

# Indirect detection constraints on a $p$ -wave model through an $s$ -wave bremsstrahlung process

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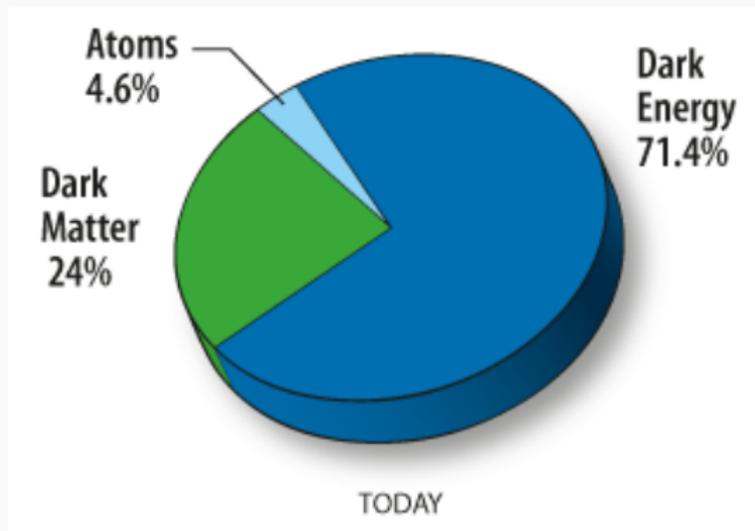
Collaboration with:

J. Dent, B. Dutta, and L. Strigari;

Phys.Rev. D99 (2019) no.8, 083003; arXiv:1901.01454

# Introduction

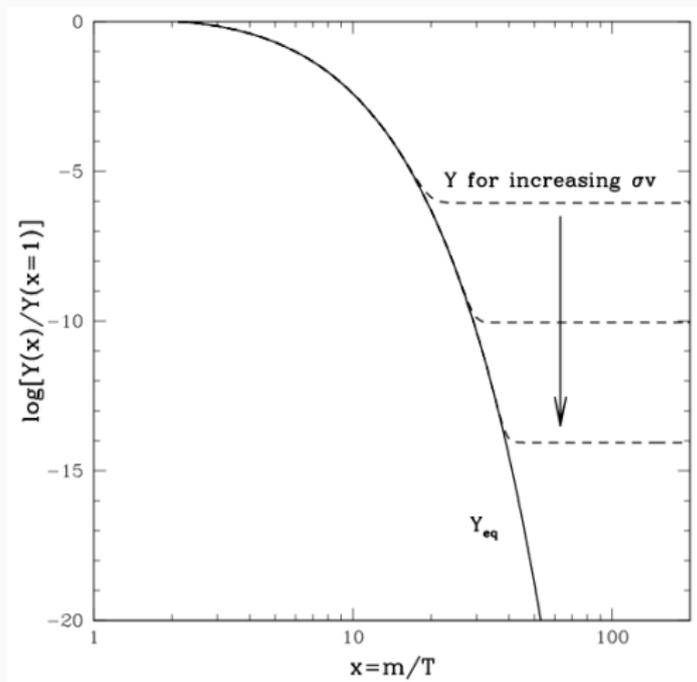
Dark matter makes up  $\sim 25\%$  of the energy density in the Universe today



[https://wmap.gsfc.nasa.gov/universe/uni\\_matter.html](https://wmap.gsfc.nasa.gov/universe/uni_matter.html)

