

Overview of Simulations for Laser-Driven Positron Sources

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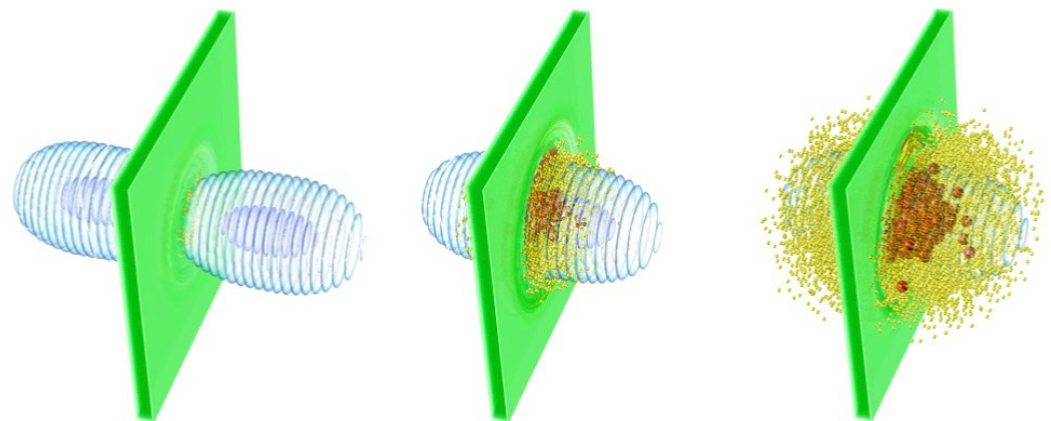
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Acknowledgments

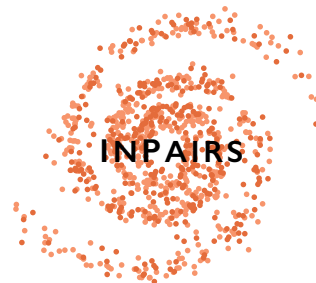
Work in collaboration with:

T. Grismayer, R. A. Fonseca, L. O. Silva (IST)

O. Klimo, S. Weber, G. Korn (ELI)

S. Meuren (Princeton)

Simulation results obtained at the IST cluster, SuperMUC (Garching), Fermi (CINECA), Salomon (Ostrava) and MareNostrum (Barcelona).



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Challenges

open questions and directions for the future

Regimes of interaction and relevant parameters

Preparing advanced computational tools

changes required to tackle the new regimes

Positrons from laser-electron beam scattering

Positrons from QED cascades

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What can we do with the next generation of lasers?

Opportunities to explore

- ▶ study the onset of QED
- ▶ study of radiation dominated regime below the QED limit
- ▶ new configurations for particle generation and acceleration
- ▶ design of efficient GeV photon sources
- ▶ study the development of self-generated e^+e^- plasmas, non-linear and collective effects

Open questions

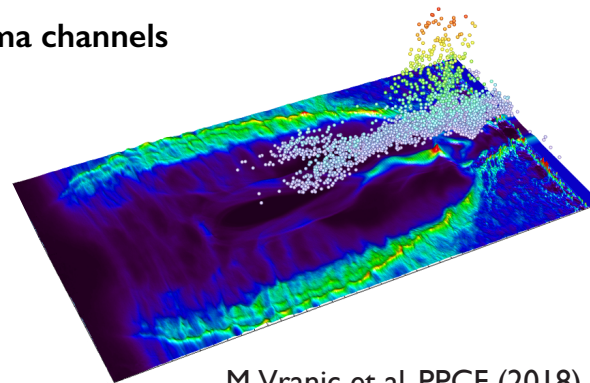
Can we produce a positron beam worthy of a collider?

Can we produce an electron-positron plasma with enough particles to be relevant for laboratory astrophysics?

New opportunities for particle acceleration

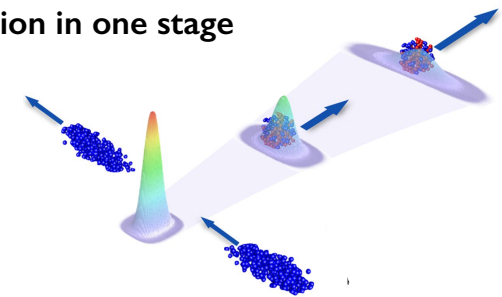
▶ Electron acceleration in plasma channels

B. Quiao et al, POP (2017)
A. Arefiev et al, POP (2016)
V. Khudik et al, POP (2016)
L. L. Ji et al, PRL (2014)
APS Robinson et al, PRL (2013)
Naseri et al, PRL (2012)
SPD Mangles et al, PRL (2005)
A. Pukhov et al, POP (1999)
M. Jirka et al, in prep.



M.Vranic et al, PPCF (2018)

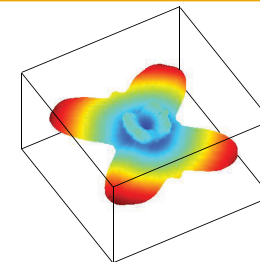
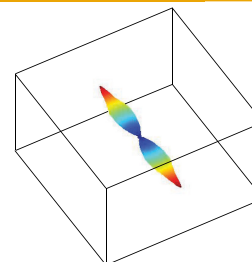
▶ Electron-positron production & acceleration in one stage



M.Vranic et al, SciRep (2018)

Design of tunable high-energy photon sources

N. Lemos, PPCF (2018); F. Albert, POP (2018)
W. Yan, Nat. Phot (2017); Gonoskov, PRX (2017)
A. Arefiev et al, PRL (2016); J. L. Martins et al, PPCF (2016)
K. Ta Phuoc, NatPhot (2012); Z. Gong, PRE (2017)
E. Esarey PRA (1992); S. Kneip, PRL (2009)



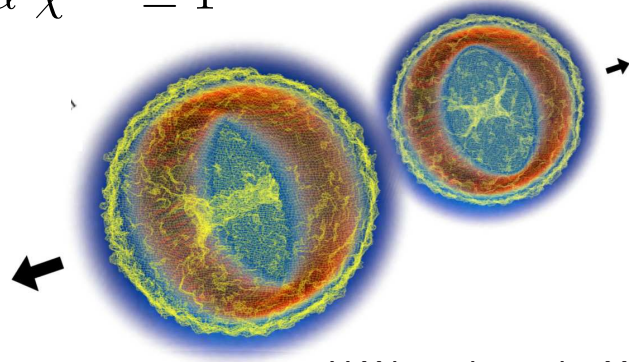
T. Grismayer et al, POP (2016)

Study of the classical and quantum radiation reaction

M. Tamburini, NIMR (2011); Neitz & DiPiazza, PRL (2013); Ilderton and Torgrimsson, PLB (2013); Zhidkov, PRSTAB (2014); M. Vranic, PRL (2014); T. Blackburn, PRL (2014); S. Yoffe, NJP (2015); M. Vranic, CPC (2016); M. Vranic, NJP (2016); C. Ridgers, JPP (2017); F. Neil, PRE (2017) and PPCF(2018); J. Cole PRX (2018); K. Poder PRX(2018);

Probing the onset of non-perturbative QED

$$\alpha \chi^{2/3} \simeq 1$$

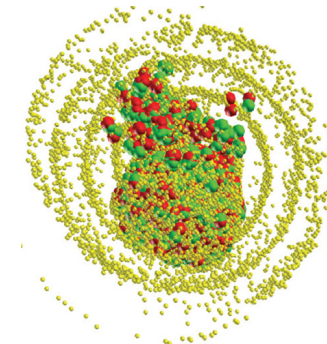


V. Yakimenko et al, arXiv:1807.09271
F. Del Gaudio et al, PRAB 22, 023402(2019)
C. Baumann et al., ArXiv:1811.03990

Evolution of self-generated e+e- plasmas

T. Bell and J. Kirk, PRL (2010)
A. Fedotov, PRL (2010)
S. Bulanov, PRL (2010)
E. Nerush, PRL (2011)
N. Elkina, PRSTAB (2011)
C. Ridgers, PRL (2012)
V. F. Bashmakov, POP (2015)
A. Gonoskov, PRE (2015)
M. Jirka, PRE (2016)
T. Grismayer, POP (2016)
X. Ribeyre et al, PRE(2016)
M. Tamburini, Sci Rep (2017)
M. Vranic, PPCF (2017)
X. Zhu, NComm (2017)

