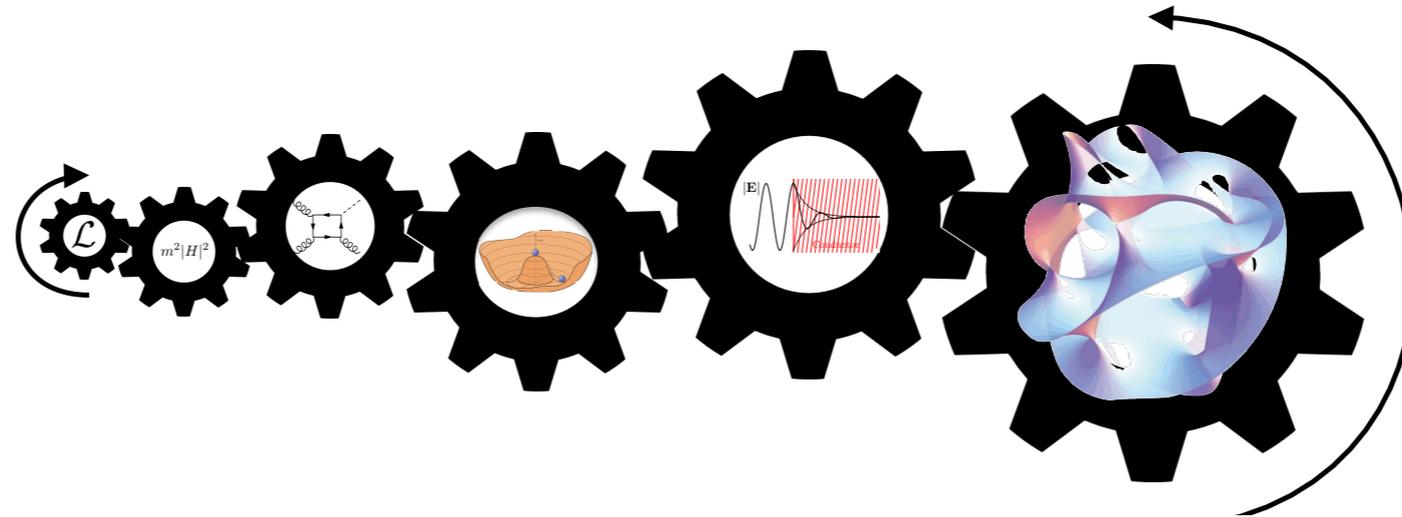


Clockworking FIMPS



Dipan Sengupta

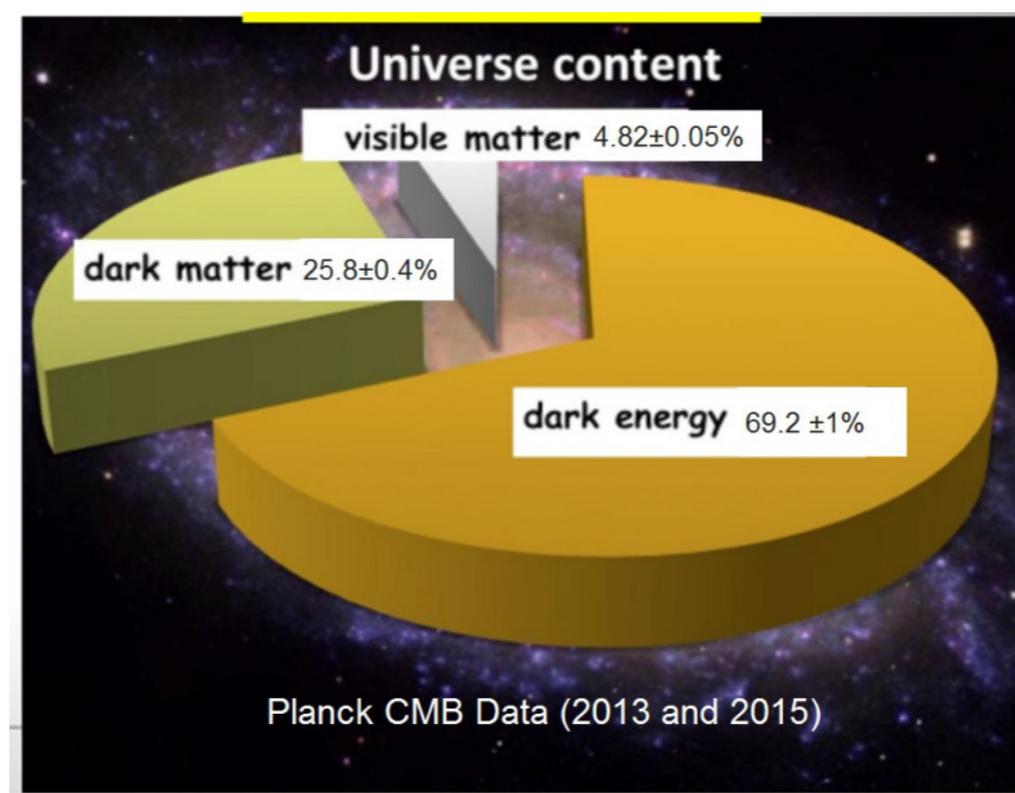
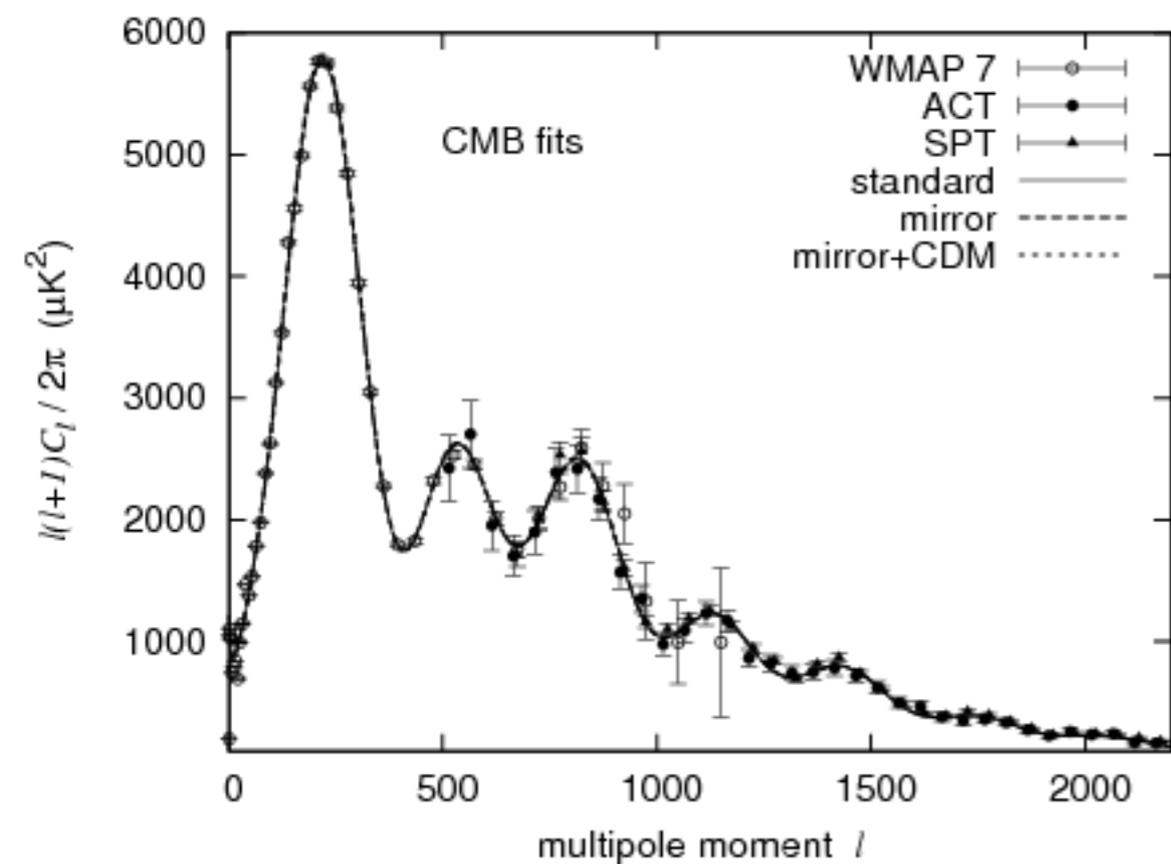
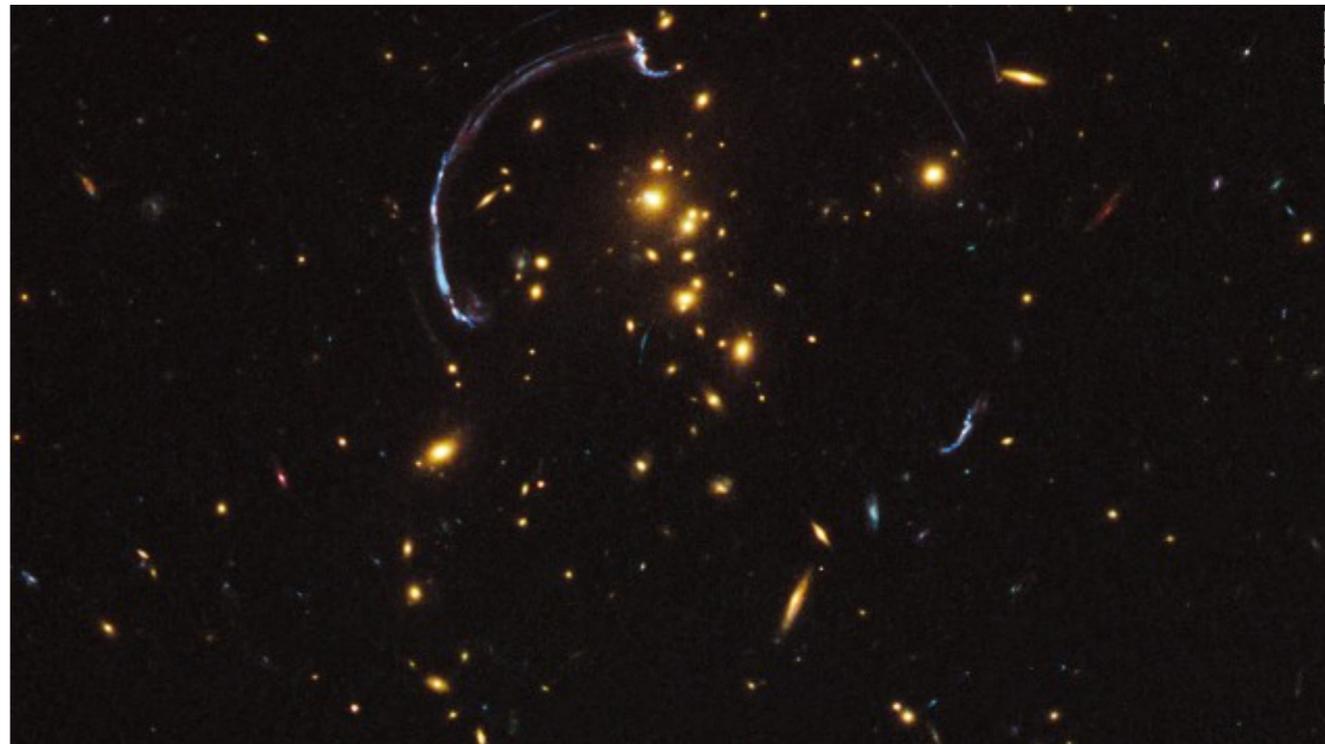
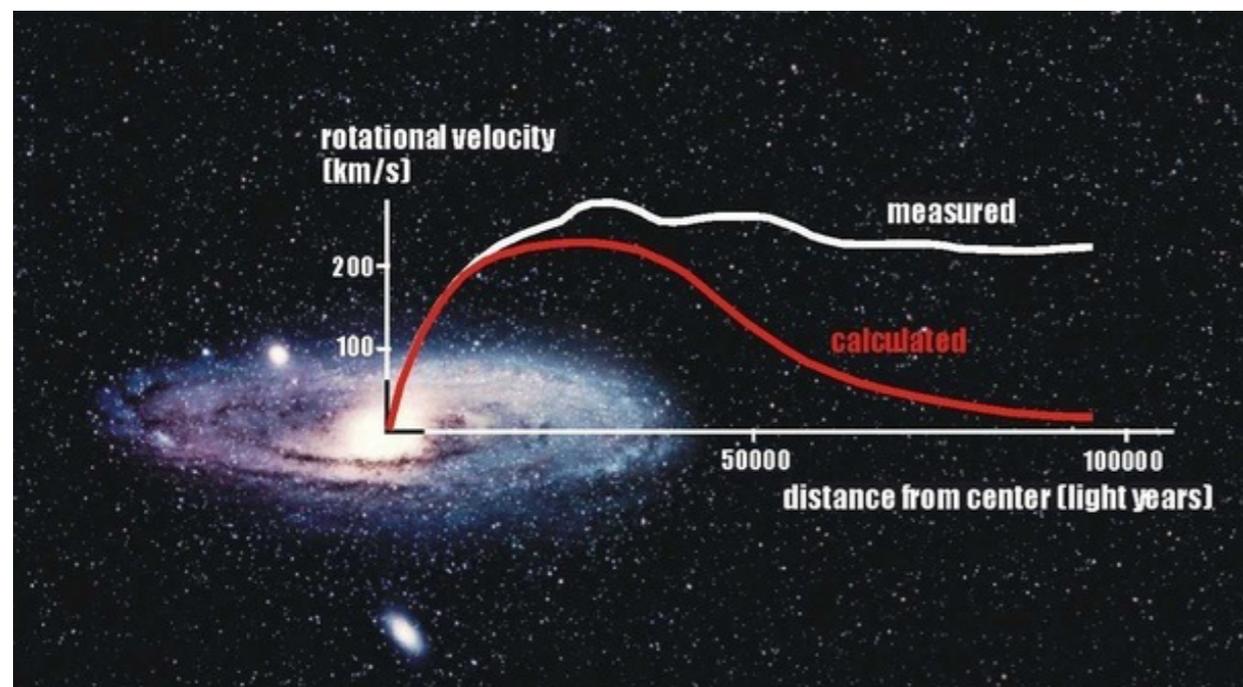