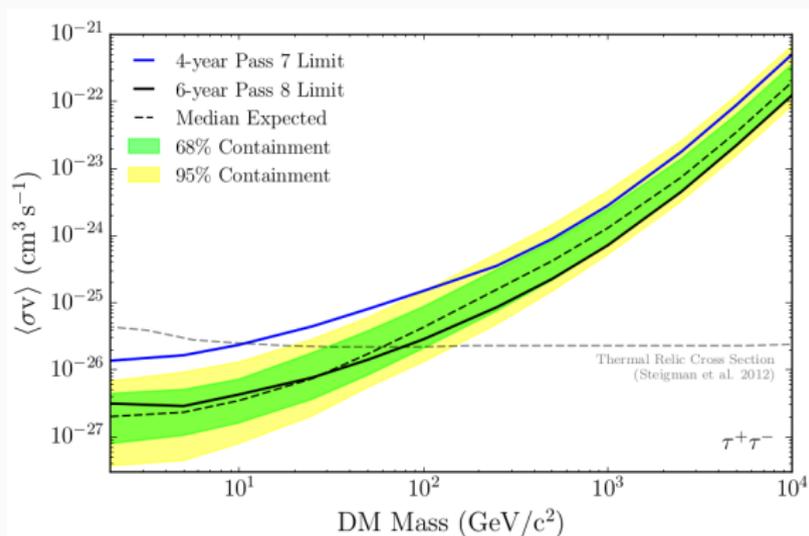
# Relic Abundance

The relic abundance is set by the thermally averaged cross section



# Fermi Limits

Strong limits are placed on numerous dark matter annihilation models ruling out thermal parameter space



## *s* and *p* wave Cross Sections

Thermally averaged cross section:

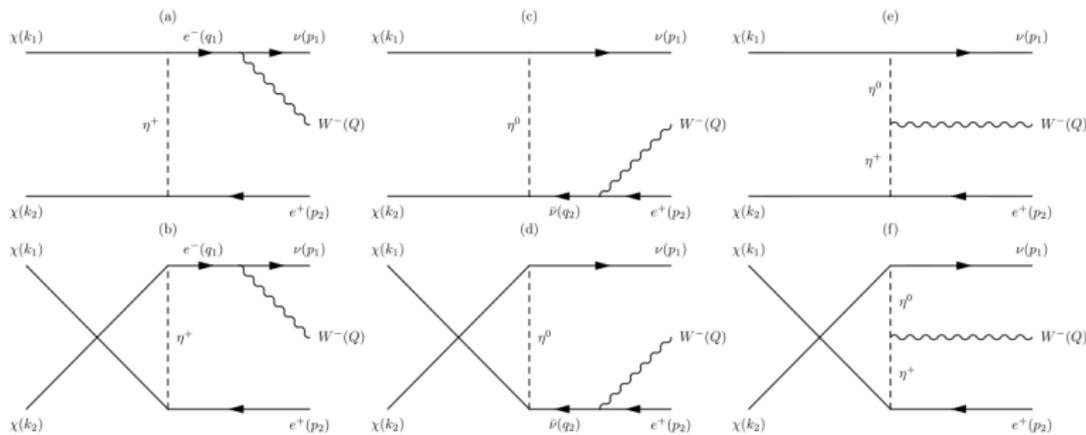
$$\langle \sigma v \rangle = a + b v^2 + \dots$$

The annihilation cross section today is the same as at freeze out for *s*-wave models, but for *p*-wave, the cross section is much lower

Due to the velocity suppression, *p*-wave models are very hard to probe

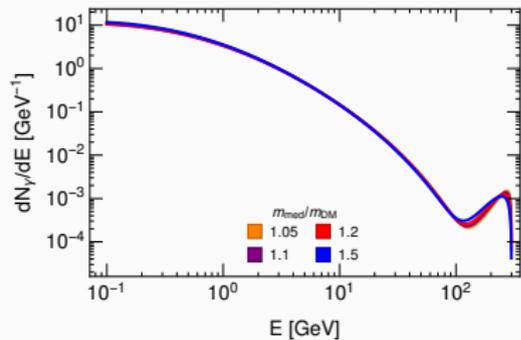
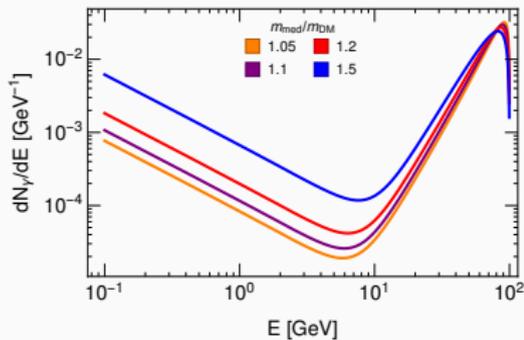
# Bremsstrahlung Gauge Boson

Through the addition of a bremsstrahlung gauge boson, velocity suppression is lifted, and an s-wave cross section is present

