

The logo for SHINE consists of the word "SHINE" in a bold, blue, sans-serif font. The letter "S" is significantly larger than the other letters. To the right of the "S", there are two small blue dots. Behind the "S" and the dots, there are several vertical, slightly curved lines in blue and orange, creating a sense of motion or energy.

SHINE

NAGI/SHINE PLANS FOR 2018 - 2024

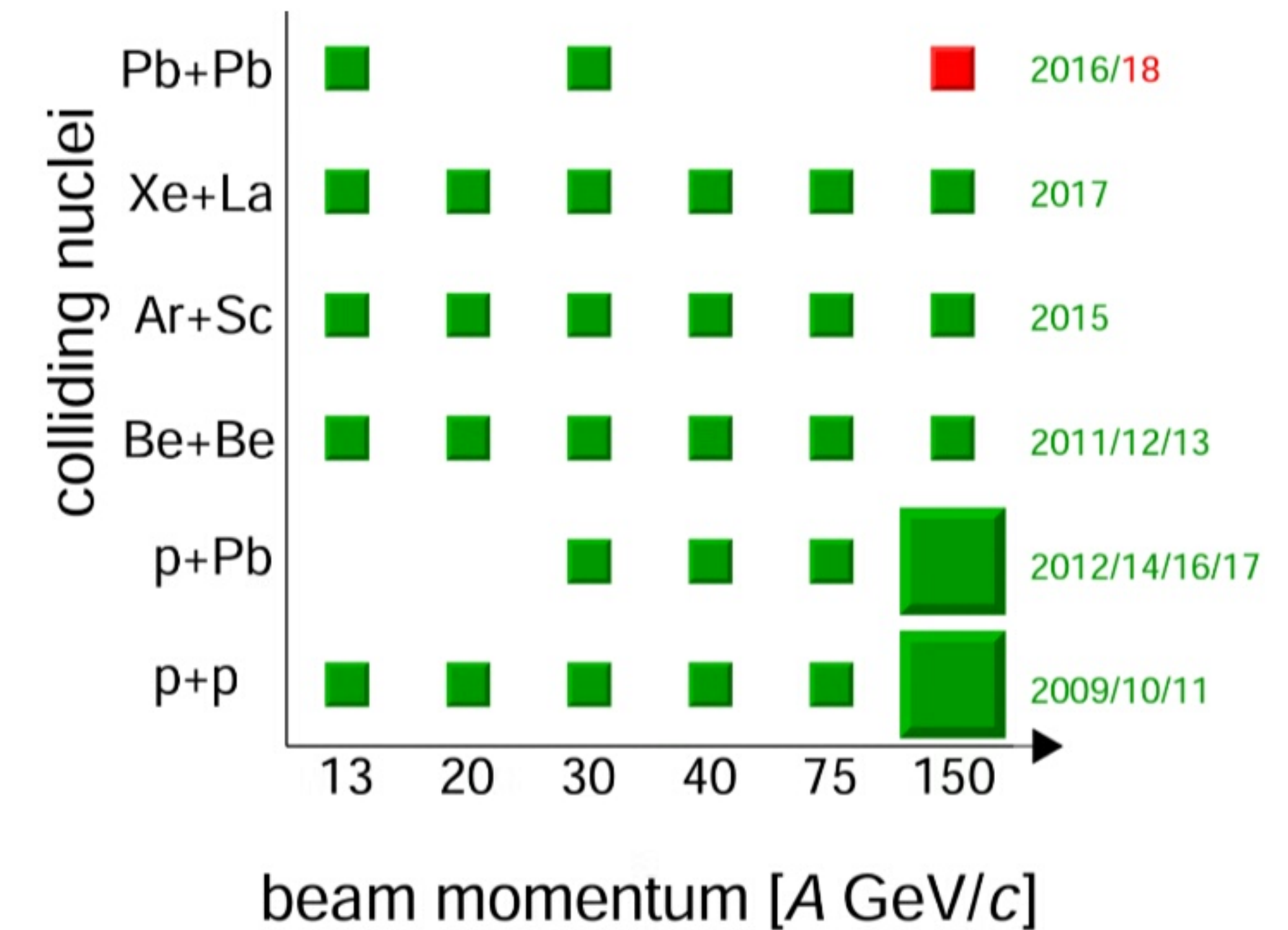
M. GAZDZICKI (FRANKFURT, KIELCE)
FOR THE NAGI/SHINE COLLABORATION

RESULTS FROM DATA RECORDED IN 2009 - 18

②

NA61/SHINE RECORDED UNIQUE DATA FOR:

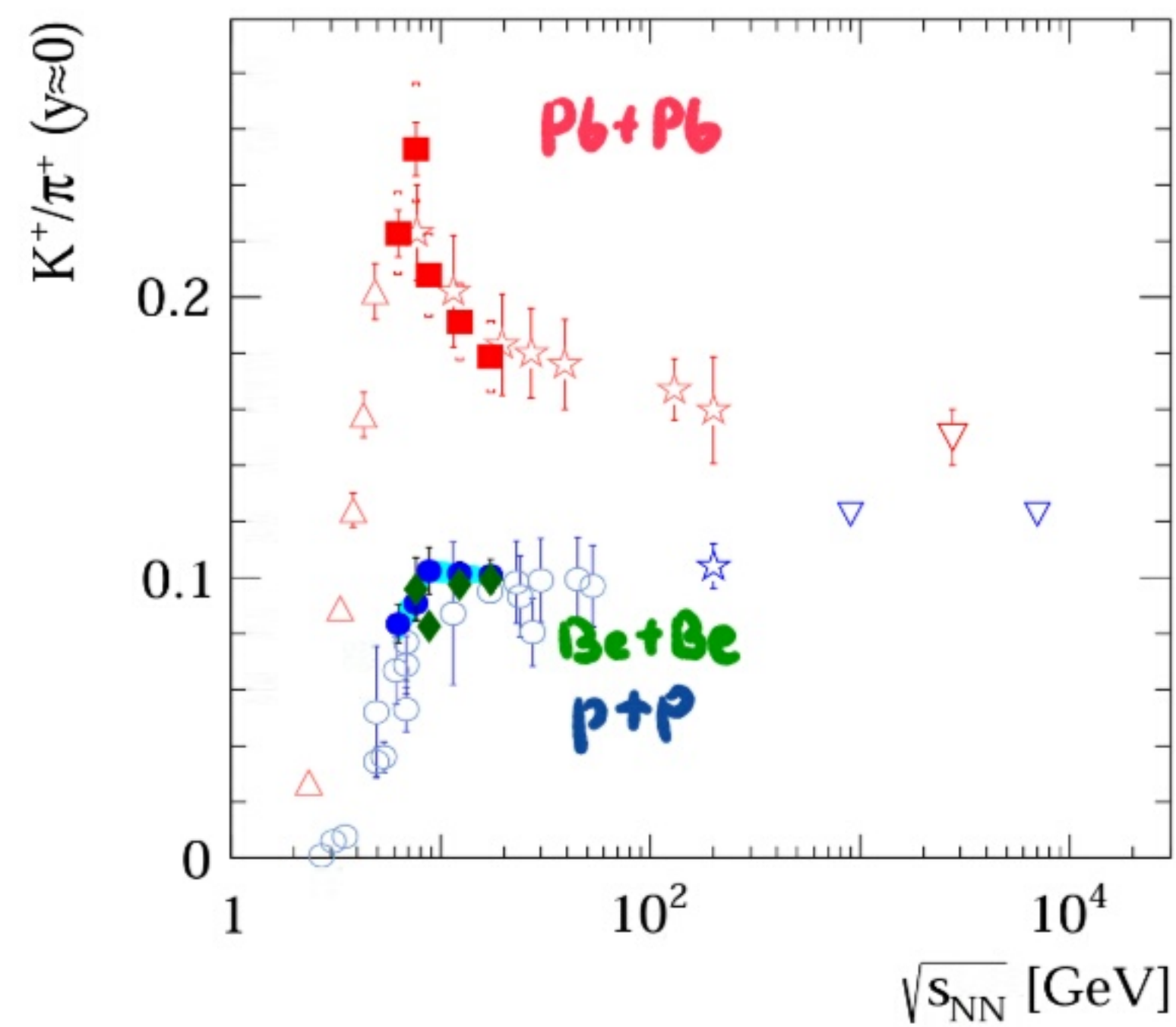
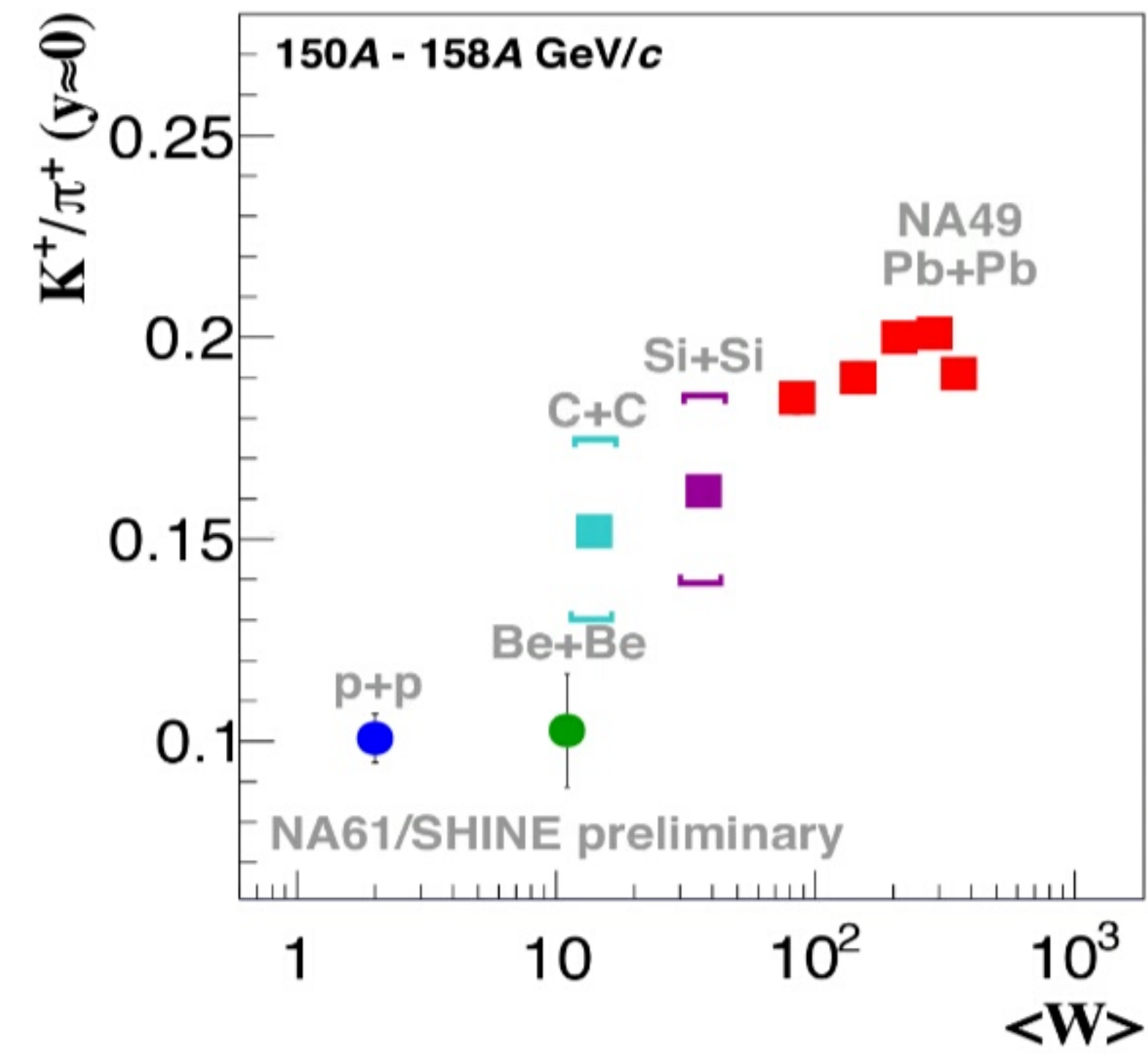
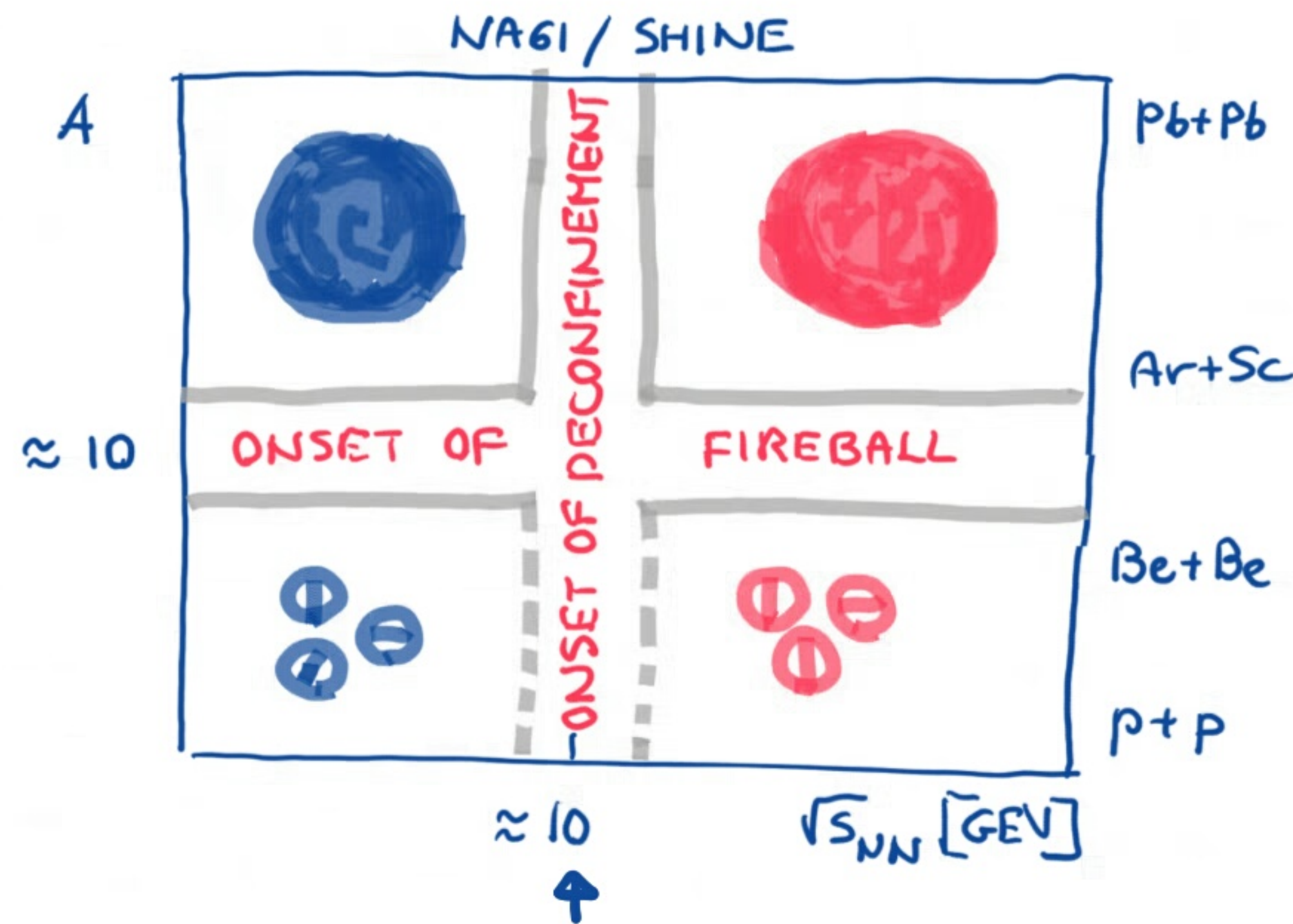
- HEAVY ION PHYSICS (ONSET OF DECONFINEMENT
ONSET OF FIREBALL, CRITICAL POINT)
- NEUTRINO PHYSICS (T2K AT J-PARC
FERMILAB EXPERIMENTS)
- COSMIC-RAY PHYSICS (EXTENSIVE AIR SHOWERS
AUGER, KASCADE)



