$$H = \sum_{i < j < k < l} J_{ijkl} \psi_i \psi_j \psi_k \psi_l$$

Quenched disorder

$$\langle \cdot \rangle \equiv \sum_{J_{ijkl}} p(J_{ijkl}) \langle \cdot \rangle_{J_{ijkl}} \quad \text{with} \quad \sigma^2(J_{ijkl}) \propto J^2$$

Some nice features

- Approximate conformal symmetry in IR  $\implies$  NAdS<sub>2</sub>/NCFT<sub>1</sub>
- Analytical treatment for  $N \rightarrow \infty$
- Explicit numerics for small  $N$  ( $|\mathcal{H}| = 2^{N/2}$ )
- Saturates bound for  $\lambda_L$

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**Not a proper Quantum Field Theory :-)**

# Vector and matrix models: an overview

	Large $D$ vector models	Large $N$ matrix models
<b>Field content</b>	$\phi_\mu$ with $\mu = 1, \dots, D$	$X^a_b$ with $a, b = 1, \dots, N$
<b>Symmetry</b>	$O(D)$	$U(N)^2$ or $U(N)$
<b>Interactions</b>	$(\vec{\phi}^2)^k$ for $k = 1, \dots$	$\text{Tr}(XX^\dagger XX^\dagger \dots), \dots$
<b>Diag. scaling</b>	$D^{V-P+\varphi} = D^{1-\ell}$	$N^{V-P+f} = N^{2-2g}$
<b>Leading</b>	cacti diagrams (auxiliary tree-level)	planar diagrams
<b>Applications</b>	cond. mat. ph., CFT, higher spin gravity	nucl. ph., QCD, string theory



## $O(D) \times U(N)^2$ model for a vector of complex matrices

Interaction vertices are  $V_B = \text{Tr} \left( X_{\mu_1} X_{\mu_2}^\dagger \cdots X_{\mu_{2s-1}} X_{\mu_{2s}}^\dagger \right)$

Usual scaling  $S = ND \left( \frac{1}{2} \text{Tr} (X_\mu X_\mu^\dagger) + \sum_B t_B V_B (X_\mu) \right)$

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Enhance 't Hooft coupling  $t_B$  for  $V_B$  as

$$t_B = \lambda_B D^{I(B)} \quad \text{with} \quad I(B) \geq 0$$

In the  $N \rightarrow \infty$  limit

$$F = \sum_{g \geq 0} N^{2-2g} F_g$$

In the  $D \rightarrow \infty$  limit ( $g$  fixed)

$$F_g = \sum_{k \geq 0} D^{1+g-k/2} F_{g,k}$$

# New large $N, D$ limit: example

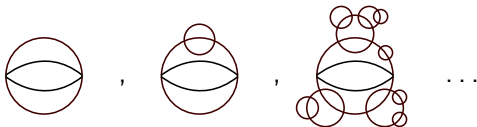
Enhanced couplings  $\implies$  more diagrams contribute order-by-order

$$\text{Tr}(X_\mu X_\mu^\dagger X_\nu X_\nu^\dagger) \implies I(B) = 0$$

$$\text{Tr}(X_\mu X_\nu^\dagger X_\mu X_\nu^\dagger) \implies I(B) = 1/2$$

$$\text{Tr}(X_\rho X_\mu^\dagger X_\rho X_\sigma^\dagger X_\mu X_\sigma^\dagger) \implies I(B) = 1$$

For  $\sqrt{D} \text{Tr}(X_\mu X_\nu^\dagger X_\mu X_\nu^\dagger)$  dominating diagrams are melons ( $\equiv$  SYK)



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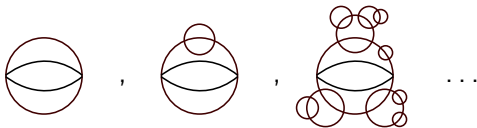
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**Intuition:** Genus  $g$  restriction kills many diagrams, so we can enhance some couplings and still have well-defined  $D \rightarrow \infty$  limit

