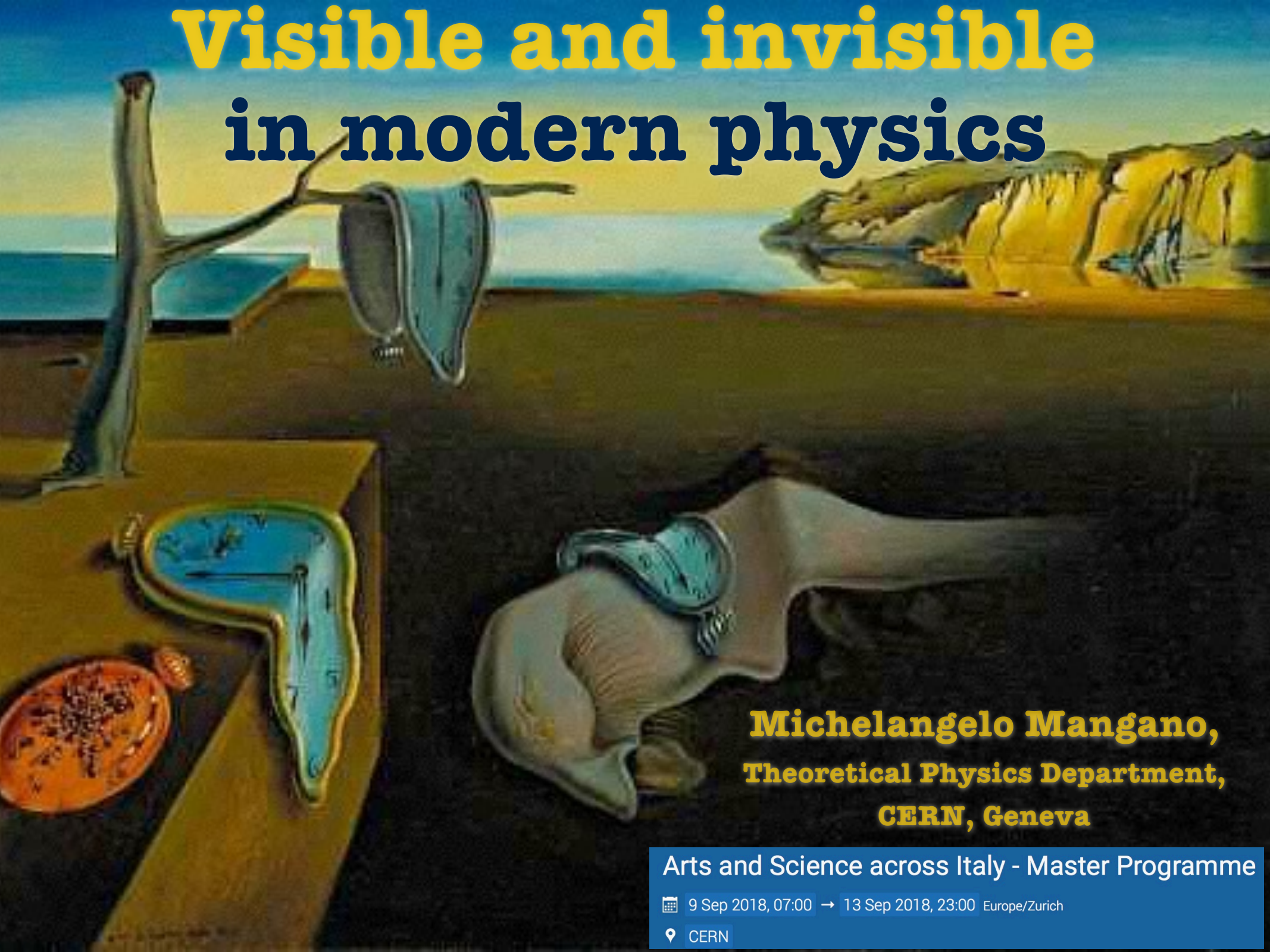


Visible and invisible in modern physics



**Michelangelo Mangano,
Theoretical Physics Department,
CERN, Geneva**

Arts and Science across Italy - Master Programme

9 Sep 2018, 07:00 → 13 Sep 2018, 23:00 Europe/Zurich

CERN

no artist better than Salvador Dali represents the inspiring power of science, and was more fascinated by the relation of visible and invisible

“In the Surrealist period I wanted to create the iconography of the interior world and the world of the marvelous, of my father Freud.

Today the exterior world and that of physics, has transcended the one of psychology.

My father today is Dr. Heisenberg.”

S.Dali, 'Anti-matter manifesto' (1958)



Hubble Ultra Deep Field
Hubble Space Telescope • Advanced Camera for Surveys

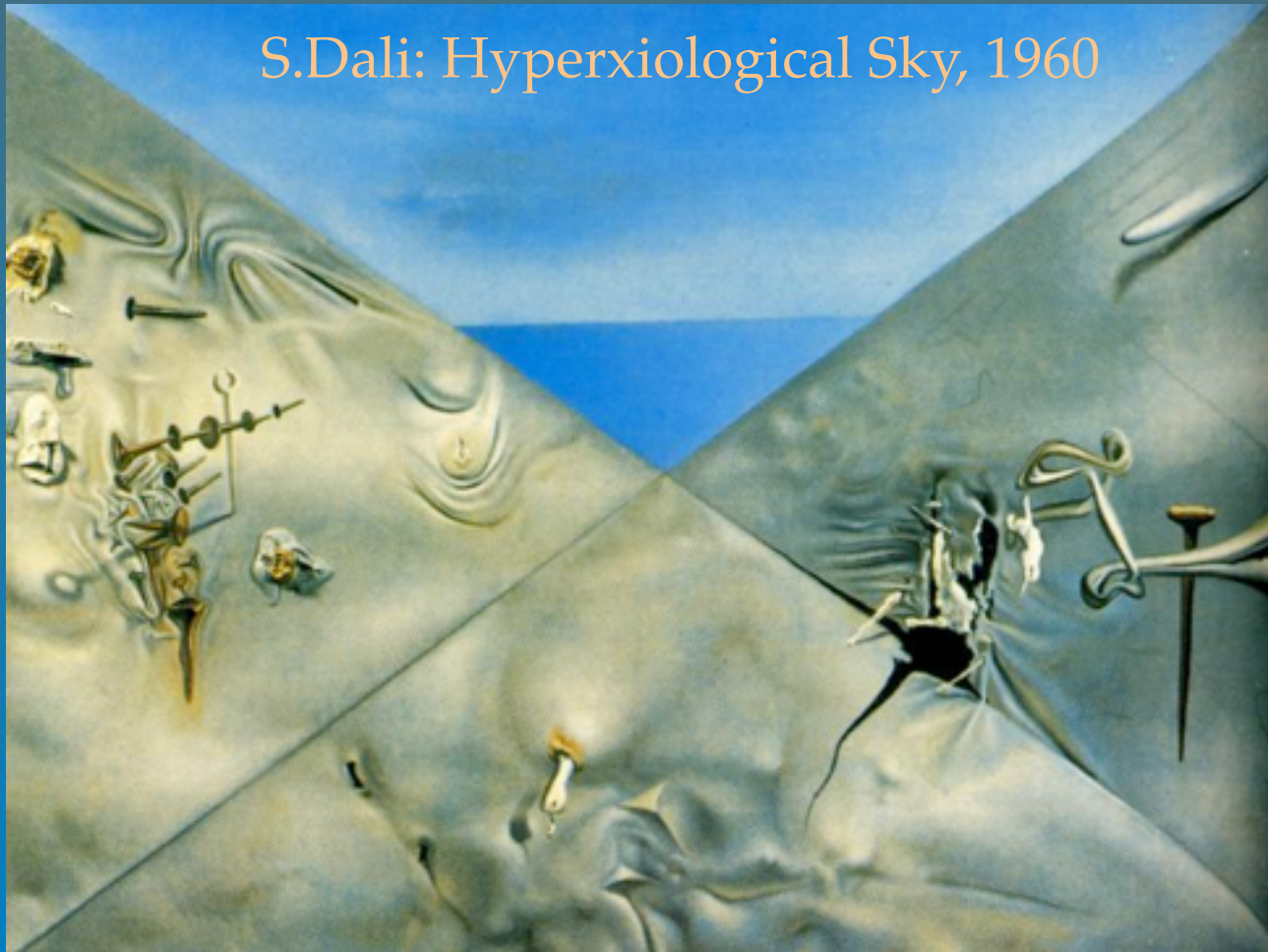
NASA, ESA, S. Beckwith (STScI) and the HUDF Team

