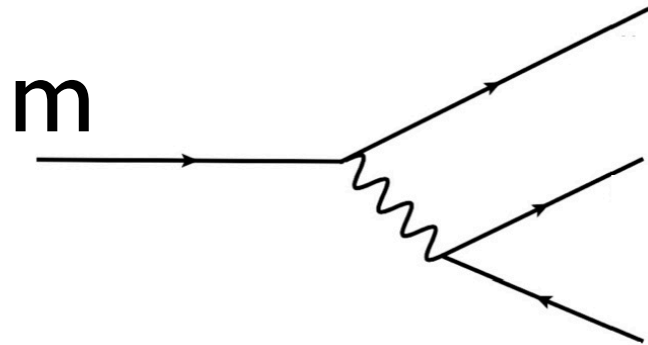




*Collab meeting
Simons Center,
August 27-31, Stony Brook*

Michelangelo L. Mangano
michelangelo.mangano@cern.ch
Theoretical Physics Department
CERN

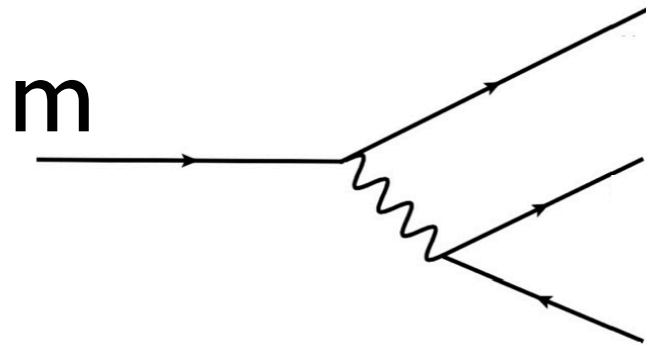
Weak lifetimes in the SM



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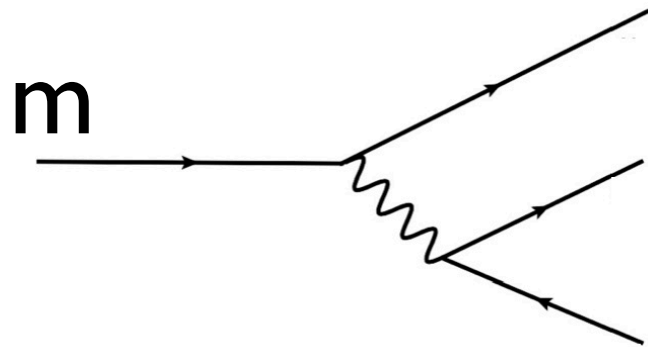


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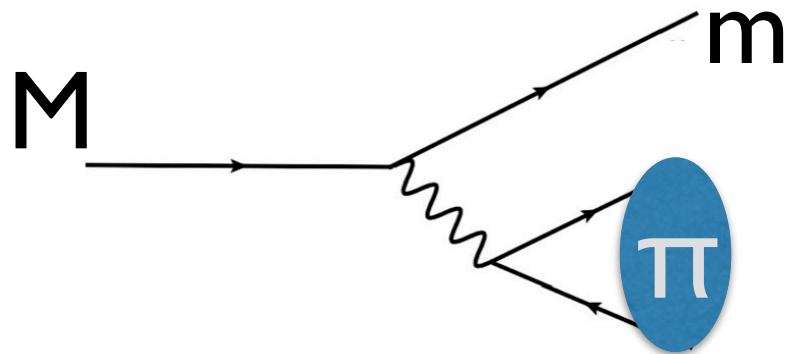
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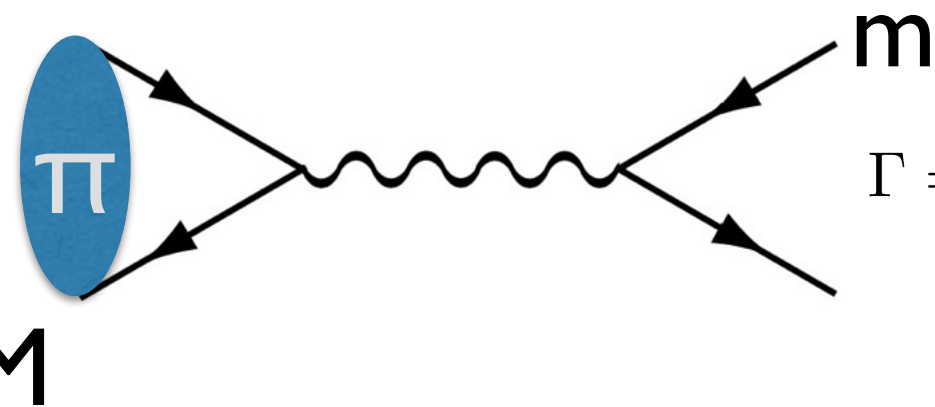
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... all occasionally corrected by CKM factors, isospin violation, etc.etc.

Lifetimes and detector technology: a tale of *anthropic** coincidences in establishing the SM

* something is “anthropic” if, were it not the way it is, we might not be here discussing...