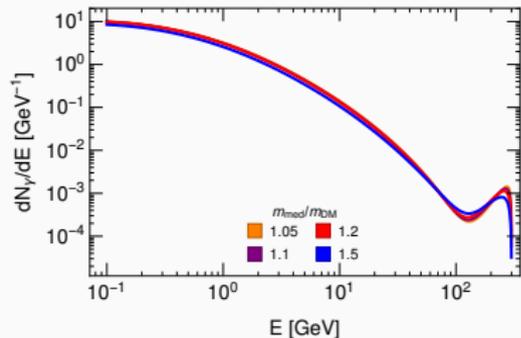
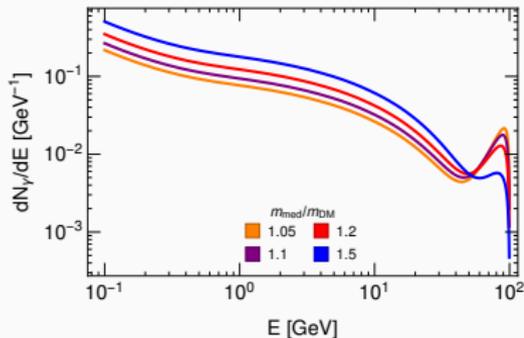


# Bremstrahlung Photon Spectra

electron



tau



# Experiments

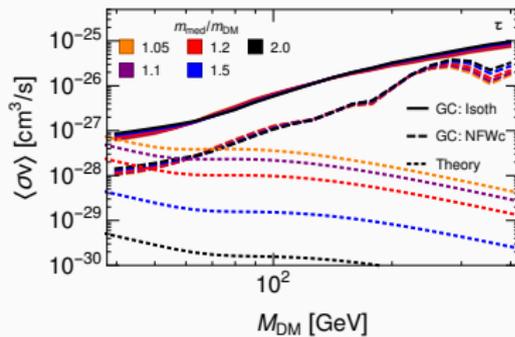
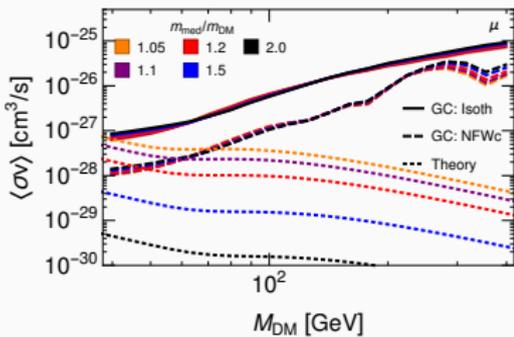
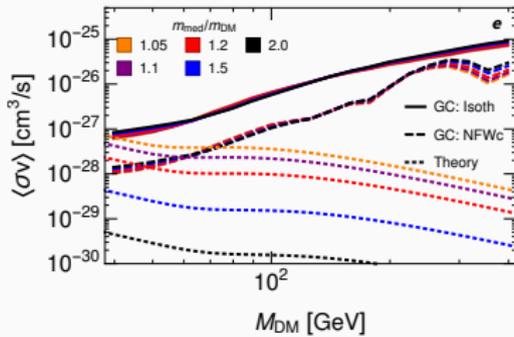
Compared  
with multiple experiments to estimate their constraining capabilities:

Fermi-LAT — Gamma-ray Spectrum  
Galactic Center  
dSph

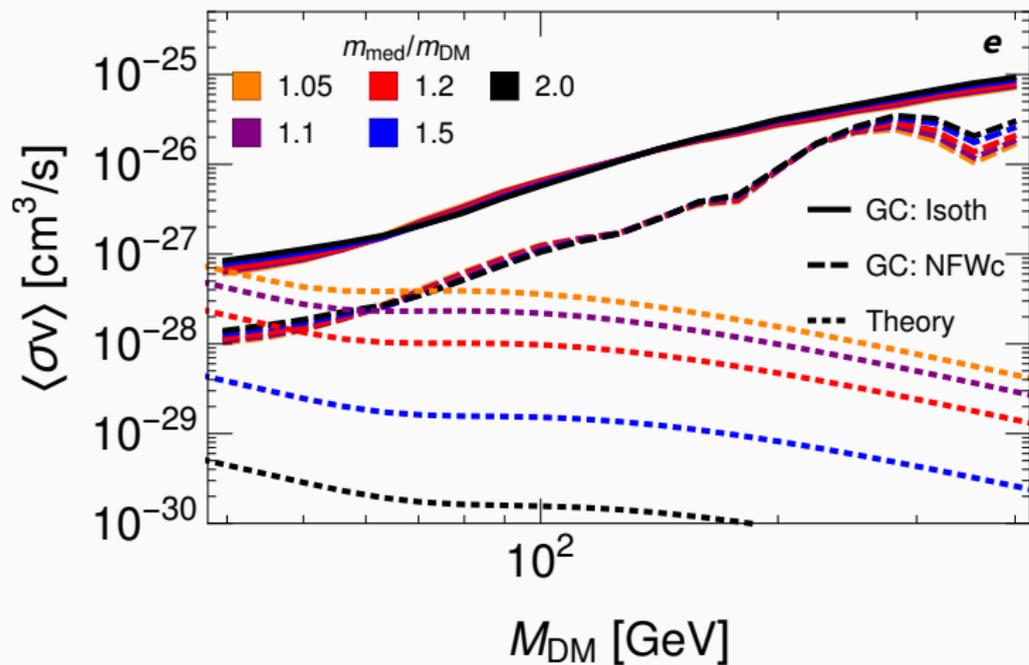
EDGES — 21 cm

Planck — CMB

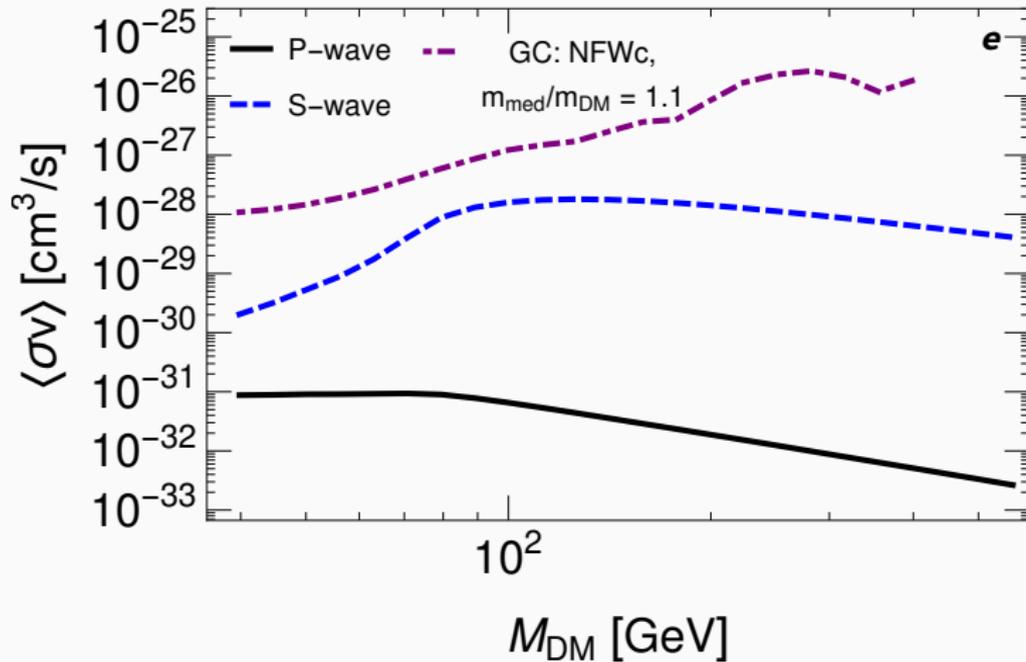
# Model Constraint



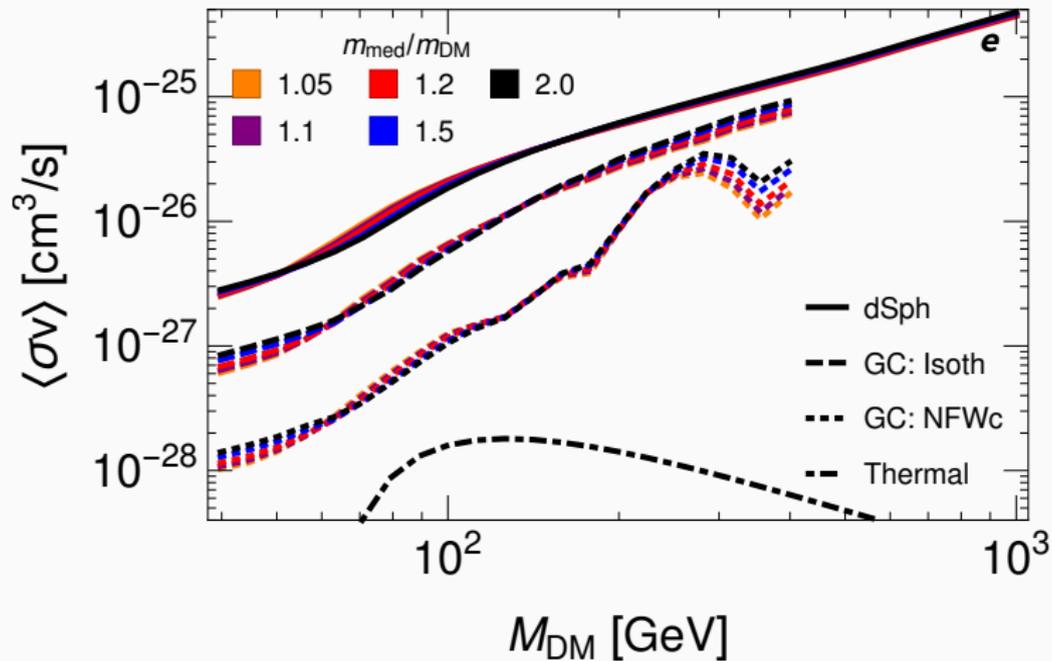
# Model Constraint



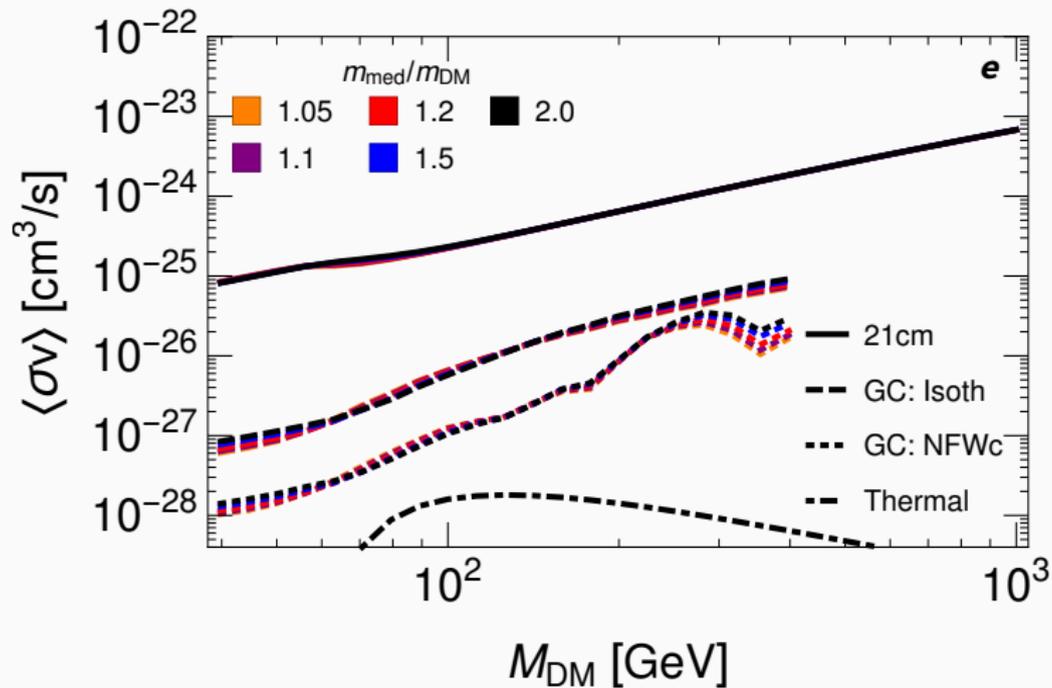
# Improvement over Base $p$ -wave Model



# Thermal