

Heavy neutrinos at NA62

based on JHEP 1807 (2018) 105

Jan Hajer

Centre for Cosmology, Particle Physics and Phenomenology — Université catholique de Louvain
in collaboration with Marco Drewes, Juraj Klaric, Gaia Lanfranchi

XXVII International Workshop on Deep Inelastic Scattering and Related Subjects

Disclaimer

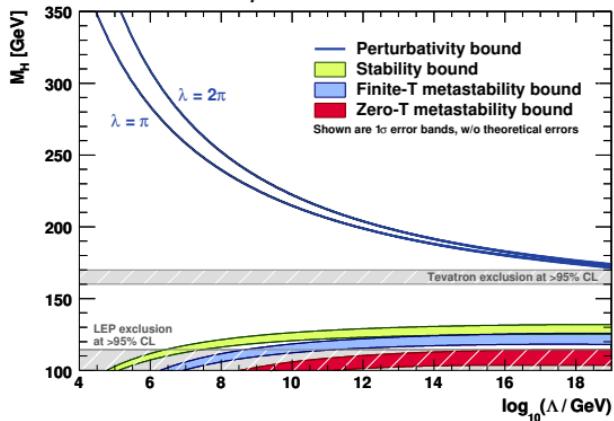
- ▶ I am not an experimentalist.
- ▶ I do not speak on behalf of the NA62 collaboration.
- ▶ Nonetheless, we have published estimates on heavy neutrinos at NA62.
- ▶ One member of our collaboration is an experimentalist at NA62.

Hierarchy problem

Higgs mass

[Ellis et al. 2009]

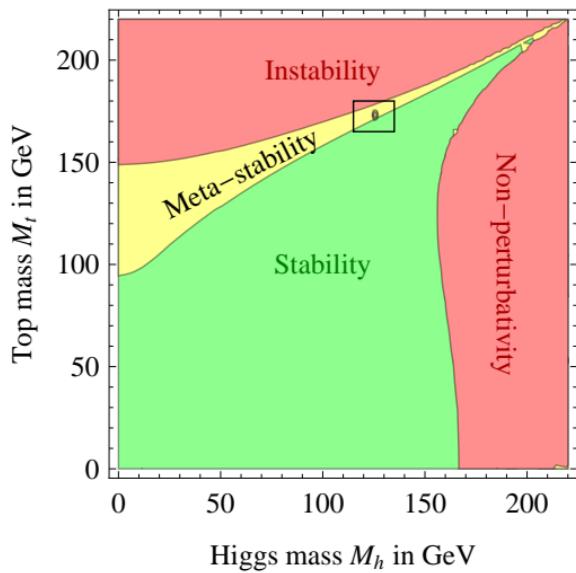
Only in a small window of m_h can the SM be extended to the M_P



Higgs-top conspiracy

[Degrassi et al. 2012]

Their mass values happen to put us into a meta-stable vacuum



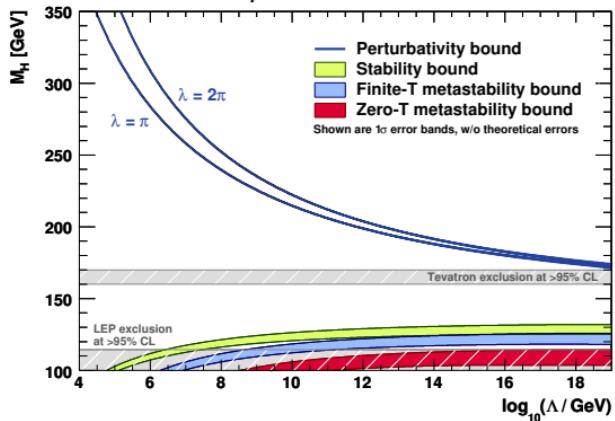
The SM is technically natural up to M_P

Hierarchy problem

Higgs mass

[Ellis et al. 2009]

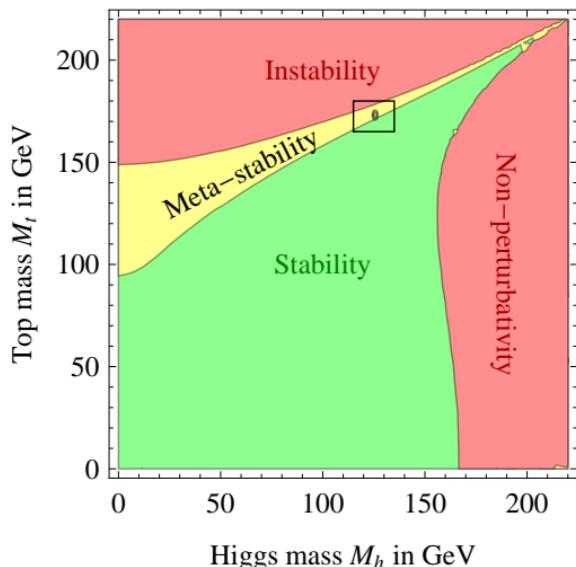
Only in a small window of m_h can the SM be extended to the M_P



Higgs-top conspiracy

[Degrassi et al. 2012]

Their mass values happen to put us into a meta-stable vacuum



New Physics

- ▶ Heavy NP easily destabilizes the Higgs mass
- ▶ NP at the electroweak scales circumvents this problem
- ▶ must be coupled feebly enough to have evaded detection

The SM is technically natural up to M_P

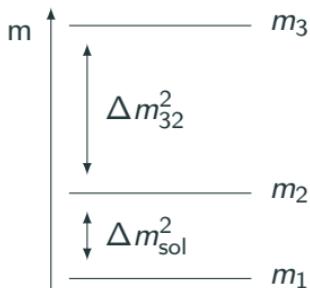
Relatively light and feebly coupled NP leads to long lived particles

Neutrino masses

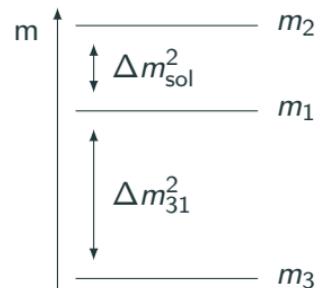
Masses differences

- ▶ There are two measured mass differences for potentially three light neutrino masses m_1, m_2, m_3
- ▶ one neutrino may still be massless $m_{\text{lightest}} = 0$
- ▶ The smaller “solar” mass difference is given by $\Delta m_{\text{sol}}^2 = m_2^2 - m_1^2$

Normal ordering



Inverted ordering



Parameter

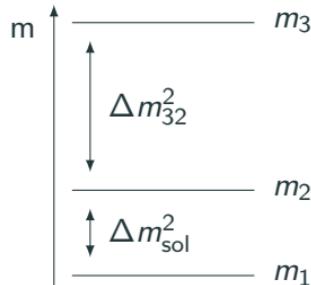
Variables	Normal Ordering	Inverted Ordering
m_1^2	m_{lightest}^2	$m_{\text{lightest}}^2 - \Delta m_{32}^2 - \Delta m_{\text{sol}}^2$
m_2^2	$m_{\text{lightest}}^2 + \Delta m_{\text{sol}}^2$	$m_{\text{lightest}}^2 - \Delta m_{32}^2$
m_3^2	$m_{\text{lightest}}^2 + \Delta m_{31}^2$	m_{lightest}^2
larger Δm^2	$\Delta m_{31}^2 = m_3^2 - m_1^2$	$\Delta m_{32}^2 = m_3^2 - m_2^2$

Neutrino masses

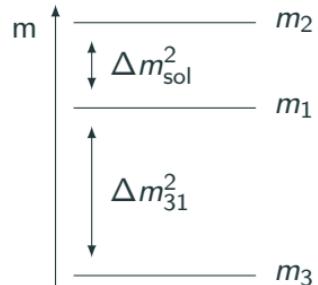
Masses differences

- ▶ There are two measured mass differences for potentially three light neutrino masses m_1, m_2, m_3
- ▶ one neutrino may still be massless $m_{\text{lightest}} = 0$
- ▶ The smaller “solar” mass difference is given by $\Delta m_{\text{sol}}^2 = m_2^2 - m_1^2$

Normal ordering



Inverted ordering



Best fit values from the NuFIT 3.1 release by the ν -fit collaboration

Variables	Normal Ordering	Inverted Ordering
(smaller) Δm_{sol}^2	$7.40 \times 10^{-5} \text{ eV}^2$	$7.40 \times 10^{-5} \text{ eV}^2$
larger Δm^2	$2.515 \times 10^{-3} \text{ eV}^2$	$-2.483 \times 10^{-3} \text{ eV}^2$
$\sin^2 \theta_{12}$	0.307	0.307
$\sin^2 \theta_{13}$	0.02195	0.02212
$\sin^2 \theta_{23}$	0.565	0.572