T. Grismayer, PRE (2017)
M. Lobet et al, PRAB (2017)
I. Kostyukov et al, PAST (2018)



T. Grismayer et al, POP (2016)

Challenges

open questions and directions for the future

Regimes of interaction and relevant parameters

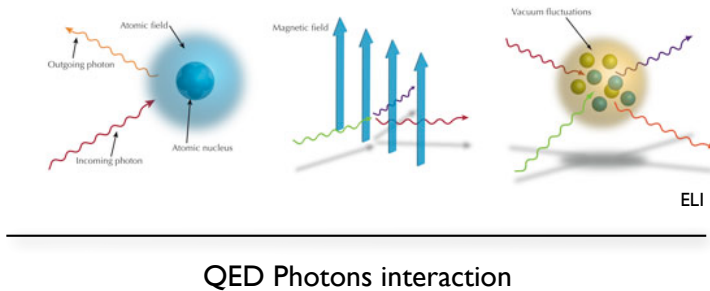
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Regimes of laser interaction with a plasma



Near-future facilities

- ▶ Pulse duration : 20-150 fs
- ▶ Intensity $\sim 10^{21} - 10^{25} \text{ W/cm}^2$
- ▶ Extreme acceleration regime

New facilities open possibilities to explore exotic physics.

Normalised vector potential a_0

- ▶ non relativistic
 $a_0 \ll 1 \quad I \sim 10^{18} \text{ W/cm}^2$
- ▶ weakly nonlinear, relativistic
 $a_0 \sim 1 \quad I \sim 10^{18} \text{ W/cm}^2$
- ▶ relativistic, nonlinear
 $a_0 \sim 10 \quad I \sim 10^{20} \text{ W/cm}^2$
- ▶ quantum
 $a_0 \sim 1000 \quad I \sim 10^{24} \text{ W/cm}^2$

$$a_0 = \frac{eA}{mc^2}$$

$$a_0 \approx 0.8 \times 10^{-9} \sqrt{I \left[\frac{\text{W}}{\text{cm}^2} \right]} \lambda [\mu\text{m}]$$

The onset of quantum effects is controlled by χ

Highest value of χ is obtained for relativistic particles counter-propagating with a laser

Unity is achieved when particle feels $E=Es$ in its own rest frame

$$E_s = \frac{m^2 c^3}{e \hbar} \quad \chi = \frac{\sqrt{(p_\mu F^{\mu\nu})^2}}{E_s mc} \quad \chi_\gamma = \frac{\sqrt{(k_\mu F^{\mu\nu})^2}}{E_s mc}$$

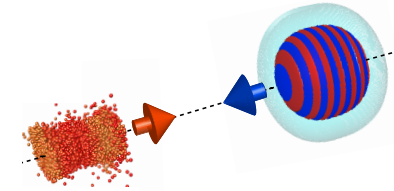
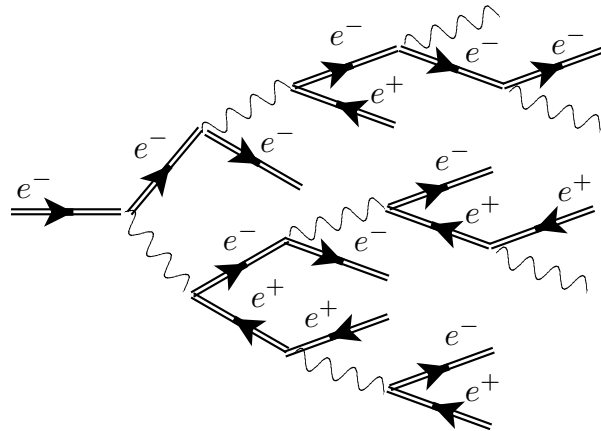
Classical: $\chi \ll 1$

$$\chi_e = \frac{1}{E_s} \sqrt{\left(\gamma \vec{E} + \frac{\vec{p}}{mc} \times \vec{B}\right)^2 - \left(\frac{\vec{p}}{mc} \cdot \vec{E}\right)^2}$$

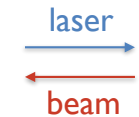
QED: $\chi \simeq 1$

QED cascade


$$\chi \gtrsim 1$$



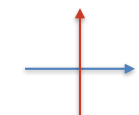
Counter-propagation


 $\chi \approx 2 \gamma_0 a_0 \times 2 \times 10^{-6}$

Co-propagation


 $\chi \approx \frac{a_0}{2\gamma_0} \times 2 \times 10^{-6}$

Interaction at 90 deg.


 $\chi \approx \gamma_0 a_0 \times 2 \times 10^{-6}$

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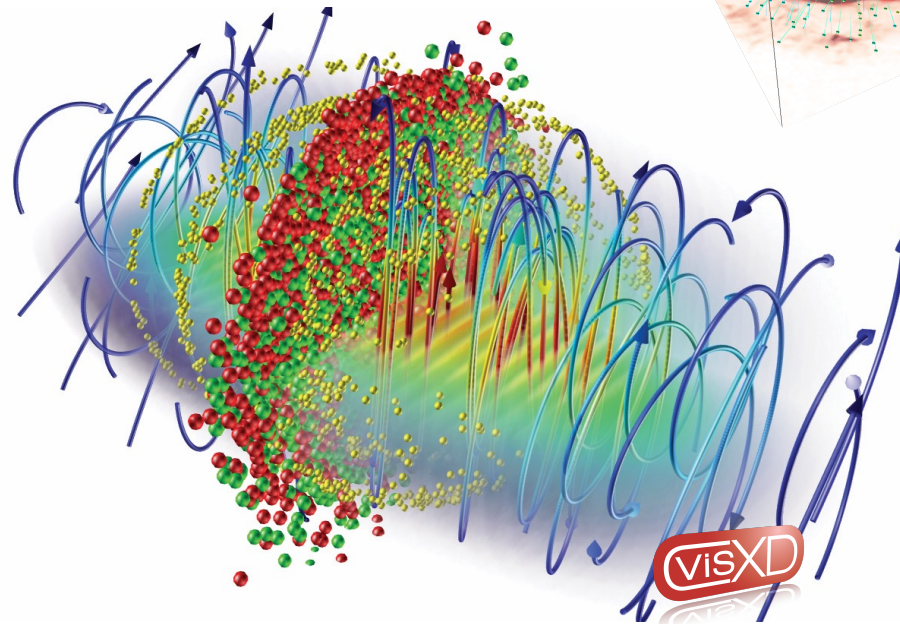
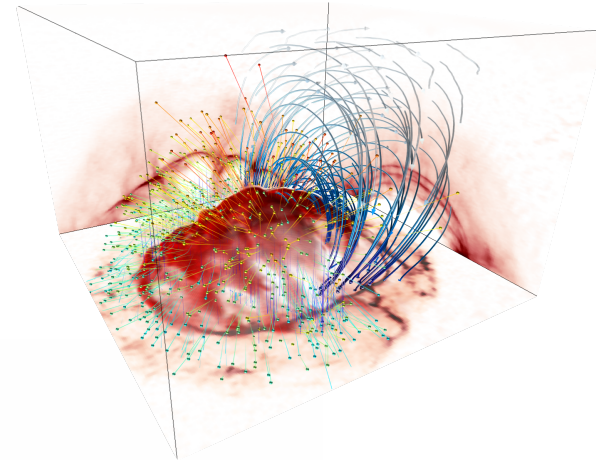
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osiris framework

- Massively Parallel, Fully Relativistic Particle-in-Cell (PIC) Code
- Visualization and Data Analysis Infrastructure
- Developed by the osiris.consortium
⇒ UCLA + IST



code features

- Scalability to ~ 1.6 M cores
- SIMD hardware optimized
- Parallel I/O
- **Dynamic Load Balancing**
- **QED module**
- **Particle merging**
- GPGPU & Xeon Phi support



UCLA

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<http://plasm asim.physics.ucla.edu/>

TOP500 List - June 2018; OSIRIS is supported by 9 out of top 10 machines

Rank	System	Cores	R_{\max} [TFlop/s]	R_{peak} [TFlop/s]	Power [kW]	OSIRIS support
1	Summit, United States	2282544	122300.0	187,659.3	8806	Full (CUDA)
2	Sunway TaihuLight, China	10649600	93014.6	125435.9	15371	No
3	Sierra, United States	1572480	71610.0	119193.6	-	Full (CUDA)
4	Tianhe-2 (MilkyWay-2), China	3120000	33862.7	54902.4	17808	Full (KNC)
5	ABCI, Japan	391680	19880.0	32576.6	1649	Full (CUDA)
6	Piz Daint, Switzerland	361760	19590.0	25326.3	2272	Full (CUDA)
7	Titan, United States	560640	17590.0	27112.5	8209	Full (CUDA)
8	Sequoia, United States	1572864	17173.2	20132.7	7890	Full (QPX)
9	Trinity, United States	979968	14137.3	43902.6	3844	Full (KNL)
10	Cori, United States	622336	14014.7	27880.7	3939	Full (KNL)

What developments are necessary?

