



SISSA
40!



10.1088/1748-0221/13/12/P12018

A **GPU-based** framework for multi-variate analysis in particle physics

25 Jan 2019, 12:10 (CET) @503-1-001, 12'+3'

Xuefeng Ding^{1,2}

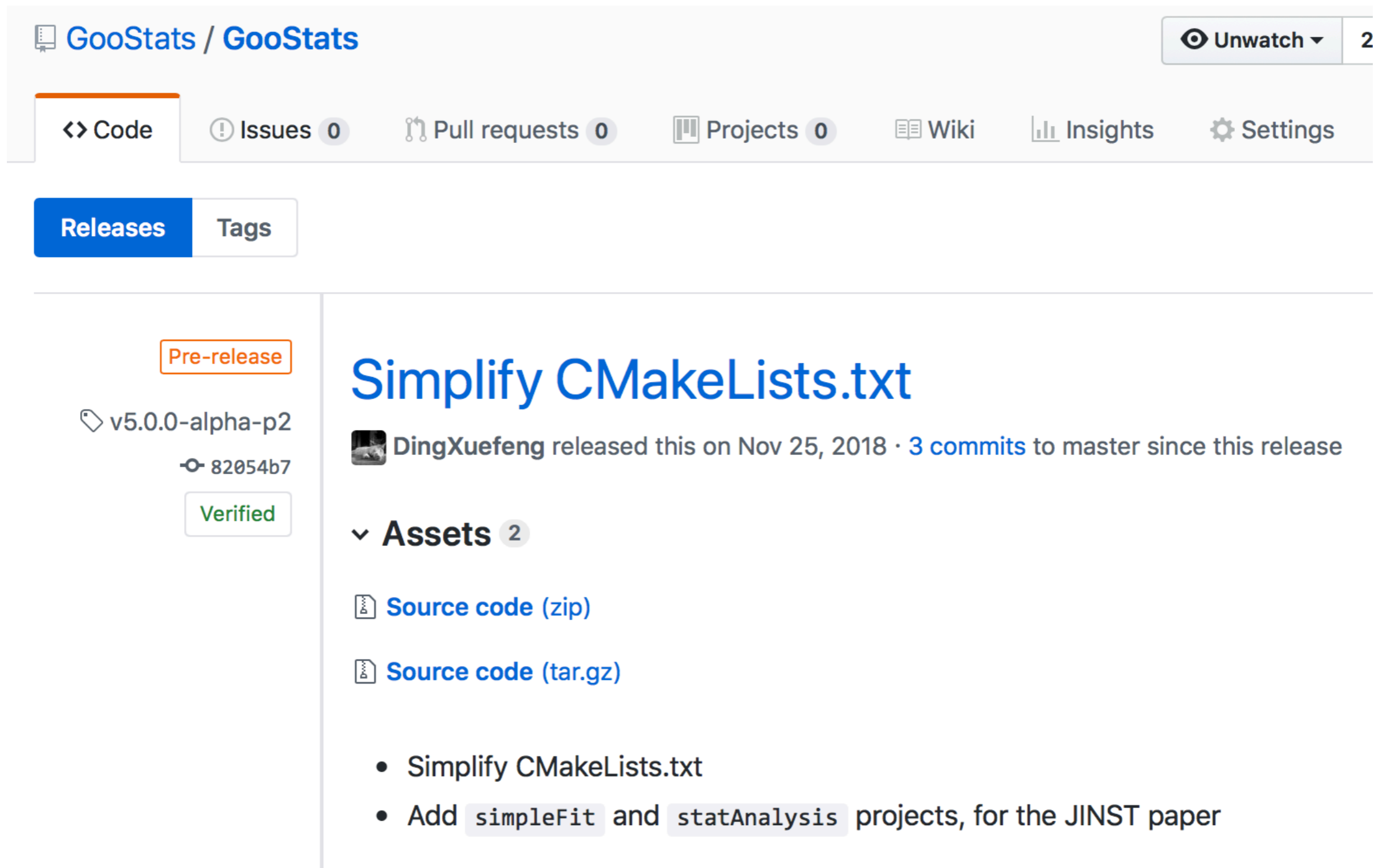
1. Gran Sasso Science Institute, L'Aquila, Italy
2. INFN Sezione di Milano, Milan, Italy

PHYSTAT-nu 2019

@ CERN, Geneva, Switzerland 25 January 2019

- The GooStats framework
- Motivation: Borexino spectrum fit
- Package design
- Systematic uncertainty estimation
- Conclusion

<https://github.com/GooStats/GooStats.git>



GooStats / GooStats Unwatch 2

Code Issues 0 Pull requests 0 Projects 0 Wiki Insights Settings

Releases Tags

Pre-release

v5.0.0-alpha-p2
82054b7
Verified

Simplify CMakeLists.txt

DingXuefeng released this on Nov 25, 2018 · 3 commits to master since this release

Assets 2

- Source code (zip)
- Source code (tar.gz)

- Simplify CMakeLists.txt
- Add `simpleFit` and `statAnalysis` projects, for the JINST paper

What and Why?

- A convenient GPU multivariate analysis framework
- Easy spectrum fitting

```
./fit toyMC.cfg out exposure=500
```

```
FIT PARAMETERS
  gaus = 14.14 #pm 0.21 (day#timeskt)^{-1}
  fbkg = 0.588 #pm 0.016 (day#timeskt)^{-1}
  sdn = 0.991 #pm 0.014
  ^{1}v = 0 [Fixed]
  sigmaT = 0 [Fixed]
  gaus Epeak = 5.000 #pm 0.014 [p] 4.90 #pm 0.50

  chi^2 = 112.7
  chi^2/N-DOF = 1.1739
  p-value = 0.117
  Minimized -2Ln(Likelihood) = 714.53
  Likelihood p-value = 0.540 ± 0.050
```

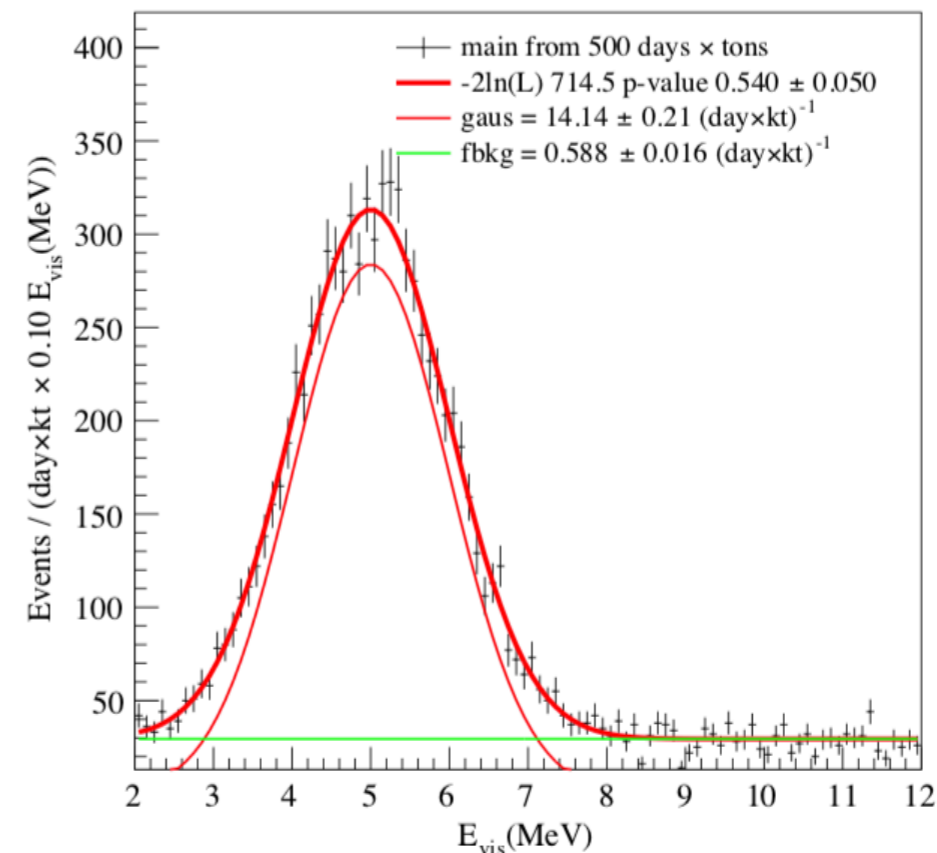


Figure 7. Left: screen shot of fit result summary. Right: produced figure in the file format of pdf.