Heavy Neutrinos

Low scale seesaw (type I) [Minkowski 1977; Mohapatra et al. 1980; Yanagida 1980; Schechter et al. 1980]

Three right handed neutrinos

$$\mathcal{L}_{\nu_R} = -\frac{1}{2}\bar{\nu_R^c} M_{ij} \nu_{Rj} - y_{ai} \bar{\ell}_a \varepsilon \phi \nu_{Ri} + \text{h.c.}$$

M_{ij} Majorana mass;

y_{ai} Yukawa coupling

Electroweak symmetry breaking

Dirac mass: $m_{ai} = v y_{ai}$

Seesaw mechanism

$m_\nu = -m_{ai} M_{ij}^{-1} m_{bj}^T = -\theta_{ai} M_{ij} \theta_{bj}^T$, $\theta_{ai} = m_{aj} M_{ij}^{-1}$, produces tiny masses for the left handed neutrinos

Small mixing into mass eigenstates

$\nu \simeq U_\nu^\dagger (\nu_L - \theta \nu_R^c)$, $N \simeq \nu_R + \theta^T \nu_L^c$, the PNMS matrix U_ν diagonalises the m_ν

The neutrino is special

Three Generations of Matter (Fermions) spin $\frac{1}{2}$		
I	II	III
mass \rightarrow charge \rightarrow name \rightarrow		
Left $\frac{2}{3}$ u up	Right $\frac{2}{3}$ c charm	Right $\frac{2}{3}$ t top
Left $\frac{-1}{3}$ d down	Right $\frac{-1}{3}$ s strange	Right $\frac{-1}{3}$ b bottom
Quarks		
Left 0 eV ν_e electron neutrino	Left 0 eV ν_μ muon neutrino	Left 0 eV ν_τ tau neutrino
Leptons		
Left 0.511 MeV e^- electron	Left 105.7 MeV μ^- muon	Left 1.777 GeV τ^- tau

May explain

- ▶ Neutrino oscillation data
- ▶ Neutrino masses
- ▶ Baryogenesis via Leptogenesis
- ▶ Dark matter

Low scale seesaw (type I) [Minkowski 1977; Mohapatra et al. 1980; Yanagida 1980; Schechter et al. 1980]

Three right handed neutrinos

$$\mathcal{L}_{\nu_R} = -\frac{1}{2}\bar{\nu}_{Ri}^c M_{ij} \nu_{Rj} - y_{ai} \bar{\ell}_a \epsilon \phi \nu_{Ri} + \text{h.c.}$$

M_{ij} Majorana mass;

y_{ai} Yukawa coupling

Electroweak symmetry breaking

Dirac mass: $m_{ai} = v y_{ai}$

Seesaw mechanism

$$m_\nu = -m_{ai} M_{ij}^{-1} m_{bj}^T = -\theta_{ai} M_{ij} \theta_{bj}^T, \quad \theta_{ai} = m_{aj} M_{ij}^{-1},$$

produces tiny masses for the left handed neutrinos

Coupling of N_i to the SM

$$\mathcal{L} \supset -\frac{m_W}{v} \overline{N} \theta_a^* \gamma^\mu e_{La} W_\mu^+ - \frac{m_Z}{\sqrt{2}v} \overline{N} \theta_a^* \gamma^\mu \nu_{La} Z_\mu - \frac{M}{v} \theta_a h \overline{\nu}_{L\alpha} N + \text{h.c.} .$$

The neutrino is special

Three Generations of Matter (Fermions) spin $\frac{1}{2}$			
I	II	III	
mass \rightarrow charge \rightarrow name \rightarrow	2.4 MeV u up	1.27 GeV c charm	171.2 GeV t top
Quarks	Left: $\frac{2/3}{-1/3}$ d down	Left: $\frac{2/3}{-1/3}$ s strange	Left: $\frac{2/3}{-1/3}$ b bottom
Leptons	Left: 0 eV e electron neutrino	Left: 0 eV ν_μ muon neutrino	Left: 0 eV ν_τ tau neutrino
	Right: 0.511 MeV e electron	Right: 105.7 MeV μ muon	Right: 1.777 GeV τ tau

SM is symmetric under $U(1)_{B-L}$

Majorana mass M_{ij} breaks this symmetry

The $B - L$ symmetry is restored

- ▶ in the limit of $M_{ij} \rightarrow 0$
- ▶ if ν_{Ri} form pseudo Dirac pairs $\nu_{Ri} + \nu_{Rj}^c$

SM is symmetric under $U(1)_{B-L}$

Majorana mass M_{ij} breaks this symmetry

The $B - L$ symmetry is restored

- ▶ in the limit of $M_{ij} \rightarrow 0$
- ▶ if ν_{Ri} form pseudo Dirac pairs $\nu_{Ri} + \nu_{Rj}^c$

Mass matrix

$$M_{ij} = M \begin{pmatrix} 1 - \mu & 0 & 0 \\ 0 & 1 + \mu & 0 \\ 0 & 0 & \mu' \end{pmatrix}$$

Yukawa coupling

$$y_{ai} = \begin{pmatrix} y_e + \epsilon_e & i(y_e - \epsilon_e) & \epsilon'_e \\ y_\mu + \epsilon_\mu & i(y_\mu - \epsilon_\mu) & \epsilon'_\mu \\ y_\tau + \epsilon_\tau & i(y_\tau - \epsilon_\tau) & \epsilon'_\tau \end{pmatrix}$$

$B - L$ violating parameter

$\epsilon, \epsilon', \mu, \mu'$ are small

- ▶ Almost mass degenerate pseudo dirac pair
- ▶ lighter $\mathcal{O}(\text{keV})$ dark matter candidate

- ▶ pseudo Dirac pair with coupling $\mathcal{O}(y)$
- ▶ Dark matter candidate with coupling $\mathcal{O}(\epsilon')$

SM is symmetric under $U(1)_{B-L}$

Majorana mass M_{ij} breaks this symmetry

The $B - L$ symmetry is restored

- ▶ in the limit of $M_{ij} \rightarrow 0$
- ▶ if ν_{Ri} form pseudo Dirac pairs $\nu_{Ri} + \nu_{Rj}^c$

Mass matrix

$$M_{ij} = M \begin{pmatrix} 1 - \mu & 0 & 0 \\ 0 & 1 + \mu & 0 \\ 0 & 0 & \mu' \end{pmatrix}$$

$B - L$ violating parameter

$\epsilon, \epsilon', \mu, \mu'$ are small

- ▶ Almost mass degenerate pseudo dirac pair
- ▶ lighter $\mathcal{O}(\text{keV})$ dark matter candidate

Yukawa coupling

$$y_{ai} = \begin{pmatrix} y_e + \epsilon_e & i(y_e - \epsilon_e) & \epsilon'_e \\ y_\mu + \epsilon_\mu & i(y_\mu - \epsilon_\mu) & \epsilon'_\mu \\ y_\tau + \epsilon_\tau & i(y_\tau - \epsilon_\tau) & \epsilon'_\tau \end{pmatrix}$$

- ▶ pseudo Dirac pair with coupling $\mathcal{O}(y)$
- ▶ Dark matter candidate with coupling $\mathcal{O}(\epsilon')$

Abbreviation

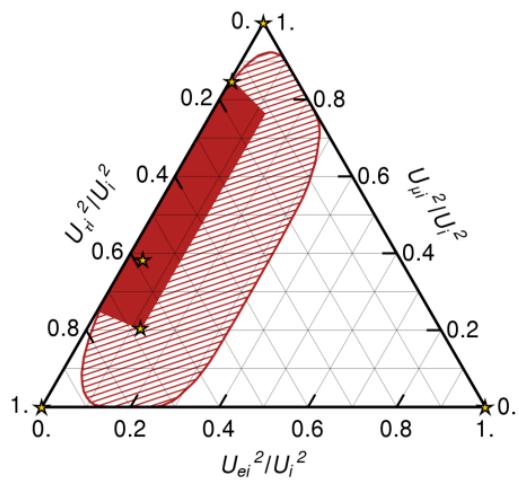
$$U^2 = \sum_a U_a^2, \quad U_a^2 = \sum_i U_{ai}^2, \quad U_{ai}^2 = |\theta_{ai}|^2,$$

