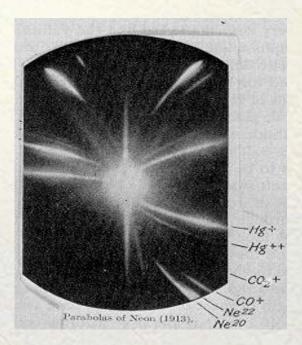
# The Morinaga-Ikeda Nucleus

Neon (New in Greek) Challenge to Mean Field, Cluster and Collective Models

<sup>20</sup>Ne is the fifth most abundant nucleus in the universe...and yet unknown



First discovery of isotopes of stable atoms by J. J. Thompson (1913)

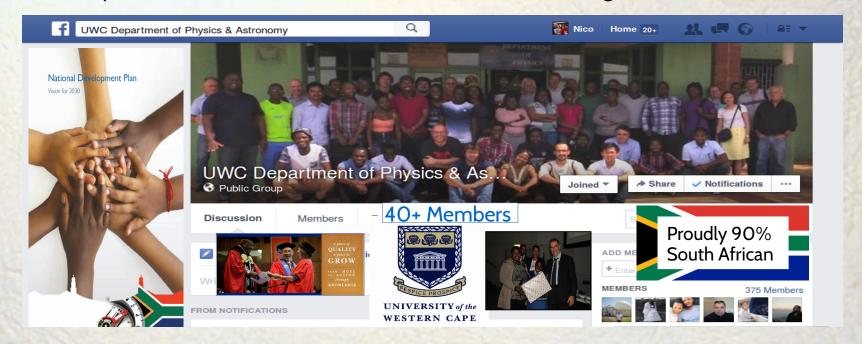


Senamile Masango, Craig Mehl and collaborators
ISOLDE Workshop Dec 2018



# 2018-2019 Fundamental Nuclear Physics & Applications @ UWC

Experiments @ UWC, iThemba LABS, HIE-SOLDE, Washington, MLL & TRIUMF



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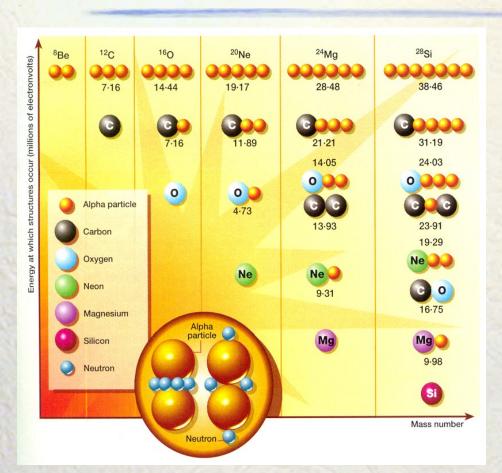
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Postdocs: Kushal Kapoor, Rakesh Dube, Joash Ongori

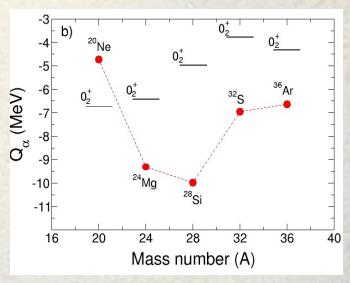
https://nuclear.uwc.ac.za

# The Morinaga-Ikeda Nucleus

 $\alpha$  clusters can emerge gradually as the internal energy of the nucleus increases and are fully realized at the  $\alpha$  threshold



``20Ne is probably the best example of a clustered nucleus" (Martin Freer)



Only self-conjugate nucleus with  $\alpha$  threshold below second  $0^+$  excitation

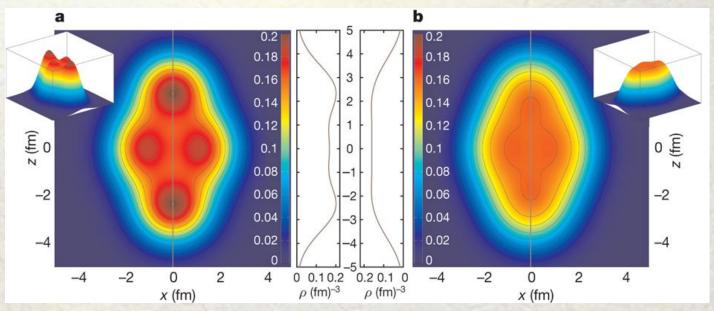
H. Morinaga, Phys. Rev. 101, 254 (1956) Ikeda, K., N. Takigawa, and H. Horiuchi, Prog. Theor. Phys. Suppl. E68, 464 (1968)