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UNIVERSITY

Based on
A.Goudelis (LPTHE, Paris), K. Mohan (MSU, U.S.A)
arXiv: 1807.06442, JHEP-xxxx

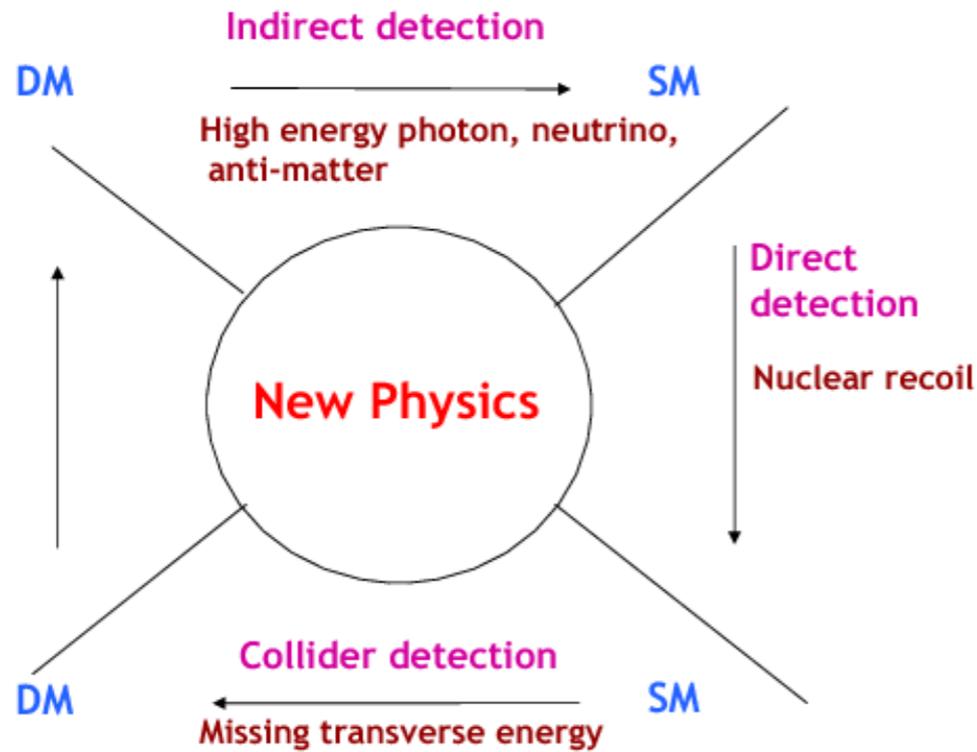
- **Revisiting Freeze-out.**
- **A general introduction to Freeze-in.**
- **Hierarchy problem and clockworks**
- **Introduction to the clockwork set up**
- **Clockwork scalars and fermions.**
- **Models of scalar and fermionic FIMPs.**

DM : Evidence



$$\Omega h^2|_{\text{exp}} = 0.1188 \pm 0.0010$$

Weakly Interacting Massive Particle(WIMPs)



DM in thermal equilibrium with bath particles in early universe

$$n = g \left(\frac{mT}{2\pi} \right)^{3/2} e^{-m/T}$$

$$\frac{dn_\chi}{dt} + 3Hn_\chi = \langle \sigma v \rangle [n_\chi^2 - n_{eq}^2]$$

Boltzmann equation for DM

$$\langle \sigma_{\text{ann}} v_{\text{rel}} \rangle(x) = \frac{4x}{K_2(x)^2} \int_1^{+\infty} d\bar{s} \sqrt{\bar{s}} \cdot (\bar{s} - 1) \cdot K_1(2x\sqrt{\bar{s}}) \cdot \sigma_{\text{ann}} \quad x \equiv \frac{M_X}{T}$$

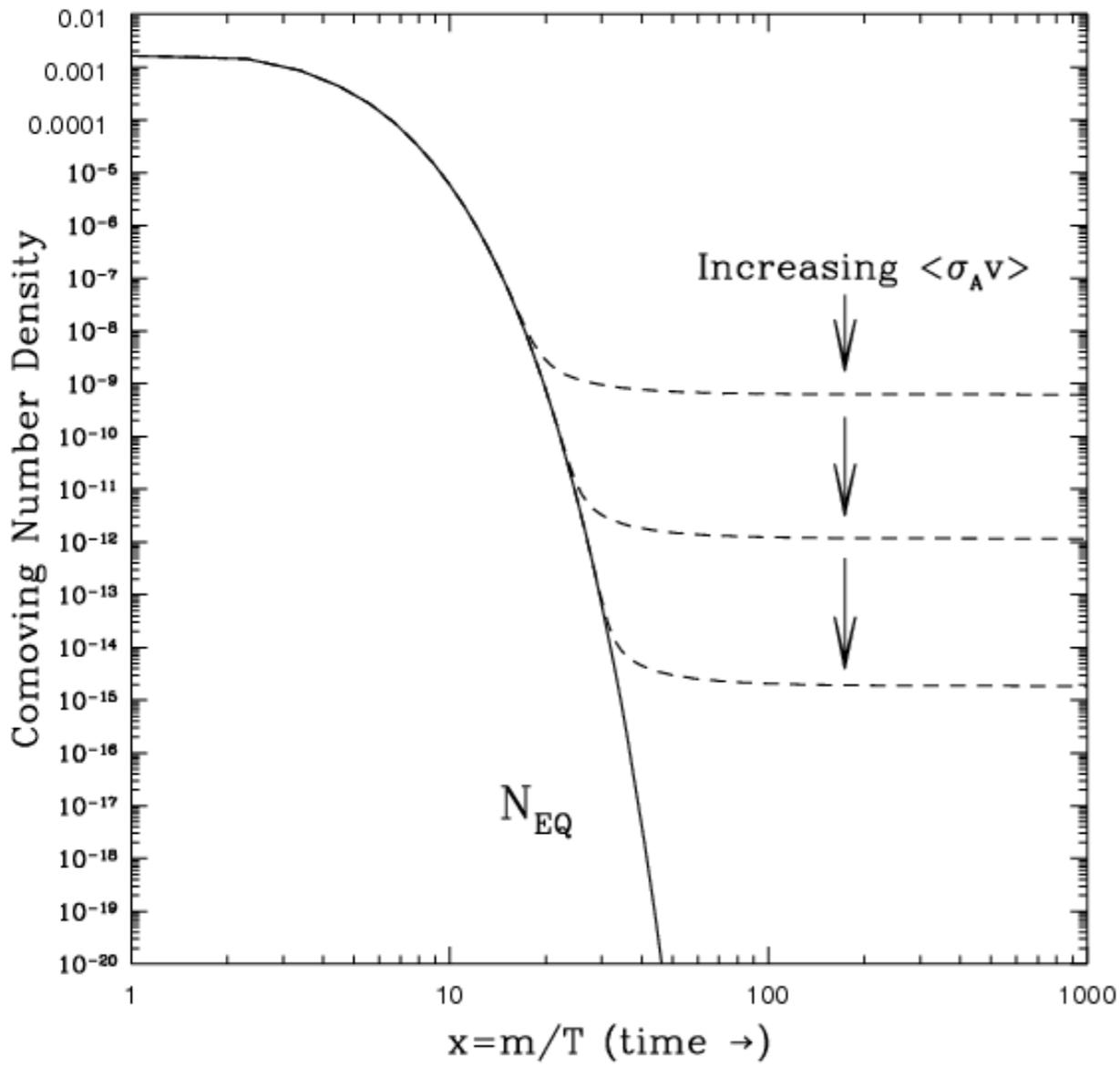
$$\begin{aligned} \Omega_X &\equiv \frac{\rho_{X,0}}{\rho_{\text{crit},0}} \\ &= \frac{M_X n_{X,0}}{3M_{\text{pl}}^2 H_0^2} = \frac{M_X N_{X,0} s_0}{3M_{\text{pl}}^2 H_0^2} = M_X N_X^\infty \frac{s_0}{3M_{\text{pl}}^2 H_0^2} \end{aligned}$$

$$\Omega_\chi h^2 = 0.1 \frac{x_f}{28} \frac{\sqrt{g_{\text{eff}}}}{10} \frac{2 \cdot 10^{-26} \text{cm}^3 \text{s}^{-1}}{\langle \sigma_{\chi\chi} v_{\text{ann}} \rangle}$$