# Quartic models for fermionic matrices

Fermionic matrices in  $0 + 1$  dimensions

$$(\psi_\mu^\dagger)^a_b = (\psi_{\mu a}^b)^\dagger \quad \text{with} \quad \left\{ \psi_{\mu b}^a, (\psi_\nu^\dagger)^c_d \right\} = \frac{1}{ND} \delta_{\mu\nu} \delta_d^a \delta_b^c$$

- **Model I:**  $O(D) \times U(N)^2$

$$H_{\text{I}} = ND \operatorname{Tr} \left( m \psi_\mu^\dagger \psi_\mu + \frac{1}{2} \lambda \sqrt{D} \psi_\mu \psi_\nu^\dagger \psi_\mu \psi_\nu^\dagger \right)$$

- **Model II:**  $O(D) \times U(N)$

$$H_{\text{II}} = ND \operatorname{Tr} \left\{ m \psi_\mu^\dagger \psi_\mu + \frac{1}{2} \sqrt{D} \left( \lambda \psi_\mu^\dagger \psi_\nu \psi_\mu \psi_\nu + \lambda^* \psi_\mu^\dagger \psi_\nu^\dagger \psi_\mu^\dagger \psi_\nu \right) \right\}$$

**Melonic-dominated models  $\implies$  Physics similar to SYK**

## Fermionic perturbation theory

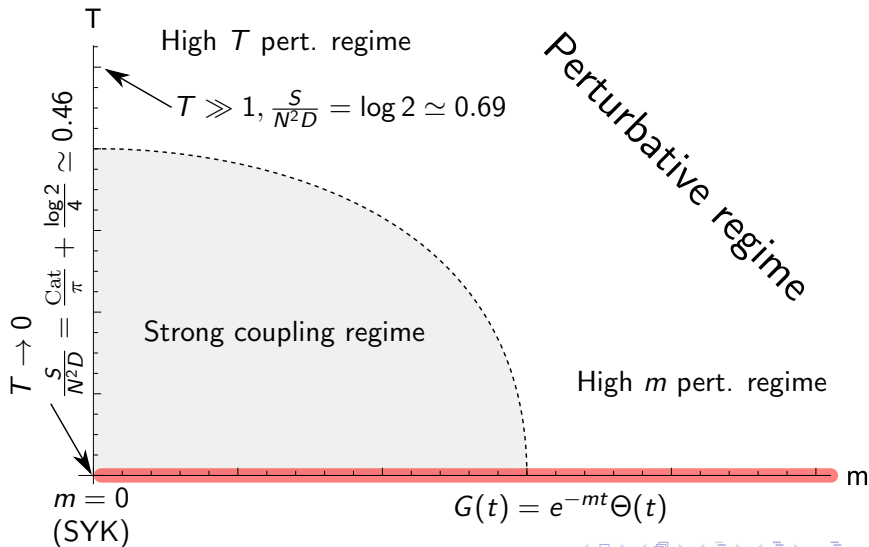
- $m \gg \lambda \implies$  Exp. around decoupled fermionic oscillators
- $T \gg \lambda \implies$  Non-standard (SYK-like) perturbation theory

