



We still believe in supersymmetry

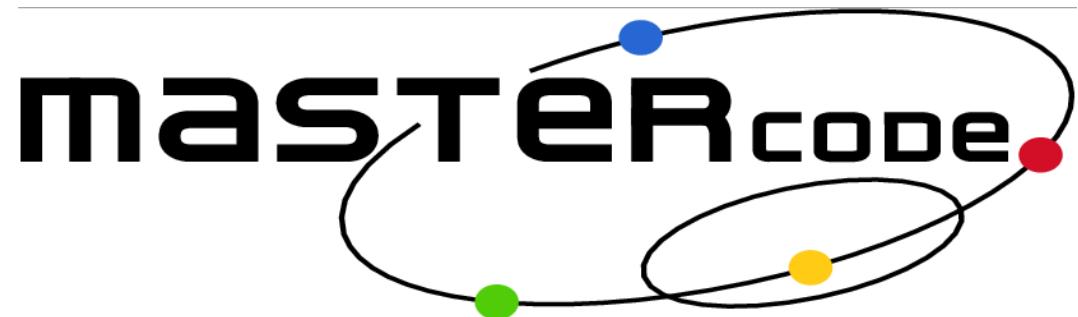
You must be joking

# Where is SUSY?

*Sven Heinemeyer, IFT/IFCA (CSIC, Madrid/Santander)*

Corpus Cristi, 05/2019

1. Introduction
2. The MasterCode
3. GUT based predictions
4. pMSSM11 predictions
5. Conclusions



# 1. Introduction

Some “recent” measurements:

- top quark mass
- Higgs boson mass
- Higgs boson “couplings”
- Dark Matter (properties)

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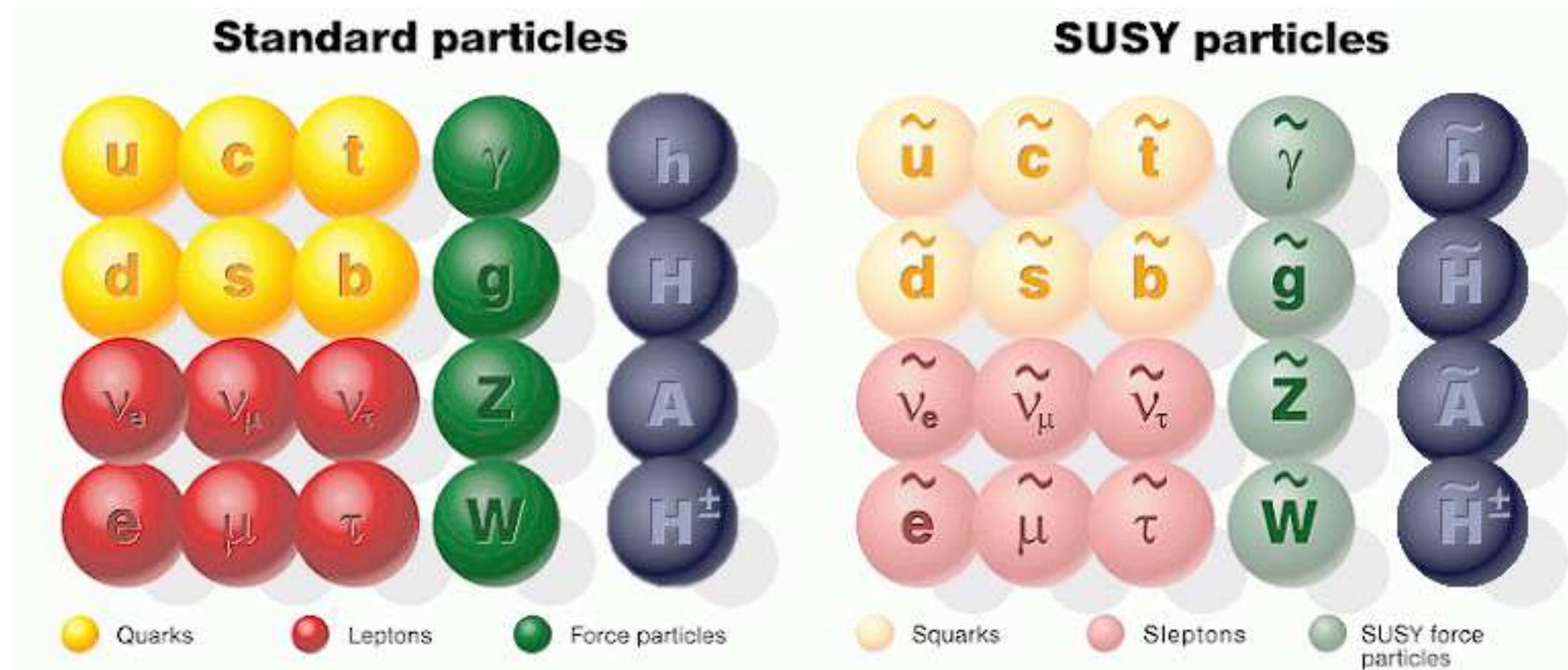
Simple SUSY models predicted correctly:

- top quark mass
- Higgs boson mass
- Higgs boson “couplings”
- Dark Matter (properties)

⇒ good motivation to look at SUSY!

# The Minimal Supersymmetric Standard Model (MSSM)

Superpartners for Standard Model particles



Problem in the MSSM: more than 100 free parameters

Nobody(?) believes that a model describing nature has so many free parameters!

## GUT based models: 1.) CMSSM (sometimes wrongly called mSUGRA):

⇒ Scenario characterized by

$$m_0, m_{1/2}, A_0, \tan\beta, \text{sign } \mu$$

$m_0$  : universal scalar mass parameter

$m_{1/2}$  : universal gaugino mass parameter

$A_0$  : universal trilinear coupling

$\tan\beta$  : ratio of Higgs vacuum expectation values

$\text{sign}(\mu)$  : sign of supersymmetric Higgs parameter

} at the GUT scale

⇒ particle spectra from renormalization group running to weak scale

⇒ Lightest SUSY particle (LSP) is the lightest neutralino ⇒ DM!

## Other GUT based models:

2.) NUHM1: CMSSM + 1 scalar mass parameter

$m_0$ ,  $m_{1/2}$ ,  $A_0$ ,  $\tan\beta$ , sign  $\mu$  and  $M_A$

3.) NUHM2: CMSSM + 2 scalar mass parameters

$m_0$ ,  $m_{1/2}$ ,  $A_0$ ,  $\tan\beta$ ,  $\mu$  and  $M_A$

4.) SU(5): CMSSM + 3 scalar mass parameters

$m_5$ ,  $m_{10}$ ,  $m_{1/2}$ ,  $A_0$ ,  $\tan\beta$ ,  $m_{H_u}$ ,  $m_{H_d}$

5.) mAMSB: different mechanism for SUSY breaking

$m_{3/2}$ ,  $m_0$ ,  $\tan\beta$ , sign( $\mu$ )

6.) sub-GUT: CMSSM, but unification at lower scale

$m_0$ ,  $m_{1/2}$ ,  $A_0$ ,  $\tan\beta$ , sign  $\mu$  and  $M_{\text{in}}$

7.) ...

⇒ wide variety of models covered!

## Problem: We cannot be sure about the SUSY-breaking mechanism

- ⇒ it is possible that with the CMSSM, NUHM, SU(5), mAMSB, sub-GUT we missed the “correct” mechanism
- ⇒ hint: strong connection between colored and uncolored sector tension between low-energy EW effects and (colored) LHC searches

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## Solution: investigate also the “general MSSM”

⇒ 11 parameters are manageable ⇒ pMSSM11

- squark mass parameters:  $m_{\tilde{q}_{1,2}} =: m_{\tilde{q}}$ ,  $m_{\tilde{q}_3}$
- slepton mass parameter(s):  $m_{\tilde{l}}$ ,  $m_{\tilde{\tau}}$
- gaugino masses:  $M_1$ ,  $M_2$ ,  $M_3$
- trilinear coupling:  $A$
- Higgs sector parameters:  $M_A$ ,  $\tan \beta$
- Higgs mixing parameter:  $\mu$

## 2. The MasterCode

⇒ collaborative effort of theorists and experimentalists

[*Bagnaschi, Borsato, Buchmüller, Cavanaugh, Chobanova, Citron, Costa, De Roeck, Dolan, Ellis, Flächer, SH, Isidori, Liu, Lucio, Martinez Santos, Olive, Sakurai, Weiglein*]

Über-code for the combination of different tools:

- Über-code original in Fortran, now re-written in C++
- tools are included as **subroutines**
- **compatibility** ensured by collaboration of authors of “MasterCode” and authors of “sub tools” **/SLHA(2)**
- sub-codes in Fortran or C++

⇒ evaluate observables of one parameter point consistently with various tools

[cern.ch/mastercode](http://cern.ch/mastercode)

## Data we have:

- Higgs boson mass/couplings/... (LHC)  $\Rightarrow$  FeynHiggs

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- SUSY searches (LHC)  $\Rightarrow$  own re-cast (fast-lim approach)

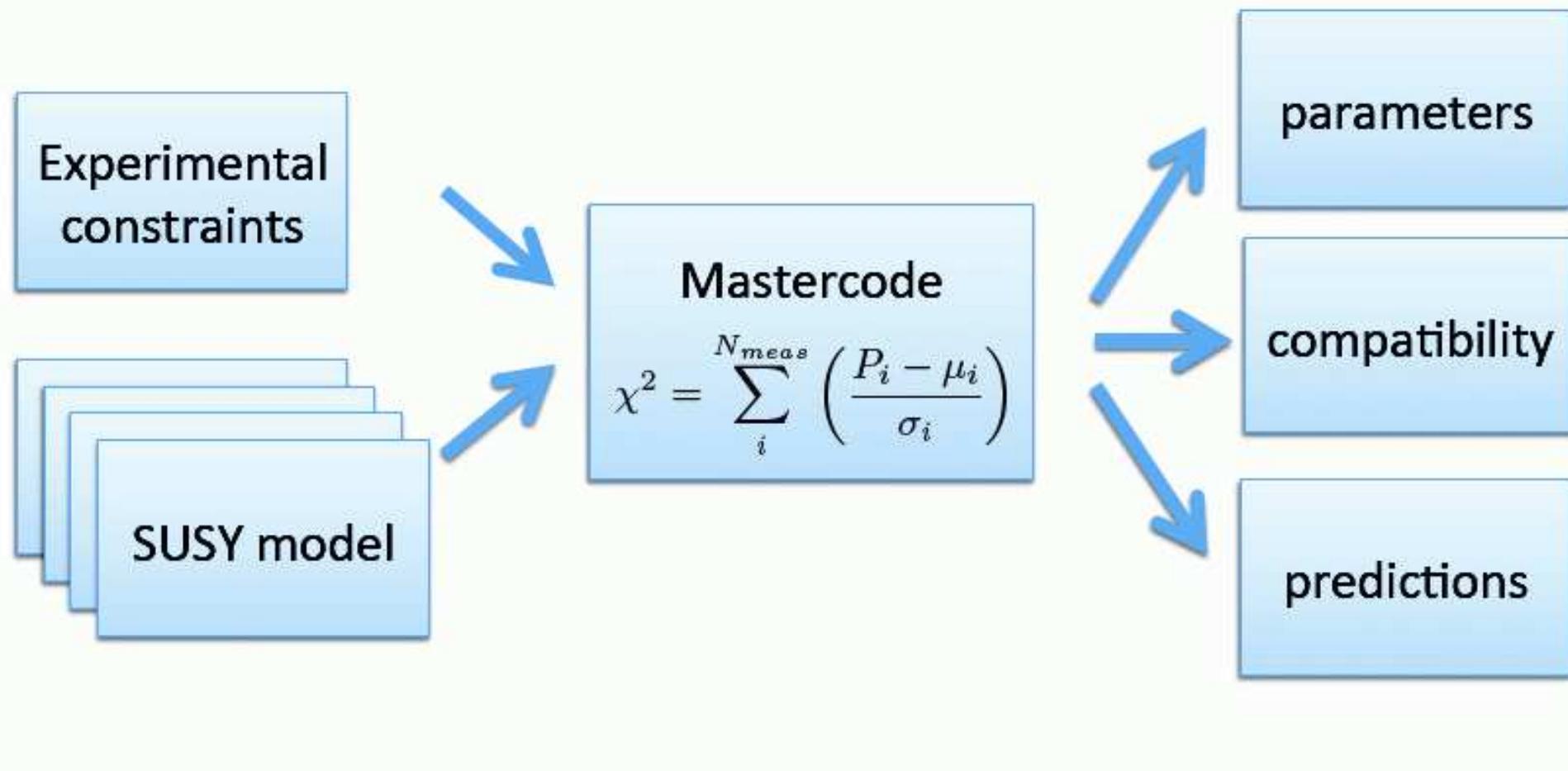
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- SUSY searches (LHC)  $\Rightarrow$  own re-cast (fast-lim approach)
- electroweak precision data  $\Rightarrow$  FeynWZ, FeynHiggs
- flavor data  $\Rightarrow$  SuperIso, SuFla
- astrophysical data (DM properties)  $\Rightarrow$  MicrOMEGAs, SSARD

The  $\chi^2$  evaluation:



## Global fits of SUSY



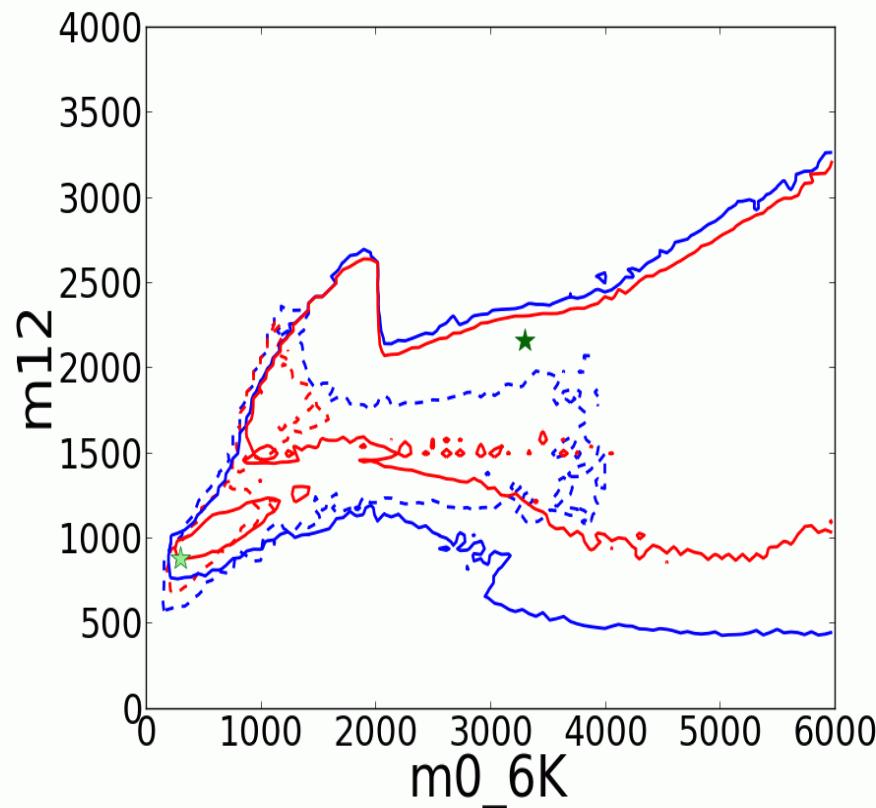
### 3. GUT based predictions

$m_0$ - $m_{1/2}$  plane including LHC 20/fb:

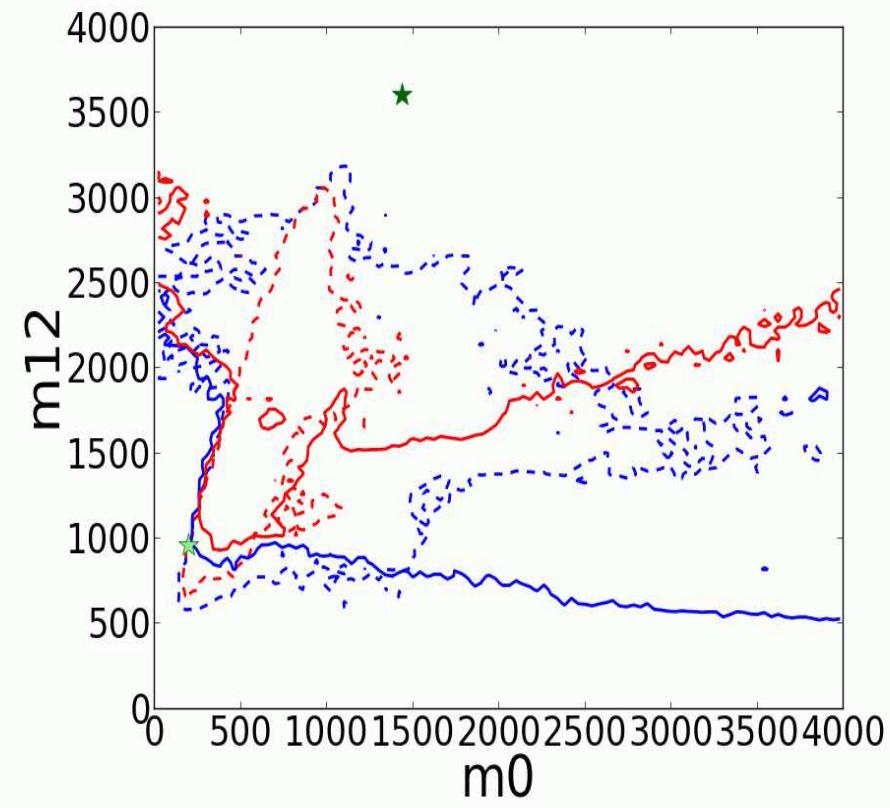


[2013]

CMSSM

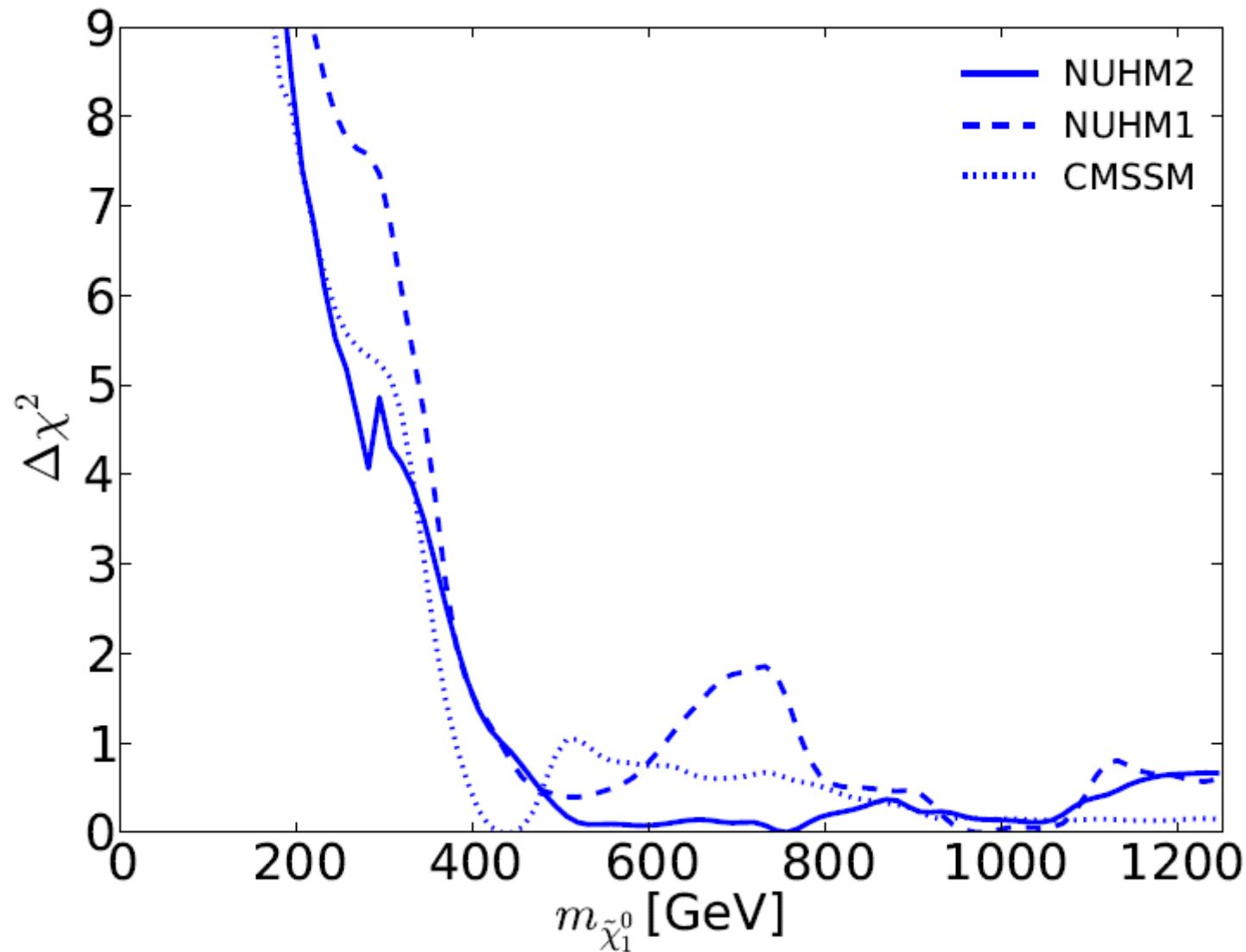


NUHM1



dotted: LHC 5/fb 7 TeV, solid: LHC 20/fb 8 TeV

⇒ very high masses favored!



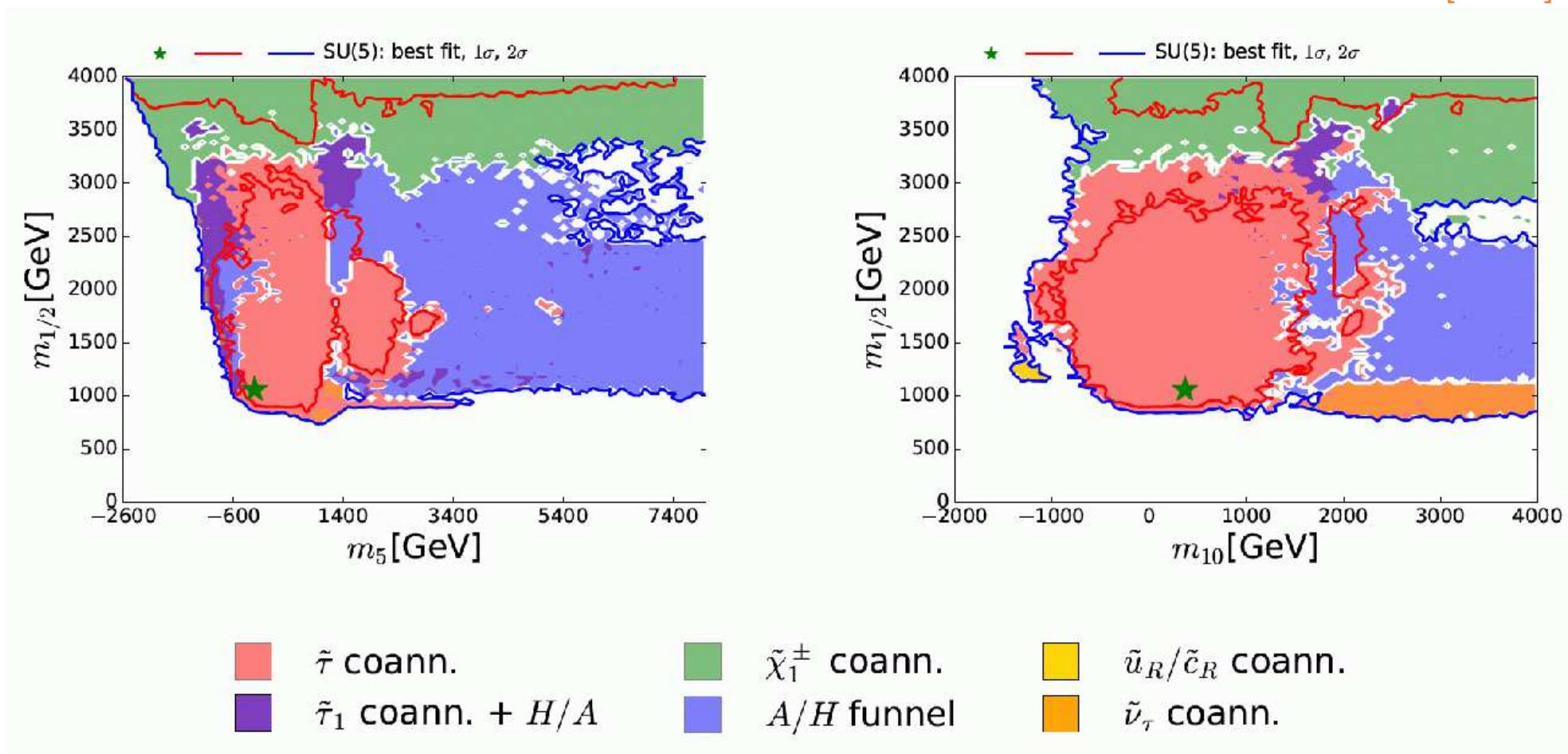
⇒ only very large values are favored

# Results in the SU(5)



[2016]

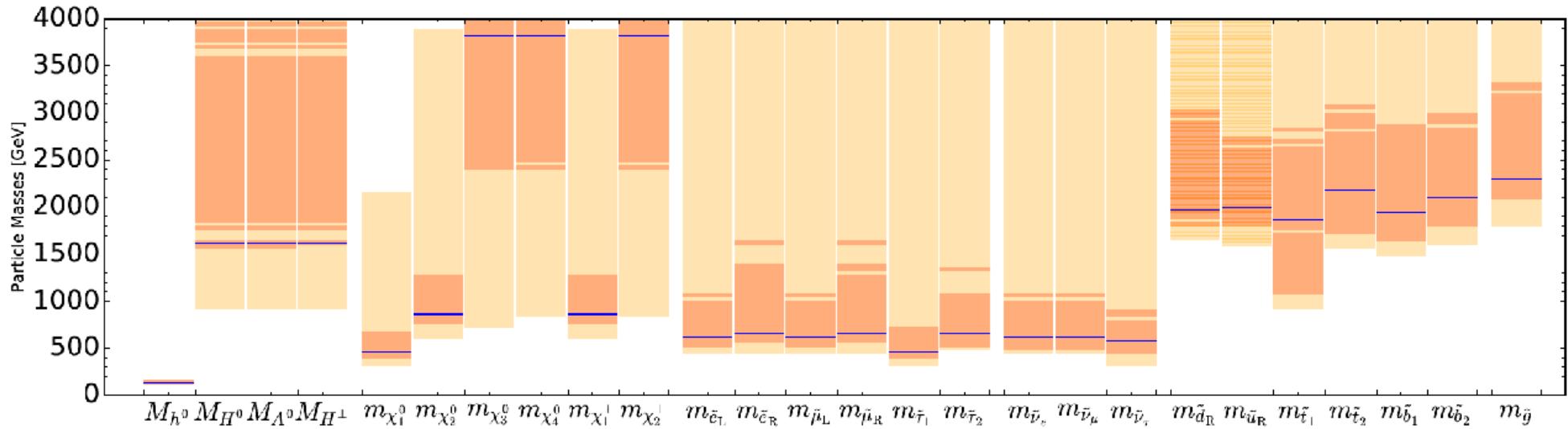
## Dark Matter annihilation mechanism:



$\Rightarrow \tilde{u}_R/\tilde{c}_R/\tilde{\nu}_\tau$  co-ann. possible  $\Rightarrow$  but  $\tilde{\tau}_1$  co-ann. dominant!

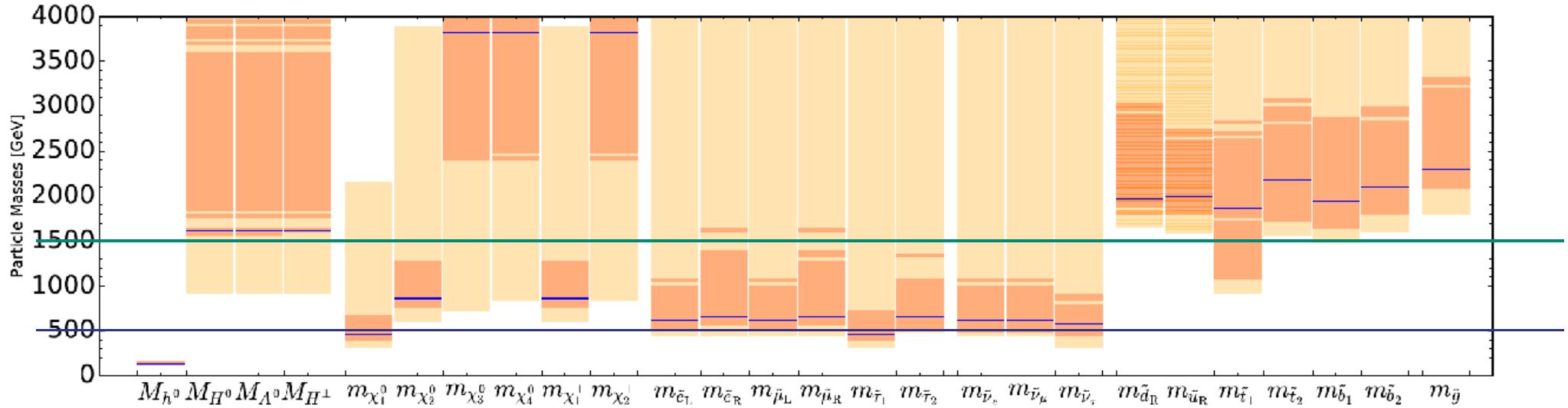
## SU(5) prediction: best-fit masses

[2016]



- ⇒ high colored masses
- ⇒ lower electroweak masses  
partially with not too large  $1\sigma$  ranges
- ⇒ clear prediction for future experiments?