- Speed up: around 200 times

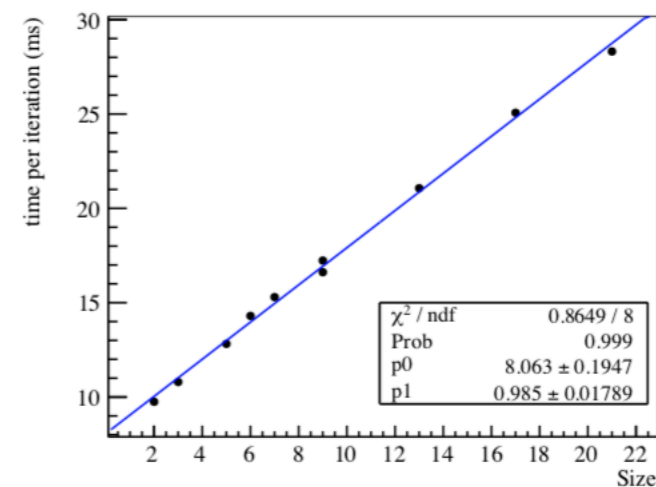
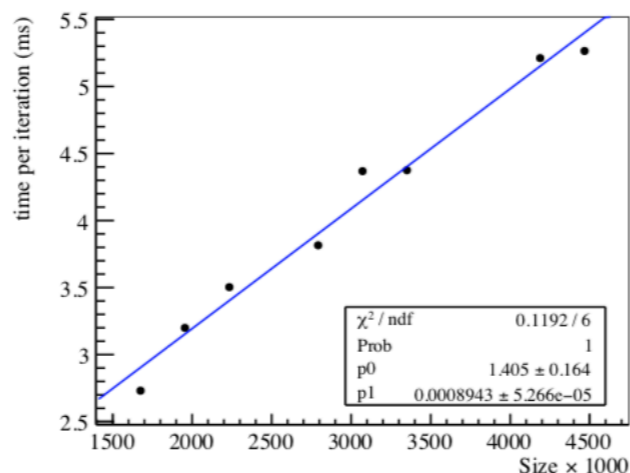
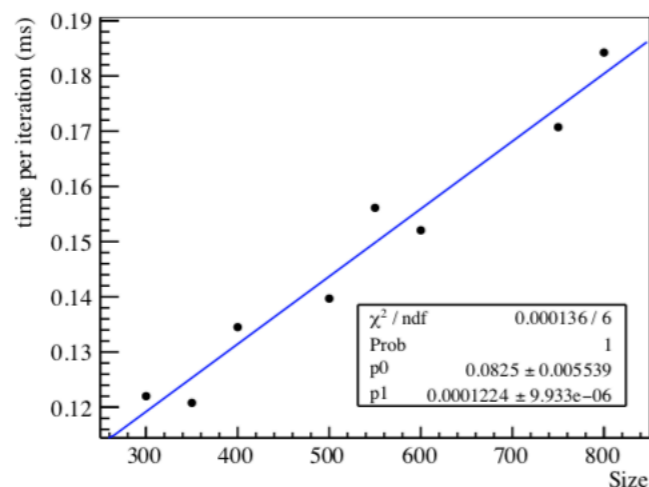
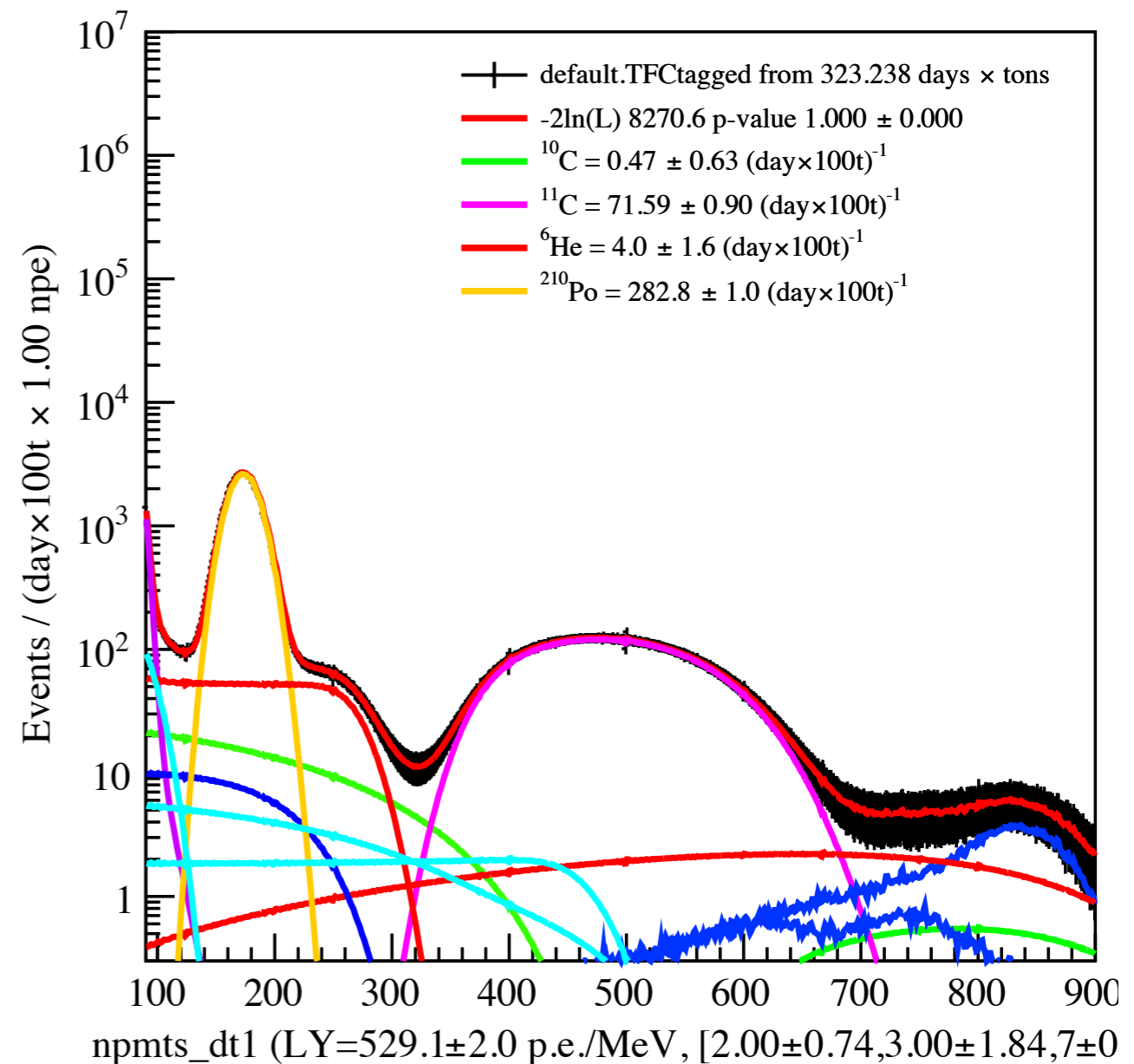
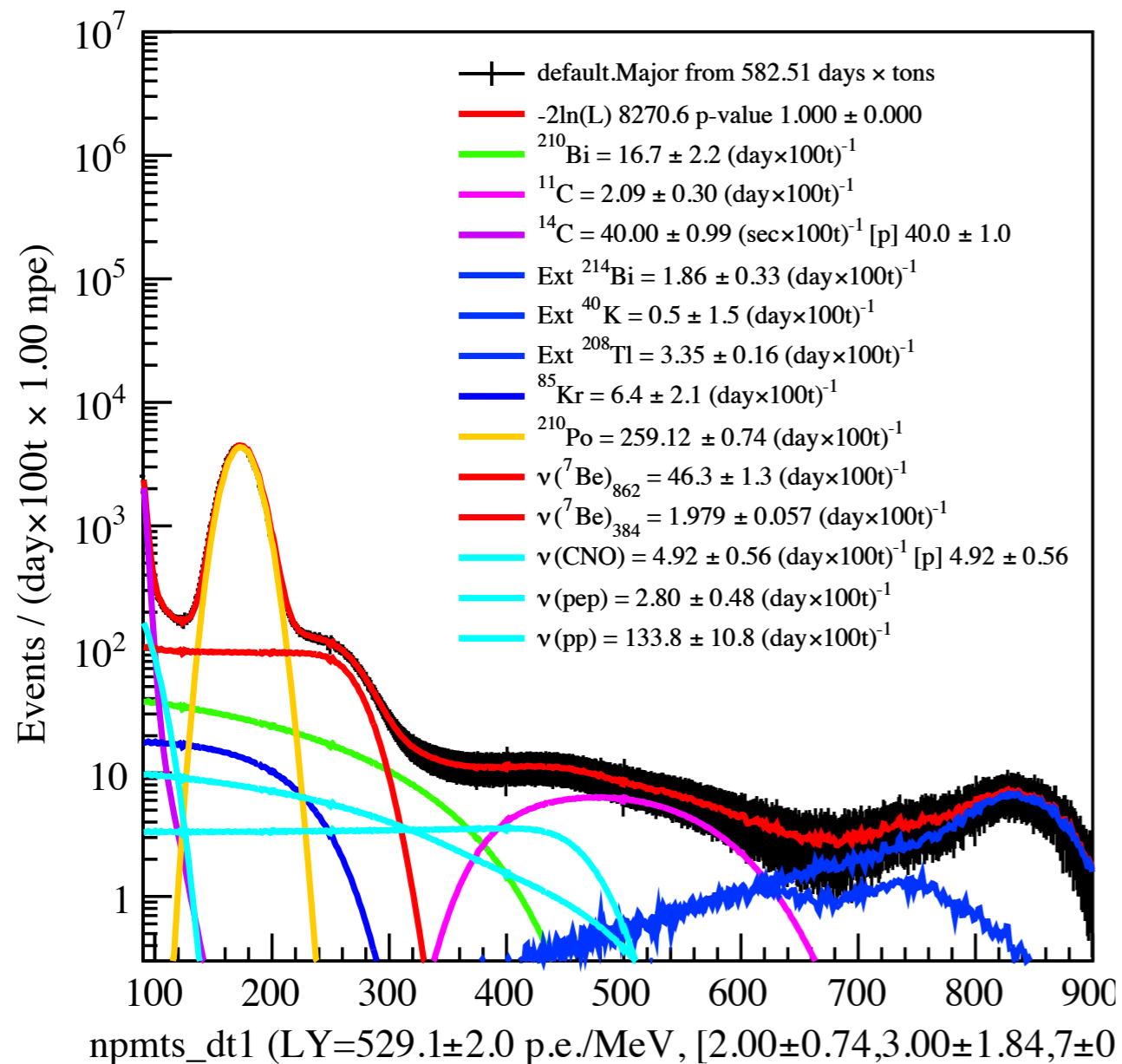


