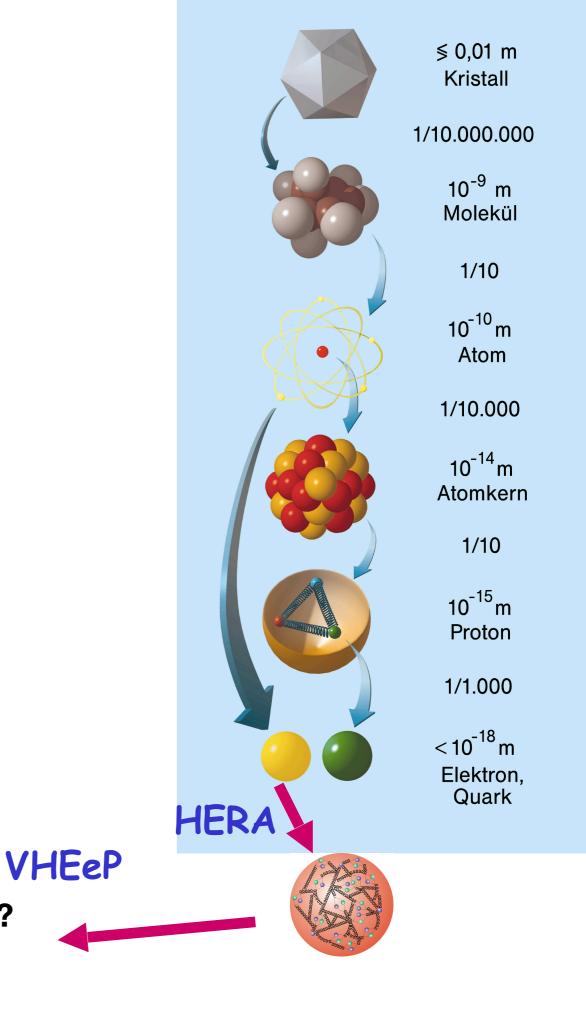
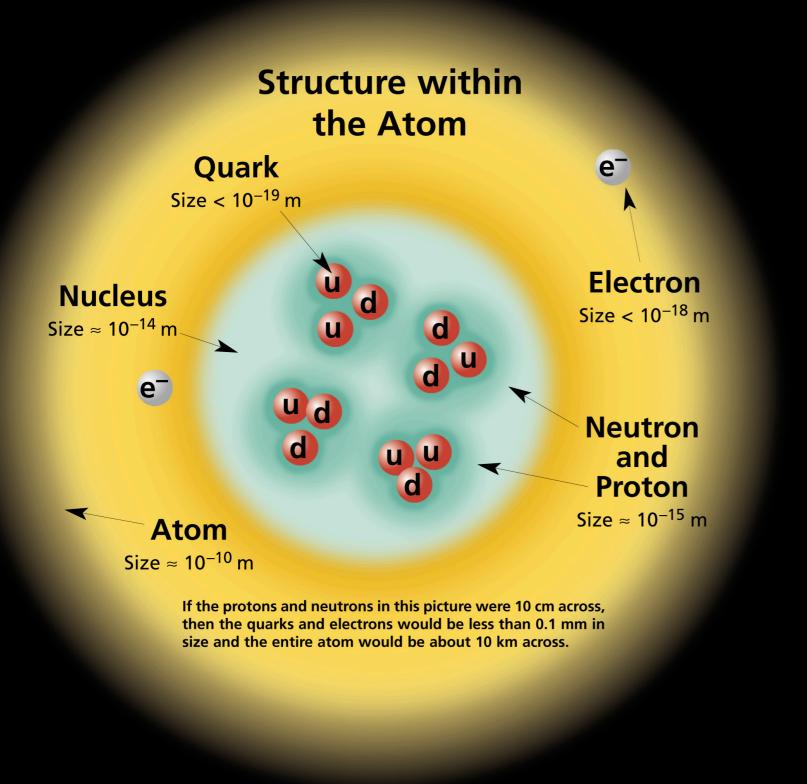
<u>Physics case</u> <u>high energy e/p and e/A collider</u>

Allen Caldwell Max Planck Institute for Physics

> Alegro Workshop CERN March 26, 2019



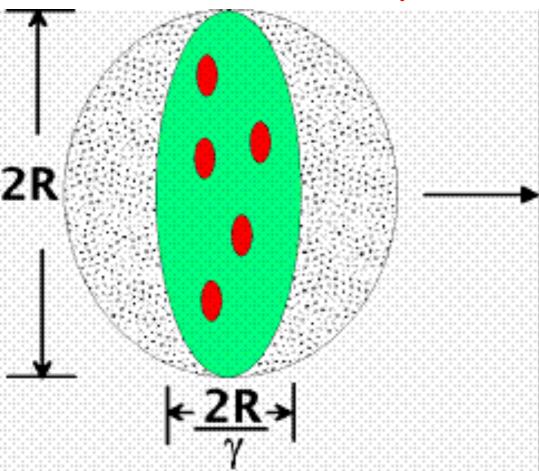
fundamental state of QCD matter? relation to gravity? color confinement?



To the outside world, the proton is made of three quarks, two 'u' and a 'd'.

Other quarks which have been discovered: s,c,b,t

Feynman's Parton Model



p

Imagine a very fast moving nucleon. Feynman asked us to think of it as a collection of point-like constituents- the partonswhich are behaving incoherently.



All partons moving parallel to the nucleon. Assume massless and no significant transverse momentum. Parton carries fractional momentum $\sum_{i} x_i = 1$

Interactions between partons in this frame time-dilated, so wavefunction not changing during the time of the interaction.

Kinematics

Proton
k
q=k-k'

Р

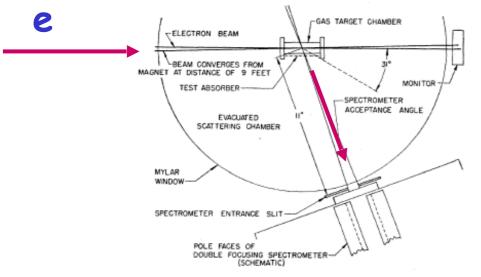
s=(k+P)² Q²=-(k-k')² W²=(q+P)² CM energy squared virtuality v*P CM energy squared

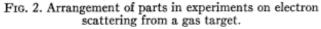
 $x \approx \frac{Q^2}{W^2}$

 W^2 Transverse distance scale probed: $b \approx \frac{\hbar c}{Q}$

> $b_{min}=0.4 \text{ fm}$ McAllister, Hofstadter Ee=188 MeV Bloom et al. 10 GeV0.05 fm 0.007 fm CERN, FNAL fixed target 500 GeV HERA 54 TeV 0.0007 fm LHeC/PEPIC 0.0002 fm 900 TeV **VHEeP** 45000 TeV 0.00003 fm