## SU(5) prediction: best-fit masses



ILC:  $\sqrt{s} = 1000$  GeV  $\Rightarrow$  only few EW particles possibly accessible

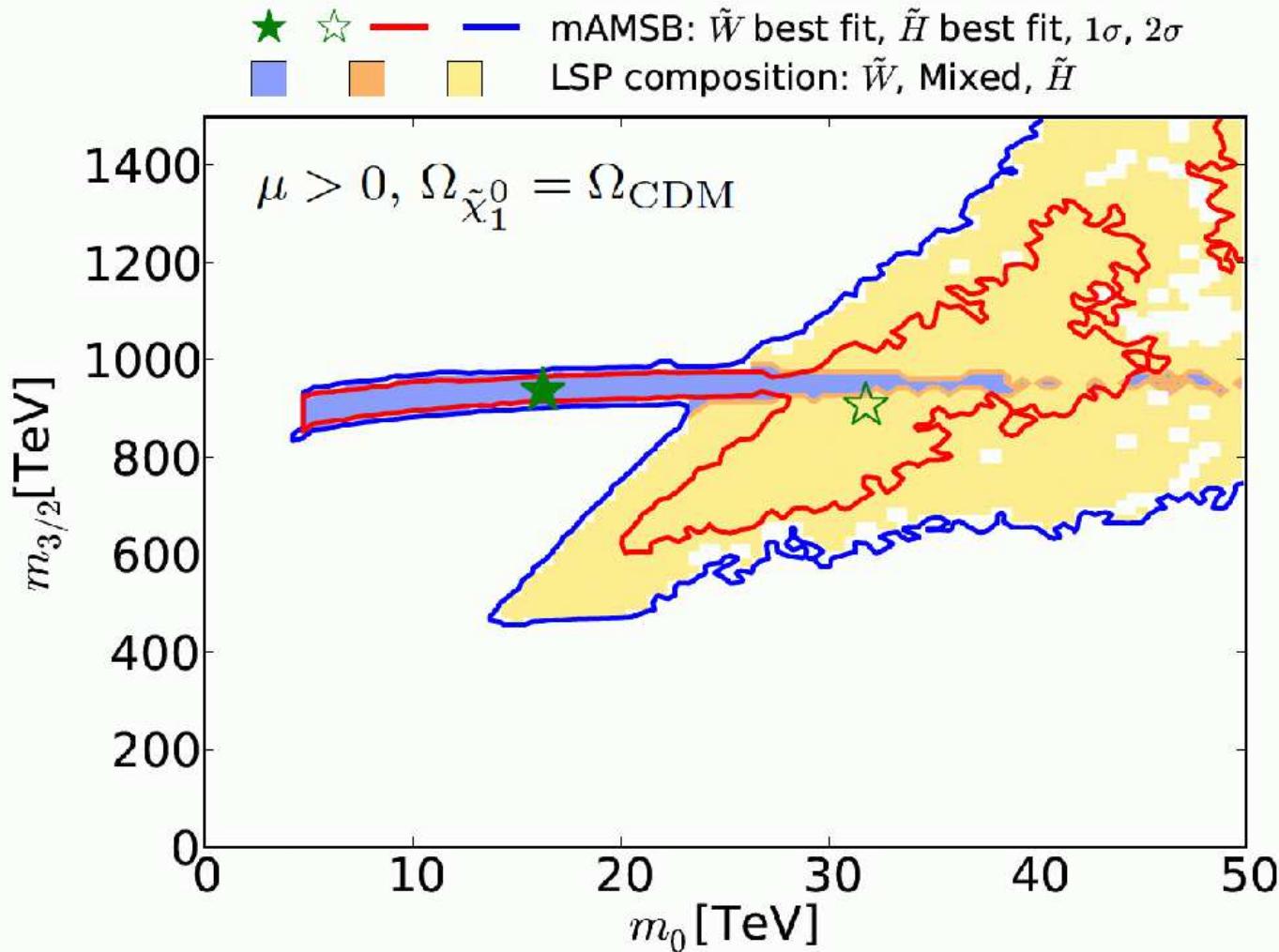
CLIC:  $\sqrt{s} = 3000$  GeV  $\Rightarrow$  pair production of many SUSY particles “likely”  
 $\Rightarrow$  no access to colored particles

## Results in the mAMSB



[2016]

### Dark Matter composition:



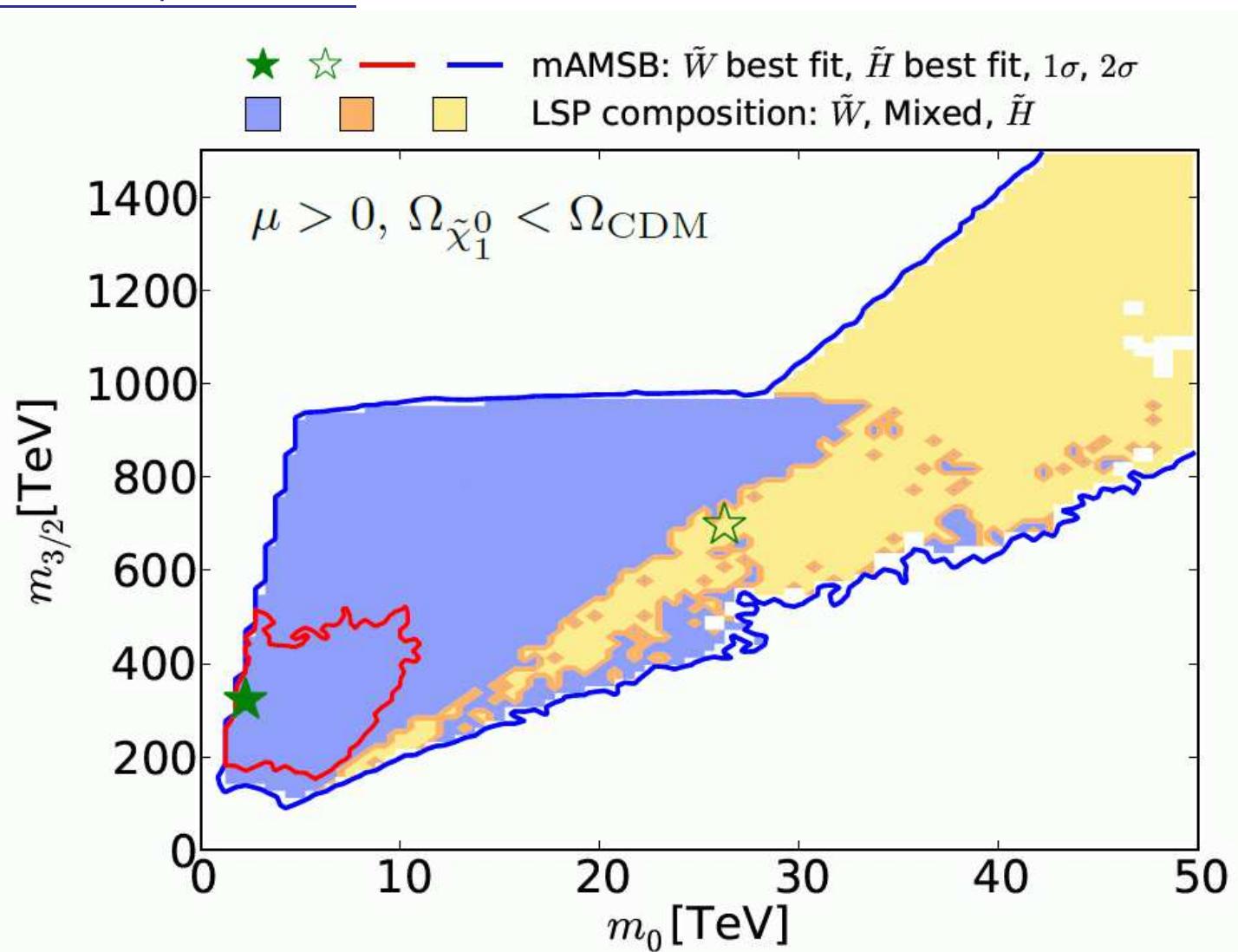
$\Rightarrow m_{\tilde{\chi}_1^0} \sim 2.9 \pm 0.1$  TeV (wino),  $\sim 1.1 \pm 0.02$  TeV (higgsino)

## Results in the mAMSB



[2016]

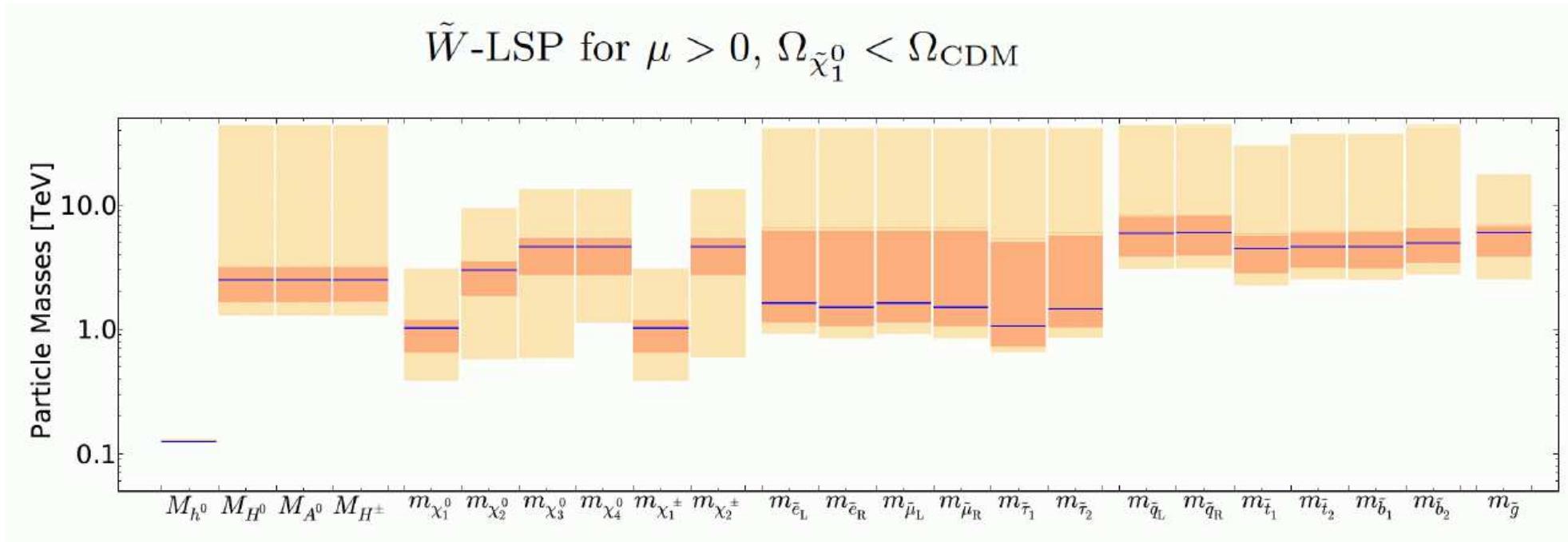
### Dark Matter composition:



⇒ very relaxed limits ⇒ lower masses

## mAMSB prediction: best-fit masses (wino)

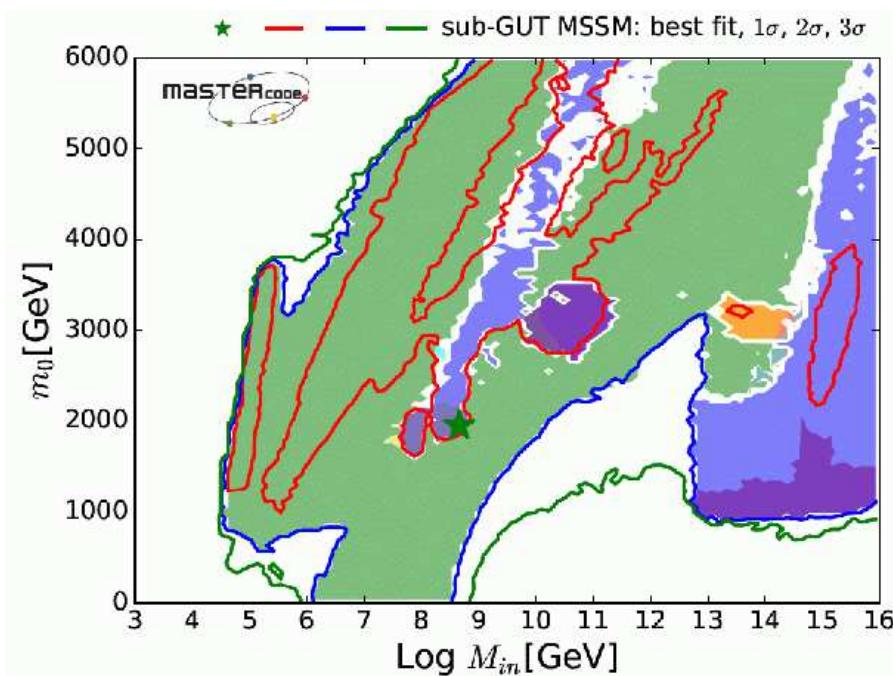
[2016]



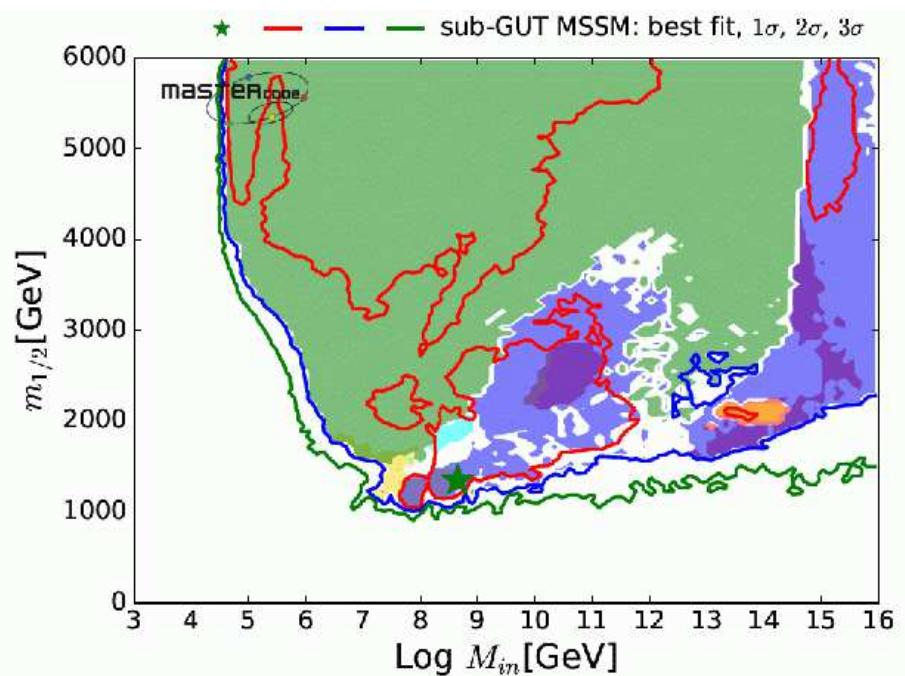
- ⇒ high colored masses
- ⇒ lower electroweak masses  
partially with not too large  $1\sigma$  ranges
- ⇒ clear prediction for future experiments?
- ⇒ Higgsino case even heavier . . .