STScI-PRC04-07a

S.Dali: Ciel, 1931



S.Dali: Hyperxiological Sky, 1960



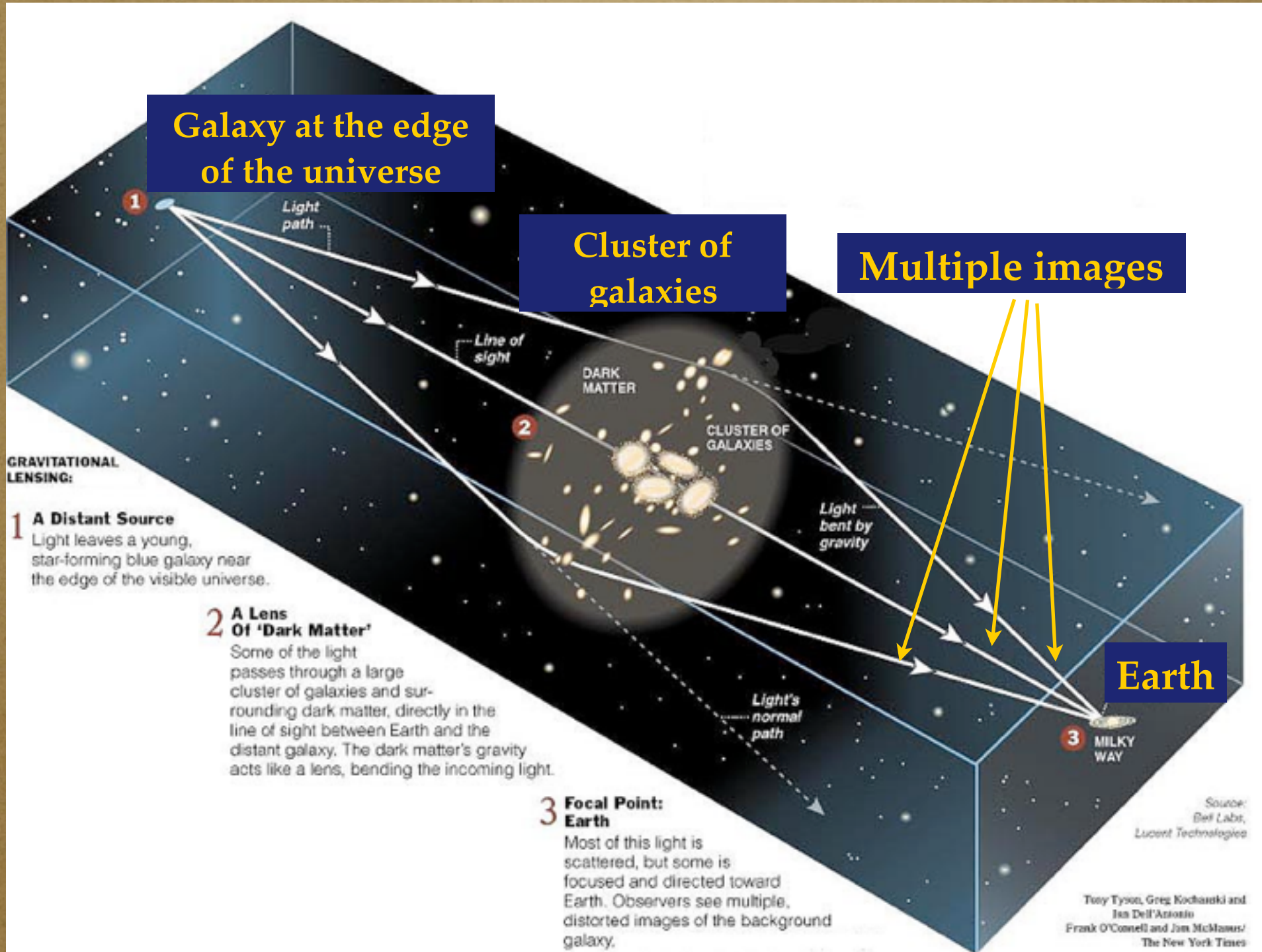


Hubble Ultra Deep Field
Hubble Space Telescope • Advanced Camera for Surveys

NASA, ESA, S. Beckwith (STScI) and the HUDF Team

STScI-PRC04-07a

Gravitational lensing



Galaxy at the edge of the universe

Cluster of galaxies

Multiple images

Earth

GRAVITATIONAL LENSING:

1 A Distant Source

Light leaves a young, star-forming blue galaxy near the edge of the visible universe.

2 A Lens Of 'Dark Matter'

Some of the light passes through a large cluster of galaxies and surrounding dark matter, directly in the line of sight between Earth and the distant galaxy. The dark matter's gravity acts like a lens, bending the incoming light.

3 Focal Point: Earth

Most of this light is scattered, but some is focused and directed toward Earth. Observers see multiple, distorted images of the background galaxy.

Source: Bell Labs, Lucent Technologies

Tony Tyson, Greg Kochanski and Jan Dell'Antonio
Frank O'Connell and Jim McClasus/
The New York Times

Galaxy cluster Abell 2218



Credits: European Space Agency, NASA, J.-P. Kneib (Observatoire Midi-Pyrénées) and R. Ellis (Caltech)

The shape and intensity of lensed images requires the presence of much more mass than what's **visible** in those galaxies !



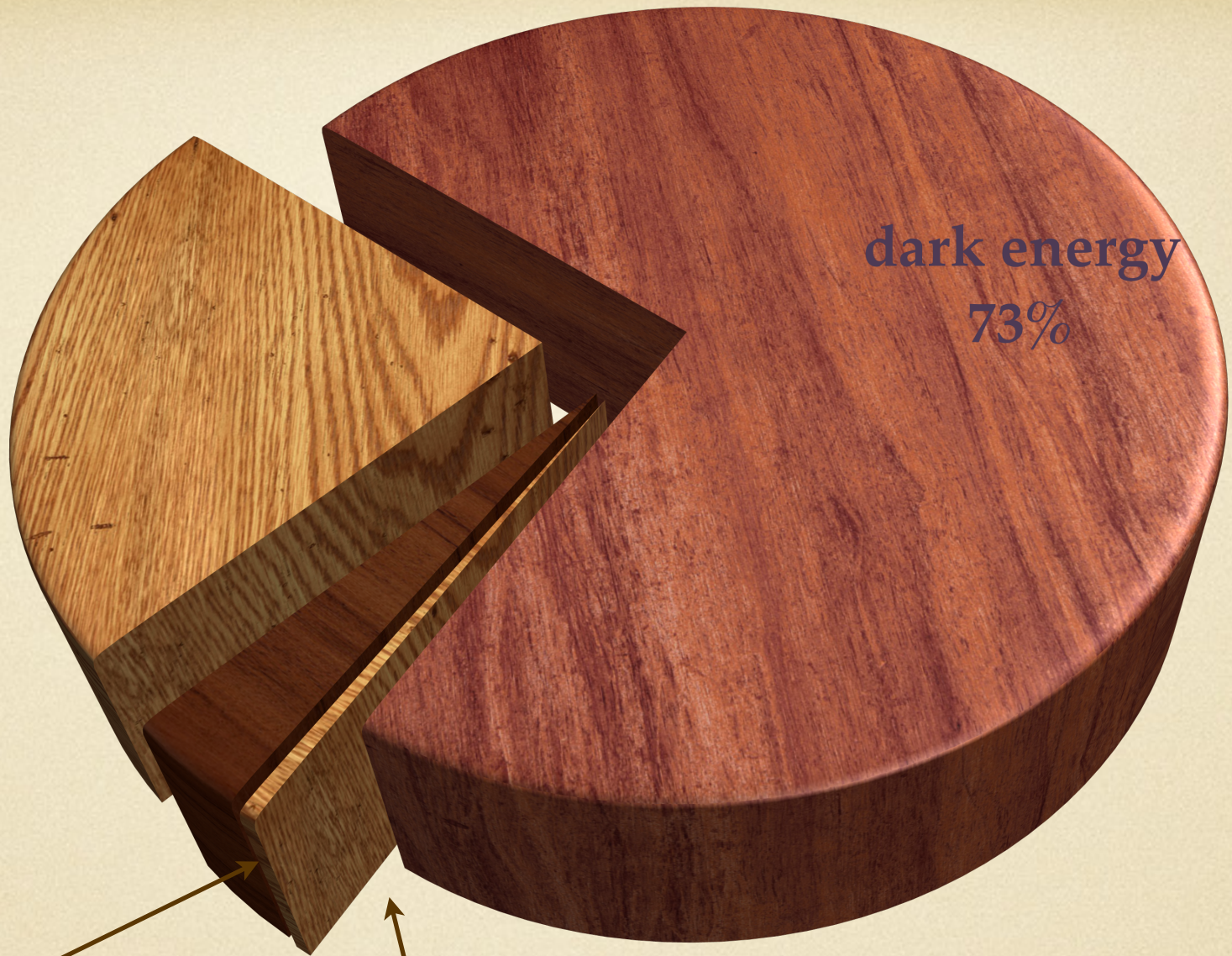
Invisible dark matter

dark matter
23%

dark energy
73%

non-luminous ordinary
matter (e.g. planets,
dead stars, dust, etc),
~4%

luminous matter (stars, gas) ~0.5%



For something to be
declared invisible, we
must know it's there, and
if we know it's there, it's
not *truly* invisible any
longer



S.Dali, Surrealist Composition with Invisible Figures, 1936

Proving the existence of the invisible, namely providing evidence that there is something where there appears to be nothing, turning the invisible into visible, is one of the main drivers of scientific progress.

It is a very basic process, that moves us from the realm of magic and superstition to the domain of science.

