

**B-decay Discrepancies:  
How the Picture Changed After Moriond 2019**

Diego Guadagnoli  
CNRS, LAPTh Annecy

# **B-decay Discrepancies: How the Picture Changed After Moriond 2019**

Diego Guadagnoli  
CNRS, LAPTh Annecy

## ***Overall message***

*The TH picture has evolved while, remarkably, staying coherent – in spite of all the constraints*

*Based on work with*

*J. Aebischer, W. Altmannshofer, M. Reboud, P. Stangl and D. M. Straub*

## Data review

*4 groups of interesting datasets, w/ different challenges*

①  *$b \rightarrow s \mu\mu$  BR data < SM*

*Challenge:  $B \rightarrow$  light meson f.f.'s*

## Data review

*4 groups of interesting datasets, w/ different challenges*

**1**  *$b \rightarrow s \mu\mu$  BR data < SM*

*Challenge:  $B \rightarrow$  light meson f.f.'s*

**2**  *$B \rightarrow K^* \mu\mu$  angular data*

*Challenge: charm loops*

## Data review

4 groups of interesting datasets, w/ different challenges

- 1  $b \rightarrow s \mu\mu$  BR data  $<$  SM  
Challenge:  $B \rightarrow$  light meson f.f.'s
- 2  $B \rightarrow K^* \mu\mu$  angular data  
Challenge: charm loops
- 3  $b \rightarrow s \mu\mu$  /  $b \rightarrow s ee$  ratios  
Challenge: (mostly) stats

## Data review

4 groups of interesting datasets, w/ different challenges

- 1  $b \rightarrow s \mu\mu$  BR data  $<$  SM  
Challenge:  $B \rightarrow$  light meson f.f.'s
- 2  $B \rightarrow K^* \mu\mu$  angular data  
Challenge: charm loops
- 3  $b \rightarrow s \mu\mu$  /  $b \rightarrow s ee$  ratios  
Challenge: (mostly) stats
- 4  $b \rightarrow c \tau\nu$  /  $b \rightarrow c \ell\nu$  ratios  
Challenge: stats + syst

## Basic TH considerations

- ①  $(b \rightarrow s \mu\mu \text{ BR data} < \text{SM})$
- +
- ②  $(B \rightarrow K^* \mu\mu \text{ angular data})$

## Basic TH considerations

- ①  $(b \rightarrow s \mu\mu \text{ BR data} < \text{SM})$
- +
- ②  $(B \rightarrow K^* \mu\mu \text{ angular data})$



*may be alleviated by  
more conservative TH  
assumptions*



## Basic TH considerations

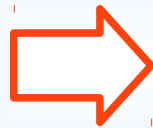
- ①  $(b \rightarrow s \mu\mu \text{ BR data} < \text{SM})$
- +
- ②  $(B \rightarrow K^* \mu\mu \text{ angular data})$



*may be alleviated by  
more conservative TH  
assumptions*

*But*

- ① + ② + ③
- $(R_{K^{(*)}})$



*Explicable (quantitatively)  
w/ two semi-leptonic operators*



*substantial improvement  
w.r.t. SM alone*

## Basic TH considerations

- ①  $(b \rightarrow s \mu\mu \text{ BR data} < \text{SM})$
- +
- ②  $(B \rightarrow K^* \mu\mu \text{ angular data})$



*may be alleviated by  
more conservative TH  
assumptions*

*But*

- ① + ② + ③
- $(R_{K^{(*)}})$



*Explicable (quantitatively)  
w/ two semi-leptonic operators*



*substantial improvement  
w.r.t. SM alone*

*And*

- ① + ② + ③ + ④
- $(R_{D^{(*)}})$



*Explicable (quantitatively)  
w/ single-mediator  
simplified models*

## Data updates

*Our analysis includes the following data updates*

**(a)**  $R_K$  update from LHCb Run1 + 1/3 of Run2

$$R_K \cong 0.85 (1 \pm 7\%)$$



$$2.5\sigma$$

## Data updates

*Our analysis includes the following data updates*

**(a)  $R_K$  update from LHCb Run1 + 1/3 of Run2**

$$R_K \cong 0.85 (1 \pm 7\%) \quad \Rightarrow \quad 2.5\sigma$$

**(b)  $R_{K^*}$  update from Belle**

$$R_{K^*} \cong 0.90 (1 \pm 30\%) \quad \Rightarrow \quad (\text{compatible w/ LHCb's } R_{K^*})$$

## Data updates

*Our analysis includes the following data updates*

**(a)  $R_K$  update from LHCb Run1 + 1/3 of Run2**

$$R_K \cong 0.85 (1 \pm 7\%) \quad \Rightarrow \quad 2.5\sigma$$

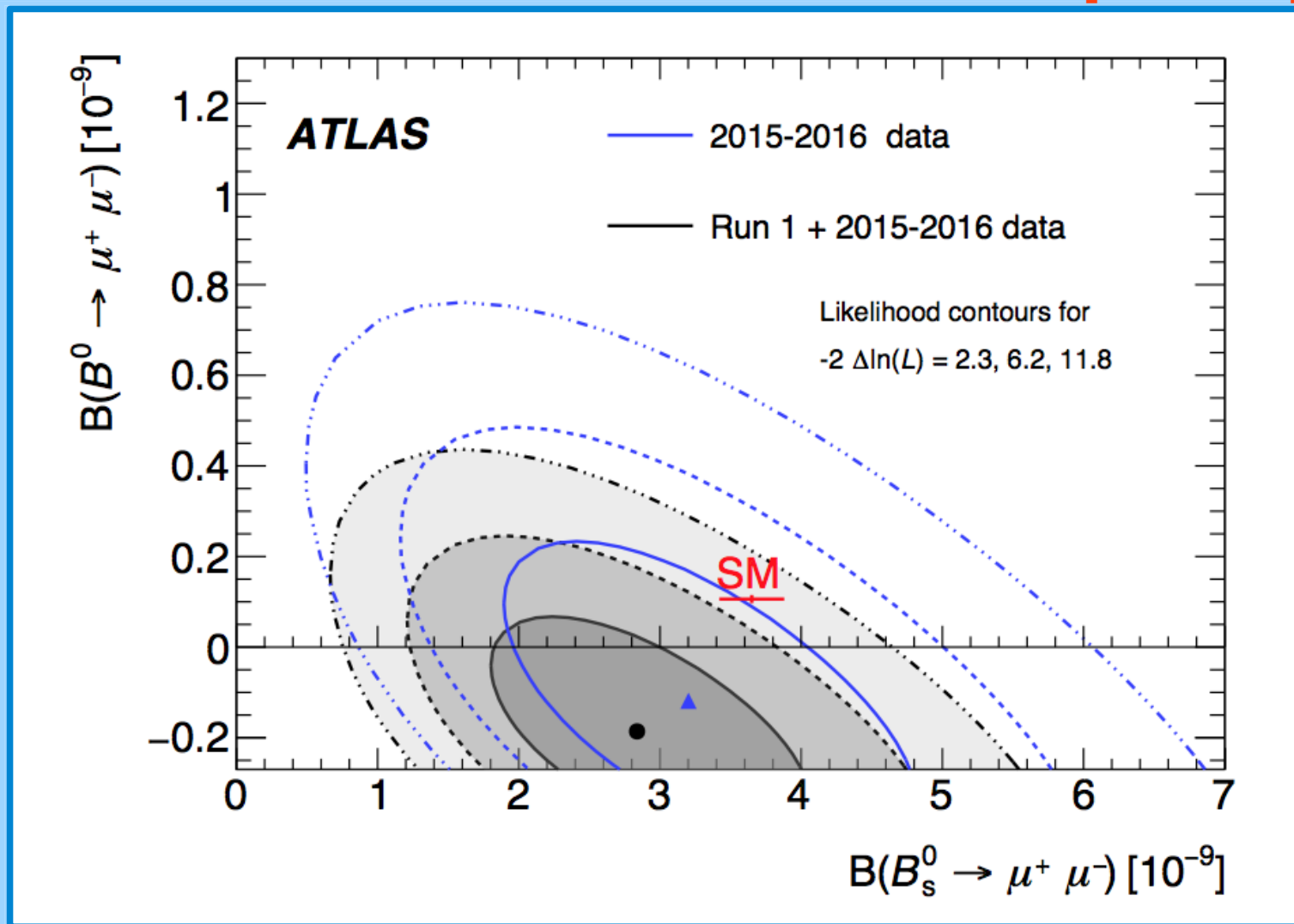
**(b)  $R_{K^*}$  update from Belle**

$$R_{K^*} \cong 0.90 (1 \pm 30\%) \quad \Rightarrow \quad (\text{compatible w/ LHCb's } R_{K^*})$$

**(c)  $B_s \rightarrow \mu\mu$  from ATLAS**

# Data updates

[1812.03017]



## Data updates

Our analysis includes the following data updates

**(a)  $R_K$  update from LHCb Run1 + 1/3 of Run2**

$$R_K \cong 0.85 (1 \pm 7\%) \quad \Rightarrow \quad 2.5\sigma$$

**(b)  $R_{K^*}$  update from Belle**

$$R_{K^*} \cong 0.90 (1 \pm 30\%) \quad \Rightarrow \quad (\text{compatible w/ LHCb's } R_{K^*})$$

**(c)  $B_s \rightarrow \mu\mu$  from ATLAS**

**(d)  $\Lambda_b \rightarrow \Lambda \ell\ell$  :  $A_{FB}$  and BR from LHCb**

**(e)  $b \rightarrow s \gamma$  &  $b \rightarrow s g$  dipole transitions**

**EW-scale**

**Effective-Theory picture**



## **$b \rightarrow s$ EFT picture**

- *One starts from the following Hamiltonian*

$$H(\bar{b} \rightarrow \bar{s} \mu \mu) = -\frac{4G_F}{\sqrt{2}} V_{tb}^* V_{ts} \frac{\alpha_{\text{em}}}{4\pi} \left[ \bar{b}_L \gamma^\lambda s_L \cdot \left( C_9^{(\mu)} \bar{\mu} \gamma_\lambda \mu + C_{10}^{(\mu)} \bar{\mu} \gamma_\lambda \gamma_5 \mu \right) \right]$$

## **b → s EFT picture**

- One starts from the following Hamiltonian

$$H(\bar{b} \rightarrow \bar{s} \mu \mu) = -\frac{4G_F}{\sqrt{2}} V_{tb}^* V_{ts} \frac{\alpha_{\text{em}}}{4\pi} \left[ \bar{b}_L \gamma^\lambda s_L \cdot \left( C_9^{(\mu)} \bar{\mu} \gamma_\lambda \mu + C_{10}^{(\mu)} \bar{\mu} \gamma_\lambda \gamma_5 \mu \right) \right]$$

*About equal size & opposite sign  
in the SM (at the  $m_b$  scale)*

## **b** → **s** EFT picture

- One starts from the following Hamiltonian

$$H(\bar{b} \rightarrow \bar{s} \mu \mu) = -\frac{4G_F}{\sqrt{2}} V_{tb}^* V_{ts} \frac{\alpha_{\text{em}}}{4\pi} \left[ \bar{b}_L \gamma^\lambda s_L \cdot \left( C_9^{(\mu)} \bar{\mu} \gamma_\lambda \mu + C_{10}^{(\mu)} \bar{\mu} \gamma_\lambda \gamma_5 \mu \right) \right]$$

About equal size & opposite sign  
in the SM (at the  $m_b$  scale)

$(V - A) \times (V - A)$  interaction

## **b** → **s** EFT picture

- One starts from the following Hamiltonian

$$H(\bar{b} \rightarrow \bar{s} \mu \mu) = -\frac{4G_F}{\sqrt{2}} V_{tb}^* V_{ts} \frac{\alpha_{\text{em}}}{4\pi} \left[ \bar{b}_L \gamma^\lambda s_L \cdot \boxed{C_9^{(\mu)}} \bar{\mu} \gamma_\lambda \mu + \boxed{C_{10}^{(\mu)}} \bar{\mu} \gamma_\lambda \gamma_5 \mu \right]$$

About equal size & opposite sign  
in the SM (at the  $m_b$  scale)

$(V - A) \times (V - A)$  interaction

- The best-performing BSM scenarios to explain the data involve

$$O_9 \propto \bar{b}_L \gamma^\lambda s_L \cdot \bar{\mu} \gamma_\lambda \mu \quad O_{10} \propto \bar{b}_L \gamma^\lambda s_L \cdot \bar{\mu} \gamma_\lambda \gamma_5 \mu$$

## **b** → **s** EFT picture

- One starts from the following Hamiltonian

$$H(\bar{b} \rightarrow \bar{s} \mu \mu) = -\frac{4G_F}{\sqrt{2}} V_{tb}^* V_{ts} \frac{\alpha_{\text{em}}}{4\pi} \left[ \bar{b}_L \gamma^\lambda s_L \cdot \boxed{C_9^{(\mu)}} \bar{\mu} \gamma_\lambda \mu + \boxed{C_{10}^{(\mu)}} \bar{\mu} \gamma_\lambda \gamma_5 \mu \right]$$

About equal size & opposite sign  
in the SM (at the  $m_b$  scale)

$(V - A) \times (V - A)$  interaction

- The best-performing BSM scenarios to explain the data involve

$$O_9 \propto \bar{b}_L \gamma^\lambda s_L \cdot \bar{\mu} \gamma_\lambda \mu \quad O_{10} \propto \bar{b}_L \gamma^\lambda s_L \cdot \bar{\mu} \gamma_\lambda \gamma_5 \mu$$

- Specifically, either  $O_9$  alone,
- or  $O_9 - O_{10}$   $\Rightarrow$  again,  $(V - A) \times (V - A)$   
well-suited to UV-complete models