## Results in sub-GUT

[2017]



$\tilde{\chi}_1^\pm$ coann.	$\tilde{\tau}_1$ coann.
$A/H$ funnel	focus point
$\tilde{t}_1$ coann.	$\tilde{t}_1$ coann. + $H/A$ funnel



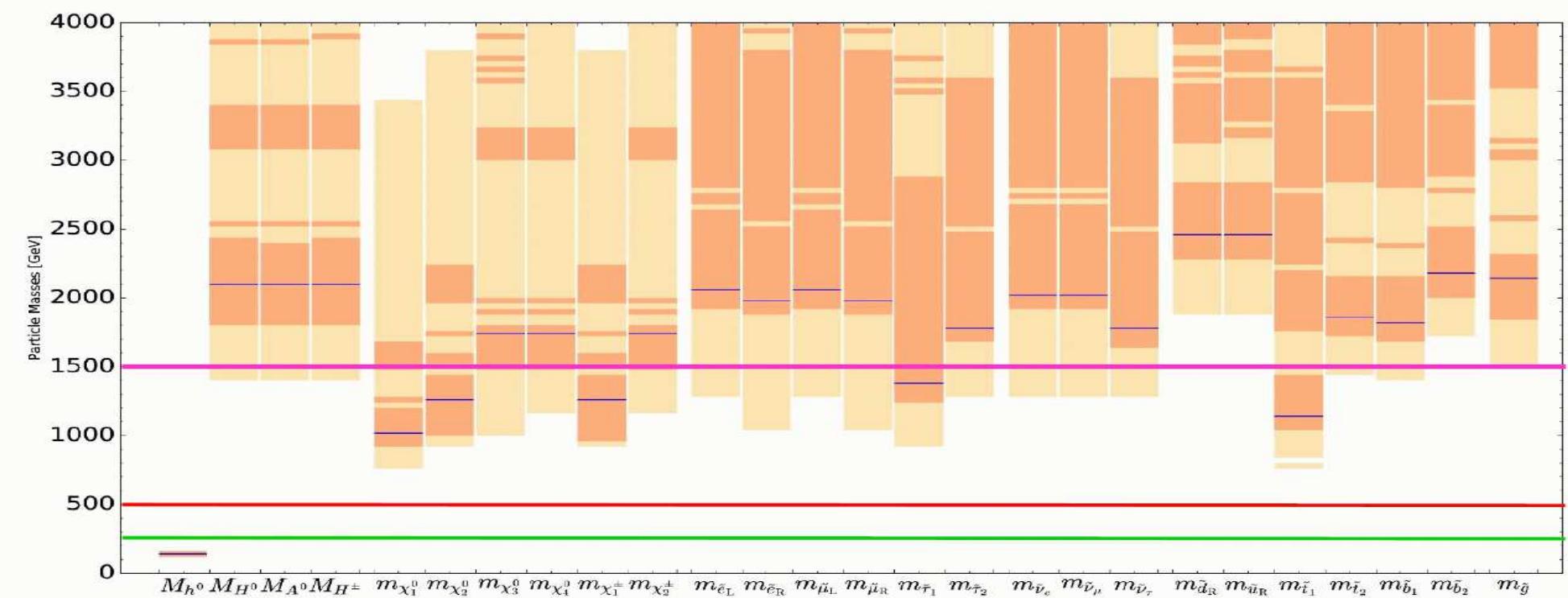
$\tilde{\tau}_1 + \tilde{t}_1$ coann.
$\tilde{t}_1 + \tilde{\chi}_1^\pm$ coann.
$\tilde{\tau}_1$ coann. + $\tilde{t}_1$ coann. + $H/A$

⇒ low  $M_{in}$  possible/favored

⇒ mainly due to  $\text{BR}(B_s \rightarrow \mu^+ \mu^-)$

## sub-GUT prediction: best-fit masses

[2017]



⇒ high colored masses  
⇒ high electroweak masses

ILC:  $\sqrt{s} = 1000 \text{ GeV}$  ⇒ nothing

CLIC:  $\sqrt{s} = 3000 \text{ GeV}$  ⇒ pair production of few SUSY particles  
⇒ no access to colored particles

## Intermediate summary (simplified):

- data: Higgs, LHC searches, DM measurements/searches, EW, flavor
- GUT based models exhibit a heavy spectrum
- very difficult for the LHC
- ILC has to be “lucky” (I did not discuss its great Higgs/EW capabilities)
- CLIC has some particles in reach
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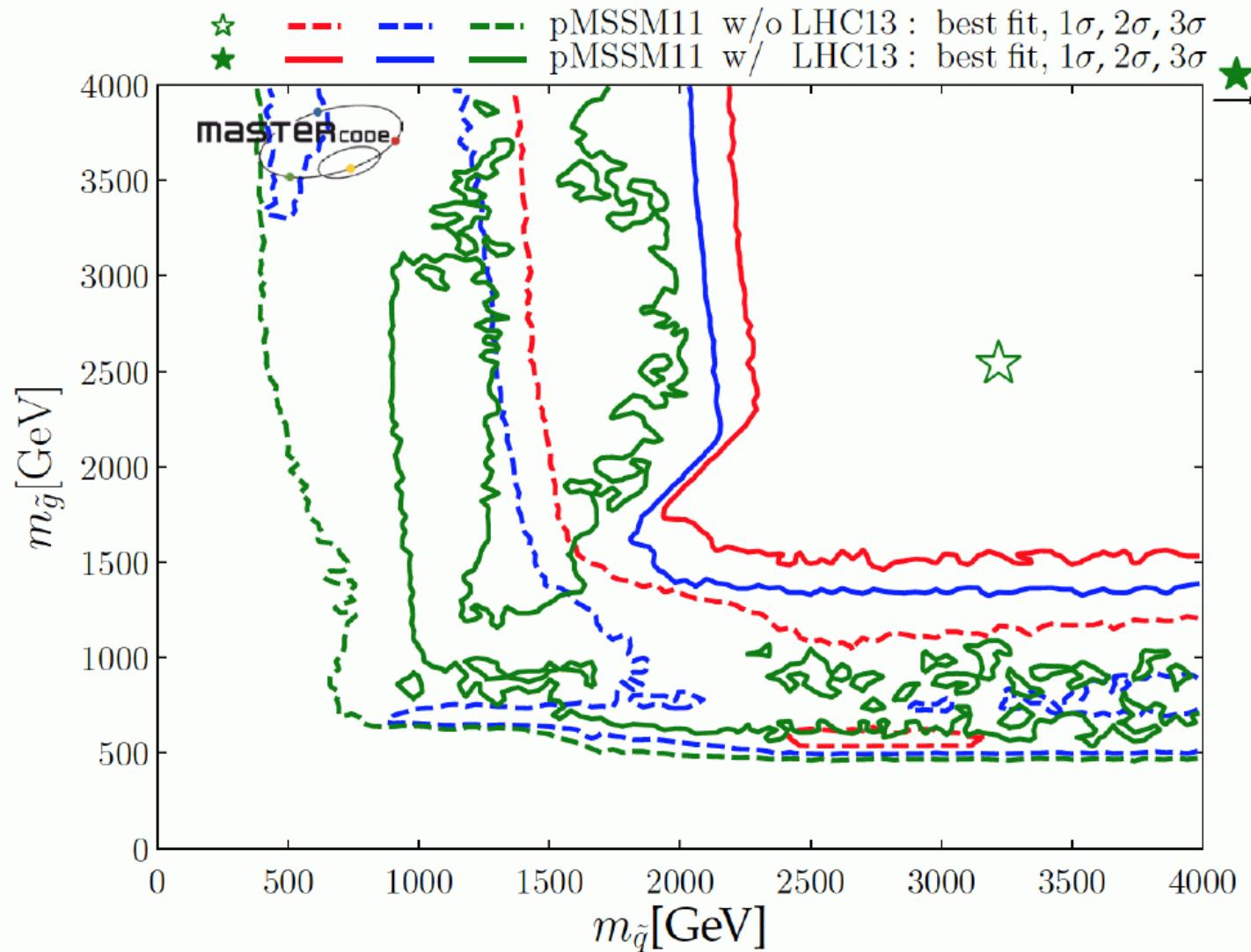
Let's look at the more general pMSSM11!

## 4. pMSSM11 predictions

Parameter	Range	Number of segments
$M_1$	( -4 , 4 ) TeV	6
$M_2$	( 0 , 4 ) TeV	2
$M_3$	( -4 , 4 ) TeV	4
$m_{\tilde{q}}$	( 0 , 4 ) TeV	2
$m_{\tilde{q}_3}$	( 0 , 4 ) TeV	2
$m_{\tilde{l}}$	( 0 , 2 ) TeV	1
$m_{\tilde{\tau}}$	( 0 , 2 ) TeV	1
$M_A$	( 0 , 4 ) TeV	2
$A$	( -5 , 5 ) TeV	1
$\mu$	( -5 , 5 ) TeV	1
$\tan \beta$	( 1 , 60)	1
Total number of boxes		384

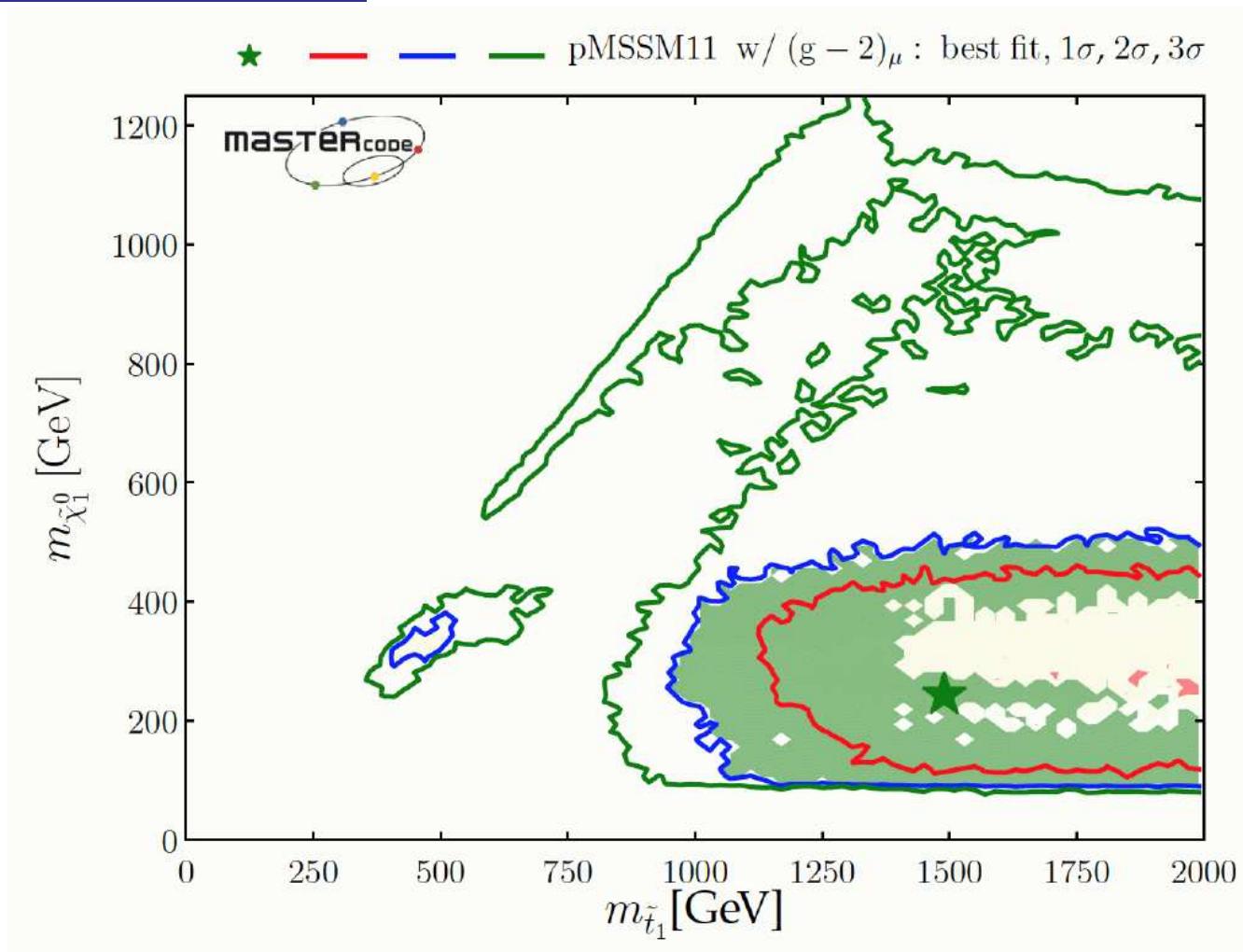
# pMSSM11: Going from 8 TeV to 13 TeV (and adding latest DM limits)

[2017]



