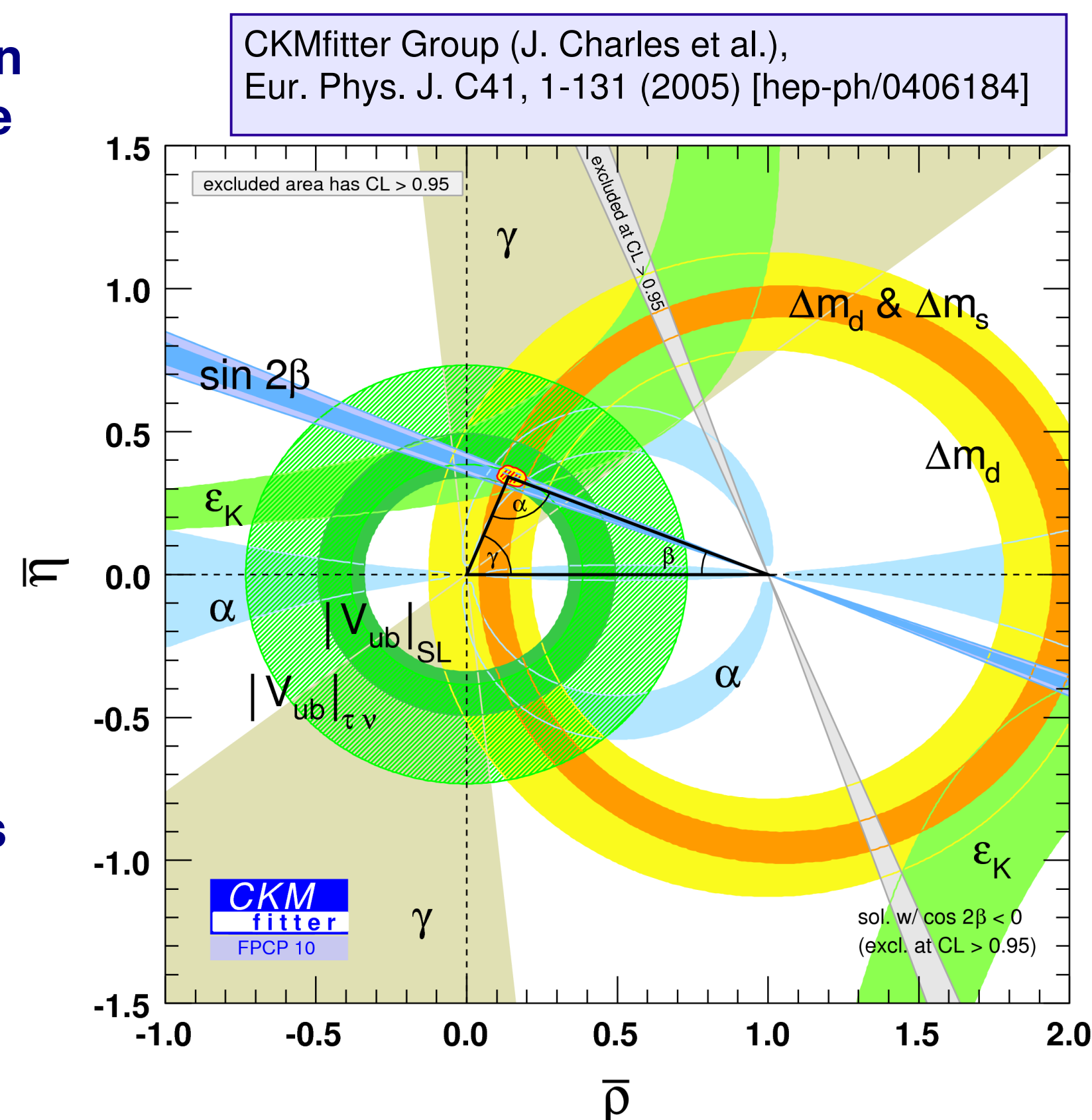


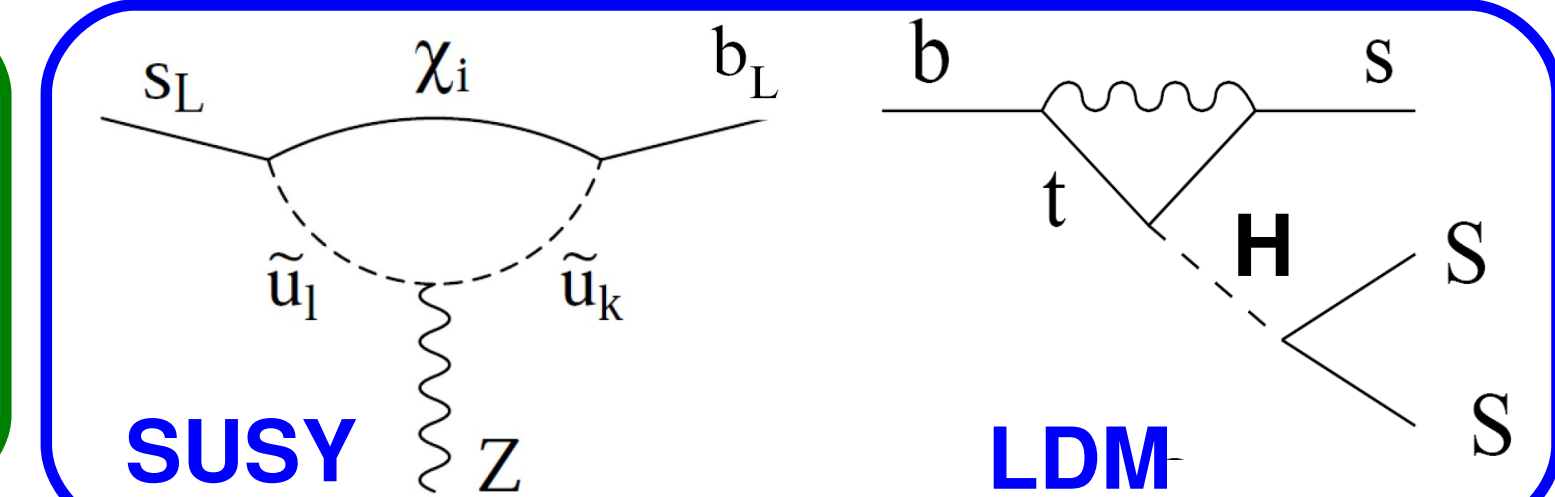
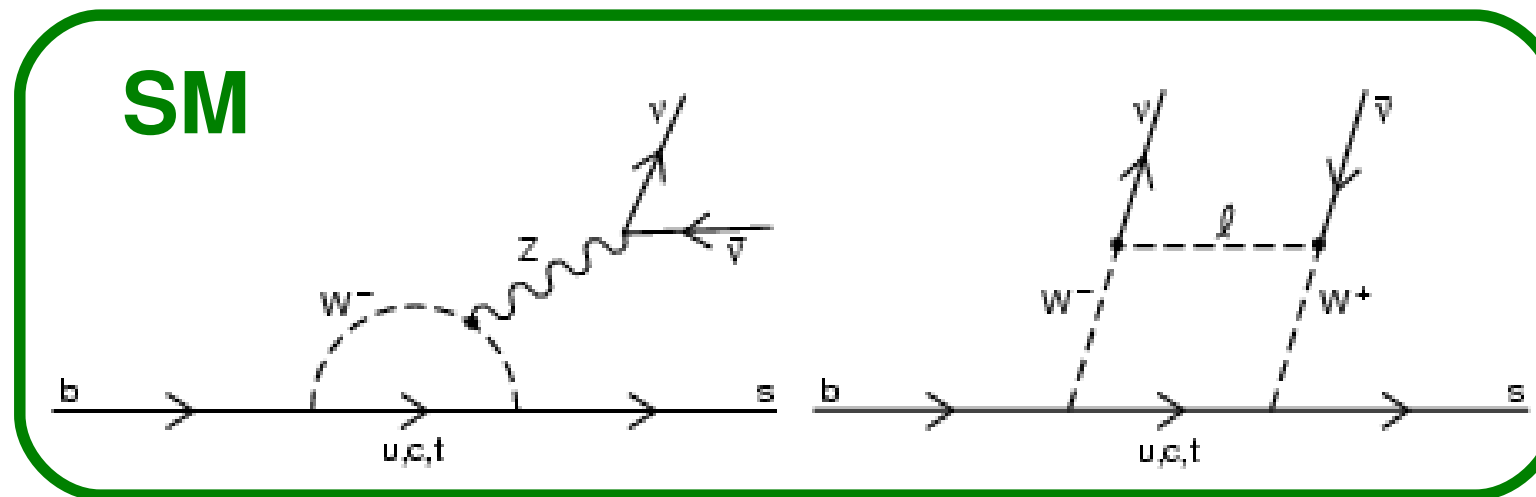
General Introduction to Rare decays

- The Standard Model (SM) has been able to explain coherently (almost) all the experimental evidence of electroweak and strong interactions
- Cabibbo-Kobayashi-Maskawa (CKM) picture has been able to explain all the measurements in the flavor sector
- BUT everybody is eager for New Physics (NP)**
- Currently two complementary approaches:
 - Explore unopened energy frontiers (LHC)
 - Measure precisely virtual processes which can test high energy scales (flavor factories)
- Rare decays are prolific ground and clean probes for this:
 - If a suppressed decay is observed above SM prediction \Rightarrow clean sign of NP
 - If an UL is set \Rightarrow NP scenarios are constrained



The $B \rightarrow K^* \nu \bar{\nu}$ and $B \rightarrow K \nu \bar{\nu}$ system

- Electroweak penguin (loop diagrams) radiated processes ($b \rightarrow s$):
 - Flavor changing neutral current (FCNC) prohibited in SM at tree level
 - Sensitive to Higgs, SUSY particles, light dark matter (LDM)

