

# Proton decay and $n\bar{n}$ oscillations with SK(-Gd), DUNE, Hyper-K, and JUNO

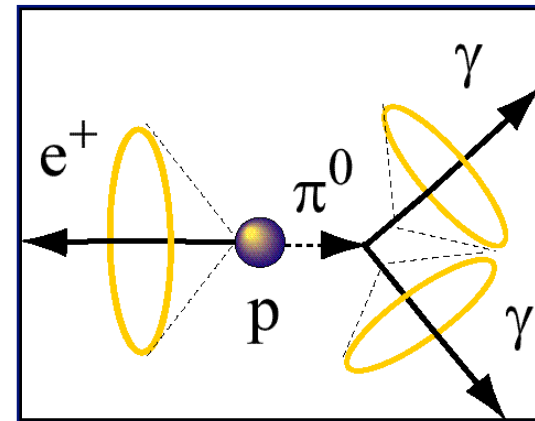
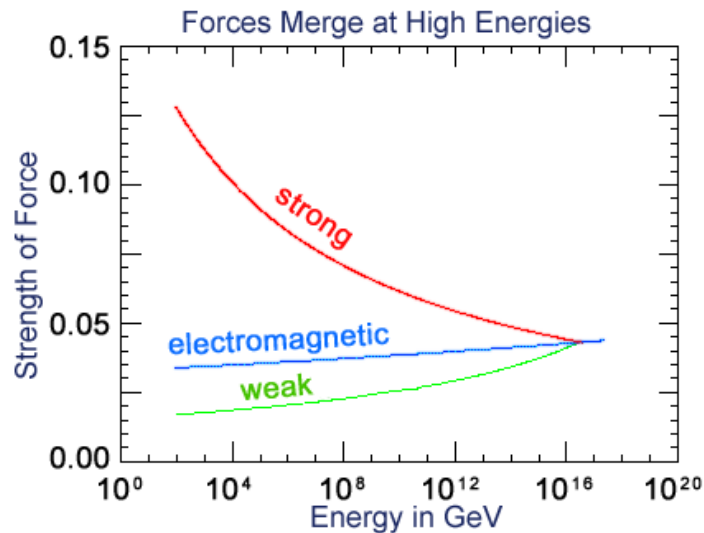
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*European Neutrino Town meeting @ CERN  
22-24 October 2018*

# Motivation

- Proton decay observation would prove the Grand Unification of elementary particles.
- Neutrino mass/mixings/CPV and proton decays could be related to each other at very high energy physics (GUTs).

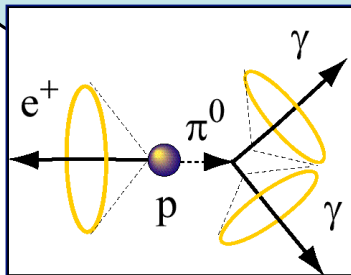
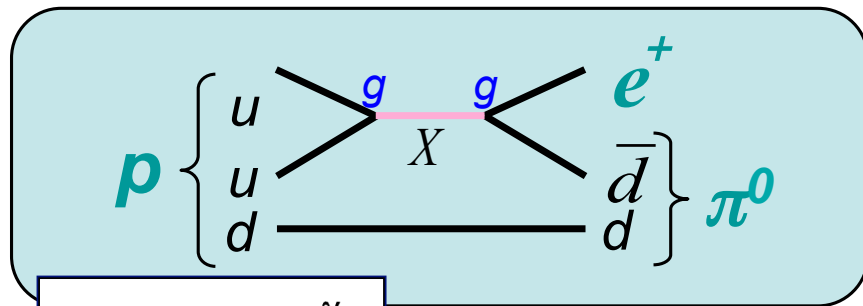


We are in an exciting era because large neutrino detectors (JUNO, DUNE, Hyper-K) are planned to start operation near future. They are also good proton decay detectors!

# Various proton decay modes

- Two prominent decay modes

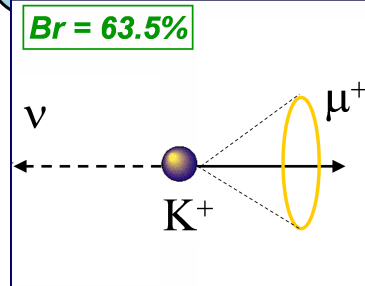
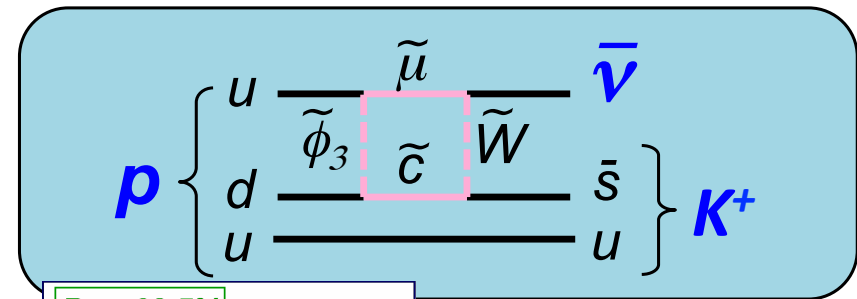
Mediated by gauge bosons



$p \rightarrow e^+ \pi^0$

$$\Gamma(p \rightarrow e^+ \pi^0) \sim \frac{g^4 m_p^5}{M_X^4}$$

SUSY mediated



$Br = 63.5\%$

$p \rightarrow \bar{\nu} K^+$

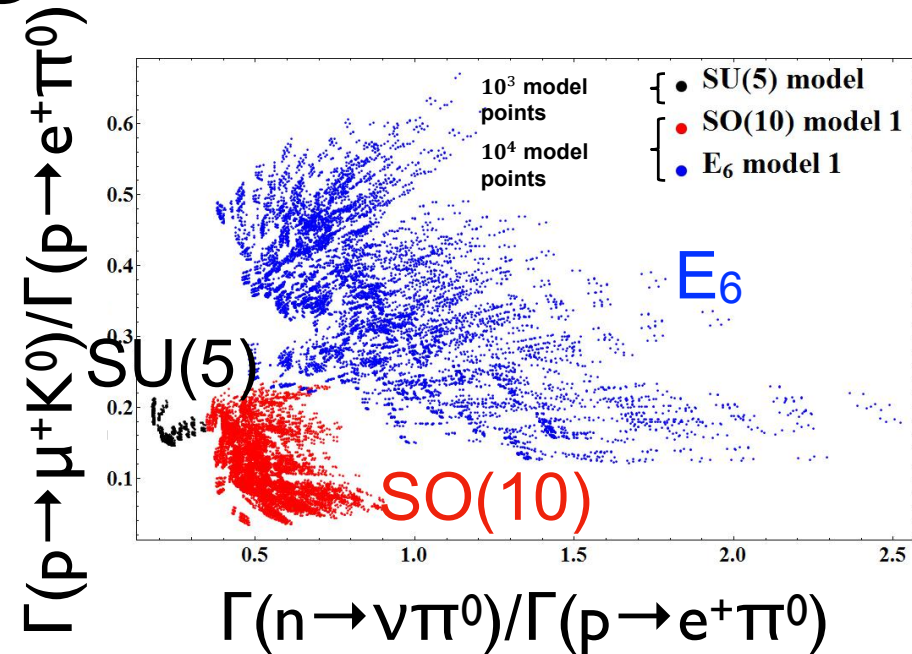
$$\Gamma(p \rightarrow \bar{\nu} K^+) \sim \frac{\tan^2 \beta \times m_p^5}{M_{\tilde{q}}^2 \times M_3^2}$$

- Variety of predictions and no SUSY in LHC

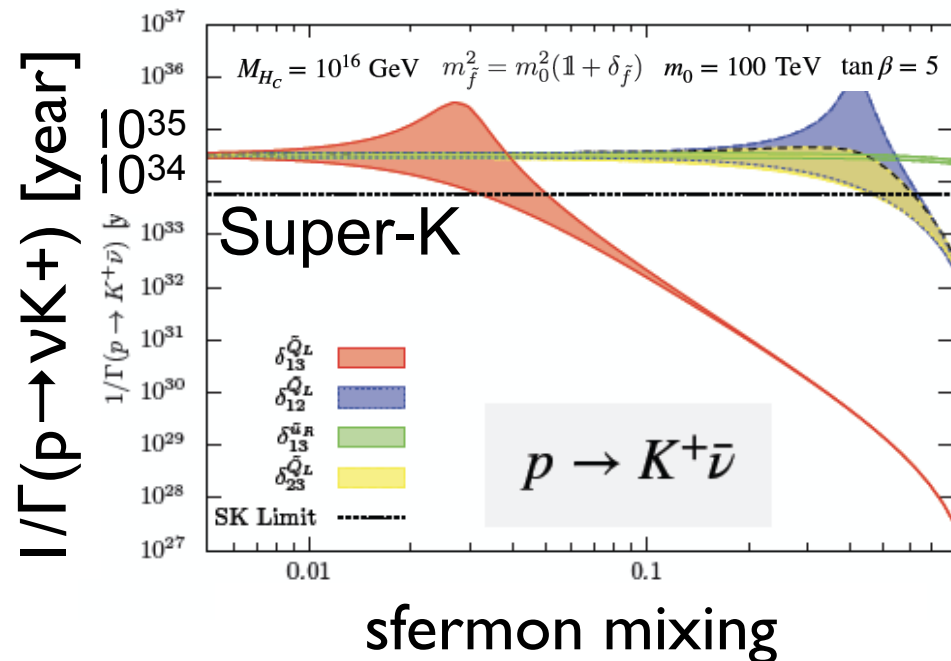
→ We need to experimentally pursue both and other decay modes

# Strong cases

- We could identify details of unification picture, e.g. gauge group and other symmetries
  - $\Gamma(n \rightarrow \nu \pi^0) / \Gamma(p \rightarrow e^+ \pi^0)$  depends on SU(5), SO(10), E<sub>6</sub> (Y. Muramatsu)



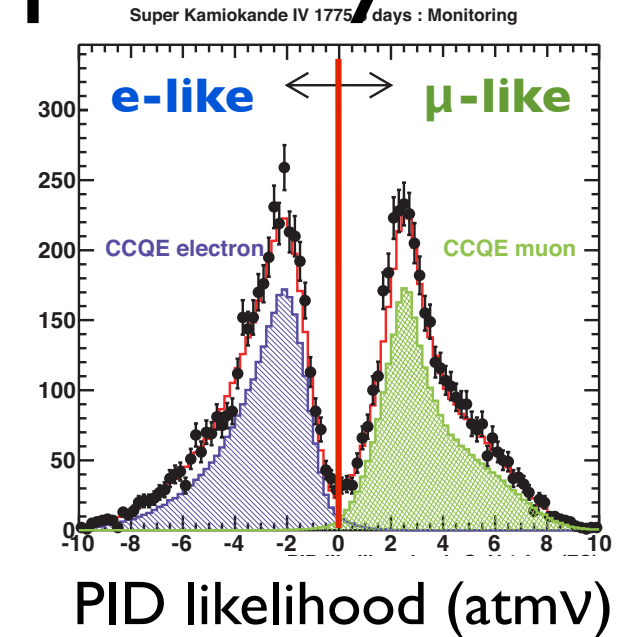
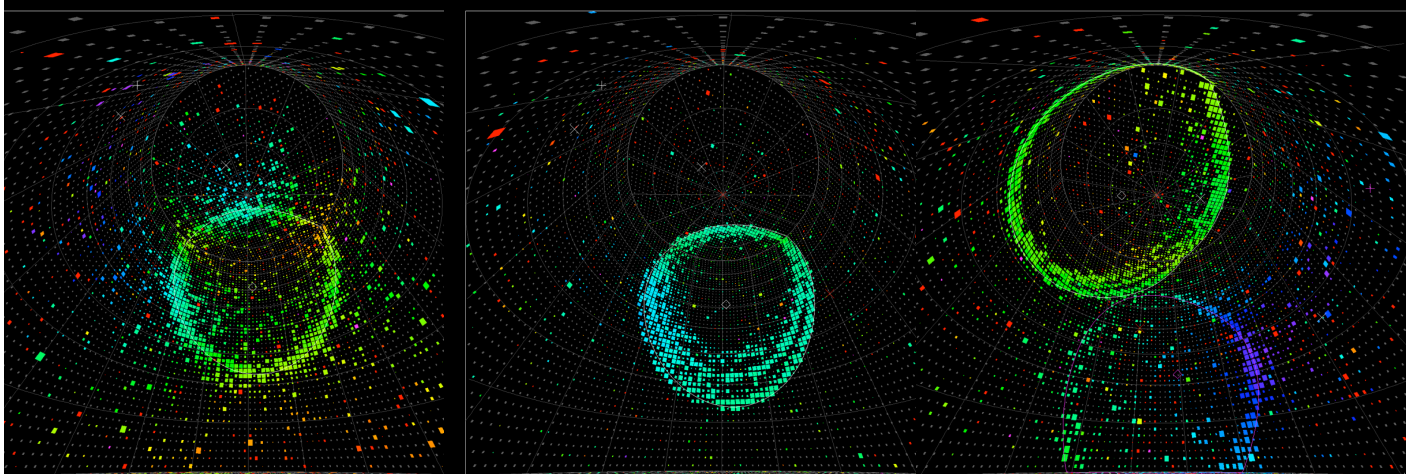
- P-decay Br. ratio could tell us flavor structure of SUSY particles.
  - Decay branches depends on the size of sfermion mixing. (N.Nagata and S.Shirai, JHEP 1403, 049 (2014))