The ratio U_a^2 / U_2

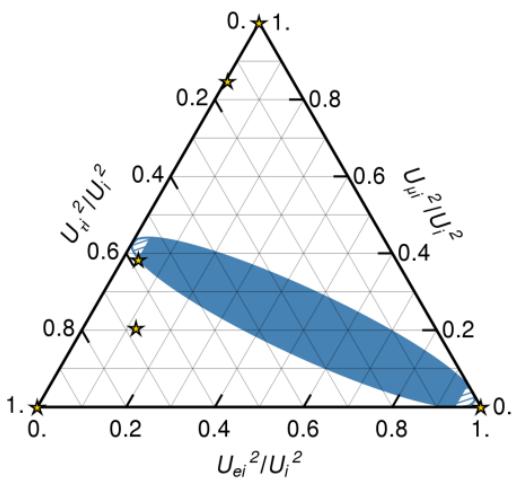
becomes independent of other heavy neutrino parameter

Allowed range of U_a^2/U^2

Normal Ordering



Inverted Ordering



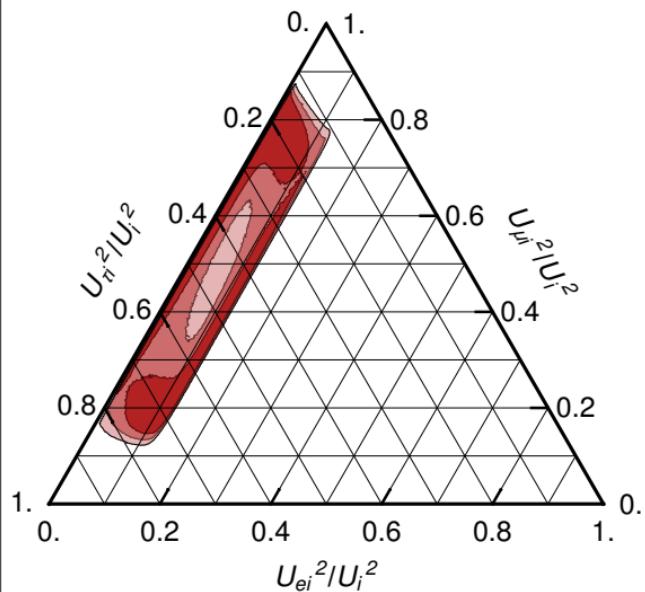
- ▶ for arbitrary parameter choices (hashed region)
- ▶ in the symmetric limit (filled region).
- ▶ The stars mark our benchmark scenarios.

Benchmark scenarios U_a^2/U^2 in %

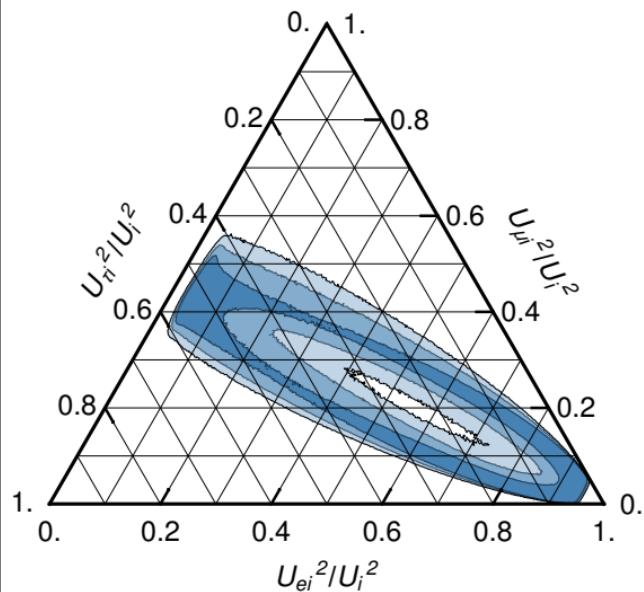
	$U_{ei}^2 : U_\mu^2 : U_\tau^2$
A	0.530 : 84.7 : 14.7
B	12.0 : 20.5 : 67.5
C	3.65 : 38.3 : 58.0
D	100 : 0 : 0
E	0 : 100 : 0
F	0 : 0 : 100

