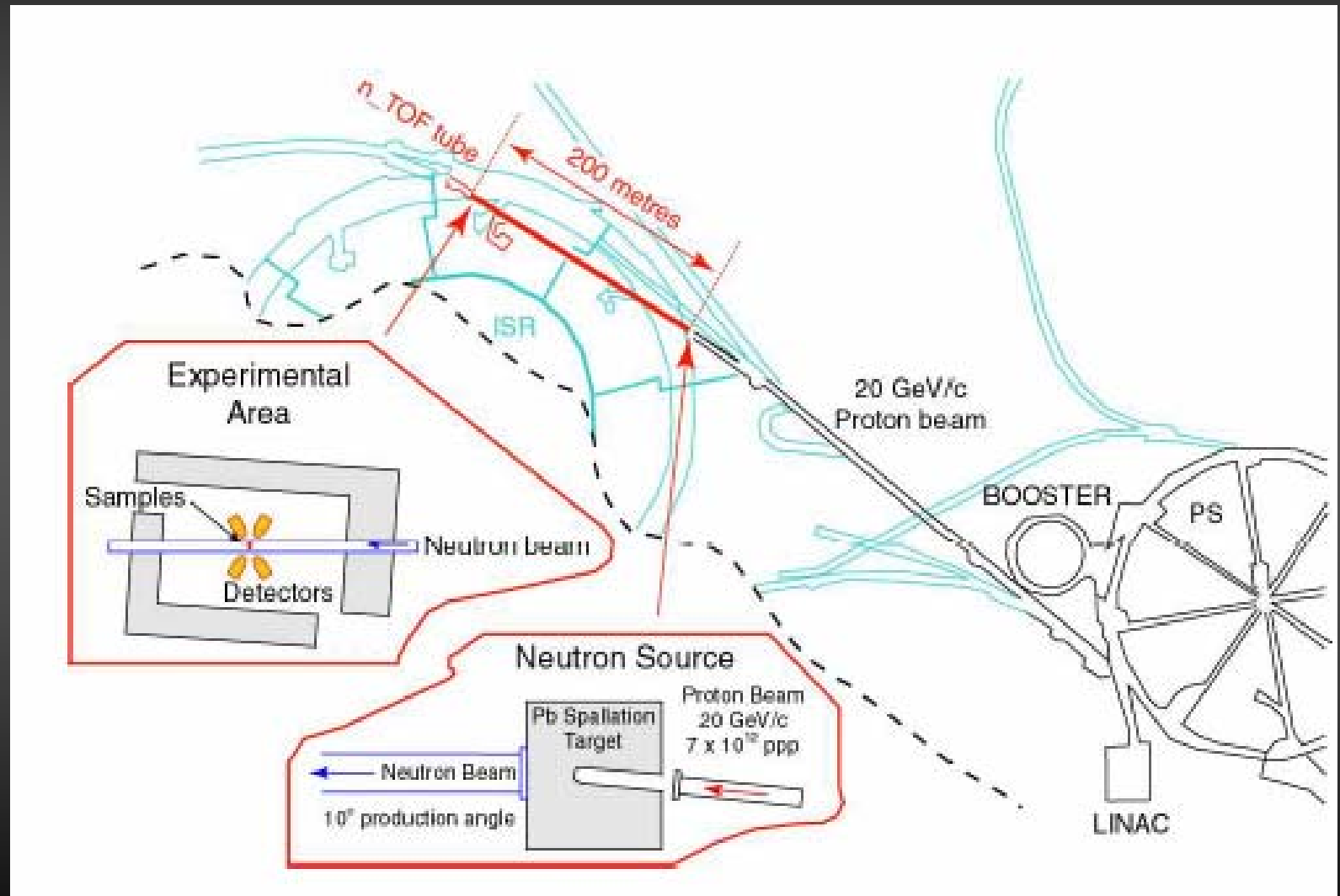


# The Physics Programme at n\_TOF (in 2007 and beyond)

Alberto Mengoni  
IAEA , Vienna

- Yesterday
- Today
- Tomorrow (n\_TOF-Phase 2)

# The n\_TOF facility at CERN



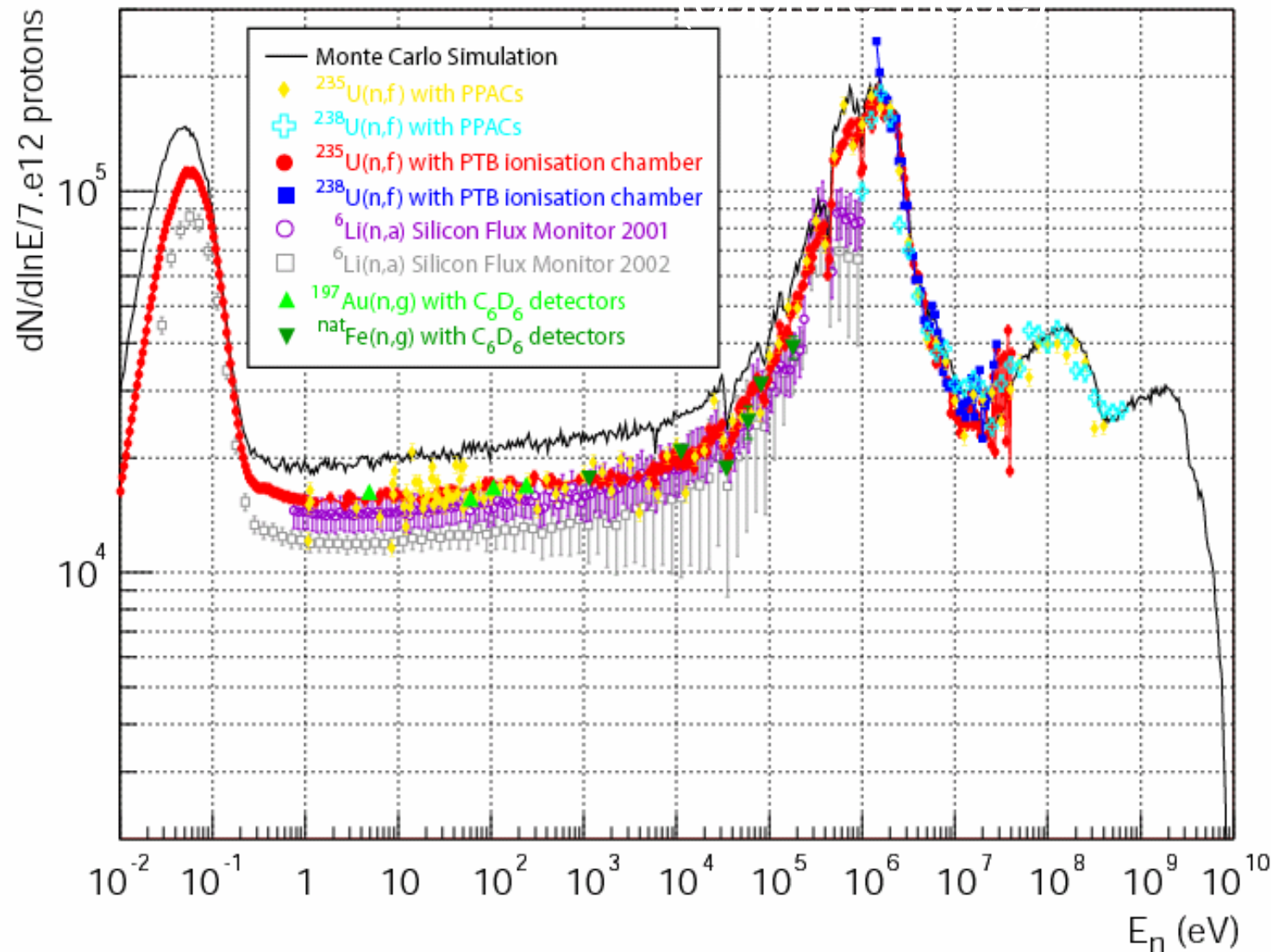
# n\_TOF basic parameters

proton beam momentum	20 GeV/c
intensity (dedicated mode)	$7 \times 10^{12}$ protons/pulse
repetition frequency	1 pulse/2.4s
pulse width	6 ns (rms)
n/p	300
lead target dimensions	80x80x60 cm <sup>3</sup>
cooling & moderation material	H <sub>2</sub> O
moderator thickness in the exit face	5 cm
neutron beam dimension in EAR-1 (capture mode)	2 cm (FWHM)

