



Bounding the Higgs-boson width at the HL-LHC through interference effects

DIS 2019 in Torino

9 April 2019

WG3: Higgs and BSM Physics in Hadron Collisions



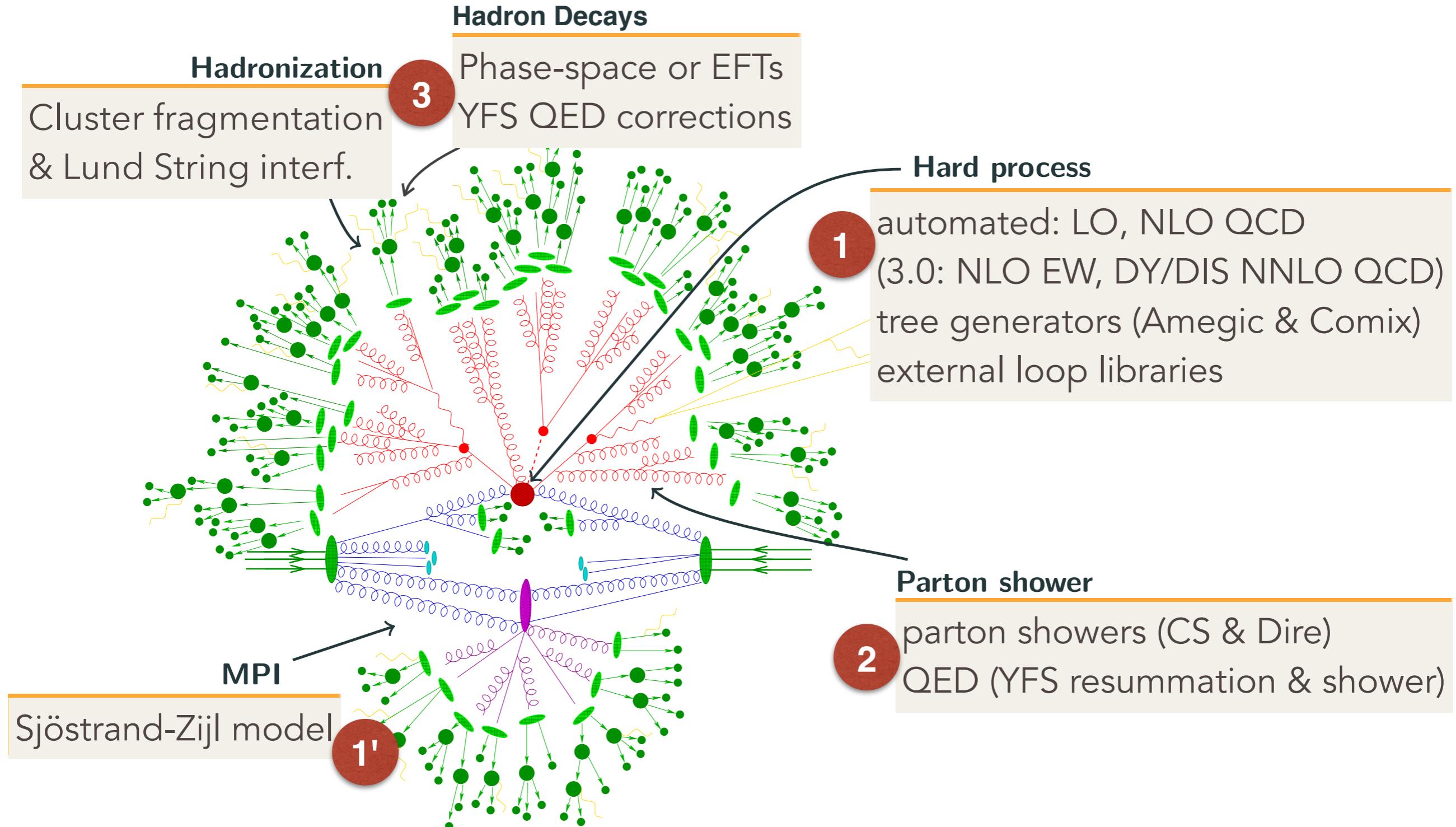
Enrico Bothmann

Lance Dixon, Stefan Höche, Silvan Kuttimalai



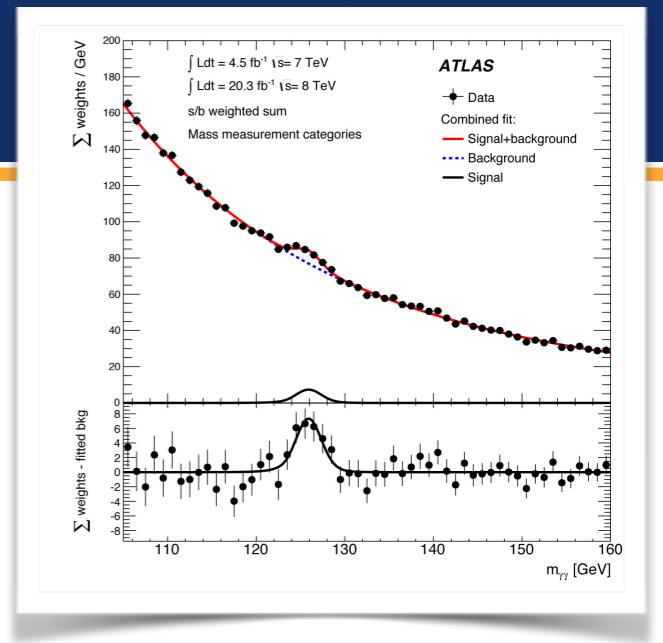
SHERPA overview

[Gleisberg et al. 0811.4622]



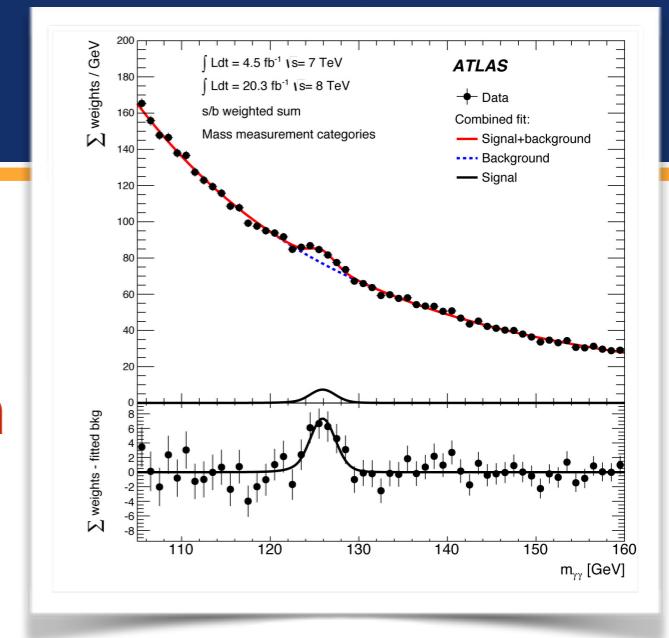
Motivation

- ▶ LHC: Higgs mostly produced through gluon fusion
- ▶ $H \rightarrow \gamma\gamma$ decay of high relevance due to clean final state
 - ▶ measure mass peak M_H with good accuracy
 - ▶ can not measure width directly, $\Gamma_H < O(10^{-2}) \times \text{exp. resolution } \sigma_{\text{res}}$
- ▶ Higgs might couple to unknown states $\sim \Gamma_H > \Gamma_H^{\text{SM}}$
- ▶ on-shell signal cross section $\sim g^2/\Gamma_H \sim$ coupling-width degeneracy
- ▶ how to break degeneracy to measure Γ_H without assumptions on g ?
 - ▶ complement with off-shell measurements (measure with precision $\Gamma_H = 4 \pm 1$ MeV, but need to assume that couplings do not vary with scale)
[ATLAS 1902.00134]
 - ▶ stay on-shell, use interference-induced rate change (bound $\sim 8-22 \Gamma_H/\Gamma_H^{\text{SM}}$)
[Campbell et al. 1704.08259]
 - ▶ stay on-shell, use interference-induced peak shift/deformation

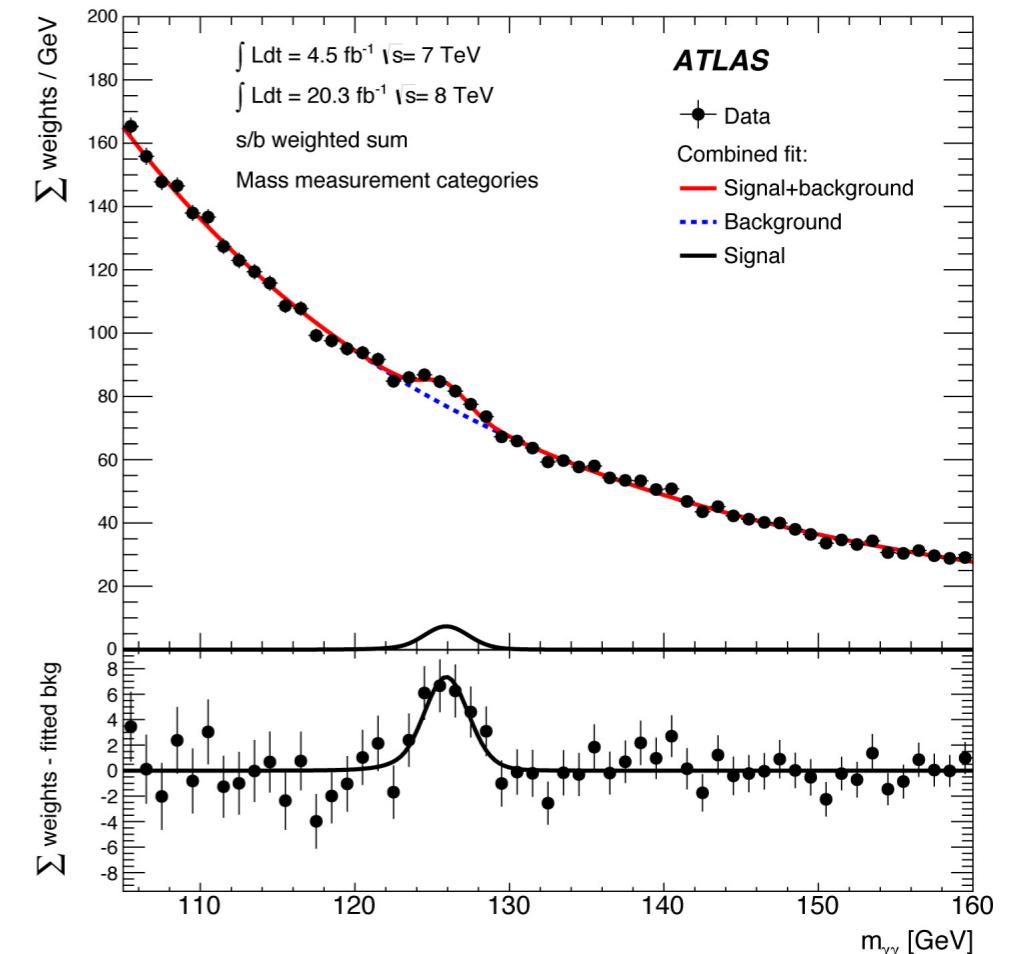


Motivation

- ▶ use interference-induced peak shift/deformation
- ▶ estimates for shift-based bounds at fixed-order already exist, but are they robust?
[Dixon, Li 1305.3854 (2013)]
- ▶ calculate particle-level prediction and use realistic analysis cuts
 - ▶ Sherpa S-MC@NLO with implementation of NLO calculation for interference terms
[Dixon, Li 1305.3854 (2013)]
 - ▶ background 0j@NLO, $\leq 3j$ @LO CKKW-L
 - ▶ estimate (HL-)LHC reach using toy experiments
 - ▶ new: explore possibility of a direct fit to the distorted line shape



Overview: Mass shift through interference

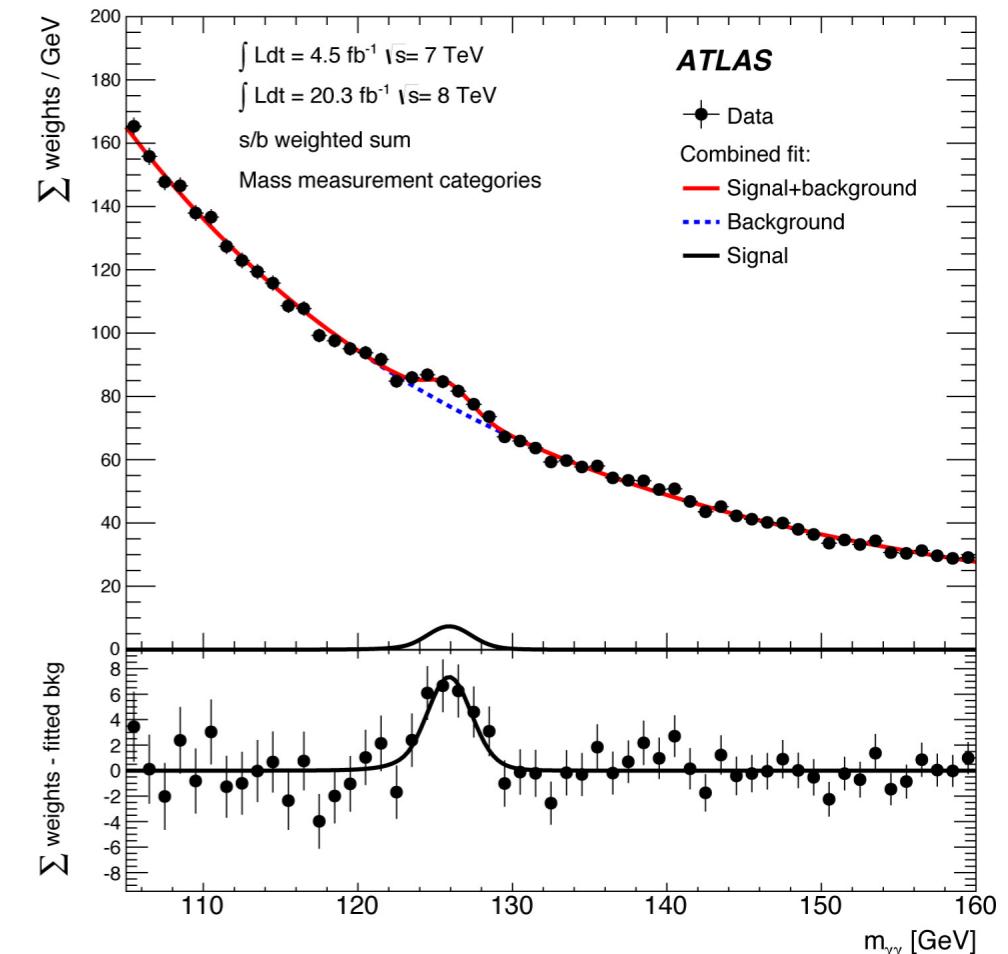


Overview: Mass shift through interference

[Martin 1208.1533 (2012)]

observation: interference of $gg \rightarrow H \rightarrow \gamma\gamma$
with $gg \rightarrow$ quark loop $\rightarrow \gamma\gamma$
 \Rightarrow smeared Higgs mass peak in $m_{\gamma\gamma}$ shifts:

$$\Delta M_H = -150 \text{ MeV} \quad (\text{LO})$$



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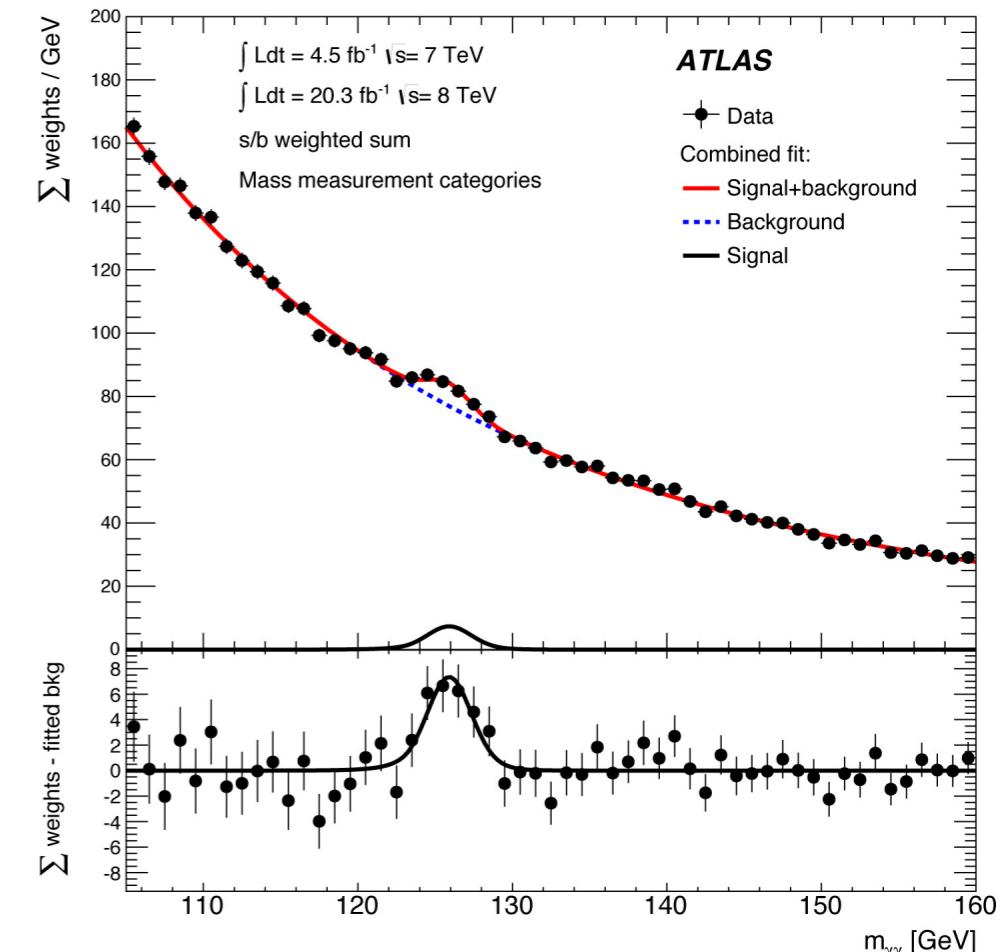
$$\Delta M_H = -150 \text{ MeV} \quad (\text{LO})$$

$$\sim 30 \times \Gamma_H^{\text{SM}} \quad (4 \text{ MeV})$$

$$\sim 0.1 \times \sigma_{\text{res}} \quad (1.7 \text{ GeV})$$

$$\sim 2.5 \times m_H^{\gamma\gamma} \text{ uncert.} \quad (0.4 \text{ GeV at } 36 \text{ fb}^{-1} \text{ 13 TeV})$$

