

# On analytic solutions of 3+1 Relativistic Ideal Hydrodynamical Equations

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# Outline

- Basics of Relativistic Ideal Hydrodynamics and its applicability to Heavy Ion Collisions
- Reduction from 3+1 problem to 1+1 problem by embedding
- Flow profile of the embedding solutions and their physical interpretations
- Possible connections with heavy ion collisions

# Basics of relativistic hydrodynamics

Conservation equation  $T^{mn}{}_{;n} = 0$  (1)

Constitutive equation  $T^{mn} = (\epsilon + p) u^m u^n - p g^{mn} = w u^m u^n - p g^{mn}$  (2)

Equation of state  $p = p(\epsilon)$  (3)

Assumption: local equilibrium

If there is a conserved charge  $(n u^m)_{;m} = 0$  (4)

Thermodynamical relations  $\epsilon + p = Ts$  (5)

$$d\epsilon = T ds, dp = s dT$$

(1), (2), (5) lead to conservation of entropy

$$(s u^m)_{;m} = 0$$

no dissipative term  
respect time reversion

# Applicability of RIHD to HIC

## Phenomenology of Heavy Ion Collisions

QGP produced at heavy ion collisions is believed to be strongly coupled.

Lower bound for general strongly coupled gauge theory:

$$\frac{\eta}{s} \geq \frac{1}{4\pi}$$

Kovtun, Son, Starinets  
2004

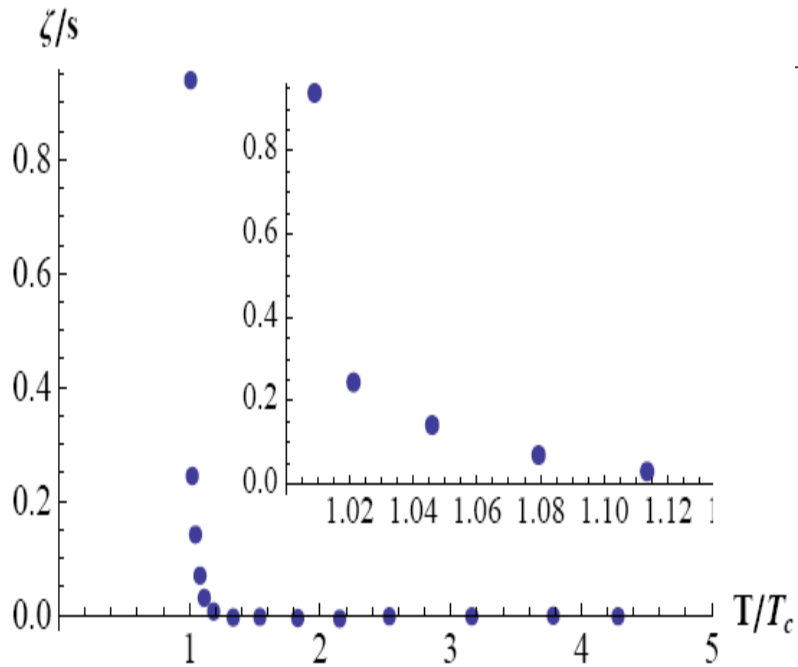
Even lower value for bulk viscosity at  $T > 1.1T_c$