# n\_TOF beam characteristics

## ■ the neutron flux

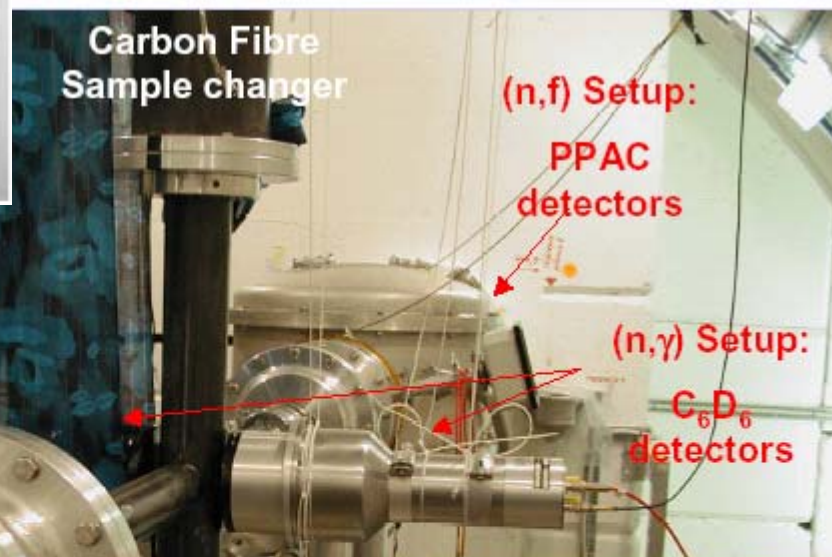
2<sup>nd</sup> collimator  $\phi=1.8$  cm



Performance Report  
CERN-INTC-2002-037, January 2003  
CERN-SL-2002-053 ECT

# The real world

- $n$ \_TOF commissioned in 2001-2002



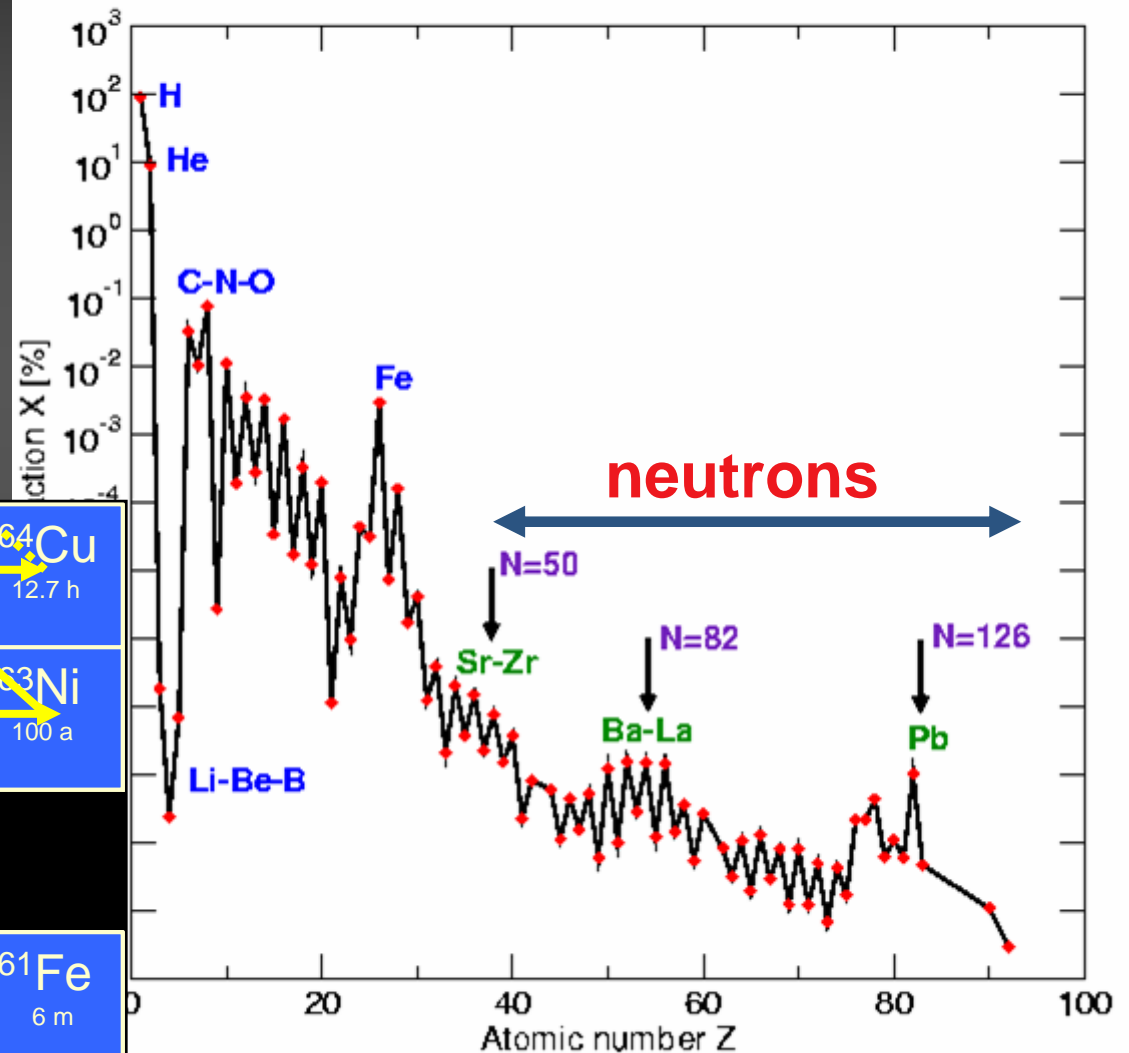
# Objectives of the activity at n\_TOF (n\_TOF Phase-2, April 2005)

1. Cross sections relevant for Nuclear Astrophysics
2. Measurements of neutron cross sections relevant for Nuclear Waste Transmutation and related Nuclear Technologies
3. Neutrons as probes for fundamental Nuclear Physics & other applications

# Nucleosynthesis: the s-process

- ½ of the elements above Fe are produced by the s-process
- The astrophysical sites of the s-process are:
  - He burning in intermediate/massive stars
  - Low-mass AGB's
- There exists a direct correlation between the neutron capture cross section and the abundance ( $\sigma(n, \gamma) \cdot N = const.$ )
- The neutron capture cross sections are key ingredients for s-process nucleosynthesis

Solar system elemental abundances



## The canonical s-process

Cu			<b>62Cu</b> 9.74 m	63Cu 69.17	64Cu 12.7 h	
Ni		60Ni 26.223	61Ni 1.140	62Ni 3.634	63Ni 100 a	
Co		<b>58Co</b> 70.86 d	59Co 100	60Co 5.272 a	61Co 1.65 h	
Fe	56Fe 91.72	57Fe 2.2	58Fe 0.28	59Fe 44.503 d	60Fe 1.5 10 <sup>6</sup> a	61Fe 6 m

Yellow arrows indicate the s-process path: 56Fe → 57Fe → 58Fe → 59Fe → 60Fe → 61Fe → 60Co → 61Co → 62Ni → 63Ni → 64Cu.

# Nuclear waste: TRU

(1000 MW<sub>e</sub> LWR)

	Cm 238 2,4 h	Cm 239 3 h	Cm 240 27 d	Cm 241 32,8 d	Cm 242 162,94 d	Cm 243 29,1 a	Cm 244 18,10 a	Cm 245 8500 a	Cm 246 4730 a
Am 236 ? 3,7 m	Am 237 73,0 m	Am 238 1,63 h	Am 239 11,9 h	Am 240 50,8 h	Am 241 432,2 a	Am 242 141 a	Am 243 7370 a	Am 244 26 m	Am 245 2,05 h
Pu 235 25,3 m	Pu 236 2,858 a	Pu 237 45,2 d	Pu 238 87,74 a	Pu 239 2,411 · 10 <sup>4</sup> a	Pu 240 6563 a	Pu 241 14,35 a	Pu 242 3,750 · 10 <sup>5</sup> a	Pu 243 4,956 h	Pu 244 8,00 · 10 <sup>7</sup> a
Np 234 4,4 d	Np 235 396,1 d	Np 236 22,5 h	Np 237 2,144 · 10 <sup>6</sup> a	Np 238 2,117 d	Np 239 2,355 d	Np 240 7,22 m	Np 241 13,9 m	Np 242 2,2 m	Np 243 1,85 m
U 233 1,592 · 10 <sup>5</sup> a	U 234 0,0055	U 235 0,7200	U 236 2,342 · 10 <sup>7</sup> a	U 237 6,75 d	U 238 99,2745	U 239 23,5 m	U 240 14,1 h		U 242 16,8 m
Pa 232 1,31 d	Pa 233 27,0 d	Pa 234 1,17 m	Pa 235 34,2 m	Pa 236 9,1 m	Pa 237 8,7 m	Pa 238 2,3 m			
Th 231 25,5 h	Th 232 100	Th 233 22,3 m	Th 234 24,10 d	Th 235 7,1 m	Th 236 37,5 m	Th 237 5,0 m			

244Cm  
1.5 Kg/yr

241Am: 11.6 Kg/yr  
243Am: 4.8 Kg/yr

239Pu: 125 Kg/yr

237Np: 16 Kg/yr

LLFP  
76.2 Kg/yr

LLFP

source: Actinide and Fission Product Partitioning and Transmutation – NEA (1999)



