

# Standard Model and Beyond

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29/1/2018

- **Standard Model is defined by**
  - **4-dimension QFT (Invariant under Poincare group).**
  - **Symmetry: Local  $SU(3)_C \times SU(2)_L \times U(1)_Y$ .**
  - **Particle content (Point particles):**
  - **3 fermion (quark and Lepton) Generations.**
  - **No Right-handed neutrinos ? Massless Neutrinos.**
- **Symmetry breaking: one Higgs doublet.**
- **No candidate for Dark Matter.**
- **SM does not include gravity.**

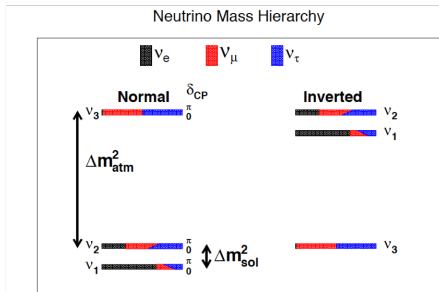
# Evidence for Physics Beyond the SM

## 1- Neutrino Mass:

- In the SM, quarks and electrons acquire masses through Yukawa couplings :  $\mathcal{L}_{Yuk} \sim \bar{Q}_L \phi U_R$ .
- Neutrinos remain massless because there are no RH  $\nu$  in the SM.
- However, it has proven experimentally (from neutrino oscillations) that  $m_\nu \neq 0$ .

parameter	best fit value $\pm 1\sigma$	$3\sigma$ range
$\sin^2 \theta_{12}$	$0.304^{+0.012}_{-0.012}$	(0.270, 0.344)
$\theta_{12}$ (degrees)	$33.48^{+0.77}_{-0.74}$	(31.30, 35.90)
$\sin^2 \theta_{23}$	$[0.451^{+0.001}_{-0.001}]$ or $0.577^{+0.027}_{-0.035}$	(0.385, 0.644)
$\theta_{23}$ (degrees)	$[42.2^{+0.1}_{-0.1}]$ or $49.4^{+1.6}_{-2.0}$	(38.4, 53.3)
$\sin^2 \theta_{13}$	$0.0219^{+0.0010}_{-0.0011}$	(0.0188, 0.0251)
$\theta_{13}$ (degrees)	$8.52^{+0.20}_{-0.21}$	(7.87, 9.11)
$\delta_{CP}$ (degrees)	$251^{+67}_{-59}$	(0, 360)
$\Delta m_{21}^2 \times 10^{-5} \text{ eV}^2$	$7.50^{+0.19}_{-0.17}$	(7.03, 8.09)
(normal) $\Delta m_{31}^2 \times 10^{-3} \text{ eV}^2$	$+2.458^{+0.046}_{-0.047}$	(+2.325, +2.599)
(inverted) $\Delta m_{32}^2 \times 10^{-3} \text{ eV}^2$	$-2.448^{+0.047}_{-0.047}$	(-2.590, -2.307)

- Needs a mechanism to give  $\nu$  masses...



# Evidence for Physics Beyond the SM

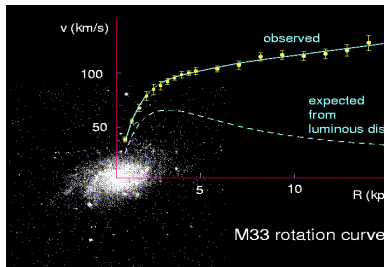
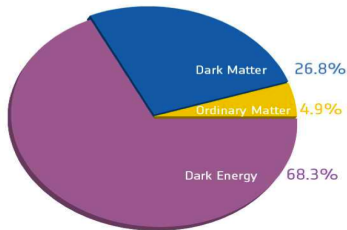
## 2- Dark Matter:

- Most astronomers, cosmologists and particle physicists are convinced that 90% of the mass of the Universe is due to some non-luminous matter, called 'Dark Matter/Energy'.

