









Aspen Center for Physics 29 March 2019

Expectations for sensitivity to BSM physics at the HL-LHC

Laura Jeanty

on behalf of the authors of the HL-LHC Working Group Reports



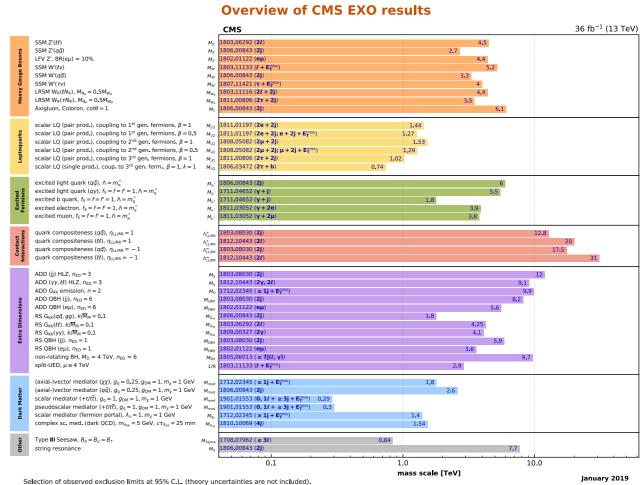
Overview

- Excellent performance of LHC and experiments in Run 1 and Run 2 lead to a plethora of SM measurements and searches for new physics
 - No signifiant deviation from SM expectation in measurements
 - No clear indication for new physics phenomena so far
- Upgrade of LHC and experiments to fully exploit potential and push discovery sensitivity significantly higher

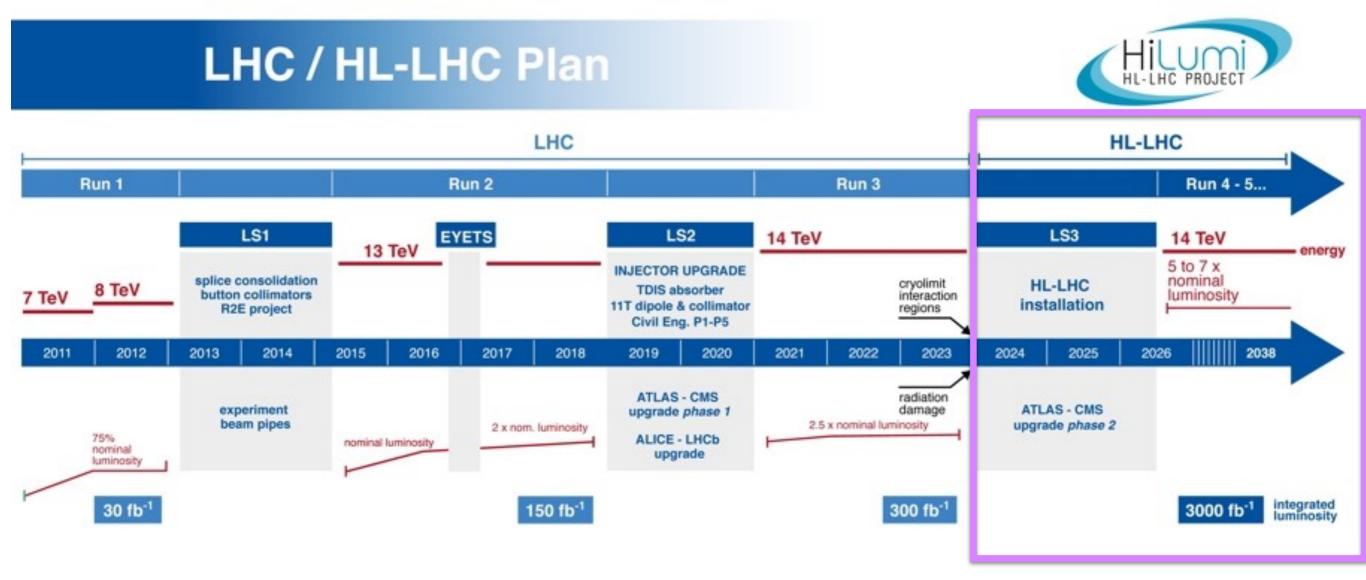
Current supersymmetry limits from ATLAS

ATLAS Preliminary ATLAS SUSY Searches* - 95% CL Lower Limits $\sqrt{s} = 7, 8, 13 \text{ TeV}$ Mode e, μ, τ, γ Jets $E_{\tau}^{\text{miss}} \int \mathcal{L} dt [\text{fb}^{-1}]$ Reference $\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{g}\tilde{\chi}_{\perp}^{0}$ 2-6 jets 1712.02332 1712.02332 7-11 jets 4 jets Yes 1708.02794 1706.03731 $\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\bar{t}\tilde{\chi}_{\perp}^{0}$ 3 b 4 jets 1708.09266, 1711.03301 1708.09266 1706.03731 $\tilde{b}_1\tilde{b}_1, \tilde{t}_1\tilde{t}_1, M_2 = 2 \times M$ 709.04183, 1711.11520, 1708.0324 709.04183, 1711.11520, 1708.0324 506.08616, 1709.04183, 1711.11520 1709.04183, 1711.11520 1709.04183, 1711.11520 1709.04183, 1711.11520 0.48-0.84 $\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow c\tilde{\chi}_1^0 / \tilde{c}\tilde{c}, \tilde{c} \rightarrow c\tilde{\chi}_1^0$ $\tilde{Y}_{1}^{\pm}\tilde{Y}_{2}^{0}$ via WZ. 1403.5294, 1806.02293 1712.06119 1706.07875 1708.07875 Yes Yes 36.1 36.1 1803.02762 1712.06119 $\hat{H}\hat{H}, \hat{H}\rightarrow h\hat{G}/Z\hat{G}$ 0.29-0.88 1806.04030 1804.03602 Stable & R-hadron 1606.05129 GMSB, $\hat{\mathcal{X}}_{1}^{0} \rightarrow \gamma \hat{C}$, long-lived $\hat{\mathcal{X}}_{1}^{0}$ 1409.5542 $6 < ct(\hat{\xi}_1^0) < 1000 \text{ mm}, m(\hat{\xi}_1^0) = 1 \text{ Tel}$ 1504.05162 $\tilde{X}_{1}^{\pm}\tilde{X}_{1}^{\mp}/\tilde{X}_{2}^{0} \rightarrow WW/Z\ell\ell\ell\ell\nu\nu$ 1804.03568 ATLAS-CONF-2018-003 m(8°) "200 GeV, bino-lik ATLAS-CONF-2018-003 ATLAS-CONF-2018-003 2 jets + 2 b 1710.07171 Mass scale [TeV]

Current exotics limits from CMS



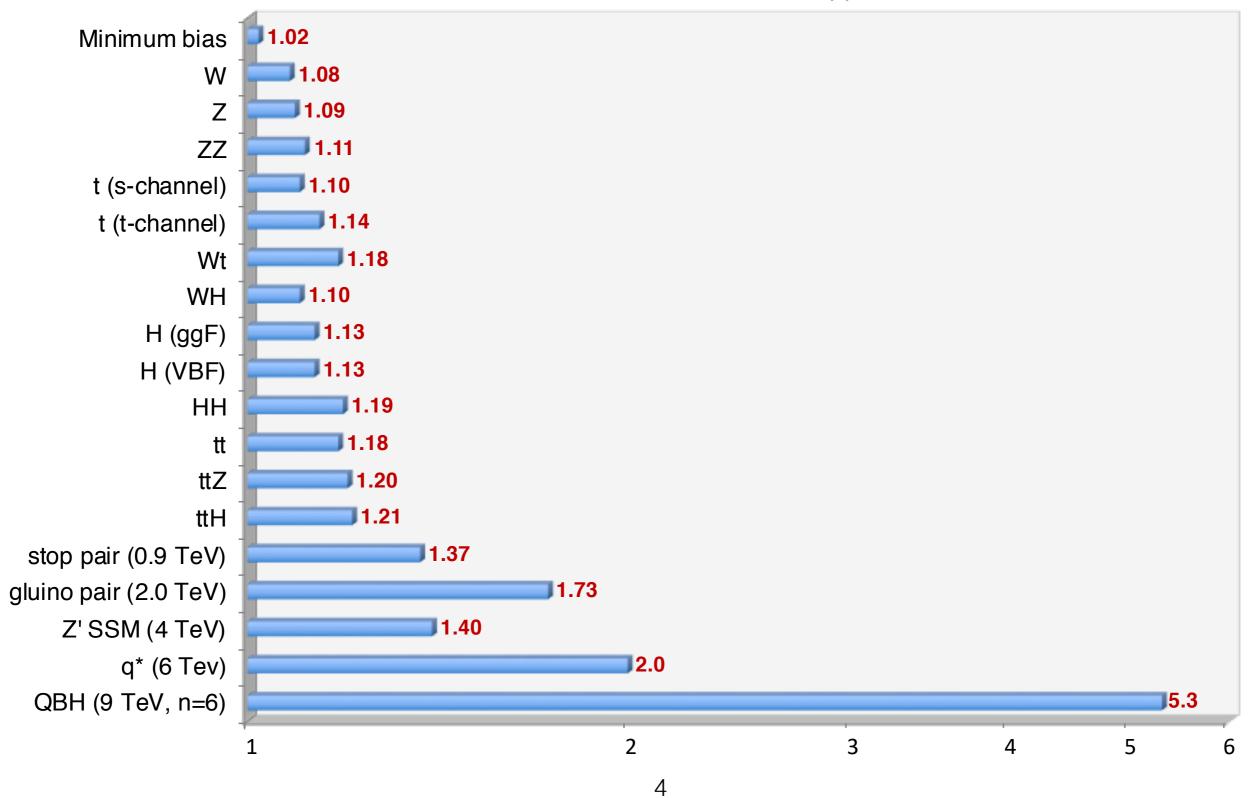
Setting the stage



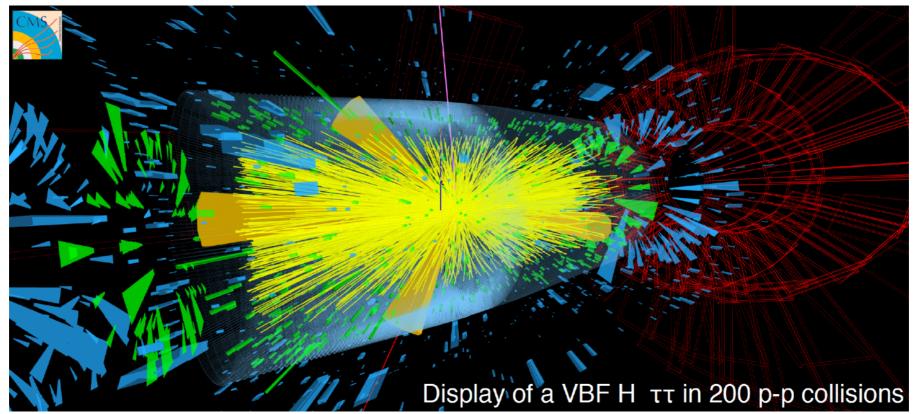
- High Luminosity LHC (HL-LHC) outlook
 - 3000 fb⁻¹ at center of mass energy of 14 TeV
 - Goals include deeper understanding of Higgs boson, precision measurements in QCD, EWK, Higgs, B physics, and probing physics beyond the standard model in both direct searches and precision measurements

Setting the stage

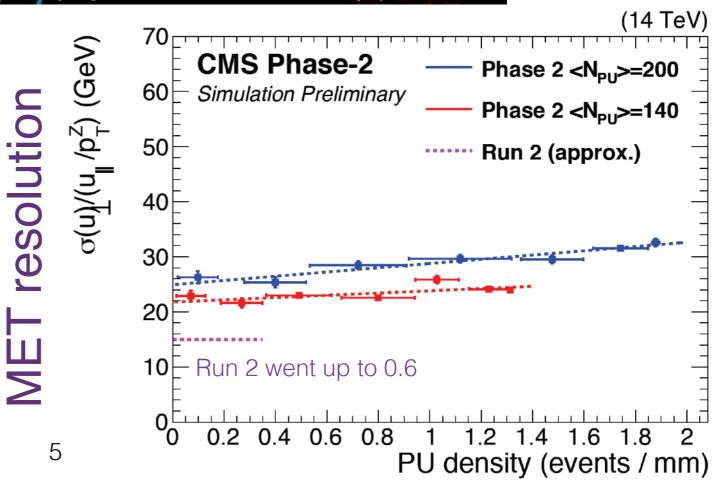
14 TeV / 13 TeV inclusive pp cross-section ratio



Challenges and opportunities: high (instantaneous) luminosity



- Average of 200 collisions per crossing
- Detector, trigger, reconstruction, performance, and analysis challenges