#### How atomic nuclei cluster

# Ground state of <sup>20</sup>Ne may already present a cluster structure

#### Fully microscopic description based on energy-density functionals (EDFs)

Ebran, Khan, Nikšić & Vretenar, Nature 487, 341 (2012)



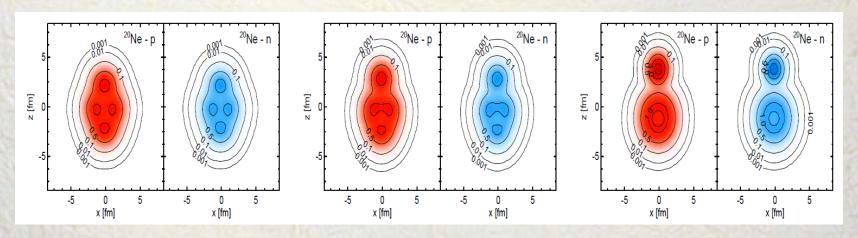
Self-consistent ground-state densities of <sup>20</sup>Ne: a) relativistic DD-ME2 and b) Skyrme SLy4

Assumption: The mixing with  $\alpha$  clusters congurations may have a pronounced effect on the shape of mean-field states at lower excitation energies

# The Morinaga-Ikeda Nucleus

The mean-field state has a sizable cluster character (agrees with largest EO strength in nature)

- $Q_s(2^+) = -0.126$  eb standard mean-field calculations (HFB)
- $Q_s(2^+) = -0.153$  eb including cluster configurations
- $Q_s(2^+) = -0.23(3)$  eb from reorientation-effect measurements

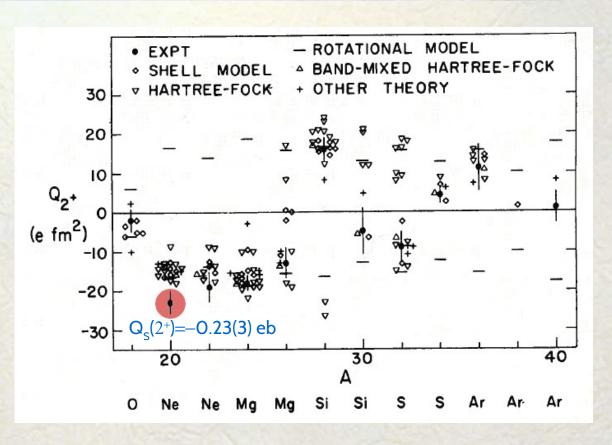


 $\alpha$  clusters can emerge gradually as the internal energy of the nucleus is increased

Thanks to Thomas Neff, Fermionic-Molecular Dynamic Calculations of <sup>20</sup>Ne

# <sup>20</sup>Ne: a 45 year-old puzzle

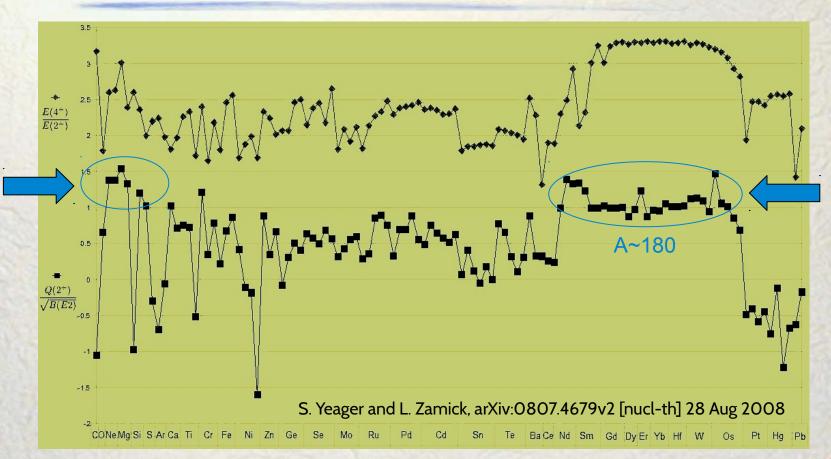
ALL calculations (mean field (HFB, shell model), rotor, cluster models)
fail to describe the spectroscopic quadrupole moment of the first 2+ state in 20Ne