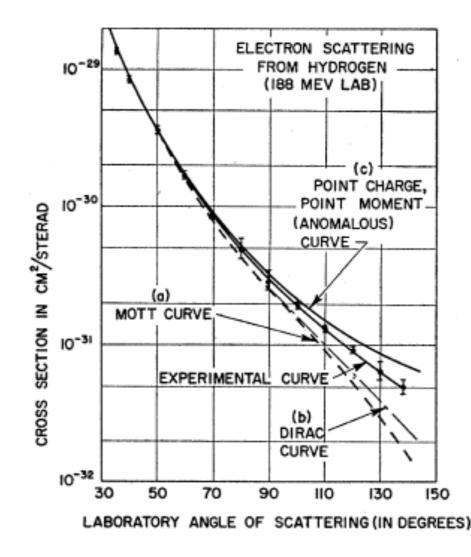
McAllister and Hofstadter 1956





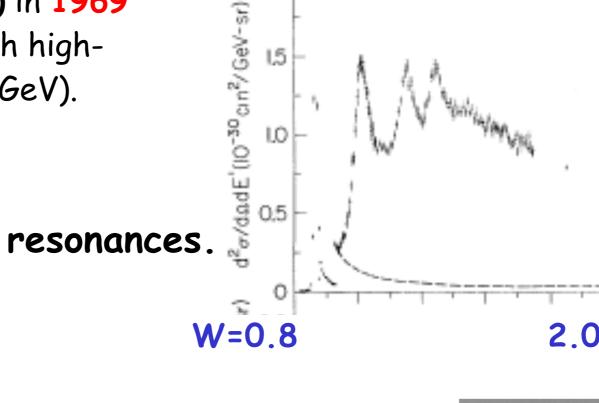
188 MeV and 236 MeV electron beam from linear accelerator at Stanford

Conclusion: radius of proton 0.7-0.8 fm

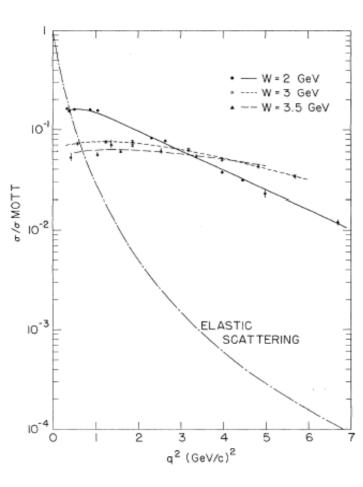


Bloom et al. (SLAC-MIT group) in 1969 performed an experiment with highenergy electron beams (7-18 GeV).

At small W, see proton and resonances.



2.0



The data did not fall with Q² as expected for elastic scattering on an extended nucleus. Rather, data looks like elastic scattering on a point charge (Mott).

SEARCH & DISCOVERY

(a)

FRIEDMAN, KENDALL AND TAYLOR WIN NOBEL PRIZE FOR FIRST QUARK EVIDENCE

dall of MIT and Rich



The 1990 Nobel laureates join hands at a SLAC fest follo

Scaling observed Predicted by Bjorken (1967). Implies electron scattering on point-like objects.

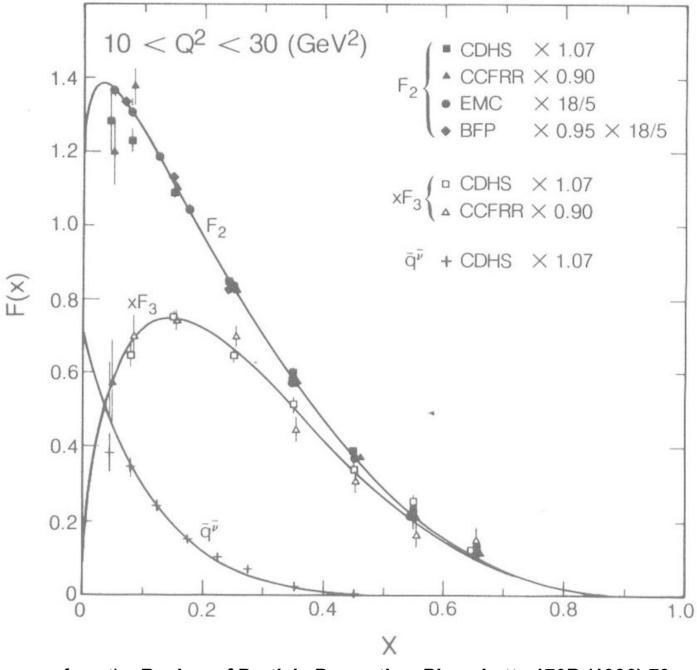
After discovery of quarks, a period of inelastic lepton-nucleon studies ensued:

Beams: electron, muon, neutrino

Targets: p, D, heavy targets

Main results: There is something else in the proton than quarks – gluons !

 $\int x \sum \left[\overline{q}(x) + q(x) \right] dx < 1$



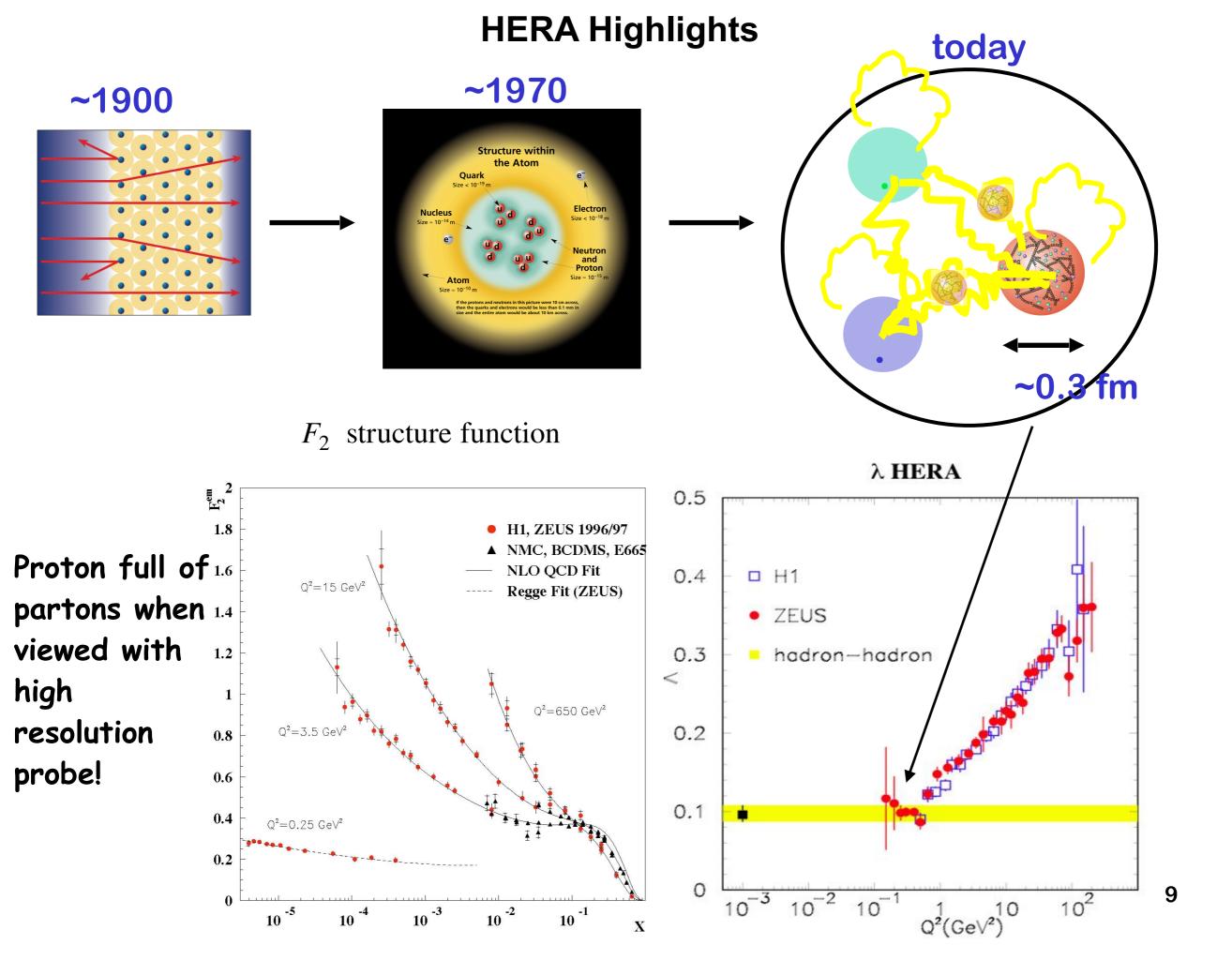
from the Review of Particle Properties, Phys. Lett., 170B (1986) 79.