DATA ANALYSIS WILL CONTINUE FOR AT LEAST FIVE YEARS

UNIQUENESS OF HEAVY ION PHYSICS AT THE CERN SPS

ONSET OF DECONFINEMENT AND ONSET OF FIREBALL:



TWO ONSETS AT SPS

UNIQUENESS OF HEAVY ION PHYSICS AT THE CERN SPS:

④

CRITICAL POINT:

- SECOND ORDER PHASE TRANSITION → SCALE INVARIANCE →
- CHARACTERISTIC DEPENDENCE OF FLUCTUATIONS ON SIZE δ OF SUBDIVISION INTERVALS OF MOMENTUM SPACE Δ
 $M = \Delta/\delta$ - NUMBER OF INTERVALS

$$F_2(M) \equiv \frac{\sum_{i=1}^M \langle N_i (N_i - 1) \rangle}{\sum_{i=1}^M \langle N_i \rangle^2}$$

WHERE N_i - PARTICLE NUMBER IN BIN i ,
 $\langle \dots \rangle$ - AVERAGING OVER EVENTS

AT CRITICAL POINT POWER LAW DEPENDENCE IS EXPECTED

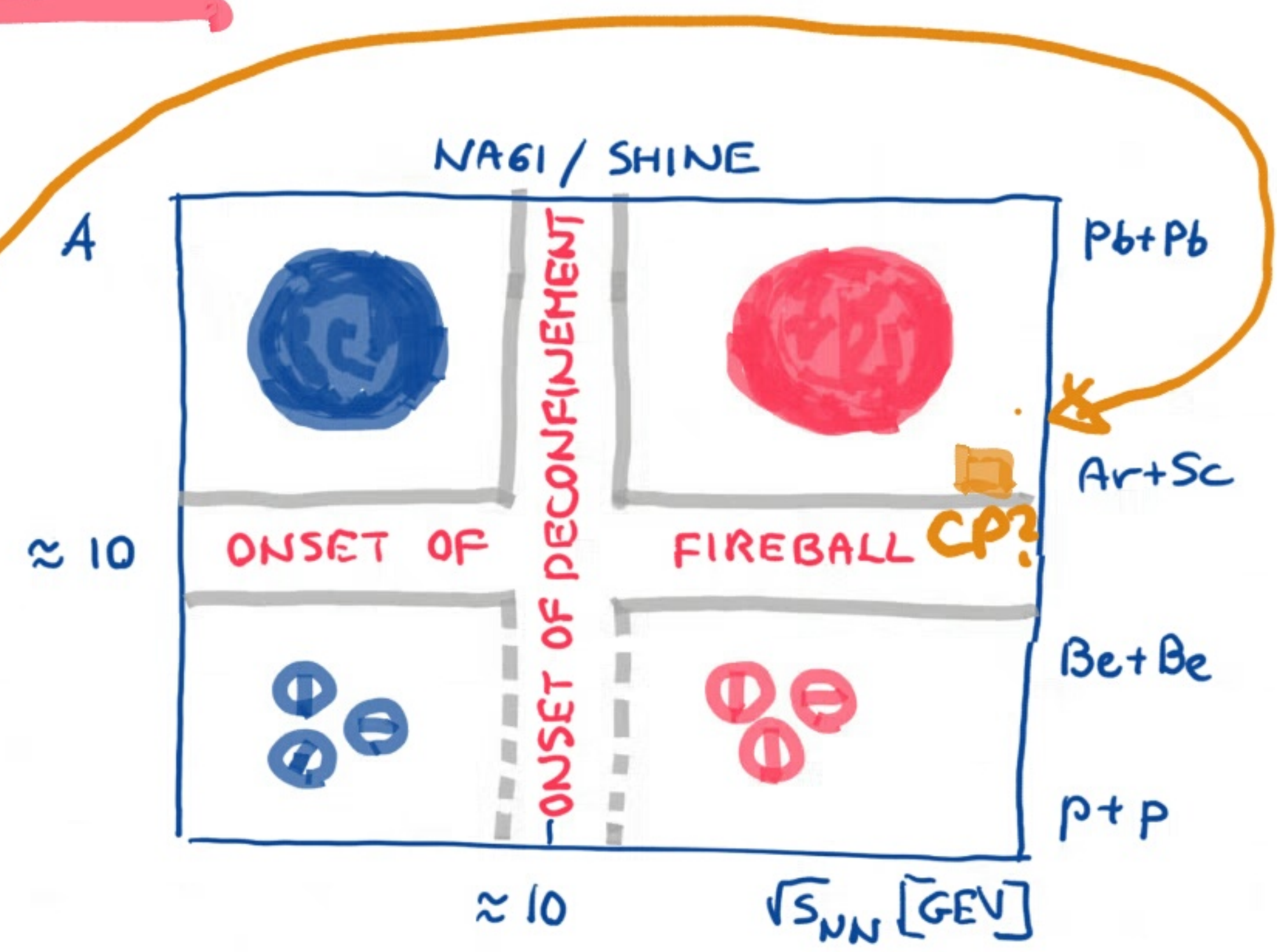
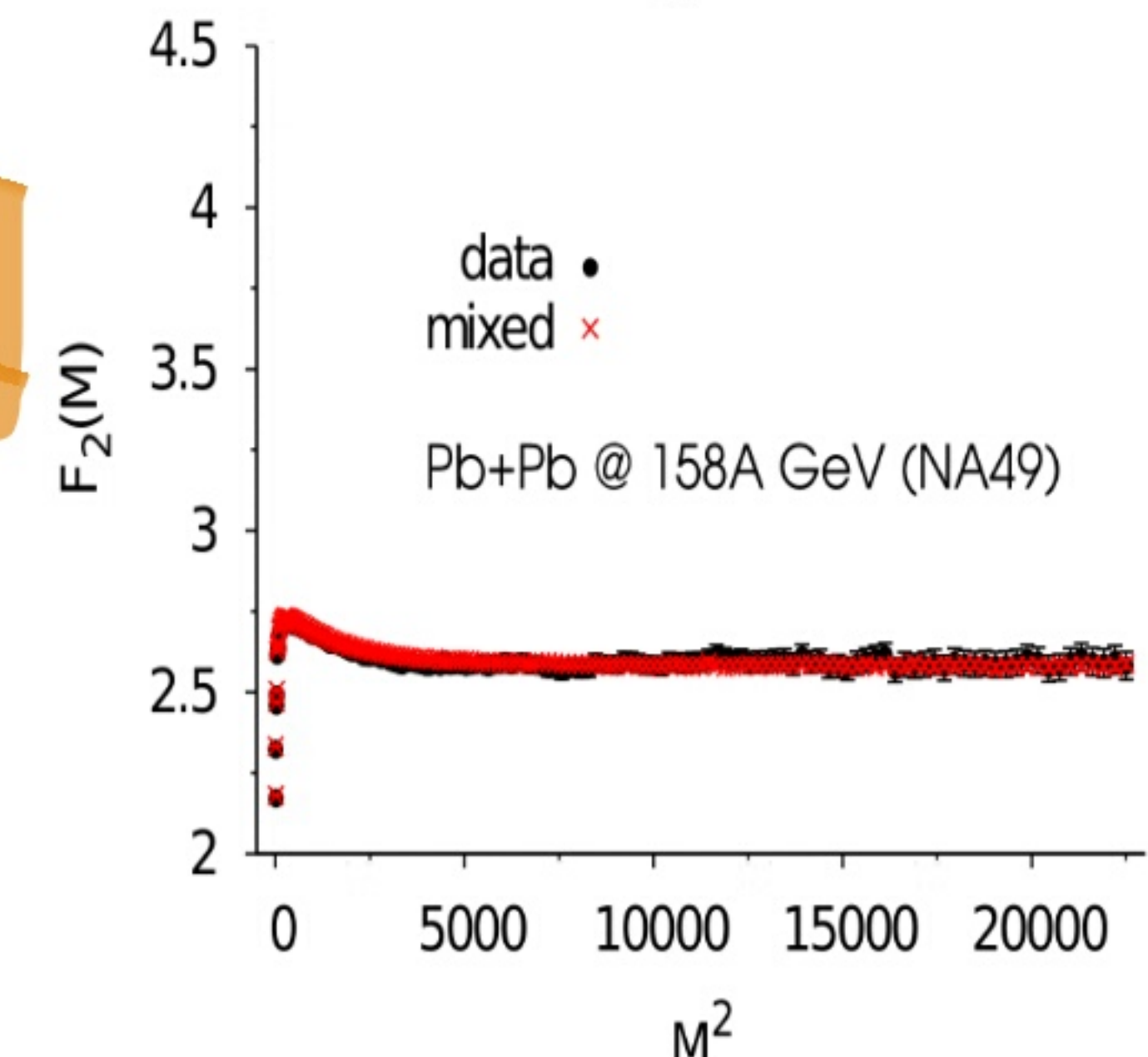
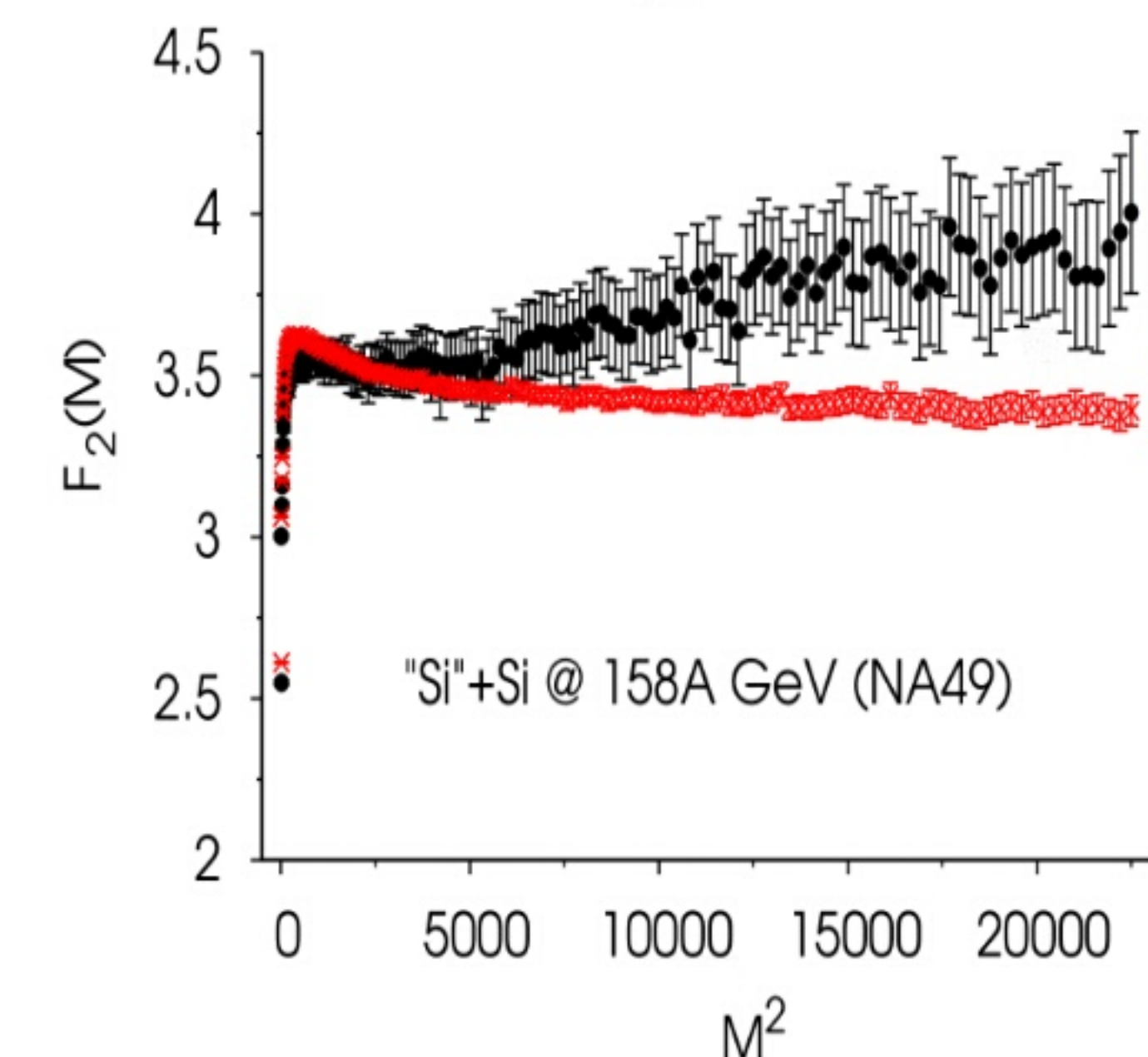
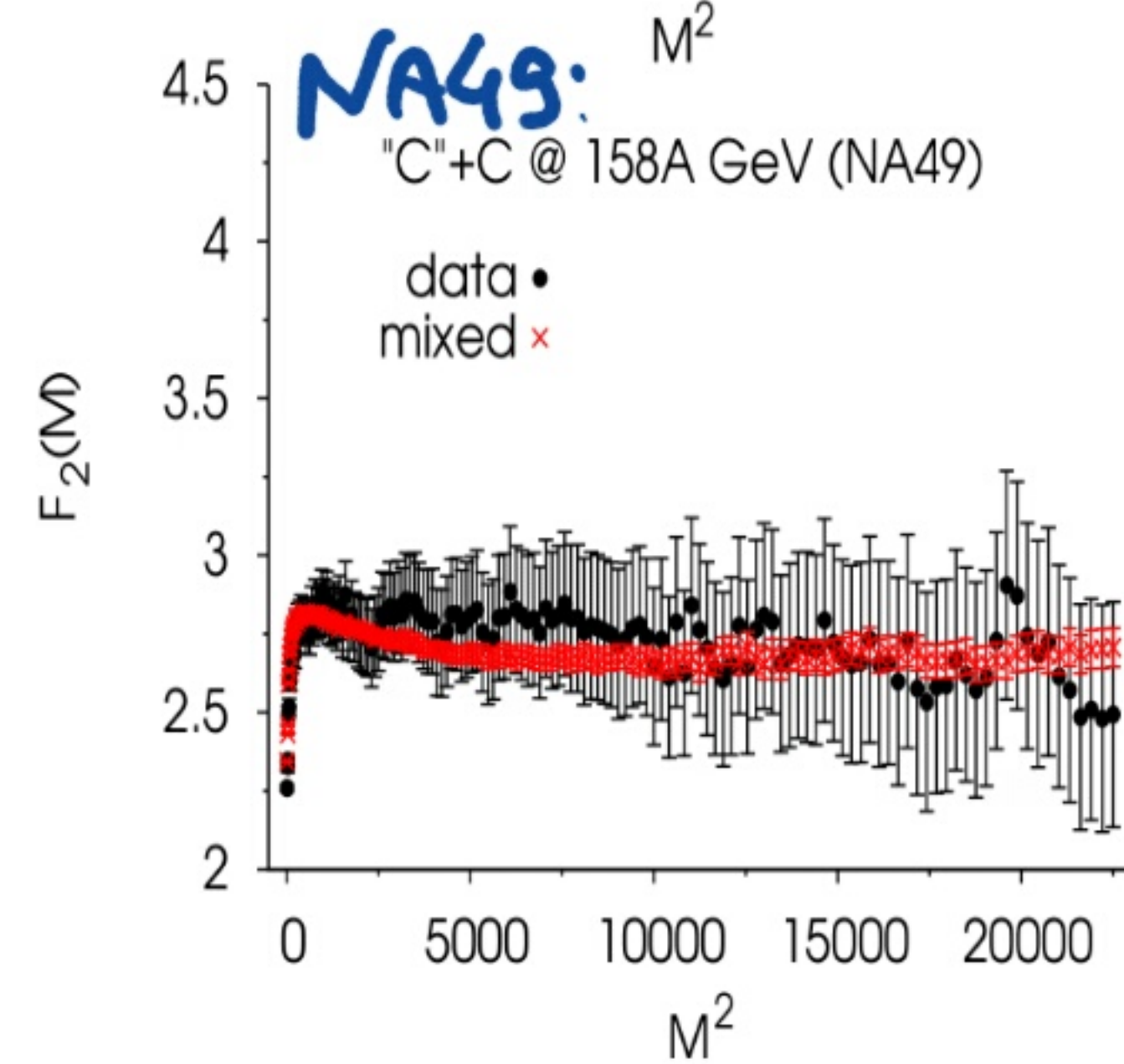
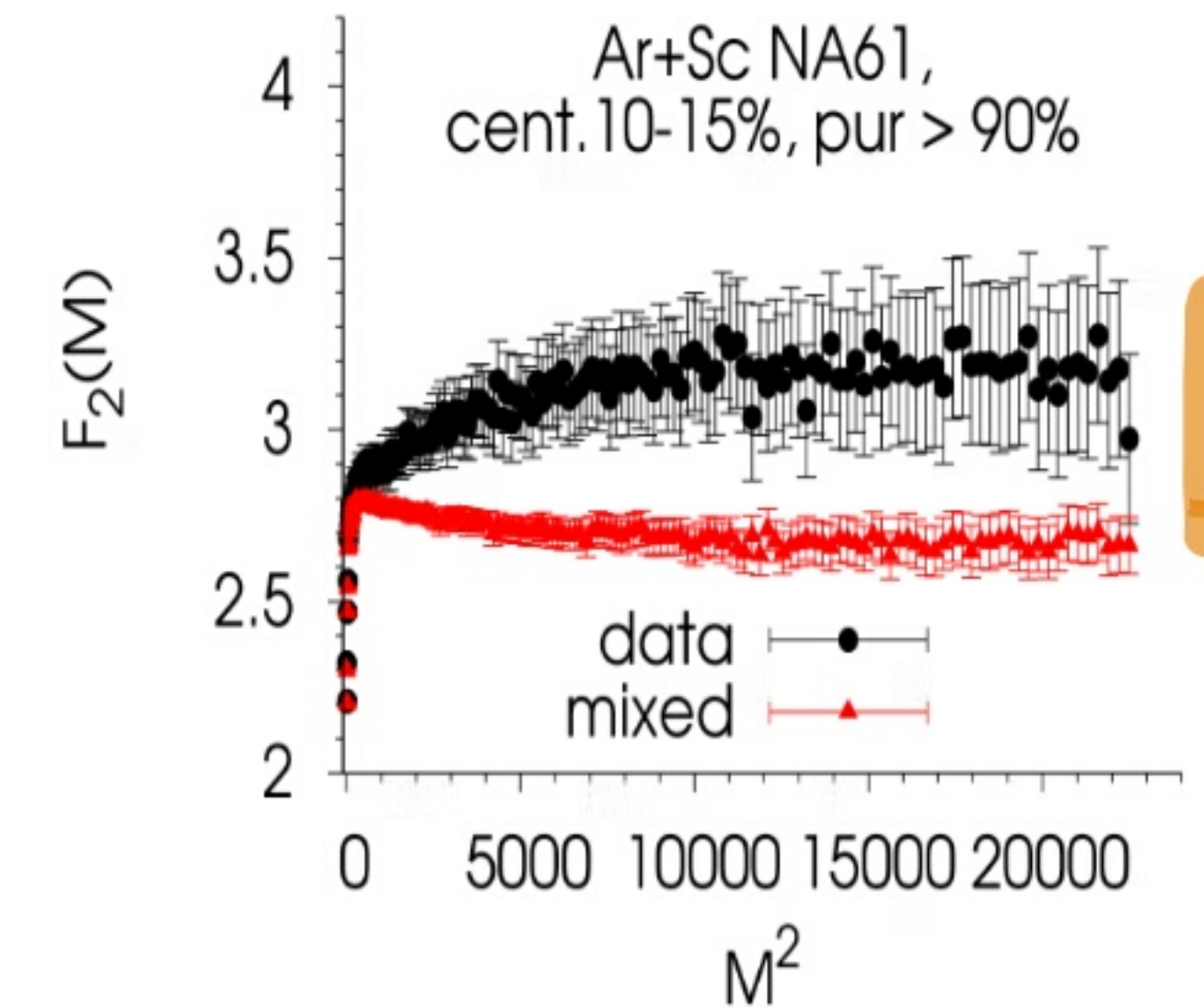
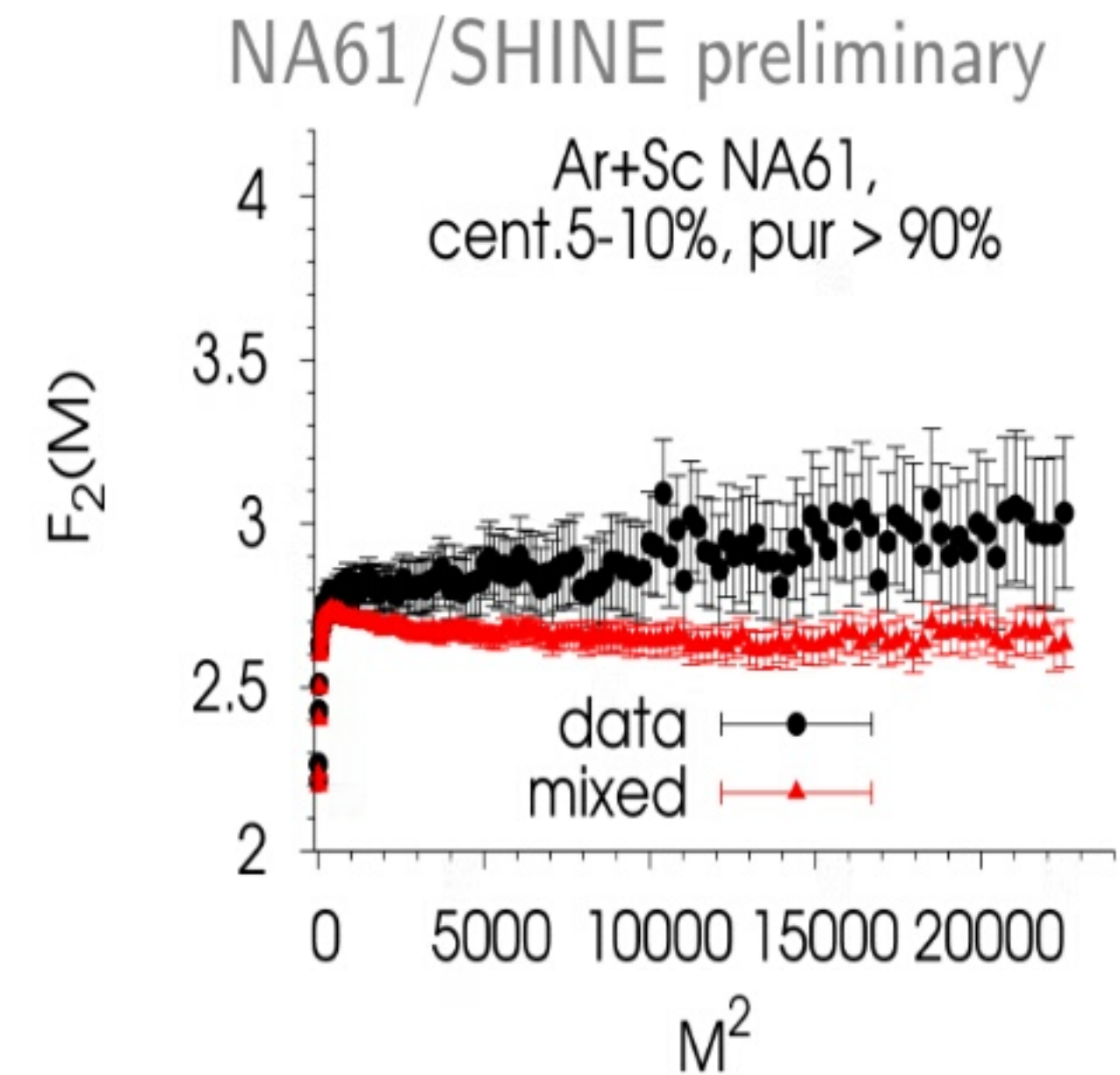
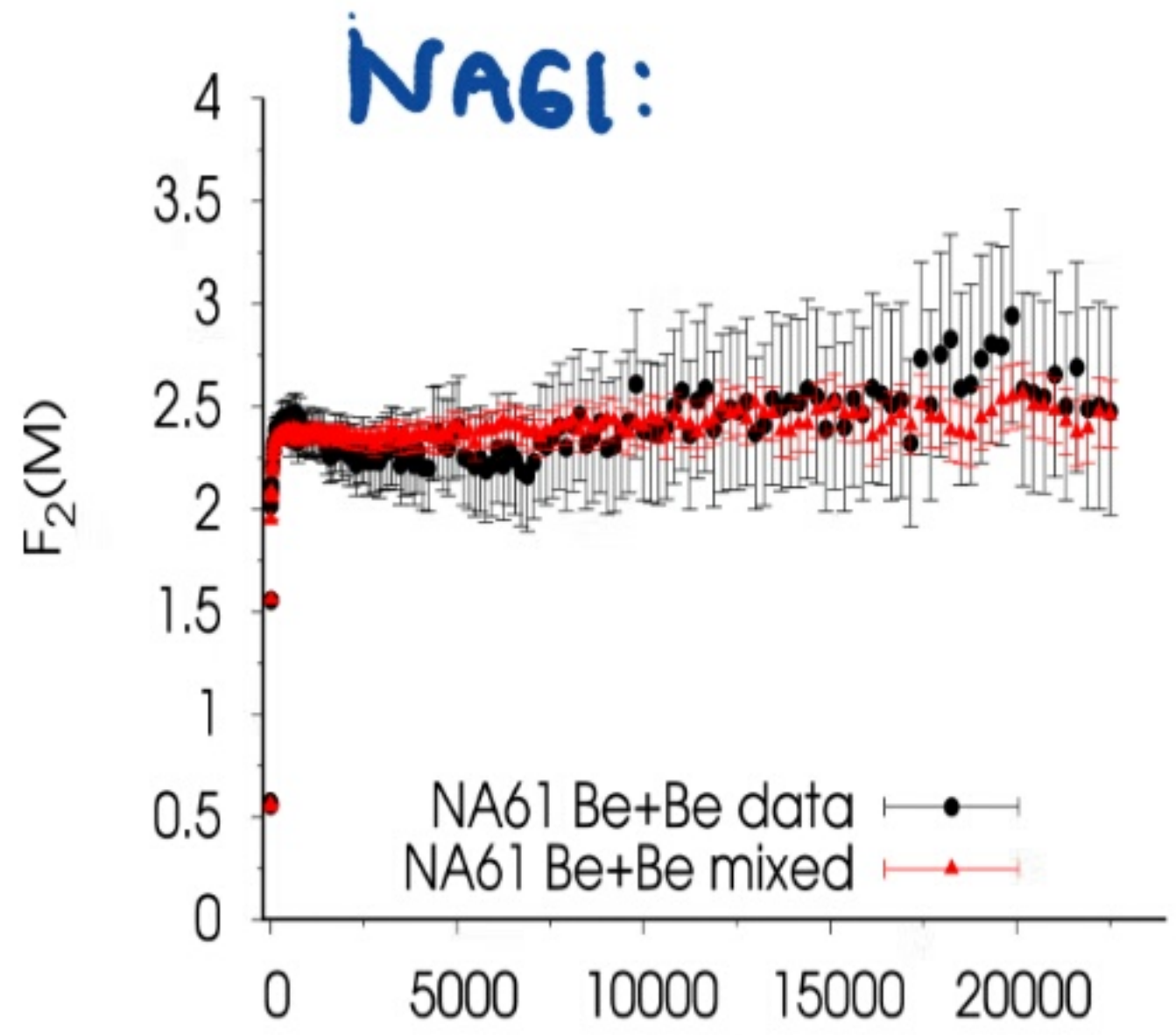
$$F_2(M) = F_2(\Delta) \cdot M^{\phi_2}$$