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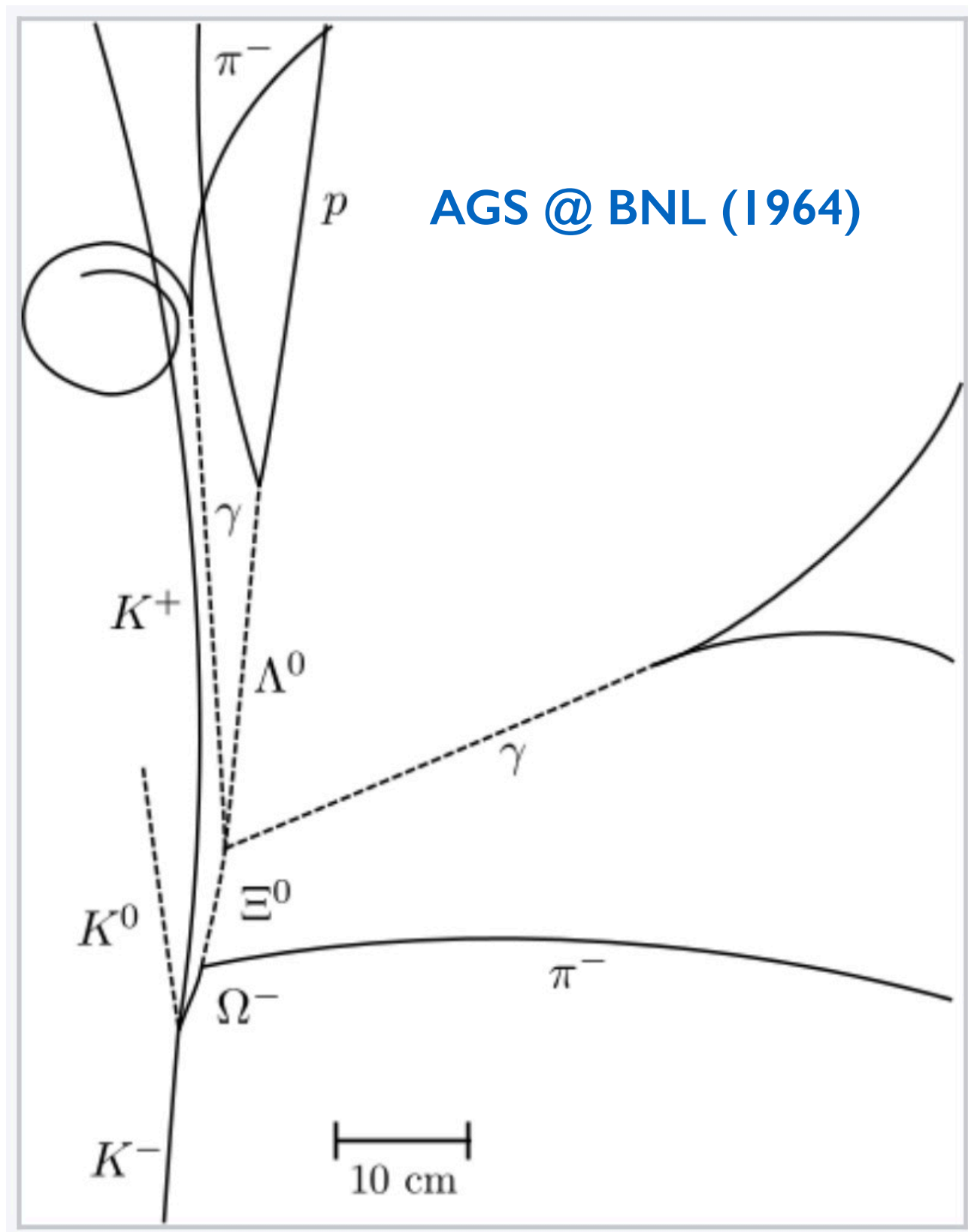
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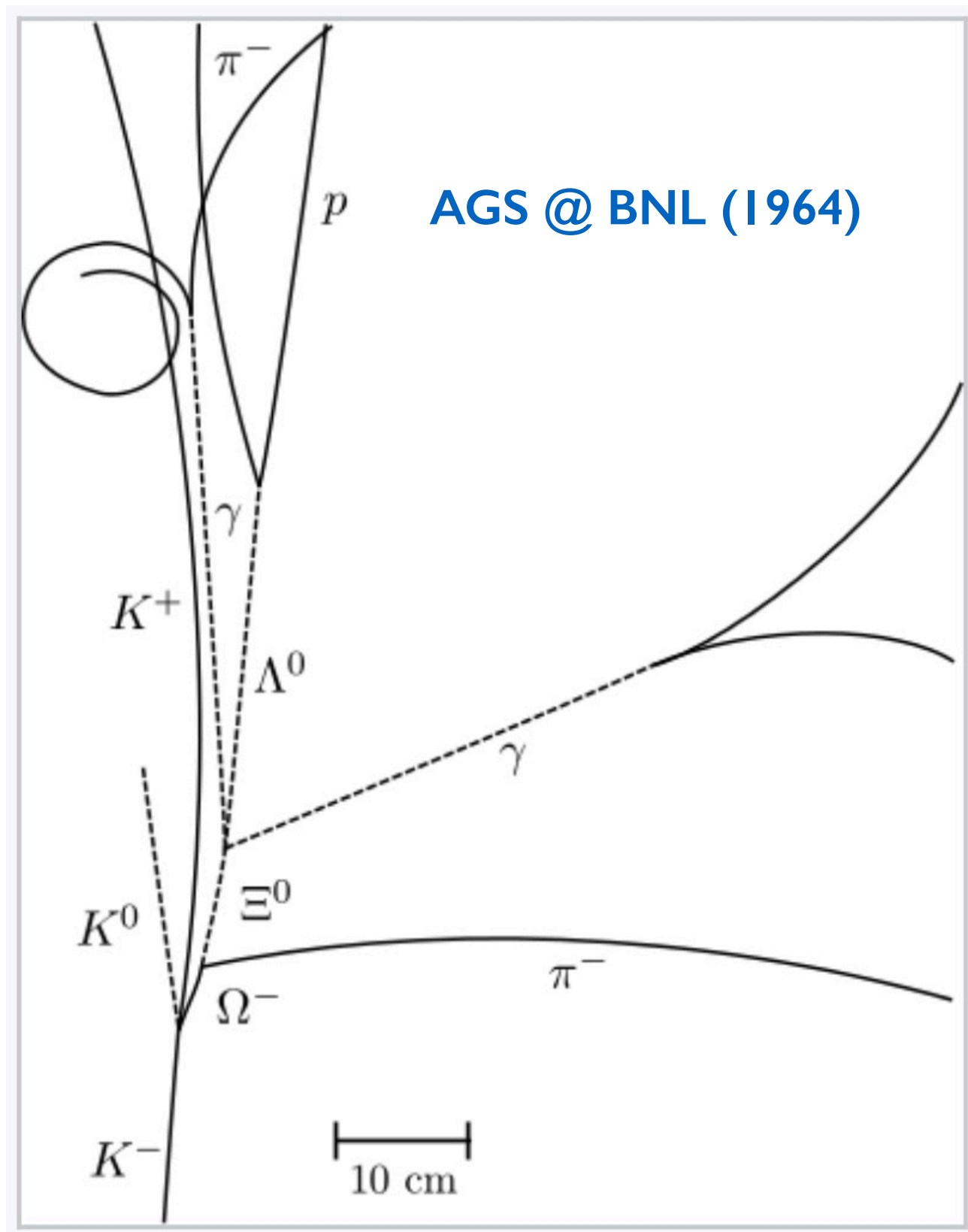
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 - first evidence for P violation
- for $E \sim \text{few GeV}$, $\tau(K^+) \sim \text{few } 10^8 \text{ m}$
 - make kaon beams!
 - study $S=2,3$ hadrons
 - establish $SU(3)_F$ and quark model!

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Hyperon	Quark Content	Decay Modes	Lifetime
Λ	uds	$p\pi^-, n\pi^0$	$0.26ns$
Σ^+	uus	$p\pi^0, n\pi^+$	$0.80ns$
Σ^0	uds	$\Lambda\gamma$	$7 \times 10^{-20}s$
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NB: The lifetime of neutral K's has also been critical to allow the discovery and measurements of CP violation (beams, regeneration,)

Accidental stability of lifetimes across flavours

$$\frac{\Gamma(\pi \rightarrow \mu\nu)}{\Gamma(K \rightarrow \mu\nu)} \sim \frac{\overset{\sim 0.7}{f_\pi^2} \overset{\sim 0.3}{m_\pi} \overset{\sim 0.2}{[1 - (m_\mu/m_\pi)^2]^2} \overset{\sim 20}{\cos^2 \theta_C}}{\overset{\sim 0.7}{f_K^2} \overset{\sim 0.3}{m_K} \overset{\sim 0.2}{[1 - (m_\mu/m_K)^2]^2} \overset{\sim 20}{\sin^2 \theta_C}} \sim 1 \quad \Rightarrow \langle \tau_K \rangle \sim 0.5 \langle \tau_\pi \rangle$$

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Through the yrs, nature provided us lifetimes adapted to the detector technologies we had available and could afford!

