Constraining the Sea Quark Distributions Through W[±] Cross Section Ratio Measurements at STAR

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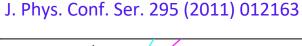


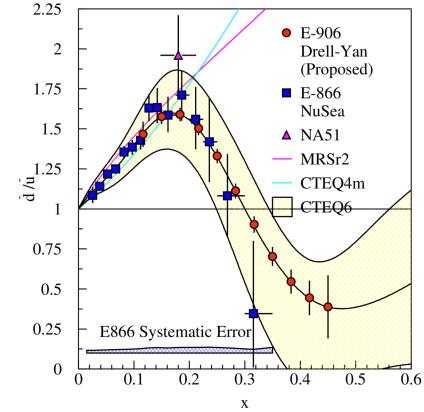




Motivation

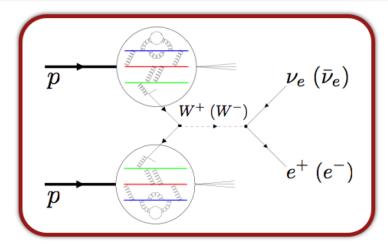
- \circ Unpolarized $\frac{\overline{d}}{\overline{u}}$ distribution can be probed via Drell-Yan production.
- E-866 suggests a trend where the $\frac{\overline{d}}{\overline{u}}$ ratio appears to be decreasing at large-x.
- The SeaQuest (E-906) will probe the sea quark distribution using Drell-Yan at higher x and lower Q² than E-866.
- More direct and indirect data are needed at high-x to help constrain the sea quark distributions.
- New measurements from different experiments can provide data at different Q² and from different scattering processes.
 - This will allow for understanding different systematic effects and also serve as a cross check of our understanding of the physics.







W Boson Production Through p+p Collisions



W bosons are sensitive to quark/anti-quark
distributions. They can be accessed via the W
leptonic decay channels in proton + proton collisions

$$\triangleright u + \bar{d} \rightarrow W^+ \rightarrow e^+ + \nu$$

$$\triangleright d + \bar{u} \rightarrow W^- \rightarrow e^- + \bar{\nu}$$

- The charged W cross section ratio
 - \rightarrow is proportional (at LO) to the $\frac{\bar{d}}{u}$ ratio
 - can be used to constrain the sea quark distributions

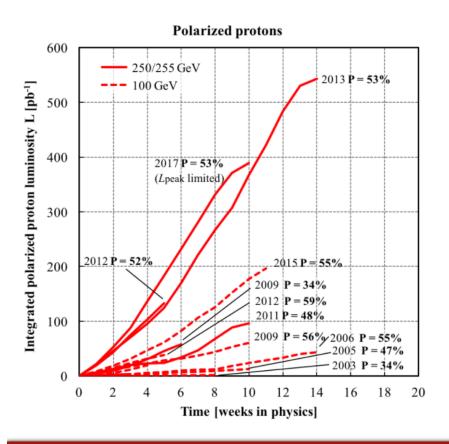
$$\frac{\sigma_{W^{+}}}{\sigma_{W^{-}}} \approx \frac{u(x_{1}) \, \bar{d}(x_{2}) + u(x_{2}) \bar{d}(x_{1})}{d(x_{1}) \bar{u}(x_{2}) + d(x_{2}) \bar{u}(x_{1})}$$

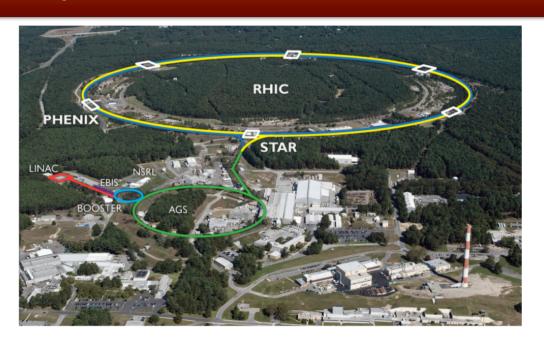
$$\frac{\sigma_{W^+}}{\sigma_{W^-}} = \left(\frac{N_O^+ - N_B^+}{N_O^- - N_B^-}\right) \left(\frac{\epsilon^-}{\epsilon^+}\right)$$

- +/- is positron/electron from W leptonic decay
- N_O is number of observed W events
- O N_B is number of background events
- ϵ is the W detection efficiency