# 1-Wilson-coeff. picture

Compare w/ [Algueró *et al.*; Alok *et al.*; Ciuchini *et al.*; Kowalska *et al.*]

Coeff.	best fit	$1\sigma$	$2\sigma$	pull
$C_9^{bs\mu\mu}$	-0.95	[-1.10, -0.79]	[-1.26, -0.63]	$5.8\sigma$
$C_9^{\prime bs\mu\mu}$	+0.09	[-0.07, +0.24]	[-0.23, +0.39]	$0.5\sigma$
$C_{10}^{bs\mu\mu}$	+0.73	[+0.59, +0.87]	[+0.46, +1.01]	$5.6\sigma$
$C_{10}^{\prime bs\mu\mu}$	-0.19	[-0.30, -0.07]	[-0.41, +0.04]	$1.6\sigma$
$C_9^{bs\mu\mu} = C_{10}^{bs\mu\mu}$	+0.20	[+0.05, +0.35]	[-0.09, +0.51]	$1.4\sigma$
$C_9^{bs\mu\mu} = -C_{10}^{bs\mu\mu}$	-0.53	[-0.62, -0.45]	[-0.70, -0.36]	$6.5\sigma$

# 1-Wilson-coeff. picture

Compare w/ [Algueró *et al.*; Alok *et al.*; Ciuchini *et al.*; Kowalska *et al.*]

Coeff.	best fit	$1\sigma$	$2\sigma$	pull
$C_9^{bs\mu\mu}$	-0.95	[-1.10, -0.79]	[-1.26, -0.63]	5.8 $\sigma$
$C_9^{\prime bs\mu\mu}$	+0.09	[-0.07, +0.24]	[-0.23, +0.39]	0.5 $\sigma$
$C_{10}^{bs\mu\mu}$	+0.73	[+0.59, +0.87]	[+0.46, +1.01]	5.6 $\sigma$
$C_{10}^{\prime bs\mu\mu}$	-0.19	[-0.30, -0.07]	[-0.41, +0.04]	1.6 $\sigma$
$C_9^{bs\mu\mu} = C_{10}^{bs\mu\mu}$	+0.20	[+0.05, +0.35]	[-0.09, +0.51]	1.4 $\sigma$
$C_9^{bs\mu\mu} = -C_{10}^{bs\mu\mu}$	-0.53	[-0.62, -0.45]	[-0.70, -0.36]	6.5 $\sigma$

- Two scenarios stand out:  $C_9$  alone or  $C_9 = -C_{10}$  ( $\mu\mu$ -channel only)

# 1-Wilson-coeff. picture

Compare w/ [Algueró *et al.*; Alok *et al.*; Ciuchini *et al.*; Kowalska *et al.*]

Coeff.	best fit	$1\sigma$	$2\sigma$	pull
$C_9^{bs\mu\mu}$	-0.95	[-1.10, -0.79]	[-1.26, -0.63]	5.8 $\sigma$
$C_9^{\prime bs\mu\mu}$	+0.09	[-0.07, +0.24]	[-0.23, +0.39]	0.5 $\sigma$
$C_{10}^{bs\mu\mu}$	+0.73	[+0.59, +0.87]	[+0.46, +1.01]	5.6 $\sigma$
$C_{10}^{\prime bs\mu\mu}$	-0.19	[-0.30, -0.07]	[-0.41, +0.04]	1.6 $\sigma$
$C_9^{bs\mu\mu} = C_{10}^{bs\mu\mu}$	+0.20	[+0.05, +0.35]	[-0.09, +0.51]	1.4 $\sigma$
$C_9^{bs\mu\mu} = -C_{10}^{bs\mu\mu}$	-0.53	[-0.62, -0.45]	[-0.70, -0.36]	6.5 $\sigma$

- Two scenarios stand out:  $C_9$  alone or  $C_9 = -C_{10}$  ( $\mu\mu$ -channel only)
- $C_9 = -C_{10}$  now better than  $C_9$  alone chiefly because of  $B_s \rightarrow \mu\mu$



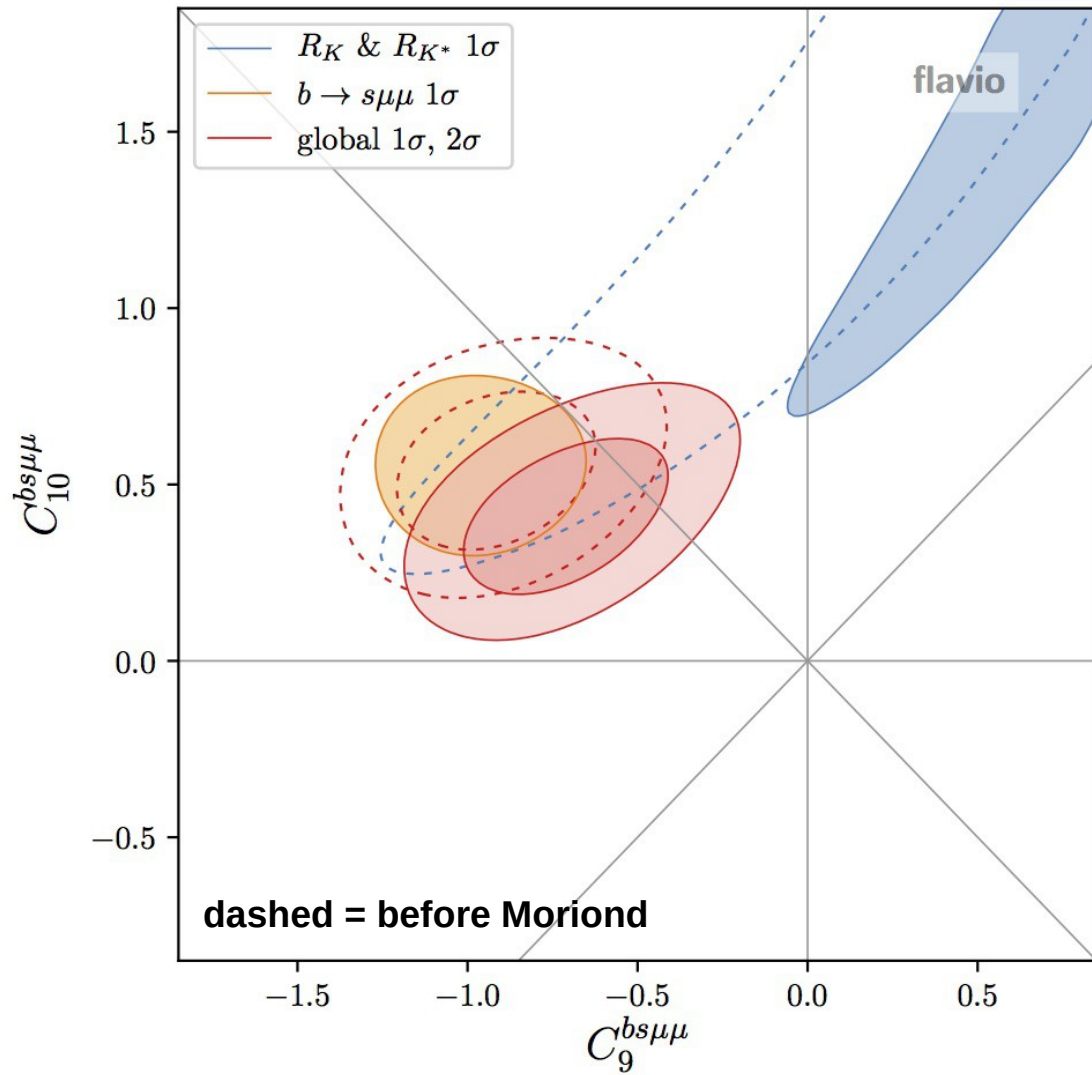
# 1-Wilson-coeff. picture

Compare w/ [Algueró *et al.*; Alok *et al.*; Ciuchini *et al.*; Kowalska *et al.*]

Coeff.	best fit	$1\sigma$	$2\sigma$	pull
$C_9^{bs\mu\mu}$	-0.95	[-1.10, -0.79]	[-1.26, -0.63]	5.8 $\sigma$
$C_9^{\prime bs\mu\mu}$	+0.09	[-0.07, +0.24]	[-0.23, +0.39]	0.5 $\sigma$
$C_{10}^{bs\mu\mu}$	+0.73	[+0.59, +0.87]	[+0.46, +1.01]	5.6 $\sigma$
$C_{10}^{\prime bs\mu\mu}$	-0.19	[-0.30, -0.07]	[-0.41, +0.04]	1.6 $\sigma$
$C_9^{bs\mu\mu} = C_{10}^{bs\mu\mu}$	+0.20	[+0.05, +0.35]	[-0.09, +0.51]	1.4 $\sigma$
$C_9^{bs\mu\mu} = -C_{10}^{bs\mu\mu}$	-0.53	[-0.62, -0.45]	[-0.70, -0.36]	6.5 $\sigma$

- Two scenarios stand out:  $C_9$  alone or  $C_9 = -C_{10}$  ( $\mu\mu$ -channel only)
- $C_9 = -C_{10}$  now better than  $C_9$  alone chiefly because of  $B_s \rightarrow \mu\mu$
- $C_{10}$  alone also ok, but  $B \rightarrow K^* \mu\mu$  unresolved

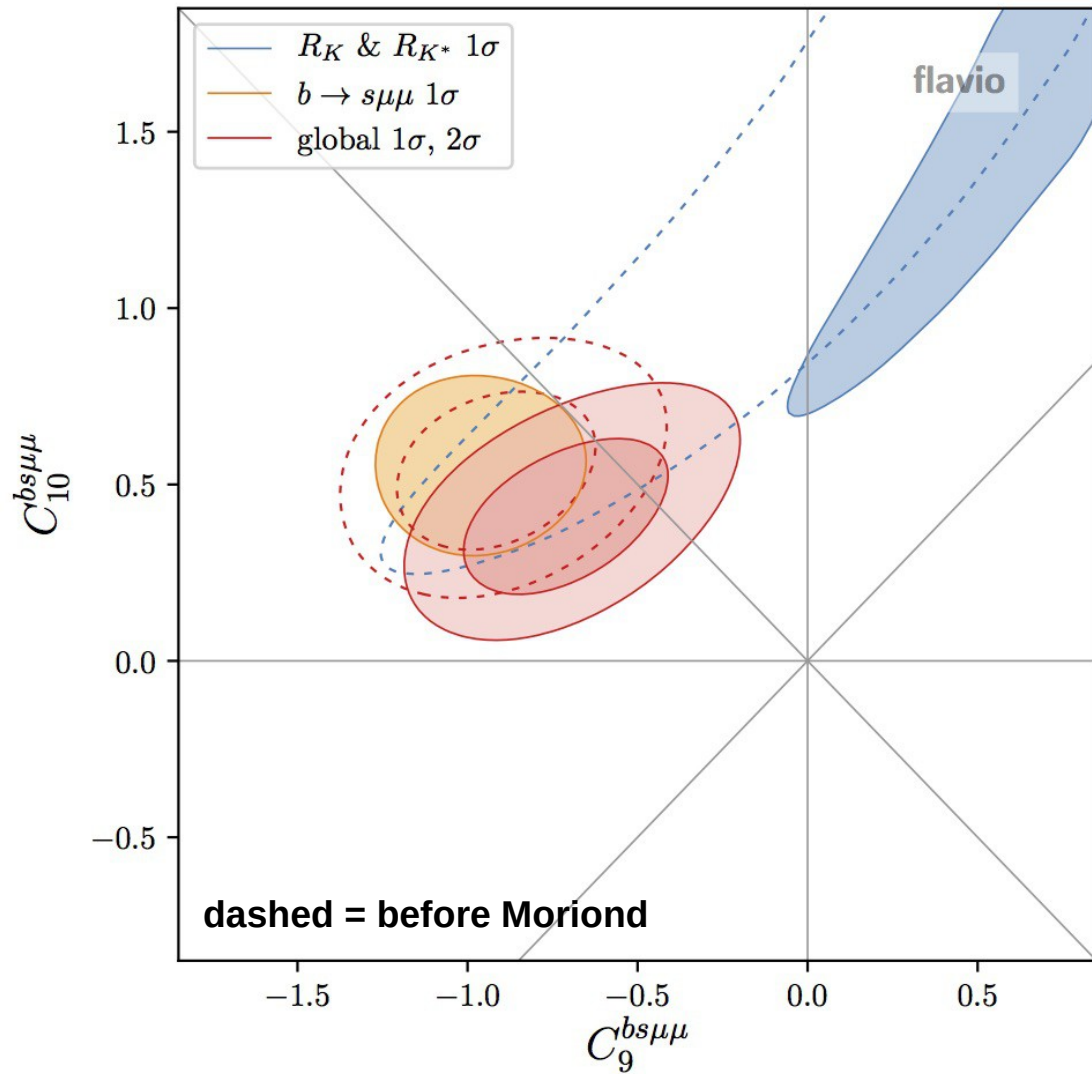
## 2-Wilson-coeffs. picture



### Main points

- $R_{K^{(*)}}$  &  $b \rightarrow s \mu \mu$  in perfect agreement before Moriond