Table 1. Comparison of fitting time between GooStats and original software used by the Borexino collaboration. T_{tot} : total execution time. N : the number of iterations taken to converge in MINUIT. T_{it} : average execution time per iteration. Speed up: $T_{\text{it}}(\text{CPU})/T_{\text{it}}(\text{GPU})$.

Type	CPU			GPU			speed up
	AMD Opteron(TM) Processor 6238			nVidia Tesla K20m			
Size	T_{tot} (s)	N	T_{it} (ms)	T_{tot} (s)	N	T_{it} (ms)	
400	27.6	1128	24.4	0.181	1346	0.135	181
350	29.4	1331	22.1	0.156	1294	0.121	183
300	22.5	1239	18.2	0.243	1995	0.122	149

Motivation: Borexino spectrum fit

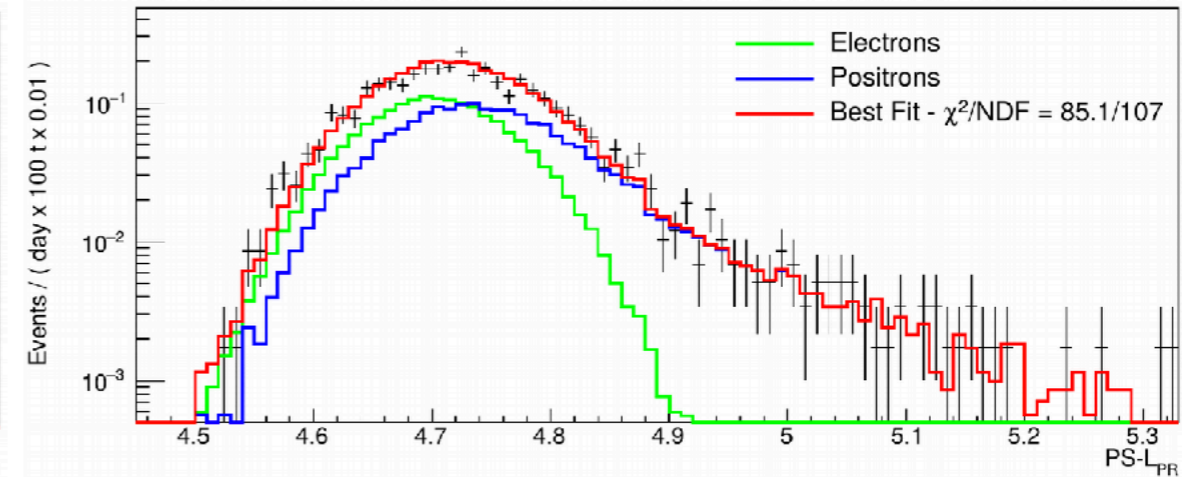
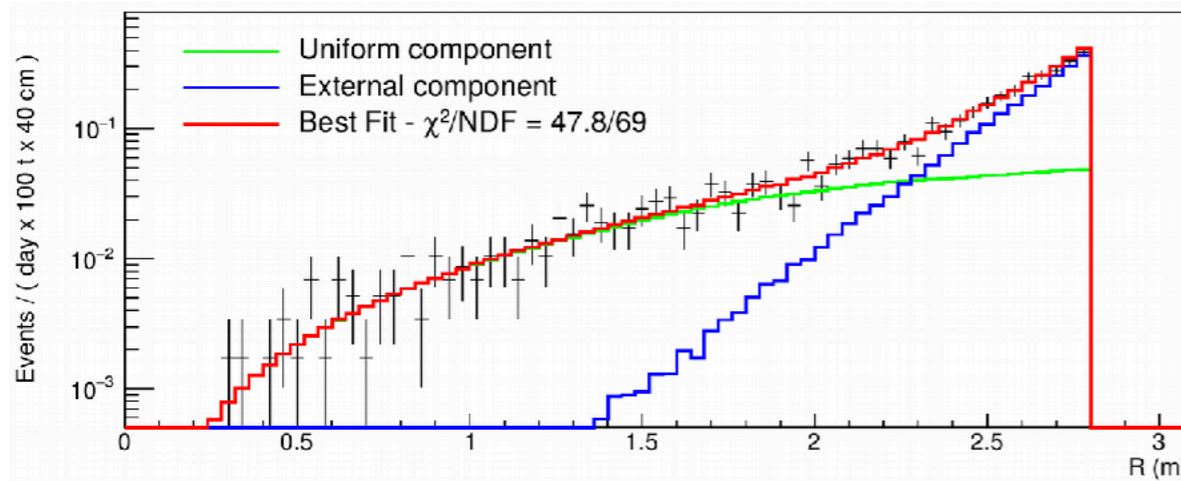
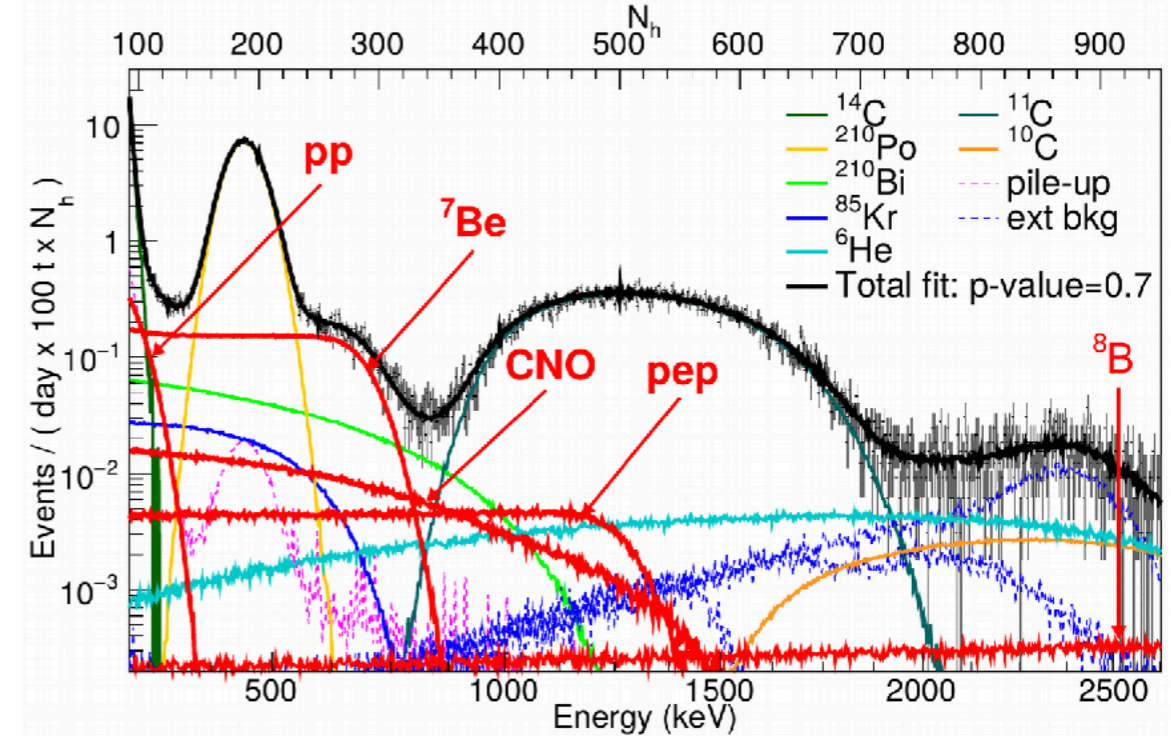
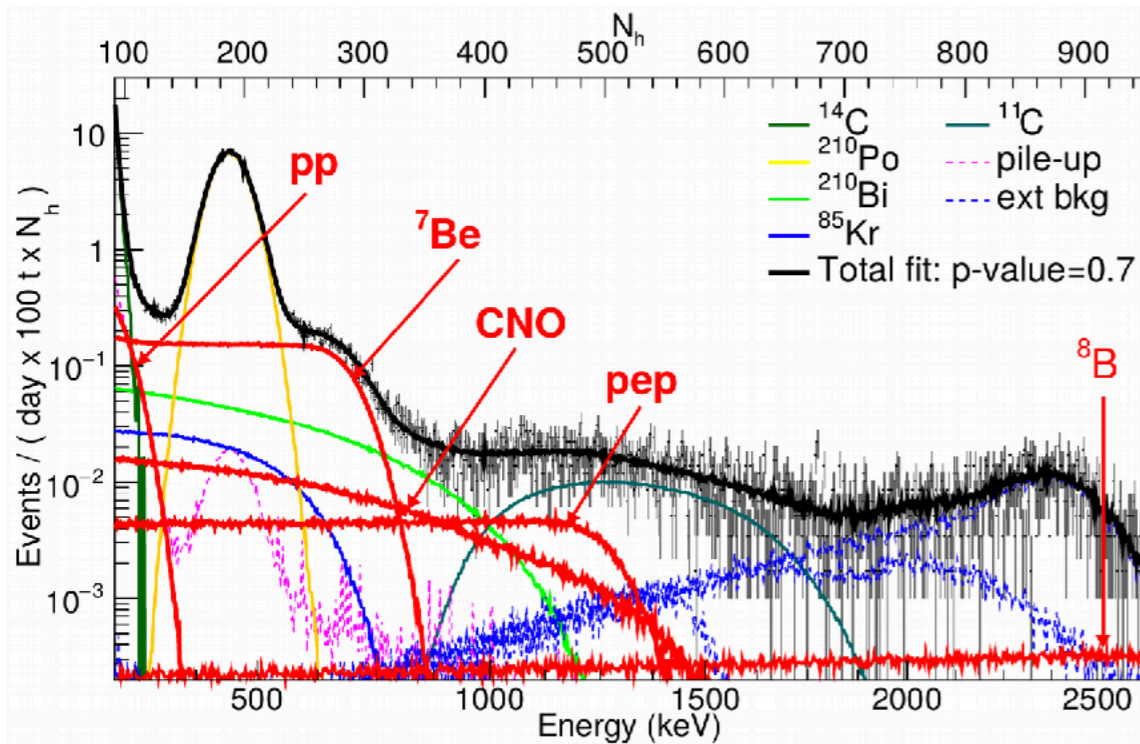
- Motivation: spectral-fitting tool for Borexino analysis