- RHIC is the world's first polarized hadron collider
- Over the past several years luminosity at RHIC has steadily increased

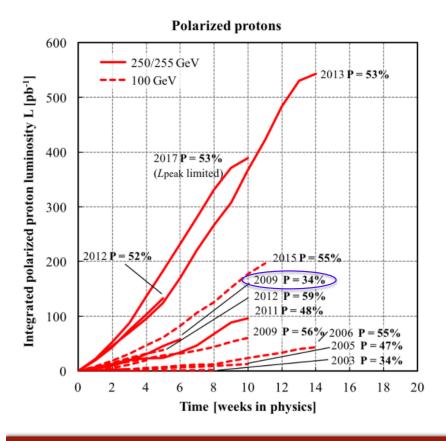


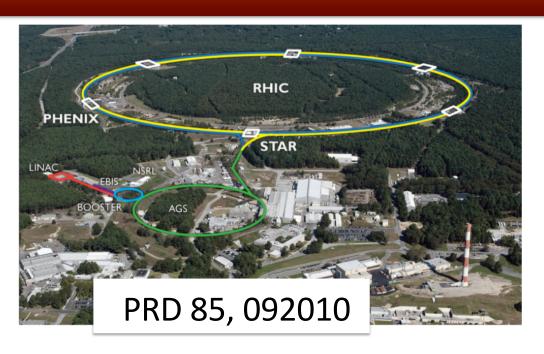


Run	\sqrt{s} (GeV)	Sampled Luminosity (pb ⁻¹)
9	500	10
11	500	25
12	510	75
13	510	250
17	510	350



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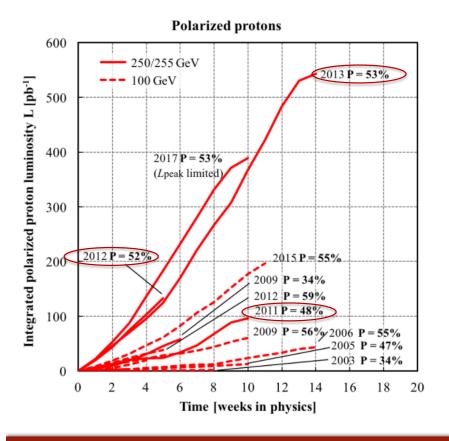


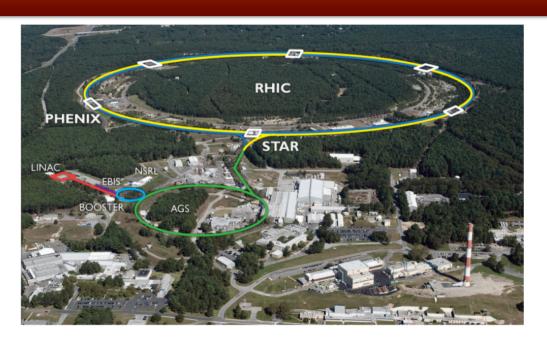


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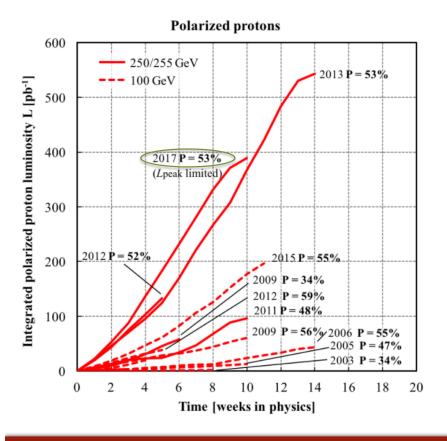