[ATLAS 1806.00242]



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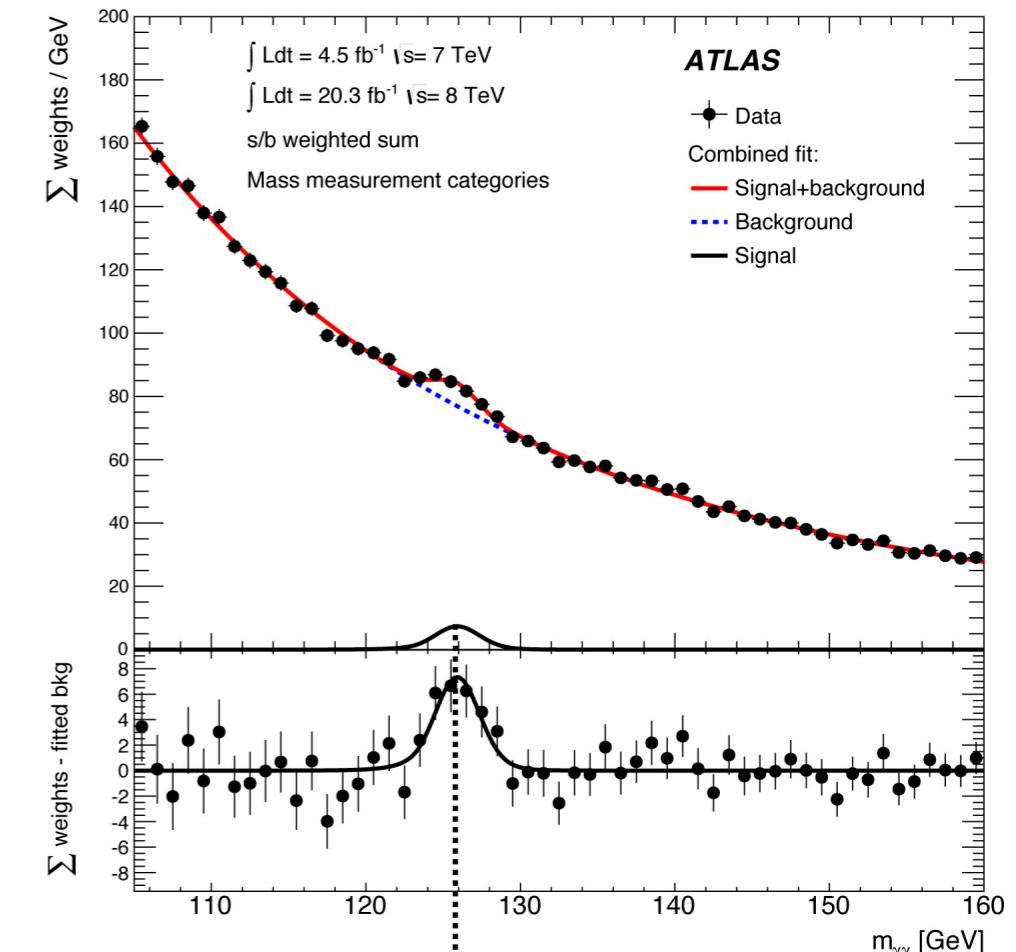
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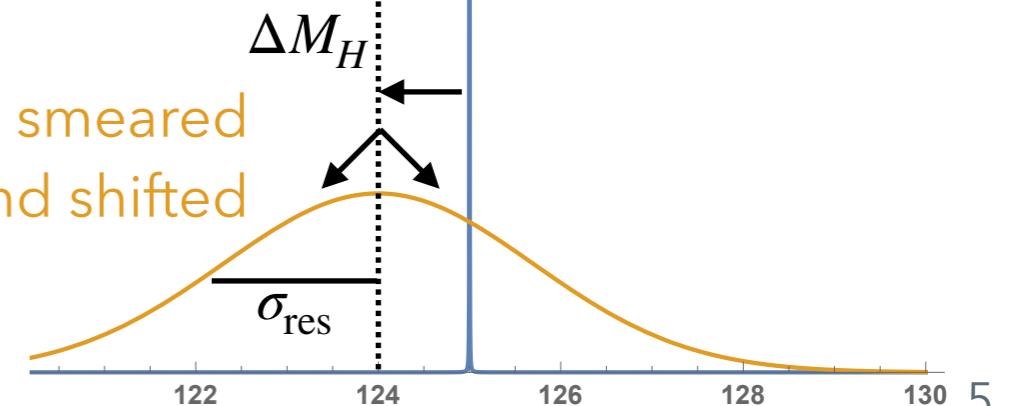
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[ATLAS 1806.00242]



1: Breit-Wigner peak

2: smeared and shifted



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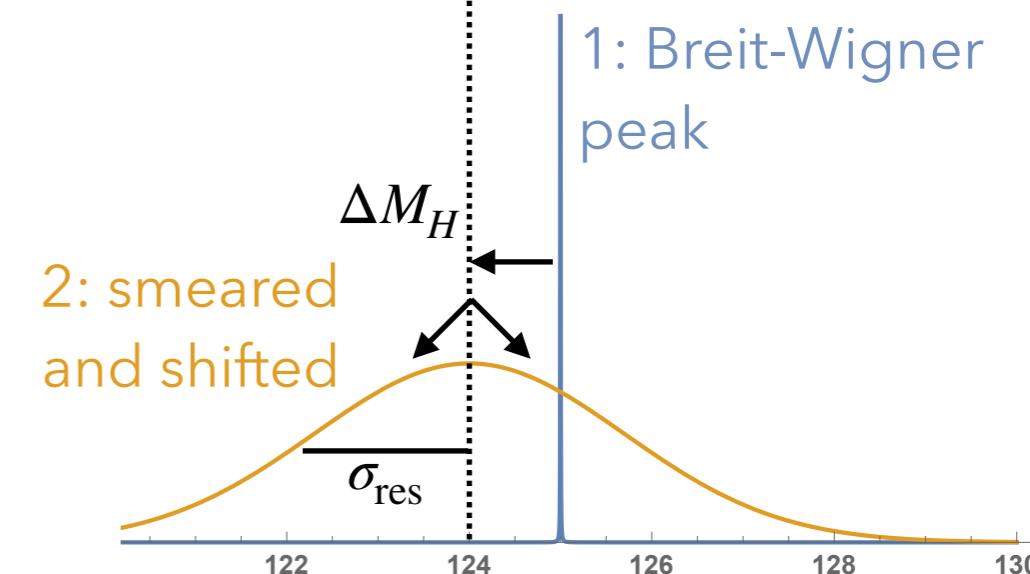
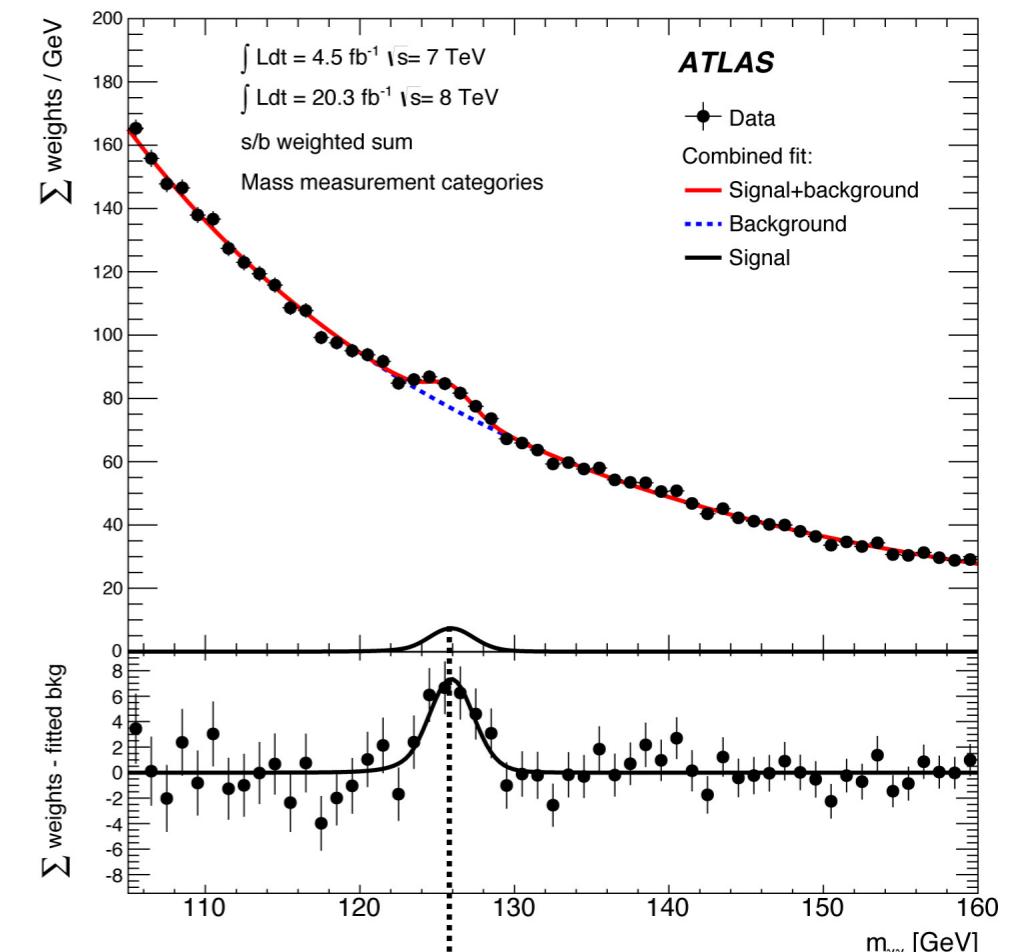
[ATLAS 1806.00242]

[Dixon, Li 1305.3854 (2013)]

$$\Delta M_H = -70 \text{ MeV} \quad (\text{NLO})$$

(reduced due to large signal K factor)

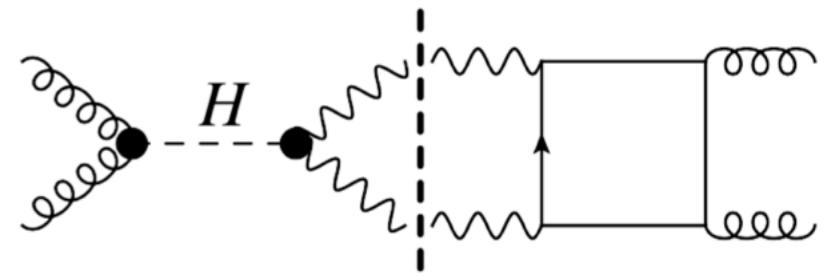
observation: fixing signal event yield gives Γ_H bound independent from further assumptions on couplings and/or decay modes



Why is the peak shifting?

[Martin 1208.1533 (2012)]

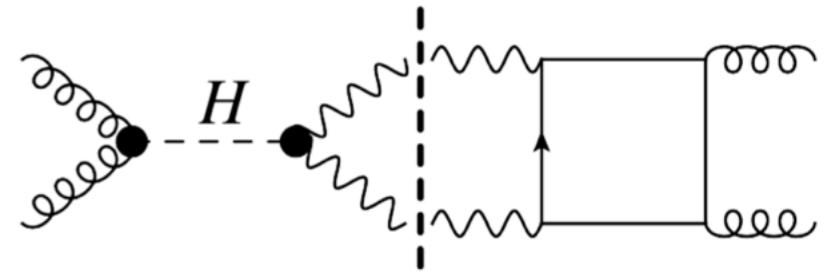
$$\mathcal{M} = \frac{1}{m^2 - m_H^2 + i m_H \Gamma_H} \frac{\mathcal{A}_S}{\sqrt{\pi}} + \frac{\mathcal{A}_B}{\sqrt{\pi}}$$



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$$\hat{\sigma}_S = \text{Re}\{\mathcal{L}\} \frac{1}{2m_H^2} \int d\Phi \frac{|\mathcal{A}_S|^2}{m_H \Gamma_H},$$

$$\hat{\sigma}_I = \text{Re}\{\mathcal{L}\} \frac{1}{2m_H^2} \int d\Phi 2 \text{Im}\{\mathcal{A}_S \mathcal{A}_B^*\},$$

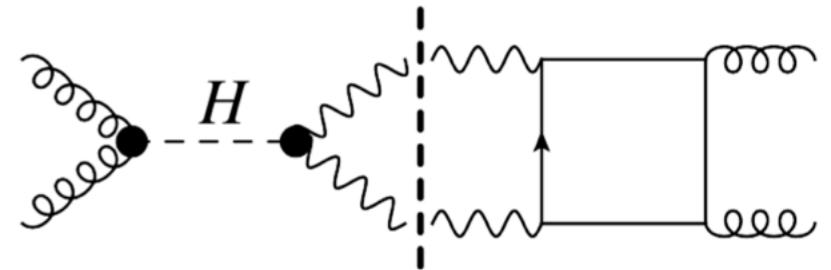
$$\hat{\sigma}_R = \text{Im}\{\mathcal{L}\} \frac{1}{2m_H^2} \int d\Phi 2 \text{Re}\{\mathcal{A}_S \mathcal{A}_B^*\}.$$

$$\mathcal{L} = \frac{1}{\pi} \frac{m_H \Gamma_H + i(m_{\gamma\gamma}^2 - m_H^2)}{(m_{\gamma\gamma}^2 - m_H^2)^2 + m_H^2 \Gamma_H^2}$$

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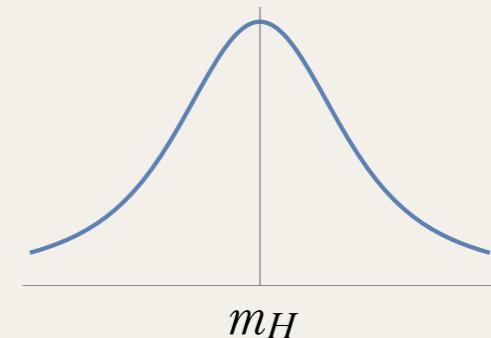
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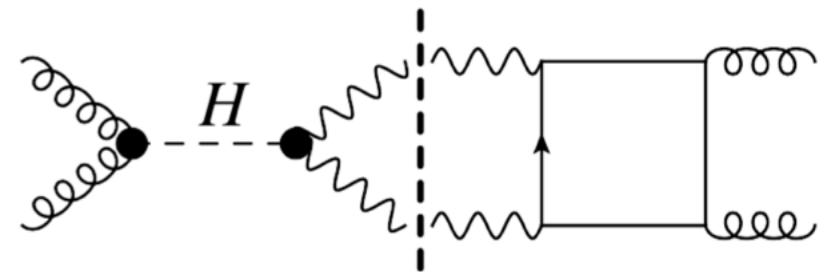
qualitative
contribution shapes
after smearing:



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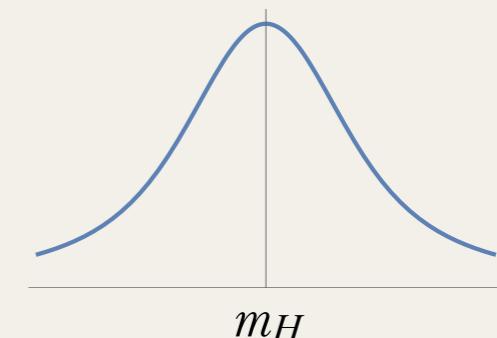
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qualitative contribution shapes after smearing:



→ sum gives distorted+shifted peak

Why should we care?