$$M : f(E) \mapsto g(\text{charge}) = \int_0^{E_{\text{end}}} dE \cdot f(E) \cdot \text{RPF} [\text{charge}; \mu(E), \text{var}(\mu)]$$

- Analytical shape of spectrum of mono-energetic events
 - **Momentum based approximation**
 - Match the average (energy scale + non-linearity model)
 - Match the variance (energy resolution model)
 - ... (—> simplified)
 - More: “Mask”, “pile-up” etc...
- We can simplify because
 - Borexino response is simple: small FV in center, low energies => no irregular tail
 - We are not sensitive.. => small systematics
 - **Fit full MC to get the bias introduced in simplification**

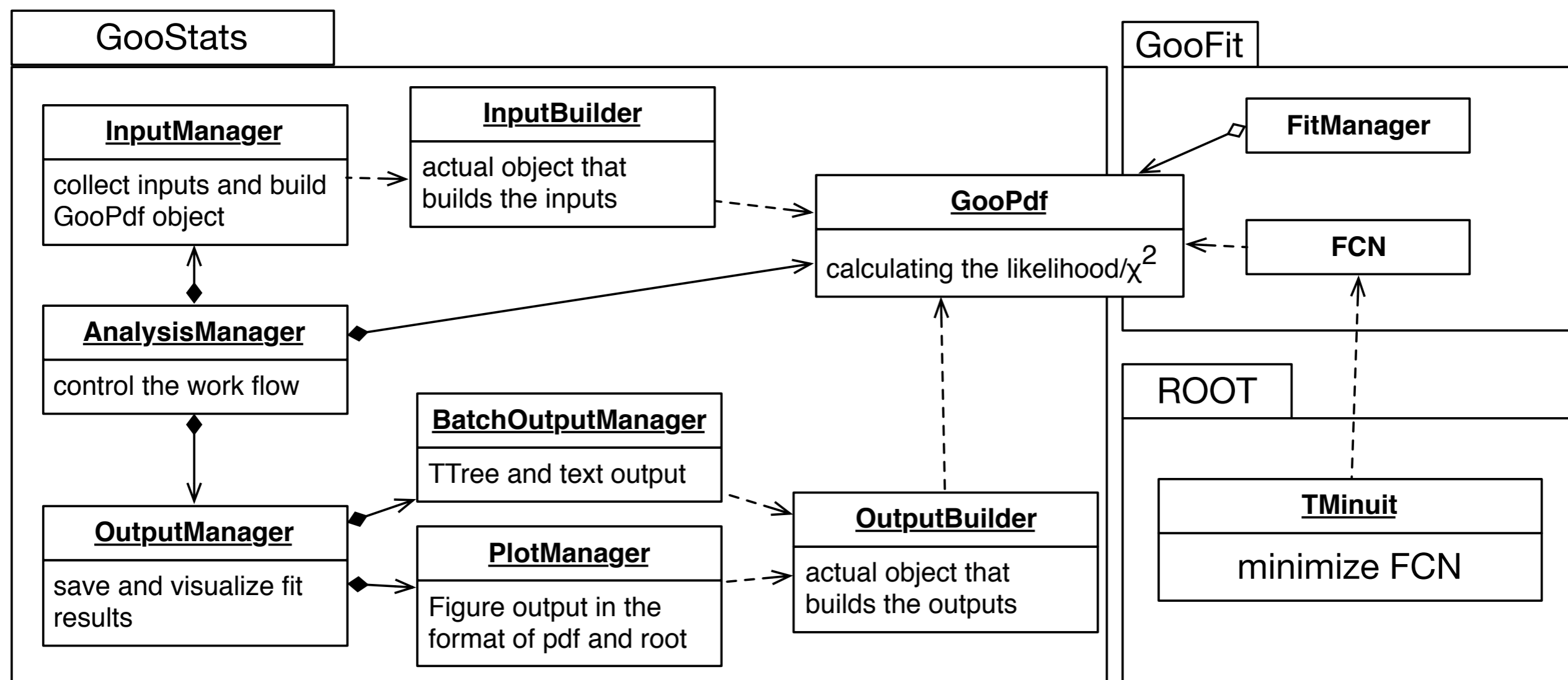
Multivariate Likelihood (Davini method)



$$\mathcal{L}_{MV}(\vec{\theta}) = \mathcal{L}_{\text{TFC-sub}}(\vec{\theta}) \cdot \mathcal{L}_{\text{TFC-tagged}}(\vec{\theta}) \cdot \mathcal{L}_{\text{RD}}(\vec{\theta}) \cdot \mathcal{L}_{\text{PS}}(\vec{\theta})$$