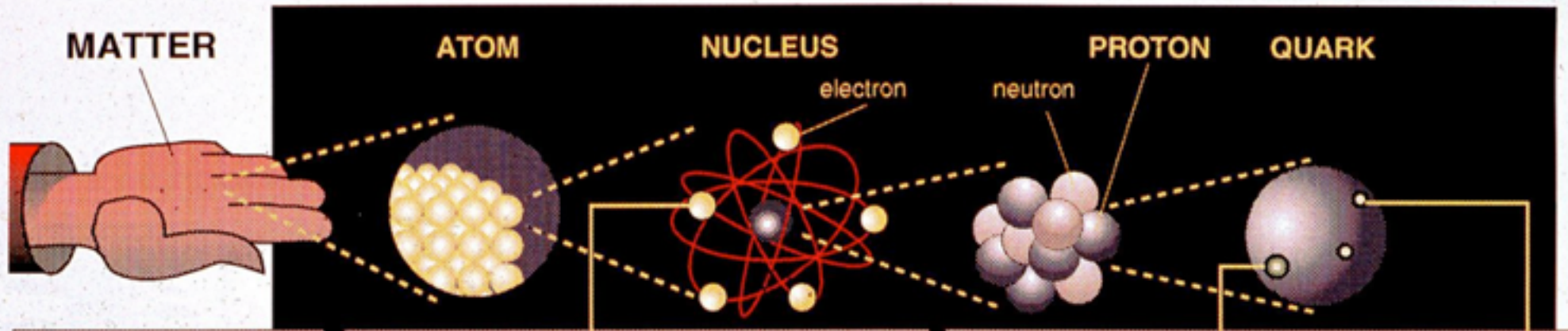


Invisible Harp, 1934

Establishing the **nature** of the invisible is the really crucial step

Goal of modern physics: to unveil the invisible, give it substance, and explore its consequences on the universe

The Standard Model of particle physics



<p>ALL ORDINARY MATTER BELONGS TO THIS GROUP.</p>	<h2>LEPTONS</h2>		<h2>QUARKS</h2>	
	<p>electron Electric charge -1. Responsible for electricity and chemical reactions</p>	<p>electron neutrino Electric charge 0. Rarely interacts with other matter.</p>	<p>up Electric charge $+2/3$. Protons have 2 up quarks Neutrons have 1 up quark</p>	<p>down Electric charge $-1/3$. ... and one down quark. ... and two down quarks.</p>
<p>THESE PARTICLES EXISTED JUST AFTER THE BIG BANG.</p>	<p>muon A heavier relative of the electron.</p>	<p>muon neutrino Created with muons when some particles decay.</p>	<p>charm A heavier relative of the up.</p>	<p>strange A heavier relative of the down.</p>
	<p>tau Heavier still.</p>	<p>tau neutrino Not yet observed directly.</p>	<p>top Heavier still, recently observed.</p>	<p>bottom Heavier still.</p>

ANTIMATTER
Each particle also has an antimatter counterpart ... sort of a mirror image.



BOSONS

force carriers
spin = 0, 1, 2, ...

Unified Electroweak spin = 1		
Name	Mass GeV/c ²	Electric charge
γ photon	0	0
W^-	80.39	-1
W^+	80.39	+1
W bosons		
Z^0 Z boson	91.188	0

Strong (color) spin = 1		
Name	Mass GeV/c ²	Electric charge
g gluon	0	0

Properties of the Interactions

The strengths of the interactions (forces) are shown relative to the strength of the electromagnetic force for two u quarks separated by the specified distances.

Property	Gravitational Interaction	Weak Interaction (Electroweak)	Electromagnetic Interaction	Strong Interaction
Acts on:	Mass – Energy	Flavor	Electric Charge	Color Charge
Particles experiencing:	All	Quarks, Leptons	Electrically Charged	Quarks, Gluons
Particles mediating:	Graviton (not yet observed)	W^+ W^- Z^0	γ	Gluons
Strength at $\left\{ \begin{array}{l} 10^{-18} \text{ m} \\ 3 \times 10^{-17} \text{ m} \end{array} \right.$	10^{-41} 10^{-41}	0.8 10^{-4}	1 1	25 60

BOSONS

force carriers
spin = 0, 1, 2, ...

Unified Electroweak spin = 1		
Name	Mass GeV/c ²	Electric charge
γ photon	0	0
W^-	80.39	-1
W^+ W bosons	80.39	+1
Z^0 Z boson	91.188	0

Strong (color) spin = 1		
Name	Mass GeV/c ²	Electric charge
g gluon	0	0

EW symmetry breaking spin=0		
H higgs	125	0

Properties of the Interactions

The strengths of the interactions (forces) are shown relative to the strength of the electromagnetic force for two u quarks separated by the specified distances.

Property	Gravitational Interaction	Weak Interaction (Electroweak)	Electromagnetic Interaction	Strong Interaction
Acts on:	Mass – Energy	Flavor	Electric Charge	Color Charge
Particles experiencing:	All	Quarks, Leptons	Electrically Charged	Quarks, Gluons
Particles mediating:	Graviton (not yet observed)	W^+ W^- Z^0	γ	Gluons
Strength at $\left\{ \begin{array}{l} 10^{-18} \text{ m} \\ 3 \times 10^{-17} \text{ m} \end{array} \right.$	10^{-41} 10^{-41}	0.8 10^{-4}	1 1	25 60

what? how?

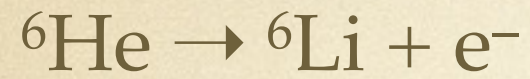
- Is nature built out of fundamental building blocks?
- If so, what are they?
- How do they interact?
- How do they determine the properties of the Universe?

The description of the natural phenomena that emerges from the Standard Model agrees **quantitatively** with great precision with all phenomena that we **see** around ourselves, and that we **measure** in the laboratory

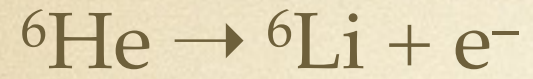
The Standard Model provides the underlying explanation of all nuclear, chemical and electrical phenomena, the atomic structure of elements, the electrical / mechanical / thermal behaviour of metals, semiconductors, etc. etc. etc.

A few anecdotes on the
role of the invisible in the
discovery and
understanding of
fundamental particles and
interactions

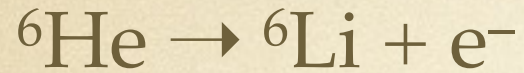
Late 1920's, a mystery of the invisible



Late 1920's, a mystery of the invisible



Late 1920's, a mystery of the invisible



If this were all that happens, energy conservation would demand that the energy of the emitted electron be the same for each decay

$$\text{Energy}[e^{-}] = \text{Mass}[{}^6\text{He}]c^2 - \text{Mass}[{}^6\text{Li}]c^2$$

Late 1920's, a mystery of the invisible



If this were all that happens, energy conservation would demand that the energy of the emitted electron be the same for each decay

$$\text{Energy}[e^{-}] = \text{Mass}[{}^6\text{He}]c^2 - \text{Mass}[{}^6\text{Li}]c^2$$



But in some decays the electron is very slow.....

Late 1920's, a mystery of the invisible



If this were all that happens, energy conservation would demand that the energy of the emitted electron be the same for each decay

$$\text{Energy}[e^{-}] = \text{Mass}[{}^6\text{He}]c^2 - \text{Mass}[{}^6\text{Li}]c^2$$



But in some decays the electron is very slow.....

Late 1920's, a mystery of the invisible



If this were all that happens, energy conservation would demand that the energy of the emitted electron be the same for each decay

$$\text{Energy}[e^{-}] = \text{Mass}[{}^6\text{He}]c^2 - \text{Mass}[{}^6\text{Li}]c^2$$



... while in others it is very fast

Late 1920's, a mystery of the invisible



If this were all that happens, energy conservation would demand that the energy of the emitted electron be the same for each decay

$$\text{Energy}[e^{-}] = \text{Mass}[{}^6\text{He}]c^2 - \text{Mass}[{}^6\text{Li}]c^2$$



... while in others it is very fast

Enter neutrinos ...

After months of speculations, including the possibility that the principle of energy conservation be violated in microscopic quantum phenomena, Wolfgang Pauli proposes the existence of the **invisible neutrino**



Wolfgang Pauli



S.Dali, The invisible
man, 1929

Enter neutrinos ...

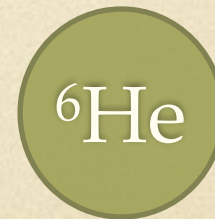
After months of speculations, including the possibility that the principle of energy conservation be violated in microscopic quantum phenomena, Wolfgang Pauli proposes the existence of the **invisible neutrino**



Wolfgang Pauli



S.Dali, The invisible man, 1929



Enter neutrinos ...

After months of speculations, including the possibility that the principle of energy conservation be violated in microscopic quantum phenomena, Wolfgang Pauli proposes the existence of the **invisible neutrino**



Wolfgang Pauli



S.Dali, The invisible man, 1929



Enter neutrinos ...

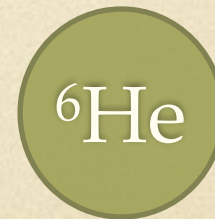
After months of speculations, including the possibility that the principle of energy conservation be violated in microscopic quantum phenomena, Wolfgang Pauli proposes the existence of the **invisible neutrino**



Wolfgang Pauli



S.Dali, The invisible man, 1929



Enter neutrinos ...