R.H. Spear, Phys. Rep. 73, 369 (1981)

# The quadrupole moment ratio: 1 for rotors, 0 for vibrators

$$\frac{Q_0(S)}{Q_0(B)} = -\frac{7}{2}\sqrt{\frac{5}{16\pi}}\frac{Q(2^+)}{\sqrt{B(E2)}} = -1.1038705\frac{Q(2^+)}{\sqrt{B(E2)}}$$



<sup>20</sup>Ne is one of the few known cases in the nuclear chart where  $Q_S(2^+)_{B(E2)} = -0.165(5)$  eb and  $Q_S(2^+)_{RE} = -0.23(3)$  eb disagree at the  $2\sigma$  confidence level

# Spectroscopic Quadrupole Moment of the first 2+ state in 20Ne

"A new measurement is desirable." R. H. Spear (1981)

#### 4.2. The nucleus <sup>20</sup>Ne (table 4.2)

There are only 3 RECE measurements, and none of them is recent. The paucity of data is largely due to the difficulty of producing negative Ne ions and the consequent unavailability of Ne beams from tandem accelerators. Two of the measurements [20, 24] included in their analysis data obtained at disturbingly high bombarding energies (corresponding to distances of closest approach as small as approximately 4 fm in each case), and none of them demonstrated convincingly that their maximum bombarding energies were safe. In addition, the analysis of Nakai et al. required assumptions about the electric quadrupole moments of the target nuclei, and the error quoted for the result of Schwalm et al. is unusually large. A new measurement is desirable. The adopted value is obtained by simply weighting the RECE results in inverse proportion to the squares of their stated errors.

R.H. Spear, Phys. Rep. 73, 369 (1981)

# Spectroscopic Quadrupole Moment of the first 2+ state in 20 Ne

"A new measurement is desirable." R. H. Spear (1981)

# Separation between nuclear surfaces, S<sub>min</sub> > 6.5 fm

# Experimental determinations of $Q_{2^+}$ for <sup>20</sup>Ne ( $E_x = 1.634 \,\text{MeV}$ )

Author [ref.]	Year	$Q_{2^+}$ (e fm <sup>2</sup> )	Projectile/Target	Method	Details	s <sub>min</sub> (fm)
Nakai [20]	1970	-24±3	<sup>20</sup> Ne/ <sup>120</sup> Sn, <sup>130</sup> Te, <sup>148</sup> Sm	RECE	$P\gamma C; \theta = 90^{\circ}, 160^{\circ}$	3.8
Horikawa [21]	1971	$-17 \pm 1$	e/Ne	(e, e')	Assumes rotational model	<u> </u>
Schwalm [22]	1972	$-23 \pm 8$	<sup>32</sup> S/ <sup>20</sup> Ne	RECE	DByLS(0°); gas target	7.1
Rebel [23]	1972	$-13 \pm 1$	⁴He/ <sup>20</sup> Ne	CCEIS	$E = 104 \mathrm{MeV}$ ; assumes rotational model	
Olsen [24]	1974	$-20 \pm 5$	<sup>20</sup> Ne/Pt, Au	RECE	γAD	4.2 ,
Gross [25]	1978	$-27 \pm 3$	<sup>20</sup> Ne/ <sup>208</sup> Pb	CCEIS	E = 131 MeV; assumes rotational model	<b>O</b> .