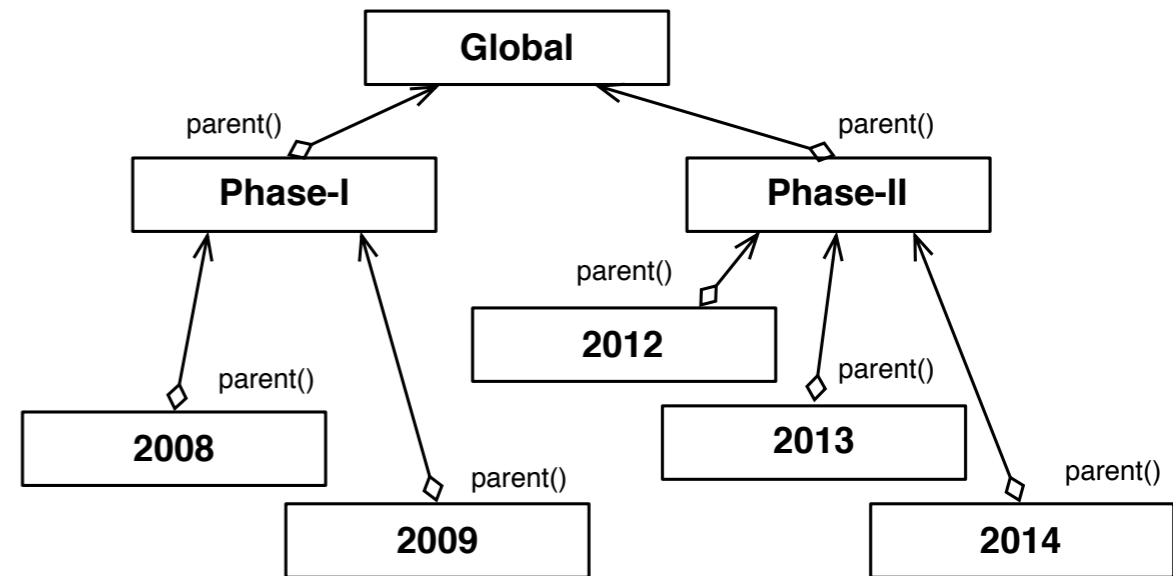
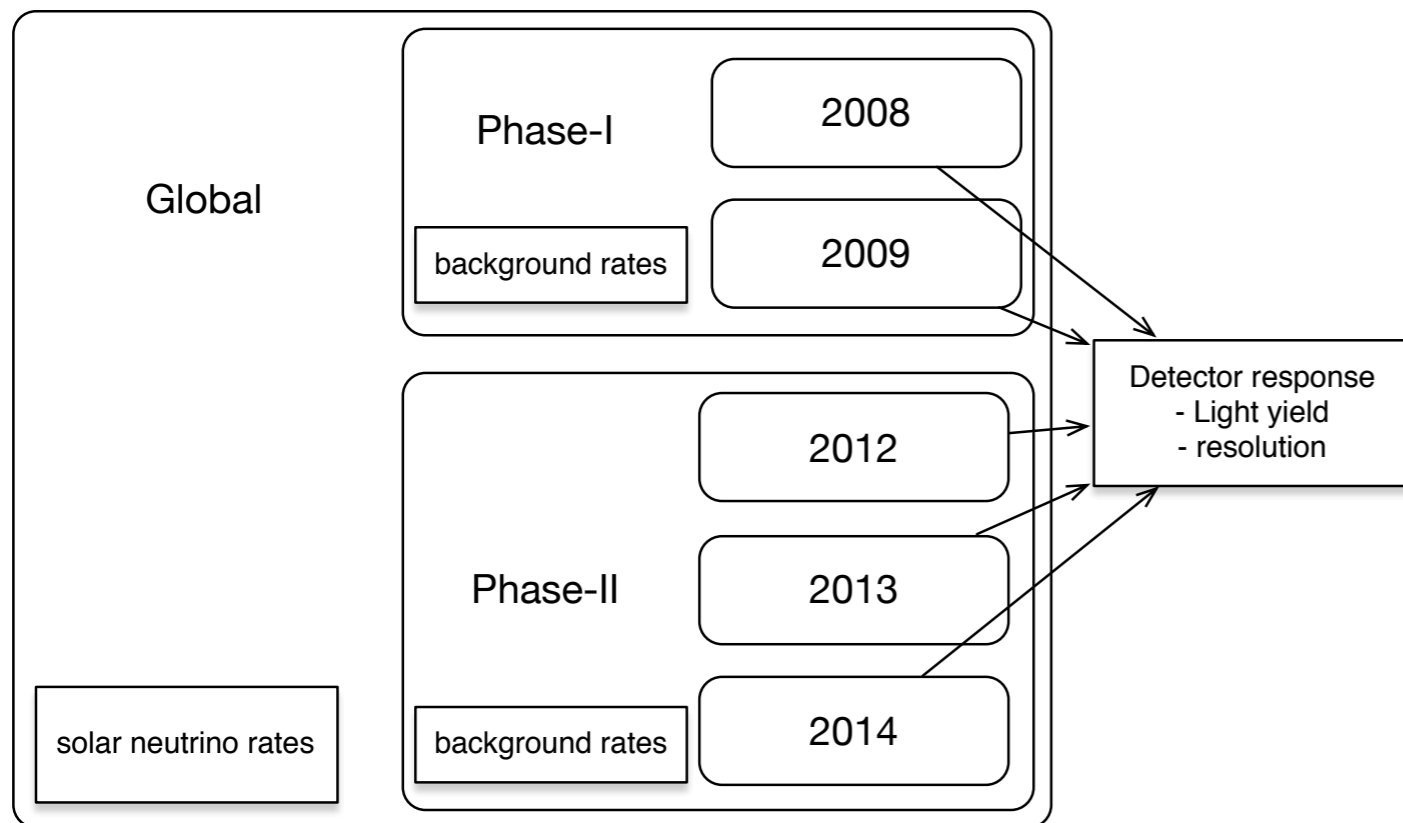
- Scaling factor introduced to remove bias.

- **Middle Layer** between **GooFit** and **User module**



Parameter Synchronization

- Synchronize fit parameters in joint analysis
- Supported by tree type internal data structure