## Capture

$^{151}\text{Sm}$

$^{204,206,207,208}\text{Pb}$ ,  $^{209}\text{Bi}$

$^{232}\text{Th}$

$^{24,25,26}\text{Mg}$

$^{90,91,92,94,96}\text{Zr}$ ,  $^{93}\text{Zr}$

$^{139}\text{La}$

$^{186,187,188}\text{Os}$

$^{233,234}\text{U}$

$^{237}\text{Np}$ ,  $^{240}\text{Pu}$ ,  $^{243}\text{Am}$

## Fission

$^{233,234,235,236,238}\text{U}$

$^{232}\text{Th}$

$^{209}\text{Bi}$

$^{237}\text{Np}$

$^{241,243}\text{Am}$ ,  $^{245}\text{Cm}$

# n\_TOF experiments 2002-4

## 10 groups of experiments endorsed by the INTC and approved by the RB:

- n\_TOF-01: Facility proposal
- n\_TOF-02: Commissioning
- n\_TOF-03: Sm & Mg
- n\_TOF-04: Os
- n\_TOF-05: Pb & Bi
- n\_TOF-06: Th-cycle fission
- n\_TOF-08: Zr & La
- n\_TOF-09: MA fission
- n\_TOF-10: MA capture (TAC)

## Capture

$^{151}\text{Sm}$

$^{204,206,207,208}\text{Pb}$ ,  $^{209}\text{Bi}$

$^{232}\text{Th}$

$^{24,25,26}\text{Mg}$

$^{90,91,92,94,96}\text{Zr}$ ,  $^{93}\text{Zr}$

$^{139}\text{La}$

$^{186,187,188}\text{Os}$

$^{233,234}\text{U}$

$^{237}\text{Np}$ ,  $^{240}\text{Pu}$ ,  $^{243}\text{Am}$

## Fission

$^{233,234,235,236,238}\text{U}$

$^{232}\text{Th}$

$^{209}\text{Bi}$

$^{237}\text{Np}$

$^{241,243}\text{Am}$ ,  $^{245}\text{Cm}$

# n\_TOF experiments 2002-4

28 weeks/yr (average)

483 effective 8hr-shft/yr

$1.3e19$  protons/yr

Problems during Phase-1 runs:

2<sup>nd</sup> collimator alignment (minor)

beam-requests (minor)

end of 2004 run increased activity in the cooling (major, to be discussed later)