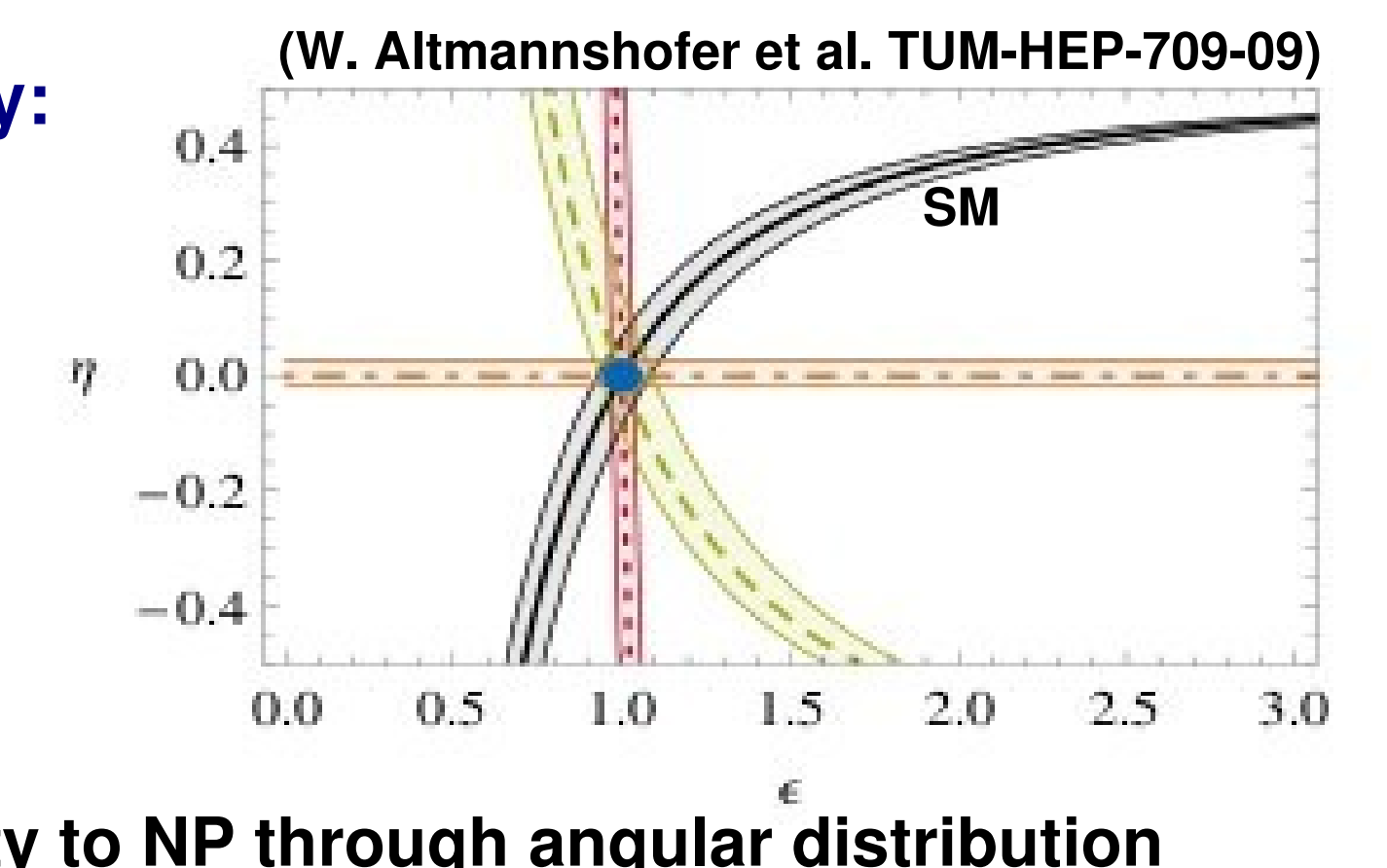


- $b \rightarrow s \nu \bar{\nu}$ model independent phenomenology:

- $BR(B \rightarrow K \nu \bar{\nu}) = 4.5 \times 10^{-6} (1 - 2\eta) \epsilon^2$
- $BR(B \rightarrow K^* \nu \bar{\nu}) = 6.8 \times 10^{-6} (1 + 1.31\eta) \epsilon^2$
- $FL(B \rightarrow K^* \nu \bar{\nu}) = 0.54 (1 + 2\eta) / (1 + 1.31\eta)$

$$\frac{d\Gamma}{d\cos\theta} \propto \frac{3}{4} (1 - \langle F_L \rangle) \sin^2\theta + \frac{3}{2} \langle F_L \rangle \cos^2\theta$$