After months of speculations, including the possibility that the principle of energy conservation be violated in microscopic quantum phenomena, Wolfgang Pauli proposes the existence of the **invisible neutrino**



e^{-}

${}^6\text{Li}$

ν

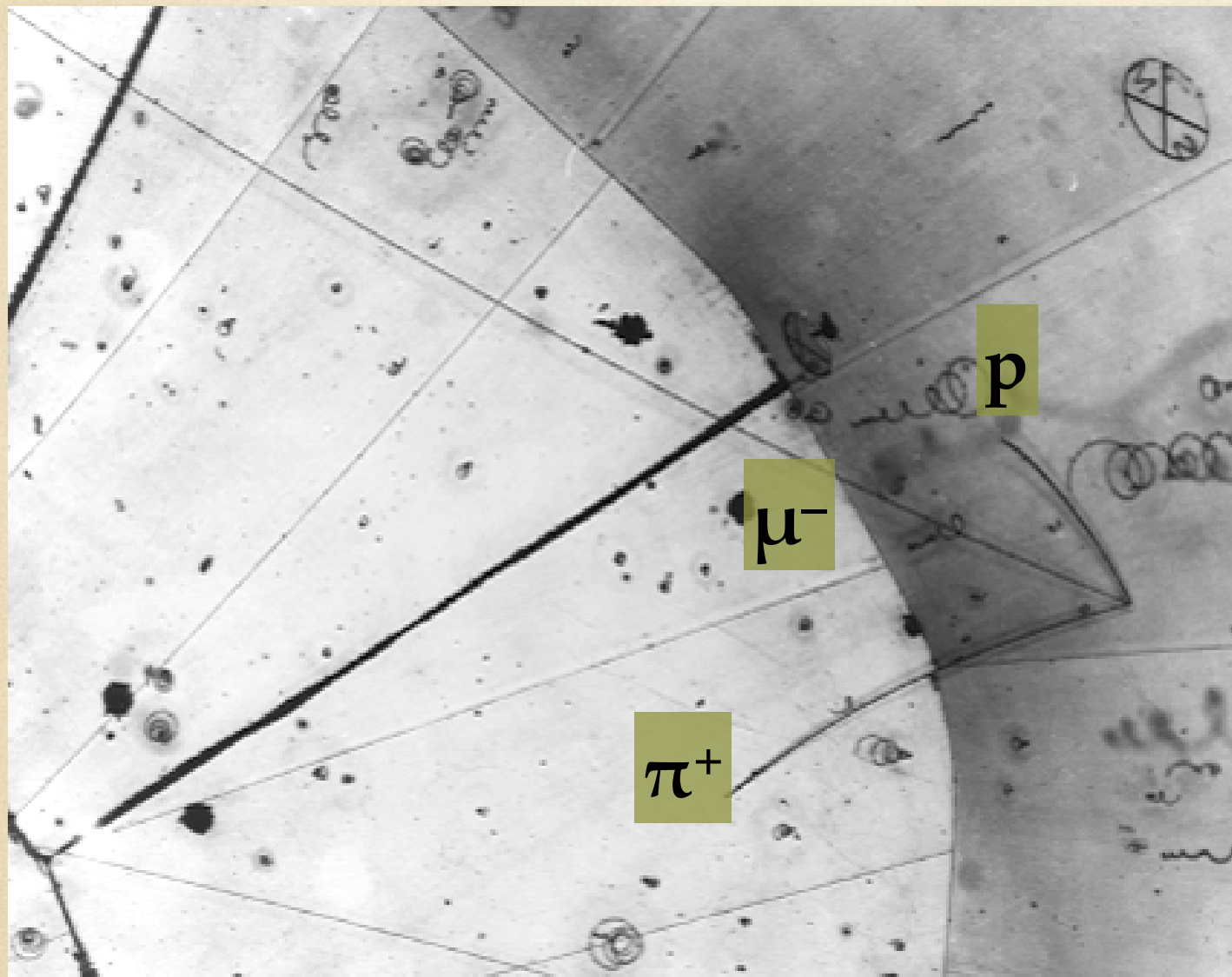


Wolfgang Pauli



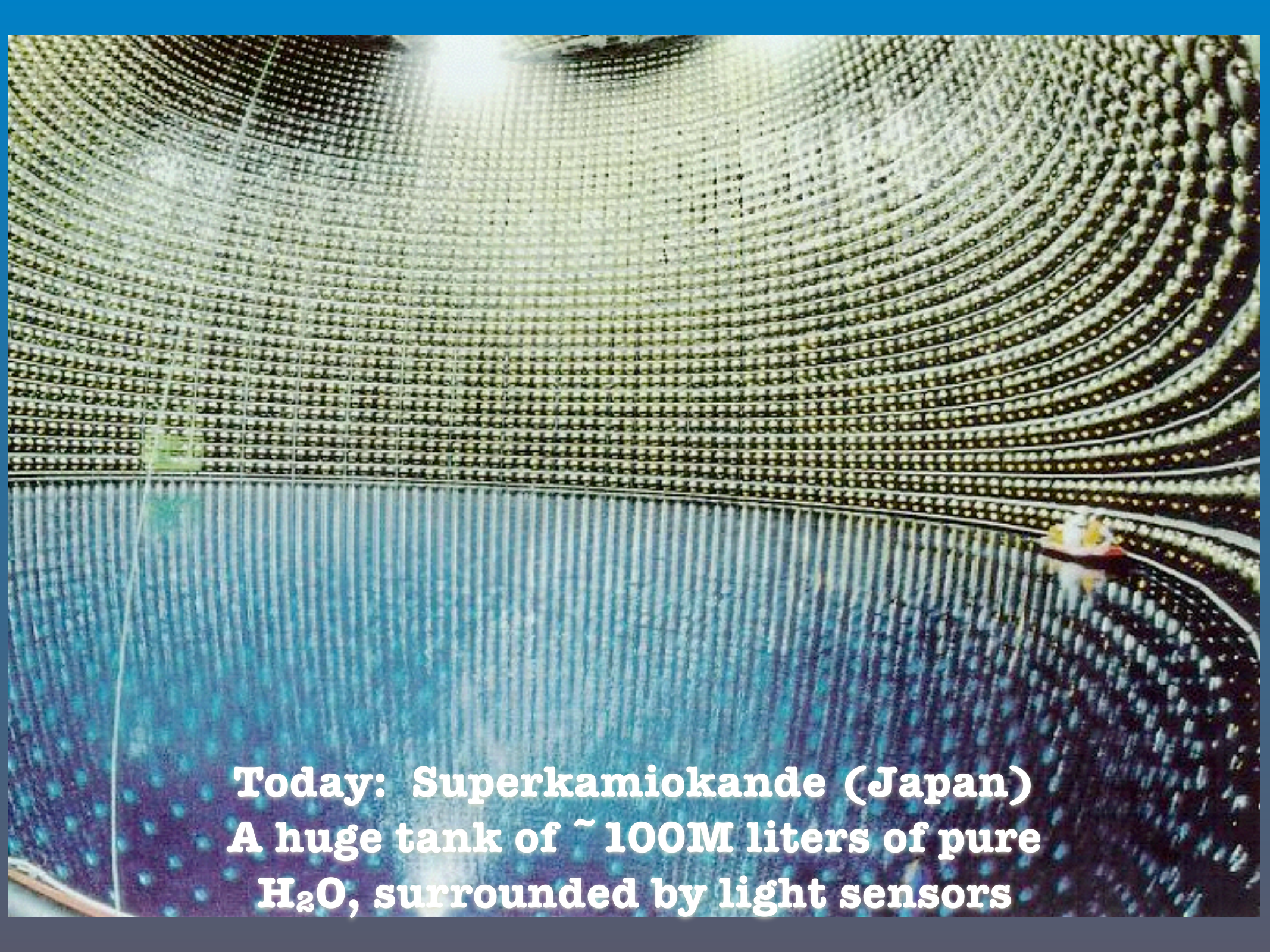
S.Dali, The invisible man, 1929

It took more than 20 years for a neutrino to be directly observed!