WOSIEK (1988)
BIALAS, PESZANSKI
SATZ
ANTONIDU, DIAKONDS, KAPOYANIS

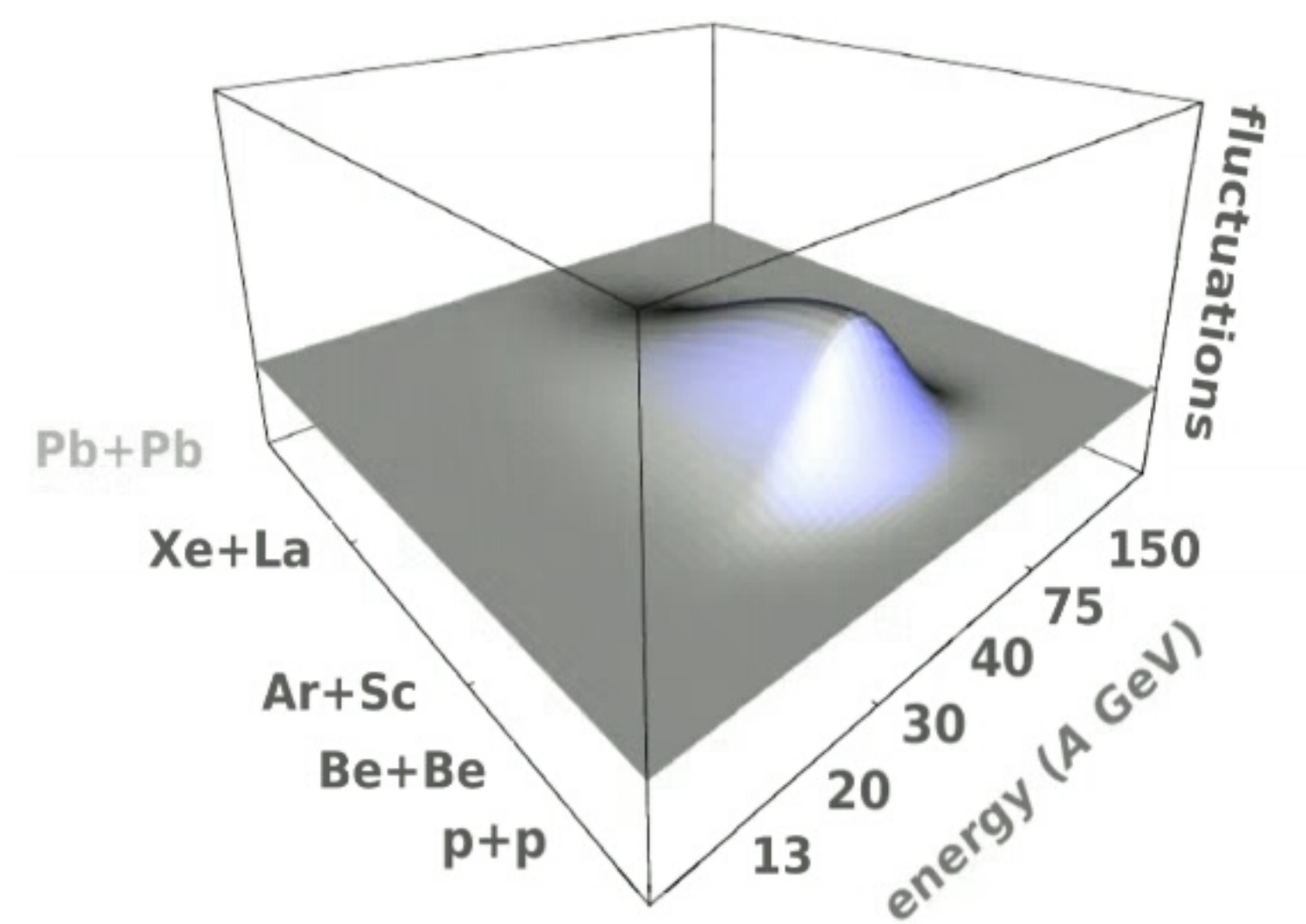
UNIQUENESS OF HEAVY ION PHYSICS AT THE CERN SPS:

INDICATION OF CP IN Ar+Sc AT 150A GeV/c:

(5)



HOPE FOR
"HILL OF FLUCTUATIONS"



↑ FLUCTUATIONS OF PROTON NUMBER: DATA, MIXED EVENTS ↑

REQUESTED NEW MEASUREMENTS (2021-2024)

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HEAVY ION PHYSICS:

- OPEN CHARM MEASUREMENTS IN Pb+Pb AT SPS

REFERENCE MEASUREMENTS:

- NUCLEAR FRAGMENTATION CROSS-SECTION FOR COSMIC RAY EXPERIMENTS
- HADRON PRODUCTION FOR NEUTRINO EXPERIMENTS

RECENT DOCUMENTS:

- March 21, 2018, Addendum 10:

*Study of Hadron-Nucleus and Nucleus-Nucleus Collisions at the CERN SPS:
Early Post-LS2 Measurements and Future Plans,*
CERN-SPSC-2018-008, SPSC-P-330-ADD-10

- June 5, 2018, Addendum 11:

Reply to the SPSC questions on Addendum CERN-SPSC-2018-008,
CERN-SPSC-2018-019, SPSC-P-330-ADD-11

OPEN CHARM MEASUREMENTS IN Pb+Pb AT SPS: MOTIVATION

⑦

- Q1: WHAT IS THE MECHANISM OF CHARM PRODUCTION ?
- Q2: HOW DOES THE ONSET OF DECONFINEMENT IMPACT CHARM PRODUCTION ?
- Q3: HOW DOES THE FORMATION OF QGP IMPACT J/ψ PRODUCTION ?