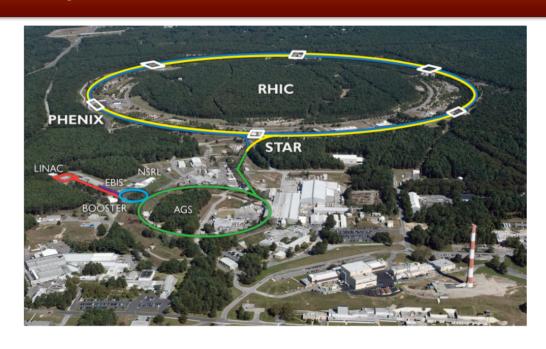


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Solenoid Tracker At RHIC

 \circ Calorimetry system with 2π coverage

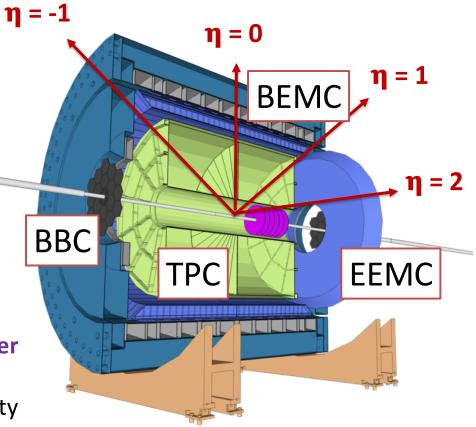
■ Barrel electromagnetic calorimeter (BEMC), $-1 < \eta < 1$

■ Endcap electromagnetic calorimeter (EEMC), $1 < \eta < 2$



• Provides tracking and particle ID $|\eta| < 1.3$

- Zero degree counter (ZDC), beam-beam counter (BBC), and vertex position detector (VPD)
 - Provides minimum bias trigger and luminosity monitors





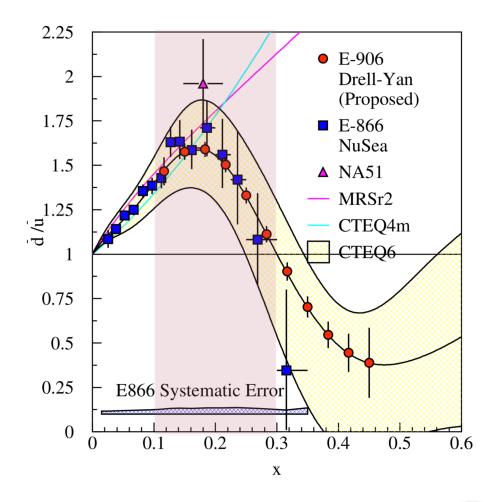
STAR Kinematics

Approximate kinematic range at STAR mid-rapidity (TPC + BEMC)

$$\triangleright$$
 0.1 < x < 0.3 for -1 < η < 1

- o For collision energies of $\sqrt{s} = 500 \text{ GeV}$ and $\eta = 0$, $(x_1 \approx x_2)$
 - $x = M_W/\sqrt{s} = 0.16$
- O Good complementarity to LHC (\sqrt{s} = 14 TeV) which probes much lower x

$$x = M_W/Vs = 5.7 \times 10^{-3} (x_1 \approx x_2)$$





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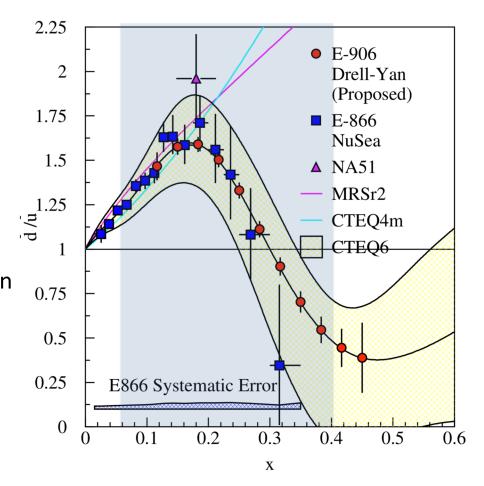
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O Good complementarity to LHC (\sqrt{s} = 14 TeV) which probes much lower x

$$\rightarrow$$
 x = M_W/ \sqrt{s} = 5.7×10⁻³ (x₁ ≈ x₂)

In STAR the **EEMC** could be used to obtain a more forward η -bin (1.1 < η < 2) which would extend the x reach of STAR

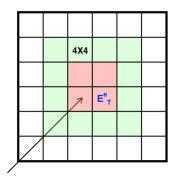
$$\triangleright$$
 0.06 < x < 0.4 for -2 < η < 2