Adopted value:  $-23 \pm 3 e$  fm<sup>2</sup>

R.H. Spear, Phys. Rep. 73, 369 (1981)

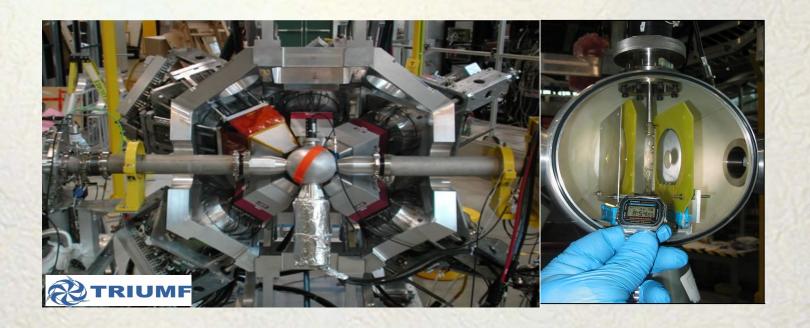
<sup>20</sup>Ne: a 45 year-old puzzle

Forget the N<sub>p</sub>N<sub>n</sub> scheme, <sup>20</sup>Ne is the most deformed nucleus in the *sd* shell and beyond!

	$Q(2_1^+)$ [e fm <sup>2</sup> ]	$B(E2)$ $[e^2 \text{ fm}^4]$	$ Q_0(S)/Q_0(B) $	$E(4_1^+)/E(2_1^+)$
<sup>20</sup> Ne	(-23 (3))	340 (30)	1.38 (19)	2.600
<sup>22</sup> Ne	-19(4)	230 (10)	1.38 (28)	2.634
$^{24}$ Mg	-16.0(8)	432 (12)	0.85 (4)	3.012
$^{28}$ Si	+16(3)	326 (12)	0.98 (18)	2.596
$^{32}$ S	-15(2)	300 (13)	0.96 (13)	2.000
$^{36}Ar$	+11(6)	340 (40)	0.66 (36)	2.240
40Ar	+1 (4)	330 (39)	0.06 (24)	1.980

S. J. Q. Robinson, A. Escuderos, L. Zamick, P. von Neumann-Cosel, A. Richter and R. W. Fearick, Phys. Rev. C 73, 037306 (2006)

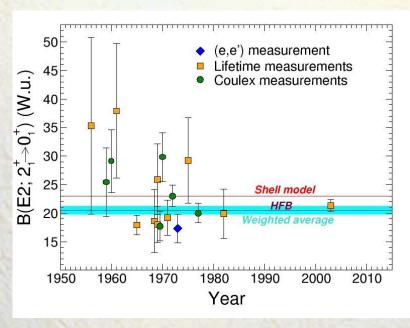
<sup>110</sup>Pd(<sup>20</sup>Ne, <sup>20</sup>Ne\*)<sup>110</sup>Pd\* @ 64.7 MeV (extremely safe with S<sub>min</sub> = 9.3 fm @ 49.1°)



Particle-gamma coincidence measurement

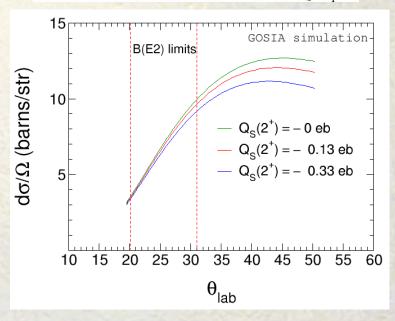
6 TIGRESS detectos + silidon S3 detector covering forward angles [20.9°, 49.1°]

<sup>110</sup>Pd(<sup>20</sup>Ne, <sup>20</sup>Ne\*)<sup>110</sup>Pd\* @ 64.7 MeV (extremely safe with S<sub>min</sub> = 9.3 fm @ 49.1°)



Previous measurements

$$\sigma_{E2} = \sigma_R[k_1(\vartheta, \xi)B(E2)(1 + k_2(\vartheta, \xi)Q_S(2_1^+))].$$

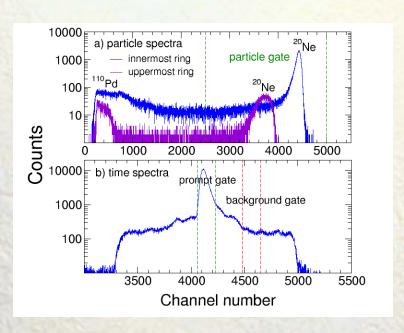


New Coulex measurement < 3% second-order contribution

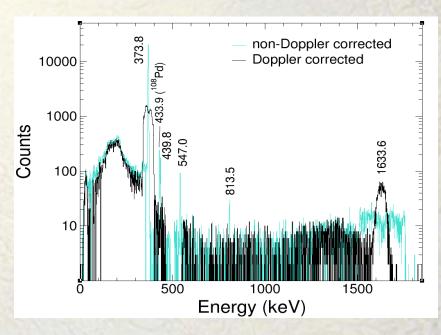
Particle-gamma coincidence measurement

6 TIGRESS detectos + silidon S3 detector covering forward angles [20.9°, 49.1°]

<sup>110</sup>Pd(<sup>20</sup>Ne, <sup>20</sup>Ne\*) <sup>110</sup>Pd\* @ 64.7 MeV