## Capture

$^{151}\text{Sm}$

$^{204,206,207,208}\text{Pb}$ ,  $^{209}\text{Bi}$

$^{232}\text{Th}$

$^{24,25,26}\text{Mg}$

$^{90,91,92,94,96}\text{Zr}$ ,  $^{93}\text{Zr}$

$^{139}\text{La}$

$^{186,187,188}\text{Os}$

$^{233,234}\text{U}$

$^{237}\text{Np}$ ,  $^{240}\text{Pu}$ ,  $^{243}\text{Am}$

## Fission

$^{233,234,235,236,238}\text{U}$

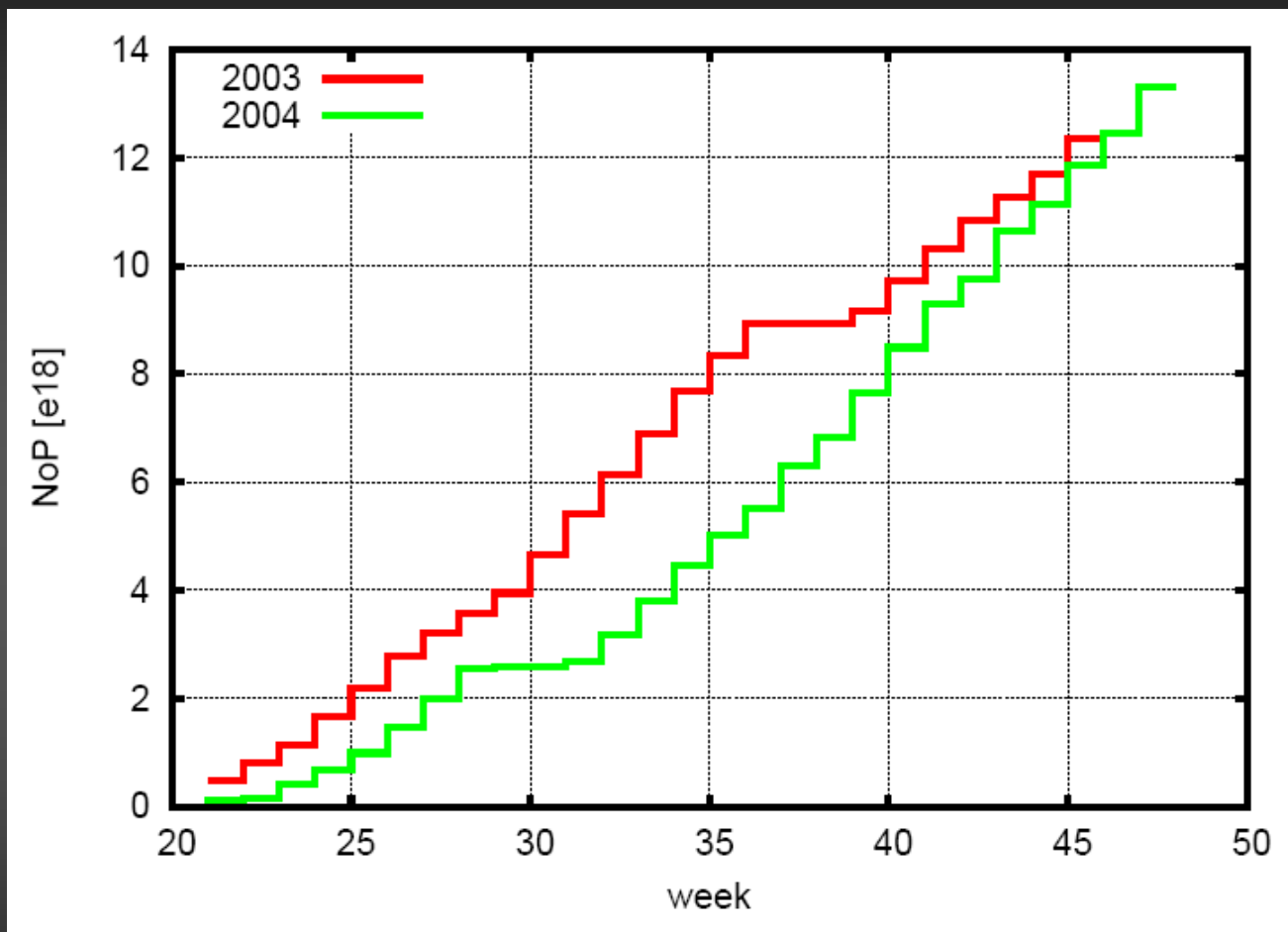
$^{232}\text{Th}$

$^{209}\text{Bi}$

$^{237}\text{Np}$

$^{241,243}\text{Am}$ ,  $^{245}\text{Cm}$

# n\_TOF experiments 2002-4



# The n\_TOF Collaboration

U.Abbondanno<sup>14</sup>, G.Aerts<sup>7</sup>, H.Álvarez<sup>24</sup>, F.Alvarez-Velarde<sup>20</sup>, S.Andriamonje<sup>7</sup>, J.Andrzejewski<sup>33</sup>, P.Assimakopoulos<sup>9</sup>, L.Audouin<sup>5</sup>, G.Badurek<sup>1</sup>, P.Baumann<sup>6</sup>, F. Bečvář<sup>31</sup>, J.Benlliure<sup>24</sup>, E.Berthoumieux<sup>7</sup>, F.Calviño<sup>25</sup>, D.Cano-Ott<sup>20</sup>, R.Capote<sup>23</sup>, A.Carrillo de Albornoz<sup>30</sup>, P.Cennini<sup>4</sup>, V.Chepel<sup>17</sup>, E.Chiaveri<sup>4</sup>, N.Colonna<sup>3</sup>, G.Cortes<sup>25</sup>, D.Cortina<sup>24</sup>, A.Couture<sup>29</sup>, J.Cox<sup>29</sup>, S.David<sup>5</sup>, R.Dolfini<sup>15</sup>, C.Domingo-Pardo<sup>21</sup>, W.Dridi<sup>7</sup>, I.Duran<sup>24</sup>, M.Embida-Segura<sup>20</sup>, L.Ferrant<sup>5</sup>, A.Ferrari<sup>4</sup>, R.Ferreira-Marques<sup>17</sup>, L.Fitzpatrick<sup>4</sup>, H.Frais-Koelbl<sup>3</sup>, K.Fujii<sup>13</sup>, W.Furman<sup>18</sup>, C.Guerrero<sup>20</sup>, I.Goncalves<sup>30</sup>, R.Gallino<sup>36</sup>, E.Gonzalez-Romero<sup>20</sup>, A.Goverdovski<sup>19</sup>, F.Gramegna<sup>12</sup>, E.Griesmayer<sup>3</sup>, F.Gunsing<sup>7</sup>, B.Haas<sup>32</sup>, R.Haight<sup>27</sup>, M.Heil<sup>8</sup>, A.Herrera-Martinez<sup>4</sup>, M.Igashira<sup>37</sup>, S.Isaev<sup>5</sup>, E.Jericha<sup>1</sup>, Y.Kadi<sup>4</sup>, F.Käppeler<sup>8</sup>, D.Karamanis<sup>9</sup>, D.Karadimos<sup>9</sup>, M.Kerveno<sup>6</sup>, V.Ketlerov<sup>19</sup>, P.Koehler<sup>28</sup>, V.Konovalov<sup>18</sup>, E.Kossionides<sup>39</sup>, M.Krtička<sup>31</sup>, C.Lamboudis<sup>10</sup>, H.Leeb<sup>1</sup>, A.Lindote<sup>17</sup>, I.Lopes<sup>17</sup>, M.Lozano<sup>23</sup>, S.Lukic<sup>6</sup>, J.Marganec<sup>33</sup>, L.Marques<sup>30</sup>, S.Marrone<sup>13</sup>, P.Mastinu<sup>12</sup>, A.Mengoni<sup>4</sup>, P.M.Milazzo<sup>14</sup>, C.Moreau<sup>14</sup>, M.Mosconi<sup>8</sup>, F.Neves<sup>17</sup>, H.Oberhummer<sup>1</sup>, S.O'Brien<sup>29</sup>, M.Oshima<sup>38</sup>, J.Pancin<sup>7</sup>, C.Papachristodoulou<sup>9</sup>, C.Papadopoulos<sup>40</sup>, C.Paradela<sup>24</sup>, N.Patronis<sup>9</sup>, A.Pavlik<sup>2</sup>, P.Pavlopoulos<sup>34</sup>, L.Perrot<sup>7</sup>, R.Plag<sup>8</sup>, A.Plompen<sup>16</sup>, A.Plukis<sup>7</sup>, A.Poch<sup>25</sup>, C.Pretel<sup>25</sup>, J.Quesada<sup>23</sup>, T.Rauscher<sup>26</sup>, R.Reifarth<sup>27</sup>, M.Rosetti<sup>1</sup>, C.Rubbia<sup>5</sup>, G.Rudolf<sup>6</sup>, P.Rullhusen<sup>16</sup>, J.Salgado<sup>30</sup>, L.Sarchiapone<sup>4</sup>, C.Stephan<sup>5</sup>, G.Tagliente<sup>13</sup>, J.L.Tain<sup>21</sup>, L.Tassan-Got<sup>5</sup>, L.Tavora<sup>30</sup>, R.Terlizzi<sup>13</sup>, G.Vannini<sup>35</sup>, P.Vaz<sup>30</sup>, A.Ventura<sup>11</sup>, D.Villamarin<sup>20</sup>, M.C.Vincente<sup>20</sup>, V.Vlachoudis<sup>4</sup>, R.Vlastou<sup>40</sup>, F.Voss<sup>8</sup>, H.Wendler<sup>4</sup>, M.Wiescher<sup>29</sup>, K.Wisshak<sup>8</sup>

40 Research Institutions

120 researchers

# n\_TOF Collaboration structure & operation

## n\_TOF Collaboration Board Chairman, Spokesperson

- 23 n\_TOF teams
- all teams participates in the development, commissioning, and operation of detectors, daq, etc.
- all teams participated in all the experiments/measurements
- 16 partners in the n\_TOF-ND-ADS EC Project (FP5)  
2.4 MEUR investment in equipments  
(6.3 MEUR total budget)  
run over the period 2001-2004  
(completed)

## n\_TOF Collaboration Board Chairman & Spokesperson

- Teams involved in individual proposals/experiment (with individual spokesperson)
  - Fe, Ni, Zn & Se proposal
  - Fission measurements proposal
  - n-scattering proposal
  - ...
- Common M&O budget
- Some team involved in EC/FP6 projects (EFNUDAT, EUROTRANS)

# The n\_TOF-Ph2 experiments 2007 and beyond

## Capture measurements

Mo, Ru, Pd stable isotopes

r-process residuals calculation  
isotopic patterns in SiC grains

Fe, Ni, Zn, and Se (stable isotopes)  
 $^{79}\text{Se}$

s-process nucleosynthesis in massive stars  
accurate nuclear data needs for structural materials

$A \approx 150$  (isotopes varii)

s-process branching points  
long-lived fission products

$^{234,236}\text{U}$ ,  $^{231,233}\text{Pa}$

Th/U nuclear fuel cycle

$^{235,238}\text{U}$

standards, conventional U/Pu fuel cycle

$^{239,240,242}\text{Pu}$ ,  $^{241,243}\text{Am}$ ,  $^{245}\text{Cm}$

incineration of minor actinides

(\*) endorsed by INTC, spokespersons: JL Tain (Valencia), M Heil (GSI, Darmstadt)

# The n\_TOF-Ph2 experiments 2007 and beyond

## Fission measurements

MA	ADS, high-burnup, GEN-IV reactors
$^{235}\text{U}(n,f)$ with $p(n,p')$	new $^{235}\text{U}(n,f)$ cross section standard
$^{234}\text{U}(n,f)$	study of vibrational resonances at the fission barrier

## Other measurements

$^{147}\text{Sm}(n,\alpha)$ , $^{67}\text{Zn}(n,\alpha)$ , $^{99}\text{Ru}(n,\alpha)$ $^{58}\text{Ni}(n,p)$ , other $(n,lcp)$	p-process studies gas production in structural materials
Al, V, Cr, Zr, Th, $^{238}\text{U}(n,lcp)$	structural and fuel material for ADS and other advanced nuclear reactors
He, Ne, Ar, Xe	low-energy nuclear recoils (development of gas detectors)
$n+D_2$	neutron-neutron scattering length

# n\_TOF: resume activities in 2007

**All teams involved in Phase-1 expressed interest to continue the activities for n\_TOF Phase-2**

- Lol, January 2005
- Budget for M&O allocated by funding agencies
- new MoU, draft ready December 2006

**2.5 years after last neutron beam delivered to EAR-1**

**New PhD students need data to work on(\*)**

**EC I3 projects in FP6 (EFNUDAT, EUROTRANS) running**

(\*) so far: 13 PhD Thesis (completed), 6 in preparation (4 to be completed by 2007), 2 starting in 2007

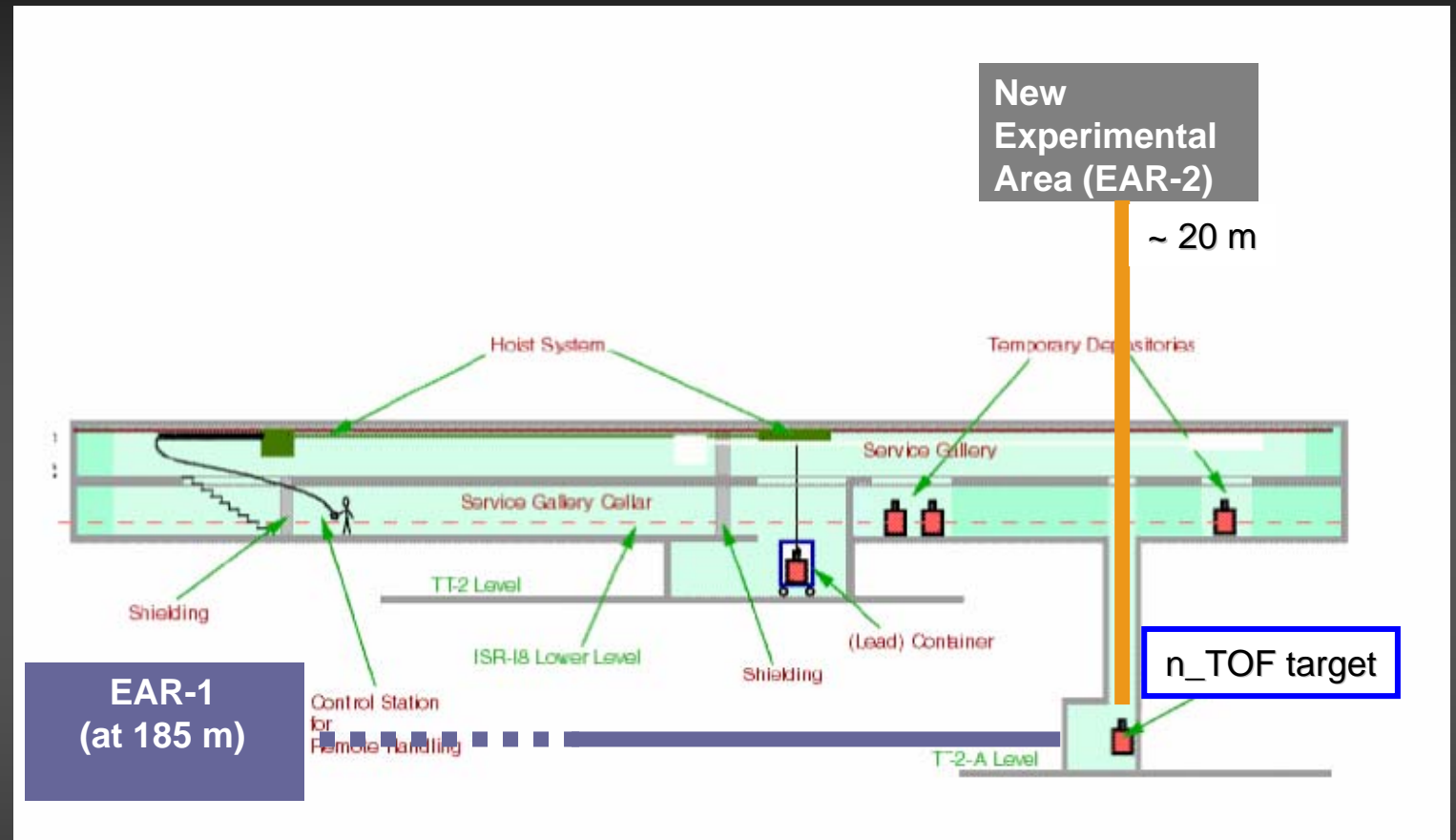


# **n\_TOF: resume activities in 2007!**

## **3 Questions:**

- 1. New target design, construction & installation**
- 2. Use of radioactive samples in EAR-1**
- 3. Perspectives for further developments & optimizations of the facility outlined (EAR-2)**