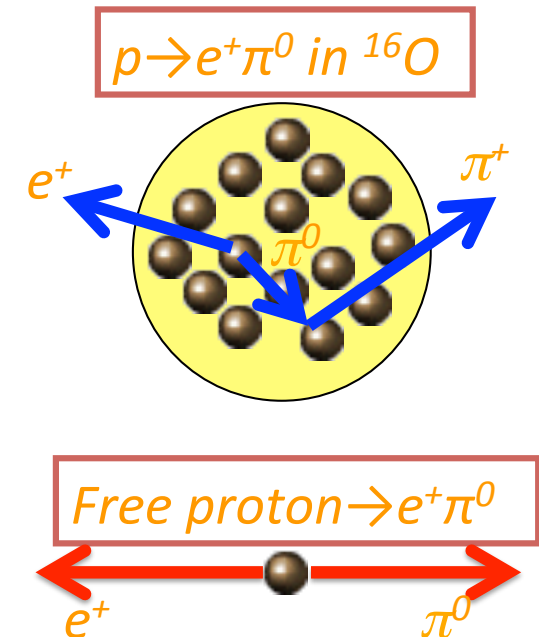
# Ring-imaging water Cherenkov detectors

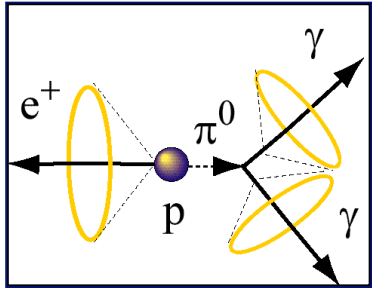
## Super-K / SK-Gd / Hyper-K

# Water Ch. detector for $p$ -decays



- **High mass is possible** (Super-K 22kton, Hyper-K 190kton  $\sim 8\times$ Super-K)
  - $p \rightarrow e^+ \pi^0$ ,  $\nu K^+$ , and **more** can be searched with high sensitivities
- **Excellent & well-proven performance**
  - Good ring-imaging capability at  $\sim 1$  GeV
  - Excellent particle ID (e or  $\mu$ ) capability  $> 99\%$
  - Energy resolution for e and  $\mu \sim 3\%$
- **Free protons are available**
  - No Fermi motion, nuclear effect
  - High efficiency & good S/N separation

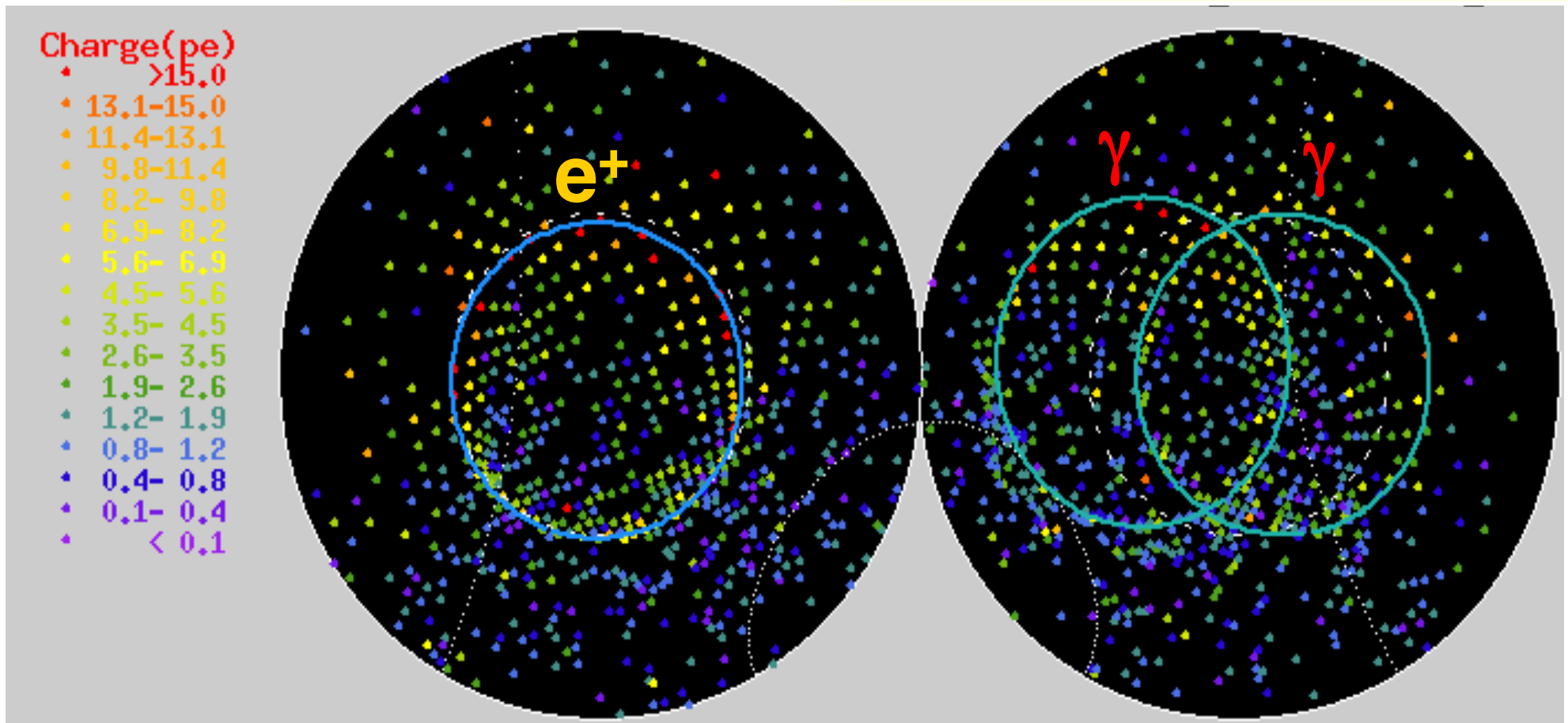




# $p \rightarrow e^+ + \pi^0$ searches

## Super-K cut

- 2 or 3 Cherenkov rings
- All rings are showering
- $85 < M_{\pi^0} < 185 \text{ MeV}/c^2$  (3-ring)
- No decay electron
- $800 < M_{\text{proton}} < 1050 \text{ MeV}/c^2$   
 $P_{\text{total}} < 250 \text{ MeV}/c$

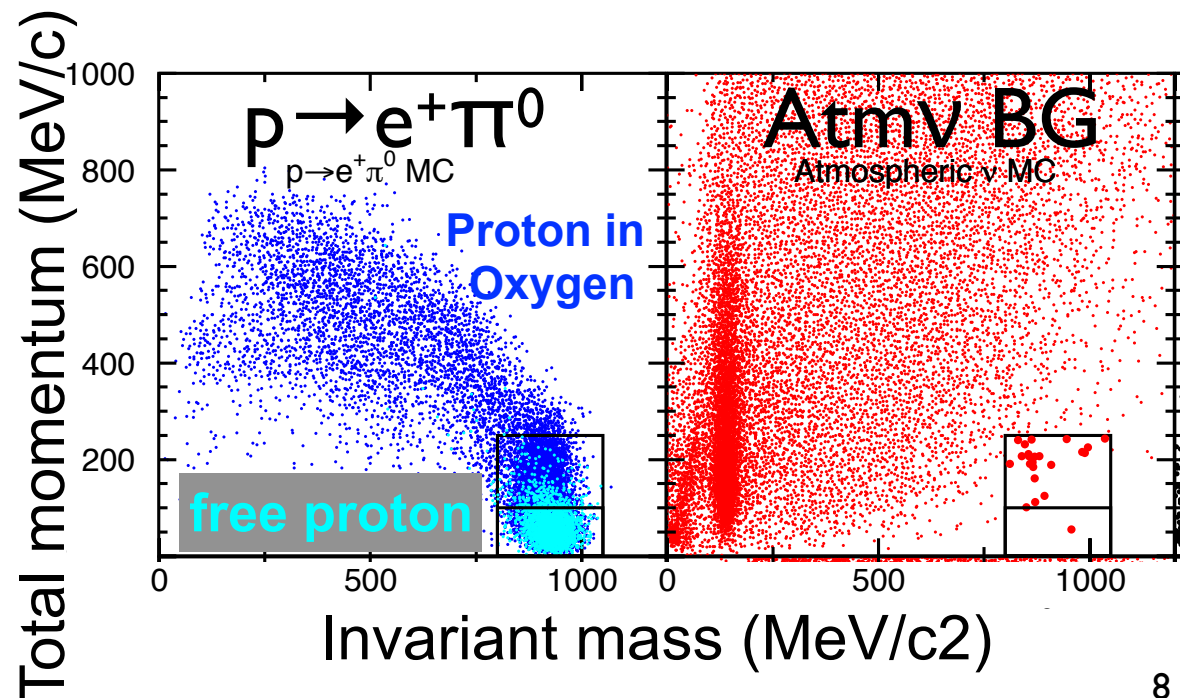
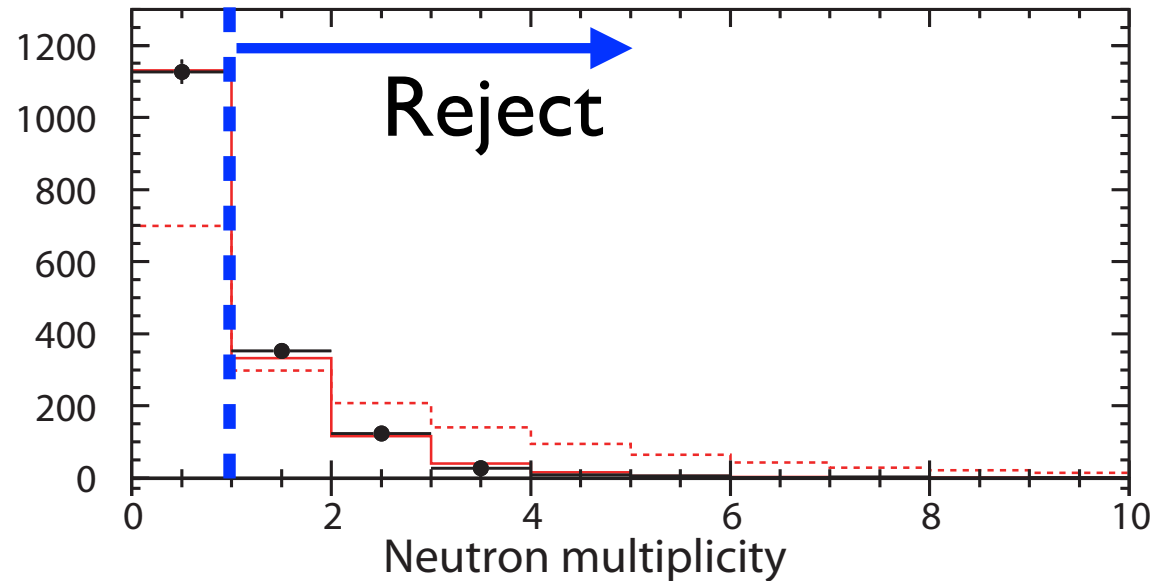


SK-II (half PMT) forward-backward display for  $p \rightarrow e^+ + \pi^0$  7

# BG reduction by neutron-tag & tighter $P_{\text{tot}}$ cut

- Shiozawa@NNN00 workshop  
 - PRD95, 012004 (2017)

- SK-IV w/ new electronics can tag neutrons by  $n+p \rightarrow d+2.2\text{MeV}\gamma$
- Atmospheric neutrino BG is reduced by 40%
- Two regions of  $P_{\text{tot}}$  to enhance discovery reach
  - $P_{\text{tot}} < 100\text{MeV}/c$  for free proton decays
  - $P_{\text{tot}} < 250\text{ MeV}/c$  for  $^{16}\text{O}$