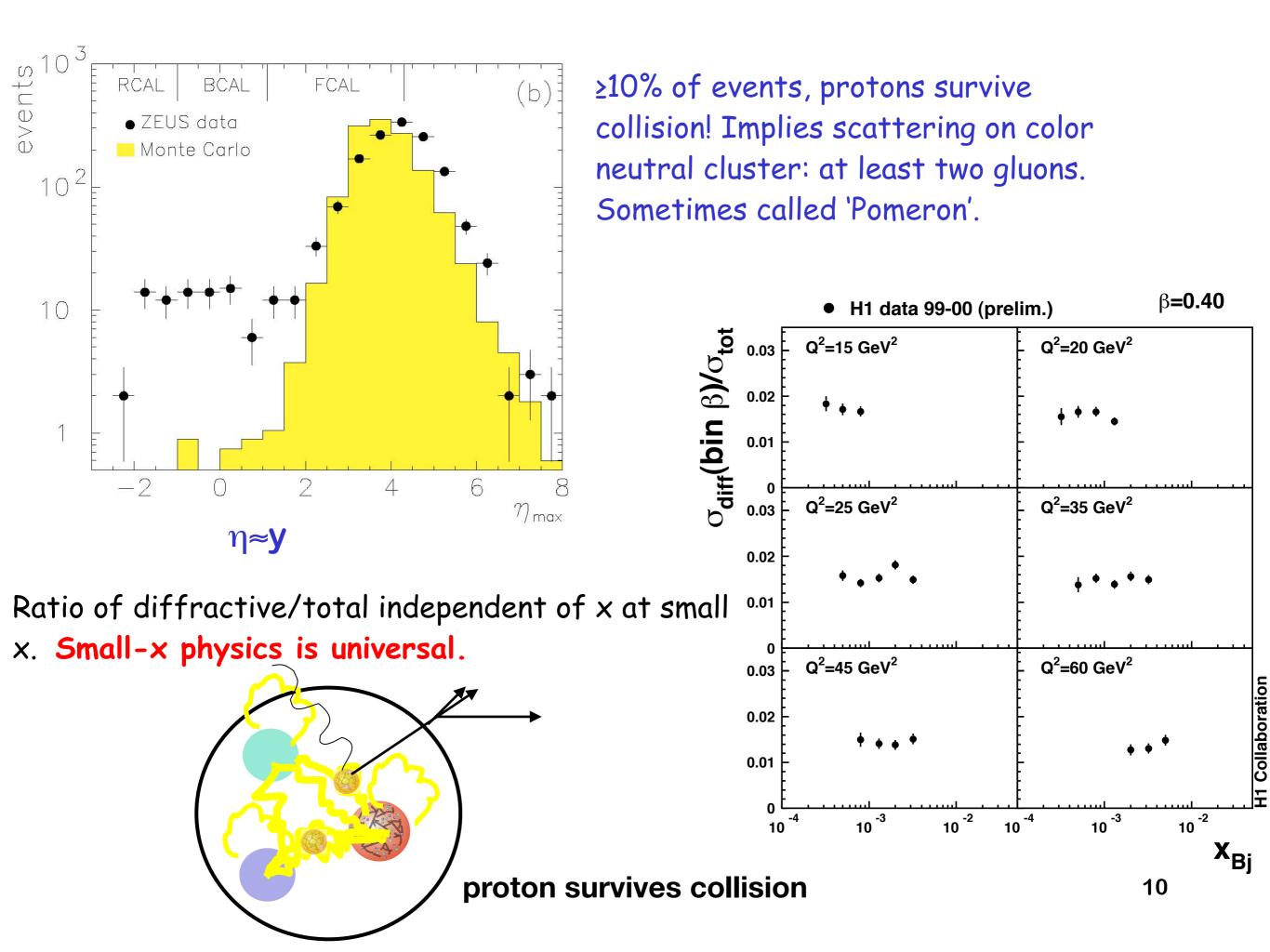
QCD was established as the correct description of the strong interactions.



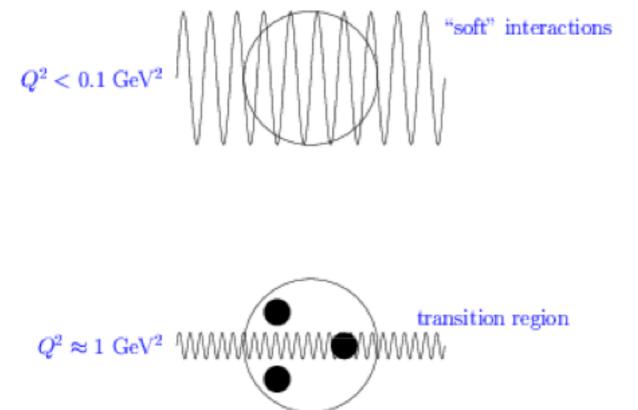
HERA





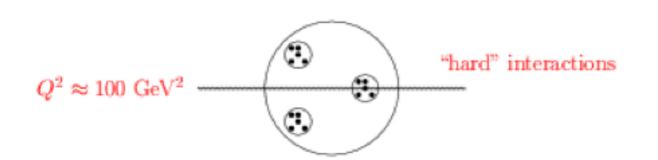


Higher energy allows finer probe



Not sensitive to details of proton structure

Start to see quarks – the valence quarks



See that valence quarks have complicated structure

What's next ?