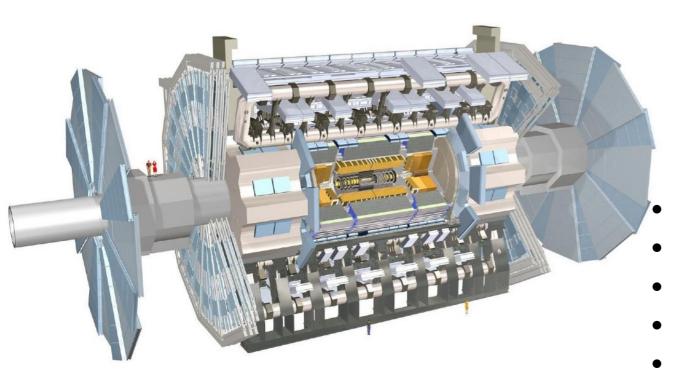


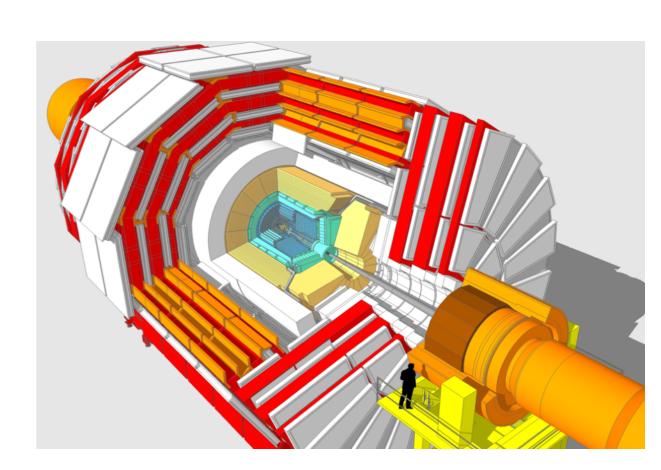
Upgraded detectors for HL-LHC

 Extensive detector upgrades to handle increased data rate, occupancy, pileup, and integrated luminosity

ATLAS upgrades include:

- Trigger and DAQ at L1 and HLT (10 kHz)
- Improved muon coverage and trigger
 - Upgraded calorimeters and trigger
- New inner tracker, coverage to $|\eta| < 4$
- High-granularity timing detector in endcap





CMS upgrades include:

Trigger and DAQ at L1 and HLT (7.5 kHz)

Improved muon coverage and trigger

Precise MIP timing layer in barrel and endcap

New inner tracker, coverage to $|\eta| < 4$

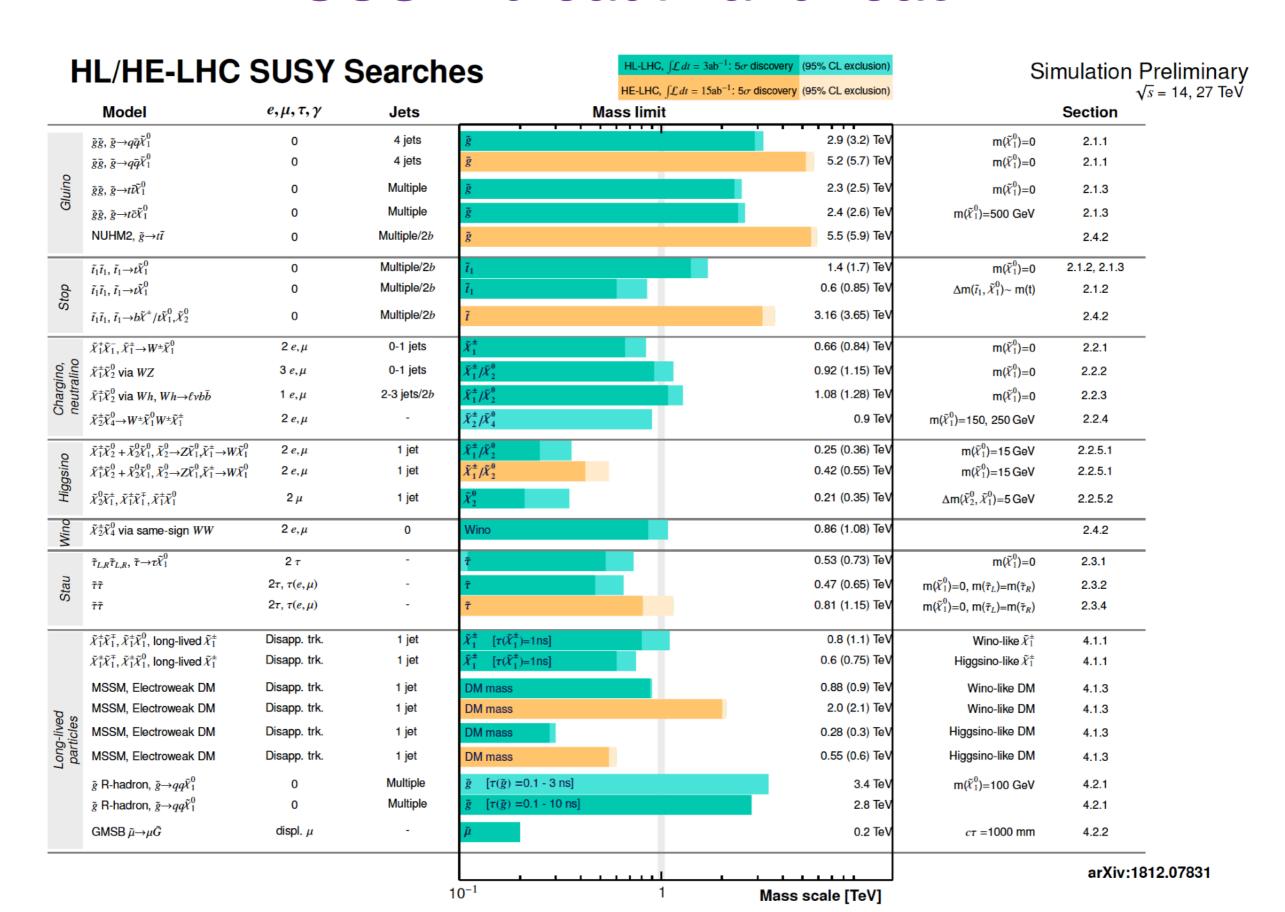
High-granularity endcap calorimeter

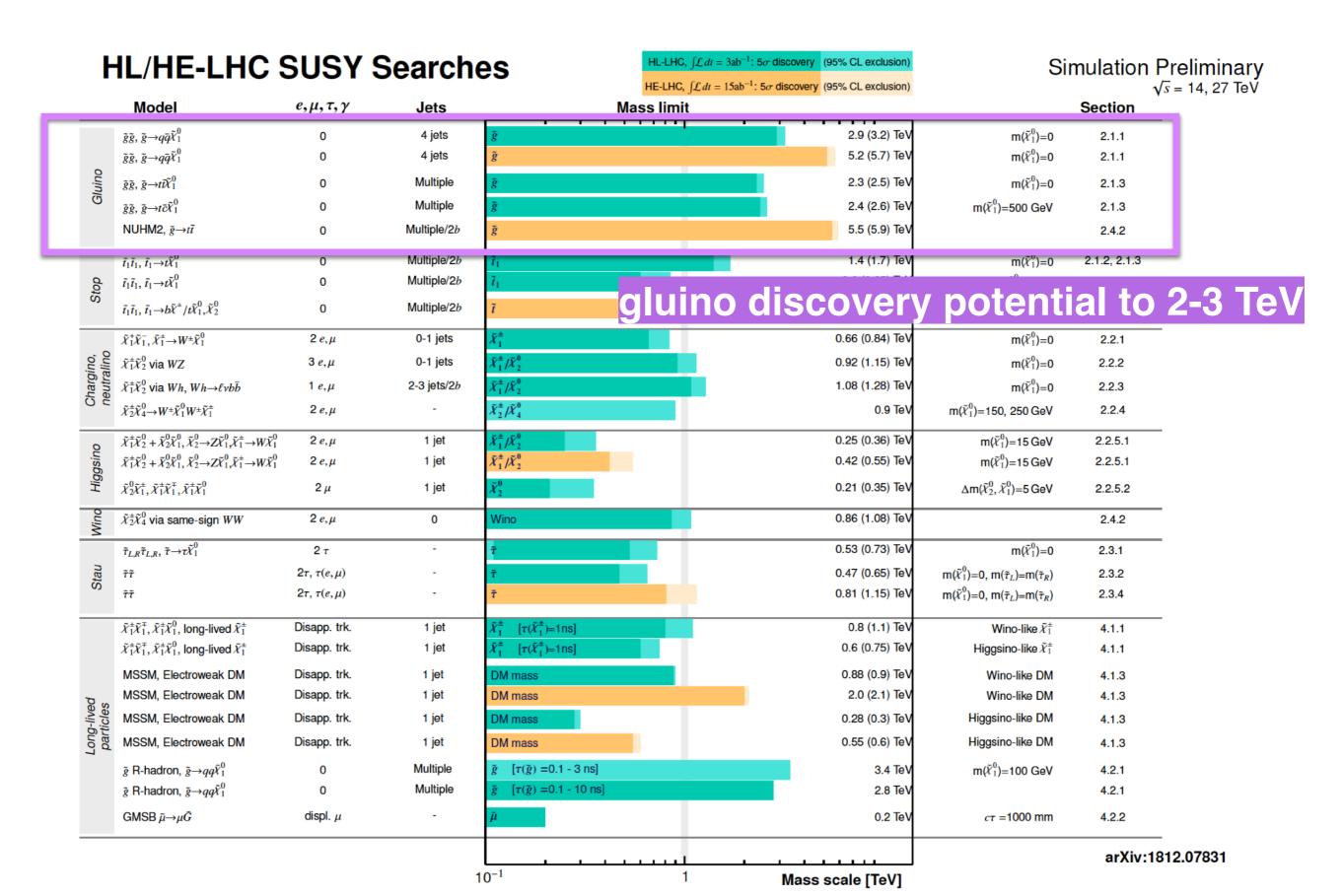
Overview of HL-LHC Working Group

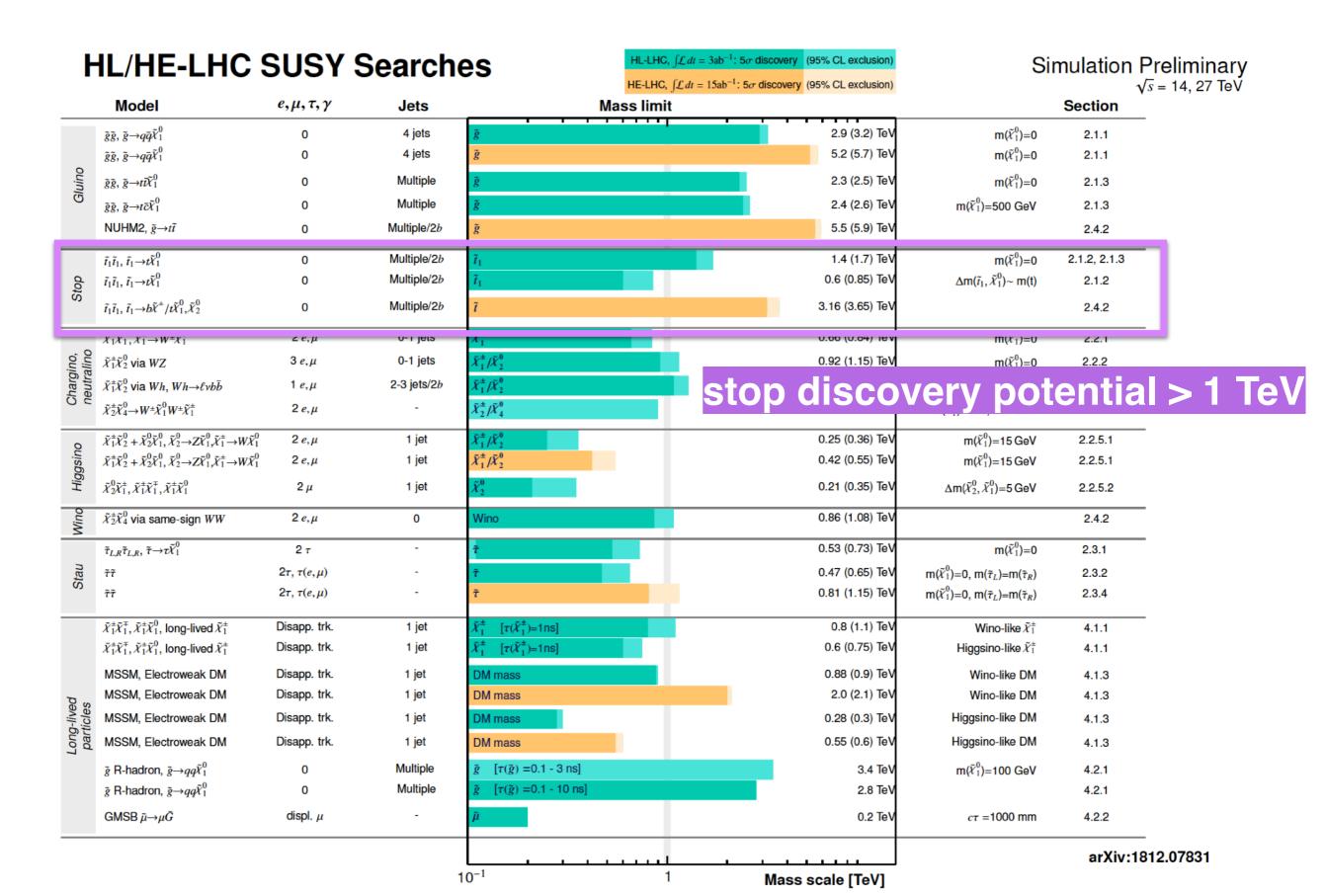
Nov 2017
Kick-offJune 2018
PlenaryOct 2018
PlenaryNov 2018
First YR draftDec 2018
Publication

