Searching for Dark Matter at Cosmic Dawn



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Based on PRD 92 (2015) Nature 557 (2018) PRL 121 (2018) with Yacine Ali-Haimoud Cora Dvorkin Avi Loeb Ely Kovetz

CMB



z = 1100

Image: ESA

Cosmic Dawn



z = 1100 z = 20

Image: ESA

Cosmic Dawn

Reionization



z = 1100 z = 20 z = 6

Image: ESA

CMB

CosmicEarth andDawnReionizationTelescopes



z = 1100z = 20z = 6 \mathcal{Z}^{\cdot}

Image: ESA

CMB

What have we learned about DM?

See Cora's and Neelima's talks



z = 1100z = 20z=6 Z^{\perp}

Image: ESA

What can we learn?

Cosmic Dawn



z = 1100z = 20z = 6z = 0

Image: ESA

Outline

Introduction to 21-cm cosmology

- EDGES and dark matter

- Other DM searches with 21 cm

z = 1100



 $I_{\nu} \propto T_{\rm CMB} \nu^2$



 $I_{\nu} \propto T_{\rm CMB} \nu^2$

(@ 21 cm)





 $T_S < T_{
m cmb}$ Absorption

$T_S > T_{ m cmb}$ Emission



$T_S < T_{\rm cmb}$ Absorption z = 1100 $z \approx 20$ Australia Н Η Н **EDGES** $\Delta I_{\nu} \propto T_{21} \nu^2$ $I_{\nu} \propto T_{\rm CMB} \nu^2$





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Bowman et al. Nature 2018

Thermal Decoupling (from CMB)



Thermal Decoupling (from CMB)











Thermal Decoupling (from CMB)





Can DM explain EDGES?

Requirements

$$n_{\chi} \ge n_b \quad \rightarrow \quad m_{\chi} \le 6 \,\mathrm{GeV}$$



JBM, Kovetz, Ali-Haimoud 1509.00029

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$$\sigma_{\chi b} \propto v^{-4}$$

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$$\begin{array}{ll} {\rm JBM+\ 2015}\\ {\rm New\ Interaction} & {\rm Barkana\ 2018}\\ & \left(m_{\phi} < 10\,{\rm keV}\right) \end{array}$$

Image: NASA/CXC/CfA/STScI ESO/WFI

Limits on millicharged particles





JBM and Loeb 1802.10094





JBM and Loeb 1802.10094



21-cm Global Signal = CMB Monopole

21-cm fluctuations



21-cm Fluctuations = CMB Anisotropies

21-cm fluctuations




JBM, Kovetz, Ali-Haimoud PRD 2015



JBM, Kovetz, Ali-Haimoud PRD 2015



21-cm fluctuations



JBM, Dvorkin and Loeb PRL 2018







-Annihilating WIMPs

Liu & Slatyer, 2018, D'Amico+ 2018 Lopez-Honorez+ 2016 ...

-Exotic radio excess

Feng and Holder 2018

$$|T_{21}| \sim \frac{T_{\rm cmb} + T_{\rm extra}}{T_S}$$

-Annihilating WIMPs

Liu & Slatyer, 2018, D'Amico+ 2018 Lopez-Honorez+ 2016 ...

-Exotic radio excess

DM to dark photons -> photons







 $f_{\rm dm} \lesssim {\rm few\%}$ $\epsilon/m_{\chi} \sim 10^{-5} {\rm MeV}^{-1}$



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Thanks!



Bowman et al. Nature 2018 EDGES (Experiment to Detect the Global EoR Signature)



Bowman et al. Nature 2018 EDGES (Experiment to Detect the Global EoR Signature)



Bowman et al. Nature 2018 EDGES (Experiment to Detect the Global EoR Signature)



JBM, Kovetz and Ali-Haimoud 2015





JBM and Loeb 2017



Murata and Tanaka 2014

Galactic Foregrounds (CMB)



Chuzhoy and Kolb (2009)



 $\rho_{\rm DM} = 0.3 \pm 0.1 \, {\rm GeV \, cm^{-3}}$

Bovy and Tremaine (2012)

Chuzhoy and Kolb (2009)



 $ho_{\rm DM} = 0.3 \pm 0.1 \, {\rm GeV \, cm^{-3}}$ Bovy and Tremaine (2012) However: $\rho_B \sim 10^{-3} \rho_{\rm dm} v_{\rm MW}^2$

Fifth-force cooling?



Fifth-force constraints



Knapen, Lin, Zurek 2017

 m_{ϕ}









DM-baryon relative velocities



$$\left\langle V_{\rm rel}^2 \right\rangle^{1/2} \approx 30 \,\rm km \, s^{-1}$$

Tseliakhovich and Hirata PRD 2010



Tseliakhovich and Hirata PRD 2010



JBM, Dvorkin and Loeb PRL 2018




-Annihilating WIMPs

Liu & Slatyer, 2018, Also D'Amico+ 2018 Lopez-Honorez+ 2016 ...

-Exotic DM decays

Light DM to radio photons

Fraser et al. 2018

-Annihilating WIMPs

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