TO ANSWER ONE NEEDS TO KNOW:

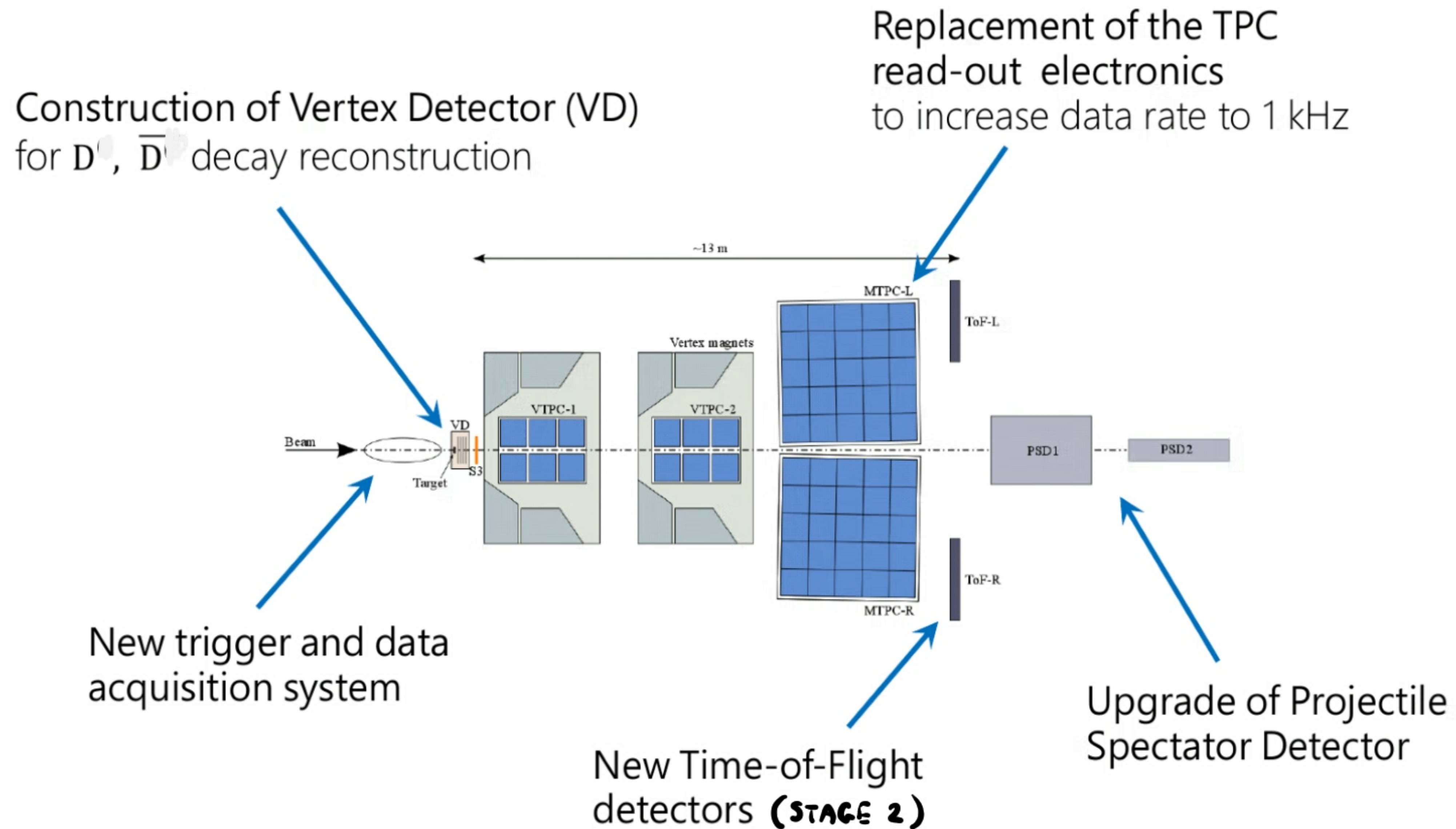
MEAN NUMBER OF CHARM QUARK PAIRS PRODUCED
IN THE FULL PHASE SPACE, $\langle c\bar{c} \rangle$, IN Pb+Pb COLLISIONS

UP TO NOW NO CORRESPONDING EXPERIMENTAL DATA

ONLY NA61/SHINE CAN PERFORM NEEDED MEASUREMENTS
IN THE NEAR FUTURE

DETECTOR UPGRADES (2018-2021)

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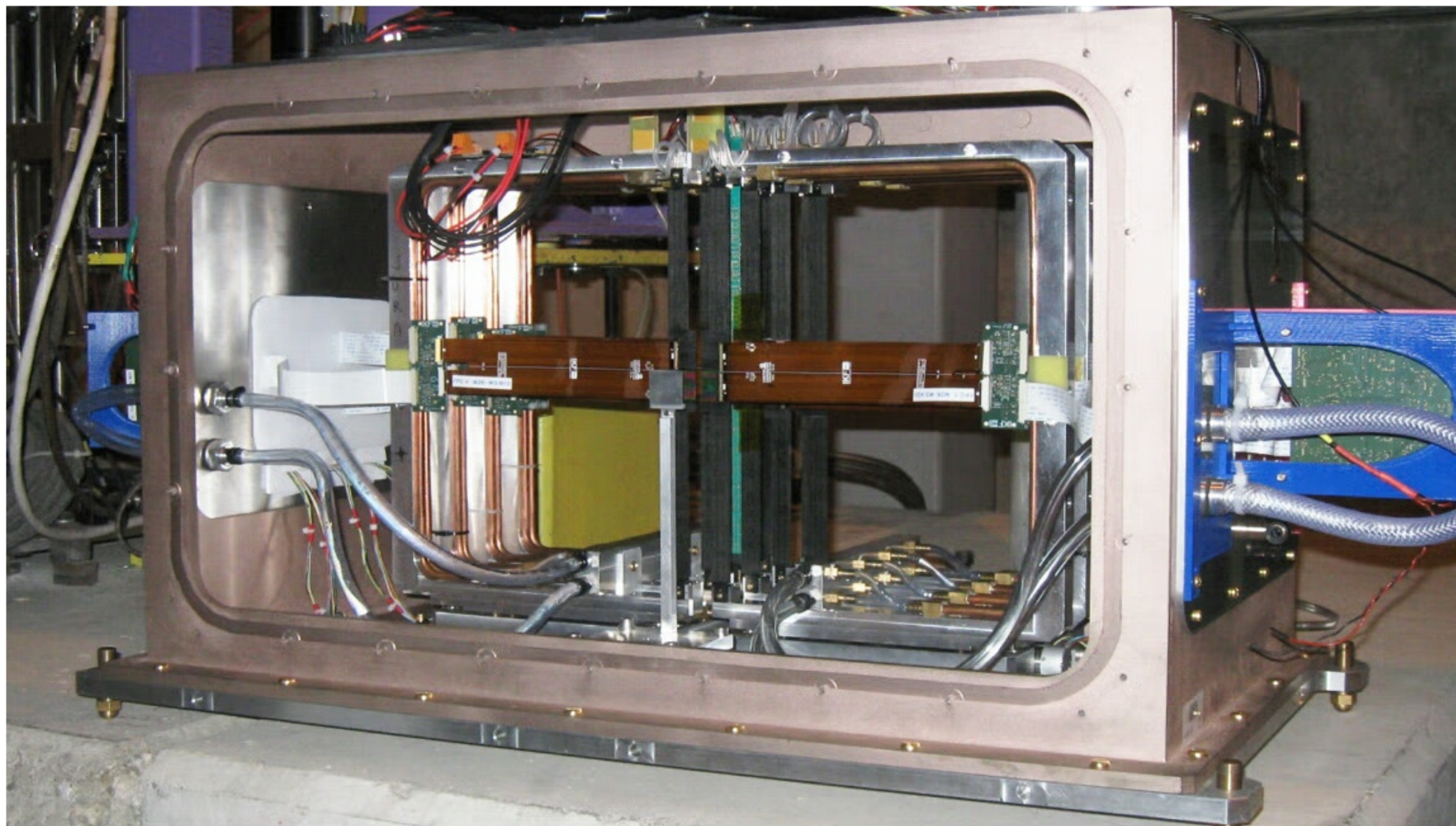
IN TOTAL: $\approx 1M$ CHF (STAGE 1, HARDWARE ONLY)

MODERATE COSTS THANKS TO COLLABORATION WITH ALICE (TPC, VD), CBM (PSD), (NICA (TOF))

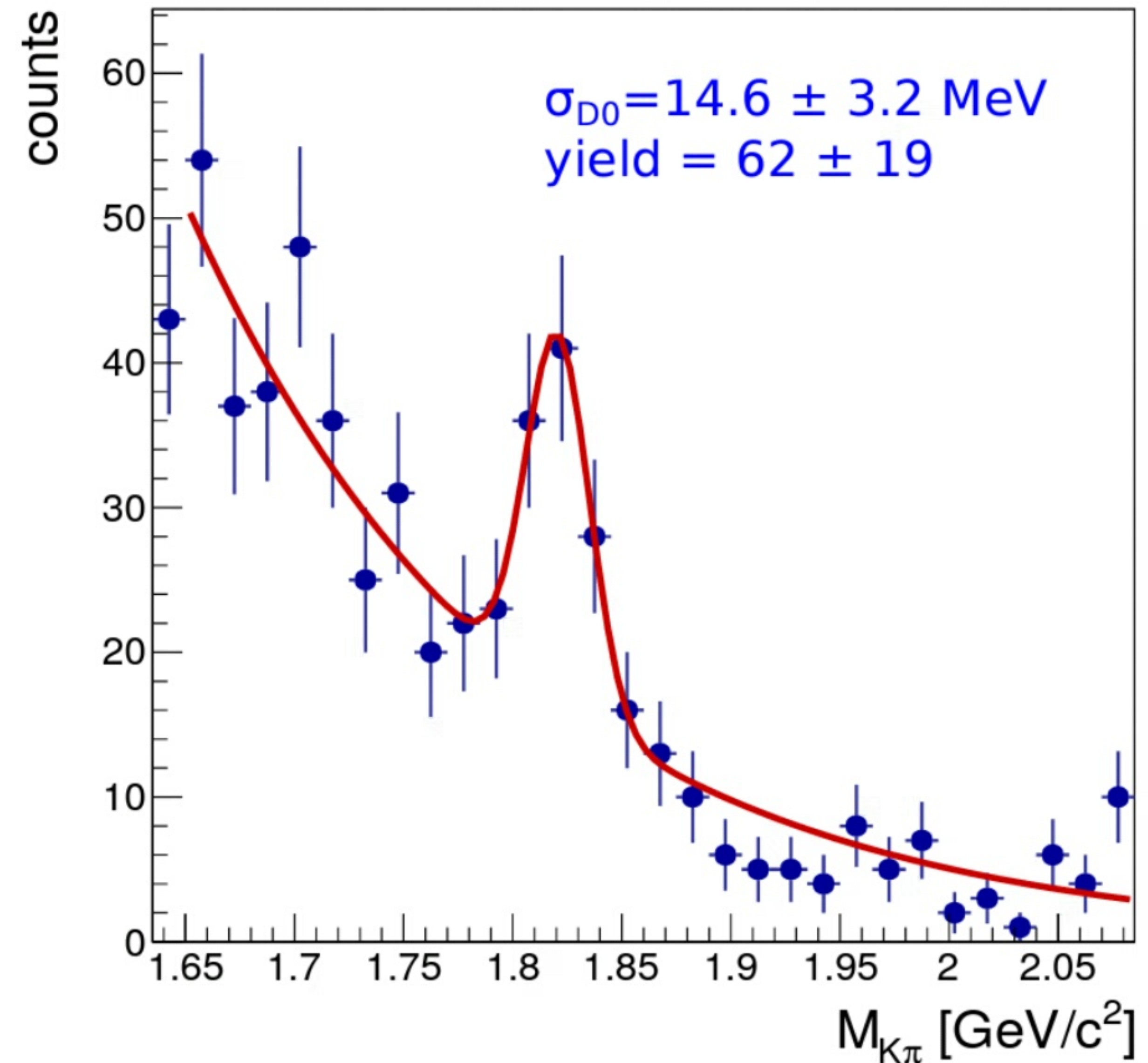
RESULTS FROM THE 2016 TEST (Pb+Pb AT 150A GEV/c)

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2016-2018:
SMALL ACCEPTANCE VERTEX DETECTOR