# The second n\_TOF beam line & EAR-2



Flight-path length : ~20 m  
at 90° respect to p-beam direction  
expected neutron flux enhancement: ~ 100  
drastic reduction of the  $t_0$  flash

# EAR-2: Optimized sensitivity

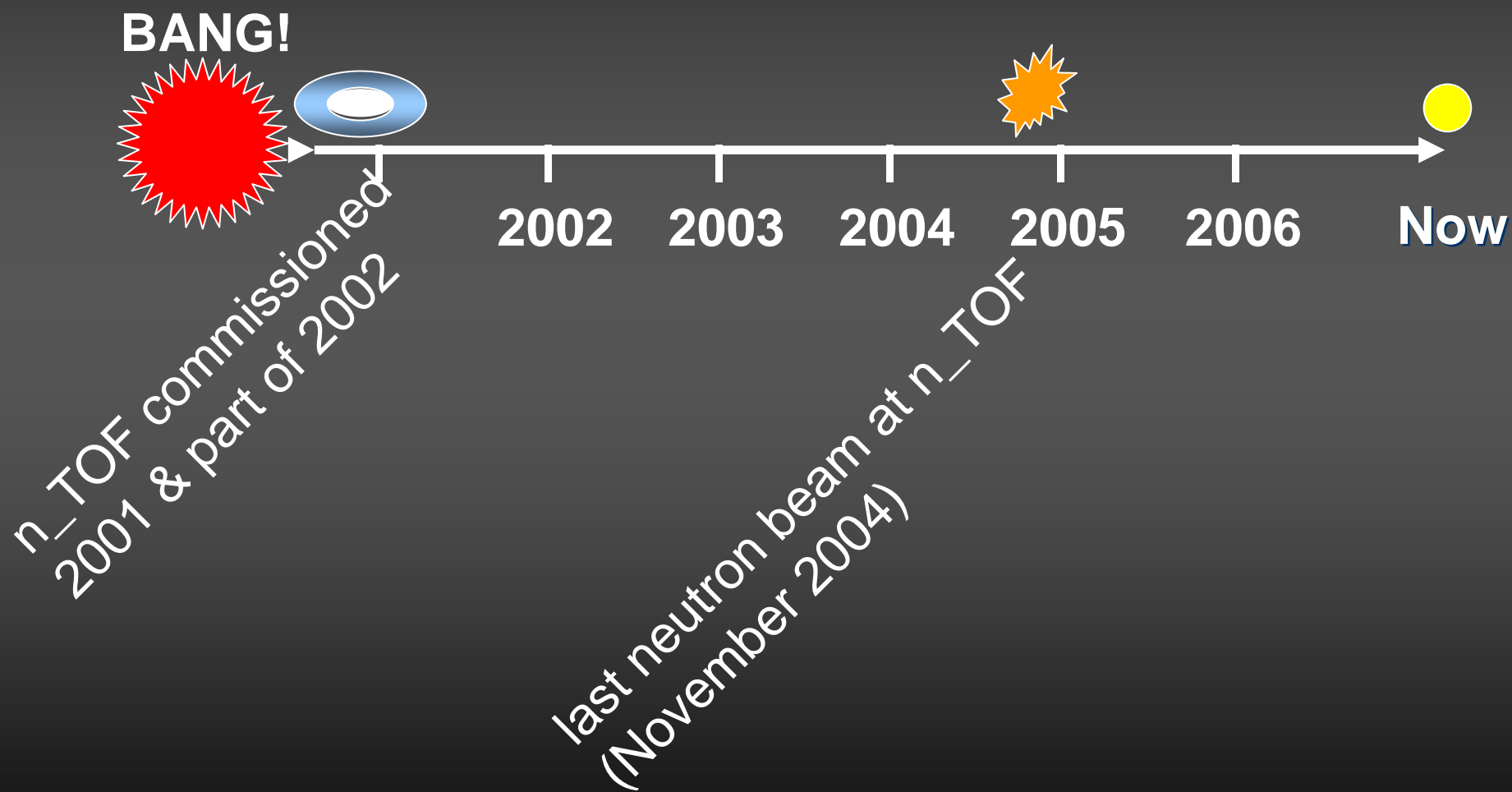
Improvements (ex: $^{151}\text{Sm}$ case)	consequences for sample mass
■ sample mass / 3 s/bkgd=1	✓ 50 mg
■ use $\text{BaF}_2$ TAC $\epsilon \times 10$	✓ 5 mg
■ use $\text{D}_2\text{O}$ $\Phi_{30} \times 5$	■ 1 mg
■ use 20 m flight path $\Phi_{30} \times 100$	■ 10 $\mu\text{g}$

boosts sensitivity by a factor of 5000 !