New Physics in QCD

$${\cal L} = - {1 \over 4} F_{\mu
u} F^{\mu
u} - {n_f g^2 heta \over 32 \pi^2} F_{\mu
u} ilde{F}^{\mu
u} + ar{\psi} (i \gamma^\mu D_\mu - m e^{i heta\,' \gamma_5}) \psi$$

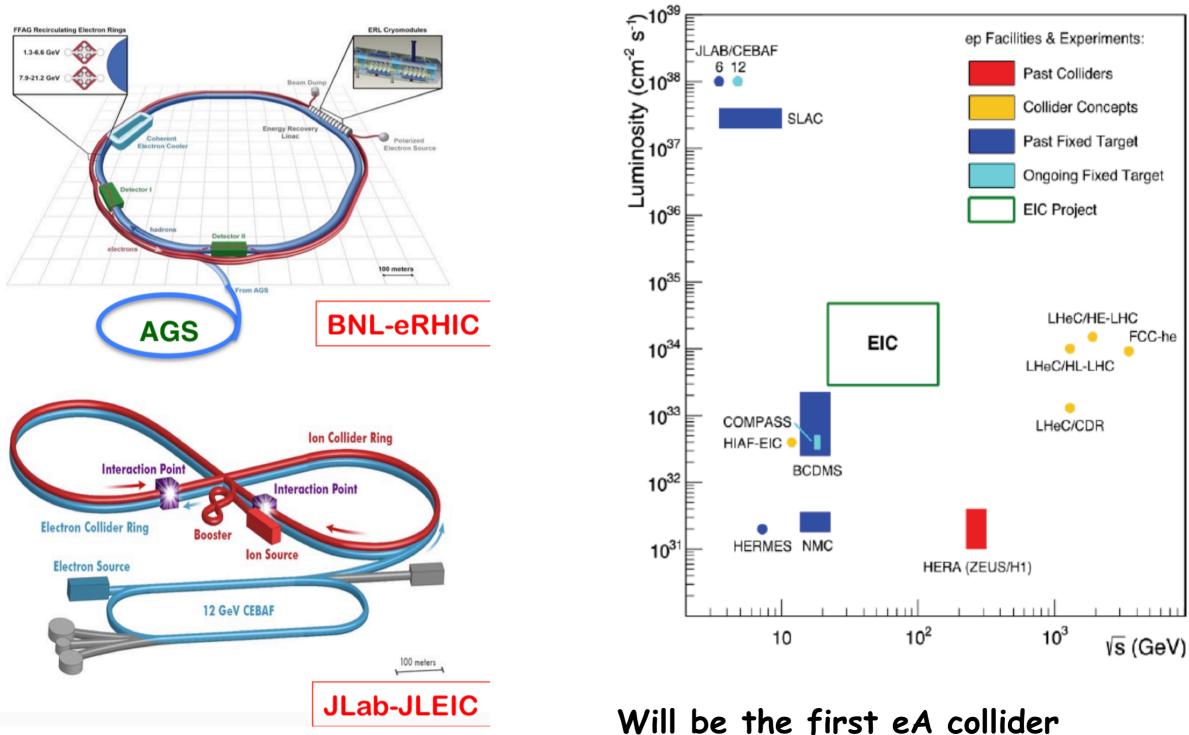
color confined ? mass of the proton = 1 GeV ? spin of the proton = 1/2 ? asymptotic freedom ? The interesting features of QCD are not apparent from the QCD Lagrangian - emergent features

Fundamental to understand how they emerge:

- most of the visible mass of the universe comes from QCD binding energy
- QCD is at the heart of everything (photons, electrons,...). What does matter look like in the high energy limit ?
- confinement of color at 1fm distance scale and spin structure - result of complicated multi particle nonlinear interactions

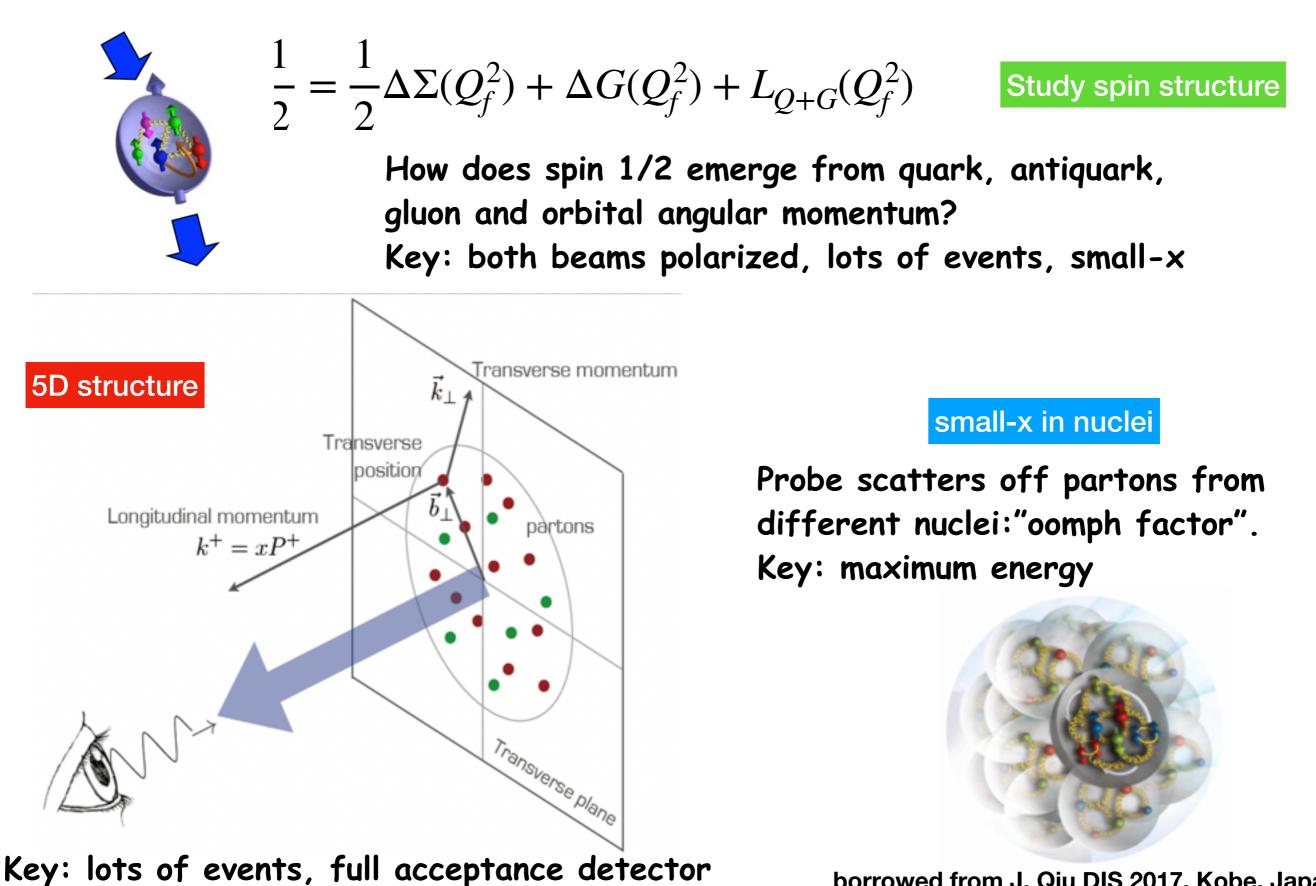
New ep/A colliders tasked with bringing insight into these important questions