## 2-Wilson-coeffs. picture



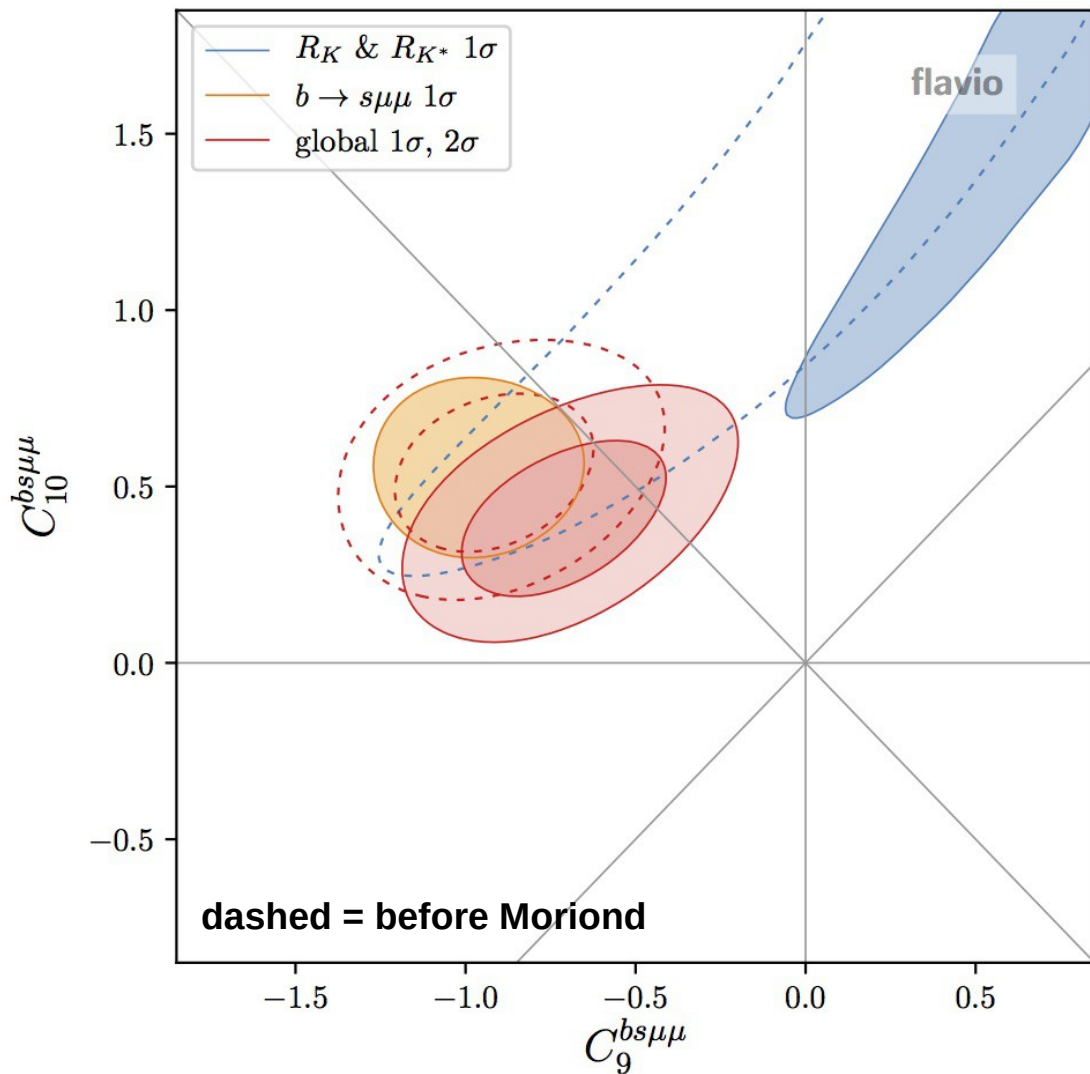
### Main points

- $R_{K^{(*)}}$  &  $b \rightarrow s \mu \mu$  in perfect agreement before Moriond
- now slight tension (in  $C_9$  dir.)




but see later for a UV interpretation

## 2-Wilson-coeffs. picture



### Main points

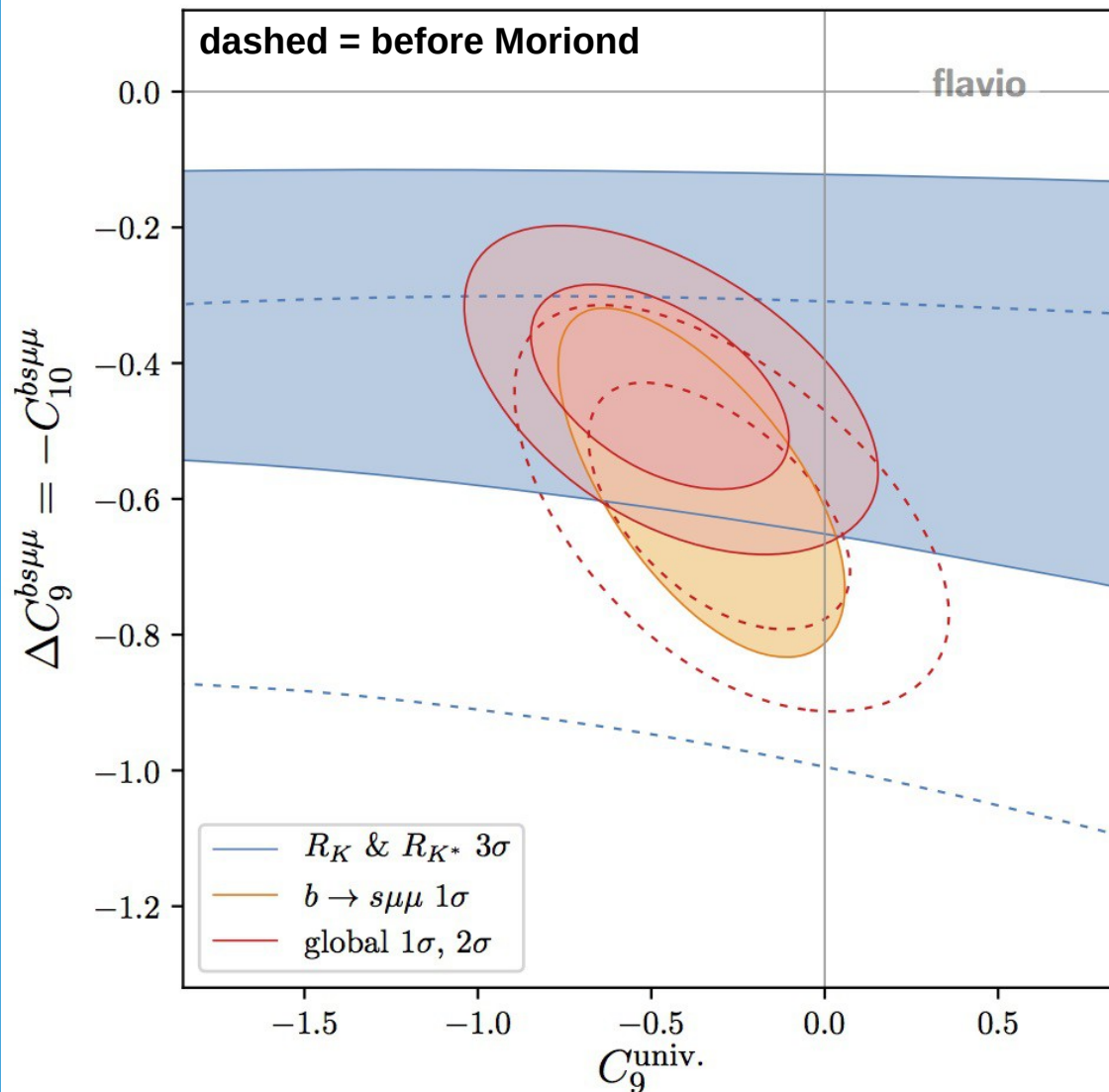
- $R_{K^{(*)}}$  &  $b \rightarrow s\mu\mu$  in perfect agreement before Moriond
- now slight tension (in  $C_9$  dir.)  

 but see later for a UV interpretation
- also (not visible) slight tension between  $R_K$  &  $R_{K^*}$ 
  - would be accommodated by RH quark currents, e.g.  $C_9'$
  - but such shift would not accommodate  $B_s \rightarrow \mu\mu$

## Univ. vs. non-univ. Wilson coeffs.

- Note: a  $C_9^{univ.}$  component would shift  $b \rightarrow s \mu\mu$  data but not  $R_{K^{(*)}}$

# Univ. vs. non-univ. Wilson coeffs.

- Note: a  $C_9^{univ.}$  component would shift  $b \rightarrow s \mu\mu$  data but not  $R_{K^{(*)}}$



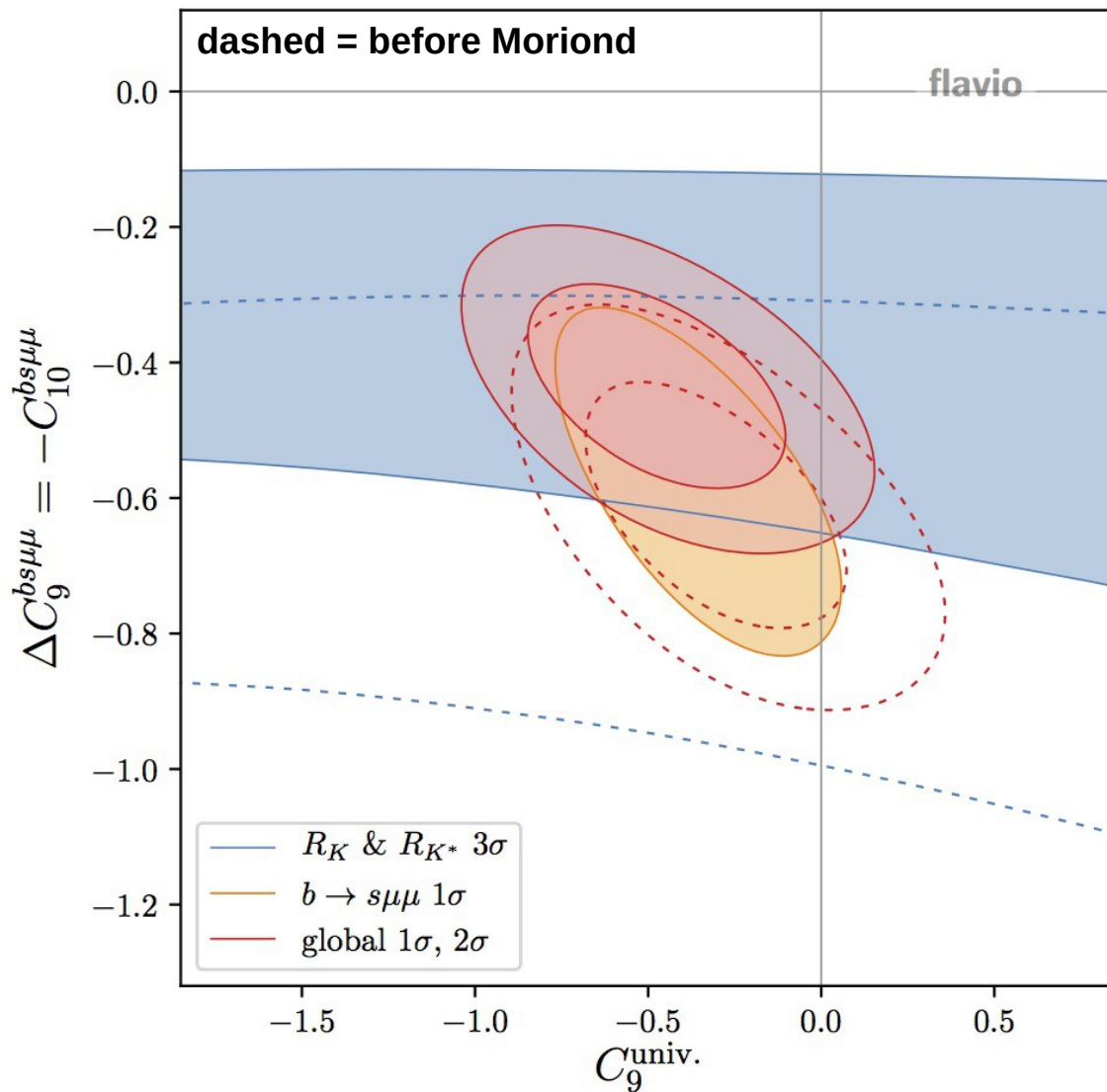
## Notes

y-axis:  $\mu$ -specific shift in  $C_9 = -C_{10}$

x-axis: additional, lepton-univ. shift in  $C_9$  only

# Univ. vs. non-univ. Wilson coeffs.

- Note: a  $C_9^{univ.}$  component would shift  $b \rightarrow s \mu\mu$  data but not  $R_{K^{(*)}}$