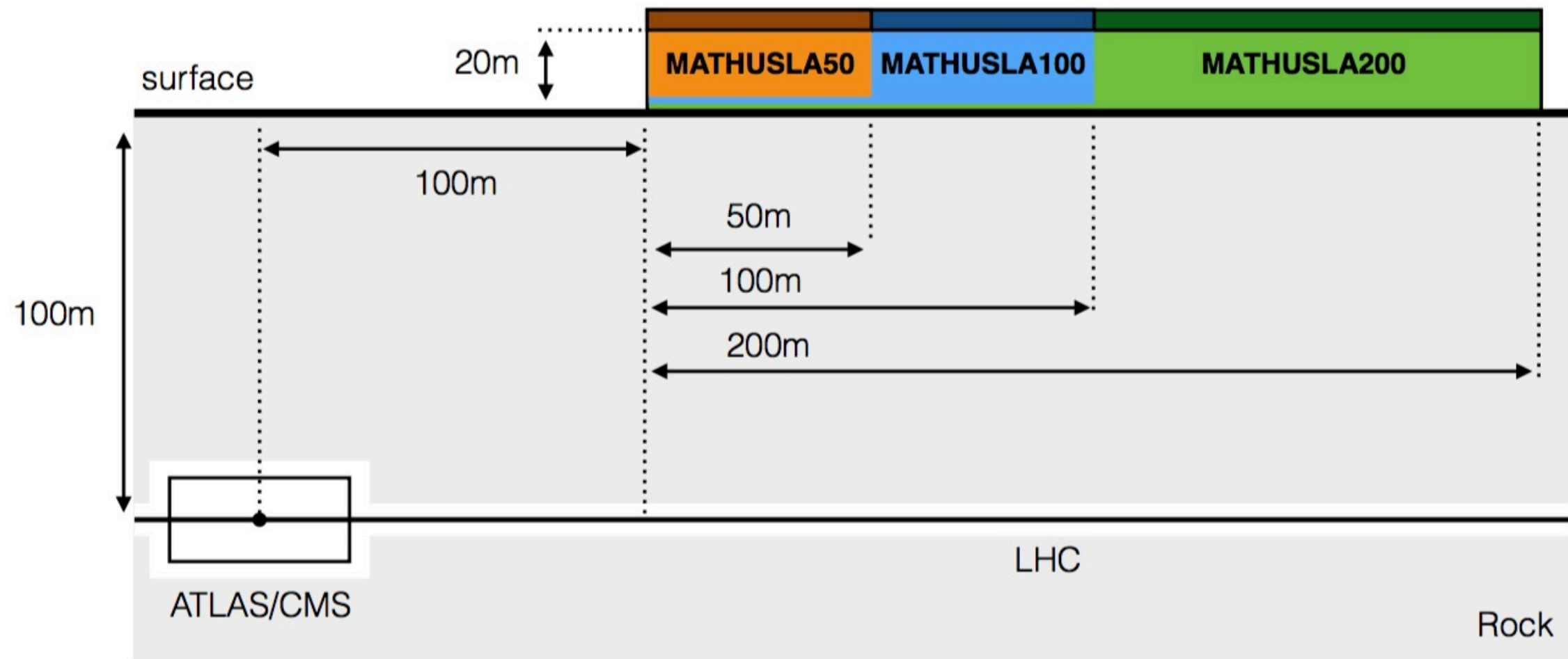
MATHUSLA

MAssive Timing Hodoscope
for ULtra-Stable Neutral PArticles

 tracking system
 air-filled LLP decay volume

Decay Volume:

MATHUSLA50: 50m x 50m x 20m
MATHUSLA100: 100m x 100m x 20m
MATHUSLA200: 200m x 200m x 20m



Will nature be kind to us, once more, in establishing the BSM ??

from my standard FCC motivational talk:

The physics potential (the “case”) of a future facility for HEP should be weighed against criteria such as:

(1) the **guaranteed deliverables**:

- knowledge that will be acquired independently of possible discoveries (*the value of “measurements”*)

(2) the **exploration potential**:

- target broad and well justified BSM scenarios *but guarantee sensitivity to more exotic options*
- exploit both direct (large Q^2) and indirect (precision) probes

(3) the potential to provide conclusive **yes/no** answers to relevant, broad questions.

Mathusla does not pretend to be a facility, and doesn't need to optimally fulfill these criteria. But proving it does, helps build a case in view of the competition for Beyond-colliders projects at CERN (SHIP, ...)