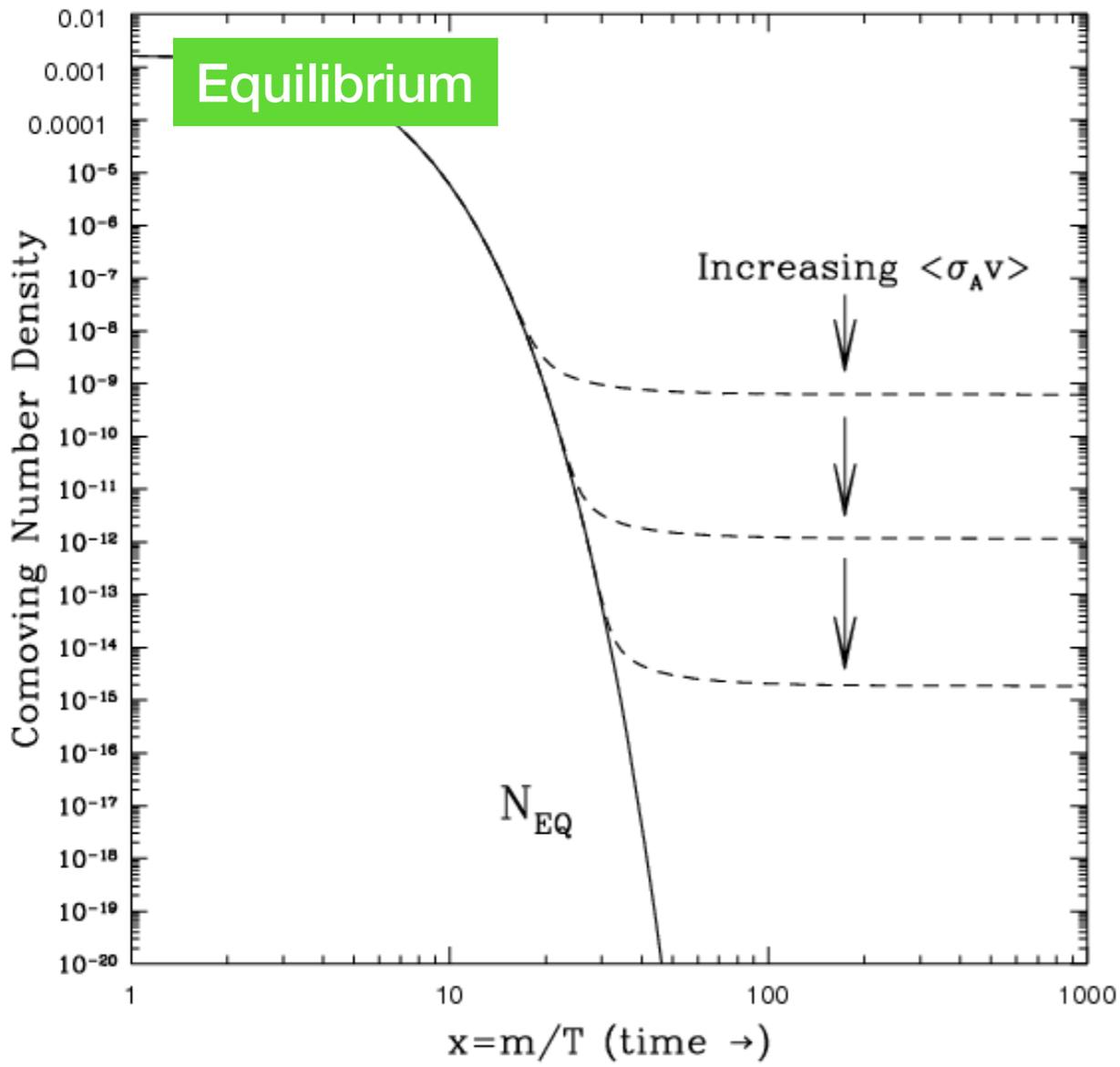
Interaction Strength : Weak Scale -> A BSM theory that also solves Hierarchy problem