## Notes

y-axis:  $\mu$ -specific shift in  $C_9 = -C_{10}$

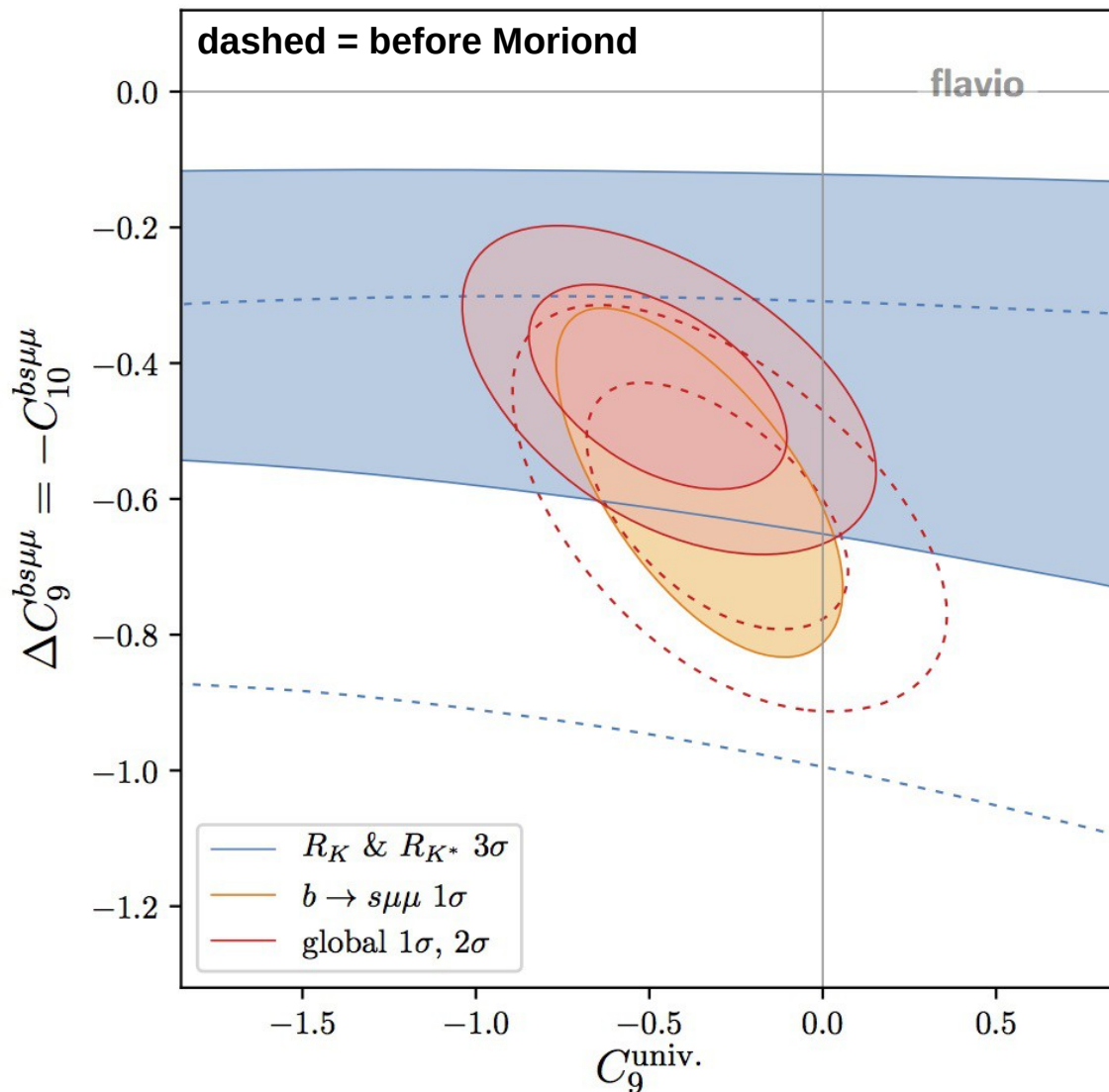
x-axis: additional, lepton-univ. shift in  $C_9$  only

- Post-Moriond data tend to prefer  $C_9^{univ.} \neq 0$



# Univ. vs. non-univ. Wilson coeffs.

- Note: a  $C_9^{univ.}$  component would shift  $b \rightarrow s \mu\mu$  data but not  $R_{K^{(*)}}$



## Notes

y-axis:  $\mu$ -specific shift in  $C_9 = -C_{10}$

x-axis: additional, lepton-univ. shift in  $C_9$  only

- Post-Moriond data tend to prefer  $C_9^{univ.} \neq 0$
- This suggests a well-defined interpretation within SMEFT [Crivellin-Greub-Müller-Saturnino]



**Going above the EW scale  
without introducing new d.o.f.:  
The SM EFT**

## SMEFT basics

- If NP is at a scale  $\Lambda \gg M_{EW}$  with nothing new in between



Effects below  $\Lambda$  are described by ops. constructed with SM fields, and invariant under the full SM group:  $SU(3)_c \times SU(2)_L \times U(1)_Y$

This defines the SMEFT

## SMEFT basics

- If NP is at a scale  $\Lambda \gg M_{EW}$  with nothing new in between



Effects below  $\Lambda$  are described by ops. constructed with SM fields, and invariant under the full SM group:  $SU(3)_c \times SU(2)_L \times U(1)_Y$

This defines the SMEFT

- After defining a (non-redundant) op. basis for SMEFT

[B. Grzadkowski et al., JHEP 2010]

contributions to muonic  $C_9 = -C_{10}$  or  $C_9^{univ.}$  can come from:

## SMEFT basics

- If NP is at a scale  $\Lambda \gg M_{EW}$  with nothing new in between



Effects below  $\Lambda$  are described by ops. constructed with SM fields, and invariant under the full SM group:  $SU(3)_c \times SU(2)_L \times U(1)_Y$

This defines the SMEFT

- After defining a (non-redundant) op. basis for SMEFT

[B. Grzadkowski et al., JHEP 2010]

contributions to muonic  $C_9 = -C_{10}$  or  $C_9^{univ.}$  can come from:

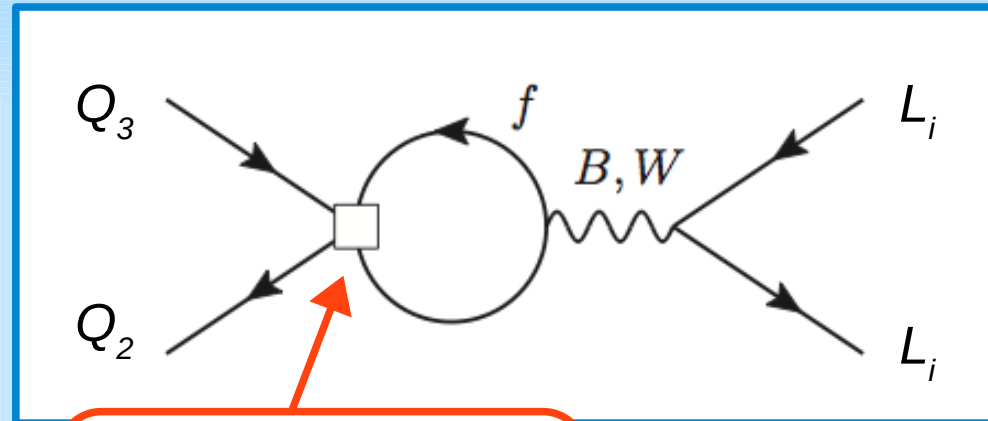
- 1 SMEFT operators directly matching onto  $O_{9,10}$

$$[O_{LQ}^{(1)}]_{2223} = \bar{L}_2 \gamma^\lambda L_2 \cdot \bar{Q}_2 \gamma_\lambda Q_3$$

$$[O_{LQ}^{(3)}]_{2223} = \bar{L}_2 \gamma^\lambda \sigma^a L_2 \cdot \bar{Q}_2 \gamma_\lambda \sigma^a Q_3$$

## SMEFT basics

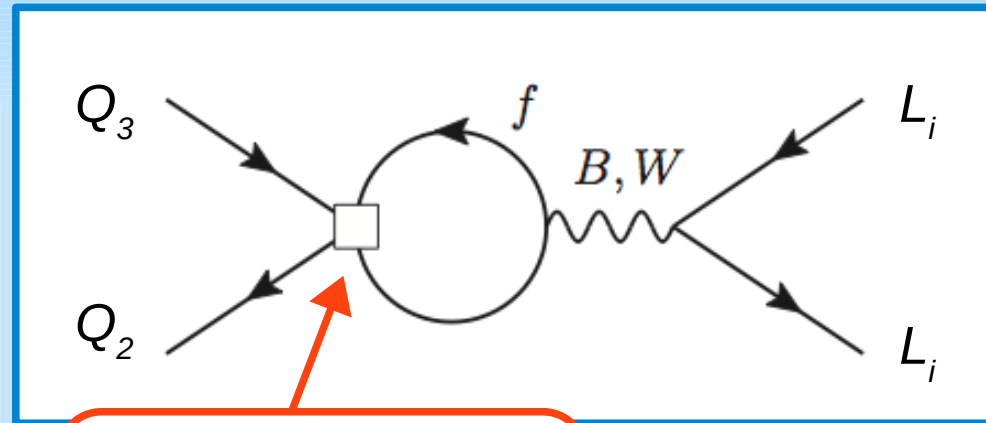
- 2 Besides, contribs. to  $C_9^{\text{univ.}}$  can come from RGE effects, in particular:



any suitable SMEFT  
4-fermion op. here

## SMEFT basics

- 2 Besides, contribs. to  $C_9^{univ.}$  can come from RGE effects, in particular:



any suitable SMEFT  
4-fermion op. here

- Case  $f = \tau$  especially interesting



Connection with “semi-tauonic” ops.,  
responsible for  $b \rightarrow c \tau \nu$

- Smoking gun of such scenario: large enhancement in  $b \rightarrow s \tau \tau$   
[B. Capdevila et al., PRL 2018]

## ***4-fermion, semi-tauonic ops.***

- $[O_{LQ}^{(3)}]_{3323} \supset \bar{\tau} \gamma_L^\lambda \nu \cdot \bar{c} \gamma_{\lambda L} b$

## 4-fermion, semi-tauonic ops.

- $[O_{LQ}^{(3)}]_{3323} \supset \bar{\tau} \gamma_L^\lambda \nu \cdot \bar{c} \gamma_{\lambda L} b$

⇒ can explain  $R_{D^{(*)}}$

⇩ also induces  $C_9^{univ.}$  w/ the right sign  
to potentially accommodate  $b \rightarrow s \mu\mu$

[Crivellin-Greub-Müller-Saturnino]



## 4-fermion, semi-tauonic ops.

- $[O_{LQ}^{(3)}]_{3323} \supset \bar{\tau} \gamma_L^\lambda \nu \cdot \bar{c} \gamma_{\lambda L} b$

⇒ can explain  $R_{D^{(*)}}$

⇩ also induces  $C_9^{univ.}$  w/ the right sign  
to potentially accommodate  $b \rightarrow s \mu \mu$   
[Crivellin-Greub-Müller-Saturnino]

- $[O_{LQ}^{(1) \text{ or } (3)}]_{2223} \supset \bar{\mu} \gamma_L^\lambda \mu \cdot \bar{s} \gamma_{\lambda L} b$

⇒ can explain  $R_{K^{(*)}}$

## 4-fermion, semi-tauonic ops.

- $[O_{LQ}^{(3)}]_{3323} \supset \bar{\tau} \gamma_L^\lambda \nu \cdot \bar{c} \gamma_{\lambda L} b$

⇒ can explain  $R_{D^{(*)}}$

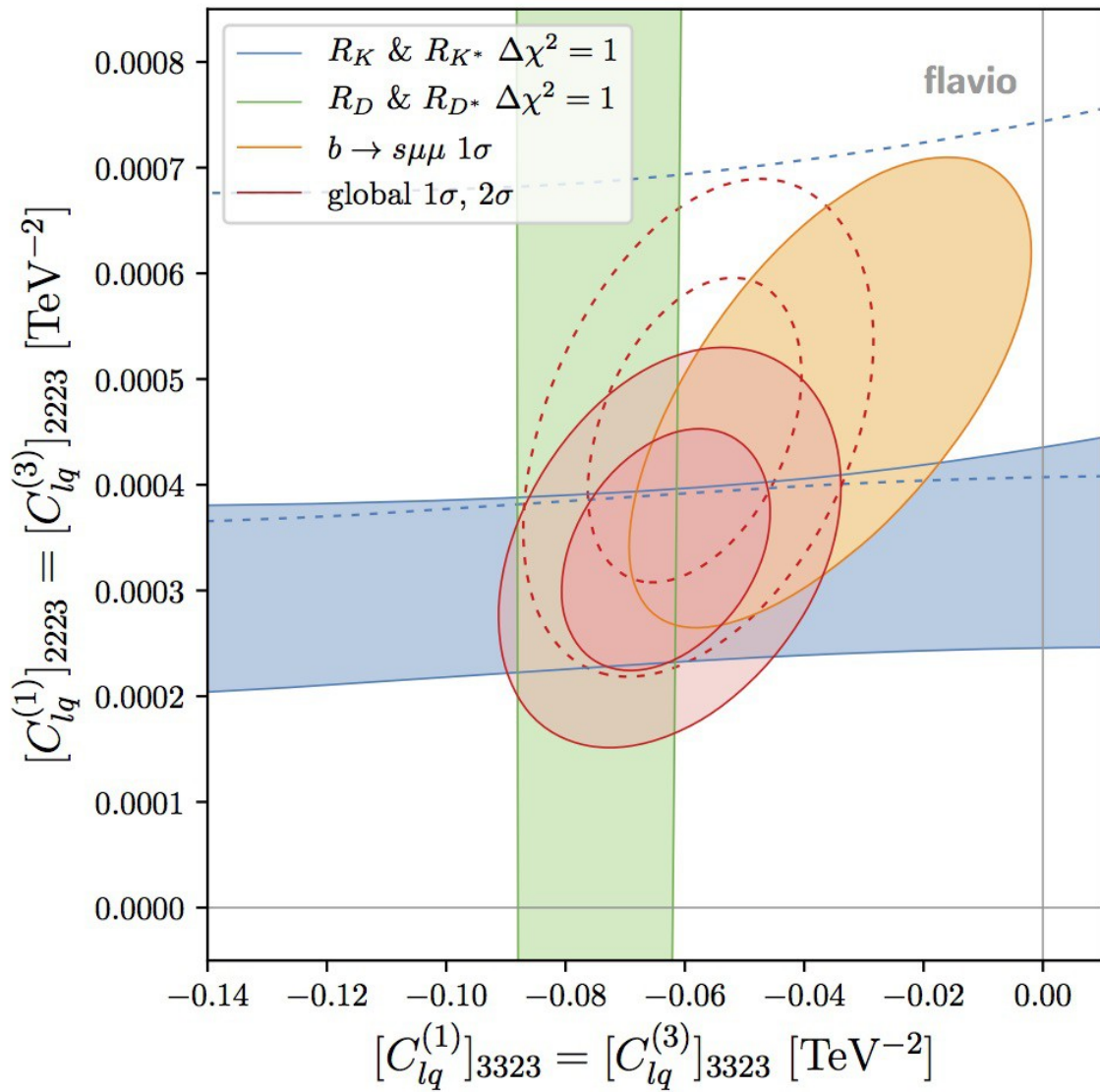
⇩ also induces  $C_9^{univ.}$  w/ the right sign  
to potentially accommodate  $b \rightarrow s \mu \mu$   
[Crivellin-Greub-Müller-Saturnino]