## Feynman diagram structure $\iff$ Schwinger-Dyson equations

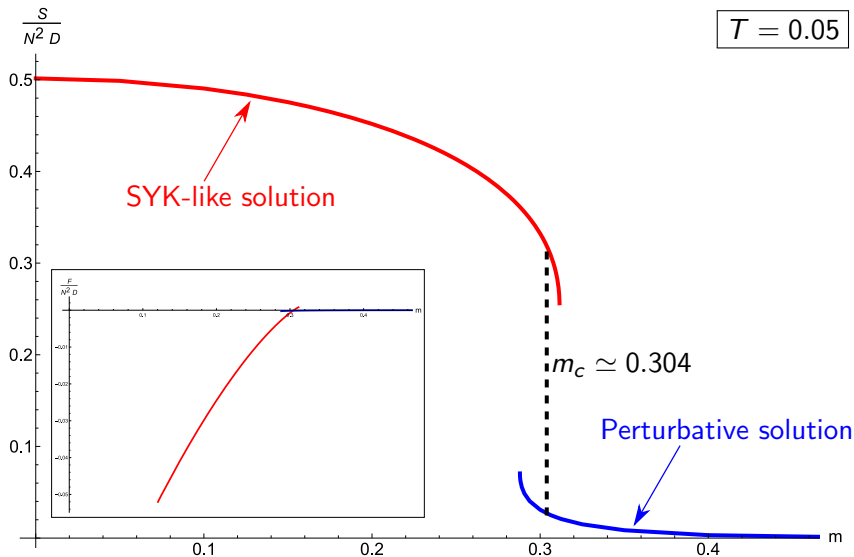


$$G_k^{-1} = m - i\omega_k + \Sigma_k \quad \Sigma(t) = \lambda^2 G(t)^2 G(-t)$$
$$G(t) = \frac{1}{\beta} \sum_k G_k e^{-i\omega_k t} \quad \omega_k = \frac{2\pi}{\beta} k \quad \text{with } k \in \mathbb{Z} + \frac{1}{2}$$