$$\frac{\zeta}{s} \leq 0.04$$

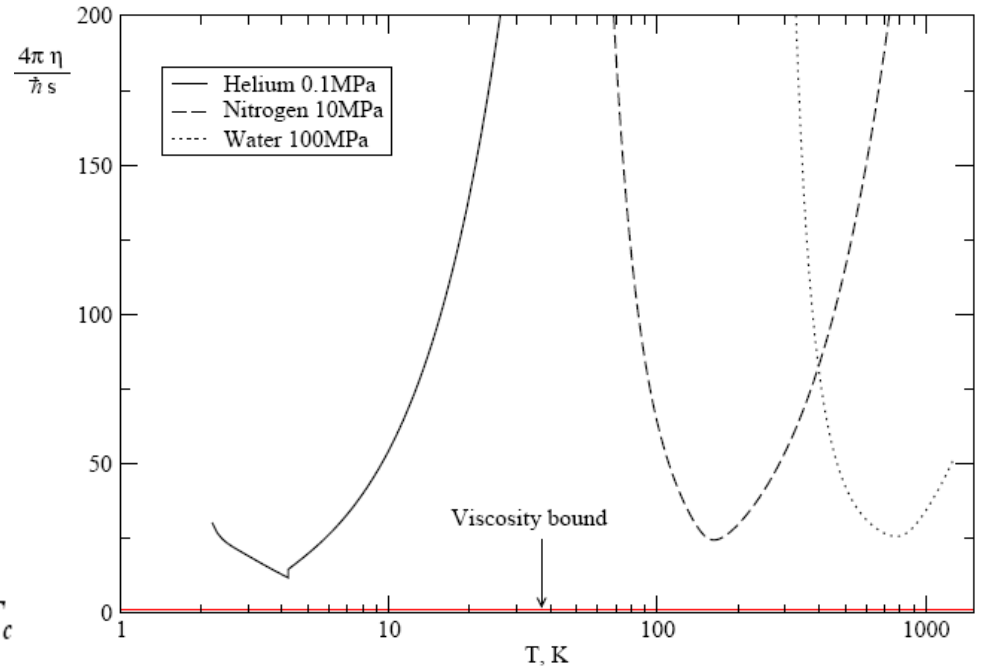
Kharzeev, Tuchin 2007

Relativistic Ideal Hydrodynamics applicable in wide region of temperature

# Bulk and Shear viscosity



QCD with lattice data



Conjecture for strongly coupled matter

# Some well known solutions to RIHD

Landau, Khalantnikov 1950s

1+1D rapidity distribution  
approximately gaussian

Hwa(1974)-Bjorken(1983)

1+1D boost invariant

Bialas, Janik, Peschanki 2007

1+1D interpolation  
between LK and HB

Biro 2000

generalization to 3+1D

Csörgő et al 2004

solution with spherical,

Nagy, Csörgő, Csanád 2008

cylindrical and ellipsoidal  
symmetries

# 2D Hubble embedding

Fluid flow  $u^\mu \rightarrow \gamma(1, v_x, v_y, v_z)$  In flat coordinate (t,x,y,z)

Solving hydrodynamical equations with specific Hubble-like transverse flow:

$$v_x = \frac{x}{t} = \tanh \eta_\perp \cos \phi_\perp, \quad v_y = \frac{y}{t} = \tanh \eta_\perp \sin \phi_\perp$$

$$t = \tau_\perp \cosh \eta_\perp, \quad z = z,$$

$$x = \tau_\perp \sinh \eta_\perp \cos \phi_\perp, \quad y = \tau_\perp \sinh \eta_\perp \sin \phi_\perp$$

$$u^m = \bar{\gamma}(1, 0, 0, \bar{v}_z)$$

$$\bar{v}_z \neq v_z$$

energy, pressure and longitudinal flow independent of  $\eta_\perp$  and  $\phi_\perp$

# scaling ansatz

Equation of state

$$p = \nu(\epsilon + p)$$

Speed of sound

$$c_s = \sqrt{\frac{\partial p}{\partial \epsilon}} = \sqrt{\frac{\nu}{1-\nu}} \leq 1, \quad 0 < \nu \leq 1/2$$

dimensionful

$$p(\tau_{\perp}, z) = \tau_{\perp}^{-\frac{a}{1-\nu}} g\left(\frac{z}{\tau_{\perp}}\right)$$

dimensionless

$$\bar{v}_z(\tau_{\perp}, z) = f\left(\frac{z}{\tau_{\perp}}\right)$$

scaling variable

$$\xi = \frac{z}{\tau_{\perp}}$$



# Symmetries of the EOM

$$f'(\xi) = -\frac{\nu(1 - f(\xi)^2)(2 - a + af(\xi)^2 - 2\xi f(\xi))}{(\nu - 1)(f(\xi) - \xi)^2 + \nu(\xi f(\xi) - 1)^2}$$

$$g'(\xi) = -\frac{g}{1 - \nu} \times \frac{-2f(\xi) + af(\xi) + 2\nu f(\xi) - 2a\nu f(\xi) + 2\xi - a\xi - 2\nu\xi + a\nu\xi + a\nu\xi f(\xi)^2}{(\nu - 1)(f(\xi) - \xi)^2 + \nu(\xi f(\xi) - 1)^2}$$

$$\xi \rightarrow -\xi, f \rightarrow -f, g \rightarrow g \quad \Leftrightarrow \tau_{\perp} \rightarrow \tau_{\perp}, z \rightarrow -z, v_z \rightarrow -v_z, p \rightarrow p$$