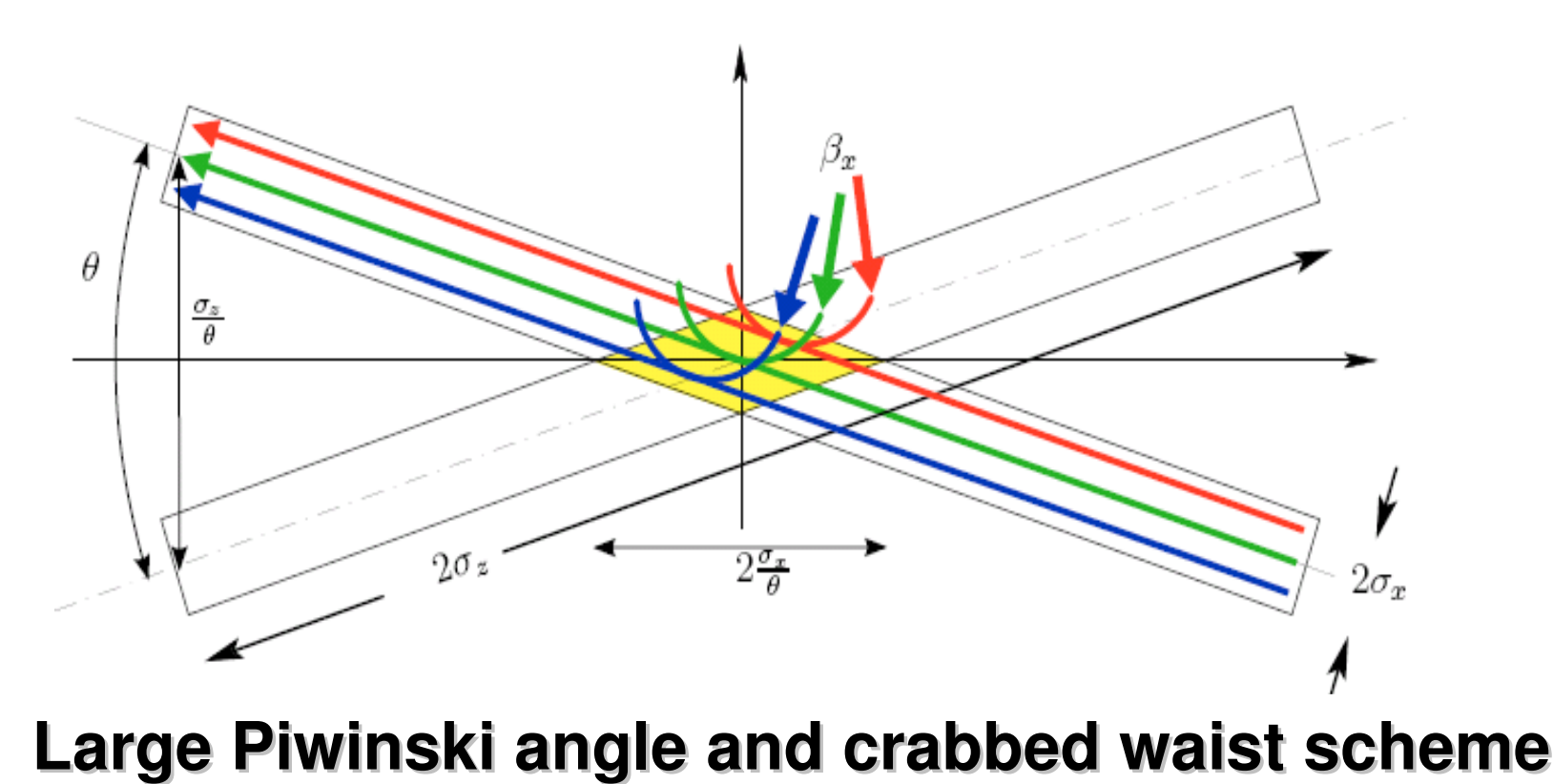
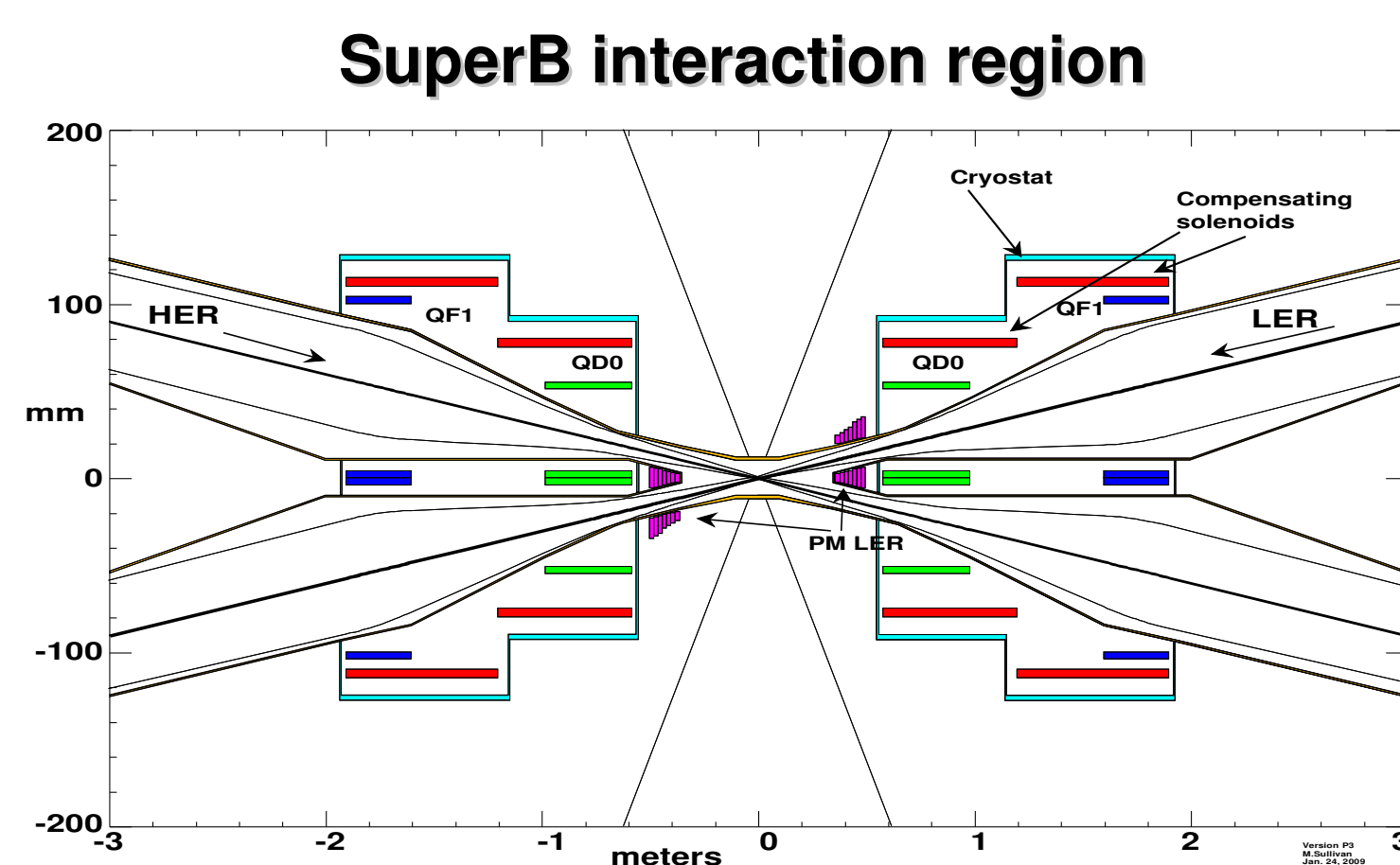
- θ = angle between:
 - K^* direction in B rest frame
 - K direction in K^* rest frame



- Good sensitivity to NP through angular distribution

What is the SuperB project?

- SuperB aims at the construction of a very high luminosity asymmetric e^+e^- flavor factory:



- Aims:
 - Operate with a reduced boost $\beta\gamma = 0.28$ (BaBar/Belle is 0.56) \Rightarrow higher hermeticity
 - Very high luminosity of $1 \times 10^{36} \text{ cm}^{-2} \text{ s}^{-1}$ or more (100 more than BaBar/Belle)
 - High reliability
 - Polarized e^- beam at interaction point $\Rightarrow \tau$ physics
 - Ability to collide at the $\Upsilon(4S)$ and at J/ψ masses $\Rightarrow B$ and D physics