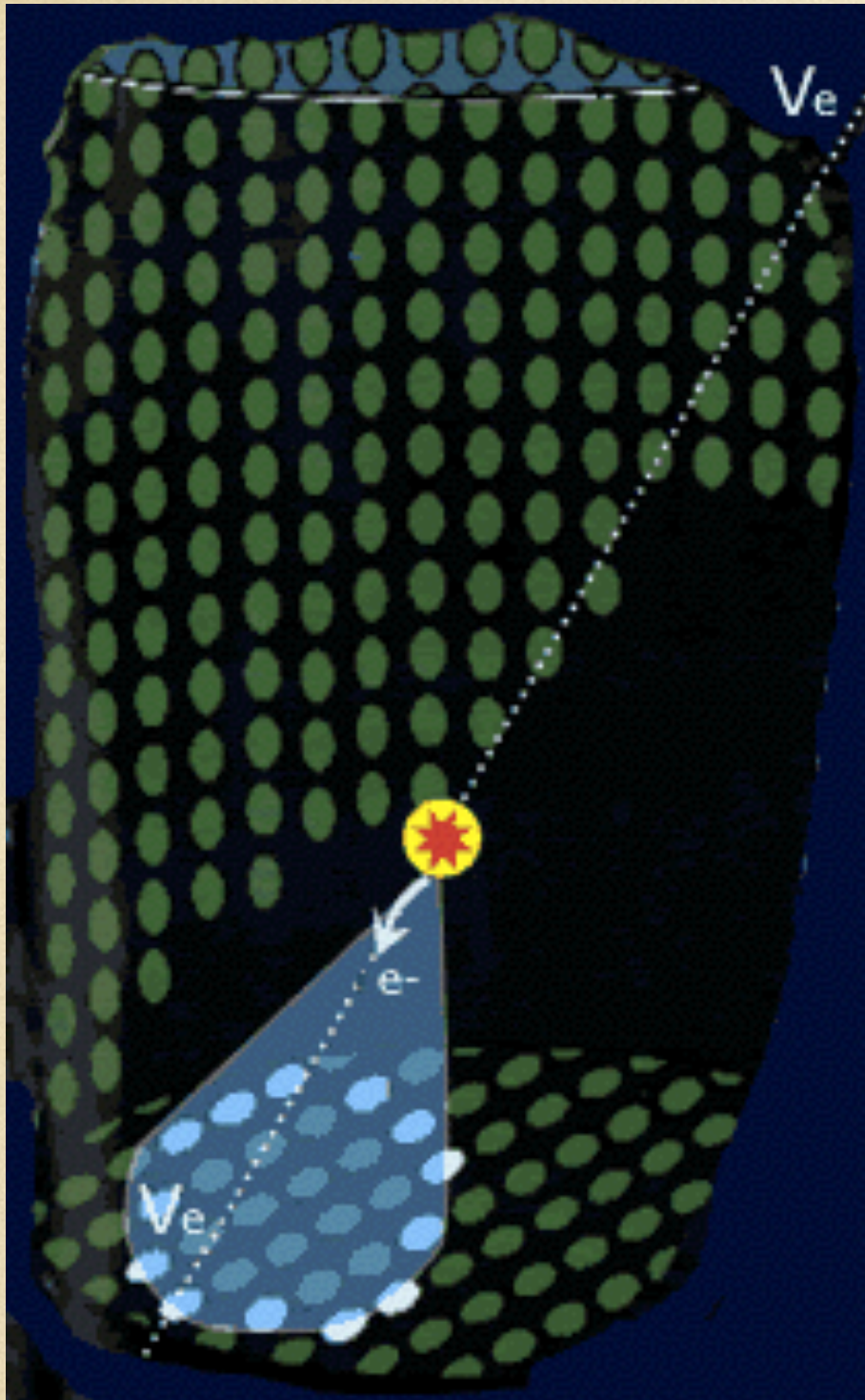


ν from a beam
of π mesons

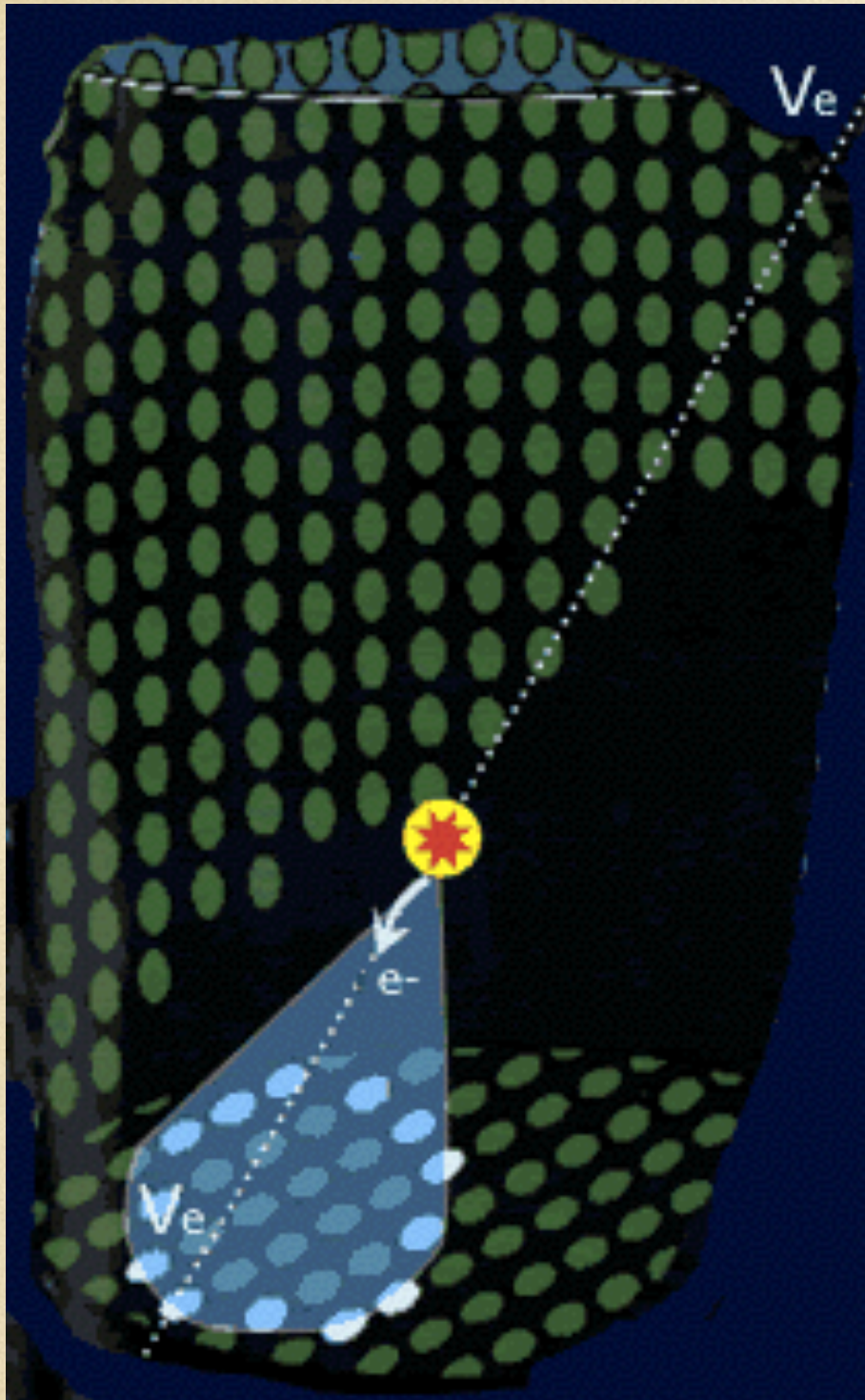
The world's first neutrino observation in a hydrogen bubble chamber. It was found Nov. 13, 1970, in this photograph from the Zero Gradient Synchrotron's 12-foot bubble chamber. The invisible neutrino strikes a proton where three particle tracks originate (lower right). The neutrino turns into a mu-meson, the long center track (extending up and left). The short track is the proton. The third track (extending down and left) is a pi-meson created by the collision.
Argonne National Laboratory

The image shows the interior of the Superkamiokande detector, a large cylindrical tank filled with pure water. The tank is surrounded by a dense array of light sensors, which are arranged in a grid pattern on the inner walls of the tank. The sensors are connected to a complex network of cables and pipes. The water is clear and blue, and the overall atmosphere is one of a large-scale scientific experiment.

Today: Superkamiokande (Japan)
A huge tank of ~ 100M liters of pure
H₂O, surrounded by light sensors