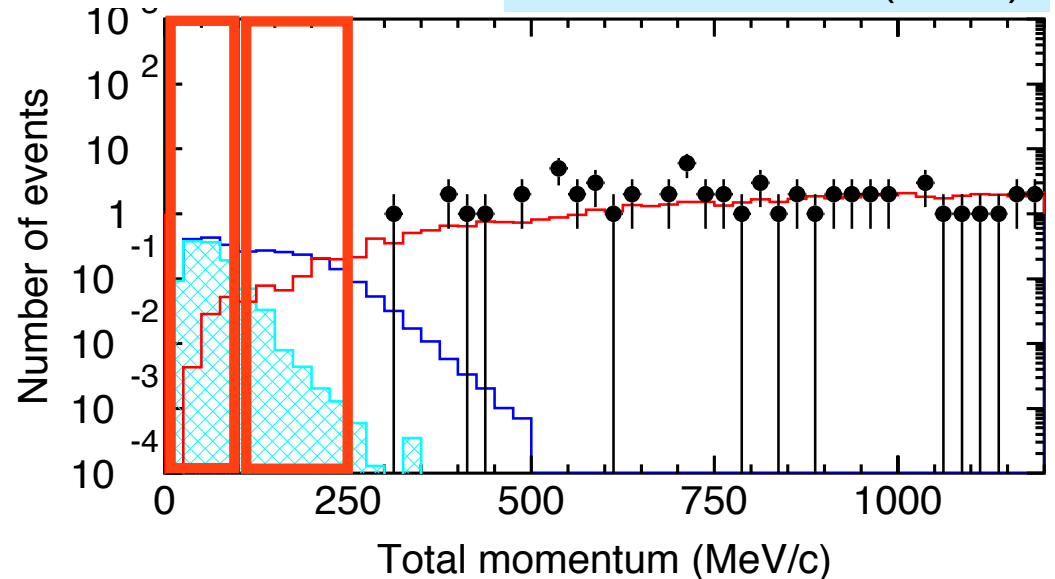
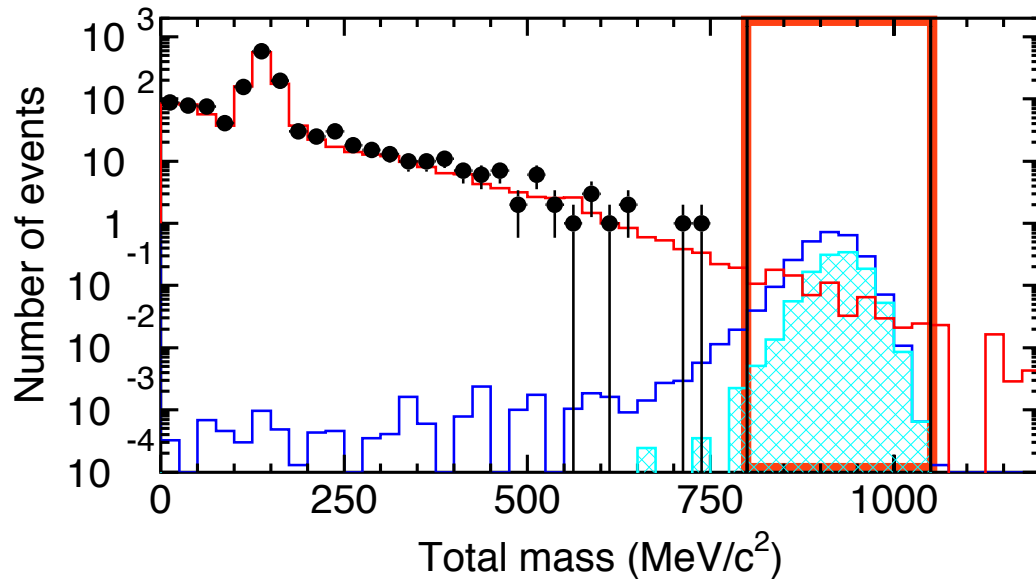


$p_{\text{tot}} < 100\text{MeV}/c$		$100 < p_{\text{tot}} < 250\text{MeV}/c$	
Sig. $\epsilon$ (%)	Bkg (/Mtyr)	Sig. $\epsilon$ (%)	Bkg (/Mtyr)
18.7	0.18	19.4	1.1



# $p \rightarrow e^+ + \pi^0$ in SK I-IV (306kt×yrs)

PRD95, 012004 (2017)



$p \rightarrow e^+ + \pi^0$

- detection efficiency = 40%

- atmospheric  $\nu$  BG = 0.61 (0.05+0.55) events in 306kton×yr

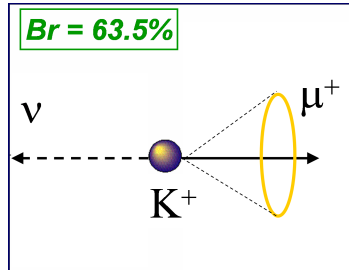
$\tau_{\text{proton}}/\text{Br} > 1.6 \times 10^{34}$  years @ 90% C.L.

► There is still discovery potential by keeping the low BG (especially for  $P_{\text{tot}} < 100 \text{ MeV/c}$ )

# $p \rightarrow \nu + K^+$ searches in SK

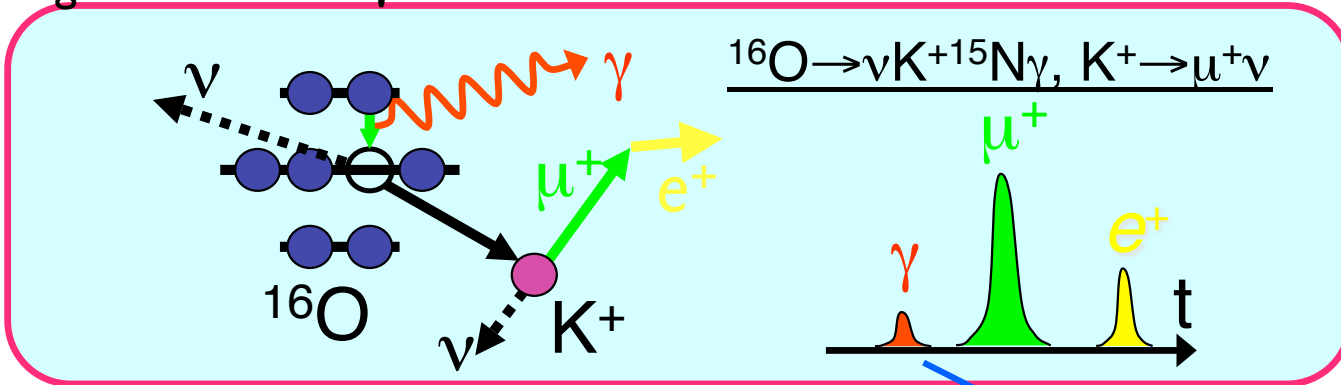
PRD90, 072005 (2014)

## (I) $K^+ \rightarrow \mu^+ \nu$ , $\mu^+ \rightarrow e^+ \nu \nu$



$K^+$  is below Cherenkov threshold  
 $\rightarrow 236 \text{ MeV}/c$   $\mu$  and muon decay electrons

Tag de-excitation  $\gamma$  from  $^{15}\text{N}^*$  to reduce BG

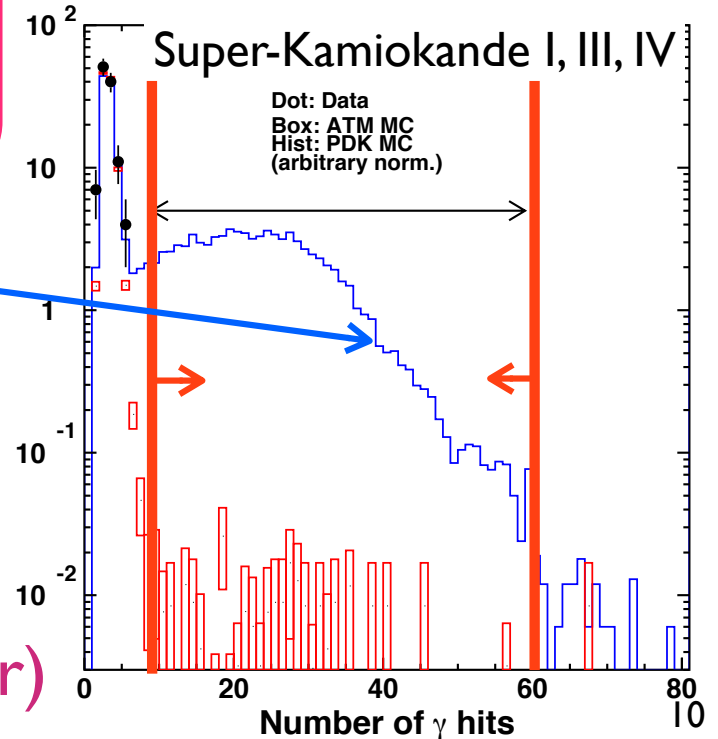
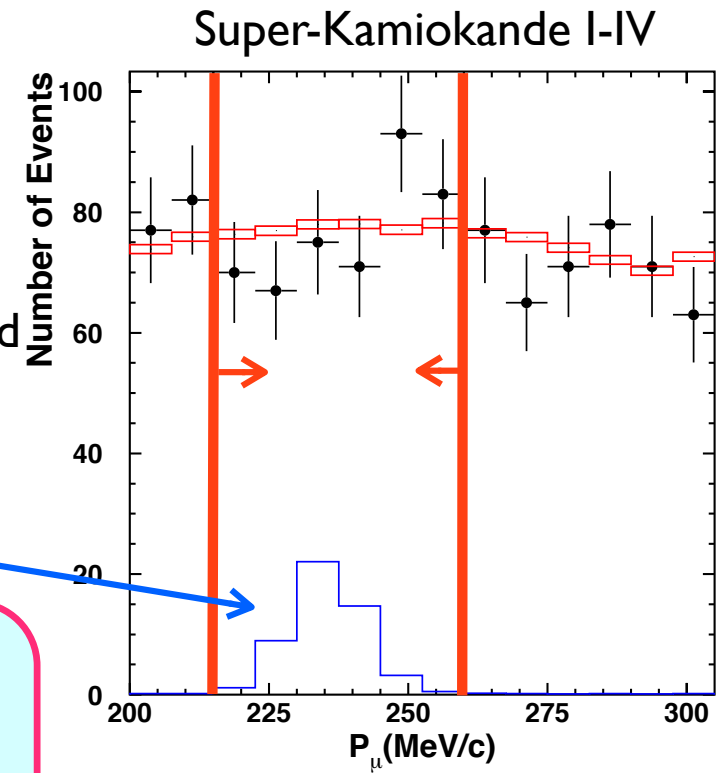


- signal efficiency = 9.1% (for SK-IV)
- Atmospheric neutrino BG 0.38/260ktyr

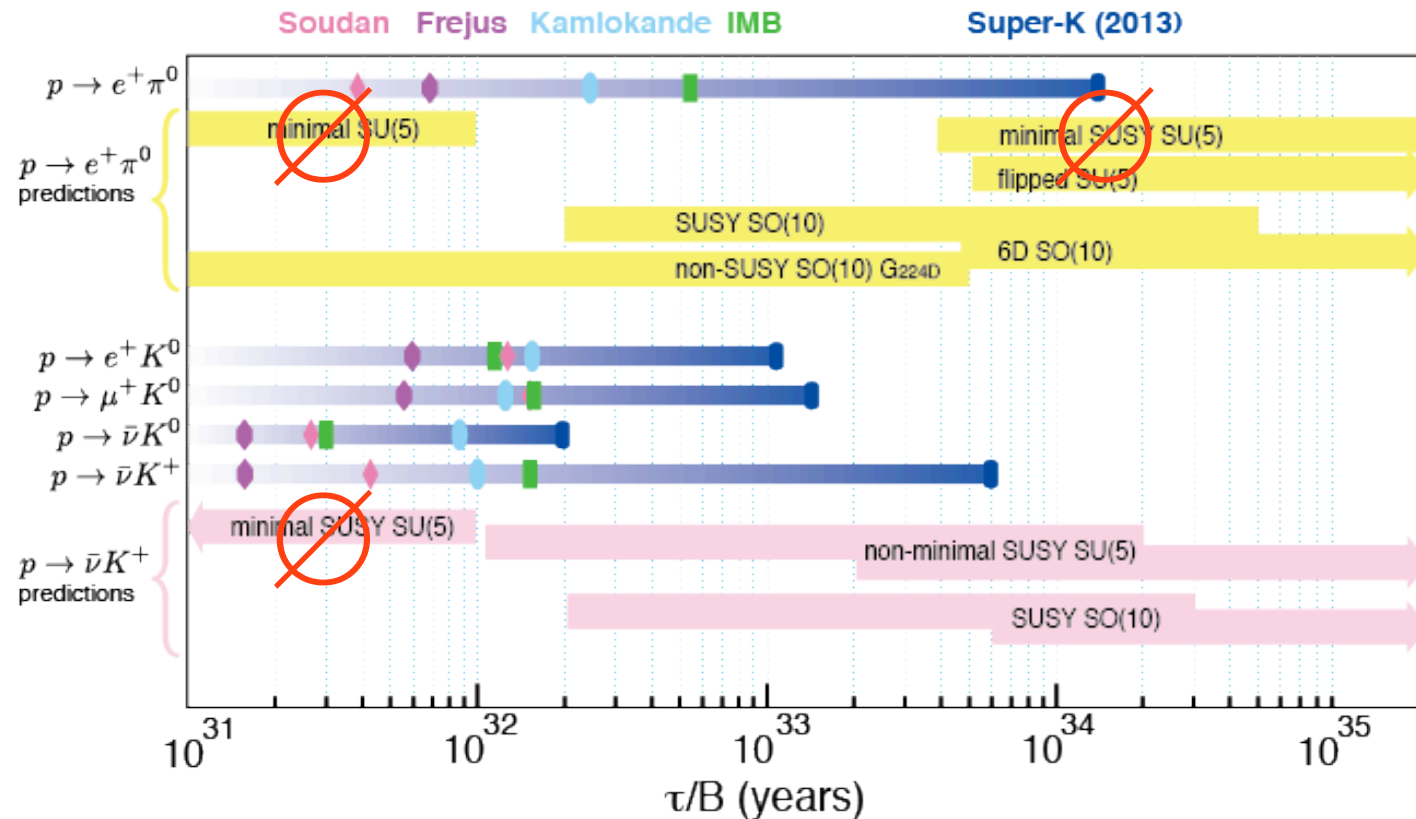
## (II) $K^+ \rightarrow \pi^+ \pi^0$

- signal efficiency = 10.0% (for SK-IV)
- Atmospheric neutrino BG 0.62/260ktyr

$\tau_{\text{proton}}/\text{Br} > 5.9 \times 10^{33} \text{ years @ 90\%CL (260ktyr)}$



# Experimental limits and models



## ▶ Super-K provides world stringent limits

▶  $\tau(p \rightarrow e^+\pi^0) > 1.6 \times 10^{34}$  years (90% C.L.)