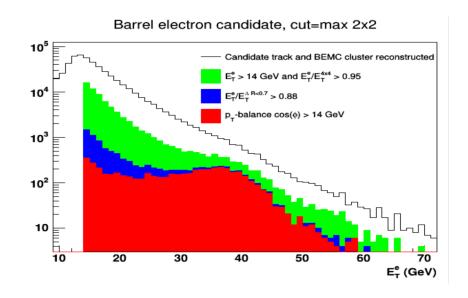


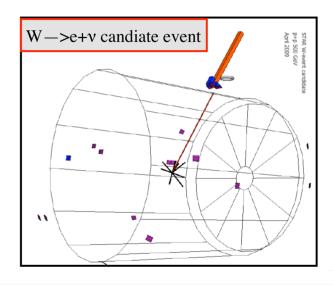
Selecting W Candidates

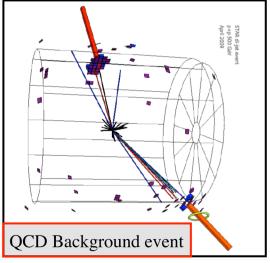
- Mid-rapidity STAR W selection criteria
 - Match p_T > 10 GeV/c track to BEMC cluster
 - Isolation ratio 1 /Isolation ratio 2
 - p_T-balance cut
 - Leads to good charge discrimination

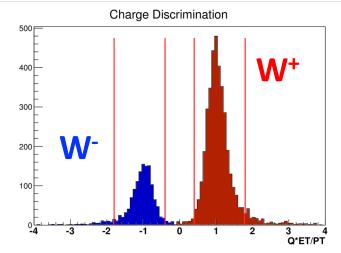


TPC track extrapolated to BEMC tower grid









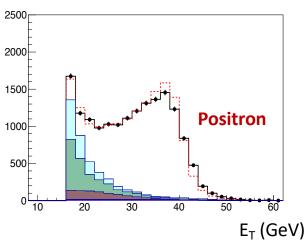


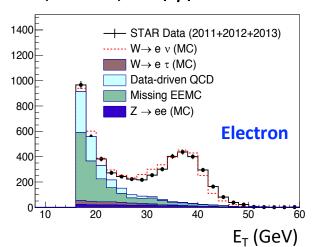
Backgrounds

W+ /W- signal and background distributions

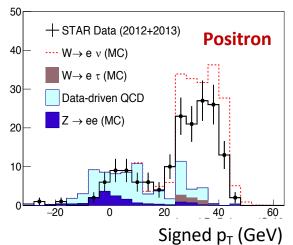
2011+2012+2013 (BEMC) $|\eta| < 1$

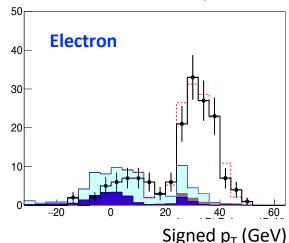
- Data-driven QCD background satisfies e^{+/-} isolation cuts
- Missing EEMC background results from backward "Jet" at non-existing calorimeter coverage for $-2 < \eta < -1.1$
- Missing EEMC background is estimated from EEMC located at $1.1 < \eta < 2$
- Electro-weak background from Z decay is estimated from PYTHIA/MC simulations.
- Small background contribution from Z decay.





2012+2013 (EEMC) $1.1 < \eta < 1.5$

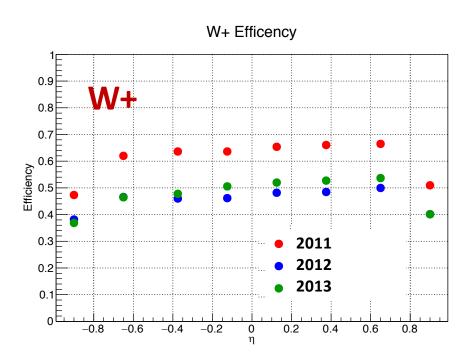


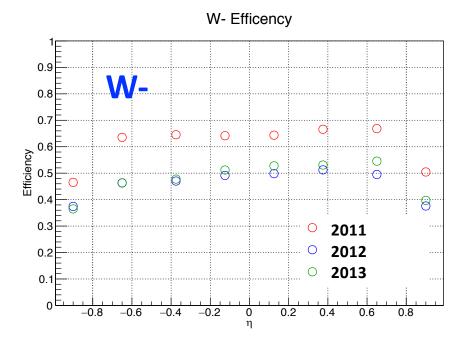




W Efficiencies

- Efficiencies computed using Pythia and GEANT.
 - 2012 and 2013 efficiencies decrease due to higher instantaneous luminosity,
 which leads to more pile-up and less efficient track reconstruction.
 - 2013 efficiencies are higher than 2012 due to new tracking algorithm (STICA).