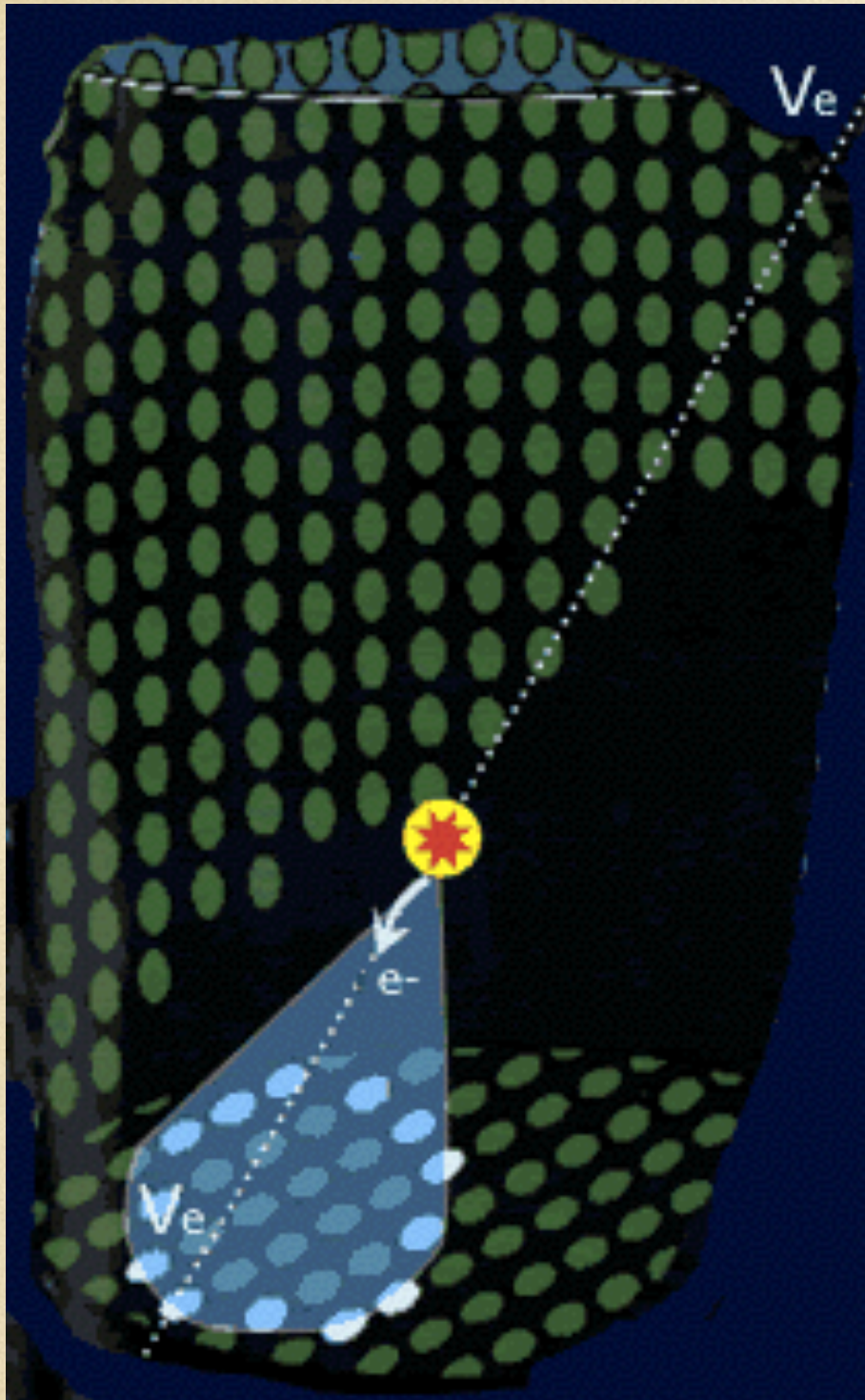


Neutrino from the collision of a cosmic ray with the earth atmosphere, or from the sun



Neutrino from the collision of a cosmic ray with the earth atmosphere, or from the sun

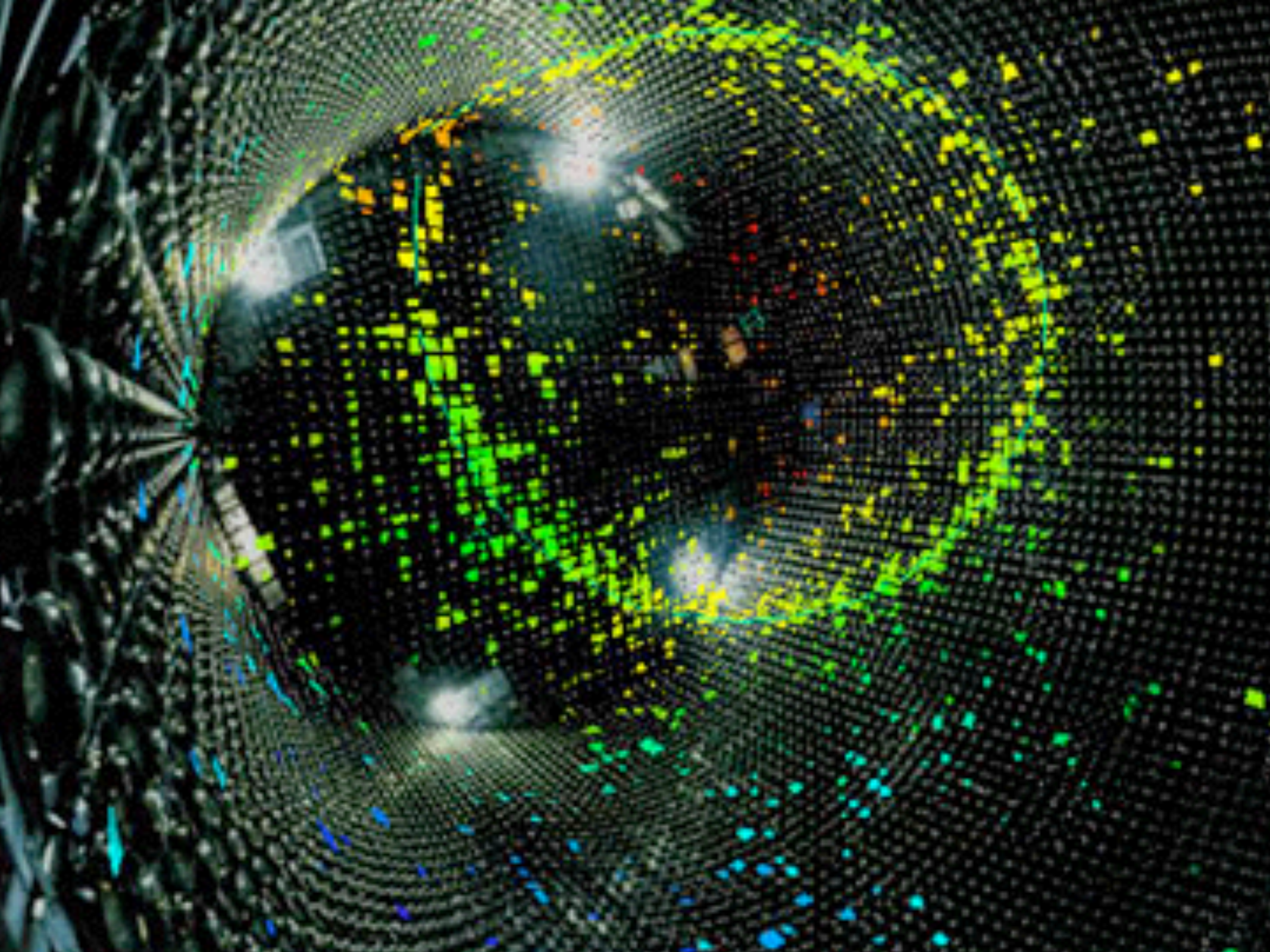
It interacts with an atom in the water, and becomes an electron



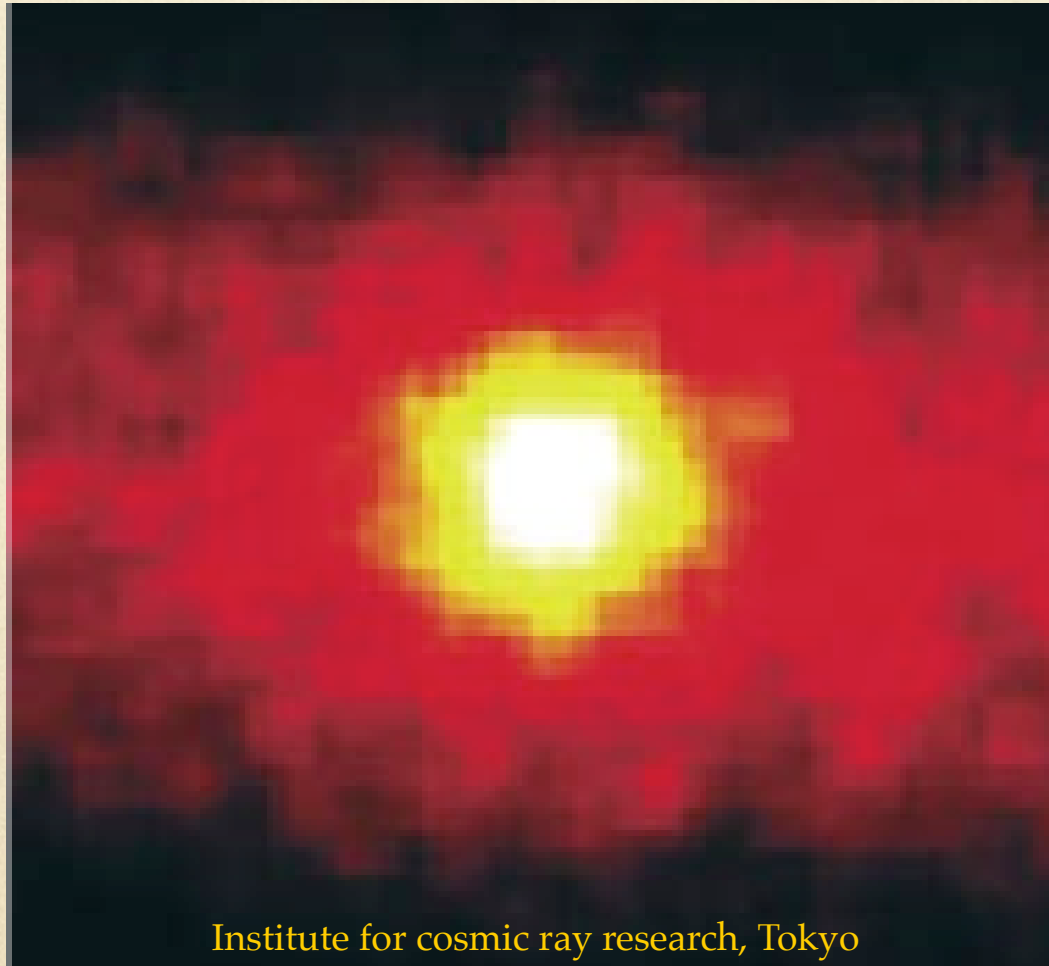
Neutrino from the collision of a cosmic ray with the earth atmosphere, or from the sun

It interacts with an atom in the water, and becomes an electron

The electron travels superluminal in water, and creates a light-bang – the luminous equivalent of an airplane supersonic bang – to be detected by the sensors on the surface of the tank



Reconstructing the neutrino direction, and mapping on the sky the position of their origin, allows to use neutrino detectors as “telescopes”: neutrino eyes!