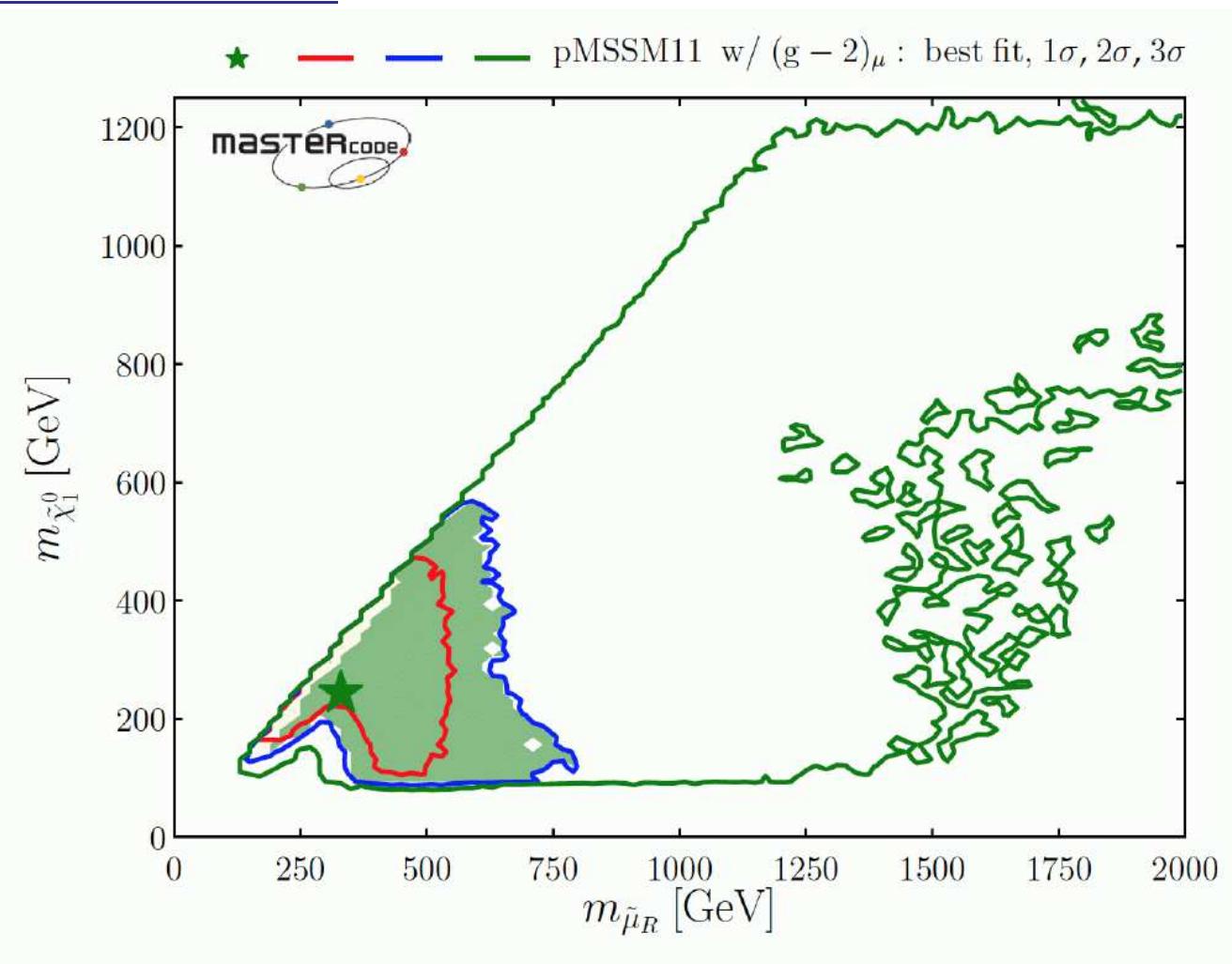
⇒ substantial move to higher masses!

⇒ notice the “nose”!

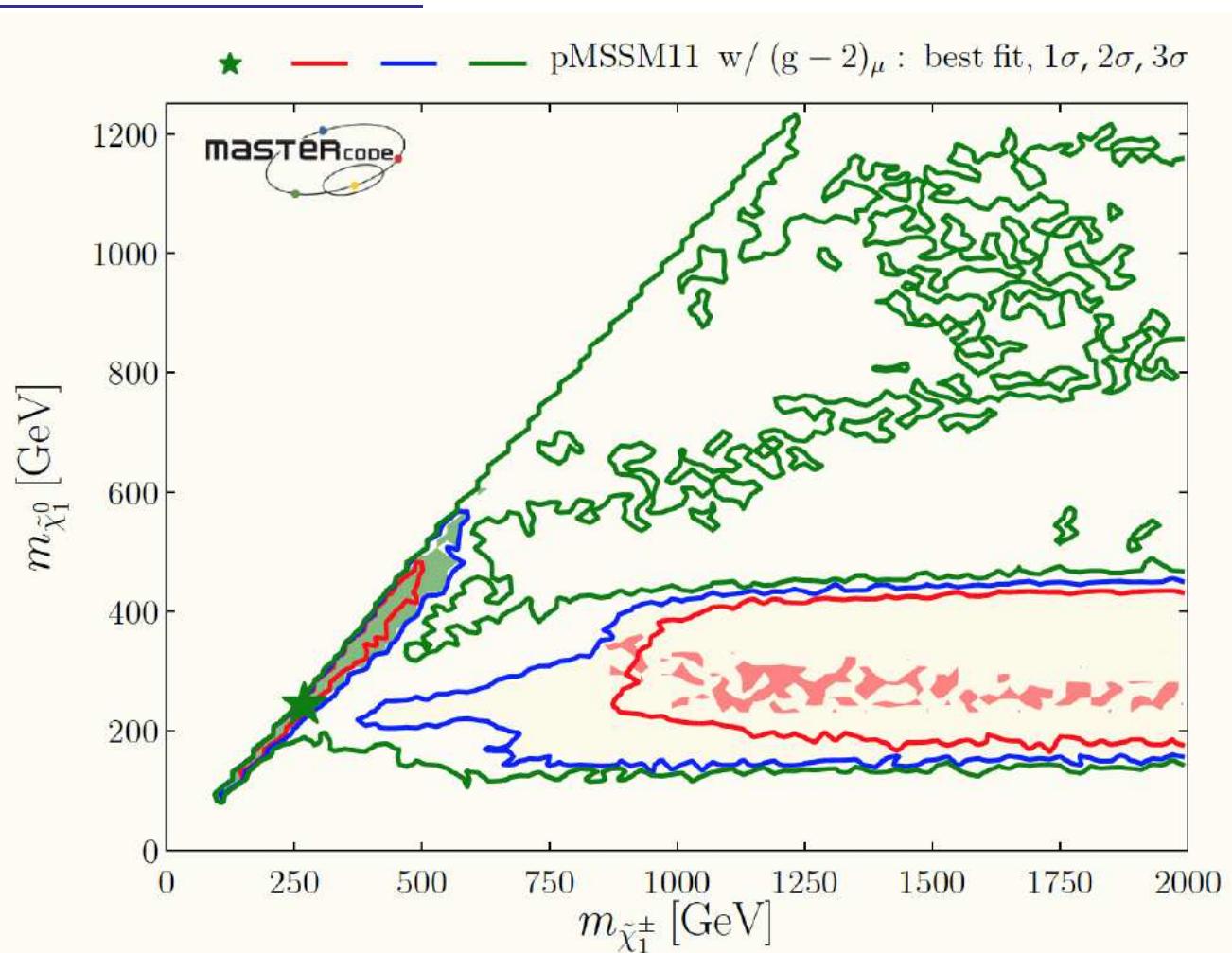


$\tilde{\chi}_1^\pm$ coann.	slep coann.	gluino coann.	stop coann.
A/H funnel	stau coann.	squark coann.	sbot coann.

⇒ high (low) stop (neutralino) masses ⇒ notice the compressed region!



⇒ all masses low!!



$\tilde{\chi}_1^\pm$ coann.	slep coann.	gluino coann.	stop coann.
A/H funnel	stau coann.	squark coann.	sbot coann.

⇒ chargino co-annihilation

⇒  $M_1 \sim M_2$

## pMSSM11: best-fit point parameters

[2017]

Parameter	With LHC 13 TeV and $(g - 2)_\mu$	
	Best fit	'Nose' region
$M_1$	0.25 TeV	- 0.39 TeV
$M_2$	0.25 TeV	1.2 TeV
$M_3$	- 3.86 TeV	- 1.7 TeV
$m_{\tilde{q}}$	4.0 TeV	2.00 TeV
$m_{\tilde{q}_3}$	1.7 TeV	4.1 TeV
$m_{\tilde{\ell}}$	0.35 TeV	0.36 TeV
$m_{\tilde{\tau}}$	0.46 TeV	1.4 TeV
$M_A$	4.0 TeV	4.2 TeV
$A$	2.8 TeV	5.4 TeV
$\mu$	1.33 TeV	- 5.7 TeV
$\tan \beta$	36	19
$\chi^2/\text{d.o.f.}$	22.1/20	24.46/20
p-value	0.33	0.22
$\chi^2(HS)$	68.01	67.97

⇒ excellent  $p$  value!

## What to conclude?

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⇒ Look at the *p* values!

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Model	Min. $\chi^2/\text{dof}$	$\chi^2$ -prob. ( $p$ -value)
CMSSM	32.8/18	11%
NUHM1	31.1/23	12%
NUHM2	30.3/22	11%
SU(5)	32.4/23	9%
mAMSB	36.5/27	11%
sub-GUT	28.9/24	23%
pMSSM11	21.0/20	33%

Which model is more likely??

What to conclude?

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⇒ pMSSM11: model with higher  $\chi^2$ -probability  
model with good ILC/CLIC prospects  
detailed LHC analysis tbd!

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detailed LHC analysis tbd!

⇒ Dark Matter prospects? ⇒ talk by S.H. later today

## 5. Conclusions

- SUSY is (still) the best-motivated BSM scenario
  - constrained models: CMSSM, NUHM, SU(5), mAMSB, sub-GUT
  - general models: pMSSM11, ...
- MasterCode: LHC, Higgs, EWPO, BPO, CDM  $\Rightarrow \chi^2$  evaluation

•

Model	Min. $\chi^2/\text{dof}$	$\chi^2$ -prob. ( $p$ -value)
GUT based models	(30 ... 33)/(18 ... 23)	$\sim 11\%$
pMSSM11	21.0/20	33%

Particle	GUT-based models	pMSSM11
gauginos	ILC CLIC	ILC CLIC
sleptons	CLIC	ILC CLIC
stops/sbottoms		CLIC
other		

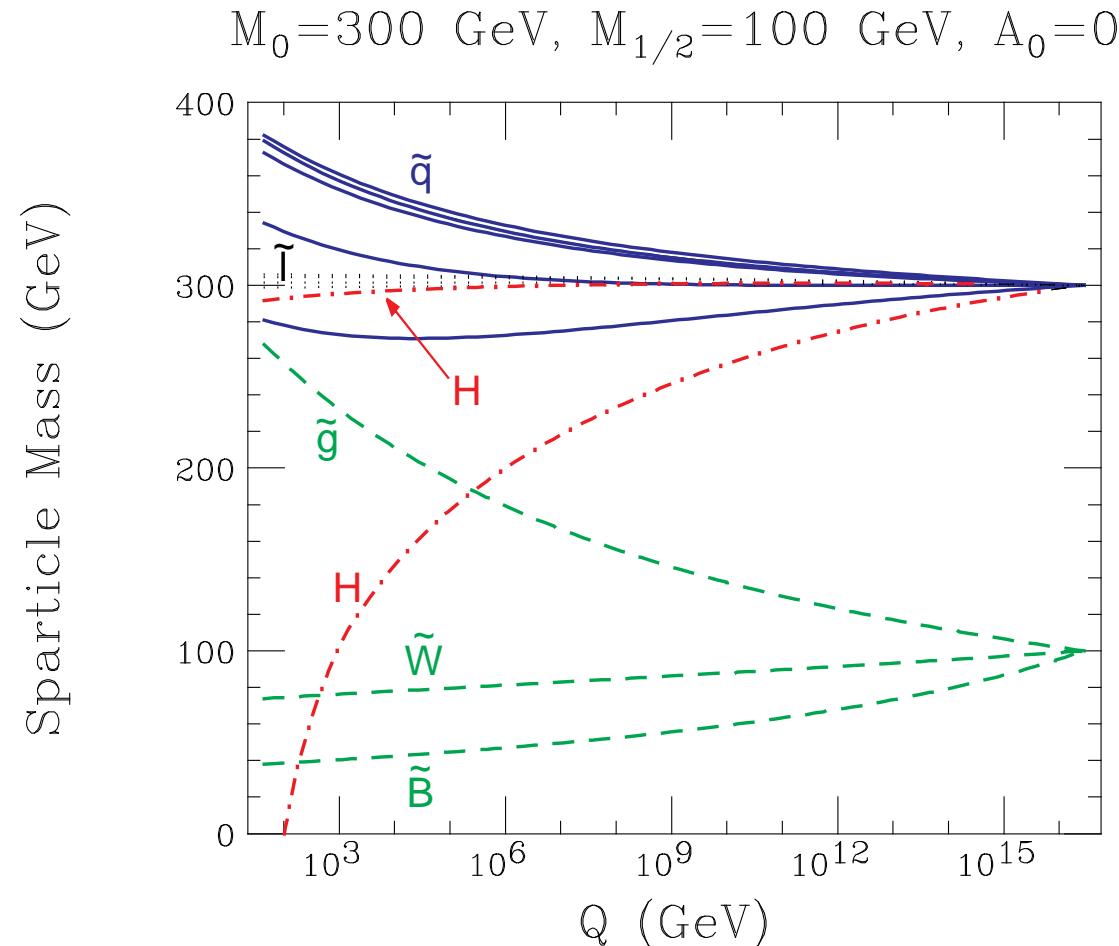
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Further Questions?