The conception of the SuperB detector

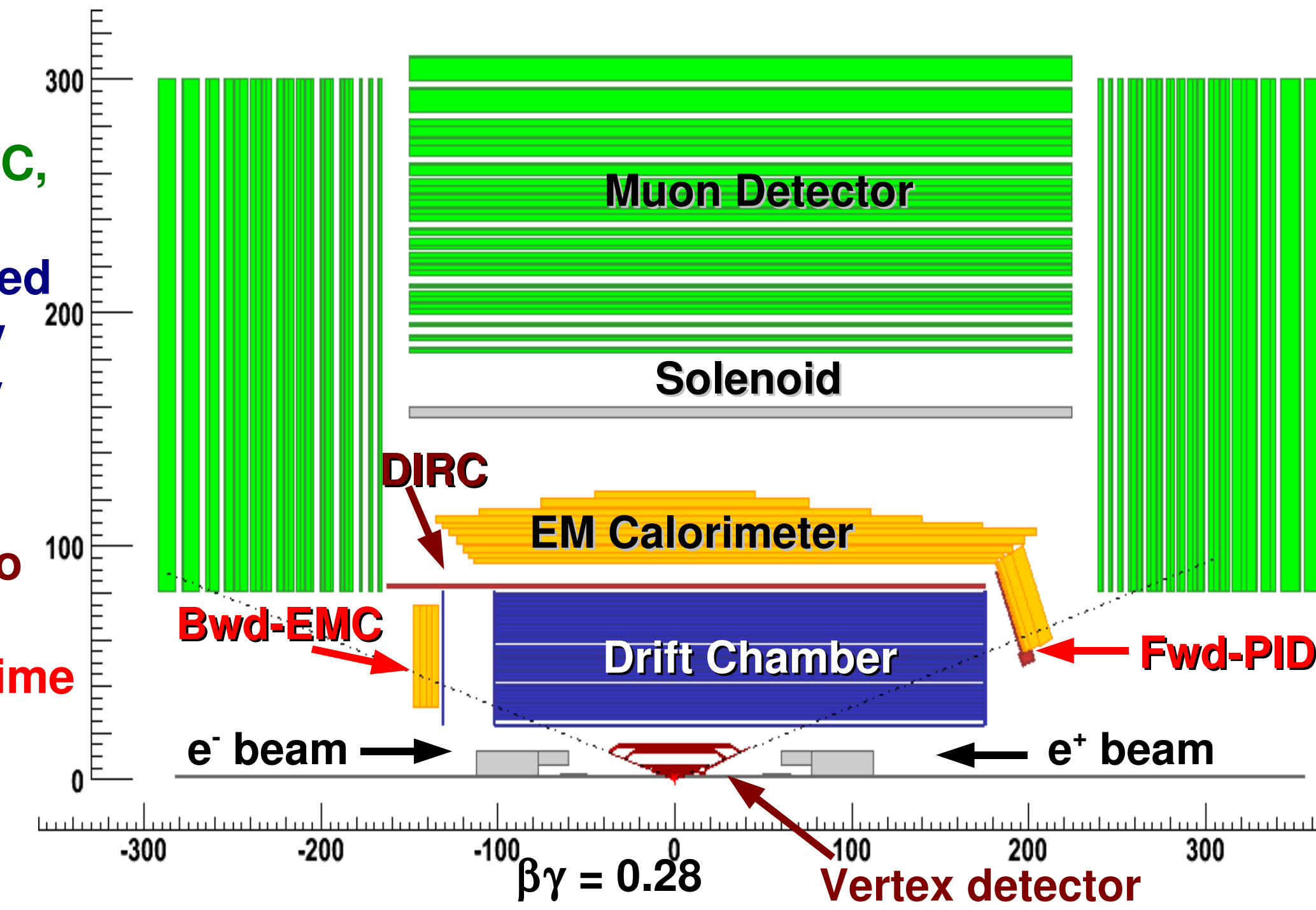
- SuperB detector concept is based on BaBar

- Silicon vertex detector (SVT) } Tracking system
- Drift Chamber (DCH)
- Cherenkov detector (DIRC) $\Rightarrow K/\pi$ identification

- Electromagnetic Calorimeter (EMC)
- Instrumented Flux Return (IFR)

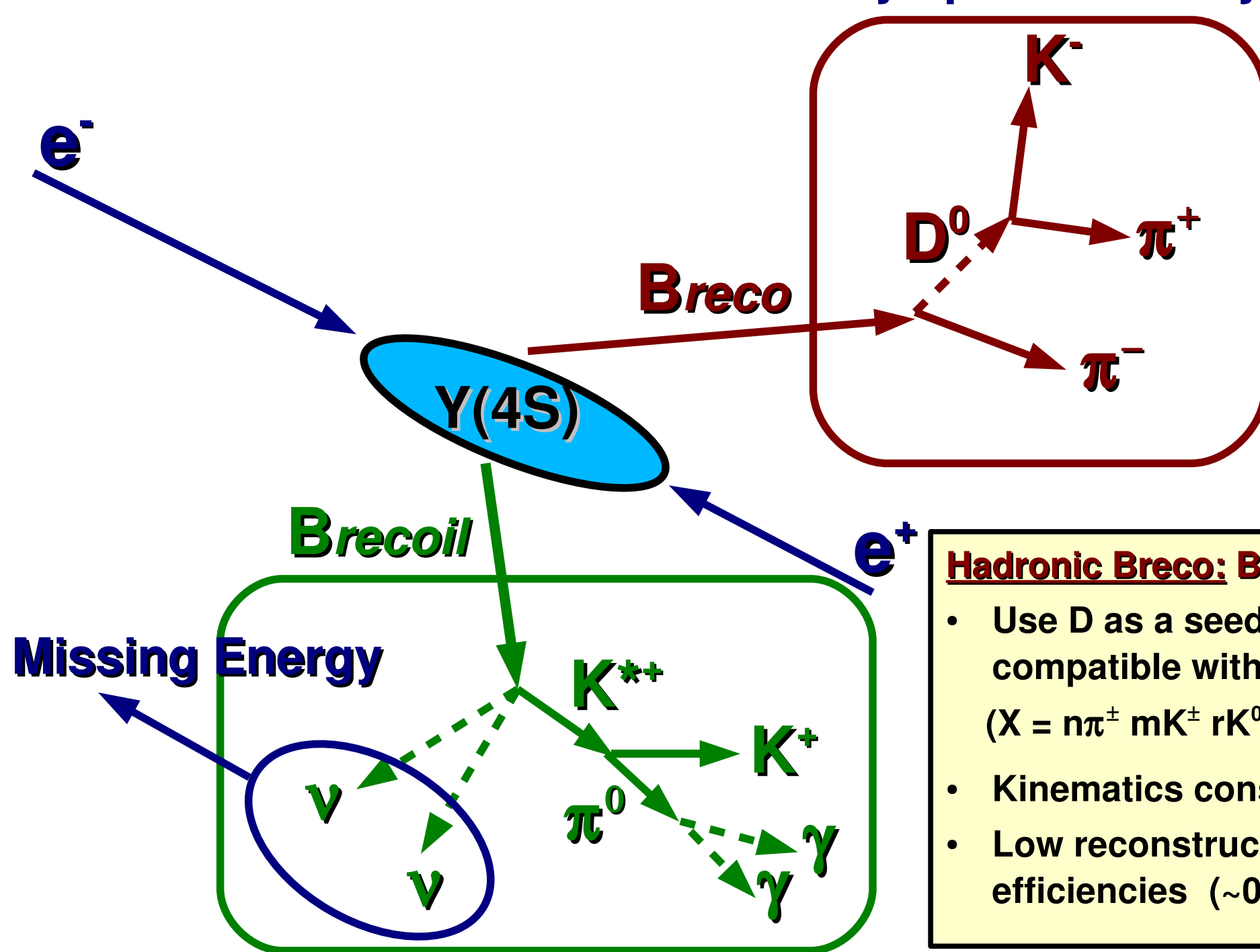
- Reuse a number of components from BaBar: flux-return-steel, Super conducting coil, barrel EMC, DIRC fused silica radiators
- Other components will be replaced with new ones which can reliably operate at the SuperB luminosity and lower boost