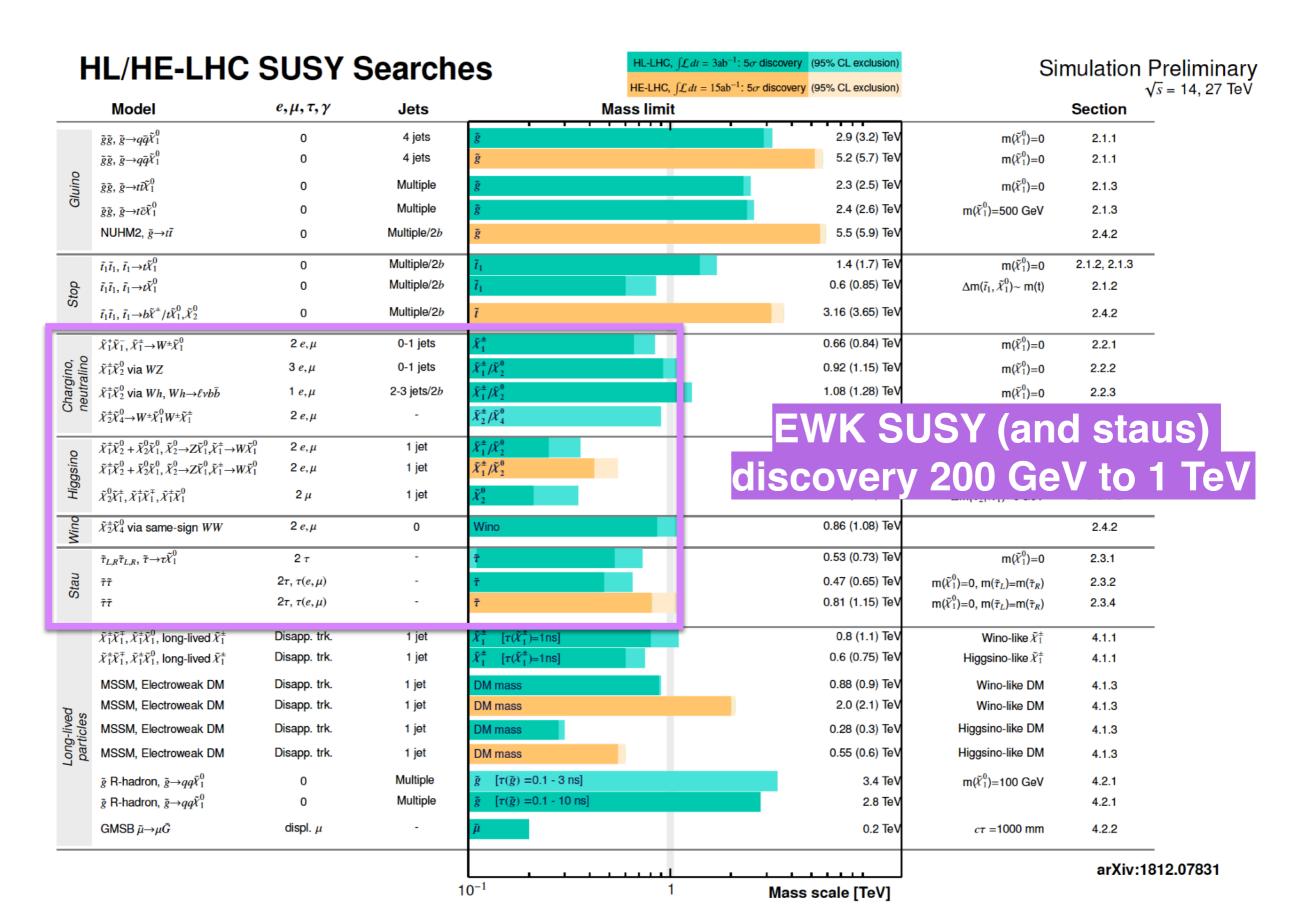
- Collaboration of theorists and experimentalists to project physics reach of HL-LHC
- HL-LHC twiki here
 - WG1: SM <u>final report</u>
 - WG2: Higgs final report
 - WG3: BSM <u>final report</u>
 - WG4: Flavor physics <u>final report</u>
 - WG5: QCD and heavy-ions <u>final report</u>
- Final <u>jamboree</u> with many interesting talks

 This talk will cover a small selection of highlights, mainly from ATLAS and CMS

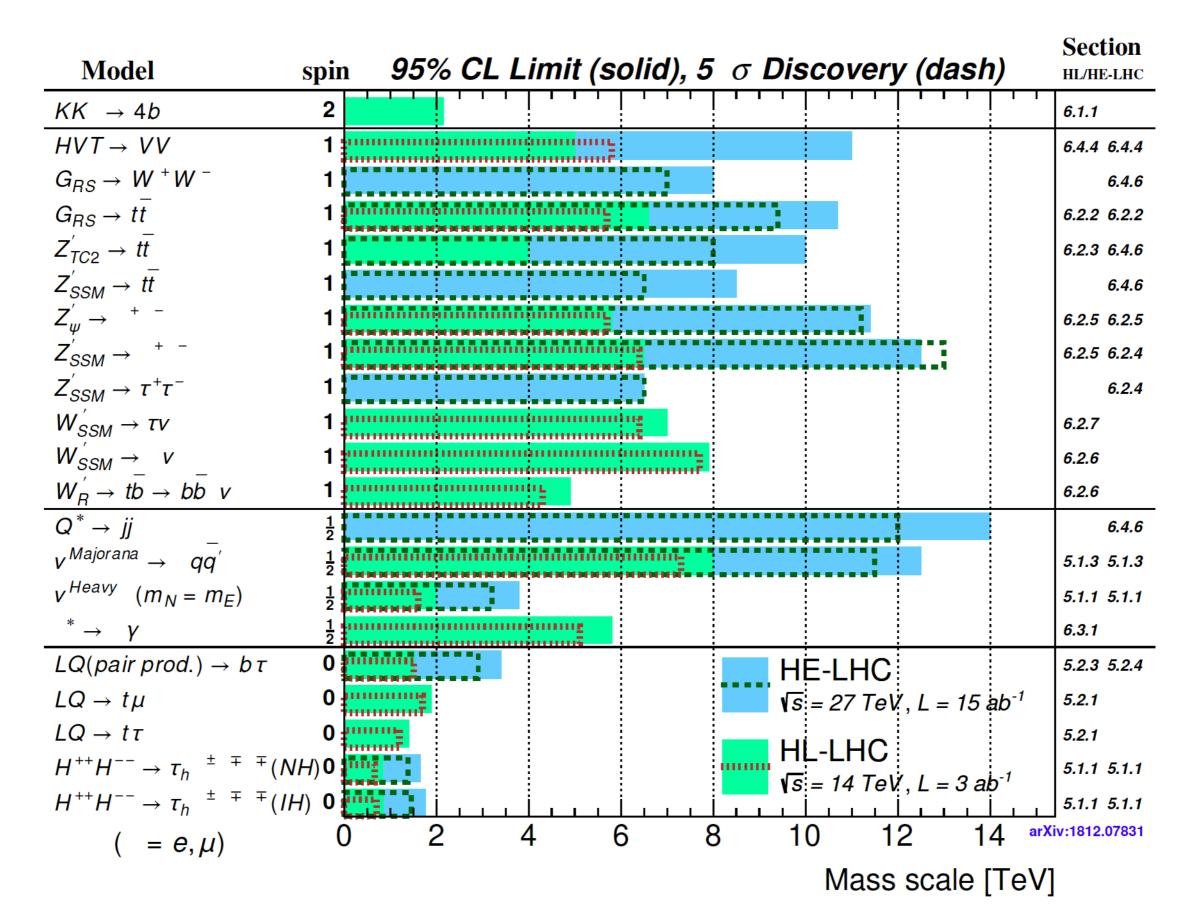




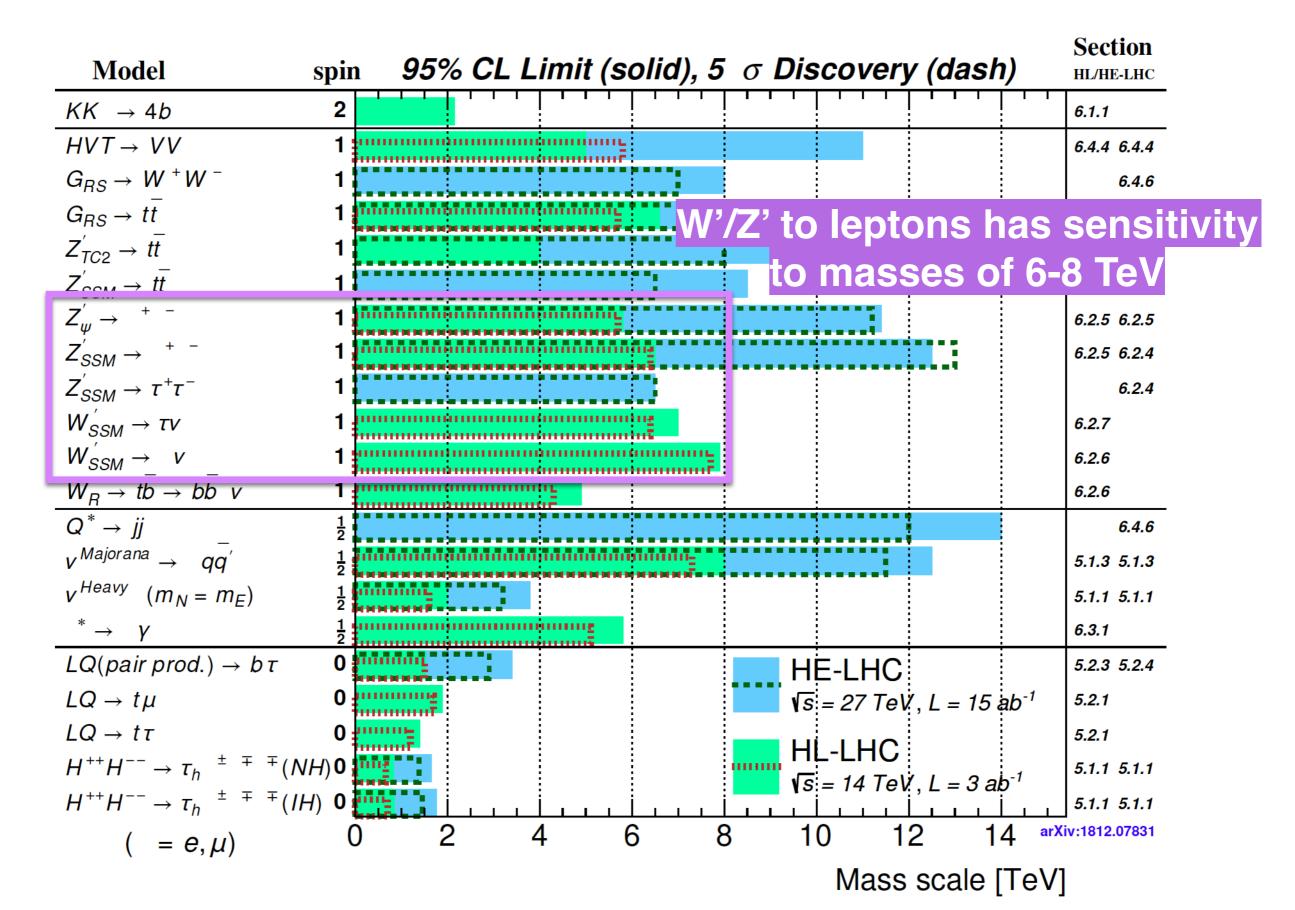




Exotics: breadth and reach



Exotics: breadth and reach

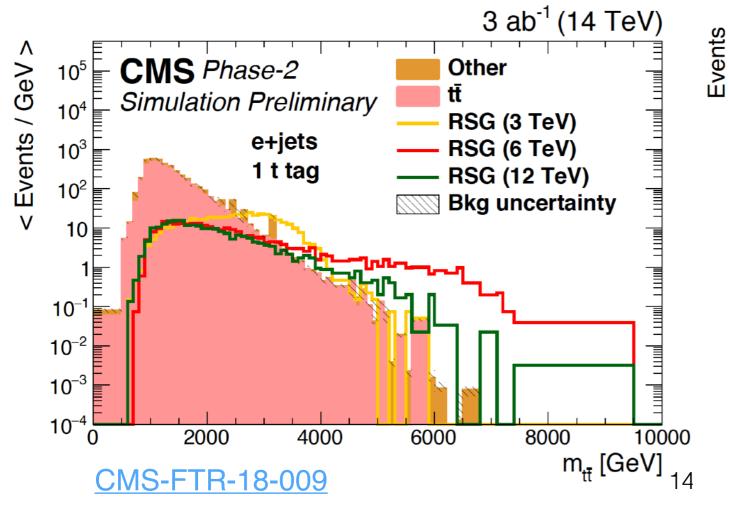


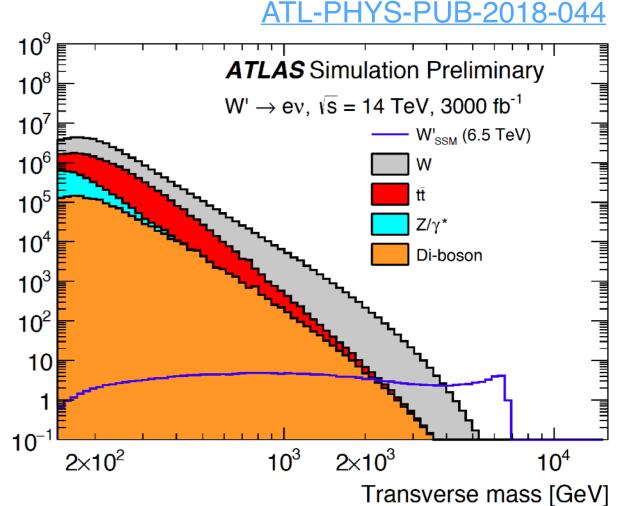
Resonance searches

clear gains from increased lumi and energy

- Continuation and expansion of the current resonance search program, across a vast variety of channels
 - requires continued excellent performance of object reconstruction and energy / mass resolution
 - high mass sensitivity is dominated by statistics
- W' discovery potential up to 7.7 TeV!
- RS Gluon discovery up to 5.7 TeV!