- $[O_{LQ}^{(1) \text{ or } (3)}]_{2223} \supset \bar{\mu} \gamma_L^\lambda \mu \cdot \bar{s} \gamma_{\lambda L} b$

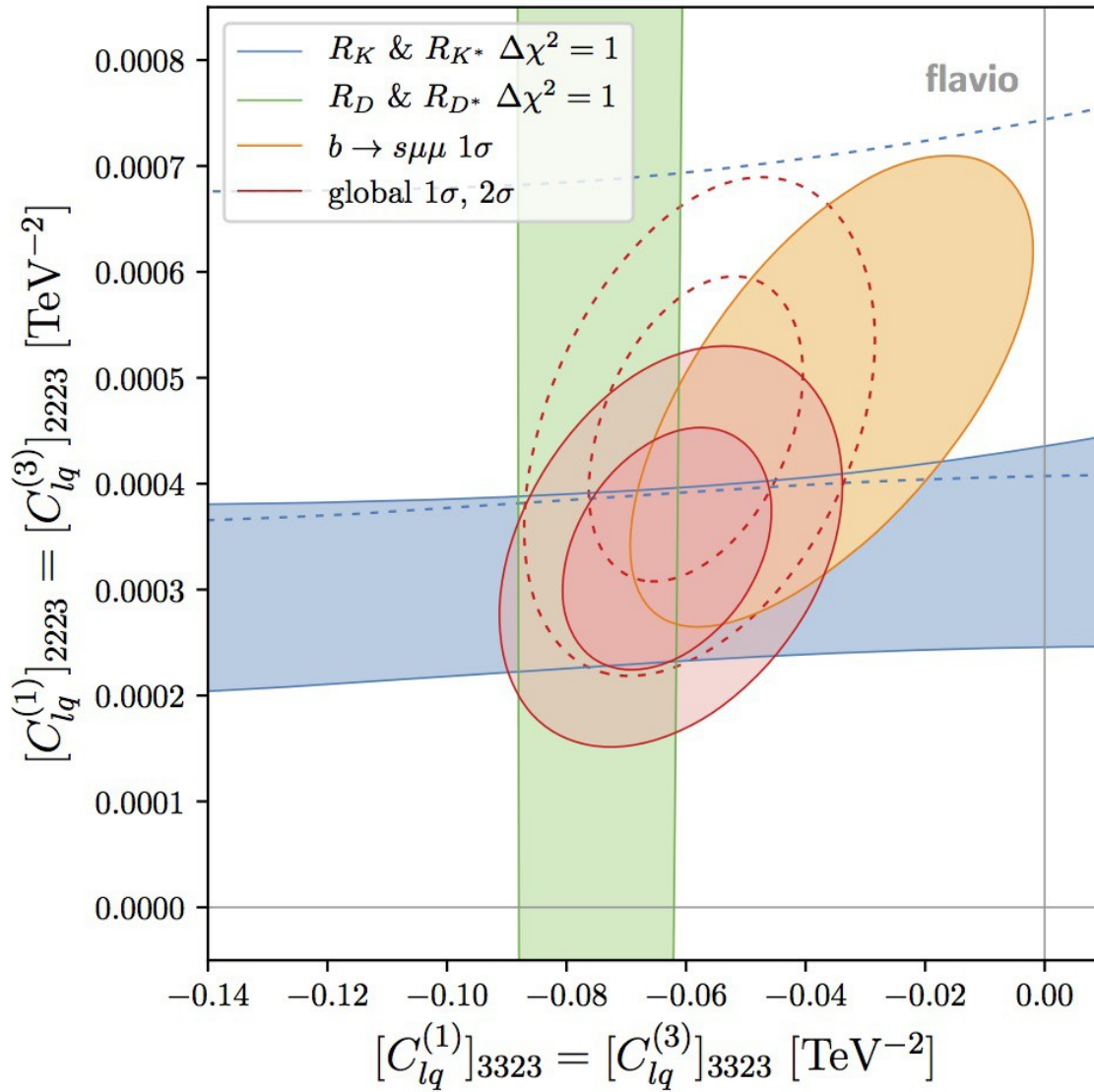
⇒ can explain  $R_{K^{(*)}}$

- *Caveat:* one must have  $[C_{LQ}^{(1)}]_{3323} \simeq [C_{LQ}^{(3)}]_{3323}$   
to avoid the  $B \rightarrow K^{(*)} \nu \nu$  constraint  
[Buras-Girrbach-Niehoff-Straub]

$$[C_{LQ}^{(1)}]_{3323} = [C_{LQ}^{(3)}]_{3323} \quad \text{vs.} \quad [C_{LQ}^{(1)}]_{2223} = [C_{LQ}^{(3)}]_{2223}$$



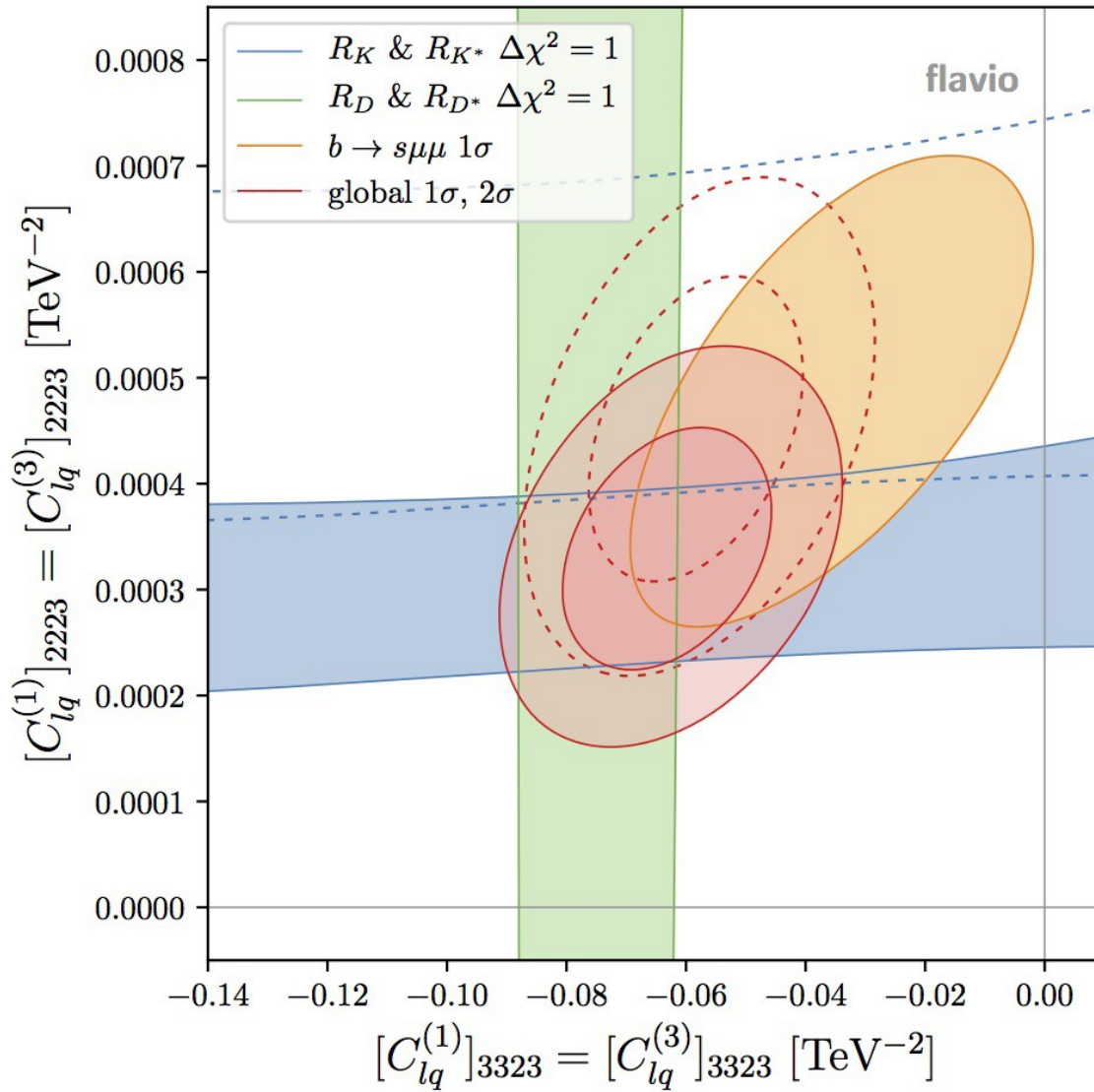
$$[C_{LQ}^{(1)}]_{3323} = [C_{LQ}^{(3)}]_{3323} \quad \text{vs.} \quad [C_{LQ}^{(1)}]_{2223} = [C_{LQ}^{(3)}]_{2223}$$



**Before Moriond (dashed)**

$R_{K^{(*)}}$  (blue) and  $b \rightarrow s\mu\mu$  (orange) were in perfect agreement (y-axis)

$$[C_{LQ}^{(1)}]_{3323} = [C_{LQ}^{(3)}]_{3323} \quad \text{vs.} \quad [C_{LQ}^{(1)}]_{2223} = [C_{LQ}^{(3)}]_{2223}$$



### Before Moriond (dashed)

$R_{K^{(*)}}$  (blue) and  $b \rightarrow s \mu\mu$  (orange) were in perfect agreement (y-axis)

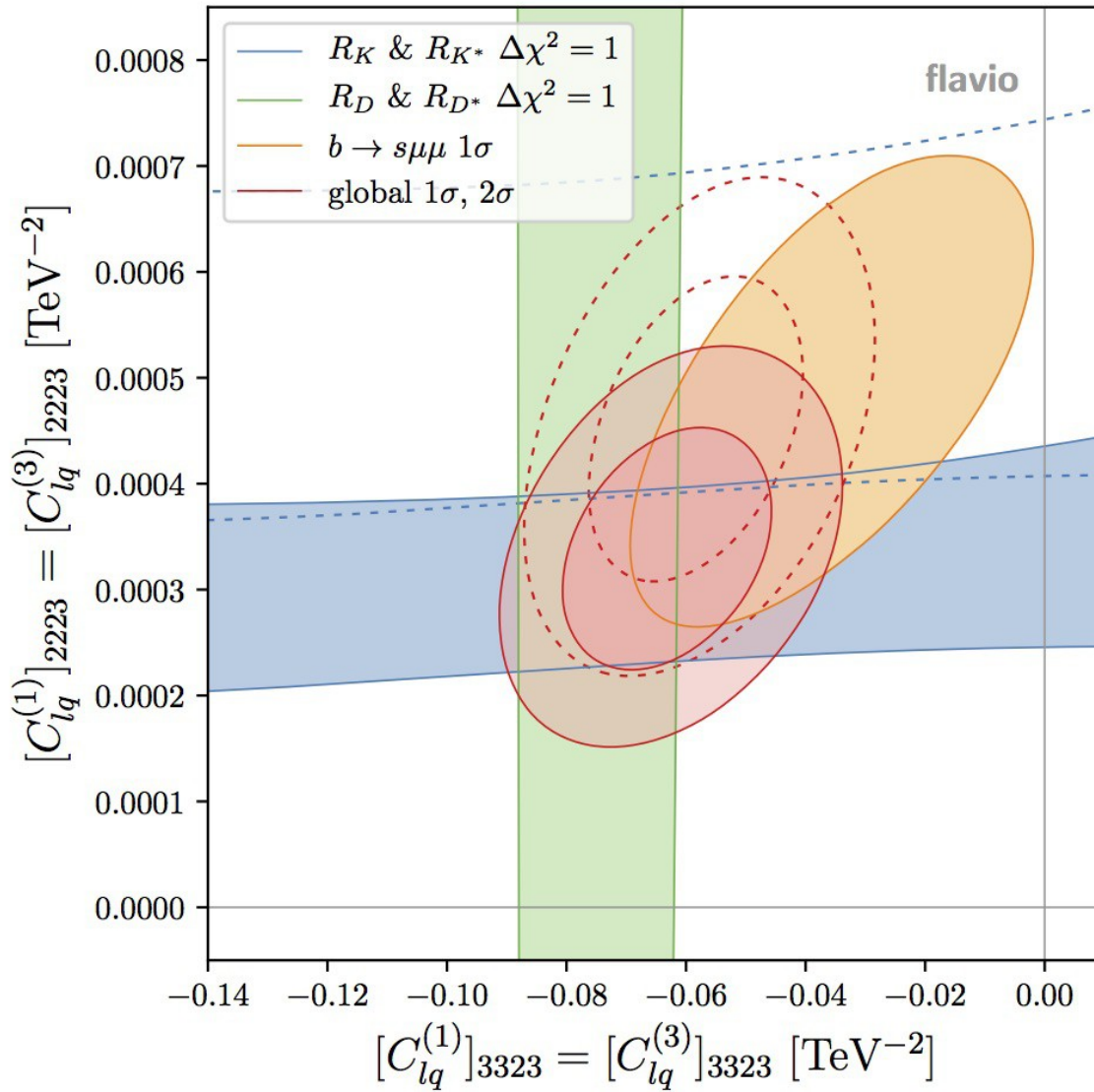


in a region close to 0 in the x-axis



$R_{D^{(*)}}$  not explained

$$[C_{LQ}^{(1)}]_{3323} = [C_{LQ}^{(3)}]_{3323} \quad \text{vs.} \quad [C_{LQ}^{(1)}]_{2223} = [C_{LQ}^{(3)}]_{2223}$$