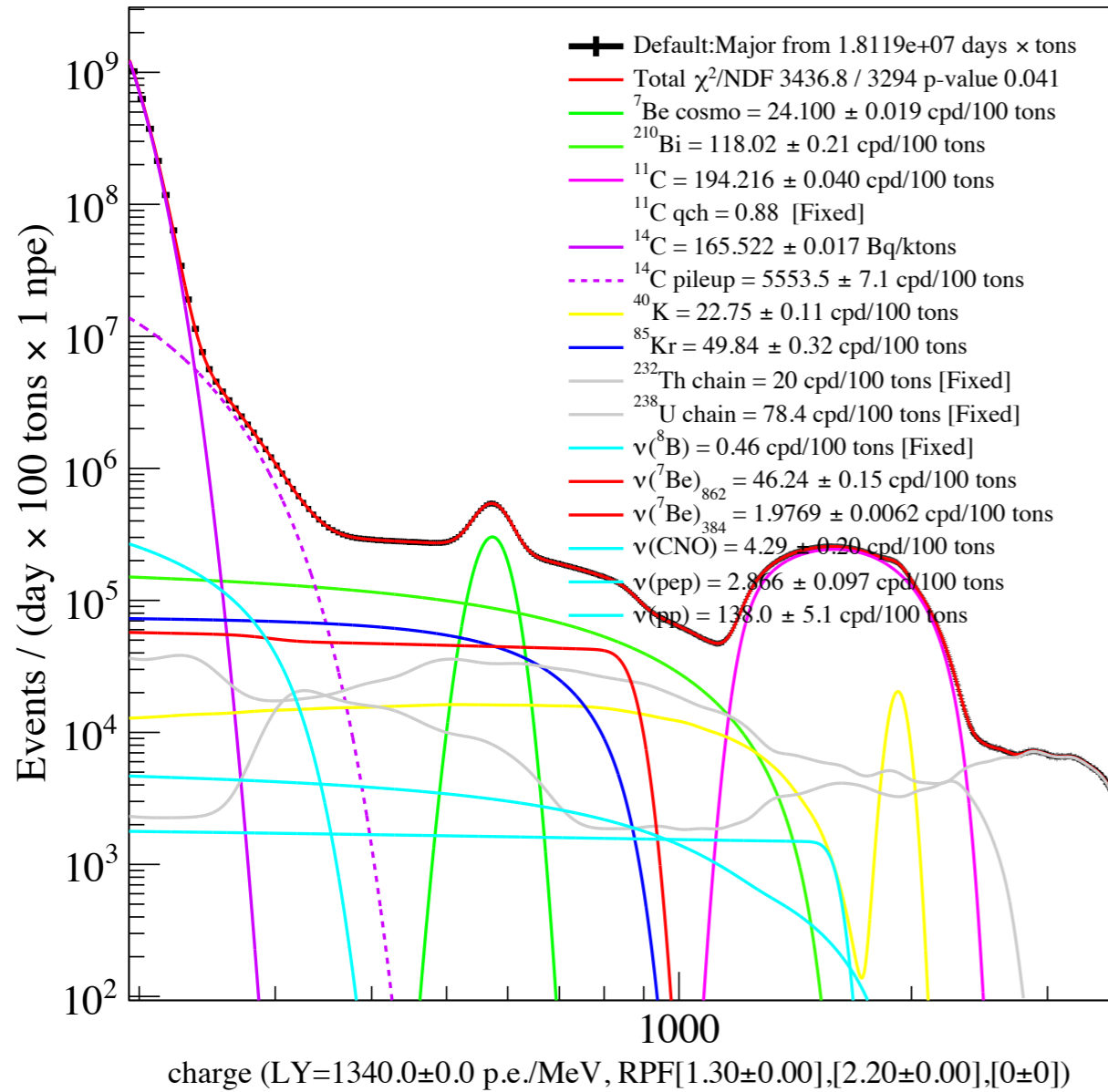
- All fit parameters saved in TTree output automatically
- More quantities can be added as lambda functions in the output builder

```
root [1] .ls
TFile**      test_tree.root
TFile*       test_tree.root
KEY: TTree   fit_result;1   Fit result of GooStats
root [2] fit_result->Show(0)
=====> EVENT:0
default.NReactor = 3.02987
default.NReactor_err = 0.0159146
default.Nbkg      = 1.02405
default.Nbkg_err  = 0.0162459
default.U235      = 0.5
default.U235_err  = 0
default.U238      = 0.2
default.U238_err  = 0
default.Pu239     = 0.1
default.Pu239_err = 0
default.U241      = 0.2
default.U241_err  = 0
default.LY        = 1300
default.LY_err    = 0
default.qc1       = 2.78788
default.qc1_err   = 0
default.qc2       = -0.528003
default.qc2_err   = 0
default.v1        = 0.3
default.v1_err    = 0
default.vT        = 5
default.vT_err    = 0
default.Reactor_dEvis = 1078.28
default.Reactor_dEvis_err = 13.2291
chi2              = 390.448
NDF               = 397
likelihood        = 1883.96
```

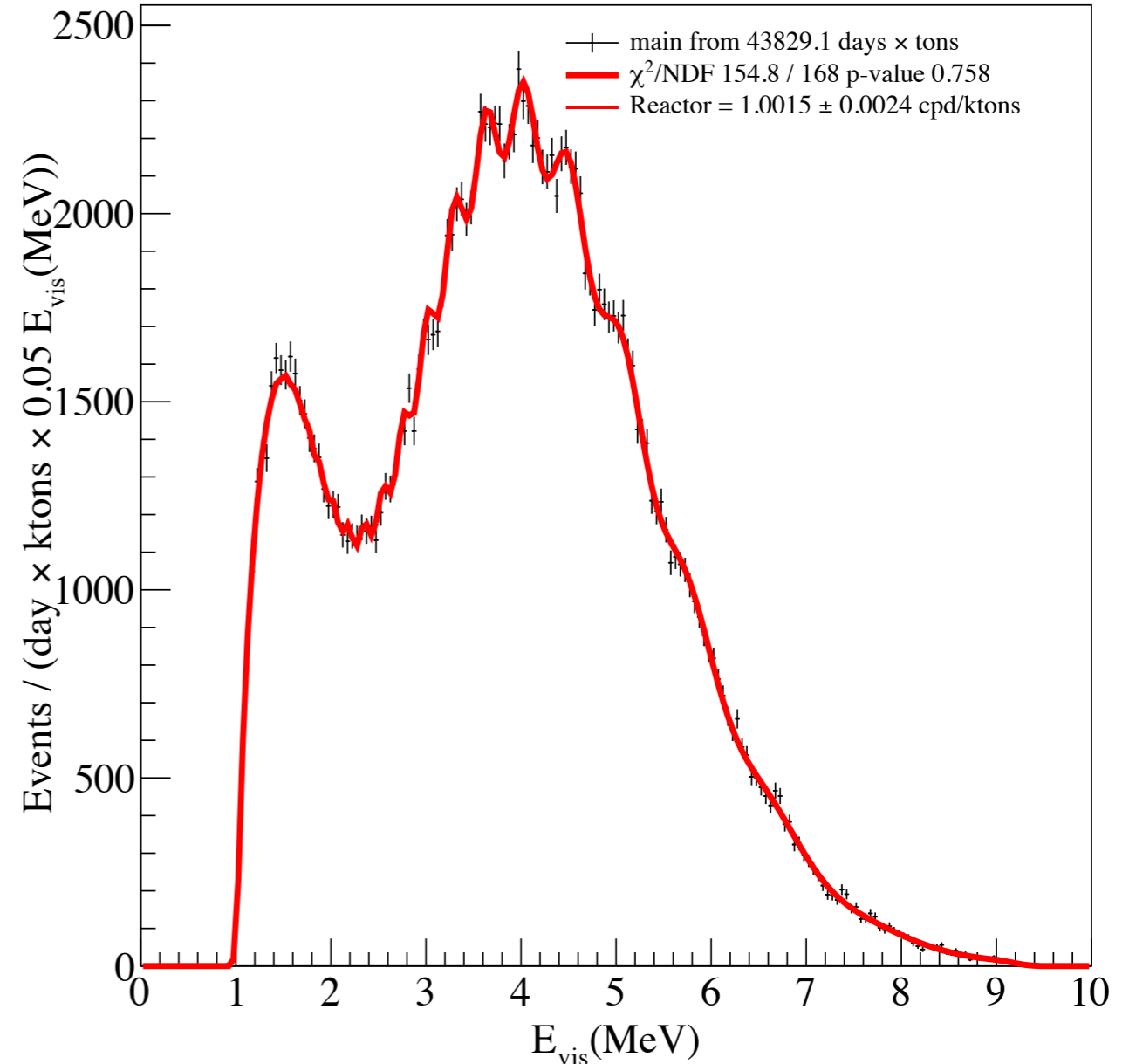
- Adding tasks of statistical analysis by registering new modules inheriting abstract classes
- A few pre-installed modules: ScanPar, DiscoveryTest..

```
PrepareData *data = new PrepareData();  
SimpleFit *fit = new SimpleFit();  
DiscoveryTest *discovery = new DiscoveryTest();  
  
ana->registerModule(inputManager);  
ana->registerModule(data);  
ana->registerModule(fit);  
ana->registerModule(discovery);  
ana->registerModule(outManager);
```