	$\sqrt{s} = 13 \text{ TeV}$		$\sqrt{s} = 14 \text{ TeV}$	
Decay	Exclusion	Discovery	Exclusion	Discovery
$Z'_{\rm SSM} \to ee$	$6.0~{ m TeV}$	$5.9~{ m TeV}$	$6.4 \mathrm{TeV}$	$6.3~{ m TeV}$
$z_{\rm SSM} \to \mu\mu$	0.0 1ev	5.4 1eV	0.8 1eV	5.7 1ev
$Z'_{ ext{SSM}} o \ell\ell$	$6.1~{ m TeV}$	6.1 TeV	$6.5~{ m TeV}$	$6.4~{ m TeV}$
$Z'_{\psi} \to ee$	5.3 TeV	5.3 TeV	5.7 TeV	$5.6~{ m TeV}$
$Z'_{\psi} ightarrow \mu \mu$	$4.9~{ m TeV}$	$4.6~{ m TeV}$	$5.2 \mathrm{TeV}$	$5.0~{ m TeV}$
$Z'_\psi o \ell \ell$	5.4 TeV	$5.4 \mathrm{TeV}$	$5.8 \mathrm{TeV}$	$5.7~{ m TeV}$

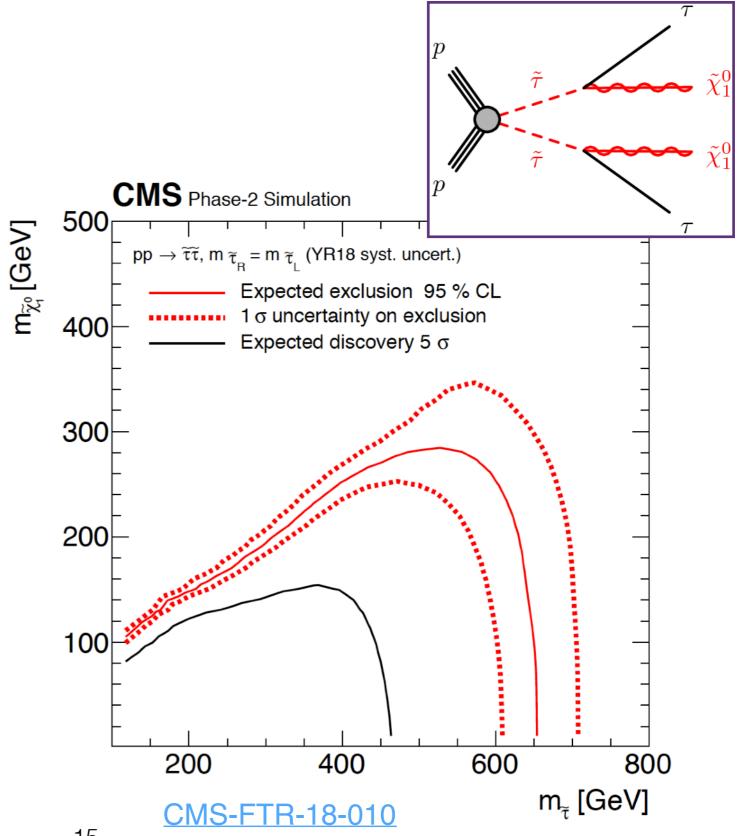




Supersymmetry: staus

new sensitivity relies on detector upgrades for taus

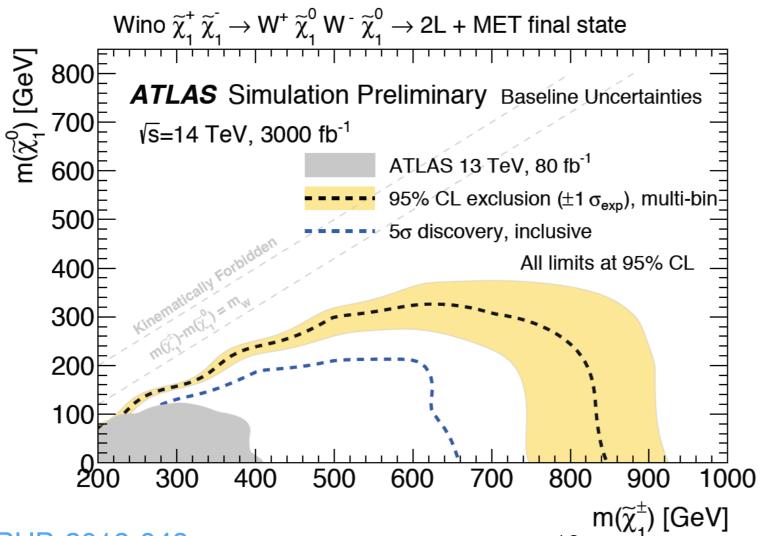
- Stau pair production, decaying to taus and neutralino, currently unconstrained by LHC
 - searched with signature of two hadronically decaying taus, and mixed hadronic / lepton signature
 - discovery sensitivity from each experiment up to 500 GeV
- Detector, trigger, and performance improvements are key
 - relies on ability to trigger taus efficiently at moderate energies
 - need to maintain low fake rake for taus at high pT

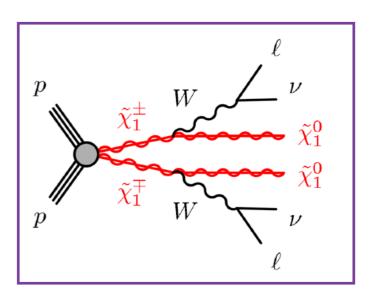


Supersymmetry: ewkinos

pushing down cross-section limits at lower masses

- Wino-like chargino states studied in multiple lepton final states
 - masses up to 650 (850) GeV can be discovered (excluded) from chargino decay to neutralino and W* or up to 900 GeV with degenerate $\tilde{\chi}_2^{\pm} \tilde{\chi}_4^0$
 - limits an order of magnitude beyond LEP and 500 GeV above current limits
 - HL-LHC will probe most of the natural EW SUSY parameter space for wino-like charginos
- Relies on vast upgrades to trigger at L1 and HLT to maintain lepton trigger thresholds!

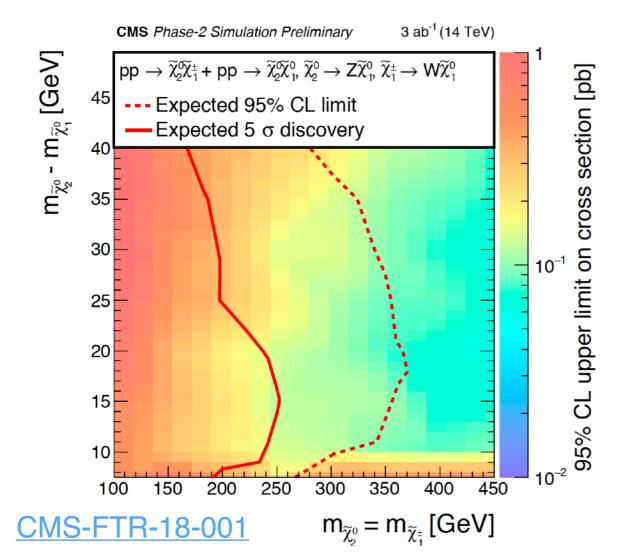


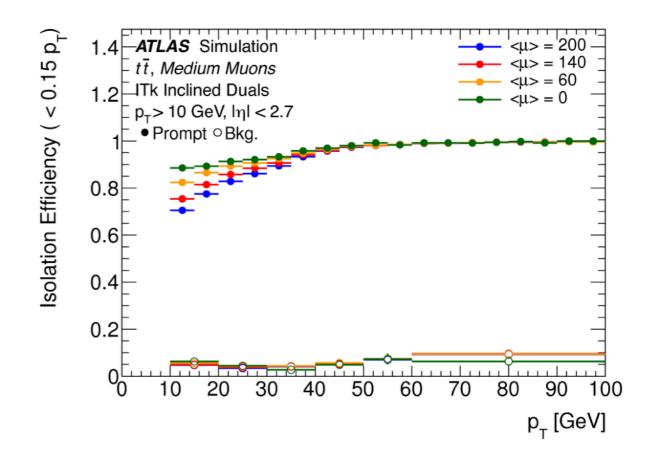


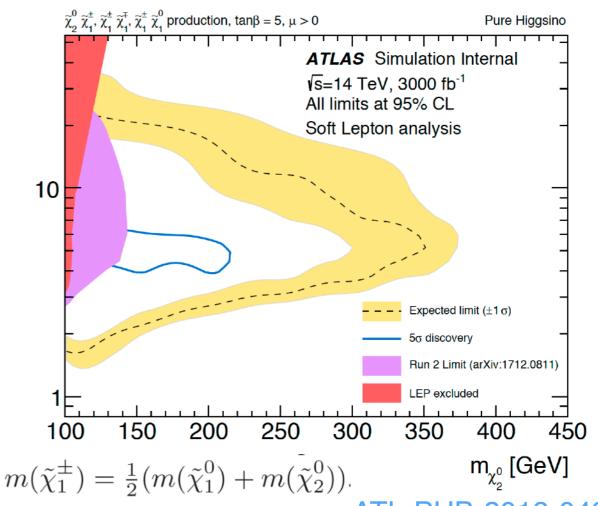
Supersymmetry: ewkinos

- The last bastion of naturalness: higgsinos and compressed EW mass spectra
 - requires efficiently triggering and reconstructing leptons down to 3-5 GeV and difficult background rejection, including isolation
 - sensitive to higgsino masses up to 350
 GeV for mass differences of > 5 GeV

 $\Delta m(\chi_2^0,\chi_1^0)$ [GeV]