→ problems of sample production and safety issues relaxed

# Time at n\_TOF

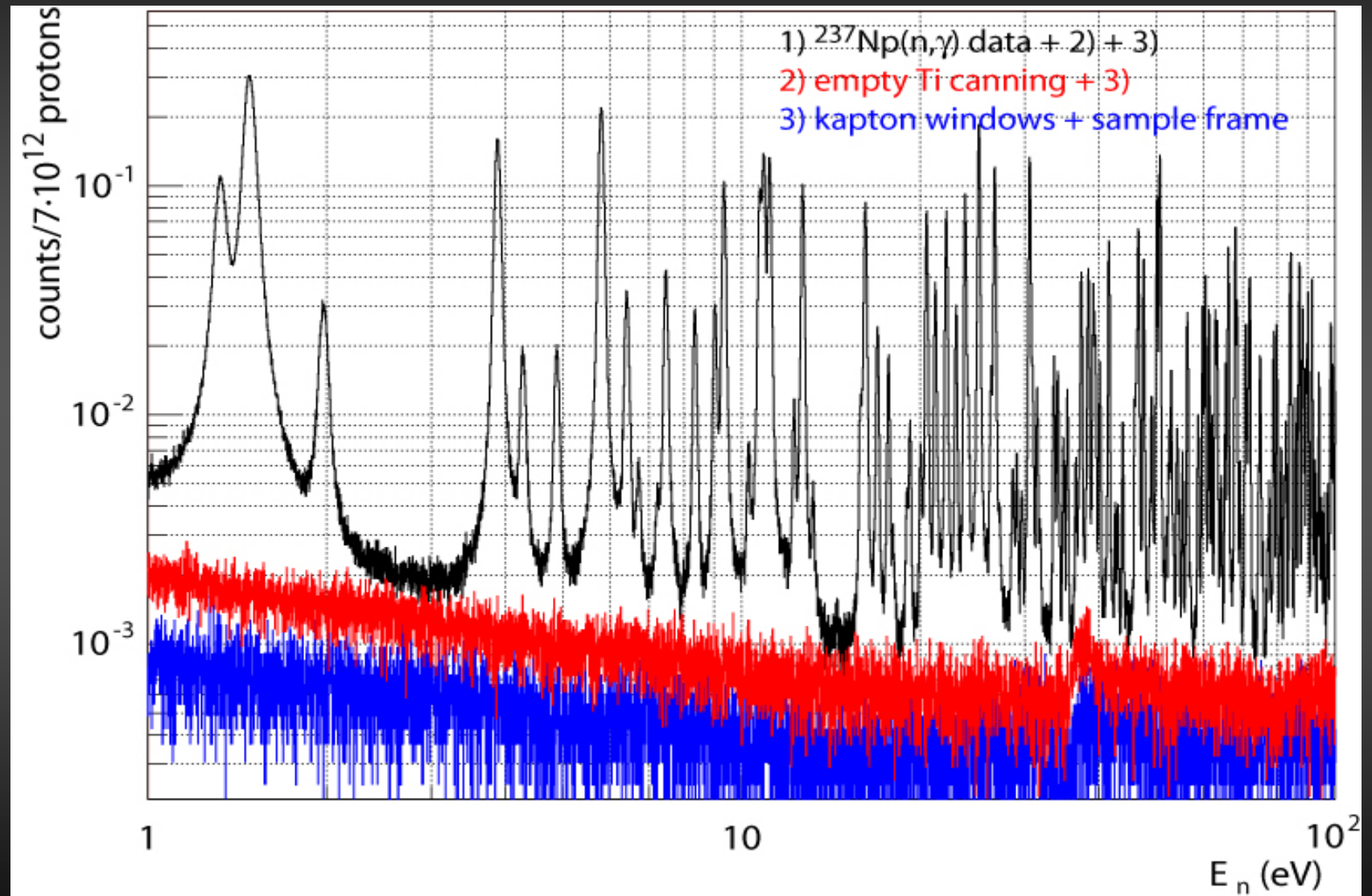


# The End



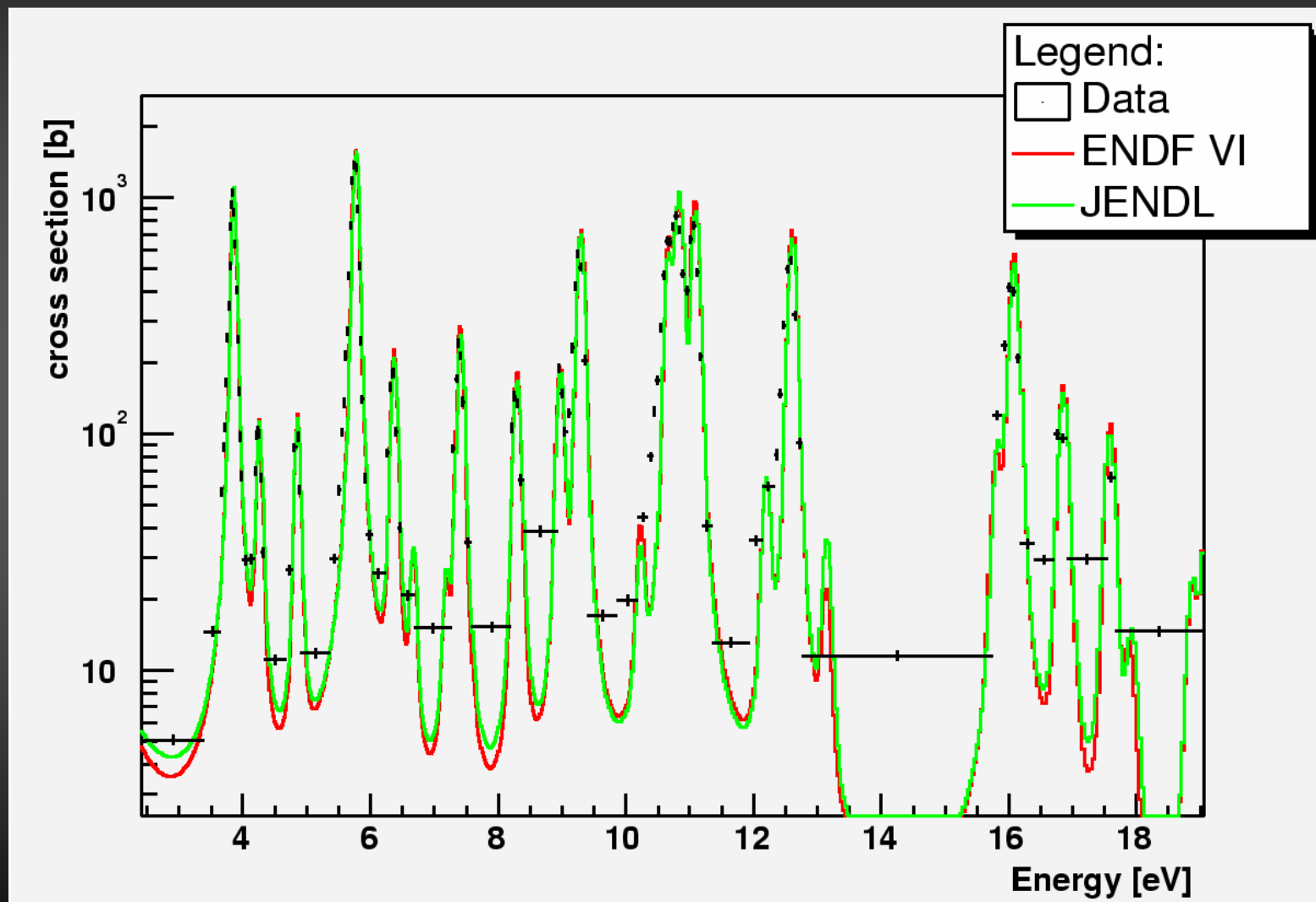
# $^{237}\text{Np}(n,\gamma)$ at n\_TOF

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# $^{237}\text{Np}(n,\gamma)$ at LANSCE

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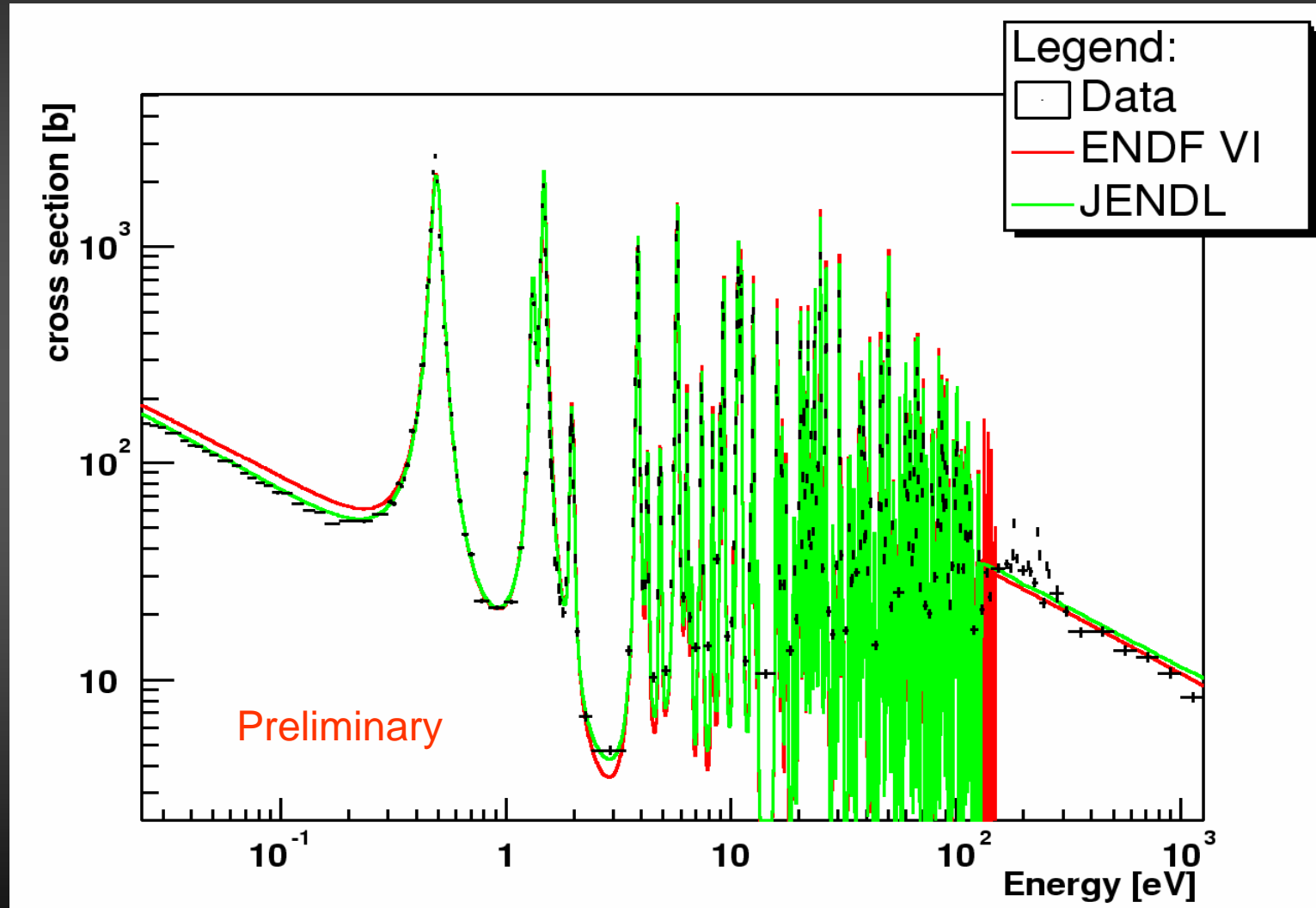
(Analysis by  
E.I. Esch and  
R. Reifarth)