The solutions should preserve parity

# Solving the equations

$$f'(\xi) = -\frac{\nu(1 - f(\xi)^2)(2 - a + af(\xi)^2 - 2\xi f(\xi))}{(\nu - 1)(f(\xi) - \xi)^2 + \nu(\xi f(\xi) - 1)^2}$$

$$t(f) \equiv \frac{f \xi(f) - 1}{f - \xi(f)}$$

$$\begin{aligned} & [(1 - f^2)(af - 2t + at)] \times \frac{dt}{df} \\ & - [(\kappa^2 - a)f + (\kappa^2 + 2 - a)t + (a - 1)ft^2 + (a - 3)t^3] = 0 \end{aligned}$$

Linear ansatz  $t(f) = \alpha f + \beta$

Nonlinear ansatz  $t(f) = \alpha f + \beta + \frac{\rho f + \lambda}{1 - f^2}$

# Solutions for general $\nu$ (EOS)

$$\bar{v}_z = 0, p \propto \tau_{\perp}^{-2/(1-\nu)}$$

2D Hubble flow(analog of Hwa-Bjorken flow)

$$\bar{v}_z = \xi, p \propto \tau_{\perp}^{-3/(1-\nu)} (1 - \xi^2)^{\frac{-3}{2(1-\nu)}}$$