Probability contours for the heavy neutrino couplings

Normal Ordering



Inverted Ordering



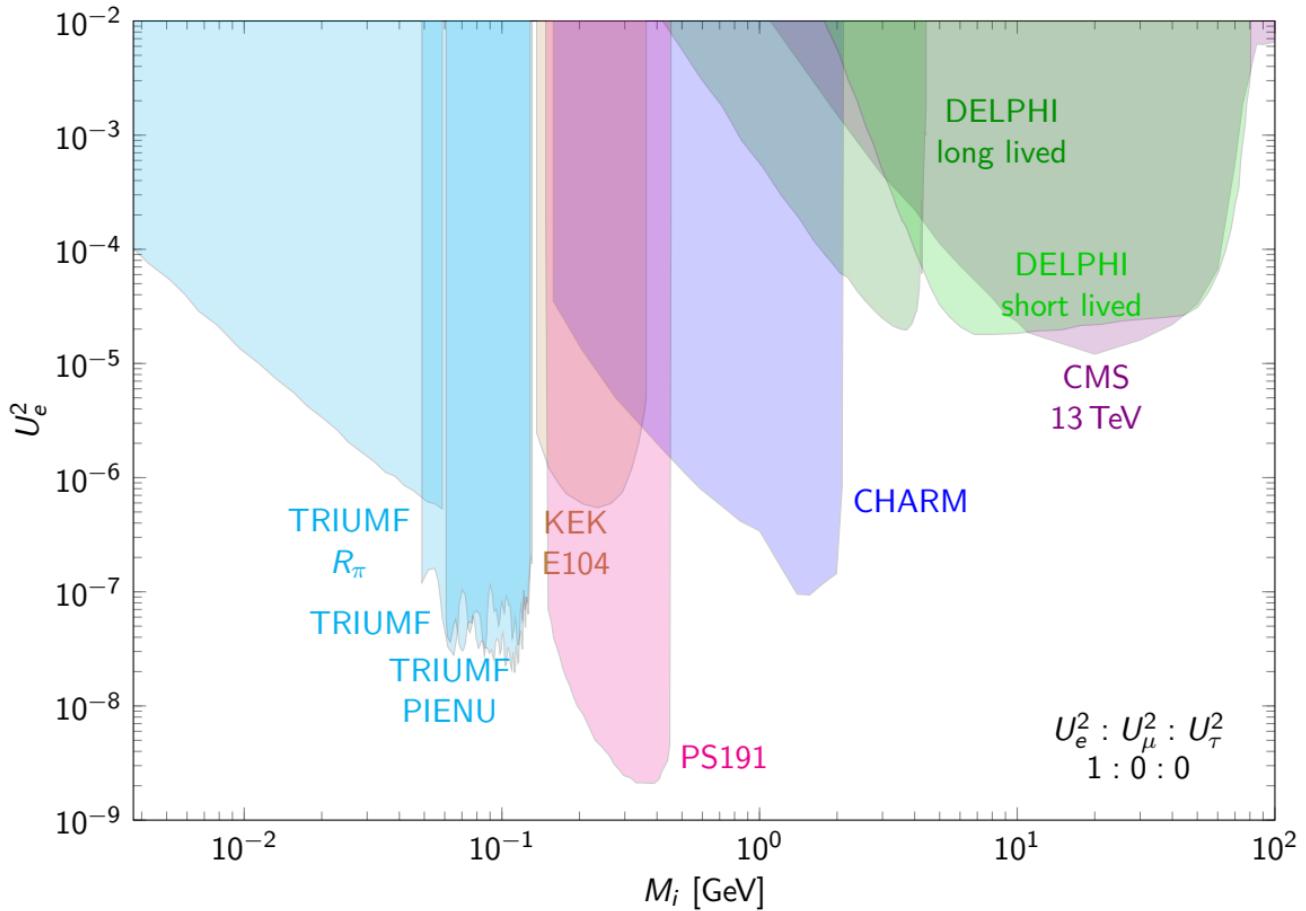
Coloured areas

consistent with neutrino oscillation data at 1σ ,
 2σ , 3σ

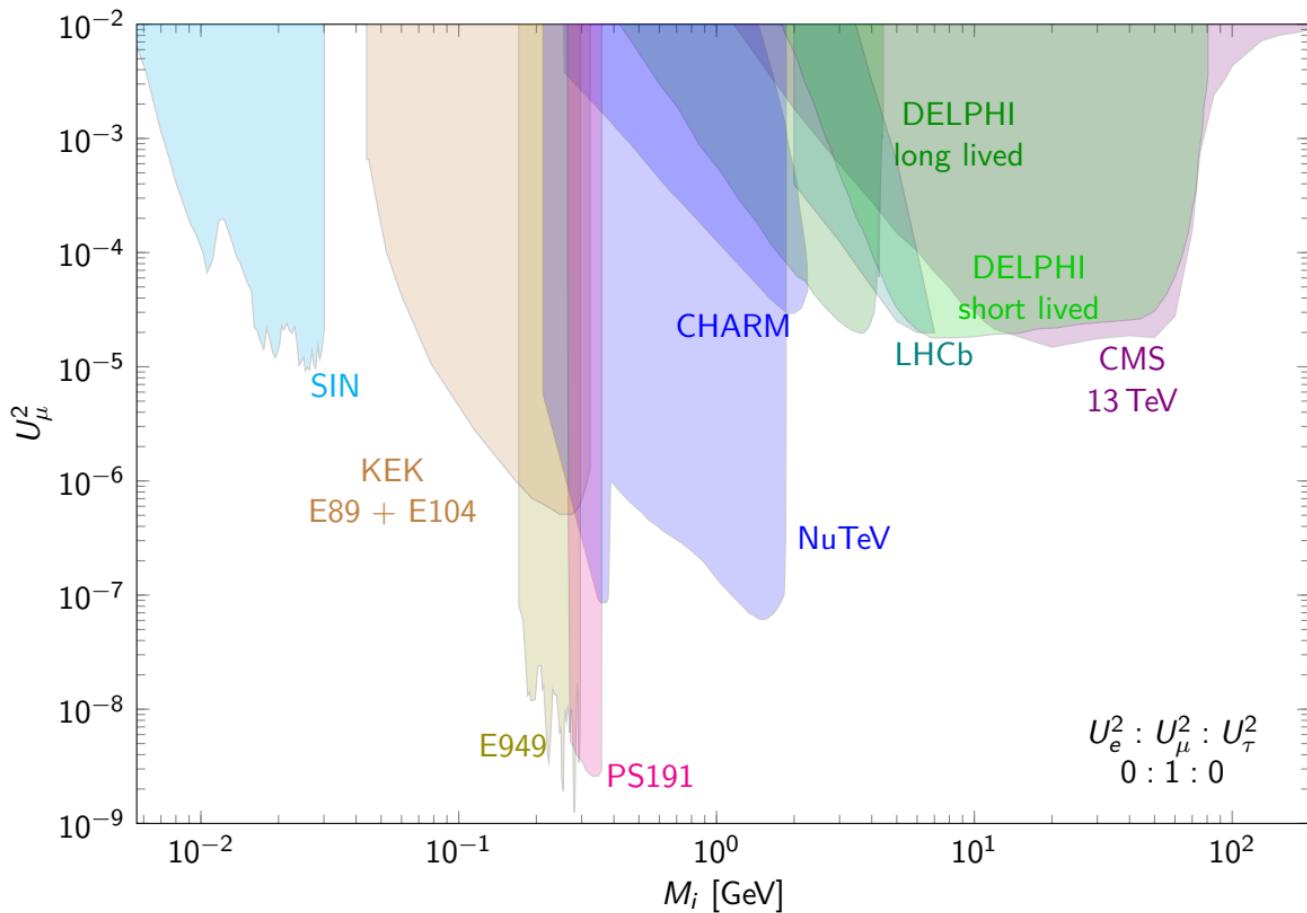
The Majorana phase is unknown

and correspond to the circular structure in the plots

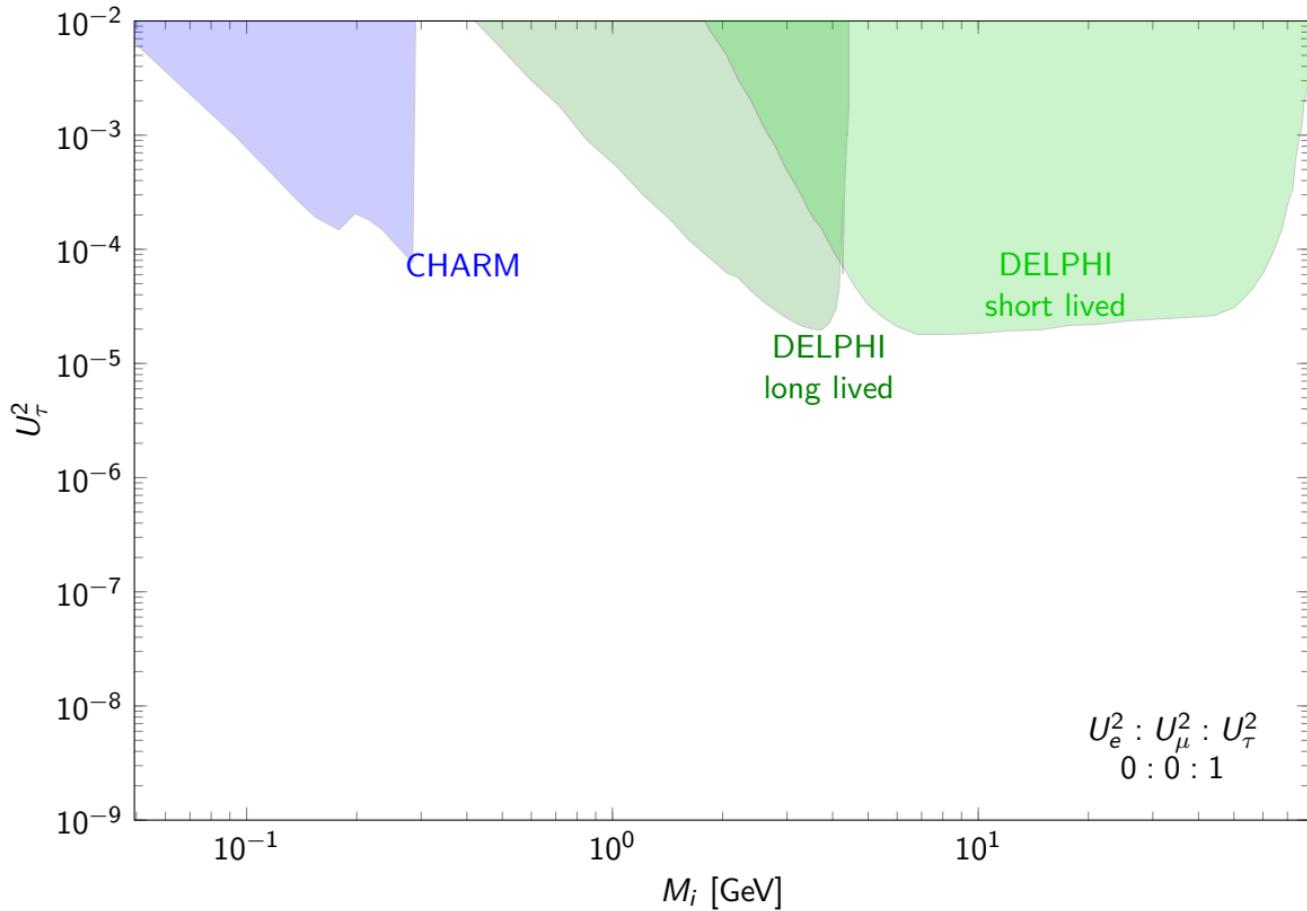
pure U_e^2



pure U_μ^2

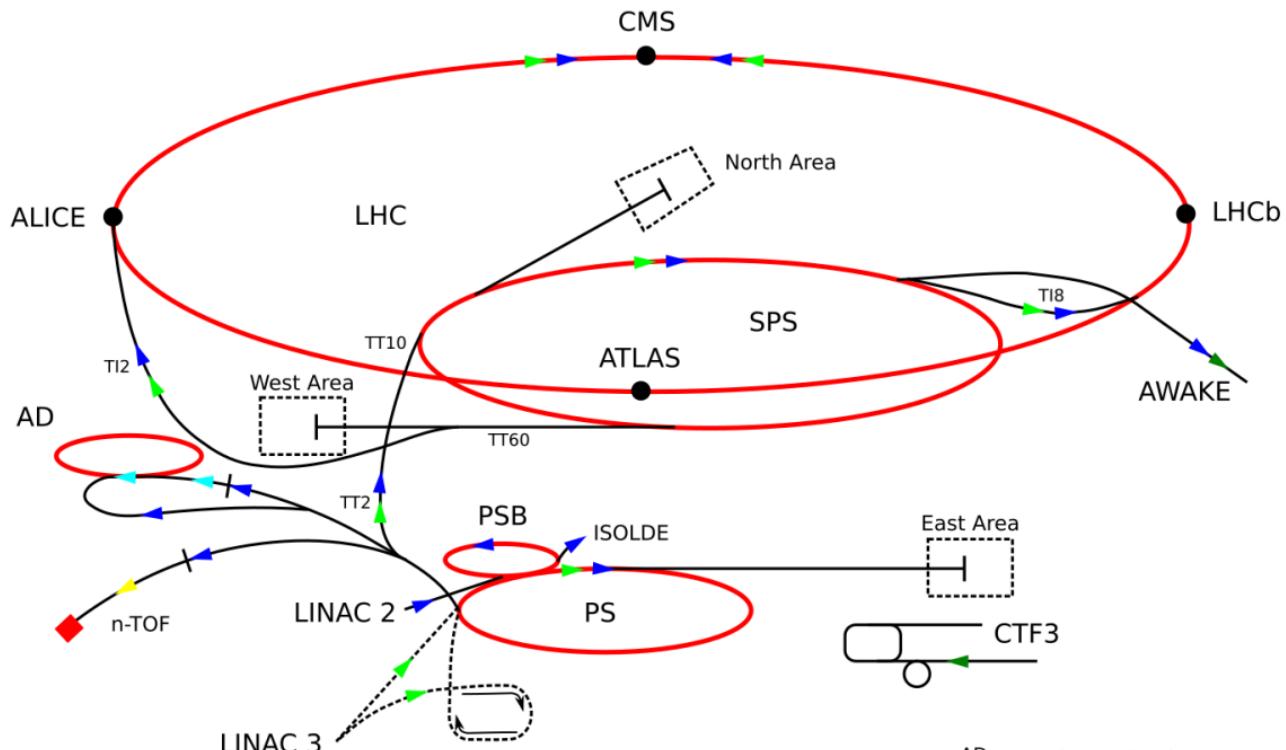


pure U_τ^2



NA62

SPS and North Area at CERN



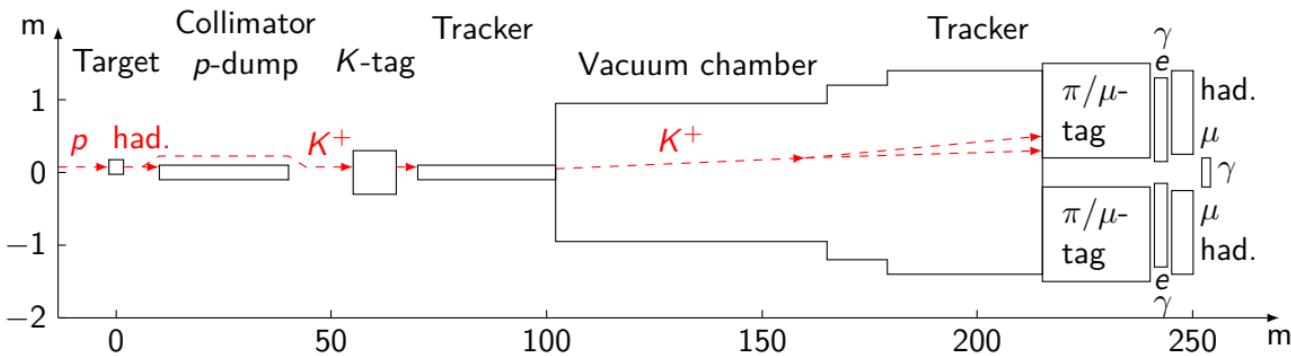
- ▶ protons
- ▶ antiprotons
- ▶ ions
- ▶ electrons
- ▶ neutrons
- ▶ neutrinos

PS Proton Synchrotron
SPS Super Proton Synchrotron
LHC Large Hadron Collider

AD Antiproton Decelerator
n-TOF Neutron Time Of Flight
AWAKE Advanced Wakefield Experiment
CTF3 CLIC Test Facility 3

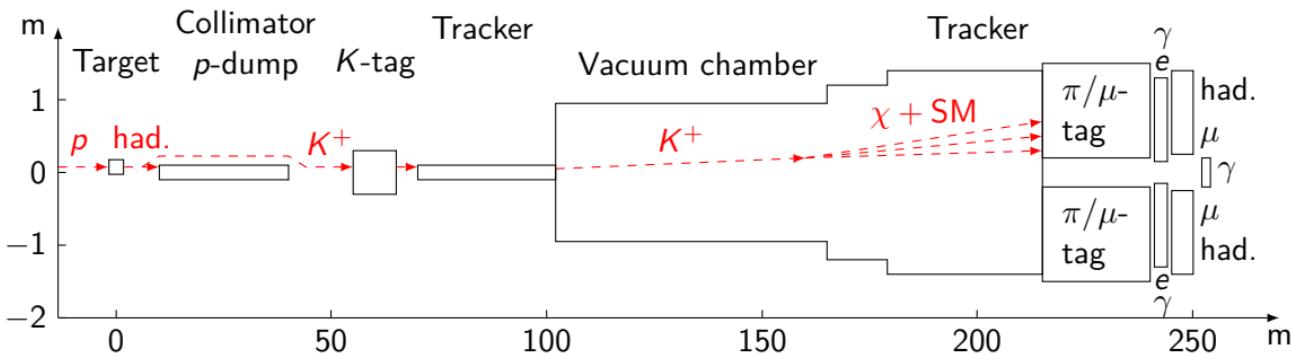
Fixed target experiment in the North Area using the CERN SPS with the goal to

- ▶ measure the very rare kaon decay $K^+ \rightarrow \pi^+ \nu \bar{\nu}$