## GUT based models: 1.) CMSSM (sometimes wrongly called mSUGRA):

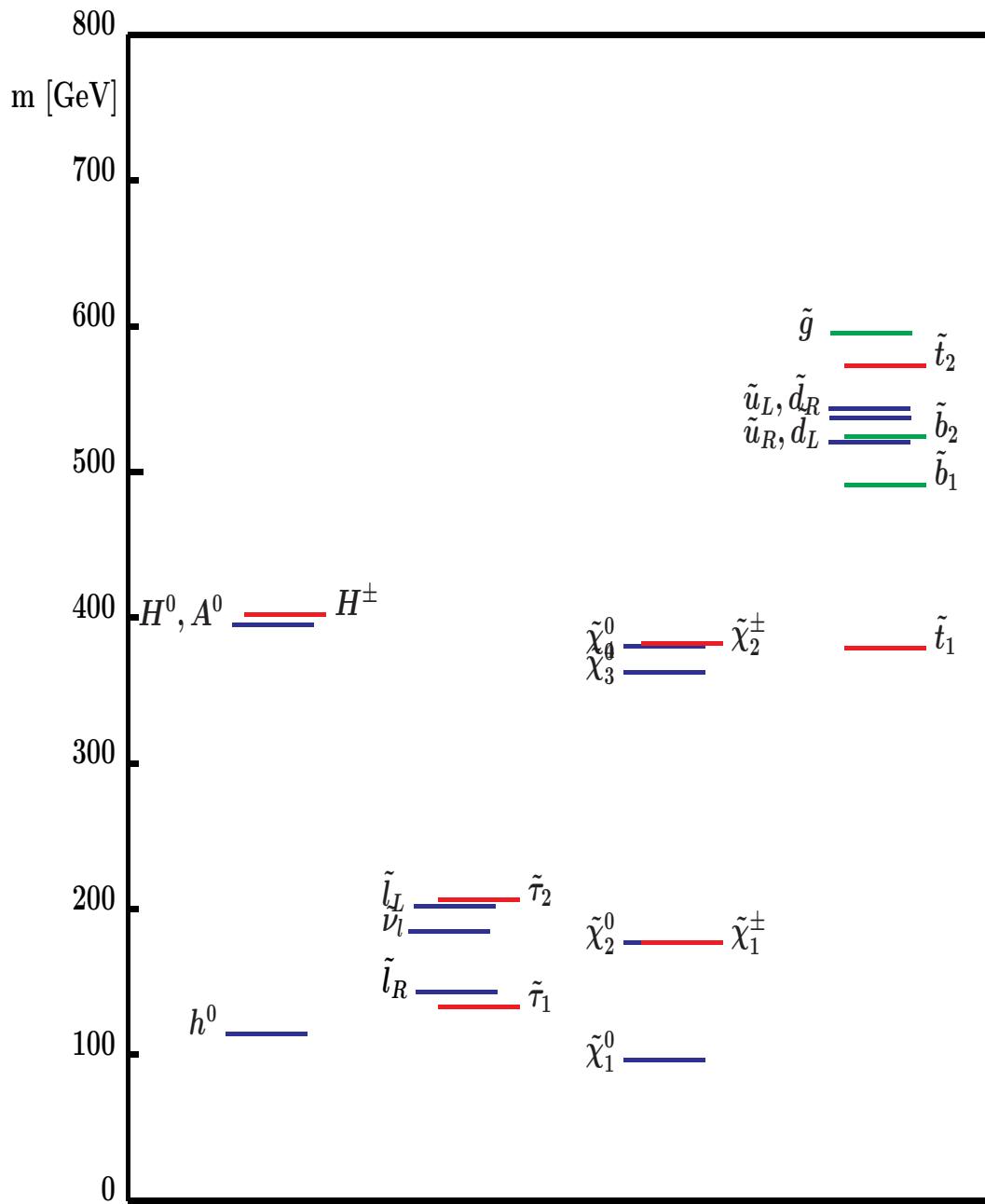
⇒ particle spectra from renormalization group running to weak scale



⇒ one parameter turns negative ⇒ Higgs mechanism for free

“Typical” CMSSM scenario  
(SPS 1a benchmark scenario):

Strong connection between  
all the sectors



## GUT based models: 2.) NUHM1: (Non-universal Higgs mass model)

Assumption: no unification of scalar fermion and scalar Higgs parameter at the GUT scale

⇒ effectively  $M_A$  as free parameters at the EW scale

⇒ Scenario characterized by

$$m_0, m_{1/2}, A_0, \tan\beta, \text{sign } \mu \text{ and } M_A$$

## GUT based models: 3.) NUHM2: (Non-universal Higgs mass model 2)

Assumption: no unification of scalar Higgs parameter at the GUT scale

⇒ effectively  $M_A$  and  $\mu$  as free parameters at the EW scale

⇒ Scenario characterized by

$$m_0, m_{1/2}, A_0, \tan\beta, \mu \text{ and } M_A$$

## GUT based models: 4.) SU(5) GUT:

Assumption I:

no unification of scalar Higgs parameter at the GUT scale

( $\Rightarrow$  effectively  $M_A$  and  $\mu$  as free parameters at the EW scale)

Assumption II:

$$(q_L, u_L^c, e_L^c)_i \in \mathbf{10}_i, (\ell_L, d_L^c)_i \in \bar{\mathbf{5}}_i$$

$\Rightarrow$  Scenario characterized by

$$m_5, m_{10}, m_{1/2}, A_0, \tan \beta, m_{H_u}, m_{H_d}$$

## GUT based models: 5.) mAMSB:

mAMSB scenario characterized by

$$m_{3/2}, m_0, \tan \beta, \text{sign}(\mu)$$

$m_{3/2} = \langle F \rangle / M_{\text{Planck}}$ : overall scale of SUSY particle masses

$m_0$ : phenomenological parameter: universal scalar mass term introduced in order to keep squares of slepton masses positive

typical feature: very small neutralino–chargino mass difference  
 $\Rightarrow \tilde{\chi}_1^\pm \rightarrow \tilde{\chi}_1^0 + \pi^\pm$  with very soft pions

## GUT based models: 6.) sub-GUT:

Based on CMSSM with unification at  $M_{\text{GUT}} \sim 2 \cdot 10^{16}$  GeV:

⇒ Scenario characterized by

$$m_0, m_{1/2}, A_0, \tan \beta, \text{sign } \mu$$

Unification is assumed at  $M_{\text{in}} \leq M_{\text{GUT}}$ :

⇒ Scenario characterized by

$$M_{\text{in}}, m_0, m_{1/2}, A_0, \tan \beta, \text{sign } \mu$$

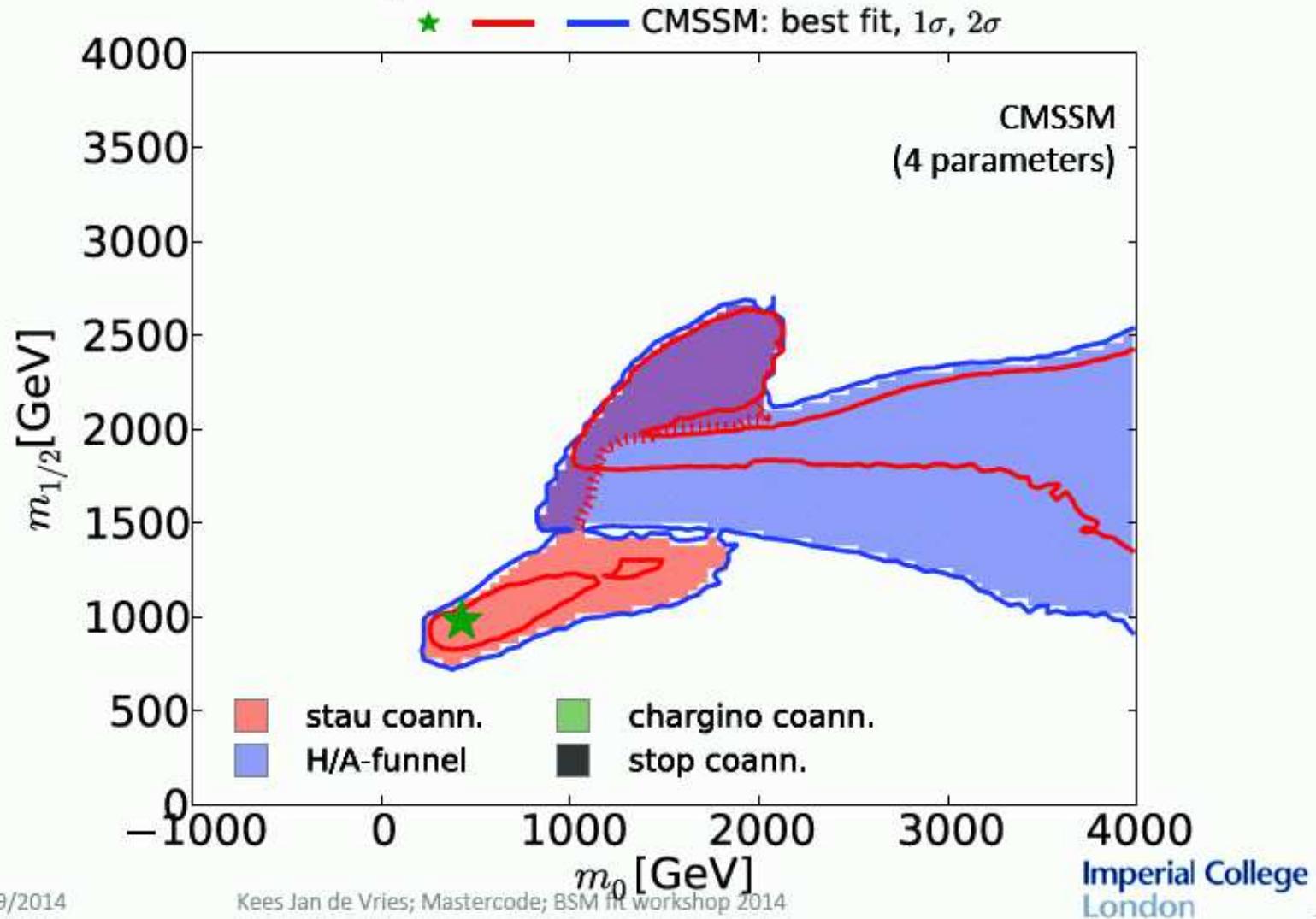
Possible realization in “mirage unification”  
warped extra dimensions

...