Adding classical radiation reaction

- ▶ Modelling electron beam slowdown in scattering configurations
- ▶ Modeling other configurations where only a fraction of electrons may be subject to RR but where this can alter qualitative behaviour

M.Vranic, PRL (2014); M.Vranic, CPC (2016); M.Vranic, PPCF (2018)

Adding quantum processes

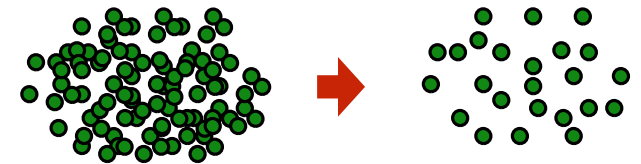
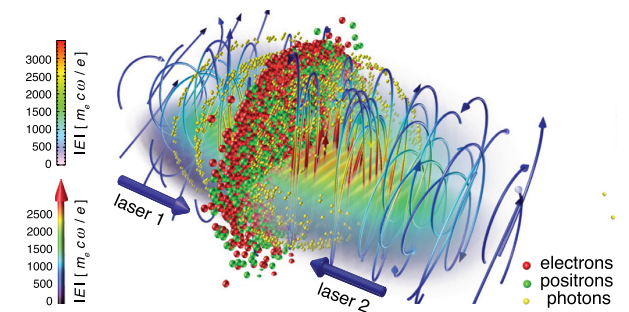
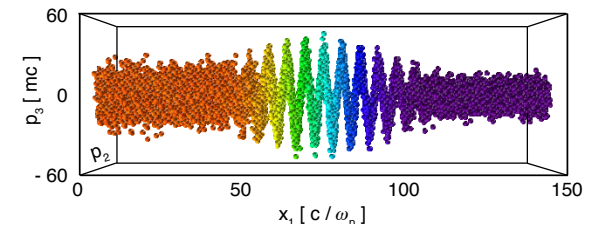
- ▶ Modelling the onset of QED, RR from quantum perspective
- ▶ Modelling e⁺e⁻ pair production, and gamma-gamma collisions
- ▶ QED cascades, nonlinear regimes where many particles are created and collective plasma dynamics can alter the background fields

M.Vranic, NJP (2016); T. Grismayer, POP (2016); T. Grismayer, PRE (2017); J. L. Martins, PPCF (2016); M.Vranic, PPCF (2017); M.Vranic, SciRep (2018); F. Del Gaudio PRAB (2019)

Adding performance improvements (particle merging, advanced load balancing schemes) + collision module for photons

- ▶ Essential for all the configurations with strong QED effects

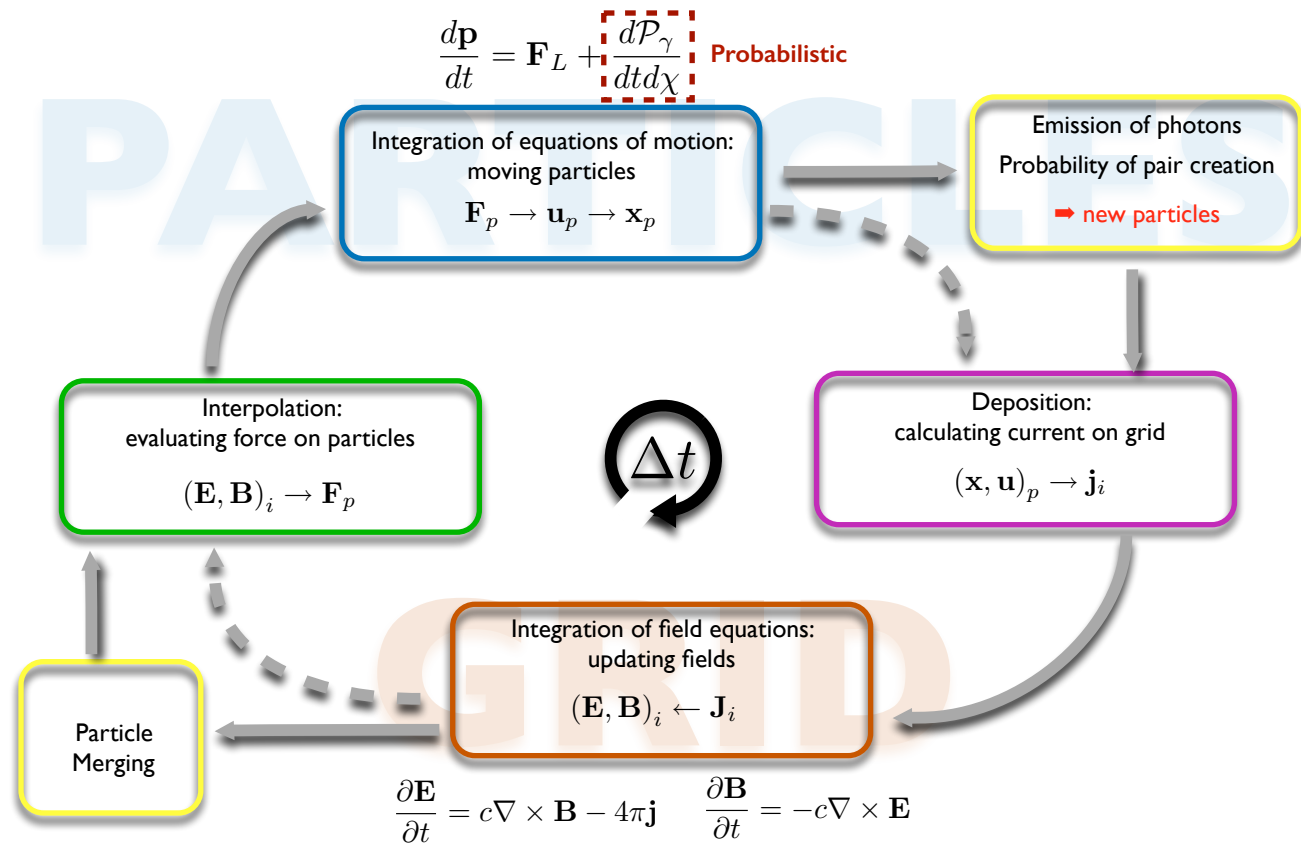
M.Vranic, CPC (2015); F. Del Gaudio, in preparation



QED loop in OSIRIS



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 tsung@physics.ucla.edu
<http://epp.tecnico.ulisboa.pt/>
<http://plasm asim.physics.ucla.edu/>



E.N Nerush et al. PRL (2011), C. P. Ridgers et al., PRL (2012), N.V. Elkina et al. PRSTAB (2011),
 A. Gonoskov et al., PRE (2015), T. Grismayer et al., POP (2016), T. Grismayer et al., PRE (2017)

Challenges

open questions and directions for the future

Regimes of interaction and relevant parameters

Preparing advanced computational tools

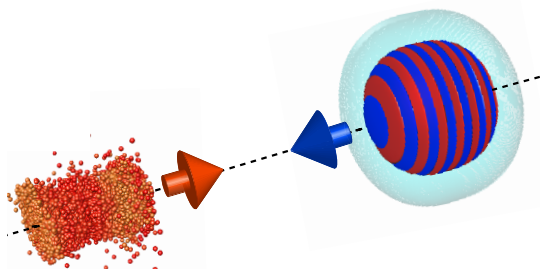
changes required to tackle the new regimes