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not sure we can expect this

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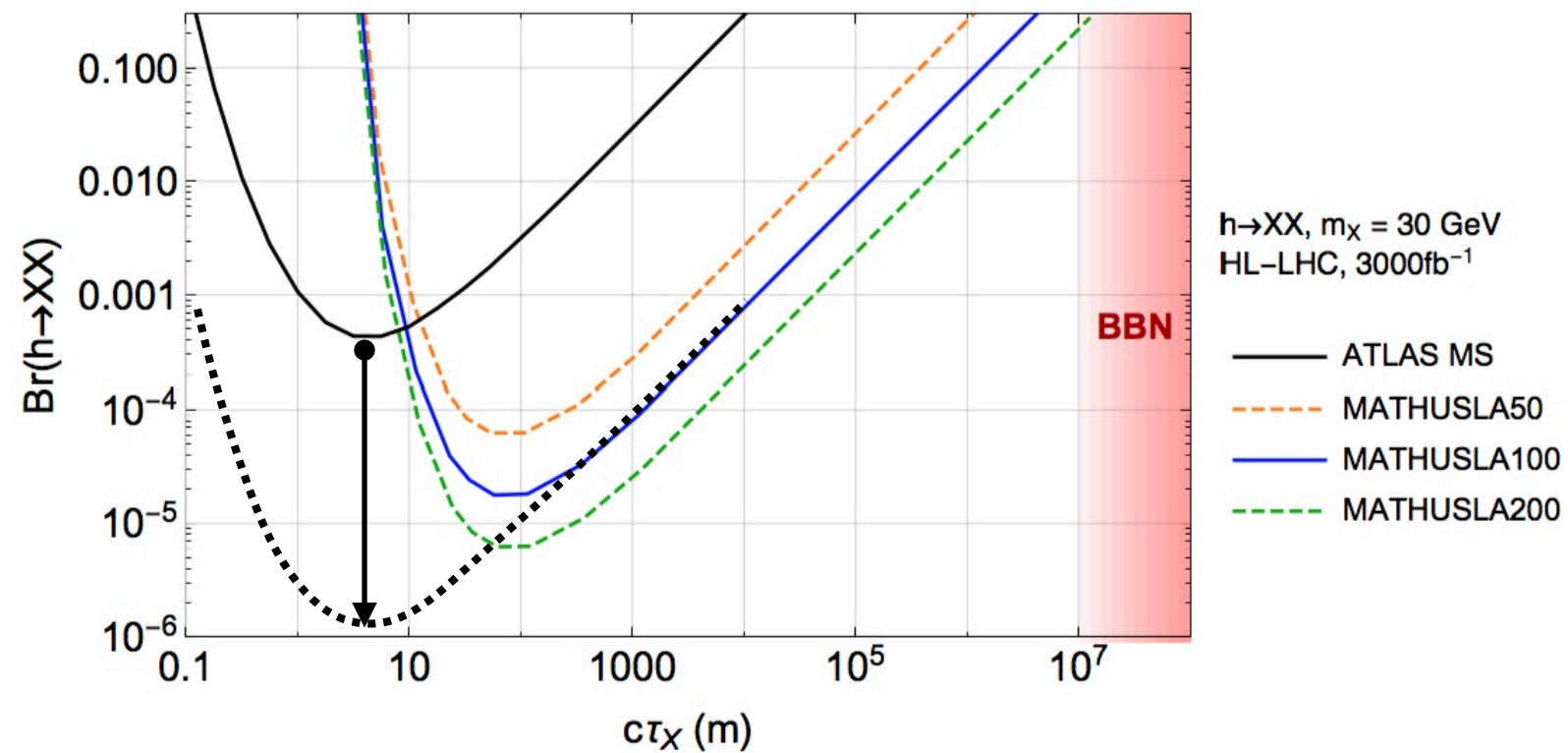
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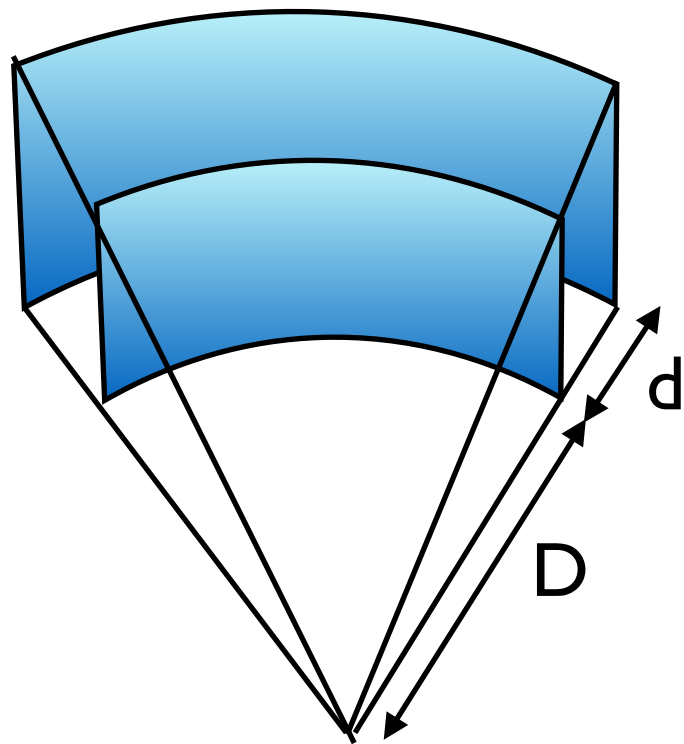
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- It would be useful to provide explicit examples of complementarity with ATLAS/CMS discovery of MET signals: under which conditions, for which class of models, covering which part of parameter space, will MATHUSLA contribute to the exploration of these MET sources?

Comparison with ATLAS, a remark



For $c\tau \gg D, d$ $N_{\text{ev}} \sim d/c\tau$, indep of D

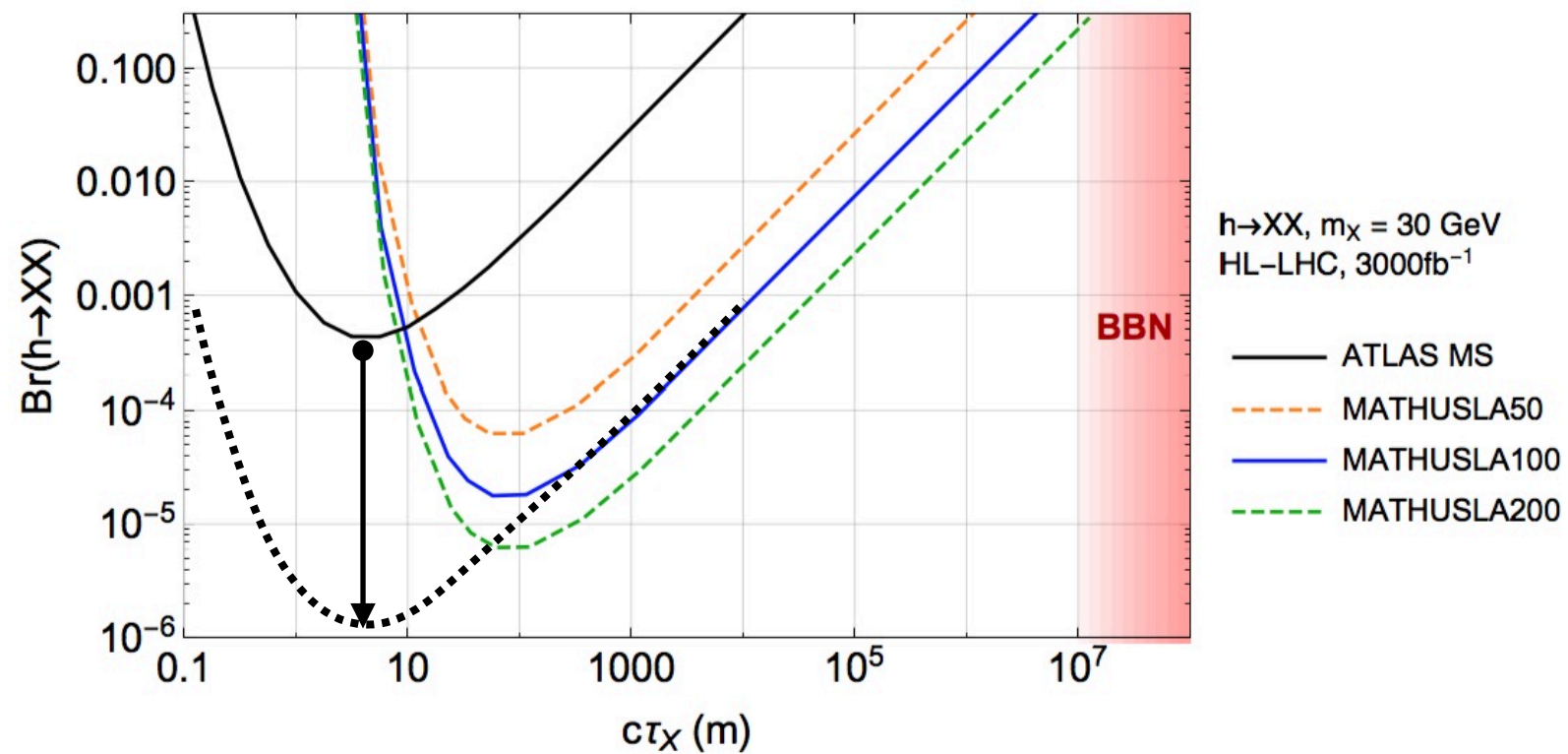


$$N_{\text{ATLAS}} \sim \Omega_{\text{ATLAS}} d_{\text{ATLAS}} \epsilon_{\text{ATLAS}}$$