Solar Neutrinos



JUNO style



Two ways of estimating σ_{sys}

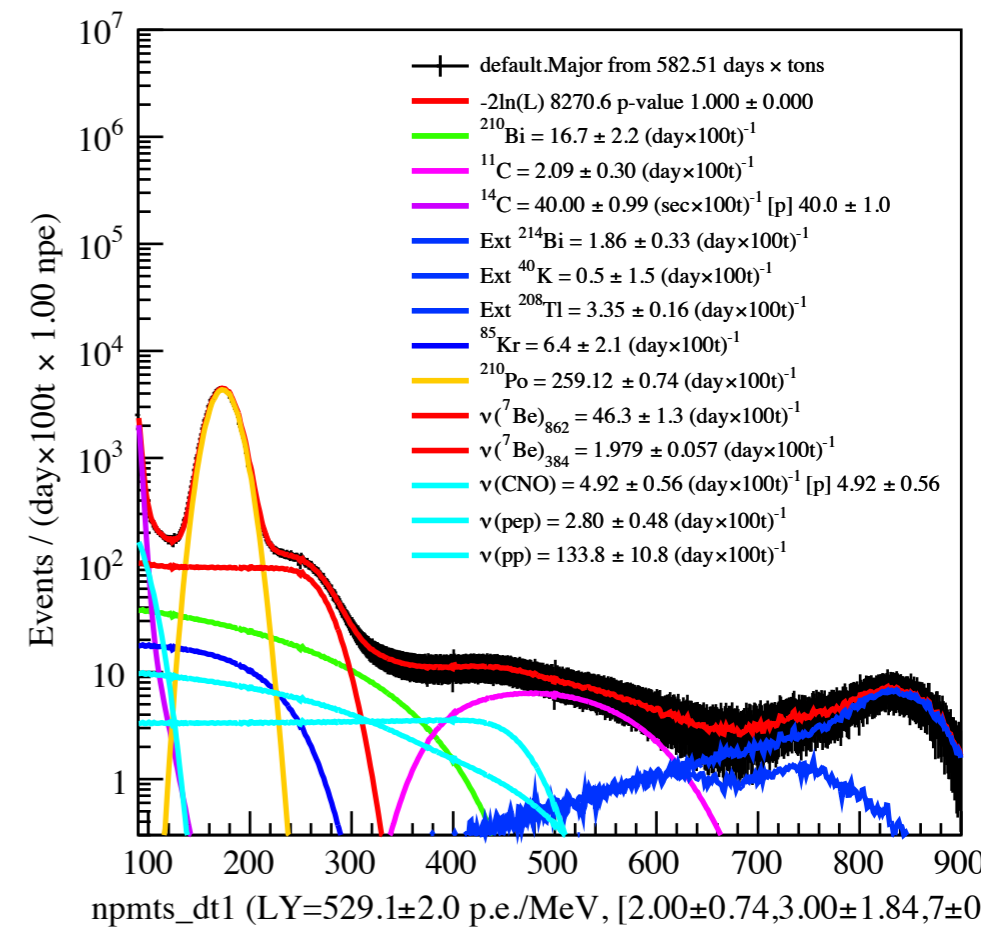


- create an ensemble of models according to experiment precision
- Fit the data with varying models through their coordinates
- Generate pseudo-experiments with varying models and take the width of distribution of best-fit

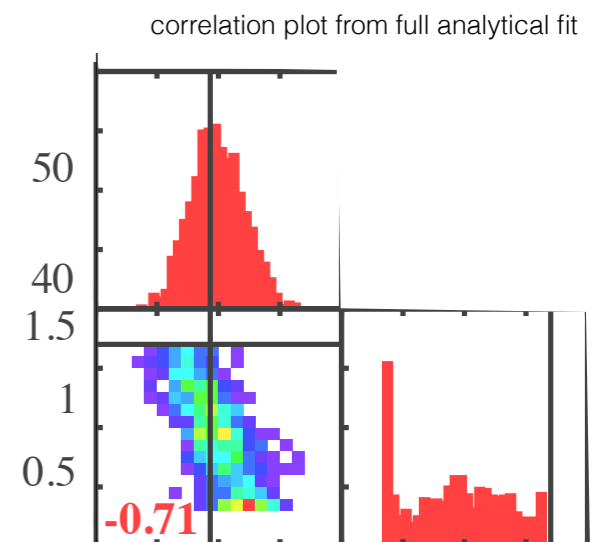
- **Full MC** during each iteration of the fit, vary **kb / absorption length spectrum etc.** **re-simulate and produce new pdf on the fly** → when one day computer is fast enough
 - ~200, 000 CPU x years per fit
- **Semi-analytical:** analytical non-linearity model + response Matrix
 - ~30 minutes per fit
- **Full analytical**
 - ~2 hours per fit

Problem with scaling response matrix

- Systematic uncertainties of LY can be included by scaling the response matrix
- This could be dangerous if the interested parameters is correlated with the resolution parameters
- When LY is scaled, the resolution should also be changed



$\nu(^7\text{Be})$
 48.5 (-2.1 +2.3)
 inj 48.00 ± 0.07
 $\beta\text{-RPF}(0)$
 0.9 (-0.4 +0.3)
 inj 1.49 ± 0.01