# Phase diagram structure

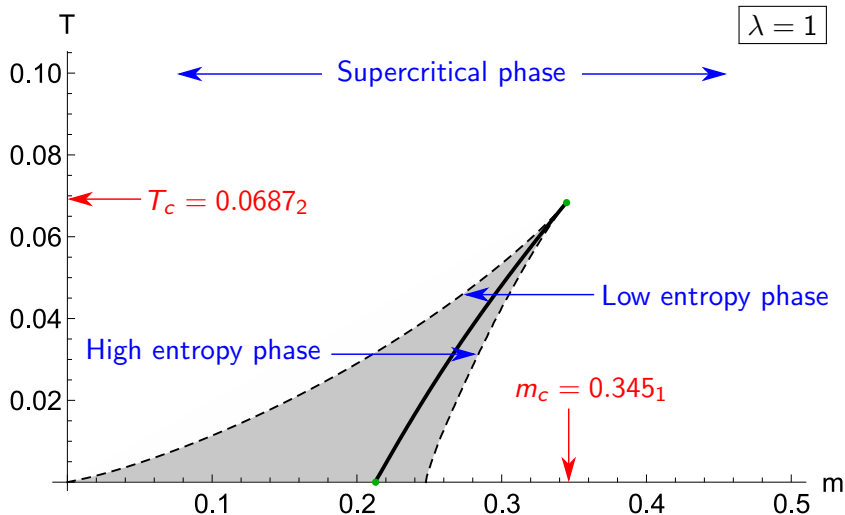


# Building the phase diagram





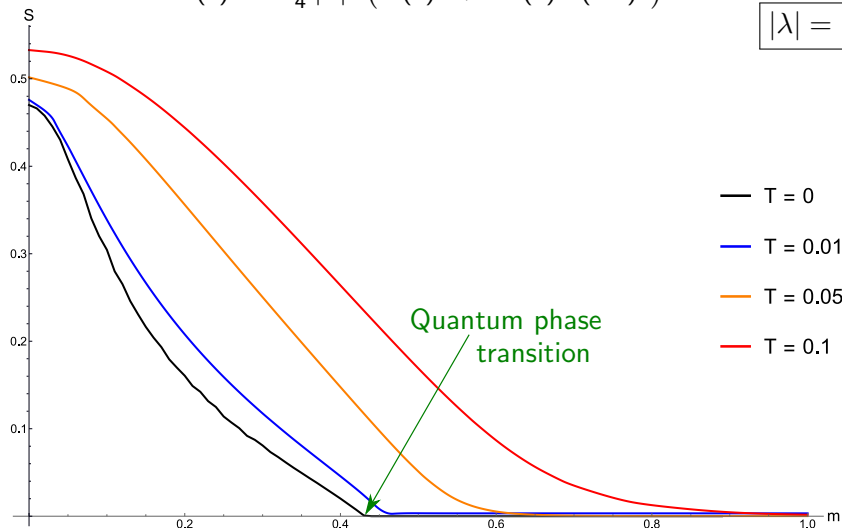
# Phase diagram: $(m, T)$ plane



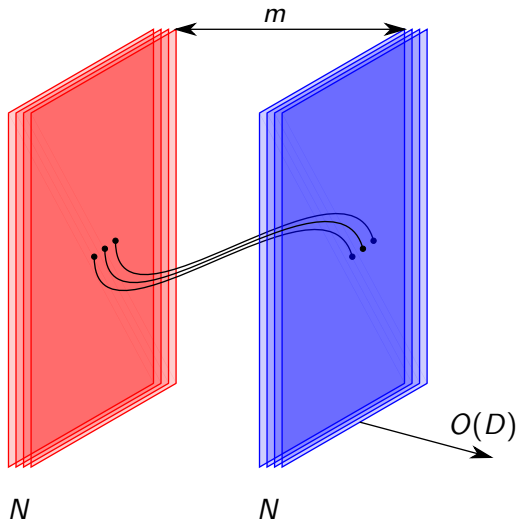
# Fermionic model II

$$\Sigma(t) = -\frac{1}{4} |\lambda|^2 (G(t)^3 + 3G(t)G(-t)^2)$$

$$|\lambda| = 1$$



## Stringy description of gravitational collapse



# Finite $N$ analysis

$\psi_{\mu b}^a$  has  $D \times N^2$  fermionic degrees of freedom



**Hilbert space is  $2^{DN^2}$  dimensional :-)**

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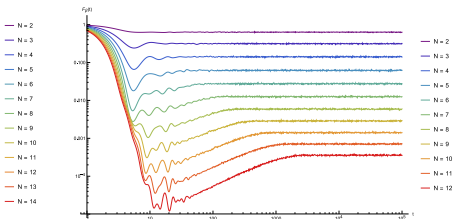
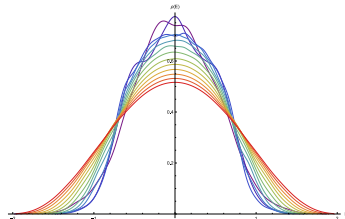


**Hilbert space is  $2^{DN^2}$  dimensional :- (**

Disordered models with  $N$  Dirac fermions ( $|\mathcal{H}| = 2^N$ )

$$\tilde{H}_I = m\chi_i^\dagger\chi^i + J_{kl}^{ij}\chi_i^\dagger\chi_j^\dagger\chi^k\chi^l$$

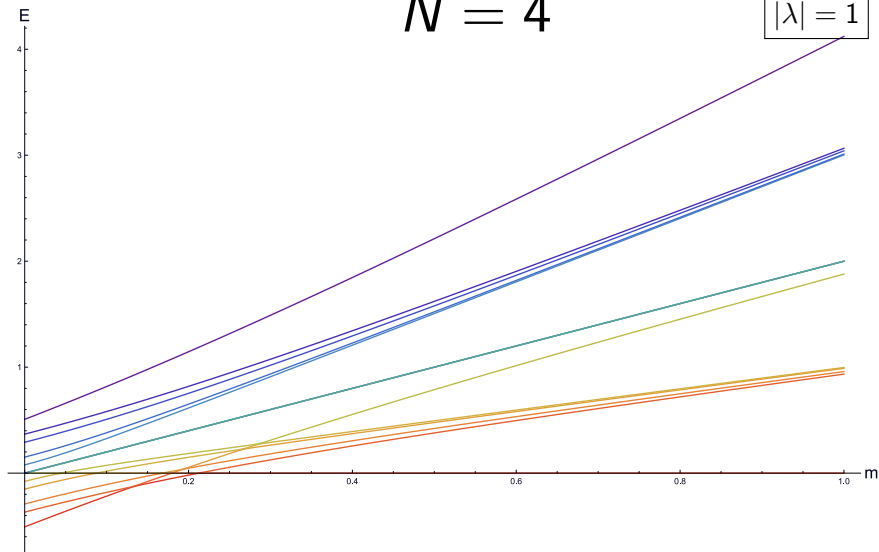
$$\tilde{H}_{II} = m\chi_i^\dagger\chi^i + J_{jkl}^i\chi_i^\dagger\chi^j\chi^k\chi^l + J'_{ijk}\chi_i^\dagger\chi_j^\dagger\chi_k^\dagger\chi^l$$