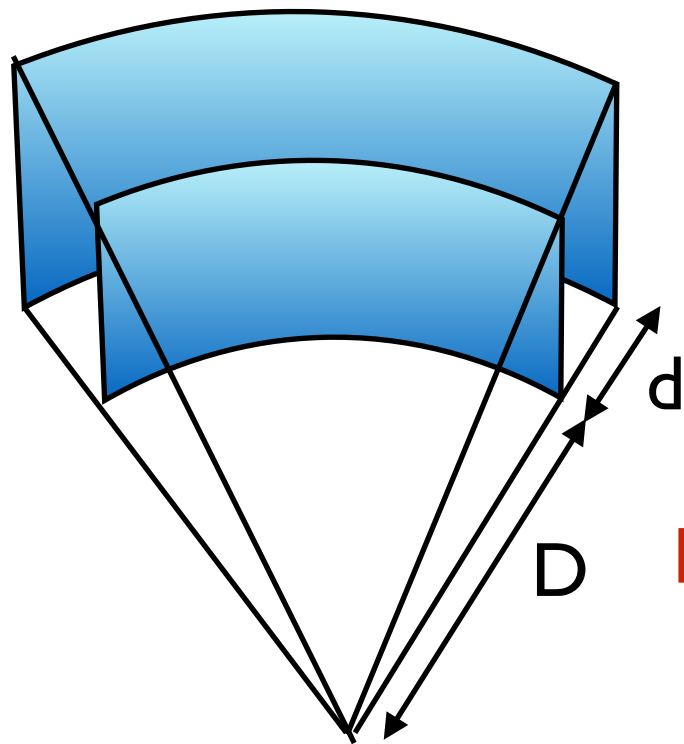
$$N_{\text{MATH}} \sim \Omega_{\text{MATH}} d_{\text{MATH}} \epsilon_{\text{MATH}} \sim I$$

solid angle

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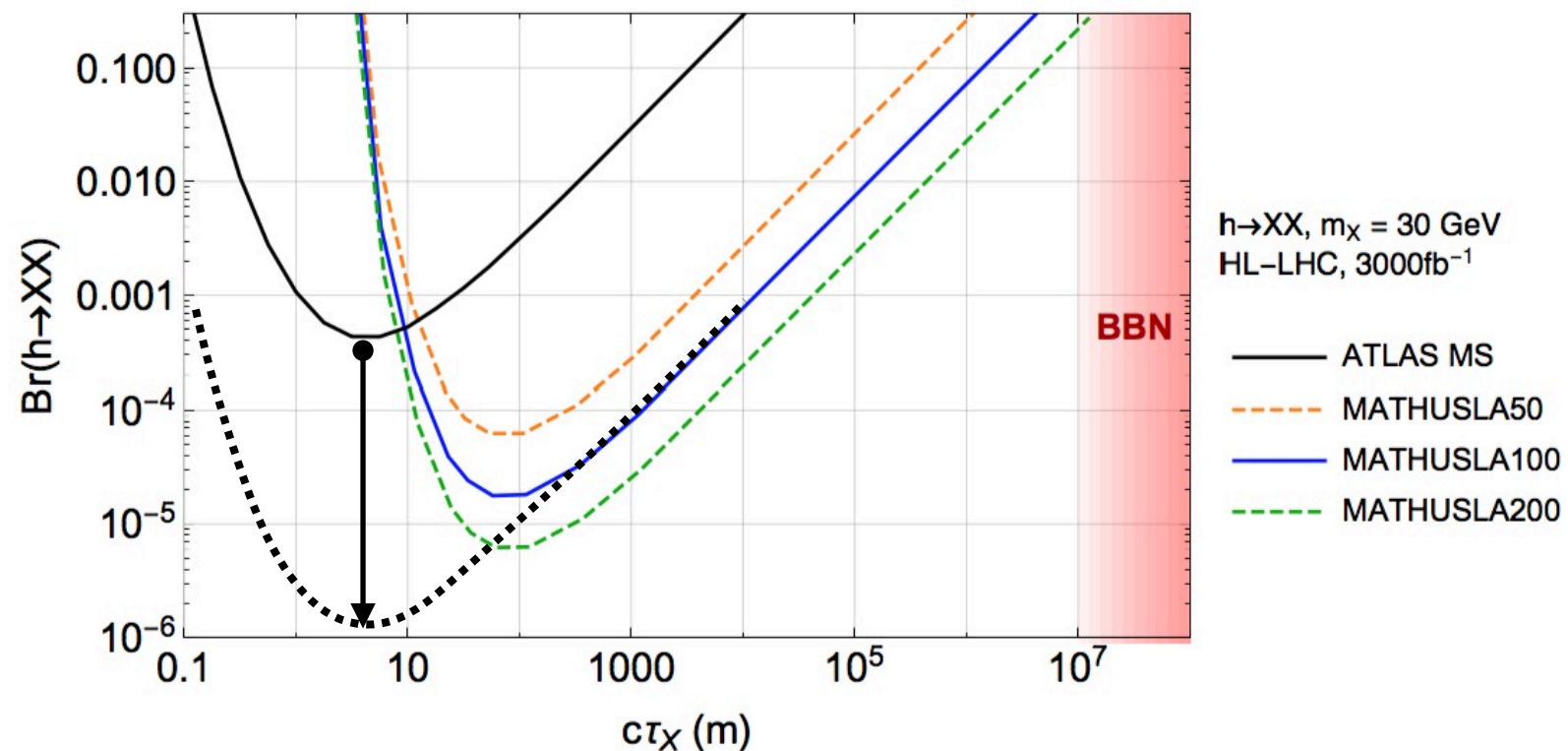
$$N_{MATH} \sim \Omega_{MATH} d_{MATH} \epsilon_{MATH} \sim l$$

$$N_{ATL} / N_{MATH} \sim \Omega_{ATLAS} / \Omega_{ATL} * d_{ATL} / d_{MATH} * \epsilon_{ATL}$$