- The velocity of rotating objects

$$v(r) = \sqrt{\frac{G M(r)}{r}}$$

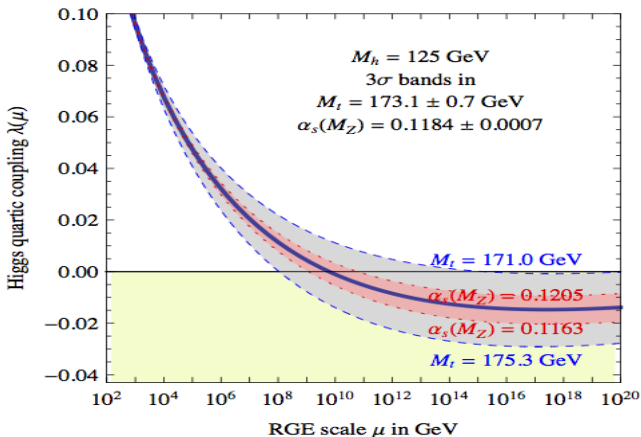
- The observation of 1000 spiral galaxies showed that away from the centre of galaxies the rotation velocities do not drop off with distance.
- The explanation for these is to assume that disk galaxies are immersed in extended DM halos.
- Dark Matter must be non-baryonic. No such candidate in the Standard Model



# Evidence for Physics Beyond the SM

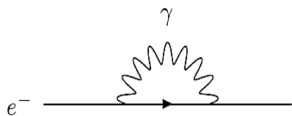
## 3- Higgs Vacuum Stability:

- Quadratic coupling evolves to zero or negative values. Recall that in SM  $M_H = \sqrt{\lambda v}$



# Evidence for Physics Beyond the SM

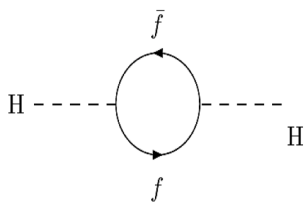
## 4- Higgs Mass Hierarchy:



$$\delta m_e = 2 \frac{\alpha_{em}}{\pi} m_e \log \frac{\Lambda}{m_e} + \dots$$

$m_e \rightarrow 0$ , Chiral Symmetry

$$\delta m_e = 0.24 m_e$$



$$\delta m_H^2(f) = -2N_f \frac{|\lambda_f|^2}{16\pi^2} [\Lambda^2 - 2m_f^2 \ln \frac{\Lambda}{m_f} + \dots]$$

$m_H^2$  no symmetry, Quadratic div.!!

$$\Lambda = 10^{19} \text{ GeV}$$

$$\delta m_H^2 \simeq 10^{30} \text{ GeV}$$

## Evidence for Physics Beyond the SM

In addition, there are a number of questions we hope will be answered:

- Electroweak symmetry breaking, which is not explained within the SM.
- Why is the symmetry group is  $SU(3) \times SU(2) \times U(1)$ ?
- Can forces be unified?
- Why are there three families of quarks and leptons?
- Why do the quarks and leptons have the masses they do?
- Can we have a quantum theory of gravity?
- Why is the cosmological constant much smaller than simple estimates would suggest?