[Dixon, Li 1305.3854 (2013)]

- BSM: factors c_g, c_γ for $Hgg, H\gamma\gamma$ couplings (SM: 1)
- let c_g, c_γ, Γ_H vary, but keep measured signal yield fixed: $\mu_{\gamma\gamma} \approx 1$

BSM parametrisation = SM \times signal yield

$$\frac{(c_g c_\gamma)^2 \sigma_S}{m_H \Gamma_H} + c_g c_\gamma \sigma_I = \left(\frac{\sigma_S}{m_H \Gamma_H^{\text{SM}}} + \sigma_I \right) \mu_{\gamma\gamma}$$

- σ_I very small, can be neglected for $\Gamma_H \lesssim 100 \Gamma_H^{\text{SM}}$

$$\Rightarrow c_g c_\gamma = \sqrt{\mu_{\gamma\gamma} \frac{\Gamma_H}{\Gamma_H^{\text{SM}}}} \quad \text{and with} \quad \Delta m_H \sim c_g c_\gamma \quad \rightarrow \quad \Delta m_H \sim \sqrt{\mu_{\gamma\gamma} \frac{\Gamma_H}{\Gamma_H^{\text{SM}}}}$$

⇒ bound on Γ_H independent from further assumptions on couplings and/or decay modes

A reference value for m_H

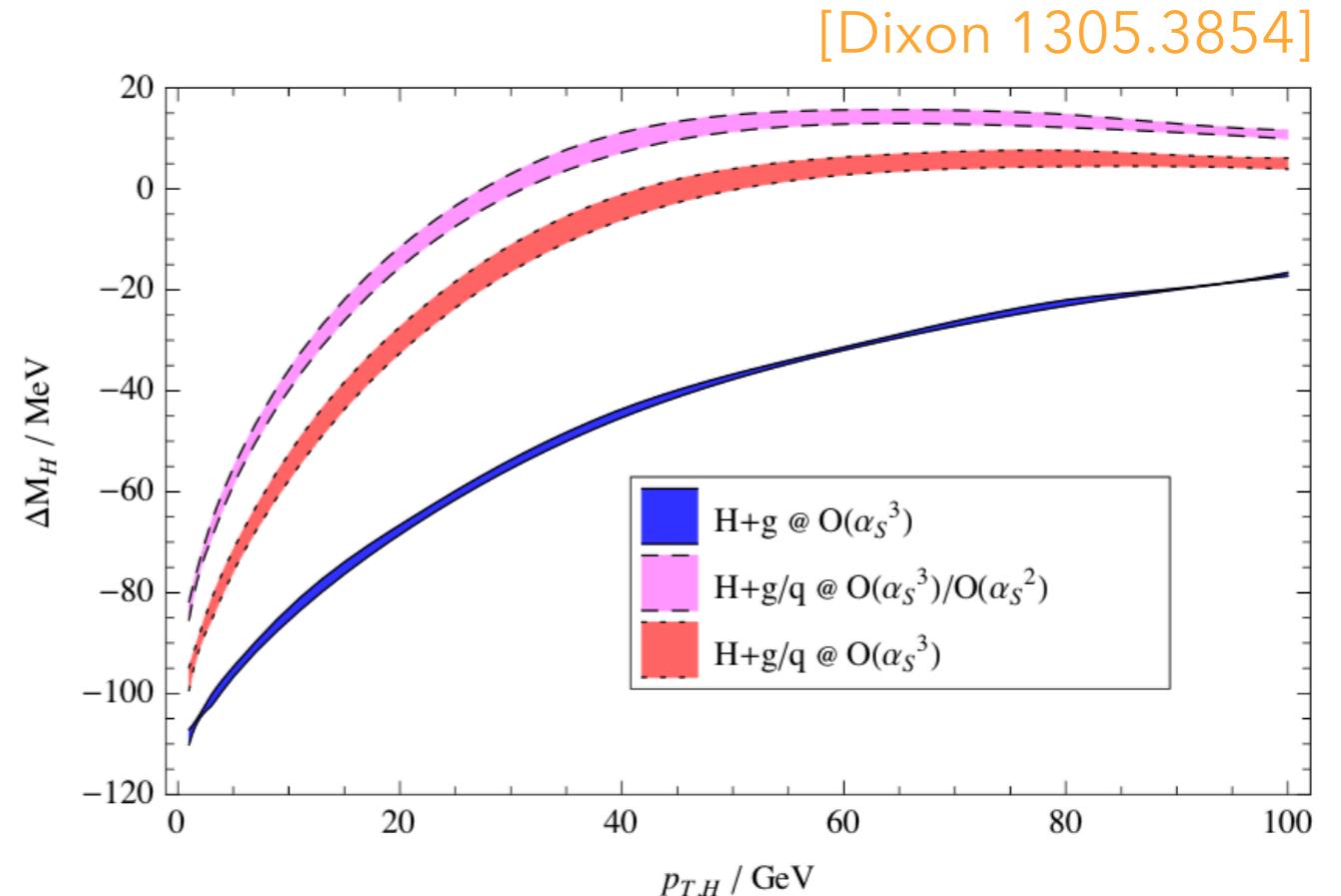
**we need a comparison value to
extract $\Delta m_H = m_H^{\text{shifted}} - m_H^{\text{actual}}$**

- ▶ $\gamma\gamma j$ has smaller relative magnitude of interference
 - ▶ ... and opposite sign of interference for qg- and gg-initiated channels \Rightarrow cancellation
- ↪ $p_{T,H}$ cut dependent mass shift

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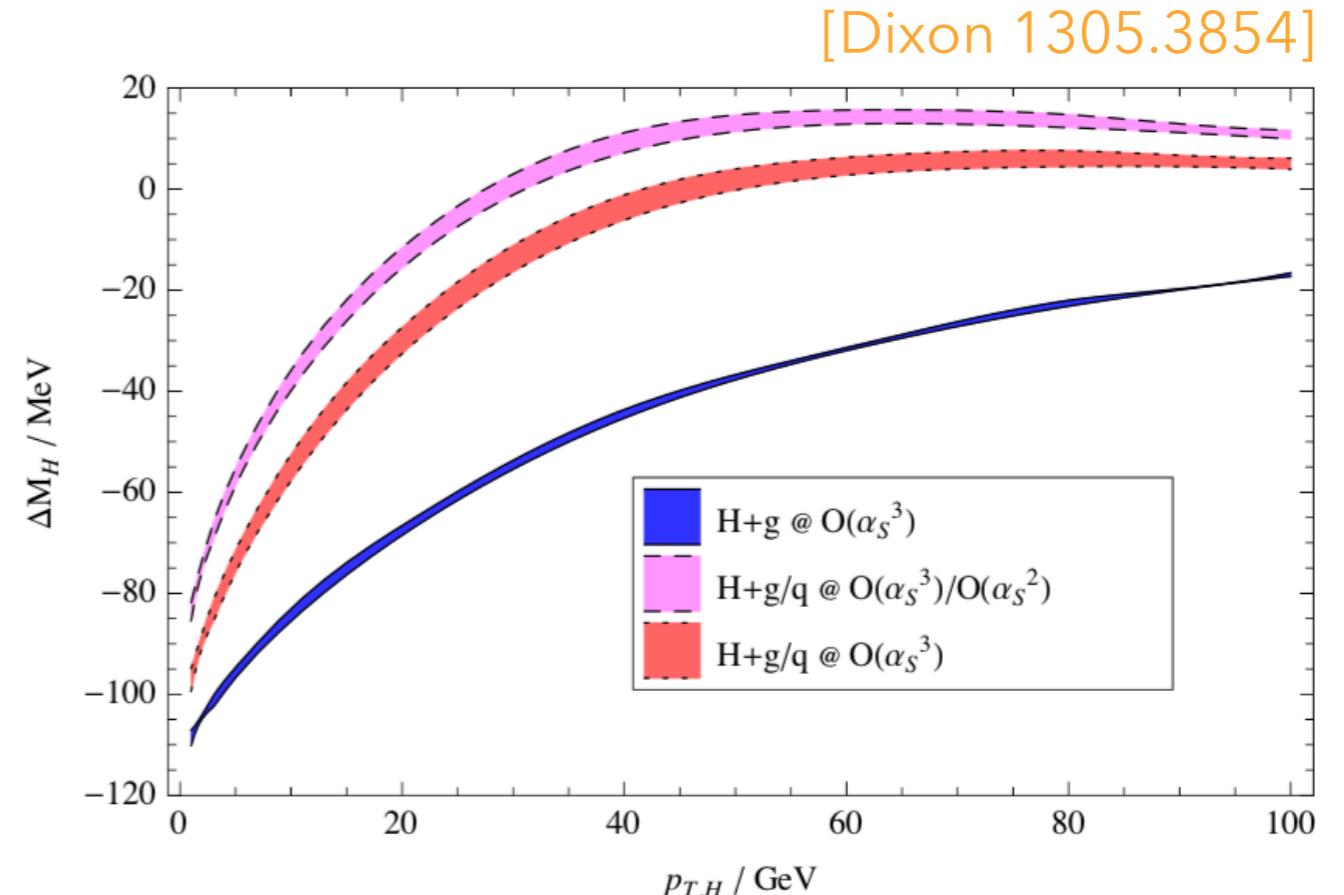
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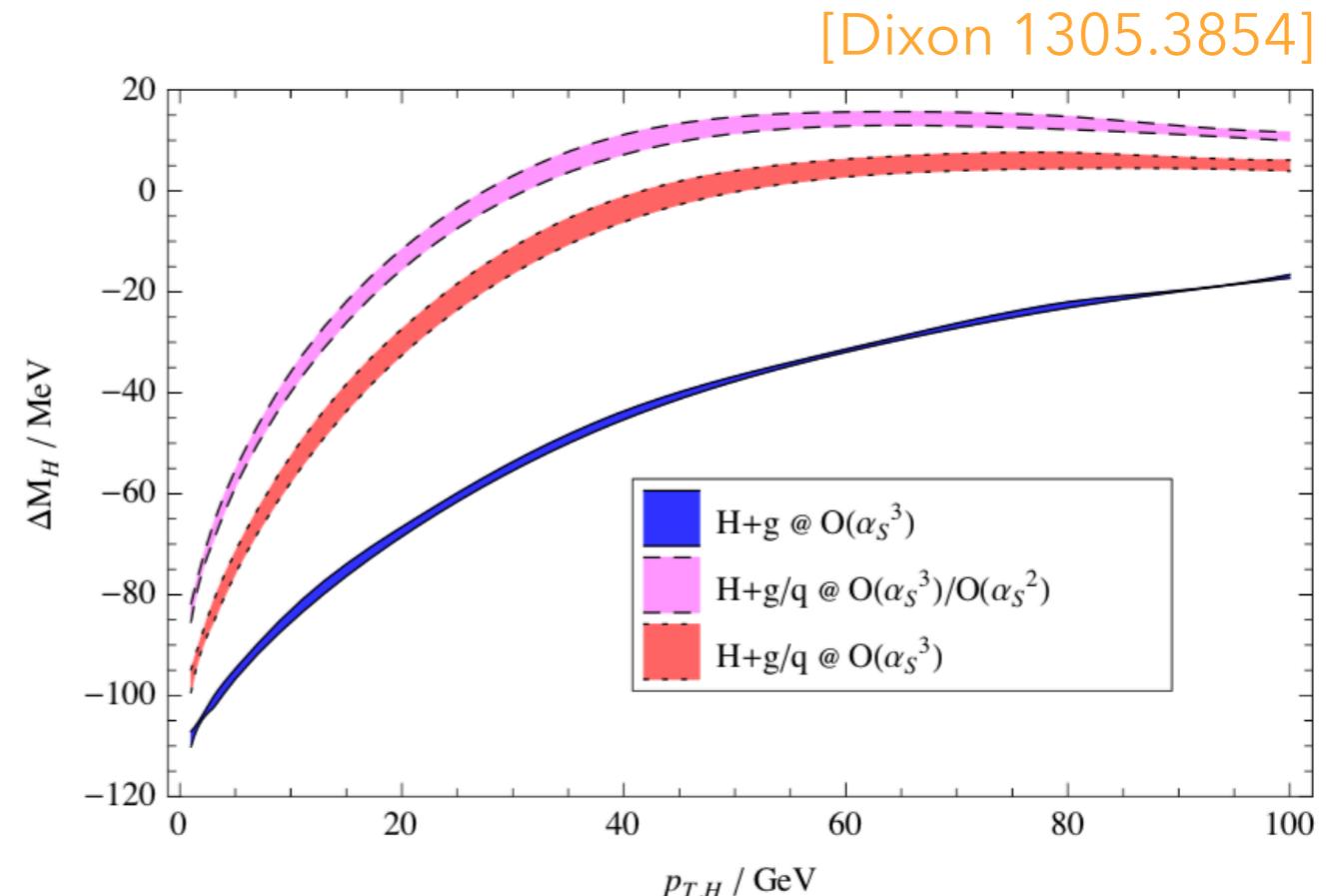


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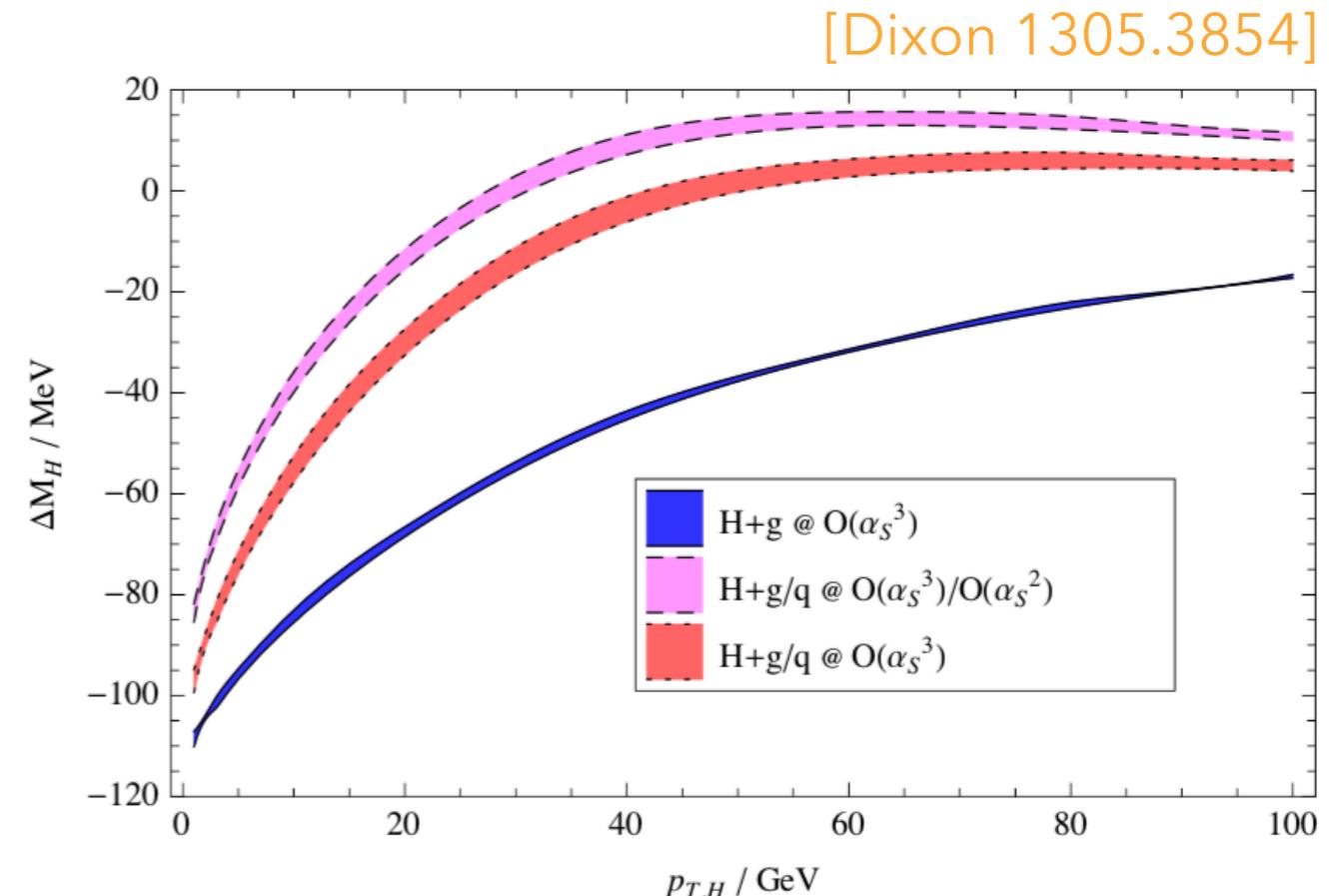