Weakly Interacting Massive Particle

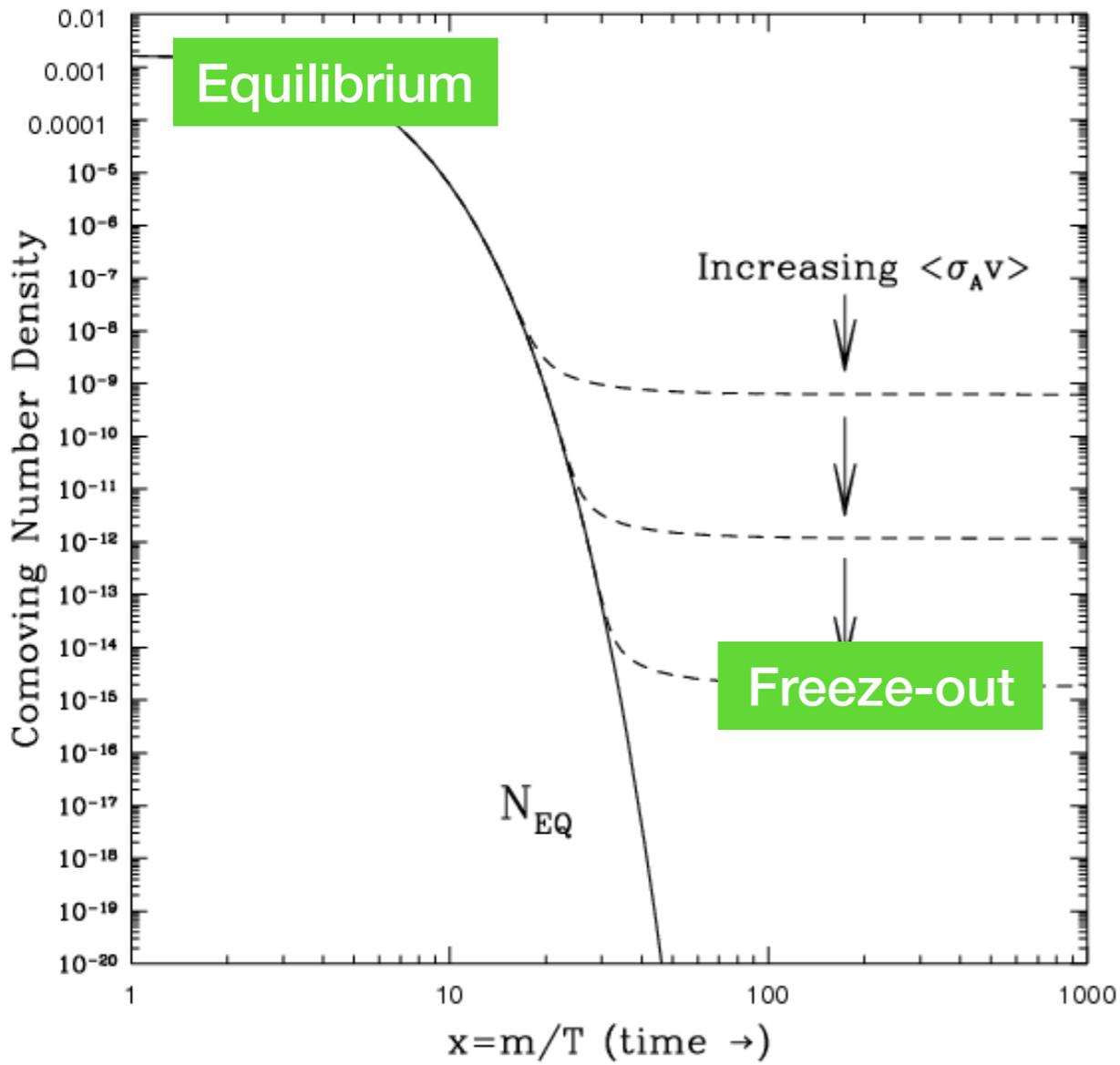
Weakly Interacting Massive Particle



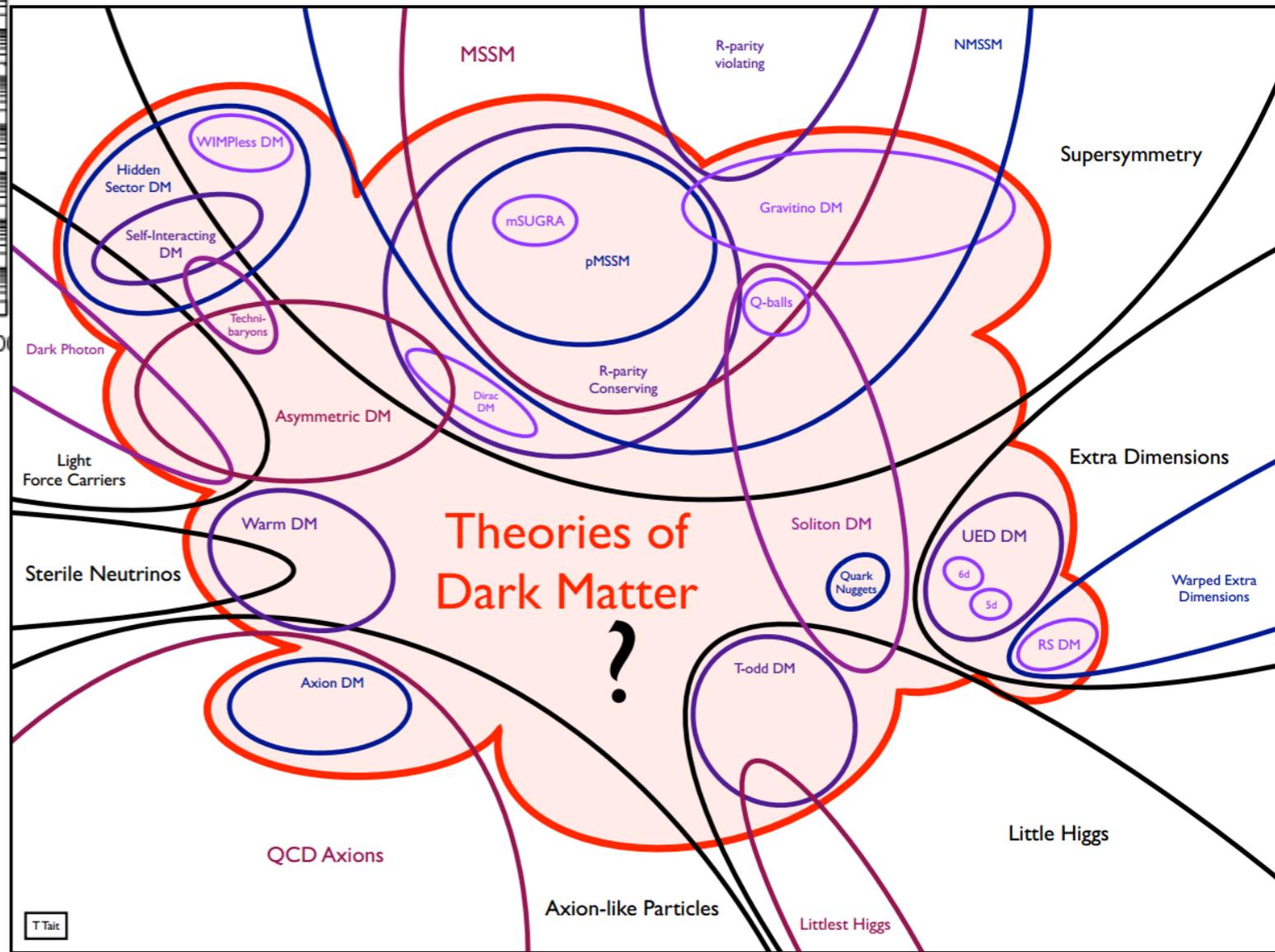
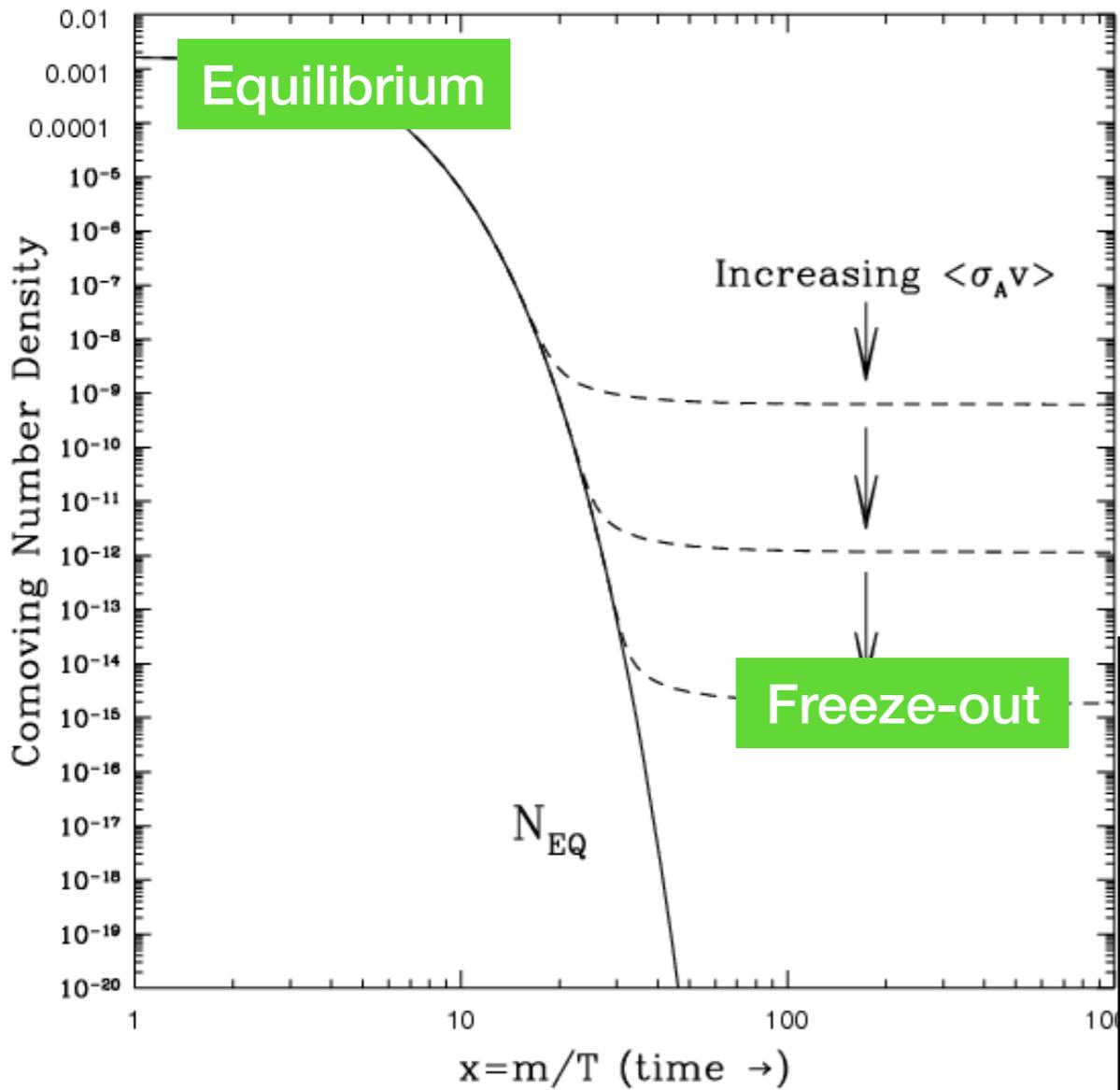
Weakly Interacting Massive Particle



Weakly Interacting Massive Particle



Weakly Interacting Massive Particle



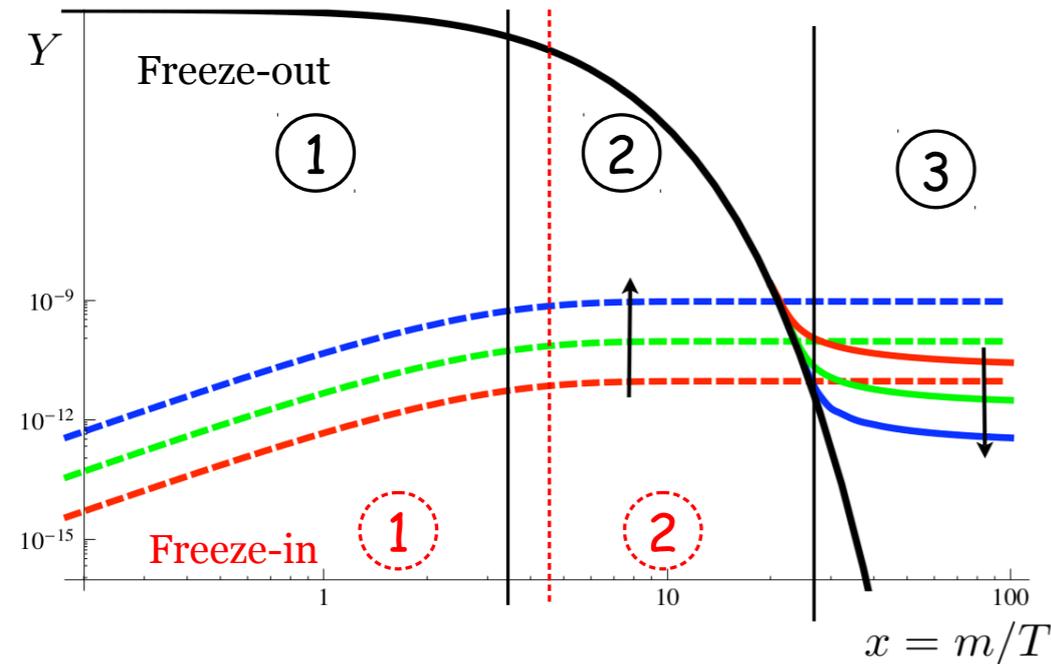
Freeze-in: general idea

arXiv:hep-ph/0106249

arXiv:0911.1120

arXiv:1706.07442...

Tweaked from, arXiv:0911.1120



Two basic premises :

- DM interacts *very* weakly with the SM.
- DM has a negligible initial density.

Assume that in reaction $A \rightarrow B$, ξ_A/ξ_B particles of type χ are destroyed/created. Integrated Boltzmann equation :

$$\dot{n}_\chi + 3Hn_\chi = \sum_{A,B} (\xi_B - \xi_A) \mathcal{N}(A \rightarrow B)$$

① DM produced from decays/annihilations of other particles.

② DM production disfavoured \rightarrow Abundance freeze-in

$$\mathcal{N}(in \rightarrow out) = \int \prod_{i=in} \left(\frac{d^3 p_i}{(2\pi)^3 2E_i} f_i \right) \prod_{j=out} \left(\frac{d^3 p_j}{(2\pi)^3 2E_j} (1 \mp f_j) \right) \times (2\pi)^4 \delta^4 \left(\sum_{i=in} P_i - \sum_{j=out} P_j \right) C_{in} |\mathcal{M}|^2$$

Freeze-in vs freeze-out

Naively, the freeze-in BE is simpler than the freeze-out one. However :

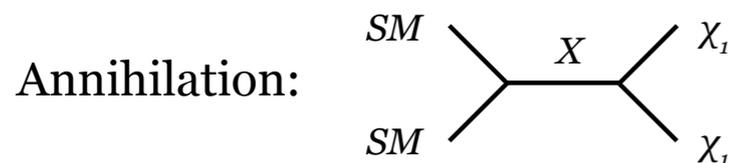
- Initial conditions:
- FO: equilibrium erases all memory.
 - FI: Ωh^2 depends on the initial conditions.

- Relevant temperature:
- FO: around $m_\chi/20$.
 - FI: several possibilities ($m_\chi/3$, $m_{\text{parent}}/3$, T_R or higher), depending on DM production channel and on nature of underlying theory.

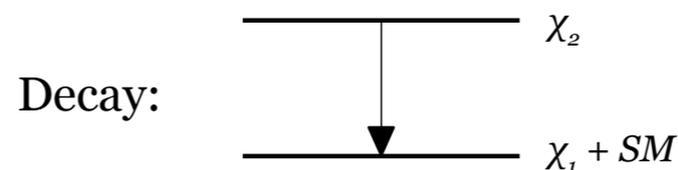
- Statistics can become important. So can very early Universe physics.
- Easily above *e.g.* T_{EWSB} → Phase transitions may occur *after* DM production

Model-building issues

What kind of couplings do we need for successful freeze-in?