▶  $\tau(p \rightarrow \nu K^+) > 5.9 \times 10^{33}$  years (90% C.L.)

▶ **already constrain the construction of many GUT models**

PRD 95, 012004 (2017)

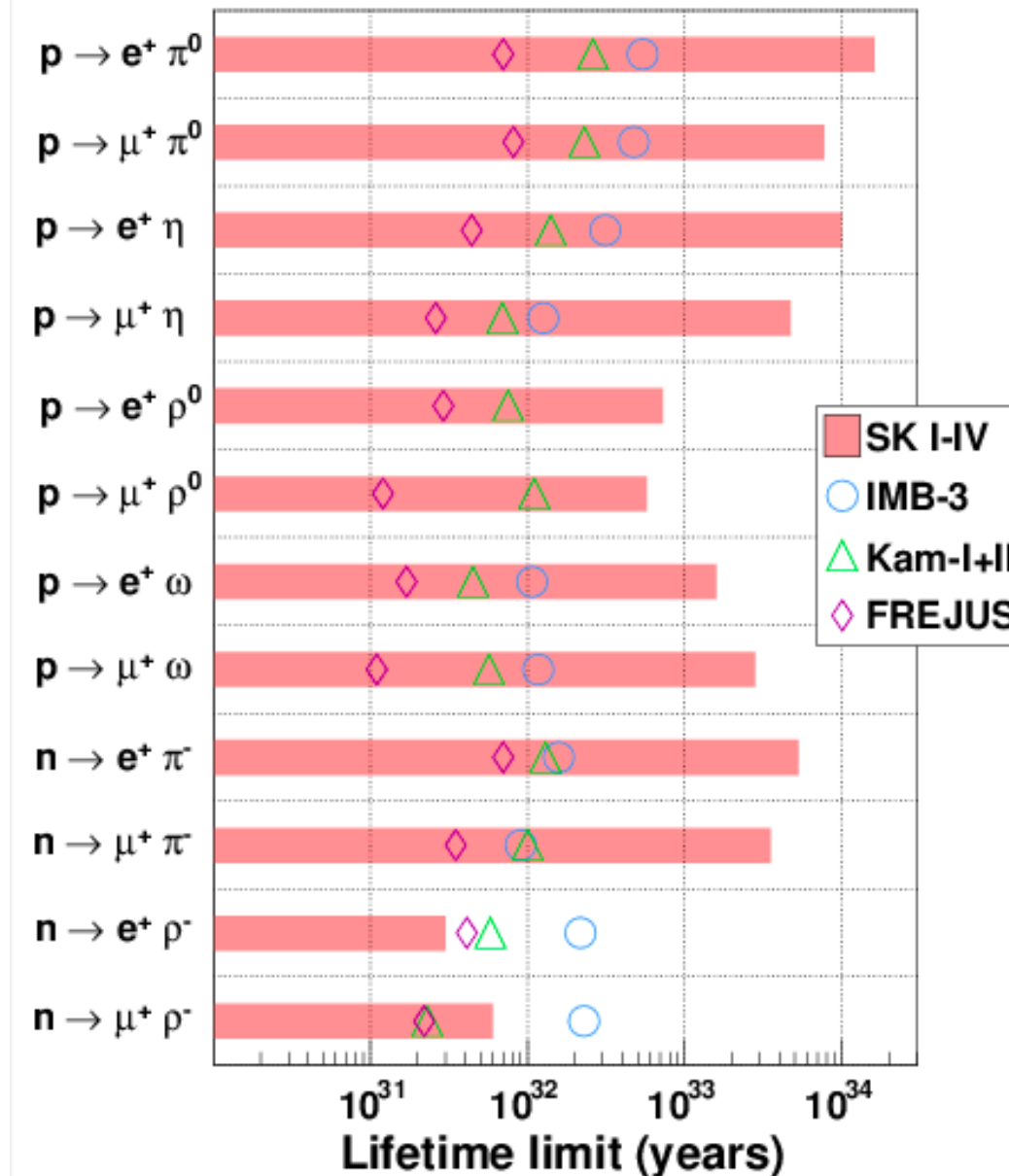
PRD 90, 072005 (2014)

## ▶ Many models predict $\tau = O(10^{34-35})$ years

▶ **Discovery could be around corner!**

# charged-lepton + meson decay

PRD 96, 012003 (2017)

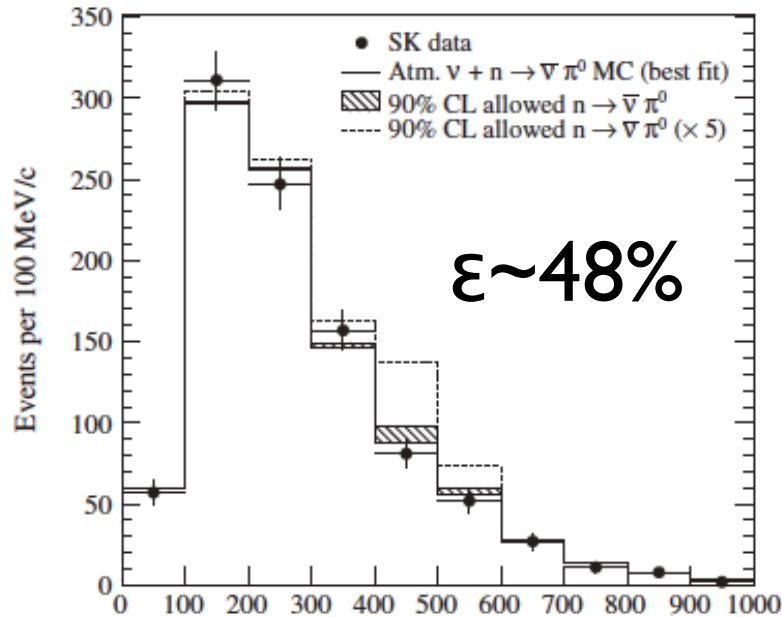


Modes	Eff.(%)		BG (SK-I+II)		Candidates SK-I+II
	SK-I	SK-II	NEUT	(NUANCE)	
$p \rightarrow e^+ \pi^0$	44.6	43.5	0.31	(0.27)	0
$p \rightarrow \mu^+ \pi^0$	35.5	34.7	0.34	(0.27)	0
$p \rightarrow e^+ \eta$	18.8	18.2	0.28	(0.29)	0
$p \rightarrow \mu^+ \eta$	12.4	11.7	0.04	(0.04)	0
$p \rightarrow e^+ \rho^0$	4.9	4.2	0.35	(0.34)	0
$p \rightarrow \mu^+ \rho^0$	1.8	1.5	0.42	(0.46)	1
$p \rightarrow e^+ \omega$	2.4	2.2	0.14	(0.29)	0
$p \rightarrow \mu^+ \omega$	2.8	2.8	0.31	(0.37)	0
$n \rightarrow e^+ \pi^-$	19.4	19.3	0.27	(0.37)	0
$n \rightarrow \mu^+ \pi^-$	16.7	15.6	0.43	(0.44)	1
$n \rightarrow e^+ \rho^-$	1.8	1.6	0.38	(0.44)	1
$n \rightarrow \mu^+ \rho^-$	1.1	0.94	0.29	(0.69)	0

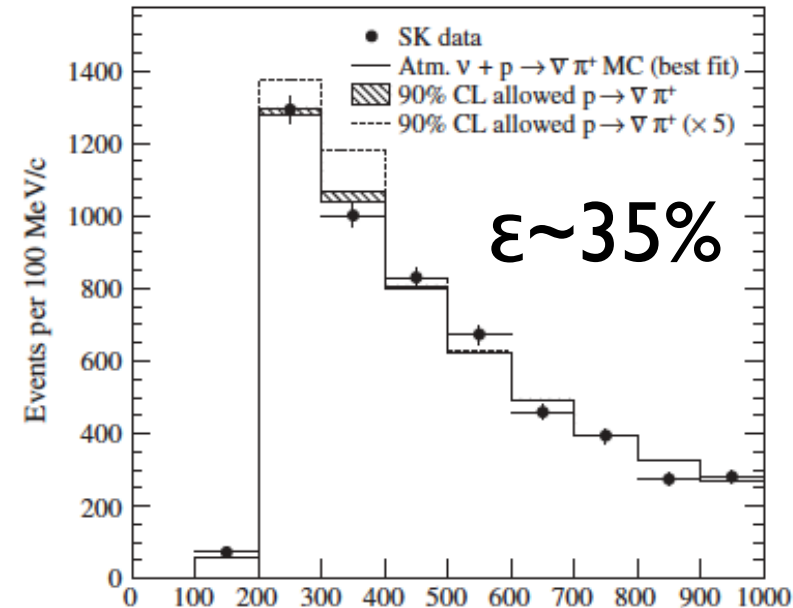
- No evidence is found in Super-K

# Searches for $n \rightarrow \nu \pi^0$ and $p \rightarrow \nu \pi^+$

- Motivated by minimal SUSY SO(10) (e.g. PLB587, 105 (2004).)



$\pi^0$  momentum (MeV/c)



$\pi^+$  momentum (MeV/c)

## $p \rightarrow \nu \pi^0$

- detection efficiency = 49%(SK-I), 44%(SK-II), 49%(SK-III)
- $\tau_p > 1.1 \times 10^{33}$  years @ 90%C.L.

## $p \rightarrow \nu \pi^+$

- detection efficiency = 35%(SK-I), 35%(SK-II), 36%(SK-III)
- $\tau_p > 3.9 \times 10^{32}$  years @ 90%C.L.

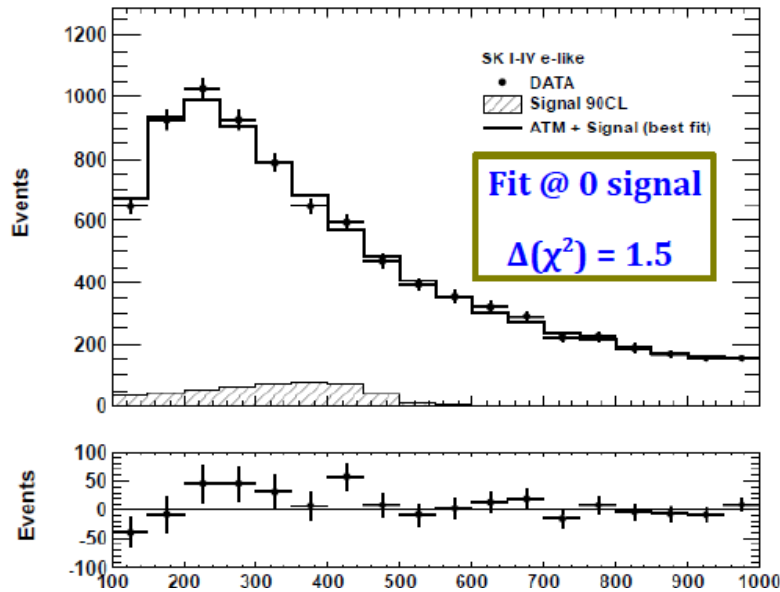
almost ruled out the prediction by PLB587, 105 (2004):

- $\tau_{p \rightarrow \nu \pi^0} = 2\tau_{p \rightarrow \nu \pi^+} < 5.7 - 13 \times 10^{32}$  years @ 90%C.L.