Particle + time spectra

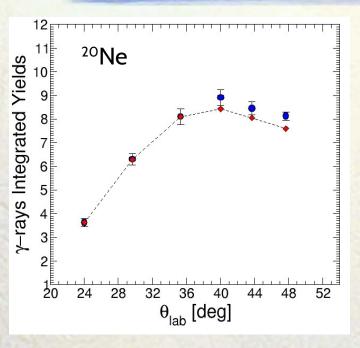


Gamma spectra

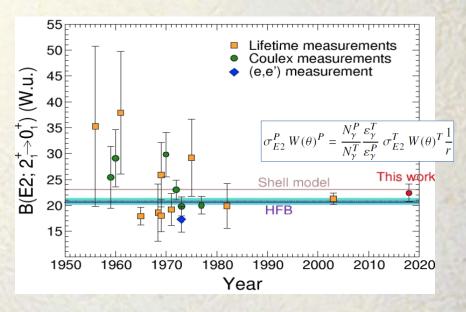
# Particle-gamma coincidence measurement

6 TIGRESS detectos + silidon S3 detector covering forward angles [20.9°, 59.1°]

<sup>110</sup>Pd(<sup>20</sup>Ne, <sup>20</sup>Ne\*)<sup>110</sup>Pd\* @ 64.7 MeV



Angular distribution (arbitrary units)



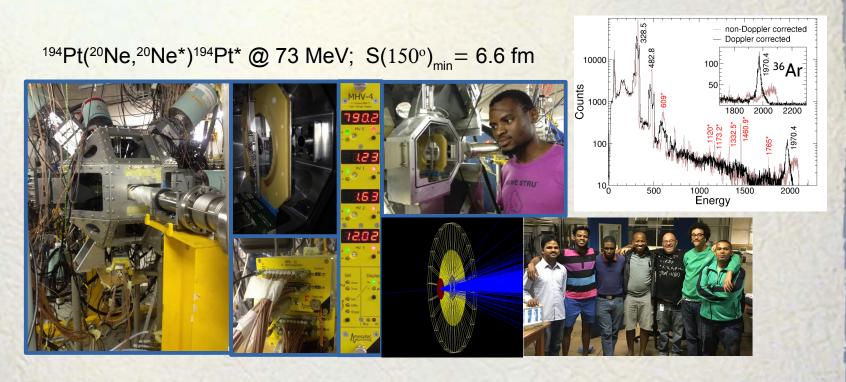
New B(E2) measurement B(E2)=22.4(1.7) W.u.

Particle-gamma coincidence measurement

6 TIGRESS detectos + silidon S3 detector covering forward angles [20.9°, 49.1°]

# New pipeline for Coulomb-excitation measurements using particle-gamma coincidences @ iThemba LABS

Two-month RE campaign April-May 2016: <sup>20</sup>Ne, <sup>32</sup>S, <sup>36</sup>Ar and <sup>40</sup>Ar

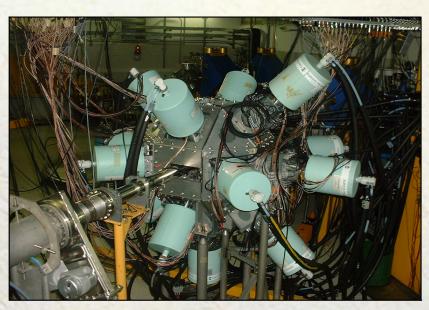


3xS3 detectors + adapters, 6xMPR-32 preamps, 1xMHV-4, feedthrough cables, computers, sorting codes, GEANT, GOSIA (Magda, Dan), <sup>20</sup>Ne, <sup>36,40</sup>Ar, <sup>32</sup>S proposals, 3 postdocs, flexible chamber+ extension (Mathis & Paul) + feedthroughs, new beam development (<sup>60-62</sup>Ni, <sup>102</sup>Ru, Ntombi, Rainer Thomae), DAQ digital system (Pete)