Institute for cosmic ray research, Tokyo

A picture of the invisible part of the Sun, namely its innermost core, where nuclear reactions take place!

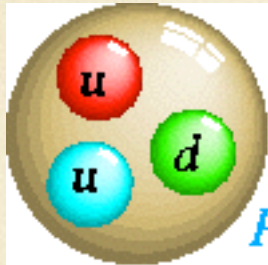
90 years after they entered in our understanding of nature, neutrinos are still among the most intriguing elements of the Standard Model (SM)

- what are the precise values of their masses?
- why is their mass so much smaller than all other SM particles?
- are neutrinos their own antiparticles?
- how many types of neutrinos are there?
- are they subject to interactions other than SM ones?
-

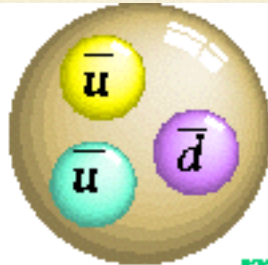
The continued exploration of neutrino properties forms one of the pillars of the future programme of experimental particle physics worldwide

“It is with π -mesons and the most
gelatinous and indeterminate
neutrinos that I want to paint the
beauty of the angles and of reality.”

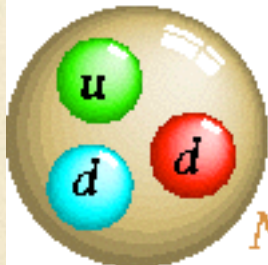
S.Dali, ‘Anti-matter manifesto’ (1958)



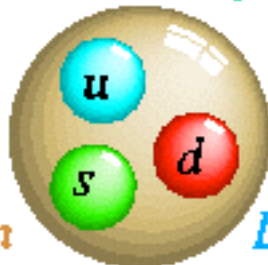
Proton



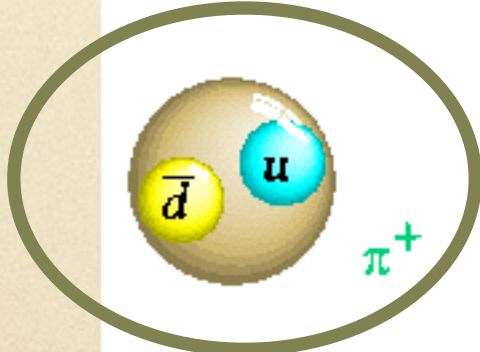
Anti-proton



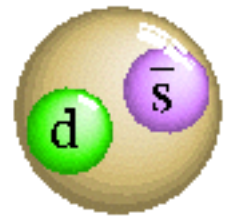
Neutron



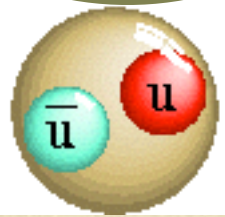
Lambda



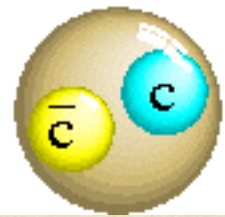
π^+



K^0



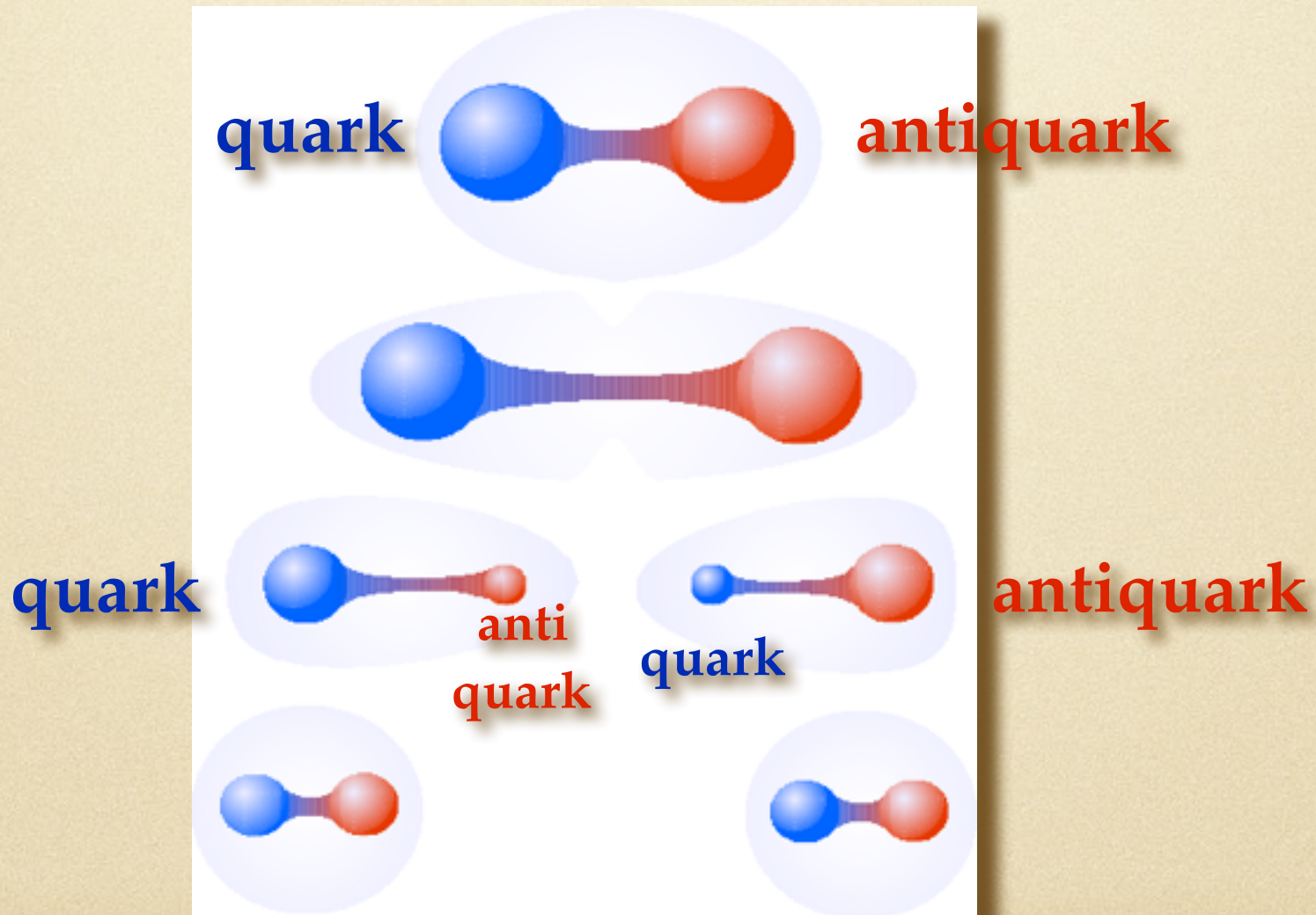
π^0



J/ψ

The ultimate invisible: quarks inside matter

If we try to pull the quarks out of a proton or a pion, the energy we need to win the strong force will eventually convert into a new quark-antiquark pair (using $E=mc^2$), and we'll be left with two pions ...
we know quarks are there, but can't get them out!