BASED ON TECHNOLOGIES DEVELOPED
FOR ALICE AND CBM



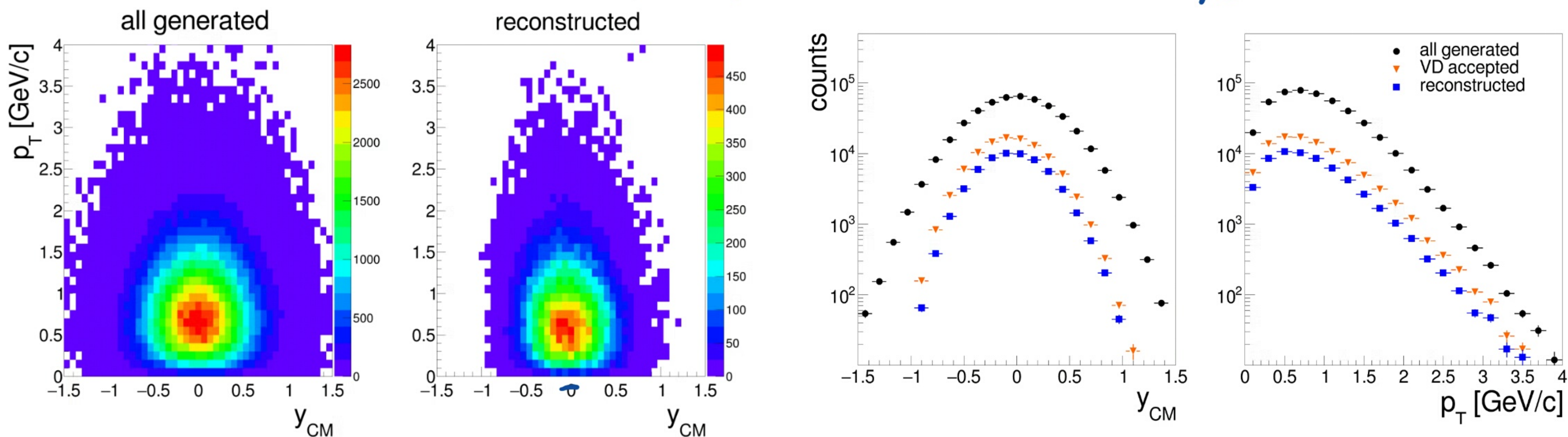
INDICATION OF $D^0 + \bar{D}^0$ PEAK

DATA STATISTICS (2021-24) AND ACCEPTANCE



Reaction	days	events	$\#(D^0 + \bar{D}^0)$	$\#(D^+ + D^-)$
Pb+Pb at 150A GeV/c	84	500M	76k	46k
Pb+Pb at 40A GeV/c	42	250M	3.6k	2.1k

$D^0 + \bar{D}^0$ ACCEPTANCE : Pb+Pb AT 150A GeV/c

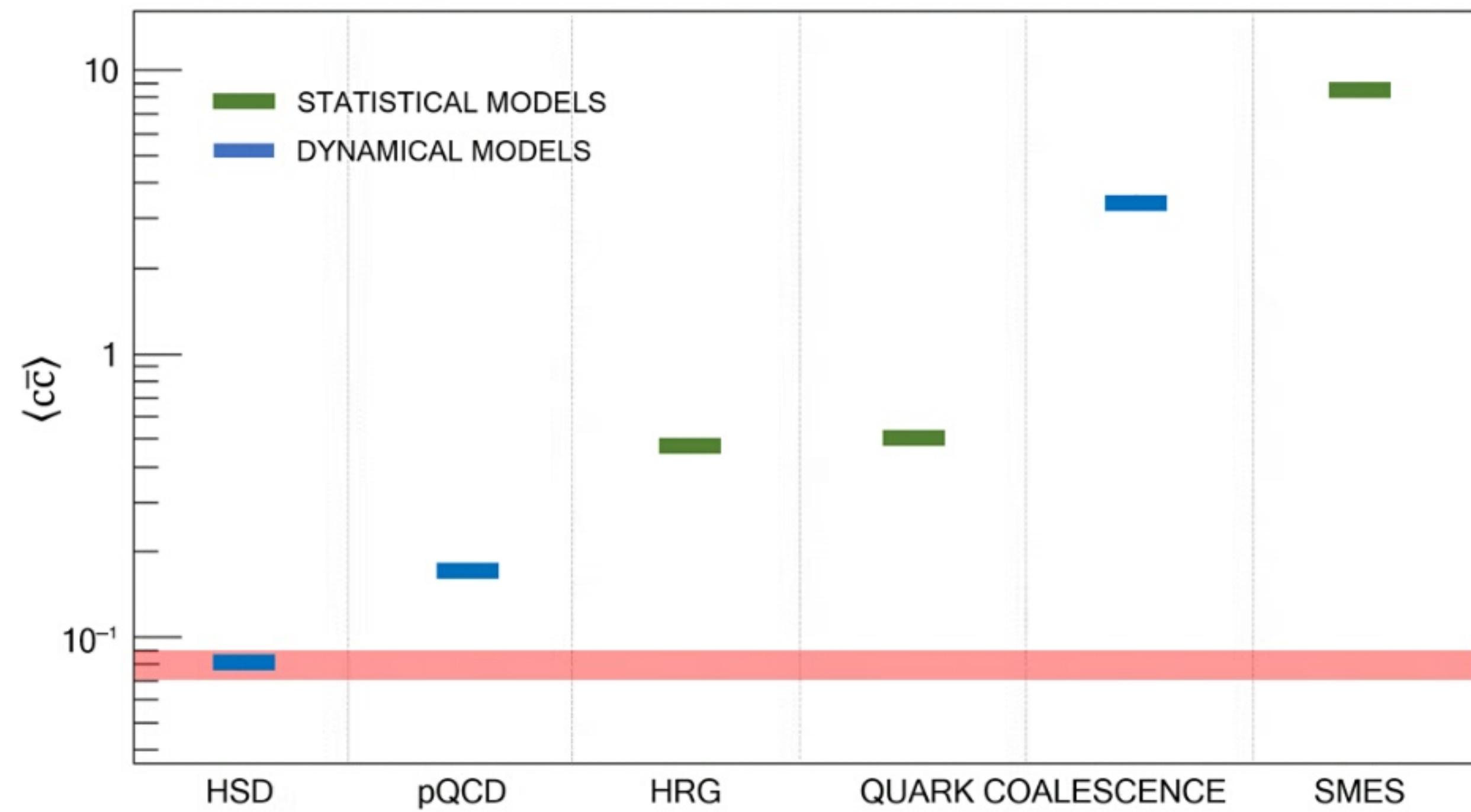


CENTRALITY	0-10%	10-20%	20-30%	30-60%	60-90%	0-90%
$\#(D^0 + \bar{D}^0)$	31k	20k	11k	13k	1.3k	76k
$\#(D^+ + D^-)$	19k	12k	7k	8k	0.8k	46k

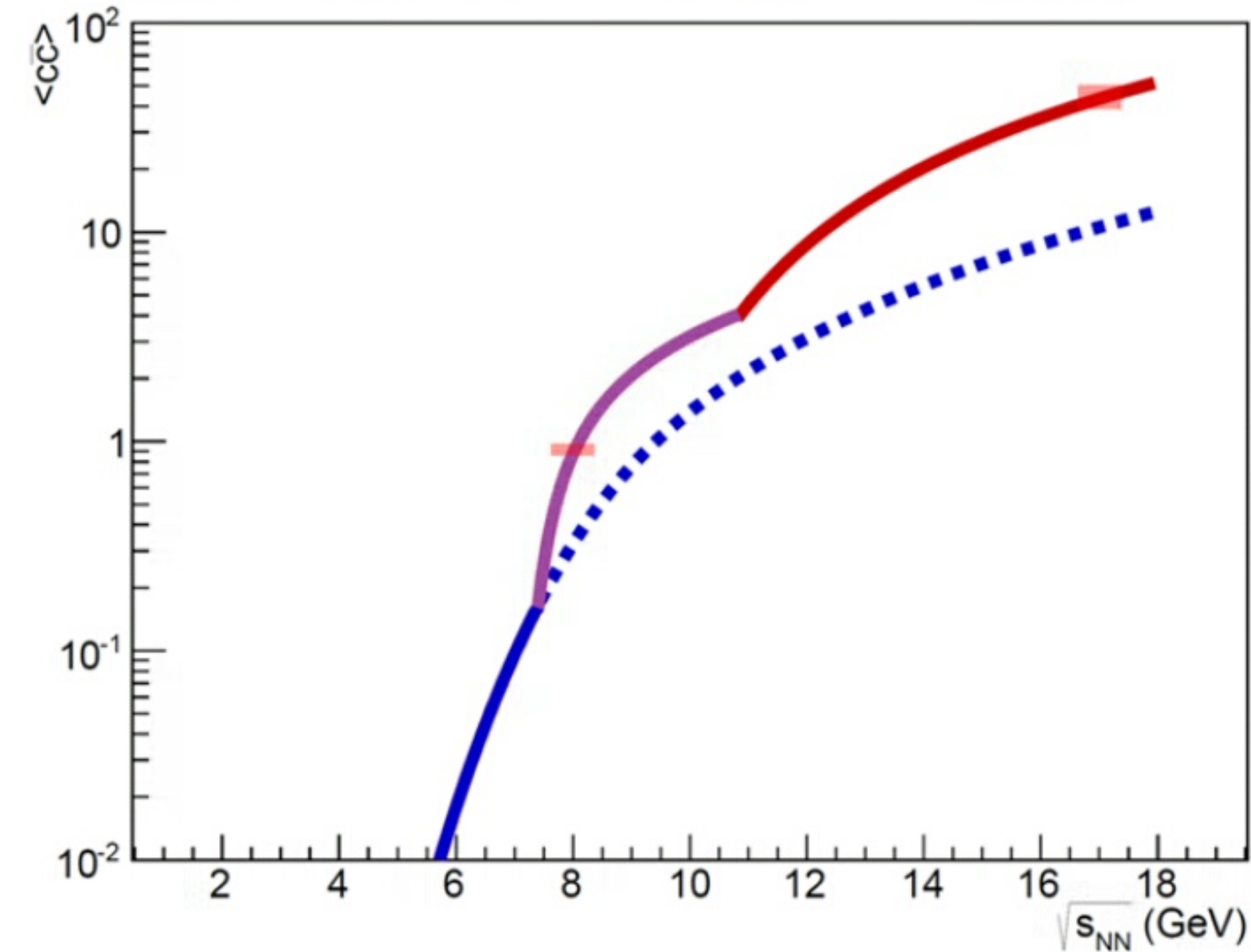
EXPECTED PHYSICS IMPACT OF NA61/SHINE $\langle c\bar{c} \rangle$ MEASUREMENTS

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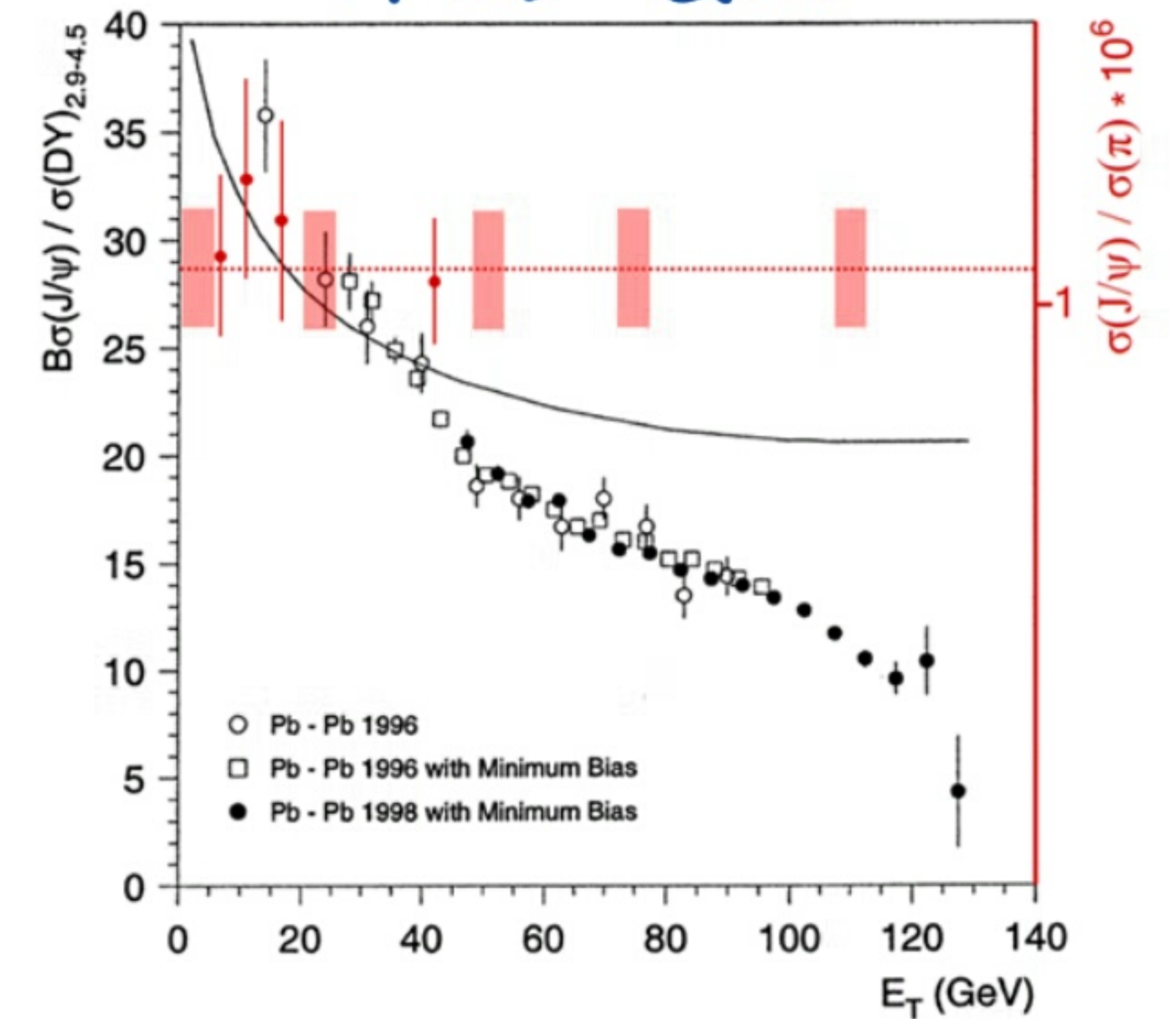
Q1: $\langle c\bar{c} \rangle$ AND MODELS