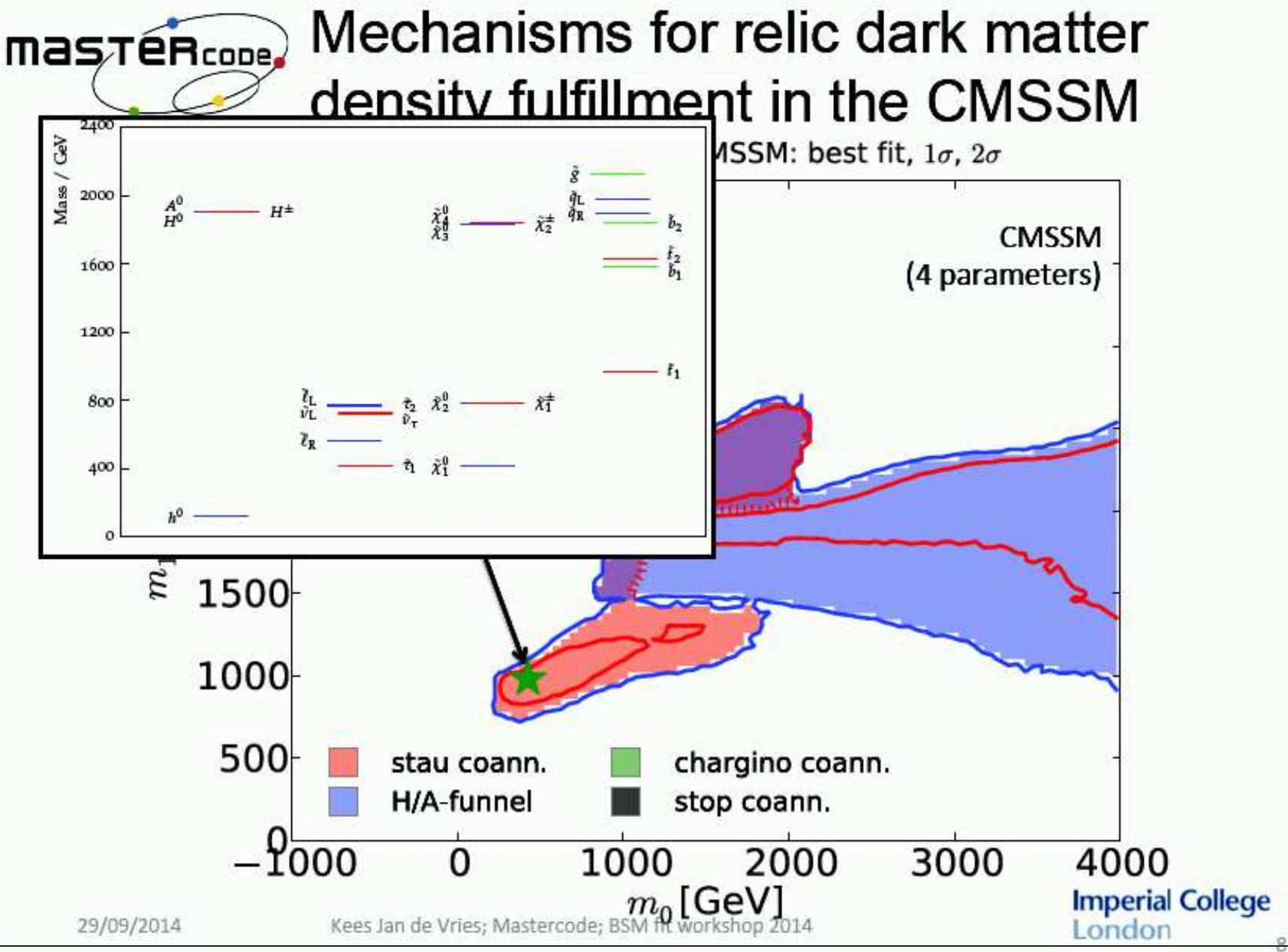
## CMSSM best-fit point prediction



## Mechanisms for relic dark matter density fulfillment in the CMSSM



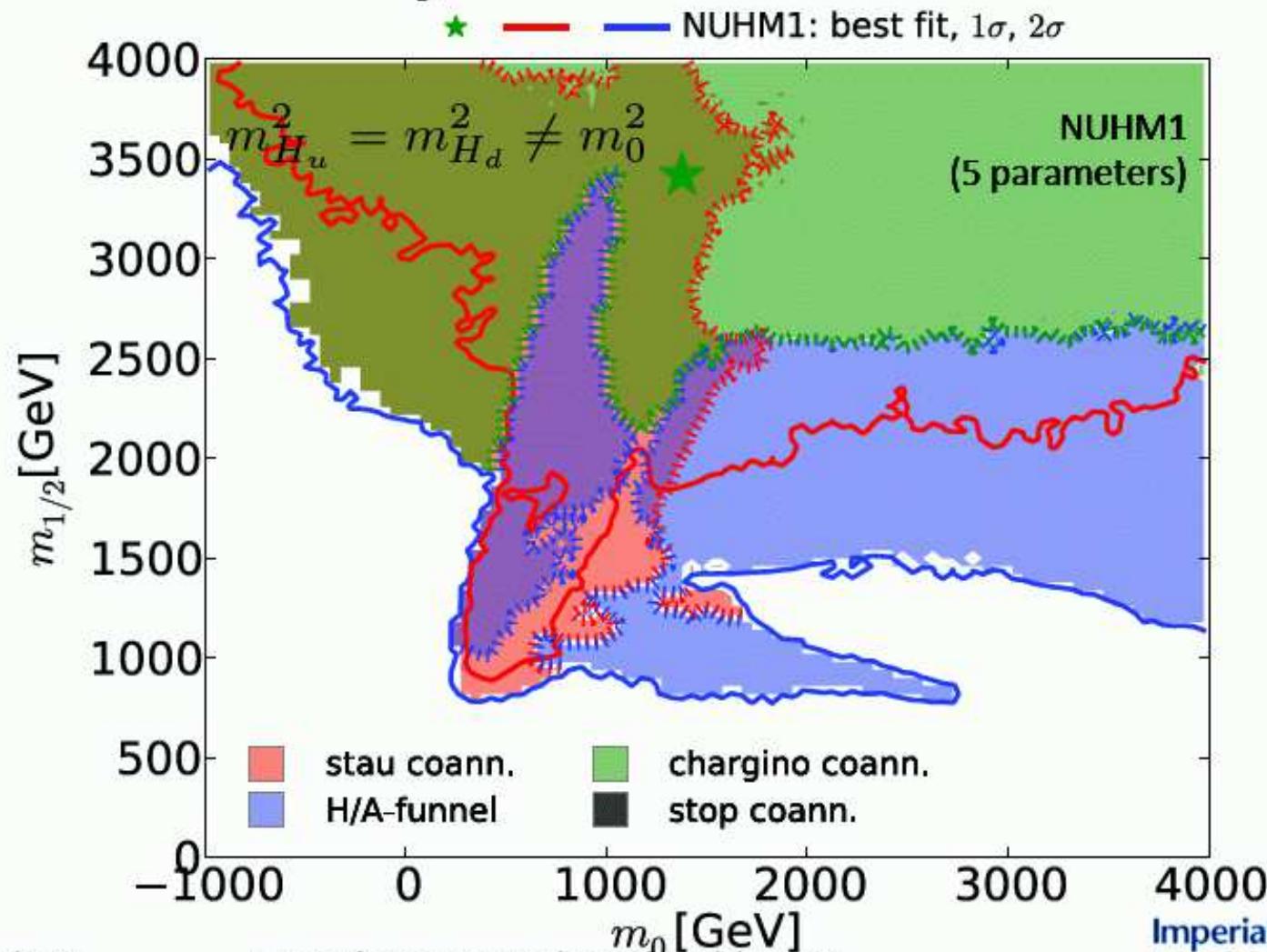
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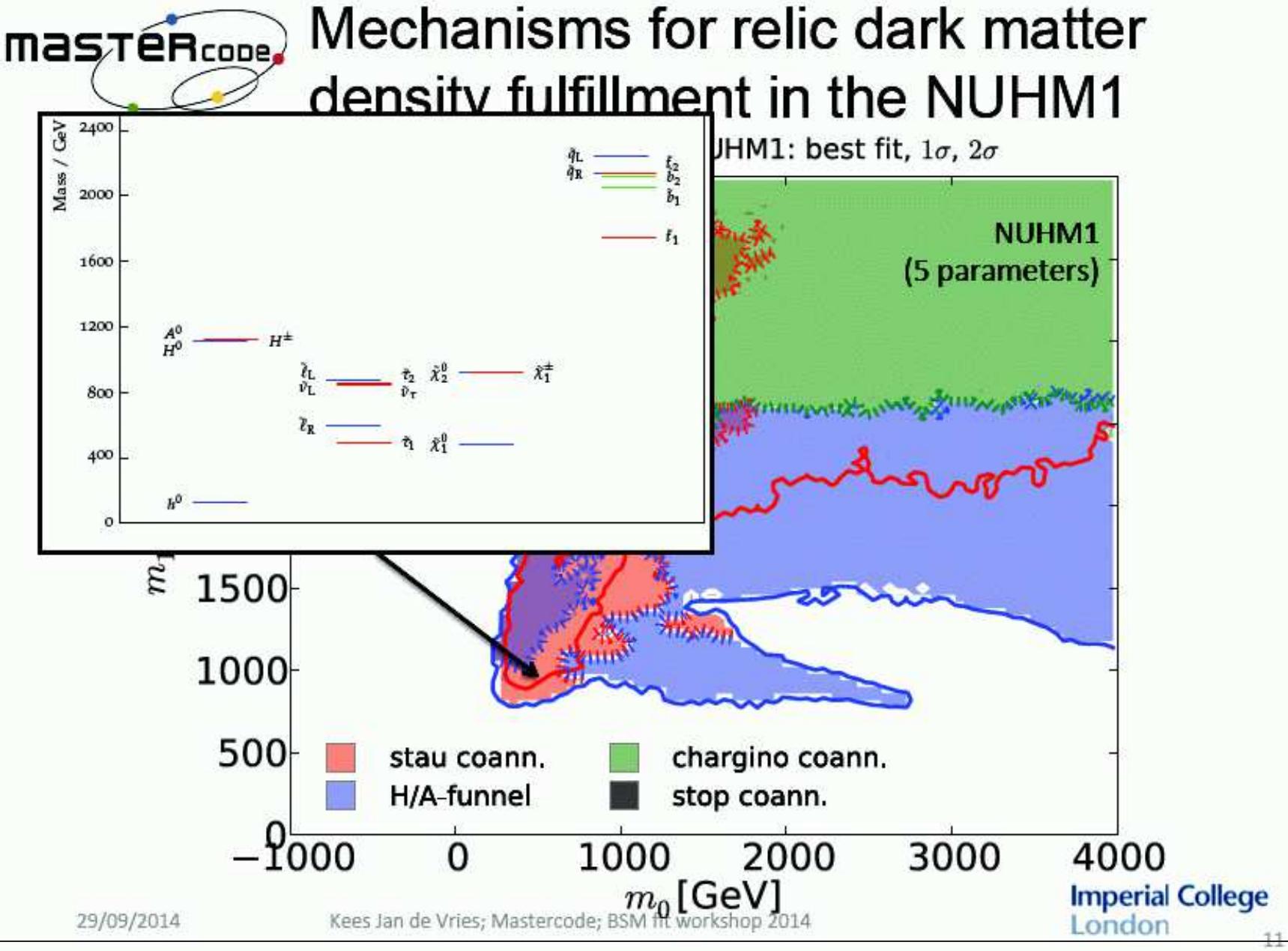
## NUHM1 best-fit point prediction