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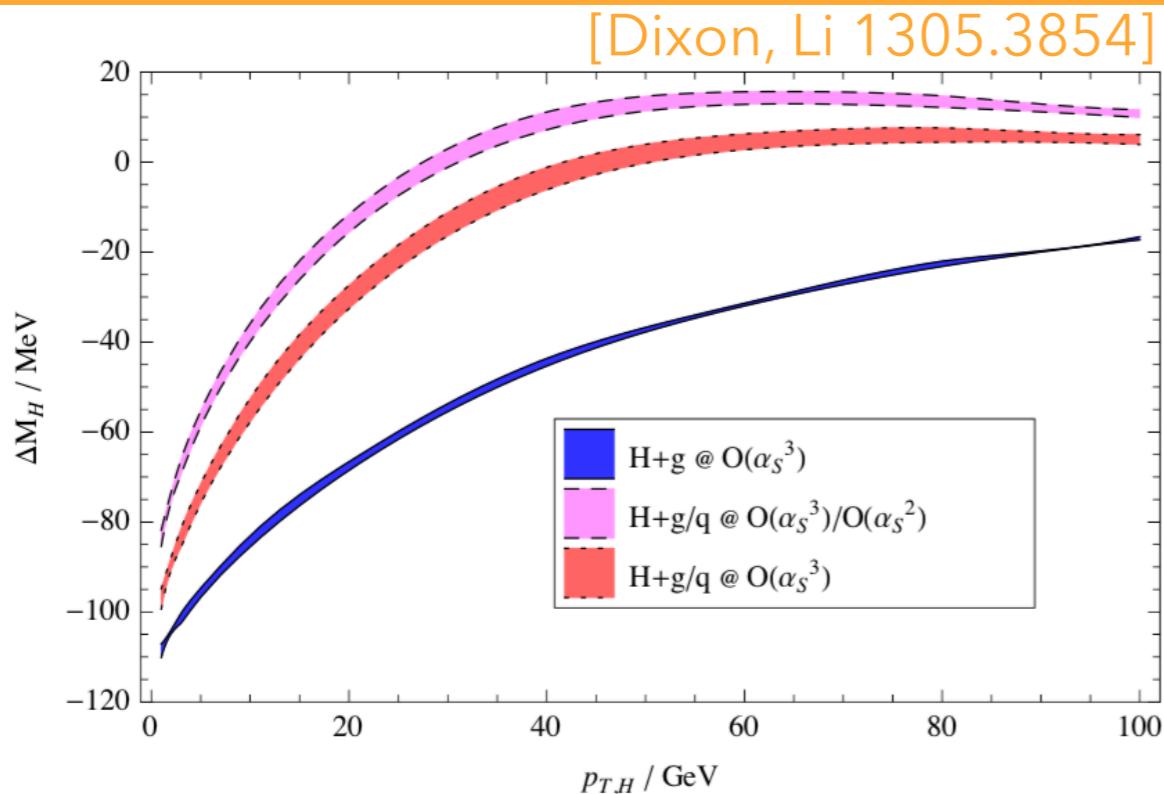
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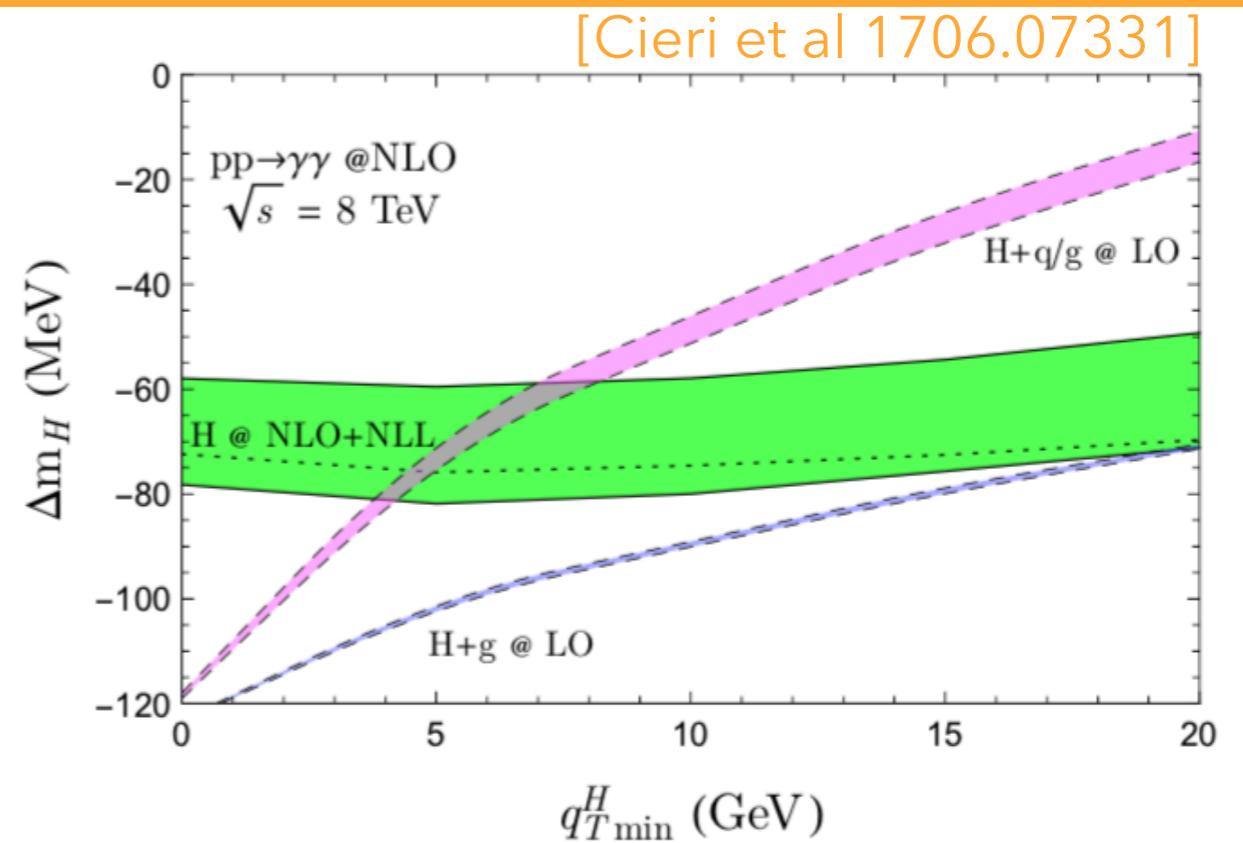
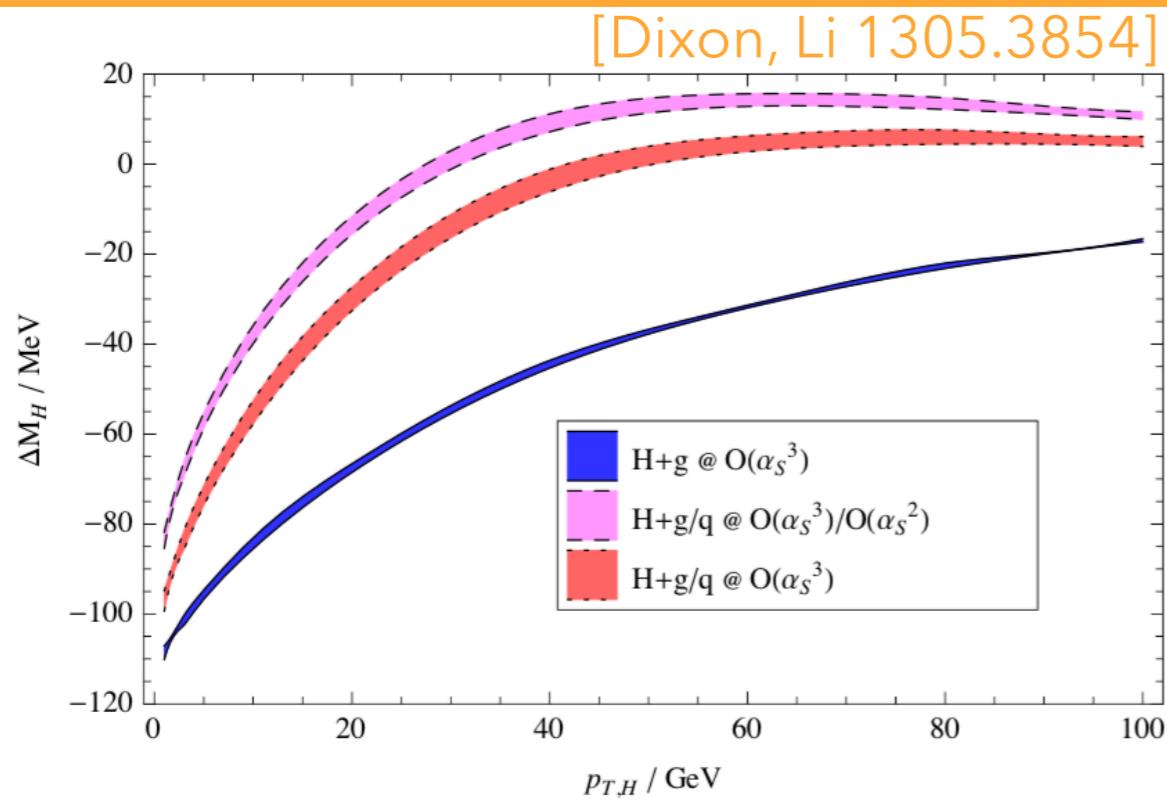


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- ▶ but fixed-order unreliable for low p_T ↪ how stable when including resummation (& hadronisation?) effects

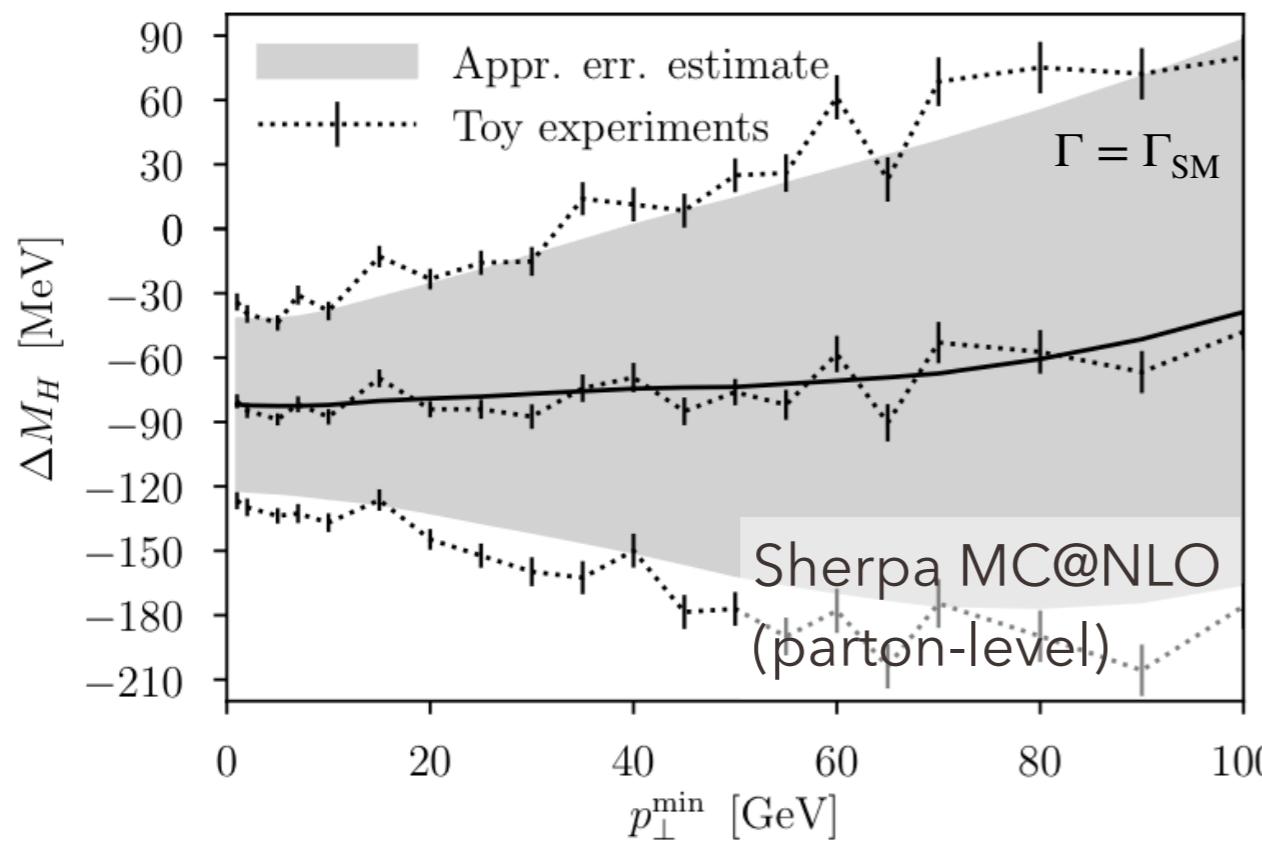
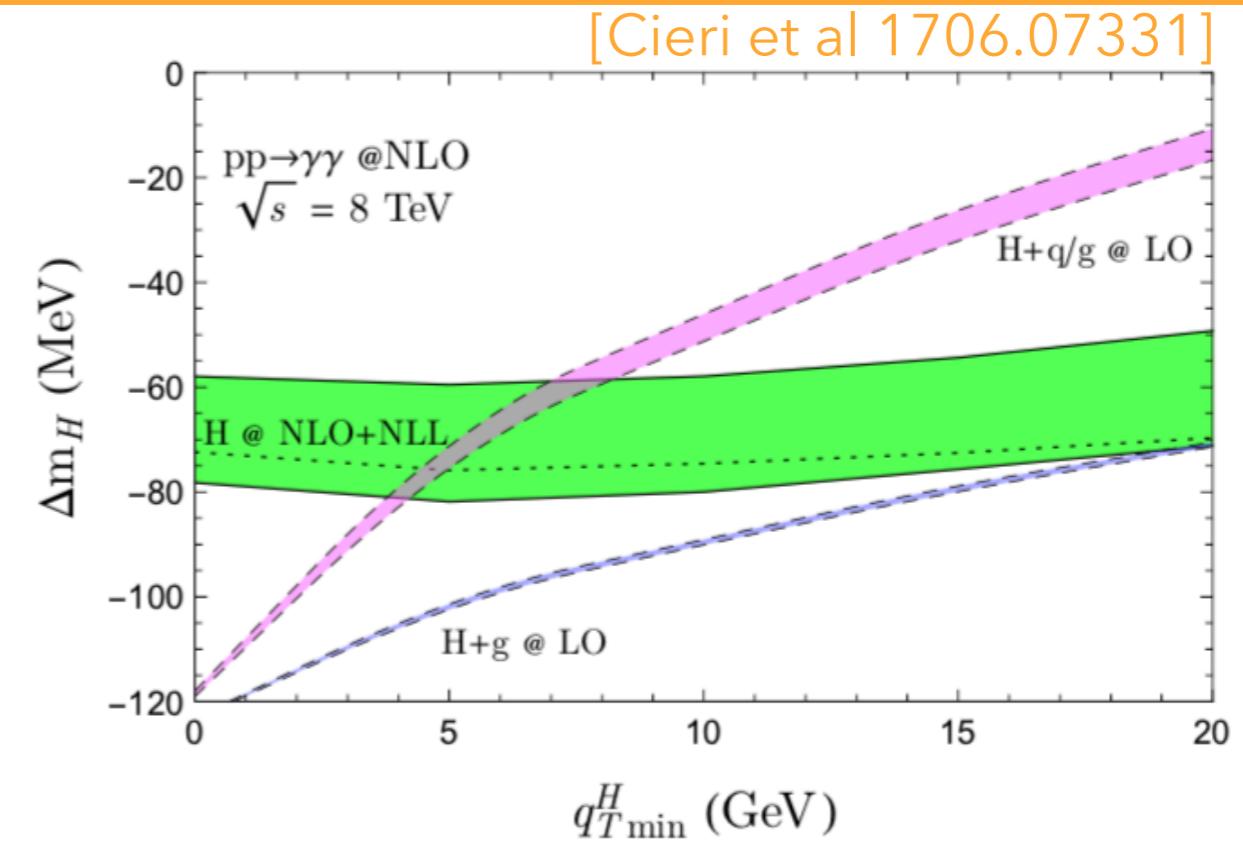
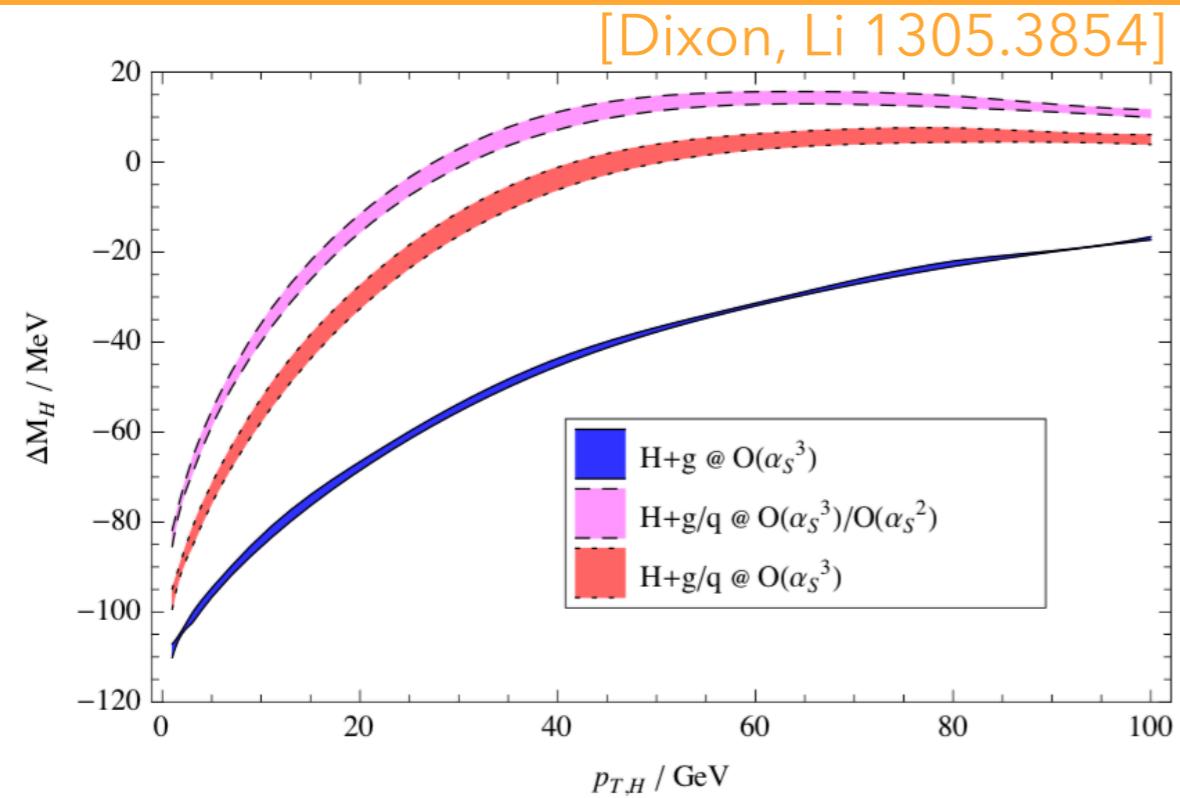
p_T extraction for fixed-order & resummed