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Positrons from QED cascades

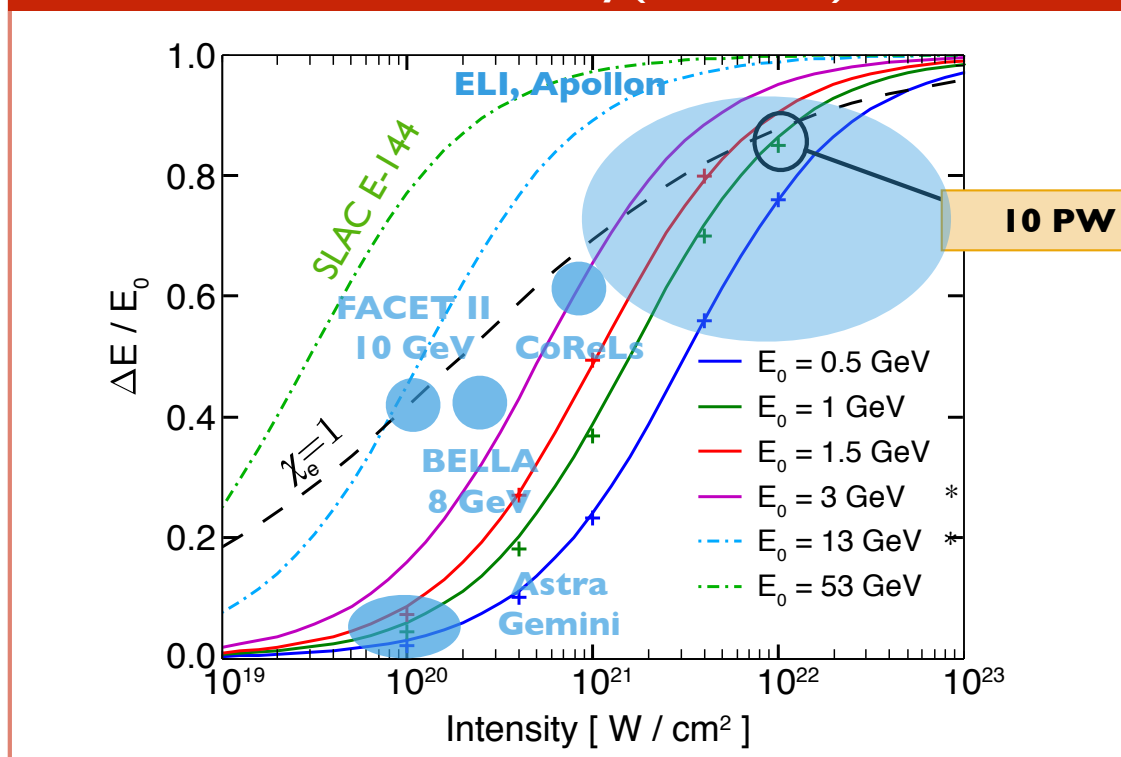
How much energy can be converted to photons in a laser - electron beam scattering?

For highly relativistic beams, most of the energy comes from the electrons (rather than the scattering laser)



$$\chi \sim \xi_e [\text{GeV}] \times \frac{a_0}{100}$$

Relative energy loss as a function of electron initial energy and the laser intensity (30 fs lasers)

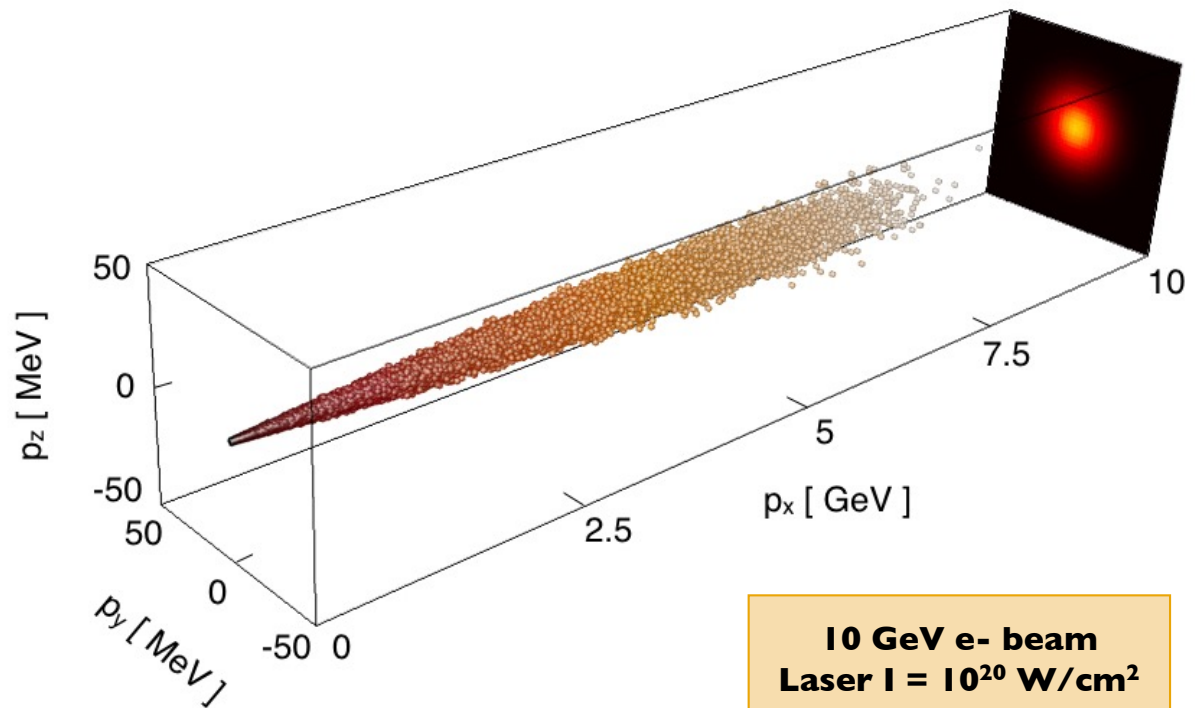


Parameters for SFQED experiment planned at FACET-II

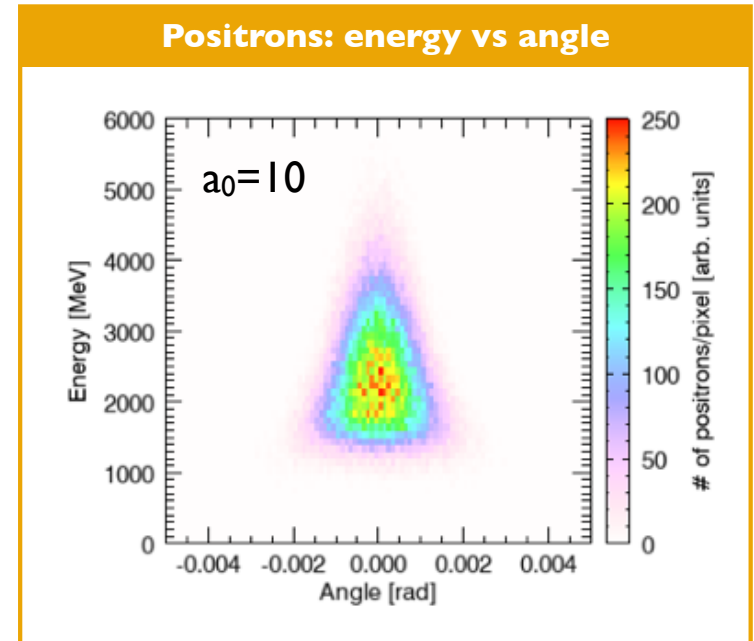
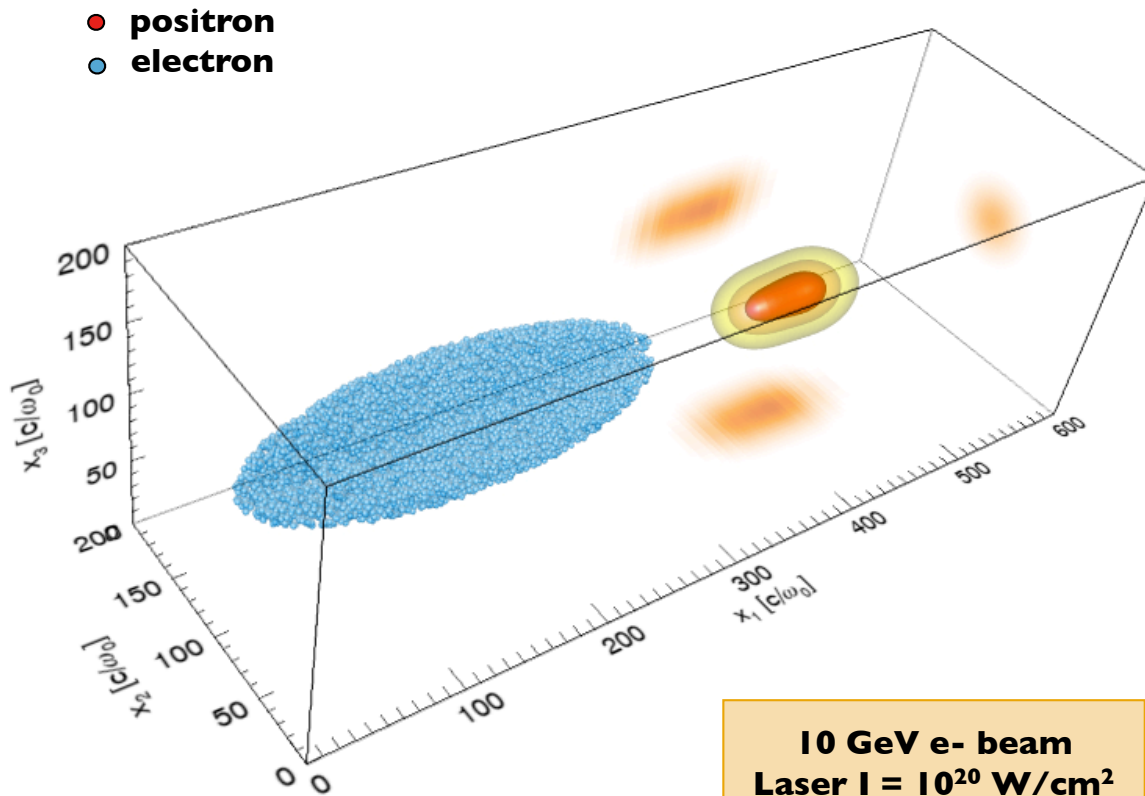
A large amount of beam energy can be converted to high-frequency photons (hard X-rays and Gamma-rays)