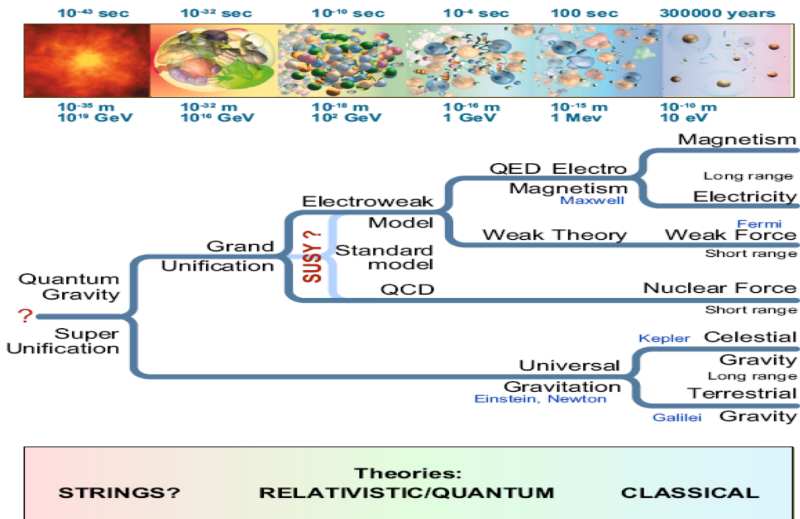
## Discovery Timeline

- 1897      electron
- 1937      muon
- 1959      neutrino (proposed in 1933)
- 1968      up, down, strange (proposed in 1964)
- 1974      charm (proposed in 1970)
- 1975      tau (proposed in 1971)
- 1977      bottom (proposed in 1964)
- 1996      top (proposed in 1964)
- 2012      Higgs (proposed in 1962)

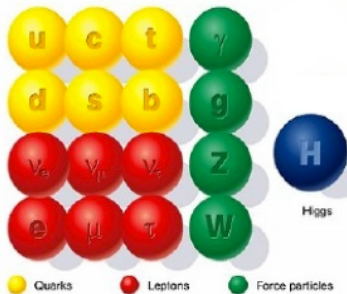


## Directions of BSM

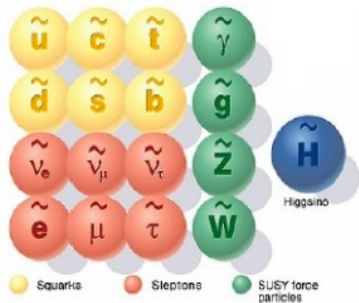
- Extension of gauge symmetry.
- Extension of Higgs Sector.
- Extension of Matter Content.
- Extension with Flavor Symmetry.
- Extension of Space-time dimensions (Extra-dimensions).
- Extension of Lorentz Symmetry (Supersymmetry).
- Incorporate Gravity (Supergravity).
- One dimension object (Superstring).



# SUPERSYMMETRY



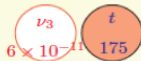
Standard particles



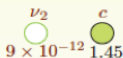
SUSY particles

## Flavour physics: who ordered that??

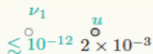
- 3 families with identical gauge quantum numbers.



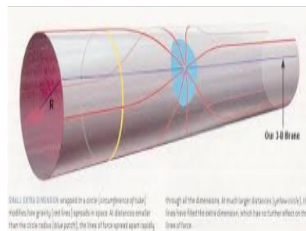
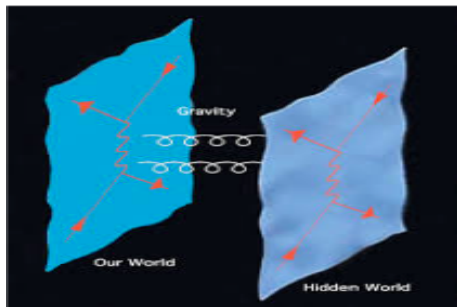
- Strong hierarchy between generations.



- Small quark, large lepton mixing angles.



# Extra Dimension and Brane World



# SM Extension with Right-handed Neutrino ( $\nu_R$ )

- Introducing the right-handed neutrinos  $\nu_R$  implies

$$\mathcal{L} = Y_\ell \bar{L} \phi e_R + Y_\nu \bar{L} \tilde{\phi} \nu_R + M \bar{\nu}_R^c \nu_R$$

Nothing prevents adding then a mass term for right-handed neutrino.

- Type I Seesaw Mechanism

$$\mathcal{L} = \frac{1}{2} (\nu_L, \nu_R) \begin{pmatrix} 0 & m_D \\ m_D^T & M \end{pmatrix} \begin{pmatrix} \nu_L \\ \nu_R \end{pmatrix}$$

- For  $m_D \ll M$ , one finds

$$m_{\nu_\ell} \simeq -m_D^2/M, \quad M_R \simeq M$$

with mixing angle:  $\tan 2\theta = 2m_D/M \ll 1$ .