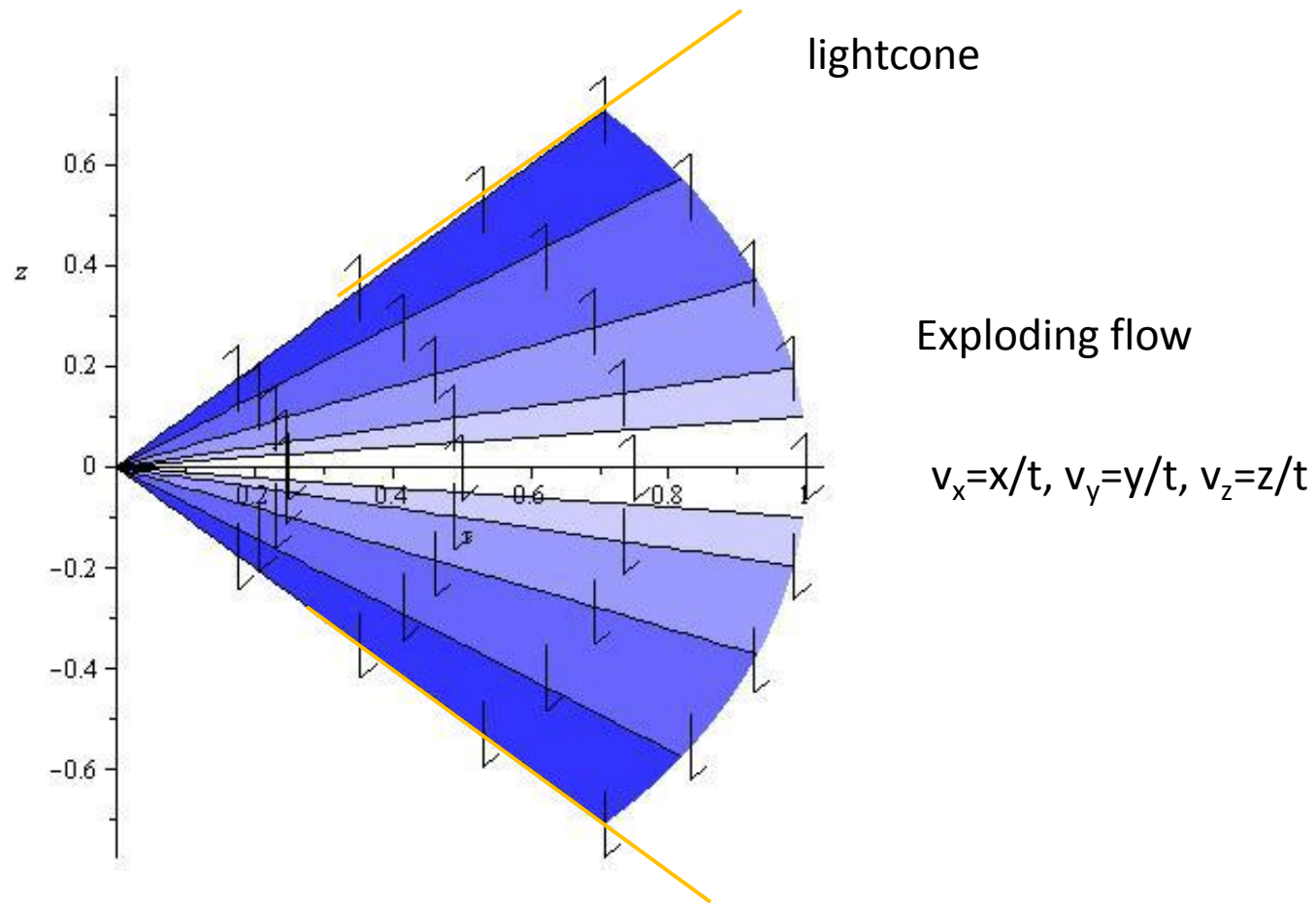
3D Hubble flow(spherical)  $|\xi| < 1$

$$\bar{v}_z = 1/\xi, p \propto \tau_{\perp}^{-1/\nu} (\xi^2 - 1)^{\frac{(3\nu-1)}{2\nu(1-\nu)}}$$