Photon source properties

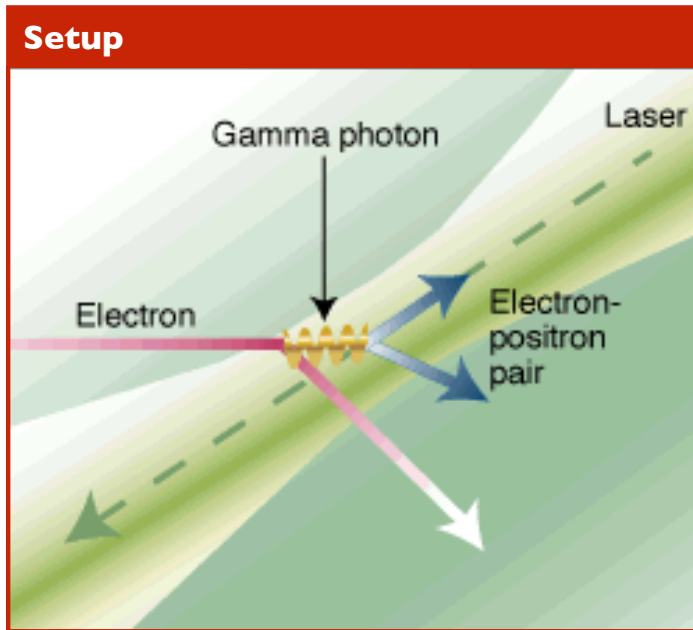
- ▶ divergence < 1 mrad
- ▶ tunable energy range (cutoff > 1 GeV)
- ▶ possible to attain very high energies (~ 10 GeV)



A fraction of radiated photons decays into electron-positron pairs



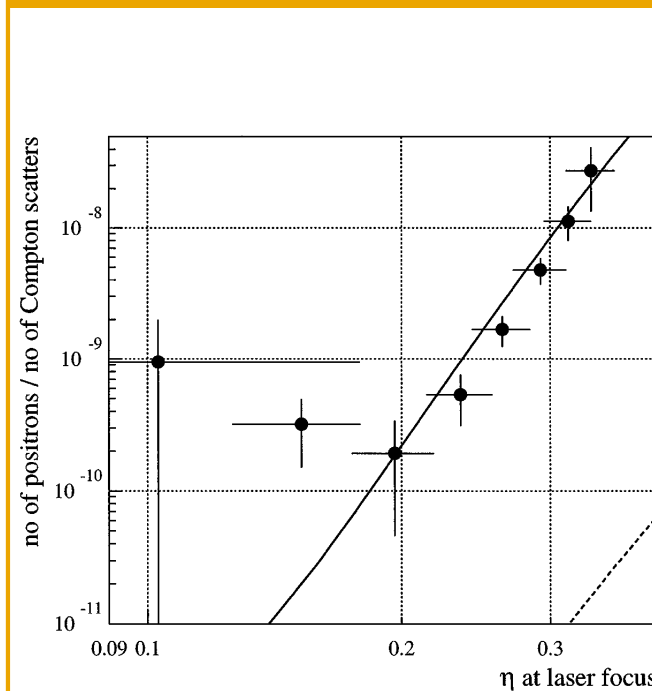
**1 nC electron beam gives
~ 0.2 pC of positrons**



Laser: 1.6 ps, 1 μ m, $I = 10^{18}$ W/cm²

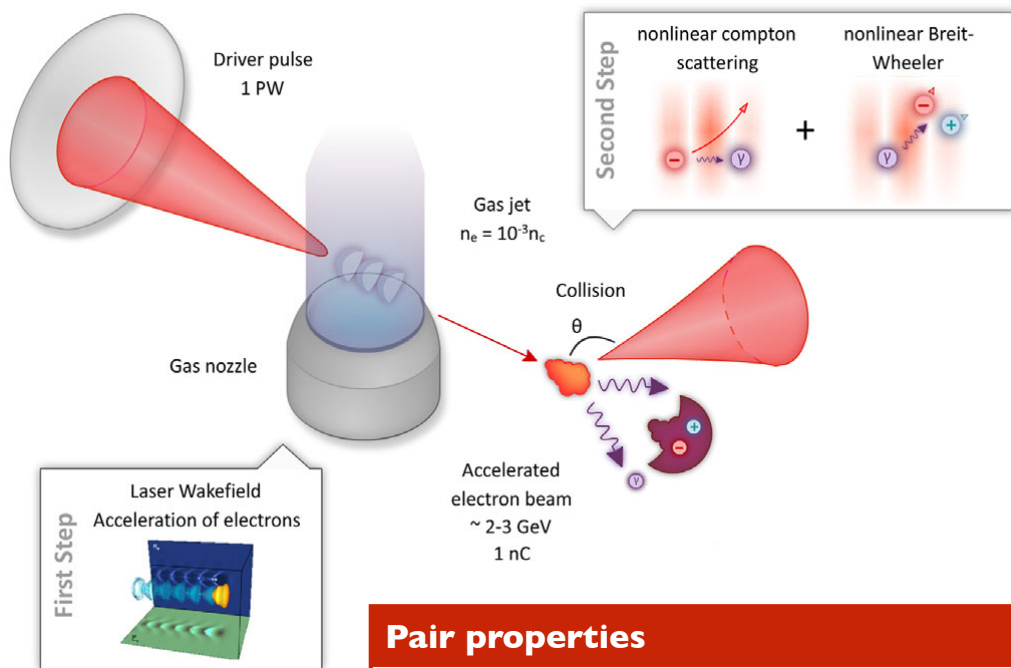
Electron energy: 46.6 GeV

~ 100 positrons



* D.L. Burke et. al., Phys. Rev. Lett.. 79, 1626-1629 (1997)