# Test of excess in e/ $\mu$ spectrum

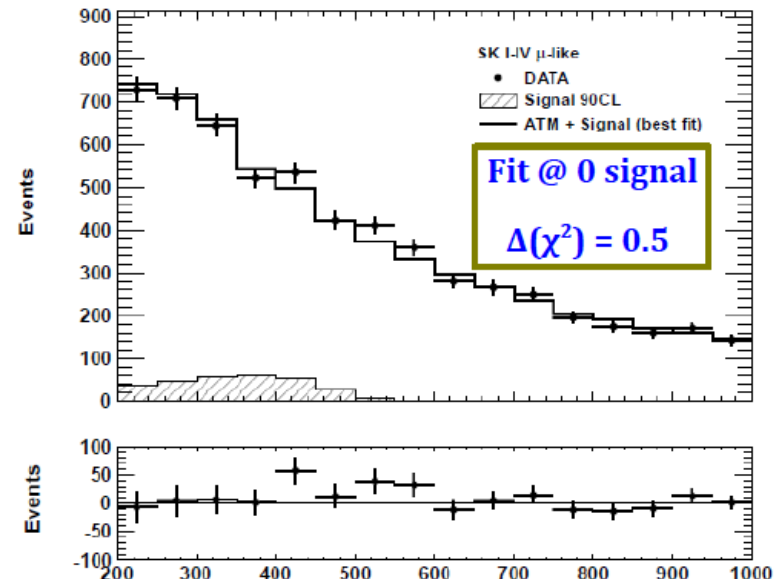
$p \rightarrow e\nu\nu$

SK I-IV



$p \rightarrow \mu\nu\nu$

SK I-IV



electron momentum (MeV/c)

muon momentum (MeV/c)

Mode	SK I-IV Sensitivity (years)	SK I-IV Limit (years)	PDG Limit (years)
$p \rightarrow e^+\nu\nu$	$2.7 \cdot 10^{32}$	$1.7 \cdot 10^{32}$	$1.7 \cdot 10^{31}$
$p \rightarrow \mu^+\nu\nu$	$2.5 \cdot 10^{32}$	$2.2 \cdot 10^{32}$	$2.1 \cdot 10^{31}$

PRL113,101801(2014)

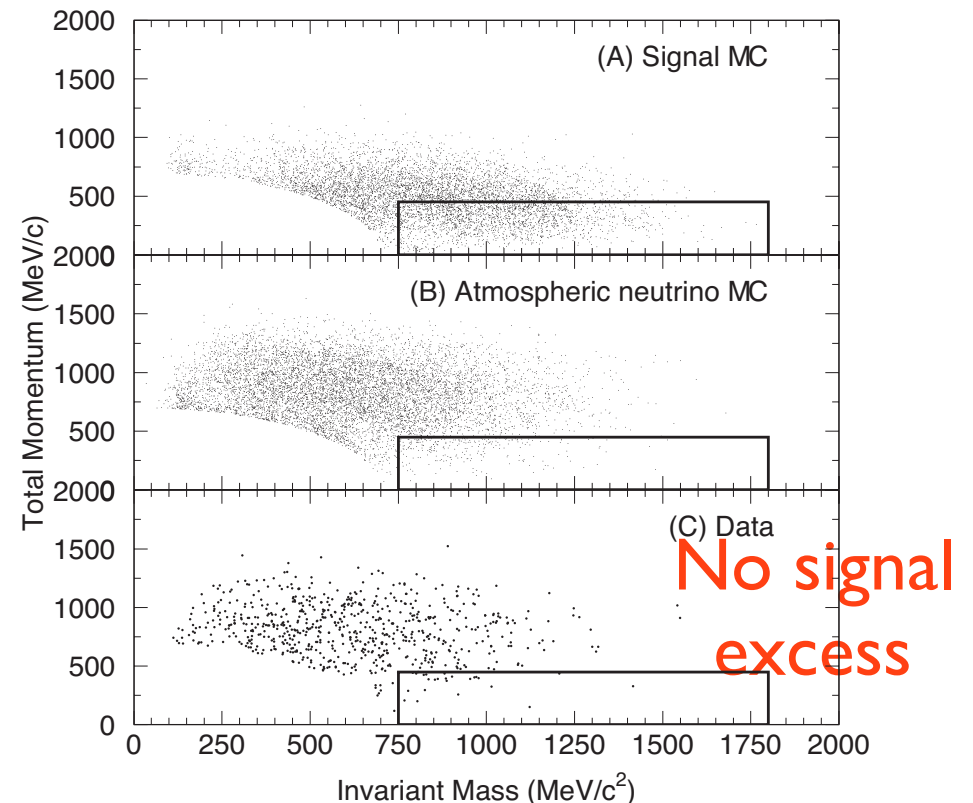
Mode	SK I-IV Sensitivity (years)	SK I-IV Limit (years)	PDG Limit (years)
$p \rightarrow e^+X$	$7.9 \cdot 10^{32}$	$7.9 \cdot 10^{32}$	—
$p \rightarrow \mu^+X$	$7.7 \cdot 10^{32}$	$4.1 \cdot 10^{32}$	—
$n \rightarrow \nu\gamma$	$5.8 \cdot 10^{32}$	$5.5 \cdot 10^{32}$	$2.8 \cdot 10^{31}$
$np \rightarrow e^+\nu$	$9.9 \cdot 10^{31}$	$2.6 \cdot 10^{32}$	$2.8 \cdot 10^{30}$
$np \rightarrow \mu^+\nu$	$1.1 \cdot 10^{32}$	$2.2 \cdot 10^{32}$	$1.6 \cdot 10^{30}$
$np \rightarrow \tau^+\nu$	$1.1 \cdot 10^{31}$	$2.9 \cdot 10^{31}$	—

# $\Delta B=2$ : $n$ - $\bar{n}$ oscillation

PRD91,072006(2015)

- $\Delta B=\Delta(B-L)=2$ , might be relevant for the matter asymmetry in the Universe
- look for multiple pions from  $n$ - $\bar{n}$  annihilation

- $\geq 2$  Cherenkov rings
- $700 < \text{Visible Energy} < 1300$  MeV
- $750 < M_{\text{tot}} < 1800$  MeV/ $c^2$
- $P_{\text{tot}} < 450$  MeV/ $c$



## $n$ - $\bar{n}$ oscillation in $^{16}\text{O}$

- detection efficiency = 12.1%
- atmospheric  $\nu$  BG = 24.1 events in 92kton $\times$ years (Super-K-I)
- observed signal = 24 events
- $T_{n-\bar{n}}(^{16}\text{O}) > 1.9 \times 10^{32}$  years @ 90% C.L.  
→  $T_{n-\bar{n}}(\text{free}) > 2.7 \times 10^8$  sec

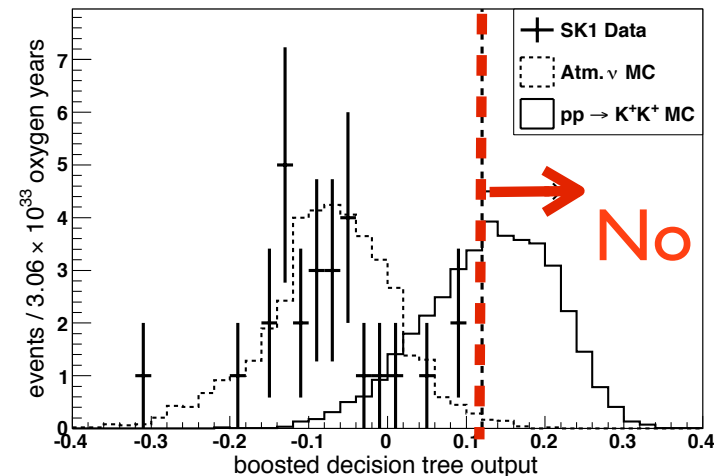
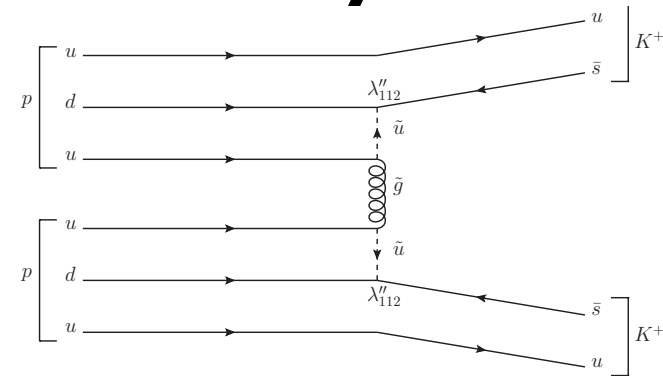
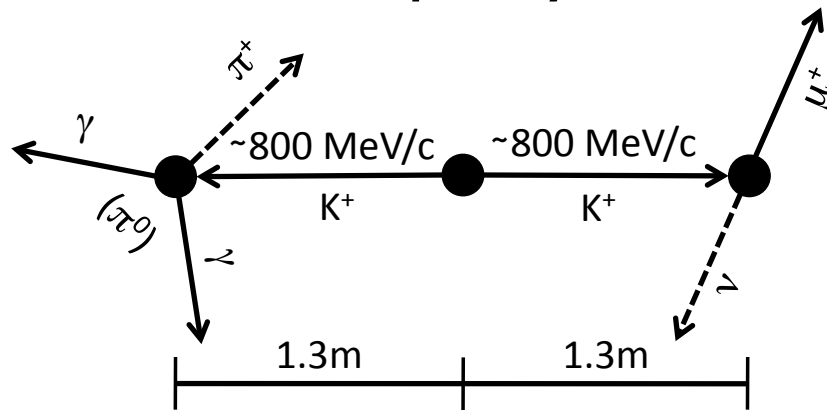
Machine learning analysis is also performed.

# $\Delta B=2$ : Dinucleon decays



- Boosted Decision Tree (BDT) by using vertex separation,  $K^+$  momentum etc.

- sensitive to R-parity-violating  $\lambda''_{112}$



- detection efficiency =  $12.6\% \pm 3.2\%$
- atmospheric  $\nu$  BG =  $0.28 \pm 0.19$  events in  $92\text{kton} \times \text{years}$  (Super-K-I)
- $\tau(^{16}\text{O}(pp) \rightarrow ^{14}\text{C}K^+K^+) > 1.7 \times 10^{32}$  years @ 90% C.L.  
 $\rightarrow \lambda''_{112} < 7.8 \times 10^{-9}$

We also searched for  $pp \rightarrow \pi^+\pi^+$ ,  $pn \rightarrow \pi^+\pi^0$ ,  $n \rightarrow \pi^0\pi^0$



# Super-K: next phase

- 2018: Tank refurbishment work to be ready for loading Gadolinium

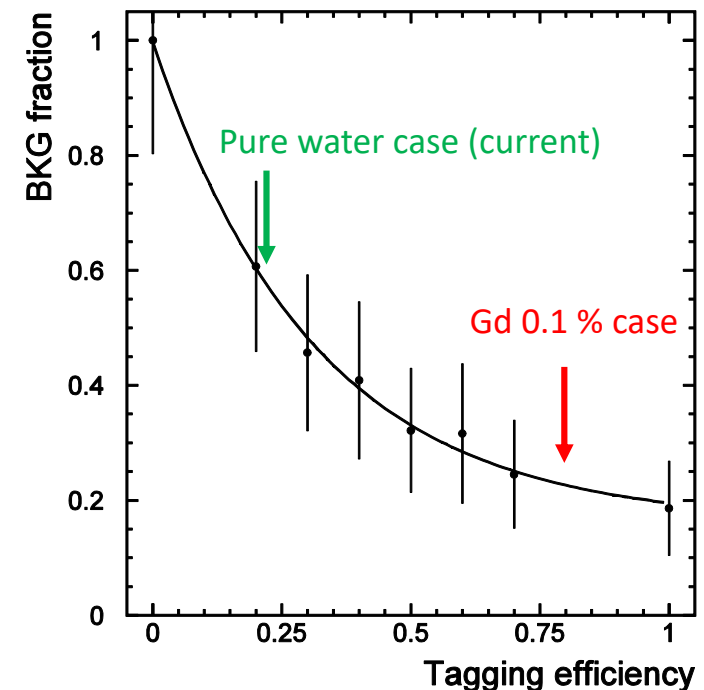


- 2020(TBD): Neutron capture 50% (0.01%w Gd) → 202X: 90% (0.1%w Gd)
- Discovery of supernova relic neutrinos ( $\bar{\nu}_e$ ) by coincidence of  $e^+ + \gamma$ s

Beacom and Vagins PRL93:171101(2004)

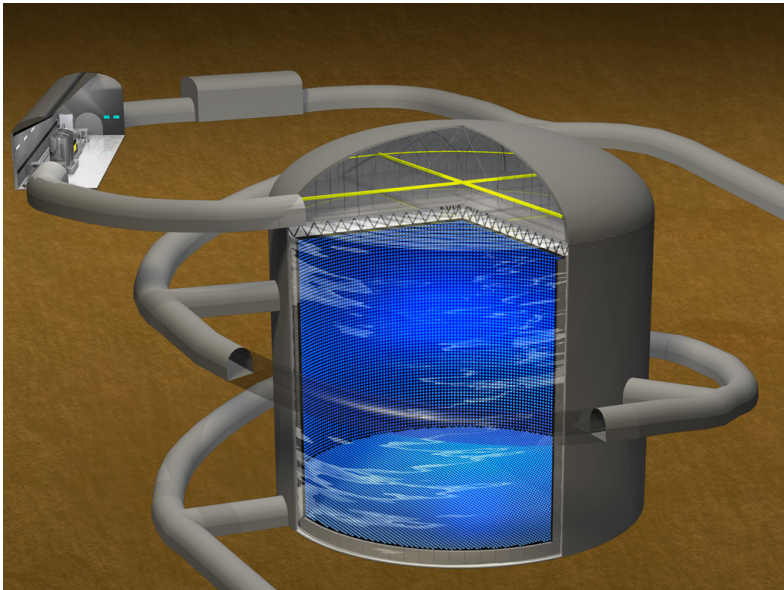
## For proton decay searches

- Further BG rate reduction by a factor of 2 for  $p \rightarrow e + \pi^0$
- Better (cleaner) searches can be performed.
- It provides opportunity to study neutron production by beam/ atmospheric neutrinos for future proton decay search program.