# Mechanisms for relic dark matter density fulfillment in the NUHM1



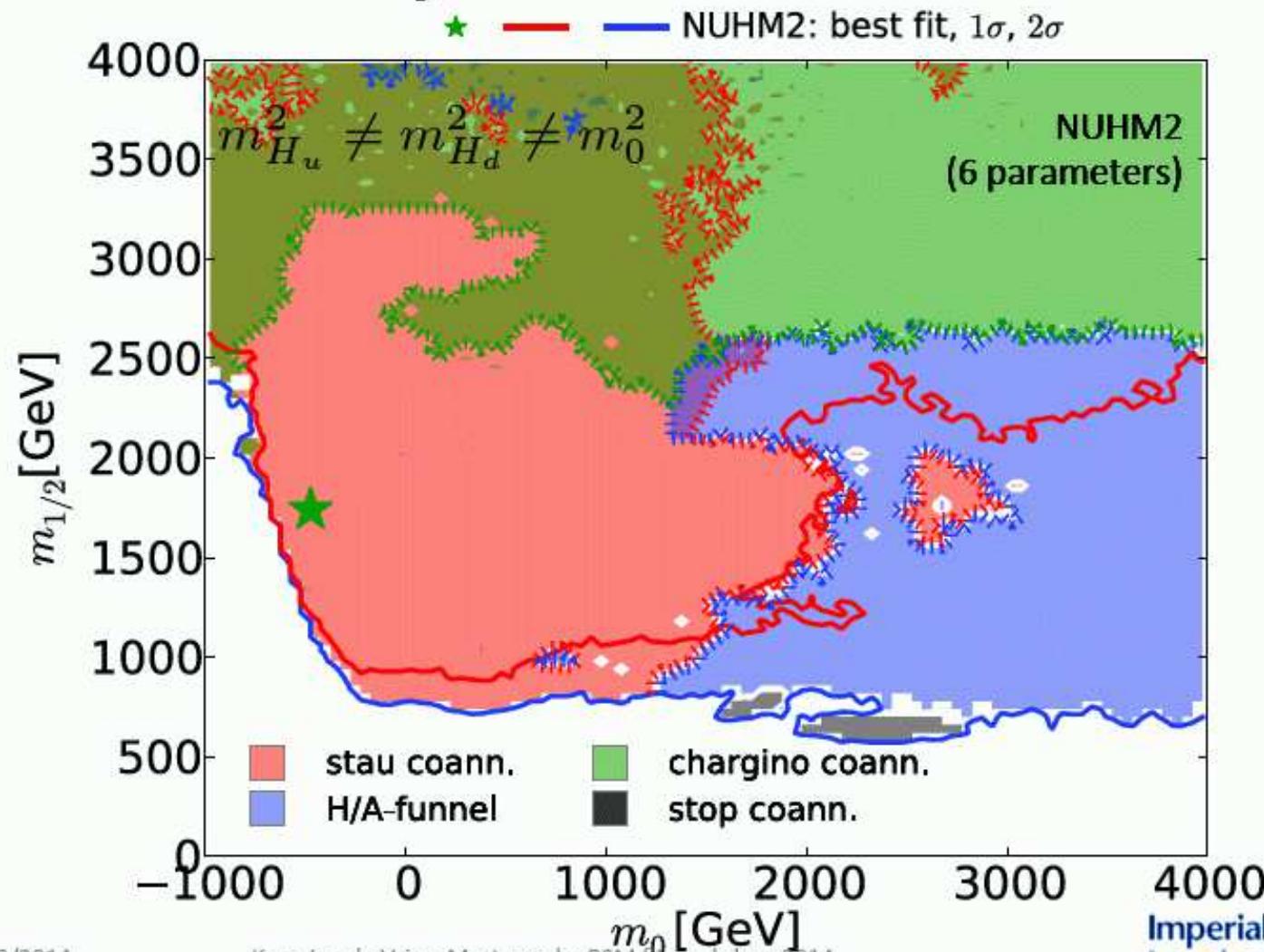
## NUHM1 best-fit point prediction



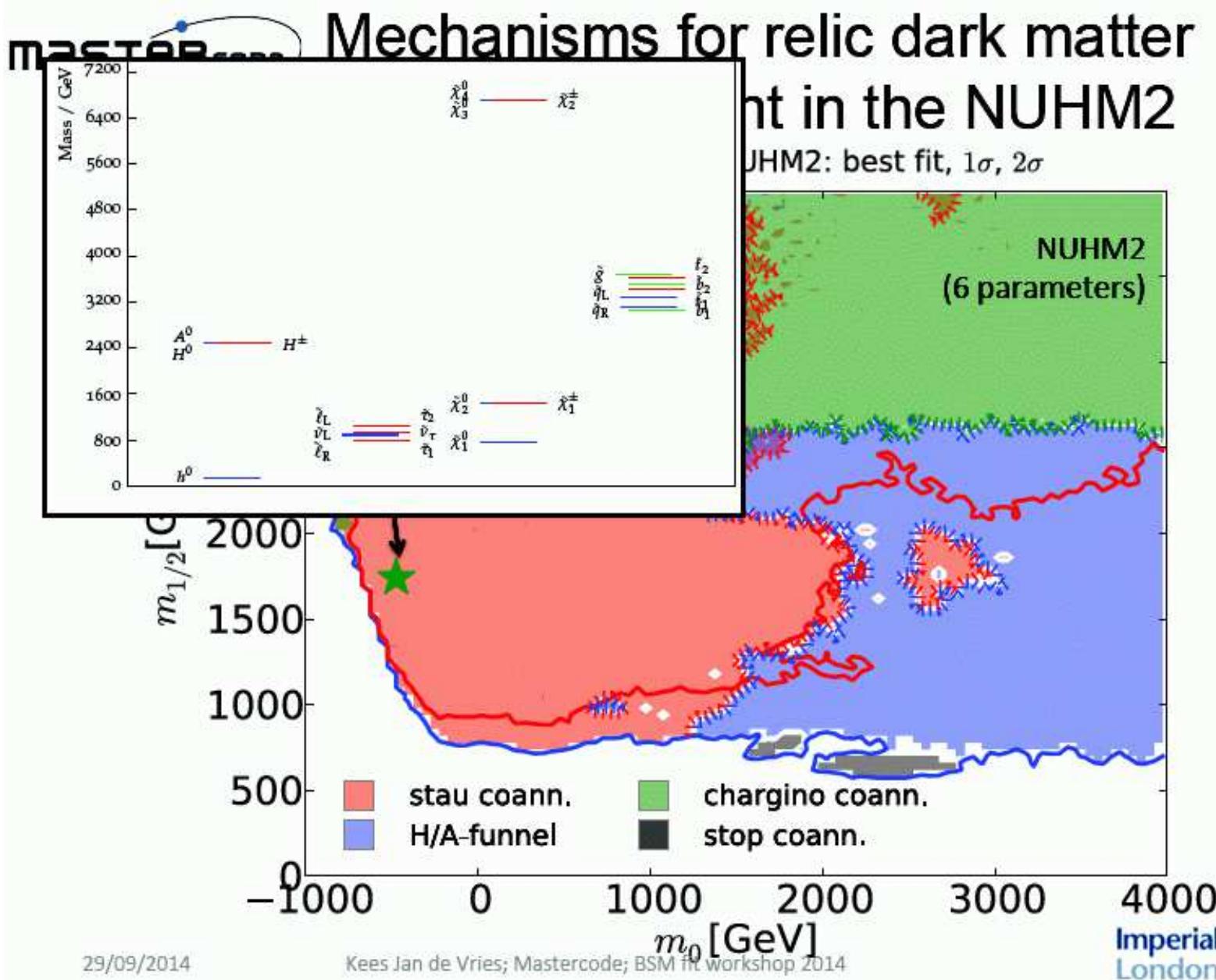
## NUHM2 best-fit point prediction



# Mechanisms for relic dark matter density fulfillment in the NUHM2

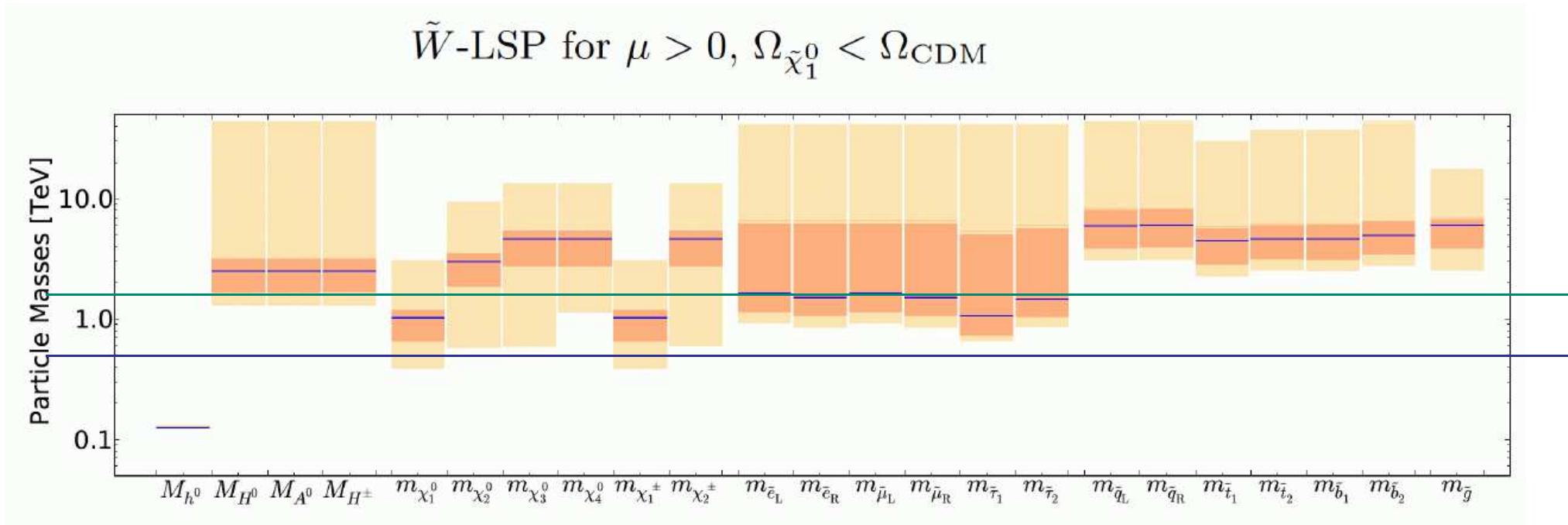


## NUHM2 best-fit point prediction



## mAMSB prediction: best-fit masses (wino)

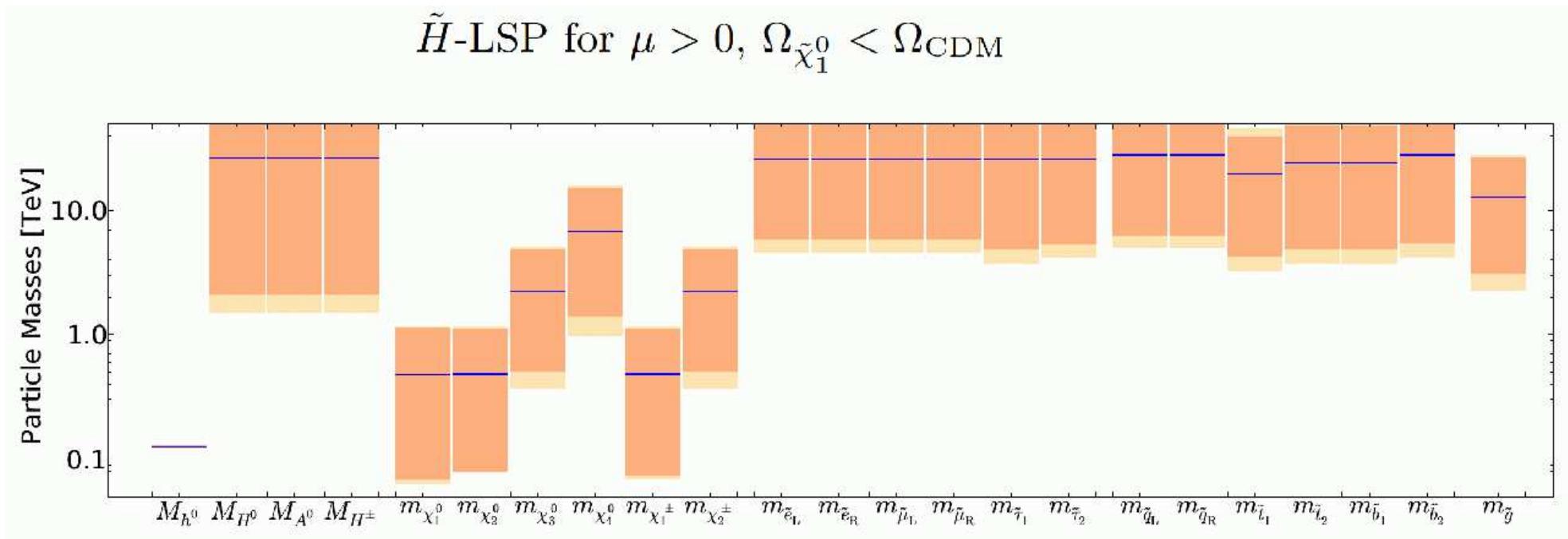
[2016]



ILC:  $\sqrt{s} = 1000$  GeV  $\Rightarrow$  bad prospects

CLIC:  $\sqrt{s} = 3000$  GeV  $\Rightarrow$  pair production of few SUSY particles “likely”  
 $\Rightarrow$  no access to colored particles

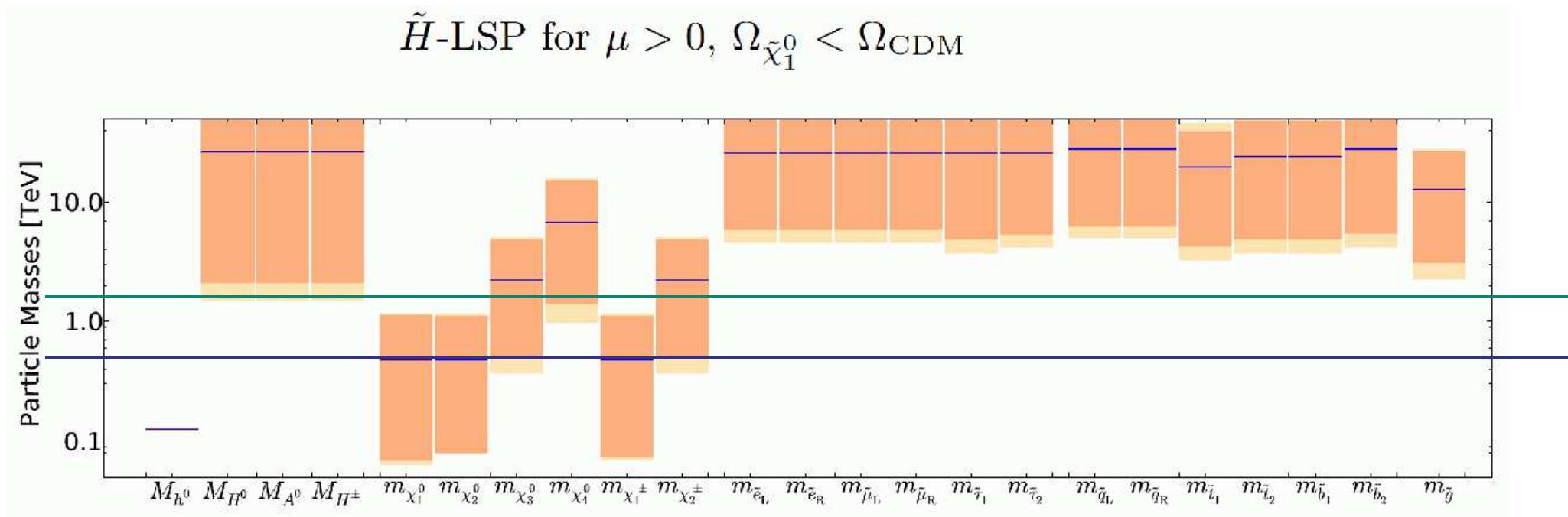
$\tilde{H}$ -LSP for  $\mu > 0$ ,  $\Omega_{\tilde{\chi}_1^0} < \Omega_{\text{CDM}}$



- ⇒ high colored masses
- ⇒ some(!) lower electroweak masses  
partially with not too large  $2\sigma$  ranges
- ⇒ clear prediction for ILC and CLIC

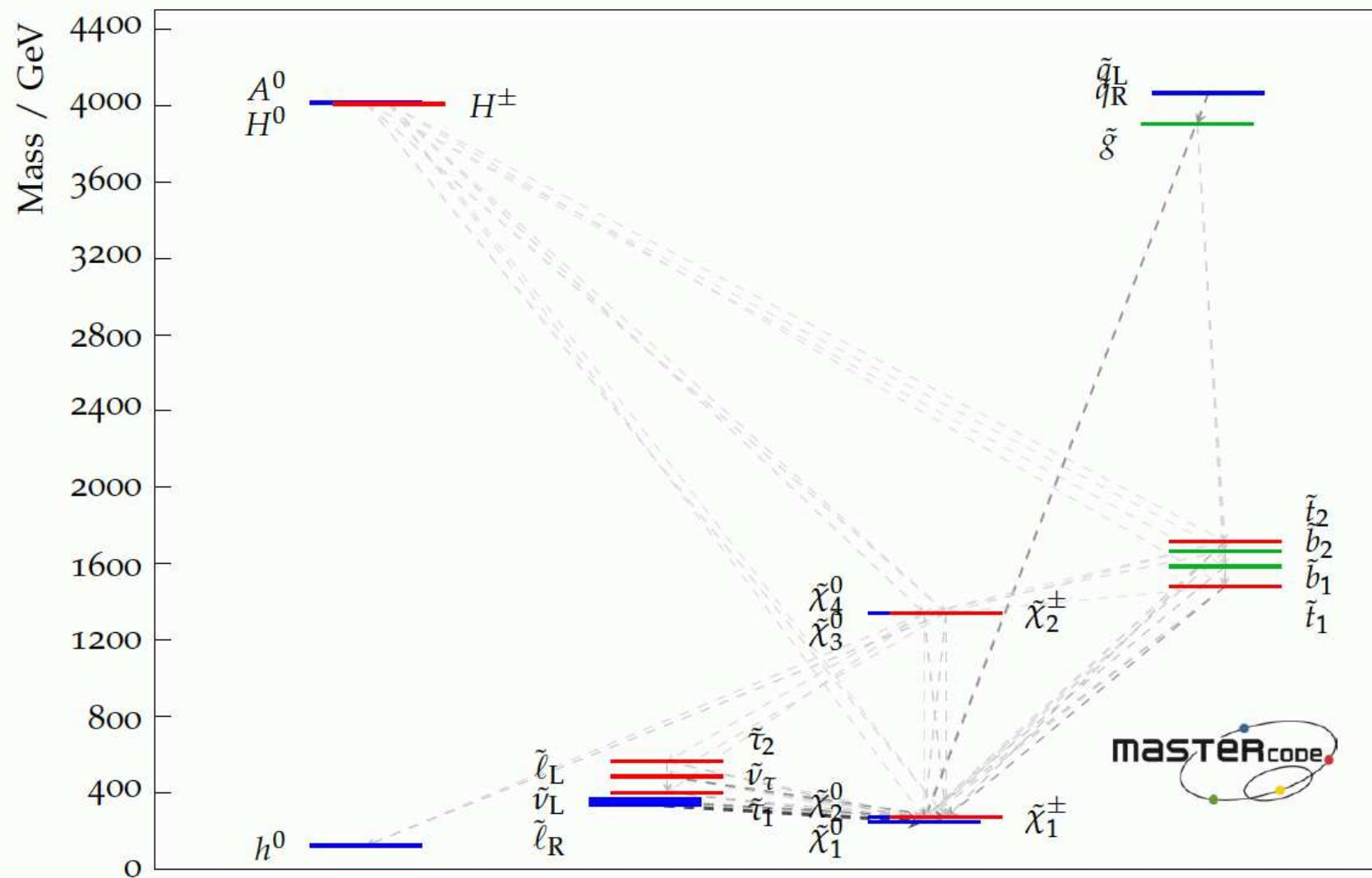
## mAMSB prediction: best-fit masses (higgsino)

[2016]

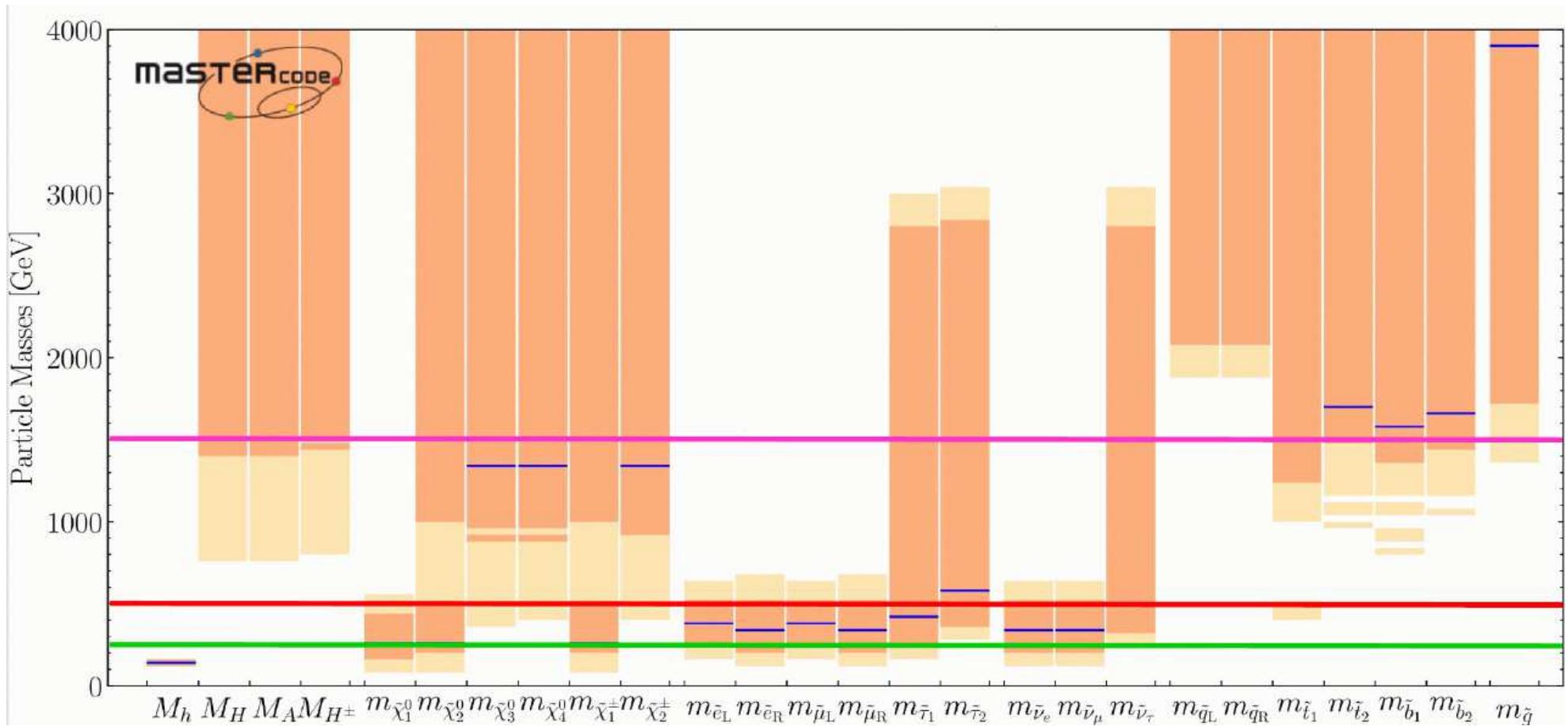


ILC:  $\sqrt{s} = 1000$  GeV  $\Rightarrow$  few EW particles possibly accessible

CLIC:  $\sqrt{s} = 3000$  GeV  $\Rightarrow$  pair production of few SUSY particles  
“guaranteed”  
 $\Rightarrow$  no access to colored particles



⇒ heavy colored, light uncolored spectrum



ILC:  $\sqrt{s} = 500 \text{ GeV} \Rightarrow$  some particles might be in reach

ILC:  $\sqrt{s} = 1000 \text{ GeV} \Rightarrow$  precision analysis of EW particle and DM easy!

CLIC:  $\sqrt{s} = 3000 \text{ GeV} \Rightarrow$  precision analysis of EW particles and DM easy!