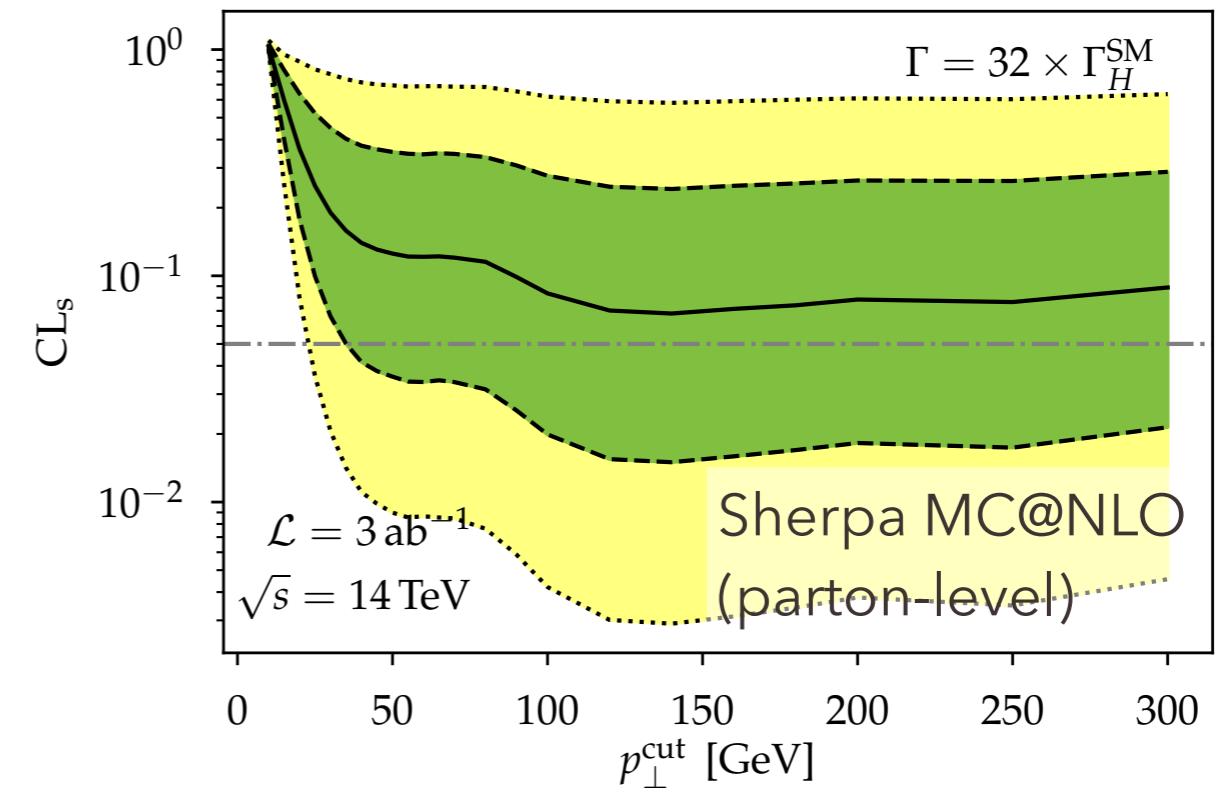
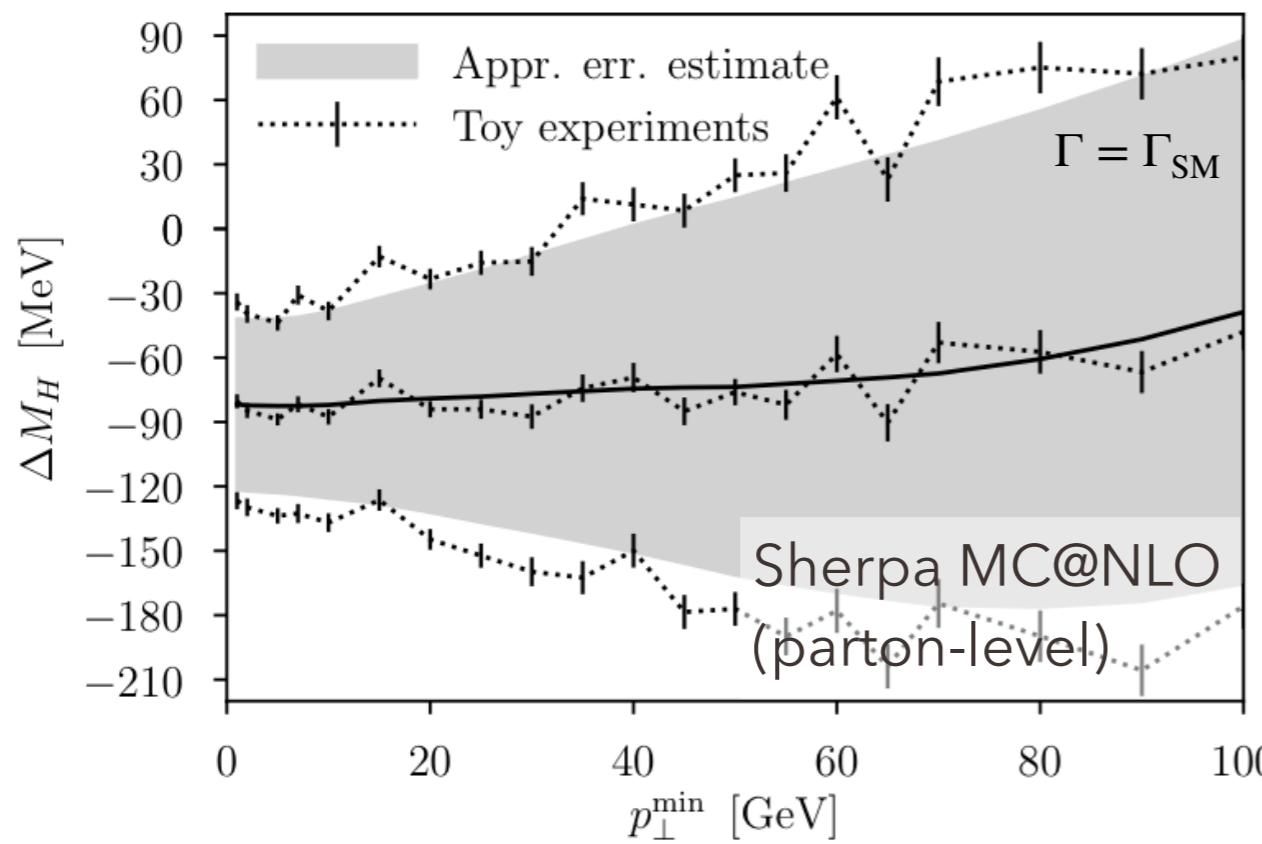
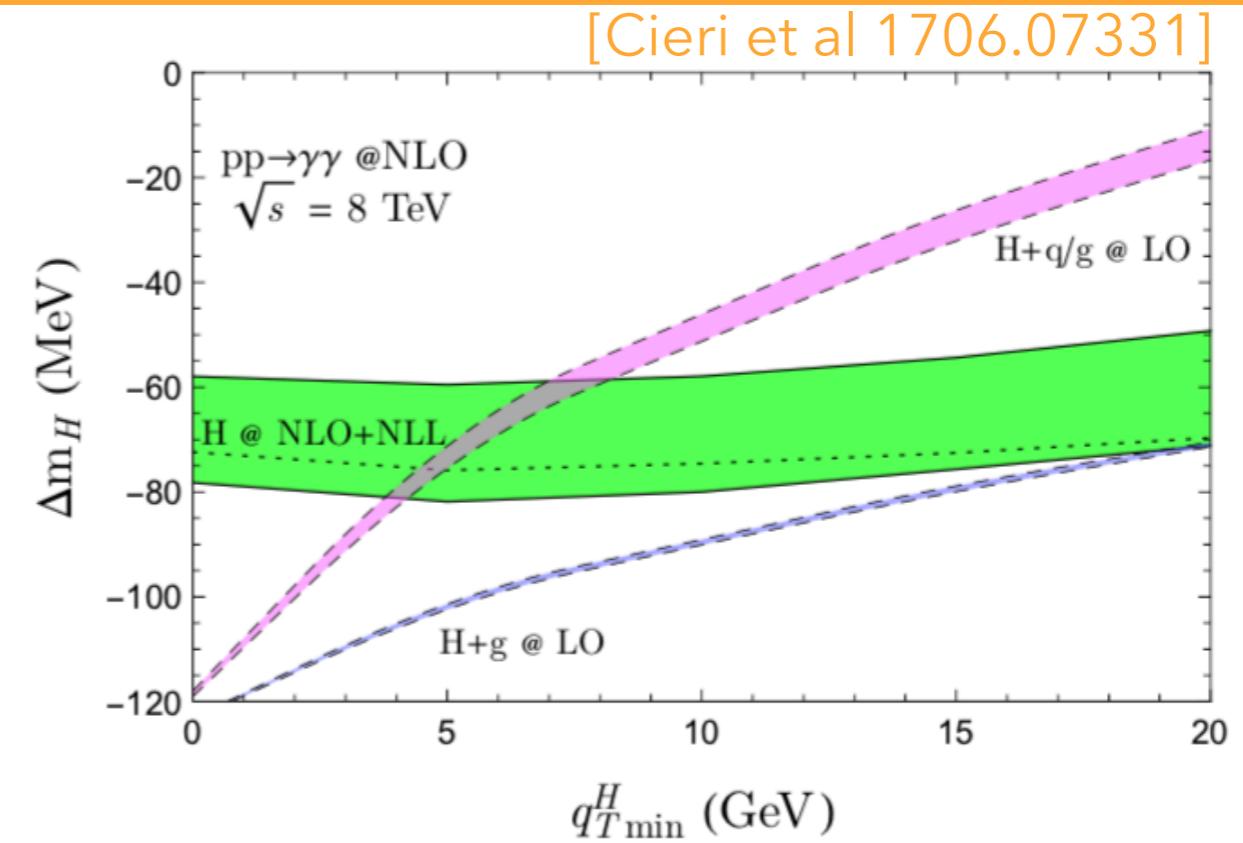
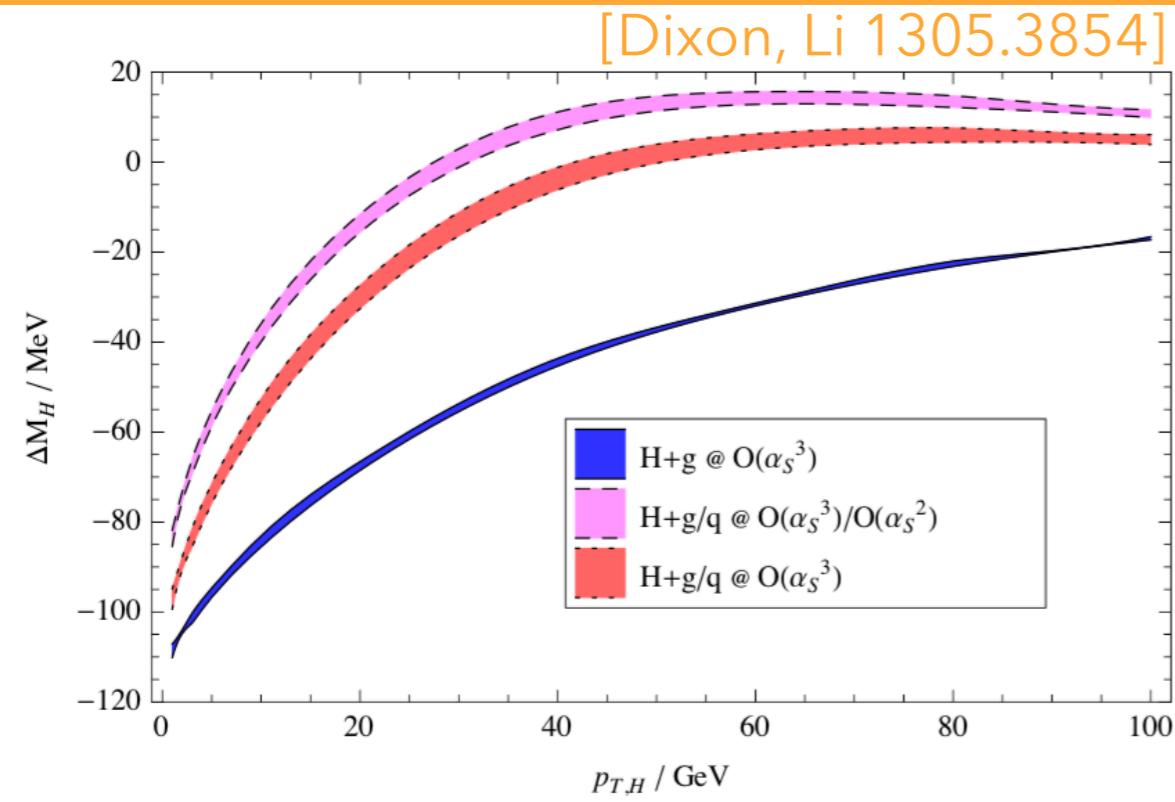
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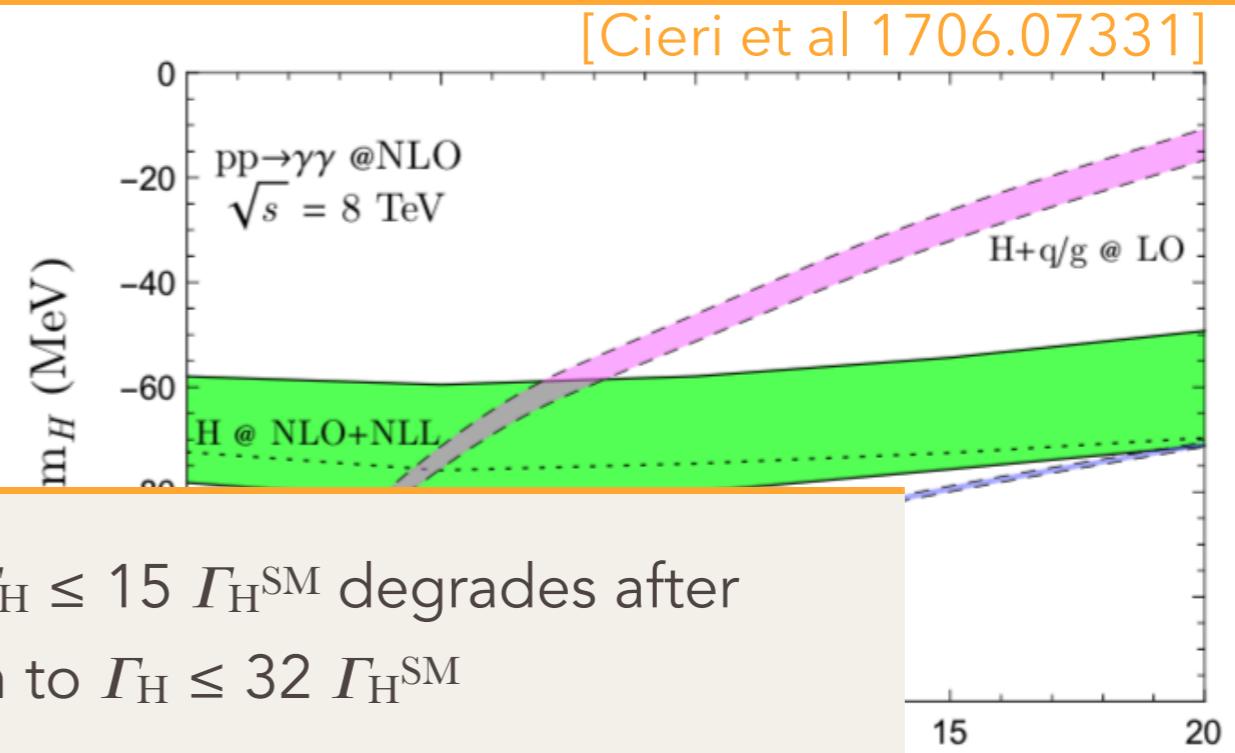
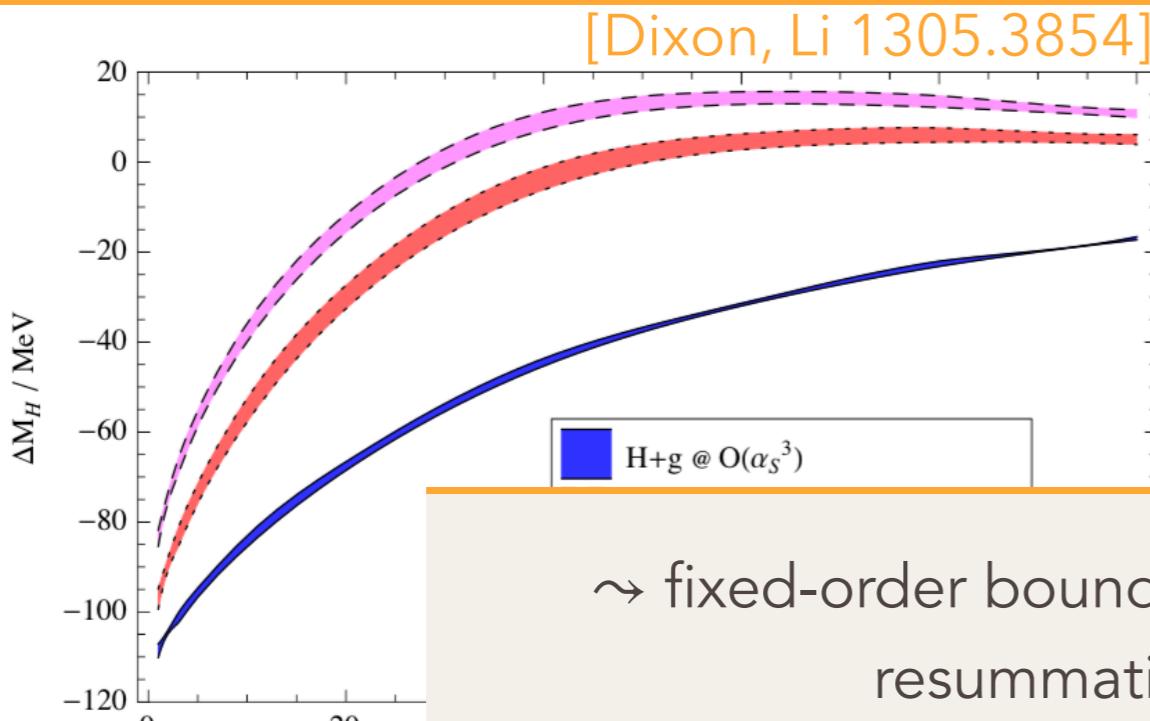
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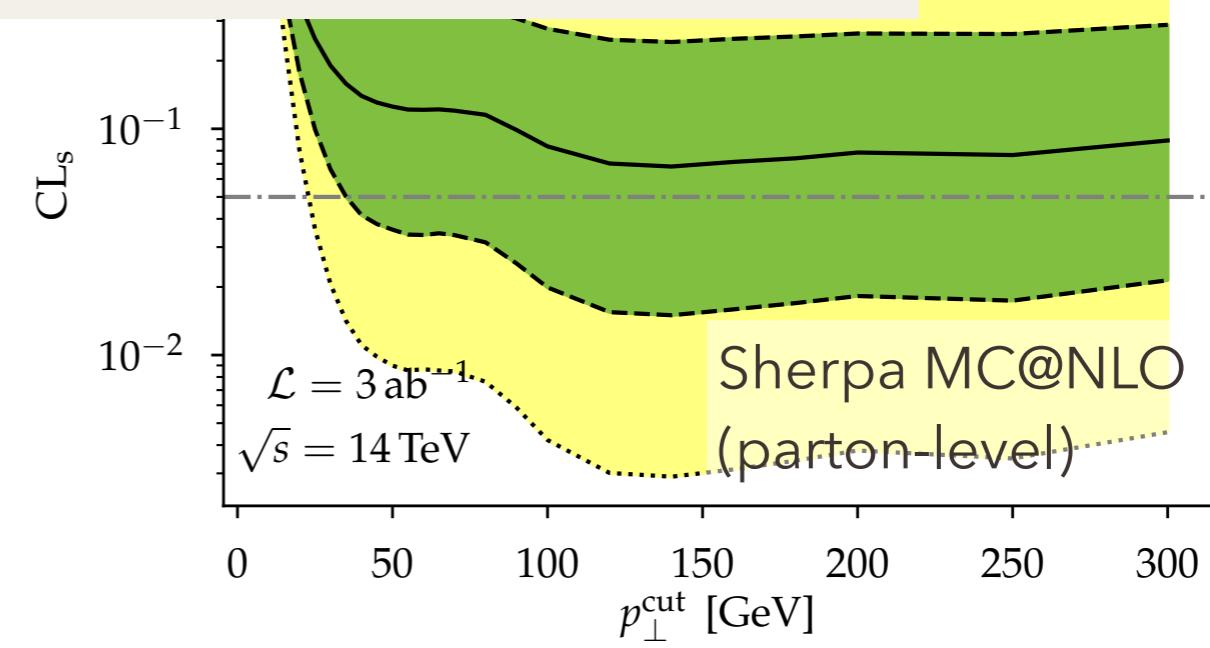
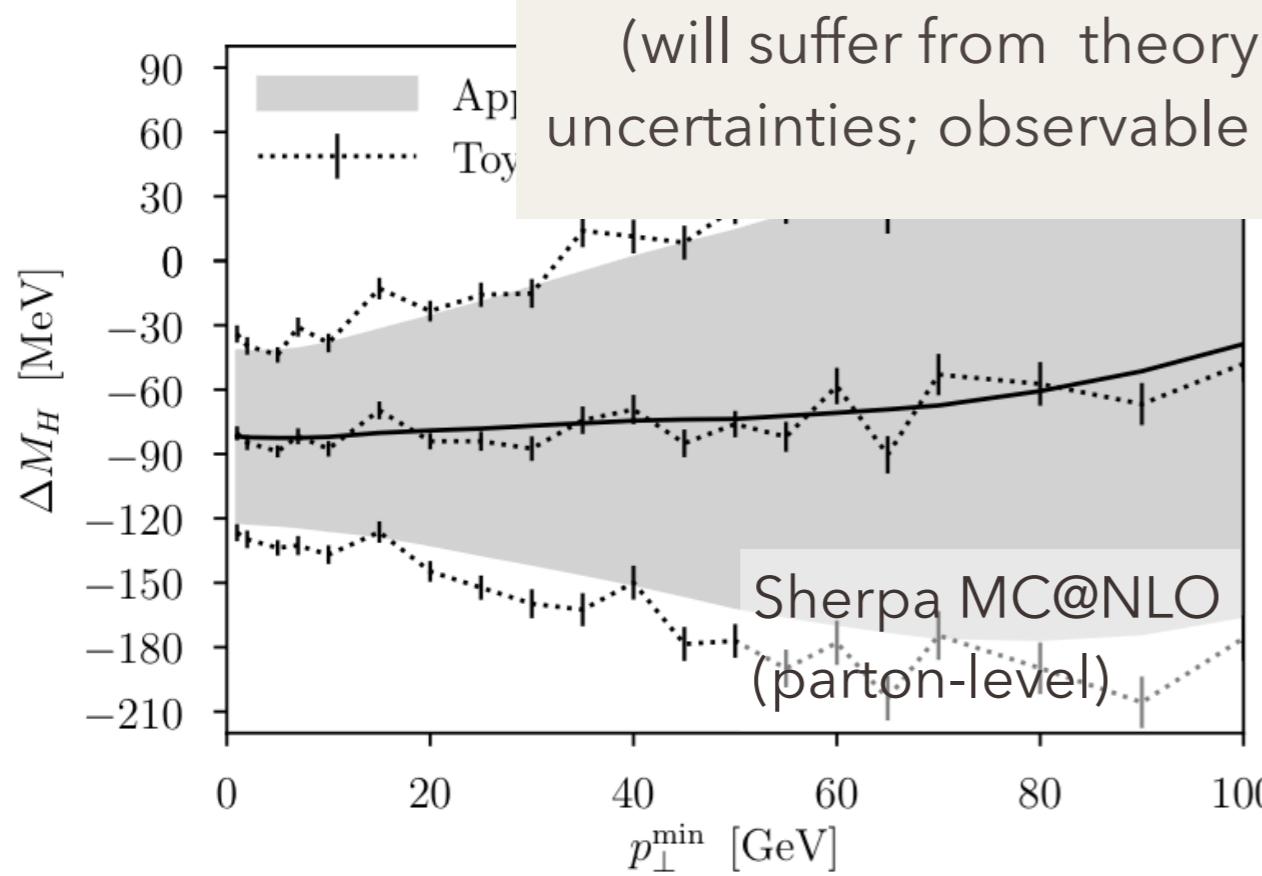
p_T extraction for fixed-order & resummed