- ▶ extract a 10 % measurement of the CKM parameter $|V_{td}|$



Fixed target experiment in the North Area using the CERN SPS with the goal to

- ▶ measure the very rare kaon decay $K^+ \rightarrow \pi^+ \nu \bar{\nu}$
- ▶ extract a 10 % measurement of the CKM parameter $|V_{td}|$

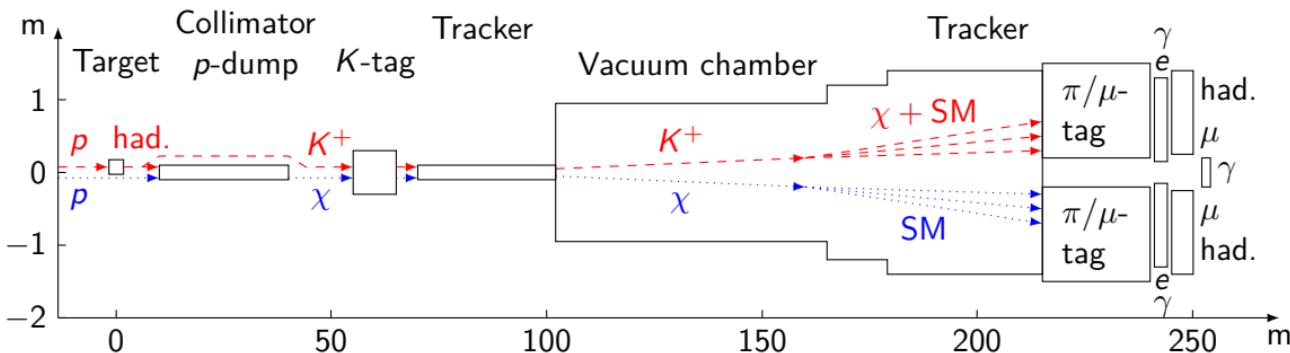


Hidden sectors at NA62

- ▶ it can also be used to search for hidden new physics χ such as a heavy neutrino
- ▶ **Target mode**
- ▶ only K^+ induced processes

Fixed target experiment in the North Area using the CERN SPS with the goal to

- ▶ measure the very rare kaon decay $K^+ \rightarrow \pi^+ \nu \bar{\nu}$
- ▶ extract a 10 % measurement of the CKM parameter $|V_{td}|$



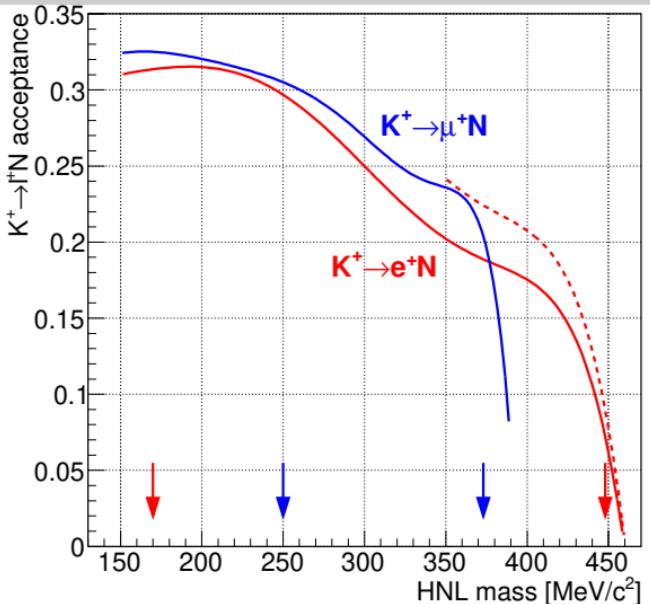
Hidden sectors at NA62

- ▶ it can also be used to search for hidden new physics χ such as a heavy neutrino
- ▶ **Target mode**
- ▶ only K^+ induced processes
- ▶ **Dump mode**
- ▶ D - and B -meson induced processes dominate