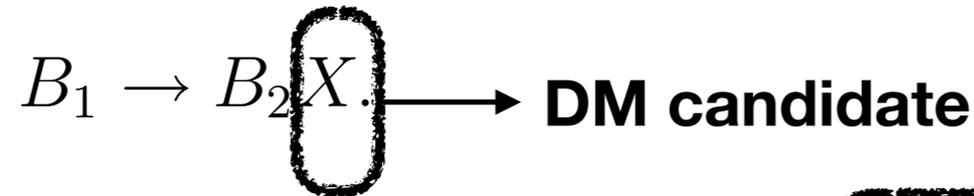
- Requires $\lambda_1 \lambda_2 \sim 10^{-10} - 10^{-12}$



- Requires $\lambda \sim 10^{-13} \times (m_{\chi_2}/m_{\chi_1})^{1/2}$

Freeze-In Dark Matter/ Feebly Interacting massive particles

$$\lambda X B_1 B_2$$



Exponentially small coupling required

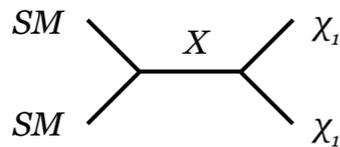
$$\Omega_X h^2|_{tot} \approx \frac{1.09 \times 10^{27}}{g_*^S \sqrt{g_*^P}} m_X \sum_i \frac{g_{B_i} \Gamma_{B_i}}{m_{B_i}^2}$$

$$\lambda \simeq 1.5 \times 10^{-13} \left(\frac{m_B}{m_X}\right)^{1/2} \left(\frac{g_*(m_B)}{10^2}\right)^{3/4} \left(\frac{g_{bath}}{10^2}\right)^{-1/2}$$

Model-building issues

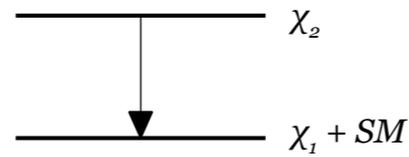
What kind of couplings do we need for successful freeze-in?

Annihilation:



• Requires $\lambda_1 \lambda_2 \sim 10^{-10} - 10^{-12}$

Decay:



• Requires $\lambda \sim 10^{-13} \times (m_{\chi_2}/m_{\chi_1})^{1/2}$

How can we justify such small numbers?

Two main ways so far:

- Scale suppression
- Symmetries

The idea of “Clockwork”-ing

What’s the clockwork mechanism?

- **clockwork mechanism** → an elegant and economical way to generate **tiny numbers**/large hierarchies X with only $\mathcal{O}(1)$ **couplings** and $\mathcal{N} \sim \log X$ **fields**
- Originally introduced in the context of relaxation models, to solve technical issues present in these [Choi, Im, '15; Kaplan, Rattazzi, '15]
- Then realized as a **framework** for model building: [Giudice, McCullough, '16]
 - low-scale invisible axions [Giudice, McCullough, '16; Farina, Pappadopulo, Rompineve, Tesi, '16;
 - **hierarchy problem** [Giudice, McCullough, '16; Giudice, McCullough, Katz, Torre, Urbano, '17;
 - inflation [Kehagias, Riotto, '16; ...]
 - neutrino physics [Ibarra, Kushwaha, Vempati, '17; ...]
 - UV/EFT relation [Craig, Garcia Garcia, Sutherland, '17; Giudice, McCullough, '17]
 - supergravity [Kehagias, Riotto, '17; Antoniadis, Delgado, Markou, Pokorski, '17]
 - **collider pheno** and cosmology [Giudice, McCullough, Katz, Torre, Urbano, '17;

Criticisms : Works only for abelian theories
Continuum solutions only for special cases
UV completions are tricky to find

See criticism by,
Craig, Garcia, Sutherland

The idea of “Clockwork”-ing

How the clockwork works (made easy)

Based on the simple observation that:

$1/2 \times 1/2 \times 1/2 \times 1/2 \times \dots \times 1/2$ can **easily** be **tiny**

Use a **chain** of N fields

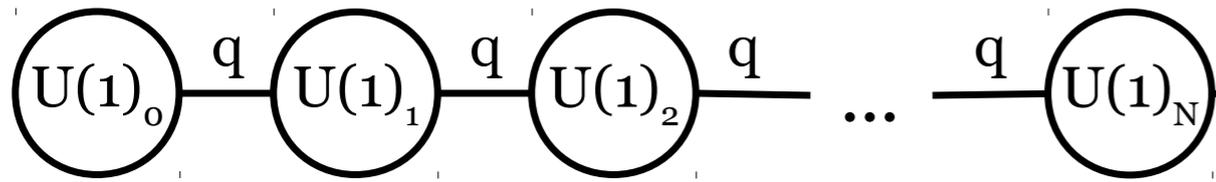
$$\phi_0 \xrightarrow{1/q} \phi_1 \xrightarrow{1/q} \phi_2 \xrightarrow{1/q} \phi_3 \xrightarrow{1/q} \dots \xrightarrow{1/q} \phi_N \text{ --- SM}$$

For **fermions** use chiral symmetries

$$R_0 \xrightarrow{m} \underbrace{L_1 \ R_1}_{qm} \xrightarrow{m} \underbrace{L_2 \ R_2}_{qm} \xrightarrow{m} \underbrace{L_3 \ R_3}_{qm} \xrightarrow{m} \dots \xrightarrow{m} \underbrace{L_N \ R_N}_{qm} \text{ --- } L_{SM}$$

$$\text{light } N \approx R_0 \implies N - L_{SM} \sim \mathbf{1/q^N}$$

A clockwork scalar



N+1 copies of U(1) global symmetry in theory space spontaneously broken down to a single U(1) at a scale f

Below f, N+1 massless Goldstones

$$U_j = e^{i\phi_j/f} \quad j = 0, \dots, N$$

Introduce N “soft” mass parameters m_i^2 that break this symmetry explicitly

Background of N spurion fields

Nearest neighbor interactions

$$Q_i = \delta_{ij} - q\delta_{i,j+1} \quad , q > 1$$

Non-diagonal charges

The unbroken generator

$$Q = \sum_{j=0}^N \frac{Q_j}{q^j}$$

A clockwork scalar

Low energy effective lagrangian

$$\mathcal{L} = -\frac{f^2}{2} \sum_{j=0}^N \partial_\mu U_j^\dagger \partial^\mu U_j + \frac{m^2 f^2}{2} \sum_{j=0}^{N-1} (U_j^\dagger U_{j+1}^q + \text{h.c.})$$

$$\mathcal{L}_{SCW} = -\frac{1}{2} \sum_{j=0}^N \partial_\mu \phi_j^\dagger \partial^\mu \phi_j + \sum_{j=0}^{N-1} \frac{m^2}{2} (\phi_j - q\phi_{j+1})^2 + \sum_{j=0}^{N-1} \frac{m^2}{24f^2} (\phi_j - q\phi_{j+1})^4 + \mathcal{O}(\phi_6)$$

$$\begin{aligned} V(\phi) &= \sum_{j=0}^{N-1} \frac{m^2}{2} (\phi_j - q\phi_{j+1})^2 + \sum_{j=0}^{N-1} \frac{m^2}{24f^2} (\phi_j - q\phi_{j+1})^4 + \mathcal{O}(\phi_6) \\ &= \frac{1}{2} \sum_{i,j=0}^N \phi_i M_{ij}^2 \phi_j + \frac{m^2}{24f^2} \sum_{i,j=0}^N (\phi_i M_{ij}^2 \phi_j)^2 + \mathcal{O}(\phi_6) \end{aligned}$$

See A.G, K.M, D.S 1807.06642 for analytic diagonalization

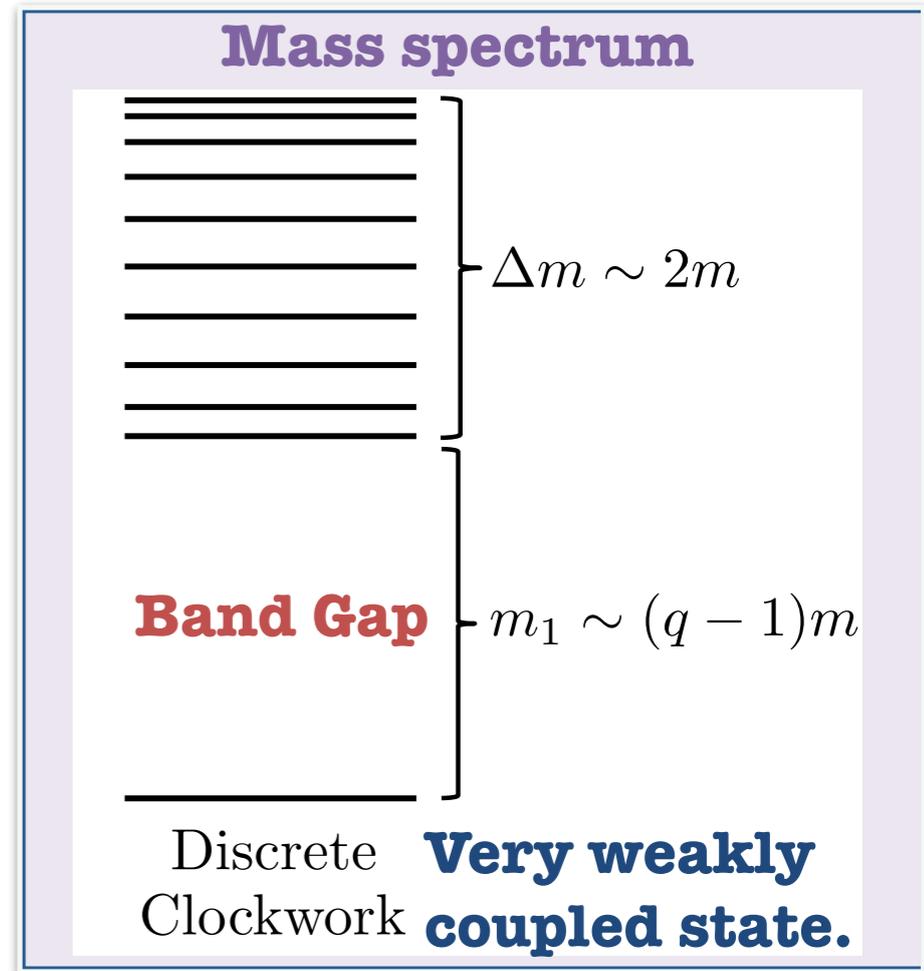
A clockwork scalar

$$\mathcal{L} = \frac{\pi_N}{16\pi^2 f} G_{\mu\nu} \tilde{G}^{\mu\nu}$$

$$\mathcal{L} = \frac{1}{16\pi^2} G_{\mu\nu} \tilde{G}^{\mu\nu} \left(\frac{a_0}{f_0} - \sum_{k=1}^N (-)^k \frac{a_k}{f_k} \right)$$

$$f_0 \equiv \frac{f q^N}{\mathcal{N}_0}, \quad f_k \equiv \frac{f}{\mathcal{N}_k q \sin \frac{k\pi}{N+1}}.$$

The clockwork axion



UV completions

- • **Parallels with RS/LED scenarios**
- **Continuum version : Linear dilaton set up.**
- **Low energy version of Little String Theory.**
- **Some interesting background geometries in the continuum**
- a well-defined continuum limit exists and selects either
 - massless field in curved **clockwork** metric $ds^2 = e^{\frac{4}{3}ky} (dx^2 + dy^2)$ [Giudice, McCullough, '16]

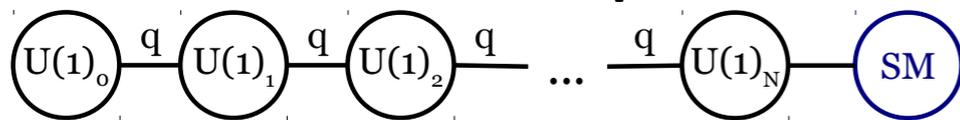
The Clockwork mechanism was initially introduced to address completely different issues. Has found many more applications (inflation, neutrinos, flavour, axions...).