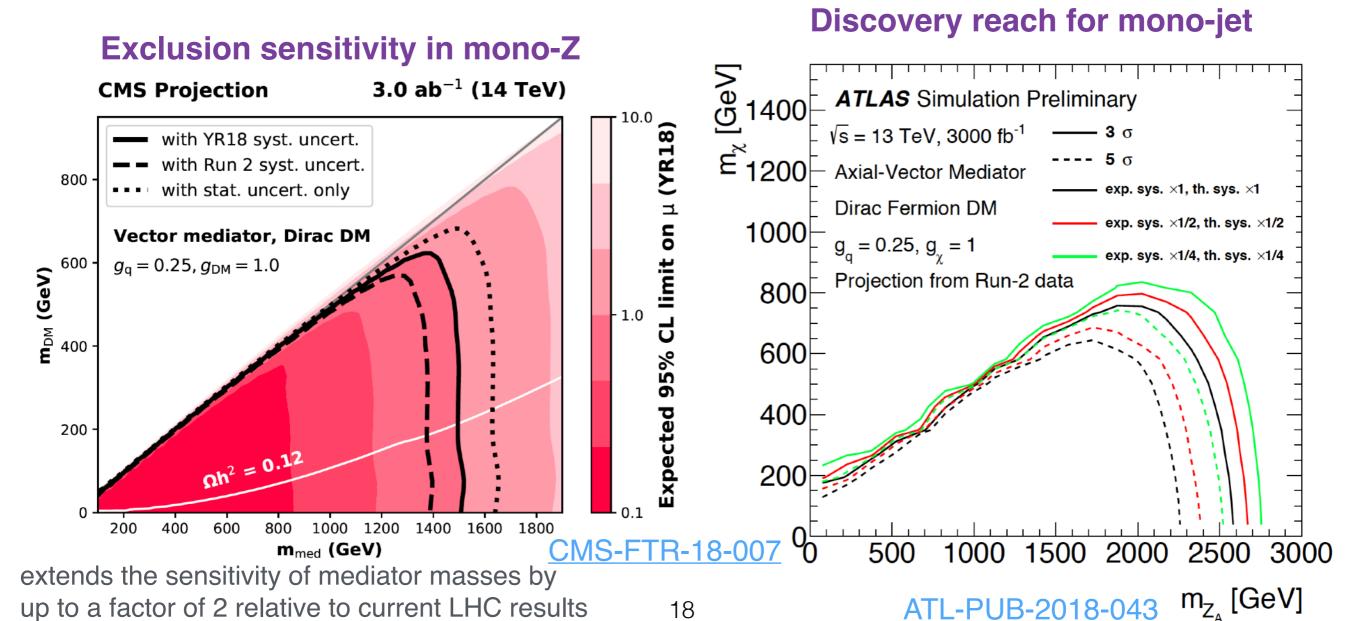




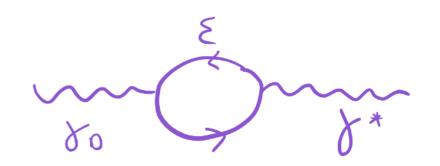
Dark Matter searches

sensitive to systematic uncertainties

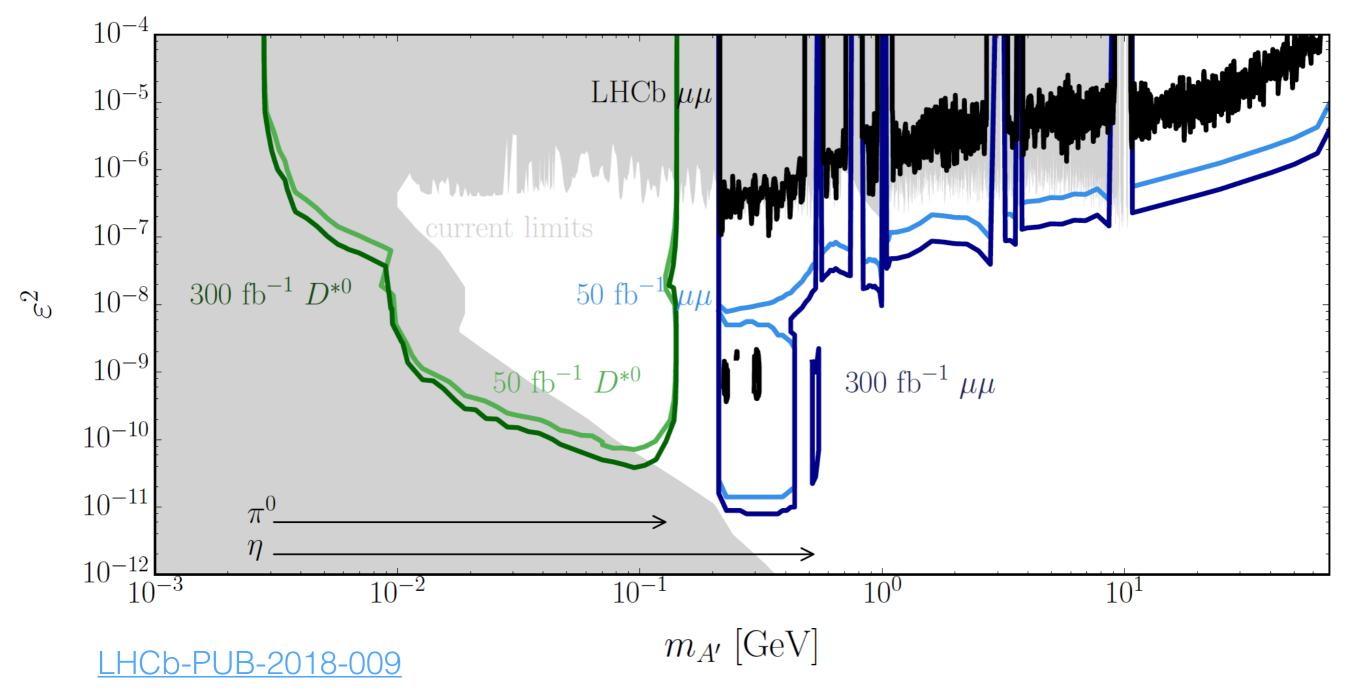
- Mono + X searches will continue to push sensitivity to DM produced in association with other objects
 - will be competitive in some regions with direct neutralino searches, sensitivity from 200 (300) GeV for higgsino (wino) neutralinos
 - quite sensitive to systematic uncertainties, largely on background (W/Z+jets, diboson) estimates



Dark photons



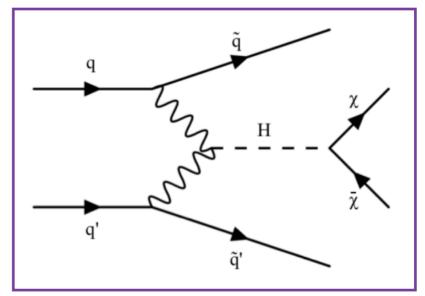
 LHCb will use a combination of prompt, long-lived, and meson analyses to probe a large part of the mass and mixing plane for dark photons

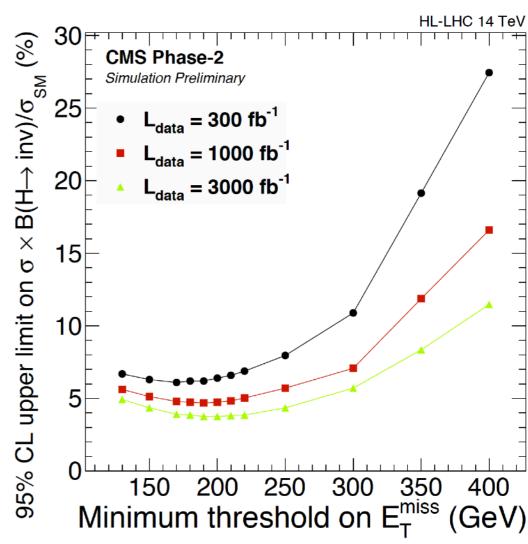


Higgs to invisible

Case study: pileup mitigation and importance of systematic uncertainties

- Higgs to invisible BR is powerful and model independent probe for light dark matter
 - current limit on BR (H—> inv.) is 0.19, mostly driven by vector boson fusion production
 - current limit has ~ equal syst. and stat.
 uncertainties
- Maintaining signal efficiency and background rejection is challenging at high pileup
 - main signal is 2 forward jets and MET
 - rejection of pileup jets is paramount
 - main backgrounds behave differently at HL-LHC relative to LHC
 - \cdot Z-> vv + jets
 - · W \rightarrow vI + jets





Higgs to invisible

- Higgs to invisible BR is powerful and model independent probe for light dark matter
 - single channels from CMS/ATLAS expected to reach 4% on BR (H—> inv.)
 - combination of experiments and channels expected to give limit to 2.5%
 - with significant work to mitigate pileup and reduce systematic and theoretical uncertainties

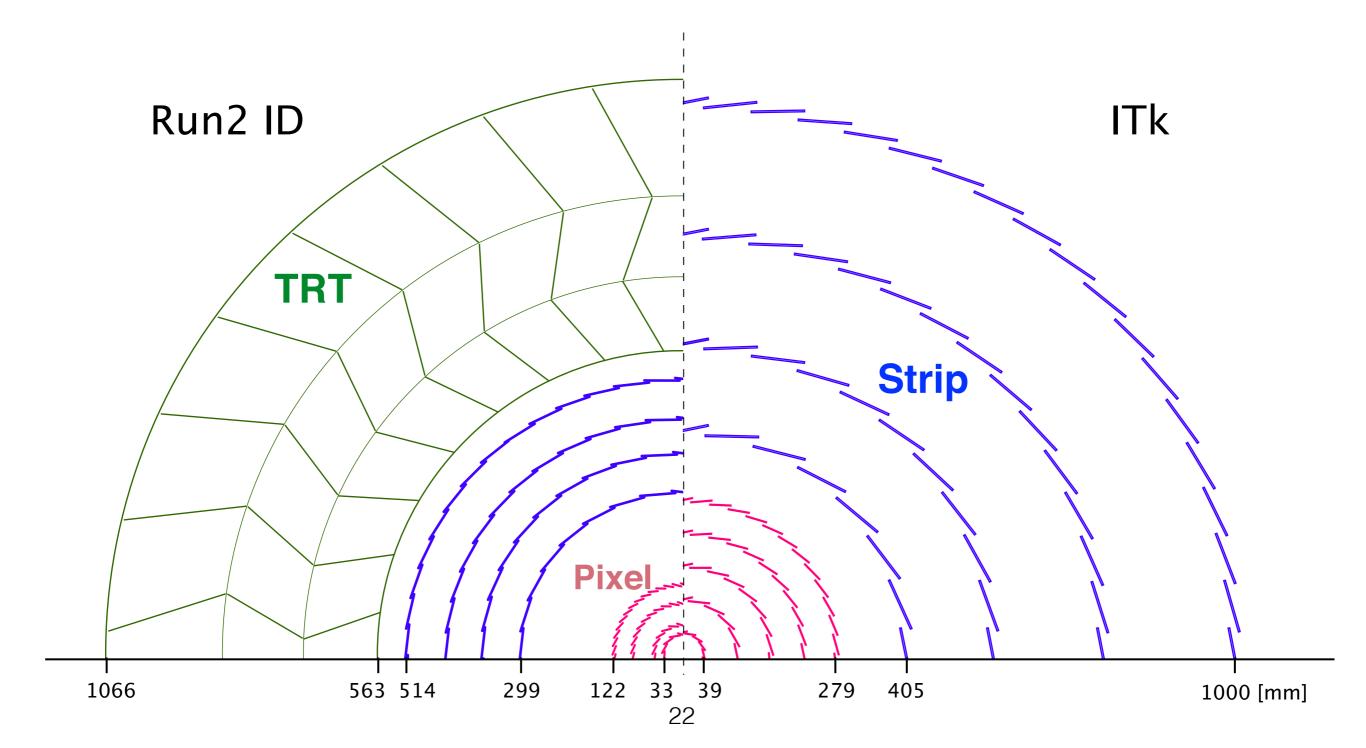
Run 2 syst — goal / assumptions for HL-LHC syst

Systematic	From Ref. [14]		This analysis	
e-ID	1%(gsf)⊕1%(idiso)		1%	
μ-ID	1%(reco)⊕1%(id)⊕0.5%(iso)		0.5%	
e-veto	0.6%(gsf)⊕1.5%(idiso)		1%	
μ -veto on QCD V+jets	5%(reco)⊕5%(id)⊕2%(iso)		2%	
μ-veto on EWK V+2jets	10%(reco)⊕10%(id)⊕6%(iso)		6%	
τ-veto	1–1.5% for QCD–EWK		0.5–0.75%	
b-tag-veto	0.1% (sig) 2% (top)		0.05% (sig) 1% (top)	
JES	14%(sig)	2%(W/W) 1%(7 / 7)	1 5%(cio) () 5%(W/V	V) 0.2%(Z/Z)
Integrated luminosity		2.5%	1%	
QCD multijet	1.5%		1.5%	
Theory on W/Z ratio		12.5%	7%	
ggH normalisation		24%	20%	

different sources of systematic uncertainty considered in Ref. [14] and for the HL-LHC setup considered in this analysis.