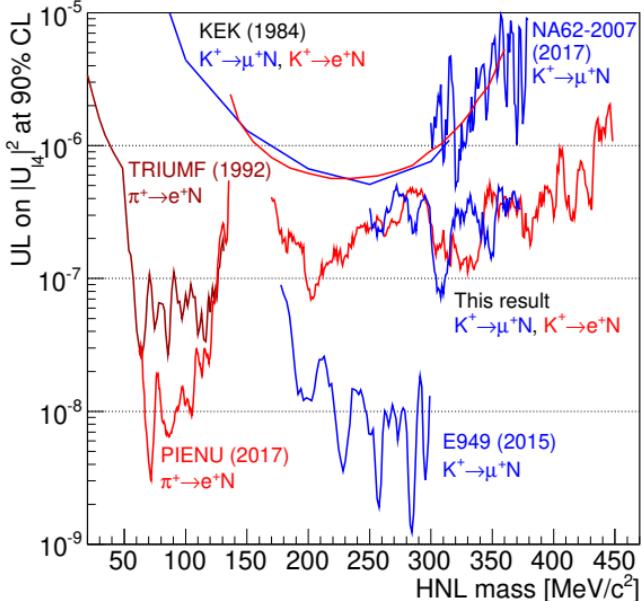
Heavy Neutrinos in the Target mode

[NA62 2017]

Acceptance of $K^+ \rightarrow \ell^+ N$



Result



5 days of operation in 2015

Heavy Neutrinos in the Dump mode

Simulation

- ▶ Toy Monte Carlo of the dump mode
- ▶ Zero background assumption

Run 3 (2021–2023)

- ▶ 10^{18} proton on target (POT)
- ▶ corresponds to about 80 days of data taking

Production of heavy neutrinos via 2×10^{15} D - and 10^{11} B -mesons

$$n_N \simeq 2N_{\text{POT}} \left(\chi_c \sum_{D_j=D^+, D^0, D_s} f_{D_j} \text{BR}(D_j \rightarrow X N_i) + \chi_b \sum_{B_k=B^+, B^0, B_s} f_{B_k} \text{BR}(B_k \rightarrow X N_i) \right) ,$$

χ production cross section normalization

f production fractions of mesons

Number of reconstructed events

$$N_{\text{obs}} = n_N \sum_{f, f' = e, \mu, \tau, \pi, K} \text{BR}(N_i \rightarrow f^+ f' - X) \mathcal{A}_i(f^+ f' - X, M_i, U_{e, \mu, \tau}^2) \varepsilon(f^+ f' - X, M_i) ,$$

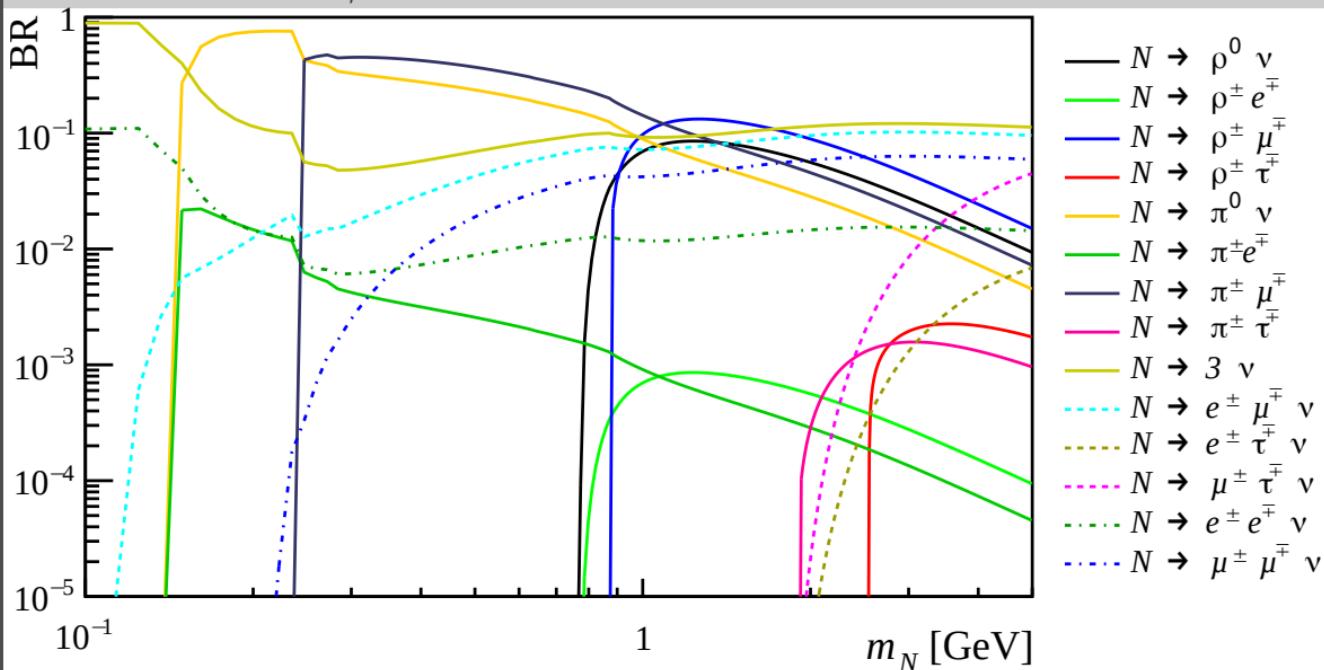
\mathcal{A}_i geometrical acceptance

ε efficiency (trigger, reconstruction, selection); assumed to be 100 %!

Branching Fractions

[Gorbunov et al. 2007]