### Before Moriond (dashed)

$R_{K^{(*)}}$  (blue) and  $b \rightarrow s \mu\mu$  (orange) were in perfect agreement (y-axis)



in a region close to 0 in the x-axis



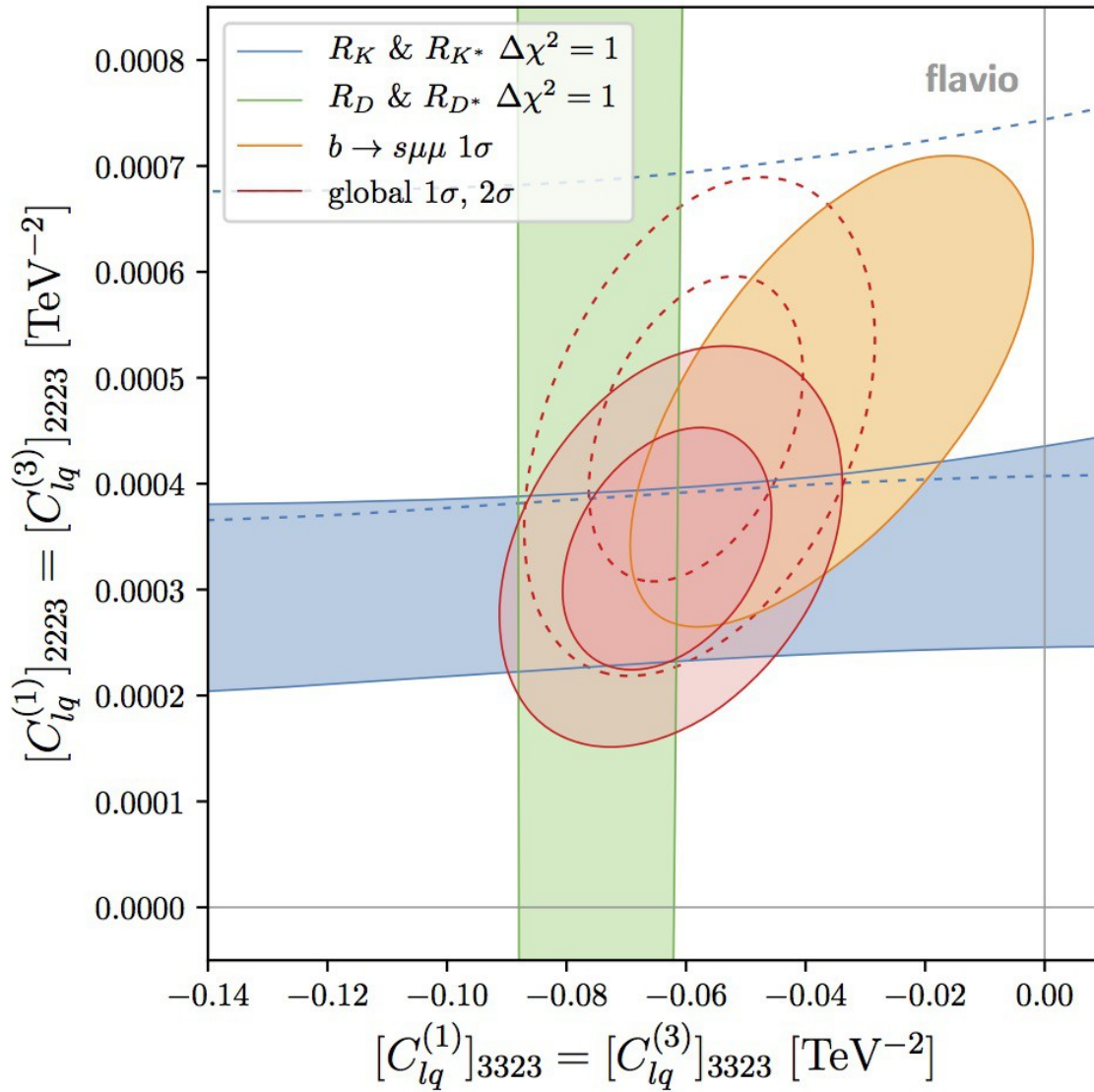
$R_{D^{(*)}}$  not explained

### After Moriond

$R_{K^{(*)}}$  and  $b \rightarrow s \mu\mu$  intersect in a region corresponding to x-axis values well below 0



$$[C_{LQ}^{(1)}]_{3323} = [C_{LQ}^{(3)}]_{3323} \quad \text{vs.} \quad [C_{LQ}^{(1)}]_{2223} = [C_{LQ}^{(3)}]_{2223}$$



### Before Moriond (dashed)

$R_{K^{(*)}}$  (blue) and  $b \rightarrow s \mu\mu$  (orange) were in perfect agreement (y-axis)



in a region close to 0 in the x-axis



$R_{D^{(*)}}$  not explained

### After Moriond

$R_{K^{(*)}}$  and  $b \rightarrow s \mu\mu$  intersect in a region corresponding to x-axis values well below 0

This region turns out to overlap substantially with the  $R_{D^{(*)}}$  region (green)

**Beyond EFTs:**

**The picture within “simplified” models**



## The $U_1$ leptoquark

- $U_1 \sim (\mathbf{3}, \mathbf{1})_{2/3}$  is the only single mediator known to yield

$$[C_{LQ}^{(1)}]_{3323} = [C_{LQ}^{(3)}]_{3323} \neq 0 \quad \&\& \quad [C_{LQ}^{(1)}]_{2223} = [C_{LQ}^{(3)}]_{2223} \neq 0$$

[Alonso-Grinstein-Martin-Camalich, Calibbi-Crivellin-Ota, 2015]

## The $U_1$ leptoquark

- $U_1 \sim (\mathbf{3}, \mathbf{1})_{2/3}$  is the only single mediator known to yield

$$[C_{LQ}^{(1)}]_{3323} = [C_{LQ}^{(3)}]_{3323} \neq 0 \quad \&\& \quad [C_{LQ}^{(1)}]_{2223} = [C_{LQ}^{(3)}]_{2223} \neq 0$$

[Alonso-Grinstein-Martin-Camalich, Calibbi-Crivellin-Ota, 2015]

- Define the couplings:

$$\mathcal{L}_{U_1} \supset g_{lq}^{ji} \bar{Q}^i \gamma^\mu L^j U_\mu + \text{h.c.}$$

## The $U_1$ leptoquark


- $U_1 \sim (\mathbf{3}, \mathbf{1})_{2/3}$  is the only single mediator known to yield

$$[C_{LQ}^{(1)}]_{3323} = [C_{LQ}^{(3)}]_{3323} \neq 0 \quad \&\& \quad [C_{LQ}^{(1)}]_{2223} = [C_{LQ}^{(3)}]_{2223} \neq 0$$

[Alonso-Grinstein-Martin-Camalich, Calibbi-Crivellin-Ota, 2015]

- Define the couplings:

$$\mathcal{L}_{U_1} \supset g_{lq}^{ji} \bar{Q}^i \gamma^\mu L^j U_\mu + \text{h.c.}$$

  $\delta R_{K(*)} |_{\text{in } \mu \text{ channel}} \propto g_{lq}^{22} \ \& \ g_{lq}^{23}$

## The $U_1$ leptoquark

- $U_1 \sim (\mathbf{3}, \mathbf{1})_{2/3}$  is the only single mediator known to yield

$$[C_{LQ}^{(1)}]_{3323} = [C_{LQ}^{(3)}]_{3323} \neq 0 \quad \&\& \quad [C_{LQ}^{(1)}]_{2223} = [C_{LQ}^{(3)}]_{2223} \neq 0$$

[Alonso-Grinstein-Martin-Camalich, Calibbi-Crivellin-Ota, 2015]

- Define the couplings:

$$\mathcal{L}_{U_1} \supset g_{lq}^{ji} \bar{Q}^i \gamma^\mu L^j U_\mu + \text{h.c.}$$

$$\begin{aligned} \Rightarrow \delta R_{K(*)} |_{\text{in } \mu \text{ channel}} &\propto g_{lq}^{22} \quad \& \quad g_{lq}^{23} \\ \delta R_{D(*)} |_{\text{in } \tau \text{ channel}} &\propto g_{lq}^{32} \quad \& \quad g_{lq}^{33} \end{aligned}$$

## The $U_1$ leptoquark

- $U_1 \sim (\mathbf{3}, \mathbf{1})_{2/3}$  is the only single mediator known to yield

$$[C_{LQ}^{(1)}]_{3323} = [C_{LQ}^{(3)}]_{3323} \neq 0 \quad \&\& \quad [C_{LQ}^{(1)}]_{2223} = [C_{LQ}^{(3)}]_{2223} \neq 0$$

[Alonso-Grinstein-Martin-Camalich, Calibbi-Crivellin-Ota, 2015]

- Define the couplings:

$$\mathcal{L}_{U_1} \supset g_{lq}^{ji} \bar{Q}^i \gamma^\mu L^j U_\mu + \text{h.c.}$$

$$\delta R_{K(*)} |_{\text{in } \mu \text{ channel}} \propto g_{lq}^{22} \ \& \ g_{lq}^{23}$$

$$\delta R_{D(*)} |_{\text{in } \tau \text{ channel}} \propto g_{lq}^{32} \ \& \ g_{lq}^{33}$$