Model Constraint



# Thermal Model Constraint

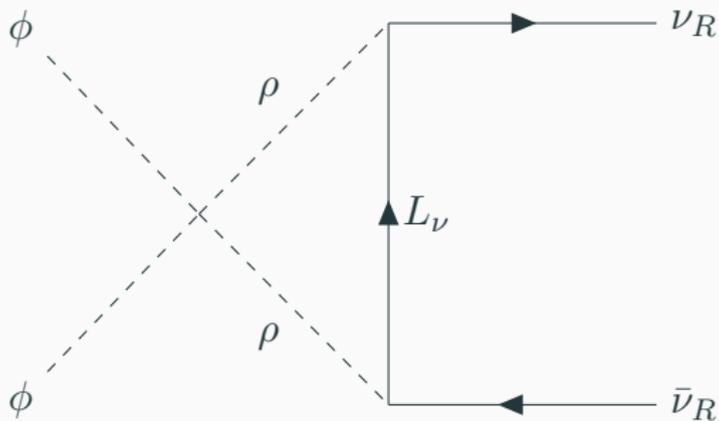


# Annihilation to Neutrinos Example

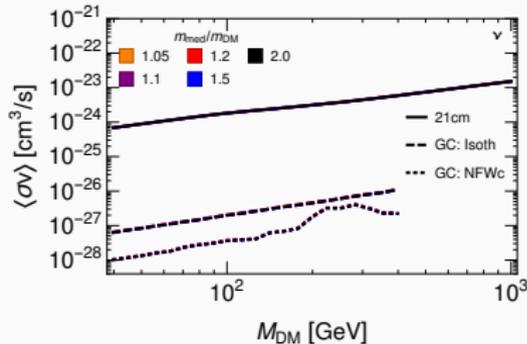
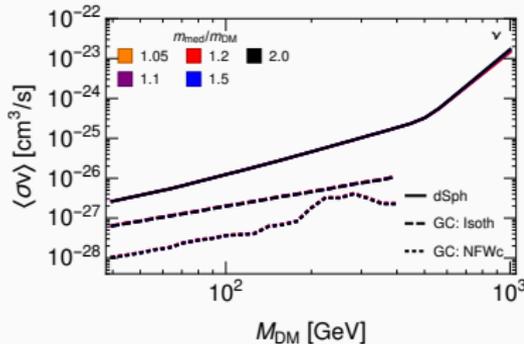
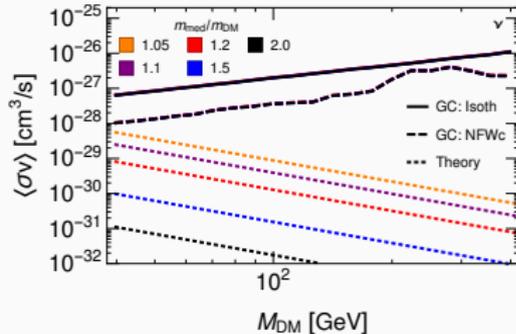
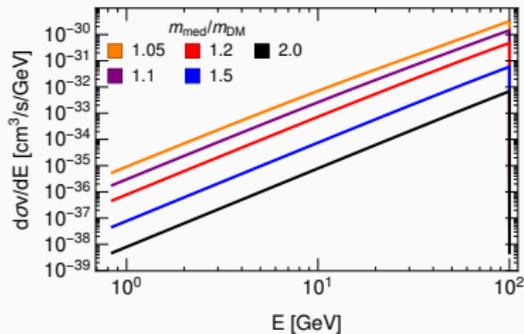
$\rho$  is a scalar doublet

$L$  is a fermion-like doublet

Photon brem is possible from the charged loop



# Annihilation to Neutrinos



# Summary

The relic abundance is set by  $p$ -wave interactions

Today, higher order  $s$ -wave channels can dominate over the simple  $p$ -wave interactions

These  $s$ -wave channels provide an avenue to investigate experimentally difficult models

Thank You!