# RE measurement @ iThemba LABS (C. Mehl, PhD)

First safe RE measurement of <sup>20</sup>Ne beams

<sup>194</sup>Pt( $^{20}$ Ne, $^{20}$ Ne $^*$ )<sup>194</sup>Pt $^*$  @ 73 MeV; S(150°)<sub>min</sub> = 6.6 fm



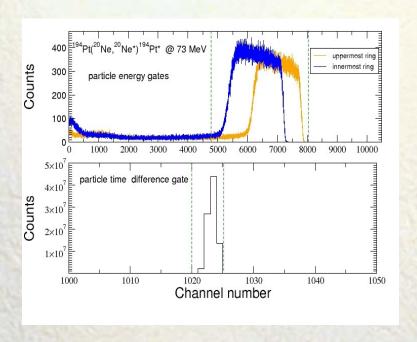


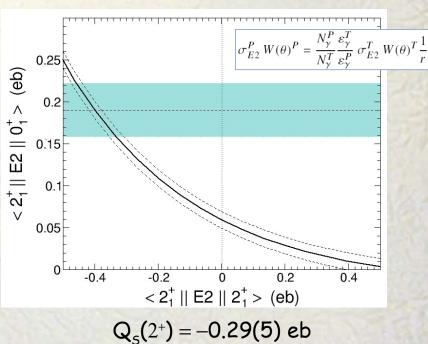
Particle-gamma coincidence technique

AFRODITE + S3 CD-type silicon @ [120°, 150°]

# RE measurement @ iThemba LABS (C. Mehl, PhD)

First safe RE measurement of <sup>20</sup>Ne beams (preliminary results)





In agreement with previous work  $Q_s(2^+) = -0.23(3)$  eb

The large static quadrupole moment in <sup>20</sup>Ne is confirmed

# Shape Coexistence back to SA through measurements using GAMKA

What about second 2+states? Generally no statistics for angular distribution Physics breakthroughs will ultimately come from both HIE-ISOLDE and iThemba LABS













GAMKA – Gamma-ray AsyMmetric spectrometer for Knowledge in Africa PI: Nico Orce

The extended angular coverage and high statistics of the new GAMKA array will allow precise lifetime and mixing ratio measurements of high-lying 2+ states

# Future work and Conclusions

- Clear discrepancy with old and modern nuclear models!
- New cluster calculations being done by Maveric, Khan et al.
- Coulomb-excitation program @ iThemba LABS has been established (new digital electronics + DAQ developed @ UWC)
- First data analysis completed for <sup>12</sup>C, <sup>20</sup>Ne, <sup>32</sup>S, <sup>36</sup>Ar, <sup>40</sup>Ar (<sup>12</sup>C published in Phys. Lett. B (2018).
- Future projects: study of nuclear shapes at low and high excitation using GAMKA (60-62Ni, 100-102Ru, etc)
- Including complementary lifetime + mixing ratio measurements (e.g., <sup>66</sup>Ge, <sup>70</sup>Se, etc).