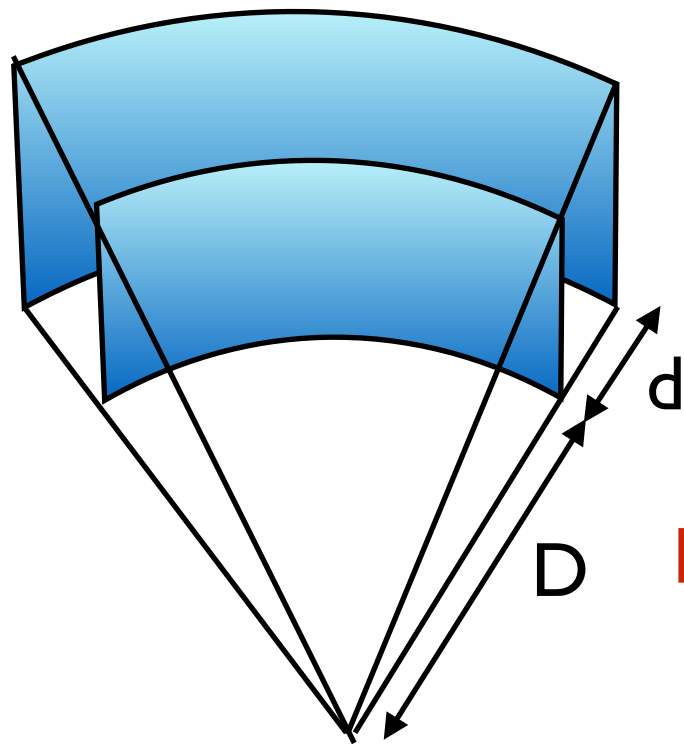
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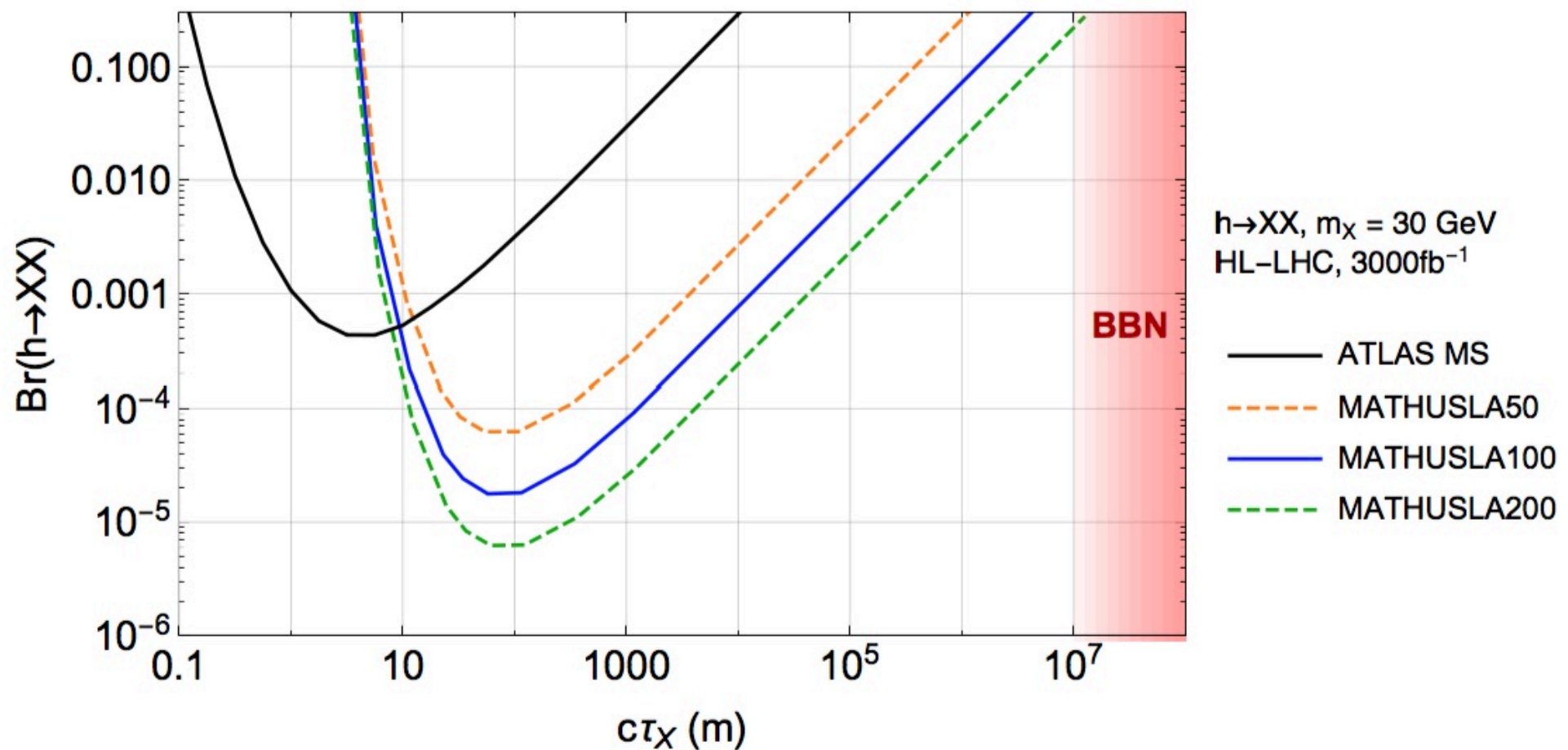
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$$\sim \epsilon_{ATL}$$