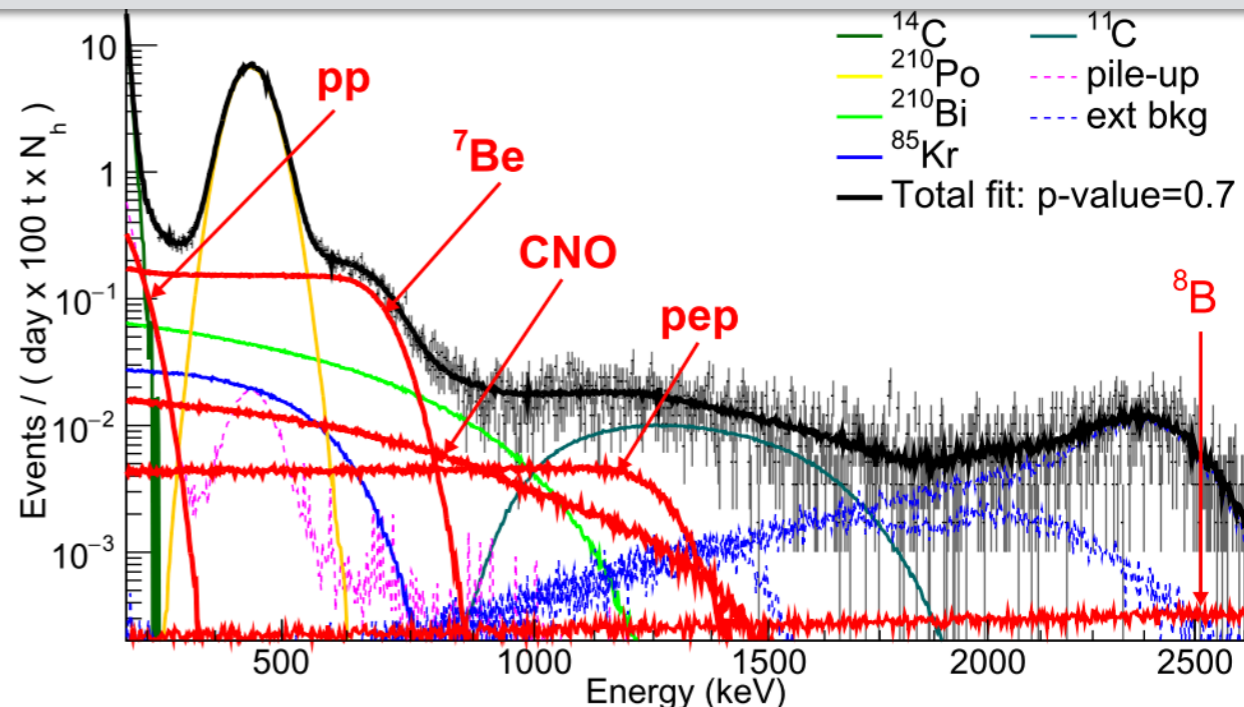


Correlation with NL/resolution

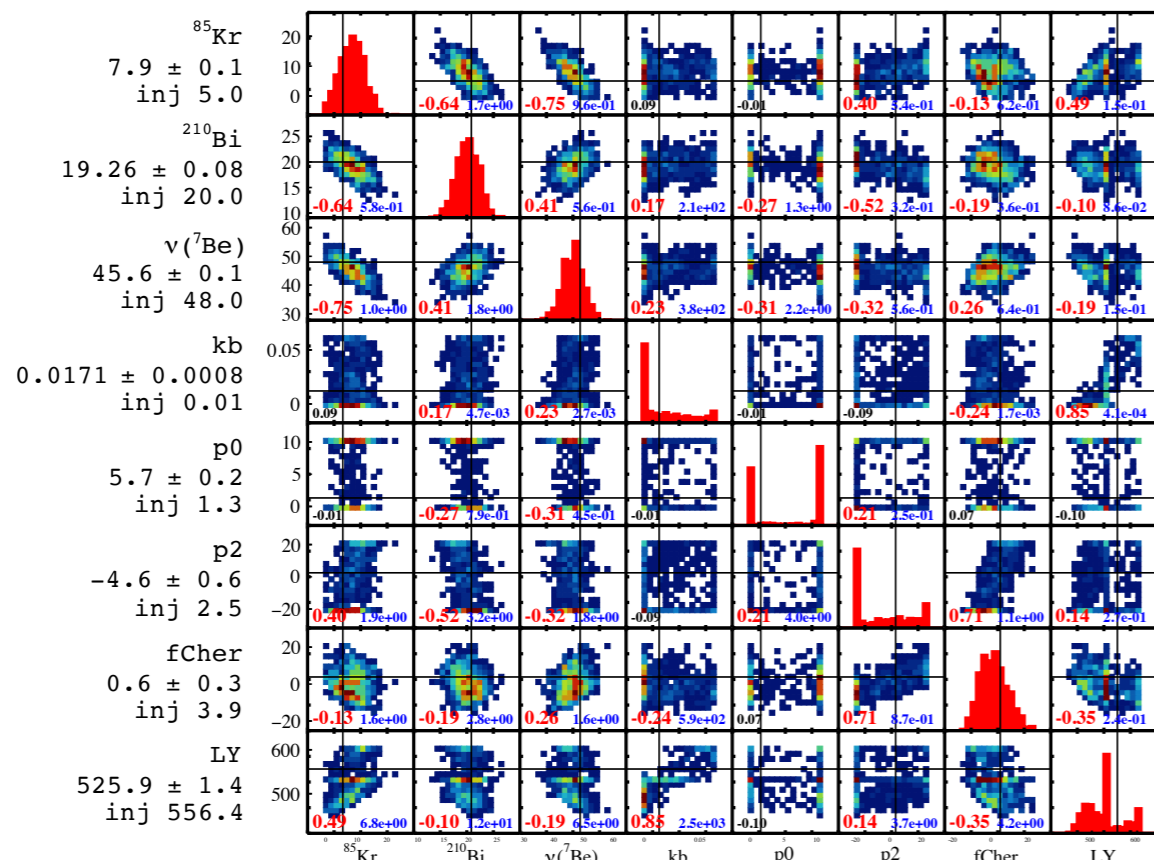


- Using full analytical response function, we can see the correlation with detector responses

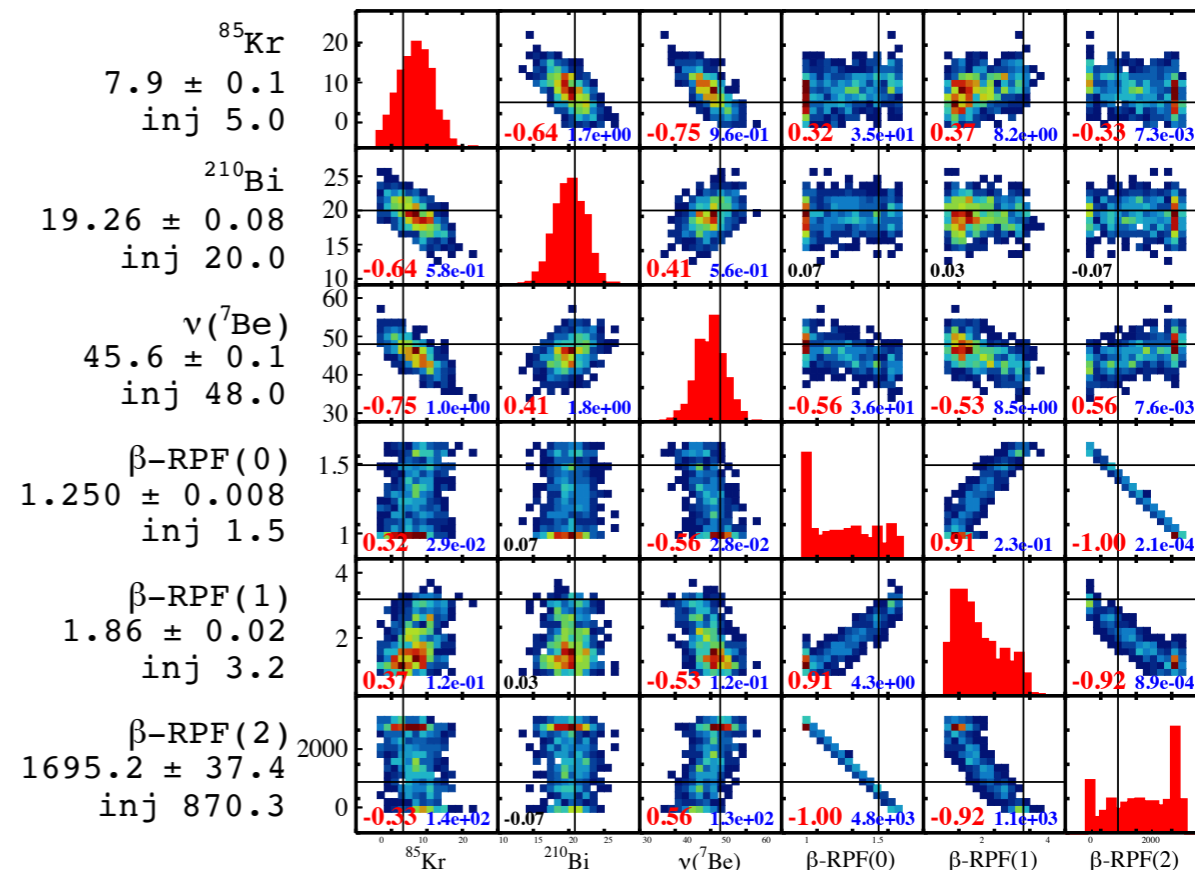
[1] M. Agostini et al., "First Simultaneous Precision Spectroscopy of ^85Kr , ^{210}Bi , and ^{210}Po Solar Neutrinos with Borexino Phase-II," pp. 1-8, Jul. 2017.



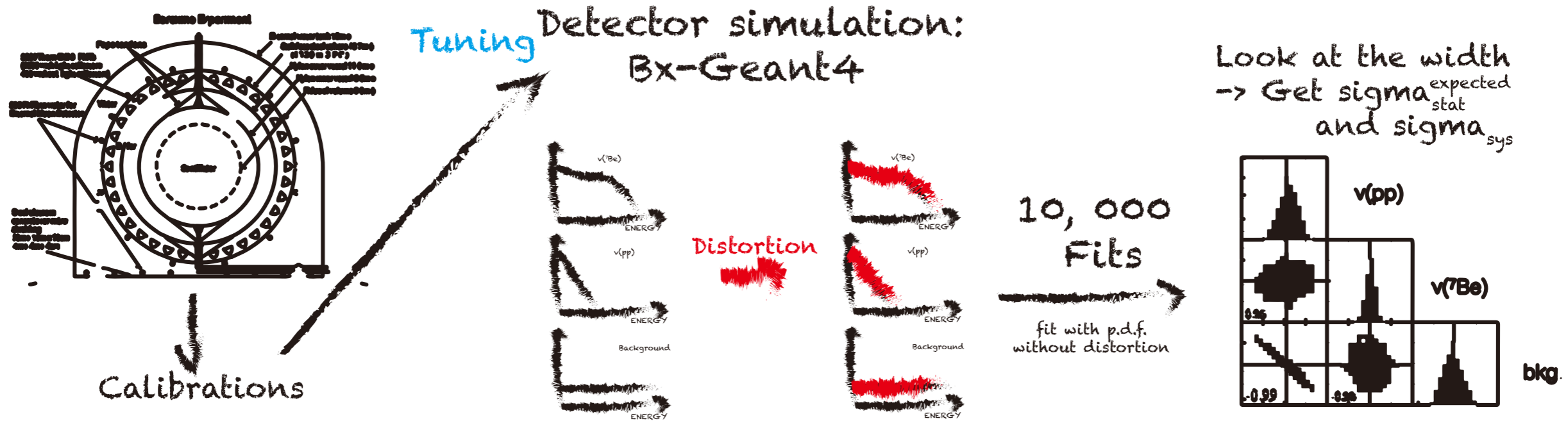
correlation with NL



correlation with res



Method 2: Monte Carlo method



- pseudo-experiment spectra without distortion —> **statistical sensitivity**
- pseudo-experiment spectra with distortion —> **statistical systematic uncertainty**

- GooStats, a GPU based multivariate analysis framework is introduced.
- Statistical analysis module is easy to be implemented as needed. TTree/figure output provided.
- Full analytical response function has advantage that it treated the NL and resolution in a coherent way when evaluating the systematic uncertainties
- The systematic uncertainties can also be evaluated using Monte Carlo method by looking at change of width of best fit with/without distortion