- Two new devices are proposed to further improve the hermeticity:
 - a forward PID device based on time of flight measurements (tTOF)
 - a backward EMC (Bwd-EMC)



The experimental technique and current status

- Most of the searches of rare B decays performed by exploiting the Recoil Technique:



Breco: full (partial) reconstruction of one B into a hadronic (semi-leptonic) final state

Brecoil: look for the signal signature, e.g. $K^{(*)}$ not accompanied by additional (charged+neutral) particles + Missing Energy

- Hadronic Breco: $B \rightarrow DX$**
- Use D as a seed to add X system compatible with B hypothesis ($X = n\pi^+ mK^+ rK^0_s q\pi^0$ and $n+m+r+q \leq 6$)
 - Kinematics constrained completely
 - Low reconstruction efficiencies ($\sim 0.4\%$)

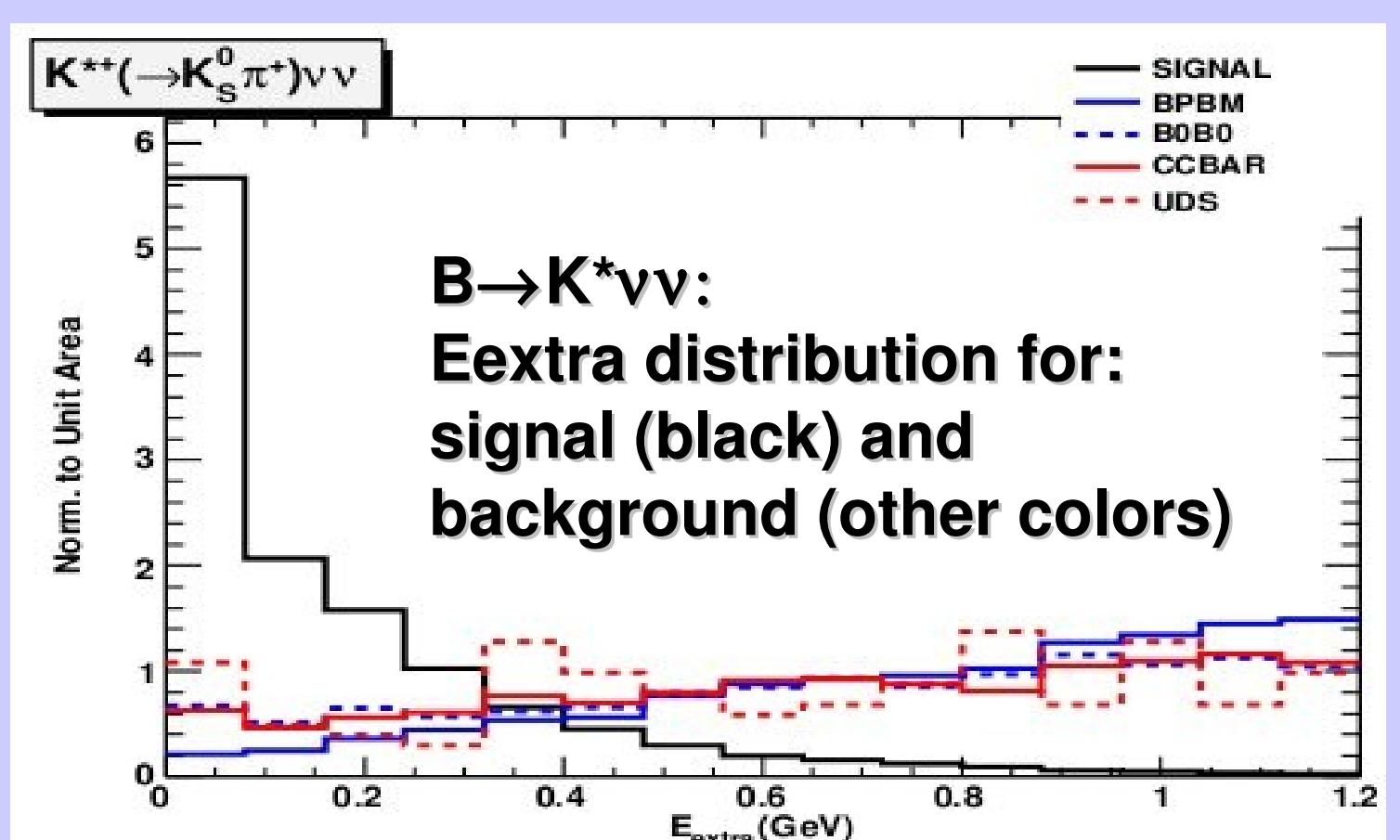
- Semi-Leptonic Breco: $B \rightarrow D\ell\nu$**
- Use D as a seed and a lepton to form a $D\ell$ pair ($\ell = e^+, \mu^+$)
 - Kinematics is unconstrained due to neutrino
 - Higher reconstruction efficiencies ($\sim 2.0\%$)