- $m_D = Y_\ell v$ , so to get  $m_{\nu_\ell} \sim O(1)eV$ , either:

$$Y_\ell \sim O(1) \quad \& \quad m_D \sim O(100)GeV \Rightarrow M \simeq 10^{13} GeV$$

OR

$$Y_\ell \sim O(10^{-6}) \quad \& \quad m_D \sim O(10^{-4})GeV \Rightarrow M \simeq 1TeV.$$

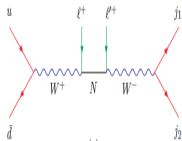
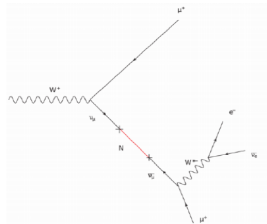
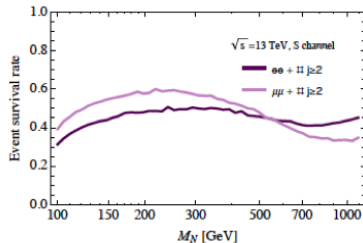


# Search for right-handed neutrinos

- The right-handed neutrino,  $N_R$  or  $\nu_R$ , mixes with the three active neutrinos.
- In terms of mass eigenstates, the gauge interaction Lagrangian is given by:

$$\mathcal{L} = -\frac{g}{\sqrt{2}}W_\mu^+ \left( \sum_{l=e}^{\tau} \sum_{m=1}^3 U_{lm}^* \bar{\nu}_m \gamma^\mu P_L l \right) - \frac{g}{\sqrt{2}}W_\mu^+ \left( \sum_{l=e}^{\tau} V_{lA}^* N_R^c \gamma^\mu P_L l \right) - \frac{g}{2 \cos \theta_W} Z_\mu \left( \sum_{l=e}^{\tau} V_{lA}^* N_R^c \gamma^\mu P_L \nu_l \right) + \text{H.c.}$$

- A signature of  $\nu_R$  signal consists of two same-sign muons, one electron and missing transverse energy .
- Or the following signal:  $p p \rightarrow \ell^\pm \ell'^\pm + j j$



- Cross section observable only if  $\nu_\ell - \nu_R$  mixing is  $> 10^{-2}$ . However even for  $M_{\nu_R} \sim 100$  GeV,  $\theta \sim 10^{-6}$  (Not observable at LHC).

## TeV Seesaw with $B - L$ forces ( $Z'$ )

- The  $B - L$  extension of the SM is based on the gauge group

$$SU(3)_C \times SU(2)_L \times U(1)_Y \times U(1)_{B-L}$$

In this model:

- Three right-handed neutrinos,  $N_R^i$ ,  $i = 1, 2, 3$ ; with  $B - L$  charge =  $-1$ .
- An extra gauge boson corresponding to  $B - L$  gauge symmetry,  $Z'$ .
- An extra SM singlet scalar,  $\chi$  with  $B - L$  charge =  $+2$ , are introduced.

S.K(2007)

Particle	$\ell_L$	$e_R$	$N_R$	$\phi$	$\chi$
$Y_{B-L}$	$-1$	$-1$	$-1$	$0$	$+2$



# Neutral Gauge Boson $Z'$

- The  $U(1)_Y$  and  $U(1)_{B-L}$  gauge kinetic mixing can be absorbed in the covariant derivative redefinition. In this basis, one finds

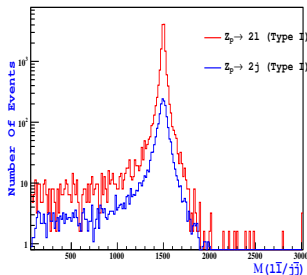
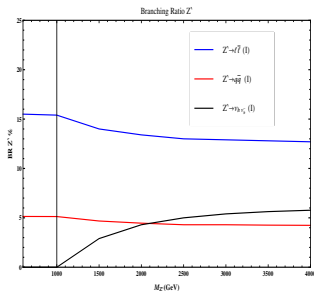
$$M_Z^2 = \frac{1}{4}(g_1^2 + g_2^2)v^2, \quad M_{Z'}^2 = g_{BL}^2 v'^2 + \frac{1}{4}\tilde{g}^2 v^2$$