# Finite $N$ spectrum

$N = 4$

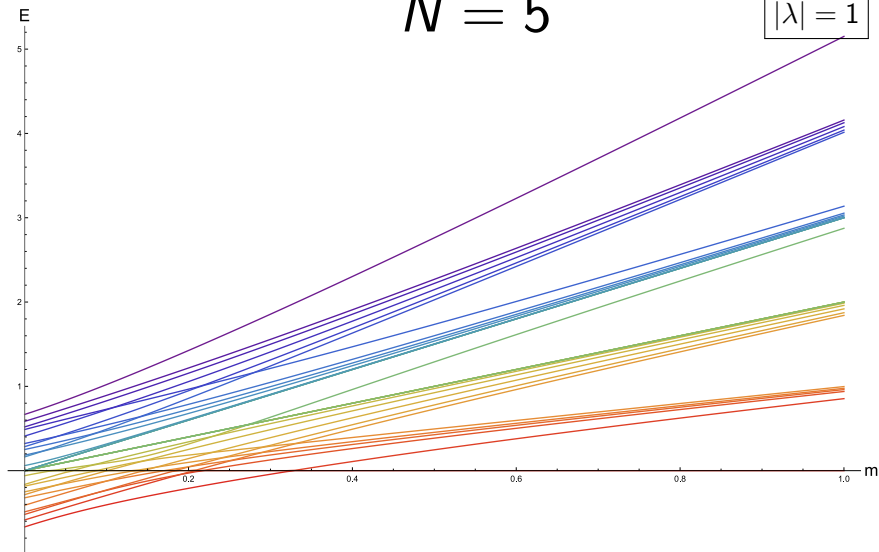
$|\lambda| = 1$



# Finite $N$ spectrum

$N = 5$

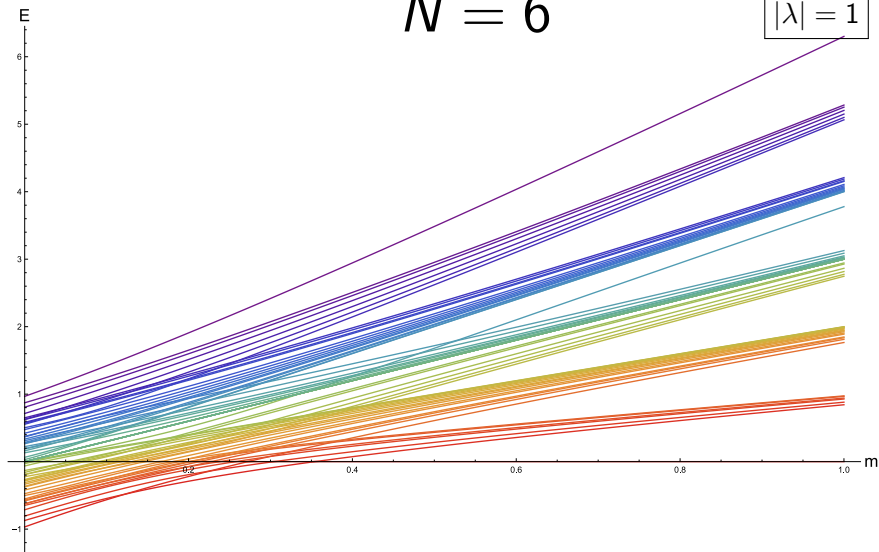
$|\lambda| = 1$



# Finite $N$ spectrum

$N = 6$

$|\lambda| = 1$

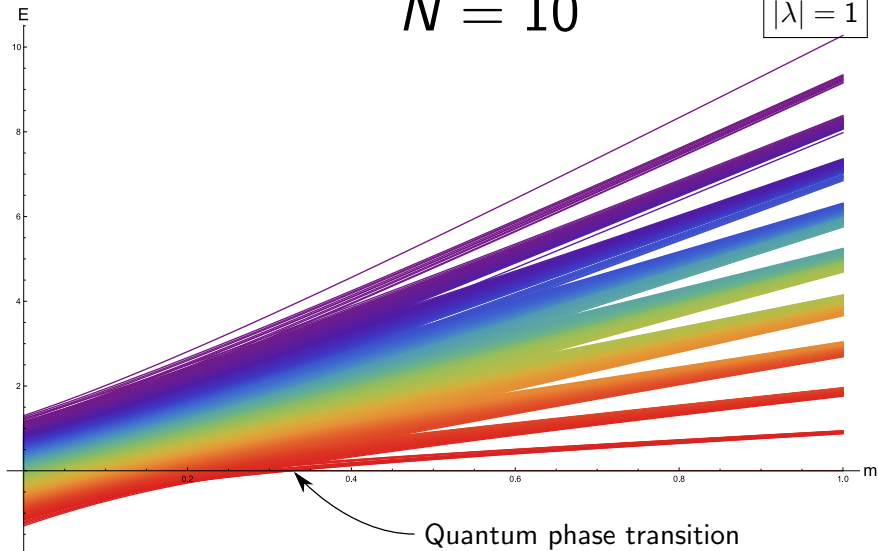




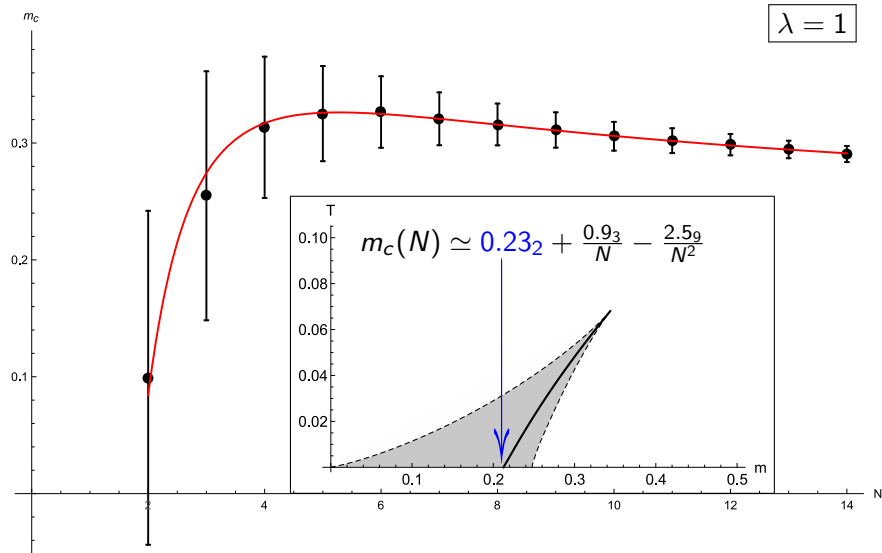
# Finite $N$ spectrum

$N = 10$

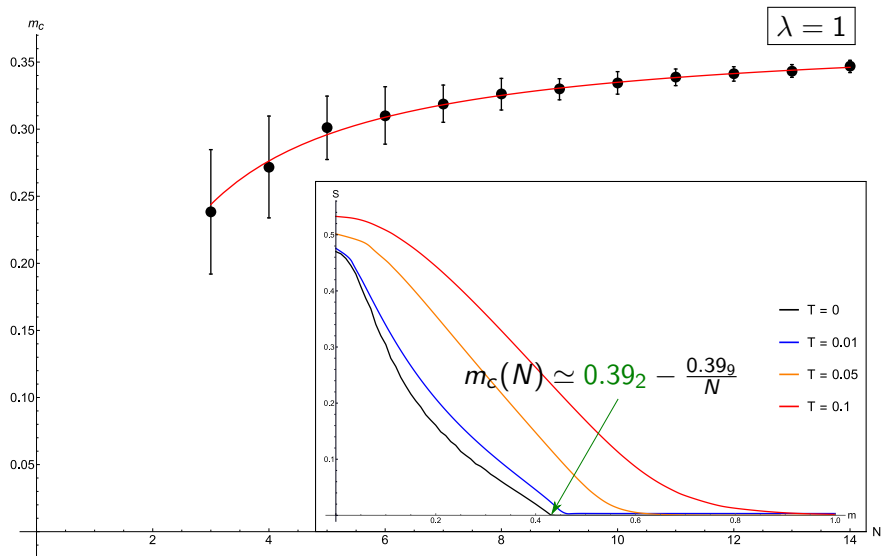
$|\lambda| = 1$