Future Collider Options - EIC



First time both e,p polarized high luminosity. Realization: late 2020s

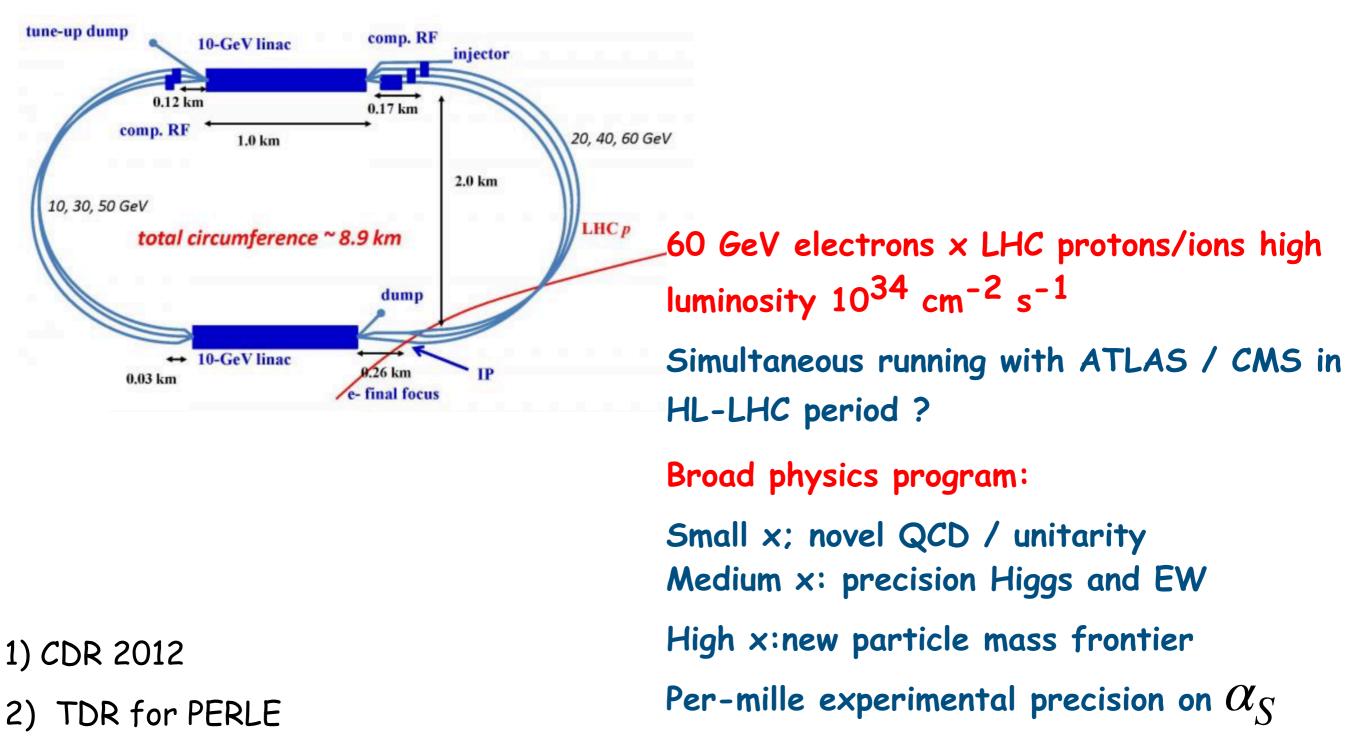
Future Collider Options - EIC



borrowed from J. Qiu DIS 2017, Kobe, Japan

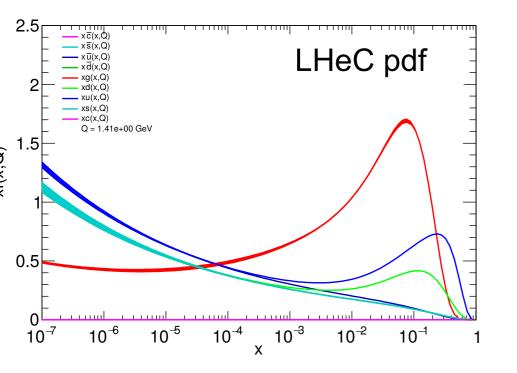
Future Collider Options - LHeC

LHeC

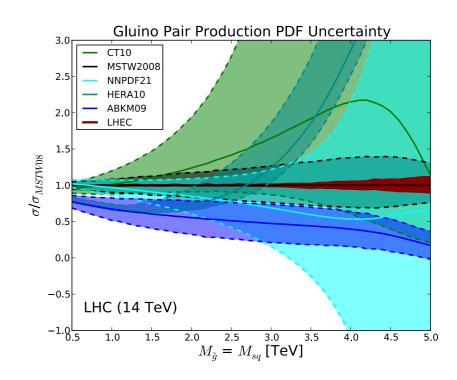


3) Further development of FCC eh

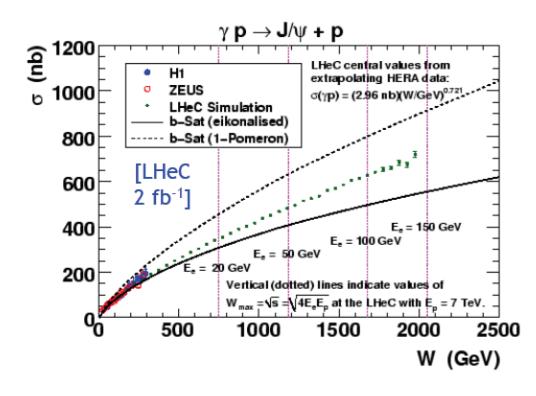
Future Collider Options - LHeC

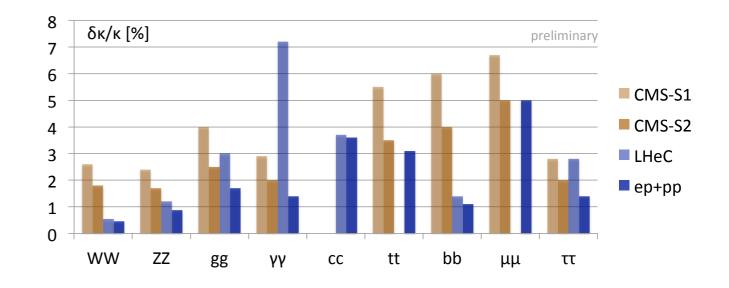


Precision proton structure functions important in recognizing new physics



small-x physics: vastly extended high lumi: exclusive observables !