For scenario A: $U_{ie}^2 : U_{i\mu}^2 : U_{i\tau}^2 = 1 : 160 : 27.8$

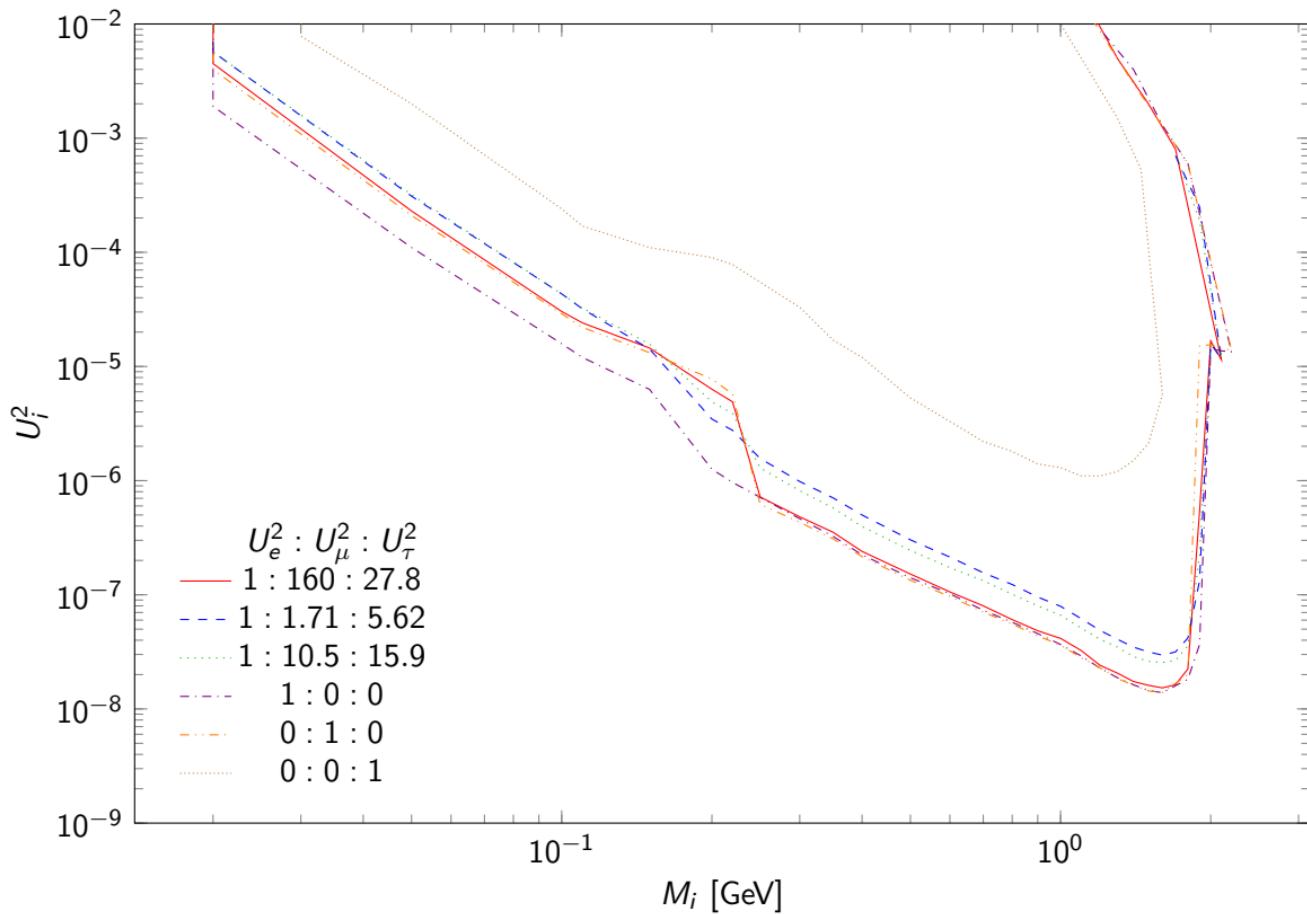


The dominant modes are

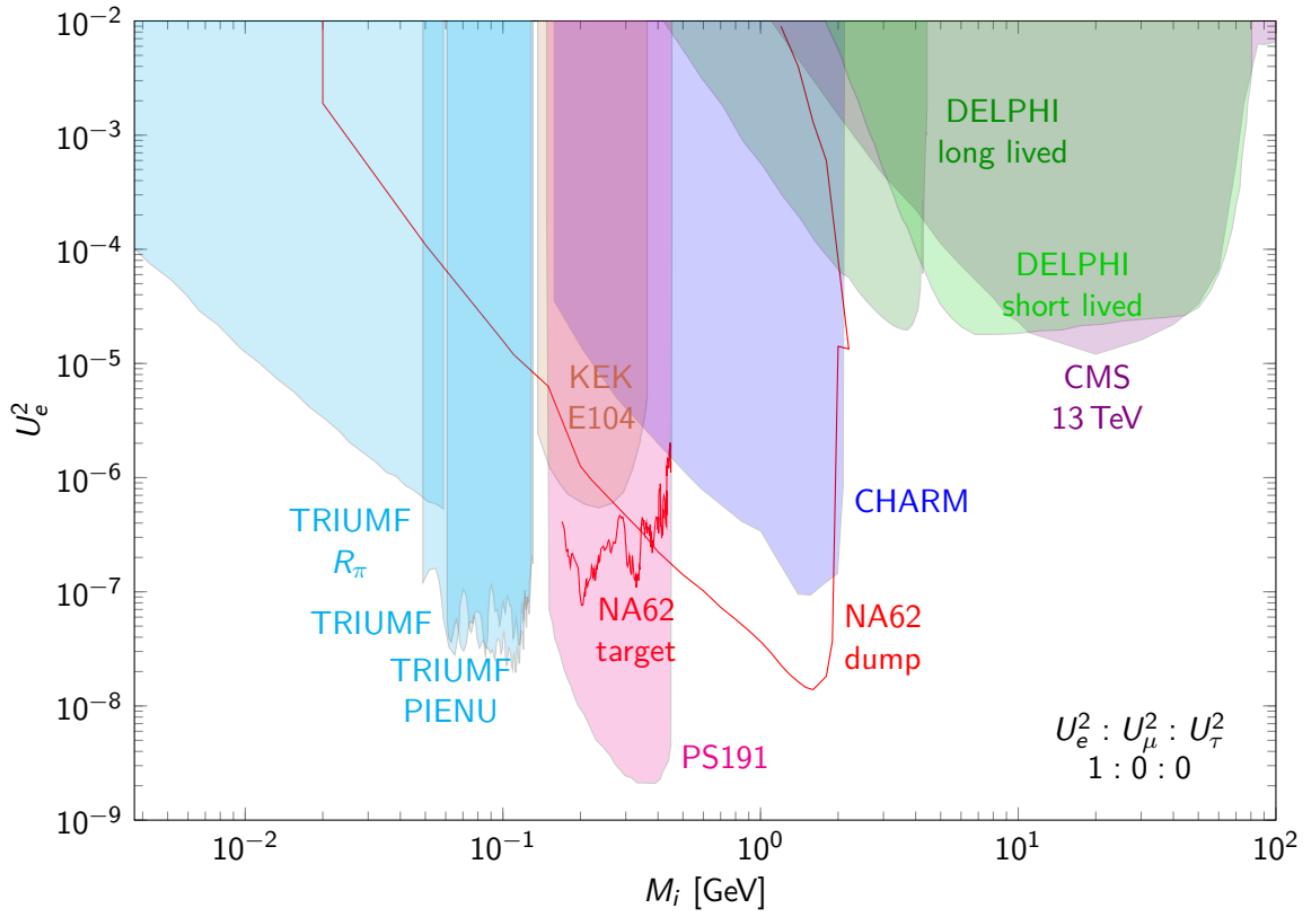
$$N_i \rightarrow 3\nu, \pi^0\nu, \pi^\pm\ell^\mp, \rho^0\nu, \rho^\pm l, \ell^+\ell^-\nu$$