Clockworking a scalar FIMP

A model of scalar FIMP

$$\mathcal{L}_{sFIMP} = -\frac{1}{2} \sum_{j=0}^N \partial_\mu \phi_j^\dagger \partial^\mu \phi_j - \frac{1}{2} \sum_{i,j=0}^N \phi_i M_{ij}^2 \phi_j - \frac{m^2}{24f^2} \sum_{i,j=0}^N (\phi_i \tilde{M}_{ij}^2 \phi_j)^2$$
$$- \kappa |H^\dagger H| \phi_n^2 + \sum_{i=0}^n \frac{t^2}{2} \phi_i^2,$$

**Diagonal mass
term for every
site**



SM coupled to the nth site

Expand the gears in the basis of Eigen-states

Identify the zero mode as the FIMP candidate

Clockworking a scalar FIMP

$$\begin{aligned}\mathcal{L}_{int} &= \kappa |H^\dagger H| \phi_n^2 \\ &= \kappa \sum_{j=0, k=0}^n O_{kl} O_{jm} a_k a_l (v^2 + 2vh + h^2)/2\end{aligned}$$

$$\phi_i \tilde{M}_{ij} \phi_j = \phi_i M_{ij} \phi_j + \kappa v^2 \delta_{in} \delta_{jn} \phi_n \phi_n$$

Modified mass matrix

FIMP mass suppressed by clockwork ~ 1 keV
(close to the warm DM)

How to raise the mass ?

Add a diagonal mass term

$$\sum_{i=0}^N \frac{t^2}{2} \phi_i^2$$

$$M_t = m \cdot \begin{bmatrix} 1 + t^2/m^2 & -q & 0 & \dots & 0 & 0 & 0 \\ -q & 1 + q^2 + t^2/m^2 & -q & \dots & 0 & 0 & 0 \\ \vdots & \vdots & \vdots & \vdots & \vdots & \vdots & \vdots \\ 0 & 0 & 0 & \dots & 1 + q^2 + t^2/m^2 & -q & 0 \\ 0 & 0 & 0 & \dots & -q & q^2 + t^2/m^2 & 0 \end{bmatrix}_{(n+1) \times (n+1)}$$

Diagonal mass term
unsuppressed by clockwork

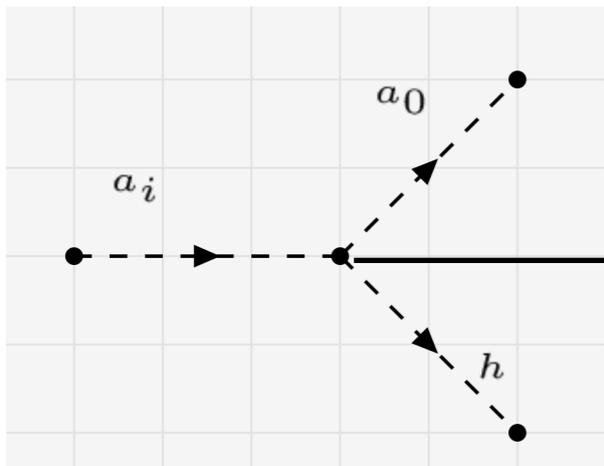


Adds a mass term to the zero
mode without spoiling the
clockwork mechanism

Spurion transformation under $t^2 f^2 (U^{q_j^\dagger} + h.c)$

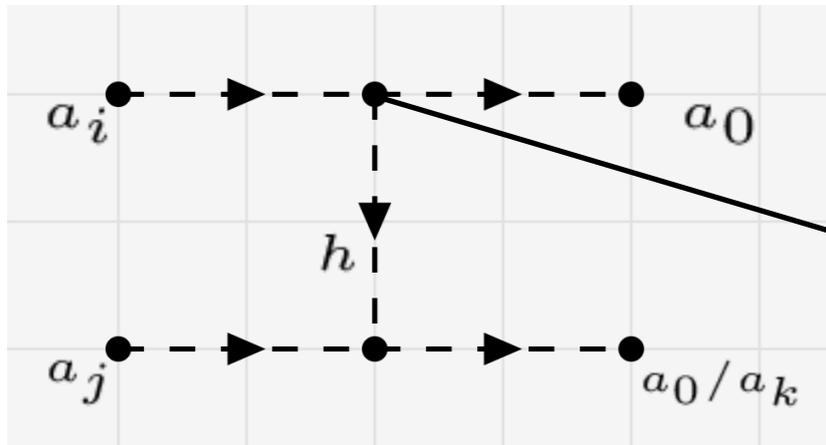
Clockworking a scalar FIMP

FIMP production

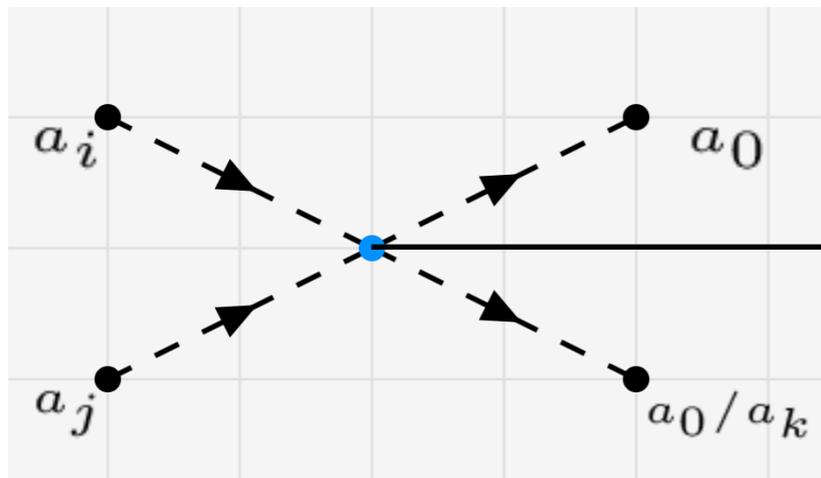


$$\frac{\kappa}{q^n} O_{jn}$$

Dominant process

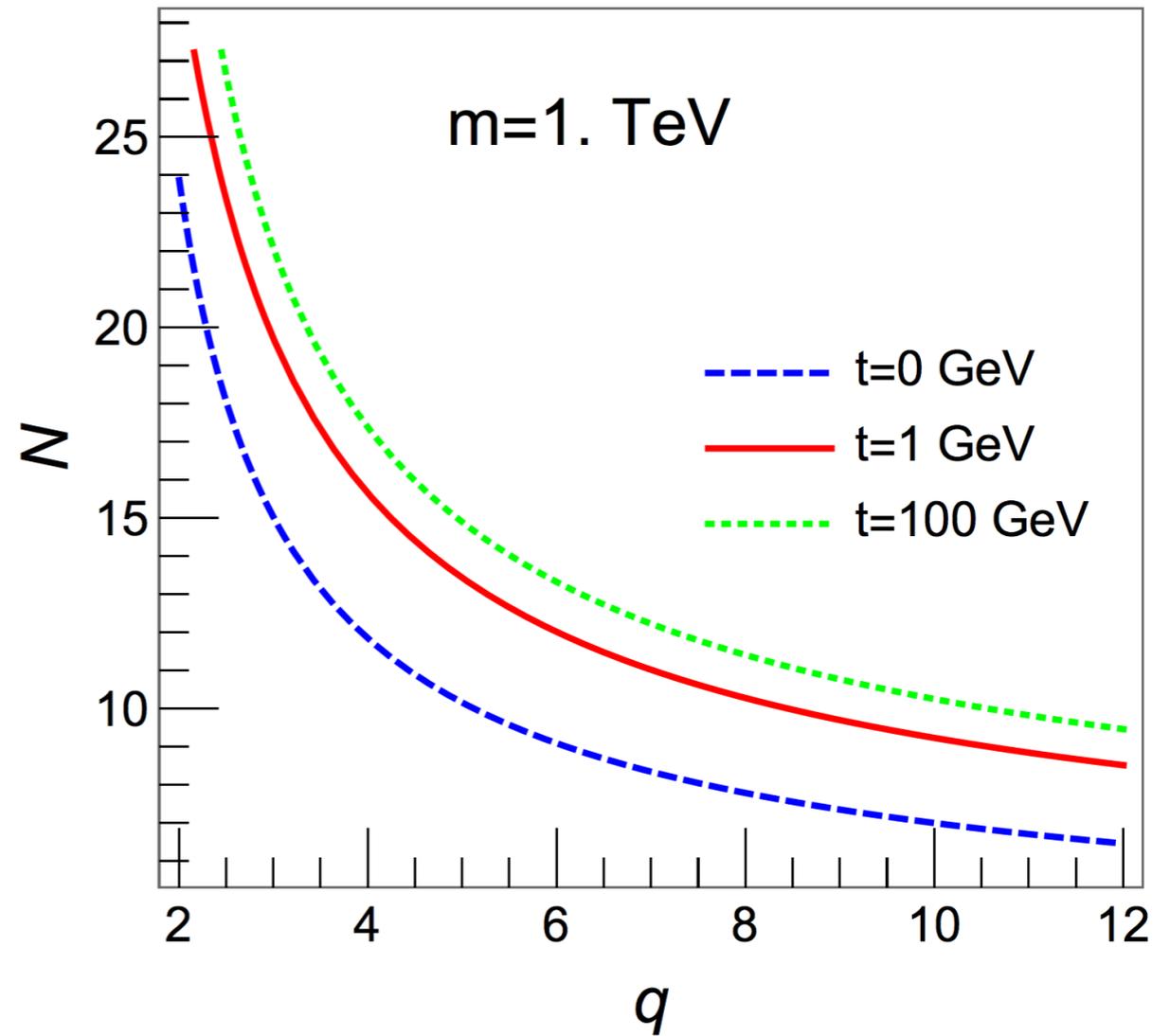
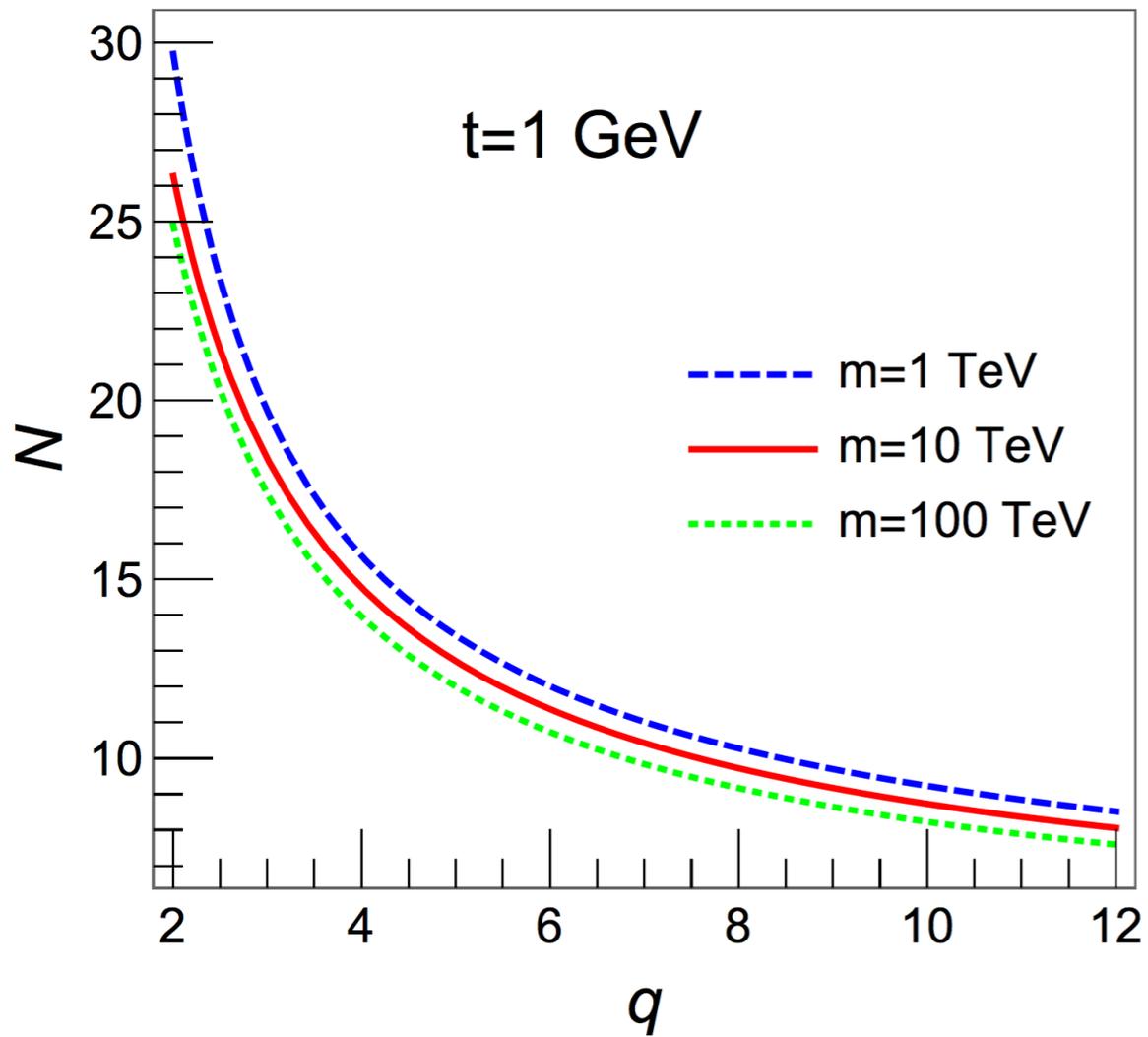