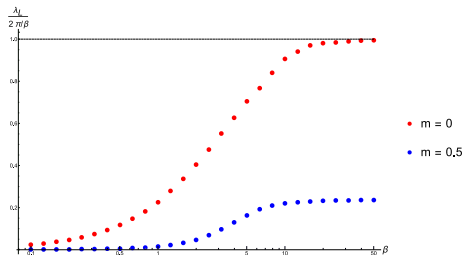
# Quantum critical mass: model I



# Quantum critical mass: model II



## 1 Real-time physics

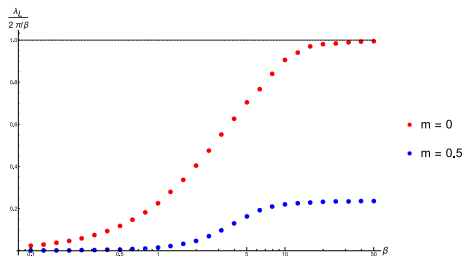


## 2 Effective description à la Landau

## 3 Supersymmetric models

## 4 Holographic picture

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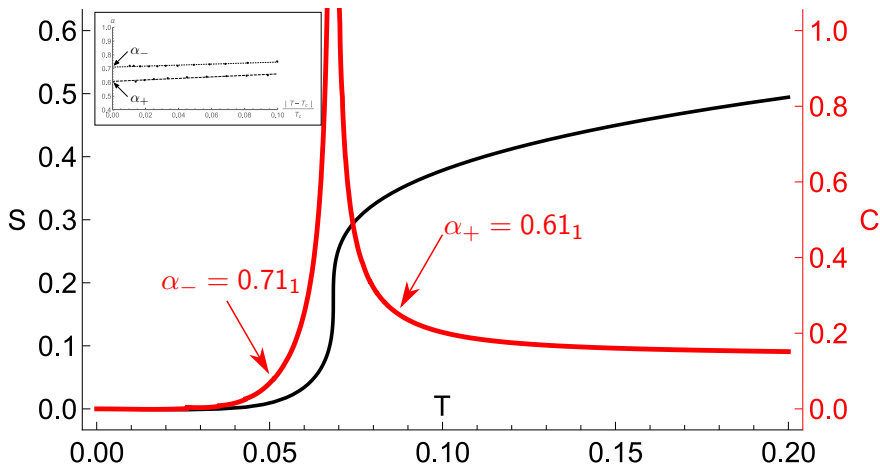