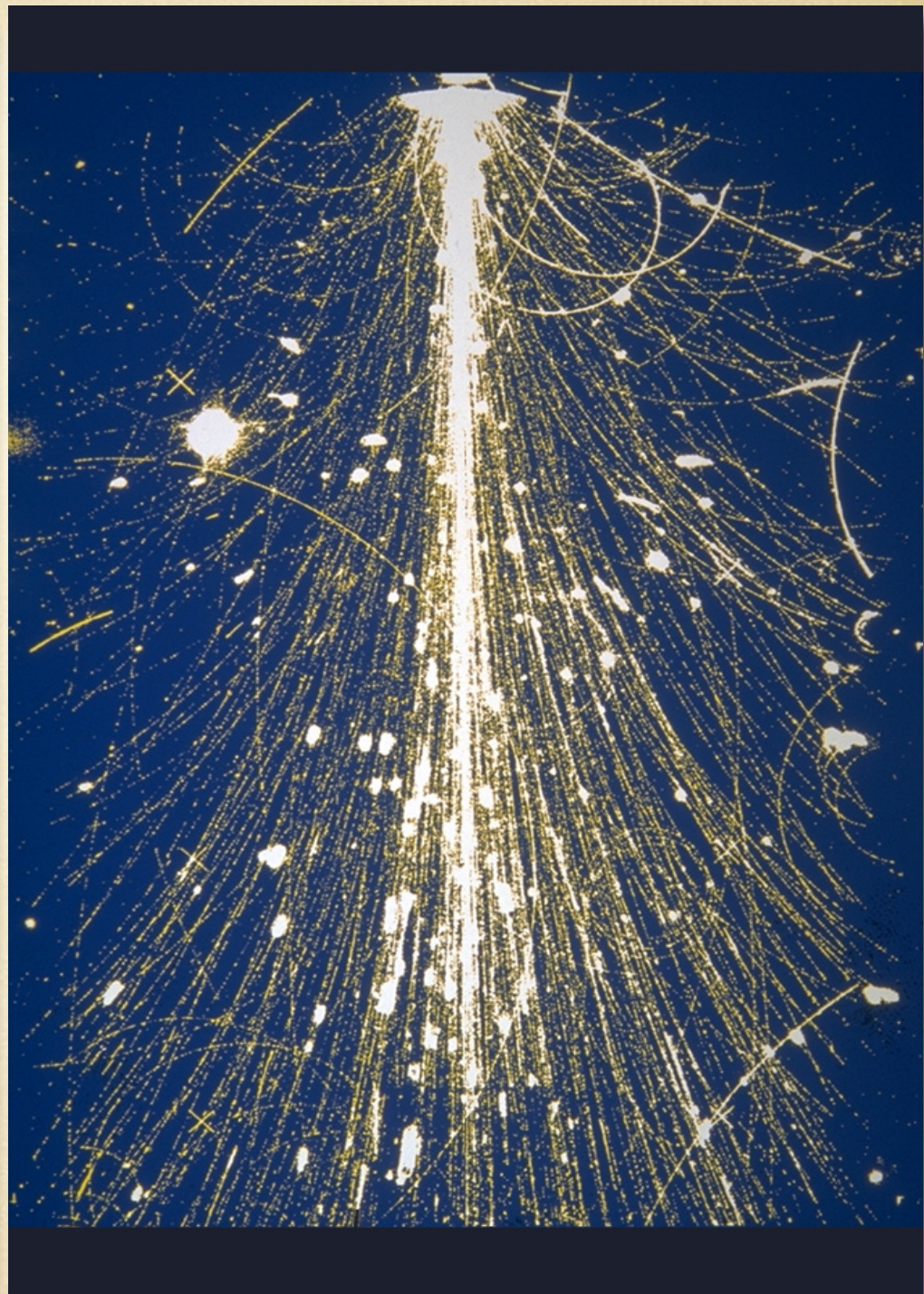




Saint surrounded by three π mesons, 1956



Saint surrounded by three π mesons, 1956



... just π mesons, in a real experiment ...

Some of the main
open questions

**What is the vacuum really
made of?**

What is the vacuum really made of?

S.Dali,
The echo of void
1935



The vacuum, and the Higgs field

We call vacuum the state of any volume of the Universe if we were to take away from it all matter and interactions from nearby matter.

The Standard Model predicts that the vacuum is occupied by a constant density field of the Higgs boson, which we cannot “take away”.

This permeates the Universe like an ether, everywhere and permanently, since about 10^{-10} seconds after the Big Bang

Interacting with this field, particles acquire their mass

Producing Higgs bosons

Like any other medium, the Higgs continuum background can be perturbed. Similarly to what happens if we bang on a table, creating sound waves, if we “bang” on the Higgs background (something achieved by concentrating a lot of energy in a small volume) we can stimulate “Higgs waves”. These waves manifest themselves as particles, the so-called Higgs bosons

**What is required is that the energy available be larger than the Higgs mass
⇒ particle accelerators !!!**

What's the origin of invisible dark matter ?



L'homme invisible, 1929

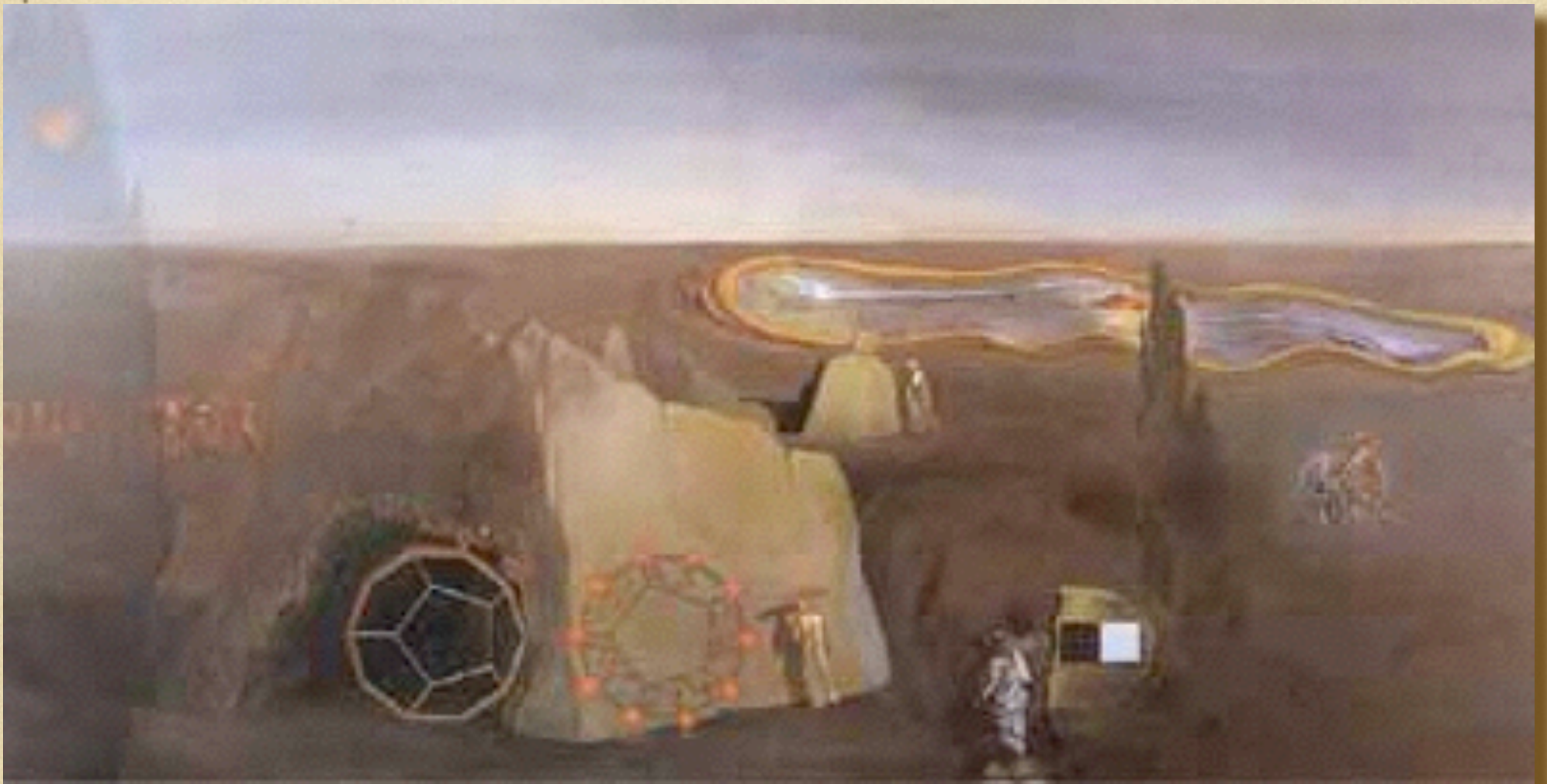
Whatever happened to the antimatter in the Universe?



Antiprotonic assumption, '56

Why do we live in 4 dimensions?

are there more, hidden and invisible dimensions of space-time?



In the search of the 4th dimension (1979)

What will the next big discovery be?

What will the next big discovery be?

- a new (set of) particle(s) (*eg the DM particle, or those predicted by Supersymmetry*)?

What will the next big discovery be?

- a new (set of) particle(s) (*eg the DM particle, or those predicted by Supersymmetry*)?
- a new interaction (*eg a new weak force, restoring the parity asymmetry of radioactive phenomena*)?

What will the next big discovery be?

- a new (set of) particle(s) (*eg the DM particle, or those predicted by Supersymmetry*)?
- a new interaction (*eg a new weak force, restoring the parity asymmetry of radioactive phenomena*)?
- a new layer of substructure inside those we consider today as elementary particles?

What will the next big discovery be?

- a new (set of) particle(s) (*eg the DM particle, or those predicted by Supersymmetry*)?
- a new interaction (*eg a new weak force, restoring the parity asymmetry of radioactive phenomena*)?
- a new layer of substructure inside those we consider today as elementary particles?
- a new conceptual paradigm (*like the discovery that the speed of light is constant, or energy levels are quantized*)?

What will the next big discovery be?

- a new (set of) particle(s) (*eg the DM particle, or those predicted by Supersymmetry*)?
- a new interaction (*eg a new weak force, restoring the parity asymmetry of radioactive phenomena*)?
- a new layer of substructure inside those we consider today as elementary particles?
- a new conceptual paradigm (*like the discovery that the speed of light is constant, or energy levels are quantized*)?
- all of the above ??

Final remarks

Final remarks

- The depth of our understanding of Nature has reached new levels, uncovering the existence, origin and the role of many *invisibles*

Final remarks

- The depth of our understanding of Nature has reached new levels, uncovering the existence, origin and the role of many *invisibles*
- Progress continues on the remaining ones, but with the understanding that ...

Final remarks

- The depth of our understanding of Nature has reached new levels, uncovering the existence, origin and the role of many *invisibles*
- Progress continues on the remaining ones, but with the understanding that ...
- there will always be a new layer of unknown invisible, the invisible hand of nature that decided that, after all, there should be **something instead of nothing**