$$\frac{\kappa}{q^n} O_{jn}$$



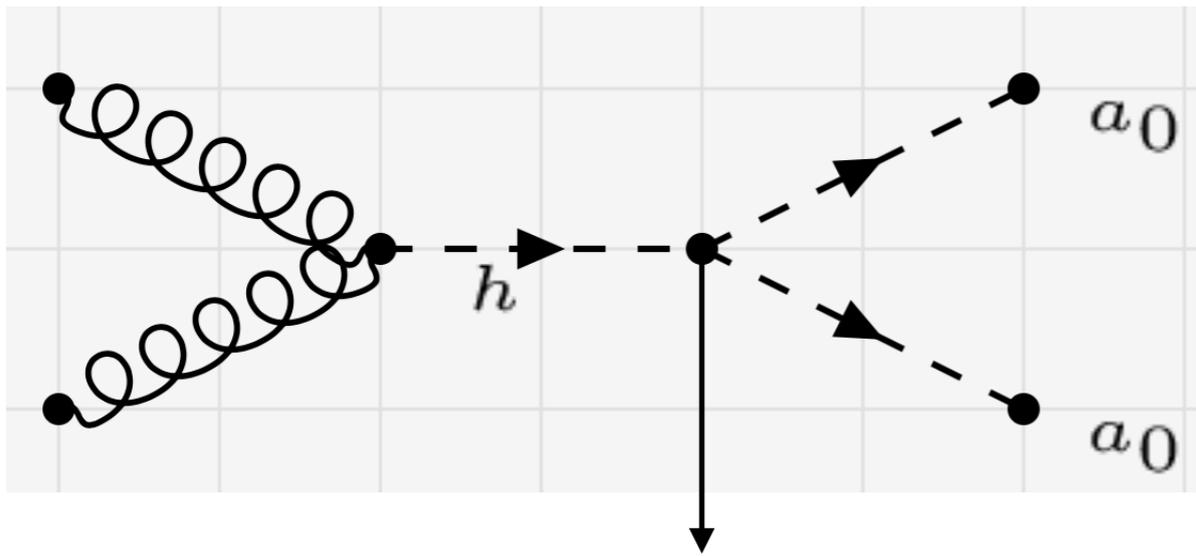
$$\frac{\kappa}{q^n} O_{jn}$$

Clockworking a scalar FIMP

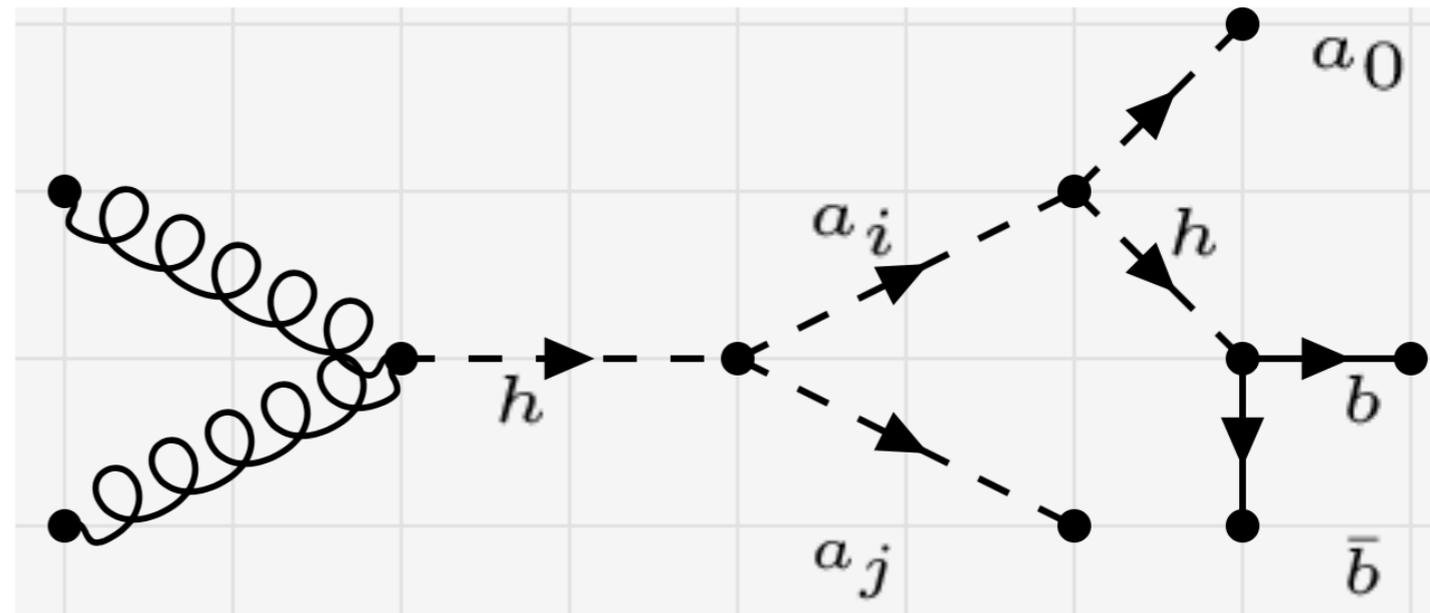


1807.06642 : A . Goudelis, K. Mohan and D.S

Phenomenology of the scalar FIMP



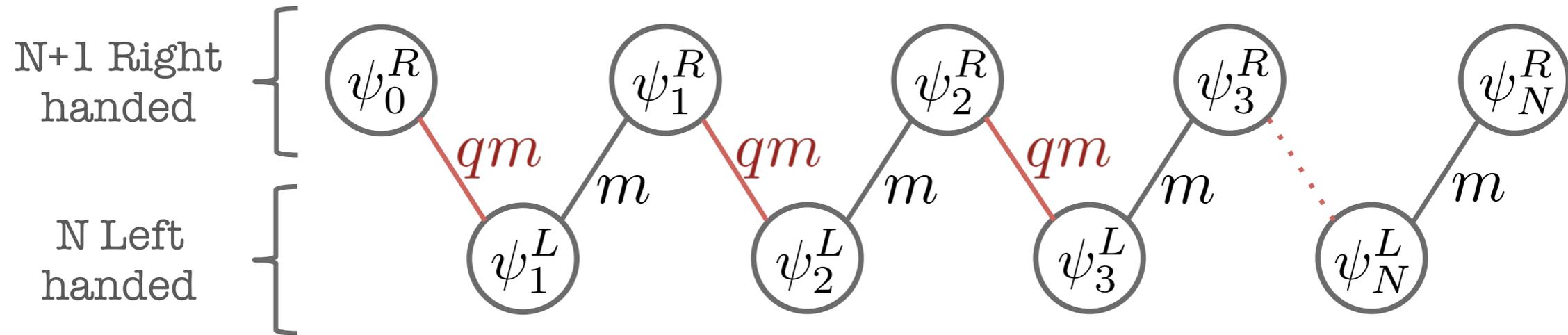
Tiny Invisible Width



Produce the higher gears with Order(α) coupling via off-shell Higgs

Displaced b-jets

A fermionic clockwork



$$\begin{aligned}
 \mathcal{L}_{FCW} &= \mathcal{L}_{kin} - m \sum_{i=0}^{N-1} (\bar{\psi}_{L,j} \psi_{R,j} - q \bar{\psi}_{L,j} \psi_{R,j+1} + h.c) - \frac{M_L}{2} \sum_{i=0}^{N-1} (\bar{\psi}_{L,i}^c \psi_{L,i}) - \frac{M_R}{2} \sum_{i=0}^N (\bar{\psi}_{R,i}^c \psi_{R,i}) \\
 &= \mathcal{L}_{kin} - \frac{1}{2} (\bar{\Psi}^c \mathcal{M} \Psi + h.c)
 \end{aligned} \tag{12}$$