Anti-Hubble flow  $|\xi| > 1$

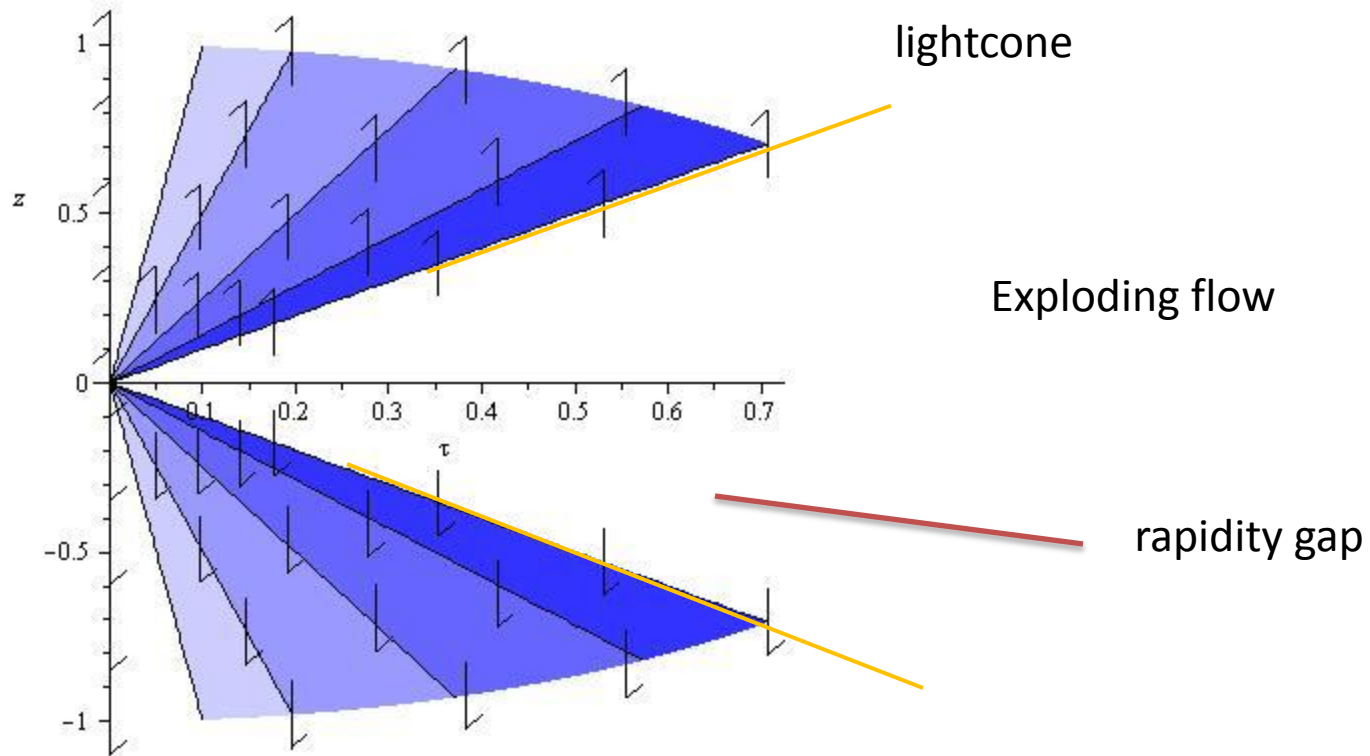
$$\xi = z / \tau_{\perp}$$

# 3D Hubble flow(spherical)



arrows indicate direction of the flow  
darkness of the color indicate the flow magnitude

# Anti-Hubble flow



# Solutions for general $\nu$ (EOS)

$$\bar{v}_z = f\left(\frac{z}{\tau_{\perp}}\right)$$

$$p = \text{constant} \times \tau_{\perp}^{-3} \frac{f^{\frac{1-3\nu}{1-\nu}}}{(1-f^2)^{\frac{2-3\nu}{1-\nu}}}$$

domain 1 and domain 3, domain 2 and domain 4 are related by parity!

$$\text{domain 1: } f = \frac{z/\tau_{\perp} + \sqrt{z^2/\tau_{\perp}^2 - 4\nu(1-\nu)}}{2(1-\nu)}$$

$$\text{with } 2\sqrt{\nu(1-\nu)} < \frac{z}{\tau_{\perp}} < 1$$

$$\text{domain 2: } f = \frac{z/\tau_{\perp} + \sqrt{z^2/\tau_{\perp}^2 - 4\nu(1-\nu)}}{2(1-\nu)}$$