$B \rightarrow K^{(*)} \nu \bar{\nu}$ Analysis strategy:

- $B \rightarrow K \nu \bar{\nu}$: look for single K^+ or K^0_s in Brecoil
- $B \rightarrow K^* \nu \bar{\nu}$: look for $K^{*(+0)}$ in Brecoil. Several modes: $K^{*+} \rightarrow K^0_s \pi^+$ and $K^* \pi^0$, $K^{*0} \rightarrow K^+ \pi^-$
- Require opposite charges of Breco and Brecoil for charged modes
- Additional cuts on other kinematic variables (K^0_s and K^+ mass) and missing energy
- Main discriminant variable:**
 $E_{\text{extra}} = \sum(\text{extra neutrals in the EMC})$

Recoil technique at B-Factories:

- search for rare decays (10^{-5}) with missing energy (Not possible at hadronic machines)
- Several benchmark channels at SuperB:
 $B \rightarrow \tau \nu$, $B \rightarrow K^{(*)} \nu \bar{\nu}$, ...



Current Status

BABAR	BELLE
SL Recoil (90 million BB pairs) ² : $BF(B^+ \rightarrow K^+ \nu \bar{\nu}) < 5.2 \times 10^{-5}$	Had Recoil (535 million BB pairs) ¹ : $BF(B^+ \rightarrow K^+ \nu \bar{\nu}) < 1.4 \times 10^{-5}$
Had Recoil (351 million BB pairs) ³ : $BF(B^+ \rightarrow K^+ \nu \bar{\nu}) < 4.2 \times 10^{-5}$	$BF(B^0 \rightarrow K^0 \nu \bar{\nu}) < 1.6 \times 10^{-4}$
Had+SL Recoil (454 million BB pairs) ⁴ : $BF(B^+ \rightarrow K^+ \nu \bar{\nu}) < 8.0 \times 10^{-5}$	$BF(B^+ \rightarrow K^{*+} \nu \bar{\nu}) < 1.4 \times 10^{-4}$
	$BF(B^0 \rightarrow K^{*0} \nu \bar{\nu}) < 3.4 \times 10^{-4}$

¹ K. F. Chen et al. [Belle Collaboration], Phys. Rev. Lett. 99, 221802 (2007).
² B. Aubert et al. [BaBar collaboration], Phys. Rev. Lett. 101801 (2005).
³ H. Kim on behalf of the BaBar collaboration, arXiv:hep-ex/08052365 (2008).
⁴ B. Aubert et al. [BaBar collaboration], Phys. Rev. D78:072007,2008