N (N+1) copies of left (right) handed fermions

Majorana Mass terms

A.G, K.M, D.S 1807.06642 for analytic diagonalization

A fermionic FIMP

$$\mathcal{L}_{fFIMP} = \mathcal{L}_{fcw} + i\bar{L}\not{D}L + i\bar{R}\not{D}R + M_D(\bar{L}R) + Y\bar{L}\tilde{H}\psi_{R,n} + \text{h.c}$$

$$L' = (l_1, l_2), R' = (r_1, r_2)$$

$$(SU(3), SU(2))_Y = (1, 2)_{-1/2}$$

$$m_\nu = \begin{matrix} & l_1 & r_1 & \chi_0 & \chi_1 & \chi_2 & \cdots & \chi_{2n} \\ \begin{matrix} l_1 \\ r_1 \\ \chi_0 \\ \chi_1 \\ \chi_2 \\ \vdots \\ \chi_{2N} \end{matrix} & \left(\begin{array}{cccccccc} 0 & M_D & vY_0 & vY_1 & vY_2 & \cdots & vY_{2N} \\ M_D & 0 & 0 & 0 & 0 & \cdots & 0 \\ vY_0 & 0 & M_0 & 0 & 0 & \cdots & 0 \\ vY_1 & 0 & 0 & M_1 & 0 & \cdots & 0 \\ vY_2 & 0 & 0 & 0 & M_2 & \cdots & 0 \\ \vdots & \vdots & \vdots & \vdots & \vdots & \vdots & \vdots \\ vY_{2N} & 0 & 0 & 0 & 0 & \cdots & M_{2N} \end{array} \right) \end{matrix}$$

$$Y_0 = Y(u_R)_N = \frac{1}{q^N} \sqrt{\frac{q^2 - 1}{q^2 - q^{-2N}}}$$

$$Y_j = Y_{(N+j)} = \frac{Y}{\sqrt{2}} (U_R)_{Nj} = \sqrt{\frac{1}{(N+1)\lambda_j}} \left[q \sin \frac{Nj\pi}{N+1} \right], \quad j = 1, \dots, N$$

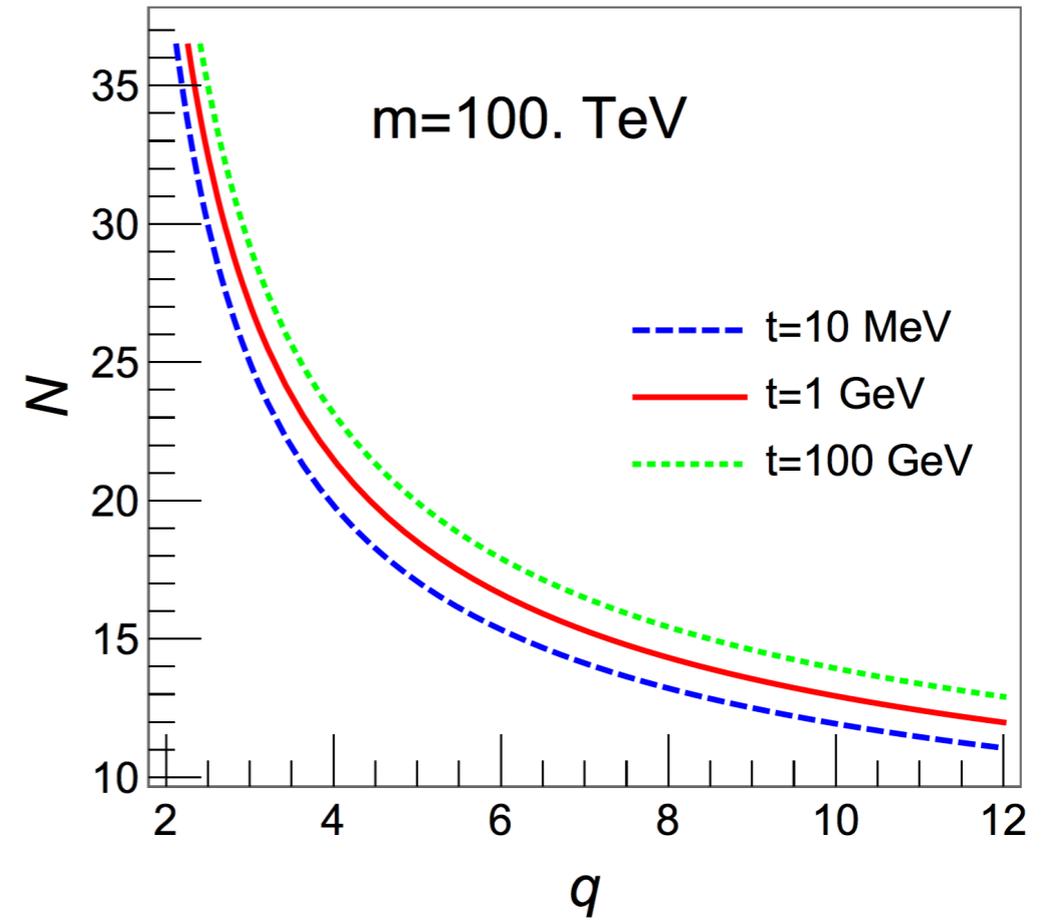
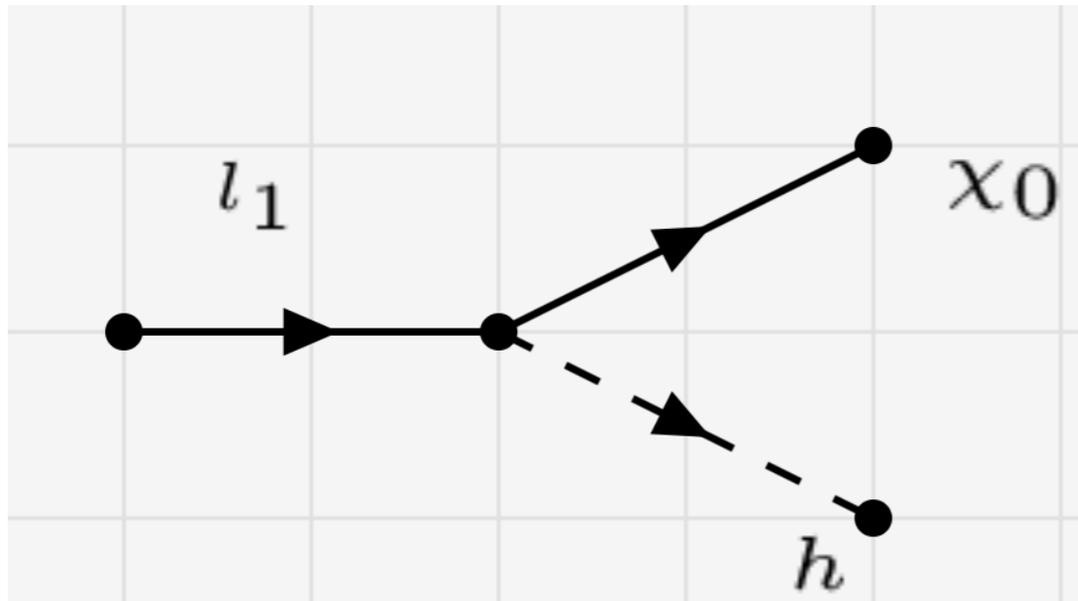
Work in the limit of $v/m \rightarrow 0$ set $m = 100$ TeV

All off-diagonal entries except the Dirac mass vanish

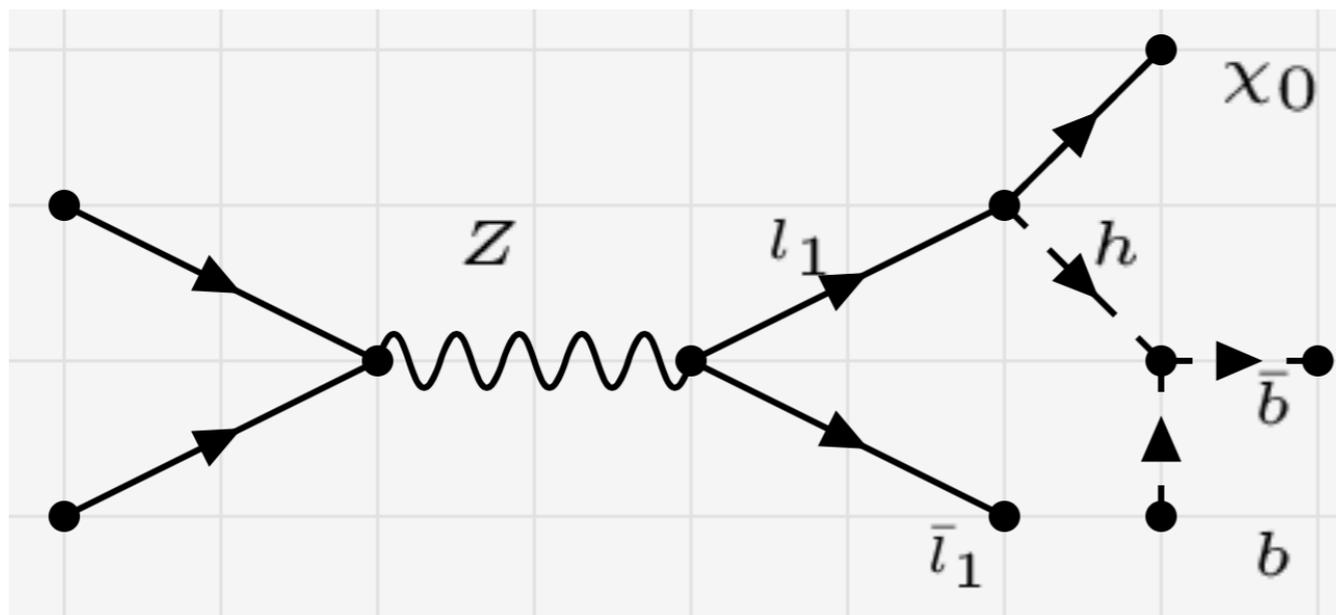
Identify the zero mode fermion of the clockwork as the FIMP candidate

A fermionic clockwork

FIMP production



LHC pheno



Freeze-in phenomenology

Can we test freeze-in? Certainly not in full generality, but

There are actually numerous handles!

Arguably, both remarks also apply to freeze-out

If there are heavier particles in the spectrum

Primordial nucleosynthesis

Long lifetimes

Displaced vertices/
kinked tracks

Shorter lifetimes, requires
tweaking.
More relevant arXiv:1705.09292

Charged track
searches @ LHC

Long lifetimes,
charged parent

Structure formation
(Lyman- α)

Long lifetimes

Mono-X searches @ LHC + new experiments

Long lifetimes, neutral parent, *cf e.g.* arXiv:1806.07396

Otherwise

Direct/Indirect
detection

In very special limits

Structure formation

If DM warm/self-
interacting

Conclusion

- **Freeze -in dark matter scenarios are alternatives to WIMP freeze outs**
- **FIMPs are out of thermal equilibrium and require very small couplings**
- **A “natural” way to generate exponentially small couplings are Clockwork mechanisms**
- **Scalar and Fermionic clockwork freeze-in models are constructed and shown to be a viable set up.**
- **Such set ups also have long lived particles as a viable signature**

Conclusion

- Freeze-in dark matter scenarios are alternatives to WIMP freeze outs
- FIMPs are out of thermal equilibrium and require very small couplings
- A “natural” way to generate exponentially small couplings are Clockwork mechanisms
- Scalar and Fermionic clockwork freeze-in models are constructed and shown to be a viable set up.
- Such set ups also have long lived particles as a viable signature