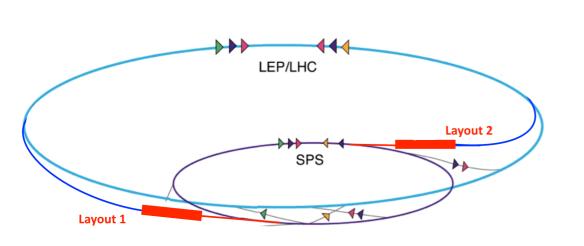




High lumi: sensitivity to Higgs couplings

eA will also be possible

Future Collider Options - PEPIC



PEPIC

Create ~50 GeV beam within 50–100 m of plasma driven by SPS protons

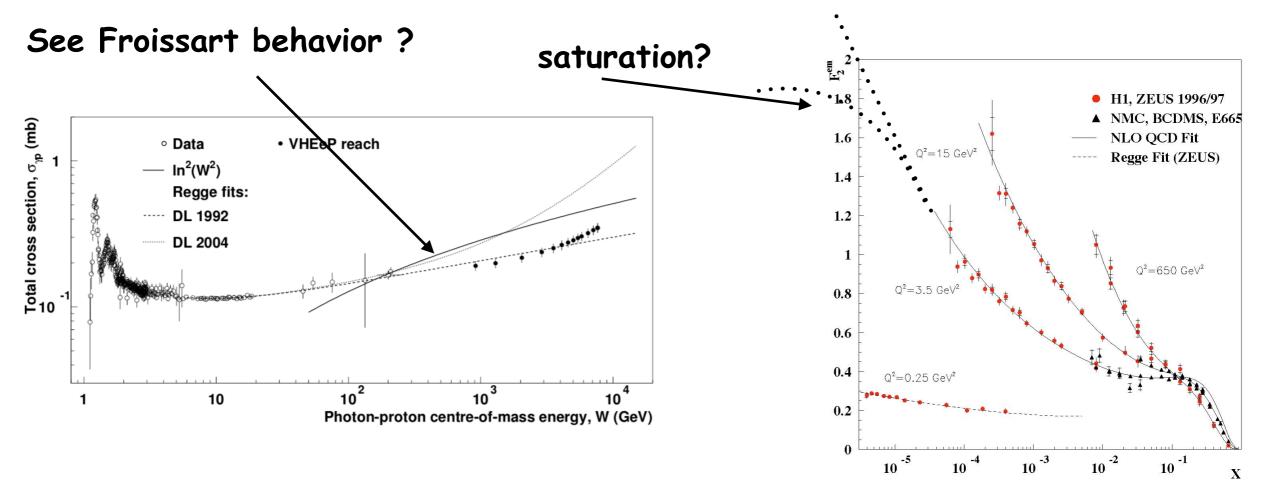
luminosity expected to be $\ll 10^{30}$ cm⁻² s⁻¹. Focus on QCD

Studied within Physics Beyond Colliders workshop @ CERN.

Relatively inexpensive first application of PWA?

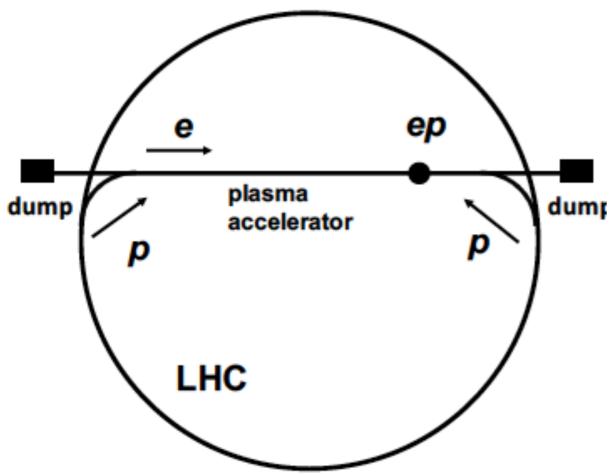
Physics: focus on QCD processes

large cross sections, growing with energy -> luminosity requirement modest



VHEeP

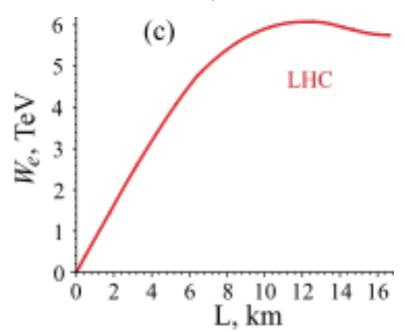
(Very High Energy electron-Proton collider)



One proton beam used for electron acceleration to then collide with other proton beam

Luminosity ~ 10²⁸ – 10²⁹ cm⁻² s⁻¹ gives ~ 1 pb–1 per year. Studies on achievable luminosity ongoing.

Electron energy from wakefield acceleration by LHC bunch



Choose $E_e = 3$ TeV as a baseline for a new collider

with $E_P = 7$ TeV yields $\int s = 9$ TeV. Can vary.

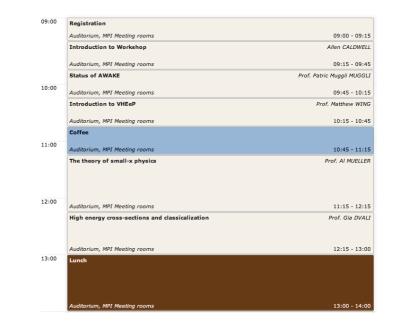
- Centre-of-mass energy ~30 higher than HERA.
- Reach in (high) Q^2 and (low) Bjorken x extended by ~1000 compared to HERA.
- Opens new physics perspectives

Prospects for a very high energy eP and eA collider

June 1,2 2017 Max Planck Institute for Physics

14:00	Applying AdS/CFT to very low x physics	Prof. Johanna ERDMENGER
	Auditorium, MPI Meeting rooms	14:00 - 14:45
	Low x synergy between DIS and ultrahigh energy neutrinos	Prof. Anna STASTO
15:00		
	Auditorium, MPI Meeting rooms	14:45 - 15:30
	Formation zone physics	leo STODOLSKY
	Auditorium, MPI Meeting rooms	15:30 - 16:00
16:00	Coffee	
	Auditorium, MPI Meeting rooms	16:00 - 16:30
17:00	Auditorium, MPI Meeting rooms	16:30 - 18:00
18:00	Reception	
		18:00 - 19:15
19:00	Auditorium, MPI Meeting rooms	18.00 - 19.15
19:00	Auditorium, MPI Meeting rooms Dinner	10.00 - 19.15

09:00	Small-x physics in ep-scattering: thoughts on results from HERA and future aspects	Prof. Jochen BARTELS
	Auditorium, MPI Meeting rooms	09:00 - 09:45
	BKFL and dipoles	Dr. Henri KOWALSKI
10:00	Auditorium, MPI Meeting rooms	09:45 - 10:15
	Color dipole at small x	Prof. Dieter SCHILDKNECHT
	Auditorium, MPI Meeting rooms	10:15 - 10:45
	Coffee	
11:00	Auditorium, MPI Meeting rooms	10:45 - 11:15
	eA physics at very high energies	Mrs. Heikki MÄNTYSAARI
	Auditorium, MPI Meeting rooms	11:15 - 12:00
12:00	Polarised eP and eA physics	Dr. Elke ASCHENAUER
	Auditorium, MPI Meeting rooms	12:00 - 12:45
	Lunch	
13:00		
	Auditorium, MPI Meeting rooms	12:45 - 13:45



	What the HERA data tell us about low-x physics	Dr. Volodymyr MYRONENKO
14:00		
	Auditorium, MPI Meeting rooms	13:45 - 14:30
	New results for VHEeP	Mr. Fearghus KEEBLE
	Auditorium, MPI Meeting rooms	14:30 - 15:00
15:00	Simulation of high energy ep / eA collisions	Dr. Simon PLAETZER
	Auditorium, MPI Meeting rooms	15:00 - 15:45
	Close out	Allen CALDWELL et al.
	Auditorium, MPI Meeting rooms	15:45 - 16:00

Mini-workshop on QCD and Gravity December 12,13 Max Planck Institute for Physics **2018**

Wednesday, December 12

- 14:15-15:15 Raju Venugopalan 'A many-body theory of QCD in the Regge limit'
- 15:30-16:00 Eran Palti 'News on Swampland'
- 16:00-16:30 Stephan Stieberger 'QCD meets Gravity'
- 16:30-17:00 Johanna Erdmenger 'AdS/CFT and very small x'
- 17:00-17:30 Angnis Schmidt-May 'News from bimetric gravity'
- 17:30-18:30 Discussion time
- 19:00- Dinner somewhere

Thursday, December 13

9:00-10:00 Gia Dvali et al 'Proof of the Axion?'