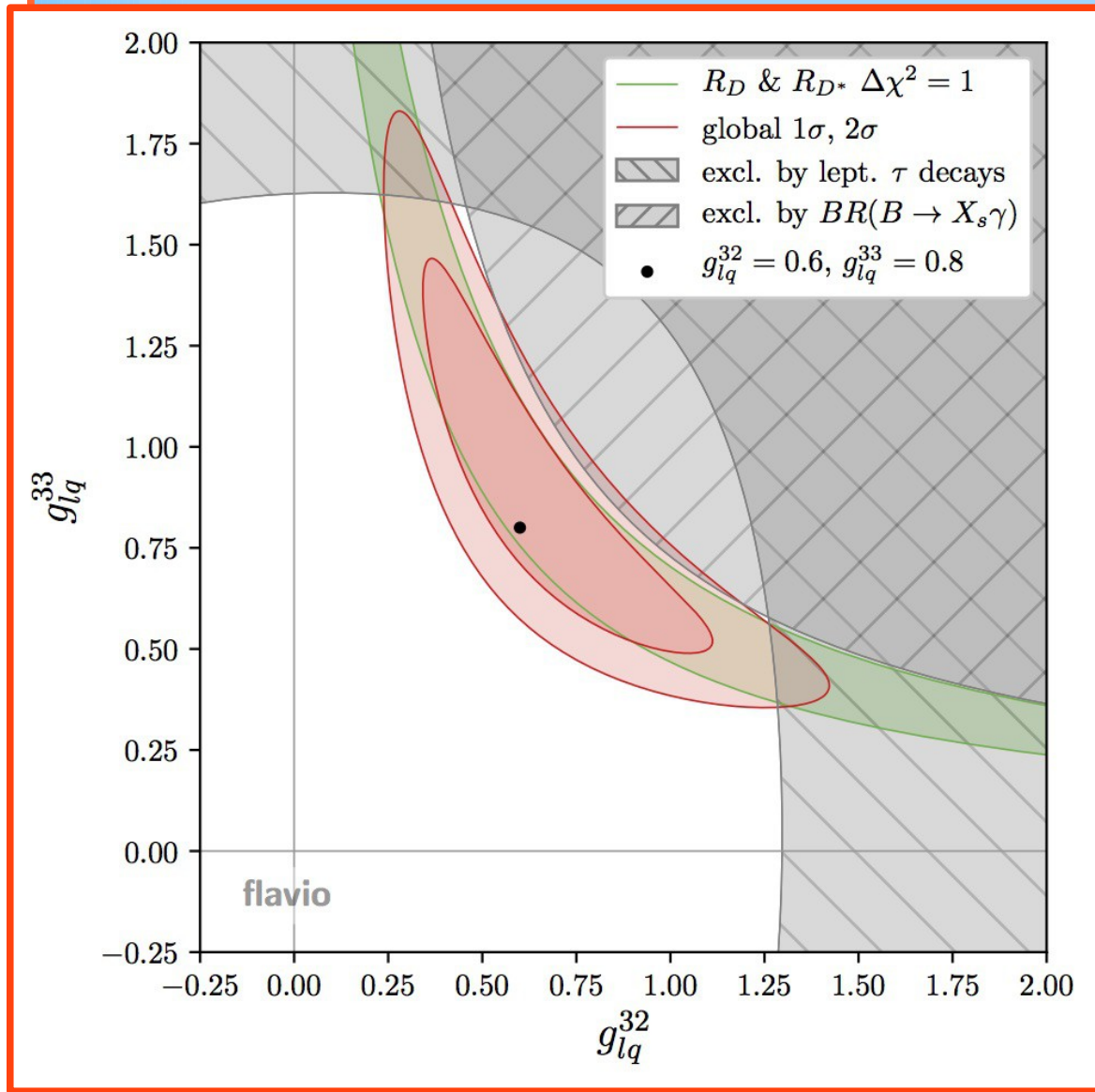


these couplings also famously constrained by

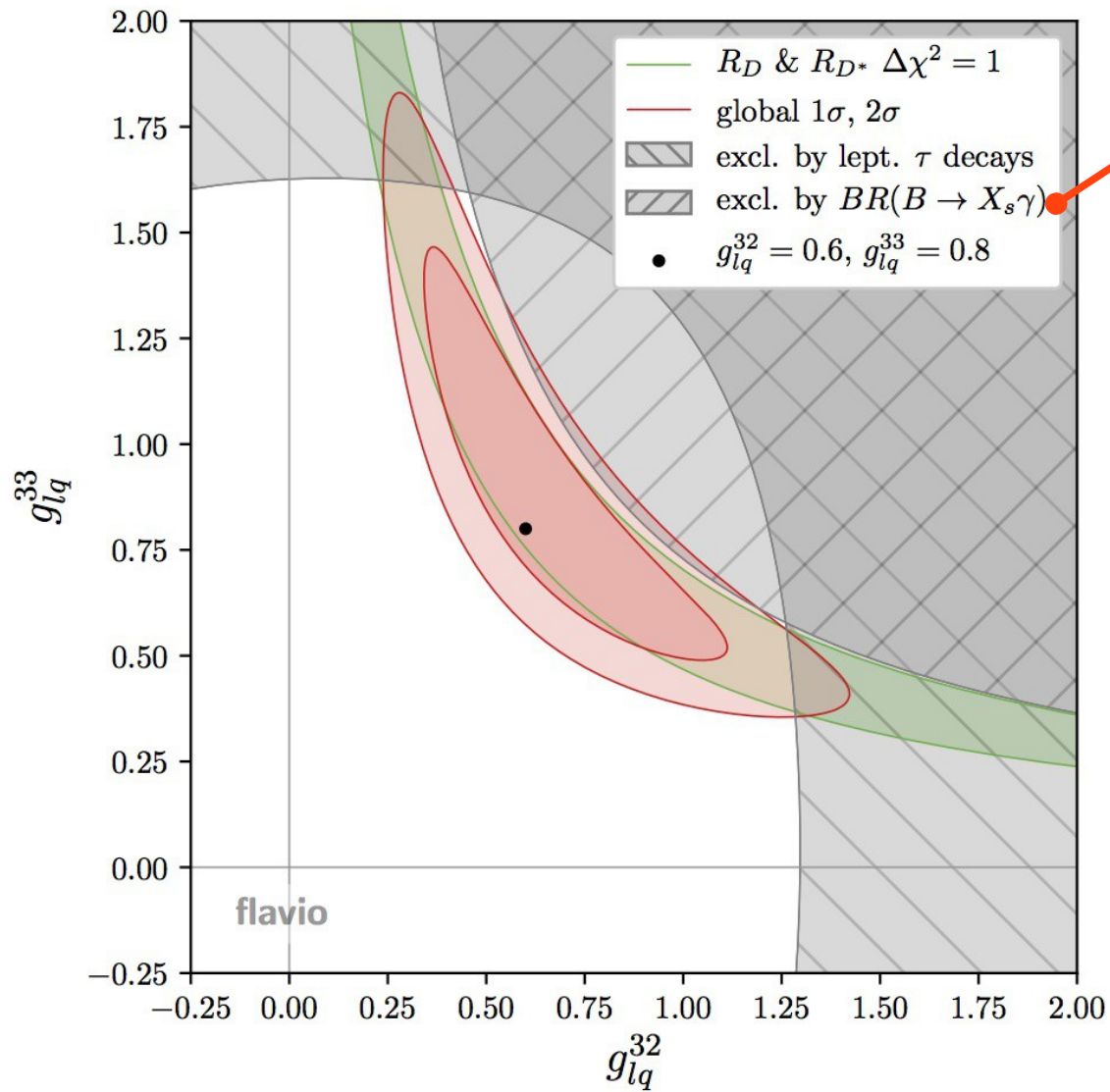
$$\tau \rightarrow \ell \nu \nu \quad [\text{Feruglio-Paradisi-Pattori}]$$

(hence far from obvious that an  $R_{D(*)}$  description achievable)

**U<sub>1</sub> LQ:  $g_{lq}^{32}$  vs.  $g_{lq}^{33}$**



**U<sub>1</sub> LQ:  $g_{lq}^{32}$  vs.  $g_{lq}^{33}$**

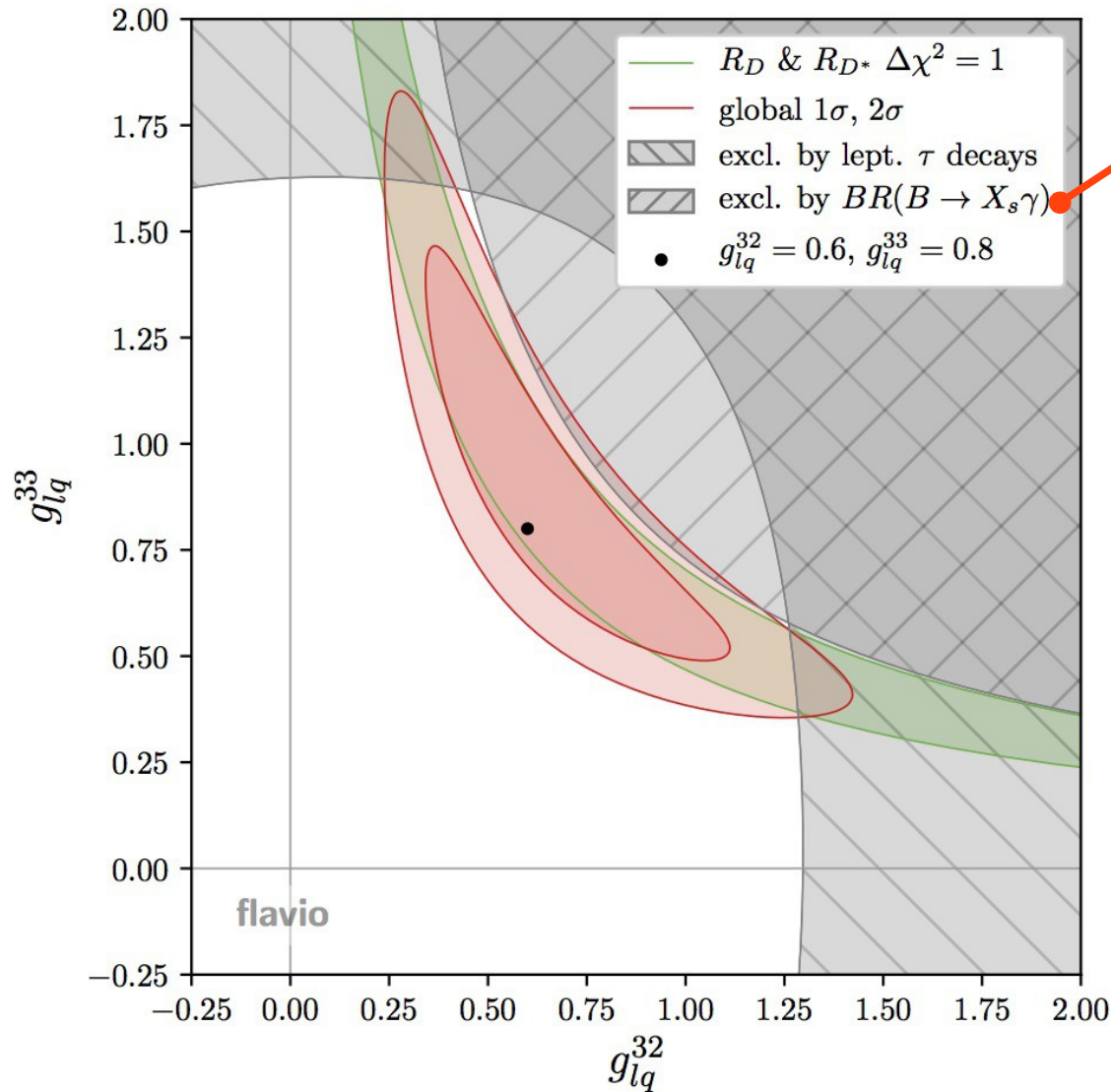


*Model-dependent constraint*

See discussion in

[Cornella-Fuentes-Isidori, 2019;  
Calibbi-Crivellin-Li, 2018;  
Bordone *et al.*, 2018]

# $U_1$ LQ: $g_{lq}^{32}$ vs. $g_{lq}^{33}$



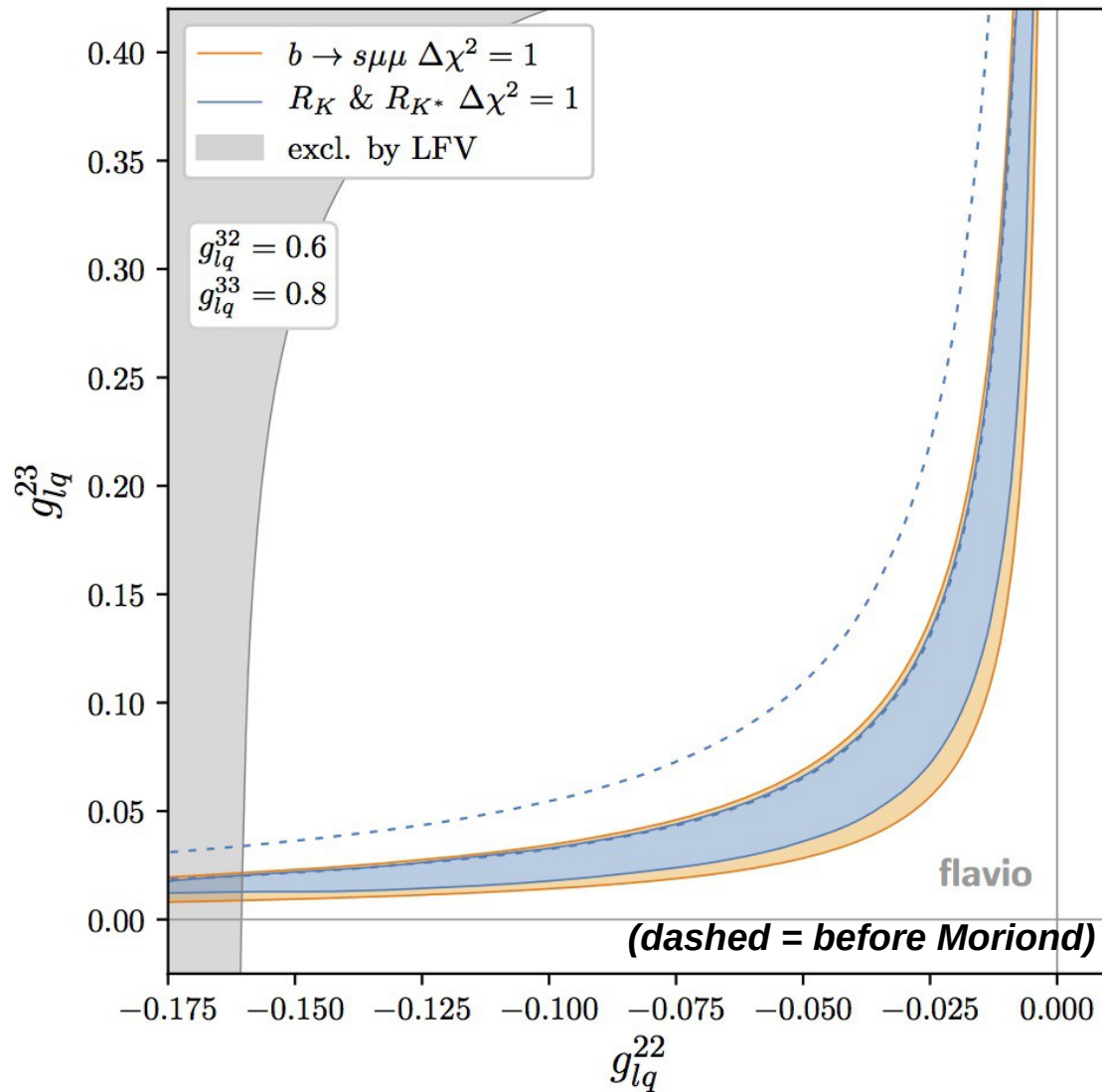
## Model-dependent constraint

See discussion in  
 [Cornella-Fuentes-Isidori, 2019;  
 Calibbi-Crivellin-Li, 2018;  
 Bordone et al., 2018]

- $R_{D^{(*)}}$  and  $\tau \rightarrow \ell \nu \nu$  select a non-trivial region
- We pick a benchmark point, then constrain the other two couplings



# $U_1$ LQ: $g_{lq}^{22}$ vs. $g_{lq}^{23}$



*The plane of muonic couplings shows that the picture works better after than before Moriond*

- *The  $R_{K^{(*)}}$  and  $b \rightarrow s \mu \mu$  regions now perfectly overlap*