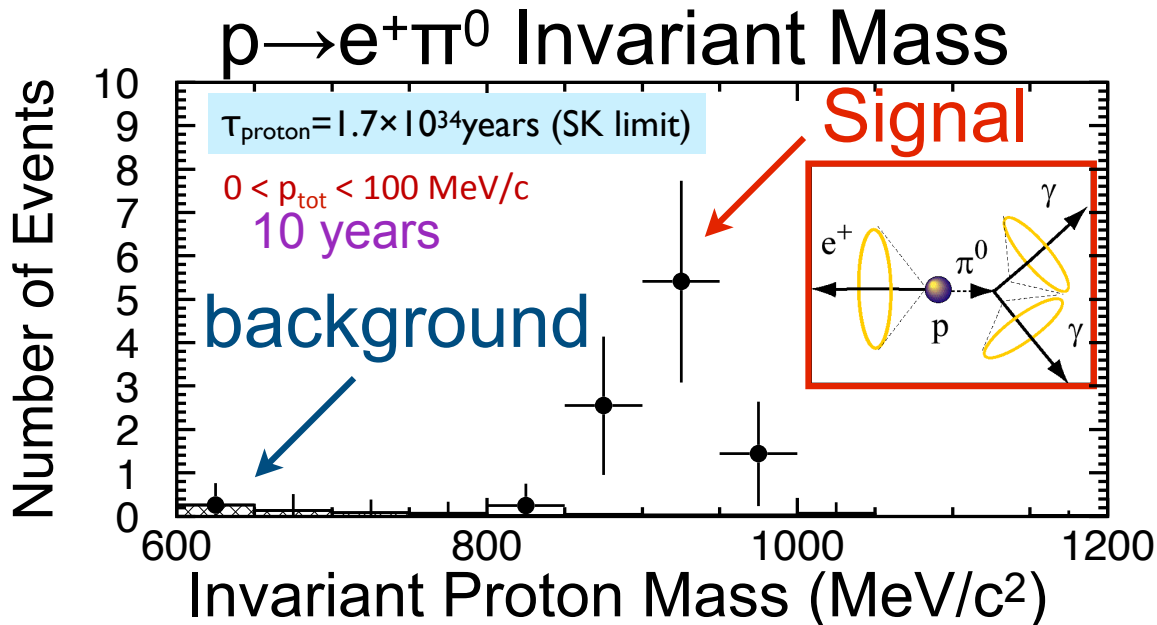


# Hyper-K: big step forward

arXiv:1805.04163



- Start operation in  $\sim 2027$  with higher mass (190kton,  $8 \times$  Super-K)
- **BG free search** possible: 0.01 BG/year for  $p_{\text{tot}} < 100 \text{ MeV}/c$
- **Invariant proton mass** would be a compelling evidence
- **Reach to  $10^{35}$  yrs**



	$p_{\text{tot}} < 100 \text{ MeV}/c$		$100 < p_{\text{tot}} < 250 \text{ MeV}/c$	
	Sig. $\epsilon(\%)$	Bkg (/ Mtyr)	Sig. $\epsilon(\%)$	Bkg (/ Mtyr)
HK	18.7	0.06	19.4	0.62

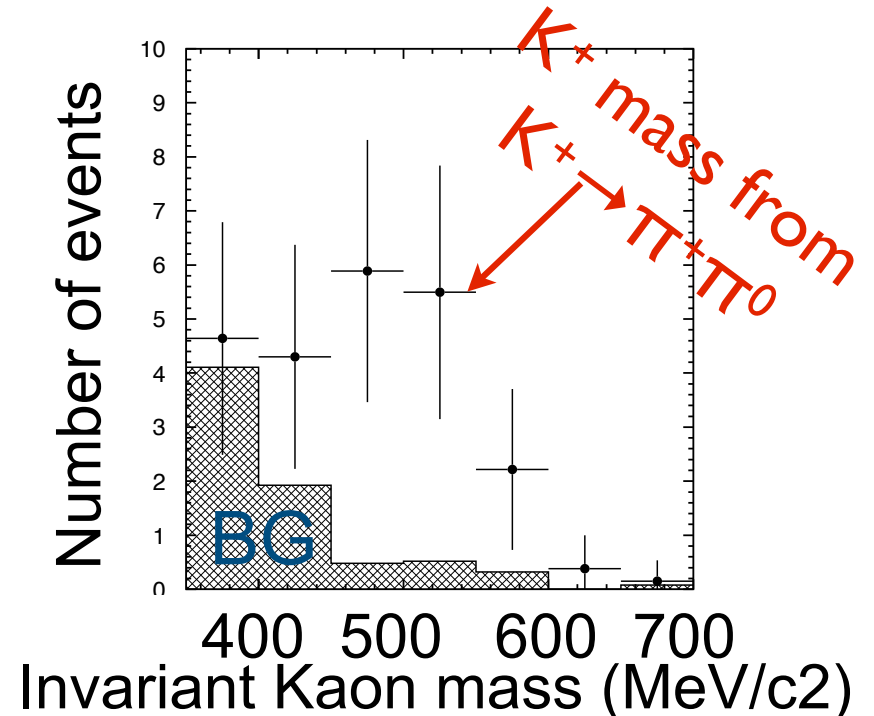
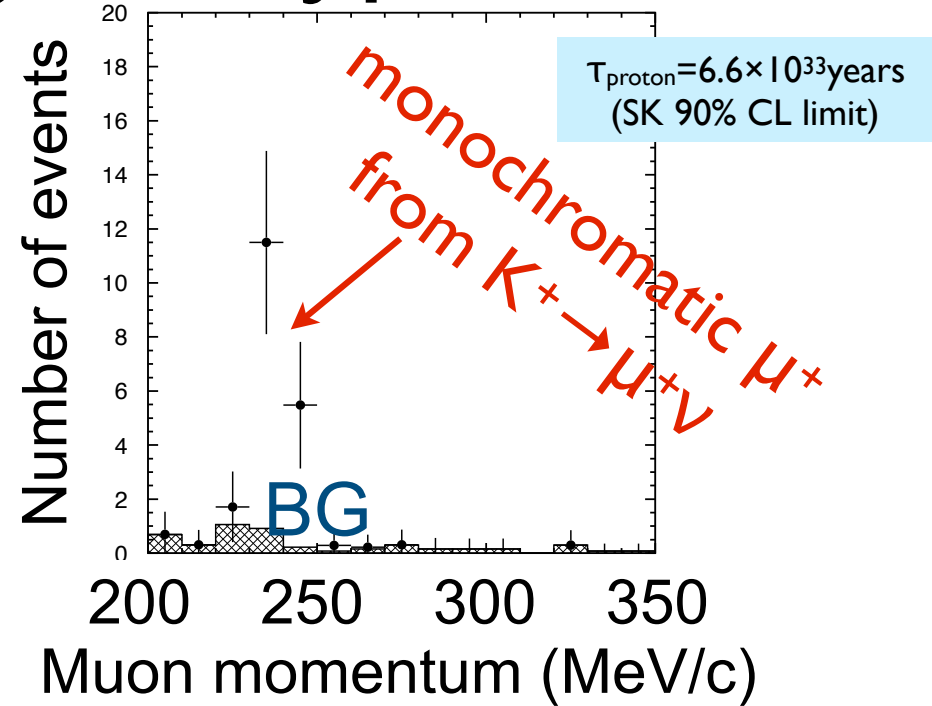
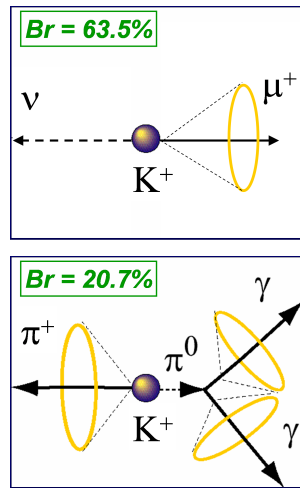
# $p \rightarrow \nu K^+$ discovery in Hyper-K

arXiv:1805.04163

- $K^+$  is invisible so signal

signature are:

- 236 MeV/c muon
- $\pi^+\pi^0$
- Discovery reach to  $3 \times 10^{34}$  years

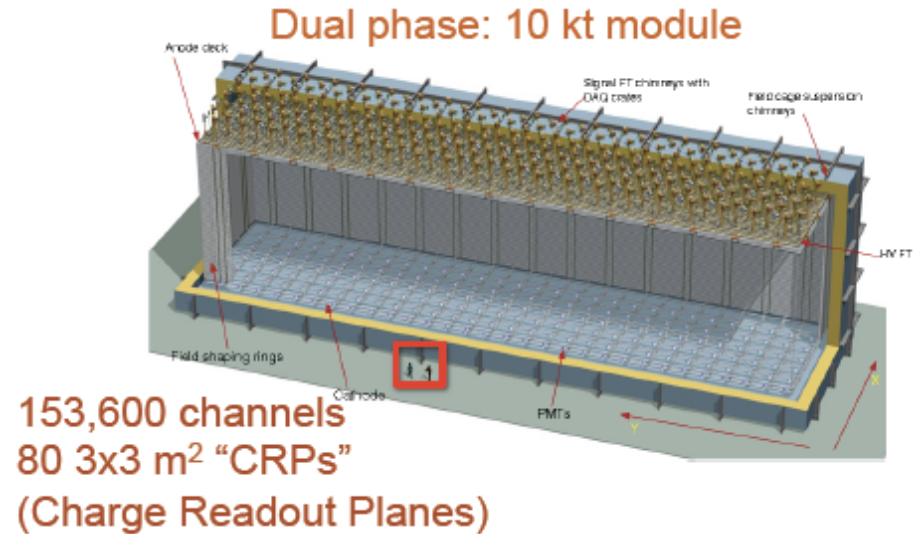
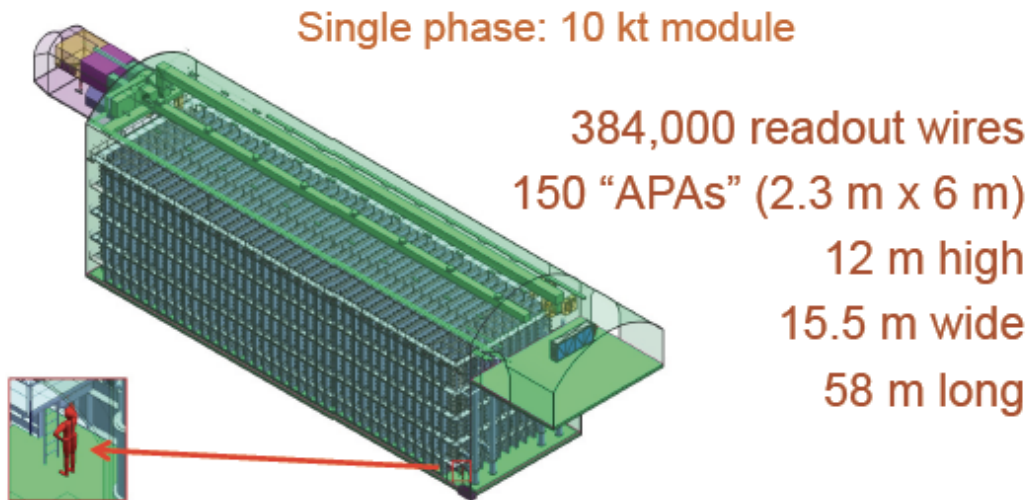


	prompt- $\gamma$ & $K^+ \rightarrow \mu^+ \nu$		$K^+ \rightarrow \pi^+ \pi^0$	
	Sig. $\epsilon(\%)$	Bkg (/ Mtyr)	Sig. $\epsilon(\%)$	Bkg (/ Mtyr)
HK	12.7	0.9	10.8	0.7