Source: J Ullman, n\_BANT workshop, CERN, March 2005



# $^{237}\text{Np}(n,\gamma)$ at LANSCE

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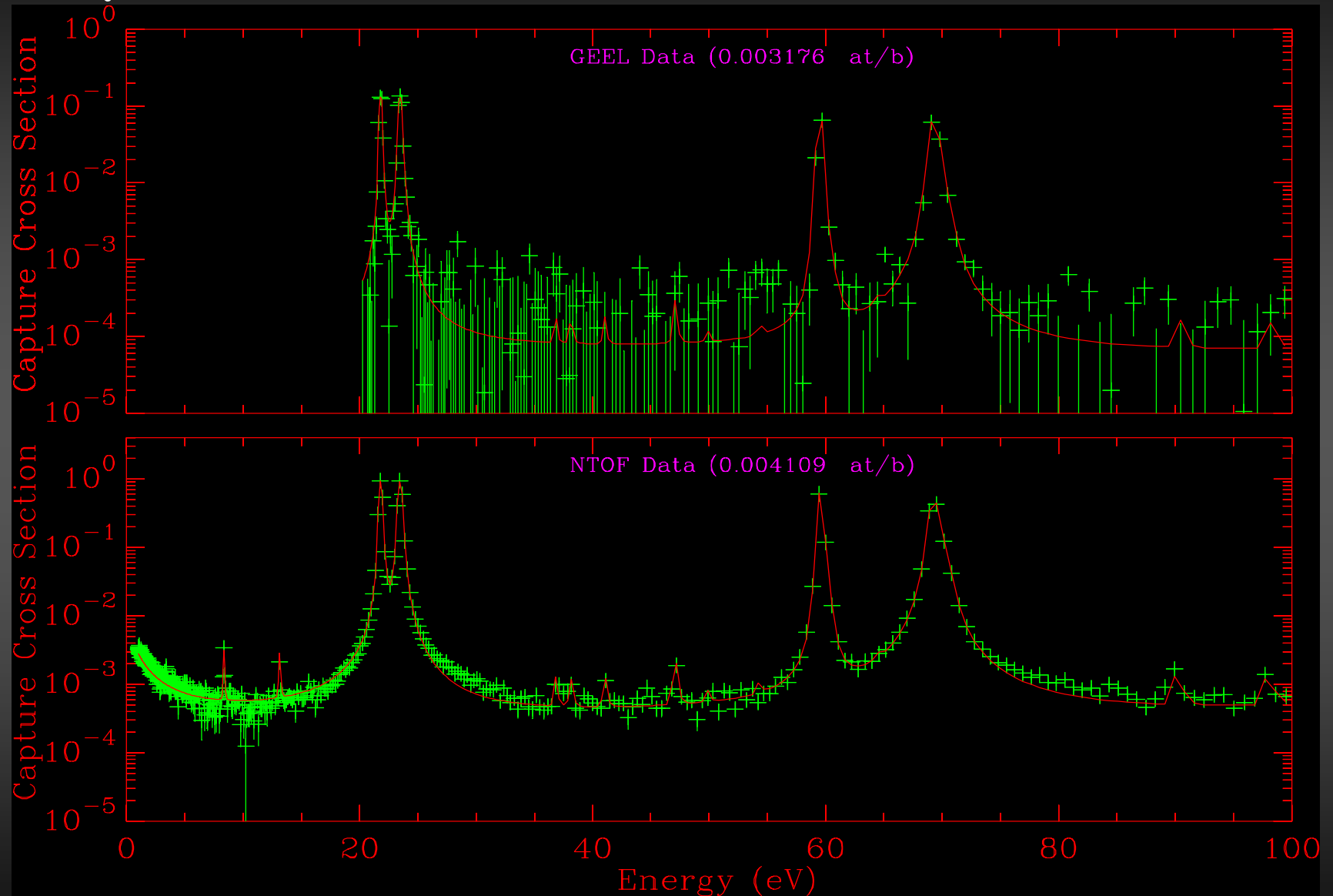


(Analysis by  
E.I. Esch and  
R. Reifarth)

Source: J Ullman, n\_BANT workshop, CERN, March 2005

# $^{232}\text{Th}(n,\gamma)$ : n\_TOF & GELINA

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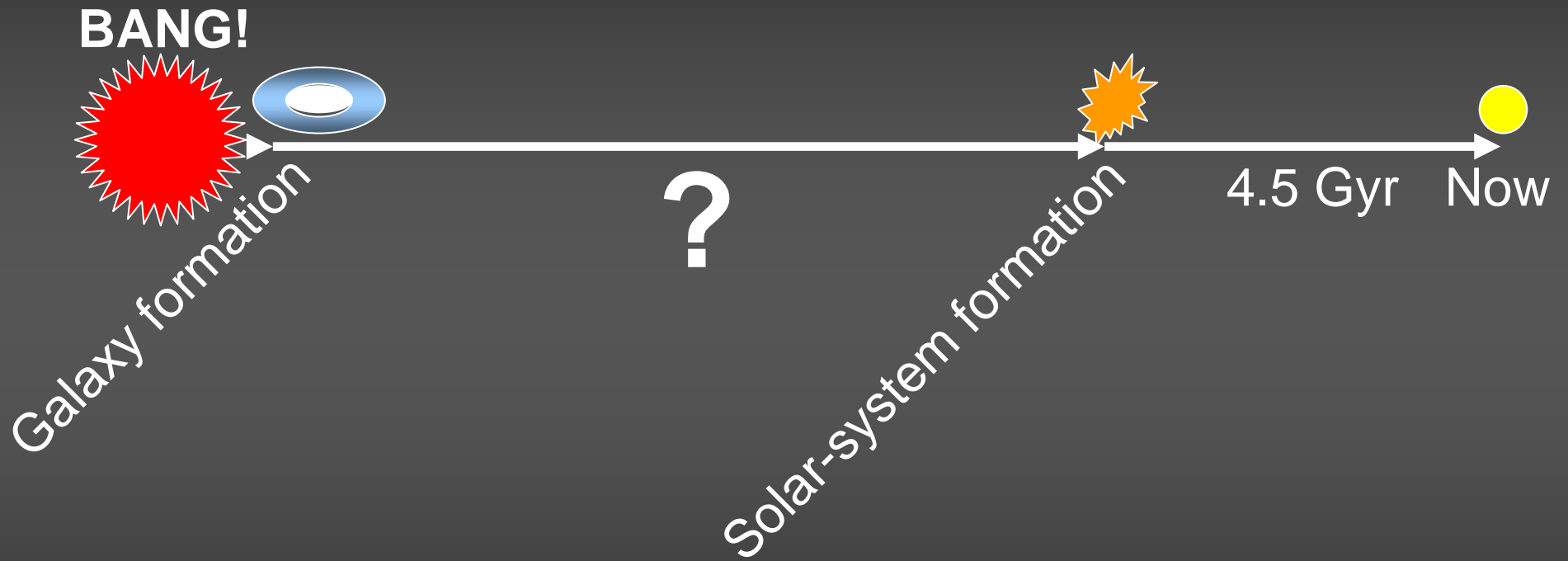


Source: L Leal, IAEA CRP meeting, December 2004

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# Time

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# Re/Os clock

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Os	Os 184 0.02	Os 185 94 d	Os 186 1.58	Os 187 1.6	Os 188 13.3	Os 189 16.1	Os 190 26.4	Os 191 15.4 d	Os 192 41.0
Re	Re 183 71 d	Re 184 38 d	Re 185 37.4	Re 186 90.64 h	Re 187 62.6 42.3x10 <sup>9</sup> a	Re 188 16.98 h	Re 189 24.3 h	Re 190 3.1 m	
W	W 182 26.3	W 183 14.3	W 184 30.67	W 185 75.1 d	W 186 28.6	W 187 23.8 h	W 188 69 d		

The  $\beta$ -decay half-life of  $^{187}\text{Re}$  is 42.3 Gy

Effect on the abundance of the decay daughter  $^{187}\text{Os}$

# Summary

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- Cosmological way (WMAP observations)