$$\text{with } \frac{z}{\tau_{\perp}} < -1$$

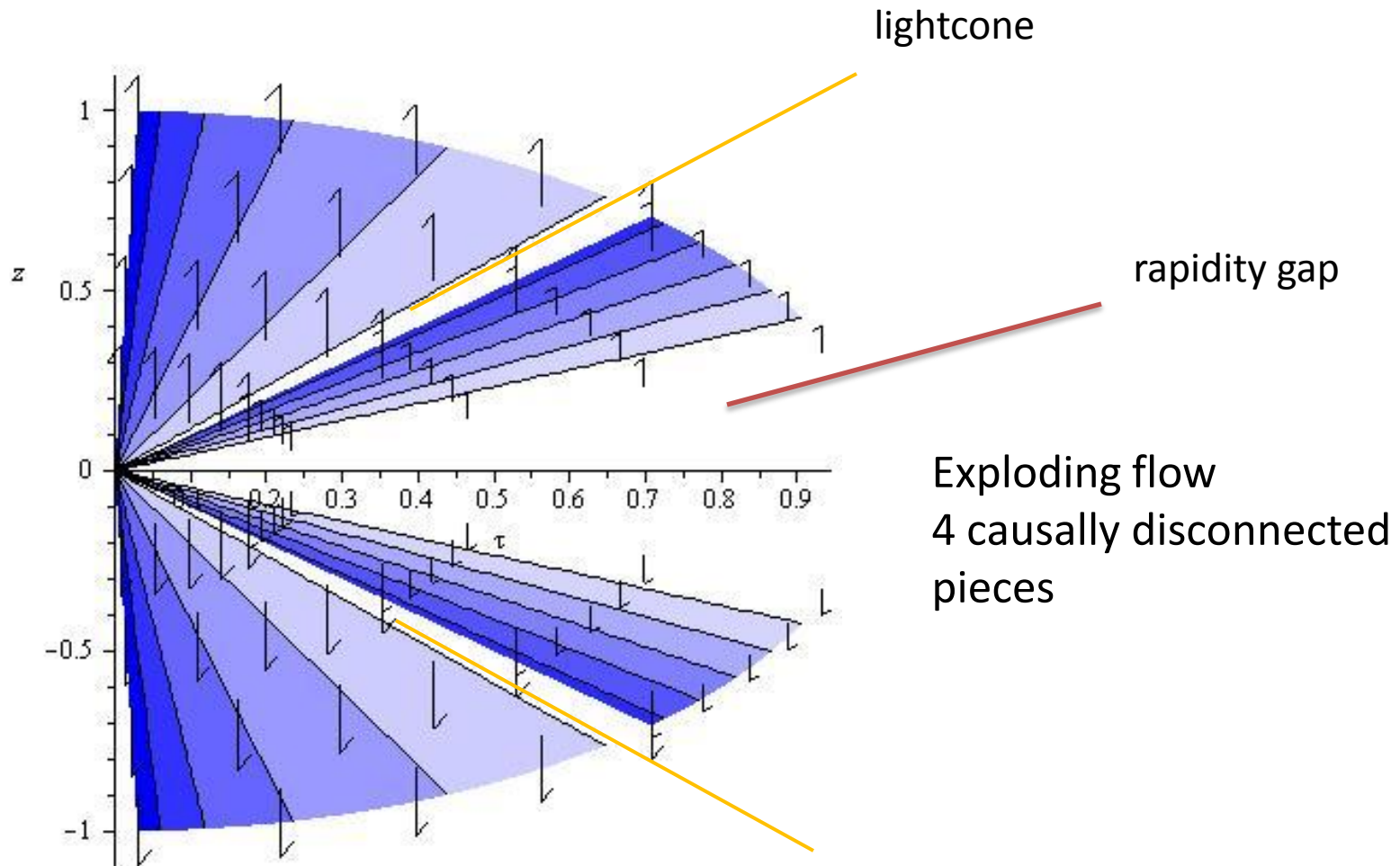
$$\text{domain 3: } f = \frac{z/\tau_{\perp} - \sqrt{z^2/\tau_{\perp}^2 - 4\nu(1-\nu)}}{2(1-\nu)}$$

$$\text{with } -1 < \frac{z}{\tau_{\perp}} < -2\sqrt{\nu(1-\nu)}$$

$$\text{domain 4: } f = \frac{z/\tau_{\perp} - \sqrt{z^2/\tau_{\perp}^2 - 4\nu(1-\nu)}}{2(1-\nu)}$$