p_T extraction for fixed-order & resummed



→ fixed-order bound $\Gamma_H \leq 15 \Gamma_H^{\text{SM}}$ degrades after
resummation to $\Gamma_H \leq 32 \Gamma_H^{\text{SM}}$



Go back to the $m_{\gamma\gamma}$ distribution?

- ▶ can we just go back to the $m_{\gamma\gamma}$ distribution and fit something that includes the shape distortion? \leadsto all data in fiducial region
- ▶ convolution of Lorentzian with Gaussian
 \Rightarrow Faddeeva function: $w(z) = e^{-z^2} \operatorname{erfc}(-iz)$

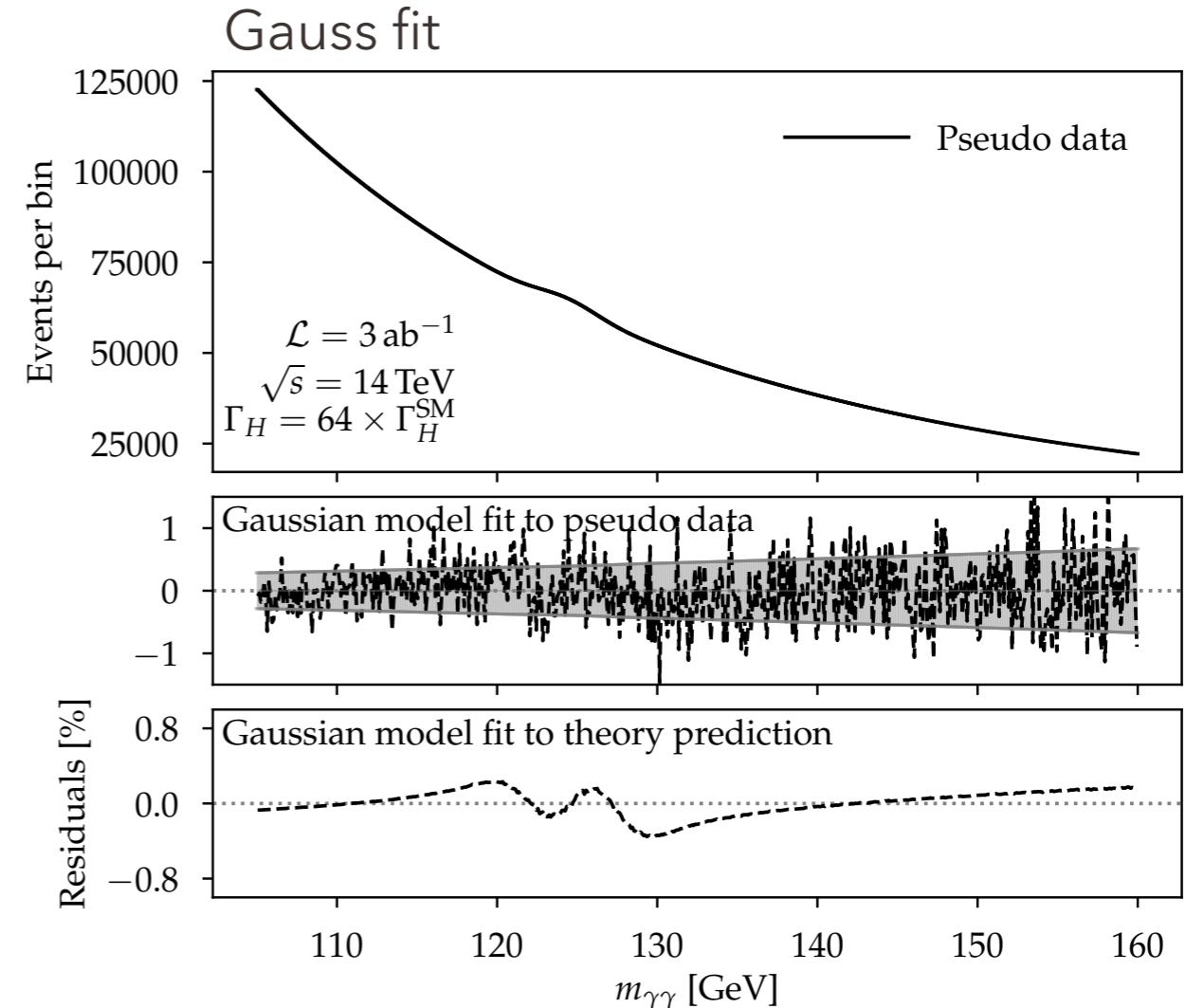
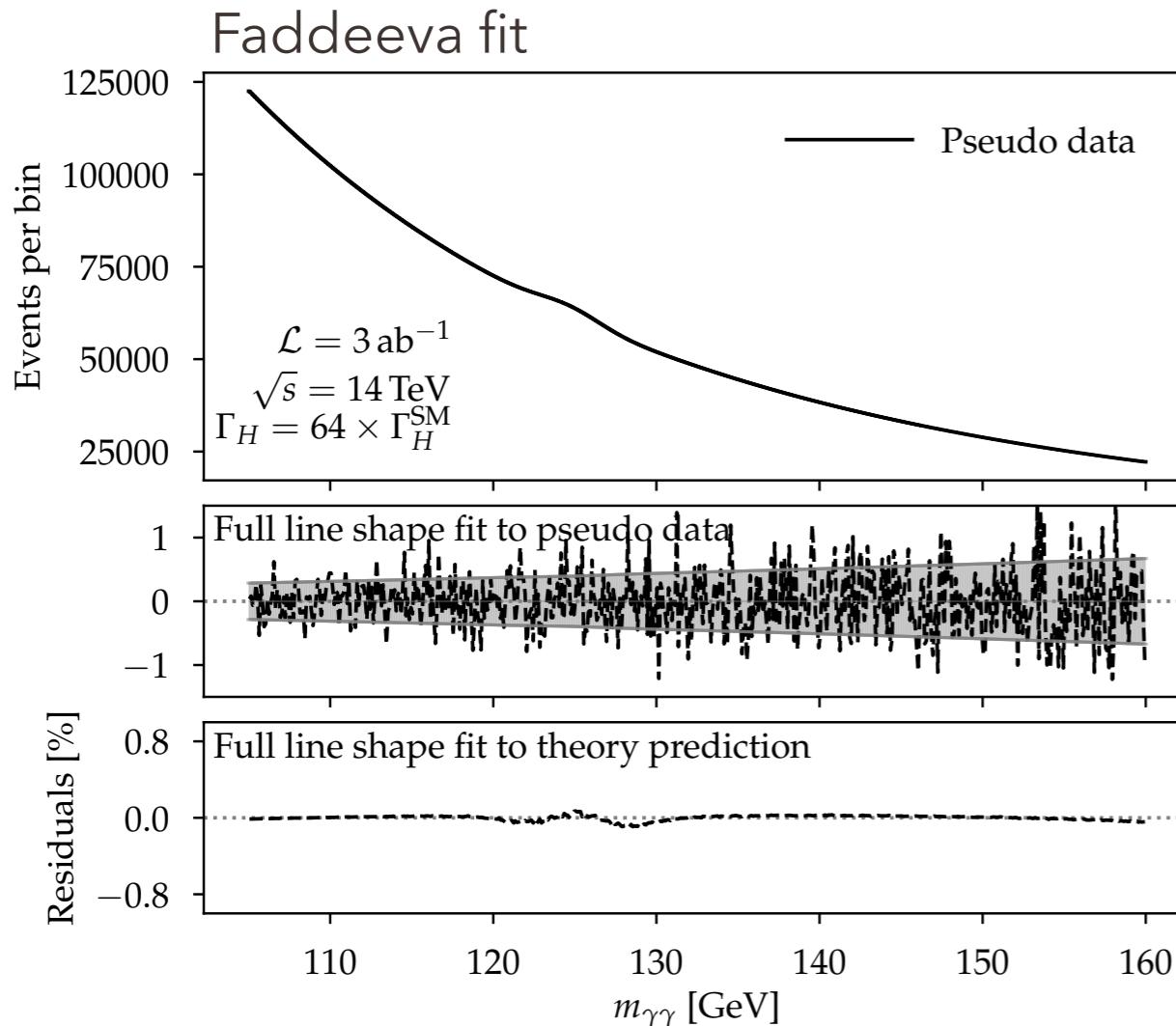
$$\mathcal{S} = \frac{w(z_-) - w(z_+)}{2\sqrt{2\pi}\sigma} \quad \text{with} \quad z_{\mp} = \frac{m_{\gamma\gamma} \mp M_H}{\sqrt{2}\sigma}, \quad M_H = \sqrt{m_H^2 - i m_H \Gamma_H}$$

$$\mathcal{F} = N \left[\frac{\operatorname{Re}\{\mathcal{S}\}}{\operatorname{Re}\{\mathcal{N}\}} + N_{RS} \frac{\operatorname{Im}\{\mathcal{S}\}}{\operatorname{Im}\{\mathcal{N}\}} \right] \quad \text{where} \quad N_{RS} = \sigma_R \left(\sigma_S c_{g\gamma} \frac{\Gamma_{H,\text{SM}}}{\Gamma_H} + \sigma_I \right)^{-1}$$

- ▶ sole theoretical input: $\sigma_R, \sigma_S, \sigma_I$
- ▶ fit to (MC) data

GOF comparison for Faddeeva vs. Gaussian

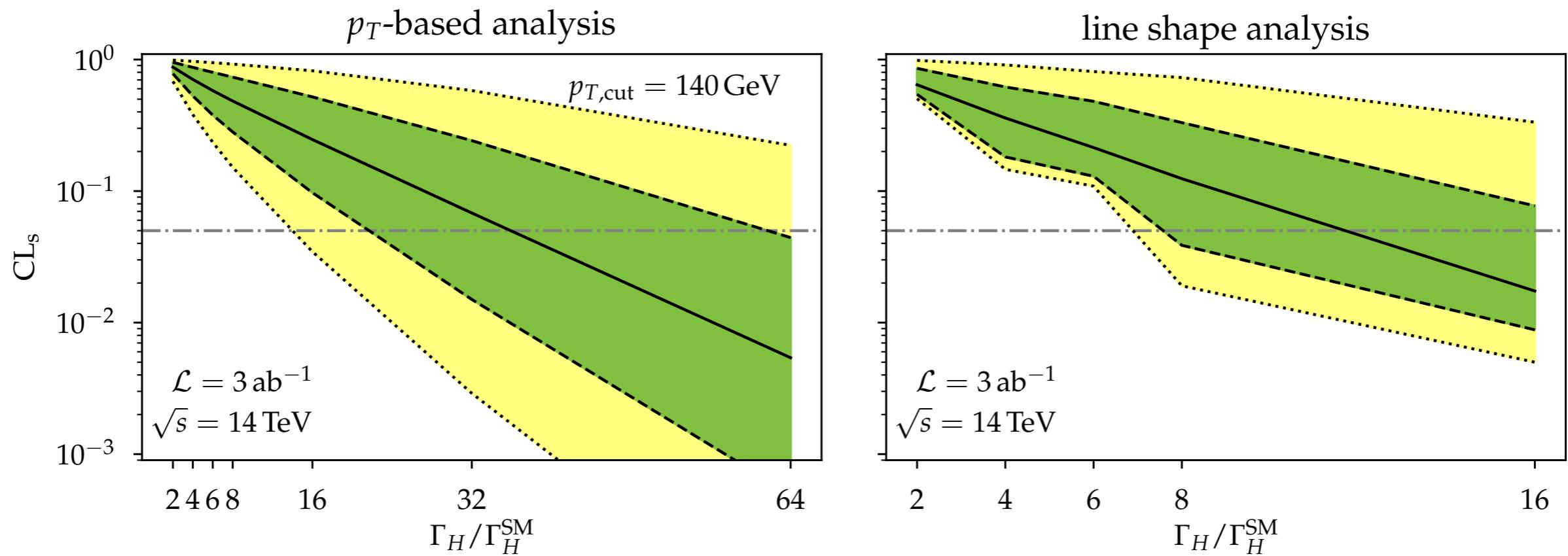
preliminary Sherpa MC@NLO (parton-level)