Q2: $\langle c\bar{c} \rangle$ AND ONSET OF DECONFINEMENT



Q3: $\langle c\bar{c} \rangle$, $\langle J/\psi \rangle$ AND QGP

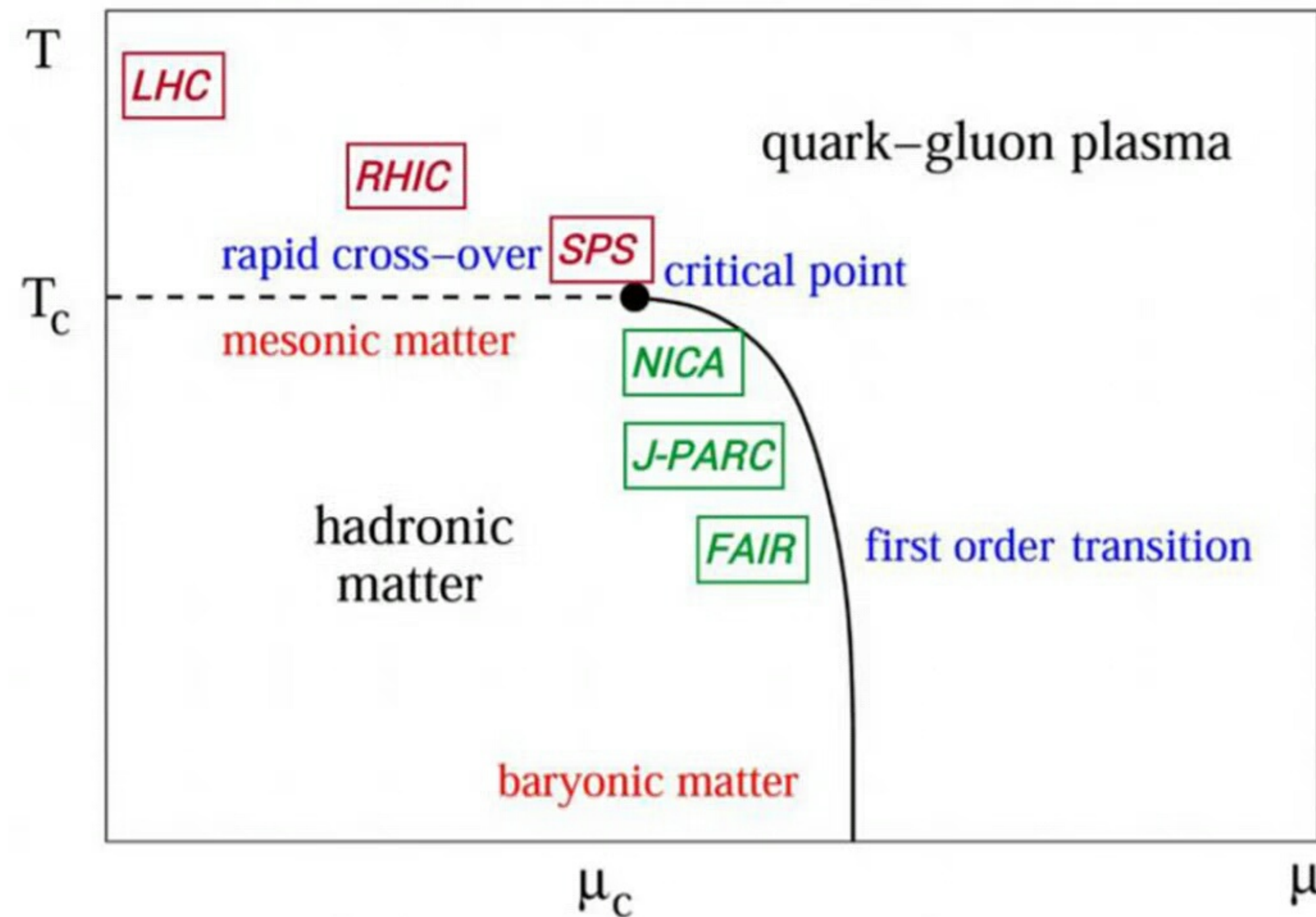


FORESEEN NA61/SHINE RESOLUTION IS SUFFICIENT TO ANSWER Q1, Q2 AND Q3

UNIQUENESS OF NAGI/SHINE PROGRAMME

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LANDSCAPE OF PRESENT AND FUTURE HEAVY ION EXPERIMENTS



LHC and RHIC at high energies ($\sqrt{s_{NN}} \geq 200$ GeV):

measurements in limited phase space due to collider geometry and kinematics

RHIC BES (3 – 39 GeV):

measurement not under consideration

NICA (< 11 GeV):

under consideration during stage 2

J-PARC (< 6 GeV) :

maybe possible after 2025

FAIR SIS-100 (< 5 GeV):

not possible at SIS-100,
planned at SIS-300 (< 7 GeV)

ONLY NAGI/SHINE IS ABLE TO MEASURE OPEN CHARM PRODUCTION IN HEAVY ION COLLISIONS IN FULL PHASE SPACE AND IN THE NEAR FUTURE

NA61/SHINE 2020+ COLLABORATION : 33 INSTITUTIONS ≈ 150 AUTHORS



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REQUESTED BEAMS

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2021: 6 WEEKS FOR DETECTOR COMMISSIONING