Challenges and opportunities: new detectors

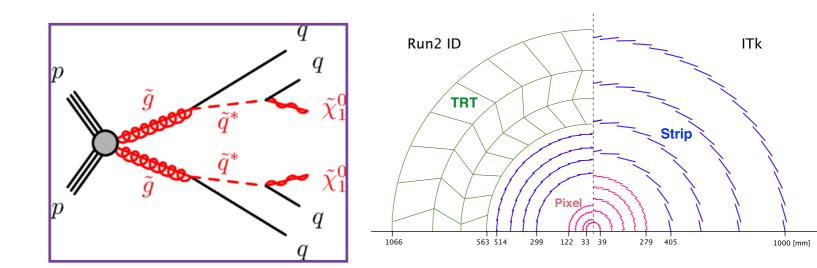
Case study: new inner detector in ATLAS and its impact on sensitivity of long-lived particle searches

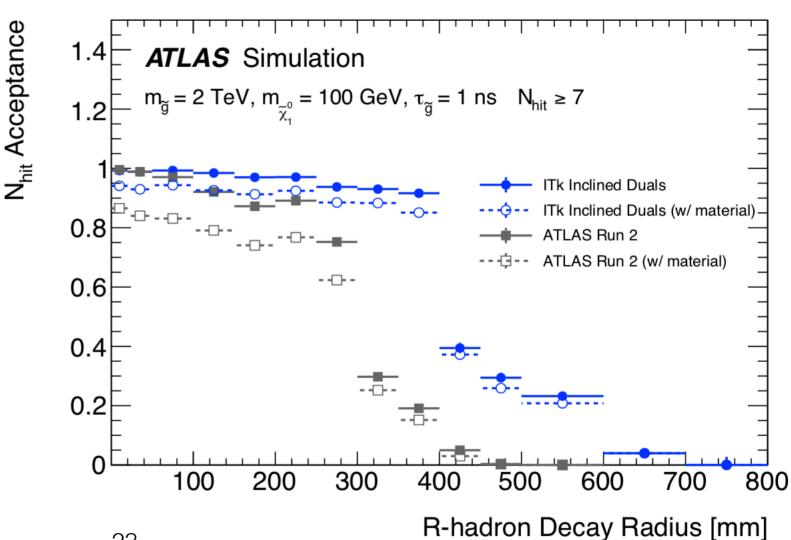


Displaced vertex signatures

- Displaced vertices in the inner detector are sensitive to long-lived particles with τ ~10 ps 10 ns
 - decaying to multiple charged particles

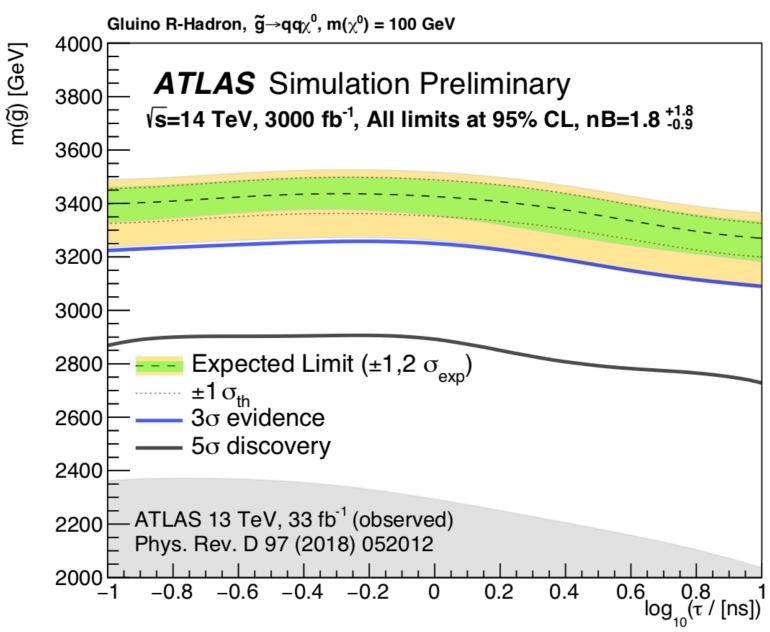
- Need a minimum number of silicon hits after decay to well-reconstruct vertex and reject background
 - reconstruction efficiency gains relative to current detector from larger silicon volume!



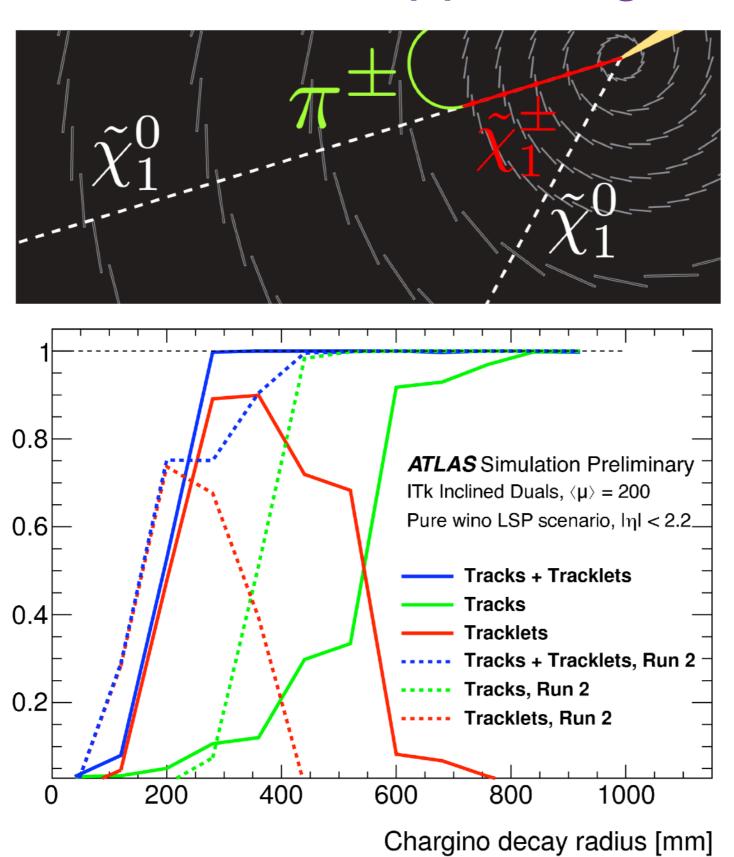


Displaced vertex signatures

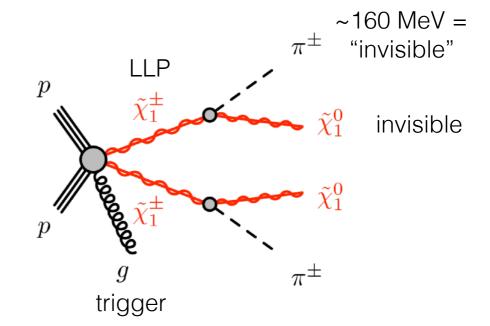
 Translates to increased relative sensitivity at larger lifetimes relative to current analysis, purely from new detector geometry



Disappearing tracks



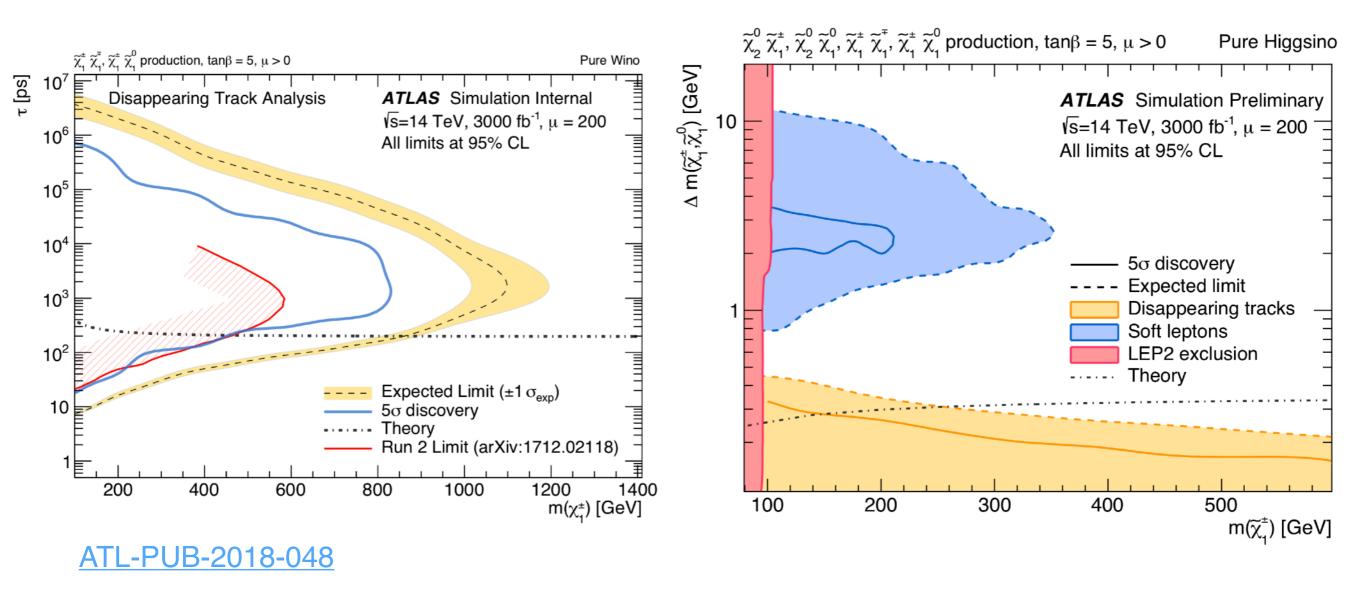
Reconstruction Efficiency



- Pure higgsino or wino LSP can give rise to long-lived chargino which decays to invisible
- Look for short track with 3 or 4 hits in pixel detector
- Efficiency of short tracks very sensitive to position of first few layers of pixel detector!

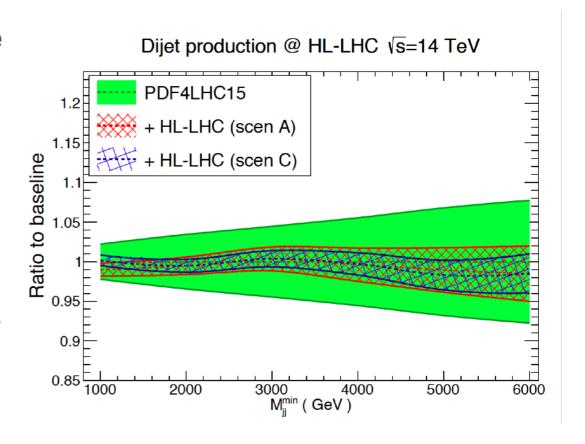
Disappearing tracks

 Sensitivity gains relative to Run 2 (3) more challenging at shorter lifetimes, solely from new detector geometry



Summary and outlook

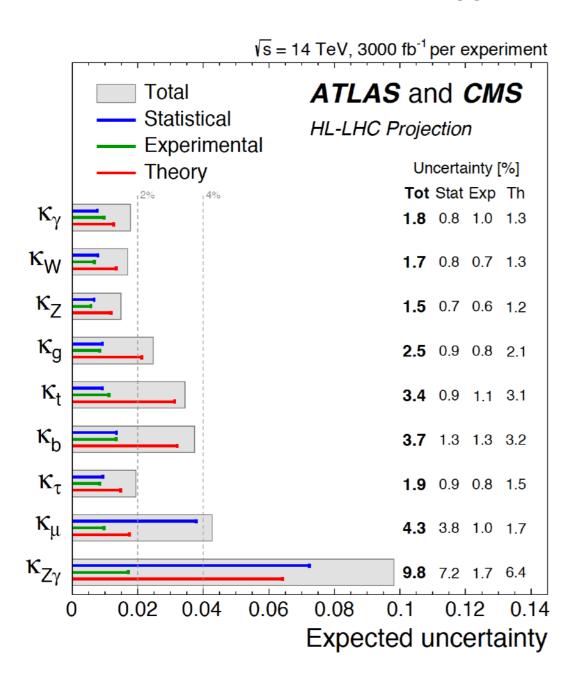
- Enormous work by community to study the potential physics reach of HL-LHC
- The work is just beginning
 - All projections rely on enormous work on many fronts — the accelerator upgrade, the detector upgrades, trigger, performance and reconstruction advancements, analysis improvements, theory improvements, PDF reduction, precise luminosity determination...
- Depending on the new physics, sensitivity (and the potential to discover!) new physics with masses from 200 GeV to 8 TeV
 - some channels have straightforward gains with higher energy, increased lumi, and upgraded detector
 - other channels will require more creative and sustained effort to improve over Run 3 sensitivity