One pair per electron in a head-on collision

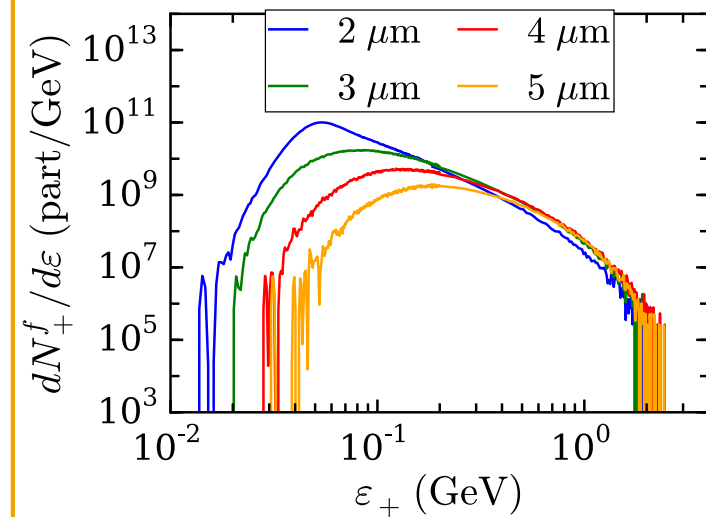


Pair properties

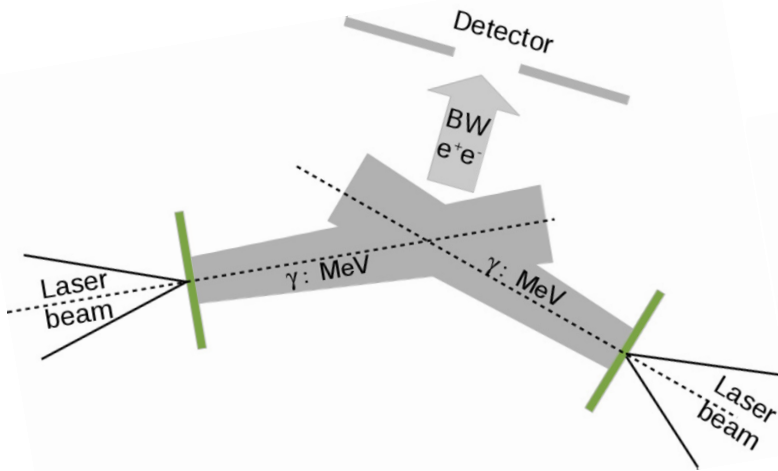
$\sim 10^{10}$ particles, max. energy 2 GeV

2-3 GeV e- beam, 1 nC
Laser of 15 fs
 $I = 10^{22}-10^{23}$ W/cm²

Positron spectrum



Configurations involving solid targets or cones

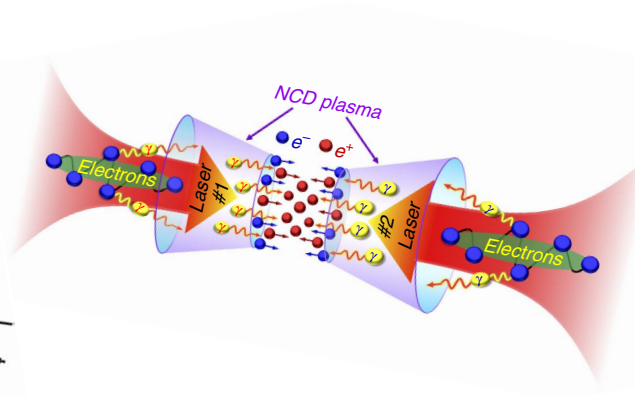


150 fs, $I = 10^{20} - 10^{21} \text{ W/cm}^2$

Breit-Wheeler pairs

$\sim 10^4$ positrons, $\sim \text{MeV}$ energy

X. Ribeyre et al., PRE 93, 013201 (2016)



3.3 fs, $I = 3 \times 10^{22} \text{ W/cm}^2$

Breit-Wheeler pairs

$\sim 10^{11}$ e+ below 2 GeV

Zhu et. al., Nat. Commun 7, 13686 (2016)

Bethe-Heitler pairs

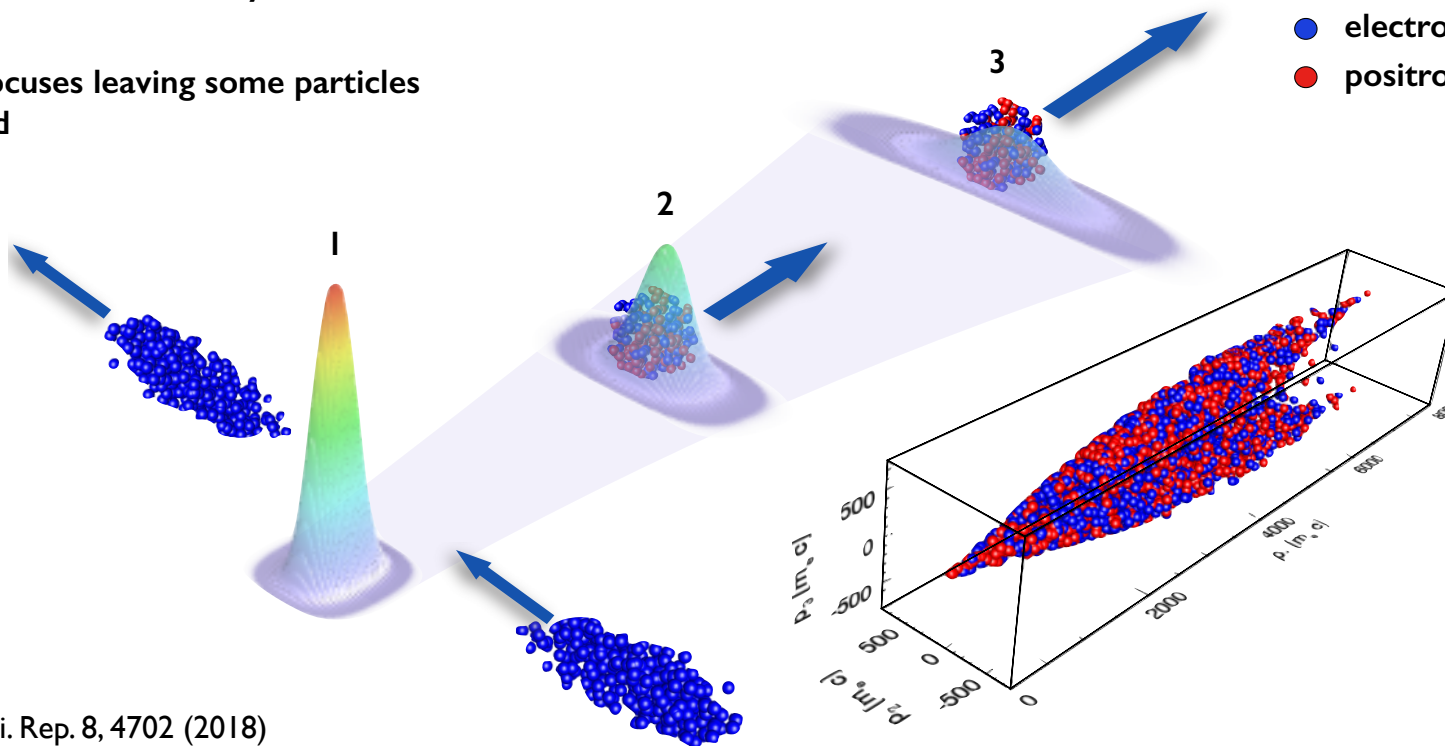
G. Sarri et al., Nat. Comm., 6, 6747 (2015)