# Acknowledgements





#### Evidence for clustering in <sup>20</sup>Ne: the Morinaga nucleus

C. V. Mehl, S. Masango, P. Marević, J. N. Orce, E. Khan, M. Djongolov, G. Hackman, M. Kumar Raju, 1,5 K. J. Abrahams, P. Adsley, 5, E. H. Akakpo, 7, G. C. Ball, R. A. Bark, J. A. Becker, D. Bucher, W. Y. Cherif, 5 D. Cline, R. Churchman, T. S. Dinoko, T. E. Drake, D. J.-P. Ebran, N. Erasmus, P. E. Garrett, D. Cline, R. Churchman, T. S. Dinoko, T. E. Drake, D. J.-P. Ebran, N. Erasmus, D. E. Garrett, D. Cline, R. Churchman, D. Churchman, D. Churchman, D. Churchman, D. Cline, R. Churchman, D. Churchman, D A. B. Haves, A. M. Hurst, P. M. Jones, Y. N. Kheswa, J. N. A. Khumalo, J. E. Lawrie, J. J. Lawrie, A. B. Haves, A. M. Hurst, P. M. Jones, Y. N. Kheswa, J. N. A. Khumalo, J. E. Lawrie, J. J. Lawrie, K. L. Malatji, <sup>1,5</sup> E. Martin Montes, <sup>1</sup> D. L. Mavela, <sup>1</sup> M. J. Mokgolobotho, <sup>1</sup> C. Ngwetsheni, <sup>1</sup> S. S. Ntshangase, <sup>7</sup> S. P. Noncolela, <sup>1,5</sup> P. Papka, <sup>6,5</sup> C. J. Pearson, <sup>4</sup> B. Rebeiro, <sup>1</sup> O. Shirinda, <sup>5</sup> M. A. Schumaker, <sup>11</sup> M. A. Stover, <sup>8</sup> C. E. Svensson, <sup>11</sup> S. Triambak, <sup>1</sup> M. Wiedeking, <sup>5</sup> S. J. Williams, <sup>4</sup> J. L. Wood, <sup>12,1</sup> C. Y. Wu, <sup>8</sup> and M. Zielińska <sup>13</sup> <sup>1</sup>Department of Physics & Astronomy, University of the Western Cape, P/B X17, Bellville ZA-7535, South Africa <sup>2</sup>CEA, DAM, DIF, F-91297 Arpaion, France <sup>3</sup>Institut de Physique Nucléaire, IN2P3-CNRS, Université Paris-Saclay, F-91406 Orsay Cedex, France <sup>4</sup>TRIUMF, 4004 Wesbrook Mall, Vancouver, BC V6T 2A3, Canada <sup>5</sup>iThemba LABS, National Research Foundation, P.O. Box 722, Somerset-West 7129, South Africa <sup>6</sup>Department of Physics, Stellenbosch University, Private Bag X1, 7602 Matieland, South Africa <sup>7</sup> University of Zululand, Private Bag X1001, KwaDlangezwa, 3886, South Africa <sup>8</sup>Lawrence Livermore National Laboratory, Livermore, CA 94550, USA <sup>9</sup>Department of Physics and Astronomy, University of Rochester, Rochester, NY 14627, USA <sup>10</sup>Department of Physics, University of Toronto, Toronto ON, M5S 1A7, Canada

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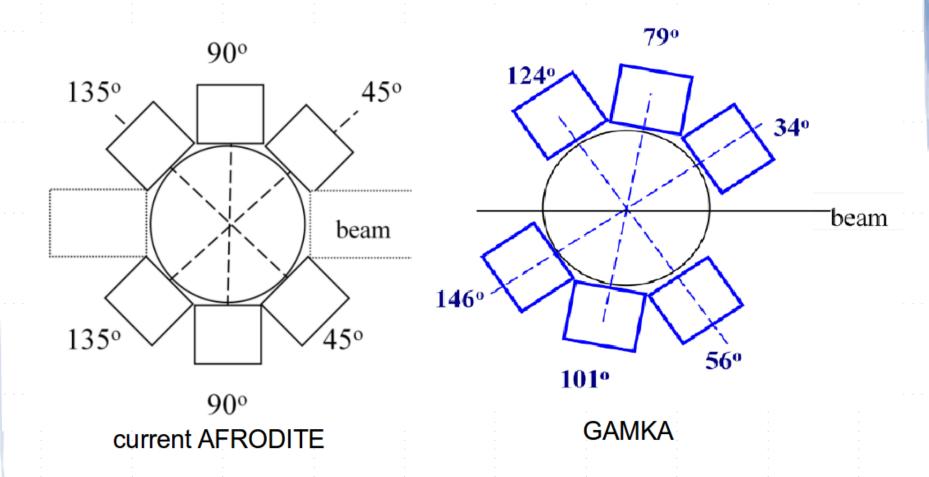
# Ngiyabonga! Imibuzo noma Imibono



**Thank you! Questions or comments** 

# GAMKA: assymetric array at iThemba LABS

How nuclei decay: Transition probabilities, lifetimes, mixing ratios,...



GAMKA – Gamma-ray AsyMmetric spectrometer for Knowledge in Africa