- The mixing angle between  $Z$  and  $Z'$ :

$$\tan 2\theta' = \frac{2\tilde{g}\sqrt{g_1^2 + g_2^2}}{\tilde{g}^2 + 16\left(\frac{v'}{v}\right)^2 g_{BL}^2 - g_2^2 - g_1^2}$$

- The decay widths of  $Z' \rightarrow f\bar{f}$  are given by

$$\begin{aligned} \Gamma(Z' \rightarrow l^+l^-) &\approx \frac{(g'' Y_{B-L}^l)^2}{24\pi} m_{Z'} \\ \Gamma(Z' \rightarrow q\bar{q}) &\approx \frac{(g'' Y_{B-L}^q)^2}{8\pi} m_{Z'} \left(1 + \frac{\alpha_s}{\pi}\right), \\ \Gamma(Z' \rightarrow t\bar{t}) &\approx \frac{(g'' Y_{B-L}^t)^2}{8\pi} m_{Z'} \left(1 - \frac{m_t^2}{m_{Z'}^2}\right) \\ &\times \left(1 - \frac{4m_t^2}{m_{Z'}^2}\right)^{1/2} \left(1 + \frac{\alpha_s}{\pi}\right) \end{aligned}$$

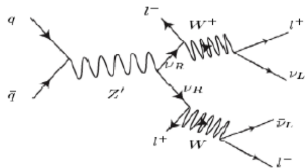


# Signature of $\nu_R$ at the LHC

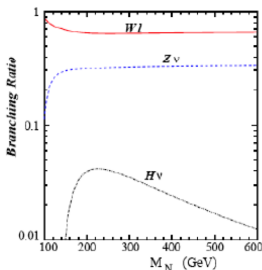
- In BLSM, seesaw effect observable at LHC even with tiny  $\nu_\ell - \nu_R$  mixings.

S.K., K. Huitu, et al (2008)

$$\mathcal{L}_{int.} \sim -g'' C_\mu [(\overline{\nu_R})_i \gamma^\mu (\nu_R)_i + b_{ij} \overline{(\nu_L)^c}_i \gamma^\mu (\nu_R)_j + h.c.] + \frac{g_2}{\sqrt{2}} [W_\mu^- l_i^+ \gamma^\mu U_{ij} (\nu_L)_j + b_{ij} W_\mu^- l_i^+ \gamma^\mu (\nu_R)_j^c + h.c.],$$



The decay modes, which go through the Higgs H or H' and Z boson, can be neglected compared to the main mode  $\nu_R \rightarrow Wl$ .



These decays are very clean with four hard leptons in the final states and large missing energy due to the associated neutrinos.

## SM Extension with Singlet Scalar

- The extension of the SM Higgs sector with singlet scalar  $\chi$  is natural in several models beyond the SM, e.g.,  $\text{SM} \times U(1)'$ .

- In this class of model, the Higgs potential is given by

$$V(\phi, \chi) = m_1 \phi^+ \phi + m_2 \chi^+ \chi + \lambda_1 (\phi^+ \phi)^2 + \lambda_2 (\chi^+ \chi)^2 + \lambda_3 (\phi^+ \phi)(\chi^+ \chi)$$

- For  $V(\phi, \chi)$  bounded from below, it is required:

$$\lambda_3 > -2\sqrt{\lambda_1 \lambda_2}, \quad \lambda_1, \lambda_2 > 0$$

For non-zero local minimum, we require

$$\lambda_3^2 < 4\lambda_1 \lambda_2$$

## Higgs Vacuum Stability in $U(1)'$ Extension of the SM

- In  $SM \times U(1)'$ , like BLSM, the physical Higgs mass eigenstates are given by

$$\begin{pmatrix} h \\ H \end{pmatrix} = \begin{pmatrix} \cos \theta & -\sin \theta \\ \sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} \phi^0 \\ \chi \end{pmatrix},$$