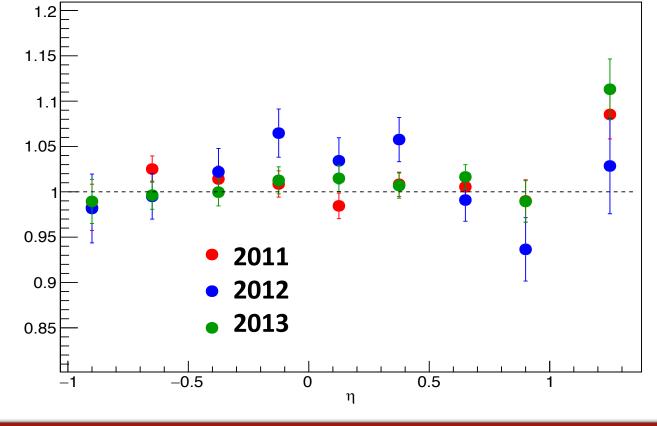
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Minimal charge dependence leads to small correction to the W cross section

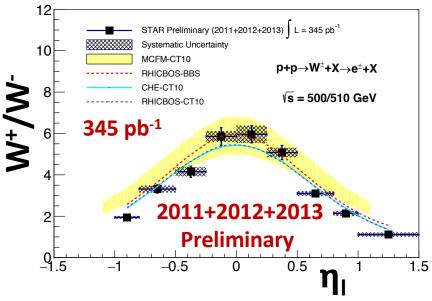
ratio.

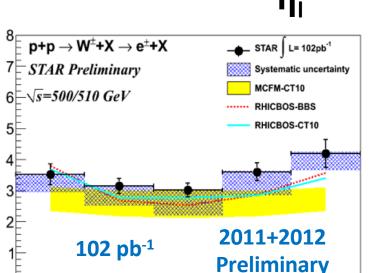
W-/W+ Efficiency





W Cross Section Ratio





- STAR 2017 W production data is expected to add 350 pb⁻¹ more data.
- Final systematic uncertainties will be reduced for W cross section ratio vs. lepton pseudorapidity compared to preliminary result.
- Impact on PDF distributions currently under investigation.
- The W boson rapidity can now also be reconstructed at STAR via its recoil (used for 2011 transverse single-spin asymmetry measurement, Phys.Rev.Lett. 116 (2016)).
- Work is ongoing to improve the systematic uncertainty associated with the reconstructed W boson rapidity.



-0.5

0.5

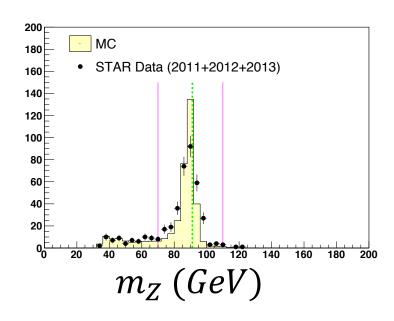
Уw

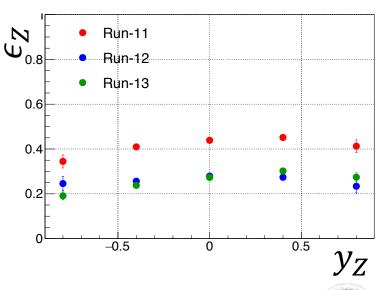
Selecting Z Candidates

 STAR is able to reconstruct Z boson candidates via their leptonic decays

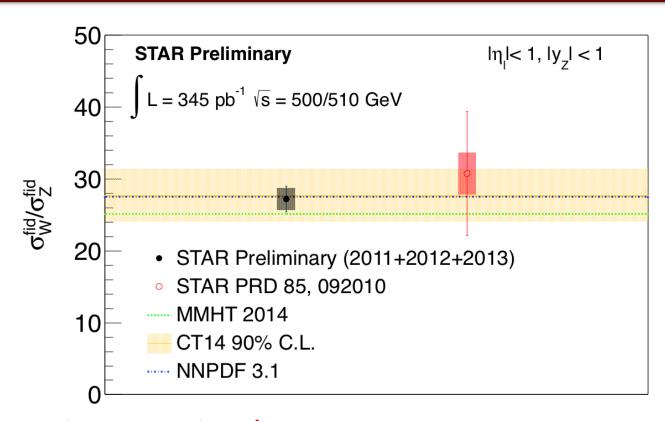
$$p + p \rightarrow \frac{Z}{\gamma^*} \rightarrow e^+ + e^- + X$$