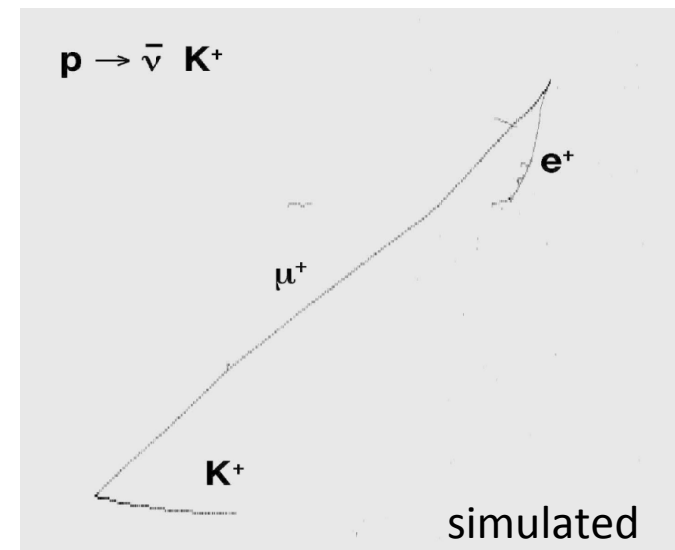
# Liquid Argon TPC DUNE

# DUNE: Liquid Argon TPC

- 4 10-kiloton liquid Argon TPC modules
- Starting data taking before 2026

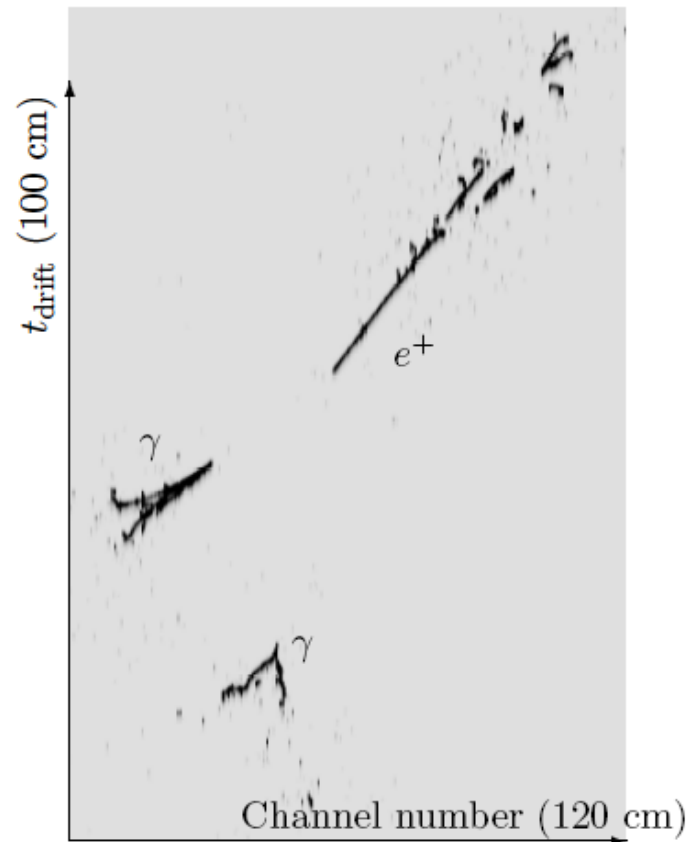


- LArTPC could identify the  $K^+$  track by higher ionization density with high efficiency.
- Single-event discovery could be possible.
- In addition, potential clean search for neutron-antineutron oscillation ( $\Delta B=2$ ) and other modes for which significant BG for water Cherenkov detectors

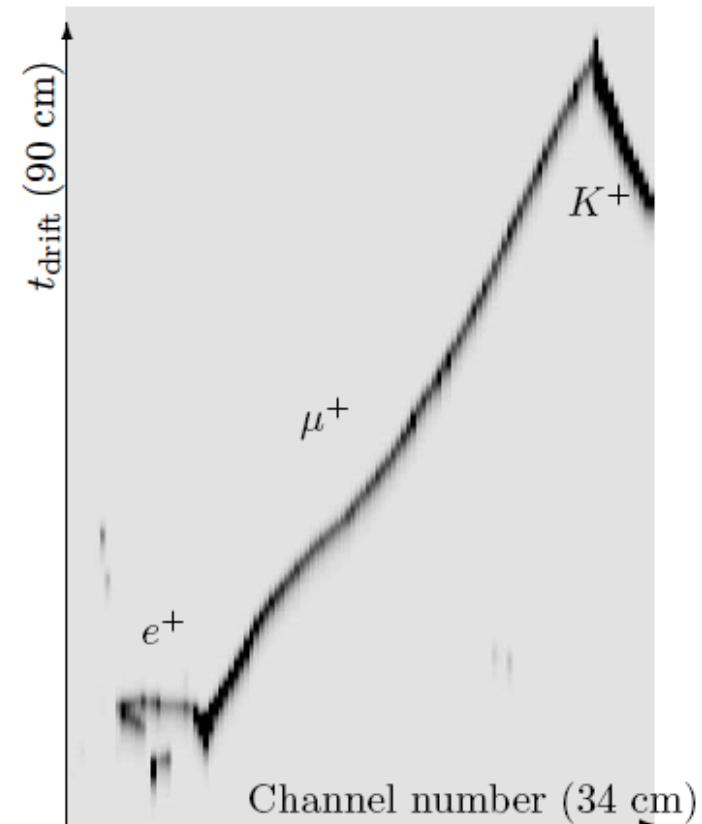


# Large Liq. Argon TPC detector

JHEP0704(2007)041; arXiv:1512.06148



$p \rightarrow e^+ + \pi^0$   
- efficiency = 45%, 1BG/Mtyr



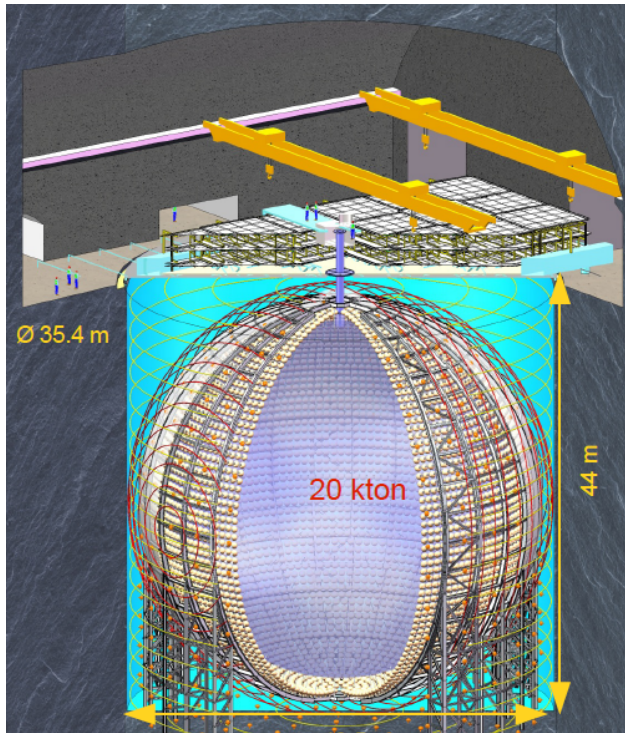
$p \rightarrow \nu + K^+$   
- efficiency = 97%, 1BG/Mtyr

Full detector simulation and automatic reconstruction  
under development

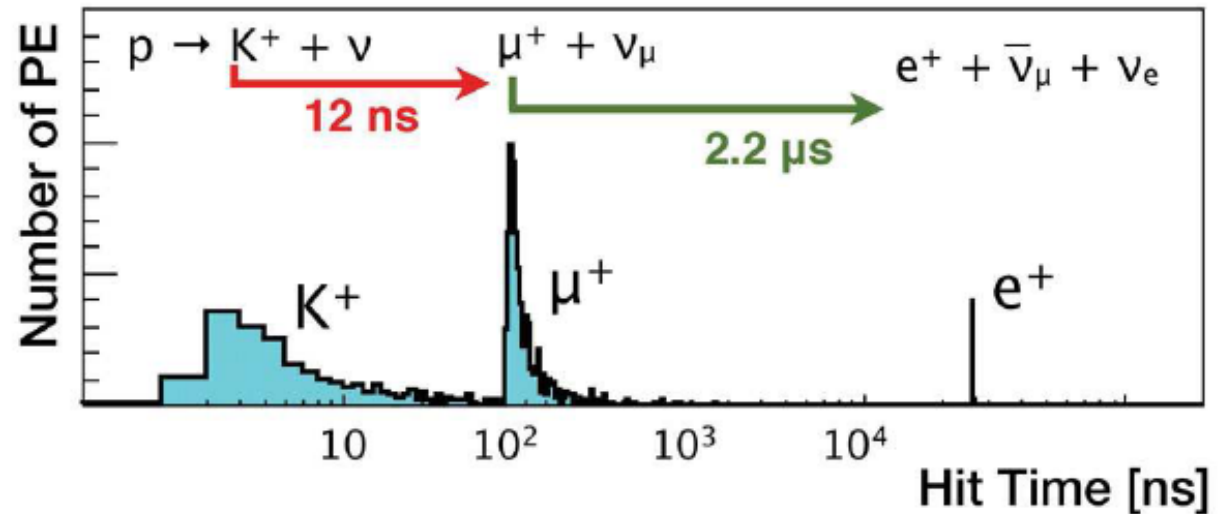
# Liquid scintillator detector JUNO

# $p \rightarrow \nu K^+$ in JUNO

arXiv:1507.05613



- 20 kiloton liquid scintillator
- Starting data taking in 2021



- Triple coincidence of  $K^+ \rightarrow \mu^+ \rightarrow e^+$  w/ well-defined time constant of (12nsec, 2.2 $\mu$ sec) and particle energies
- signal efficiency = 64% (pulse shape cut+energy cut+decay positron cut)
- Estimated backgrounds = 0.5 evt./ 10 years
- $\tau_{\text{proton}}/\text{Br}(p \rightarrow \nu K^+) = 1.9 \times 10^{34}$  years assuming zero candidates



# Summary

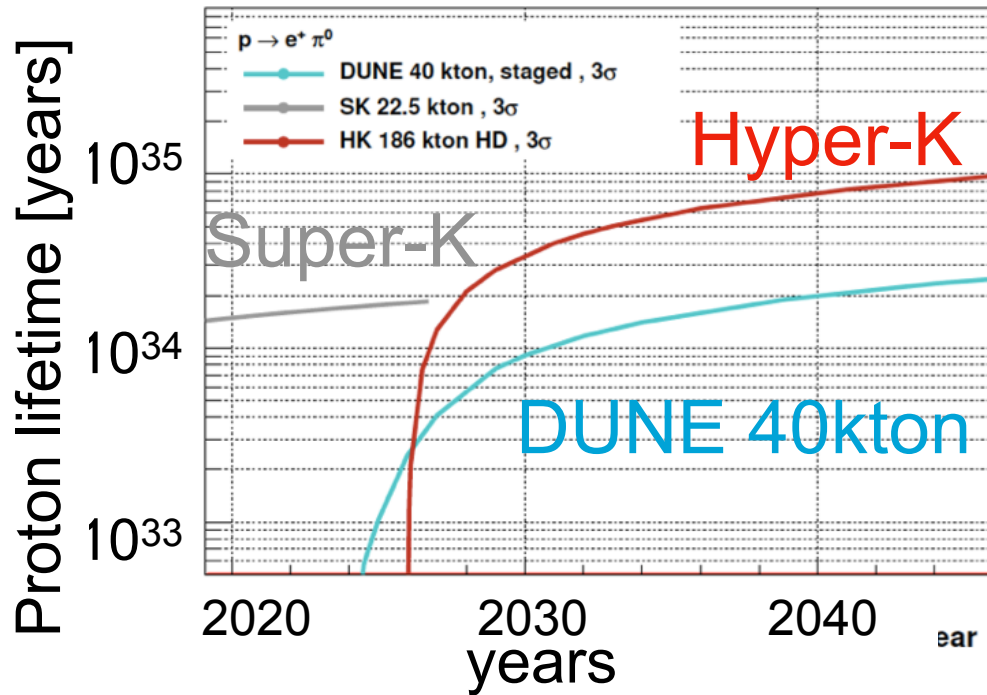
# Comparison

	Hyper-K 190 kton		DUNE 40 kton		JUNO 20 ton	
	Eff. (%)	BG (/Mt y)	Eff. (%)	BG (/Mt y)	Eff. (%)	BG (/Mt y)
$e^+\pi^0$	40	0.7	45	1	-	-
$\nu K^+$	24	1.6	97	1	64	2.5
	arXiv:1805.04163		JHEP0704(2007)041; arXiv:1512.06148		arXiv:1507.05613	