=> it is signal efficiency, not geometry, that handicaps ATLAS

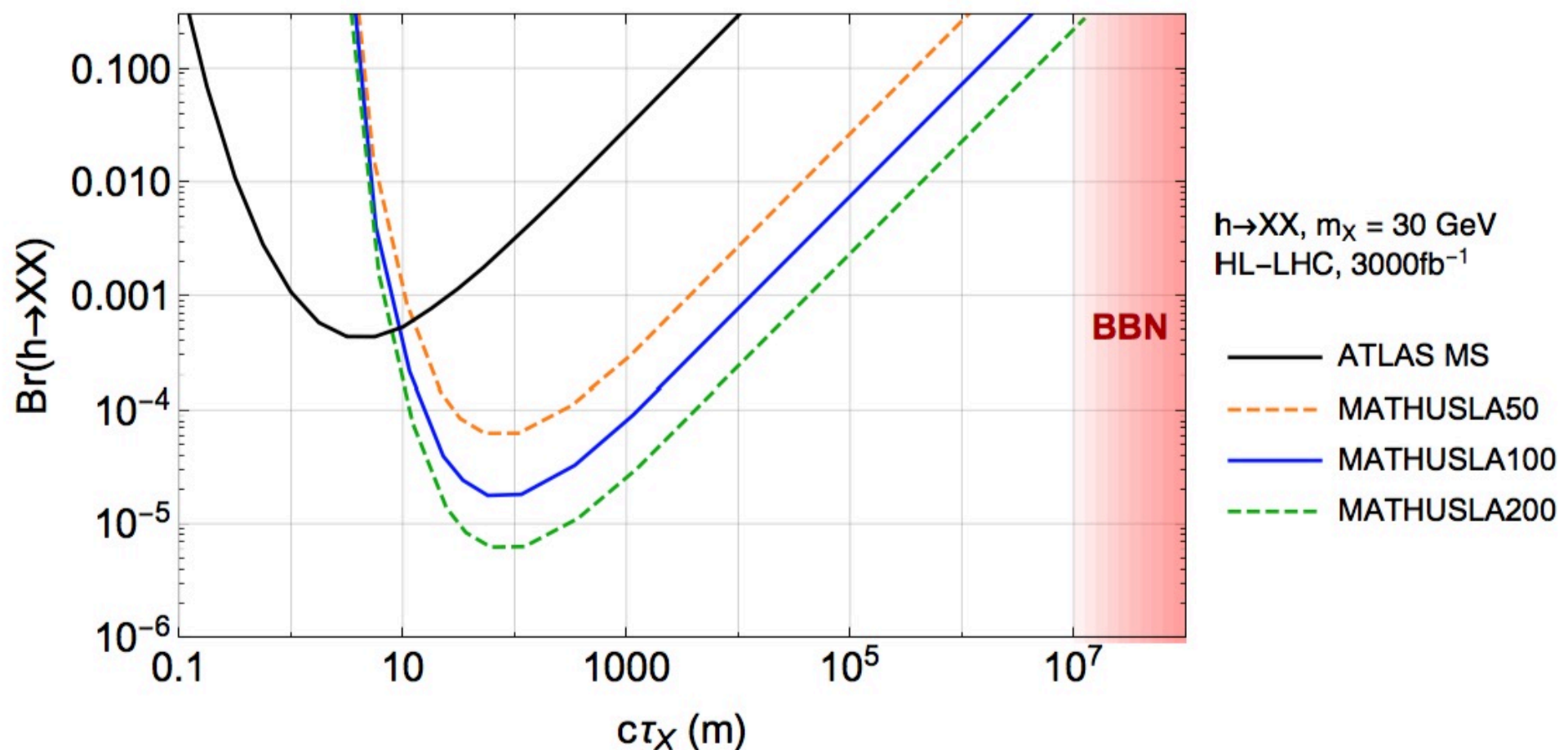
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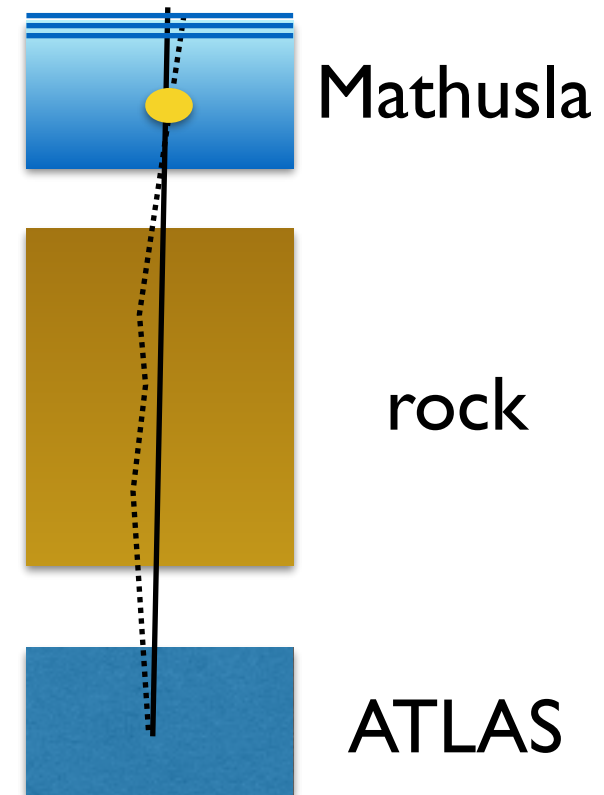


reaching the BBN limit is a good example of a relevant target for definitive confirmation/exclusion. Anything else?

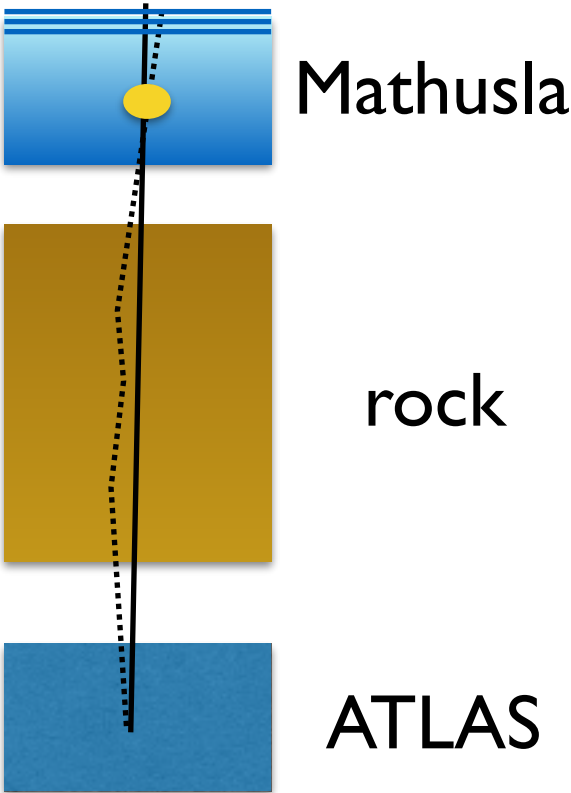
**Some general questions/remarks emerged
in the first discussion among LHCC referees**

backgrounds

Multiple scattering in the rock for low-mass dimuon pairs ?