**Takeaway: quantum black hole playground (in a computer :-)**

# Thanks!

$S(T)$  and  $C_m = T\partial S/\partial T$  at  $m = m_c$

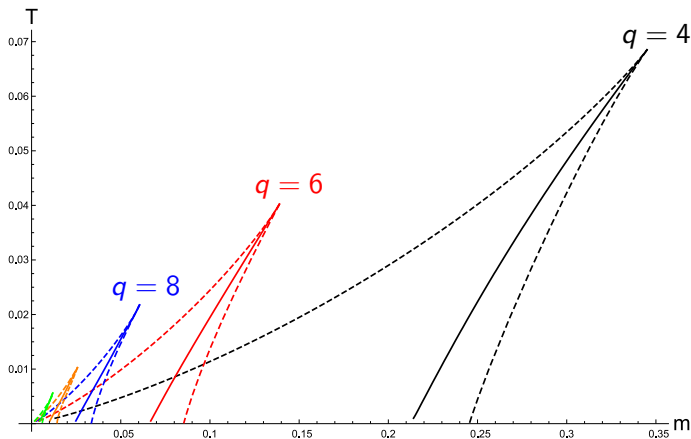
## New strongly coupled critical point

$\lambda = 1$



## $q$ -body interacting generalization of the fermionic model I

$$\Sigma(t) = \lambda^2 G(t)^2 G(-t) \implies \Sigma(t) = (-1)^{\frac{q}{2}} \lambda^2 G(t)^{\frac{q}{2}} G(-t)^{\frac{q}{2}-1}$$