- 10:00-10:30 Discussion time
- 10:30-11:00 Henri Kowalski 'BFKL analysis of HERA data'
- 11:00-11:30 Agustin Sabio Vera 'The Regge limit in QCD, SUSY and gravity'

2 workshops to discuss novel physics with very high energy eP/eA collider.

Second - focus on relation between QCD and gravity.

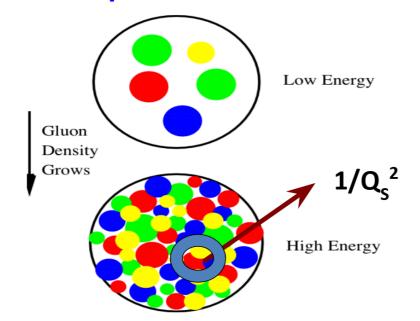
12:00-14:00 lunch and discussion

Small-x physics and color confinement. One of the key unsolved questions in physics.

How?

R. Venugopalan, Mini-workshop on QCD and Gravity **December 12,13 Max Planck Institute for Physics**

The boosted proton viewed head-on



When occupancies become large ~ $1/\alpha_s$, gluons resist further close packing by recombining and screening their color charges -- leading to gluon saturation

Emergent semi-hard scale dynamical scale $Q_s(x) >> \Lambda_{OCD}$

Asymptotic freedom! $\alpha_s(Q_s) \ll 1$ provides weak coupling window into infrared

Q^2 (GeV²) Resolution Quarks and Gluons Linear evolution **Strongly Correlated** Nonlinearevolutio **Quark-Gluon Dynamics** perturbative Ost) **High-Density** Gluon Matter non-perturbative

Hadrons

Parton Density

Pomerons

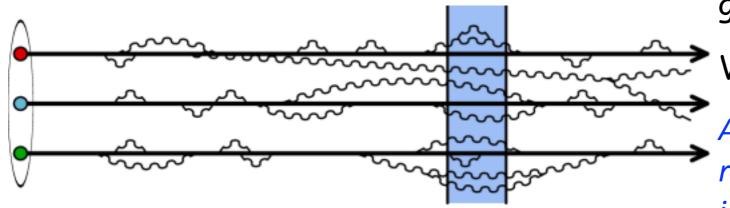
Regge trajectories

From the physics we know - these are the strongest fields that can be achieved in nature

New: perturbative approach to infrared physics! Relevant equations very similar to fundamental statistical mechanics equations. Strong overlap with quantum description of black holes.

Saturation in the QCD landscape

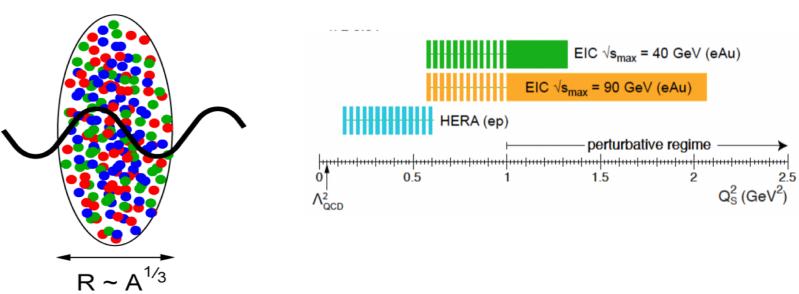
At high energy, see short-lived fluctuations due to time dilation





"oomph factor" from Nuclei

Q_s² ~ A^{1/3} since "wee" gluons couple coherently for x << A^{-1/3}



Markovian process leads to power law growth of gluon distribution at small x

Violates Froissart bound asymptotically

A fascinating equilibrium of splitting and recombination should eventually result. It is a considerable theoretical challenge to calculate this equilibrium in detail...

VHEeP

F. Wilczek, Nature (1999)

with VHEeP and eA, we will be in a region where the saturation scale is well into the perturbative region. Allows detailed probing of this new physics: high density & weak coupling !

QCD and Gravity: more than math ?

Consider: the visible mass is largely due to baryons. The mass of baryons is largely due to QCD (not the Higgs mechanism). Gravity couples to mass

S. Stieberger, Mini-workshop on QCD and Gravity December 12,13 Max Planck Institute for Physics

Can gravity be described by YM-theory ? do we see some generic or unifying structures in scattering amplitudes?

 $G: \operatorname{spin} 2$ $\gamma: \operatorname{spin} 1$

graviton as composite particle ?

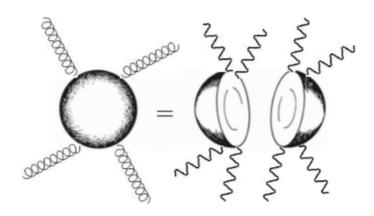
<u>Alert:</u> graviton cannot be a composite particle in a relativistic quantum field theory (Weinberg, Witten)

proof relies on the construction of a <u>conserved and Lorentz-covariant stress tensor</u>

$$T^{\mu\nu}(x) = (-g)^{-1/2} \frac{\partial}{\partial g_{\mu\nu}(x)} S[g]$$

theorem holds for any known renormalizable field-theory, e.g. QCD

However there are many ways out: massive gravity, <u>conformal field theory, string theory</u>, ...



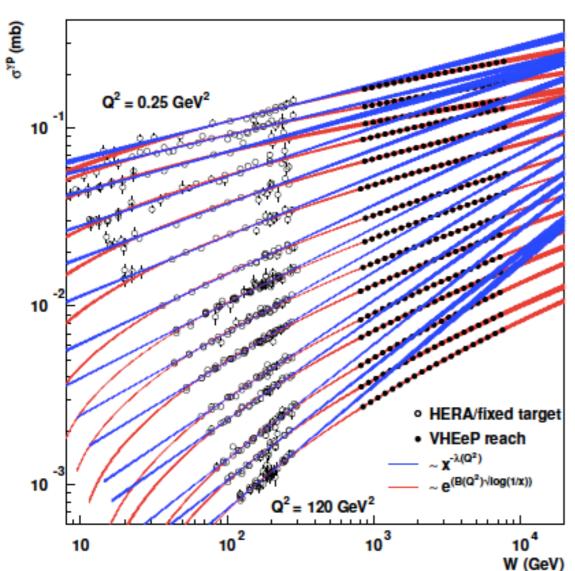
Concluding remarks

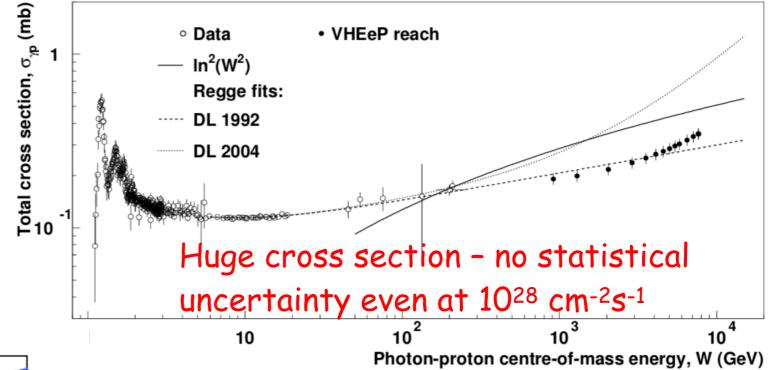
- growing set of <u>interconnections</u> between open & closed amplitudes with gauge theory and supergravity amplitudes
- indication for the existence of some gauge structure in quantum gravity

What will we measure and why will it help?