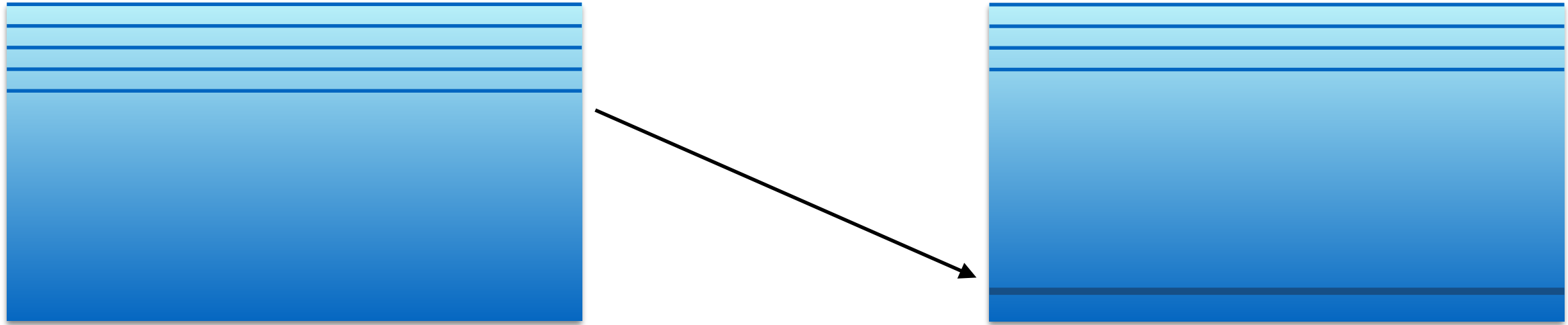
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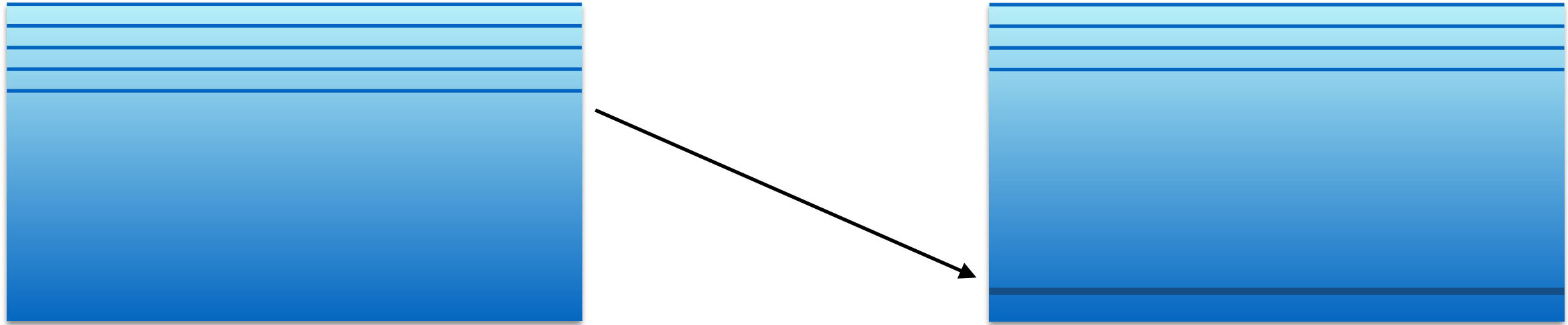
	$p_T > 100$	$p_T > 200$	$p_T > 300$
$b \rightarrow J/\psi \rightarrow \mu\mu$	$2.5 \cdot 10^6$	$8 \cdot 10^4$	$4 \cdot 10^3$
direct $J/\psi \rightarrow \mu\mu$?	?	?
$DY(\mu\mu), 5 < m(\text{GeV}) < 30$	$2.5 \cdot 10^6$	$3 \cdot 10^5$	$6 \cdot 10^4$
$b \rightarrow \mu$	$3.5 \cdot 10^7$	10^6	$5 \cdot 10^4$
$W \rightarrow \mu$	$1.5 \cdot 10^8$	$1.1 \cdot 10^7$	$1.8 \cdot 10^6$

p_T min for μ or $\mu\mu$

N_{events} in 3ab^{-1} at the IP ($|\eta| < 2.5$)

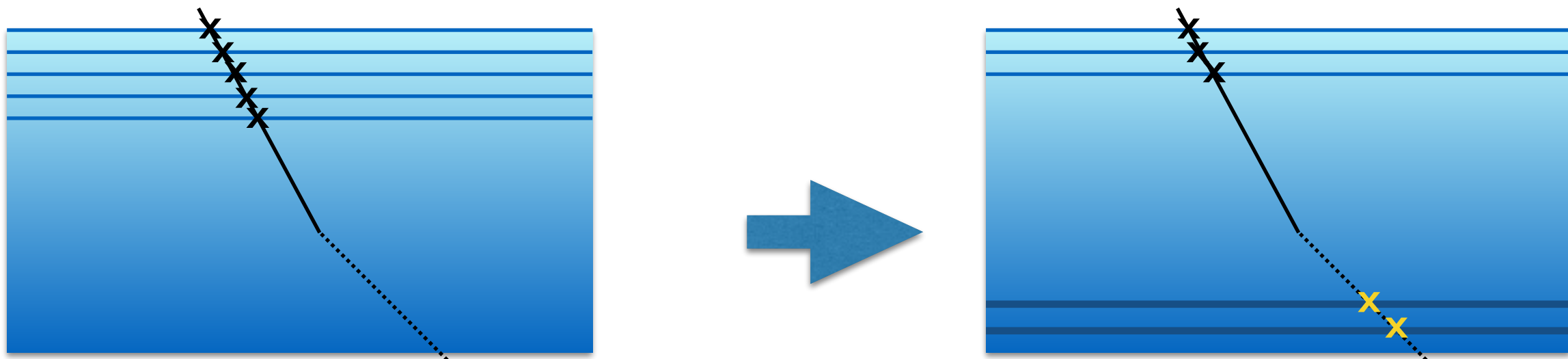


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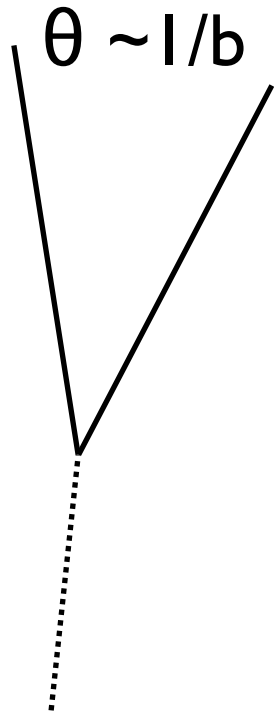


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It could also give higher acceptance to slow, heavy-LLP decays



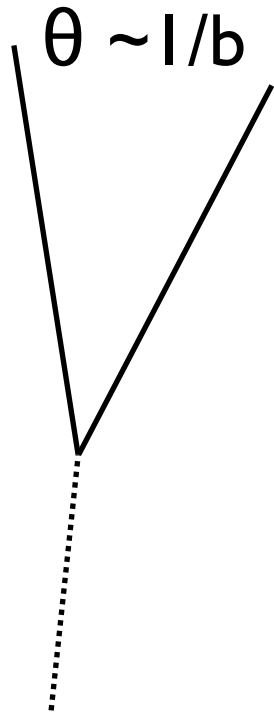
$\mu + \delta$ rays



$$\Delta \sim \theta L > \Delta x \Rightarrow \theta > 10^{-3} (\Delta x / \text{cm}) \Rightarrow b < b_{\text{max}} \sim 10^3$$

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How are the estimates of reach in mass and boost affected by the $\theta > 2^\circ$ cut required to eliminate $\mu+\delta$ -ray bgs??

Does this limit some component of the physics reach?

Baseline and upgrade scenarios

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- Define scenarios in which a technology upgrade could be more interesting than a volume upgrade (MAT100 => 200)
- Allow ambition in the plan: how much more could be gained if more resources were available?

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Other misc questions

- Test stand: when do you expect to have the first results on bg's etc?
- Are there studies ongoing on the environmental impact? In particular, how do you see the gas mixture studies for the full fledged detector, in order to avoid large consumption of greenhouse gases. 75% of CFC, with a GWP=1400, is likely not viable?
- Need a more clear itemization of costs, including electronics, gas systems, CE, etcetcetc
- Funding strategy:
 - How do you plan building the collaboration? Currently most members are penniless theorists! Resources for LHC experiments are pretty much capped already by funding agencies => engage new groups and communities (eg CRs) not currently committed to LHC programme?

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- A lot of work still remains to be done, and we'll be happy to follow and encourage your progress