- Z candidates are selected by using isolated e^{\pm} sample and requiring a pair of isolated e^{\pm} candidates to have opposite charge.
 - The invariant mass of each e^+e^- pair can be reconstructed.
 - Final Z candidates selected using an invariant mass cut of $70~GeV < m_Z < 110~GeV$
- Reconstruction of two charged tracks lead to cleaner identification of Z candidates.
- Efficiencies (ϵ_Z) and background are estimated using Pythia and GEANT
 - No background correction applied.





W/Z Cross Section Ratio



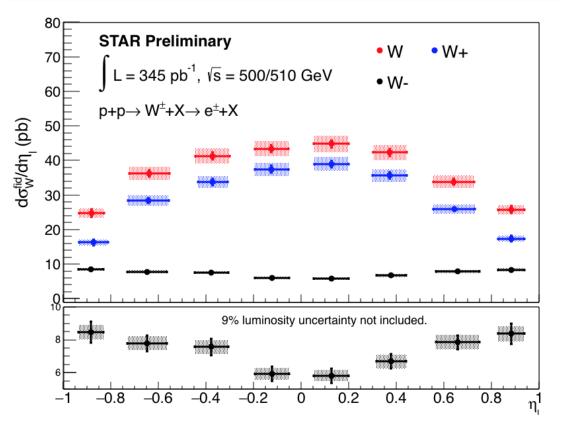
Can be used to measure the W/Z cross section ratio

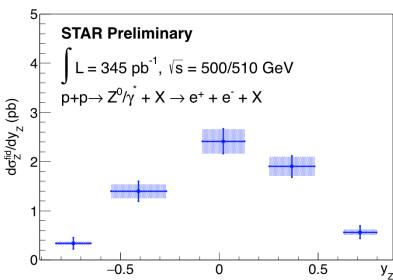
$$\frac{\sigma_W^{fid}}{\sigma_Z^{fid}} = \frac{N_O^W - N_B^W}{N_Z^O} \cdot \frac{\epsilon_Z}{\epsilon_W}, \text{ where } W \text{ is the total } W (W^+ + W^-)$$

- W/Z cross section ratio in great agreement with various PDF sets (computed with FEWZ).
- Consistent with previous STAR result based on 2009 data.
- Will help provide further constraints to PDFs.



Differential Cross Sections





 \circ Including the luminosity (L) information, one can also measure the differential cross sections

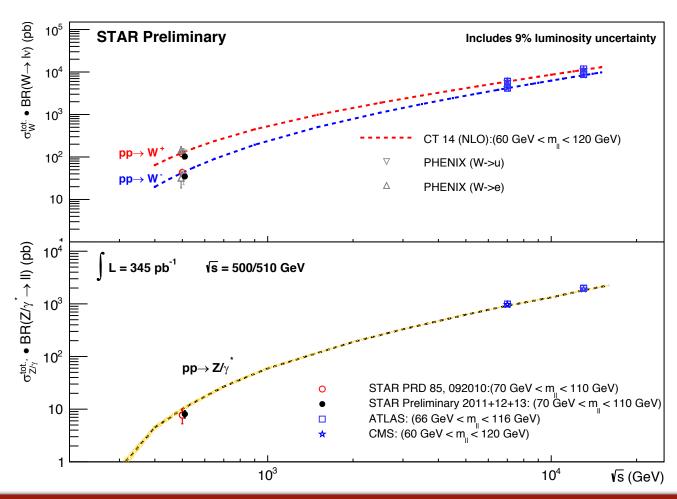
$$\frac{d\sigma^{fid}}{d\eta} = \frac{(N_O - N_B)}{L \cdot \epsilon}$$

- Work is currently being done in hopes of reducing the two dominant systematic uncertainties
 - Tracking efficiency: 5% for W's (10% for Z's)
 - Luminosity: 9%



Total Cross Sections

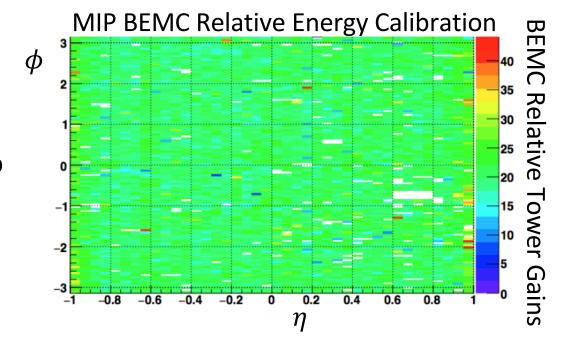
- The total cross sections can be computed from the measured fiducial cross sections by correcting for STAR acceptance.
- Acceptance correction computed using FEWZ
- \circ Preliminary results are consistent with world p + p data and theory.