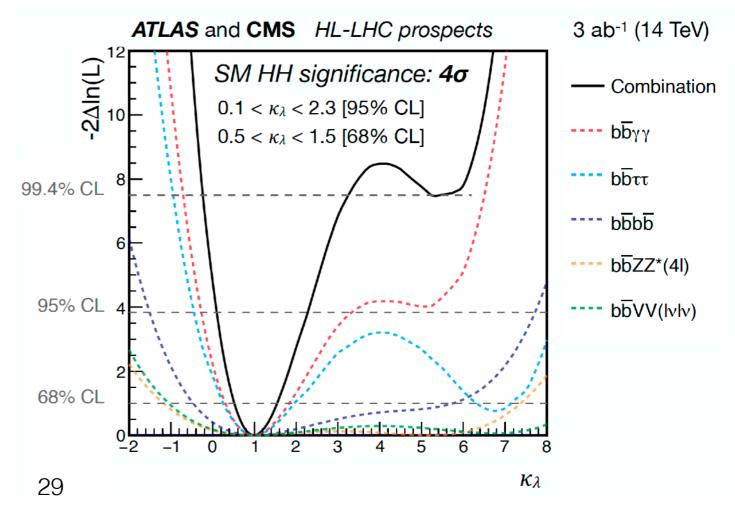
Backup

Higgs physics as a probe of BSM

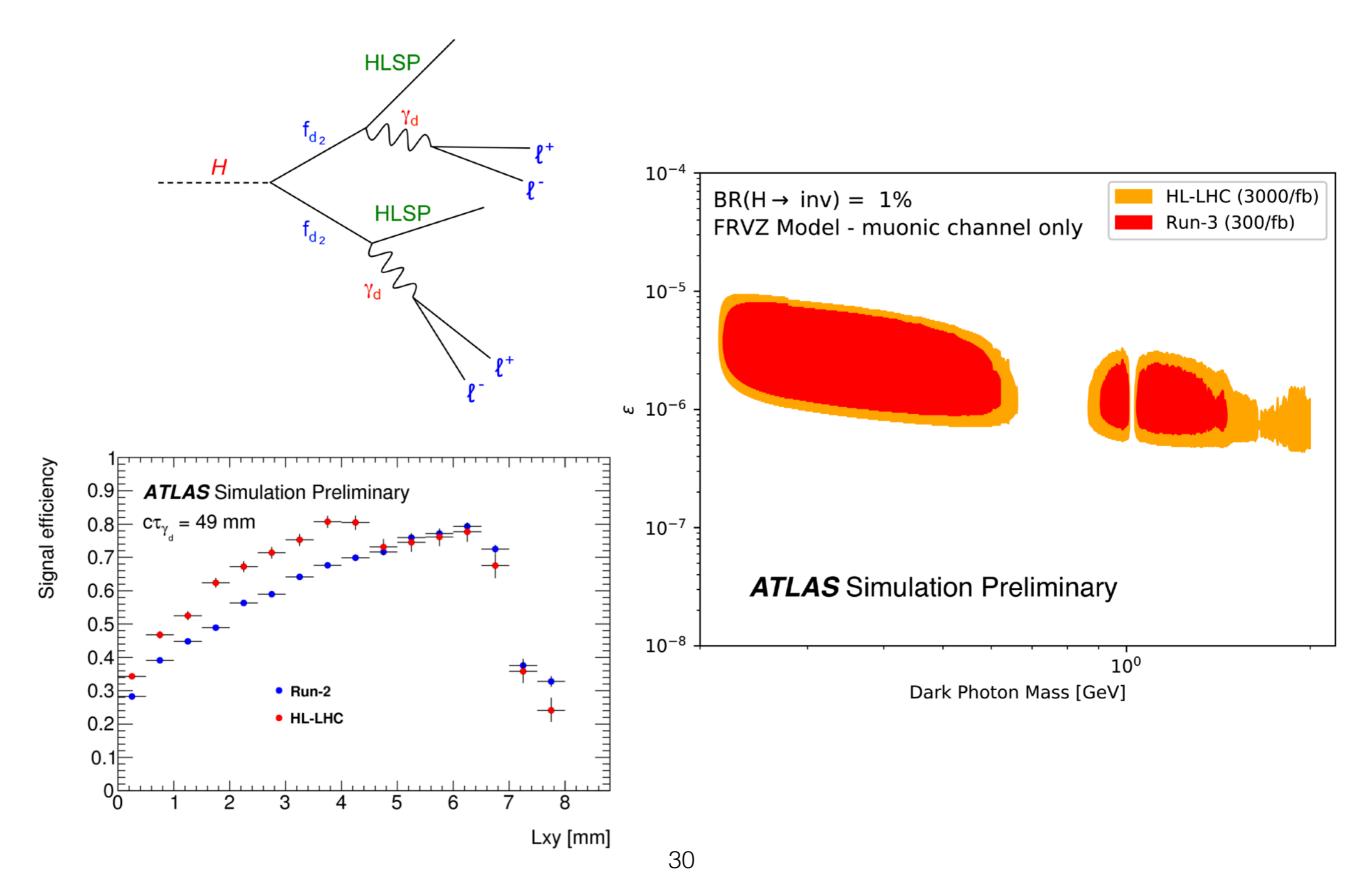
- Main Higgs couplings measured to % level precision
 - limited by theoretical systematic uncertainties
- First measurements of $H \rightarrow \mu\mu$ and $H \rightarrow \gamma Z$
- Improved constraints on higgs self-coupling



 Higgs self-coupling measured to 0.5 < κ_λ < 1.5 at 68% CL

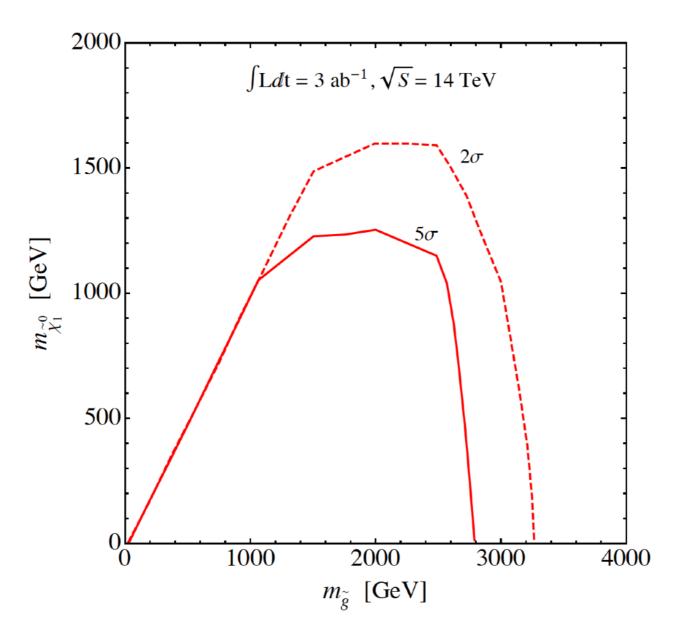


Higgs portal to dark photons



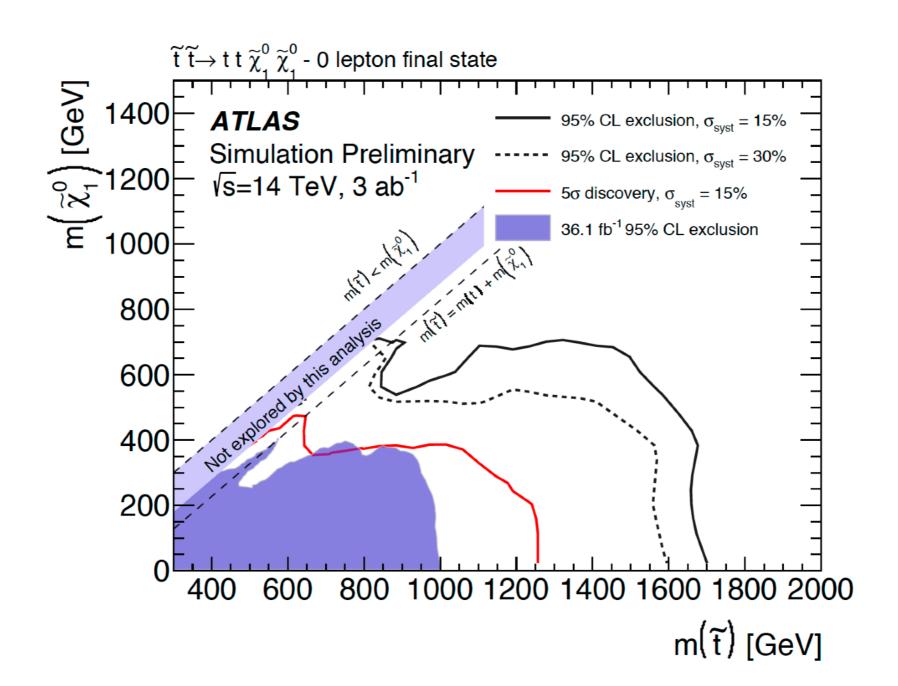
Supersymmetry: gluino and squarks

- Gluino pair production with prompt decay to quarks and neutralino
 - sensitivity to discover 2.8 TeV gluinos and exclude up to 3.2 TeV
 - nearly 1.0 TeV above current Run 2 limits



Supersymmetry: gluino and squarks

- Stop pair production, decaying to prompt top-pair and neutralinos
 - sensitivity to discover 1.25 TeV stops and exclude up to 1.7 TeV
 - nearly 700 GeV above current Run 2 limits



ATLAS SUSY Searches* - 95% CL Lower Limits

ATLAS Preliminary

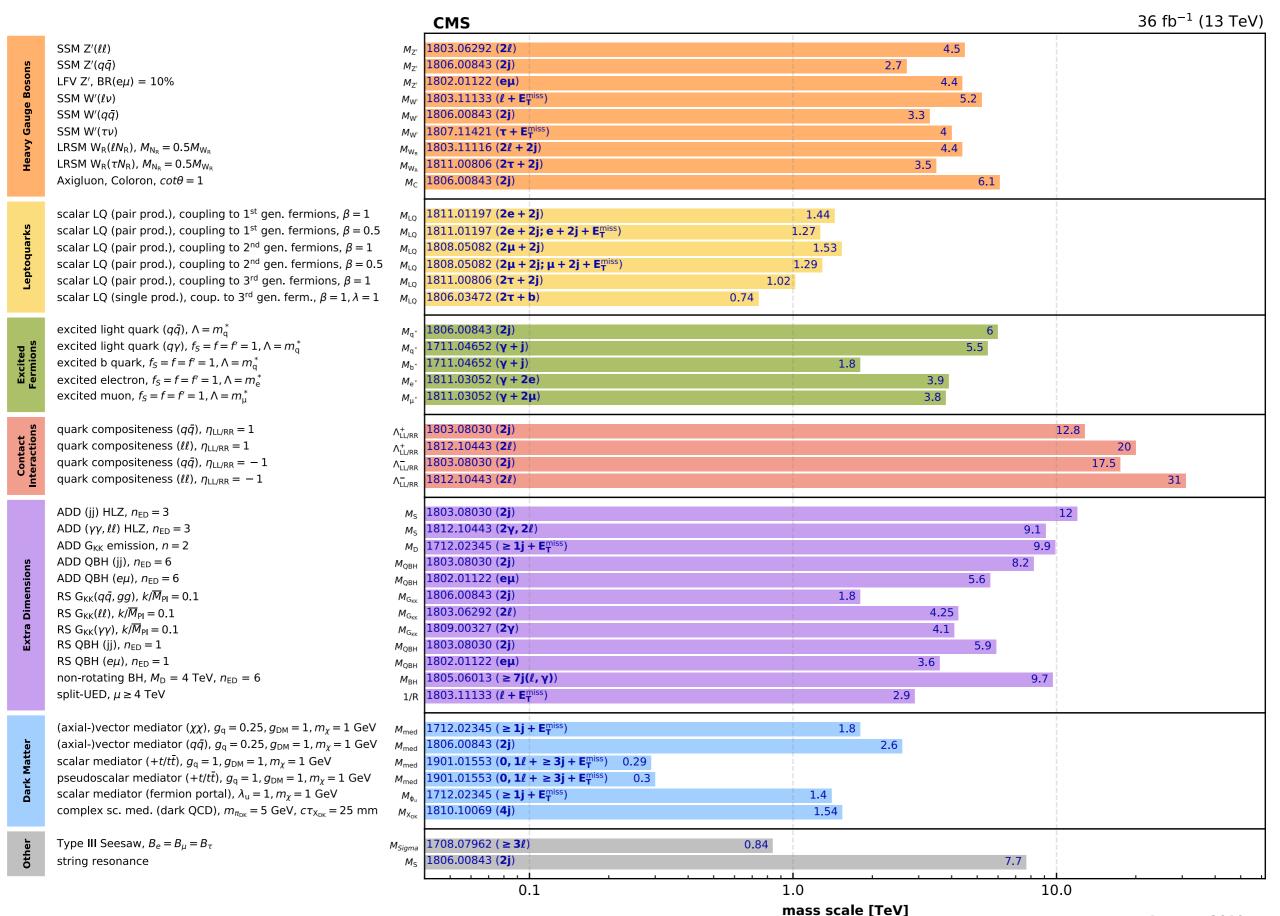
Mass scale [TeV]