$$\text{with } \frac{z}{\tau_{\perp}} > 1$$

# Solution with 4 domains



# Solutions for specific $\nu$ (EOS)

$$\bar{v}_z = f\left(\frac{z}{\tau_\perp}\right)$$

$$p = \text{constant} \times \tau_\perp^{-6} \frac{1 + 1/f}{(1/f - 1)^3}$$

$$\text{with } \nu = \frac{1}{2}$$

$$\text{domain 1: } f = \frac{1}{2 - z/\tau_\perp}$$

$$\text{with } \frac{z}{\tau_\perp} > 3$$

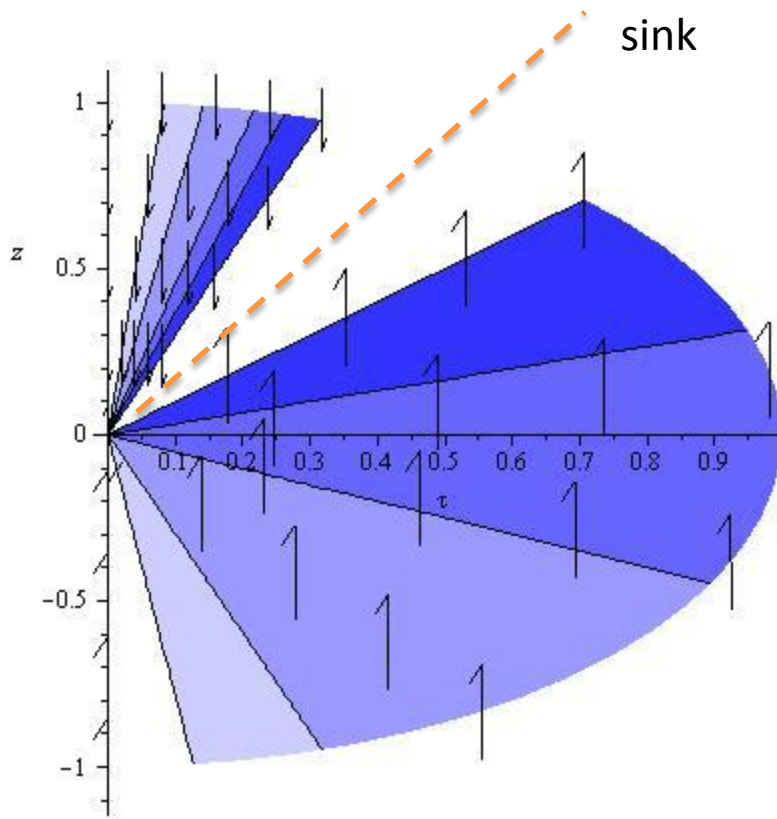
$$\text{domain 2: } f = \frac{1}{2 - z/\tau_\perp}$$

$$\text{with } \frac{z}{\tau_\perp} < 1$$

Also its party partner



# Flow with $v=1/2$



Impinging flow with a moving "sink" at  $\xi=2$

# Solutions for specific $\nu$ (EOS)

$$\bar{v}_z = f\left(\frac{z}{\tau_\perp}\right)$$

$$p = \text{constant} \times \tau_\perp^{-7/2} (1/f^2 - 1)^{-7/4}$$

$$\text{with } \nu = \frac{1}{7}$$

$$\text{domain 1: } f = f_1 \equiv \frac{2\tau_\perp}{3z} [1 - \zeta(z/\tau_\perp) - \bar{\zeta}(z/\tau_\perp)]$$

$$\text{with } \frac{z}{\tau_\perp} > 0 \text{ or } \frac{z}{\tau_\perp} < 0$$

$$\text{domain 2: } f = f_2 \equiv \frac{2\tau_\perp}{3z} \left[ 1 + \frac{1 + i\sqrt{3}}{2} \zeta(z/\tau_\perp) + \frac{1 - i\sqrt{3}}{2} \bar{\zeta}(z/\tau_\perp) \right]$$

$$\text{with } 0 < \frac{z}{\tau_\perp} < 1 \text{ or } -1 < \frac{z}{\tau_\perp} < 0$$

$$\zeta(\xi) = \left[ -1 + \frac{27}{16}\xi^2 + i\sqrt{1 - \left(1 - \frac{27}{16}\xi^2\right)^2} \right]^{1/3}$$

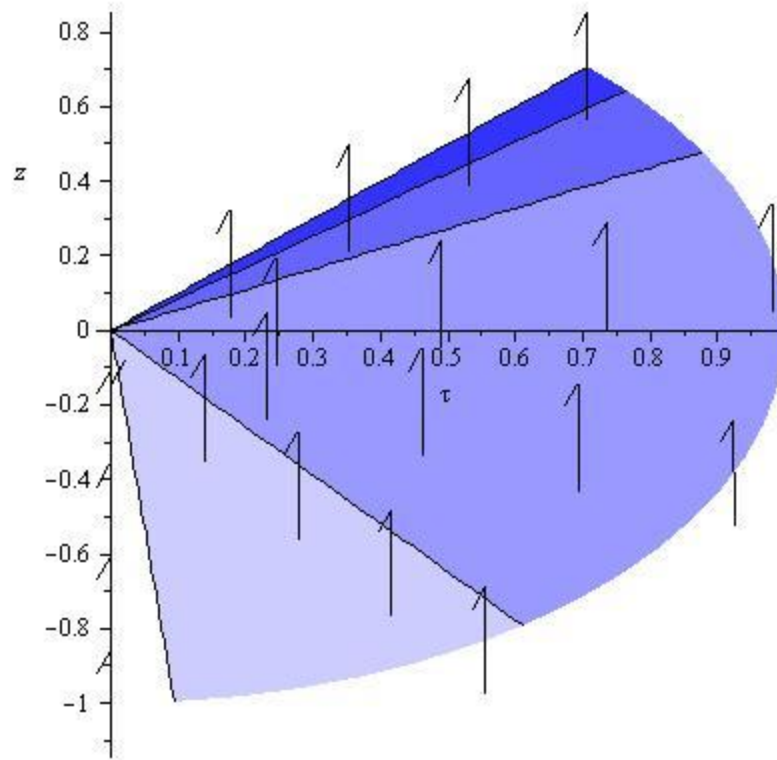
$$\bar{\zeta}(\xi) = \left[ -1 + \frac{27}{16}\xi^2 - i\sqrt{1 - \left(1 - \frac{27}{16}\xi^2\right)^2} \right]^{1/3}$$