STAR 2017 Analysis Update

- STAR 2017 p^{\uparrow} + p data set collected 350 pb⁻¹
 - Sivers function (via W,Z A_N)
 - Drell-Yan
 - W/Z cross sections and cross section ratios
- W and Z production data calibration, QA, and analysis is underway
- Offline BEMC calibration is now wrapping up
 - Initial tower QA has been done
 - Relative tower calibration completed using MIPs
 - Absolute energy calibration underway using electron E/p





Summary

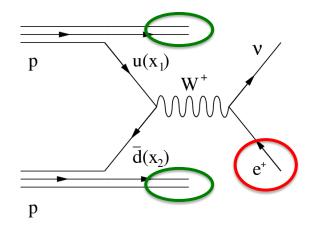
- STAR measured W+/- cross section ratio
 - A complementary measurement to SeaQuest and E-866 and LHC
 - Will help to further constrain the sea quark PDFs
- W/Z cross section ratio
 - A complementary measurement to LHC
 - Will help to constrain PDFs
- Impact of cross section ratios on the PDF distributions currently under analysis
- Preliminary W and Z differential and total cross sections were also presented
- Ongoing analysis from STAR 2017 W and Z production data will double the statistics of the STAR preliminary cross section and cross section ratio measurements.



Reconstruction W bosons

First developed at STAR for run 11 transverse single-spin asymmetry measurement of W bosons

Phys.Rev.Lett. 116 (2016)



Ingredients for the analysis

- Isolated electron
- neutrino (not measured directly)
- Hadronic recoil

W boson momentum reconstruction technique well tested at FermiLab and LHC

[CDF: PRD 70, 032004 (2004); ATLAS: JHEP 1012 (2010) 060]

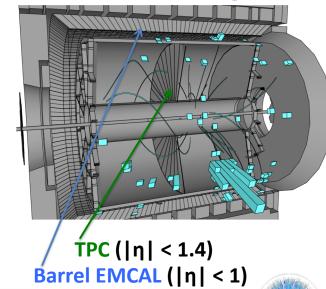
- Select events with the W-signature (STEP 1)
 - ➤ Isolated high P_T electron
- Neutrino transverse momentum is reconstructed from missing P_T (Step 2) $\vec{P}_T^{\nu} \approx -\sum \vec{P}_T^{i}$

 $i \in_{clusters}^{tracks}$

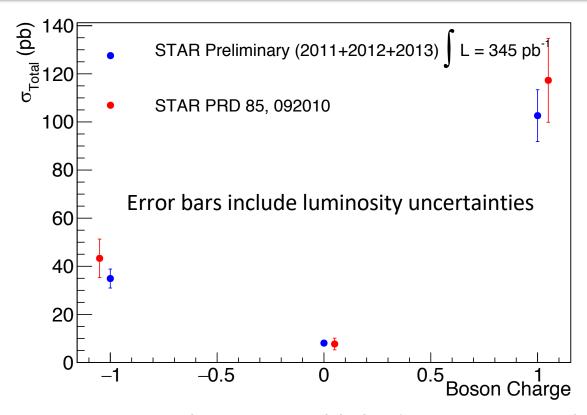
□ Neutrino's longitudinal momentum is reconstructed from the decay kinematics (Step 3)

$$M_W^2 = (E_e + E_v)^2 - (\vec{p}_e + \vec{p}_v)^2$$

The STAR detector @ RHIC



Total Cross Sections: STAR 2009 Comparison



- Measurements consistent between published cross sections based on STAR
 2009 data set and new preliminary results based on STAR 2011,2012, and 2013 data sets.
- Error bars include luminosity uncertainties
 - 2009: 13%
 - STAR Preliminary (2011+2012+2013): 9%