5 WEEKS OF PROTON BEAM AT 31 GEV/C FOR
DATA TAKING FOR NEUTRINO PHYSICS

4 WEEKS OF Pb BEAM AT 150A GEV/C FOR
OPEN CHARM MEASUREMENT

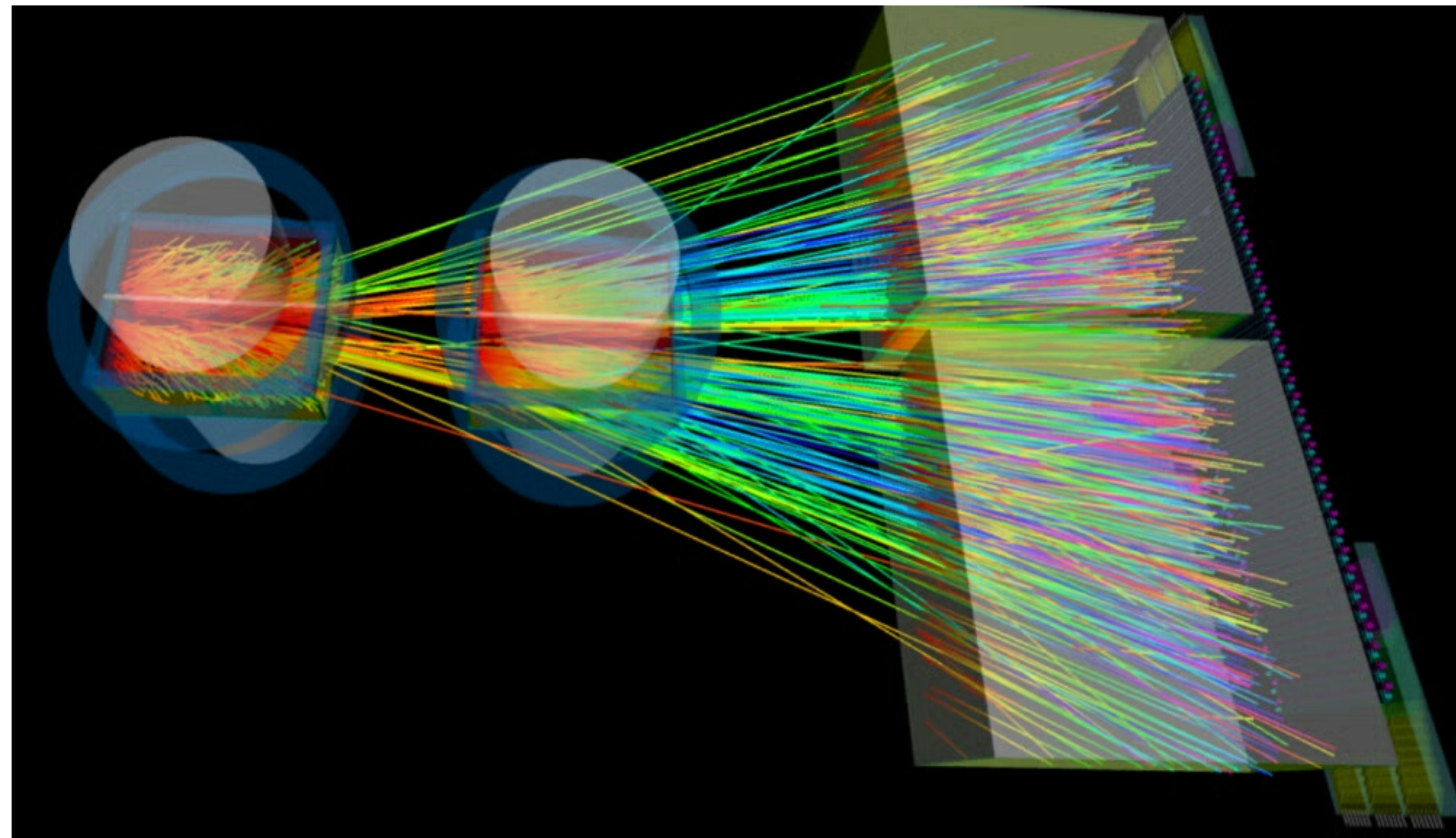
RECOMMENDED
BY SPSC

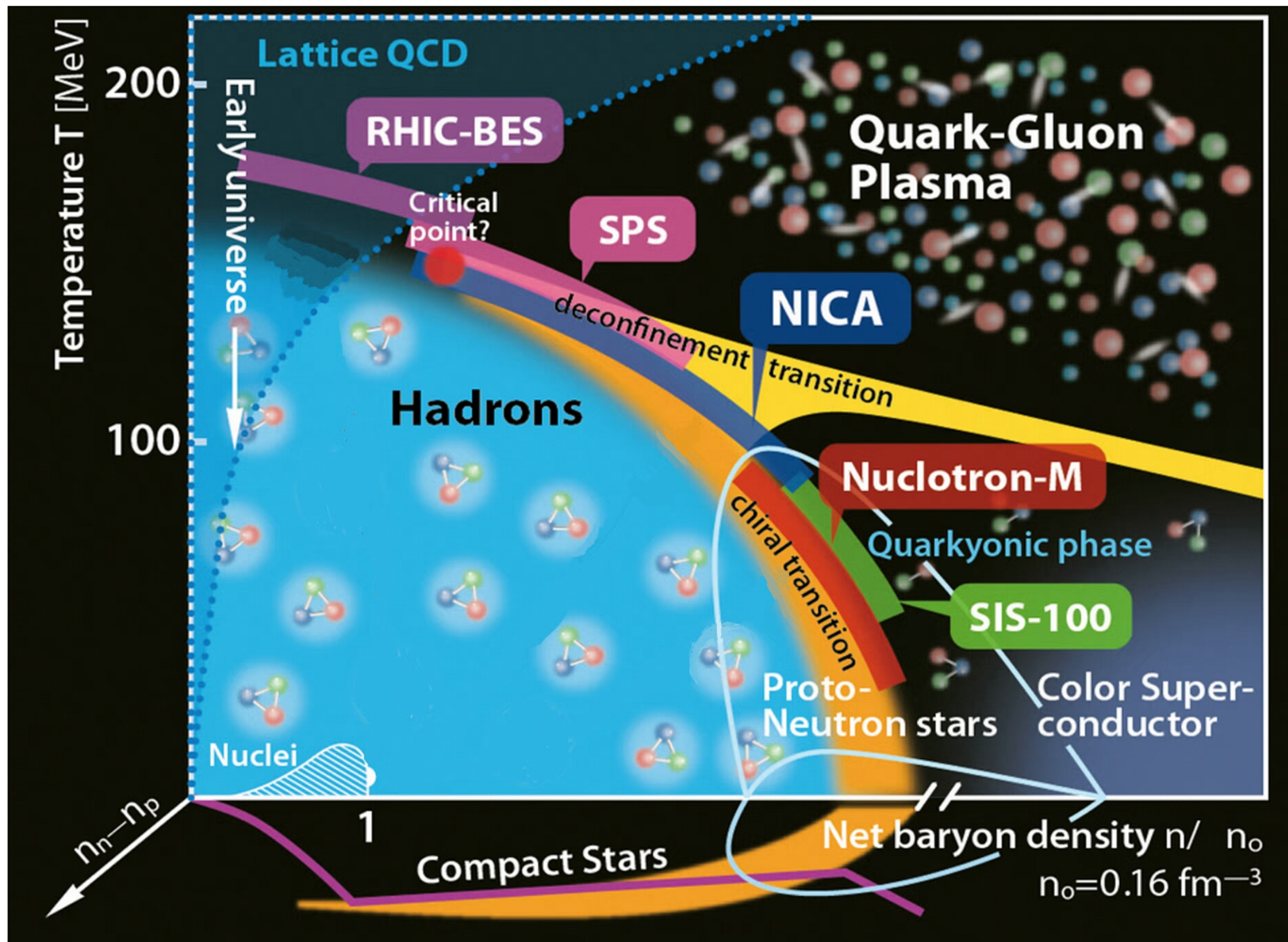
2022-24: HADRON, LIGHT ION (SECONDARY) AND Pb BEAMS
IN ACCORDANCE TO THE PROPOSED PHYSICS PROGRAM

SUMMARY

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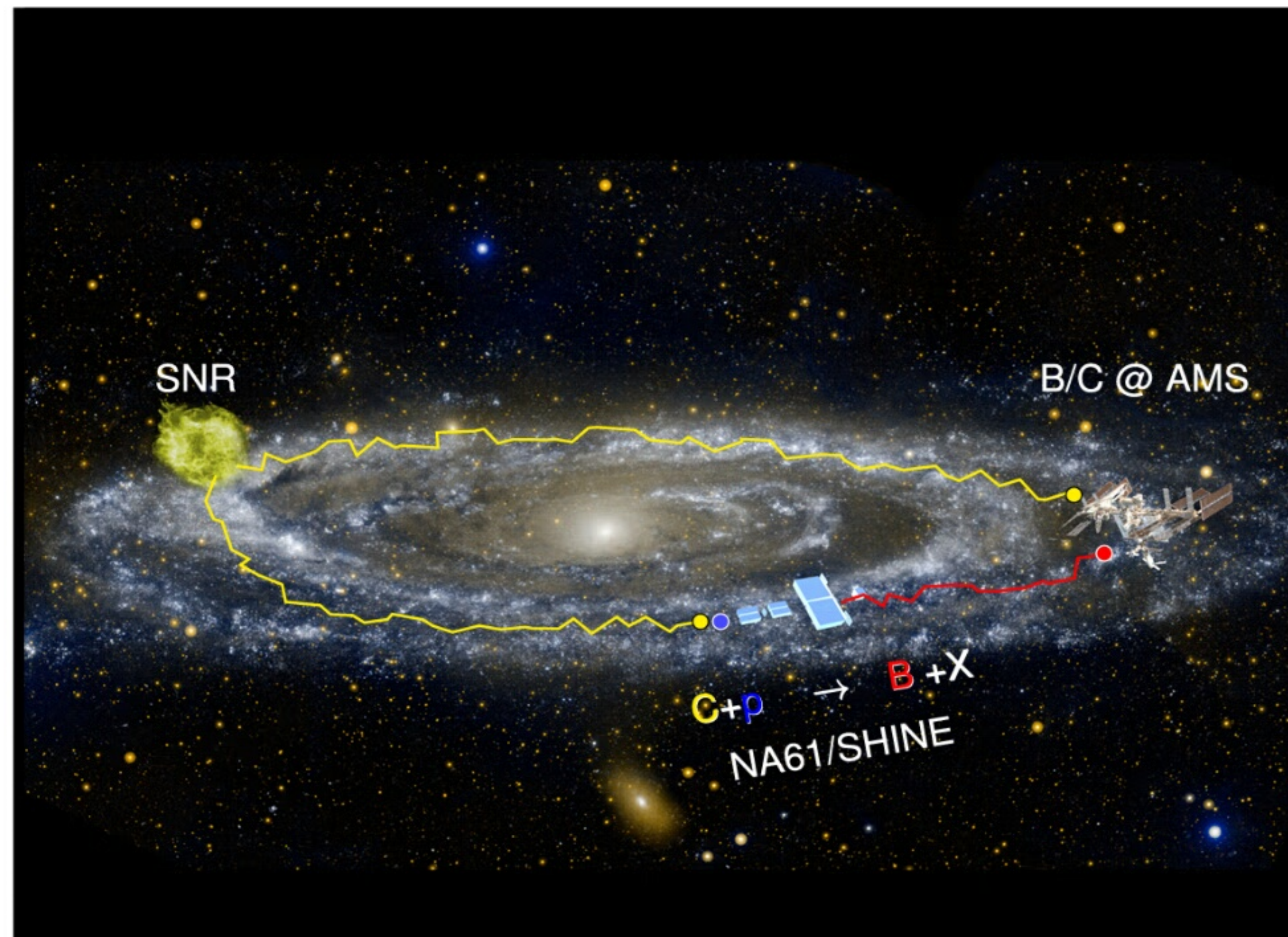
- NAGI/SHINE PLANS UNIQUE OPEN CHARM MEASUREMENTS IN Pb+Pb COLLISIONS AT SPS AS WELL AS NEW REFERENCE MEASUREMENTS FOR COSMIC RAY AND NEUTRINO EXPERIMENTS
 - DATA TAKING IN 2021 IS RECOMMENDED BY SPSC
 - DATA TAKING IN 2022-2024 | ← SPSC ← ESPP ← | PBC TOWN MEETINGS NAGI/SHINE
 - WORK ON DETECTOR UPGRADES HAS STARTED
-





REFERENCE MEASUREMENTS:

NUCLEAR FRAGMENTATION CROSS SECTION FOR COSMIC RAY EXPERIMENTS

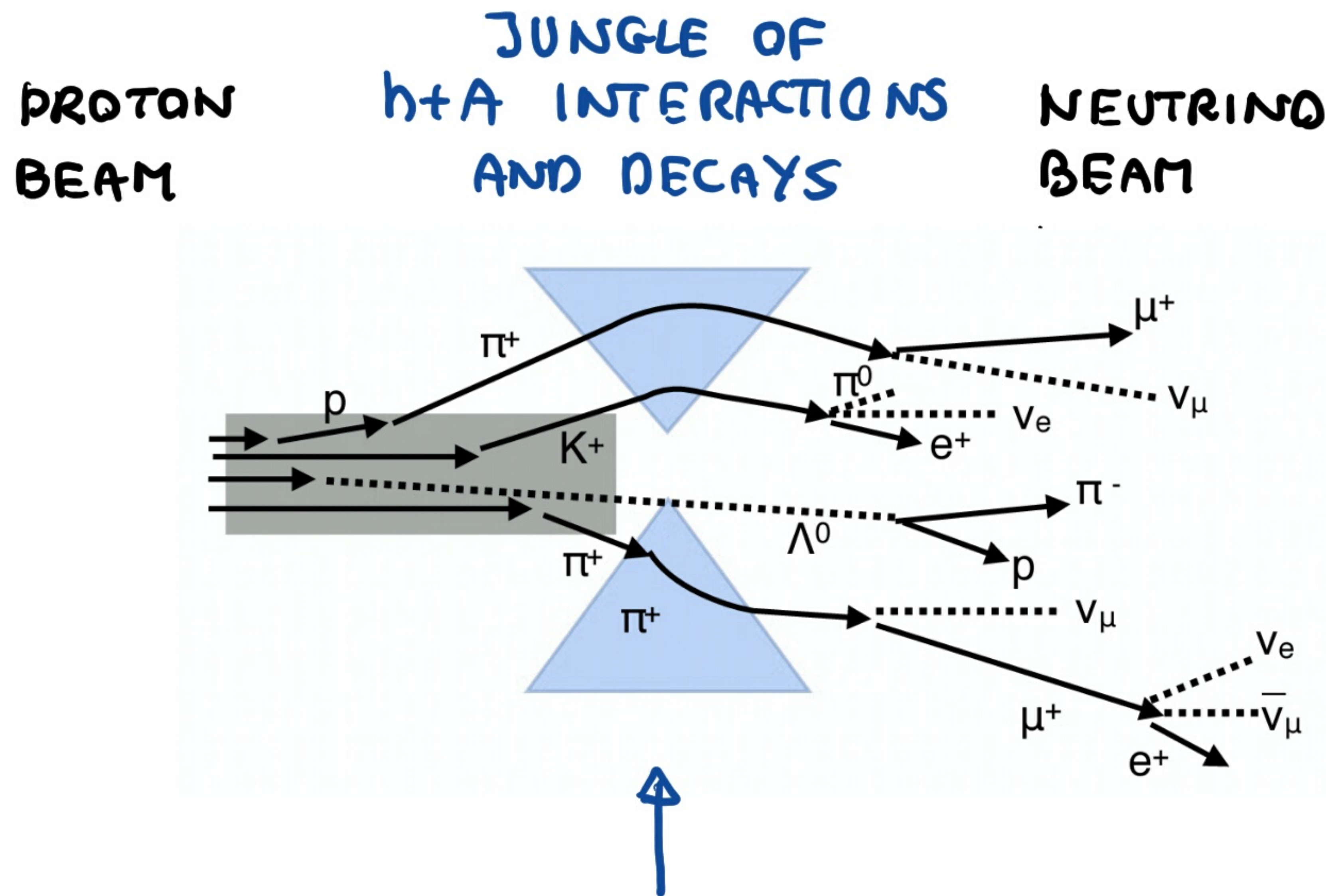


- Primary cosmic rays from supernova remnants
- Secondary cosmic rays from interactions with interstellar matter during propagation e.g.
 $^{12}\text{C} + \text{p} \xrightarrow{\text{frag.}} \text{B} + \text{X}$
 $^{12}\text{C} + \text{p} \xrightarrow{\text{frag.}} ^{11}\text{C} + \text{p} \xrightarrow{\text{decay}} \text{B} + \text{Y}$
- Primary-to-secondary ratios (e.g. B/C)
→ traversed mass density
- Unstable-to-stable ratios (e.g. $^{10}\text{Be}/^9\text{Be}$)
→ traversed distance
- Important for the understanding of origin of Galactic cosmic rays and backgrounds for DM searches

UNDERSTANDING OF COSMIC RAY PROPAGATION LIMITED BY UNCERTAINTIES OF FRAGMENTATION CROSS SECTIONS

NA61/SHINE WILL SIGNIFICANTLY REDUCE THE UNCERTAINTIES.
(FROM $\approx 20\%$ TO $\approx 0.5\%$)

REFERENCE MEASUREMENTS: HADRON PRODUCTION FOR NEUTRINO EXPERIMENTS



- Further improvement of the precision of measurements for the currently used T2K replica target,
- Measurements for a new target material (super-sialon) for T2K-II and Hyper-Kamiokande,
- Study of the possibility of measurements with beams $< 12 \text{ GeV}/c$ for improved predictions of atmospheric and accelerator ν fluxes,
- Ultimate hadron production measurements with prototypes of Hyper-Kamiokande and DUNE targets.

NAGI/SHINE WILL DECREASE SYSTEMATIC UNCERTAINTIES OF NEUTRINO FLUX (FOR T2K-II, HYPER-K FROM $\approx 10\%$ TO $\approx 3\%$)