where the mixing angle  $\theta$  is given by

$$\tan 2\theta = \frac{\lambda_3 v v'}{\lambda_1 v^2 - \lambda_2 v'^2}.$$

- The masses of light and heavy Higgs particles are given by

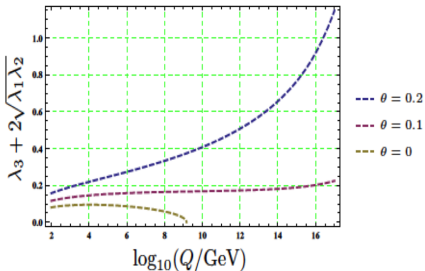
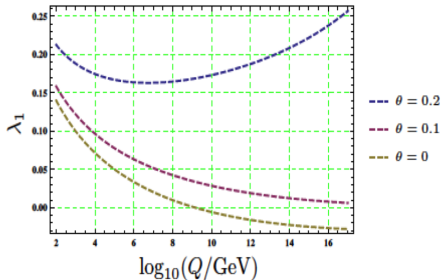
$$m_{h,H}^2 = \lambda_1 v^2 + \lambda_2 v'^2 \mp \sqrt{(\lambda_1 v^2 - \lambda_2 v'^2)^2 + (\lambda_3 v v')^2},$$

The RGEs of the scalar couplings:  $\lambda_1$ ,  $\lambda_2$  and  $\lambda_3$  in the context of  $B-L$  extension of the SM, are given by

$$\begin{aligned} \frac{d\lambda_1}{dt} = & \frac{1}{16\pi^2} \left( 24\lambda_1^2 + \lambda_3^2 + 12\lambda_1 Y_t^2 - 6Y_t^4 \right) + \frac{9}{8}g_2^4 + \frac{3}{8}g_1^4 + \frac{3}{4}g_2^2 g_1^2 + \frac{3}{4}g_2^2 \tilde{g}^2 \\ & + \frac{3}{4}g_1^2 \tilde{g}^2 + \frac{3}{8}\tilde{g}^4 - 9\lambda_1 g_2^2 - 3\lambda_1 g_1^2 - 3\lambda_1 \tilde{g}^2 \Big), \end{aligned}$$

$$\frac{d\lambda_2}{dt} = \frac{1}{8\pi^2} \left( 10\lambda_2^2 + \lambda_3^2 - \frac{1}{2} \text{Tr} [(Y_N)^4] + 48g_1'^4 + 4\lambda_2 \text{Tr} [(Y_N)^2] - 24\lambda_2 g_1'^2 \right),$$

$$\frac{d\lambda_3}{dt} = \frac{\lambda_3}{8\pi^2} \left( 6\lambda_1 + 4\lambda_2 + 2\lambda_3 + 3Y_t^2 - \frac{9}{4}g_2^2 - \frac{3}{4}g_1^2 - \frac{3}{4}\tilde{g}^2 + 2\text{Tr} [(Y_N)^2] - 12g_1'^2 + 6\frac{\tilde{g}^2 g_1'^2}{\lambda_3} \right),$$



- The RG running of the quartic coupling  $\lambda_1$  in the BLSM with type-I seesaw, for three values of the scalar mixing angle  $\theta$  for SM-like Higgs mass  $m_h = 125$  GeV.
- The evolution of the second stability condition,  $\lambda_3 + 2\sqrt{\lambda_1\lambda_2}$ , up to the GUT scale.
- At  $\theta = 0$ , the running of  $\lambda_1$  coincides with that of the SM. Hence one again finds that the Higgs potential becomes unstable at an energy scale  $\gtrsim 10^{9-10}$  GeV.
- With  $\theta \neq 0$ , initial values of  $\lambda_1$  at EW scale is larger than its value in the SM and also its scale dependence becomes rather different. One finds that with not very large mixing,  $\lambda_1$  and also  $\lambda_3 + 2\sqrt{\lambda_1\lambda_2}$  can remain positive up to the GUT scale. Hence the Higgs vacuum stability is accomplished.