→ Residuals reduced by factor > 4 by using Faddeeva function

Results for both methods

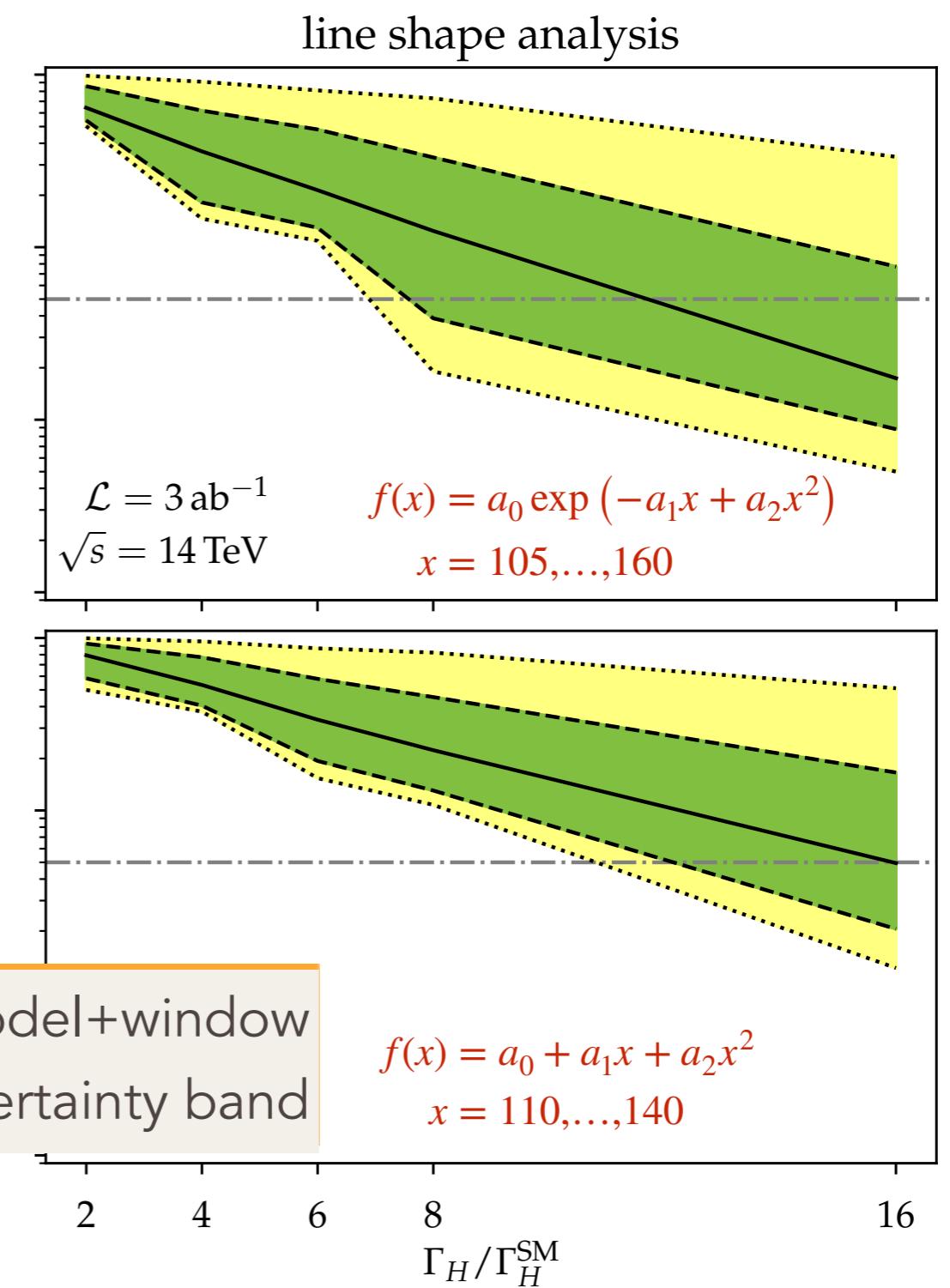
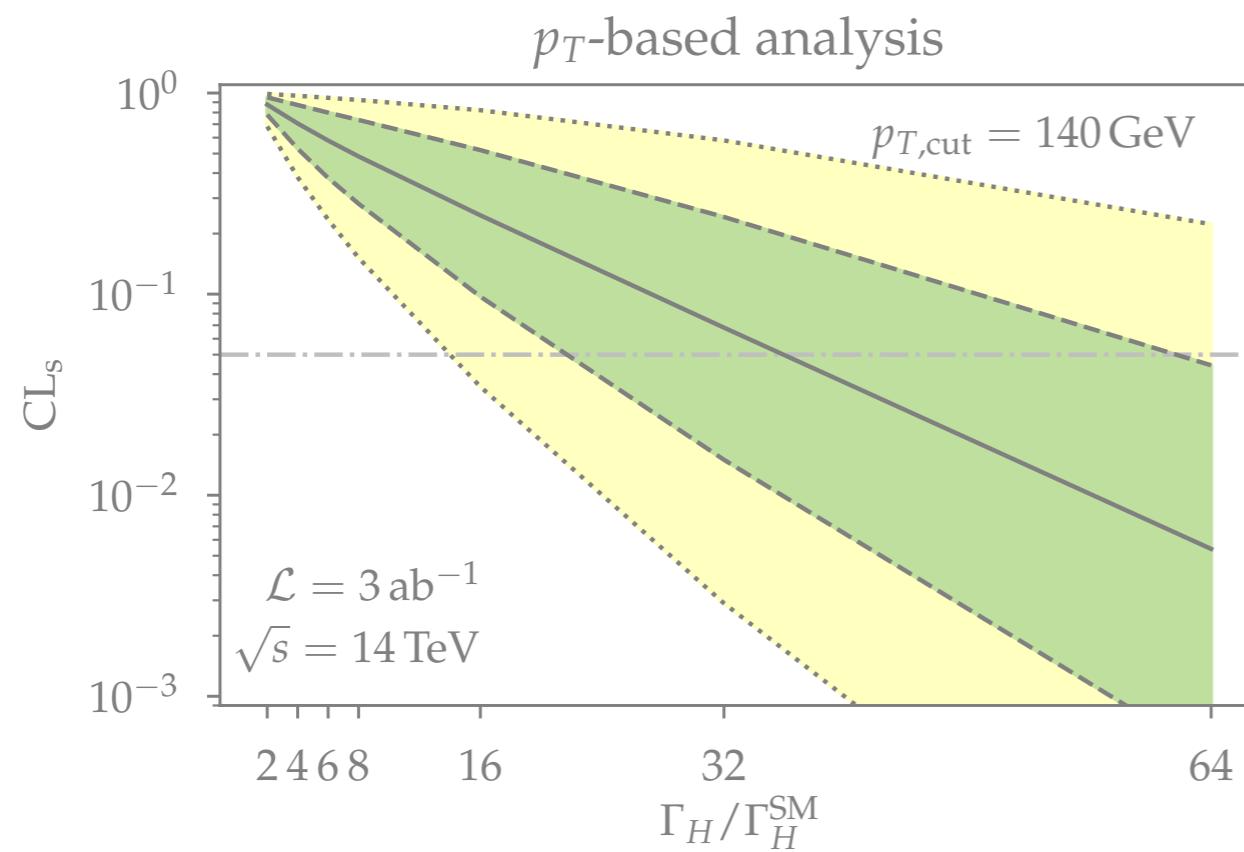
preliminary Sherpa MC@NLO (parton-level)



- ~ HL-LHC bound using p_T cut method: $\Gamma_H \lesssim 32 \Gamma_H^{\text{SM}}$
- ~ HL-LHC bound using direct-fit method: $\Gamma_H \lesssim 12 \Gamma_H^{\text{SM}}$

Line shape method: dependence of BG fit

preliminary Sherpa MC@NLO (parton-level)



→ dependence on BG fit model+window
within 1σ uncertainty band

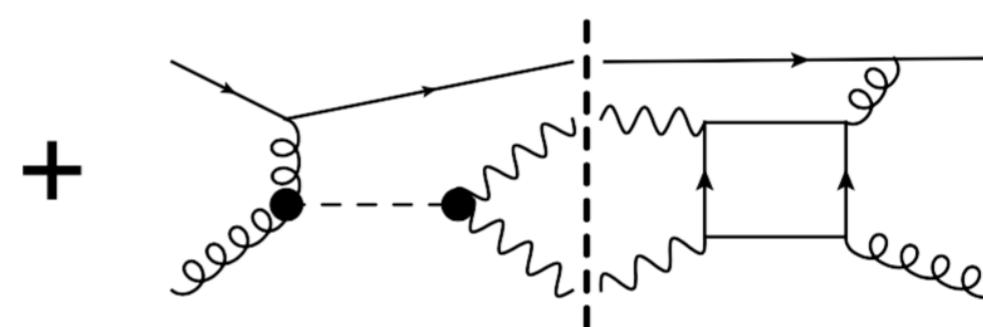
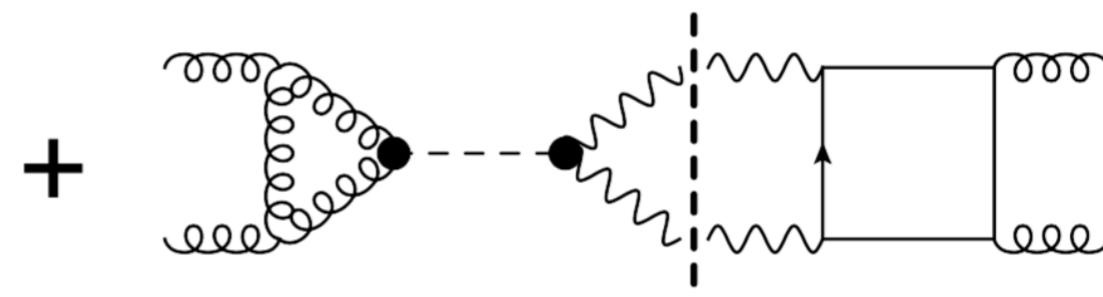
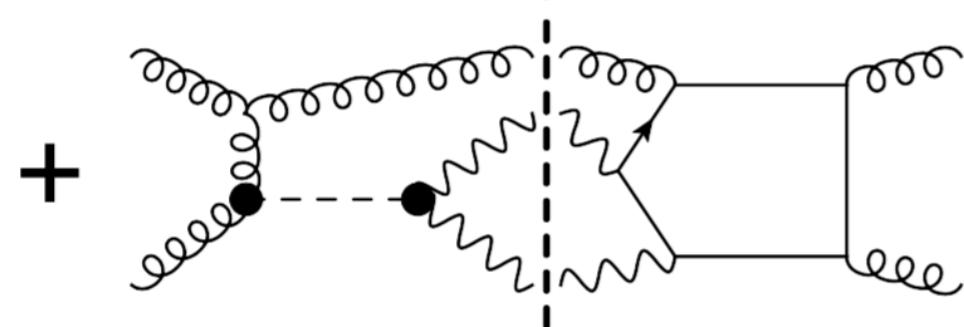
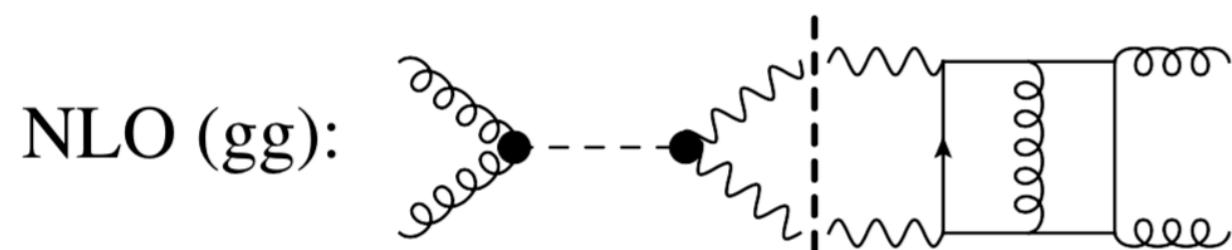
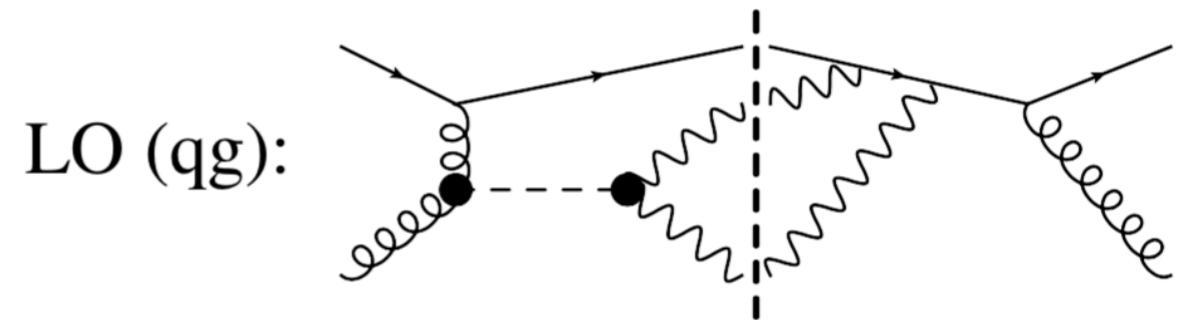
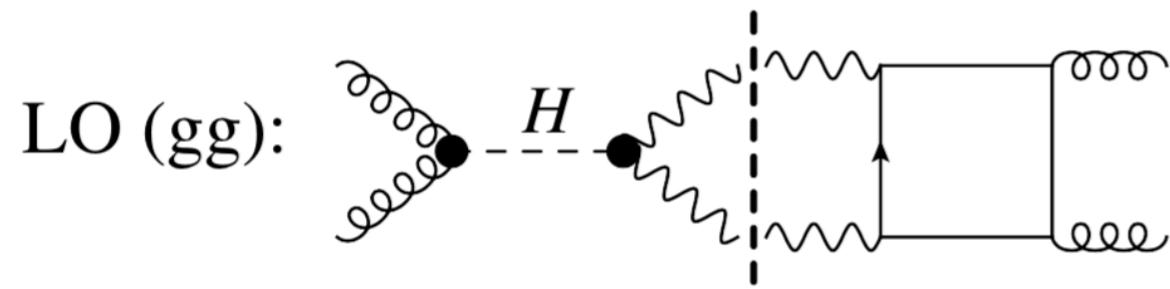
Conclusions

- ▶ interference-induced Higgs peak shift
- ▶ extract the shift \Rightarrow model-independent bound on Γ_H
- ▶ HL-LHC bounds (preliminary)
 - ▶ e.g. by comparing shift in high-/low $H p_T$
 - ▶ fixed-order bound $\Gamma_H \leq 15 \Gamma_{H}^{\text{SM}}$ degrades after resummation to $\Gamma_H \leq 32 \Gamma_{H}^{\text{SM}}$
 - ▶ or by directly fitting distorted peak in $m_{\gamma\gamma}$
 - ▶ optimistic to get $\Gamma_H \leq 12 \Gamma_{H}^{\text{SM}}$
 - ▶ TODO:
 - ▶ use Crystal-Ball function for signal fits instead of Gaussian

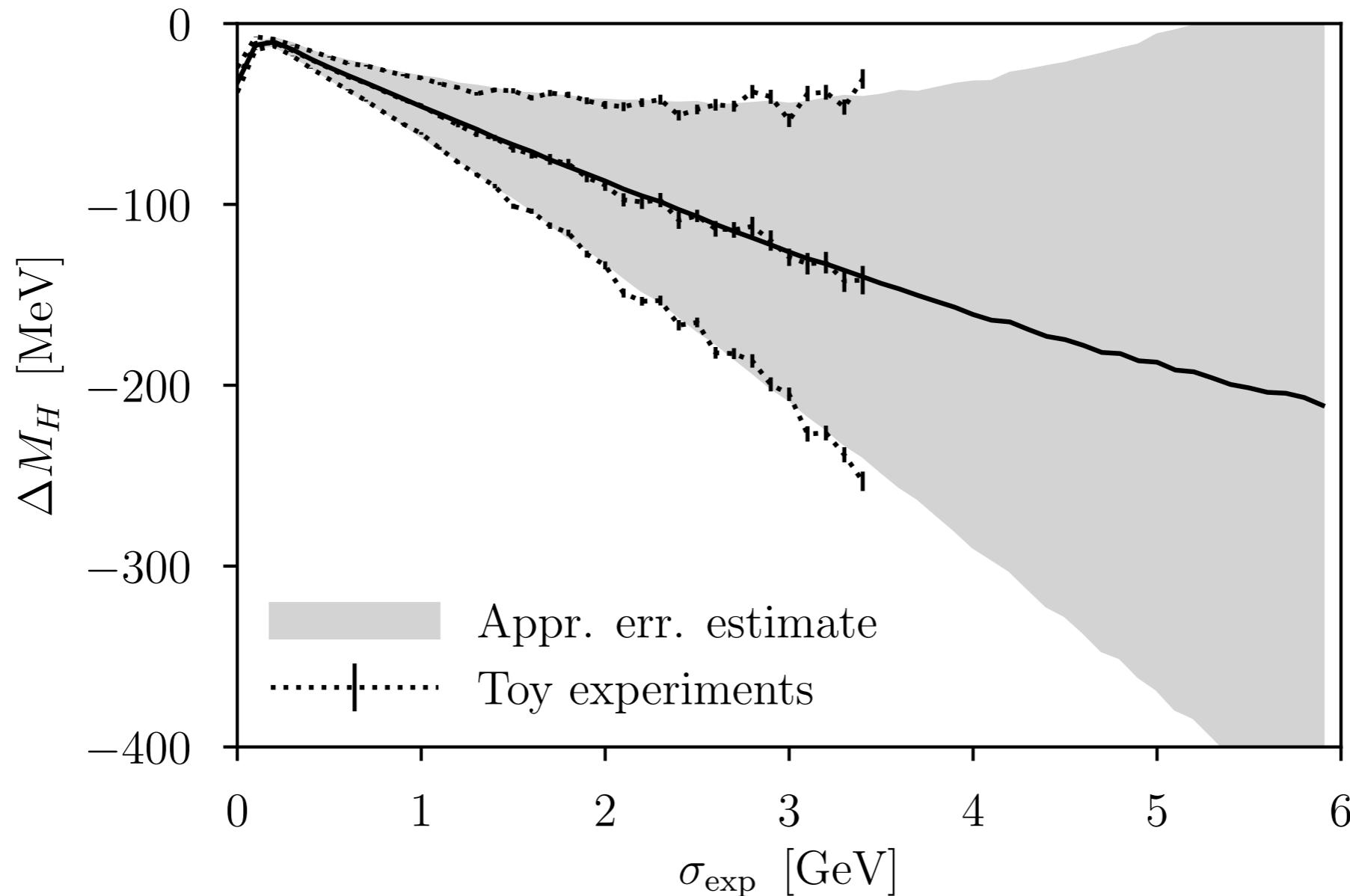
Back-up

Interference contributions

[Dixon 1305.3854]