The detector is able to reconstruct all final states having two charged tracks

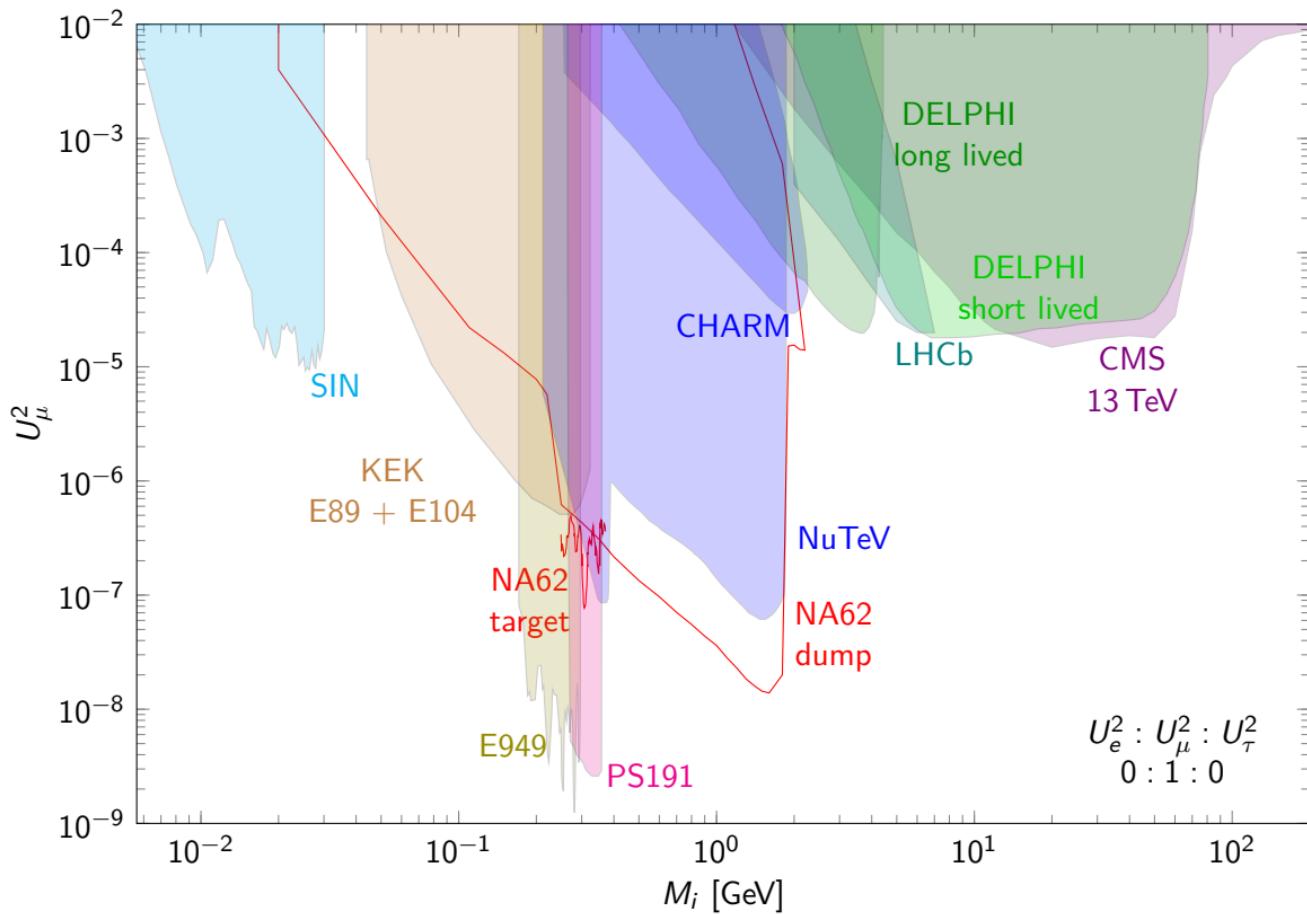
Results for NA62



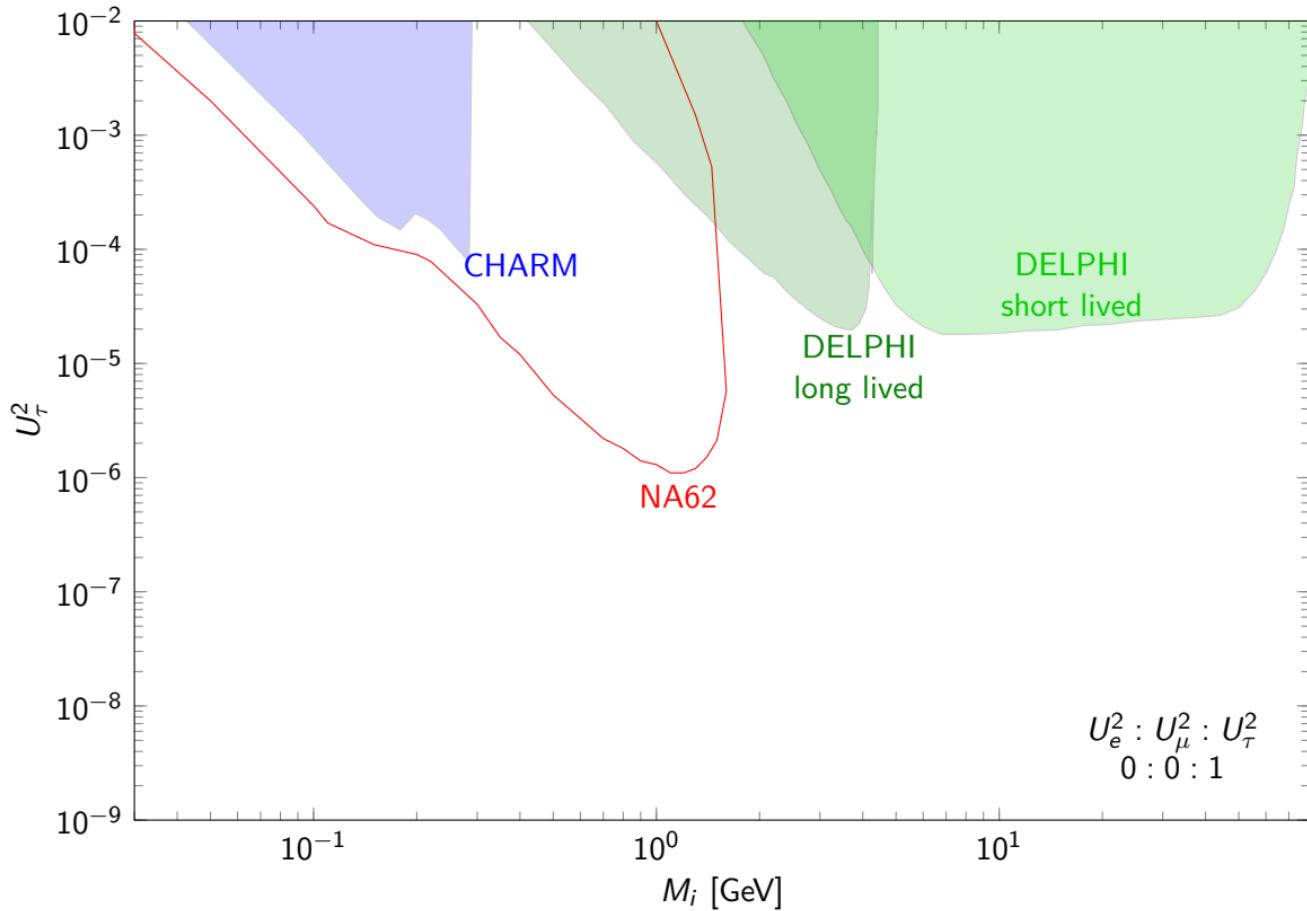
pure U_e^2



pure U_μ^2



pure U_τ^2



Conclusion

- ▶ Heavy neutrinos constitute a minimal extension to the SM
- ▶ They can easily have properties detectable at current or future experiments
- ▶ Fixed target experiments are a powerful tool to unveil hidden sectors
- ▶ Although not designed for this purpose NA62 is at the moment the leading experiment for masses between the Kaon and the D -meson mass (shown at the example of heavy neutrinos)
- ▶ There is an opportunity for fixed target experiments dedicated to new physics

References I

- M. Drewes, J. Hاجر, J. Klaric, and G. Lanfranchi. "NA62 sensitivity to heavy neutral leptons in the low scale seesaw model". *JHEP* 07, p. 105. DOI: 10.1007/JHEP07(2018)105. arXiv: 1801.04207 [hep-ph].
- J. Ellis, J. R. Espinosa, G. F. Giudice, A. Hoecker, and A. Riotto. "The Probable Fate of the Standard Model". *Phys. Lett.* B679, pp. 369–375. DOI: 10.1016/j.physletb.2009.07.054. arXiv: 0906.0954 [hep-ph]. №: CERN-PH-TH-2009-058.
- G. Degrassi et al. "Higgs mass and vacuum stability in the Standard Model at NNLO". *JHEP* 08, p. 098. DOI: 10.1007/JHEP08(2012)098. arXiv: 1205.6497 [hep-ph]. №: CERN-PH-TH-2012-134, RM3-TH-12-9.
- P. Minkowski. " $\mu \rightarrow e\gamma$ at a Rate of One Out of 10^9 Muon Decays?" *Phys. Lett.* 67B, pp. 421–428. DOI: 10.1016/0370-2693(77)90435-X. №: PRINT-77-0182 (BERN).
- R. N. Mohapatra and G. Senjanovic. "Neutrino Mass and Spontaneous Parity Violation". *Phys. Rev. Lett.* 44, p. 912. DOI: 10.1103/PhysRevLett.44.912. №: MDDP-TR-80-060, MDDP-PP-80-105, CCNY-HEP-79-10.
- T. Yanagida. "Horizontal Symmetry and Masses of Neutrinos". *Prog. Theor. Phys.* 64, p. 1103. DOI: 10.1143/PTP.64.1103. №: TU-80-208.
- J. Schechter and J. W. F. Valle. "Neutrino Masses in $SU(2) \times U(1)$ Theories". *Phys. Rev.* D22, p. 2227. DOI: 10.1103/PhysRevD.22.2227. №: SU-4217-167, COO-3533-167.
- T. Asaka and M. Shaposhnikov. "The ν MSM, dark matter and baryon asymmetry of the universe". *Phys. Lett.* B620, pp. 17–26. DOI: 10.1016/j.physletb.2005.06.020. arXiv: hep-ph/0505013 [hep-ph].

References II

- M. Shaposhnikov. "A Possible symmetry of the ν MSM". *Nucl. Phys.* B763, pp. 49–59. DOI: 10.1016/j.nuclphysb.2006.11.003. arXiv: hep-ph/0605047 [hep-ph]. №: CERN-PH-TH-2006-079.
- NA62**. "Search for heavy neutral lepton production in K^+ decays". arXiv: 1712.00297 [hep-ex]. №: CERN-EP-2017-311.
- D. Gorbunov and M. Shaposhnikov. "How to find neutral leptons of the ν MSM?" *JHEP* 10. [Erratum: JHEP11,101(2013)], p. 015. DOI: 10.1007/JHEP11(2013)101, 10.1088/1126-6708/2007/10/015. arXiv: 0705.1729 [hep-ph].