Total photoproduction cross section – energy dependence ? See approach to Froissart bound ?

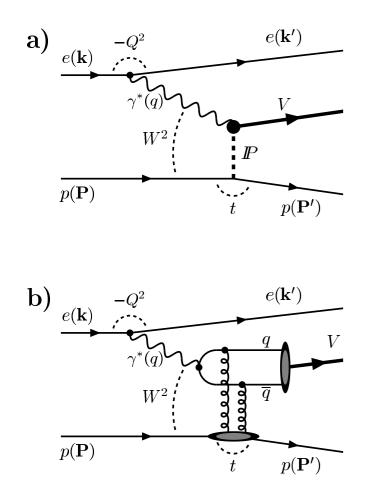
Impact on cosmic ray physics





Virtual photon cross section: unphysical extrapolation of cross sections -> observation of saturation of parton densities ?

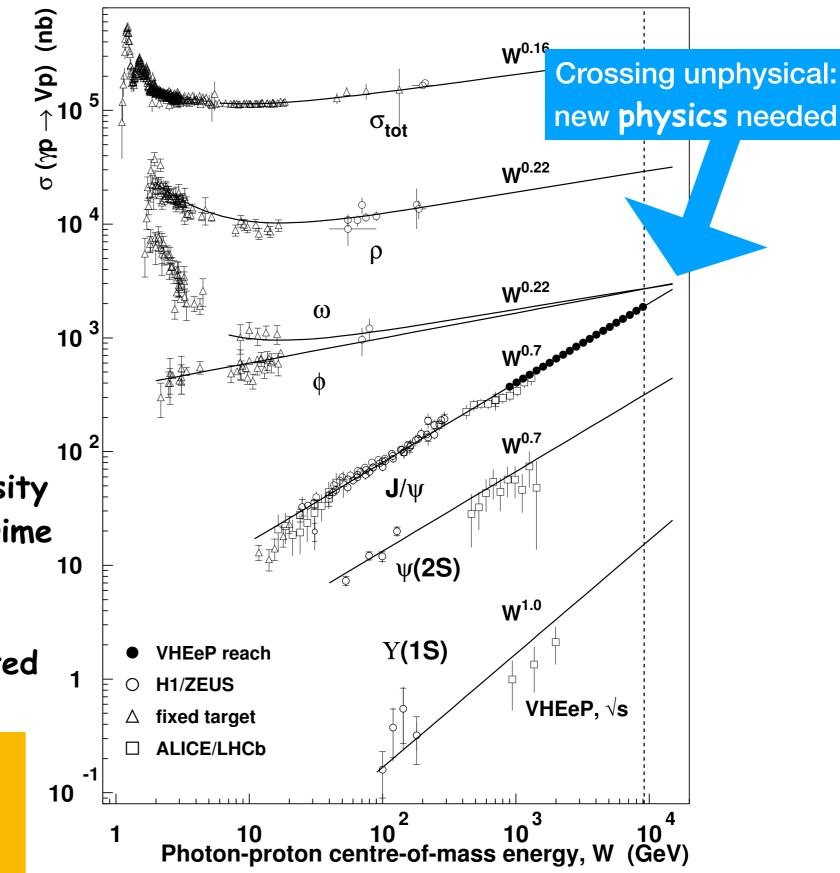
With the three orders of magnitude extension in the range at small-x, expect to see signs of the fundamental saturated regime.

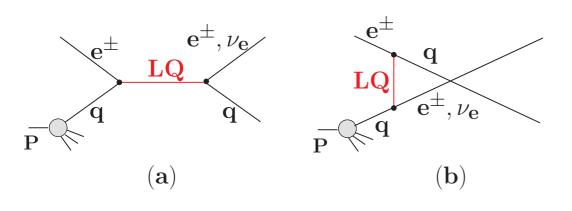


Exclusive processes: Sensitive to square of gluon density Early signs of new saturated regime

Good opportunity to see the fundamental high-energy saturated state!

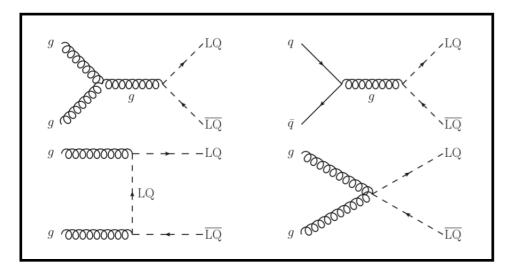
eA possibility will make this physics even more dramatic "oomph"-factor again





 e^{\pm} e^{\pm} , ν_{e} γ, Z, W q qP (c)

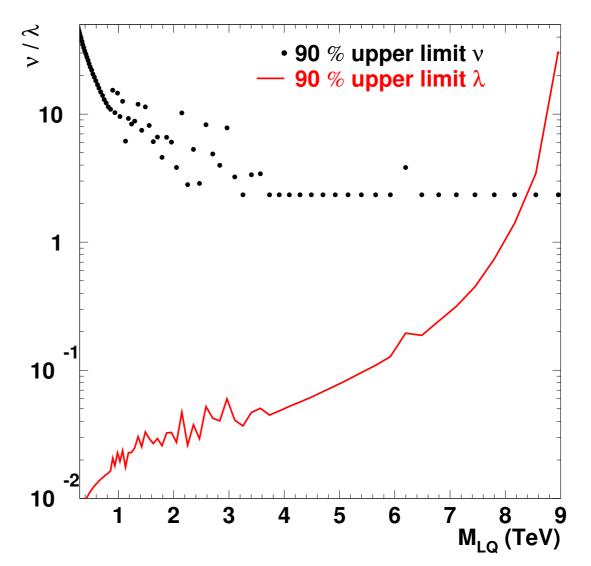
Sensitivity well beyond expected reach of LHC



At high energy - leptoquarks appear ?

tailor-made for eP Collider

Simulation study: 100 pb⁻¹



High energy electron-proton and electron-ion collider

new conceptual breakthroughs, (probably) not new particles

- where mass comes from (saturated high energy state of QCD matter).
- approach color confinement via large saturation scale
- connections to gravity may become apparent
- and lot's of 'bread-and-butter' physics