All measurements (all upper limits!) still consistent with SM expectation

Some preliminary results

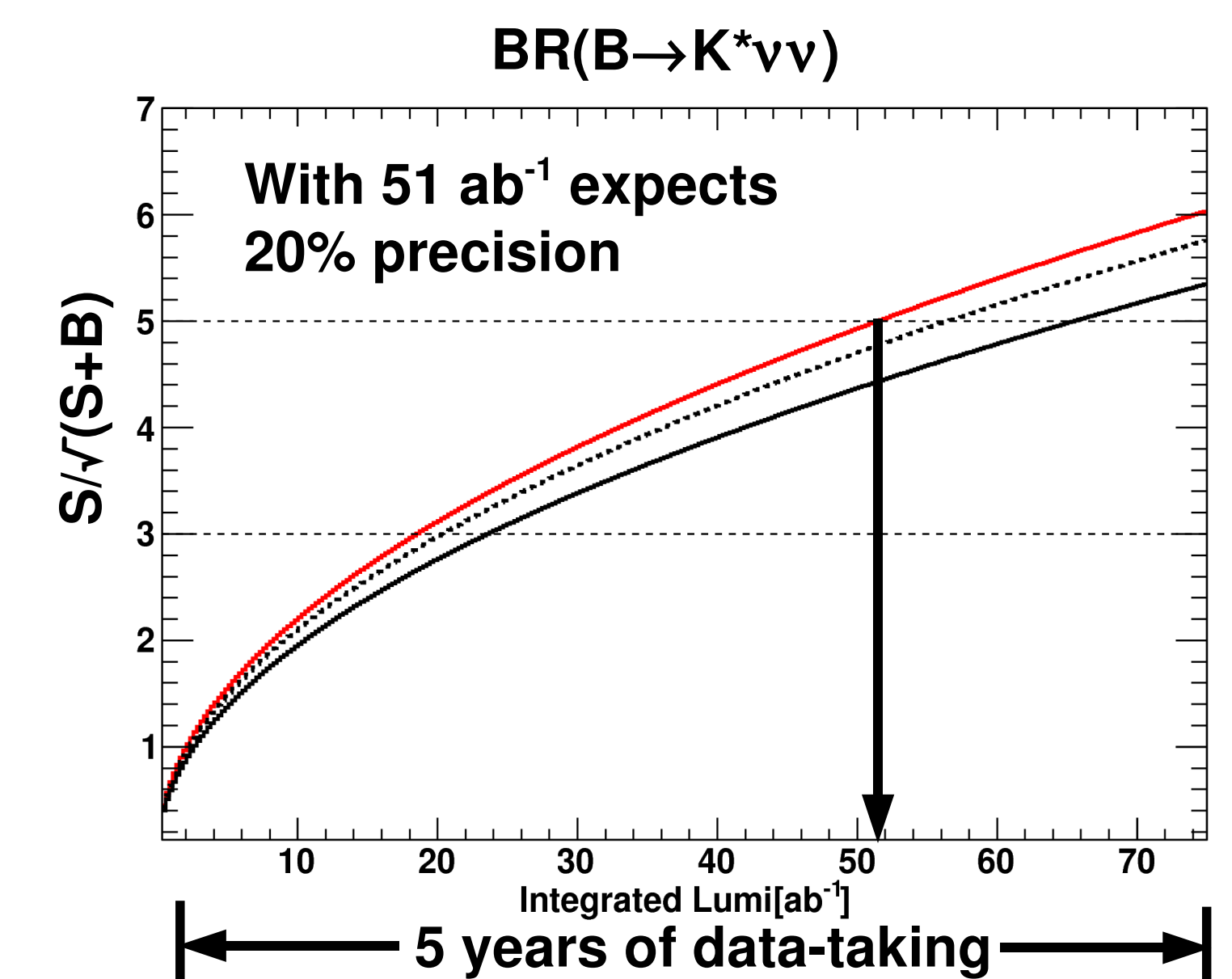
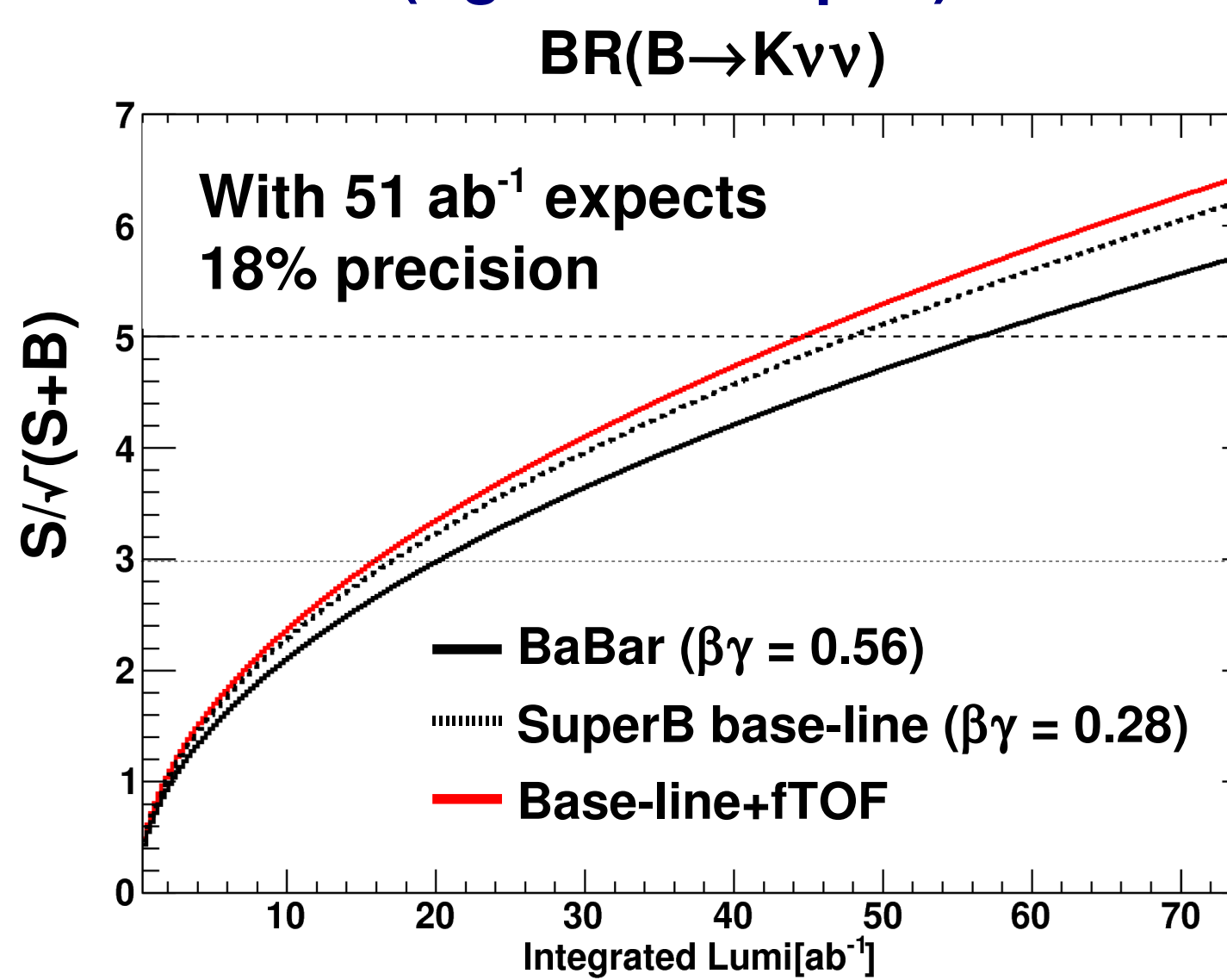
- Use SuperB fast simulation to quantify the potential on rare B decays of different Detector Geometries (DG)

- Current DG's tested:

- BaBar detector with previous boost of $\beta\gamma = 0.56$
- SuperB baseline (lower boost $\beta\gamma = 0.28$ for higher hermeticity)
- SuperB base line + tTOF

- Currently, dominant systematic is MC statistics. Systematics rescaled with luminosity
- Results: tTOF effect is to reduce statistics needed for observation by a factor of $\sim 5\%$. Expects observation (5σ) with a luminosity of

- $B \rightarrow K \nu \bar{\nu}$ (left bottom plot): $\sim 43 \text{ ab}^{-1}$
- $B \rightarrow K^* \nu \bar{\nu}$ (right bottom plot): $\sim 51 \text{ ab}^{-1}$



- Warning: no machine background included

- Right plot: constrain on ϵ - η plane for SuperB expected performances after ~ 3.4 years of data-taking (51 ab^{-1})

- Current status only set very weak constrains on NP

- With a reduced fraction of its expect total dataset SuperB will be able significantly reduce parameter space of NP
- Hopefully NP effects will be spotted after a couple years of running!