H. Chen, F. Fiuza, A. Link et al., PRL, 114, 21 (2015)

Creating an e^+e^- beam from laser - e^- scattering at 90°

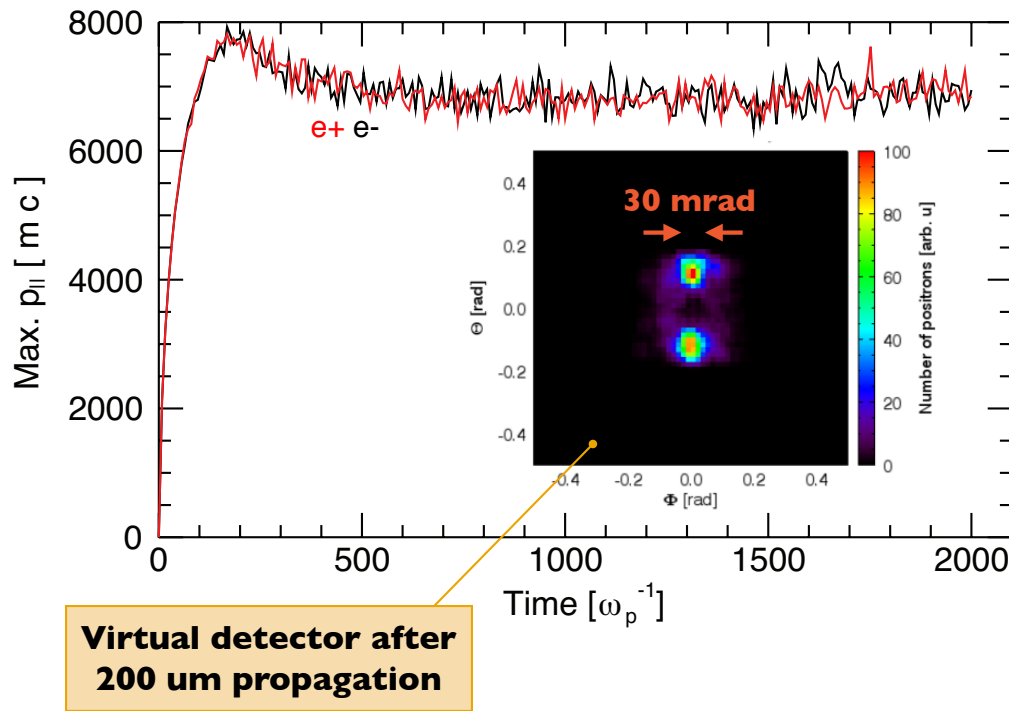
1. LWFA electrons collide with the laser; pairs are produced in the highest field region
2. e^+e^- beam is accelerated by the laser in vacuum
3. Laser defocuses leaving some particles accelerated

Pair creation and acceleration are decoupled!

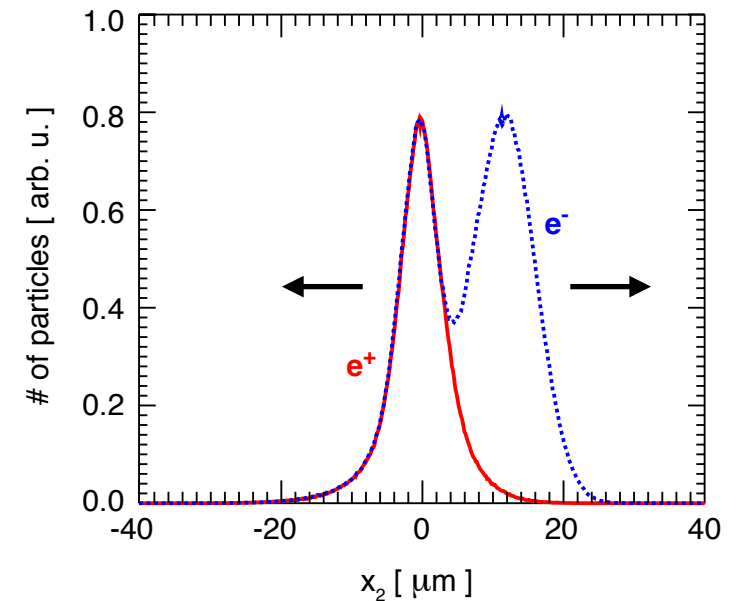


Final e^+e^- beam can be collected separately from initial electrons

~ 30 mrad beam divergence on detector for laser $a_0=600$, 50% e^+ and 50% e^-

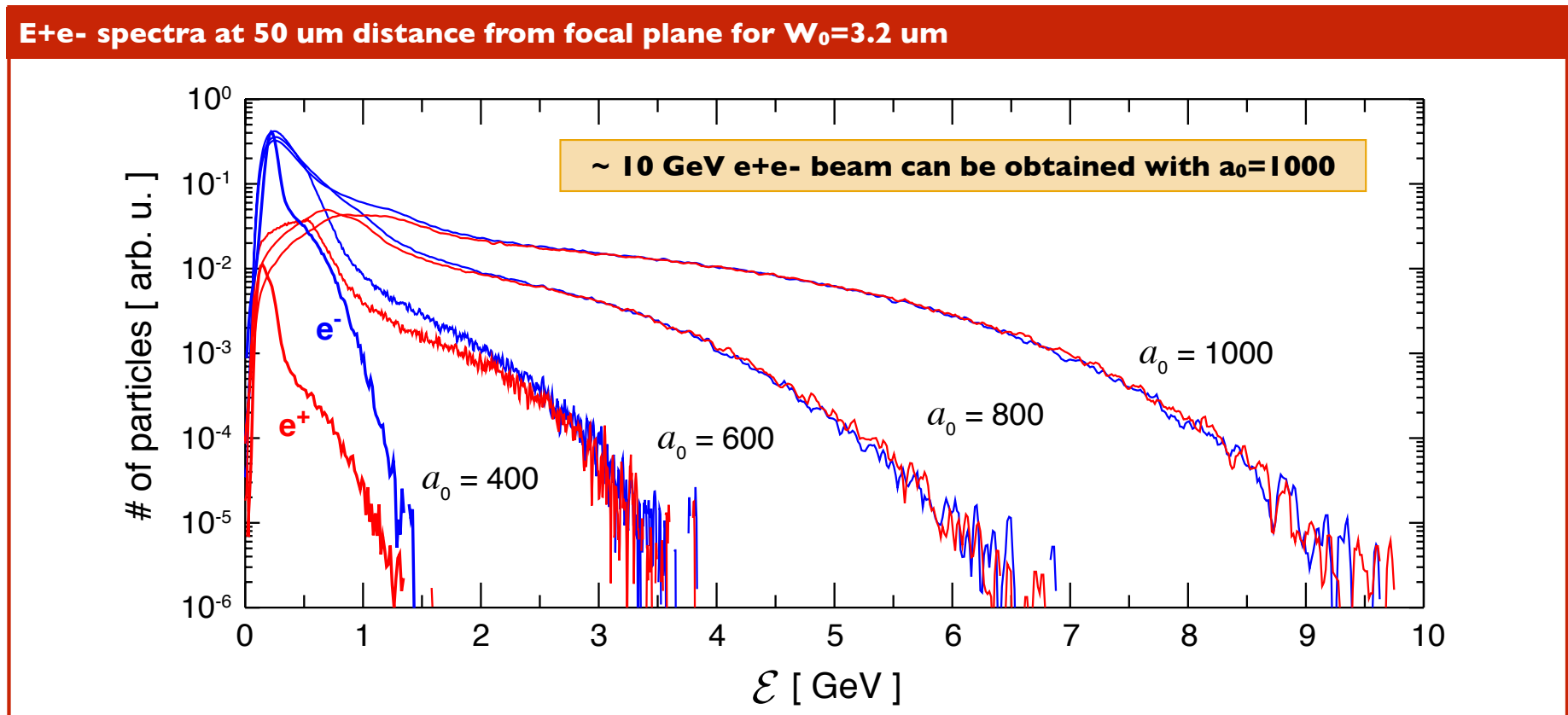


For 2 GeV e^- beam, you can get one pair per e^-



At energies > 2 GeV, e^- and e^+ have equal spectra

50 % of the positrons are above 2 GeV for $a_0=1000$



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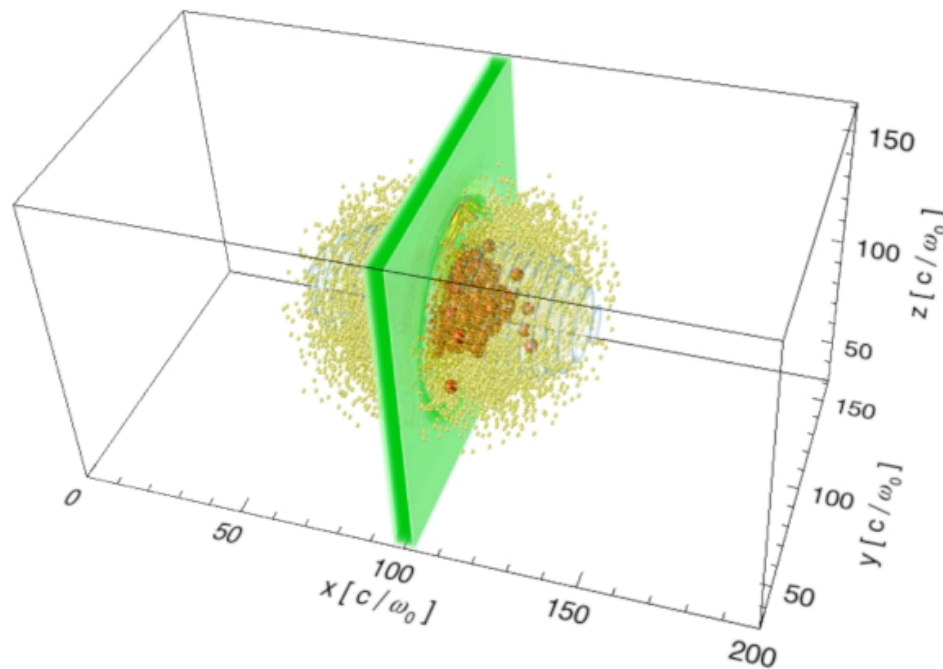
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Positrons from a hydrogen ice target