$O(D) \times U(N)^2$  purely bosonic models of Hermitian matrices

$$H_{\text{III}} = ND \text{Tr} \left( \frac{m^2}{2} X_\mu X_\mu + \frac{\lambda^3}{4} \sqrt{D} X_\mu X_\nu X_\mu X_\nu \right)$$

$$H_{\text{IV}} = ND \text{Tr} \left( \frac{m^2}{2} X_\mu X_\mu + \lambda^4 D X_\rho X_\mu X_\rho X_\sigma X_\mu X_\sigma \right)$$

Crucial differences:

- **Classical limit**  $T \gg \lambda$ :

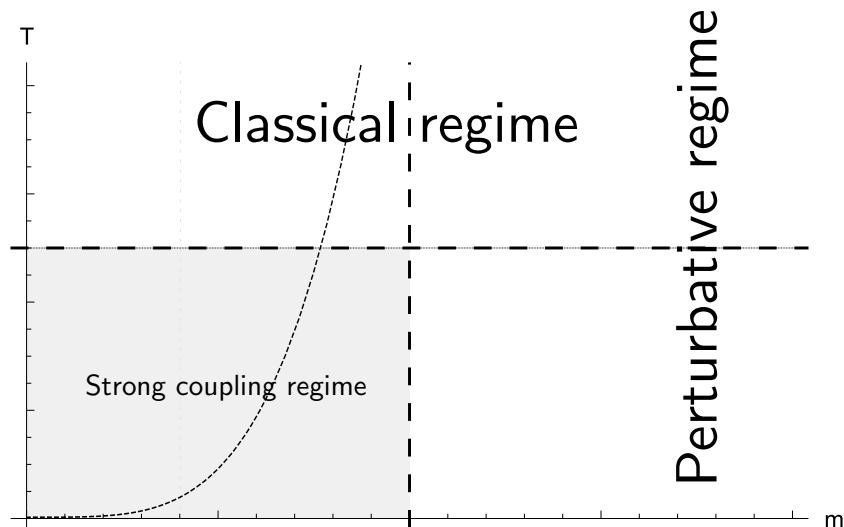
Strongly coupled zero mode  $\implies$  No high  $T$  pert. theory

(Classical eq. for zero mode of (III):  $x^{-1} = m^2 - \frac{\lambda^6}{\beta^2} x^3$ )

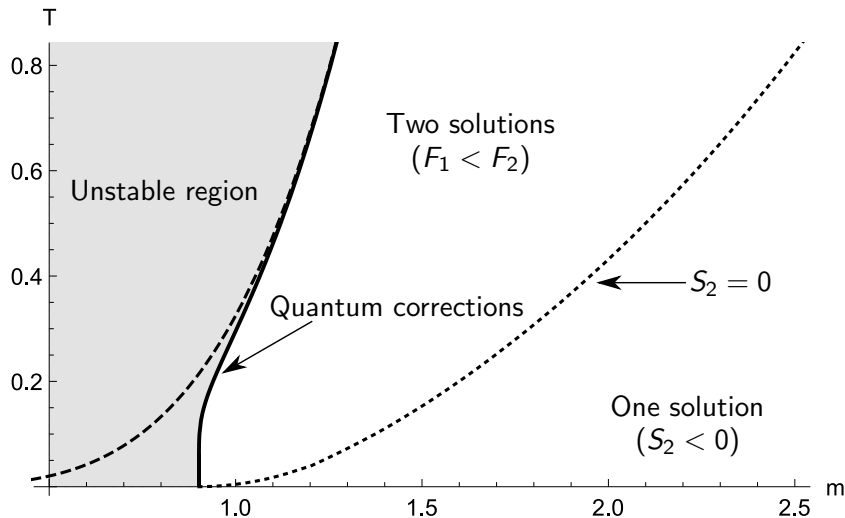
- **Instabilities:** Potential for (III) is unstable

No gap  $\implies$  IR divergences

# Phase diagram structure



# Phase diagram for (III): $(m, T)$ plane



# $(q_1, q_2) = (4, 8)$ domain wall

$$m = 0; T = 10^{-4}$$