**$13.7 \pm 0.2$  Gyr**

- Astronomical way

**$14 \pm 2$  Gyr**

- Nuclear way: Re/Os clock

**$14.9 \pm 2$  Gyr(\*)**

Th/U clock

**$14.5 \pm {}^{2.8}_{2.2}$  Gyr**

Source: Dauphas, Nature 435 (2005) 1203

(\*) 0.4 Gyr uncertainty due to x-sections

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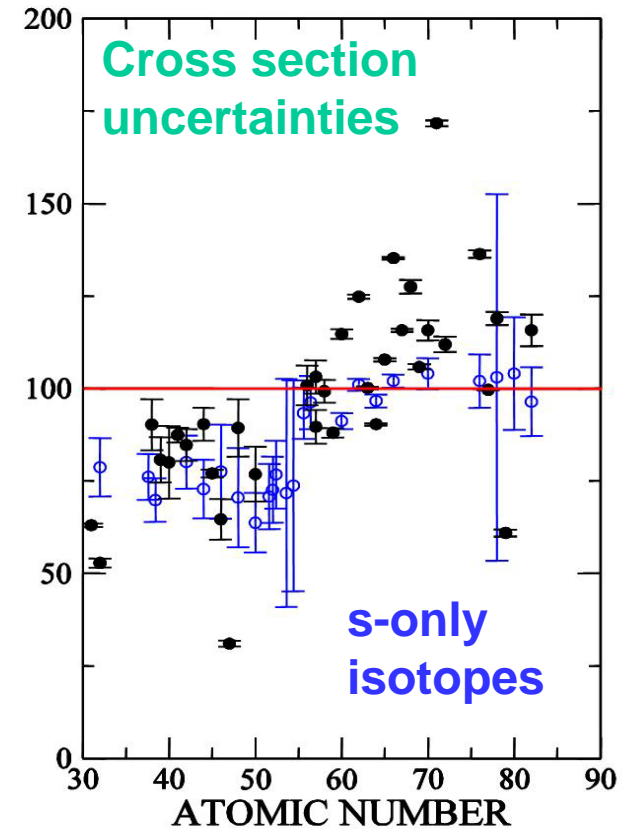
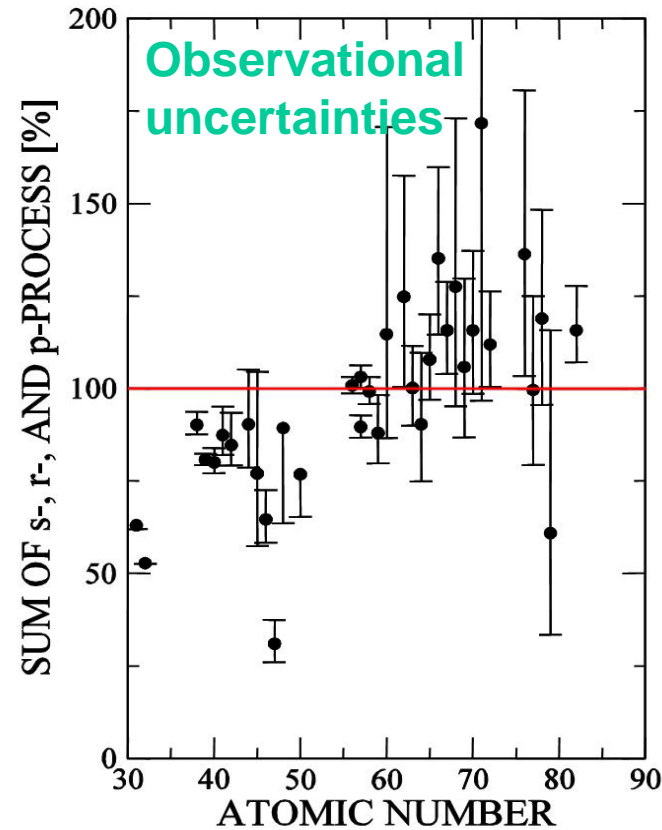
# In general:

Stellar model  
s-abundance  
calculation

Tavaglio et al. APJ521 (1999)  
Arlandini et al. APJ525 (1999)

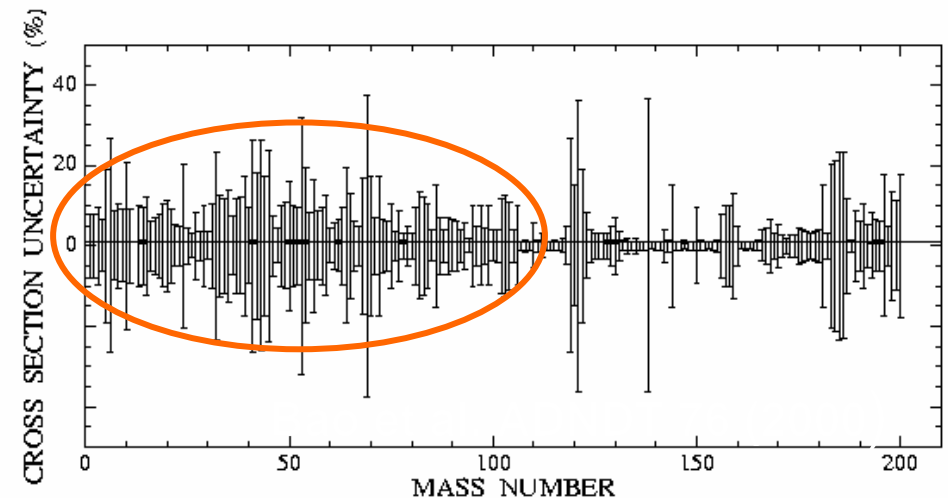
+

CS22892-052  
r-abundance



Need for more than one r-process  
or need for improved s- abundances  
(cross sections)?

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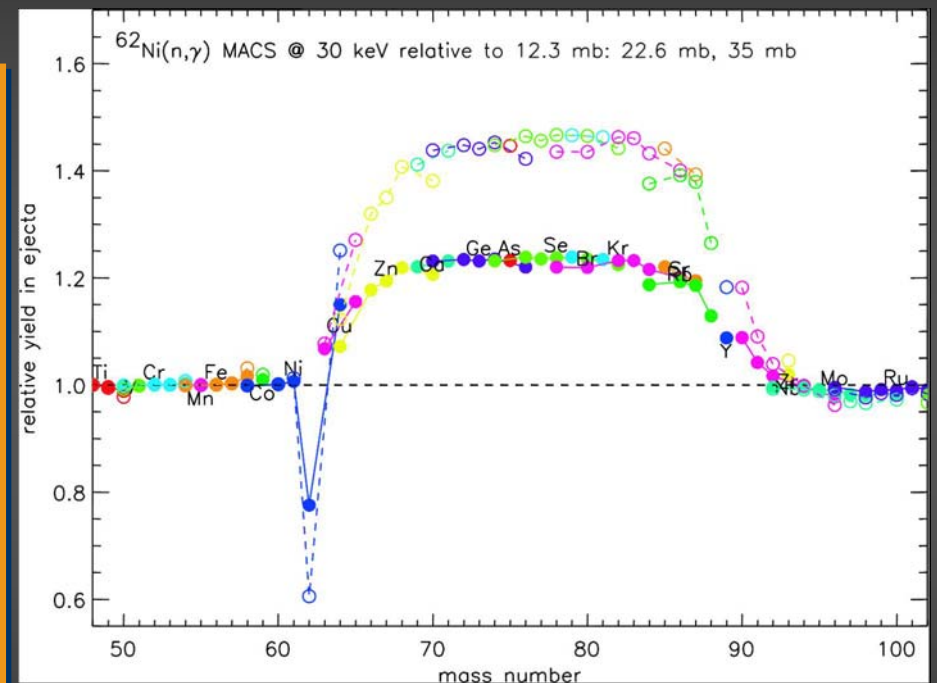


# Capture studies: Fe, Ni, Zn, and Se

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## Motivations:

- Study of the weak s-process component (nucleosynthesis up to  $A \sim 90$ )
- Contribution of massive stars (core He-burning phase) to the s-process nucleosynthesis.
- s-process efficiency due to bottleneck cross sections (Example:  $^{62}\text{Ni}$ )



## In addition:

Fe and Ni are the most important structural materials for nuclear technologies. Results of previous measurements at n\_TOF show that capture rates for light and intermediate-mass isotopes need to be revised.

# Capture studies: Fe, Ni, Zn, and Se

34	Kr 73 26 s	Kr 74 11,5 m	Kr 75 4,5 m	Kr 76 14,6 h	Kr 77 1,24 h	Kr 78 0,35	Kr 79 34,9 h	Kr 80 2,25	Kr 81 2,3 10 <sup>2</sup> a	Kr 82 11,6	Kr 83 11,5 h	Kr 84 57,0	Kr 85 4,48 h 10,76 a	Kr 86 17,3
	Br 72 10,9 s	Br 73 3,3 m	Br 74 46 m 25,4 m	Br 75 1,6 h	Br 76 1,32 s 16,0 h	Br 77 4,3 m 57,0 h	Br 78 6,46 m	Br 79 4,9 s 50,69	Br 80 4,42 h 17,6 m	Br 81 49,31	Br 82 5,1 m 35,34 h	Br 83 2,40 h	Br 84 6,0 m 31,6 m	Br 85 2,87 m
32	Se 71 4,74 m	Se 72 8,5 d	Se 73 39 m 7,1 h	Se 74 0,89	Se 75 119,64 d	Se 76 9,36	Se 77 17,5 s 7,63	Se 78 23,78	Se 79 3,9 m 6,5 10 <sup>1</sup> a	Se 80 49,61	Se 81 57,3 m 18 m	Se 82 8,73 1,08 · 10 <sup>20</sup> a	Se 83 69 s 22,4 m	Se 84 3,1 m
	As 70 53 m	As 71 65,28 h	As 72 26,0 h	As 73 80,3 d	As 74 17,77 d	As 75 100	As 76 26,4 h	As 77 38,9 h	As 78 1,5 h	As 79 8,2 m	As 80 15,2 s	As 81 34 s	As 82 14,0 s 19,1 s	As 83 13,3 s
30	Ge 69 39,0 h	Ge 70 21,23	Ge 71 11,43 d	Ge 72 27,66	Ge 73 7,73	Ge 74 35,94	Ge 75 47 s 83 m	Ge 76 7,44	Ge 77 53 s 11,3 h	Ge 78 88 m	Ge 79 39 s 19 s	Ge 80 29,5 s	Ge 81 7,6 s 7,6 s	Ge 82 4,60 s
	38		40		42		44		46		48		50	

## The <sup>79</sup>Se case

- s-process branching: neutron density & temperature conditions for the weak component.
- $t_{1/2} < 6.5 \times 10^4$  yr

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