 \sqrt{s} = 7. 8. 13 TeV July 2018 e, μ, τ, γ Jets $E_{\mathrm{T}}^{\mathrm{miss}} \int \mathcal{L} \, dt [\mathrm{fb}^{-1}]$ **Mass limit** Reference Model $\sqrt{s} = 7, 8 \text{ TeV}$ \sqrt{s} = 13 TeV $\tilde{q}\tilde{q}, \, \tilde{q} \rightarrow q\tilde{\chi}_1^0$ 2-6 jets [2x, 8x Degen.] 1.55 $m(\tilde{\chi}_{\perp}^{0})<100\,\text{GeV}$ 1712.02332 Yes 36.1 0 0.9 [1x, 8x Degen.] mono-iet 1-3 jets Yes 36.1 0.43 0.71 $m(\tilde{q})-m(\tilde{\chi}_{1}^{0})=5 \text{ GeV}$ 1711.03301 Inclusive Searches $\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q}\tilde{\chi}_1^0$ 0 2-6 jets Yes 36.1 2.0 $m(\tilde{\chi}_{\perp}^{0})$ <200 GeV 1712.02332 Forbidden 0.95-1.6 $m(\tilde{\chi}_1^0)=900 \,\text{GeV}$ 1712.02332 $3e, \mu$ 4 jets $m(\tilde{\chi}_1^0)$ <800 GeV 1706.03731 $\tilde{g}\tilde{g}, \, \tilde{g} \rightarrow q\bar{q}(\ell\ell)\tilde{\chi}_1^0$ 36.1 1.85 ee, µµ 2 jets Yes 36.1 1.2 $m(\tilde{g})-m(\tilde{\chi}_1^0)=50 \text{ GeV}$ 1805.11381 1.8 $\tilde{g}\tilde{g}, \, \tilde{g} \rightarrow qqWZ\tilde{\chi}_1^0$ 0 7-11 jets Yes 36.1 $m(\tilde{\chi}_{\perp}^{0}) < 400 \,\text{GeV}$ 1708.02794 4 jets 0.98 $3e, \mu$ 36.1 $m(\tilde{g})-m(\tilde{\chi}_1^0)=200 \text{ GeV}$ 1706.03731 $\tilde{g}\tilde{g}, \, \tilde{g} \rightarrow t\bar{t}\tilde{\chi}_1^0$ $0-1~e,\mu$ 3 b Yes 36.1 2.0 $m(\tilde{\chi}_1^0)$ <200 GeV 1711.01901 З e,μ 4 jets 1.25 $m(\tilde{g})-m(\tilde{\chi}_1^0)=300 \text{ GeV}$ 36.1 1706.03731 $\tilde{b}_1 \tilde{b}_1, \, \tilde{b}_1 \rightarrow b \tilde{\chi}_1^0 / t \tilde{\chi}_1^{\pm}$ Multiple Forbidden $m(\tilde{\chi}_{\perp}^{0})=300 \text{ GeV}, BR(b\tilde{\chi}_{\perp}^{0})=1$ 1708.09266, 1711.03301 36.1 \tilde{b}_1 0.9 Multiple 0.58-0.82 $m(\tilde{\chi}_{\perp}^{0})=300 \text{ GeV, } BR(b\tilde{\chi}_{\perp}^{0})=BR(t\tilde{\chi}_{\perp}^{\pm})=0.5$ 36.1 \tilde{b}_1 Forbidden 1708.09266 Multiple 36.1 Forbidden 0.7 $m(\tilde{\chi}_{1}^{0})=200 \text{ GeV}, m(\tilde{\chi}_{1}^{\pm})=300 \text{ GeV}, BR(t\tilde{\chi}_{1}^{\pm})=1$ \tilde{b}_1 1706.03731 $\tilde{b}_1 \tilde{b}_1, \tilde{t}_1 \tilde{t}_1, M_2 = 2 \times M_1$ Multiple 1709.04183, 1711.11520, 1708.03247 36.1 0.7 $m(\tilde{\chi}_1^0)=60 \text{ GeV}$ Multiple Forbidden 0.9 1709.04183, 1711.11520, 1708.03247 36.1 \tilde{t}_1 $m(\tilde{\chi}_{\perp}^{0})=200 \,\text{GeV}$ $\tilde{t}_1 \tilde{t}_1, \, \tilde{t}_1 \rightarrow Wb\tilde{\chi}_1^0 \text{ or } t\tilde{\chi}_1^0$ 0-2 e, µ 0-2 jets/1-2 b Yes 36.1 1506.08616, 1709.04183, 1711.11520 \tilde{t}_1 1.0 $m(\tilde{\chi}_1^0)=1 \text{ GeV}$ $\tilde{t}_1\tilde{t}_1,\,\tilde{H}\,\mathsf{LSP}$ Multiple 36.1 \tilde{t}_1 0.4-0.9 $m(\tilde{\chi}_{\perp}^{0})=150 \text{ GeV}, m(\tilde{\chi}_{\perp}^{\pm})-m(\tilde{\chi}_{\perp}^{0})=5 \text{ GeV}, \tilde{t}_{\perp} \approx \tilde{t}_{L}$ 1709.04183, 1711.11520 Multiple 36.1 $\tilde{\iota}_1$ Forbidden 0.6-0.8 $m(\tilde{\chi}_1^0)$ =300 GeV, $m(\tilde{\chi}_1^{\pm})$ - $m(\tilde{\chi}_1^0)$ =5 GeV, $\tilde{\iota}_1 \approx \tilde{\iota}_L$ 1709.04183, 1711.11520 $\tilde{t}_1\tilde{t}_1$, Well-Tempered LSP \tilde{t}_1 1709.04183, 1711.11520 Multiple 36.1 0.48-0.84 $m(\tilde{\chi}_{\perp}^{0})=150 \text{ GeV}, m(\tilde{\chi}_{\perp}^{\pm})-m(\tilde{\chi}_{\perp}^{0})=5 \text{ GeV}, \tilde{t}_{\perp} \approx \tilde{t}_{L}$ $\begin{array}{c} \mathsf{m}(\tilde{\chi}_1^0) {=} \mathsf{0} \; \mathsf{GeV} \\ \mathsf{m}(\tilde{t}_1, \tilde{c}) {-} \mathsf{m}(\tilde{\chi}_1^0) {=} \mathsf{50} \; \mathsf{GeV} \\ \mathsf{m}(\tilde{t}_1, \tilde{c}) {-} \mathsf{m}(\tilde{\chi}_1^0) {=} \mathsf{5} \; \mathsf{GeV} \end{array}$ $\tilde{t}_1 \tilde{t}_1, \, \tilde{t}_1 \rightarrow c \tilde{\chi}_1^0 / \, \tilde{c} \tilde{c}, \, \tilde{c} \rightarrow c \tilde{\chi}_1^0$ 0.85 0 2cYes 36.1 1805.01649 0.46 1805.01649 0 mono-jet Yes 36.1 0.43 1711.03301 $\tilde{t}_2\tilde{t}_2, \, \tilde{t}_2 \rightarrow \tilde{t}_1 + h$ 1-2 e, μ Yes 36.1 \tilde{t}_2 0.32-0.88 $m(\tilde{\chi}_1^0)=0$ GeV, $m(\tilde{t}_1)-m(\tilde{\chi}_1^0)=180$ GeV 1706.03986 4 b $\begin{array}{c} \tilde{\chi}_1^{\pm}/\tilde{\chi}_2^0 \\ \tilde{\chi}_1^{\pm}/\tilde{\chi}_2^0 \end{array}$ $\tilde{\chi}_1^{\pm} \tilde{\chi}_2^0$ via WZ2-3 e, μ Yes 36.1 0.6 $m(\tilde{\chi}_1^0)=0$ 1403.5294, 1806.02293 ee, $\mu\mu$ 36.1 0.17 $m(\tilde{\chi}_{\perp}^{\pm})-m(\tilde{\chi}_{\perp}^{0})=10 \text{ GeV}$ 1712.08119 ≥ 1 Yes $\tilde{\chi}_1^{\pm} \tilde{\chi}_2^0$ via Wh $\ell\ell/\ell\gamma\gamma/\ell bb$ $\tilde{\chi}_1^{\pm}/\tilde{\chi}_2^0$ Yes 20.3 0.26 $m(\tilde{\chi}_1^0)=0$ 1501.07110 $\begin{array}{c} \tilde{\chi}_1^{\pm}/\tilde{\chi}_2^0 \\ \tilde{\chi}_1^{\pm}/\tilde{\chi}_2^0 \end{array}$ $\tilde{\chi}_{1}^{\pm}\tilde{\chi}_{1}^{\mp}/\tilde{\chi}_{2}^{0}, \tilde{\chi}_{1}^{+} \rightarrow \tilde{\tau}\nu(\tau\tilde{\nu}), \tilde{\chi}_{2}^{0} \rightarrow \tilde{\tau}\tau(\nu\tilde{\nu})$ 2 τ Yes 36.1 0.76 $m(\tilde{\chi}_1^0)=0$, $m(\tilde{\tau}, \tilde{\nu})=0.5(m(\tilde{\chi}_1^{\pm})+m(\tilde{\chi}_1^0))$ 1708.07875 0.22 $m(\tilde{\chi}_{\perp}^{\pm}) - m(\tilde{\chi}_{\perp}^{0}) = 100 \text{ GeV}, m(\tilde{\tau}, \tilde{\nu}) = 0.5(m(\tilde{\chi}_{\perp}^{\pm}) + m(\tilde{\chi}_{\perp}^{0}))$ 1708.07875 $\tilde{\ell}_{L,R}\tilde{\ell}_{L,R}, \tilde{\ell} \rightarrow \ell \tilde{\chi}_1^0$ $2e, \mu$ 0 Yes 36.1 0.5 $m(\tilde{\chi}_1^0)=0$ 1803.02762 0.18 $2e, \mu$ ≥ 1 Yes 36.1 $m(\tilde{\ell})-m(\tilde{\chi}_1^0)=5 \text{ GeV}$ 1712.08119 $\tilde{H}\tilde{H}, \tilde{H} \rightarrow h\tilde{G}/Z\tilde{G}$ $\mathsf{BR}(\tilde{\chi}^0_1 \to h\tilde{G}) = 1$ 0.13-0.23 0 $\geq 3b$ 36.1 Й 0.29-0.88 1806.04030 Yes $4e, \mu$ 0 36.1 0.3 $BR(\tilde{\chi}_1^0 \to Z\tilde{G})=1$ 1804.03602 Yes \tilde{H} Direct $\tilde{\chi}_1^+ \tilde{\chi}_1^-$ prod., long-lived $\tilde{\chi}_1^{\pm}$ Disapp. trk 1 jet 0.46 Yes 36.1 Pure Wino 1712.02118 0.15 Pure Higgsino ATL-PHYS-PUB-2017-019 Stable \tilde{g} R-hadron SMP 3.2 1.6 1606.05129 Metastable \tilde{g} R-hadron, $\tilde{g} \rightarrow qq\tilde{\chi}_1^0$ Multiple $[\tau(\tilde{g}) = 100 \text{ ns}, 0.2 \text{ ns}]$ 2.4 1710.04901.1604.04520 32.8 1.6 $m(\tilde{\chi}_1^0)=100 \text{ GeV}$ GMSB, $\tilde{\chi}_1^0 \rightarrow \gamma \tilde{G}$, long-lived $\tilde{\chi}_1^0$ 0.44 2γ 20.3 $1 < \tau(\tilde{\chi}_1^0) < 3$ ns, SPS8 model 1409.5542 Yes $\tilde{g}\tilde{g}, \tilde{\chi}_{1}^{0} \rightarrow eev/e\mu v/\mu\mu v$ displ. $ee/e\mu/\mu\mu$ 20.3 1.3 $6 < c\tau(\tilde{\chi}_1^0) < 1000 \text{ mm, m}(\tilde{\chi}_1^0) = 1 \text{ TeV}$ 1504.05162 LFV $pp \rightarrow \tilde{v}_{\tau} + X, \tilde{v}_{\tau} \rightarrow e\mu/e\tau/\mu\tau$ εμ,ετ,μτ λ'_{311} =0.11, $\lambda_{132/133/233}$ =0.07 3.2 1.9 1607.08079 $\tilde{\chi}_{1}^{\pm}\tilde{\chi}_{1}^{\mp}/\tilde{\chi}_{2}^{0} \rightarrow WW/Z\ell\ell\ell\ell\nu\nu$ $4e, \mu$ 36.1 0.82 1.33 0 Yes $\tilde{\chi}_1^{\pm}/\tilde{\chi}_2^0 \quad [\lambda_{i33} \neq 0, \lambda_{12k} \neq 0]$ $m(\tilde{\chi}_1^0)=100 \text{ GeV}$ 1804.03602 Large λ"₁₁₂ $\tilde{g}\tilde{g}, \tilde{g} \rightarrow qq\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow qqq$ 4-5 large-R jets $[m(\tilde{\chi}_{1}^{0})=200 \text{ GeV}, 1100 \text{ GeV}]$ $[\lambda_{112}^{\prime\prime}=2\text{e-4}, 2\text{e-5}]$ 36.1 1.3 1.9 1804.03568 Multiple 1.05 36.1 2.0 $m(\tilde{\chi}_1^0)=200$ GeV, bino-like ATLAS-CONF-2018-003 $\tilde{g}\tilde{g}, \tilde{g} \to tbs / \tilde{g} \to t\bar{t}\tilde{\chi}_1^0, \tilde{\chi}_1^0 \to tbs$ Multiple 36.1 $\tilde{g} = [\lambda_{323}'' = 1, 1e-2]$ 1.8 2.1 $m(\tilde{\chi}_1^0)$ =200 GeV, bino-like ATLAS-CONF-2018-003 $\tilde{t}\tilde{t}, \tilde{t} \rightarrow t\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow tbs$ $\tilde{g} = [\lambda_{323}'' = 2e-4, 1e-2]$ Multiple 36.1 0.55 1.05 $m(\tilde{\chi}_1^0)=200$ GeV, bino-like ATLAS-CONF-2018-003 $\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow bs$ 2 jets + 2 b 36.7 0.42 0.61 0 $\tilde{t}_1 = [qq, bs]$ 1710.07171 $\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow b\ell$ $2e, \mu$ 2b \tilde{t}_1 0.4-1.45 BR($\tilde{t}_1 \rightarrow be/b\mu$)>20% 1710.05544 36.1

 10^{-1}

^{*}Only a selection of the available mass limits on new states or phenomena is shown. Many of the limits are based on simplified models, c.f. refs. for the assumptions made.

Overview of CMS EXO results



Challenges and opportunities: uncertainties

- Many of the HL-LHC measurements and searches rely on reducing both theoretical and experimental systematic uncertainties
- Common assumptions include
 - Precision measurements assume a factor two reduction in theoretical uncertainties
 - Continued precise determination of luminosity in a harsher environment to 2 or even 1% level requires new luminosity detectors and new analysis techniques
 - Continued SM measurements are essential to reduce PDF uncertainties by a factor of 2-4, which is already assumed in many of these results
 - New, more precise forward detectors in ATLAS and CMS, and LHCb measurements will help