Target parameters

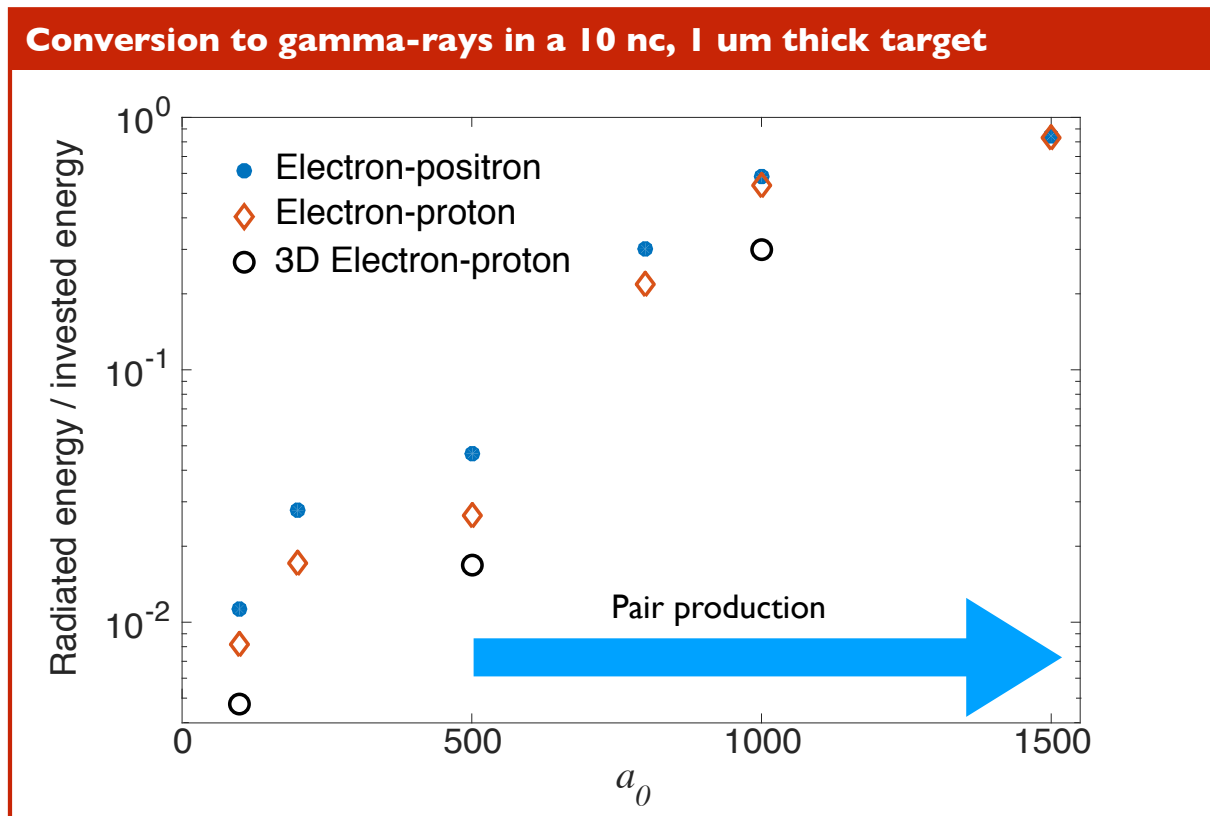
initial $n = 10$ nc
1 μ m thickness

Laser parameters

$I \sim 10^{24}$ W/cm²
30 fs, 1 μ m wavelength

With currently available targets, we could transfer more than 50% of energy to gamma-rays

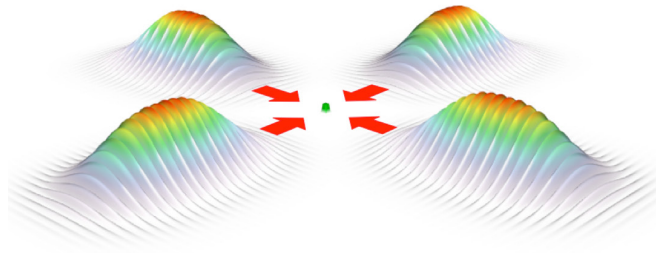
We get ~ a pair per interacting particle at $a_0=500$



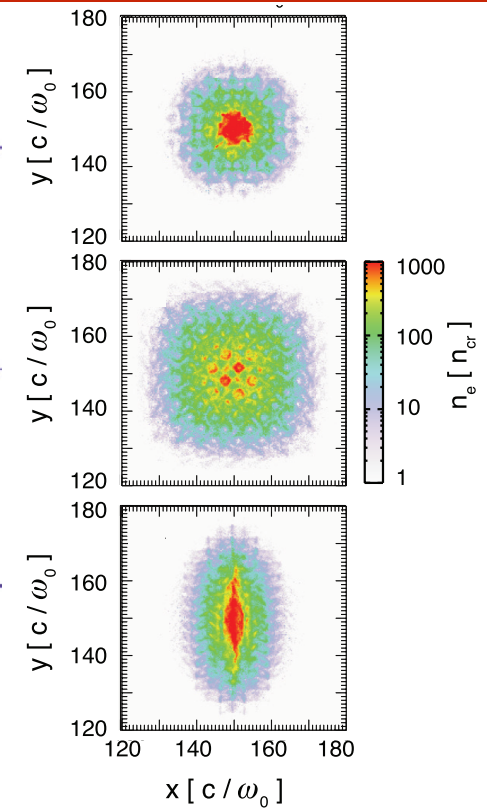
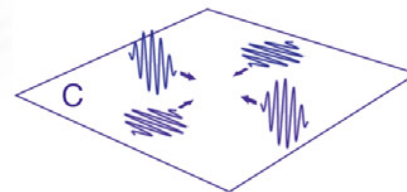
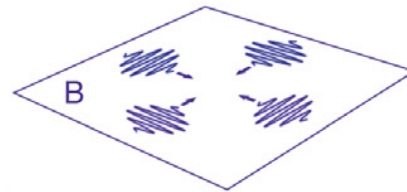
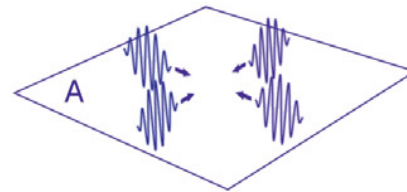
* M.Vranic et al., POP submitted (2019)

Optical traps can confine e⁺e⁻ pairs

Different polarisation combinations yield different microstructures



M. Vranic et al.,
PPCF 59, 014040 (2017)



Parameters needed to start creating \sim one e^+ per e^-

Facility	Electron energy	Laser intensity required for $\chi \simeq 1$
FACET	10 GeV	10^{20} W/cm ²
FACET*	20 GeV	2×10^{19} W/cm ²
AWAKE	50 GeV	4×10^{18} W/cm ²
Two-laser cascade		3×10^{23} W/cm ²

It is essential to include additional physics to PIC codes for modelling the next generation of laser experiments.

Performance developments are also necessary to tackle the new computational challenges associated with exponential growth of the number of particles, intrinsic load imbalance etc.

Classical vs. quantum radiation reaction can be studied in future experiments. Especially interesting is crossing the quantum threshold from radiation-dominated regime.

Scattering relativistic electrons off a laser at 90 degrees of incidence allows to both create and accelerate a neutral e^+e^- beam in a single stage.

QED cascades can create abundant plasma and lead to a very efficient energy transfer from the laser into gamma-rays.