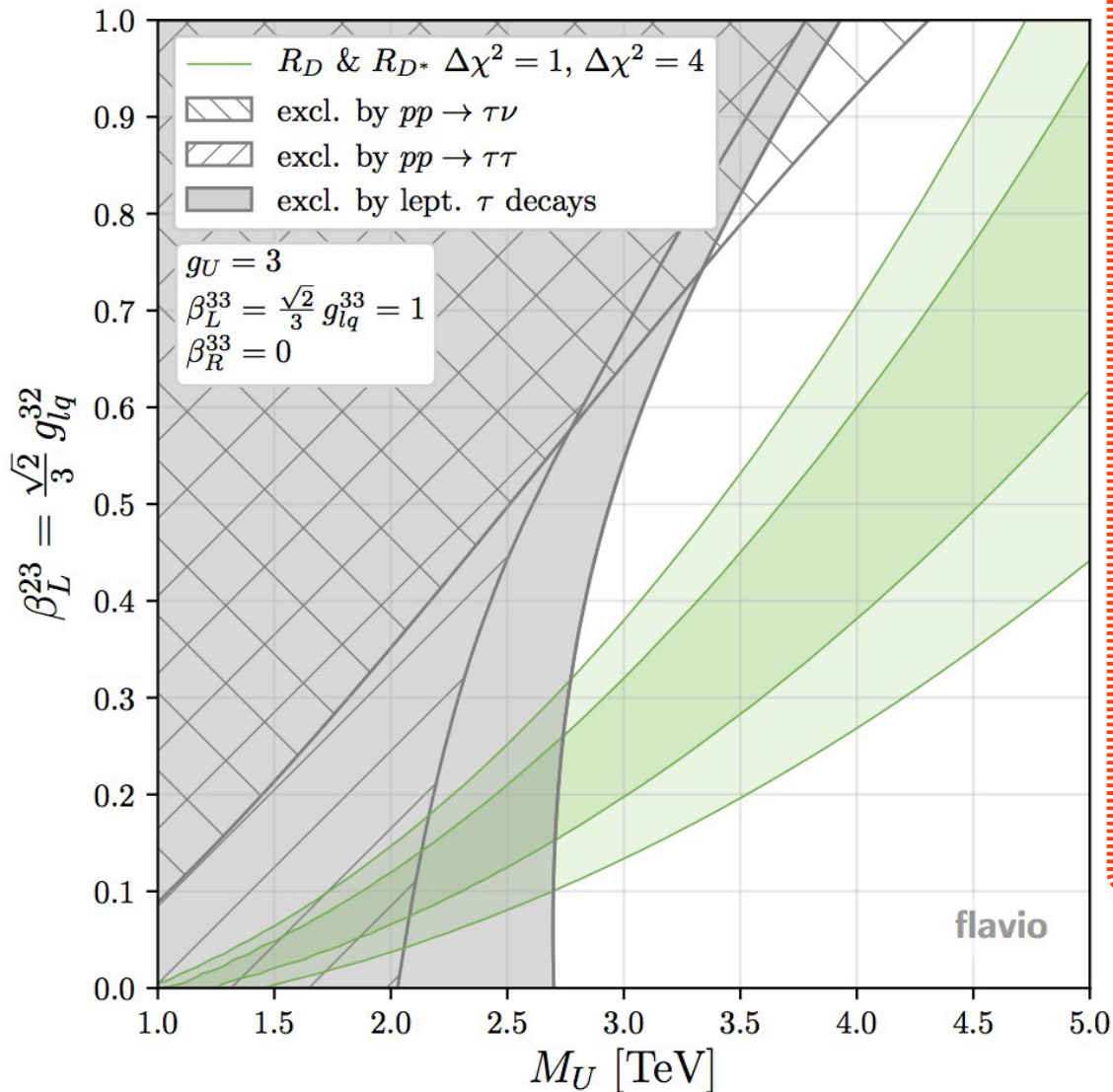
## $U_1$ LQ: direct constraints

*Aren't such tauonic couplings also constrained by direct searches?  
E.g.  $pp \rightarrow \tau\tau$  or  $\tau\nu$*

# $U_1$ LQ: direct constraints

Aren't such tauonic couplings also constrained by direct searches?

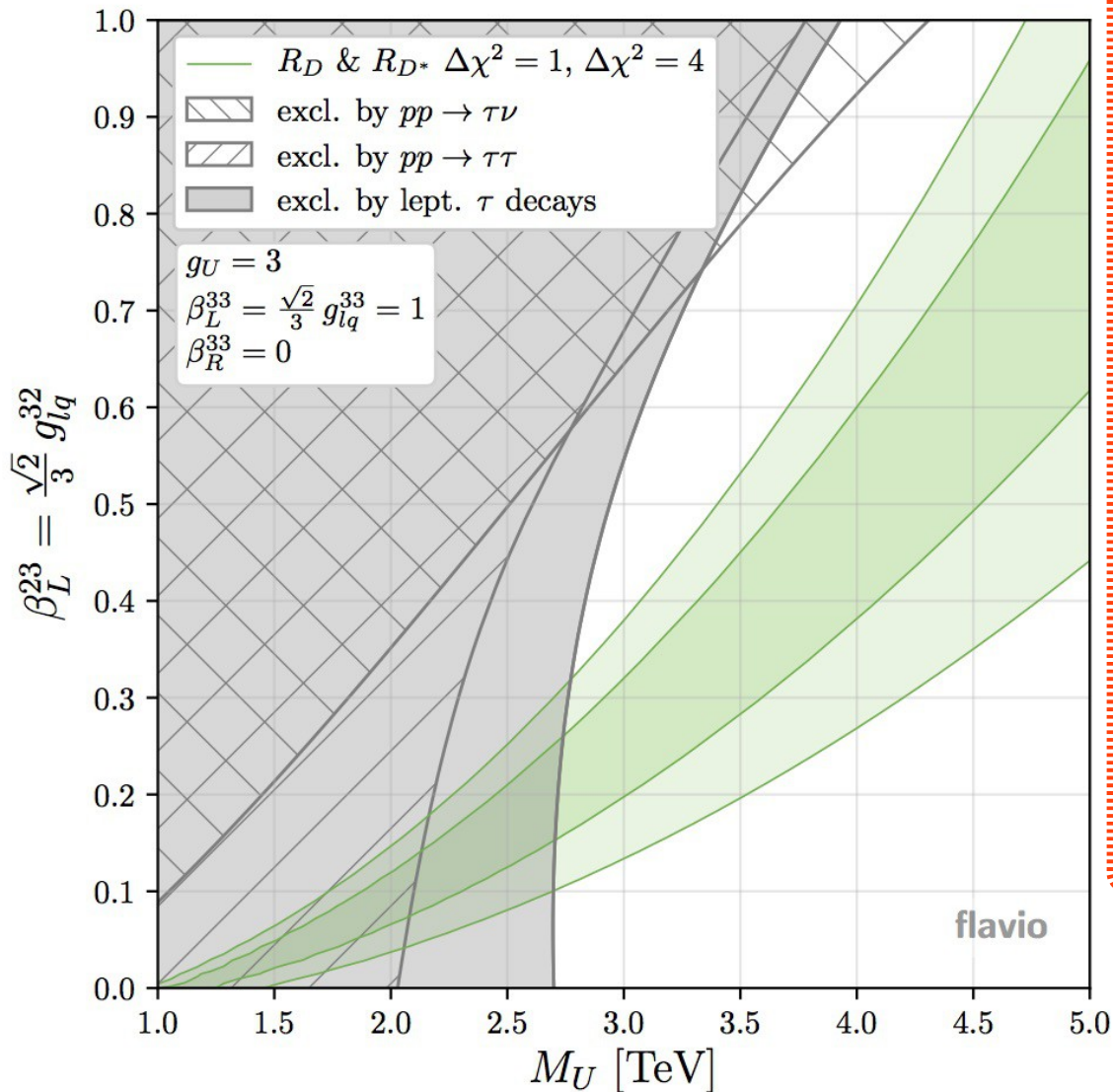
E.g.  $pp \rightarrow \tau\tau$  or  $\tau\nu$



For the sake of comparison, we tuned coupling values to those used in [Baker-Fuentes-Isidori-König]

# $U_1$ LQ: direct constraints

Aren't such tauonic couplings also constrained by direct searches?  
E.g.  $pp \rightarrow \tau\tau$  or  $\tau\nu$

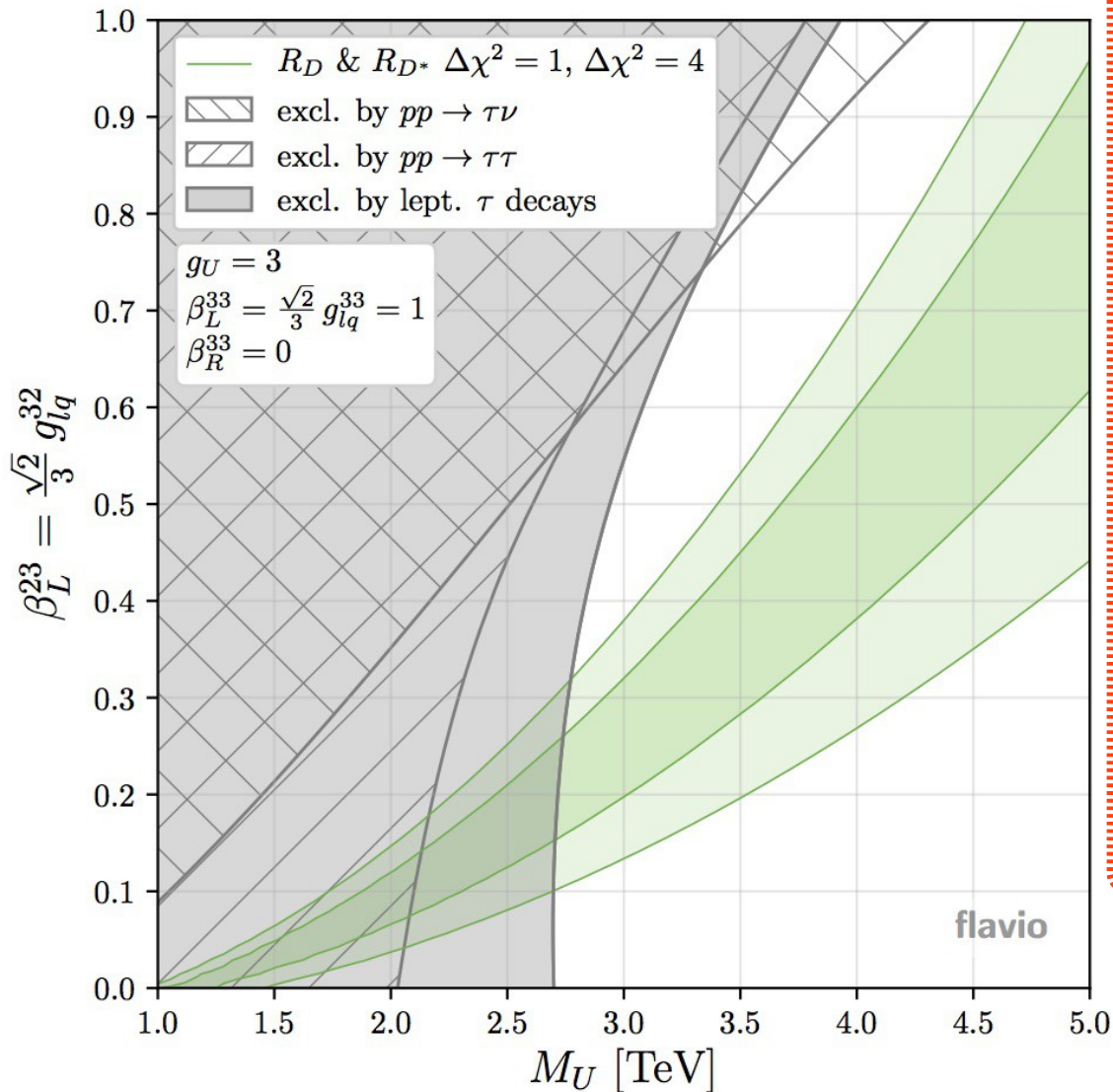


For the sake of comparison, we tuned coupling values to those used in [Baker-Fuentes-Isidori-König]

With  $g_{lq}^{3i} \neq 0$  (as required by  $R_{D^{(*)}}$ ) LUV constraints are stronger than direct ones

# $U_1$ LQ: direct constraints

Aren't such tauonic couplings also constrained by direct searches?  
 E.g.  $pp \rightarrow \tau\tau$  or  $\tau\nu$



For the sake of comparison,  
 we tuned coupling values to those  
 used in [Baker-Fuentes-Isidori-König]

With  $g_{lq}^{3i} \neq 0$  (as required by  $R_{D^{(*)}}$ )  
 LUV constraints are stronger  
 than direct ones

Lower bound  $M_U > 2.7$  TeV  
 due to the large couplings  
 chosen here (e.g.  $g_{lq}^{32} \approx 2.1$ )

I.e. it doesn't apply in general

## Conclusions

- *Semi-leptonic B-decay data still displays a clear preference for new effects in 4-f semi-leptonic operators w/ LH quarks*

## Conclusions

- *Semi-leptonic B-decay data still displays a clear preference for new effects in 4-f semi-leptonic operators w/ LH quarks*
- *The solution with muonic  $C_9 = -C_{10}$  now favoured over pure  $C_9$*   
*Welcome news from the standpoint of UV completions*



## Conclusions

- *Semi-leptonic B-decay data still displays a clear preference for new effects in 4-f semi-leptonic operators w/ LH quarks*
- *The solution with muonic  $C_9 = -C_{10}$  now favoured over pure  $C_9$*   
*Welcome news from the standpoint of UV completions*
- *Even better description of data obtained by allowing for additional  $C_9^{univ.}$*



## Conclusions

- *Semi-leptonic B-decay data still displays a clear preference for new effects in 4-f semi-leptonic operators w/ LH quarks*
- *The solution with muonic  $C_9 = -C_{10}$  now favoured over pure  $C_9$*   
*Welcome news from the standpoint of UV completions*
- *Even better description of data obtained by allowing for additional  $C_9^{univ.}$*
- *Interestingly, such effect is RG-generated from 4-f operators above the EW scale, in particular semi-tauonic ones, able to explain  $b \rightarrow c$  discrepancies*

## Conclusions

- *Semi-leptonic B-decay data still displays a clear preference for new effects in 4-f semi-leptonic operators w/ LH quarks*
- *The solution with muonic  $C_9 = -C_{10}$  now favoured over pure  $C_9$   
Welcome news from the standpoint of UV completions*
- *Even better description of data obtained by allowing for additional  $C_9^{univ.}$*
- *Interestingly, such effect is RG-generated from 4-f operators above the EW scale, in particular semi-tauonic ones, able to explain  $b \rightarrow c$  discrepancies*
- *Also interestingly, this whole picture finds a natural realization in the  $U_1$ -LQ model*