

Discussion: Transport coefficients

Topics for discussion:

- What transport coefficients do we need?
- What do we know?
- Important problems
- How to make transport coefficients easily accessible for a broad NS community: ideas, suggestions

• What transport coefficients do we need?


- Cooling of isolated NSs and X-ray transients $\longrightarrow \kappa$
- Oscillations of NSs $\longrightarrow \eta, \xi$
- Magnetic field evolution $\longrightarrow D_{ij}, \sigma$
- Vortex transport (glitches, inter-glitch dynamics) \longrightarrow **transport coefficients describing vortex-particle interactions**
- Dissipative effects during binary NS inspiral and post-merger phase (may affect GW signal) $\longrightarrow \xi, \eta, \kappa$
- ...?

• What do we know?

Crust:

- basic transport coefficients are rather well known κ, η, σ (but not in the pasta phase)

Core:

- As a rule, $n\rho e\mu$ composition, no magnetic field, effects of superfluidity are partly taken into account  $\kappa, \eta, \xi, \sigma$

D_{ij}

- diffusion coefficients in the magnetic field are only well known for a simplified nonsuperfluid model of *Yakovlev & Shalybkov (1991)*
- vortex-related transport coefficients are poorly known and controversial. The reliable result: mutual friction coefficient describing scattering of electrons off the vortex magnetic field.

Important problems

- Thermal and electrical conductivity in the pasta phases: important, e.g., for explaining magnetars
- Calculate diffusion tensor D_{ij} with the updated microphysics
- Vortex transport
- Polarization properties of $npe\mu$ -matter accounting for the Fermi-liquid effects and superfluidity/superconductivity (important for correctly describing the electron-electron and electron-proton scattering in dense Fermi-mixture).
- ...?

- How to make transport coefficients easily accessible for a broad NS community: ideas, suggestions

- Fitting formulas

(example: fits from the review by *Schmitt & Shternin 2017*)

$$\eta_{e\mu} = 8.43 \times 10^{20} \left(\frac{n_B}{n_0} \right)^{14/9} \left(\frac{T}{10^8 \text{ K}} \right)^{-5/3} \frac{x_e^2 + x_\mu^2}{(x_e^{2/3} + x_\mu^{2/3} + x_p^{2/3})^{2/3}} \frac{\text{g}}{\text{cm s}}$$

$$\eta_n = 2.15 \times 10^{17} \left(\frac{n_n}{n_0} \right)^{5/3} \left(\frac{m_n^*}{m_N} \right)^{-2} \left(\frac{T}{10^8 \text{ K}} \right)^{-2} \left[\left(\frac{m_n^*}{m_N} \right)^2 m_\pi^2 S_{\eta nn} + \left(\frac{m_p^*}{m_N} \right)^2 m_\pi^2 S_{\eta np} \right]^{-1} \frac{\text{g}}{\text{cm s}}$$

- Tables
- Publicly available codes (example: library of Alexander Potekhin)

<http://www.ioffe.ru/astro/conduct/index.html>

- ...?