# Flow with $v=1/7$

$$f = \begin{cases} f_1 & z/\tau_{\perp} < 0 \\ f_2 & 0 < z/\tau_{\perp} < 1 \end{cases}$$

$$f = \begin{cases} f_2 & -1 < z/\tau_{\perp} < 0 \\ f_1 & z/\tau_{\perp} > 0 \end{cases}$$

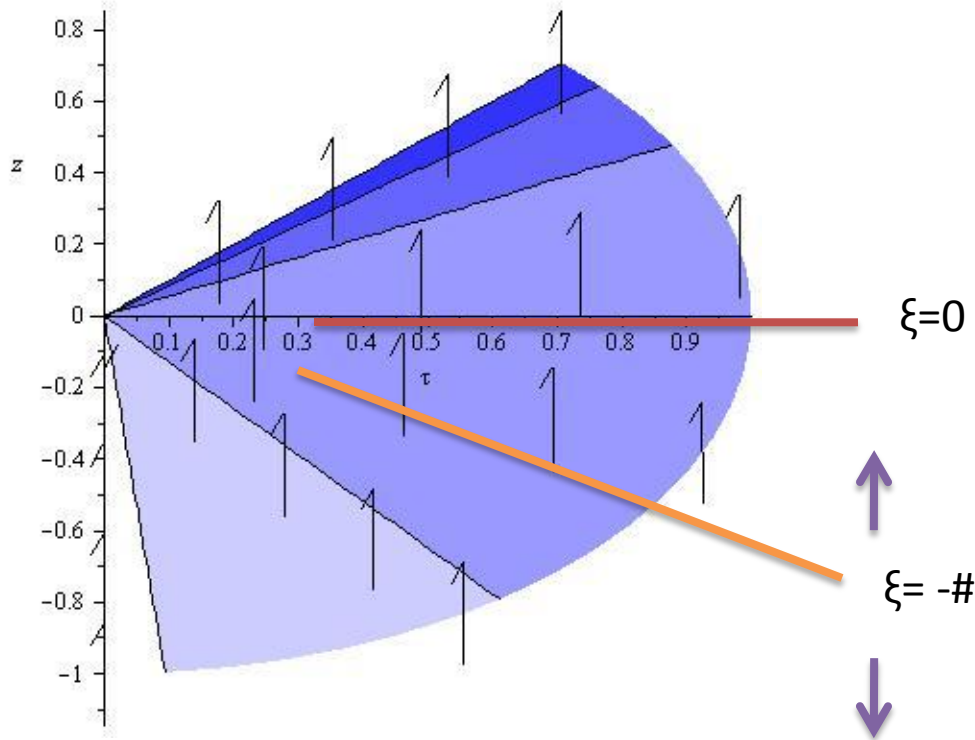
Form a parity pair



One-way shock wave

Flow reaches speed of light at  $\xi=1$

# Connection to Heavy Ion Collisions



Flow direction is observer dependent

One-way shock wave viewed from observer at  $\xi=0$   
Explosion viewed from observer at  $\xi=-\#$

A change of reference frame from  $\xi=0$  to  $\xi=-\#$  may be close to the situation of fireball explosion

# Summary

- We have found several longitudinal flow profiles based on prescribed transverse flow(embedding)
- Connections to HIC may be established by applying longitudinal boost to certain solutions
- Extension from cylindrical symmetry to ellipsoidal symmetry can be used to gain insight to elliptic flow

Thank you!