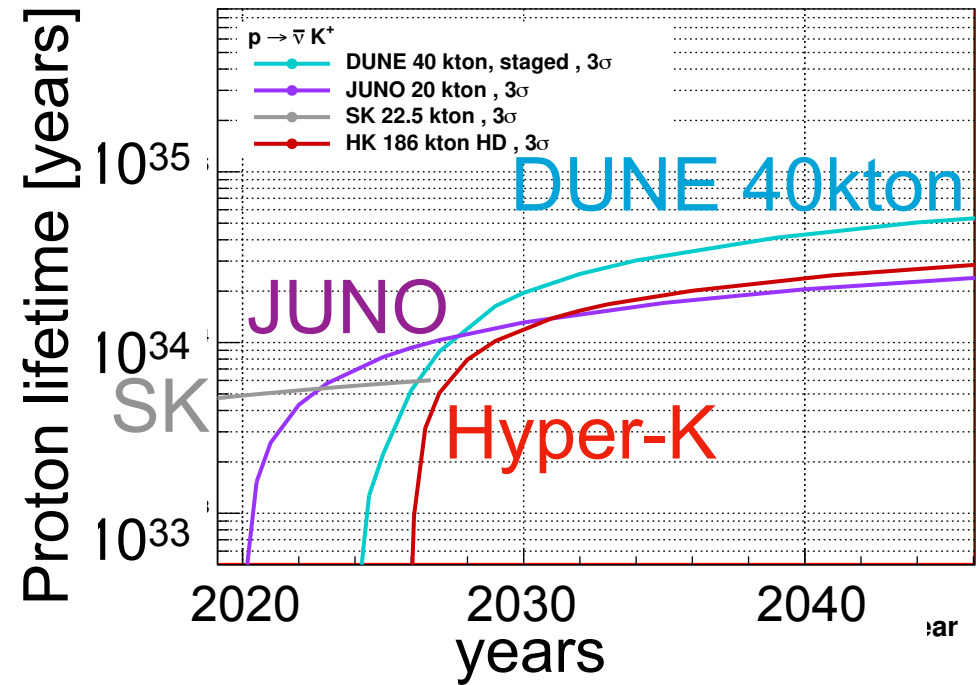
- For modes with **Kaons**, DUNE and JUNO can benefit from K identification and expected to have better S/N than water.
- For modes of “**charged lepton plus mesons**” like  $p \rightarrow e^+\pi^0$ , Hyper-K sensitivities are better by high mass.

# Proton lifetime sensitivities

$p \rightarrow e^+ \pi^0$   $3\sigma$  discovery



$p \rightarrow \nu K^+$   $3\sigma$  discovery



(Lines for DUNE and JUNO experiment have been generated based on numbers in the literature.)

$3\sigma$  discovery potential will reach to

- $1 \times 10^{35}$  years for  $p \rightarrow e^+ \pi^0$
- $5 \times 10^{34}$  years for  $p \rightarrow \nu K^+$

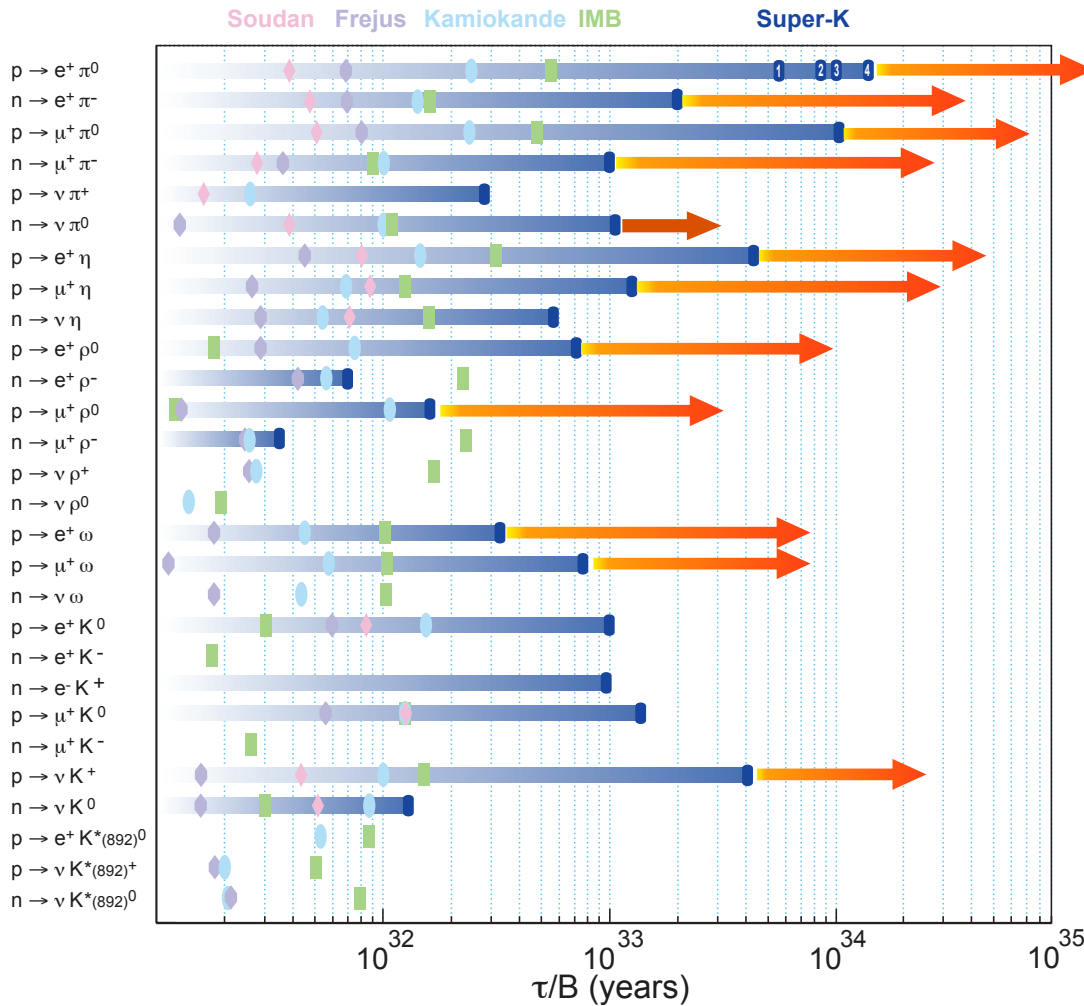


# Hyper-K's sensitivities

Improvements in many modes by a factor  $\sim 10$

Open for many decay modes including  $p \rightarrow e^+ \pi^0$ ,  $p \rightarrow \nu K^+$

Good chance for discovery!



▶  $p \rightarrow e^+ \pi^0$

- ▶  $\tau_{\text{proton}}/\text{Br} > 1 \times 10^{35}$  years @90%CL
- ▶ 5Mton  $\times$  years (9 Hyper-K years)

▶  $p, n \rightarrow (e^+, \mu^+) + (\pi, \rho, \omega, \eta)$

- ▶  $O(10^{34 \sim 35})$  years

▶ *SUSY* favored  $p \rightarrow \nu + K^+$

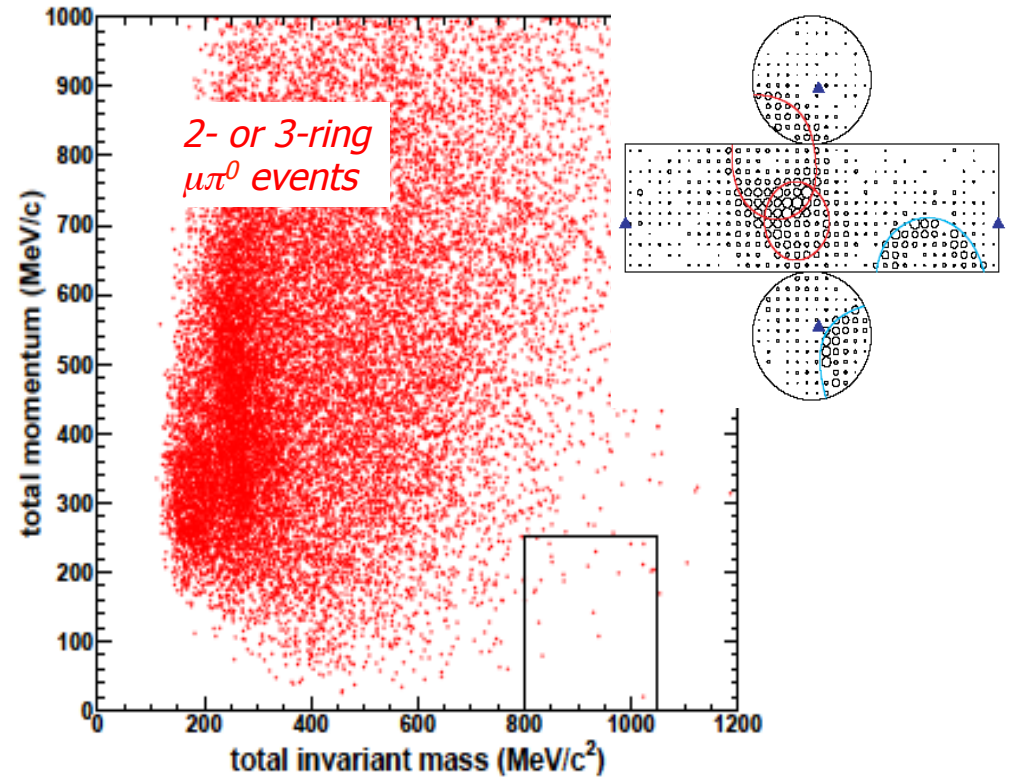
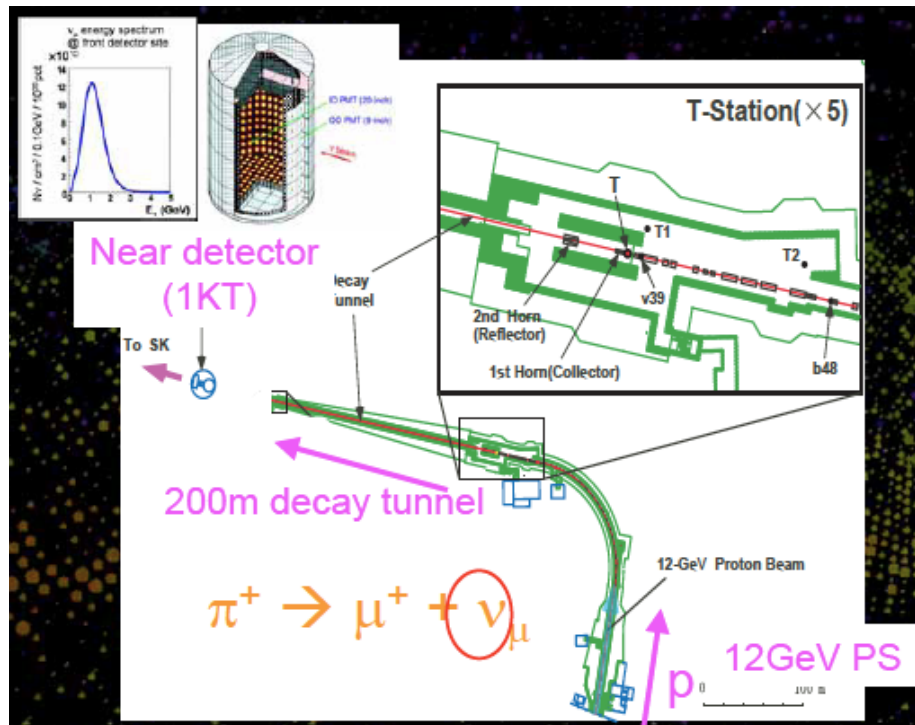
- ▶  $3 \times 10^{34}$  years

▶  $K^0$  modes,  $\nu \pi^0$ ,  $\nu \pi^+$  possible

▶ Others

- ▶ (B-L) violated modes
- ▶ radiative decays  $p \rightarrow e^+ \gamma$ ,  $\mu^+ \gamma$
- ▶ neutron-antineutron oscillations ( $|\Delta B|=2$ )
- ▶ di-nucleon decays ( $|\Delta B|=2$ )
- ▶  $pp \rightarrow XX \dots$ ,  $nn \rightarrow XX \dots$

# BG measurement by K2K beam



$\mu\pi^0$  converted to  $e\pi^0$  assuming lepton universality

K2K ( $p \rightarrow e^+ + \pi^0$  BG by  $E_\nu < 3\text{GeV}$ )  
 $1.63 + 0.42/-0.33(\text{stat.}) + 0.45/-0.51(\text{sys.})$  events/Mton-years

consistent

NEUT+SK simulation,  $E_\nu < 3\text{GeV}$   
 $1.8 \pm 0.3(\text{stat.})$  events/Mton-years