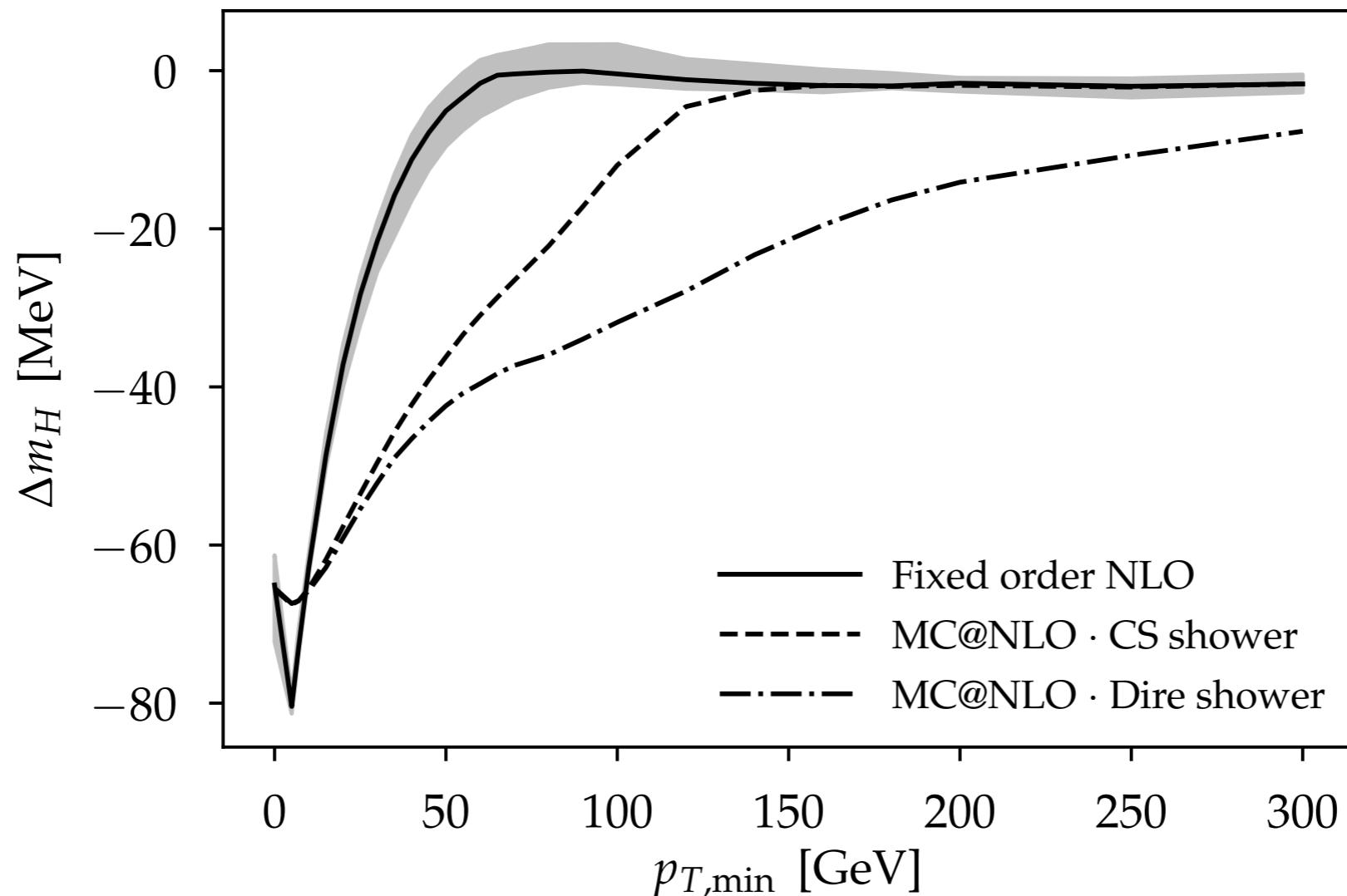


Reducing the experimental resolution



→ larger exp. uncert. gives larger shift, but at some point fitting the washed out peak comes with large uncertainties itself

Matching uncertainties

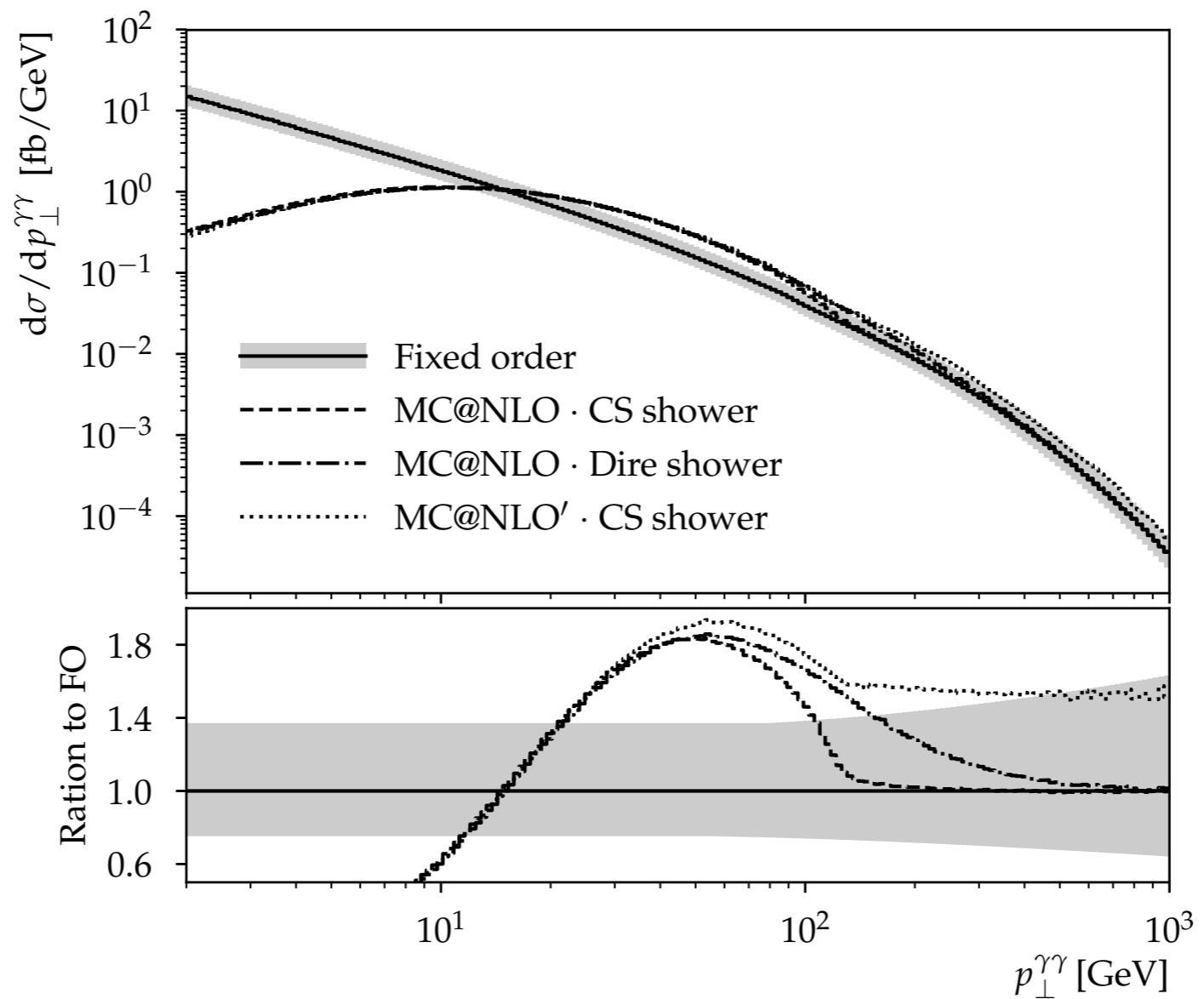


↪ large matching uncertainties, can be traced back to the large radiative corrections to the signal at NLO

NLO „fudge“ factor for real-emission events

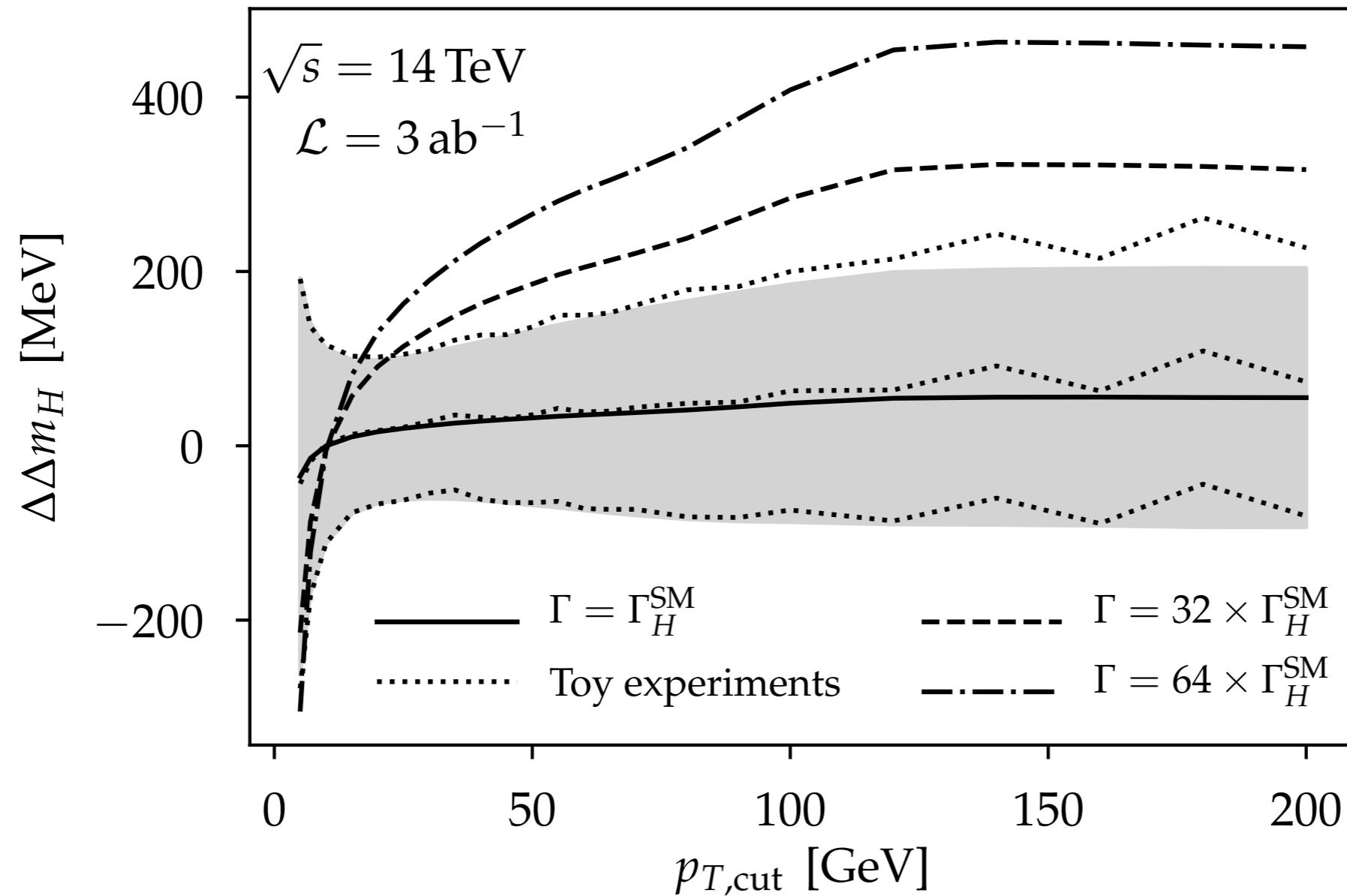
$$\left| \frac{\Gamma_a(Q^2)}{\Gamma_a(-Q^2)} \right|^2 = 1 + \frac{\alpha_s(Q^2)}{2\pi} C_a \pi^2 + \mathcal{O}(\alpha_s^2)$$

[Magnea, Sterman Phys.
Rev. D42, 4222 (1990)]



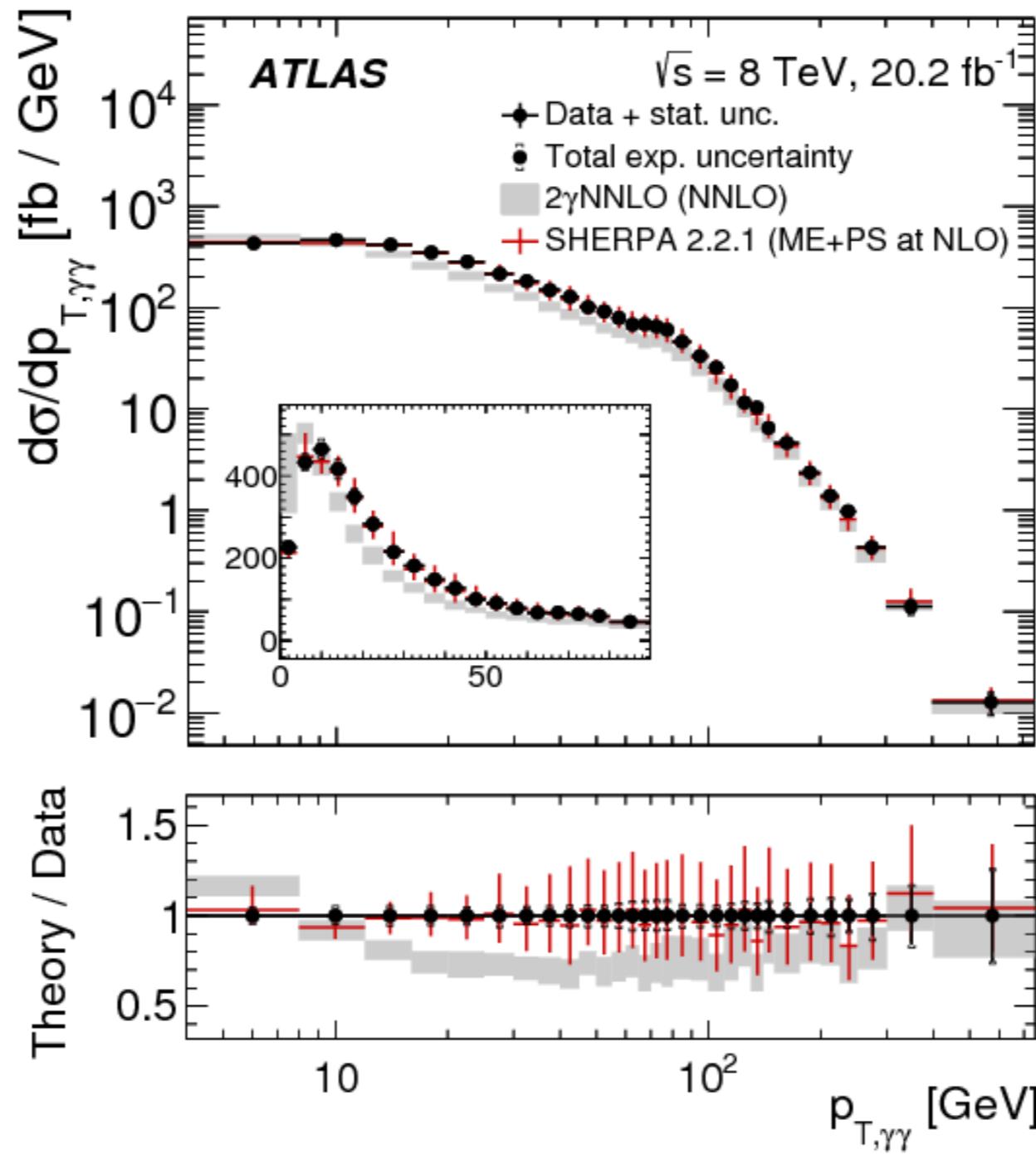
include universal higher-order corrections in all components
of the NLO calculation and subtracted the overlap

Difference of mass shift above vs. below $p_{T,\text{cut}}$



Background prediction

[ATLAS 1704.03839]



Background fit

